

# Neutrino-Nucleus Interactions

J. Carlson - LANL

- Different energy and momentum regimes and different processes (beta decay, ...)
- `Realistic` model ingredients
- $A=2$  (deuteron)
- Light Nuclei ( $A \lesssim 12$ )
- Medium/Heavy Nuclei (Ar, Pb, ...)
- Relations to matter
- Outlook

see Formaggio and Zeller, RMP, 2013  
Scholberg, ARNPS, 2012

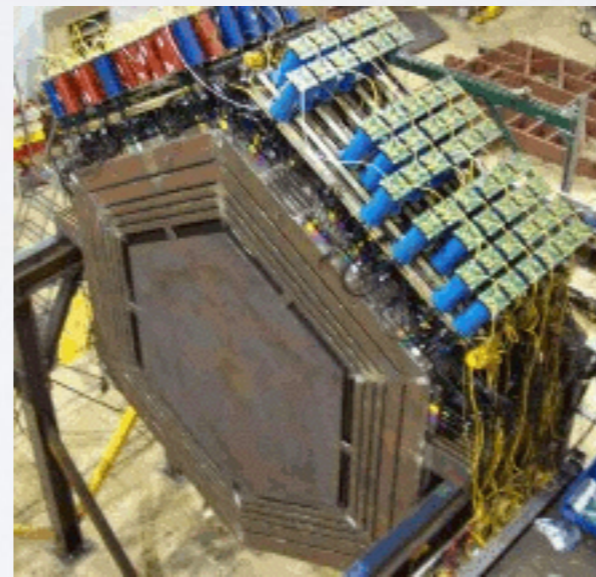
# Accelerator Neutrinos



MINOS



SuperK

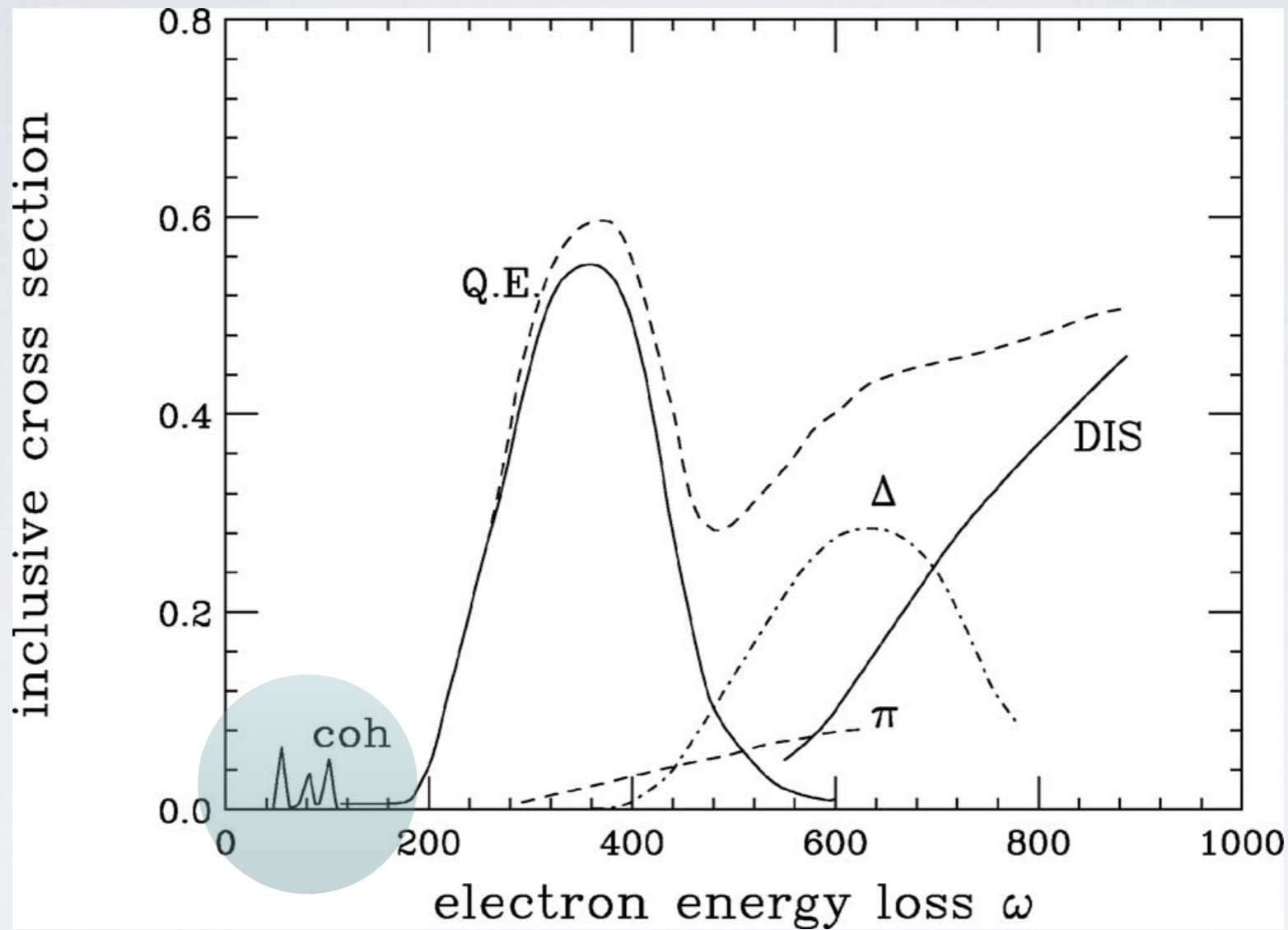


MINERva



MicroBooNE

# Accelerator Electron/Neutrino Scattering

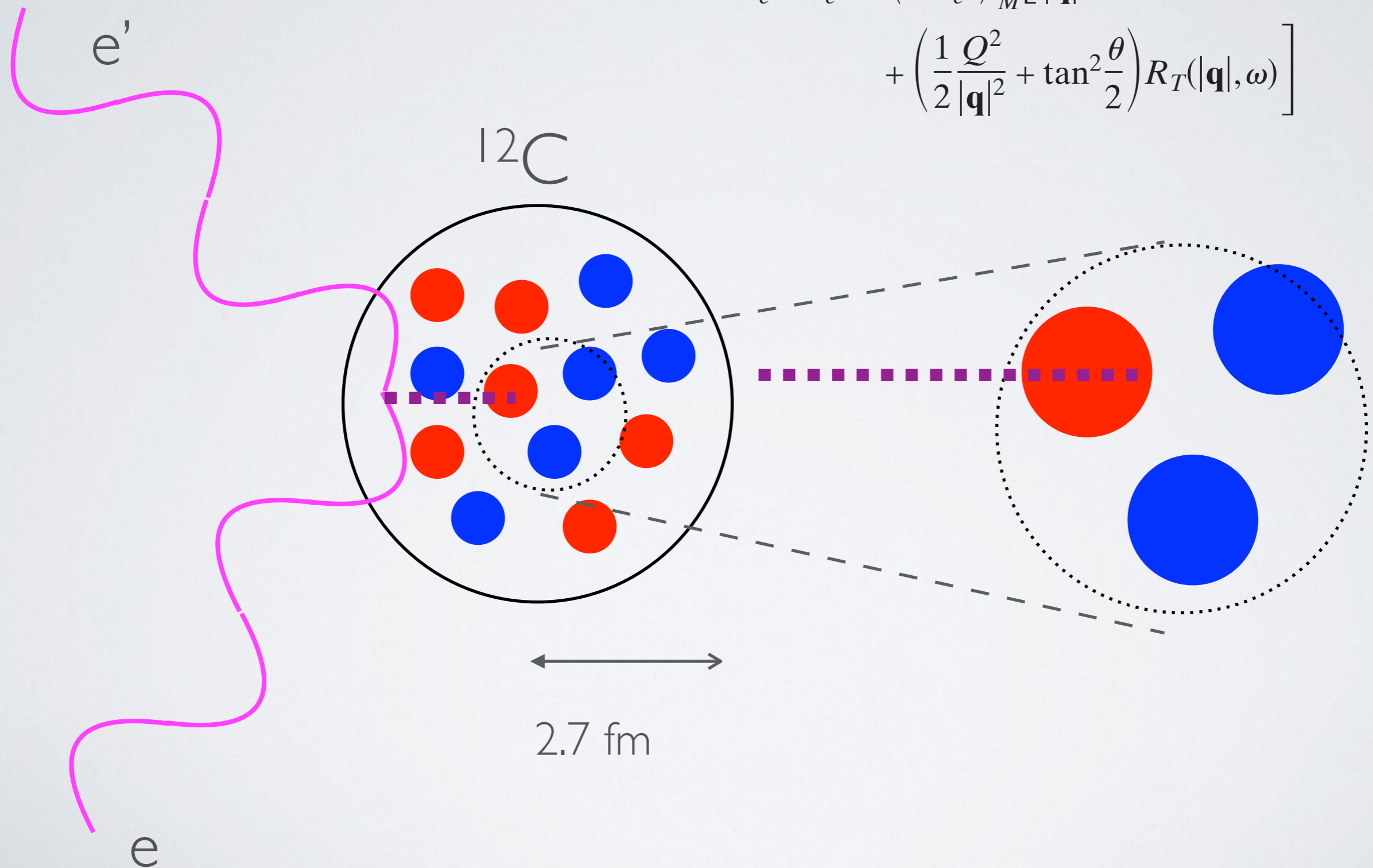


Benhar, Day, Sick, RMP 2008

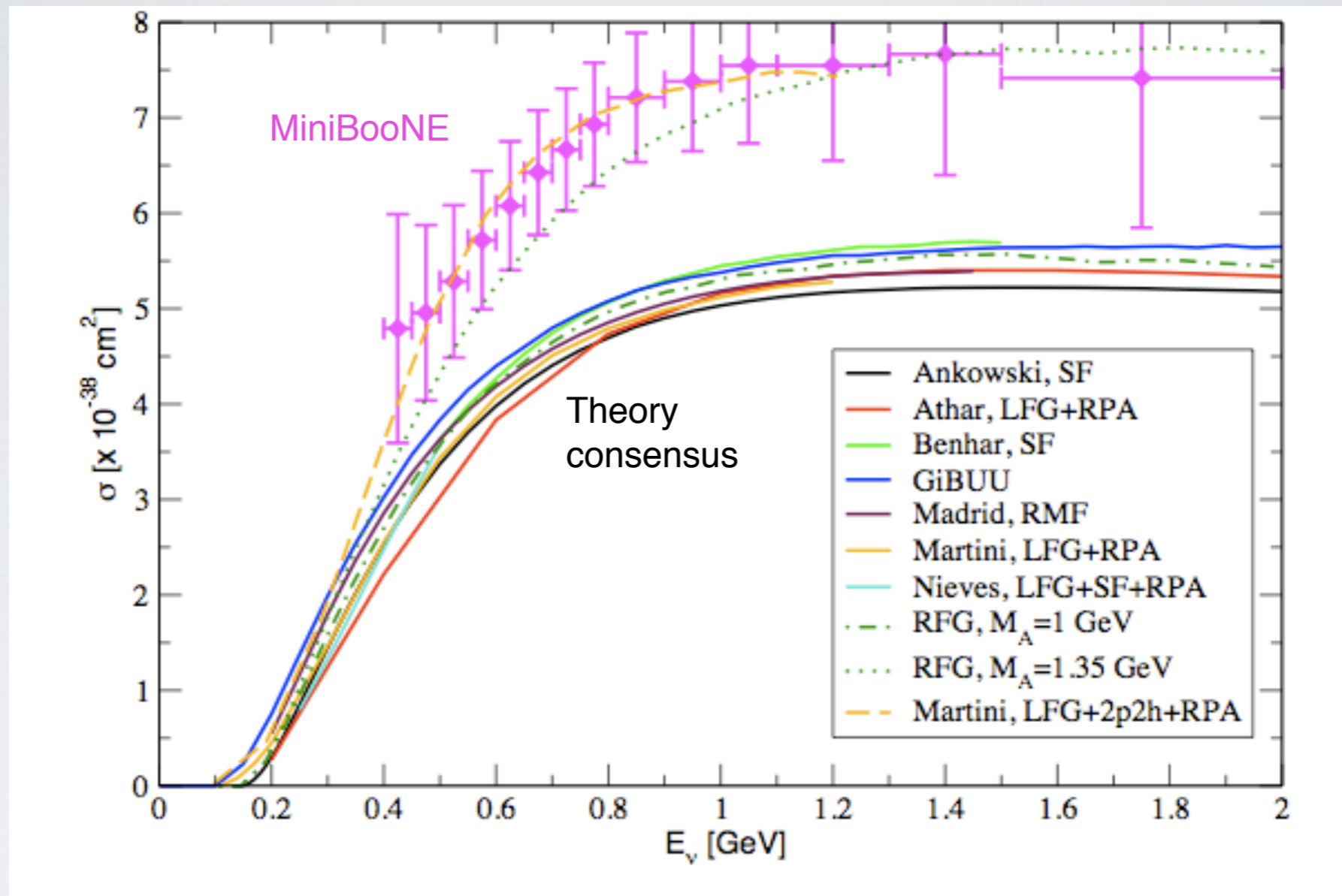
# Inclusive electron scattering at larger $q$

measure electron kinematics only

