

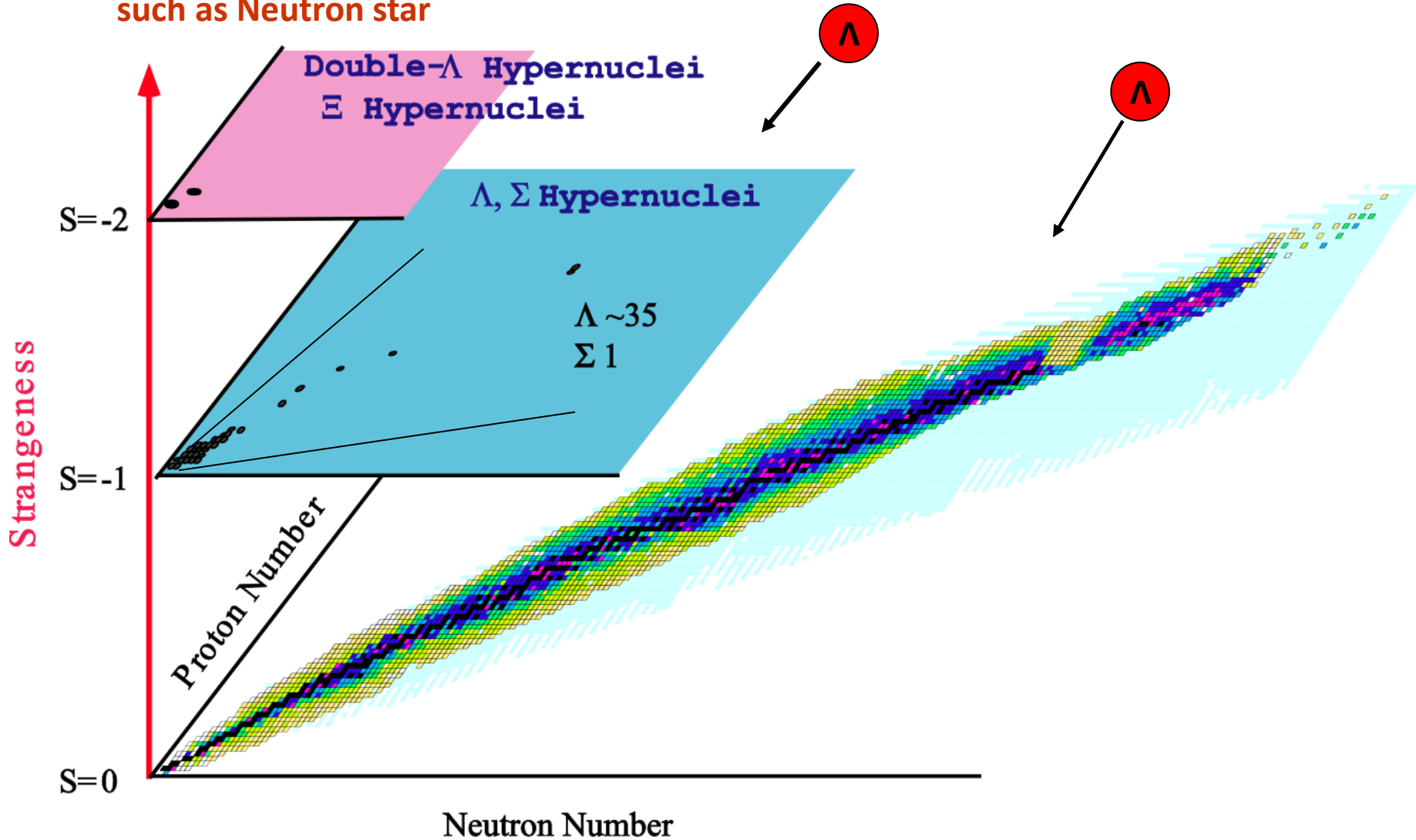
# Structure of light $\Xi$ hypernuclei with modern $\Xi N$ interaction

Emiko Hiyama (Kyushu Univ./RIKEN)

In 2017, move to Kyushu University from RIKEN  
keep my position at RIKEN for 6 years since 2019

# Nuclear chart with strangeness

Multi-strangeness system  
such as Neutron star



# Outline of my talk

- Historical Overview on hypernuclear physics
- What is the hot topics in hypernuclear physics?
  - $\Xi$  hypernuclei
- What is problem? How do we try to solve it?

# The major goal of hypernuclear physics

1) To understand baryon-baryon interactions

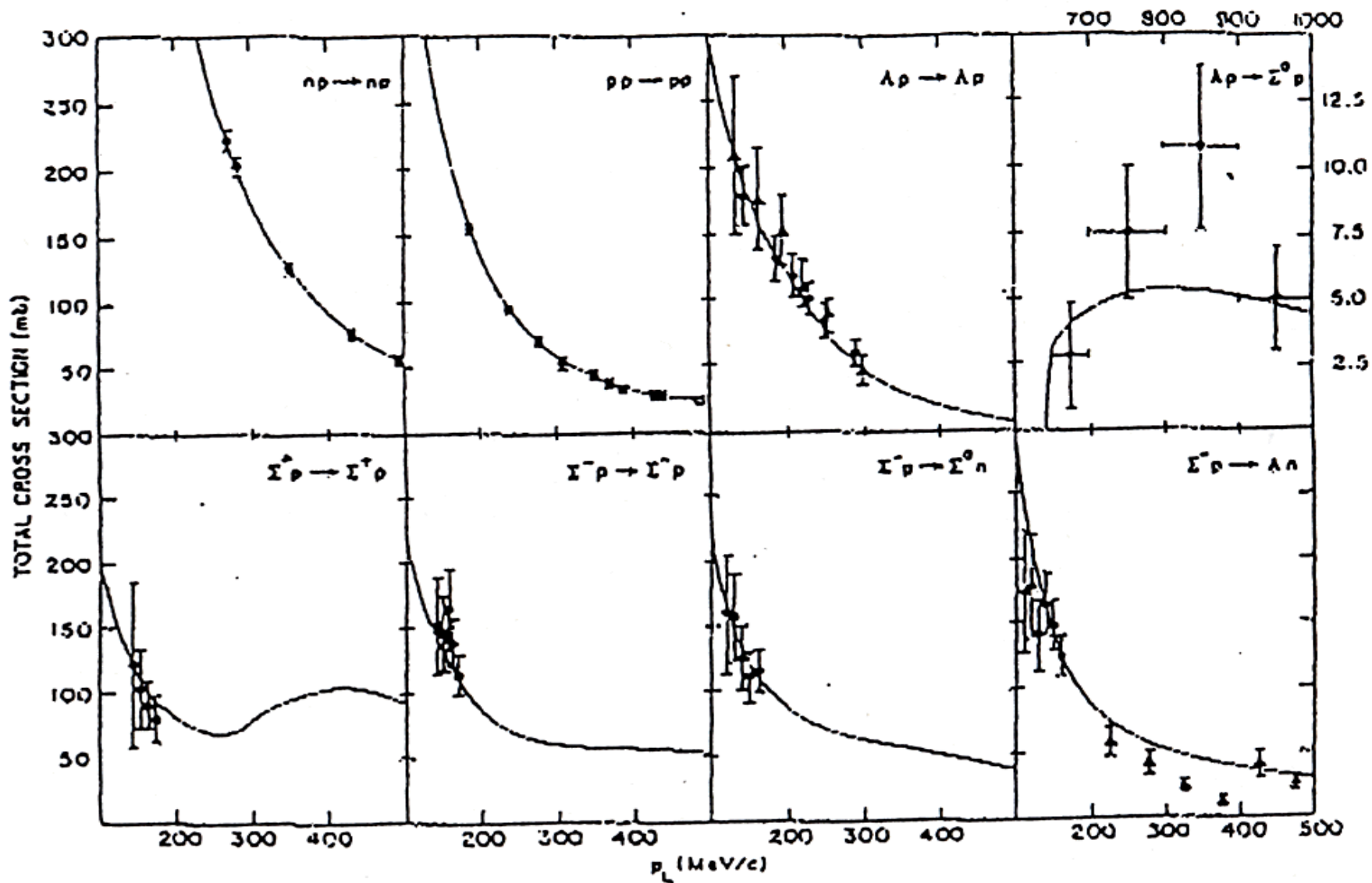
To understand the baryon-baryon interaction, two-body scattering experiment is most useful.

Fundamental and important for the study of nuclear physics

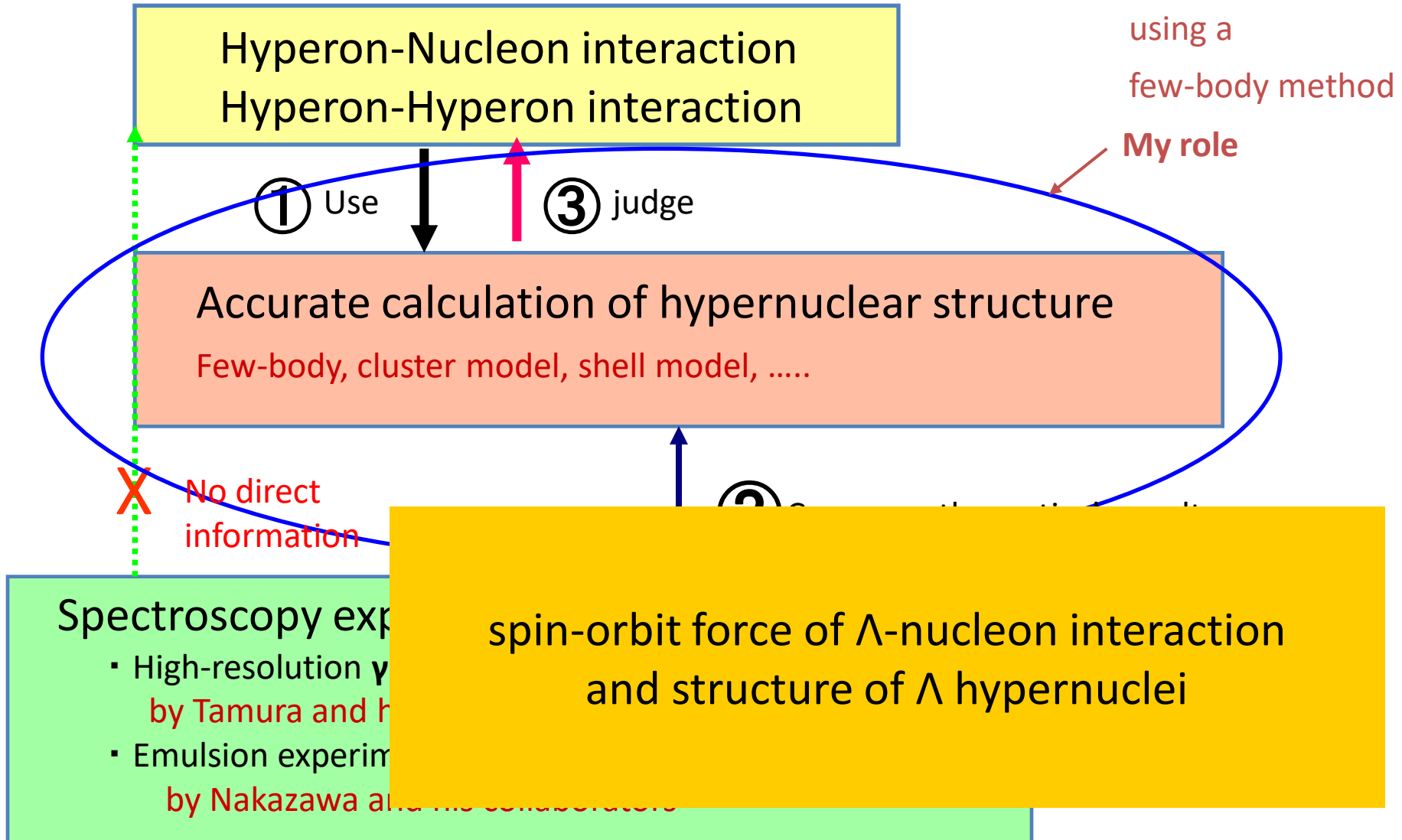
Total number of Nucleon (N) -Nucleon (N) data: 4,000

- Total number of differential cross section Hyperon (Y) -Nucleon (N) data: 40
- **NO** YY scattering data

YN and YY potential models so far proposed (ex. Nijmegen, Julich, Kyoto-Niigata) have large ambiguity.



# Strategy to determine Hyperon-Nucleon and Hyperon-Hyperon interactions from the studies of light Hypernuclear structure



Our few-body calculation method

## Gaussian Expansion Method (GEM) , since 1987,

- A variational method using Gaussian basis functions
- Take all the sets of Jacobi coordinates

Developed by Kyushu Univ. Group,  
Kamimura and his collaborators.

