

# Vortex pinning and dynamics in the neutron star crust

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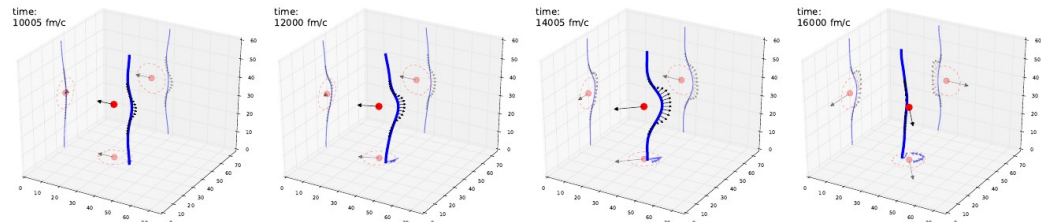
Reference: *arXiv:1606.04847*

Supported by:

- Polish National Science Center (NCN) grant under decision No. DEC-2014/13/D/ST3/01940.

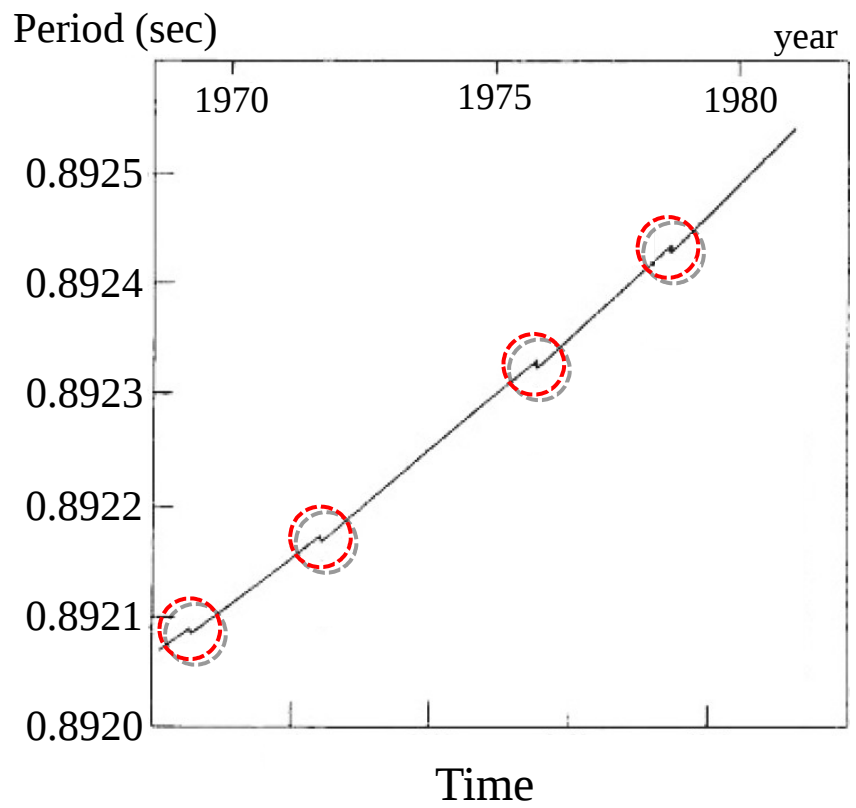
- Polish National Science Center (NCN) grant under decision No. DEC-2013/08/A/ST3/00708.

- U.S. DOE Office of Science Grant DE-FG02-97ER4101



# Glitch: a sudden increase of the rotational frequency

## Glitches in the Vela pulsar



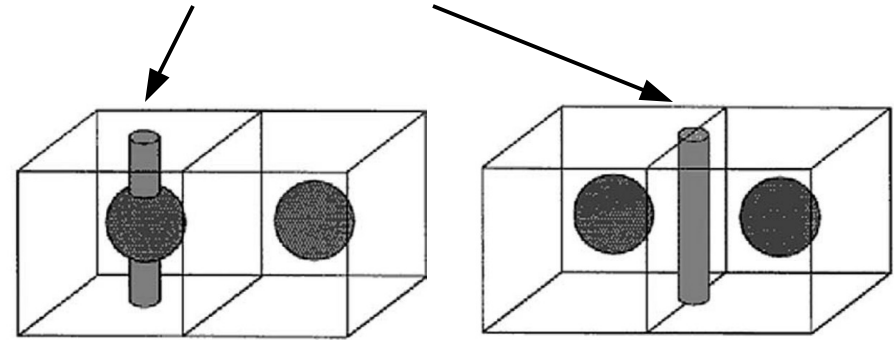
V.B. Bhatia, A Textbook of Astronomy and Astrophysics with Elements of Cosmology, Alpha Science, 2001.

First observed in 1969: V. Radhakrishnan and R. N. Manchester, Nature 222, 228–229 (1969); P. E. Reichley and G. S. Downs, Nature 222, 229–230 (1969);

## Vortex model

(P. W. Anderson and N. Itoh, Nature 256 (1975))

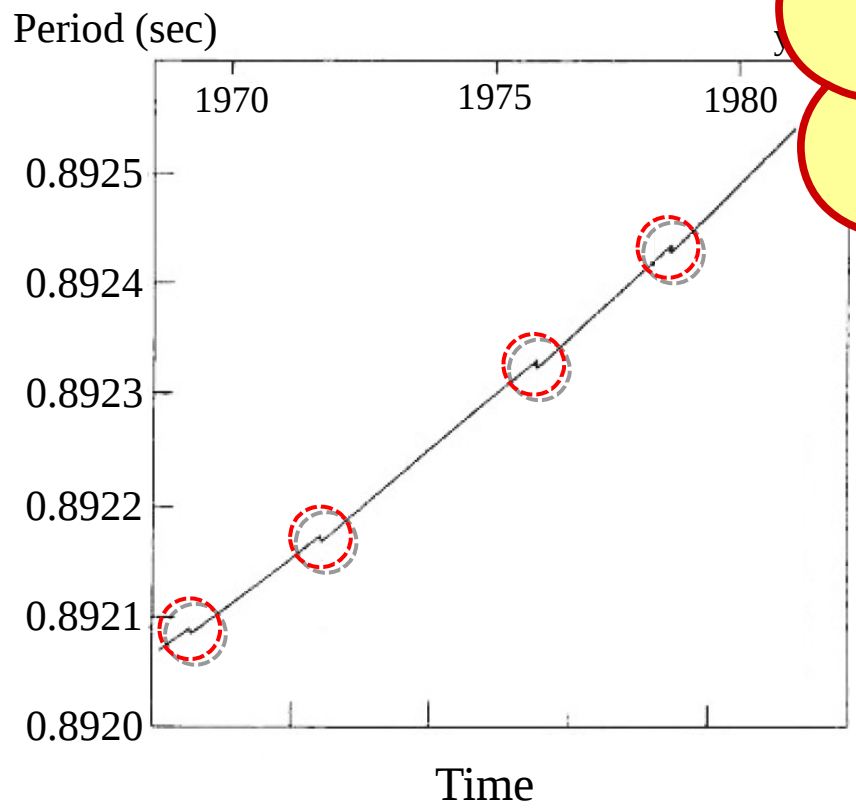
- Presently the standard picture for pulsar glitches
- Can explain: post-glitch relaxation, statistics of the glitching populations...
- Idea:
  - Superfluid interior contains quantized vortices pinned to the crustal lattice
  - Glitches are believed to occur when a large number of vortices simultaneously unpin and move outward
- Open problem: **Nuclear or Interstitial pinning**



Figs from: P. Donati et al., Nuclear Physics A 742 (2004) 363

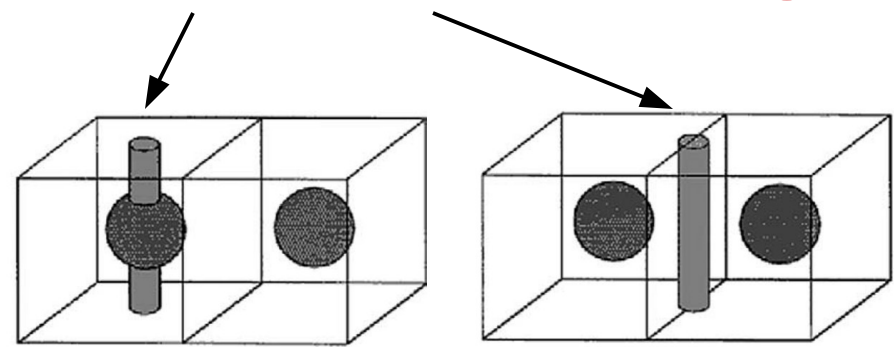
# Glitch: a sudden increase of the rotational frequency

## Glitches in the Vela pulsar



**Note:**  
 Type of pinning can affect dynamics dramatically...  
 (in case of interstitial the vortex moves around the lattice avoiding impurities)

Open problem:  
**Nuclear or Interstitial pinning**



Figs from: P. Donati et al., Nuclear Physics A 742 (2004) 363

V.B. Bhatia, A Textbook of Astronomy and Astrophysics with Elements of Cosmology, Alpha Science, 2001.

First observed in 1969: V. Radhakrishnan and R. N. Manchester, Nature 222, 228–229 (1969);  
 P. E. Reichley and G. S. Downs, Nature 222, 229–230 (1969);

*Predictions for “pinning force”:*

M.A. Alpar et al. *Astrophys.J.*213,527 (1977); 276,325 (1984)

R.I. Epstein, G.Baym, *Astrophys.J.*328,680 (1988)

R.K.Link,R.I.Epstein, *Astrophys.J.*373, 592 (1991)

Hydrodynamics  
+ GL (for pairing)