$$\frac{d^2\sigma}{d\Omega_{e'}dE_{e'}} = \left(\frac{d\sigma}{d\Omega_{e'}}\right)_M \left[ \frac{Q^4}{|\mathbf{q}|^4} R_L(|\mathbf{q}|, \omega) + \left( \frac{1}{2} \frac{Q^2}{|\mathbf{q}|^2} + \tan^2 \frac{\theta}{2} \right) R_T(|\mathbf{q}|, \omega) \right]$$



# Simple Models of QuasiElastic Neutrino Scattering



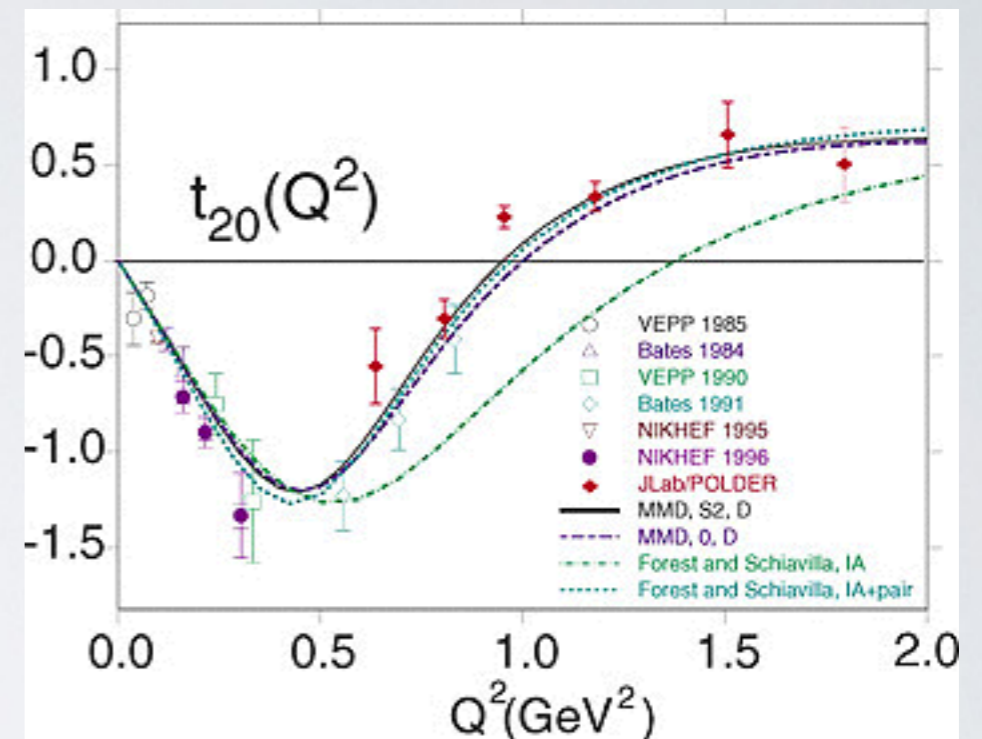
Simplest models fail at 30-40% level (too small)  
requires two-nucleon currents and correlations

# Nuclear Interactions and Currents

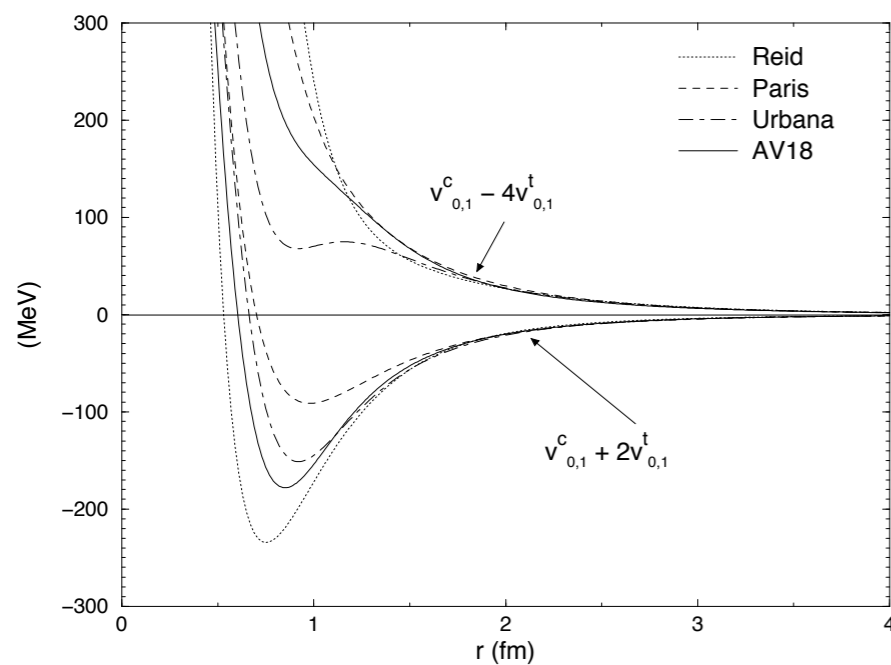
Non-relativistic nucleons w/ 2, 3-body interactions, currents

$$H = \frac{1}{2m} \sum_i p_i^2 + \sum_{i < j} V_{ij} + \sum_{i < j < k} V_{ijk}$$

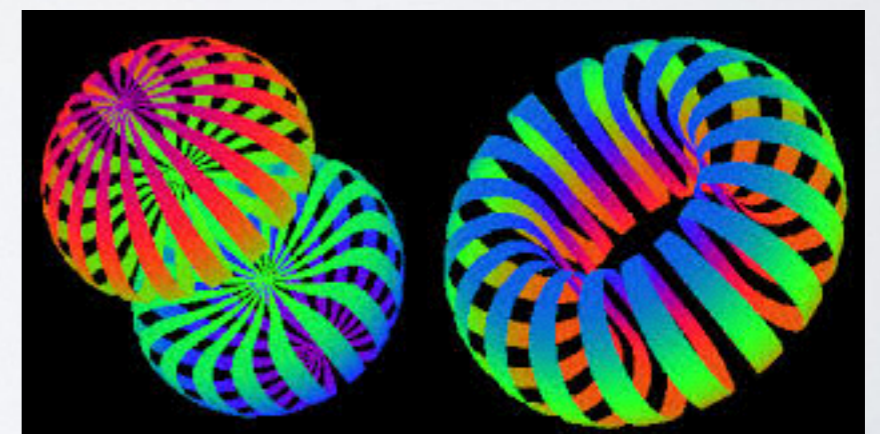
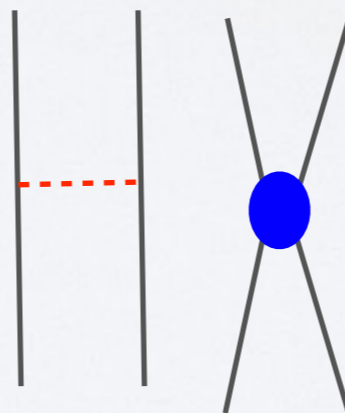
$$\mathbf{J} = \sum_i \mathbf{j}_{1;i} + \sum_{i < j} \mathbf{j}_{2;ij} + \dots$$