Review article :

E. Hiyama, M. Kamimura and Y. Kino,  
Prog. Part. Nucl. Phys. 51 (2003), 223.

**High-precision calculations** of various 3- and 4-body systems:

Exotic atoms / molecules ,  
3- and 4-nucleon systems,  
multi-cluster structure of light nuclei,

Light hypernuclei,  
3-quark systems,

In order to solve the **Schrödinger equation**, we use **Rayleigh-Ritz variational method** and we **obtain eigen value E and eigen function  $\Psi$** .

$$(\mathbf{H} - \mathbf{E}) \Psi = 0$$

Here, we expand the total wavefunction in terms of a set of  $L^2$ -integrable basis function  $\{\Phi_n : n=1, \dots, N\}$

$$\Psi = \sum_{n=1}^N C_n \Phi_n$$

The Rayleigh-Ritz variational principle leads to a generalized matrix eigenvalue problem.

$$\langle \Phi_i | \mathbf{H} - \mathbf{E} | \sum_{n=1}^N C_n \Phi_n \rangle = 0, \quad (\mathbf{i} = 1, \dots, \mathbf{N})$$

||  
 $\Psi$

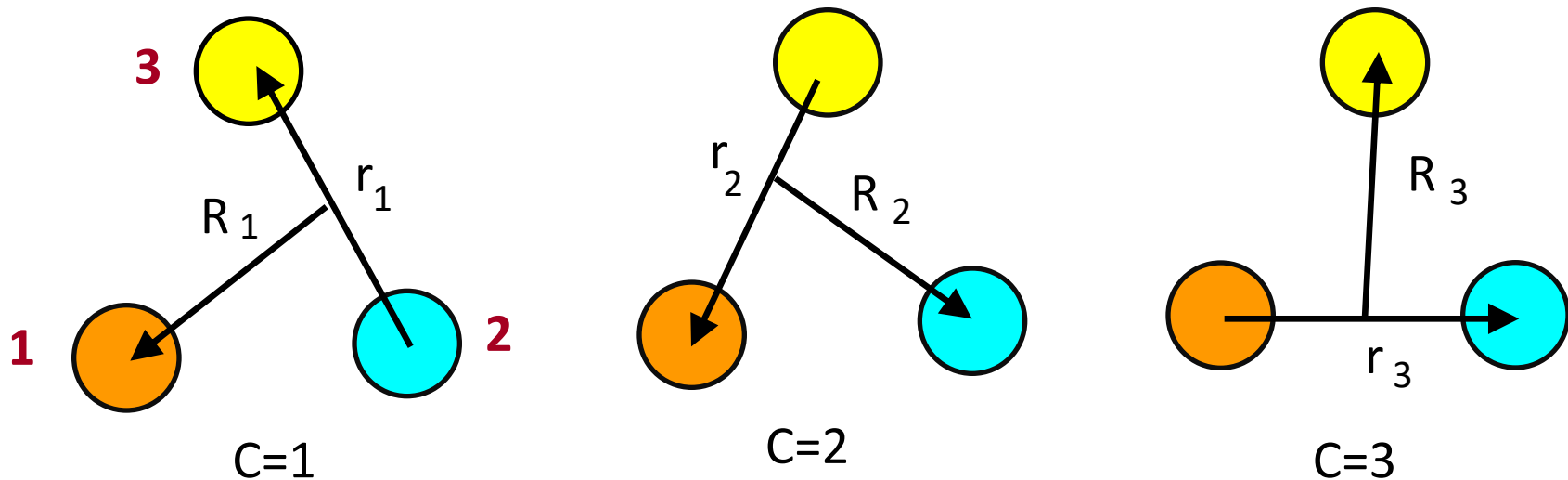
Where the energy and overlap matrix elements are given by

$$H_{in} = \langle \Phi_i | H | \Phi_n \rangle$$

$$N_{in} = \langle \Phi_i | 1 | \Phi_n \rangle$$

**Next, by solving eigenstate problem, we get eigenenergy  $E$  and unknown coefficients  $C_n$ .**

$$\left[ \begin{array}{c} (H_{in}) - E (N_{in}) \end{array} \right] \left[ \begin{array}{c} C_n \end{array} \right] = 0$$

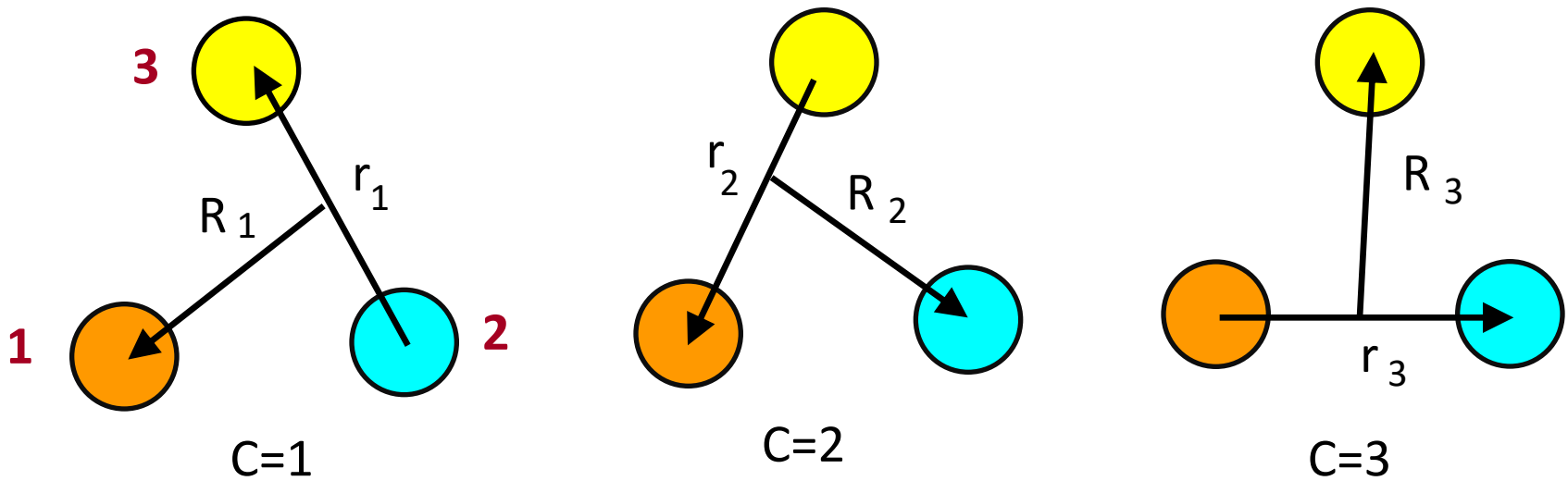


$$(H - E)\Psi_{JM} = 0$$

$$H = T + V_1(r_1) + V_2(r_2) + V_3(r_3)$$

$$T = -\frac{\hbar^2}{2\mu_{r_c}} \nabla_{r_c}^2 - \frac{\hbar^2}{2\mu_{R_c}} \nabla_{R_c}^2 \quad (c = 1, 2, \text{ or } 3)$$

$$\Psi_{JM} = \Phi_{JM}^{(1)}(r_1, R_1) + \Phi_{JM}^{(2)}(r_2, R_2) + \Phi_{JM}^{(3)}(r_3, R_3)$$



$$\Psi_{JM} = \Phi_{JM}^{(1)}(r_1, R_1) + \Phi_{JM}^{(2)}(r_2, R_2) + \Phi_{JM}^{(3)}(r_3, R_3)$$

Basis functions of each Jacobi coordinate

$$\phi_{nl}^{(c)}(r_c) Y_{lm}(\hat{r}_c), \quad \psi_{NL}^{(c)}(R_c) Y_{LM}(\hat{R}_c), \quad (c = 1, 2, 3)$$

$\downarrow$   $(\theta, \phi)$   $\downarrow$   $(\Theta, \Phi)$

$$\Phi_{JM}^{(c)}(r_c, R_c) = \sum_{nl, NL} \underbrace{A_{nl, NL}^{(c)}}_{\uparrow} \phi_{nl}^{(c)}(r_c) \psi_{NL}^{(c)}(R_c) [Y_l(\hat{r}_c) \otimes Y_L(\hat{R}_c)]_{JM}$$

Determined by diagonalizing H

# Benchmark-test calculation of the 4-nucleon bound state

Good agreement among 7 different methods

In the binding energy, r.m.s. radius and wavefunction density

H. KAMADA *et al.*

PHYSICAL REVIEW C **64** 044001

TABLE I. The expectation values  $\langle T \rangle$  and  $\langle V \rangle$  of kinetic and potential energies, the binding energies  $E_b$  in MeV, and the radius in fm.

Method	$\langle T \rangle$	$\langle V \rangle$	$E_b$	$\sqrt{\langle r^2 \rangle}$
FY	102.39(5)	-128.33(10)	-25.94(5)	1.485(3)
<b>GEM</b>	102.30	-128.20	-25.90	1.482
SVM	102.35	-128.27	-25.92	1.486
HH	102.44	-128.34	-25.90(1)	1.483
GFMC	102.3(1.0)	-128.25(1.0)	-25.93(2)	1.490(5)
NCSM	103.35	-129.45	-25.80(20)	1.485
EIHH	100.8(9)	-126.7(9)	-25.944(10)	1.486

very different techniques and the complexity of the nuclear force chosen. Except for NCSM and EIHH, the expectation values of  $T$  and  $V$  also agree within three digits. The NCSM results are, however, still within 1% and EIHH within 1.5% of the others but note that the EIHH results for  $T$  and  $V$  are

Congratulations!! I was very happy

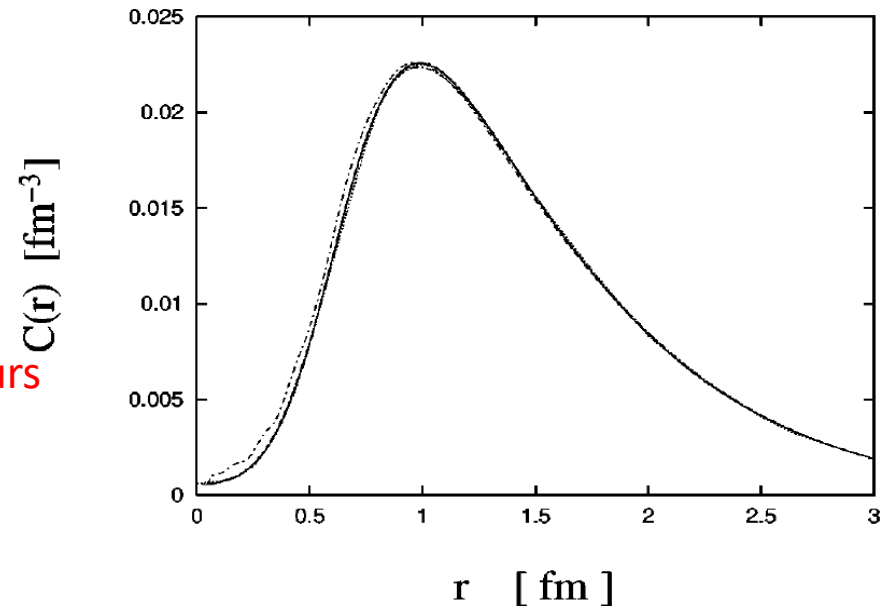
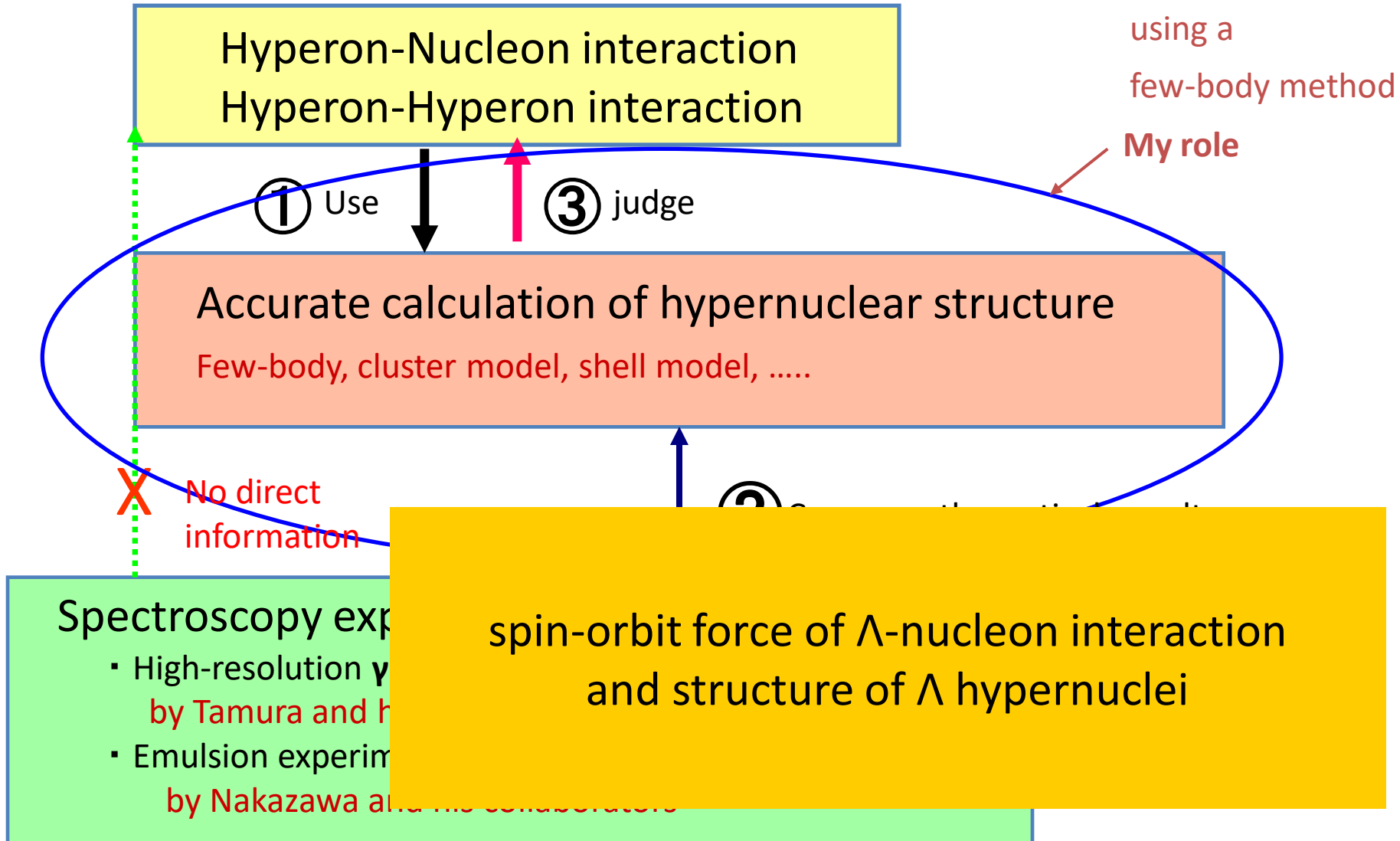


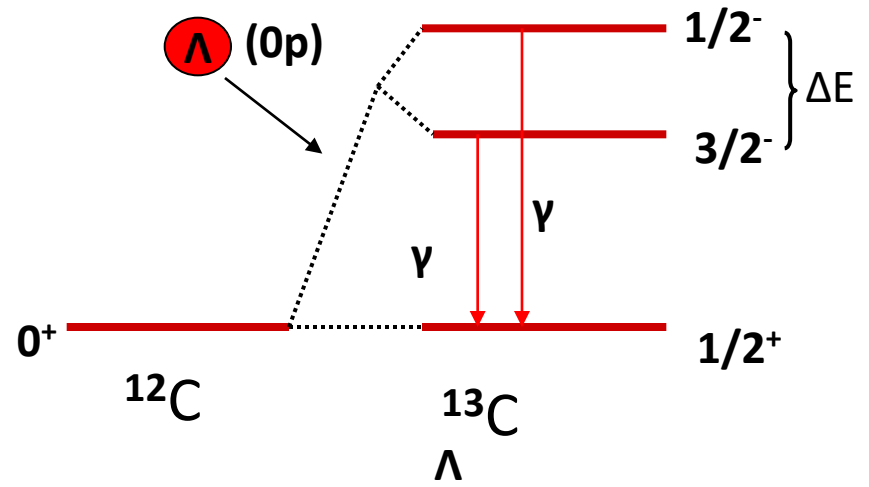
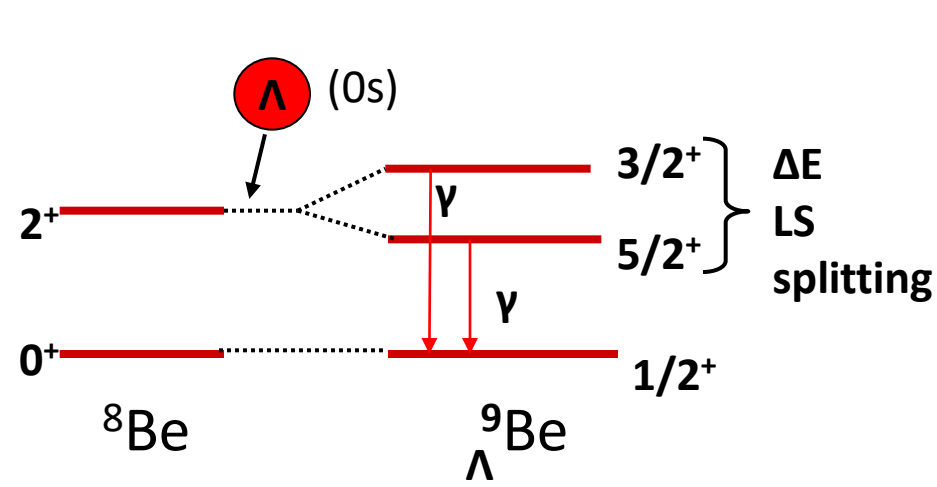
FIG. 1. Correlation functions in the different calculational schemes: EIHH (dashed-dotted curves), FY, CRCGV, SVM, HH, and NCSM (overlapping curves).