P. M. Pizzochero, L. Viverit, and R. A. Broglia,  
*Phys. Rev. Lett.* 79, 3347(1997)

P. Donati,P. Pizzochero, *Phys. Rev. Lett.* 90, 211101 (2003)

P. Donati, P.M. Pizzochero, *Nucl. Phys. A.* 742, 363 (2004)

P. Donati, P.M. Pizzochero, *Phys. Lett. B* 640, 74 (2006)

S. Seveso, P. M. Pizzochero, F. Grill, B. Haskell,  
*MNRAS* 455, 3952 (2016)

TF + LDA

P. Avogadro, F. Barranco, R. A. Broglia, and E. Vigezzi,  
*Phys. Rev. C* 75, 012805(R) (2007)

P. Avogadro, F. Barranco, R.A. Broglia, E. Vigezzi,  
*Nucl. Phys. A* 788, 130 (2007)

P. Avogadro, F. Barranco, R.A. Broglia, E. Vigezzi,  
*Nucl. Phys. A* 811, 378 (2008)

HFB

So far there is no consensus  
concerning  
the character of vortex pinning

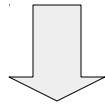
# Irrotational hydrodynamics

$$\Phi_{out}(r, \phi) = \frac{\kappa}{2\pi} \phi + A \frac{(\mathbf{r} - \mathbf{s}) \cdot (\frac{\nu}{s} \mathbf{e}_y - \mathbf{u})}{|\mathbf{r} - \mathbf{s}|^3},$$

$$\Phi_{in}(r, \phi) = B_0 + B_1(\mathbf{r} - \mathbf{s}) \cdot \frac{\kappa}{2\pi s} \mathbf{e}_y + B_2(\mathbf{r} - \mathbf{s}) \cdot \mathbf{u},$$

$$\Phi_{in}|_R = \Phi_{out}|_R,$$

$$\rho_{in} \left( \frac{\partial \Phi_{in}}{\partial r} - \vec{u} \right) |_R = \rho_{out} \left( \frac{\partial \Phi_{out}}{\partial r} - \vec{u} \right) |_R,$$



$$E = \frac{1}{2} \rho_{in} \int_{V_i} (\nabla \Phi_{in})^2 d^3 r + \frac{1}{2} \rho_{out} \int_{V - V_i - V_{vor}} (\nabla \Phi_{out})^2 d^3 r,$$

tension

effective mass

$$E = \frac{1}{4\pi} \rho_{out} \kappa^2 H \ln \left( \frac{D}{2\xi} \right) + \frac{1}{2} \left( \frac{4\pi}{3} R^3 \frac{(\rho_{out} - \rho_{in})^2}{2\rho_{out} + \rho_{in}} \right) u^2$$

$$+ \left( 2\pi R^3 \frac{\rho_{out}(\rho_{in} - \rho_{out})}{2\rho_{out} + \rho_{in}} \right) \left( \frac{\kappa}{2\pi s} \right)^2 + O(1/s^3) \quad (24)$$

vortex-nucleus interaction

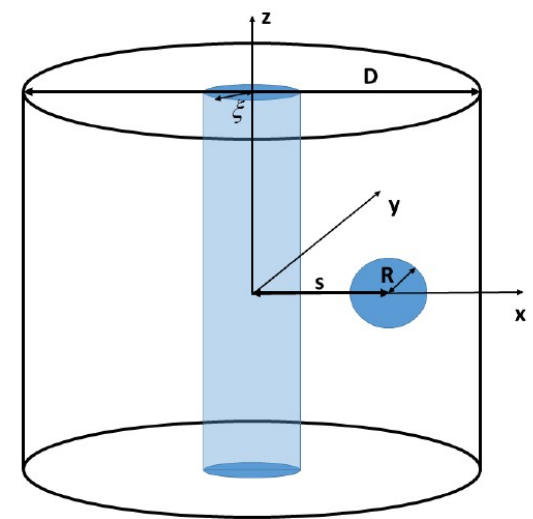


FIG. 9. (Color online) Schematic picture showing the mutual arrangement of the vortex (of radius  $\xi$ ) and the impurity (of radius  $R$ ) together with a cylinder of diameter  $D$  ( $D \rightarrow \infty$ ) defining the boundary of the system.

$$F = -\frac{\partial E}{\partial s} \propto 1/s^3$$

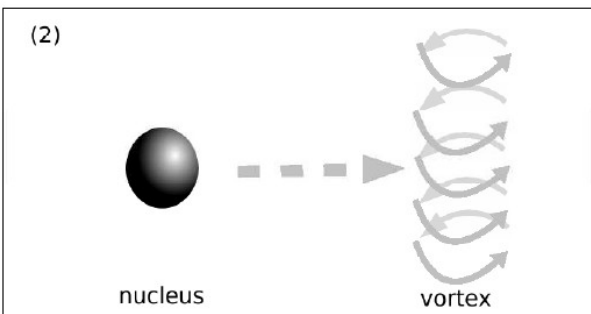
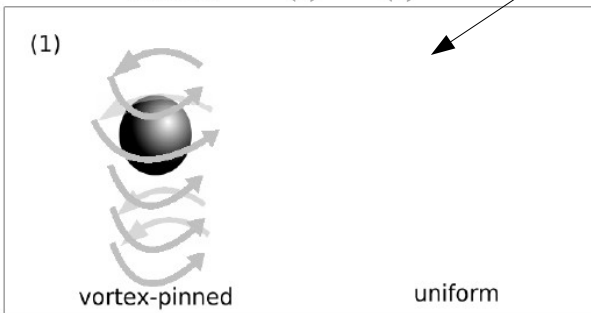
# Vortex-impurity interaction – present status

only calculations of *pinning energy*...

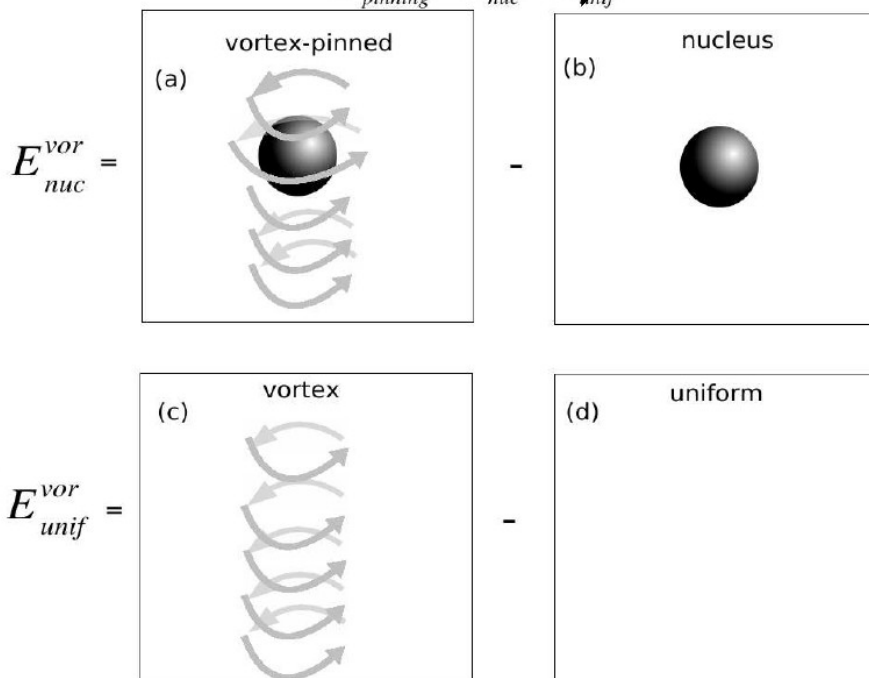
only symmetric configurations have been considered with energies of order 10,000 MeV

~1 MeV    ~10,000 MeV

$$E_{pinning} = E_{(1)} - E_{(2)}$$



$$E_{pinning} = E_{nuc}^{vor} - E_{unif}^{vor}$$



Calculations for fixed chemical potential + correction  $\Delta E = \mu \Delta N$

Fixed chemical potential or neutron background density?

Volume: tube of radius 30fm and height of 40fm

SLy4  $n_\infty = 0.026 \text{ fm}^{-3}$

Pinned	8.55	12956.57
Nucleus	8.50	12954.02
Vortex	8.55	13712.91
Uniform	8.50	13617.05
[MeV]	Chem.pot	energy

Figures and numbers from:  
 P. Avogadro, F. Barranco, R.A. Broglia, E. Vigezzi Nucl. Phys. A811, 378-412 (2008)



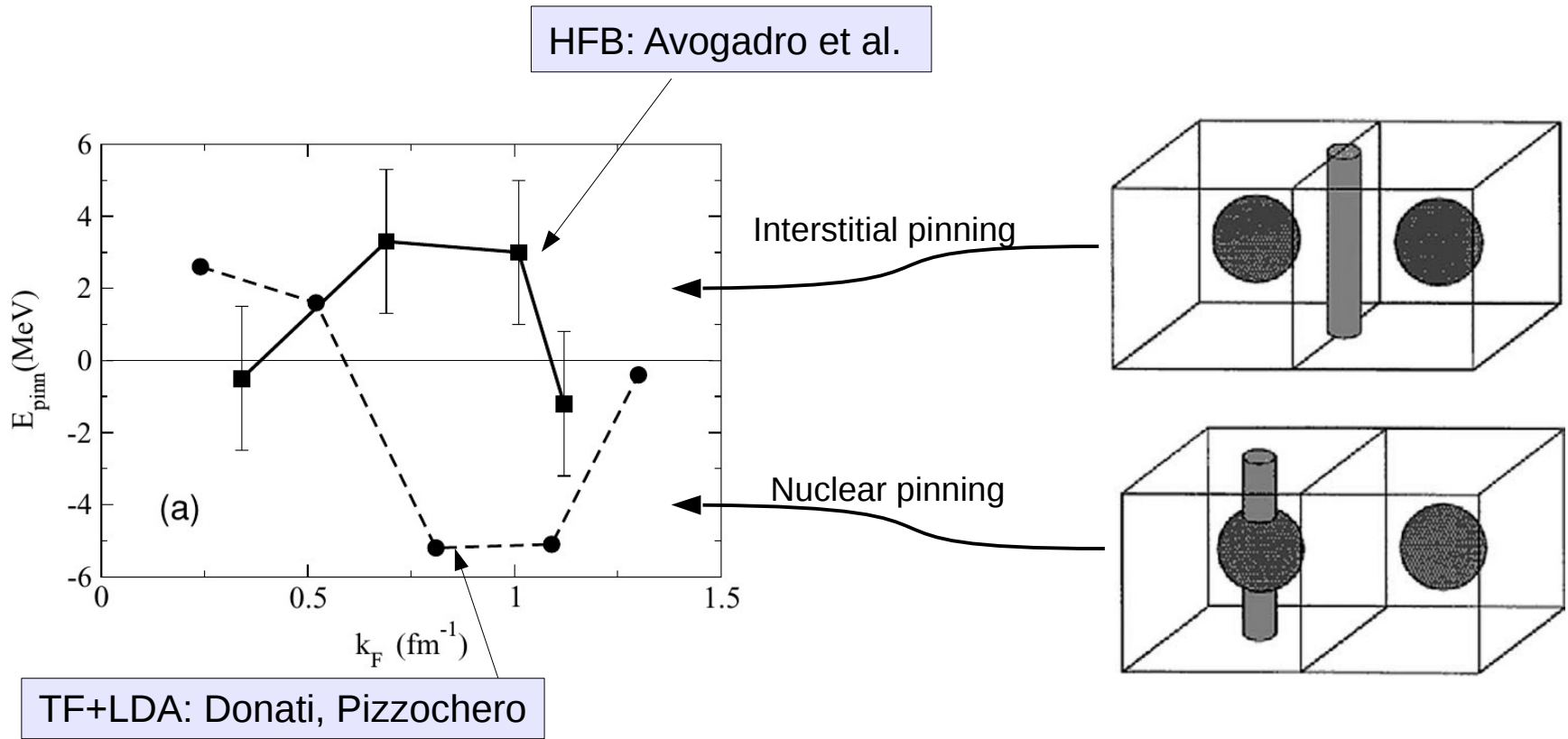
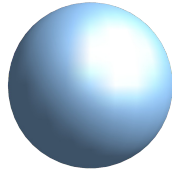
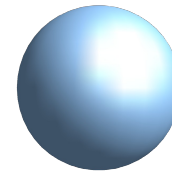