t20 experiment Jlab R. Holt



Deuteron Potential Models with Different Spin Orientations

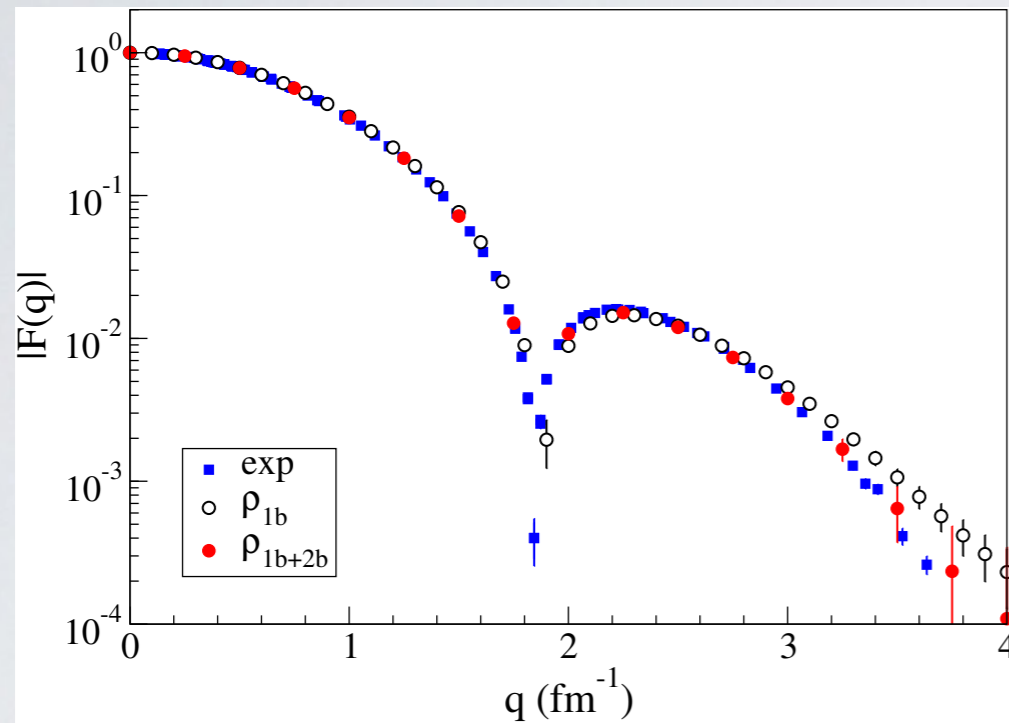


Forrest, et al, PRC 1996

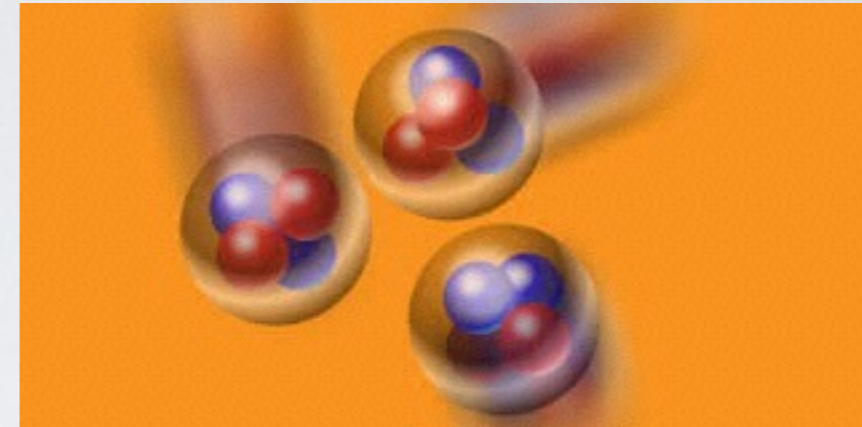
# Many measurements in EM sector

## Currents and elastic/transition form factors

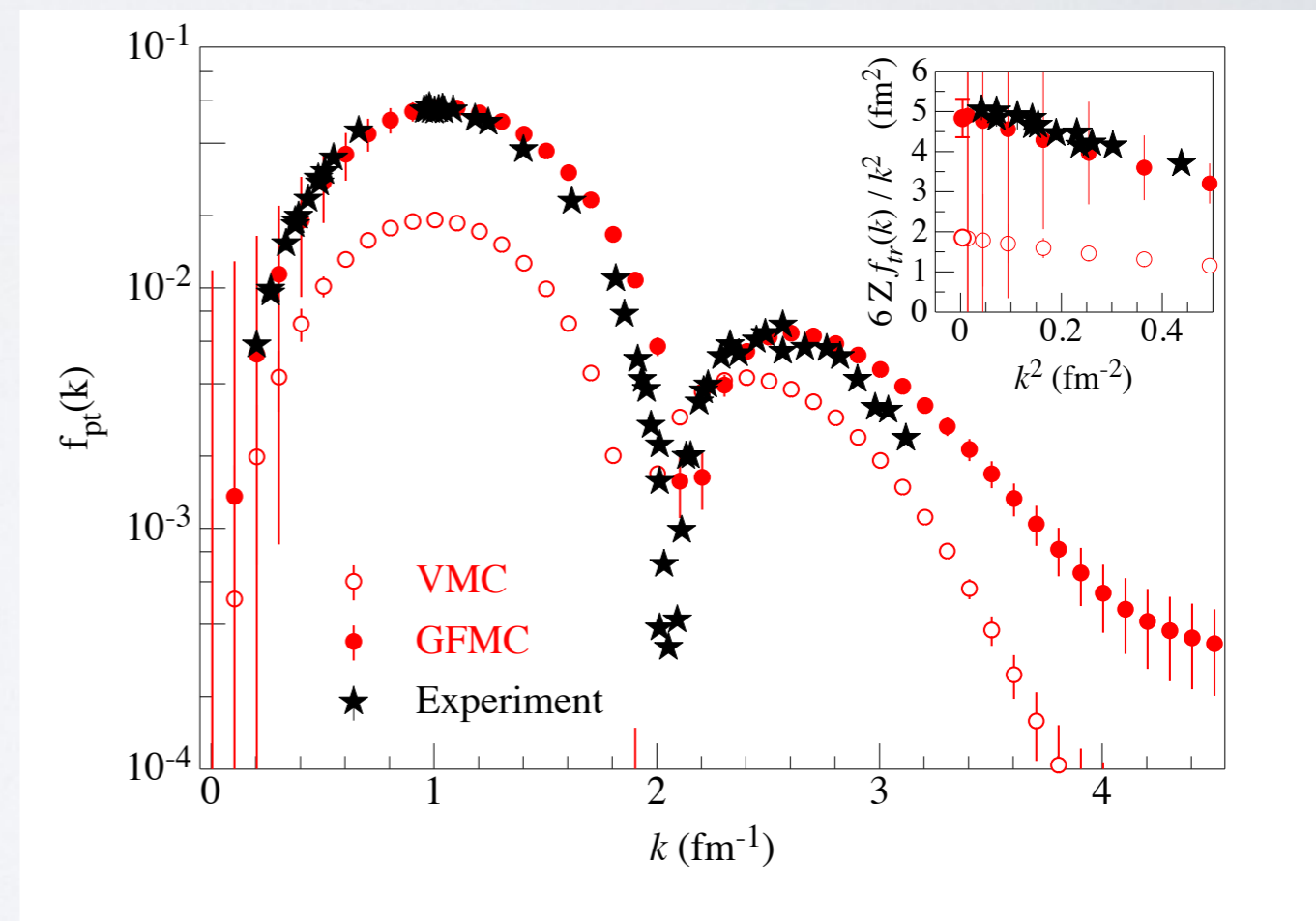
$^{12}\text{C}$  elastic form factor



2 Nucleon charge operators  
(relativistic corrections)  
are small

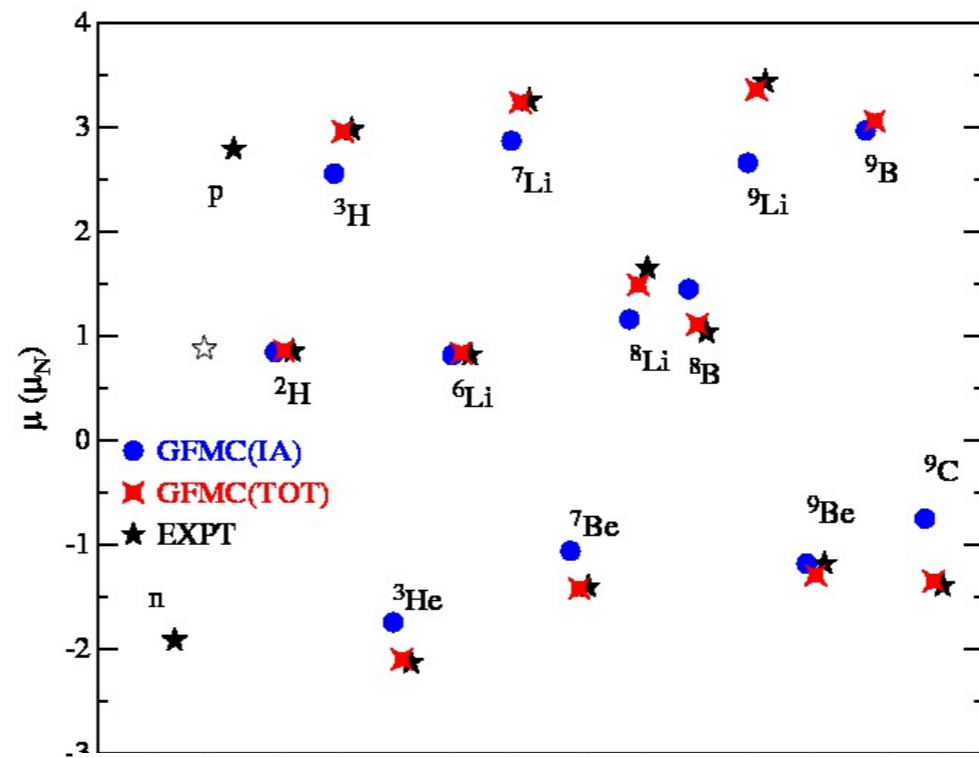
