

*Finite Energy Sum Rules
in Hadro- and Photoproduction Reactions*

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Joint Physics Analysis Center

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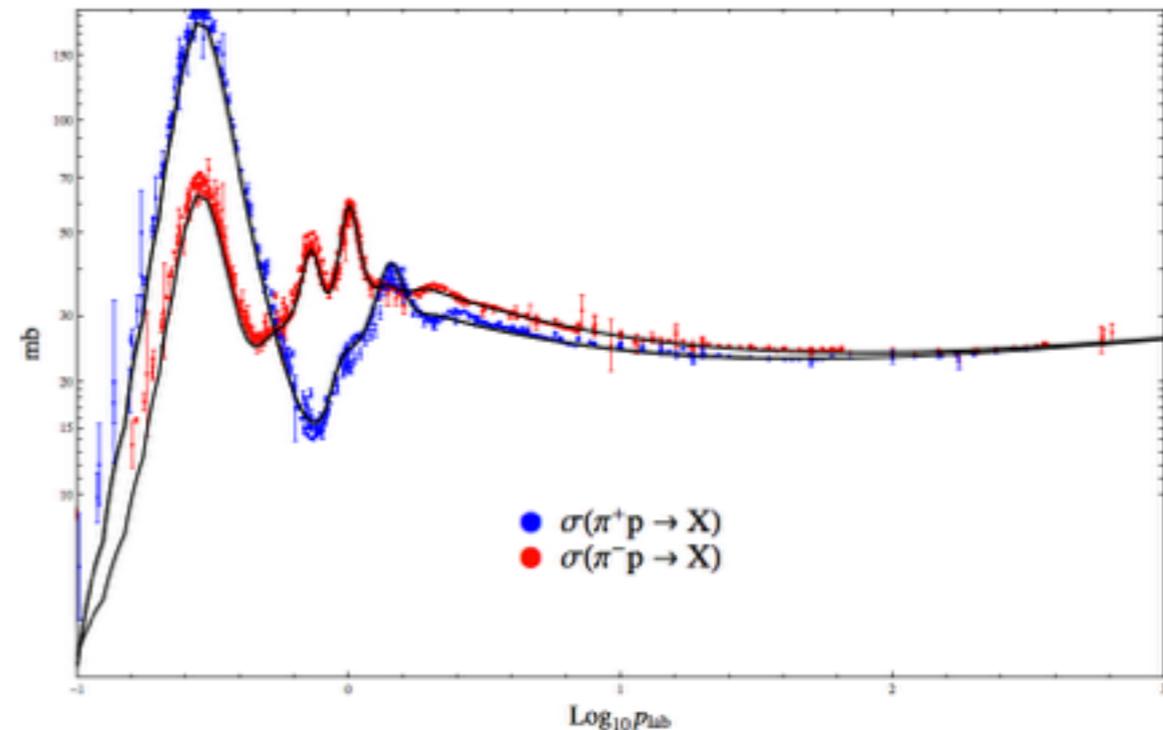
INDIANA UNIVERSITY



Outline and Motivations

- about extracting resonance properties

$g(Q^2), M, \Gamma$ from exp./lattice data



- Derivation of Finite Energy Sum Rules (FESR)

- Illustrations of FESR with

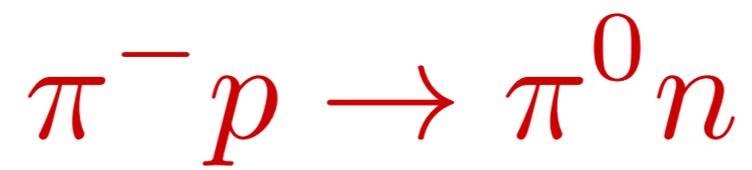
$$\pi p \rightarrow \pi p$$

$$\gamma p \rightarrow \pi p$$

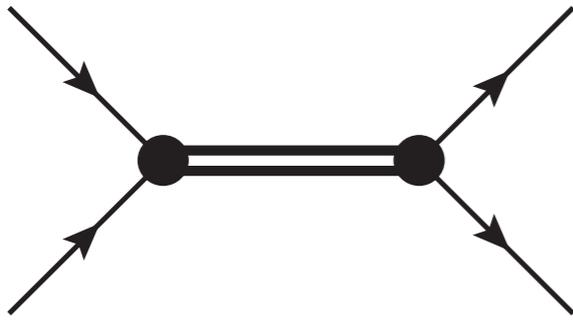
$$\gamma p \rightarrow \eta p$$

- Methodology

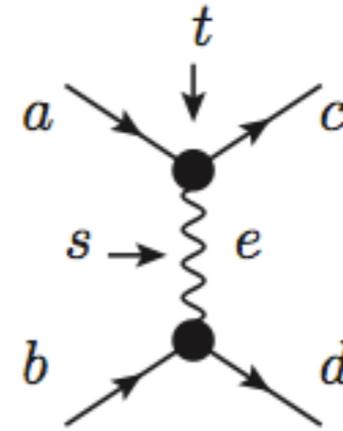
- Conclusion



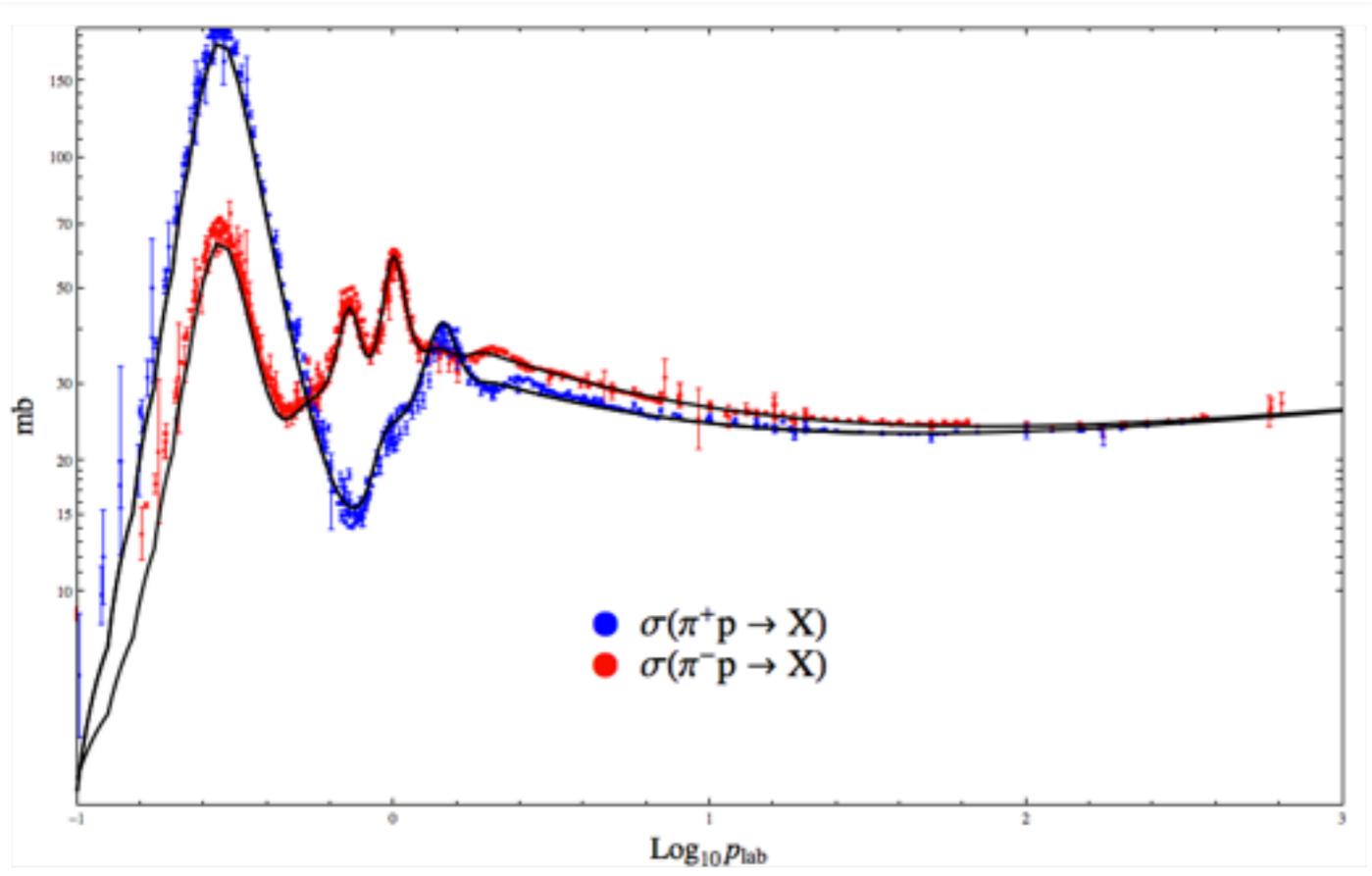
Low energy: baryon resonances

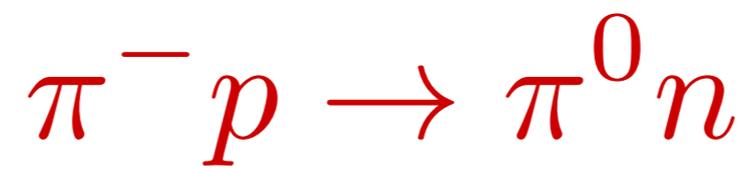


High energy: Regge exchange

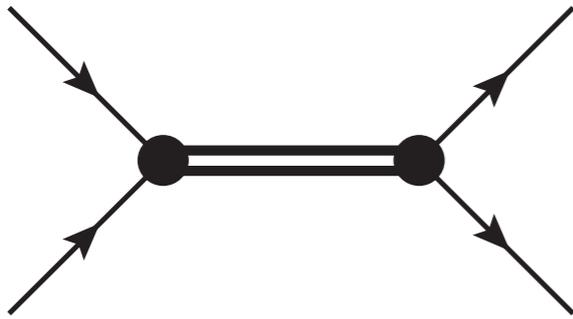


Total cross section

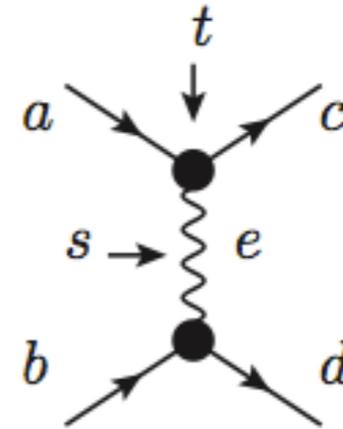




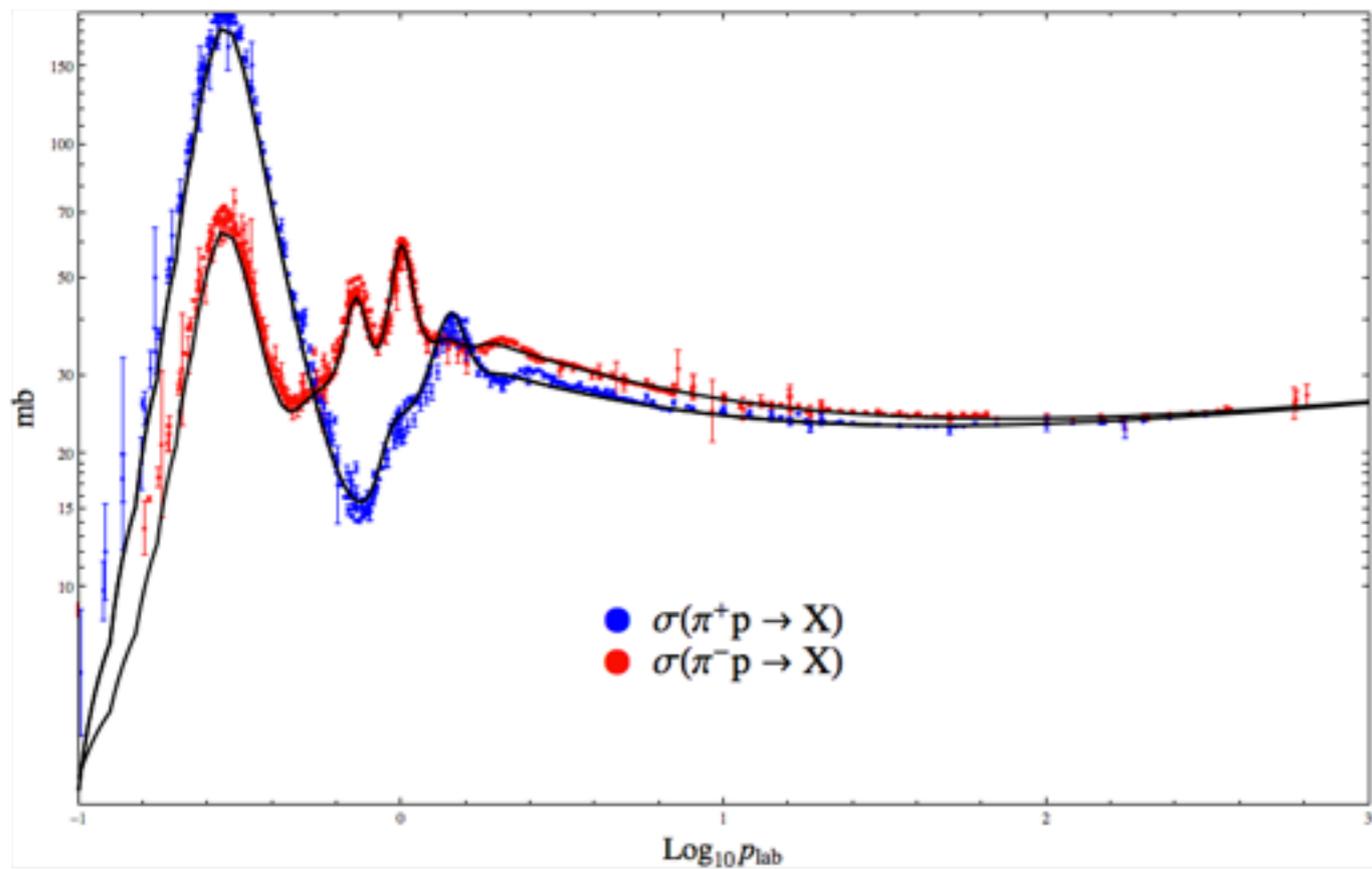
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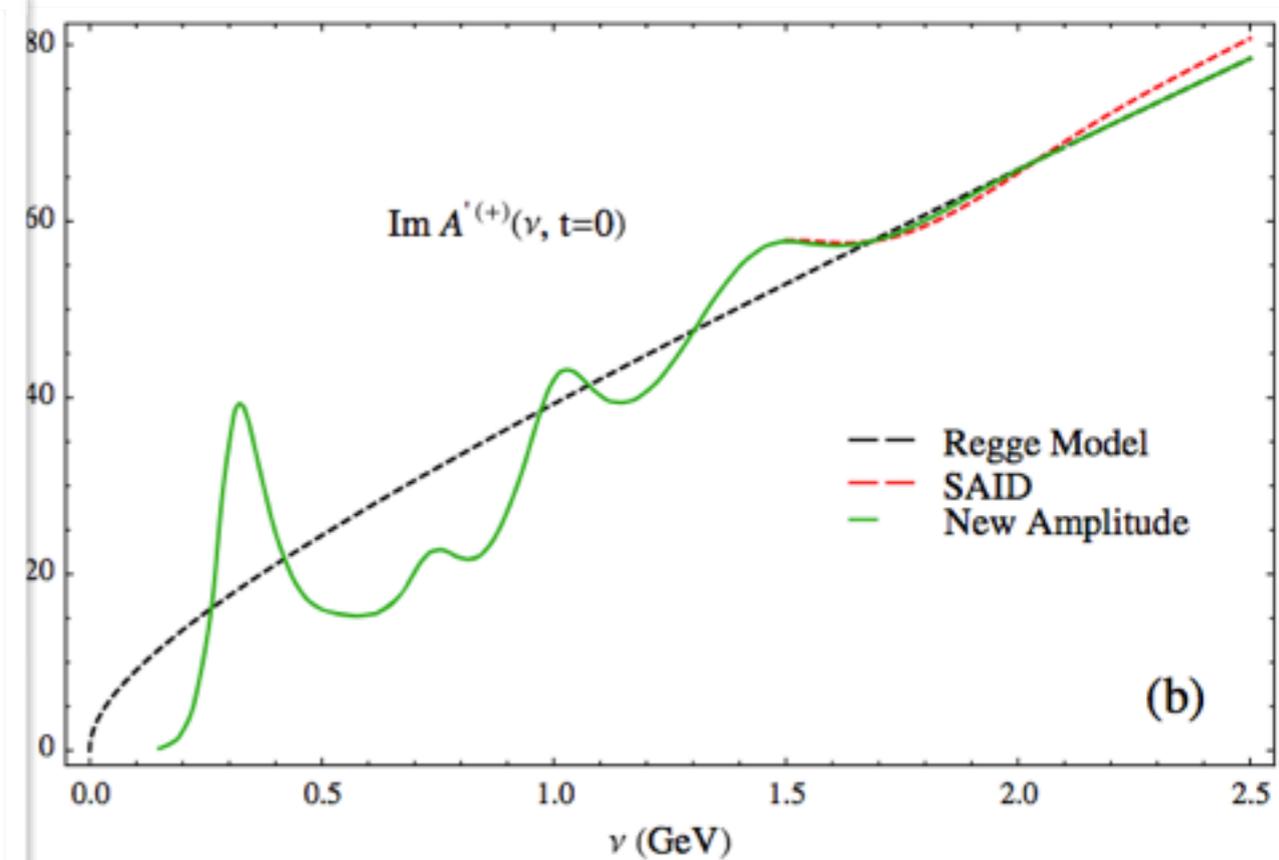
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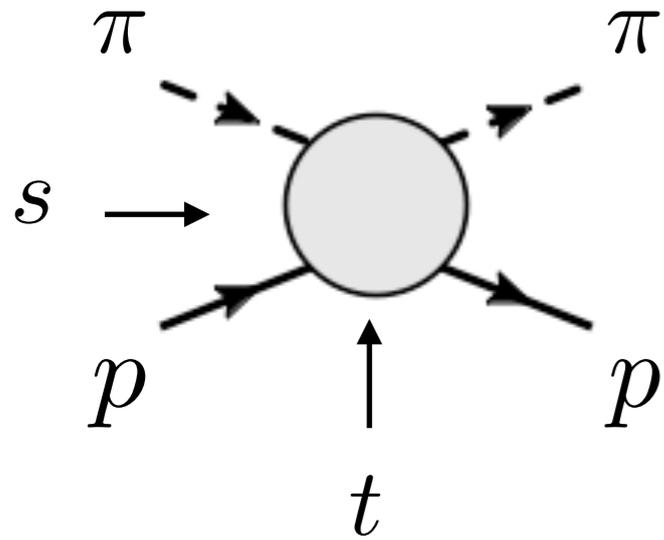
Total cross section



Amplitude

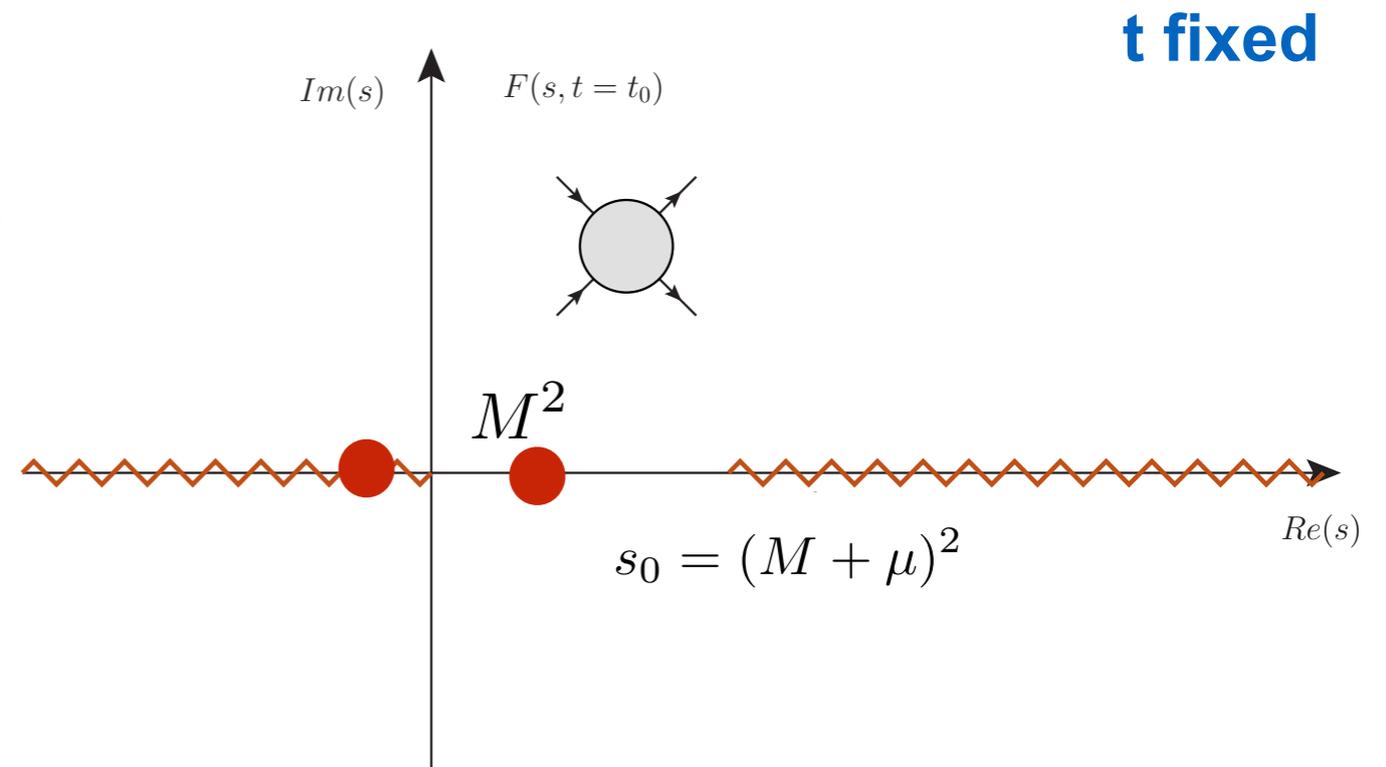


Dispersion Relation

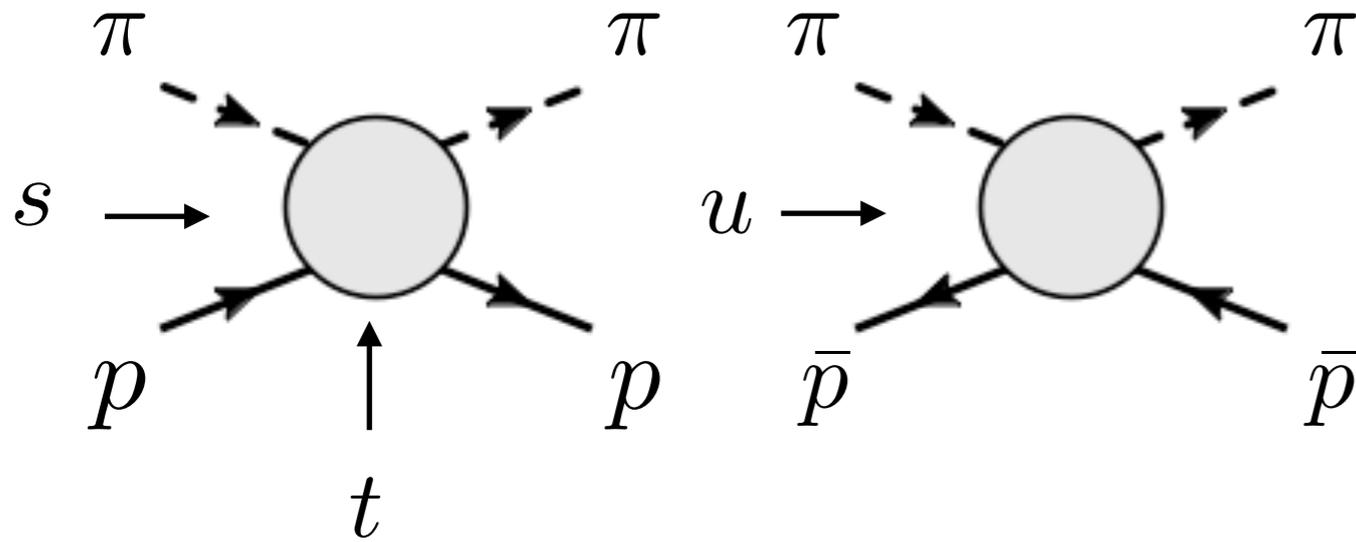


Decompose into scalar function

$$T = \bar{u}(p_4, \lambda_4) \left(A + \frac{1}{2} (\not{p}_1 + \not{p}_3) B \right) u(p_2, \lambda_2)$$



Dispersion Relation



Conjugation charge relate
 πp and $\pi \bar{p}$ scatterings

Decompose into scalar function

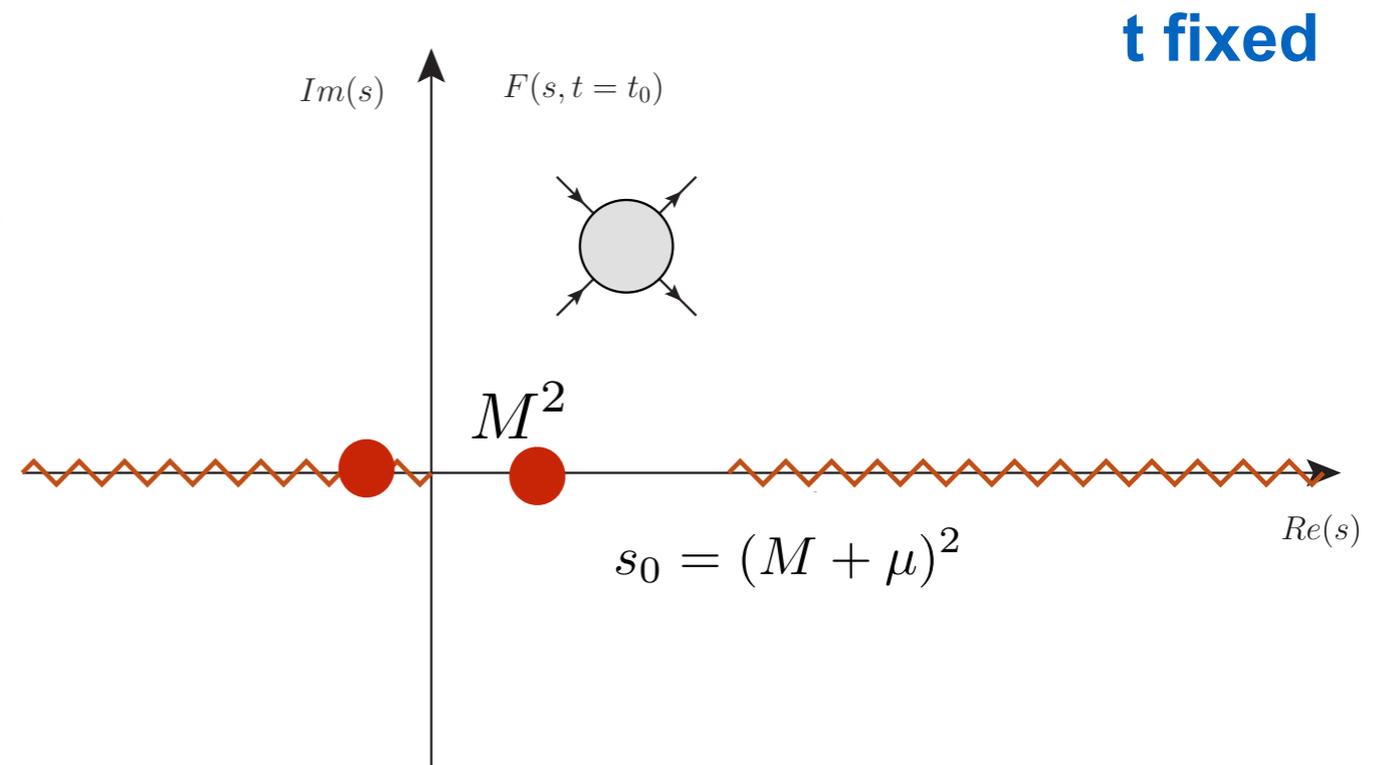
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Introduce the crossing variable

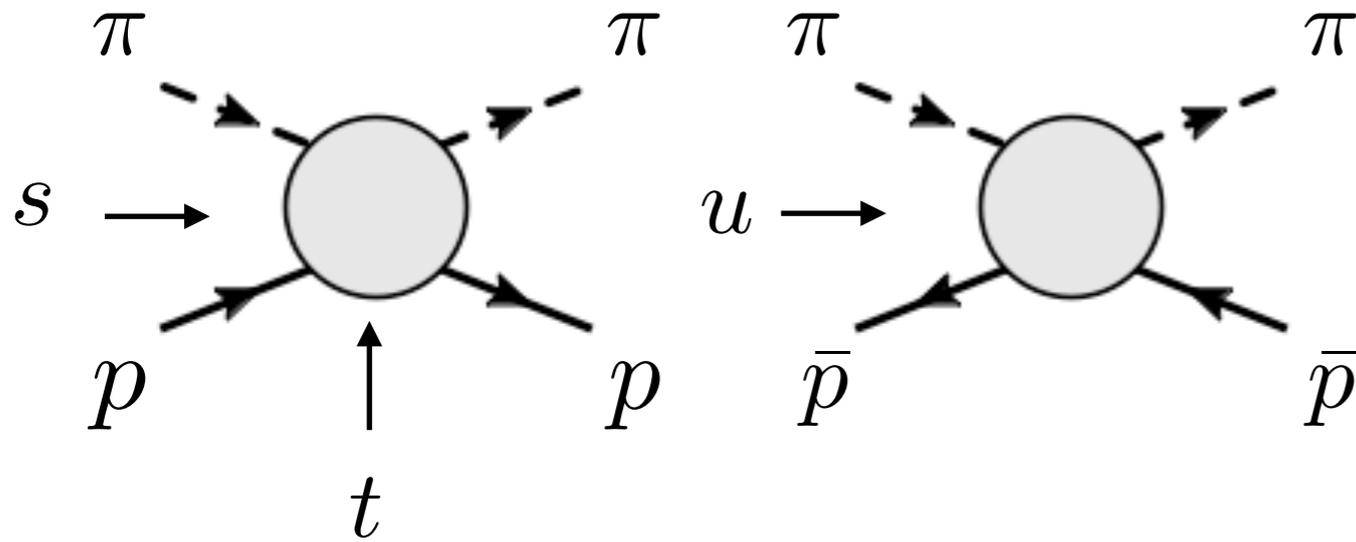
$$\nu = \frac{s - u}{2}$$

$$u(s, t) = -s - t + 2M^2 + 2\mu^2$$

and symmetrize the two cuts



Dispersion Relation



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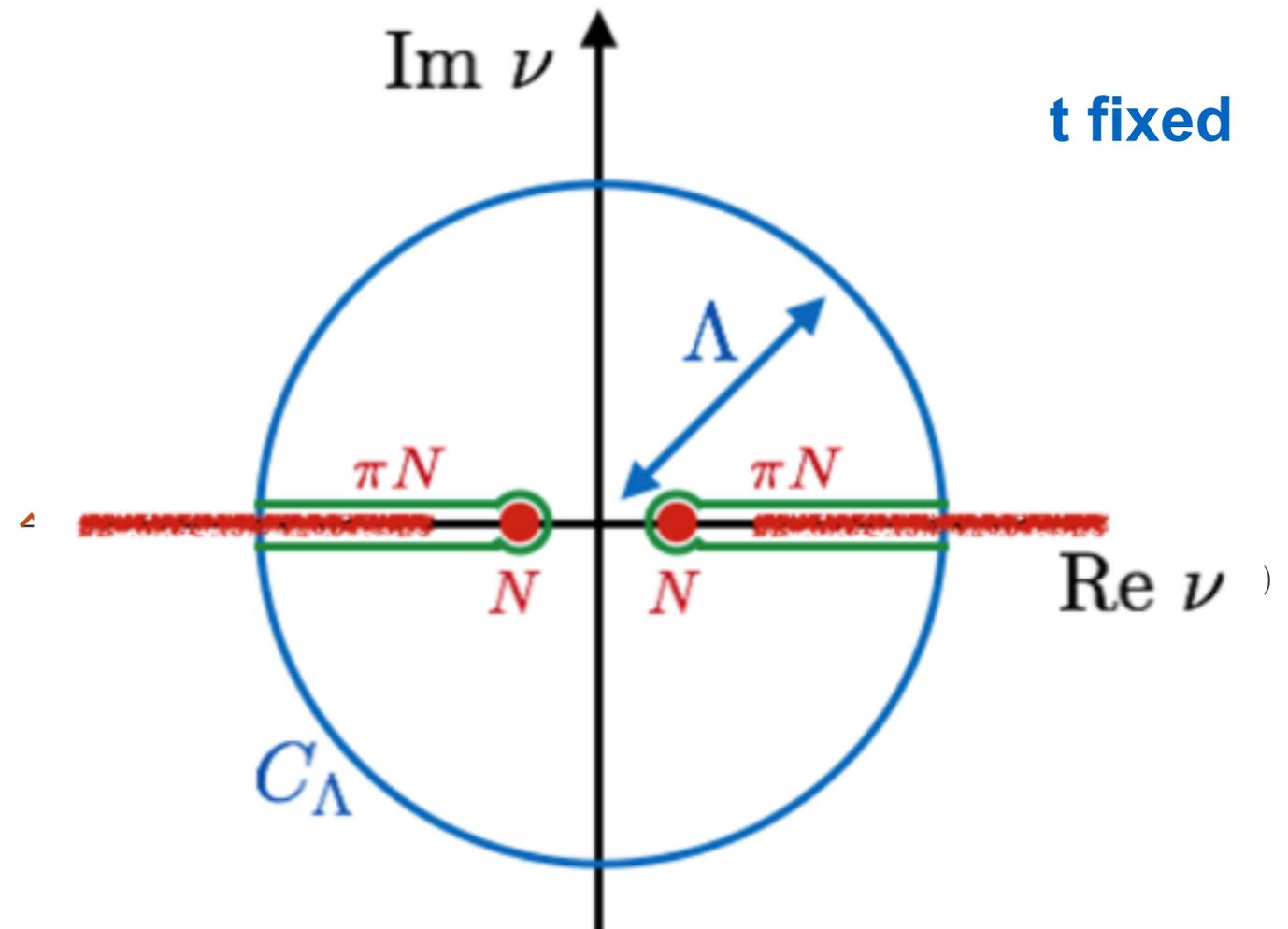
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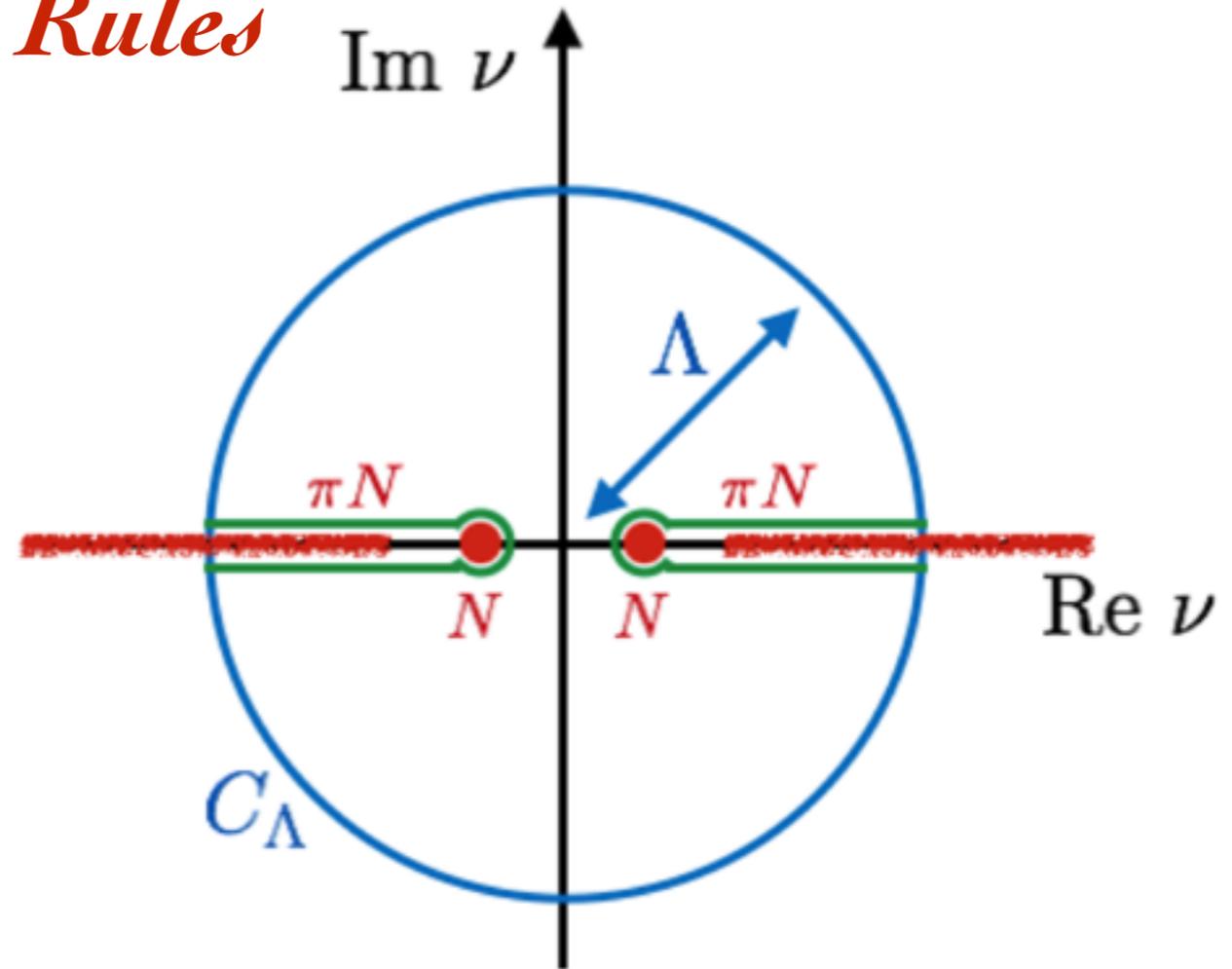
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Finite Energy Sum Rules