# Strategy to determine Hyperon-Nucleon and Hyperon-Hyperon interactions from the studies of light Hypernuclear structure



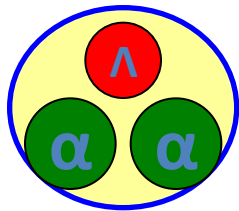
BNL-E930

BNL-E929

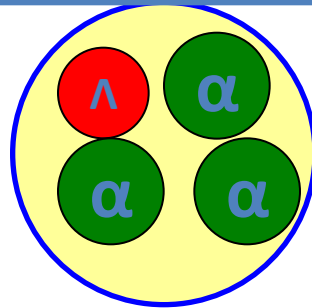


3- and 4-body calculations:

E. Hiyama, M. Kamimura, T. Motoba, T. Yamada and Y. Yamamoto  
 Phys. Rev. Lett. 85 (2000) 270.



$^9\text{Be}$   
 $\Lambda$



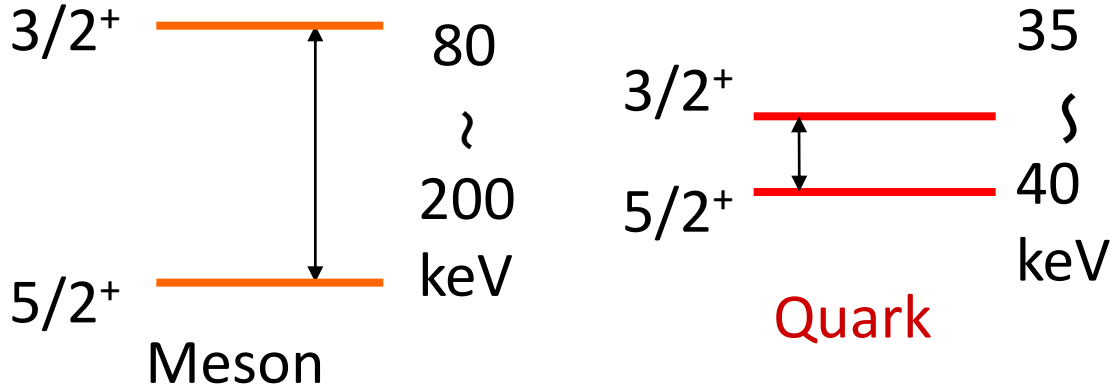
$^{13}\text{C}$   
 $\Lambda$

YN LS force

- 1) **Meson theory** : Nijmegen Model D, F, soft core'97 a – f.
- 2) **Qurak model** : Kyoto-Niigata, FSS

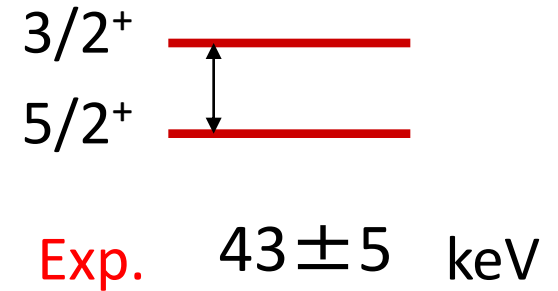
# $\Lambda N$ LS force and ${}^9\text{Be}_\Lambda$ and ${}^{13}\text{C}_\Lambda$

${}^9\text{Be}_\Lambda$



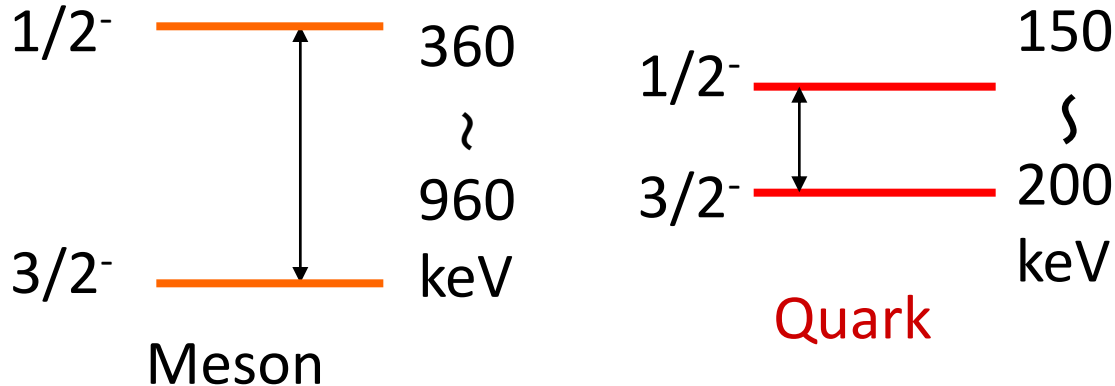
H. Akikawa et al.  
 Phys. Rev. Lett. **88** (2002) 082501;  
 H. Tamura et al. Nucl. Phys. A**754** (2005) 58c

BNL-E930

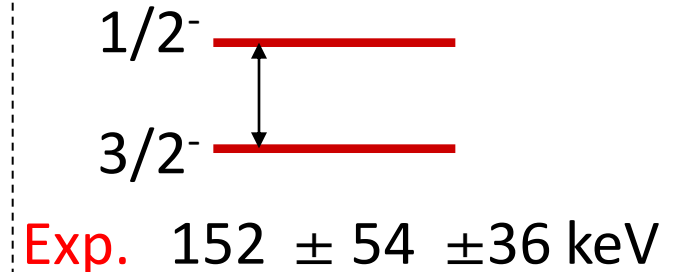


${}^{13}\text{C}_\Lambda$

Nijmegen model D,F  
 Soft core '97a-f



BNL-E929




S. Ajimura et al.  
 Phys. Rev. Lett. **86**, (2001) 4255

# Strategy to determine Hyperon-Nucleon and Hyperon-Hyperon interactions from the studies of light Hypernuclear structure

$\Lambda$ N spin-orbit potentials

Meson based :Nijmegen  
Quark based :Kyoto-Niigata

① Use  ② comparison

③ Quark based interaction gives desirable spin-orbit strength

▪ suggest to Nijmegen group to improve the spin-orbit strength

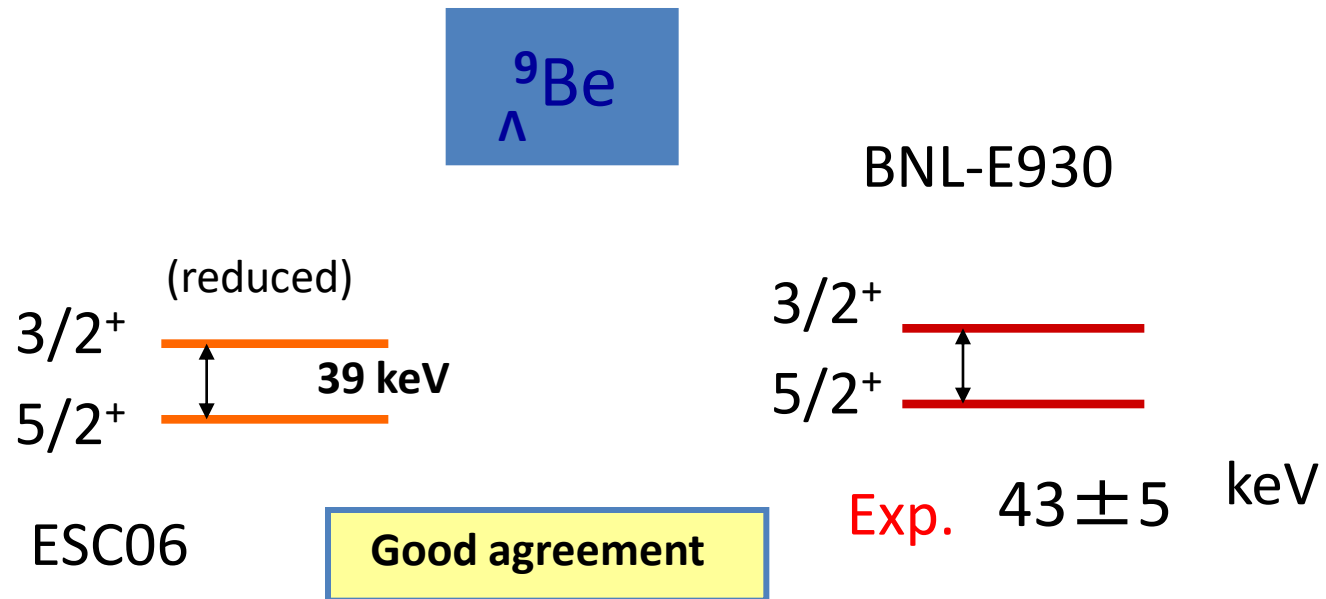
Accurate calculation of energy splitting ( ${}^9_{\Lambda}\text{Be}$  and  ${}^{13}_{\Lambda}\text{C}$ ) using the spin-orbit potentials

Spectroscopy experiments

High-resolution  $\gamma$ -ray spectroscopy experiment in  ${}^9_{\Lambda}\text{Be}$  and  ${}^{13}_{\Lambda}\text{C}$   
by Tamura and his collaborators  
by Kishimoto and his collaborators

# LS splitting in ${}^9\text{Be}_\Lambda$

Recently, a new spin-orbit  $\Lambda\text{N}$  interaction based on meson theory, extended soft core potential 06 (ESC06) by Nijmegen group



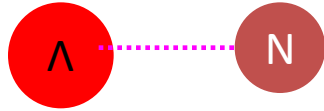
Hiyama (2007)

H. Akikawa et al.  
Phys. Rev. Lett. 88,(2002)82501;  
H. Tamura et al.  
Nucl. Phys. A754,58c(2005)

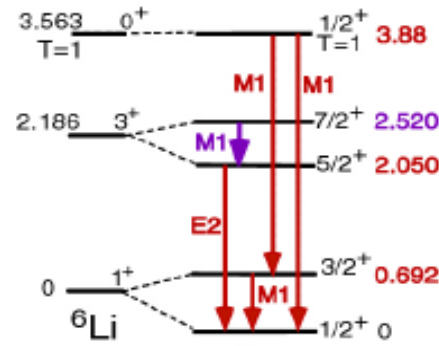
Desirable strength of  $\Lambda\text{N}$  spin-orbit force was determined.

# Hypernuclear $\gamma$ -ray data since 1998

Taken by Tamura

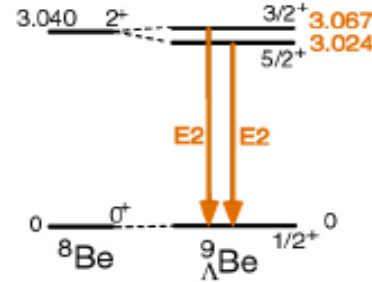


${}^7\text{Li} (\pi^+, K^+\gamma)$  KEK E419



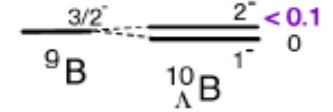
${}^7_{\Lambda}\text{Li}$  PRL 84 (2000) 5963  
 PRL 86 (2001) 1982  
 PLB 579 (2004) 258  
 PRC 73 (2006) 012501

${}^9\text{Be} (K^-, \pi^-\gamma)$  BNL E930('98)



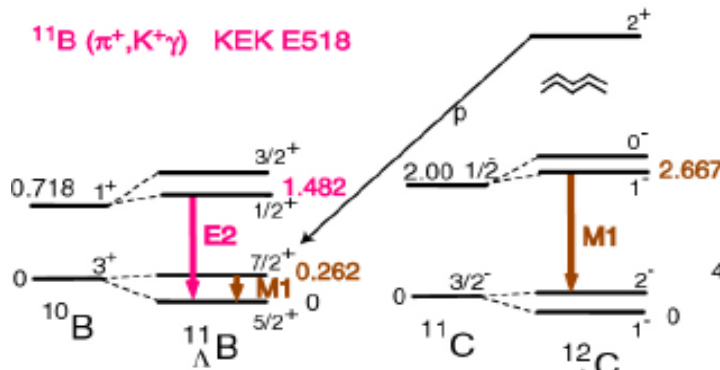
PRL 88 (2002) 082501  
 NPA 754 (2005) 58c

${}^{10}\text{B} (K^-, \pi^-\gamma)$  BNL E930('01)