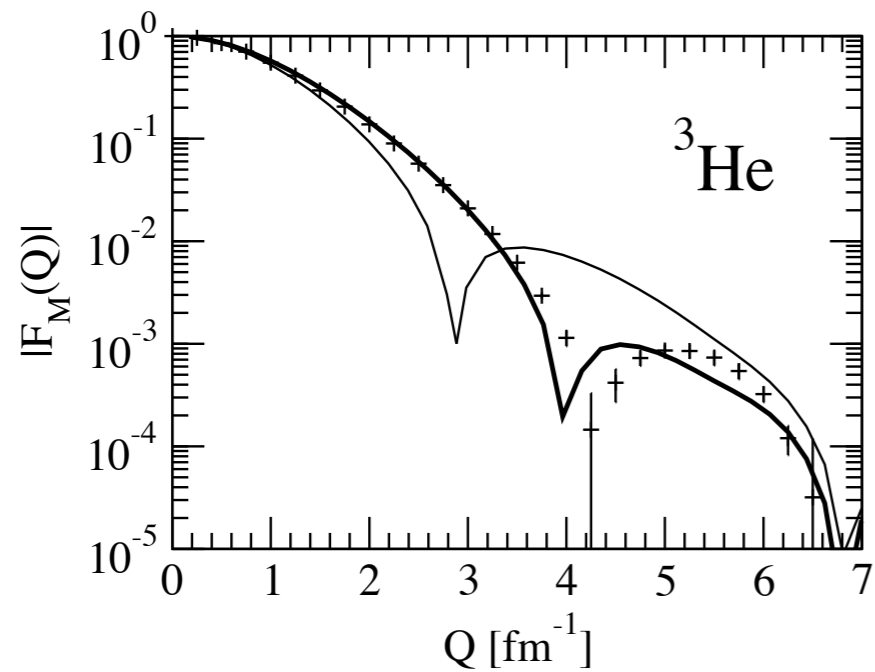
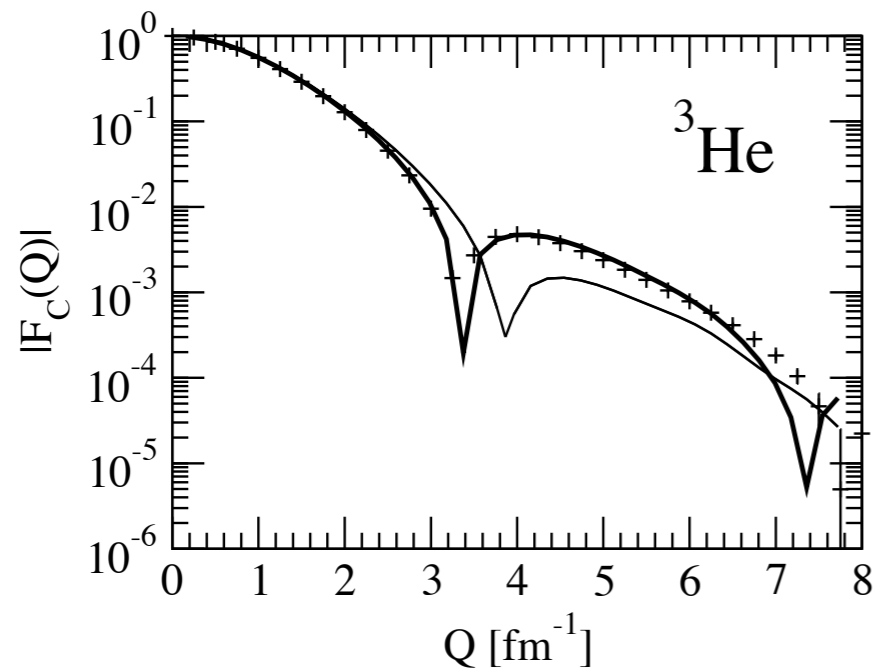


Hoyle state transition form factor

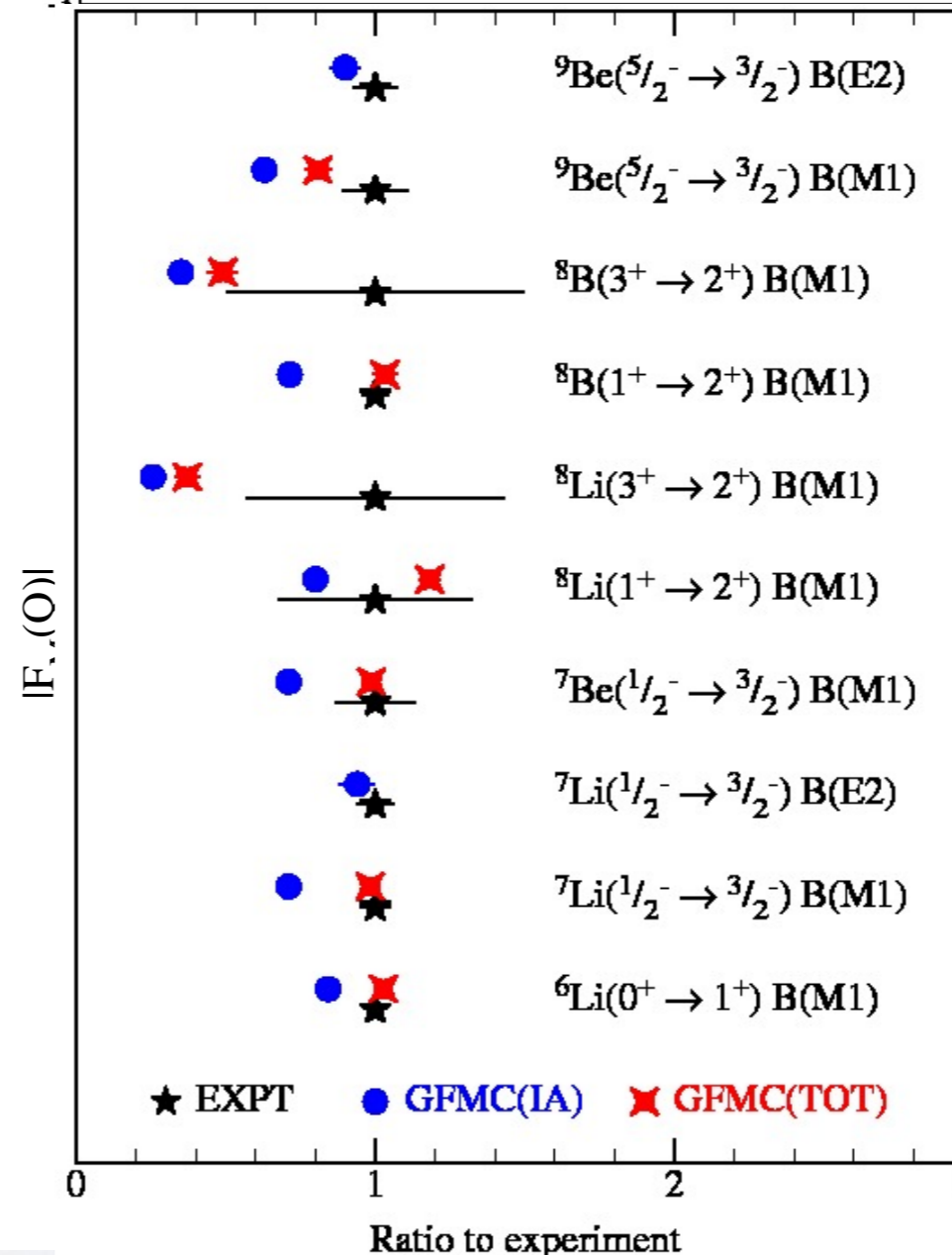


# 2-Nucleon Currents

Form Factors



Magnetic Moments



EM Transitions



**Path Integral Algorithms:**  $\Psi_0 = \exp[-H\tau] \Psi_T$

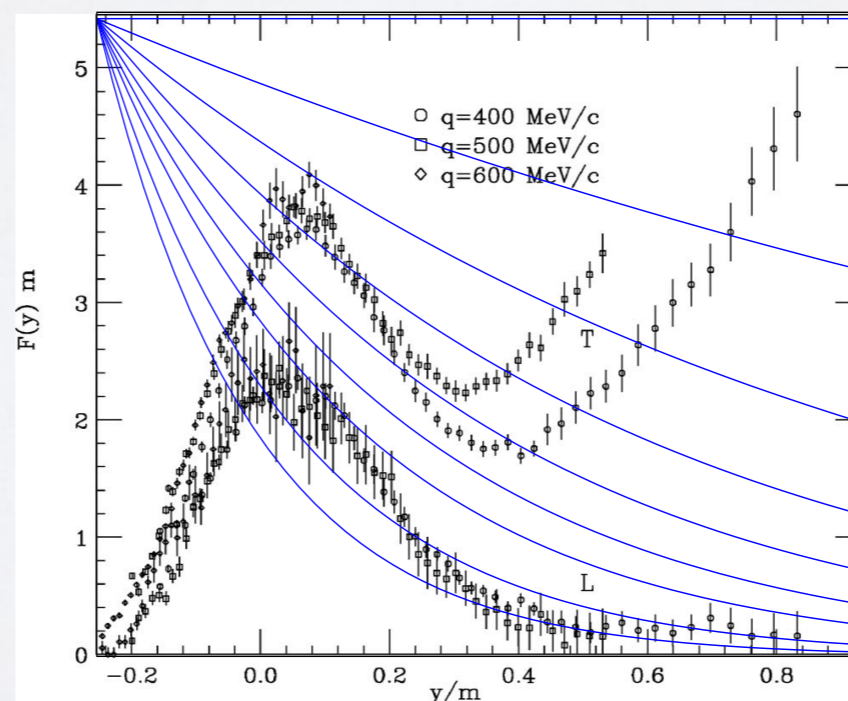
Explicit Final States  $\sigma \propto |\langle f | \mathbf{J}(\mathbf{q}) | i \rangle|^2$