Cauchy contour

$$\oint_C A(\nu, t) d\nu = 0$$

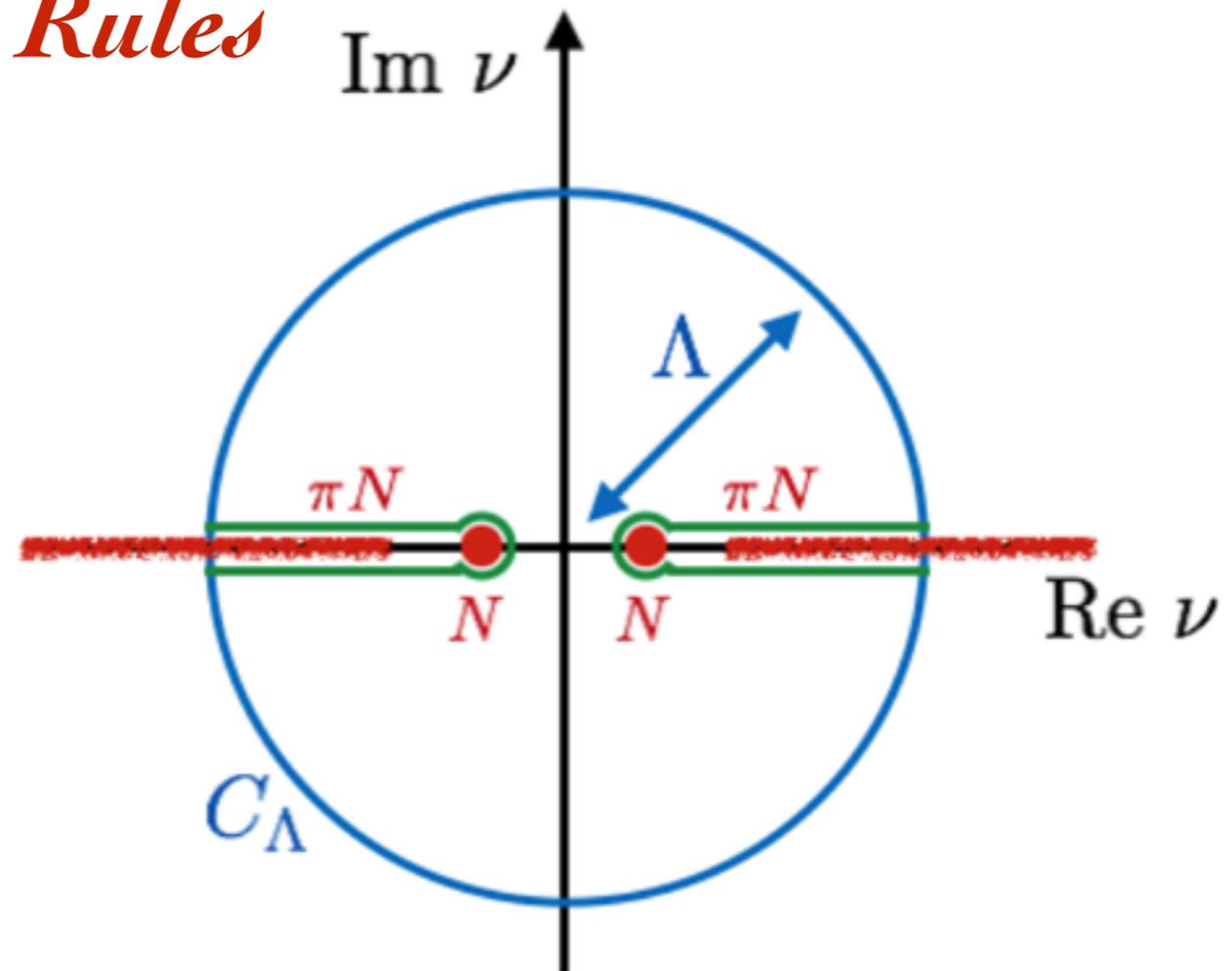


Finite Energy Sum Rules

Cauchy contour

$$\oint_C A(\nu, t) d\nu = 0$$

$$2i \int_{\nu_0}^{\Lambda} \text{Im } A(\nu, t) d\nu = - \oint_{C_\Lambda} A(\nu, t) d\nu$$



Finite Energy Sum Rules

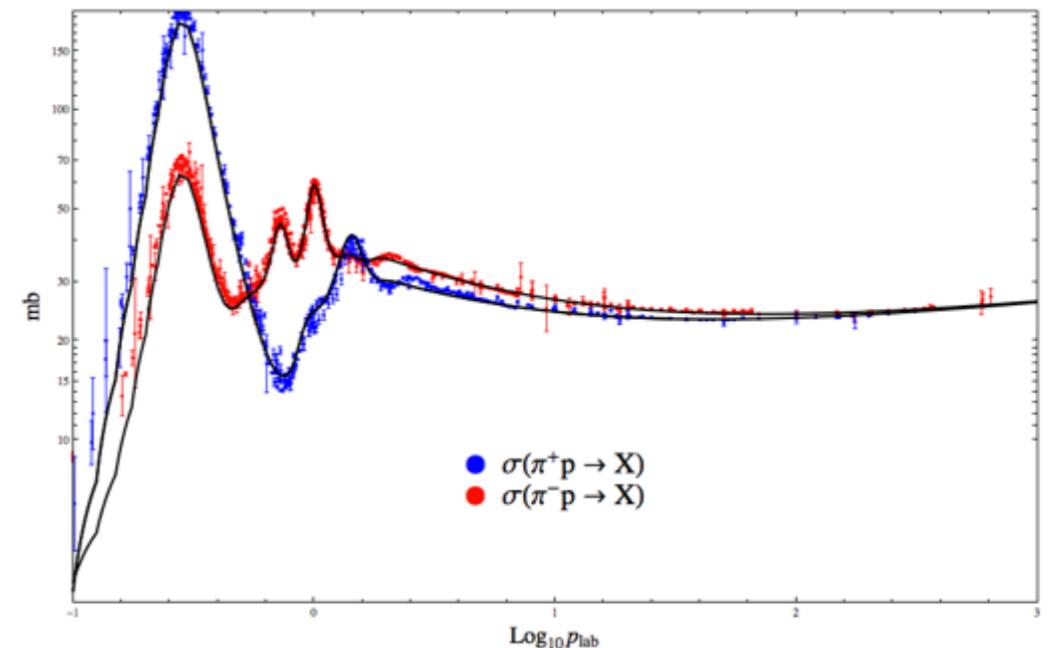
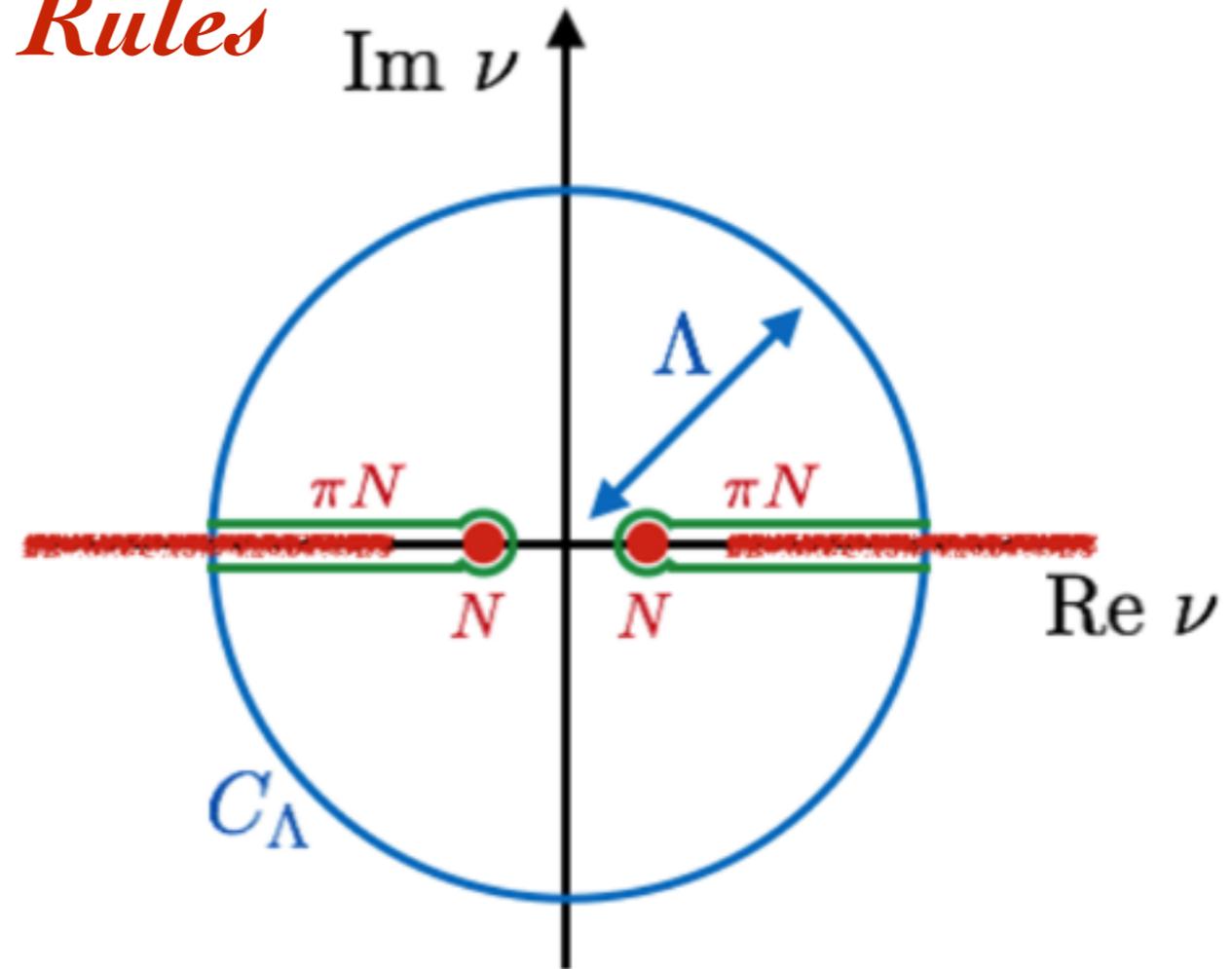
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Assume Regge form at $\nu = \Lambda$:

$$A(\nu, t) = \beta(t) \frac{\pm 1 - e^{-i\pi\alpha(t)}}{\sin \pi\alpha(t)} \nu^{\alpha(t)}$$

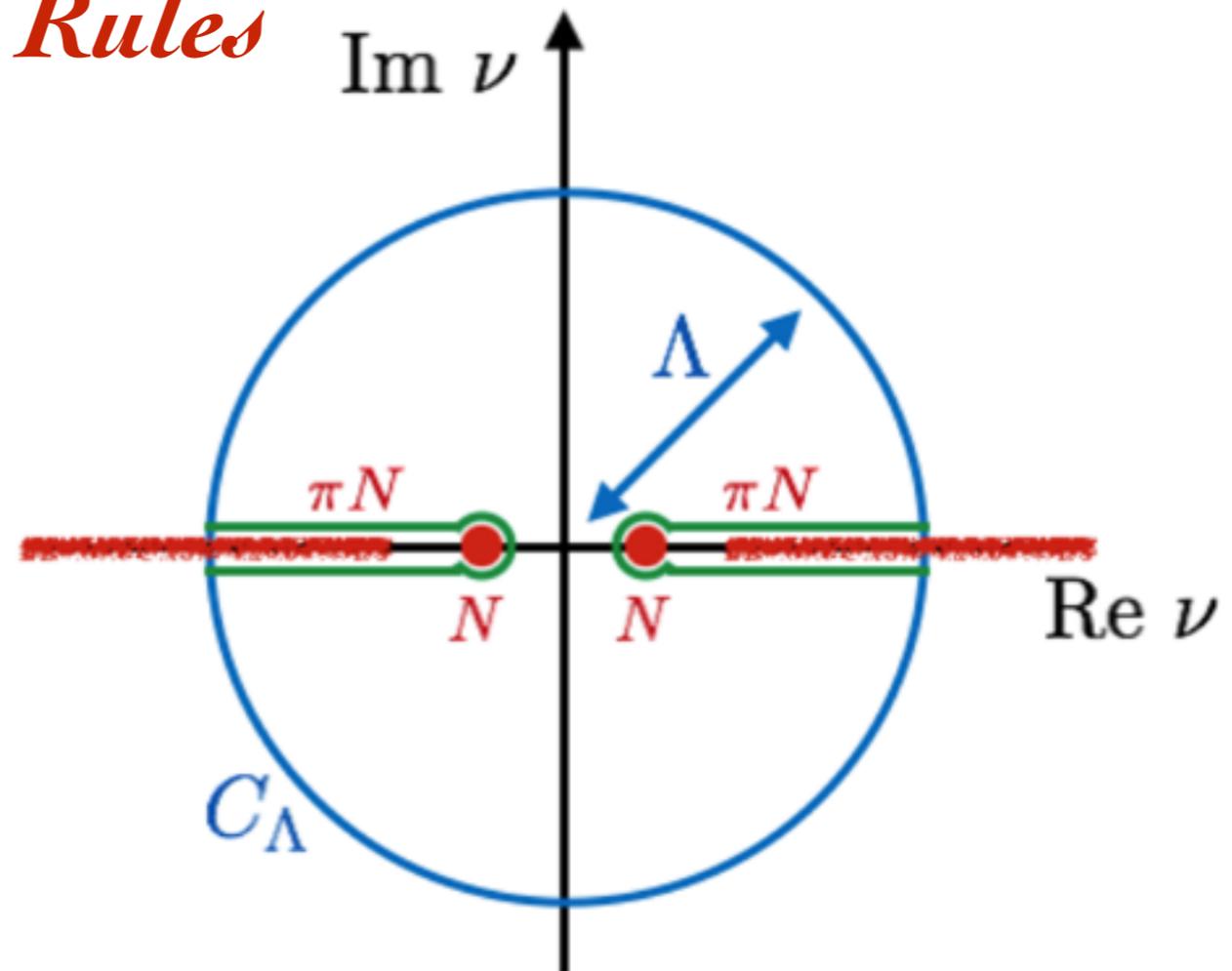


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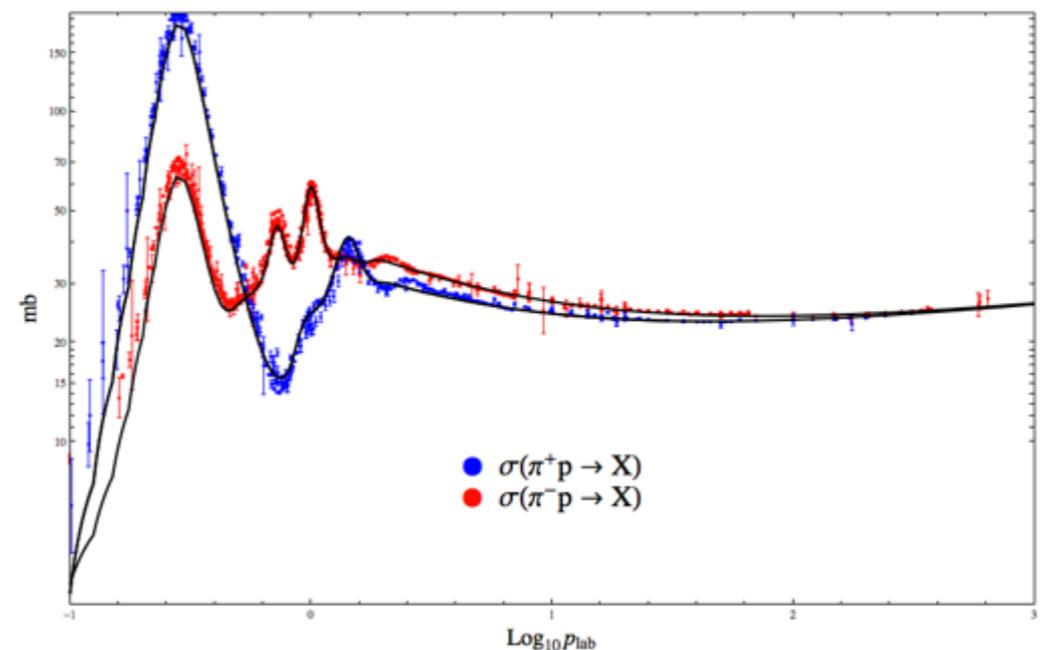


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t fixed

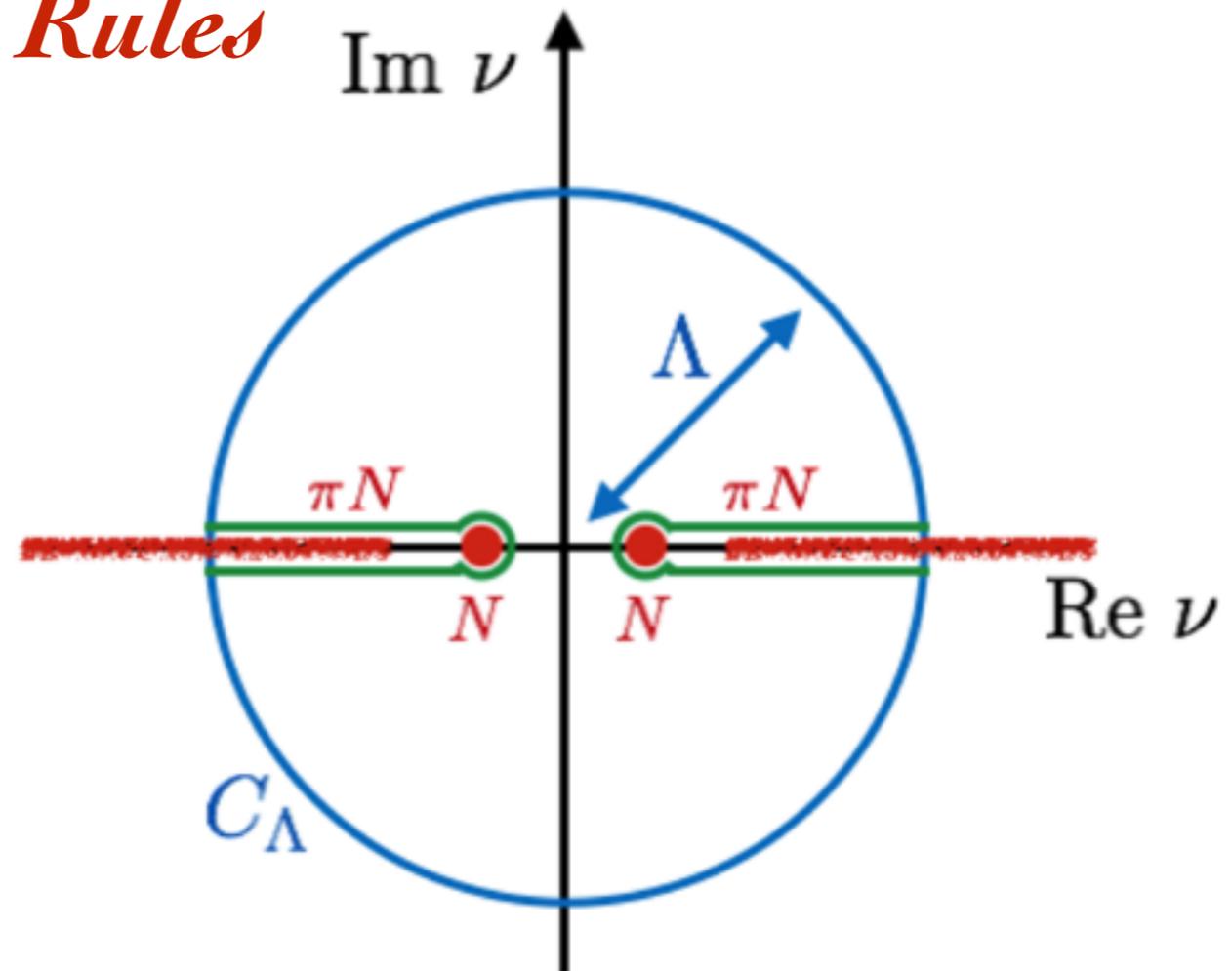


Finite Energy Sum Rules

Cauchy contour

$$\oint_C A(\nu, t) \nu^k d\nu = 0$$

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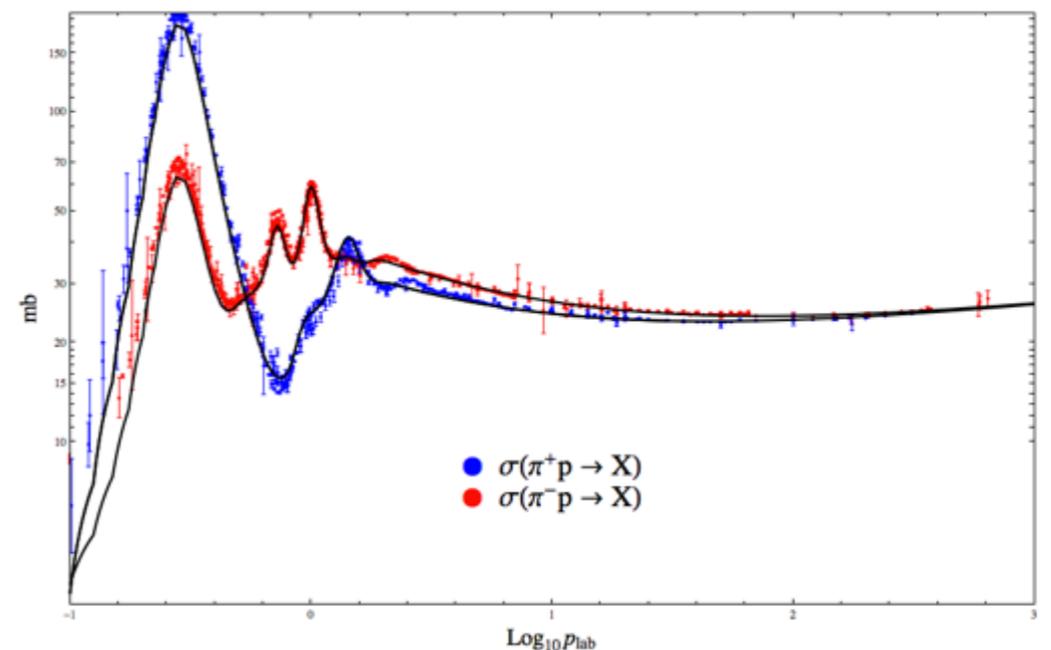


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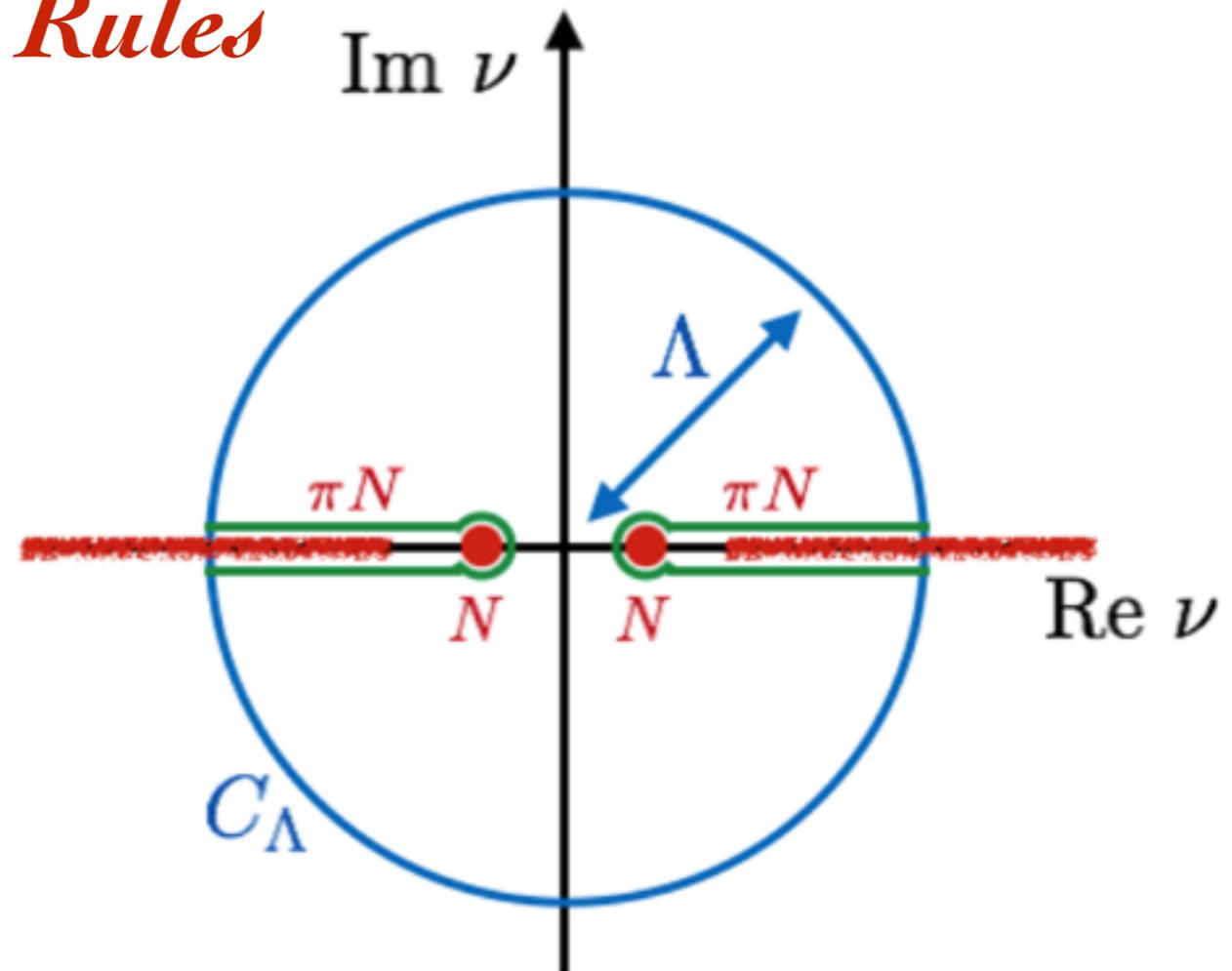


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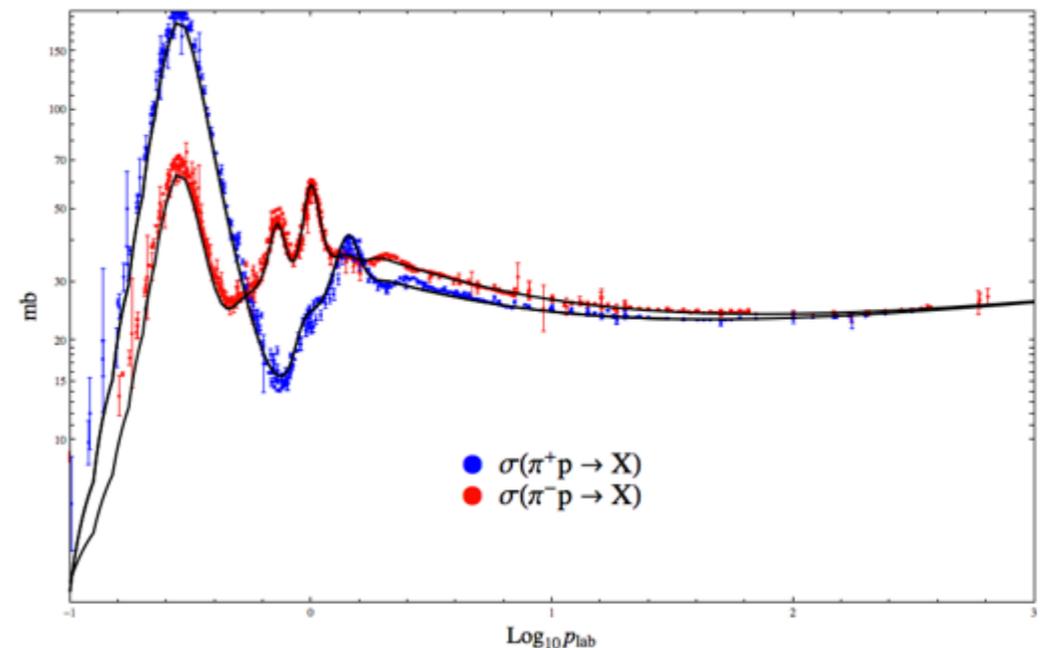