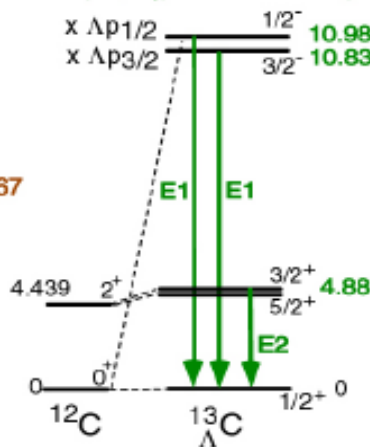
NPA 754 (2005) 58c

${}^{12}\text{C} (\pi^+, K^+\gamma)$  KEK E566



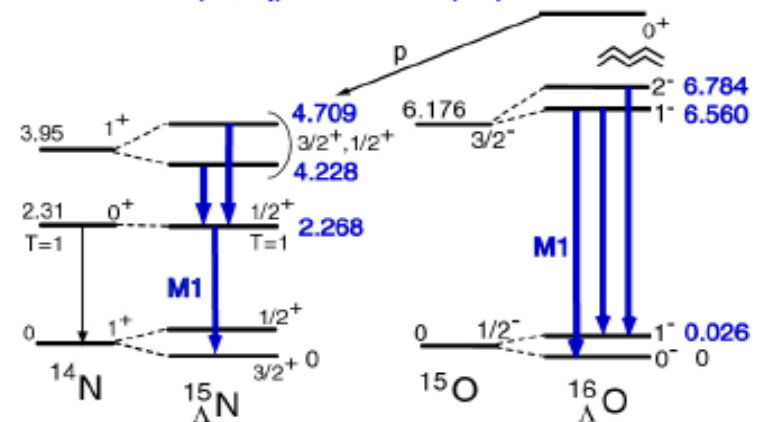
NPA 754 (2005) 58c

${}^{13}\text{C} (K^-, \pi^-\gamma)$  BNL E929 (Nal)



PRL 86 (2001) 4255  
 PRC 65 (2002) 034607

${}^{16}\text{O} (K^-, \pi^-\gamma)$  BNL E930('01)



PRL 93 (2004) 232501

$$V_{\Lambda N} = V_0 + \sigma_{\Lambda} \cdot \sigma_N V_{\sigma\sigma} + \mathbf{L} \cdot (\mathbf{s}_{\Lambda} + \mathbf{s}_N) V_{\text{SLS}} + \mathbf{L} \cdot (\mathbf{s}_{\Lambda} - \mathbf{s}_N) V_{\text{ALS}} + S_{12} V_{\text{tensor}} + \dots$$

- Millener (p-shell model),
- Hiyama (few-body)

In  $S = -1$  sector,  
what are the open questions in  $\Lambda N$  interaction?

(1) Charge symmetry breaking

(2)  $\Lambda N - \Sigma N$  coupling

SKIP

### J-PARC : Day-1 experiment

- E13 “ $\gamma$ -ray spectroscopy of light hypernuclei”  
by Tamura and his collaborators



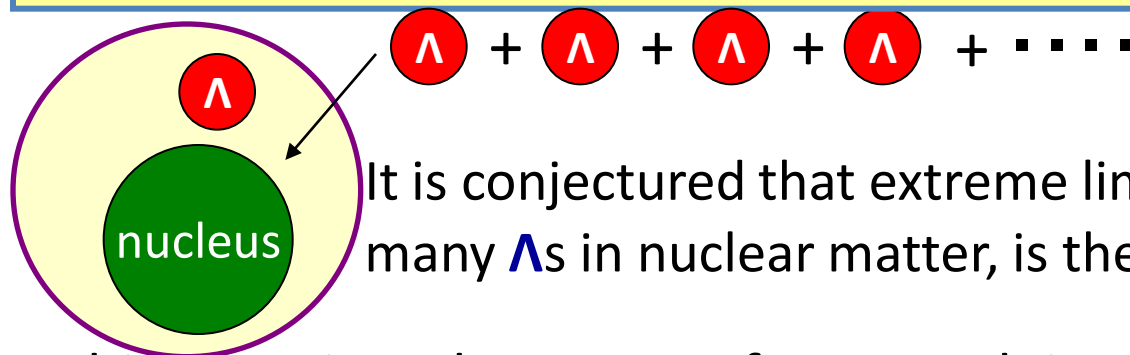
- E10 “Study on  $\Lambda$ -hypernuclei with the doubleCharge-Exchange reaction”  
by Sakaguchi , Fukuda and his collaborators



S=-2 hypernuclei  
and  
YY interaction

So far, we have discussed about single  $\Lambda$  hypernuclei.

What is the structure when one or more  $\Lambda$ s are added to a nucleus?



It is conjectured that extreme limit, which includes many  $\Lambda$ s in nuclear matter, is the **core of a neutron star**.

In this meaning, the sector of  $S=-2$  nuclei , double  $\Lambda$  hypernuclei and  $\Xi$  hypernuclei is just the entrance to the **multi-strangeness** world.

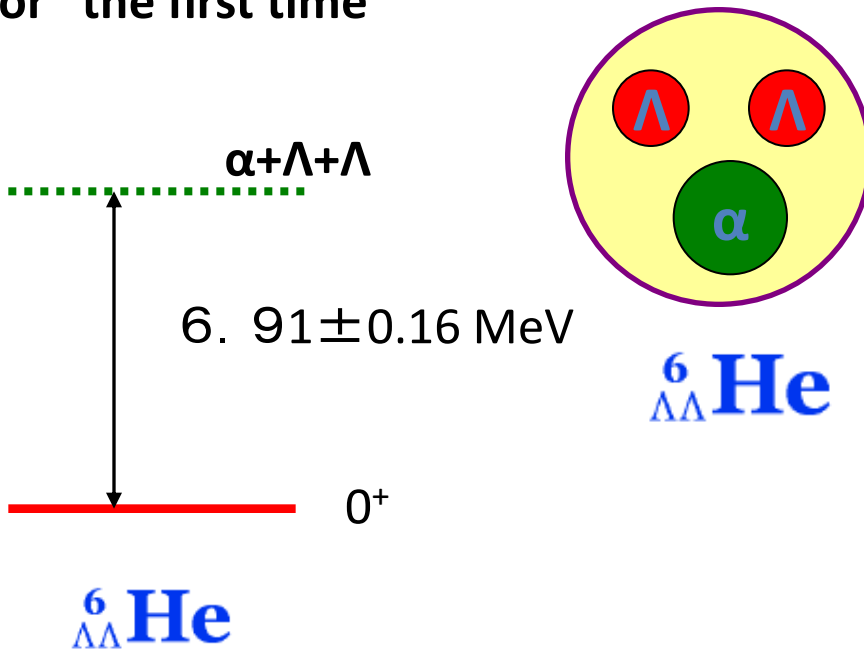
**However, we have hardly any knowledge of the  $YY$  interaction because there exist no  $YY$  scattering data.**

Then, in order to understand the  $YY$  interaction, it is crucial to study the structure of double  $\Lambda$  hypernuclei and  $\Xi$  hypernuclei. But, up to 2000, we have no observed bound double  $\Lambda$  hypernuclei.

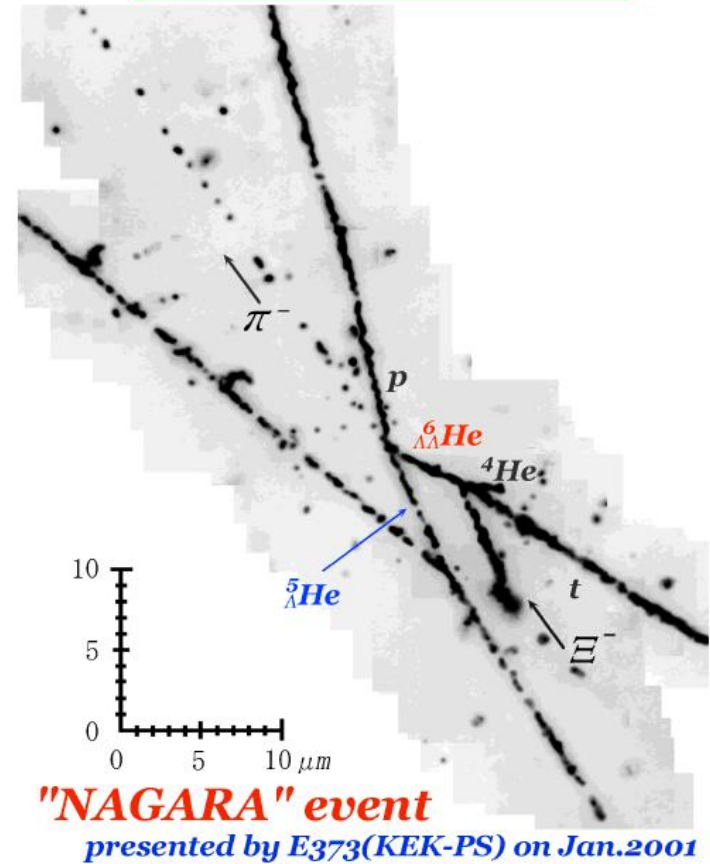
In 2001, the epoch-making data has been reported by the KEK-E373 experiment.

Observation of  ${}^6_{\Lambda\Lambda}\text{He}$

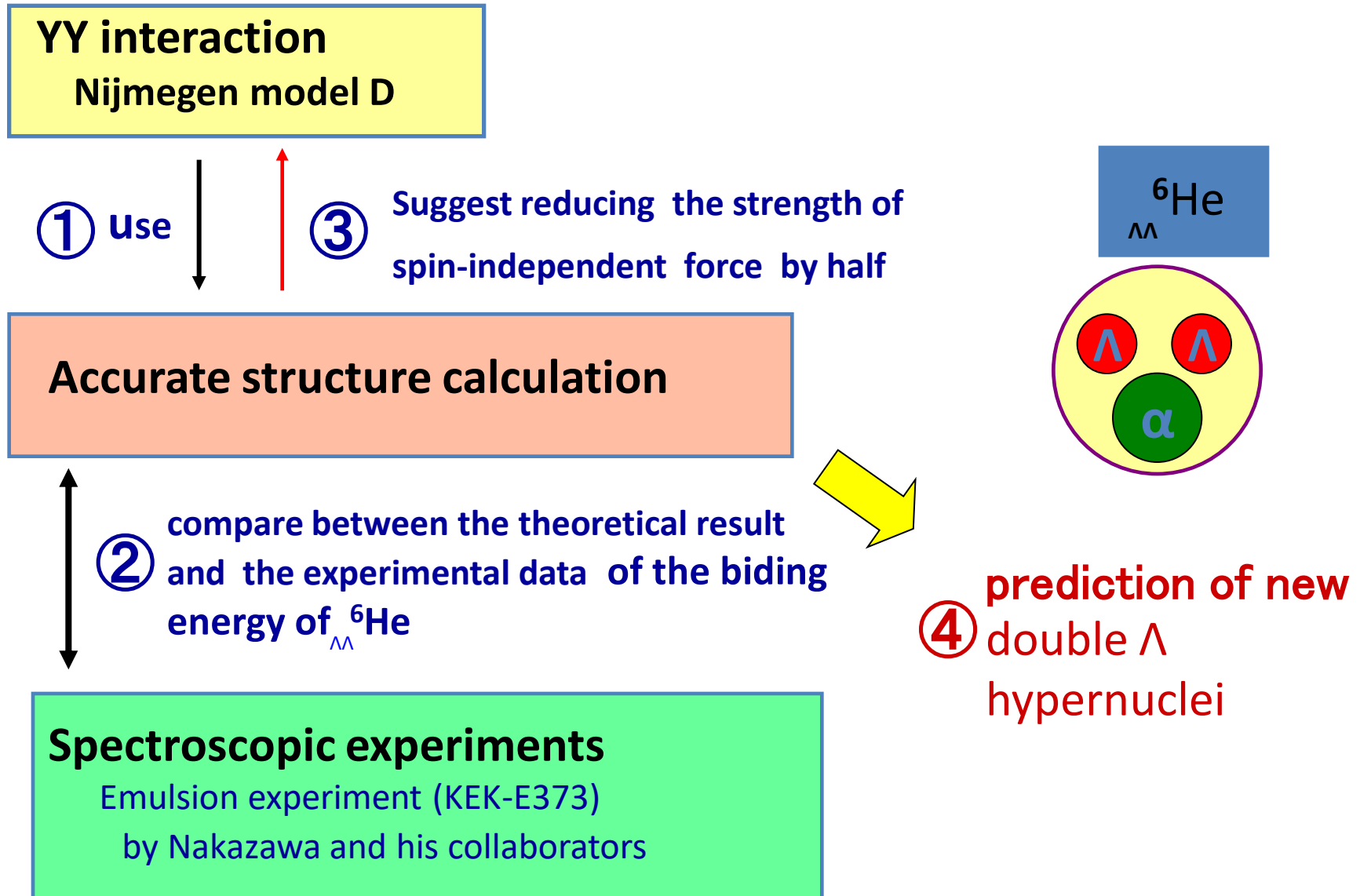
Uniquely identified without ambiguity for the first time



${}^6_{\Lambda\Lambda}\text{He}$  double-hypernucleus  
 Unique interpretation!!



# Strategy of how to determine YY interaction from the study of light hypernuclear structure

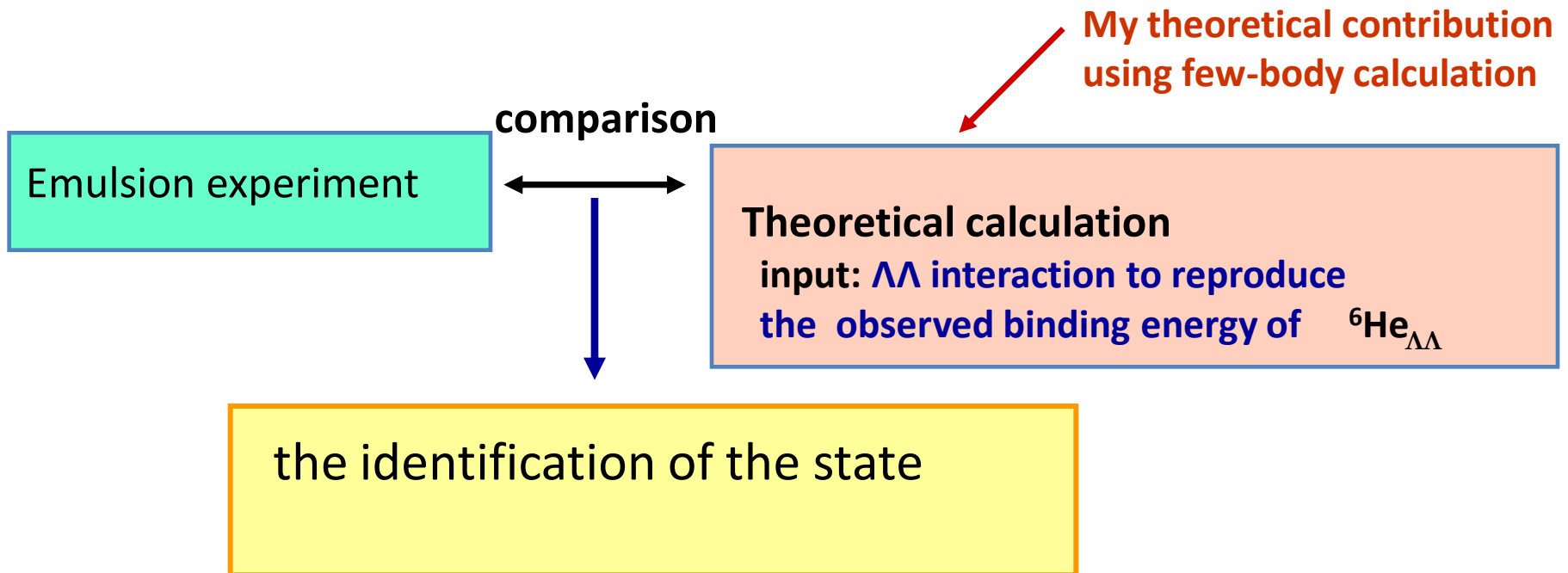


KEK-E373 experiment analysis is still in progress.