Sum Rules: ground-state observable

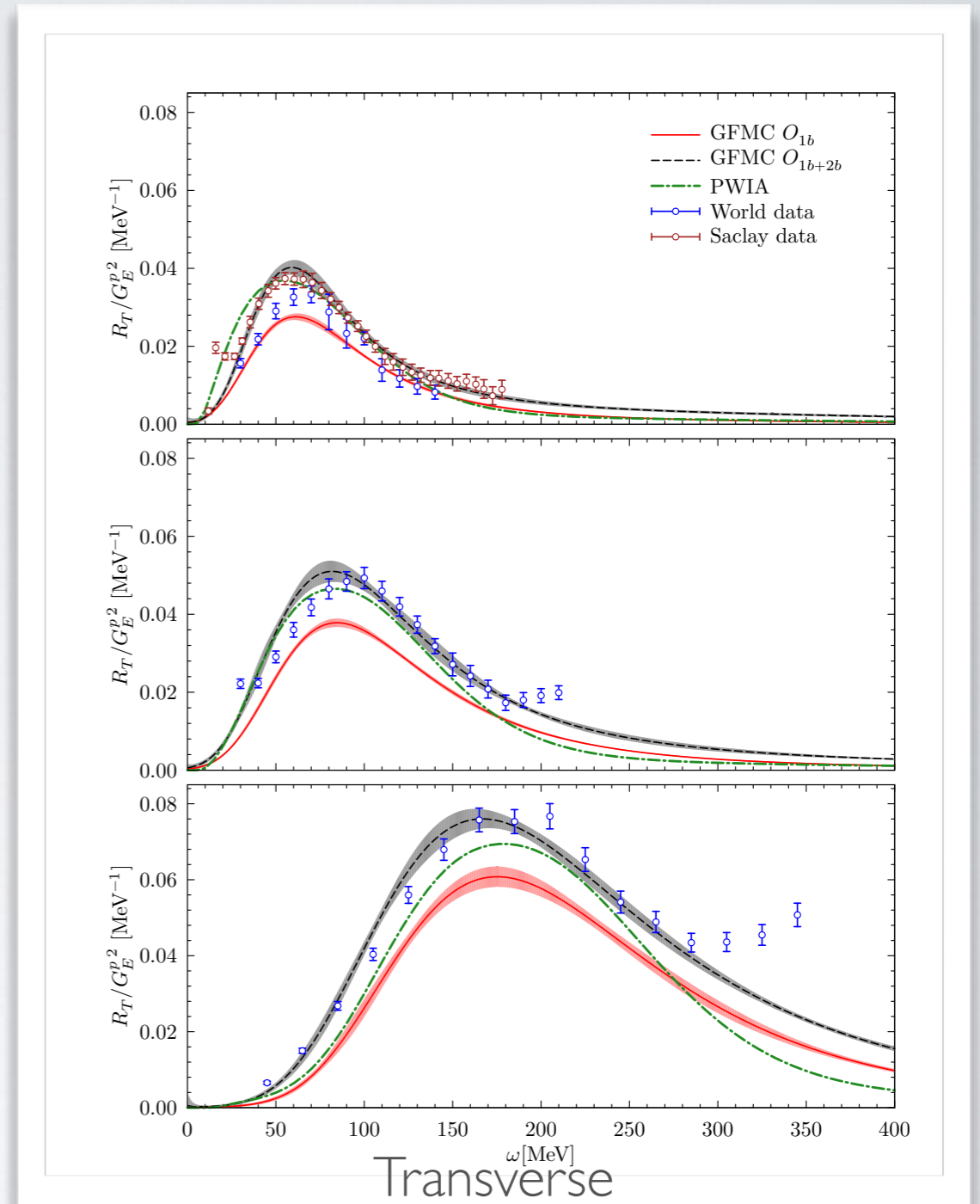
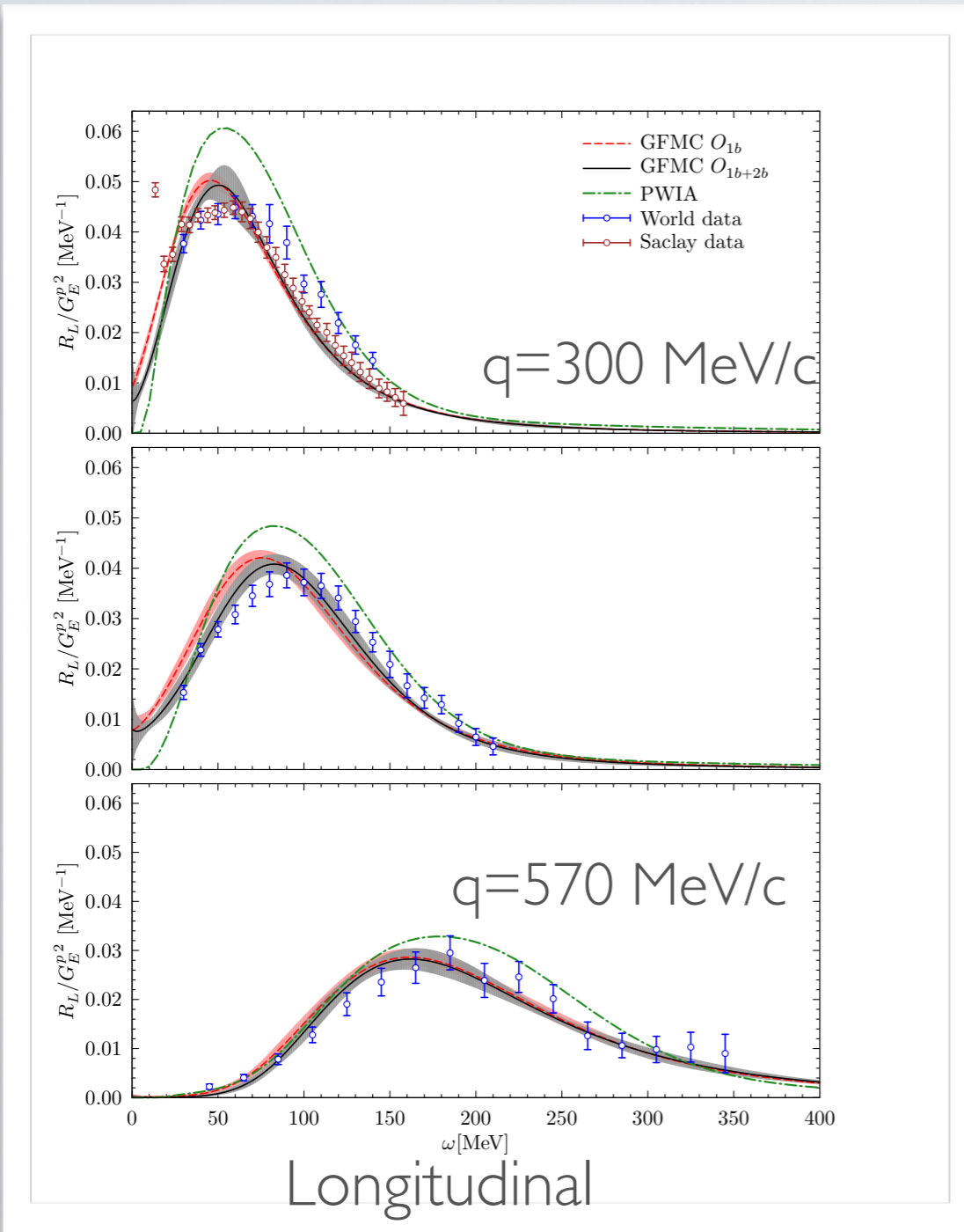
$$S(q) = \int d\omega R(q, \omega) = \langle 0 | O^\dagger(q) O(q) | 0 \rangle$$

Imaginary Time Correlations (Euclidean Response)

$$\tilde{R}(q, \tau) = \langle 0 | \mathbf{j}^\dagger \exp[-(\mathbf{H} - \mathbf{E}_0 - \mathbf{q}^2 / (2\mathbf{m}))\tau] \mathbf{j} | 0 \rangle >$$