Fig. from: P. Avogadro et al.,  
Phys. Rev. C 75, 012805(R) (2007)

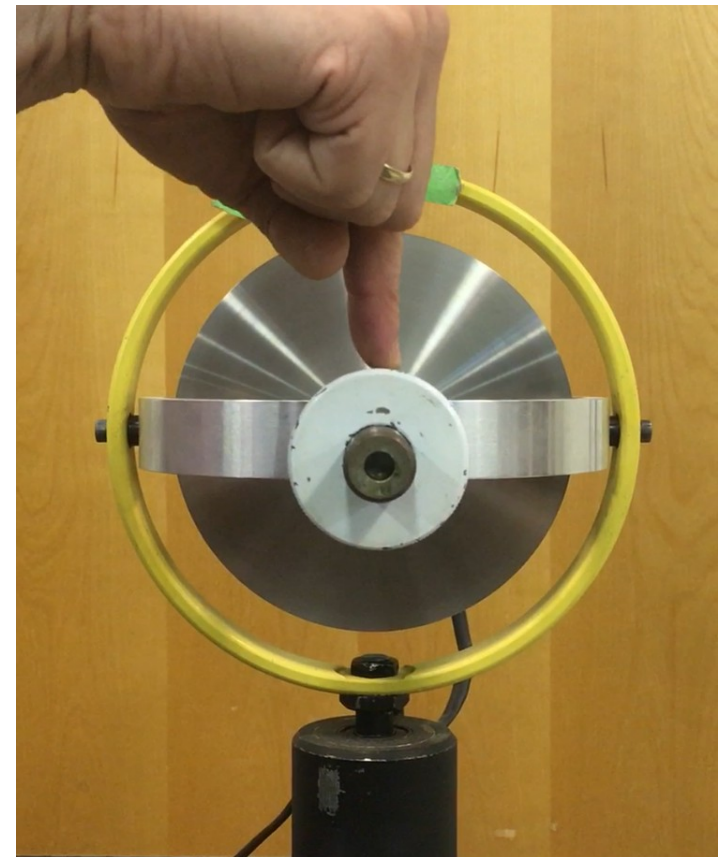
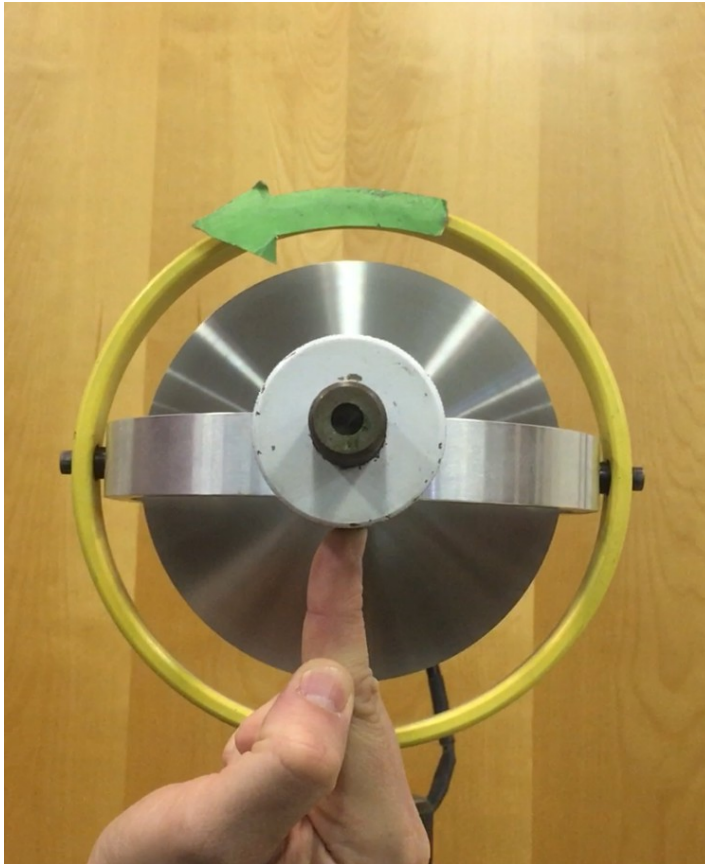
Figs from: P. Donati et al.,  
Nuclear Physics A 742 (2004) 363



*attracted by nucleus*



*repelled by nucleus*



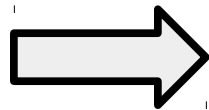
**What is response of the gyroscope when pushed?**



## Our strategy ...

Instead of solving  
static problem...

$$\hat{H}\psi = E\psi$$

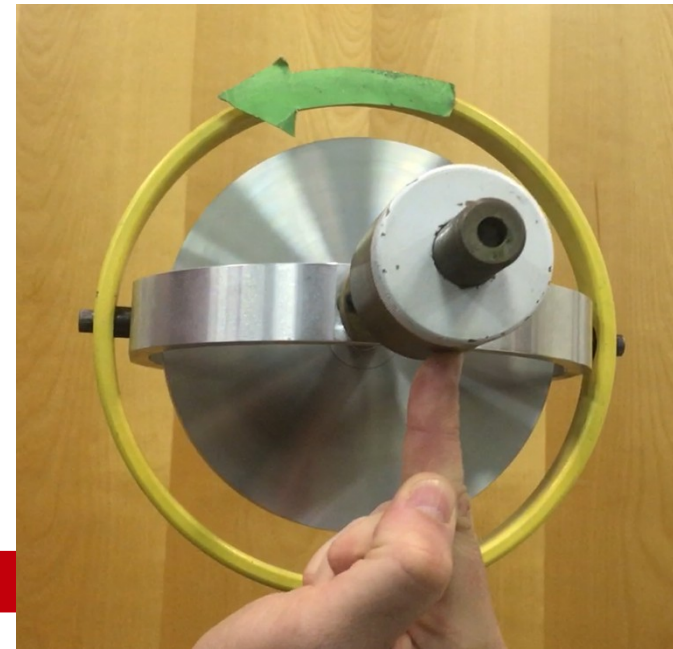


... we solve  
time-dependent problem...

$$i\hbar \frac{\partial \psi}{\partial t} = \hat{H}\psi$$

... and observing dynamics of the system  
we determine the forces.

**Unambiguous** determination of the force sign...

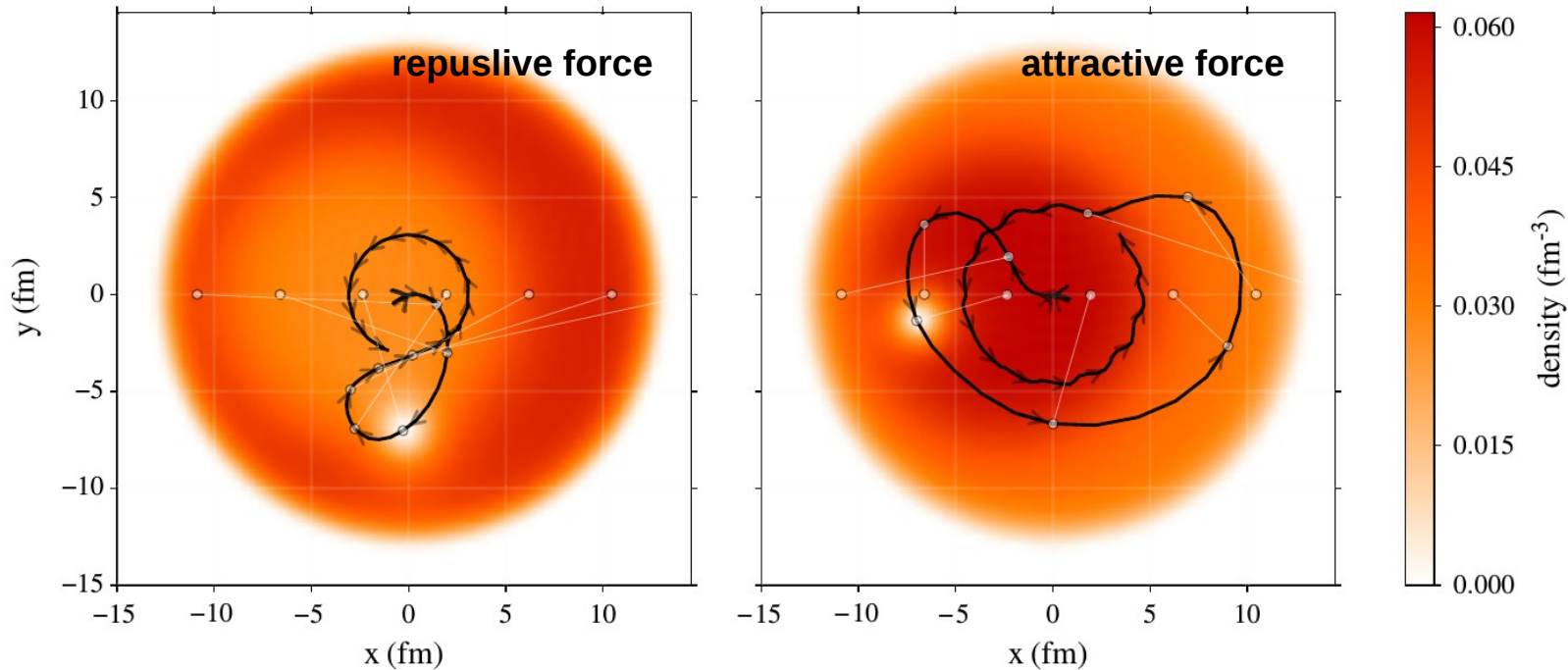


Idea introduced in:

Aurel Bulgac, Michael McNeil Forbes, and Rishi Sharma

Phys. Rev. Lett. 110, 241102 (2013)