- E07 Approved proposal at J-PARC  
“Systematic Study of double strangeness systems at J-PARC”  
by Nakazawa and his collaborators

It is difficult to determine

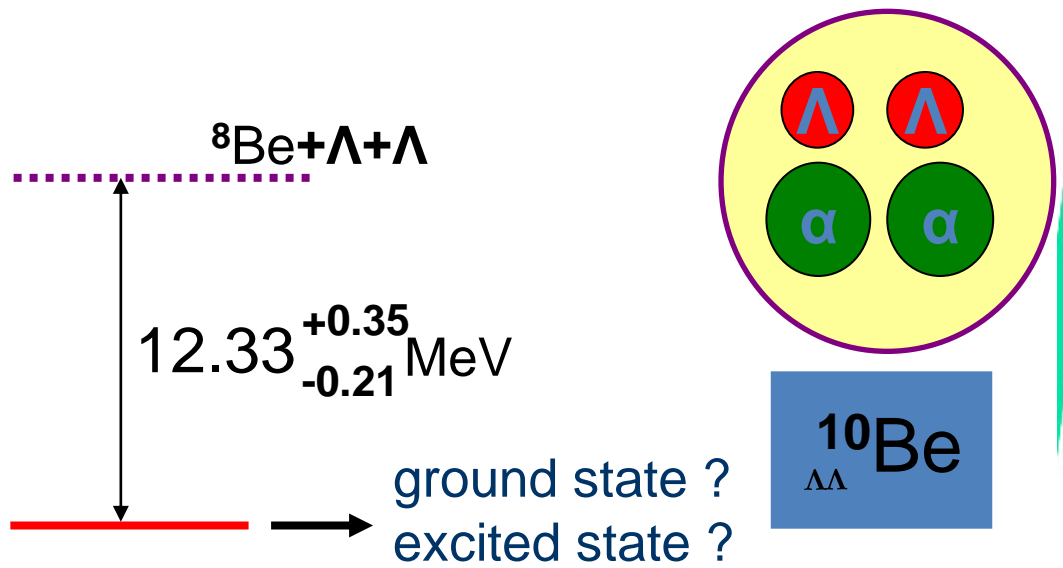
- (1) spin-parity
- (2) whether the observed state is the ground state or an excited state



Successful example to determine spin-parity of double  $\Lambda$  hypernucleus --- Demachi-Yanagi event for

$^{10}_{\Lambda\Lambda}\text{Be}$

Observation of  $^{10}_{\Lambda\Lambda}\text{Be}$  --- KEK-E373 experiment



**Demachi-yanagi event**

2001

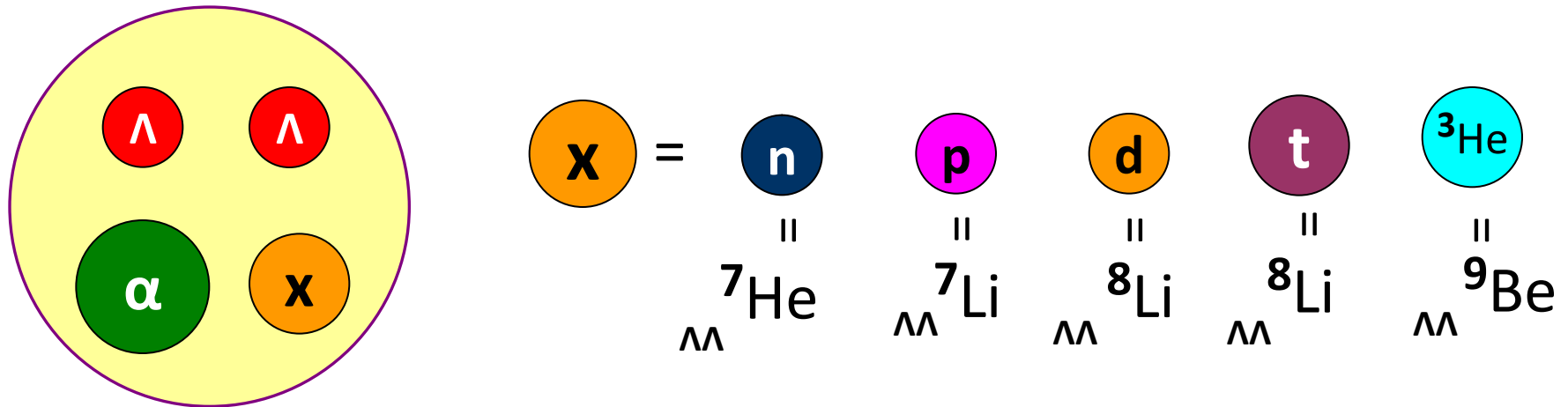
\*two body case at point A  
 $E^- + ^{12}\text{C} \rightarrow ^{10}\text{Be} + t$  or  $^{10}\text{Be}^* + t$   
 $\Delta B_{\Lambda\Lambda} : -1.14 \pm 0.19$  or  $+1.86 \pm 0.19 \text{ MeV}$   
 $B_{\Lambda\Lambda} : 12.29 \pm 0.17$  (excited) MeV  
 $15.29 \pm 0.17$  (ground) MeV

\*three body case at point A  
 1)  $E^- + ^{14}\text{N} \rightarrow ^{10}\text{B} + p + n$   
 $\Delta B_{\Lambda\Lambda} : +1.47^{+2.4}_{-0.7} \text{ MeV}$

Demachi-Yanagi event

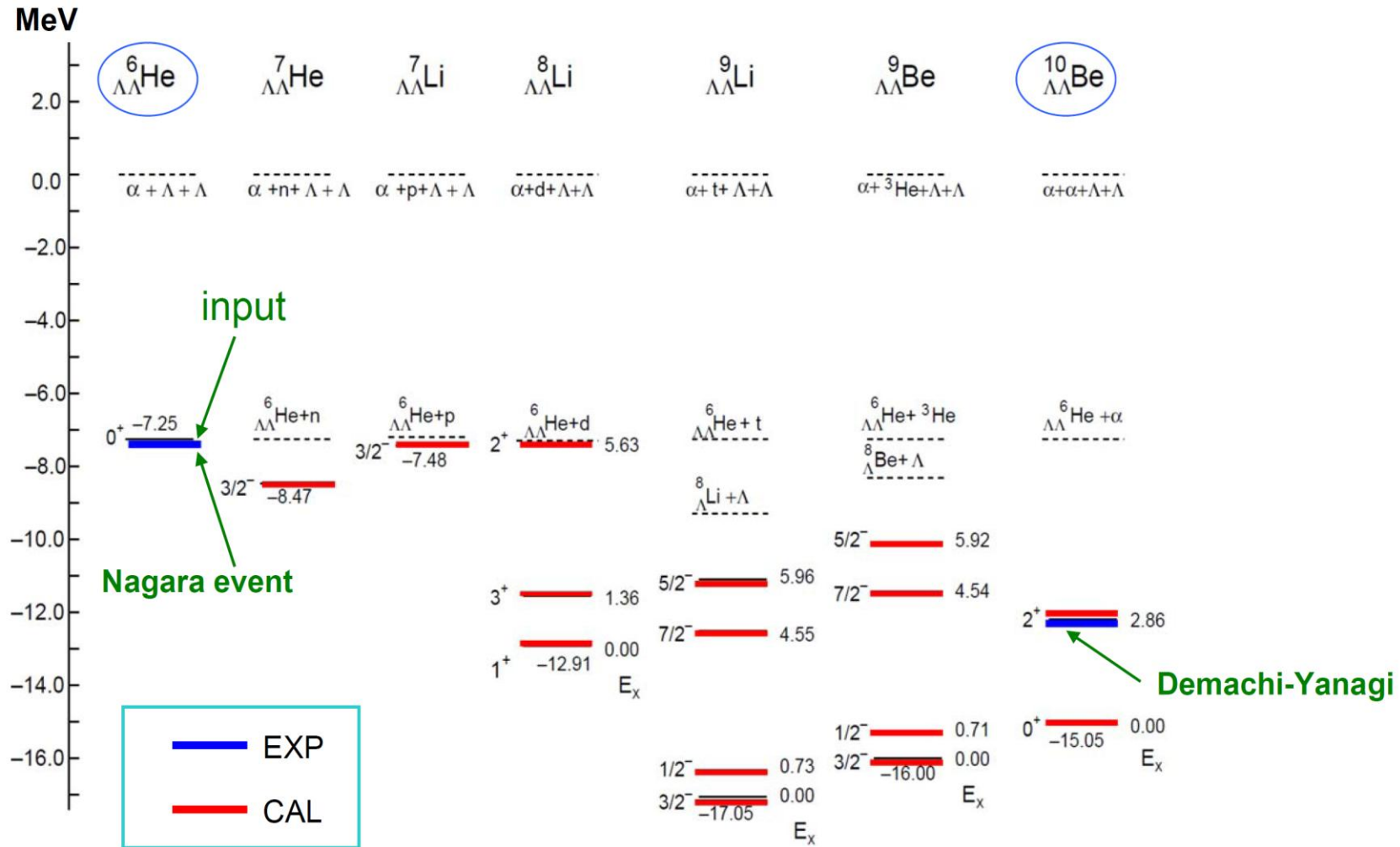
Hoping to observe new double  $\Lambda$  hypernuclei in future experiments, I predicted level structures of these double  $\Lambda$  hypernuclei within the framework of the  $\alpha+x+\Lambda+\Lambda$  4-body model.

E. Hiyama, M. Kamimura, T. Motoba, T. Yamada and Y. Yamamoto  
 Phys. Rev. C66, 024007 (2002)



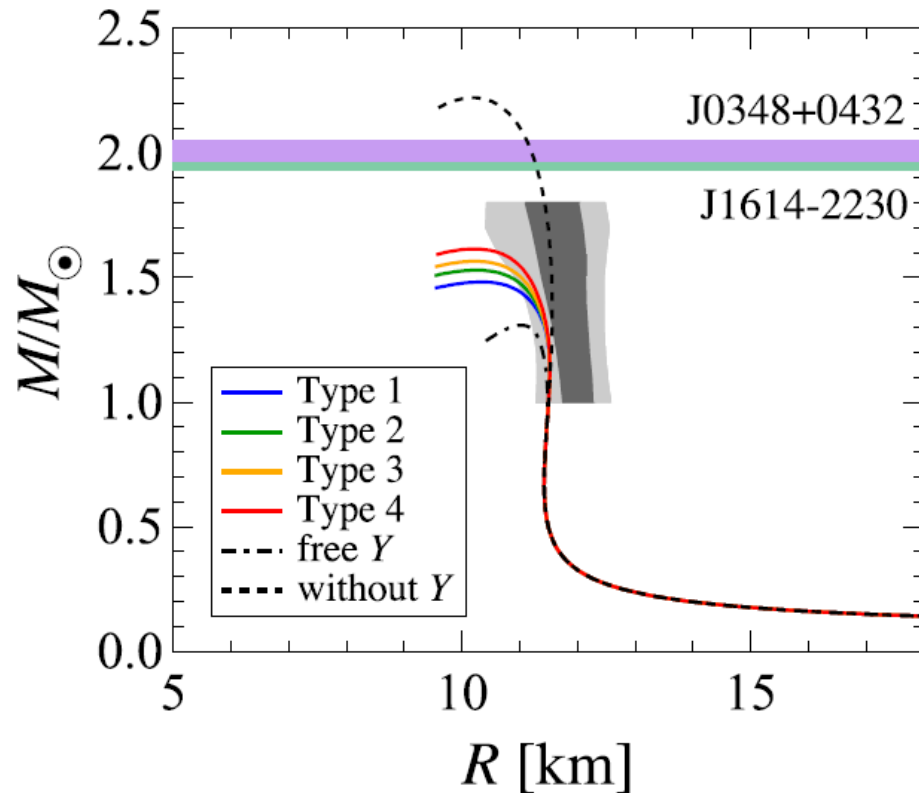
# Spectroscopy of $\Lambda\Lambda$ -hypernuclei

E. Hiyama, M. Kamimura, T. Motoba, T. Yamada and Y. Yamamoto  
 Phys. Rev. 66 (2002), 024007



I have been looking forward to having new data in this mass-number region.