# Quasi-elastic electron scattering on $^{12}\text{C}$

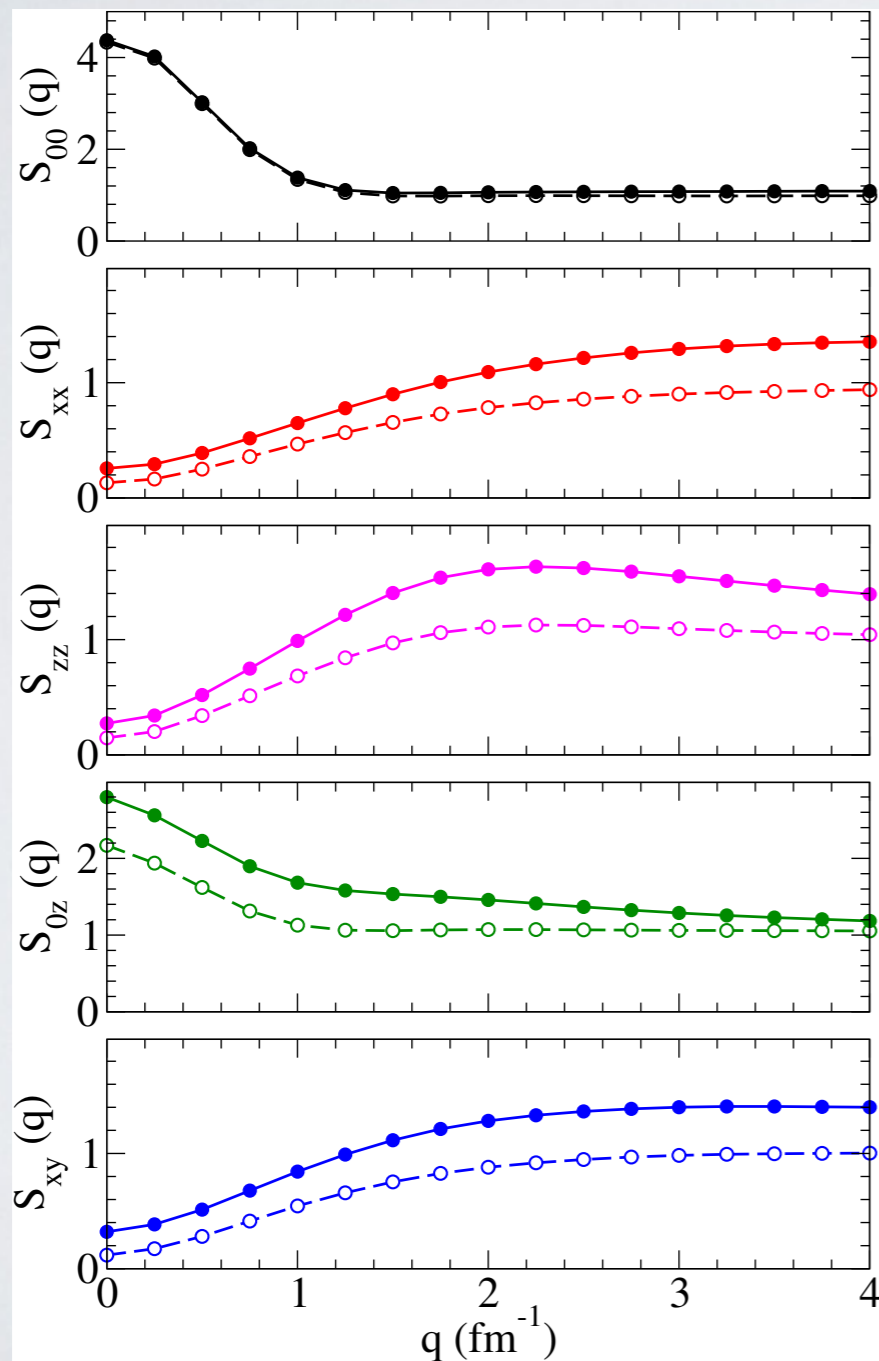


Lovato, et al, PRL 2016

Enhancement in Transverse channel  
Explicit  $0^+$ ,  $2^+$ ,  $4^+$  states important at  $q=300$  MeV/c

# Neutral Current Sum Rules in $^{12}\text{C}$

## Sum Rules



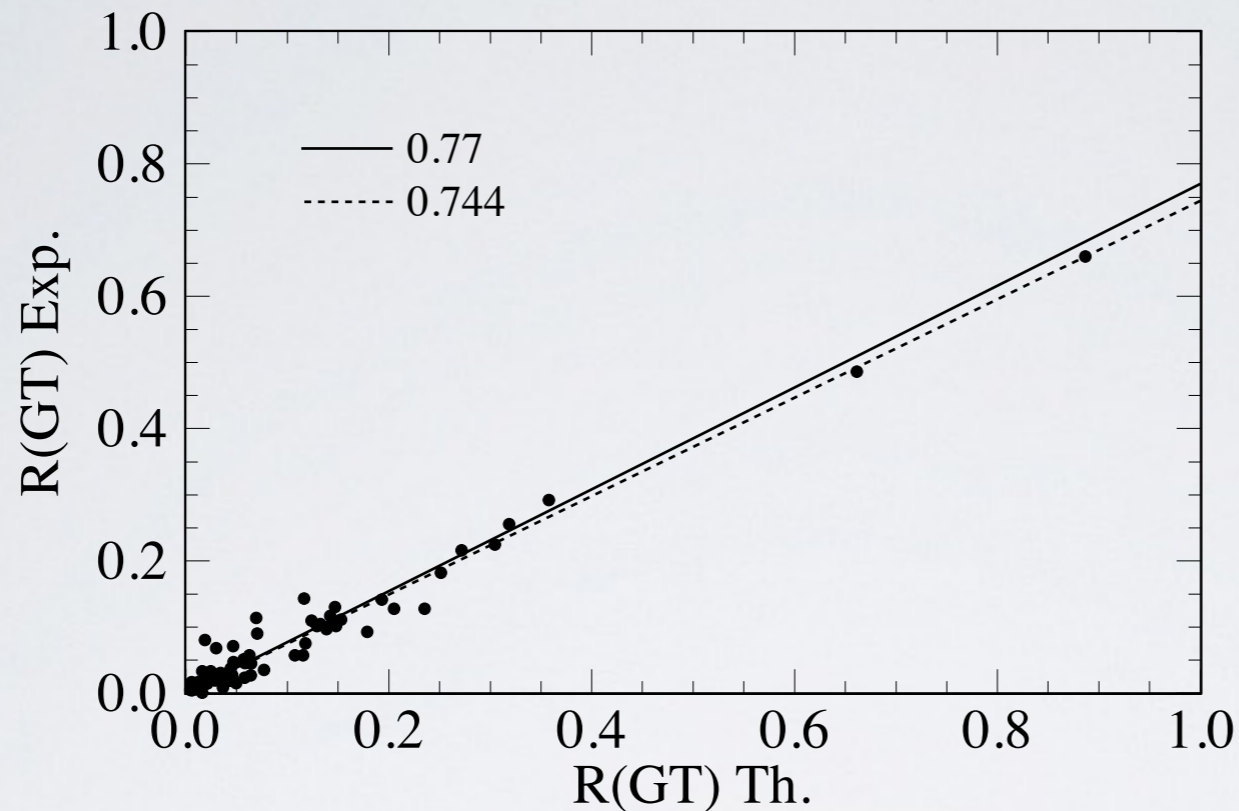
cross-section depends upon  
5 response fns

all except longitudinal response enhanced  
including axial-vector interference

sum rules - ground state expectation value

enhanced even at very low  $q$ , but strength  
inaccessible to low energy neutrinos

# Low Momentum Transfer: GT Beta Decay



Shell Model Calculations of Beta Decay typically require a quenching (reduction) of  $g_A$  by  $\sim 0.75$

**Rate reduction by 30-40%**

Martinez-Pinedo and Poves, PRC 1996

	Simple	1-Body current	1+2 current	Exp <sup>⊙</sup>
A=3	2.45	2.27	2.28*	2.28
A=6	2.4	2.15	2.19	2.2
A=7	2.58	2.29	2.39	2.4
A=10	2.45	2.06		2.34

Smaller ( $\sim 10\%$ ) quenching reproduced in light nuclei

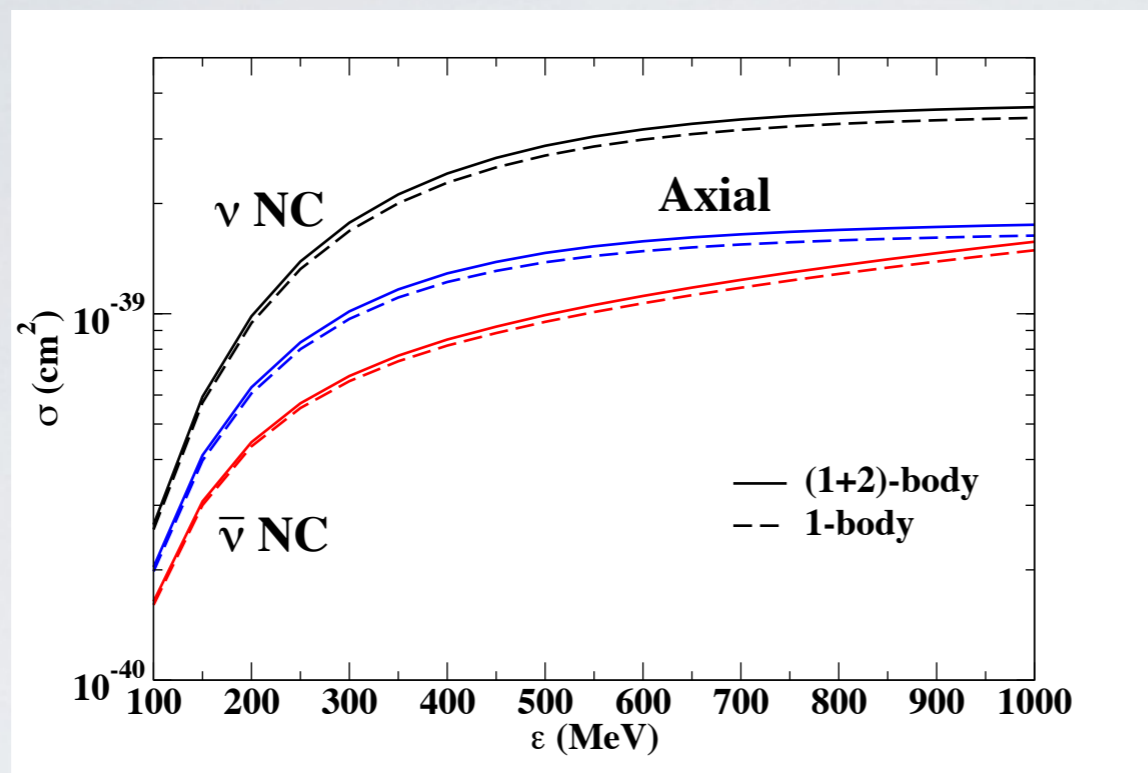
(preliminary)

# ***Astrophysical Energy Neutrinos***

- Energies up to 50 - 100 MeV
- Explicit final states and inclusive scattering measurable
- Nucleon couplings pretty well known
- What are the roles of nuclear structure, two nucleon correlations and currents ?
- Momentum transfer much less than QE, but greater than beta decay.