## 2D simulations with GP



Equation of motion:

$$\underbrace{M\ddot{\vec{r}}_v - \vec{f}_{qp}}_{\text{negligible}} = \rho_s \vec{K} \times (\dot{\vec{r}}_v - \vec{v}_s) + \vec{F}_v$$

repulsive

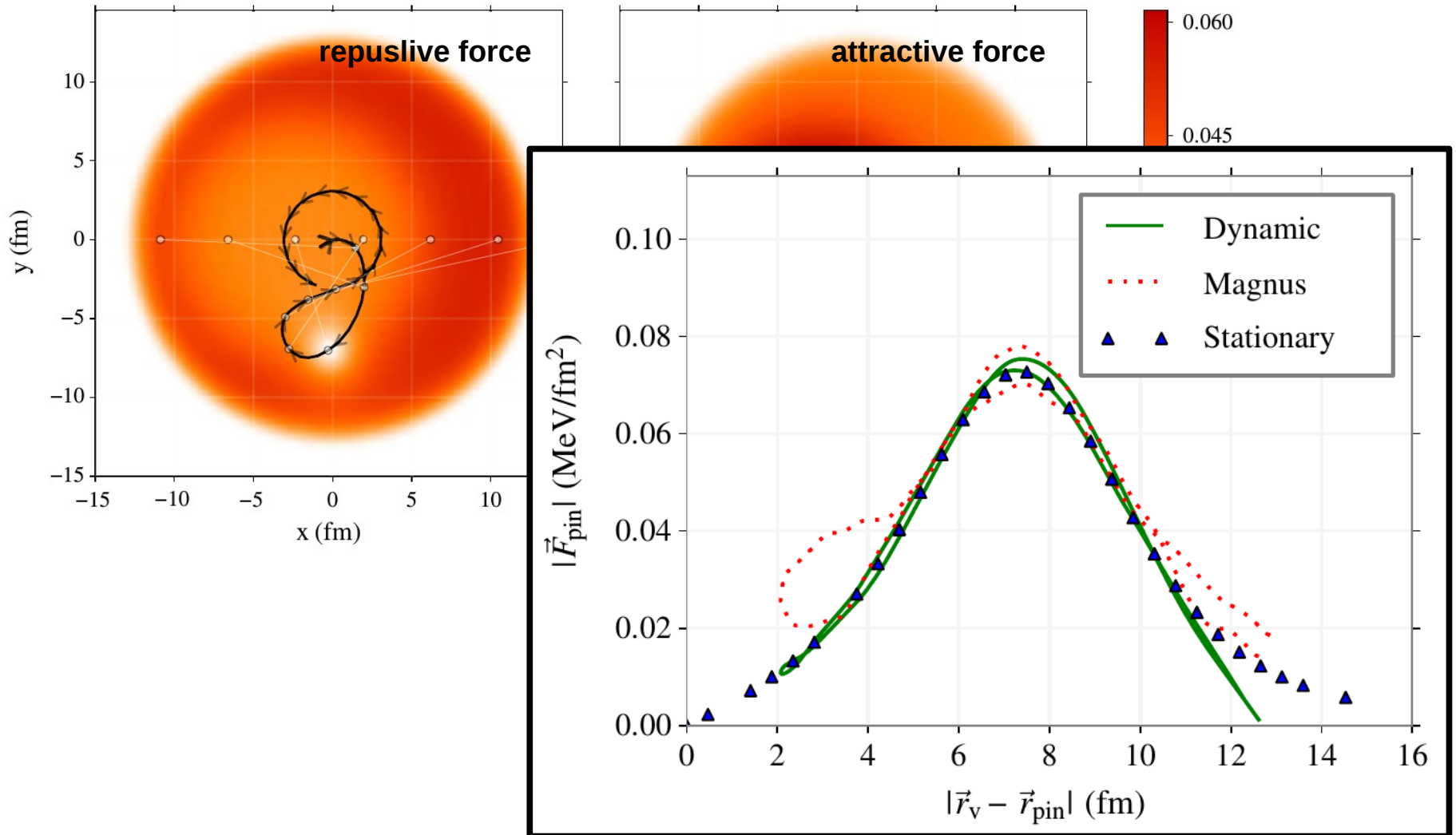
attractive

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Phys. Rev. Lett. 110, 241102 (2013)

## 2D simulations with GP

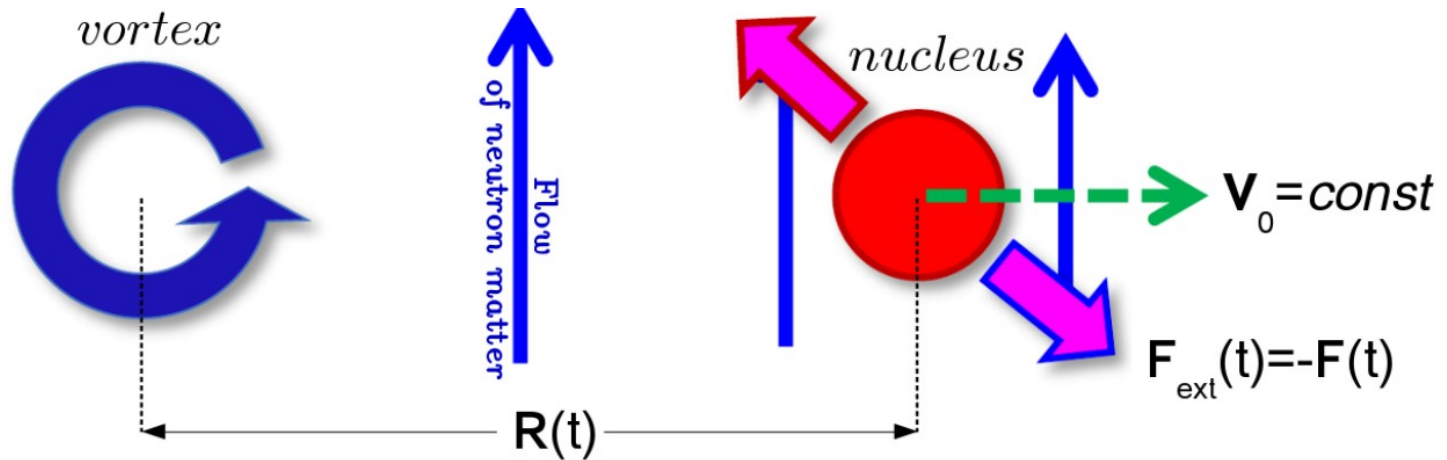


## 3D case

In 3D one should consider the equation of motion for the vortex line...

We use Newton's 3<sup>rd</sup> law and extract the force from motion of the nucleus.....

$$M \frac{d\mathbf{v}}{dt} = \mathbf{F}_{\text{tot.}} = \mathbf{F} + \mathbf{F}_{\text{ext.}} = 0 \Leftrightarrow \mathbf{v} = \text{const}$$



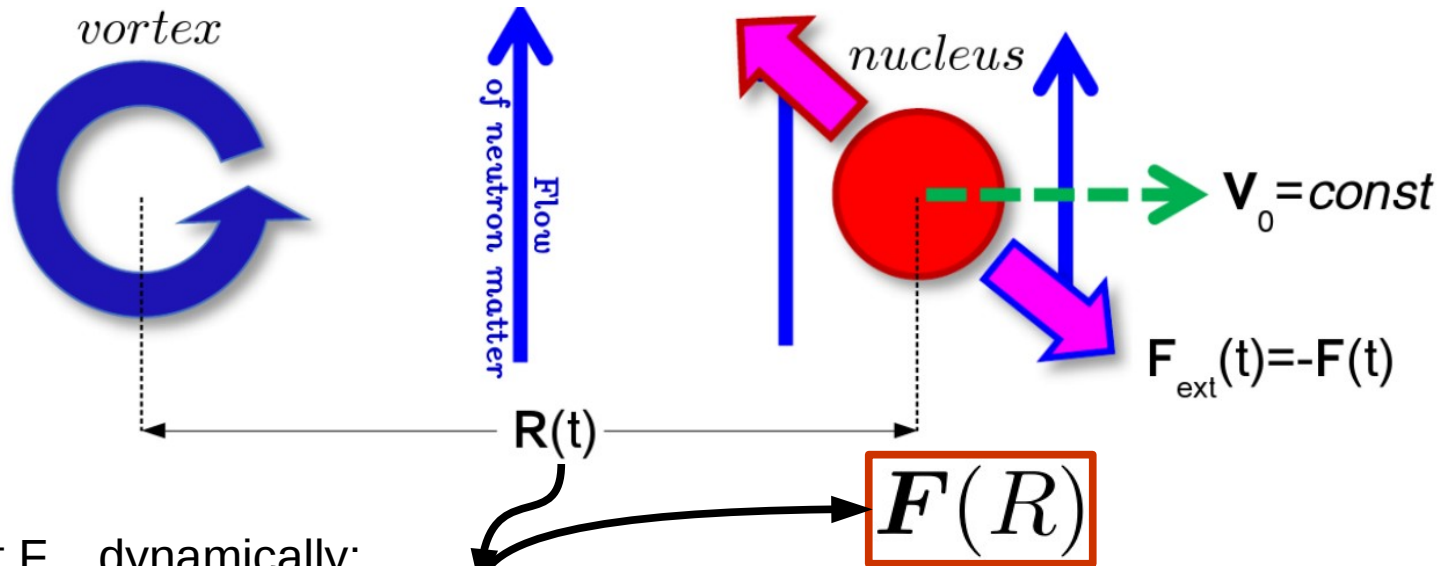
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Vortex-nucleus interaction



We construct  $\mathbf{F}_{\text{ext}}$  dynamically:

$$\mathbf{F}_{\text{ext}}(t + \Delta t) = \mathbf{F}_{\text{ext}}(t) - \alpha [\mathbf{v}(t) - \mathbf{v}_0]$$

# We performed 3D, dynamical simulations by TDDFT with superfluidity

## □ TDSLDA equations (similar to TDHFB, TD-BdG)

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_k \\ v_k \end{pmatrix} = \begin{pmatrix} h & \Delta \\ \Delta^* & -h \end{pmatrix} \begin{pmatrix} u_k \\ v_k \end{pmatrix}$$

## □ Energy density functional (EDF)

$$\mathcal{E} = \mathcal{E}_0 + \mathcal{E}_{\text{pair}}$$

$\mathcal{E}_0$  : Fayans EDF (FaNDF<sup>0</sup>) w/o LS

S.A. Fayans, JETP Letters 68, 169 (1998);