In  $\Lambda\Lambda$  interaction, what is important to study?  
odd-state of  $\Lambda\Lambda$  interaction



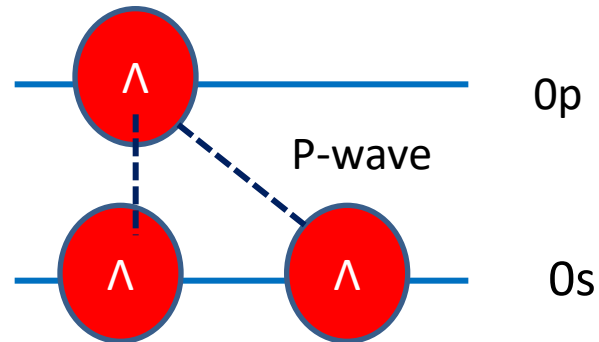
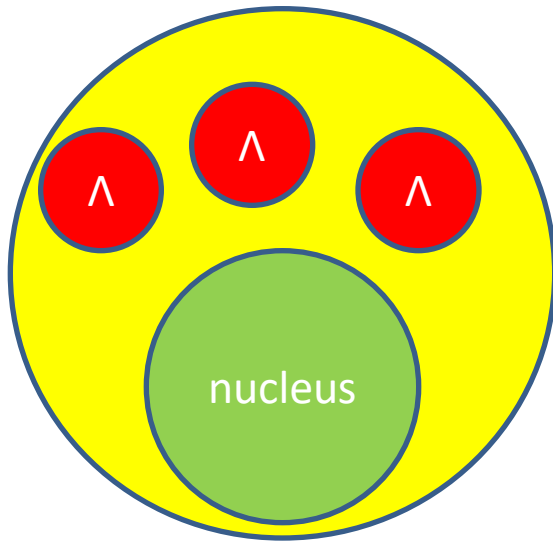
H. Togashi, E.H.,  
Y. Yamamoto, and M. Takano,  
PRC93, 035808 (2016).

TYPE1:attractive  $\rightarrow$  TYPE2:less attractive  $\rightarrow$  TYPE4:repulsive



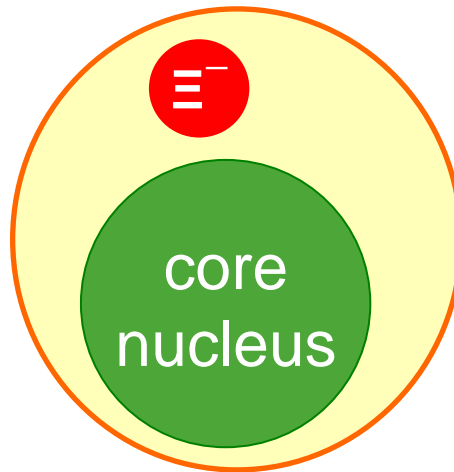
Similar with odd-state of  
 $\Lambda N$  interaction

To obtain odd-state (p-wave) interaction of  $\Lambda\Lambda$  interaction, I propose you to perform search experiment of triple  $\Lambda$  hypernucleus.



But, I do not know to produce triple  $\Lambda$  hypernuclei currently.

$\Xi$ hypernuclei



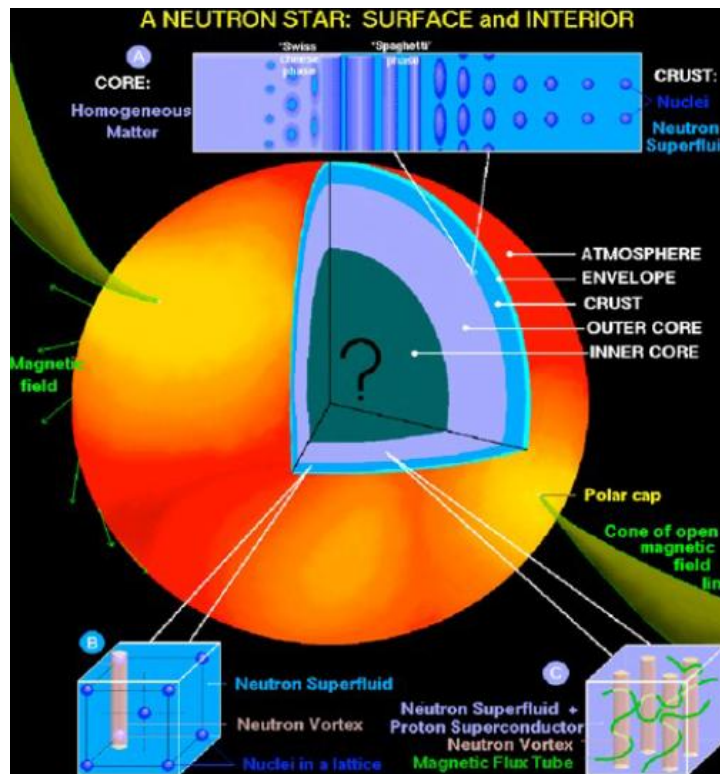
For the study of  $\Xi$ N interaction, it is important to study the structure of  $\Xi$  hypernuclei.

This is related to maximum mass of neutron star.

# Major goals of hypernuclear physics

To understand structure of few-body and many-body nuclei (with strangeness) and also life of star

## Core of neutron star



We need to understand 2 solar mass for the maximum of neutron stars. Now, it is one of hot topics in hypernuclear physics.

Question:  $\Xi$  particles are included in neutron star?

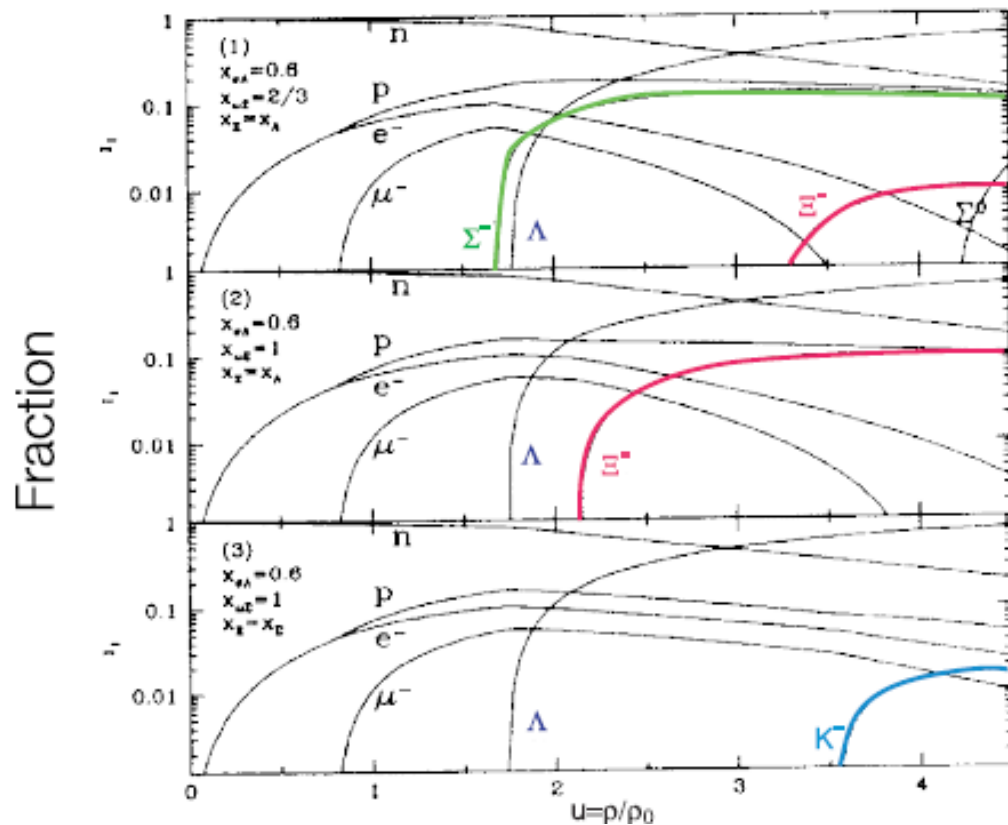
$V_{\Xi N}$ : attractive  $\Rightarrow$  Yes, include

$V_{\Xi N}$ : repulsive  $\Rightarrow$  No,

# $\Xi$ hypernuclei potential ?

- $\Lambda, \Sigma^-, \Xi^-, \Sigma^0$  in Neutron Star Core ?

Chemical Potential:  $\mu_B = m_B + \frac{k_F^2}{2m_B} + U(k_F)$



$$U_{\Sigma} < 0, U_{\Xi} < 0$$

$$U_{\Sigma} > 0, U_{\Xi} < 0$$

$$U_{\Sigma} > 0, U_{\Xi} > 0$$

- So far, we do not know whether  $\Xi N$  interaction is attractive or repulsive. It is difficult to perform  $\Xi N$  scattering data.

Then, it is important to observe find bound  $\Xi$  hypernuclei.

If there is bound  $\Xi$  bound hypernuclei.  $\Rightarrow \Xi$  — nucleus potential should be attractive.

So far, there was NO observed bound  $\Xi$  hypernucleus.

# The first evidence of a deeply bound state of $\Xi^- - ^{14}\text{N}$ system

K. Nakazawa<sup>1,\*</sup>, Y. Endo<sup>1</sup>, S. Fukunaga<sup>2</sup>, K. Hoshino<sup>1</sup>, S. H. Hwang<sup>3</sup>, K. Imai<sup>3</sup>, H. Ito<sup>1</sup>, K. Itonaga<sup>1</sup>, T. Kanda<sup>1</sup>, M. Kawasaki<sup>1</sup>, J. H. Kim<sup>4</sup>, S. Kinbara<sup>1</sup>, H. Kobayashi<sup>1</sup>, A. Mishina<sup>1</sup>, S. Ogawa<sup>2</sup>, H. Shibuya<sup>2</sup>, T. Sugimura<sup>1</sup>, M. K. Soe<sup>1</sup>, H. Takahashi<sup>5</sup>, T. Takahashi<sup>5</sup>, K. T. Tint<sup>1</sup>, K. Umehara<sup>1</sup>, C. S. Yoon<sup>4</sup>, and J. Yoshida<sup>1</sup>

<sup>1</sup>Physics Department, Gifu University, 1-1 Yanagido, Gifu 501-1193, Japan

<sup>2</sup>Department of Physics, Toho University, Funabashi 274-8510, Japan

<sup>3</sup>Advanced Science Research Center, JAEA, Tokai 319-1195, Japan

<sup>4</sup>Department of Physics, Gyeongsang National University, Jinju 660-701, Korea

<sup>5</sup>Institute of Particle and Nuclear Studies, KEK, Tsukuba 305-0801, Japan

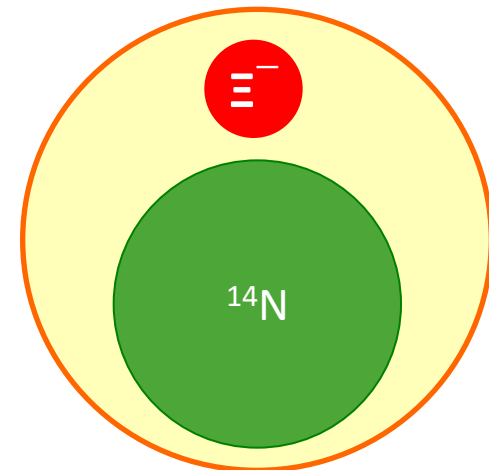
\*E-mail: nakazawa@gifu-u.ac.jp

Received October 27, 2014; Revised December 25, 2014; Accepted January 9, 2015; Published March 5, 2015

$^{14}\text{N}-\Xi^-$

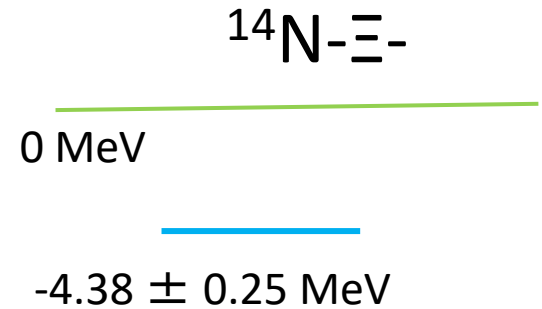
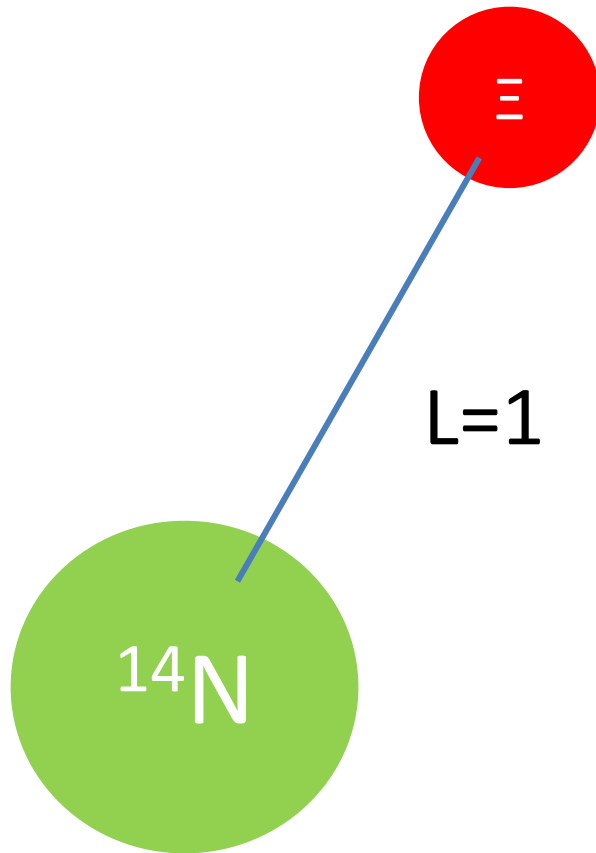
0 MeV

$-4.38 \pm 0.25$  MeV



Kiso event

In 2015, we observed bound  $\Xi$  hypernucleus, for the first time in the world. Now, we understood that  $\Xi\text{N}$  interaction should be attractive.



From this experimental data, we obtain information on the p-wave state of  $\Xi\text{N}$  interaction.

This is very important information. But, we want to obtain information on s-wave of  $\Xi\text{N}$  interaction.

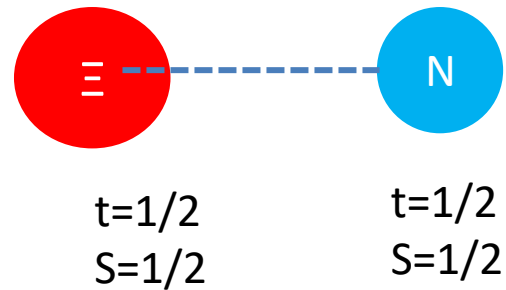
In  $\Xi$  hypernuclei,

Hot topics is

- What kinds of  $\Xi N$  interaction, we should extract information on?
- What kinds of  $\Xi$  hypernuclei are suited for this purpose?
- How do we produce these  $\Xi$  hypernuclei?