# Neutrino-Deuteron Scattering

A=2



Nakamura, et al, 2001  
Shen, et al., PRC 2012

currents fit to tritium beta decay  
small effects from 2-nucleon currents  
small (few %) model dependence

$\epsilon$ (MeV)	$\nu_l$ -NC		$\bar{\nu}_l$ -NC		$\nu_e$ -CC		$\bar{\nu}_e$ -CC	
	set I	set II	set I	set II	set I	set II	set I	set II
5	9.561(-44)	9.541(-44)	9.363(-44)	9.344(-44)	3.427(-43)	3.421(-43)	2.831(-44)	2.826(-44)
50	5.892(-41)	5.873(-41)	4.546(-41)	4.530(-41)	1.348(-40)	1.353(-40)	7.403(-41)	7.380(-41)
100	2.657(-40)	2.652(-40)	1.640(-40)	1.636(-40)	6.631(-40)	6.621(-40)	2.606(-40)	2.600(-40)

$\epsilon$ (MeV)	$\nu_l$ -NC				$\bar{\nu}_l$ -NC			
	AV18(1)	CDB(1)	AV18(1+2)	CDB(1+2)	AV18(1)	CDB(1)	AV18(1+2)	CDB(1+2)
50	5.747(-41)	5.791(-40)	5.892(-41)	5.847(-40)	4.449(-41)	4.484(-40)	4.546(-41)	4.519(-40)
100	2.577(-40)	2.597(-40)	2.657(-40)	2.638(-40)	1.604(-40)	1.617(-40)	1.640(-40)	1.633(-40)
500	2.703(-39)	2.715(-39)	2.874(-39)	2.858(-39)	9.503(-40)	9.553(-40)	9.916(-40)	9.895(-40)
1000	3.425(-39)	3.442(-39)	3.663(-39)	3.659(-39)	1.490(-39)	1.496(-39)	1.572(-39)	1.572(-39)

typically ~20%  
accuracy from  
reactor experiments

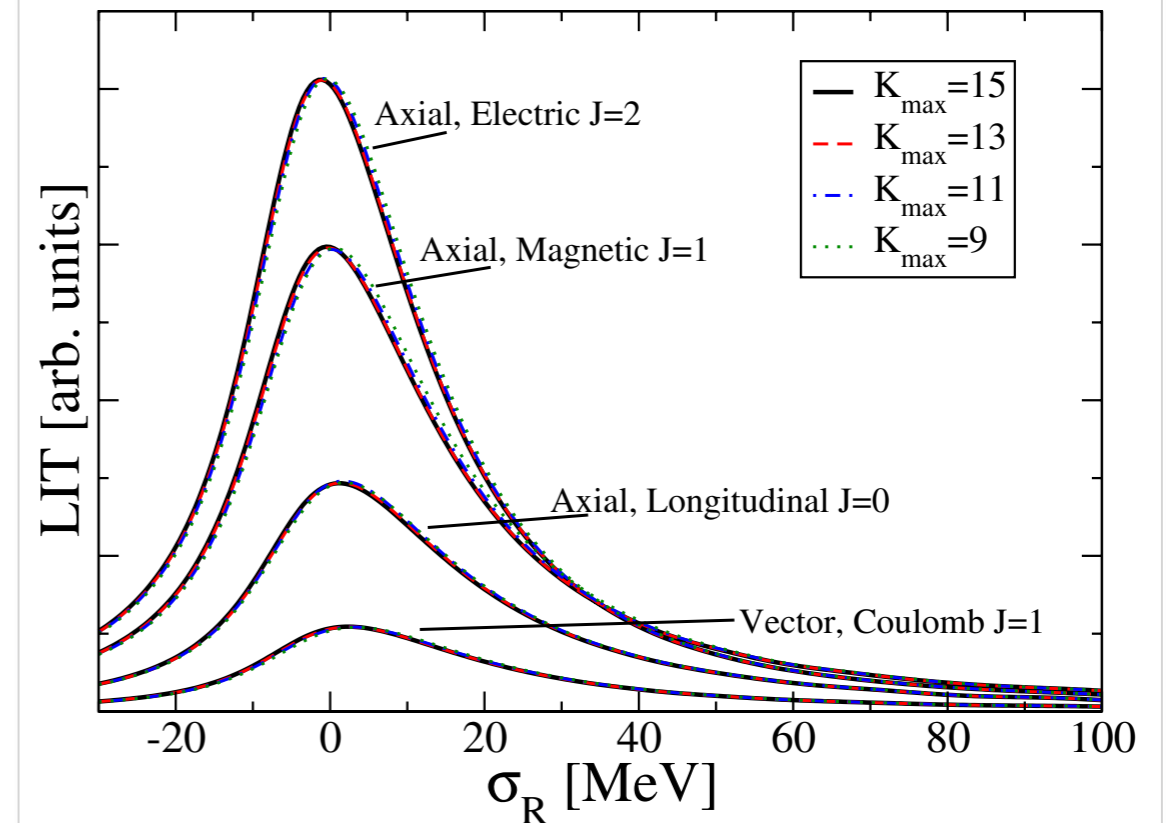
Formaggio and Zeller

# Neutrino-<sup>4</sup>He Scattering

$T$ [MeV]	$(\nu_x, \nu'_x)$	$\langle \sigma \rangle_T$ [ $10^{-42}$ cm <sup>2</sup> ]		
		$(\bar{\nu}_x, \bar{\nu}'_x)$	$(\nu_e, e^-)$	$(\bar{\nu}_e, e^+)$
2	$1.47 \times 10^{-6}$	$1.36 \times 10^{-6}$	$7.40 \times 10^{-6}$	$5.98 \times 10^{-6}$
4	$1.73 \times 10^{-3}$	$1.59 \times 10^{-3}$	$8.60 \times 10^{-3}$	$6.84 \times 10^{-3}$
6	$3.34 \times 10^{-2}$	$3.07 \times 10^{-2}$	$1.63 \times 10^{-1}$	$1.30 \times 10^{-1}$
8	$2.00 \times 10^{-1}$	$1.84 \times 10^{-1}$	$9.61 \times 10^{-1}$	$7.68 \times 10^{-1}$
10	$7.09 \times 10^{-1}$	$6.54 \times 10^{-1}$	3.36	2.71

Thermal averaged cross sections  
 Few % impact of two-nucleon currents

Fairly simple nuclei;  
 no bound excited states  
 Achievable errors  $\ll$  10%