**FP81**: B. Friedman and V. R. Pandharipande, Nucl. Phys. A 361, 502 (1981)

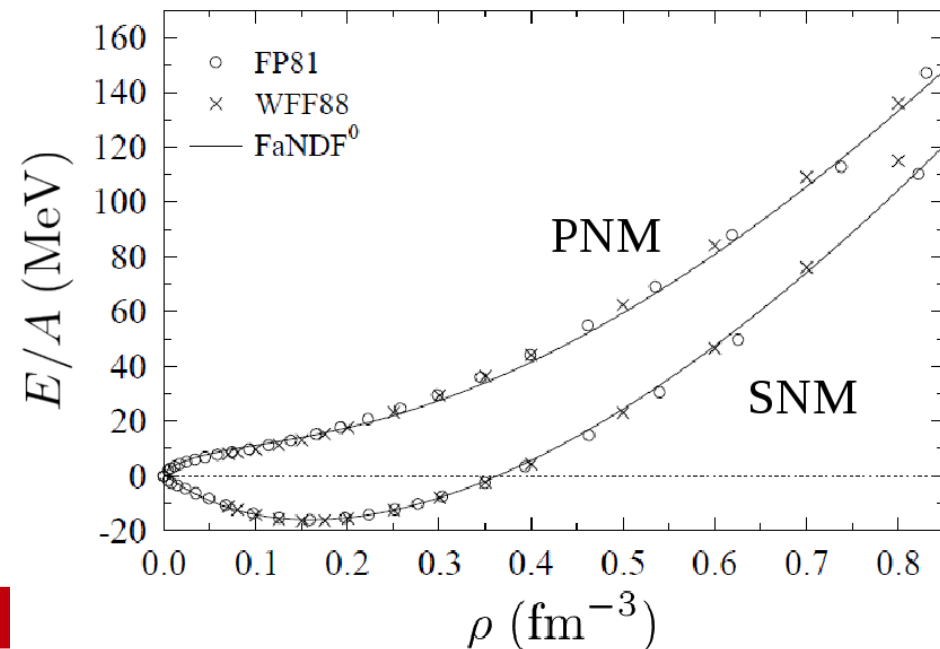
**WFF88**: R. B. Wiringa, V. Fiks, and A. Fabrocini, Phys. Rev. C 38, 1010 (1988).

## □ Potentials

$$h = \frac{\delta \mathcal{E}}{\delta n}, \quad \Delta = \frac{\delta \mathcal{E}}{\delta \nu^*}$$

$$n(\mathbf{r}) = \sum_{0 < E_k < E_c} |v_k(\mathbf{r})|^2$$

$$\nu(\mathbf{r}) = \sum_{0 < E_k < E_c} u_k(\mathbf{r})v_k^*(\mathbf{r})$$





# We performed 3D, dynamical simulations by TDDFT with superfluidity

## □ TDSLDA equations (similar to TDHFB, TD-BdG)

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## □ Energy density functional (EDF)

$$\mathcal{E} = \mathcal{E}_0 + \mathcal{E}_{\text{pair}}$$

$$\mathcal{E}_{\text{pair}}(\mathbf{r}) = g(n(\mathbf{r})) [|\nu_n(\mathbf{r})|^2 + |\nu_p(\mathbf{r})|^2]$$

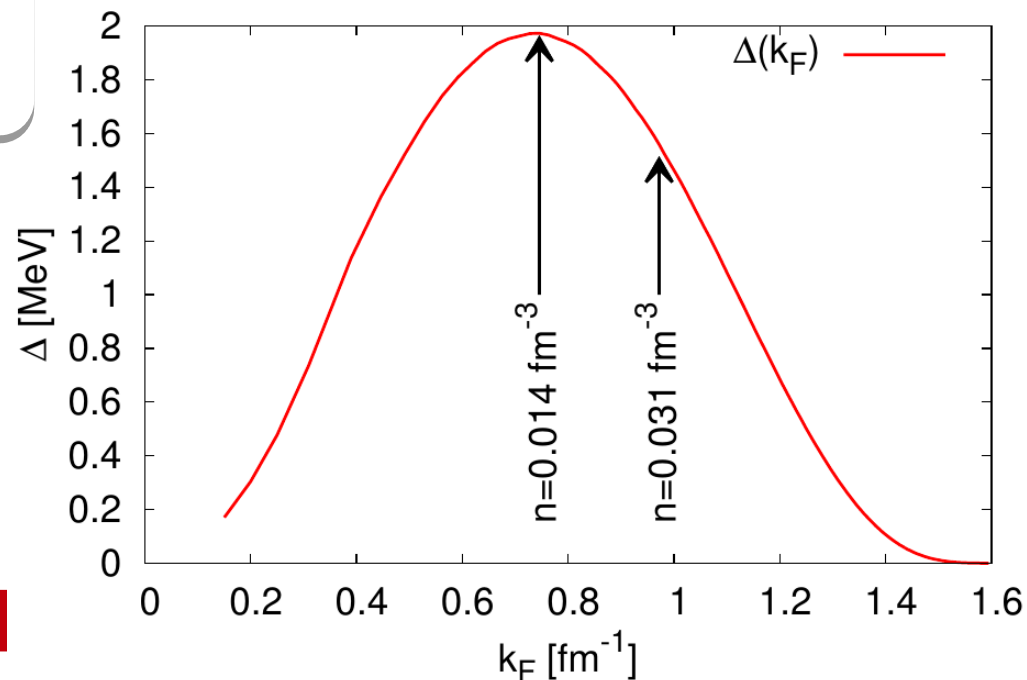
The coupling constant  $g$  is chosen to reproduce the neutron pairing gap in pure neutron matter.

## □ Potentials

$$h = \frac{\delta \mathcal{E}}{\delta n}, \quad \Delta = \frac{\delta \mathcal{E}}{\delta \nu^*}$$

$$n(\mathbf{r}) = \sum_{0 < E_k < E_c} |v_k(\mathbf{r})|^2$$

$$\nu(\mathbf{r}) = \sum_{0 < E_k < E_c} u_k(\mathbf{r}) v_k^*(\mathbf{r})$$



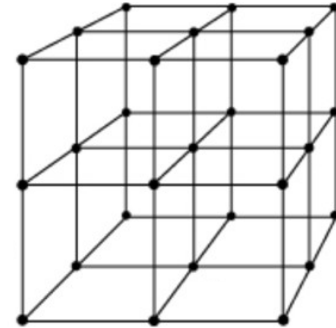
## Numerical details

Determined self-consistently in each moment

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_k(\mathbf{r}, t) \\ v_k(\mathbf{r}, t) \end{pmatrix} = \begin{pmatrix} h(\mathbf{r}, t) & \Delta(\mathbf{r}, t) \\ \Delta^*(\mathbf{r}, t) & -h(\mathbf{r}, t) \end{pmatrix} \begin{pmatrix} u_k(\mathbf{r}, t) \\ v_k(\mathbf{r}, t) \end{pmatrix}$$

HFB matrix = 200,000<sup>2</sup>

- ♦ The system is placed on a large 3D spatial lattice of size 50 x 50 x 40 with lattice spacing 1.5fm



- ★ Discrete Variable Representation (DVR) - solid framework (see for example: Bulgac, Forbes, Phys. Rev. C 87, 051301(R) (2013))
- ★ No symmetry restrictions
- ♦ Number of PDEs is of the order of the number of spatial lattice points [Typically: 10<sup>5</sup> – 10<sup>6</sup>]  
...106,000 to evolve...

## Numerical details

Determined self

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_k(\mathbf{r}, t) \\ v_k(\mathbf{r}, t) \end{pmatrix} = \begin{pmatrix} h(\mathbf{r}, t) \\ \Delta^*(\mathbf{r}, t) \end{pmatrix}$$

- ▶ The system is placed on a lattice of size  $50 \times 50 \times 40$  with lattice spacing  $a$
- ★ Discrete Variable Representation (DVR) in a solid framework (see for example: Phys. Rev. C 87, 051301(R) (2013))
- ★ No symmetry restrictions
- ▶ Number of PDEs is of the order of the spatial lattice points [Typically ...106,000 to evolve...]
- ▶ parallelization (MPI) + acceleration (GPU)



OLCF Oak Ridge Titan (DOE Contract DE-AC05-00OR22725)

NERSC Edison (DOE Contract No. DE-AC02-05CH11231)

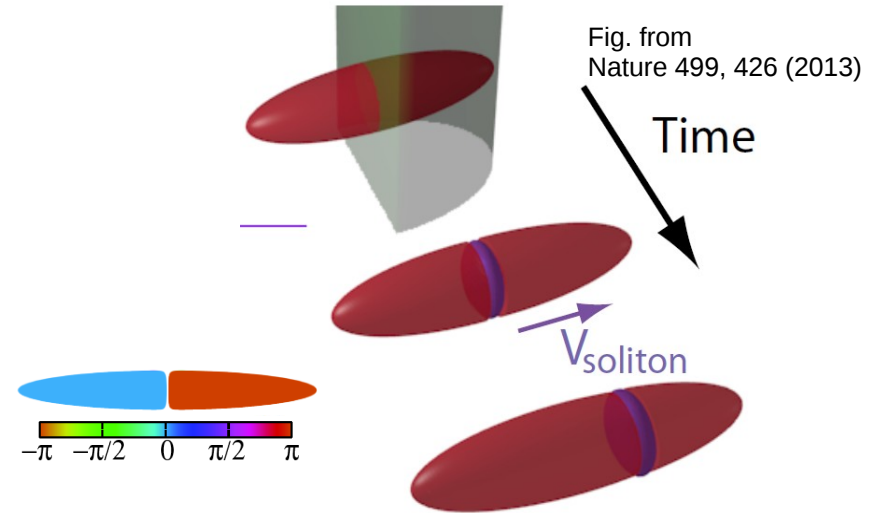


HA-PACS (PACS-VIII) system Interdisciplinary Computational Science Program in Center for Computational Sciences, University of Tsukuba

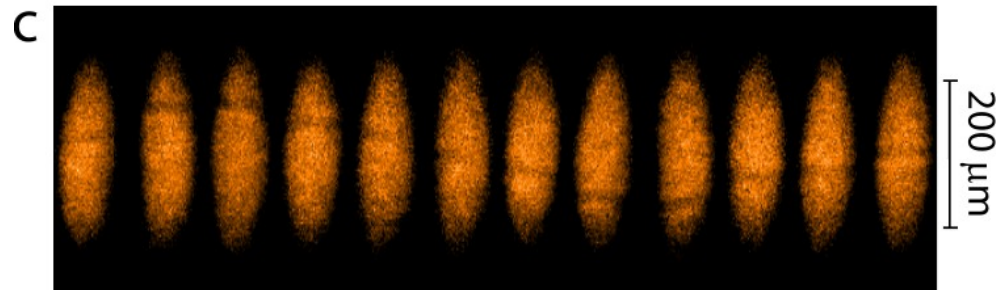
# Validation against dynamical properties of the system

MIT experiments: T. Yefsah et al., Nature 499, 426 (2013)

- ❖  ${}^6\text{Li}$  atoms near a Feshbach resonance ( $N \approx 10^6$ ) cooled in harmonic trap
- ❖ Step potential used to imprint a soliton (evolve to  $\pi$  phase shift)
- ❖ Let system evolve...
- ❖ Take picture (subtle imaging with tomography)