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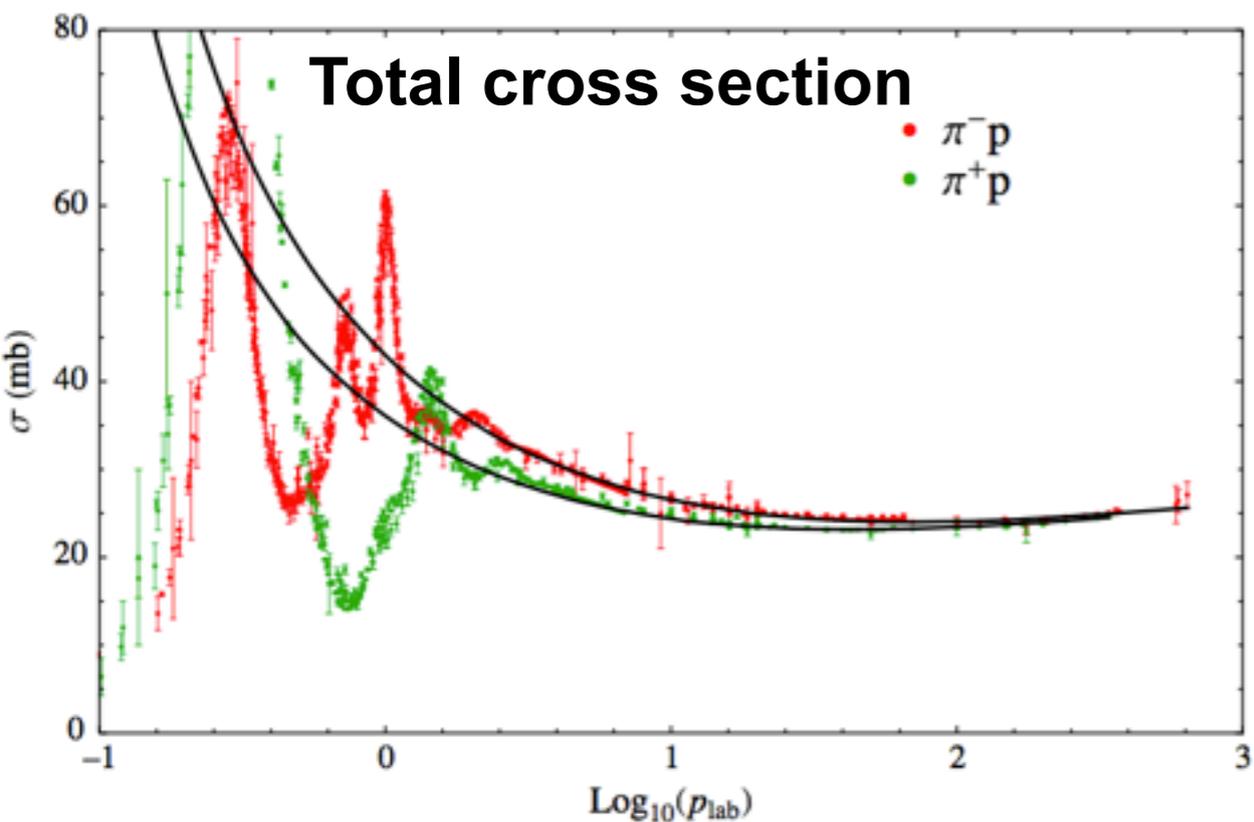
$$\frac{1}{\Lambda^k} \int_{\nu_0}^{\Lambda} \text{Im } A(\nu, t) \nu^k d\nu = \frac{\beta(t) \Lambda^{\alpha(t)+1}}{\alpha(t) + k + 1}$$

t fixed



Application to πN : High Energy Fit

VM et al (JPAC) PRD92
arXiv:1506.01764



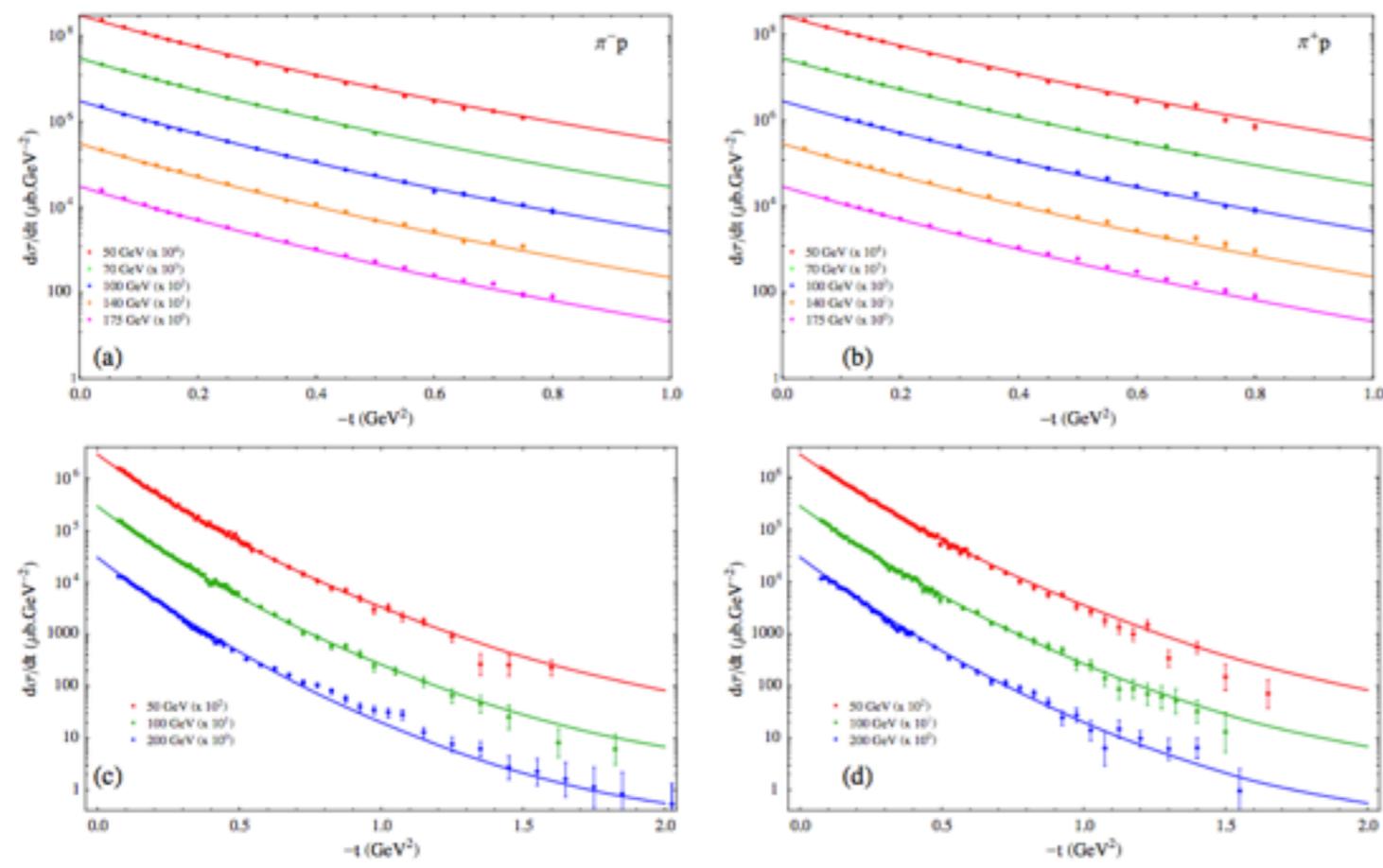
Fit to the world data on

$$\pi^\pm p \rightarrow \pi^\pm p$$

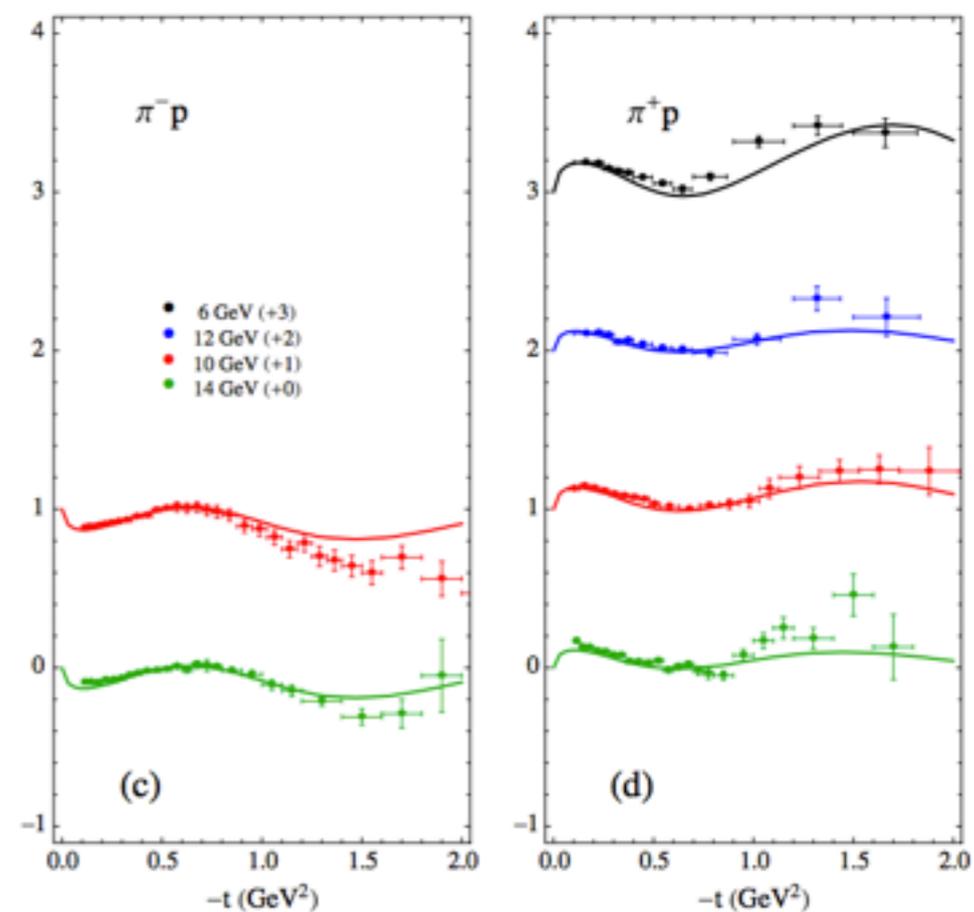
$$\pi^- p \rightarrow \pi^0 n$$

for beam energy > 2 GeV

Differential cross section



Polarization observable

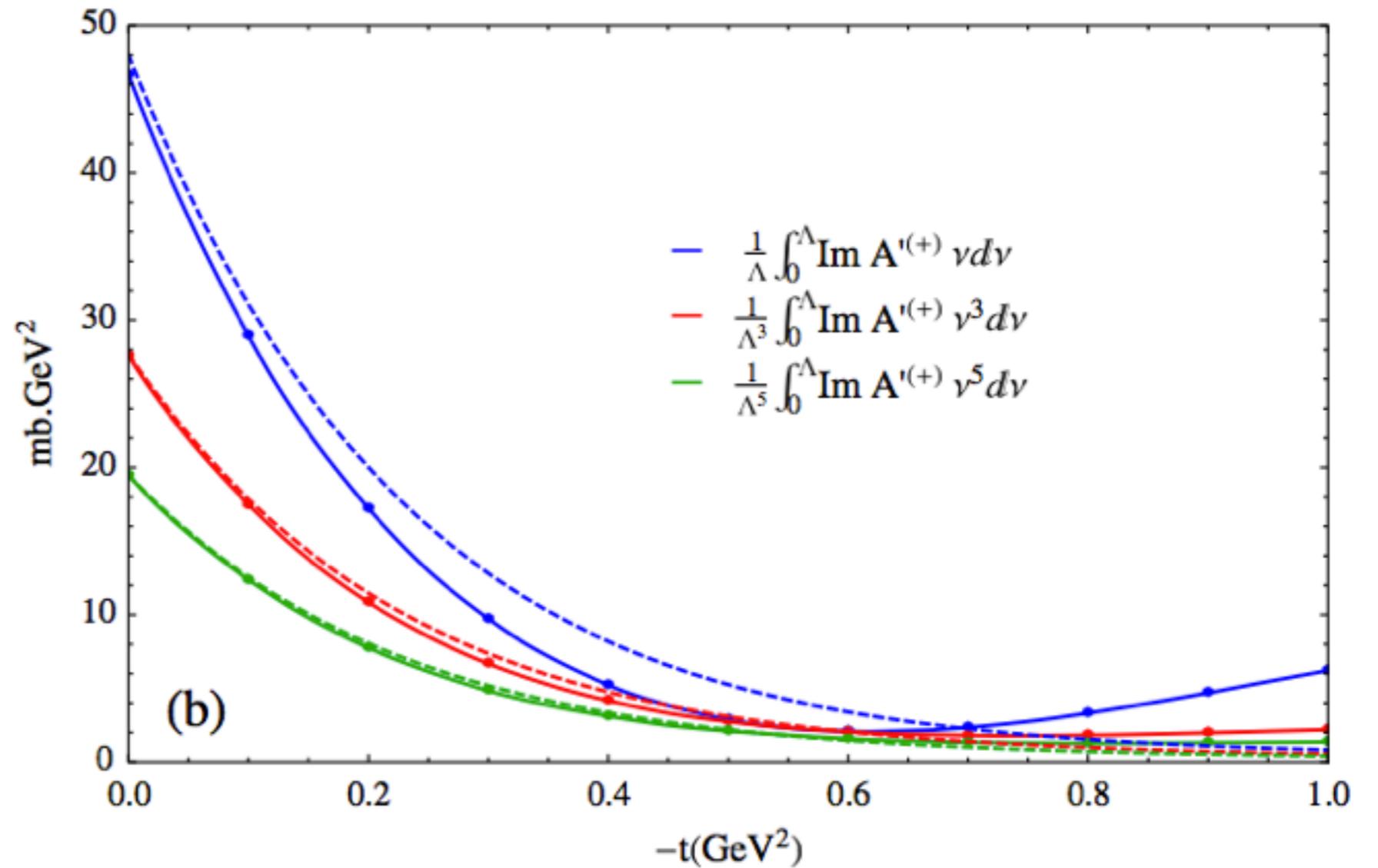


Let's compare both side of the sum rule

Solid line: SAID

Dashed line: Regge

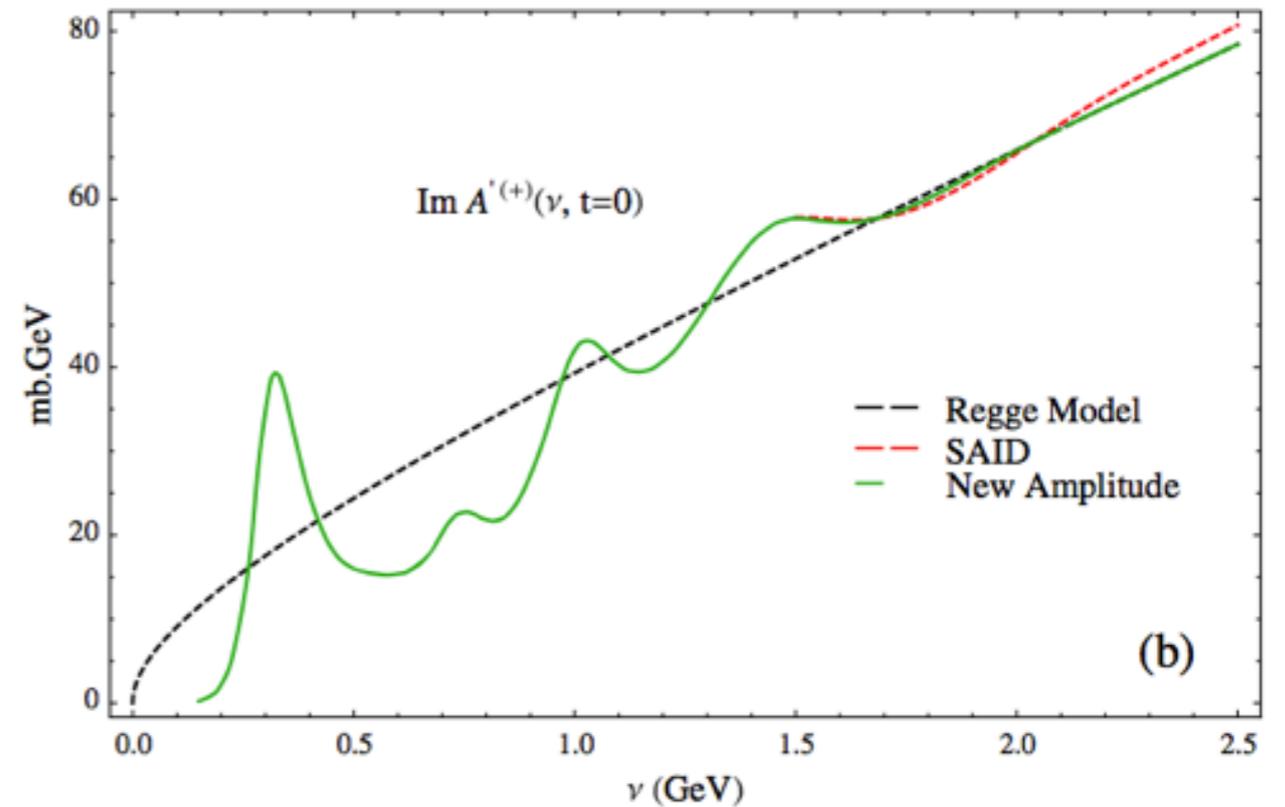
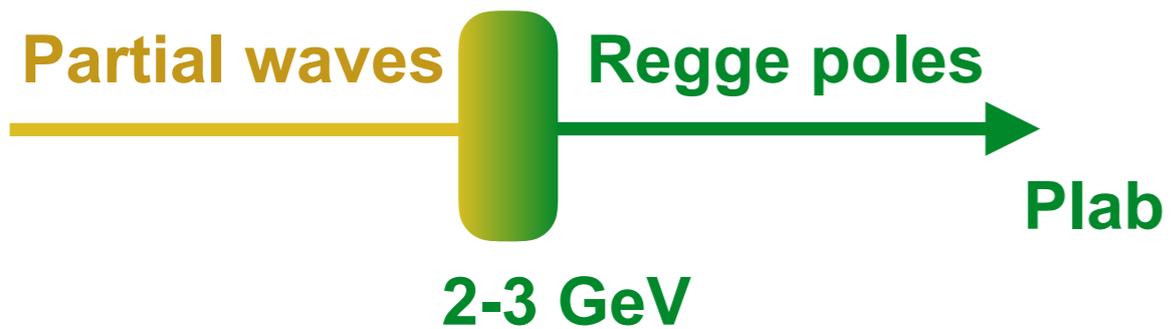
excellent match !



$$\frac{1}{\Lambda^k} \int_{\nu_0}^{\Lambda} \text{Im } A(\nu, t) \nu^k d\nu = \frac{\beta(t) \Lambda^{\alpha(t)+1}}{\alpha(t) + k + 1}$$

Checking Analyticity

Match low energy (PW)
and high energy (Regge)
imaginary parts

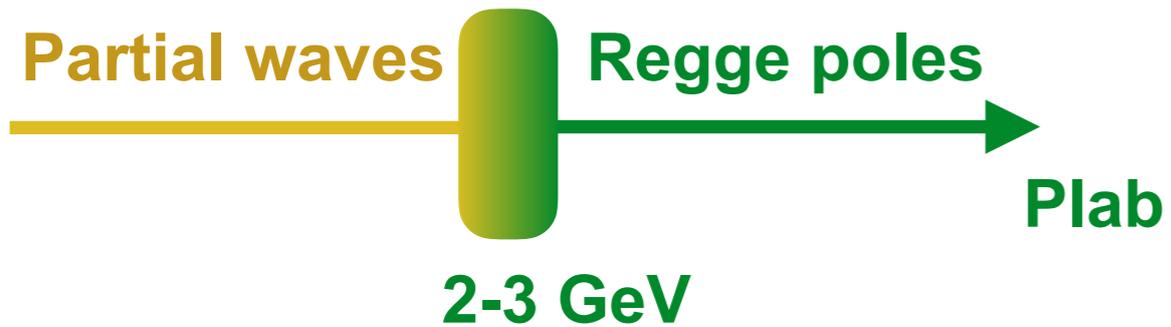


Reconstruct the real part
from the dispersion relation

$$A(\nu, t) = \frac{2}{\pi} \int_{\nu_0}^{\infty} \frac{\text{Im } A(\nu', t)}{\nu'^2 - \nu^2} \nu' d\nu'$$

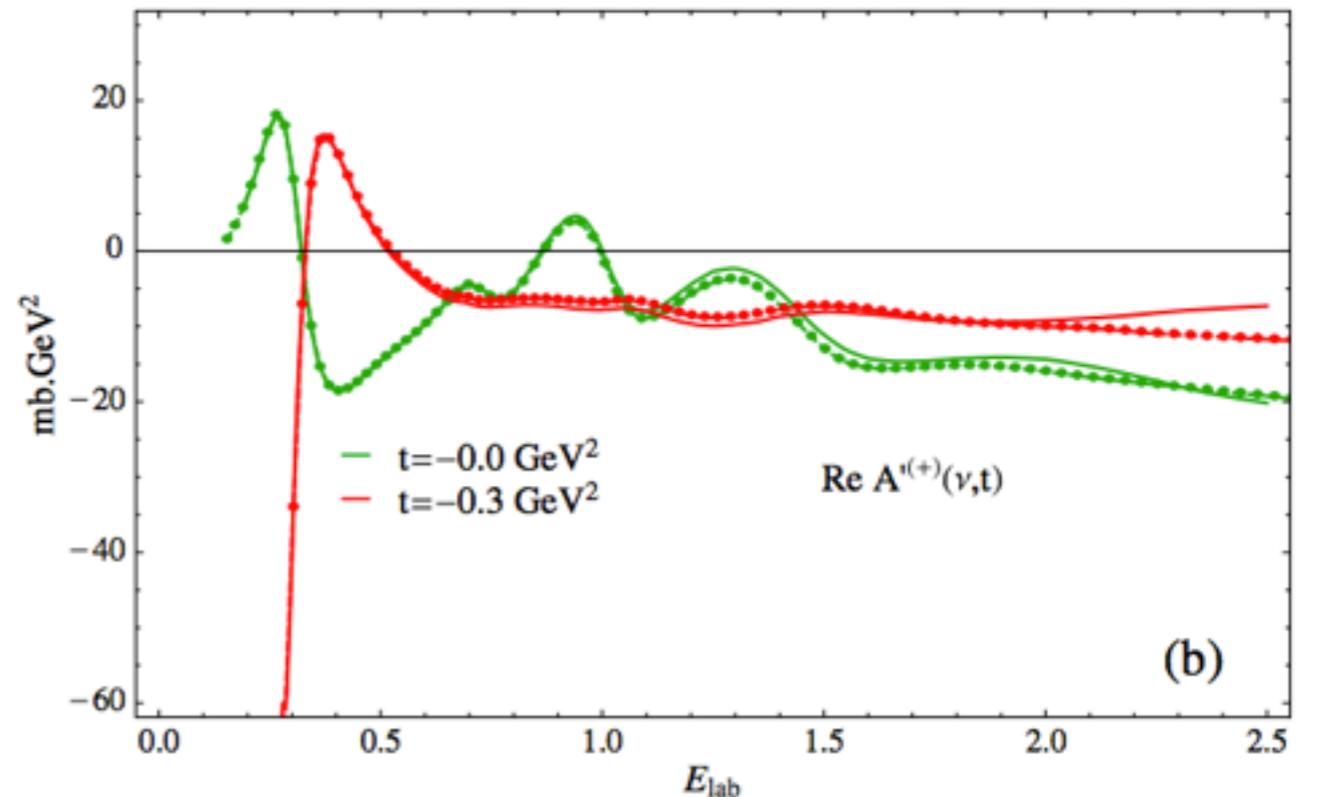
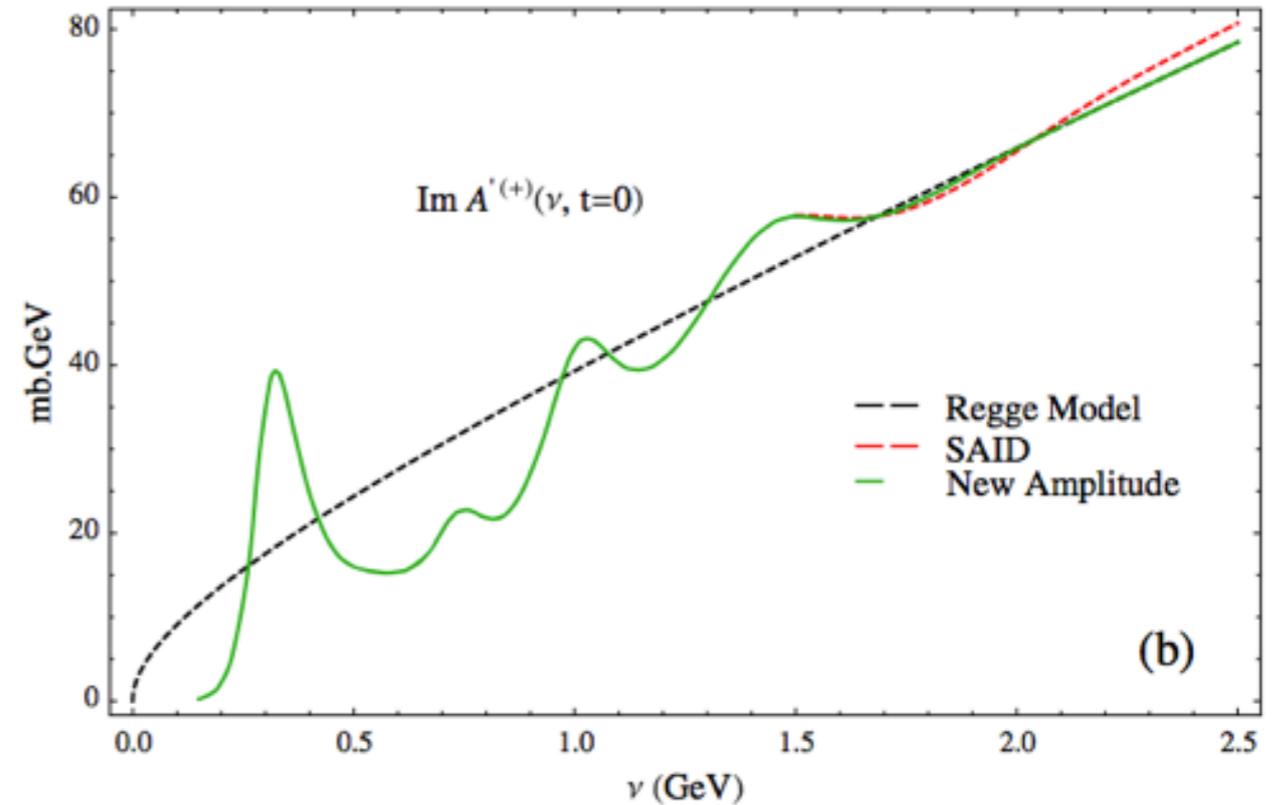
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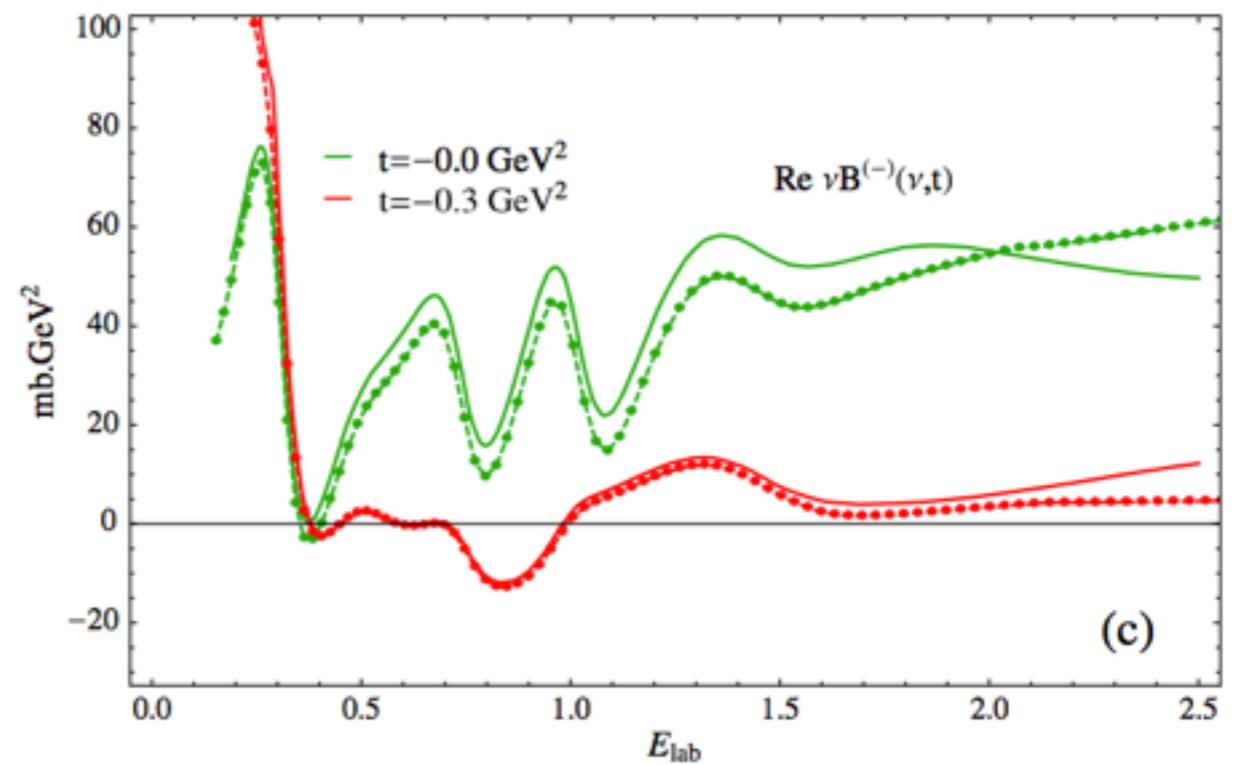
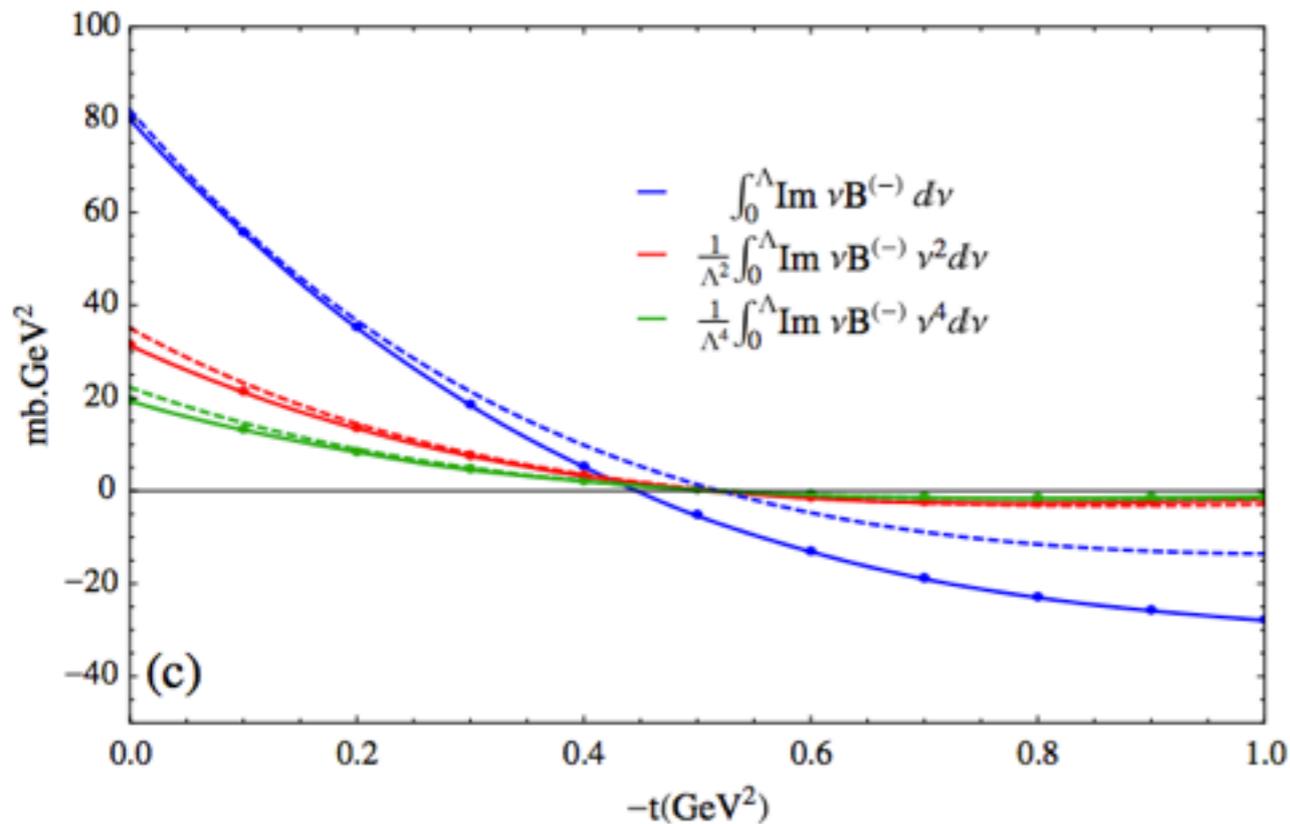
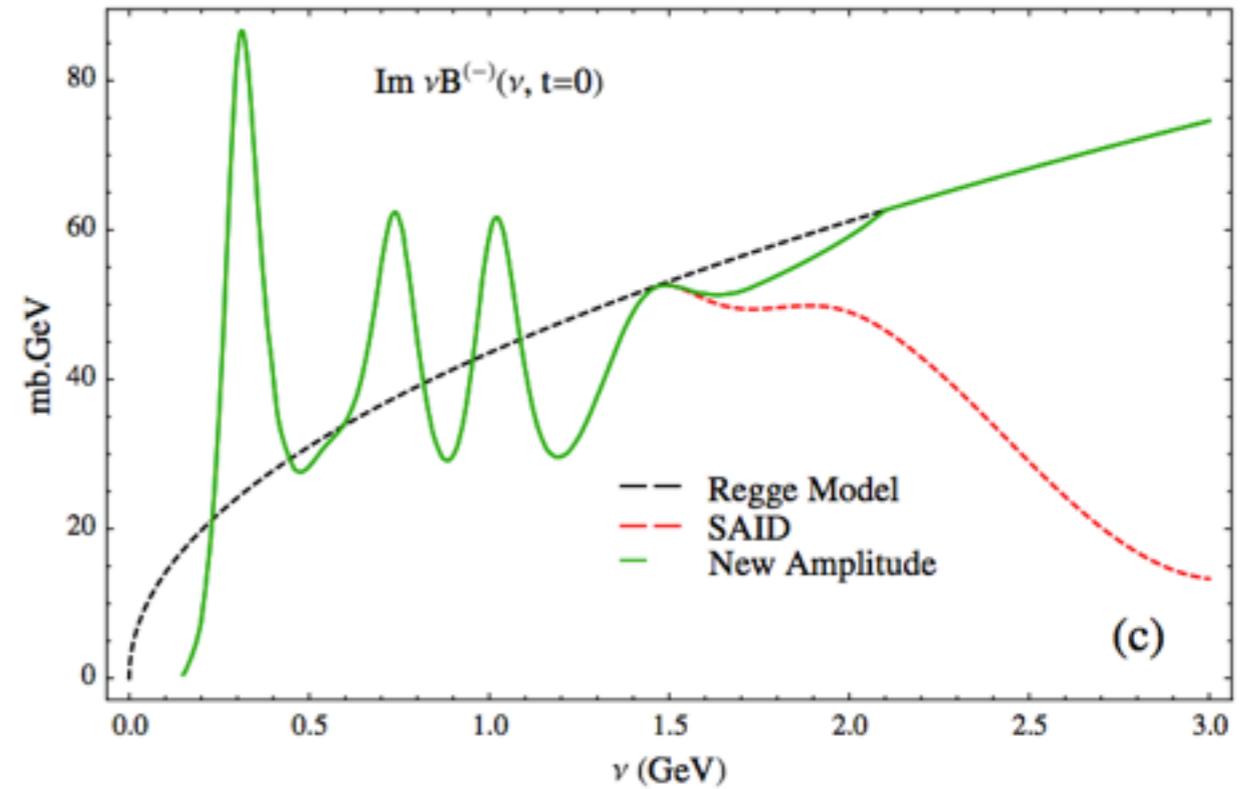
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Similar results for the other amplitude

$$T = \bar{u}(p_4, \lambda_4) \left(A + \frac{1}{2} (\not{p}_1 + \not{p}_3) B \right) u(p_2, \lambda_2)$$