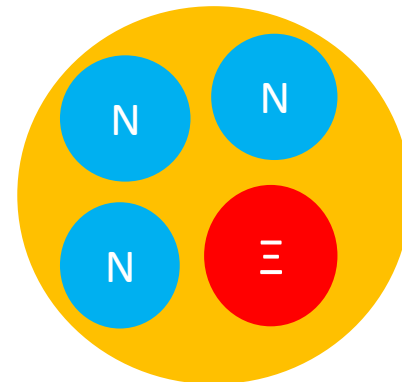
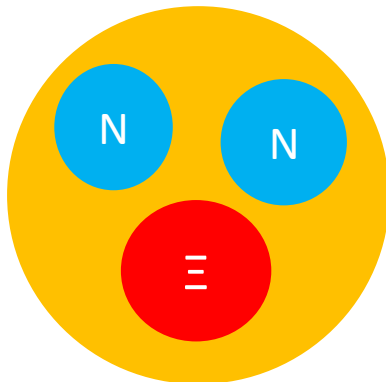
Next, we want to obtain information on s-wave of  $\Xi N$  interaction.

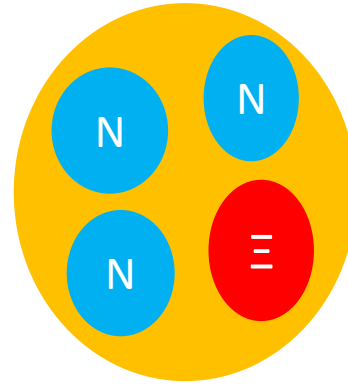
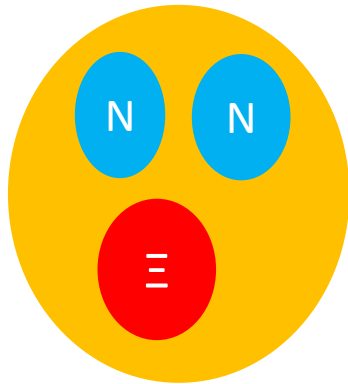
$\Xi$ N interaction:  $T=0, S=0$   
 $T=0, S=1$   
 $T=1, S=0$   
 $T=1, S=1$



We want to know which partial wave is attractive or repulsive.

The suited systems to study are s-shell  $\Xi$  hypernuclei such as  $NN\Xi$  and  $NNN\Xi$  systems.





I show my new results of these light systems.

NN interaction: AV8 potential

ΞN interaction :

Nijmegen extended soft core potential (ESC08c)

Realistic potential (only ΞN channel)

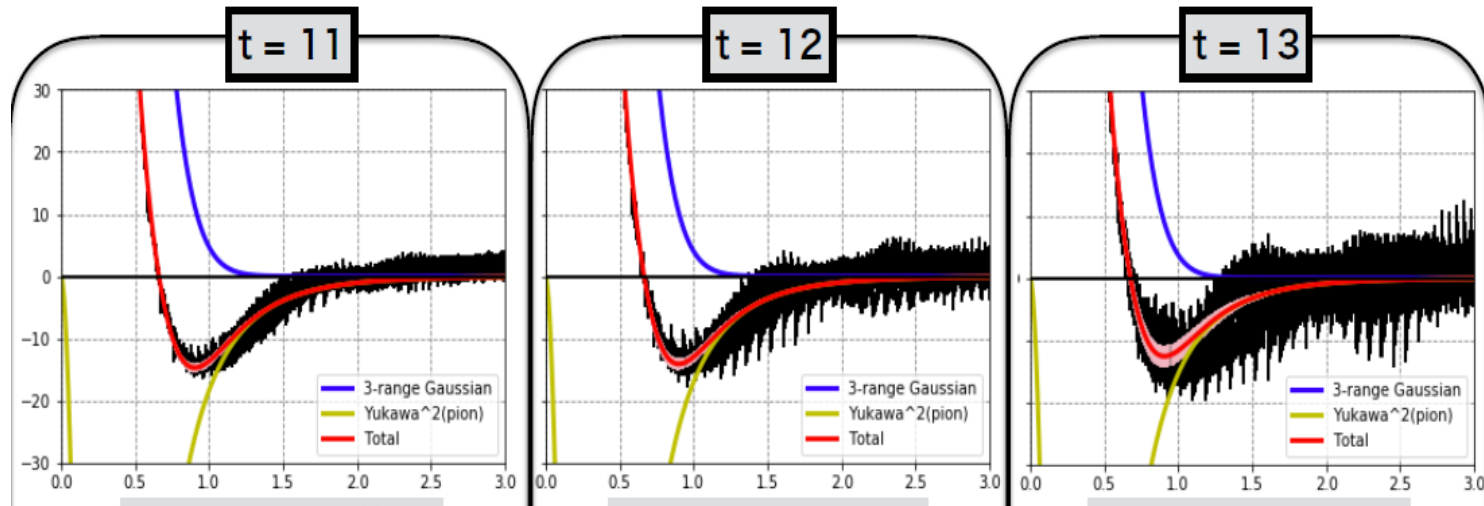
ΞN interaction by HAL collaboration (Lattice QCD calculation)

The potential was made by K. Sasaki and Miyamoto.

# HAL potential

$$V_{\Xi N} = V_0(r) + (\sigma_{\Xi} \cdot \sigma_N) V_S(r) + (\tau_{\Xi} \cdot \tau_N) V_t(r) + (\sigma_{\Xi} \cdot \sigma_N)(\tau_{\Xi} \cdot \tau_N) V_{ts}(r)$$

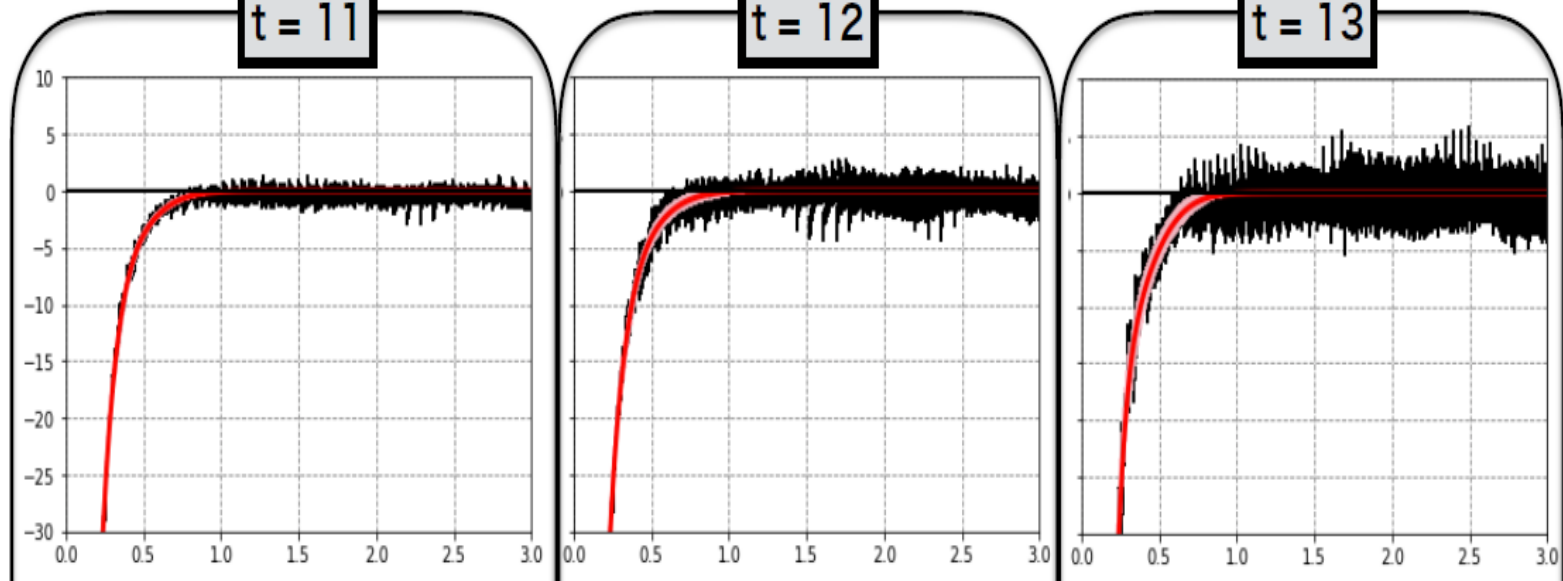
All terms are central parts only.



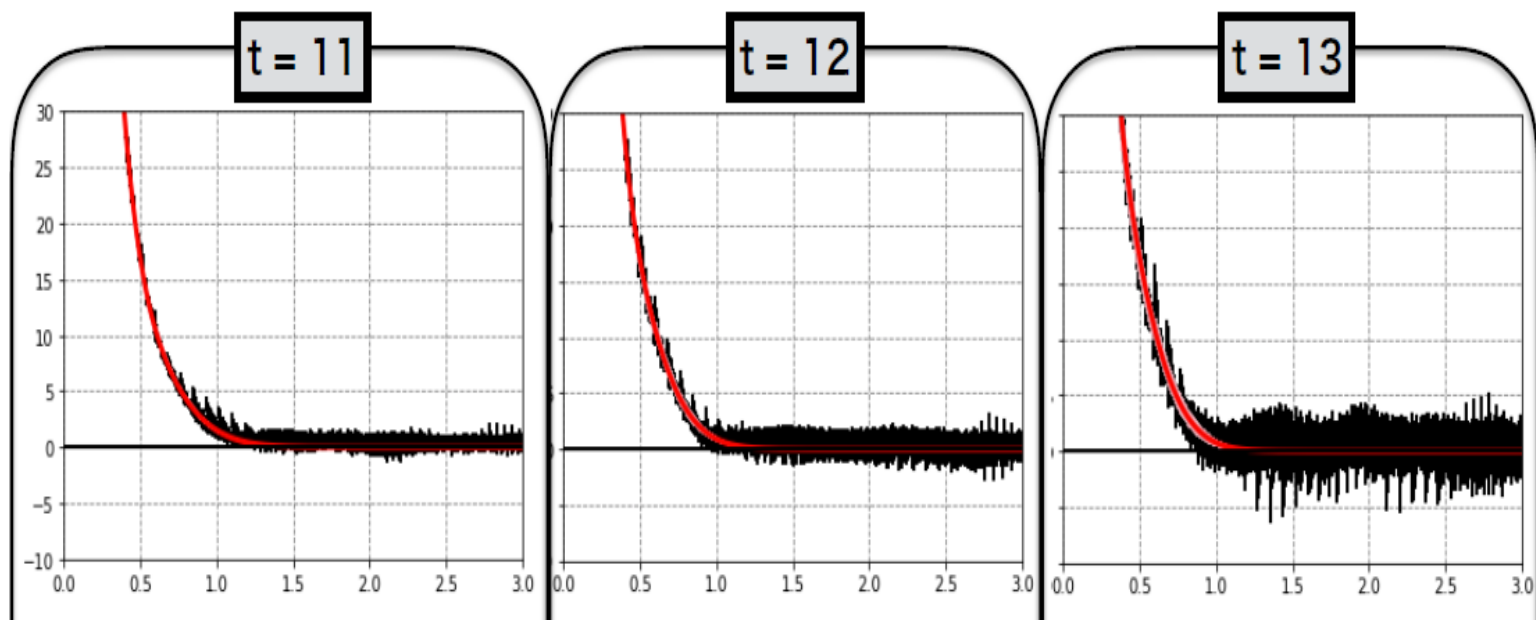
$$V_0(r)$$

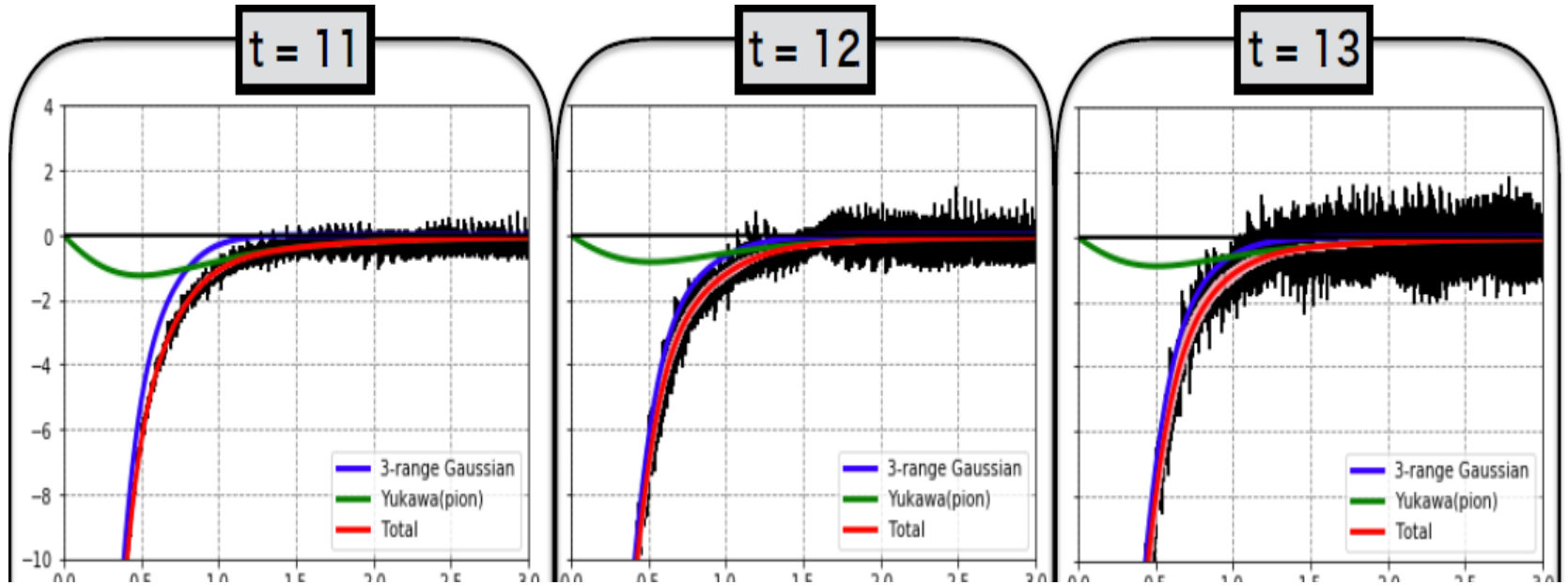
In HAL potential, the statistical errors are NOT included.

$$(\sigma_{\Xi} \cdot \sigma_N) V_S(r)$$



$$(\tau_{\Xi} \cdot \tau_N) V_t(r)$$





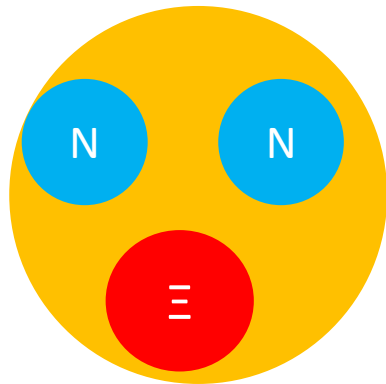
$$(\sigma_{\Xi} \cdot \sigma_N)(\tau_{\Xi} \cdot \tau_N)V_{ts}(r)$$

## Property of the spin- and isospin-components of ESC08 and HAL

V(T,S)	ESC08c	HAL
T=0, S=1	strongly attractive	Weakly attractive
T=0, S=0	weakly repulsive	Strongly attractive
T=1, S=1	strong attractive	Weakly attractive
T=1, S=0	weakly repulsive	Weakly repulsive

Although the spin- and isospin-components of these two models are very different between them.