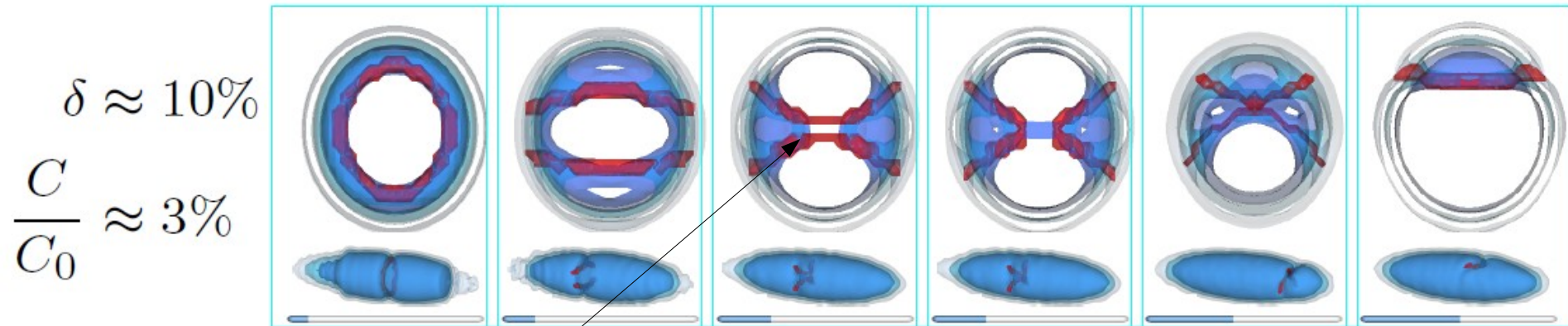
Our method predicted quantum vortex ring and corrected the initial interpretation as a heavy soliton...



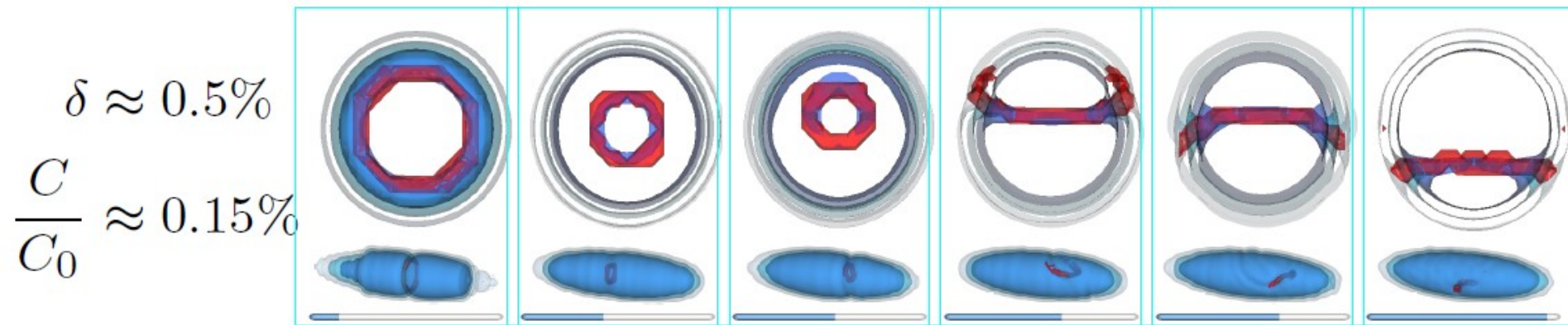
# What do fully 3D simulations reveal?

Phys. Rev. A 91, 031602 (2015)

Cold atoms 1

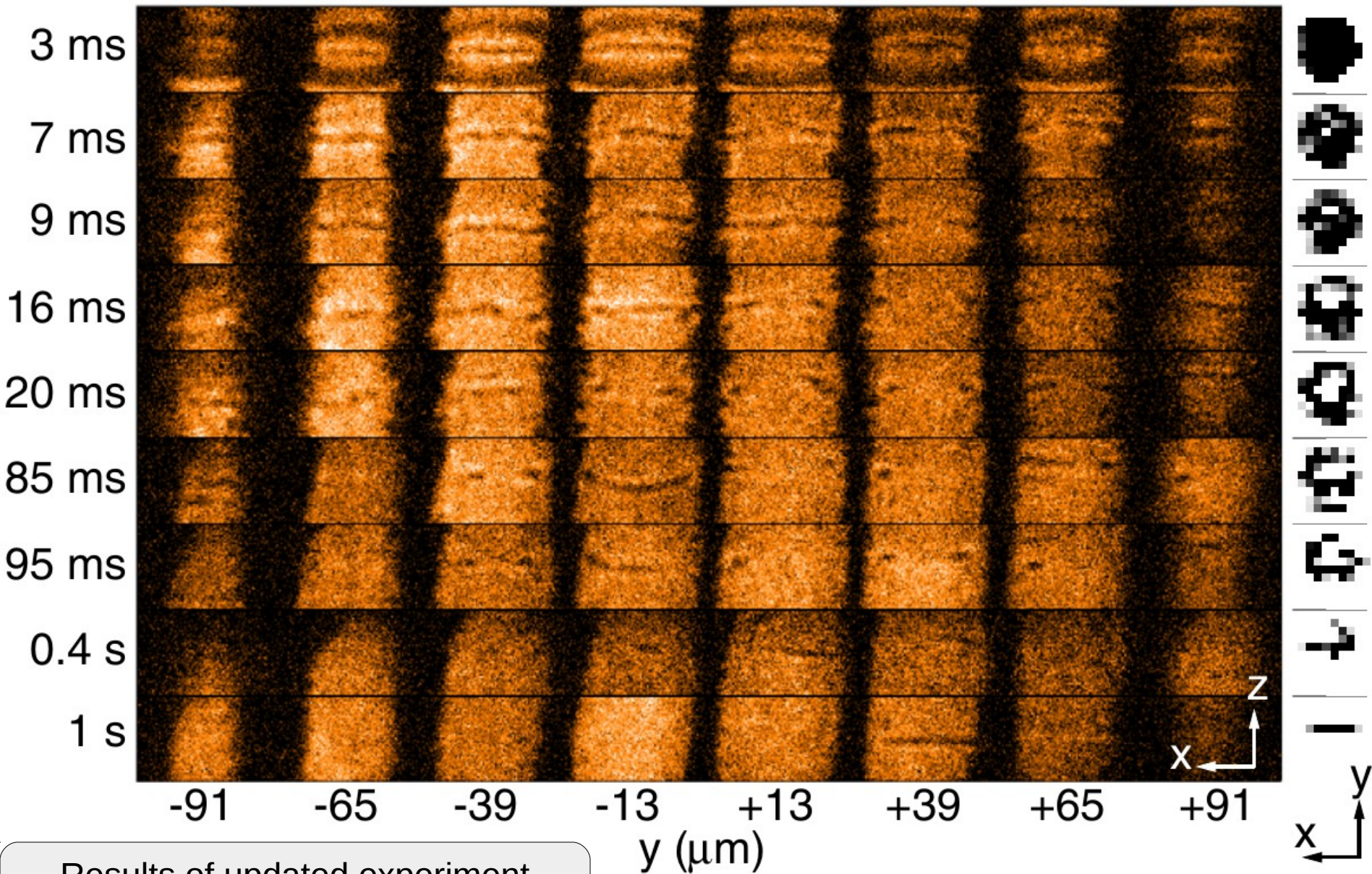


**Crossing and reconnection!**



Cold atoms 2





Results of updated experiment agree with our predictions...

Mark J. H. Ku, Biswaroop Mukherjee, Tarik Yefsah, and Martin W. Zwierlein, Phys. Rev. Lett. 116, 045304 (2016)



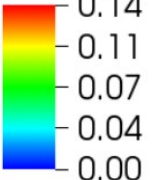
# Initial state

Self-consistent solution  
of static problem

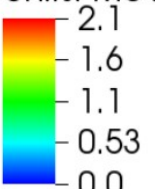
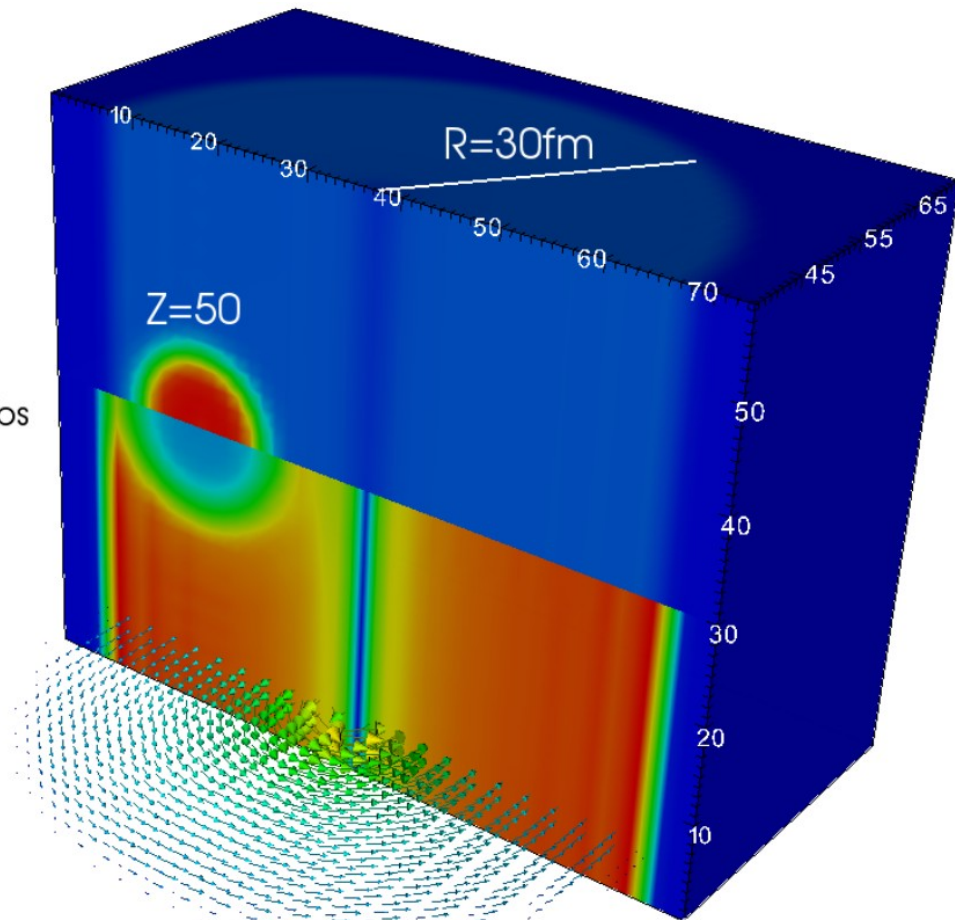
$$\begin{pmatrix} h & \Delta \\ \Delta^* & -h \end{pmatrix} \begin{pmatrix} u_k \\ v_k \end{pmatrix} = \varepsilon_k \begin{pmatrix} u_k \\ v_k \end{pmatrix}$$