Application to Pseudoscalar Photoproduction

$$\gamma N \longrightarrow \pi N$$

$$(\pm 1) \left(\pm \frac{1}{2} \right) \longrightarrow 0 \left(\pm \frac{1}{2} \right)$$

**8 helicity configurations
related by pair via parity**

4 indep. helicity configurations

use CGLN basis A_1, \dots, A_4

Isospin symmetry:

every amplitude has an isospin index (+,-,0)

$$A_{ji}^a = A^{(+)} \delta^{a3} \delta_{ji} + A^{(-)} \frac{1}{2} [\tau^a, \tau^3]_{ji} + A^{(0)} \tau_{ji}^a$$

12 indep. helicity/isospin configurations

$$\gamma p \rightarrow \pi^+ n : \sqrt{2} \left(A^{(0)} + A^{(-)} \right)$$

$$\gamma n \rightarrow \pi^- p : \sqrt{2} \left(A^{(0)} - A^{(-)} \right)$$

$$\gamma p \rightarrow \pi^0 p : A^{(+)} + A^{(0)}$$

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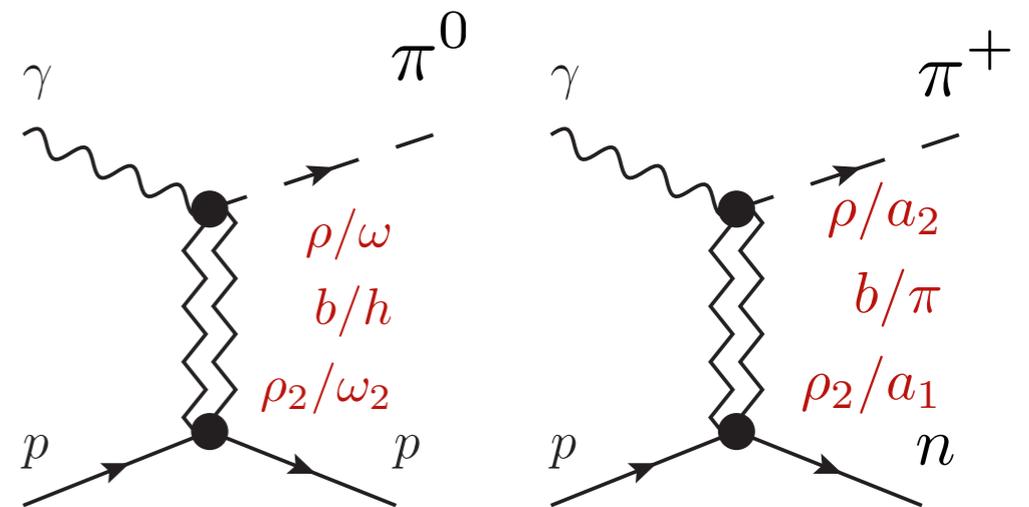
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| A_i | (0) | (+) | (-) |
|-------|----------|------------|-------|
| A_1 | ρ | ω | a_2 |
| A_2 | b | h | π |
| A_3 | ρ_2 | ω_2 | a_1 |
| A_4 | ρ | ω | a_2 |

Use CGLN basis $A_{1,2,3,4}$

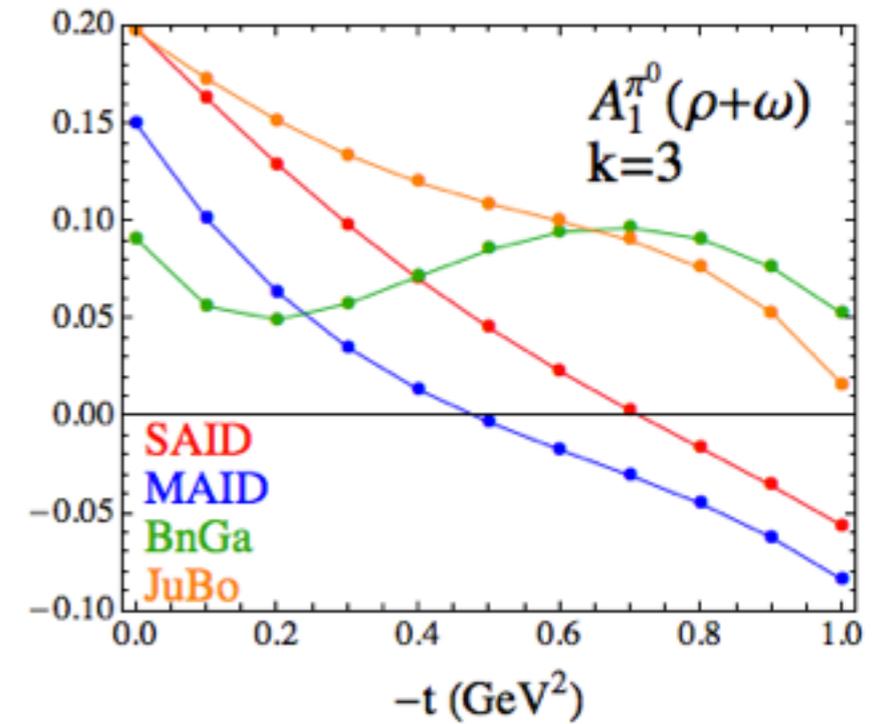
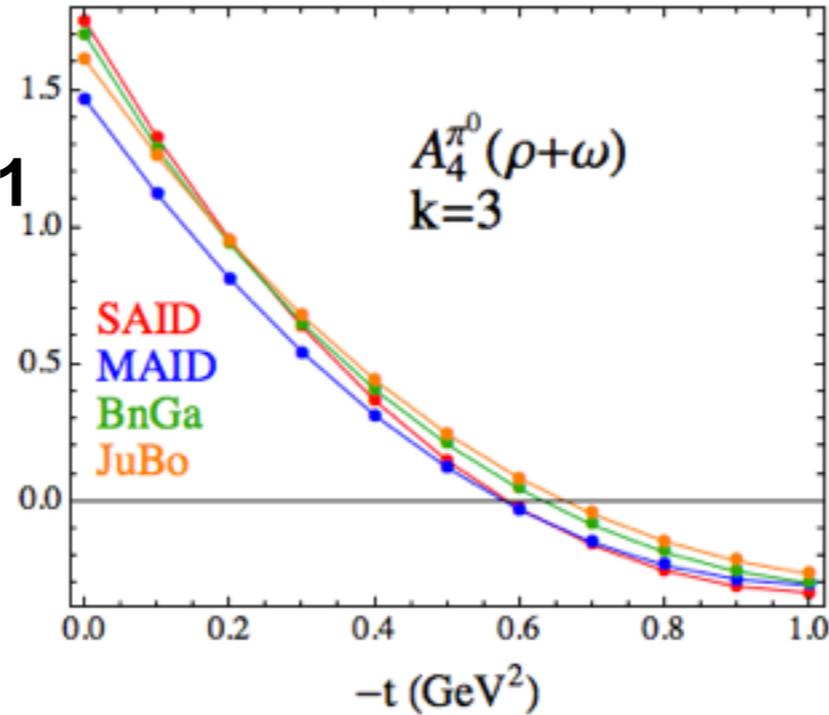


VM et al (JPAC), in preparation

Agreed on dominant A4

Variation on subdominant A1

$$\frac{1}{\Lambda^k} \int_{\nu_0}^{\Lambda} \text{Im } A(\nu, t) \nu^k d\nu = \frac{\beta(t) \Lambda^{\alpha(t)+1}}{\alpha(t) + k + 1}$$



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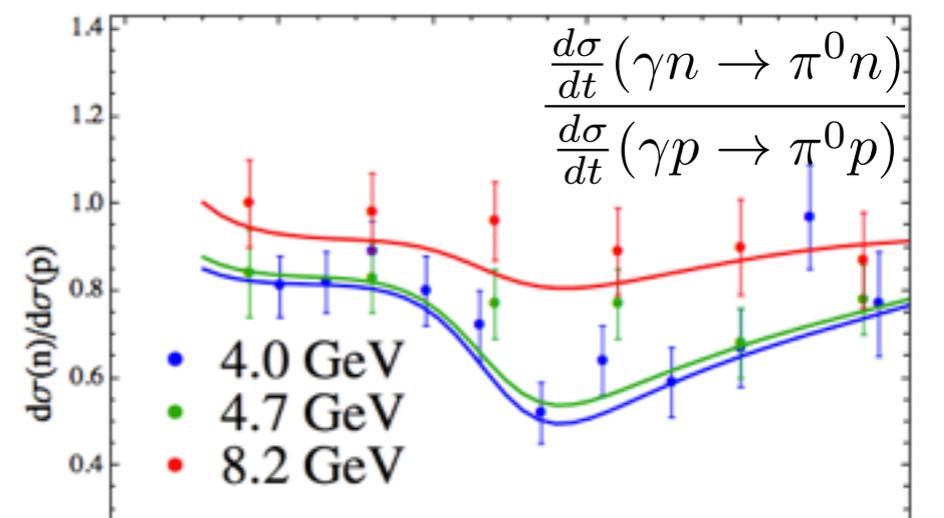
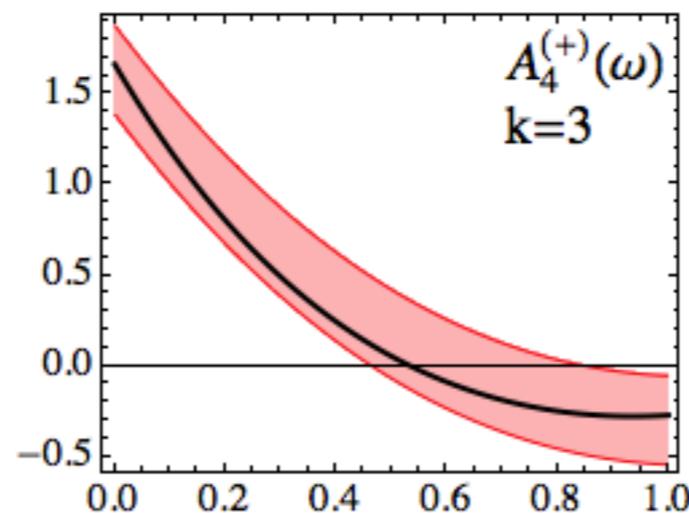
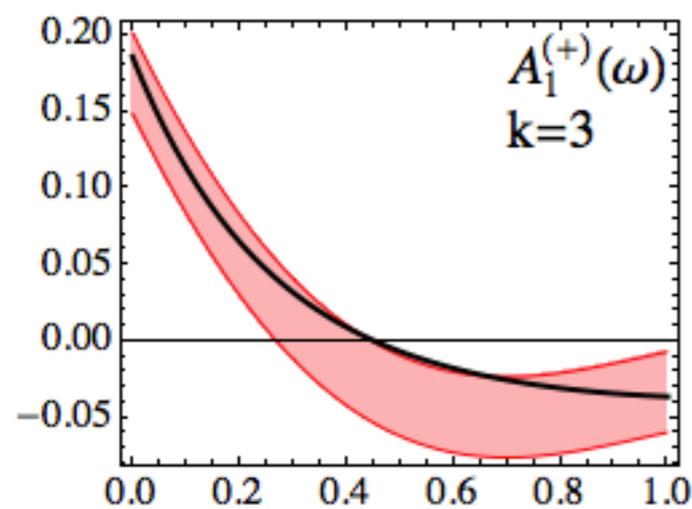
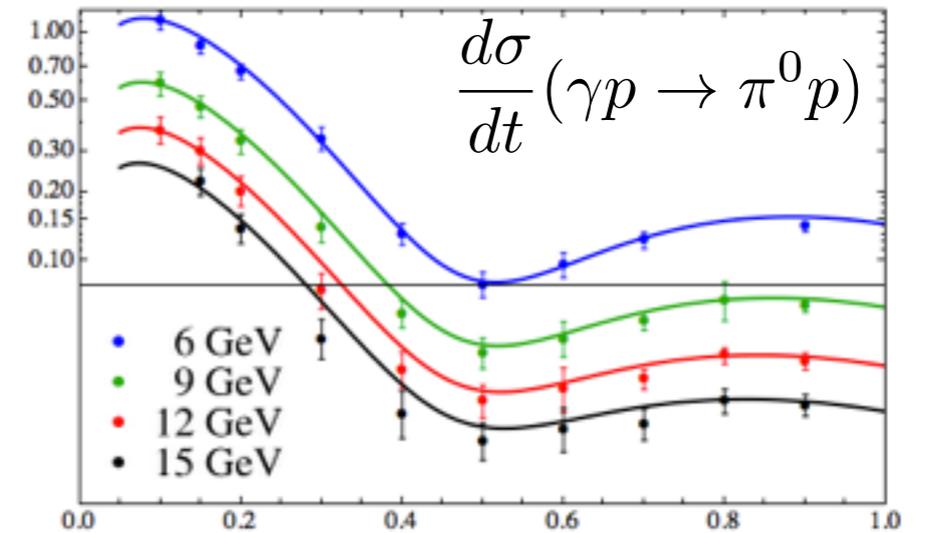
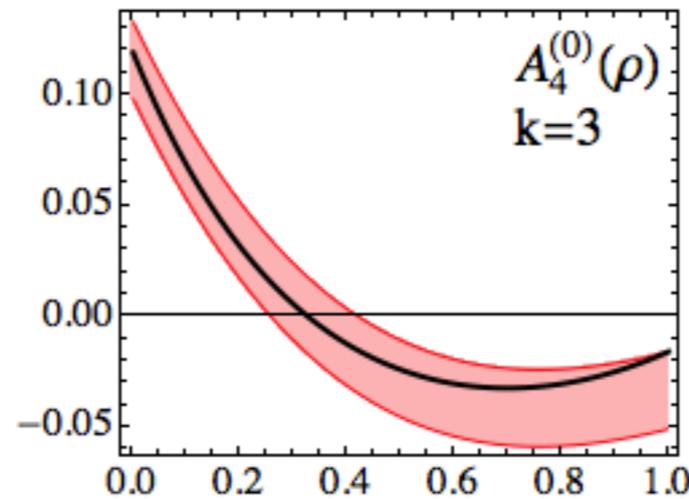
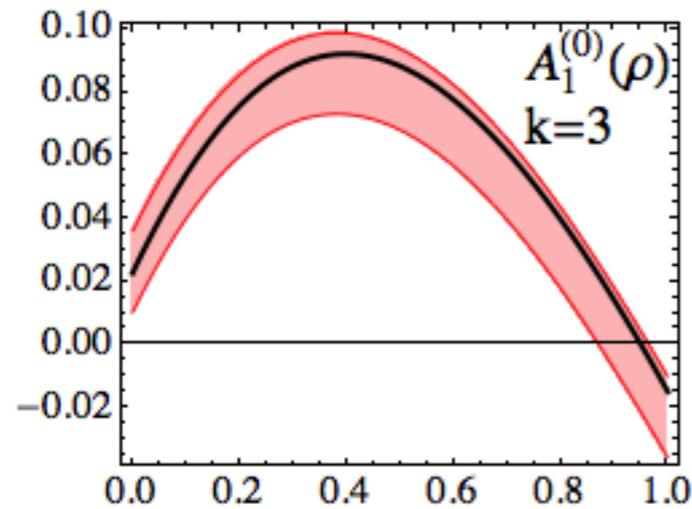
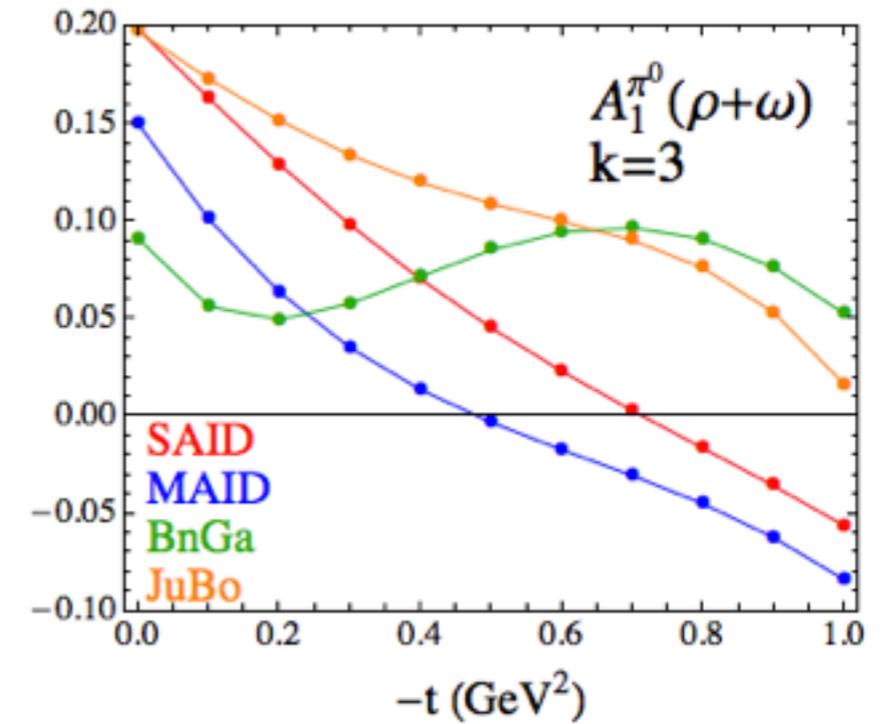
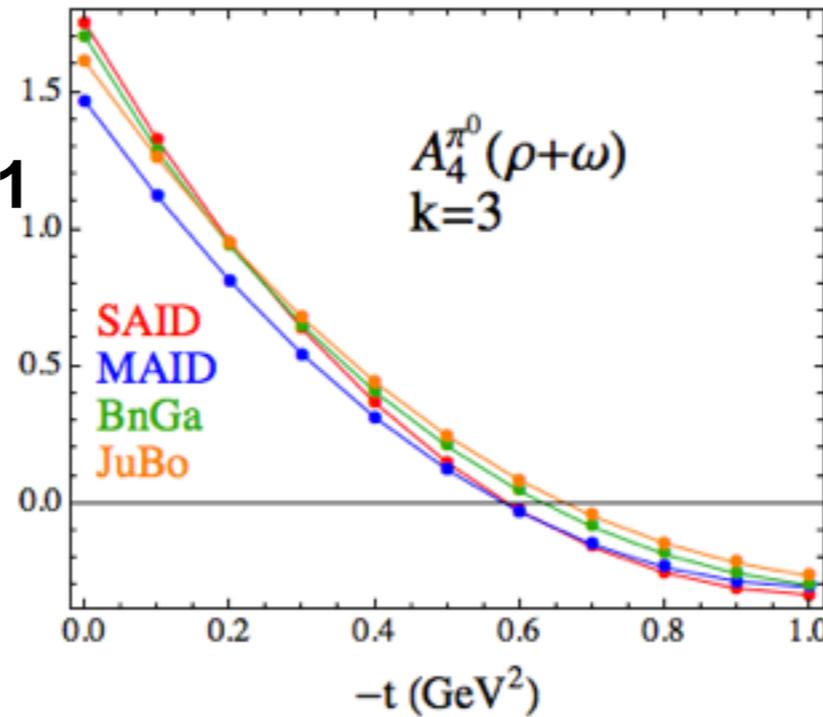


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Use SAID to fit
FESR and observable



Application to Pseudoscalar Photoproduction

$$\gamma N \rightarrow \eta N$$

$$(\pm 1) \left(\pm \frac{1}{2} \right) \rightarrow 0 \left(\pm \frac{1}{2} \right)$$

**8 helicity configurations
related by pair via parity**

**4 indep. helicity configurations
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every amplitude has an isospin index (s,v)**

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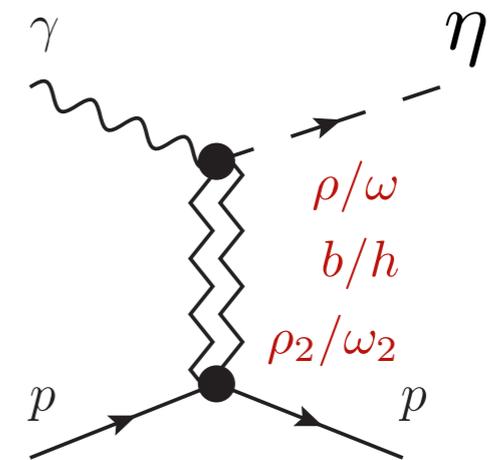
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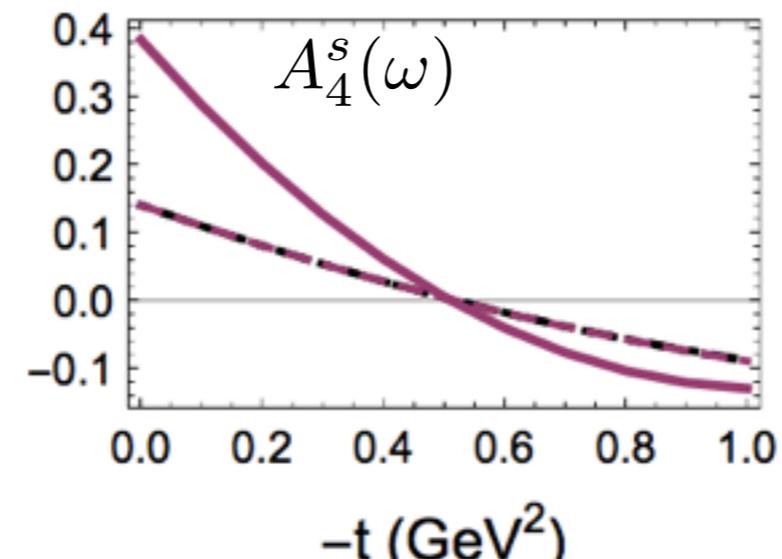
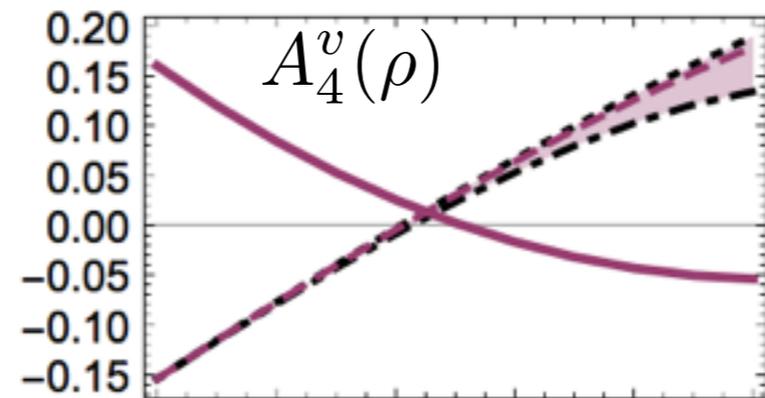
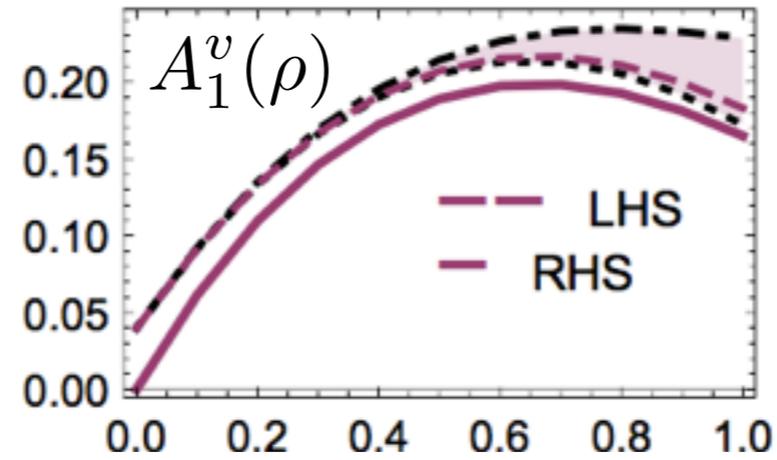
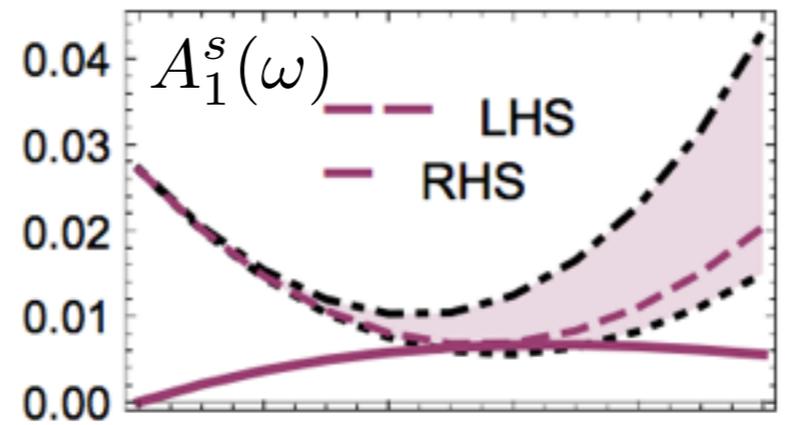
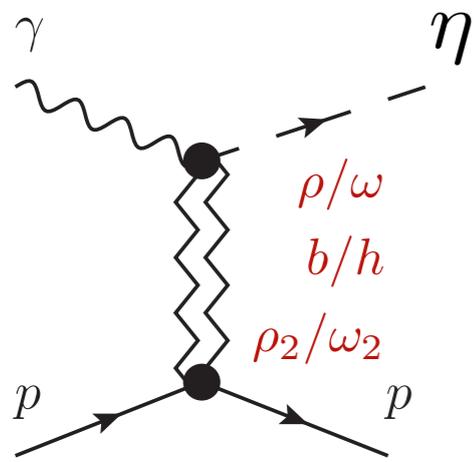
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8 helicity configurations
related by pair via parity

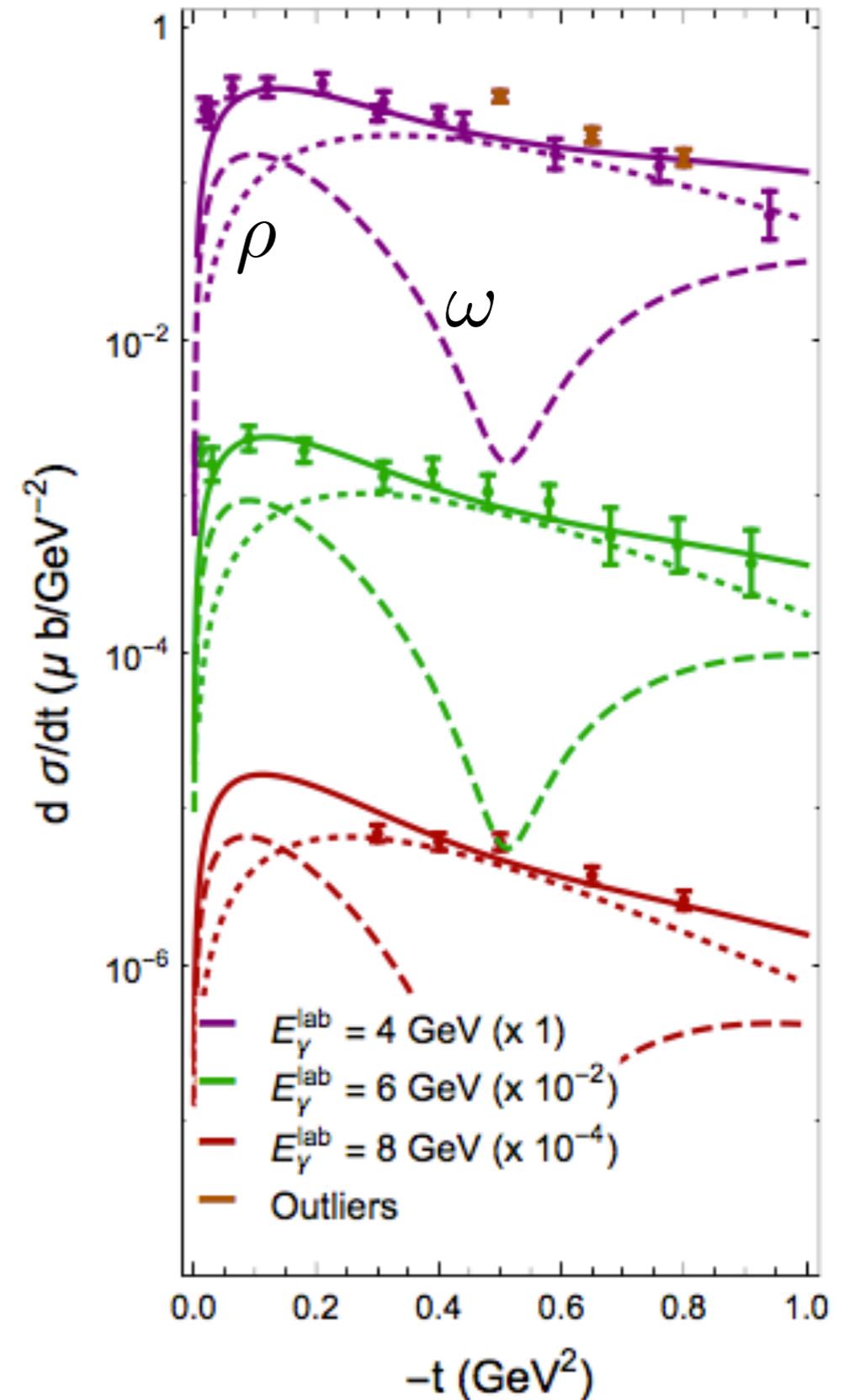
4 indep. helicity configurations
use CGLN basis A_1, \dots, A_4



| A_i | v | s |
|-------|----------|------------|
| A_1 | ρ | ω |
| A_2 | b | h |
| A_3 | ρ_2 | ω_2 |
| A_4 | ρ | ω |