It is interesting to see the difference in the energy spectra in s-shell  $\Xi$  hypernuclei.



$T=1/2, J=1/2^+$  and  $J=3/2^+$

ESC08c

0 MeV      unbound      d+ Ξ

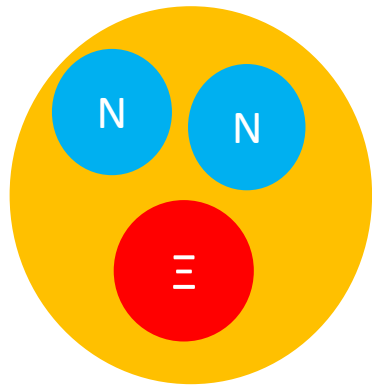
---

0 MeV      d+ Ξ

---

$J=3/2^+$       -7.29 MeV

---



$T=1/2, J=1/2^+$  and  $J=3/2^+$

HAL potential

0 MeV

$d + \Xi$

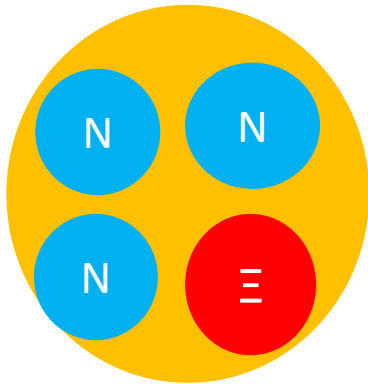
0 MeV

$d + \Xi$

$J=1/2^+$

No bound state

$J=3/2^+$



T=1 state

0 MeV

$3N+\Xi$

-3.54 MeV

$0^+$

-10.08 MeV

$1^+$

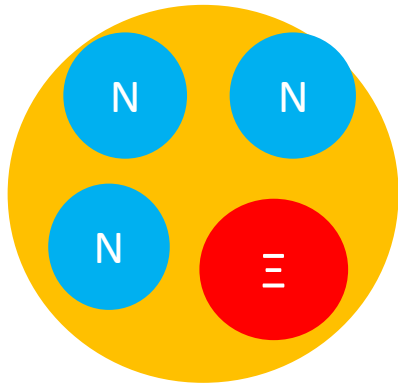
ESC08c

0 MeV

$3N+\Xi$

No bound state

HAL potential



T=0 state

In HAL potential, the statistical errors are NOT included.

0 MeV

$3N+\Xi$

0 MeV

$3N+\Xi$

$-0.05 \text{ MeV} \sim -0.5 \text{ MeV}^{1+}$

-10.18 MeV

$J=1^+$

ESC08c

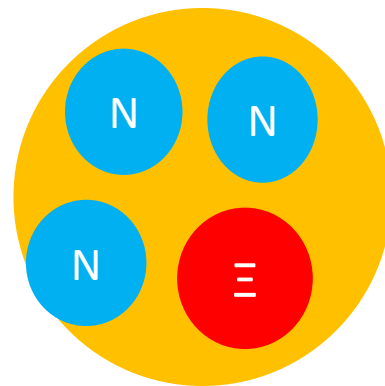
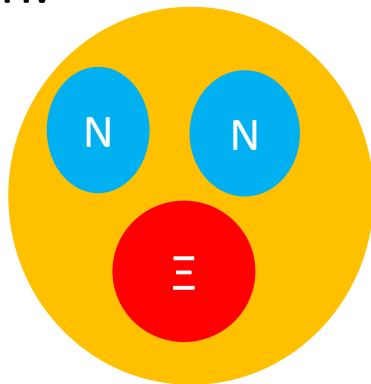
HAL potential

## Summary

To investigate spin- isospin dependence on  $NN\Xi$  and  $NNN\Xi$  system, we calculated these systems using two types  $\Xi N$  interactions, ESC08c and HAL potential.

The energies of these systems are strongly dependent on the  $\Xi N$  potential employed.

To investigate this fact, I suggest to produce these  $\Xi$  hypernuclei using  ${}^3\text{He}$  and  ${}^4\text{He}$  target by  $(K^+, K^-)$  reaction.

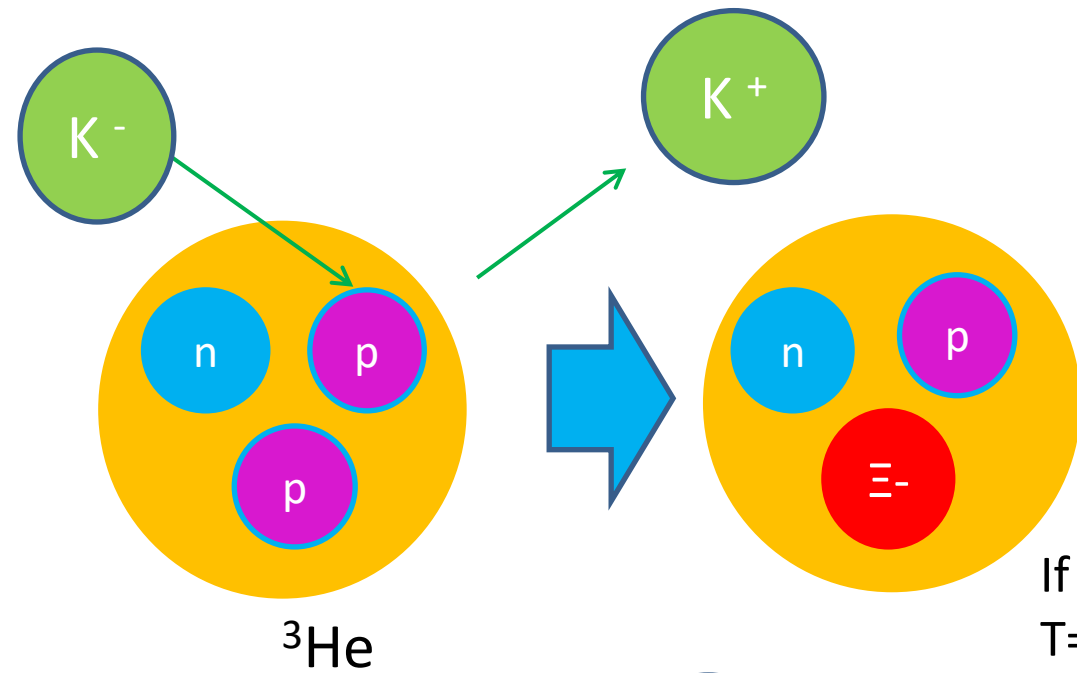


It is possible to produce  $NN\Xi$  three-body system using  ${}^3\text{He}$  target.

$T=1/2$

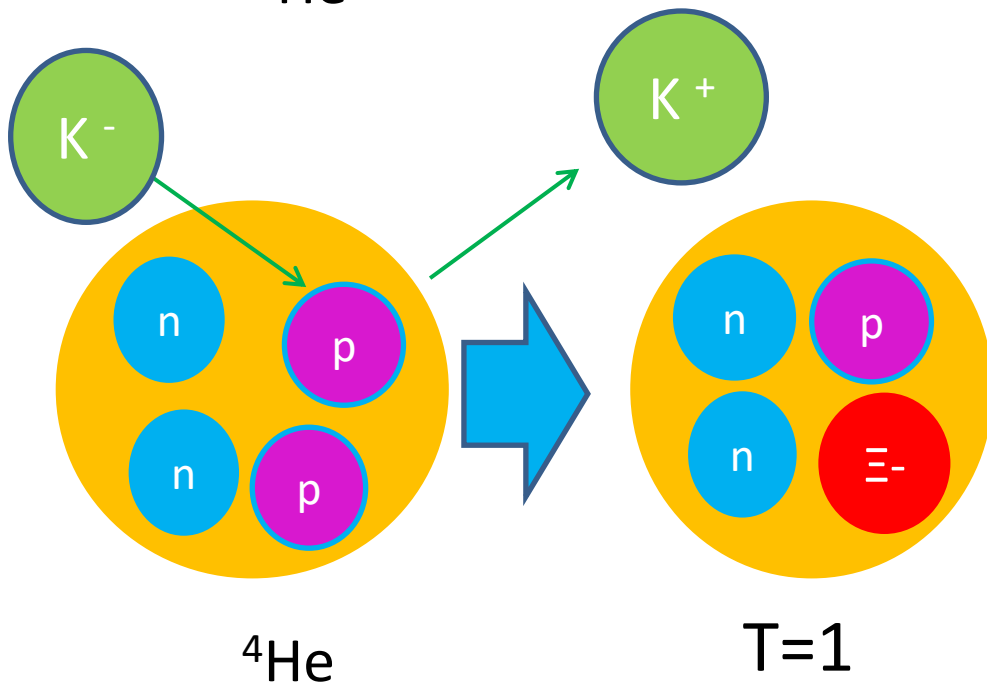
If we use Nijmegen potential,  $T=1$  four-body system becomes bound state. But, HAL potential does not produce bound state in  $T=1$  system, But produce a bound state for  $T=0$  state. How do we produce bound state with  $T=0$ ?

One possibility is to use heavy ion collision experiment.



${}^3\text{He}$

${}^4\text{He}$



$T=1$

Still, we have a question.

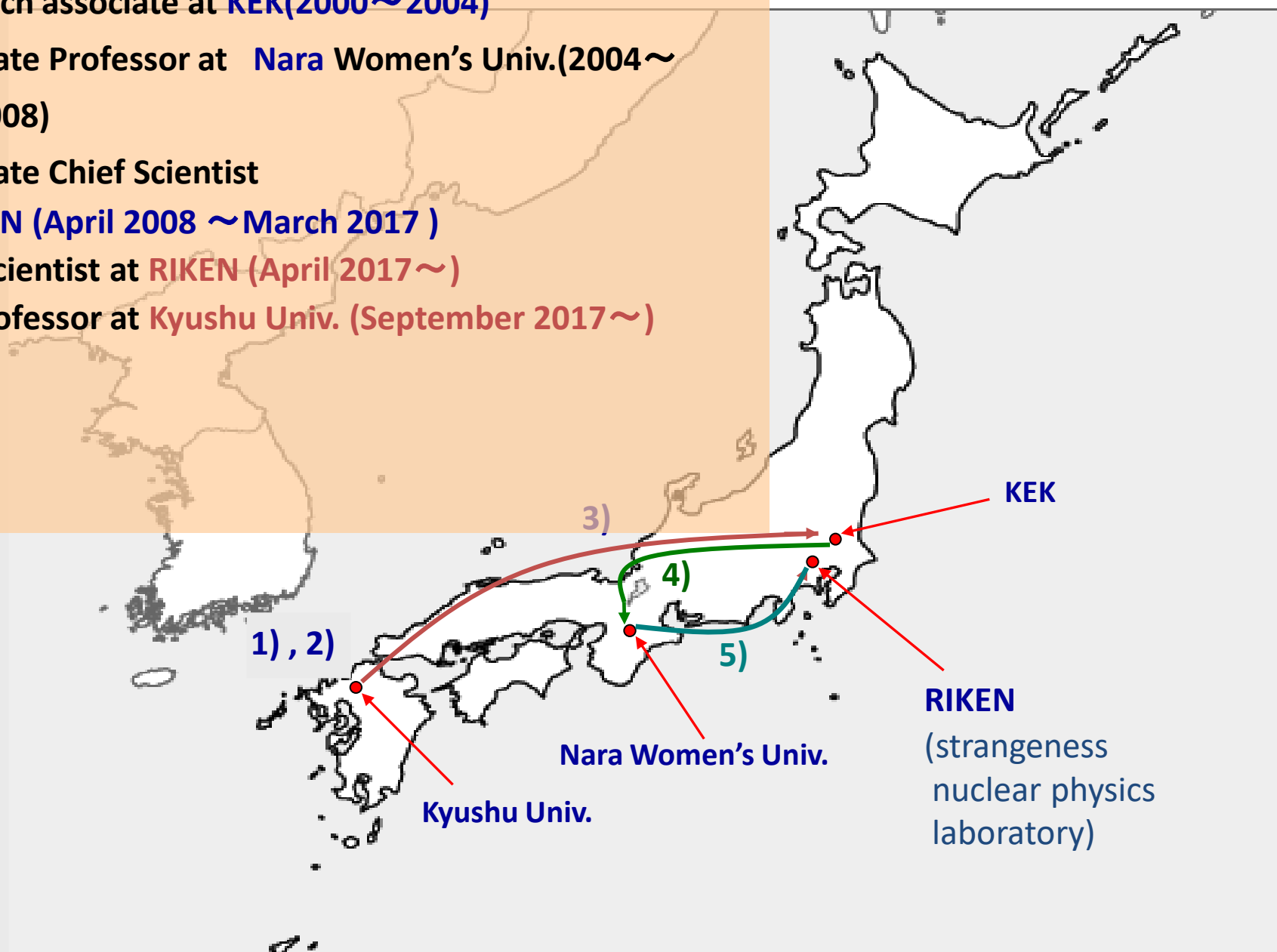
Next, what kinds of  $\Xi N$  interaction, we should obtain?

P-wave? Spin-orbit, tensor etc.

$\Lambda\Lambda$ - $\Xi N$  interaction

- Thank you !

- 1) Born in Fukuoka
- 2) Ph.D at **Kyushu Univ.**
- 3) Research associate at **KEK(2000~2004)**
- 4) Associate Professor at **Nara Women's Univ.(2004~ March 2008)**
- 5) Associate Chief Scientist at **RIKEN (April 2008 ~ March 2017 )**
- 6) Chief Scientist at **RIKEN (April 2017~)**
- 7) Full professor at **Kyushu Univ. (September 2017~)**



1) , 2)

3)

4)

5)

KEK

Nara Women's Univ.

RIKEN

(strangeness  
nuclear physics  
laboratory)

Kyushu Univ.