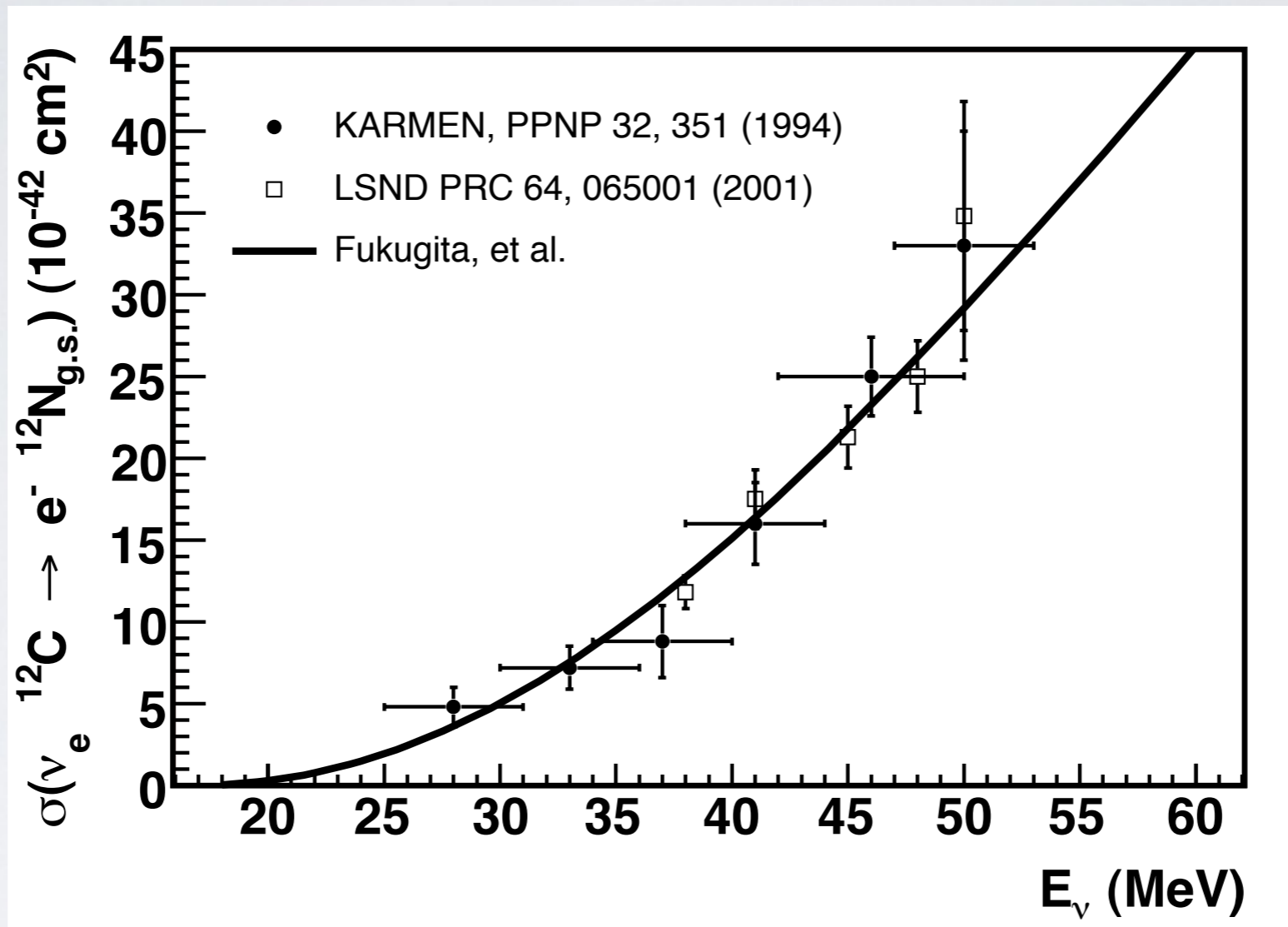
convergence in each partial wave

Gazit, Barnea, PRL 2007

**No data to compare with**

# Neutrino- $^{12}\text{C}$ Scattering

## Experiment: Karmen and LSND



$\mathbf{U}_e$  charged current to  $^{12}\text{N}$   
from muon decay at rest

theory errors estimated at  $\sim 20\%$ , Fukugita et al., 1988



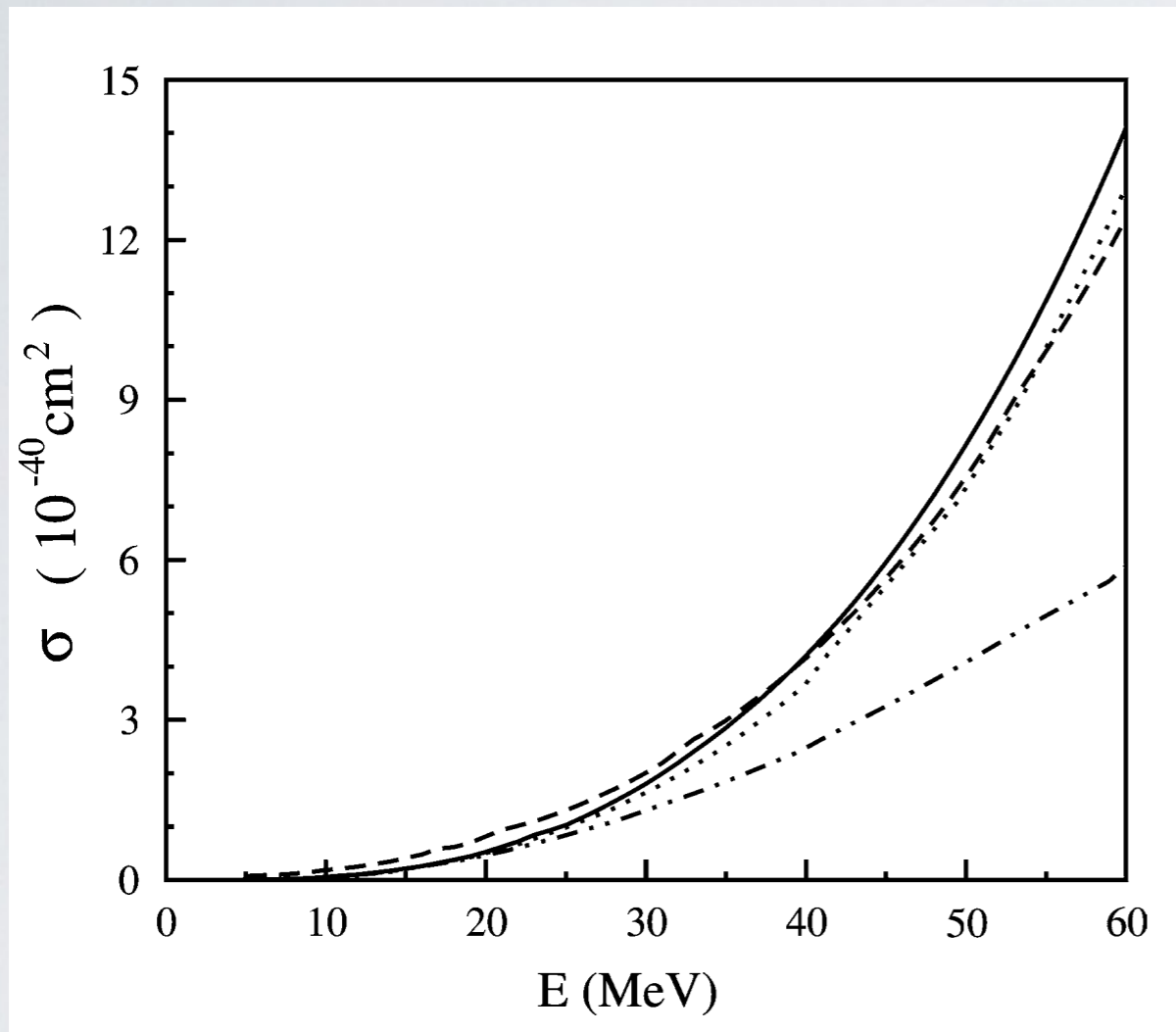
# Neutrino- $^{12}\text{C}$ Scattering

Neutrino charge current scattering  
from  $^{12}\text{C}$  (LSND/Karmen)

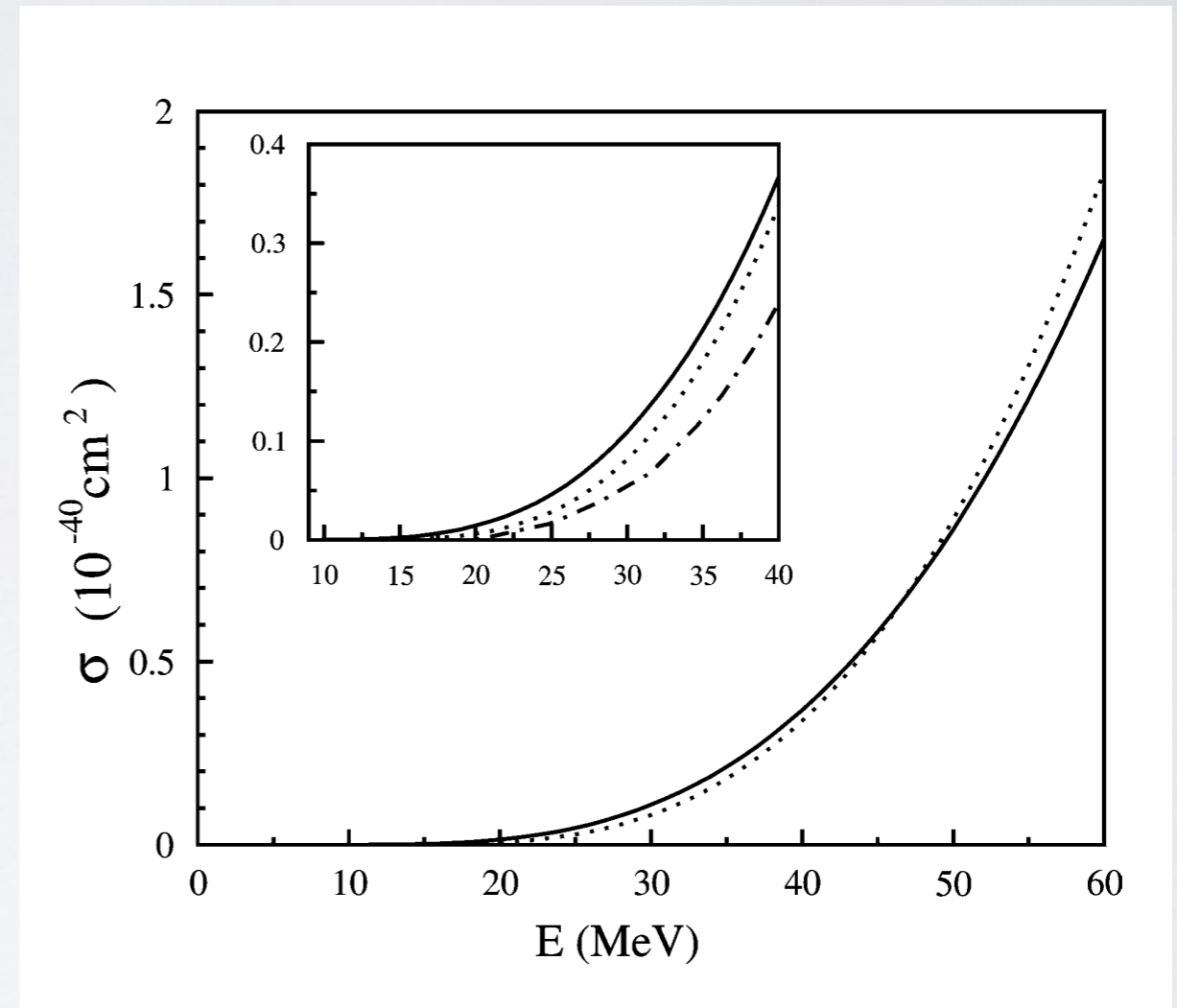
$^{12}\text{C}$	$^{12}\text{C}(\nu_e, e^-)^{12}\text{N}_{\text{g.s.}}$	Stopped $\pi/\mu$	KARMEN	$9.1 \pm 0.5(\text{stat}) \pm 0.8(\text{sys})$	9.4 [Multipole](Donnelly and Peccei, 1979)
		Stopped $\pi/\mu$	E225	$10.5 \pm 1.0(\text{stat}) \pm 1.0(\text{sys})$	9.2 [EPT] (Fukugita <i>et al.</i> , 1988).
		Stopped $\pi/\mu$	LSND	$8.9 \pm 0.3(\text{stat}) \pm 0.9(\text{sys})$	8.9 [CRPA] (Kolbe <i>et al.</i> , 1999b)
	$^{12}\text{C}(\nu_e, e^-)^{12}\text{N}^*$	Stopped $\pi/\mu$	KARMEN	$5.1 \pm 0.6(\text{stat}) \pm 0.5(\text{sys})$	5.4-5.6 [CRPA] (Kolbe <i>et al.</i> , 1999b)
		Stopped $\pi/\mu$	E225	$3.6 \pm 2.0(\text{tot})$	4.1 [Shell] (Hayes and S, 2000)
		Stopped $\pi/\mu$	LSND	$4.3 \pm 0.4(\text{stat}) \pm 0.6(\text{sys})$	
	$^{12}\text{C}(\nu_\mu, \mu^-)^{12}\text{N}_{\text{g.s.}}$	Decay in Flight	LSND	$56 \pm 8(\text{stat}) \pm 10(\text{sys})$	68-73 [CRPA] (Kolbe <i>et al.</i> , 1999b)
					56 [Shell] (Hayes and S, 2000)

Little evidence for important 2N current effects  
for 30-100 MeV neutrinos

# Neutrino - Ar Scattering



inclusive  $\nu_e$  charged current  $^{40}\text{Ar}$   
Athar, et al, 2004



anti- $\nu$  charged current to Cl

Significant differences

## *Neutrinos in Matter*

- Many studies in mean-field models, perhaps accurate enough in many cases
- Virial expansion should be accurate in hot dilute matter
- Should use same interactions/currents in nuclei and matter. More reliable constraints.
- Matter results are less directly connected to experiment for astrophysical energies, particularly for very neutron-rich matter.

***Can we identify important regimes where more accuracy is required; similarities in nuclear and matter responses?***

# Status and Outlook

- Microscopic inputs reasonably well defined (interactions, currents)
- Future inputs on one- and two-nucleon level from lattice QCD
- Accurate calculations possible in light nuclei
- Critical for accelerator neutrino energies
- What future experiments on nuclei are most valuable?
- More realistic studies of neutrinos in matter (Reddy, Schwenk, ...)