$$\gamma p \rightarrow \eta p$$



Dashed lines: MAID 2001

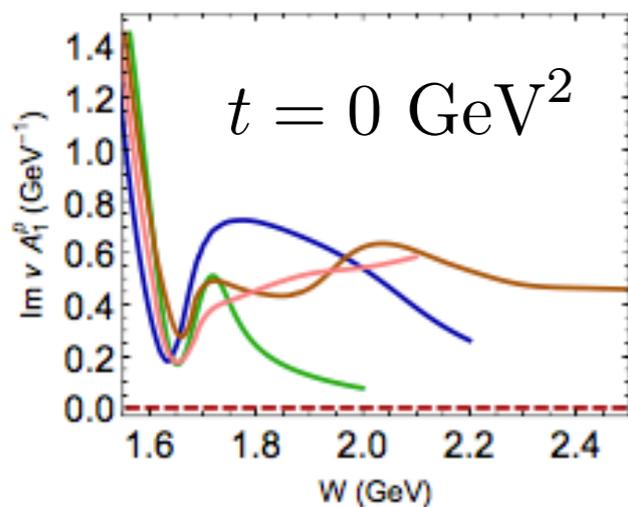
Solid lines: Regge

Not perfect agreement

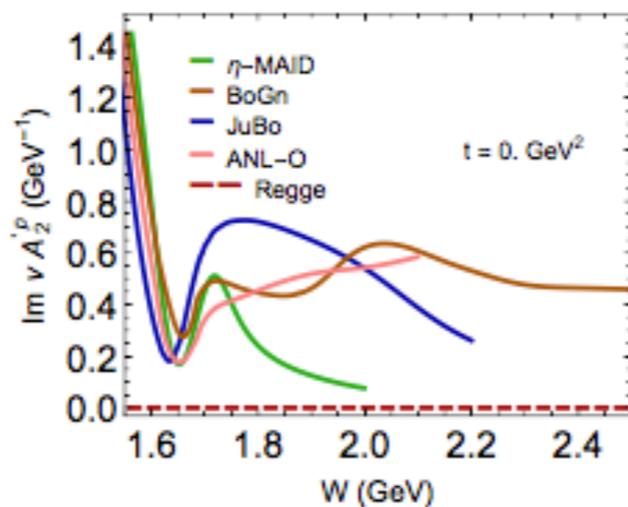
But similar features

Amplitude Comparison for $\gamma p \rightarrow \eta p$

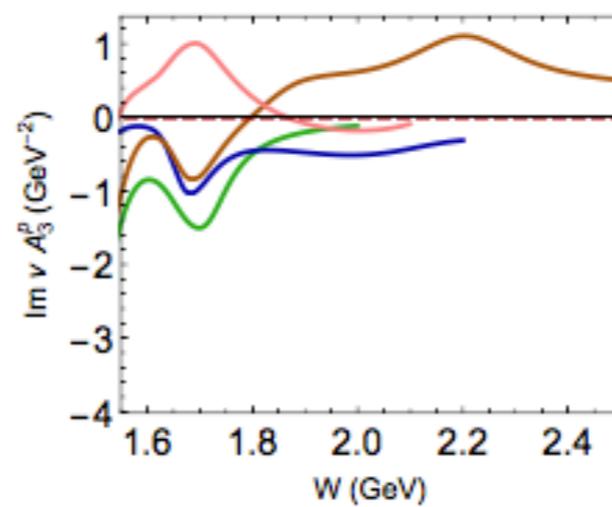
A_1



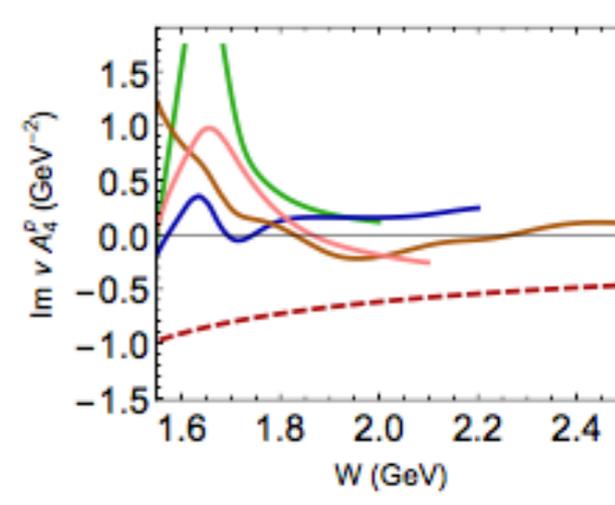
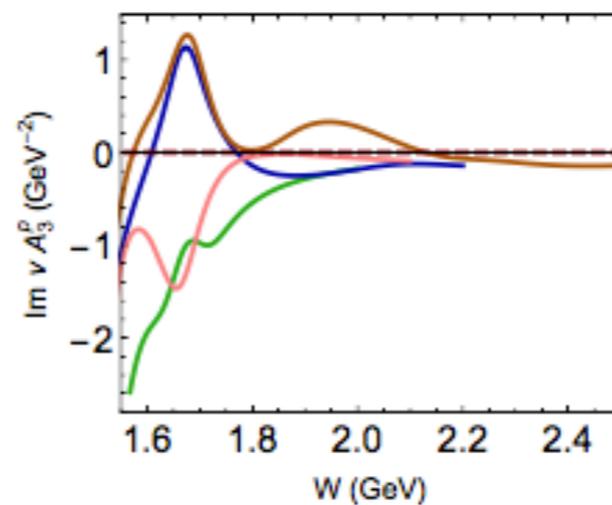
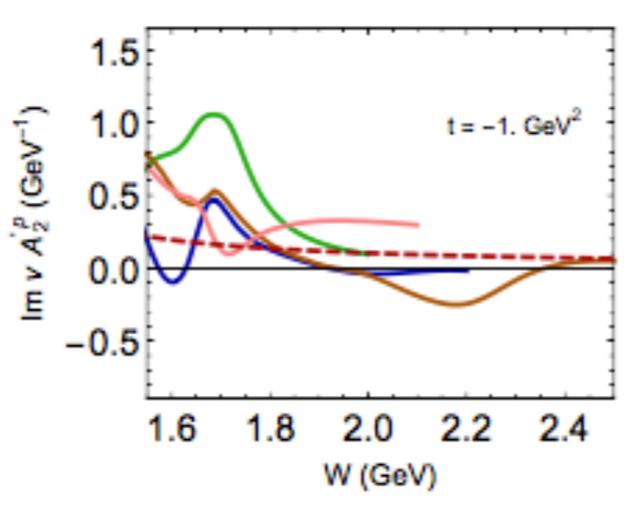
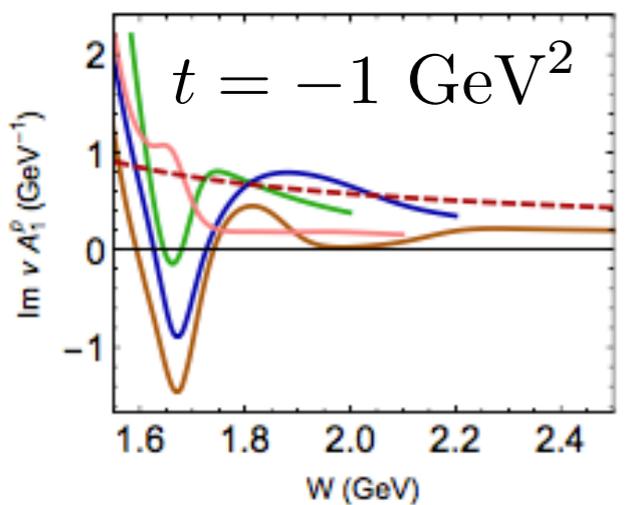
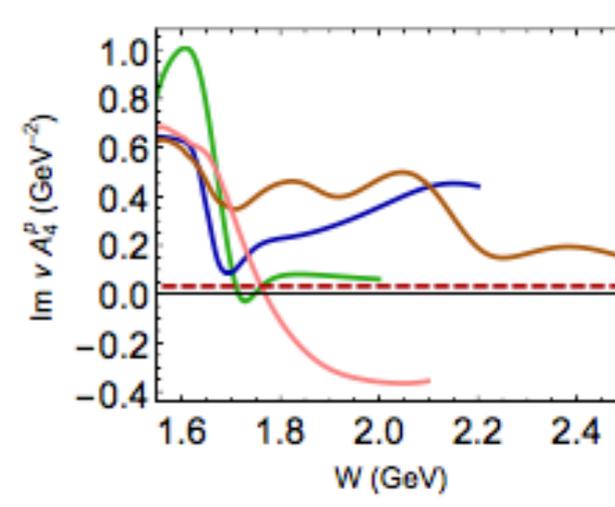
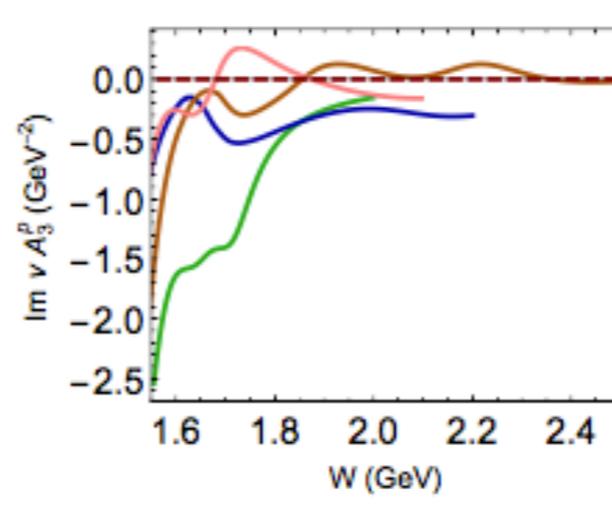
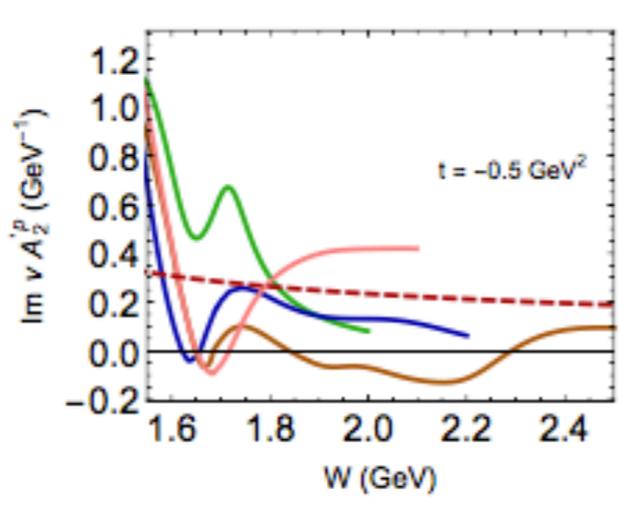
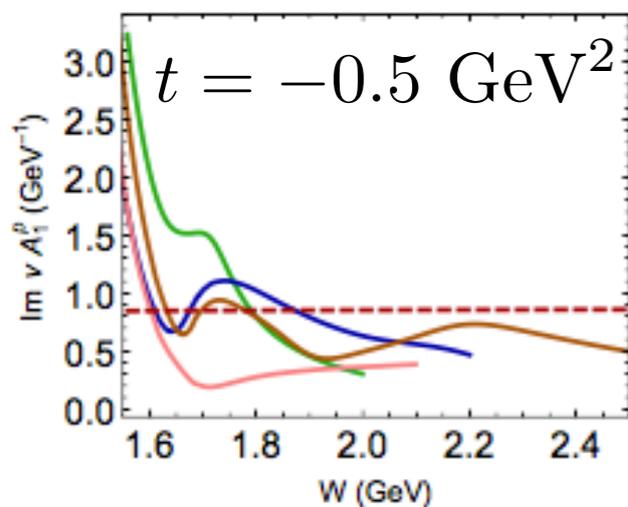
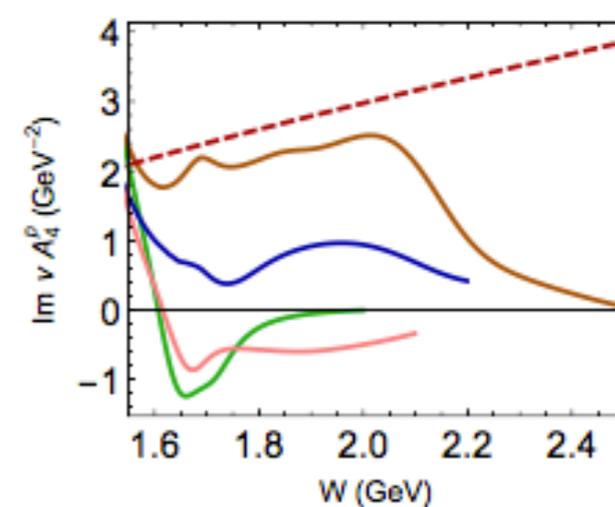
A_2



A_3



A_4



Joint Physics Analysis Center

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JPAC acknowledges support from DOE and NSF



Interactive webpage:

<http://www.indiana.edu/~jpac/index.html>



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- **Emilie Passemar** Professor
- **Tim Londergan** Professor
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- **Ina Lorenz** Postdoctoral researcher
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Bonn University

- **Misha Mikhasenko** PhD student

University of Valencia

- **Astrid Hiller Blin** PhD student

Ghent University

- **Jannes Nys** PhD student



November 2016:

- The $\gamma p \rightarrow \eta p$ page is online.

June 2016:

- The $\gamma p \rightarrow J/\psi p$ page is online.
- The πN page is online.

October 2015:

- The $\bar{K}N$ page is online.

May 2015:

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- The $\gamma p \rightarrow \pi^0 p$ page is online.
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Resources

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- **C/C++ observables:** C-code main, Input file, C-code source, C-code header, Eta-MAID 2001 multipoles
- **C/C++ minimal script to calculate the amplitudes:** C-code zip
- **Data:** Dewire , Braunschweig
- **Contact person:** Jannes Nys
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Run the code

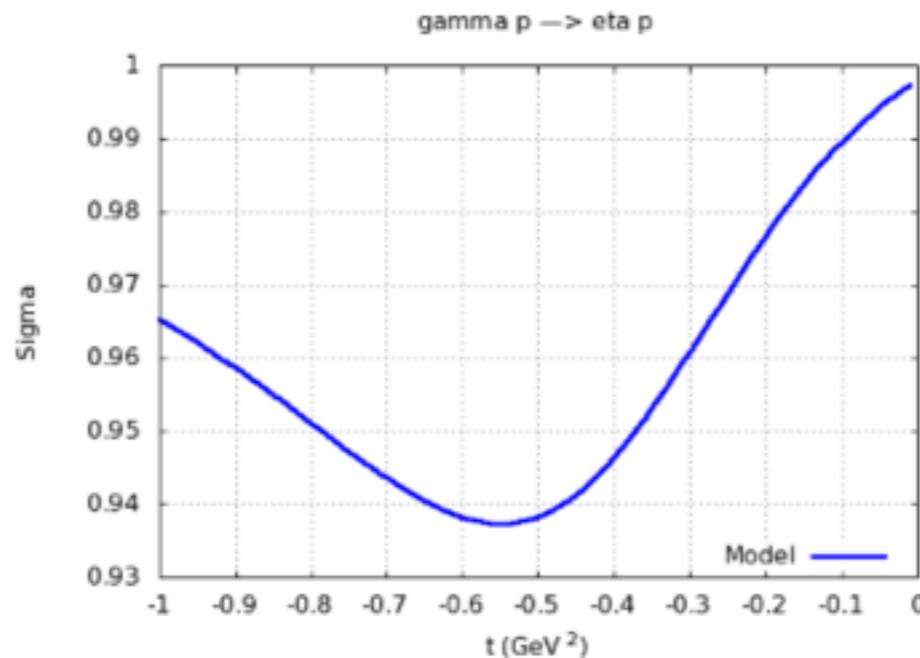
E_γ in GeV

t cos

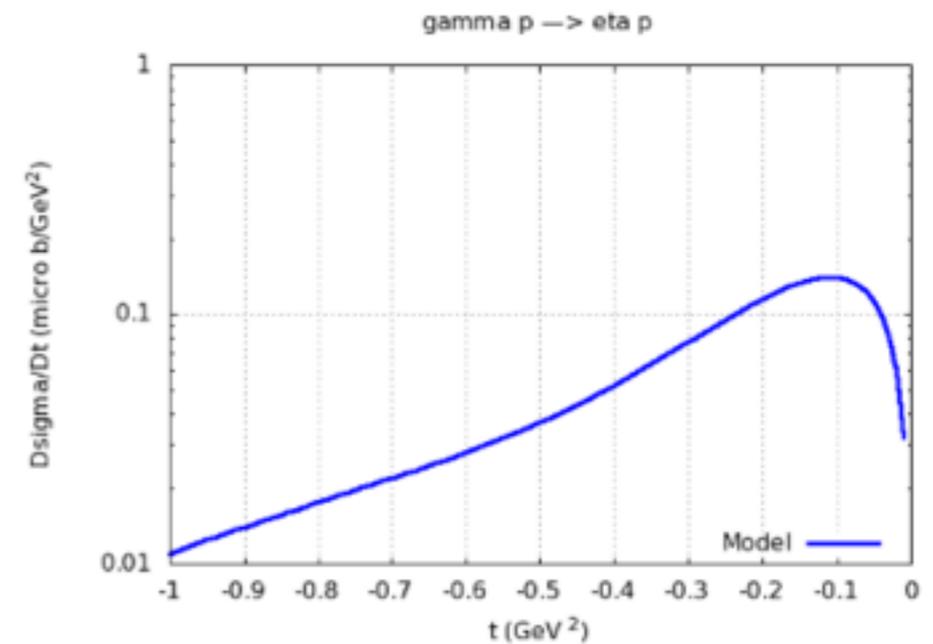
t in GeV^2 (min max step)

$\cos \theta$ (min max step)

Observable: photon beam asymmetry
 Download the the plot with $O_x=t$, the plot with $O_x=\cos$.



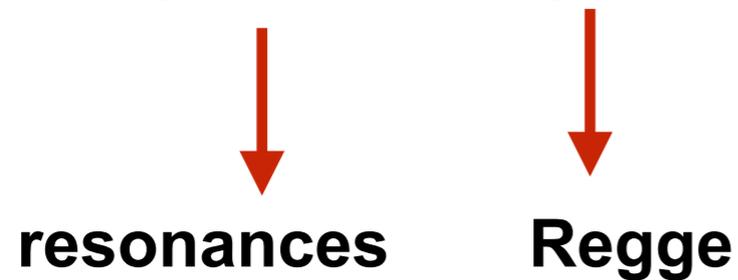
Observable: differential cross section
 Download the the plot with $O_x=t$, the plot with $O_x=\cos$.



Methodology: A

Choose one single channel: $\gamma p \rightarrow \pi^0 p$

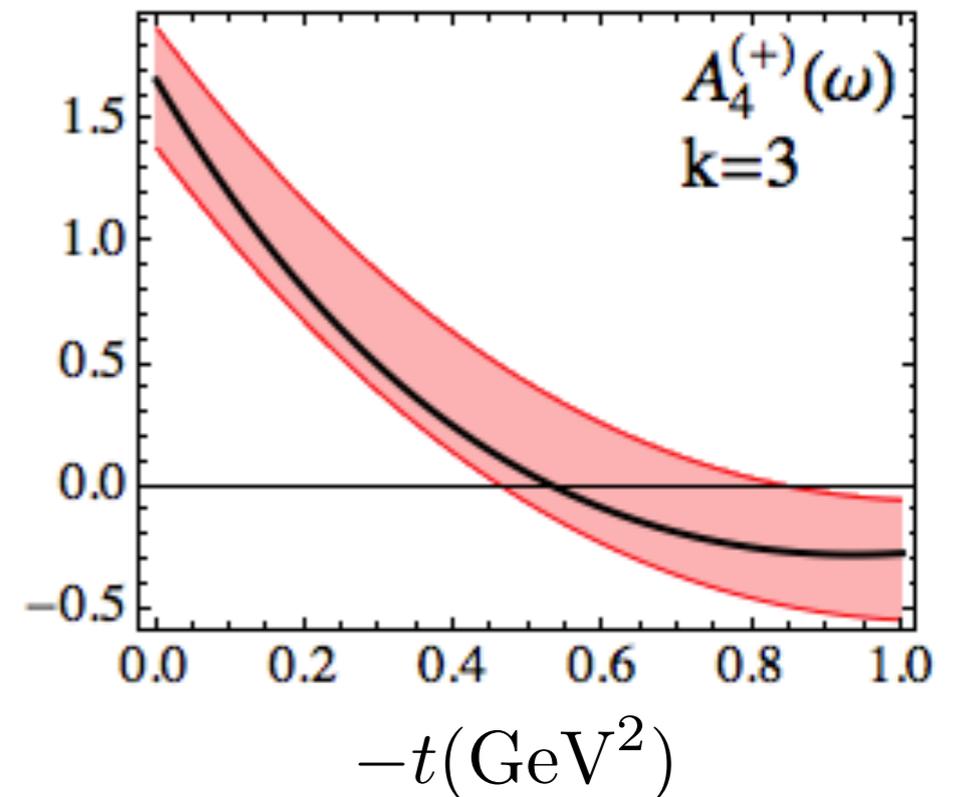
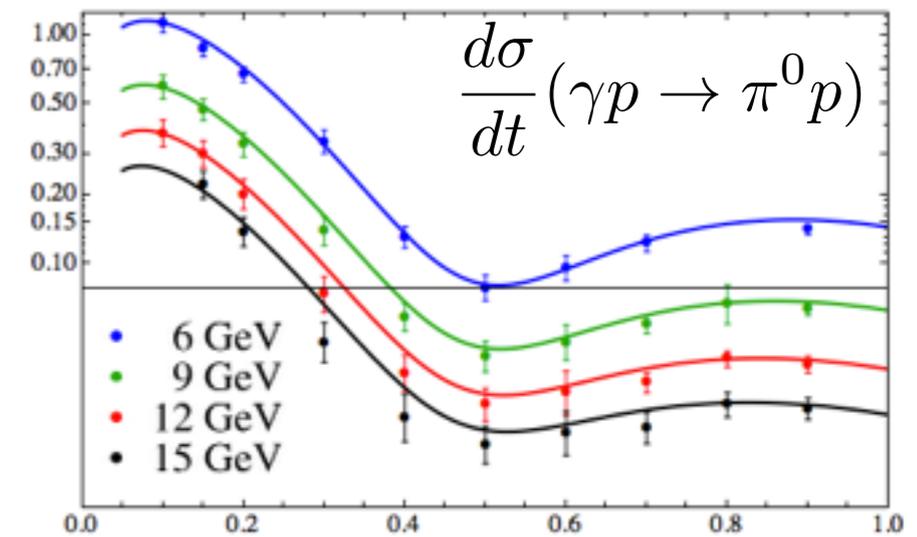
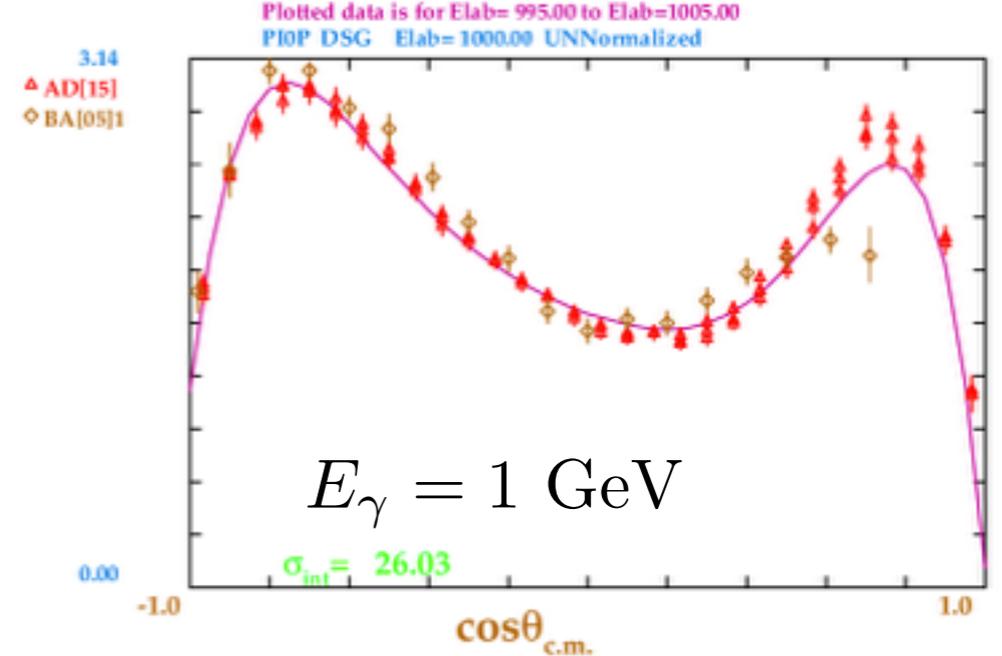
Fit independently low and high energy data



Compare amplitudes

Impose FESR as penalty function in fit

$$\frac{1}{\Lambda^k} \int_{\nu_0}^{\Lambda} \text{Im} A(\nu, t) \nu^k d\nu = \frac{\beta(t) \Lambda^{\alpha(t)+1}}{\alpha(t) + k + 1}$$



Methodology: B

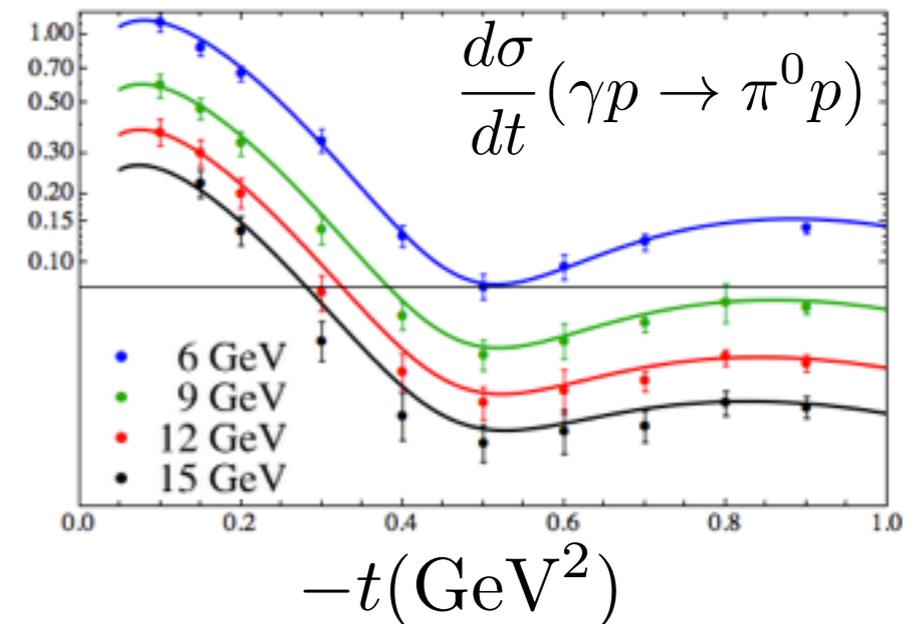
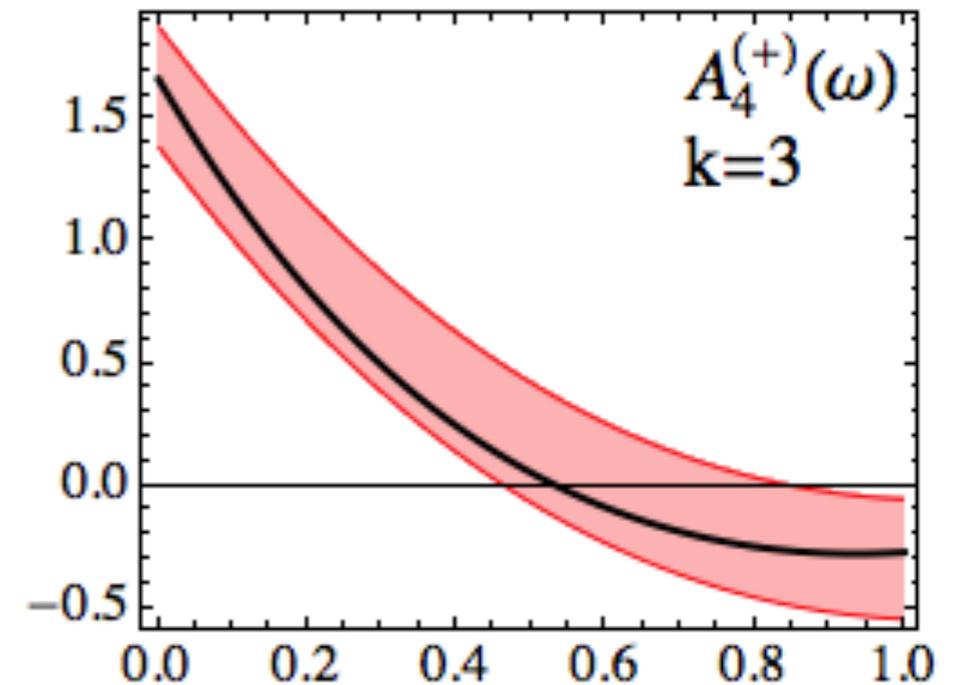
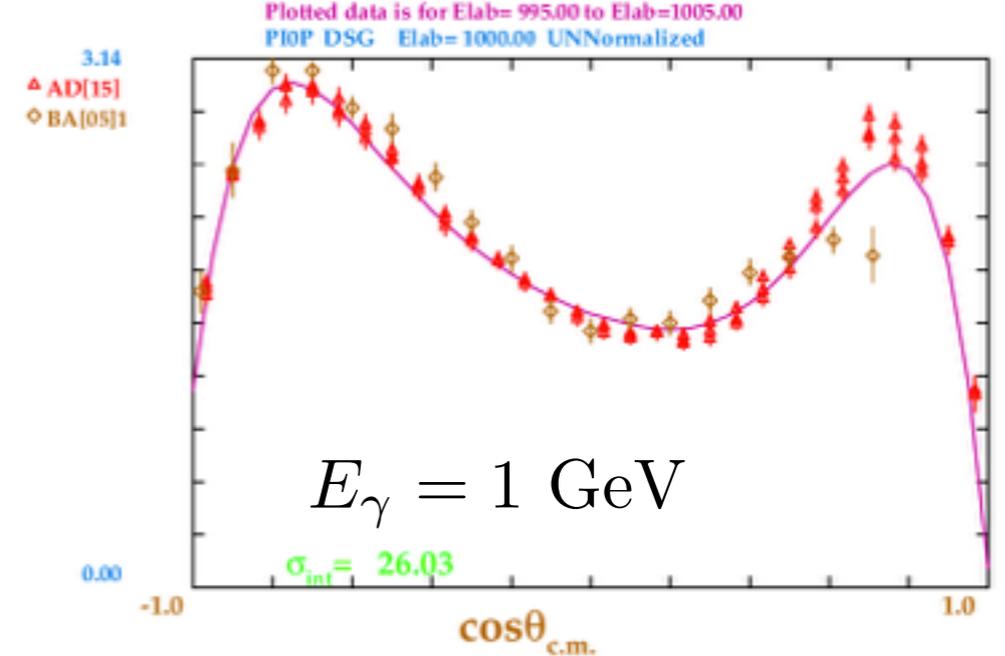
Choose one single channel: $\gamma p \rightarrow \pi^0 p$

Propose **imaginary part** (only real parameters)

Reconstruct real part from dispersion relation:

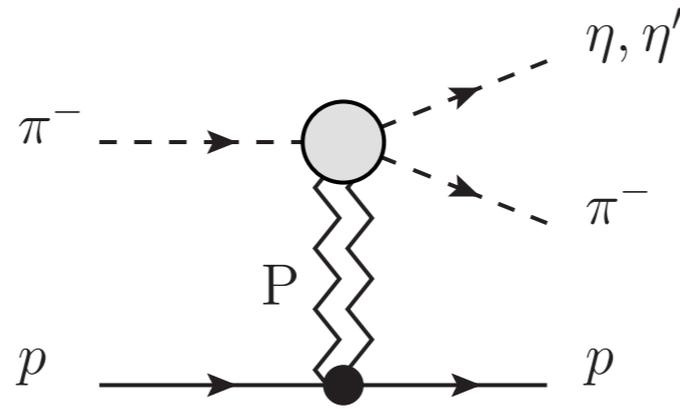
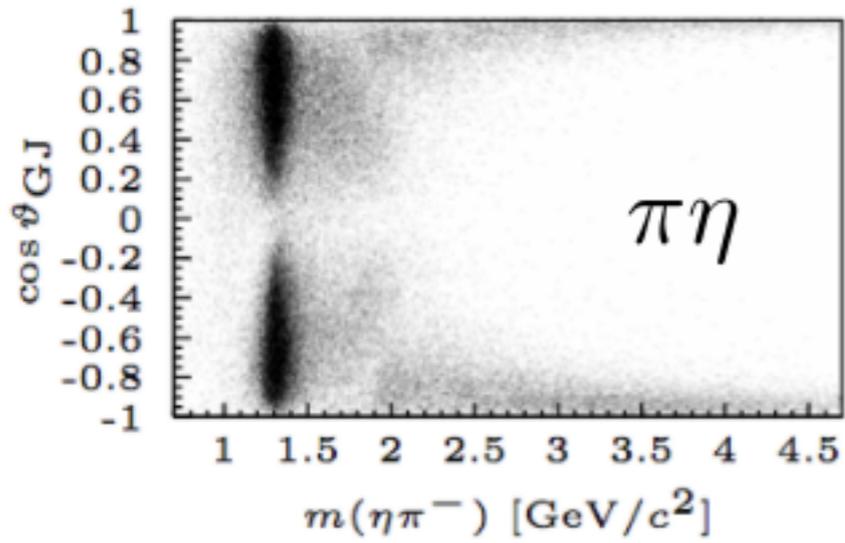
$$A(\nu, t) = \frac{2}{\pi} \int_{\nu_0}^{\infty} \frac{\text{Im } A(\nu', t)}{\nu'^2 - \nu^2} \nu' d\nu'$$

Fit all data and iterate



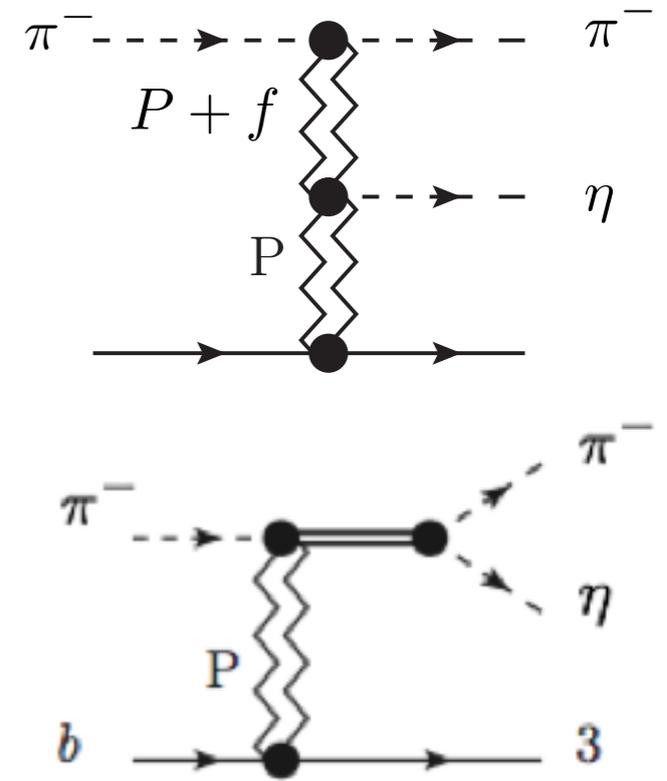
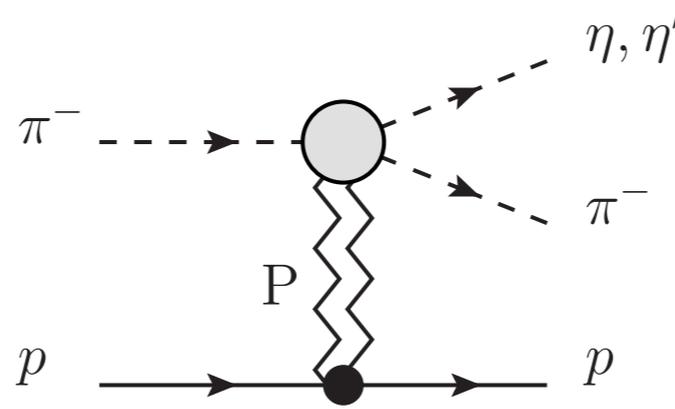
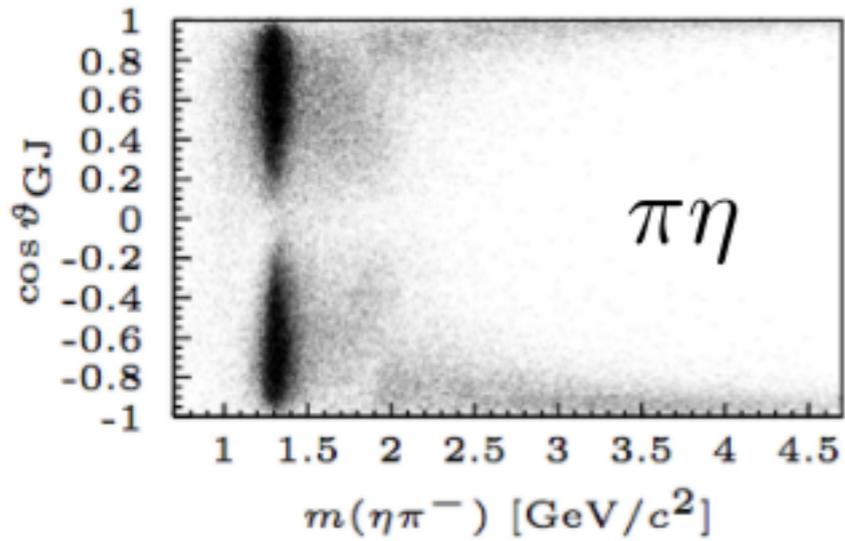
Going Further: Eta-Pi @COMPASS

COMPASS PLB740
ArXiv:1408.4286



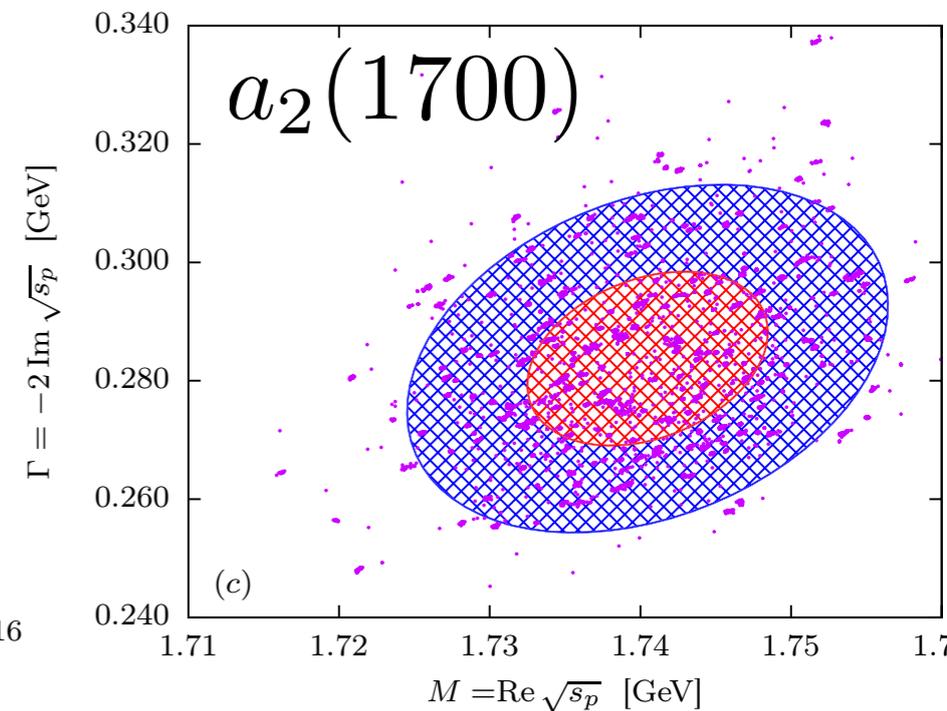
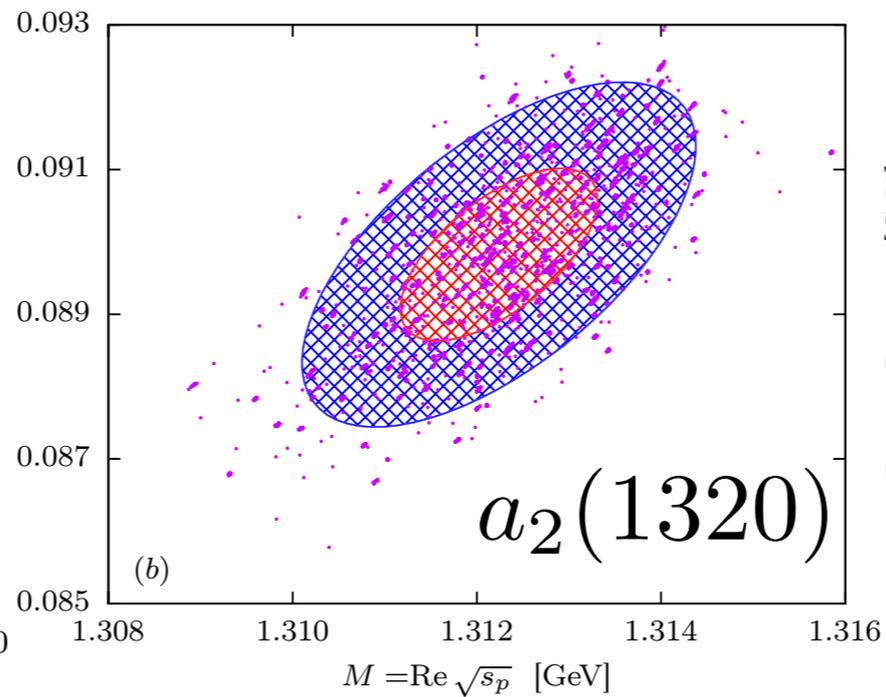
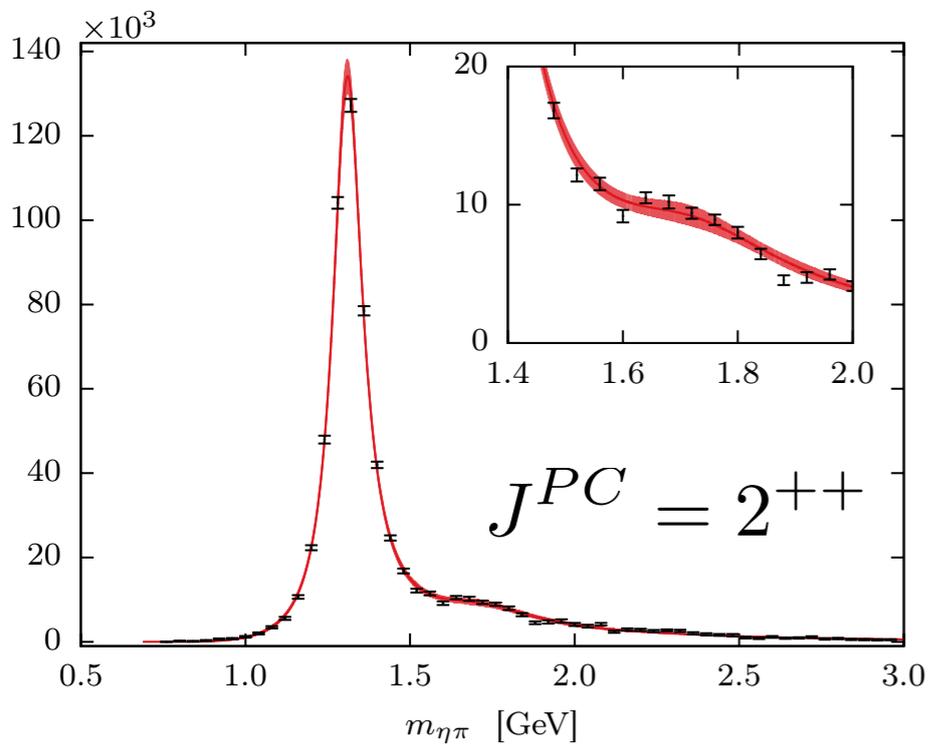
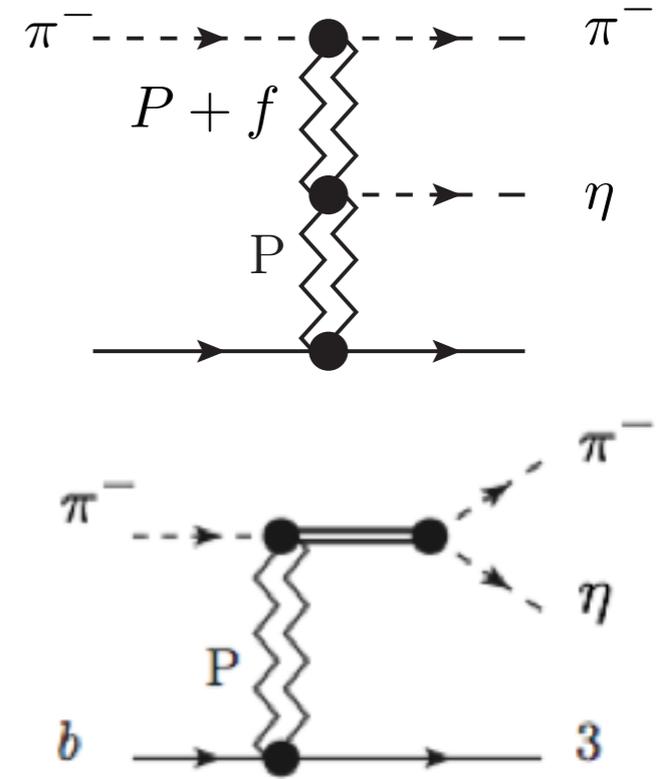
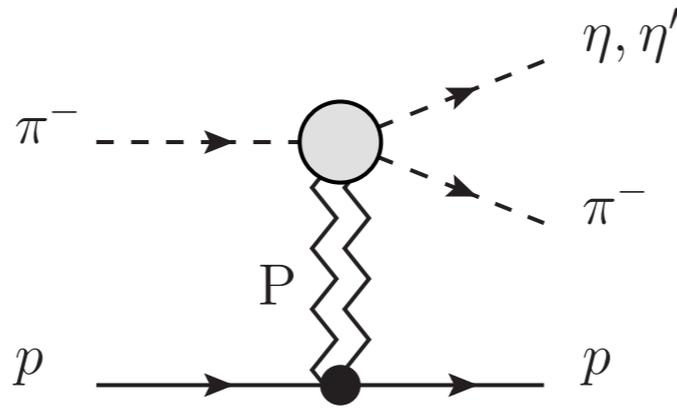
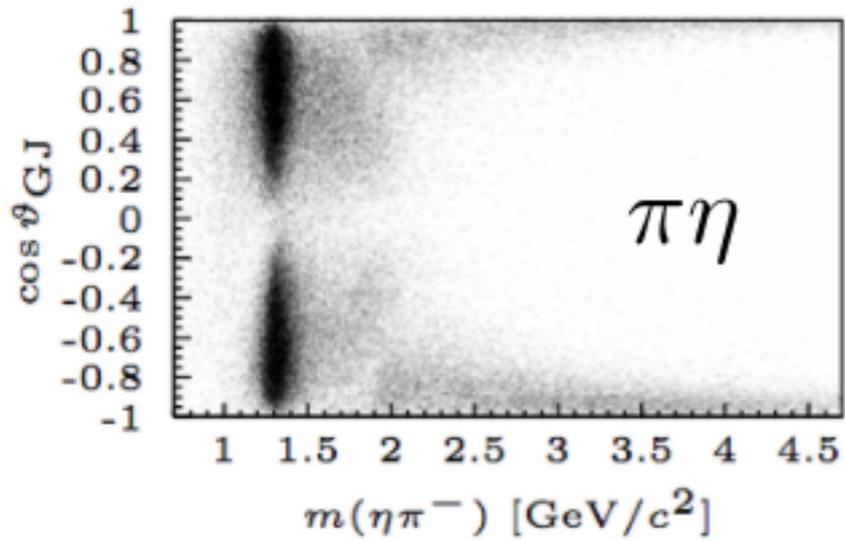
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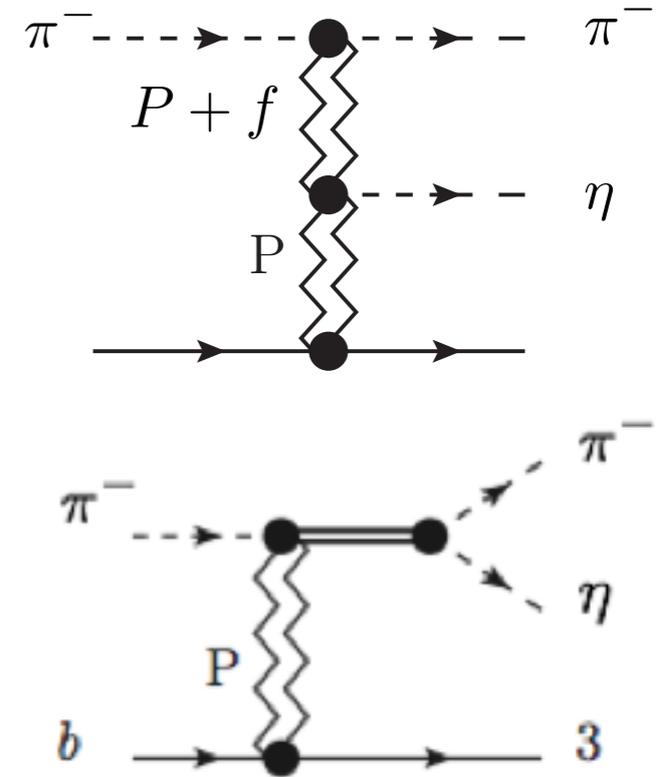
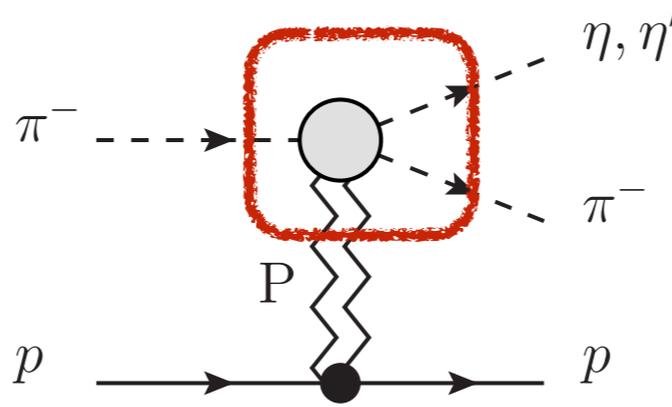
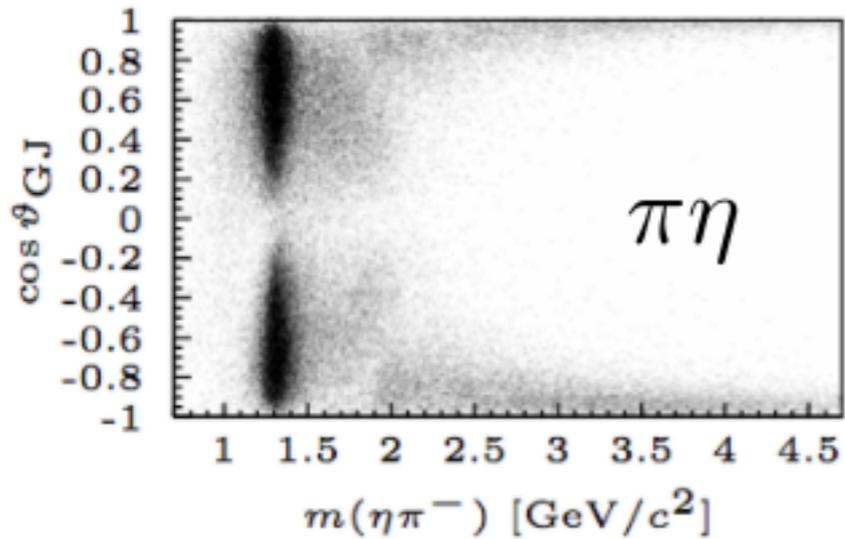
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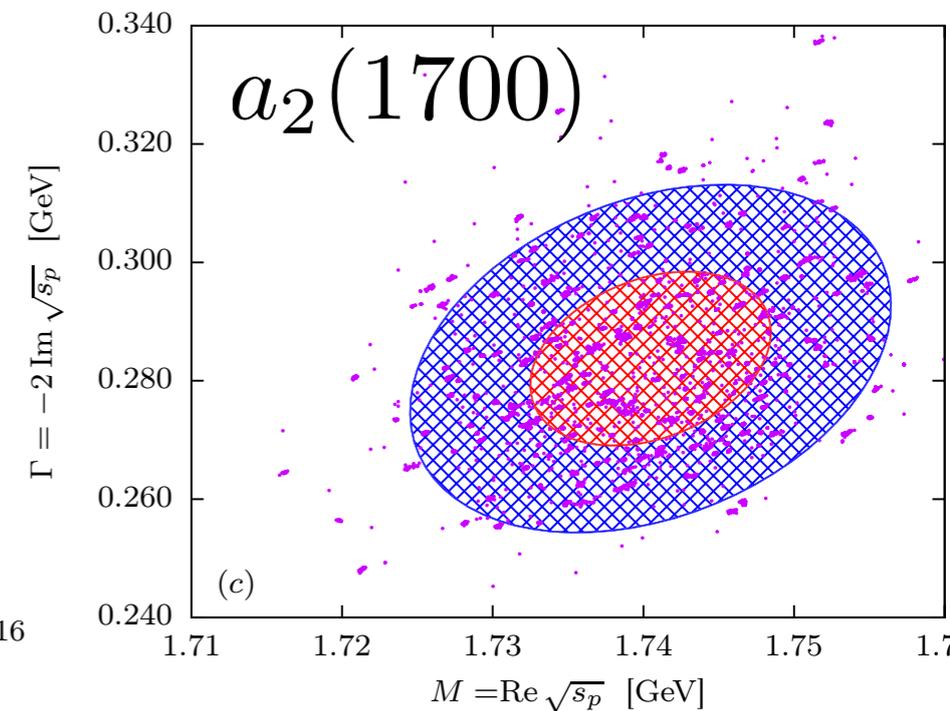
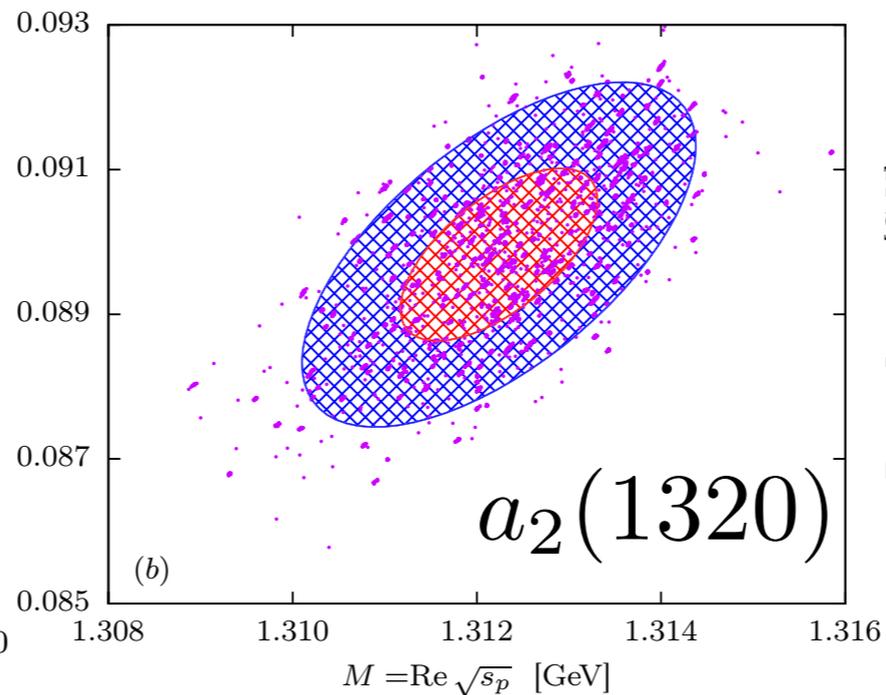
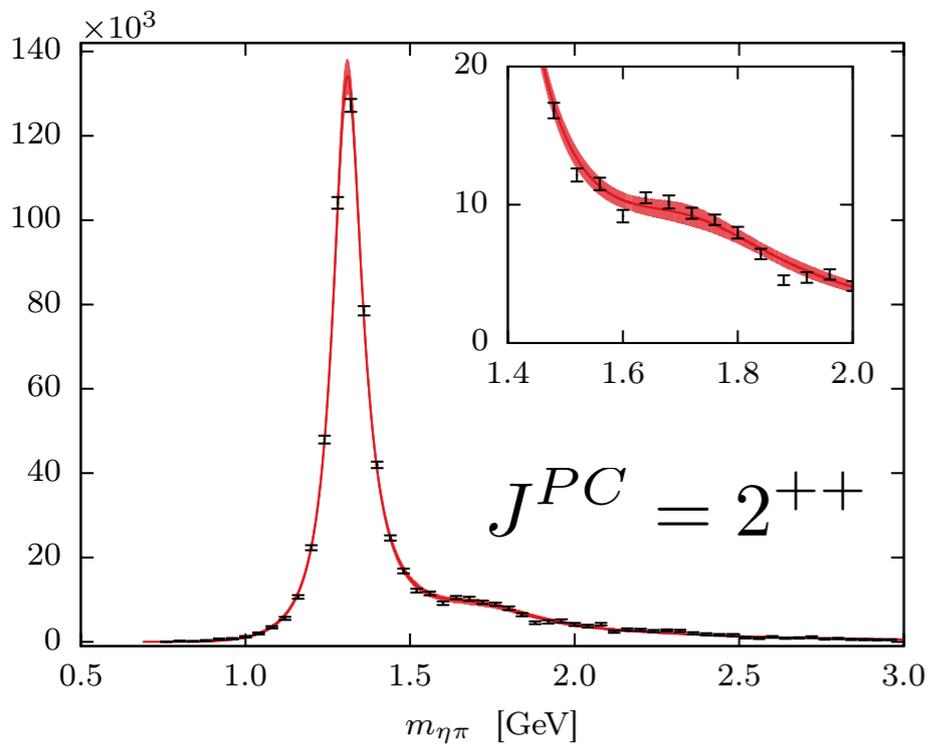


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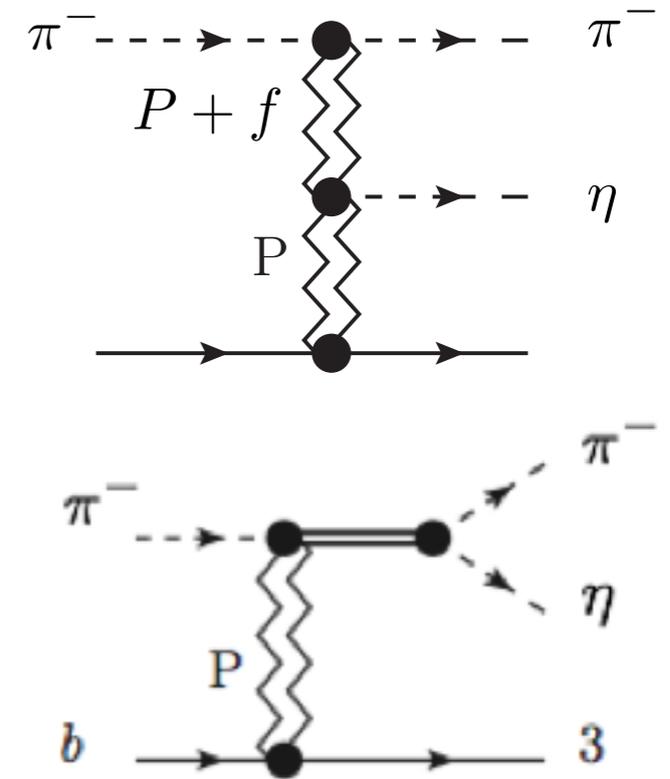
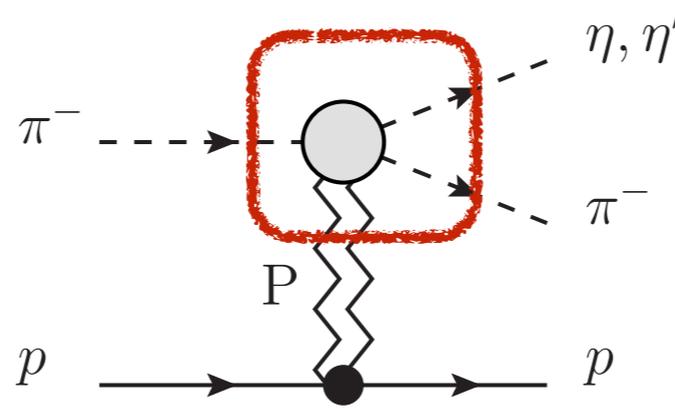
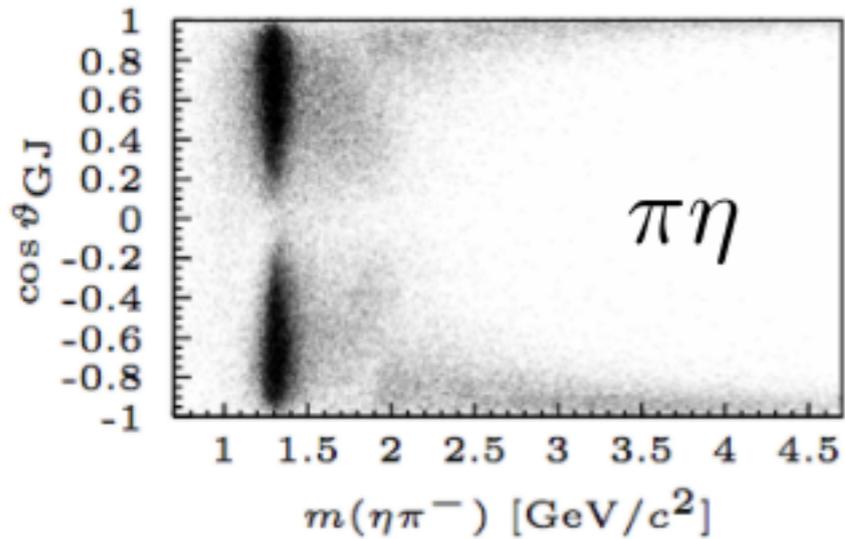


FESR with Reggeons could reduce uncertainties on pole parameters

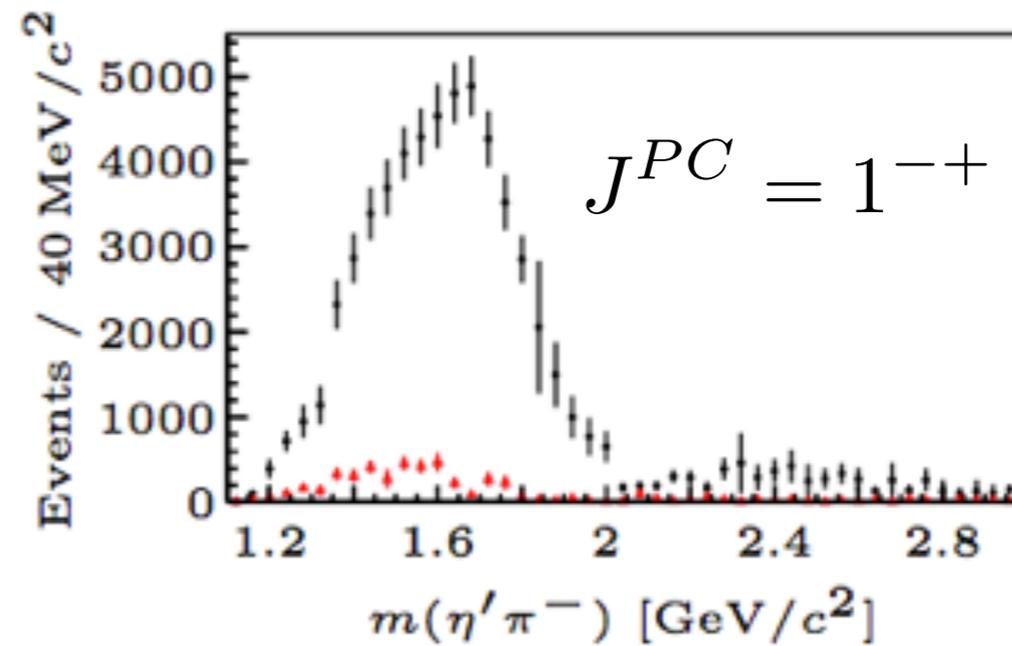
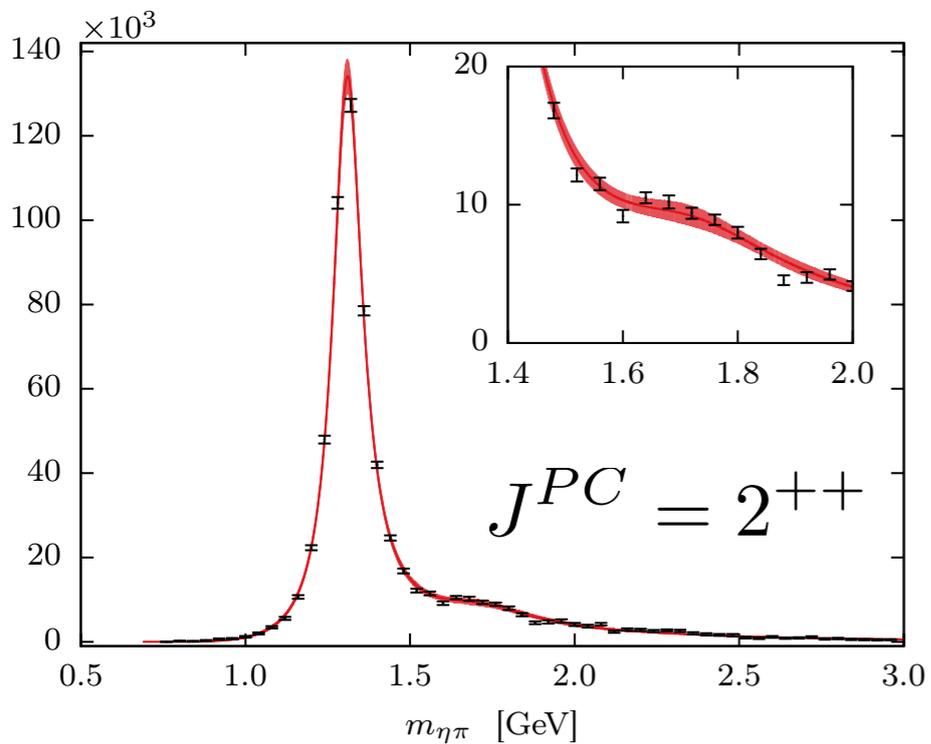


Going Further: Eta-Pi @COMPASS

COMPASS PLB740
ArXiv:1408.4286



**FESR with Reggeons could reduce uncertainties on pole parameters
constraint exotic production**



Summary: Methodology

Use constraints from analyticity: **FESR**

$$\frac{1}{\Lambda^k} \int_{\nu_0}^{\Lambda} \text{Im } A(\nu, t) \nu^k d\nu = \frac{\beta(t) \Lambda^{\alpha(t)+1}}{\alpha(t) + k + 1}$$

Great agreement for $\pi p \rightarrow \pi p$

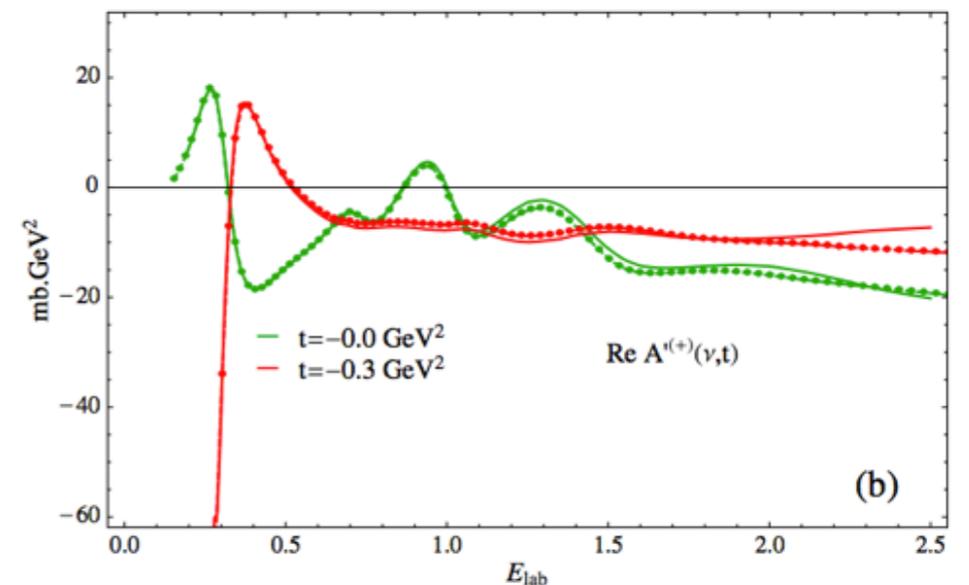
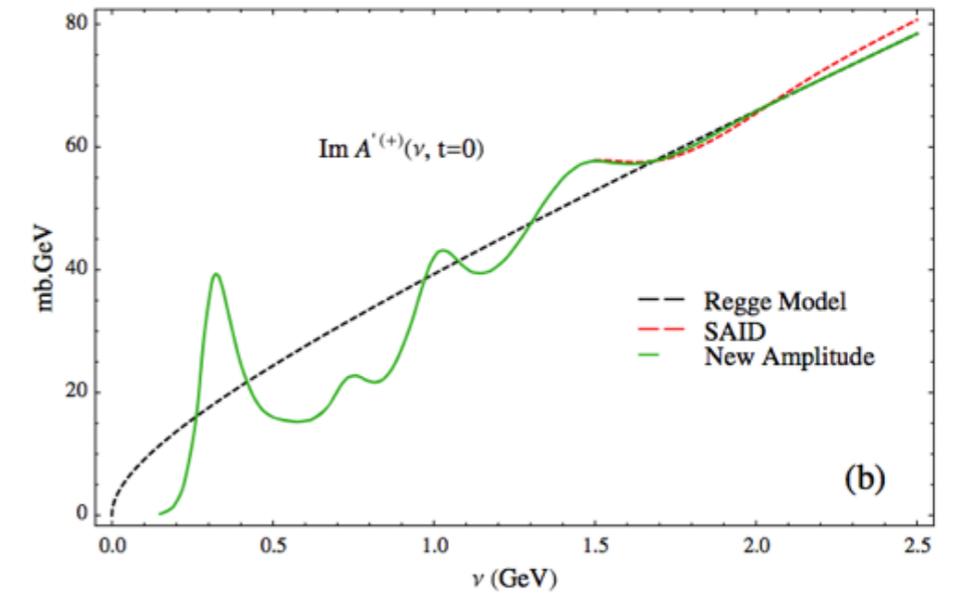
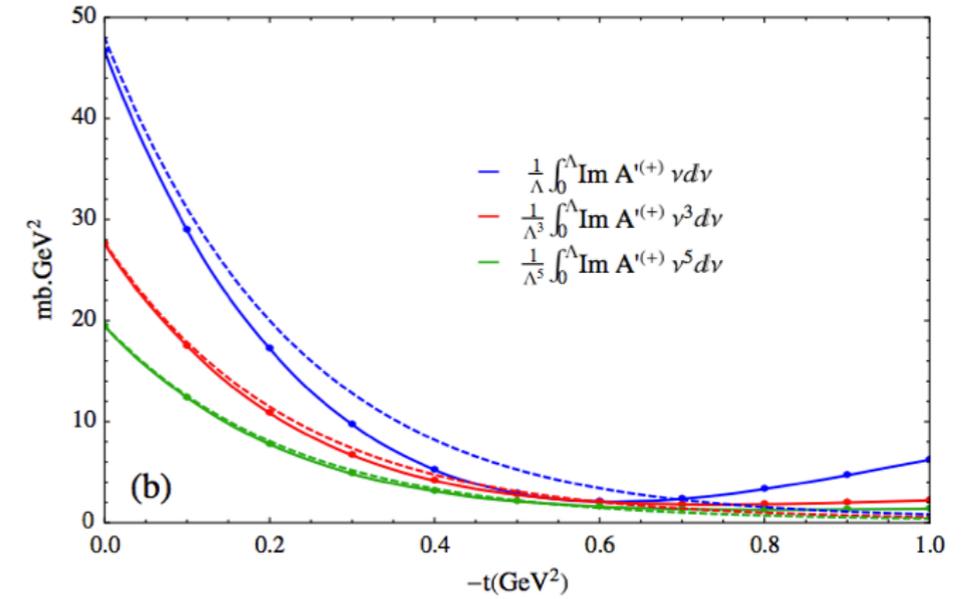
Possibly useful for $\gamma p \rightarrow \pi p$ $\gamma p \rightarrow \eta p$

to resolve ambiguities

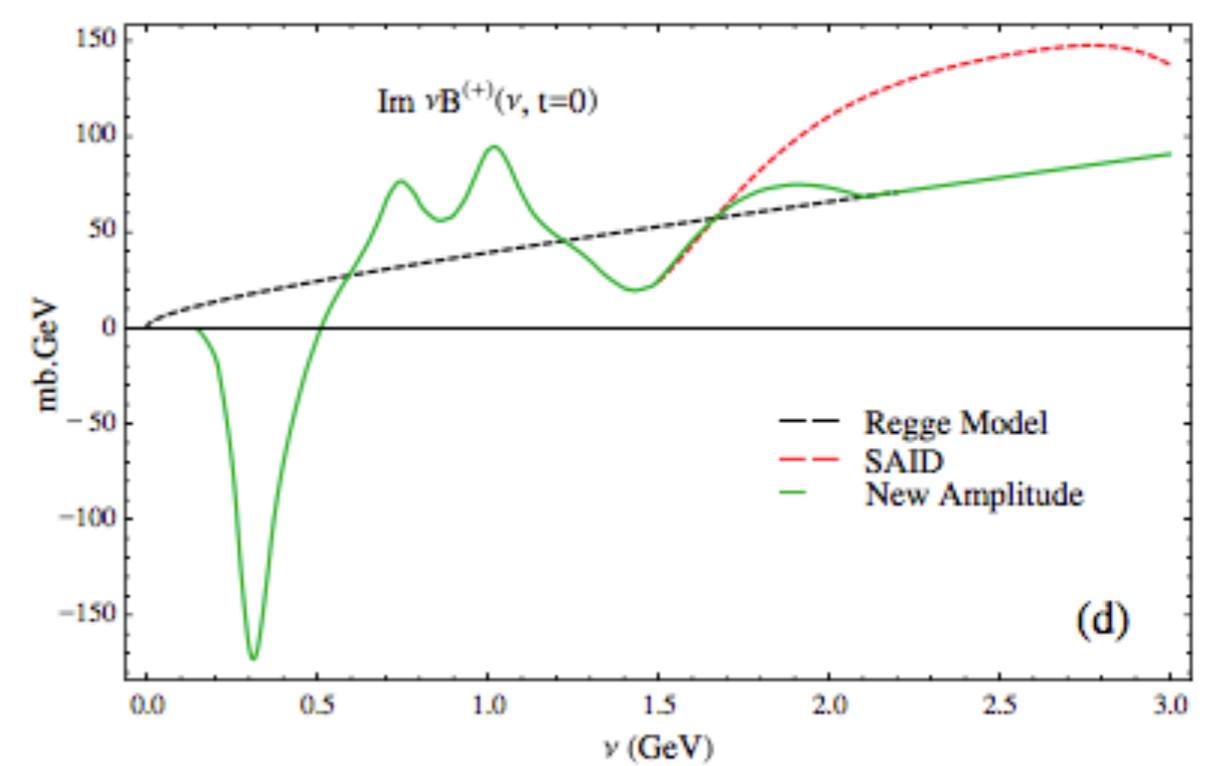
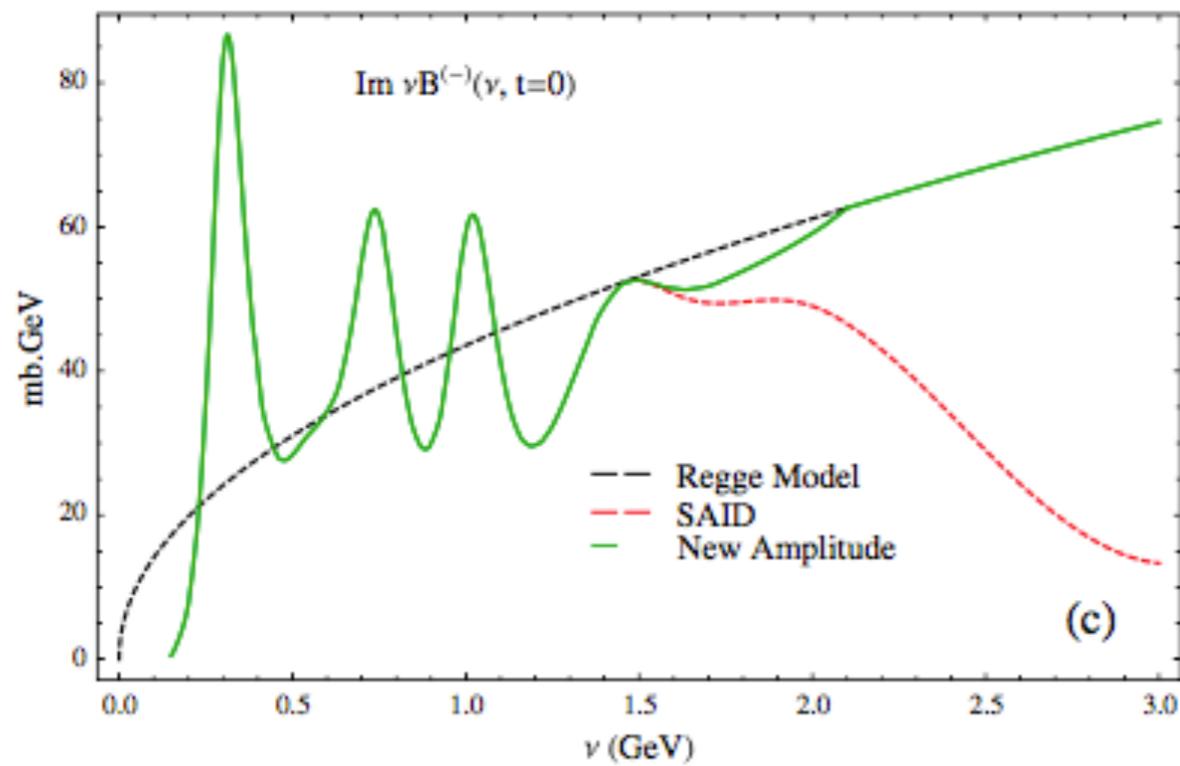
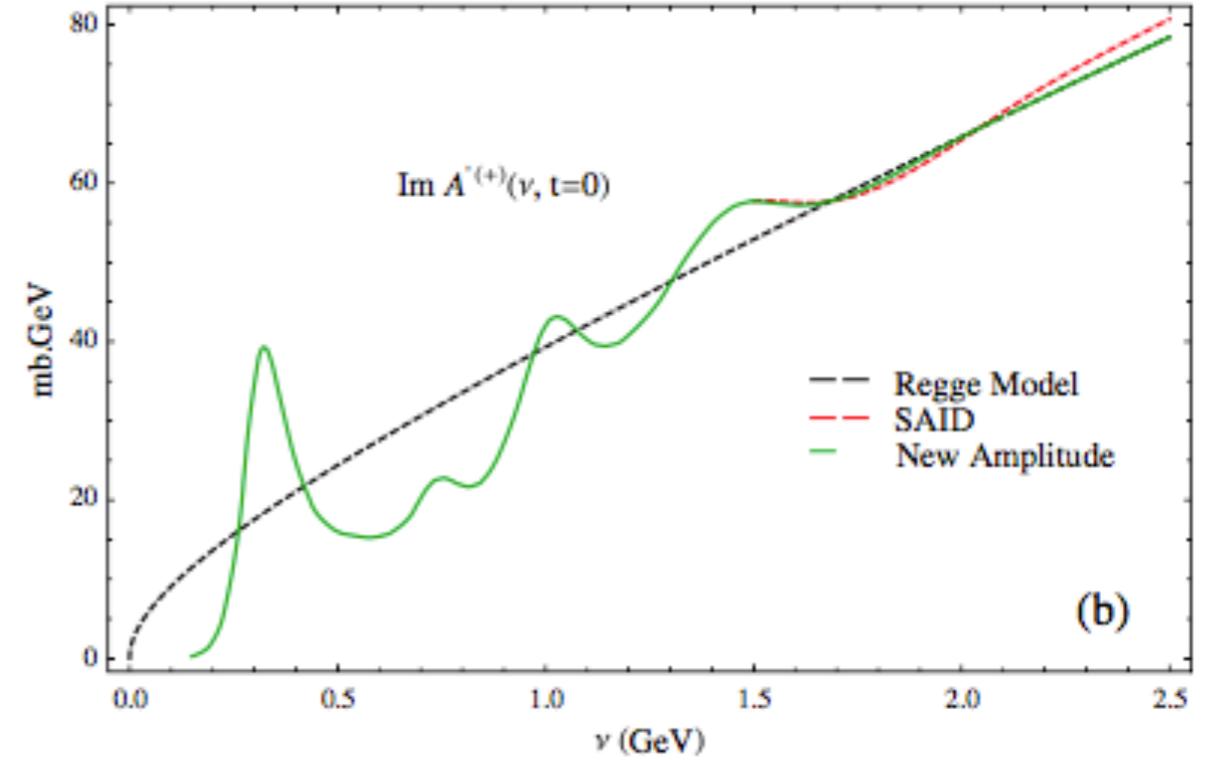
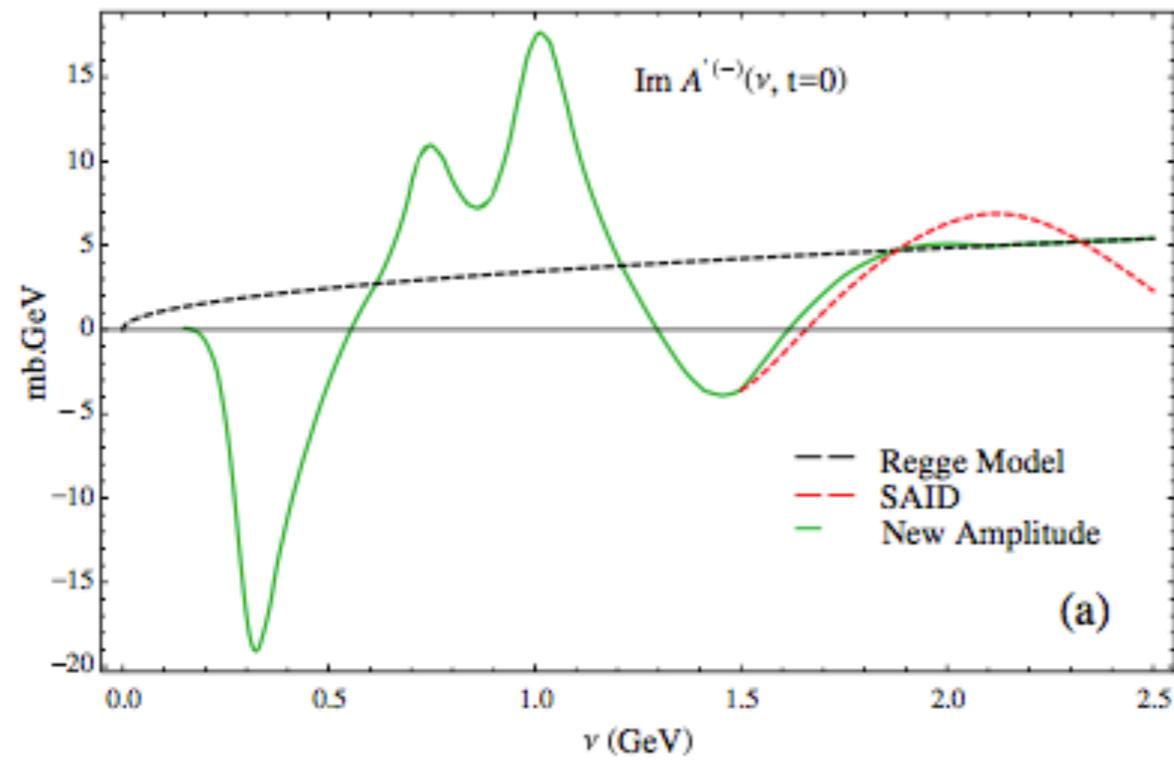
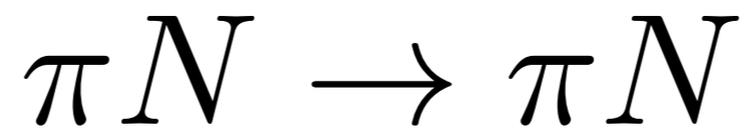
Work in progress for $KN \rightarrow KN$

VM et al (JPAC) PRD92
arXiv:1506.01764

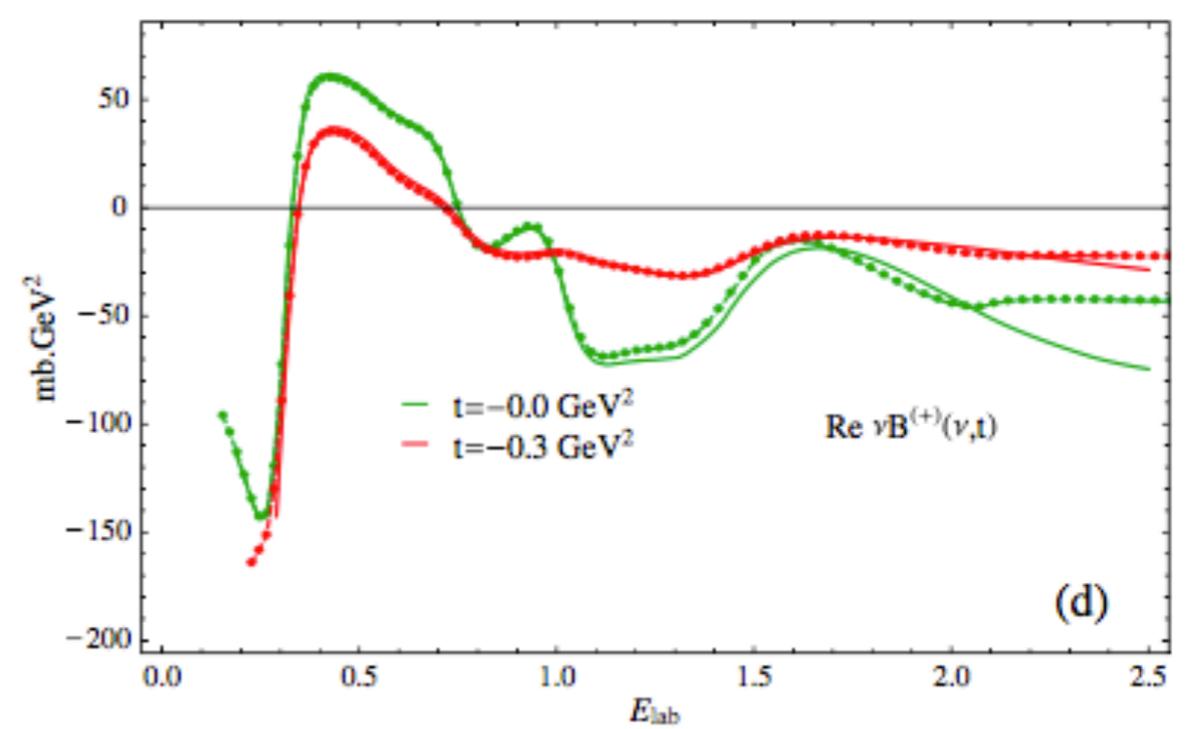
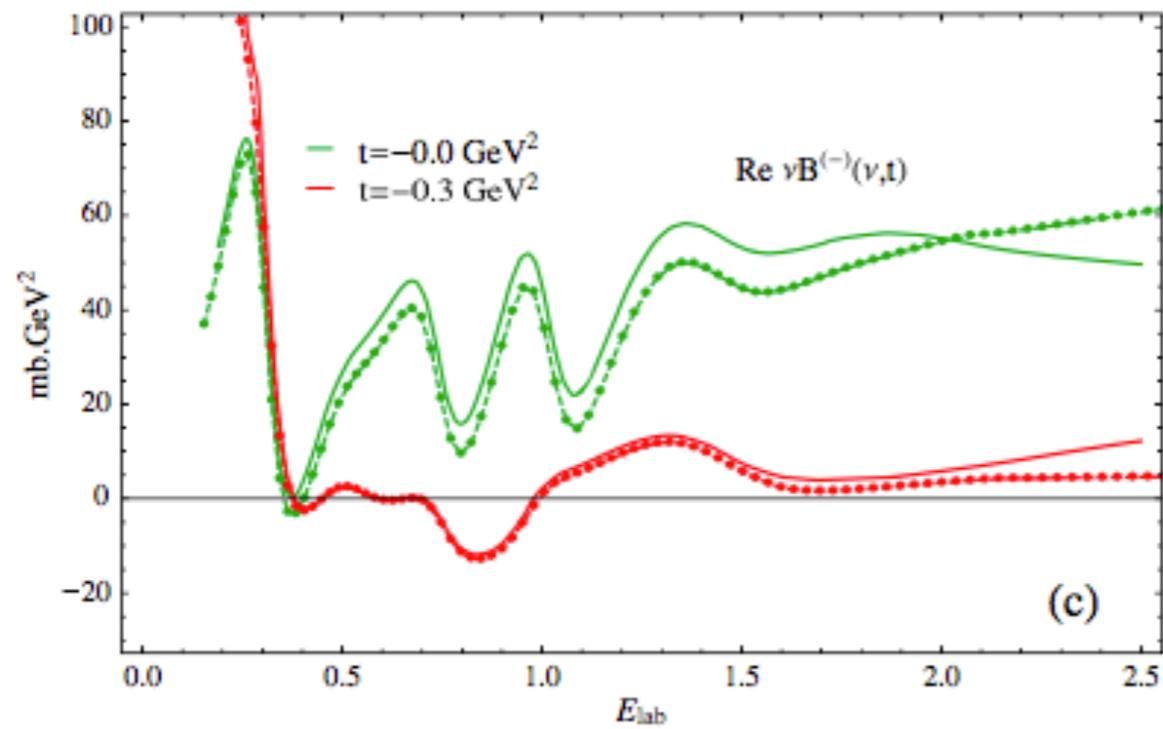
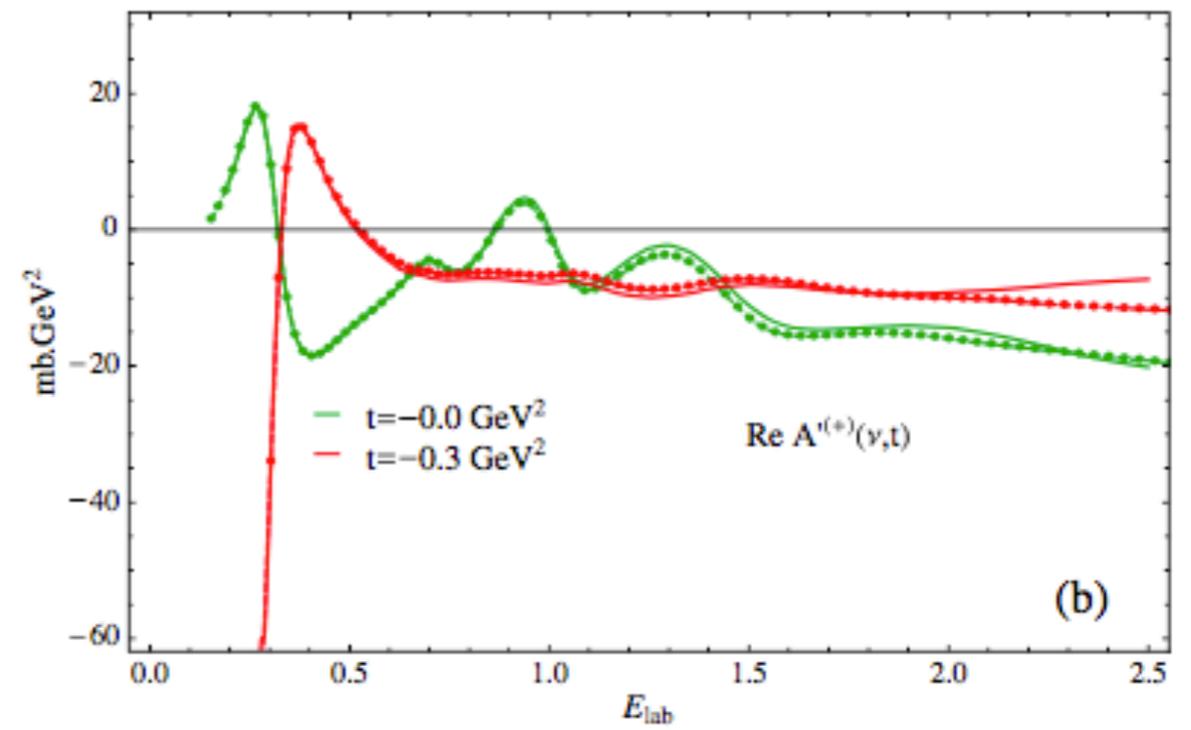
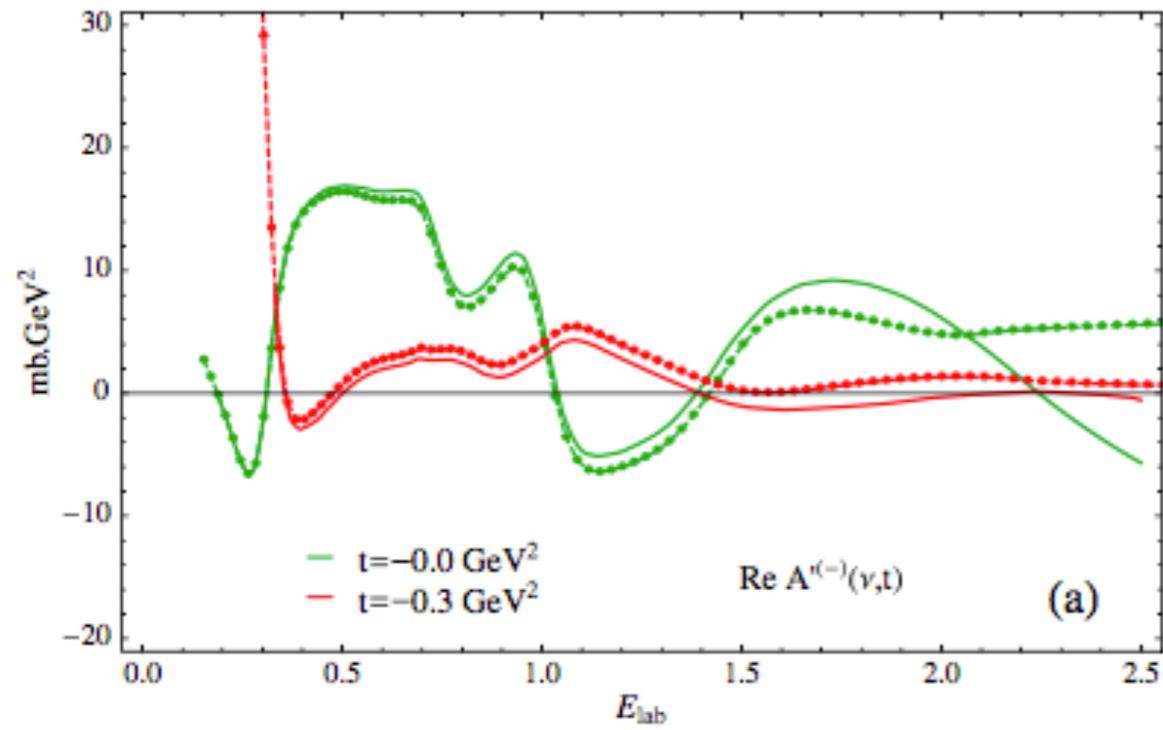
J. Nys et al (JPAC)
arXiv:1611.04658



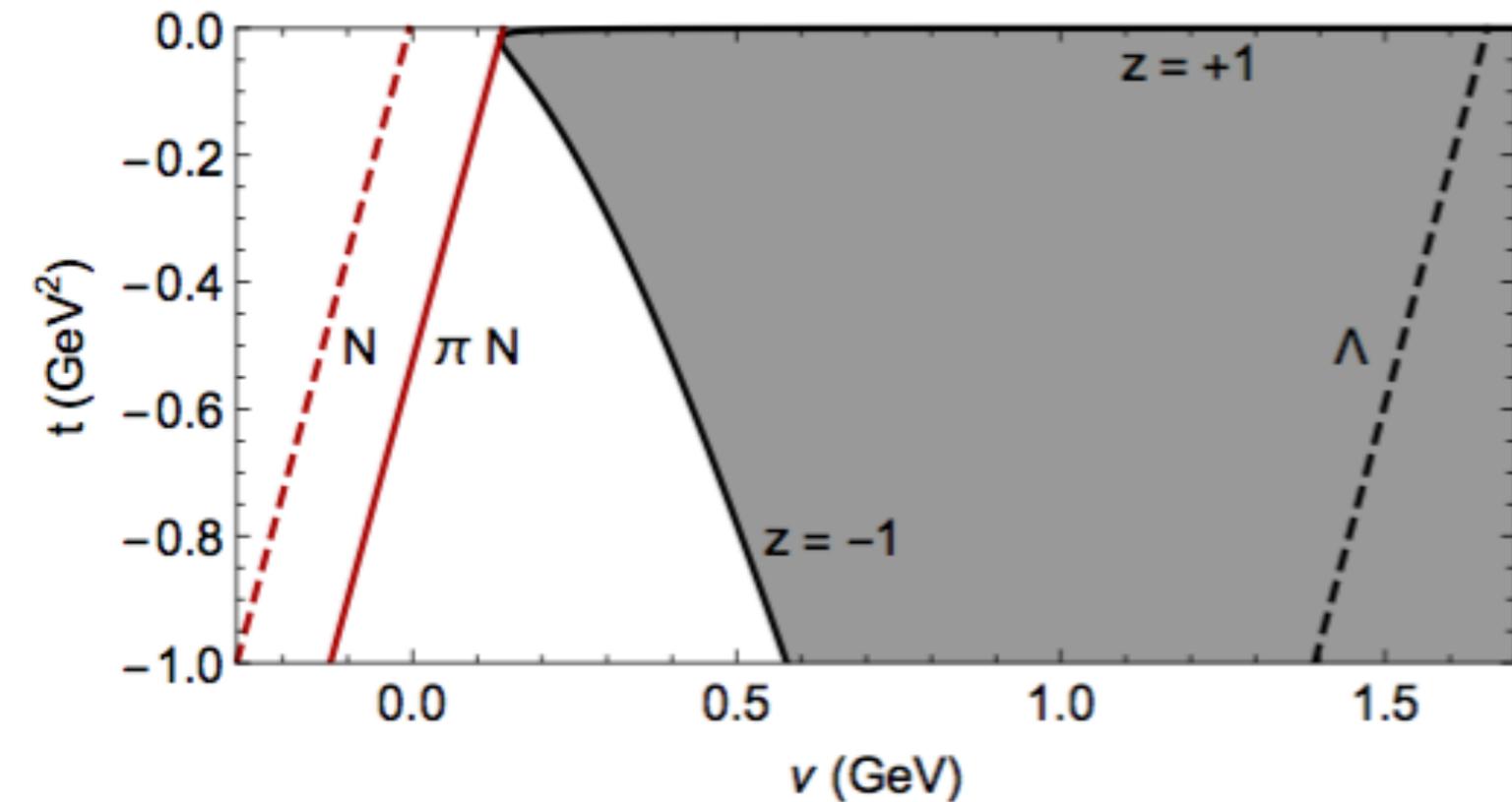
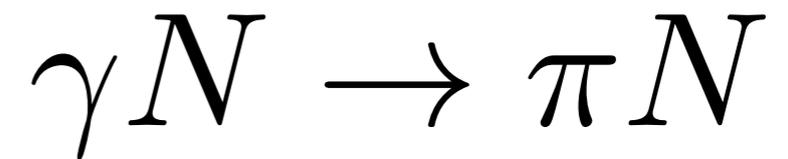
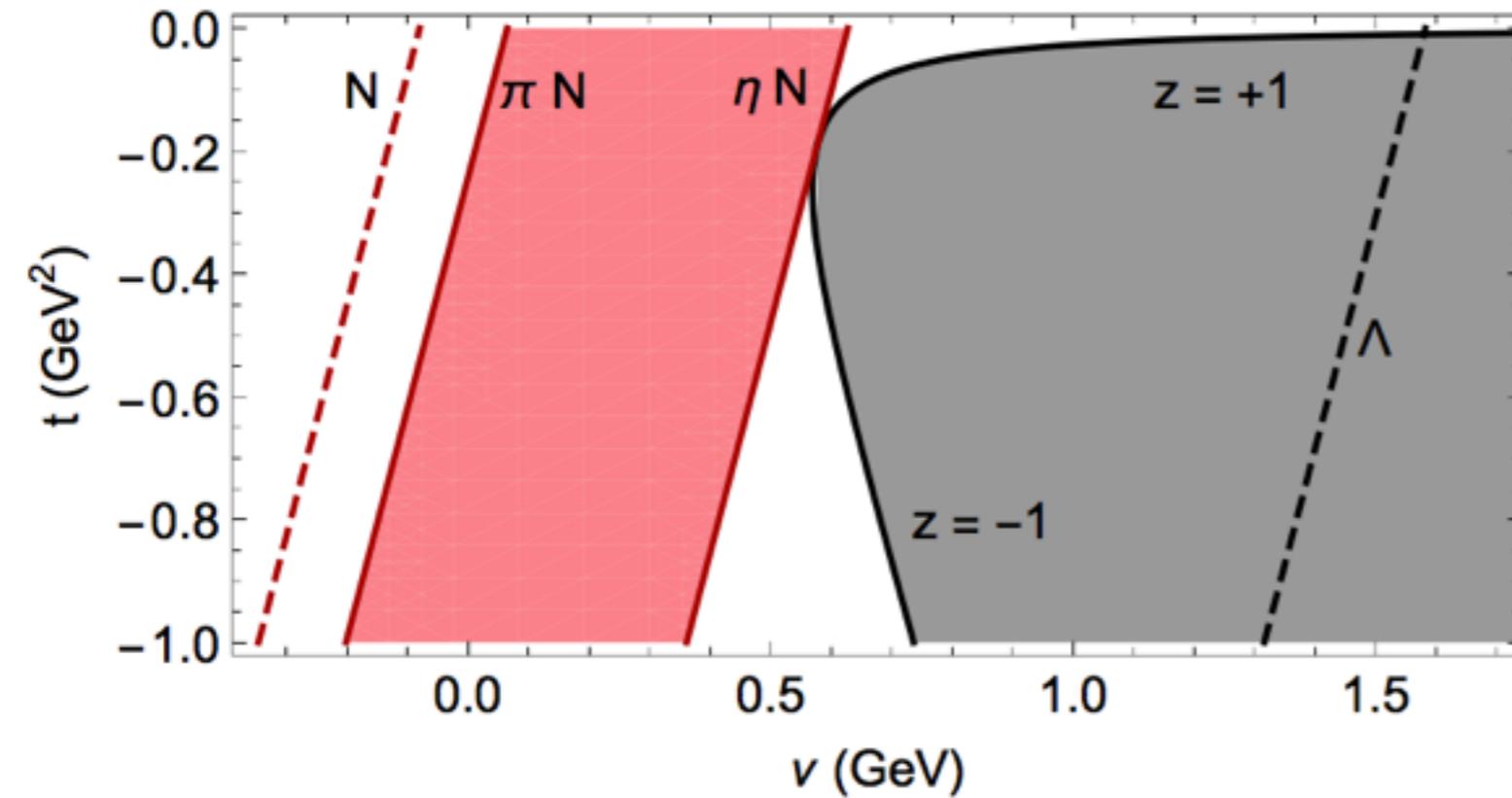
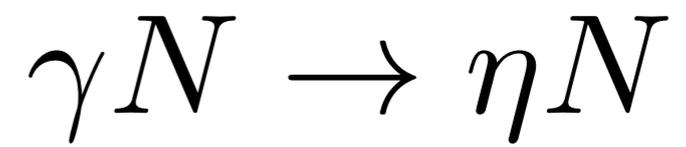
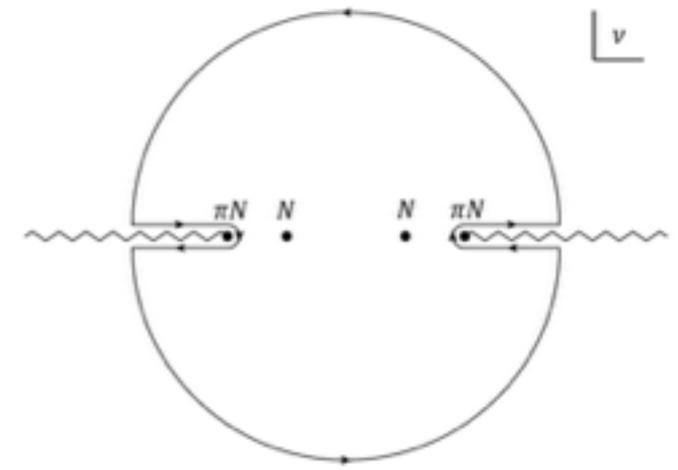
Backup Slides

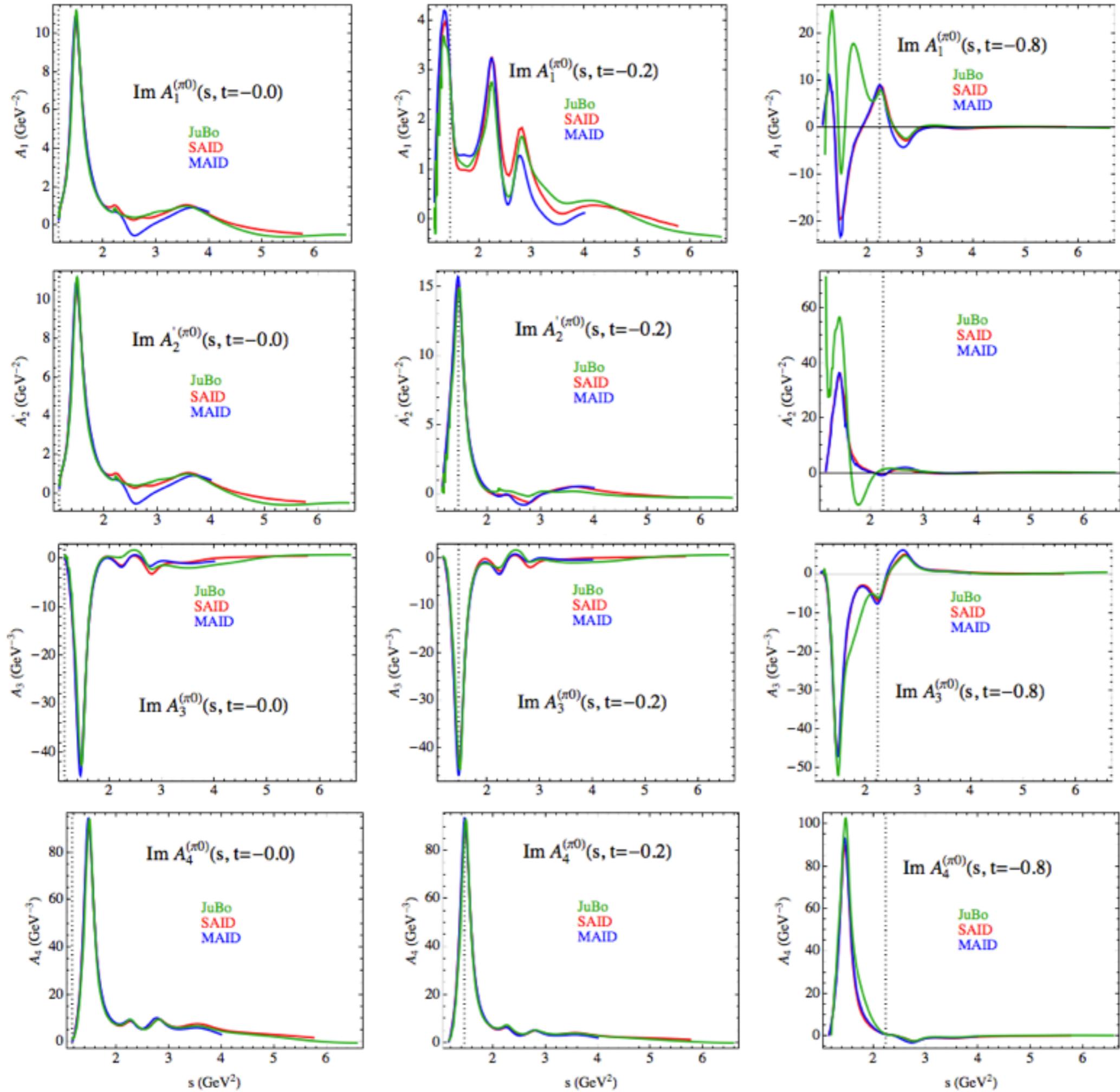


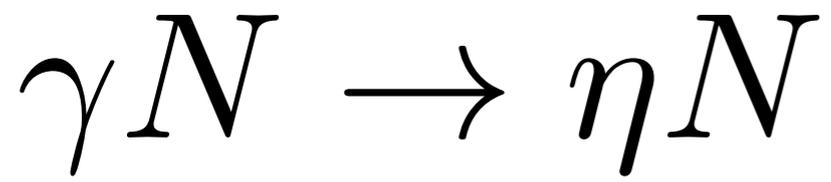
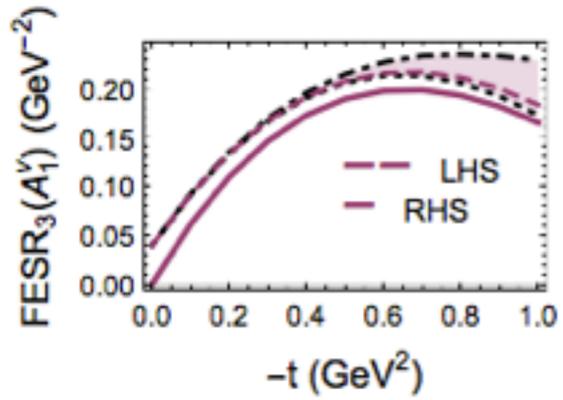
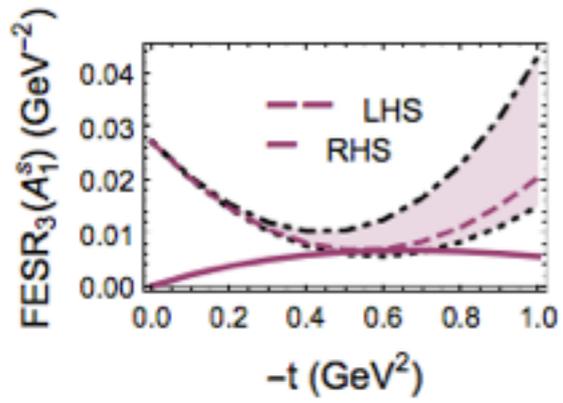
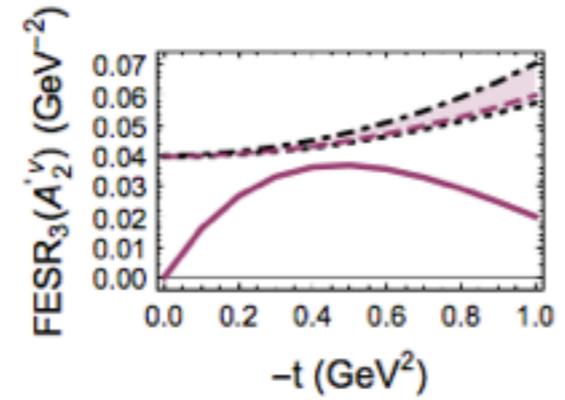
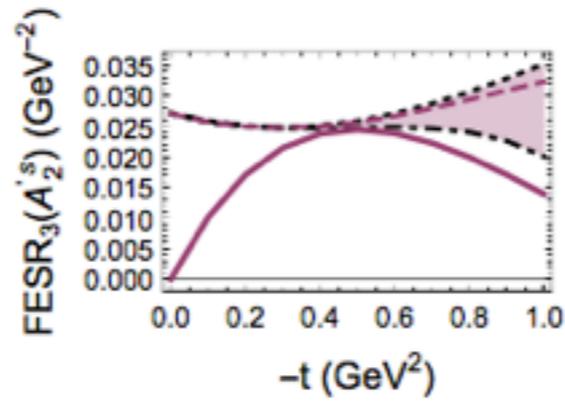
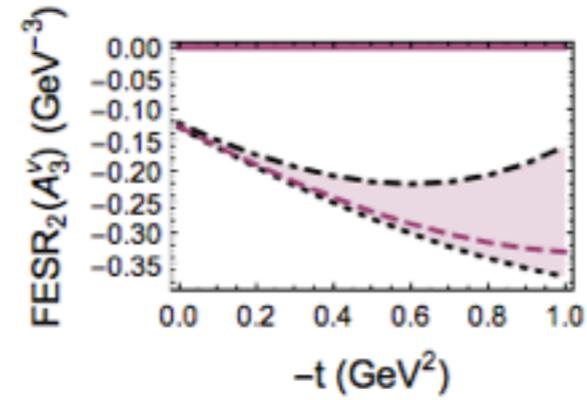
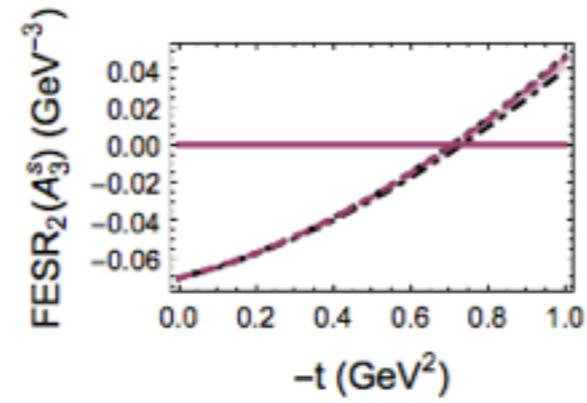
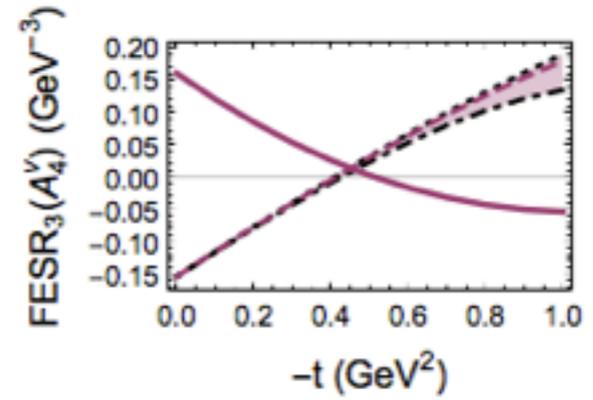
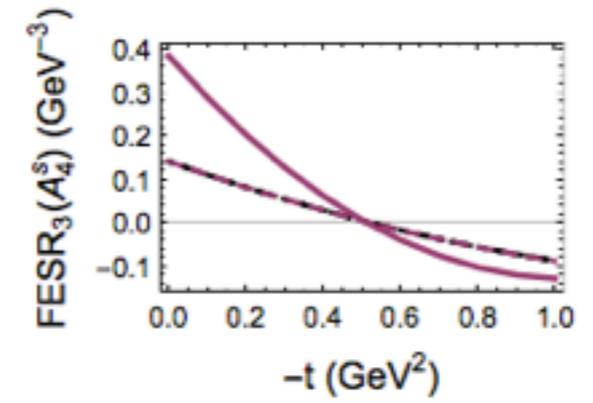
$$\pi N \rightarrow \pi N$$



Integration Region



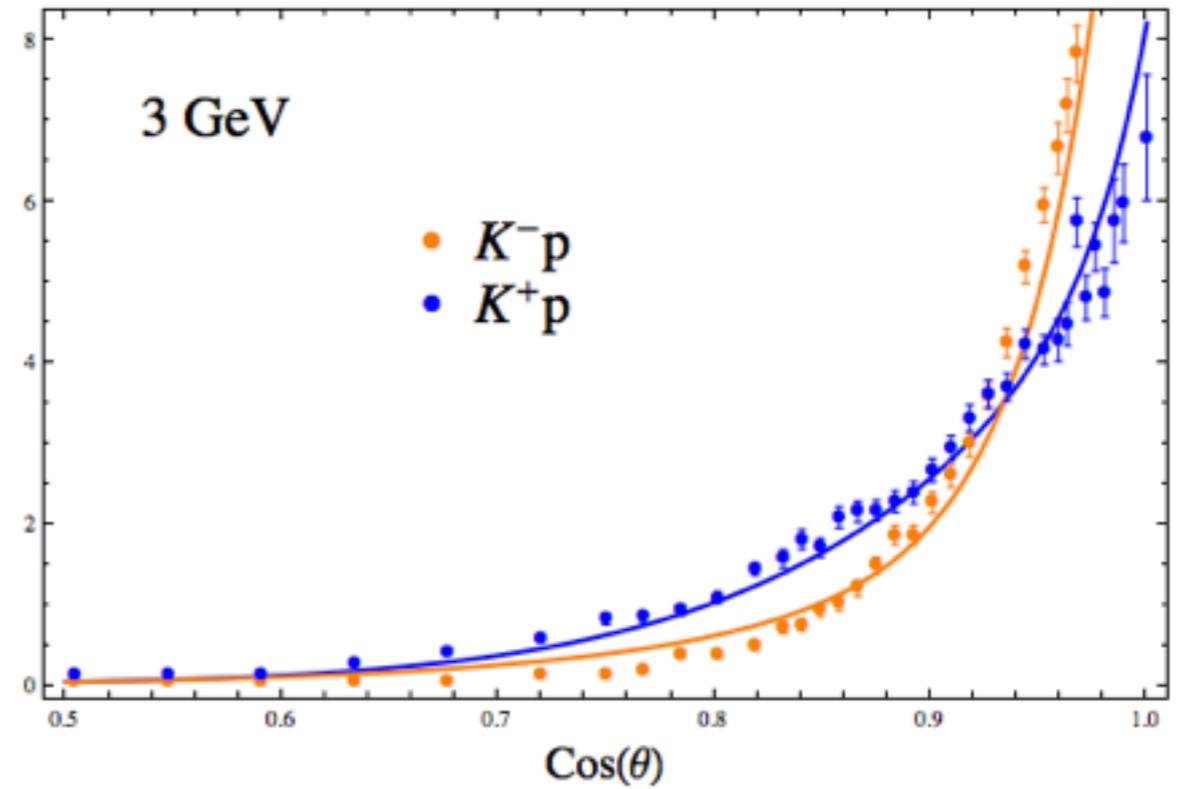
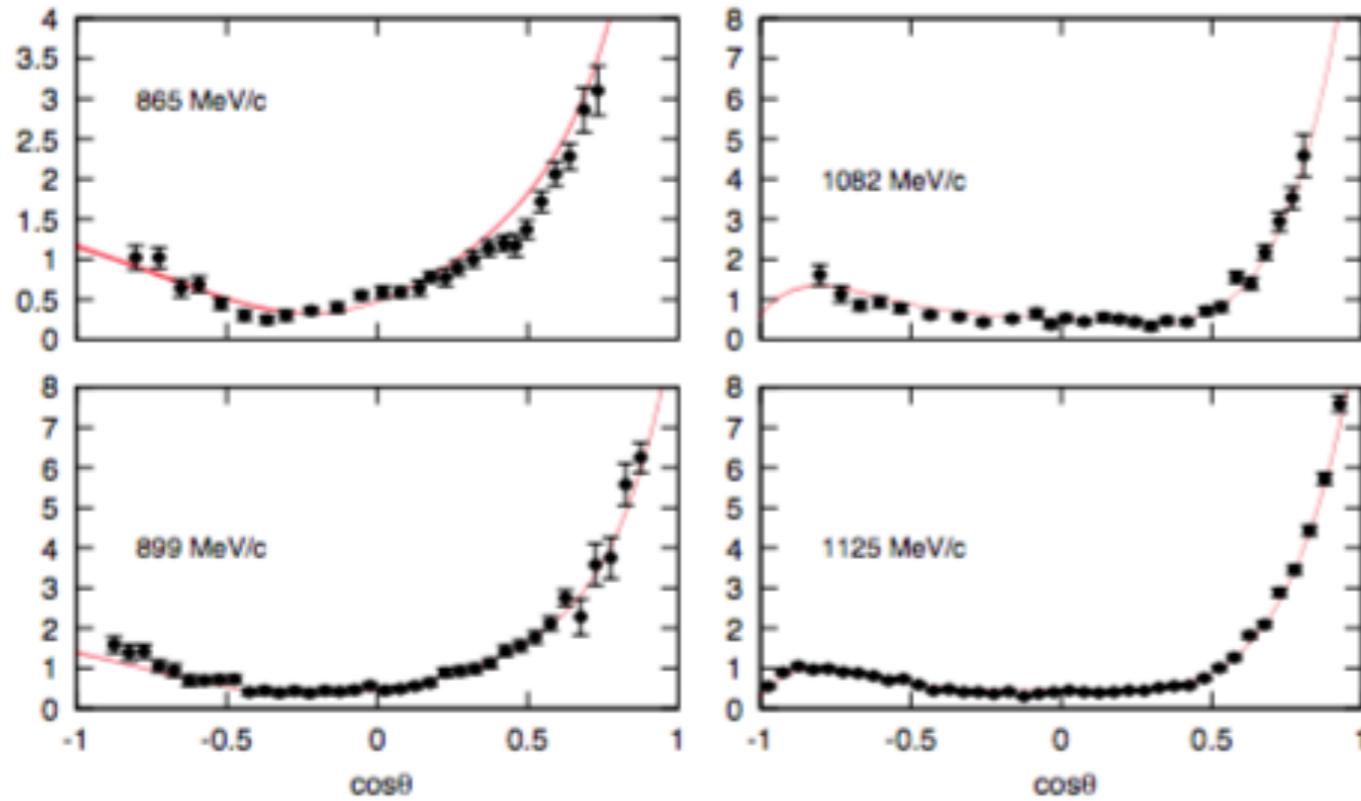



 A_1

 A_2

 A_3

 A_4


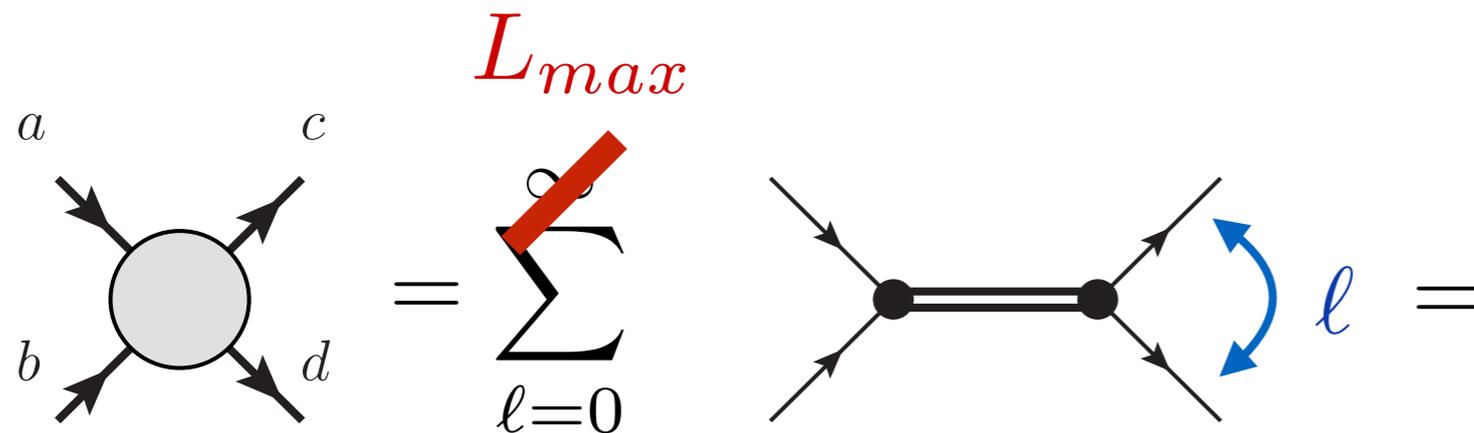
$K^- p \rightarrow K^- p$ Energy Evolution

C. Fernandez-Ramirez et al. (JPAC) ArXiv:1510:07065

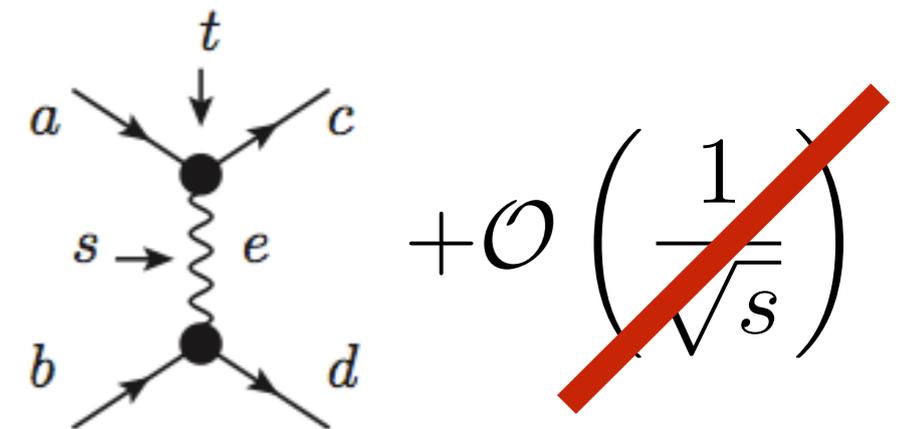
VM (unpublished)



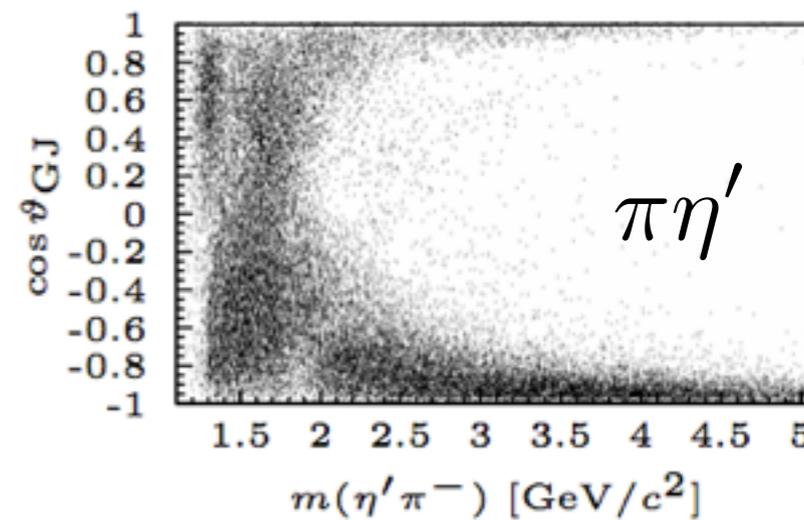
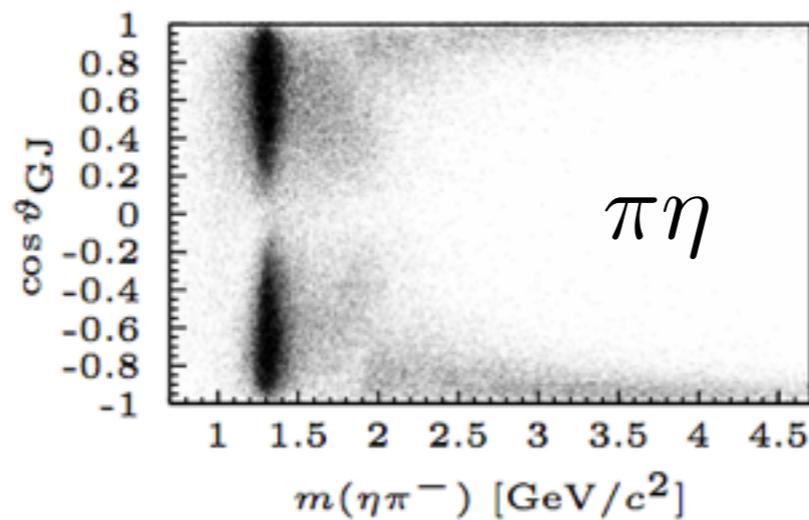
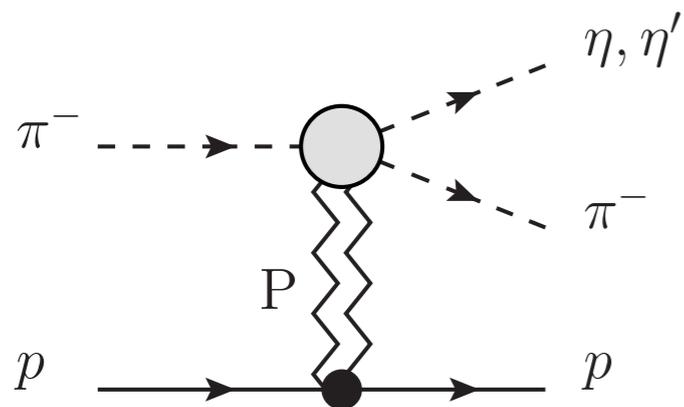
Partial wave expansion



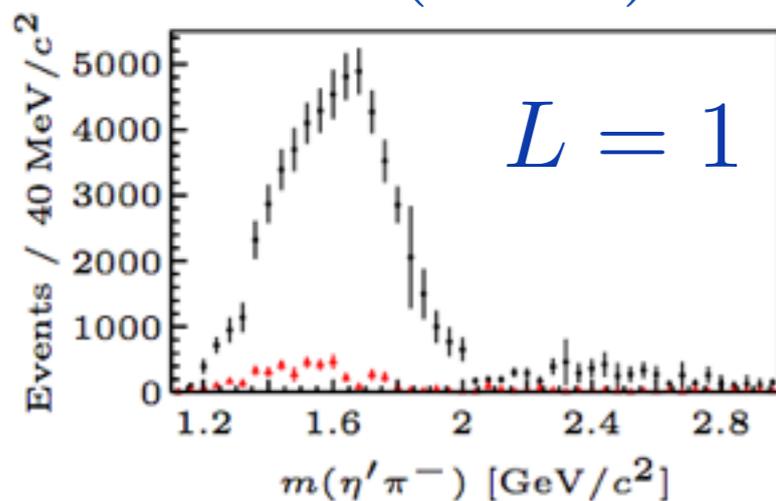
Regge pole expansion



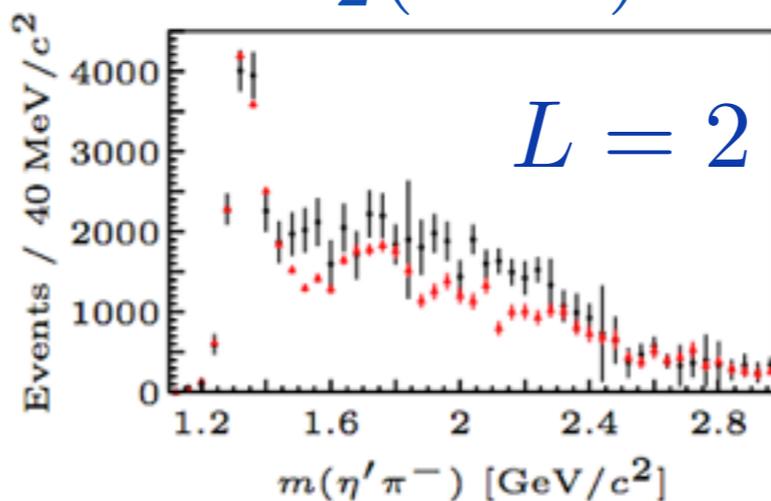
Discovering (?) New Resonances: Eta(')-Pi @COMPASS



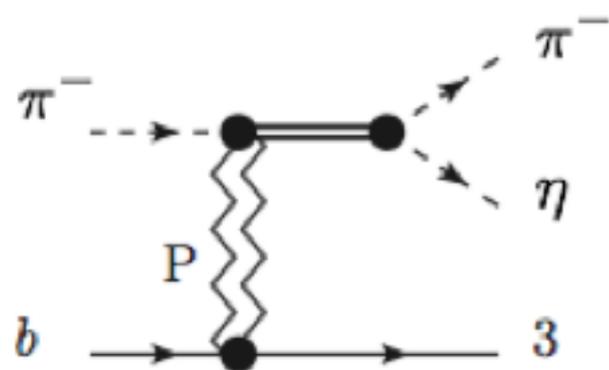
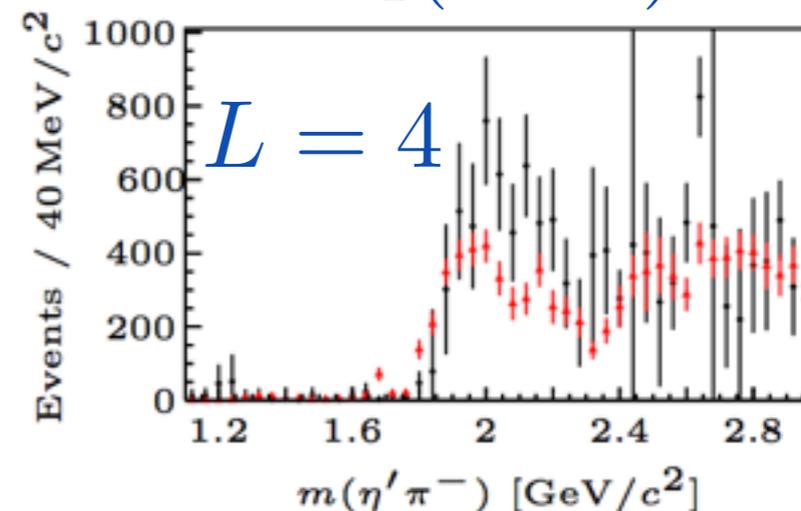
$\pi_1(1600)?$



$a_2(1320)$



$a_4(2040)$



black: $\pi\eta'$

red: $\pi\eta$ (scaled)

Resonance in angular mom. $L = 1$?



Interactive webpage:

<http://www.indiana.edu/~jpac/index.html>

$$\pi N \rightarrow \pi N$$

VM et al (JPAC)

arXiv:1506.01764

PRD92 7 074004

$$\gamma p \rightarrow \pi^0 p$$

VM et al

arXiv:1505.02321

PRD92 7 074013

$$\eta \rightarrow \pi^+ \pi^- \pi^0$$

P. Guo et al (JPAC)

arXiv:1505.01715

PRD92 5 054016

$$\begin{aligned} \omega, \phi &\rightarrow \pi^+ \pi^- \pi^0 \\ &\rightarrow \gamma^* \pi^0 \end{aligned}$$

I. Danilkin et al (JPAC)

arXiv:1409.7708

PRD91 9 094029

$$\gamma p \rightarrow K^+ K^- p$$

M. Shi et al (JPAC)

arXiv:1411.6237

PRD91 3 034007

$$KN \rightarrow KN$$

C. Fernandez-Ramirez et al (JPAC)

arXiv:1510.07065