J. Negele and D. Vautherin,  
Nucl. Phys. A 207, 298 (1973)

Pseudocolor  
Var: density  
Units: fm<sup>-3</sup>



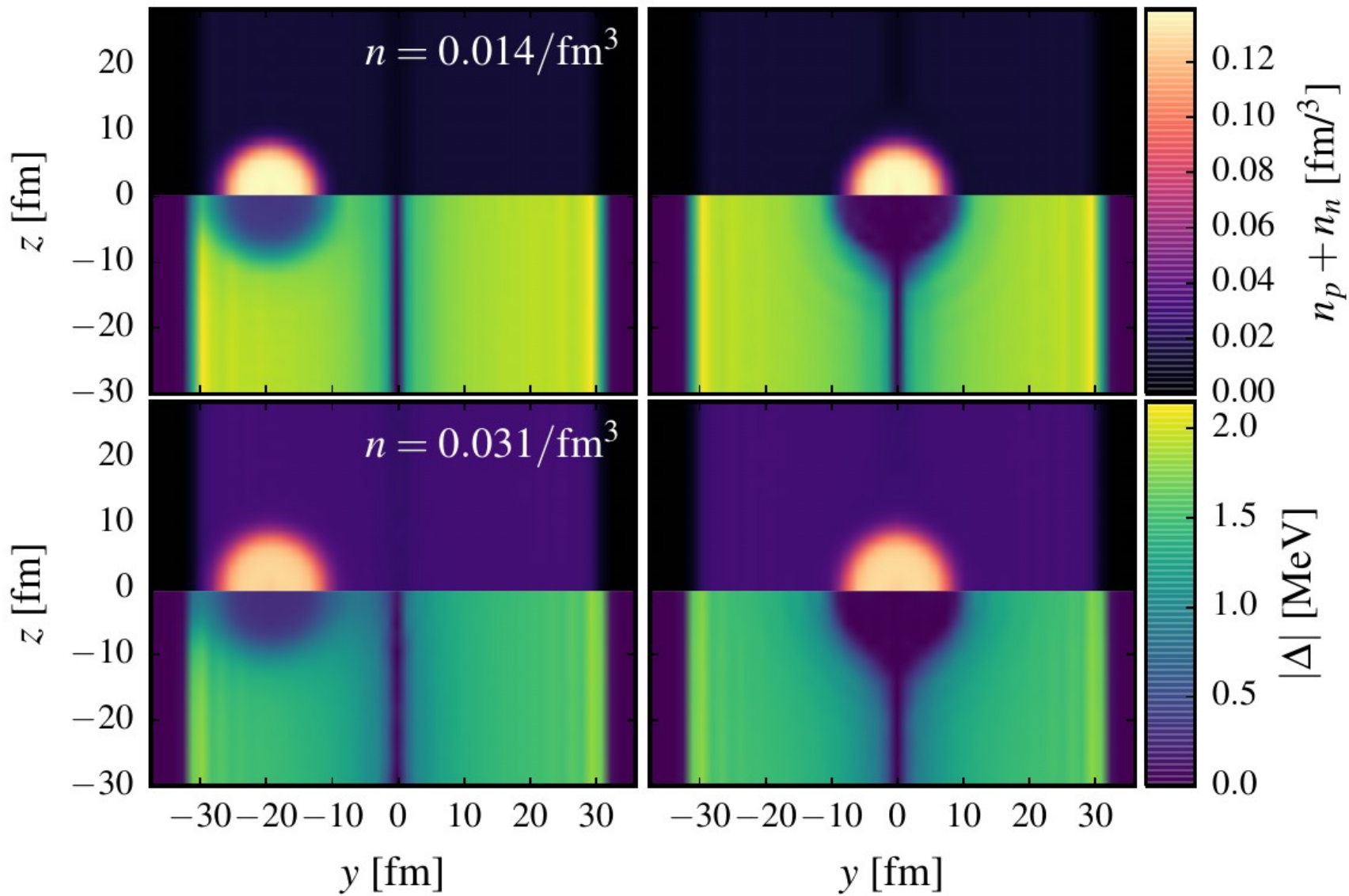
Pseudocolor  
Var: delta\_abs  
Units: MeV

75 fm × 75 fm × 60 fm  
3D lattice – no symmetry restrictions

Simulations for background density:  
 $k_F=0.75 \text{ fm}^{-1}$  and  $k_F=0.97 \text{ fm}^{-1}$

Zone	Element	Z	N	$R_{WS}$ [fm]	$\rho_b$ [g · cm <sup>-3</sup> ]	$k_{F,n}$ [fm <sup>-1</sup> ]
11	<sup>180</sup> Zr	40	140	53.6	$4.67 \cdot 10^{11}$	0.12
10	<sup>200</sup> Zr	40	160	49.2	$6.69 \cdot 10^{11}$	0.15
9	<sup>250</sup> Zr	40	210	46.4	$1.00 \cdot 10^{12}$	0.19
8	<sup>320</sup> Zr	40	280	44.4	$1.47 \cdot 10^{12}$	0.23
7	<sup>500</sup> Zr	40	460	42.2	$2.66 \cdot 10^{12}$	0.31
6	<sup>950</sup> Sn	50	900	39.3	$6.24 \cdot 10^{12}$	0.43
5	<sup>1100</sup> Sn	50	1050	35.7	$9.65 \cdot 10^{12}$	0.51
4	<sup>1350</sup> Sn	50	1300	33.0	$1.49 \cdot 10^{13}$	0.60
3	<sup>1800</sup> Sn	50	1750	27.6	$3.41 \cdot 10^{13}$	0.80
2	<sup>1500</sup> Zr	40	1460	19.6	$7.94 \cdot 10^{13}$	1.08
1	<sup>982</sup> Ge	32	950	14.4	$1.32 \cdot 10^{14}$	1.33



Static solutions:  $E_{\text{unpinned}} < E_{\text{pinned}}$

# Dragging force

external time-dependent potential couples only to protons and it is constant in space.

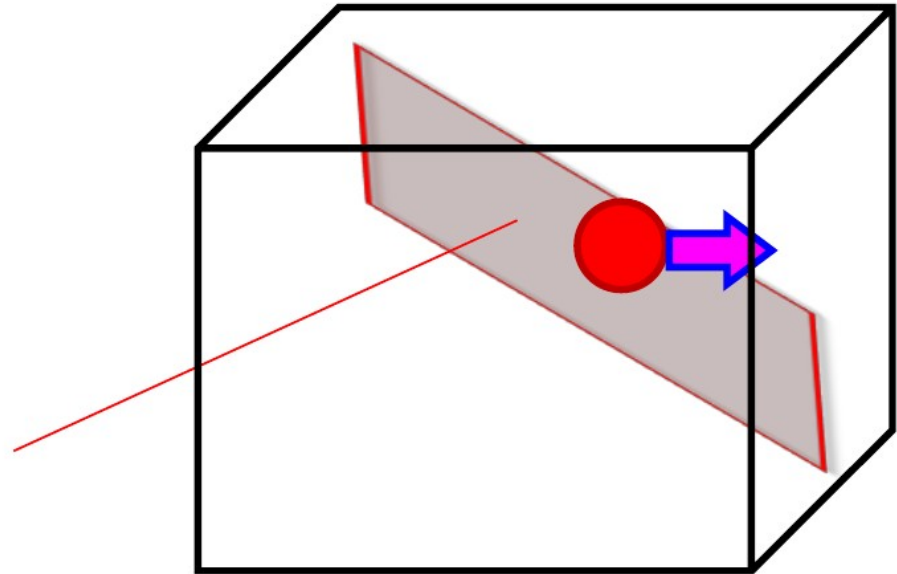
$$U_{\text{ext}}(\mathbf{r}, t) = -\frac{1}{Z} \mathbf{F}_{\text{ext}}(t) \cdot \mathbf{r}$$

$$\frac{d \langle \hat{\mathbf{p}} \rangle}{dt} = - \langle \nabla U_{\text{ext}}(\mathbf{r}, t) \rangle = \mathbf{F}_{\text{ext}}(t)$$

This force moves the center of mass of the protons together with those neutrons bound (entrained) in the nucleus without significantly modifying the internal structure of the nucleus and surrounding neutron medium

Dragging speed:

$$v_0 = 0.001c \ll v_{\text{crit}}$$



# Results

time= 0 fm/c  
 $F_m(19.1)$ = unknown  
 $Q = 28 \text{ fm}^2$

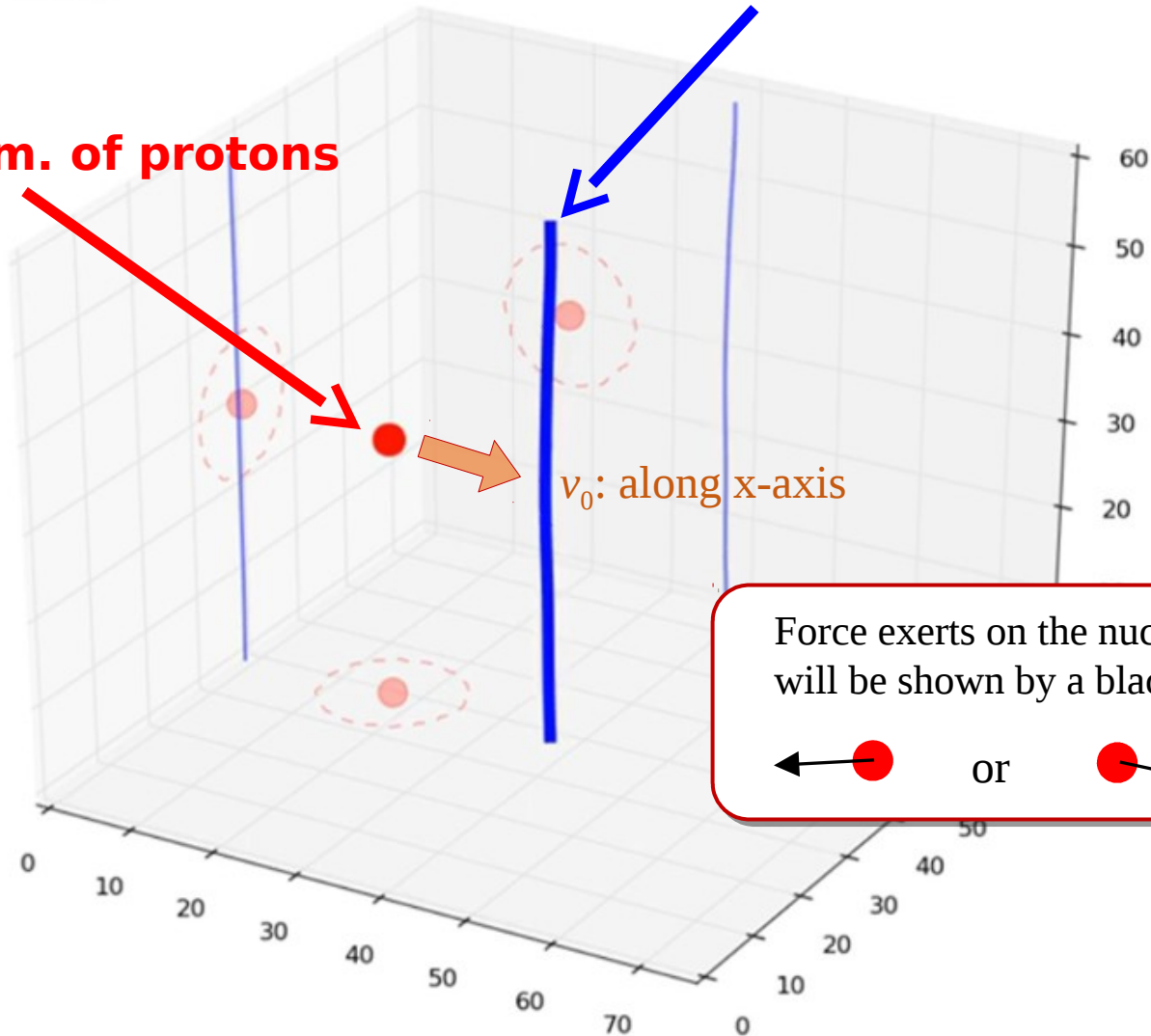
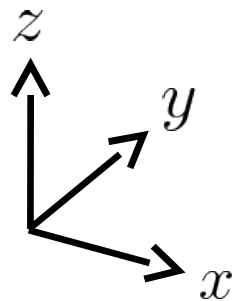
Red dot: c.m. of protons

Blue line: vortex-core

$v_0$ : along x-axis

Force exerts on the nucleus  
will be shown by a black arrow

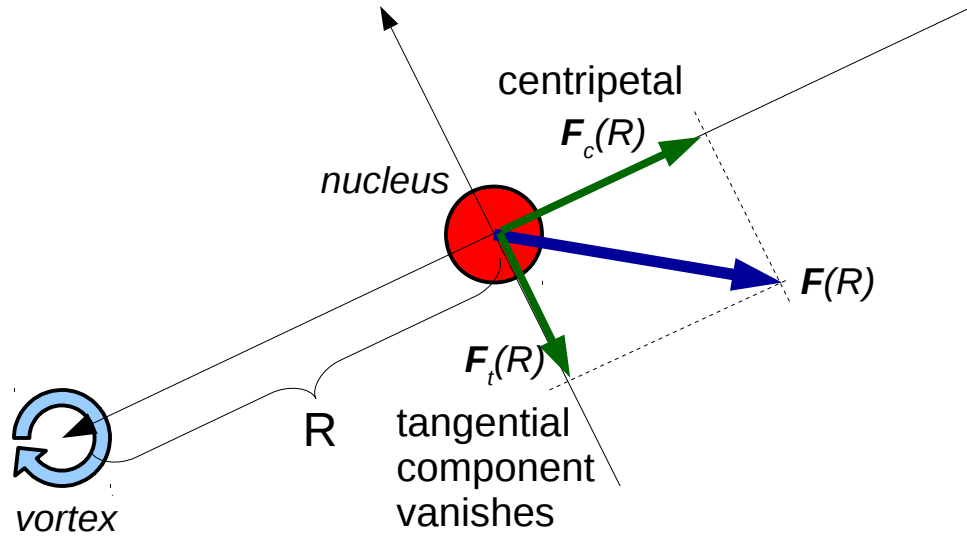
← ● or ● → ?



Movie 2

Movie 3

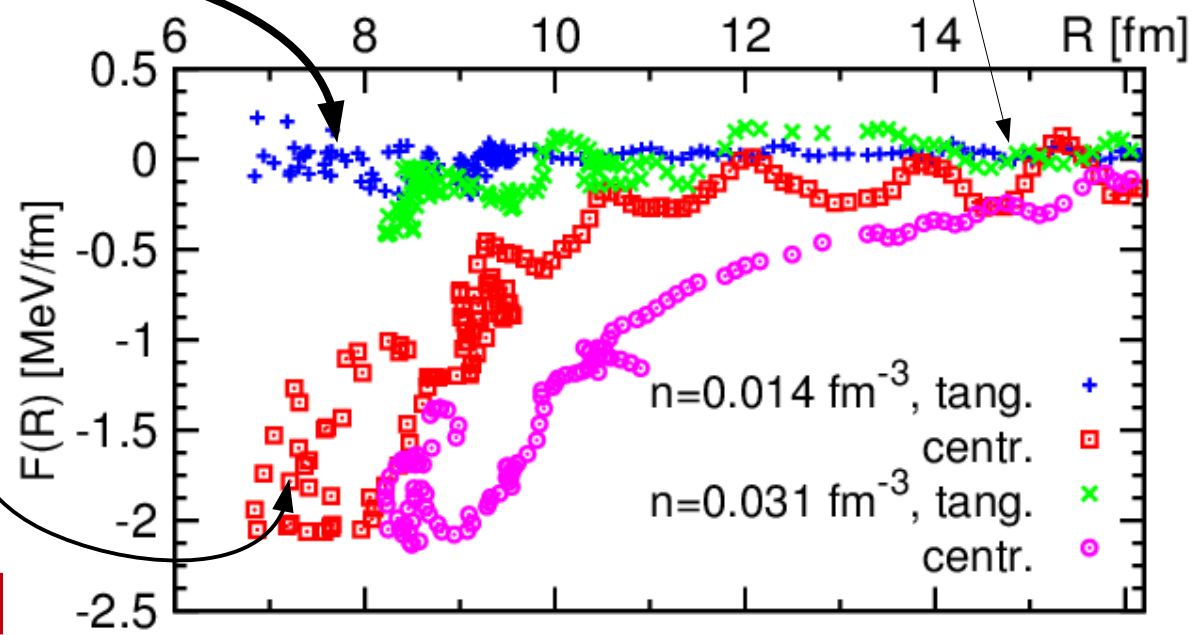
# Force decomposition



Range:  
10-15 fm

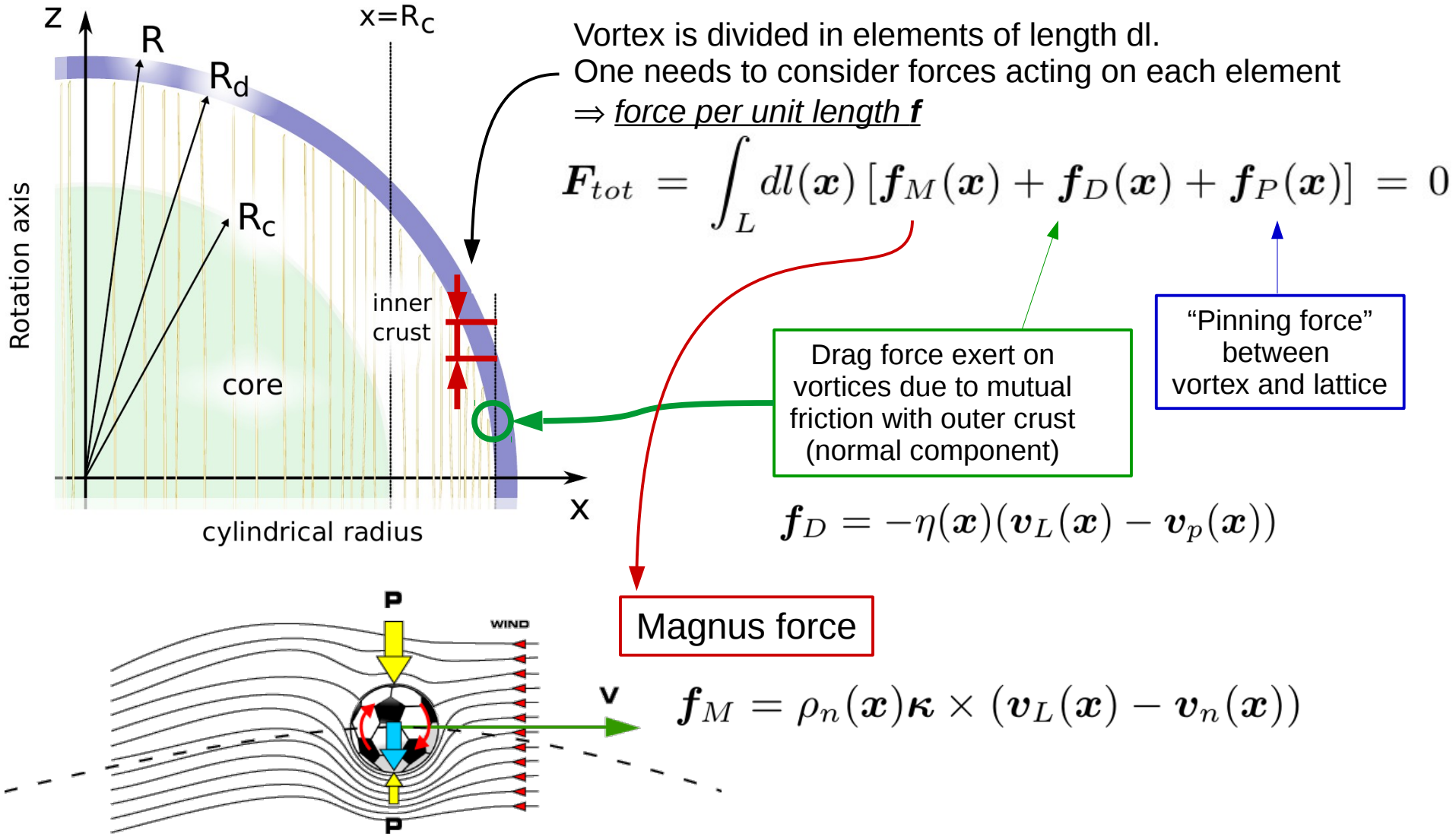
Force is  
central

At close separation  
the force is not a function  
of distance R only



# Pinning force – what is needed?

Example of hydrodynamical description (M. Antonelli, P. Pizzochero, [arXiv:1603.02838](https://arxiv.org/abs/1603.02838))



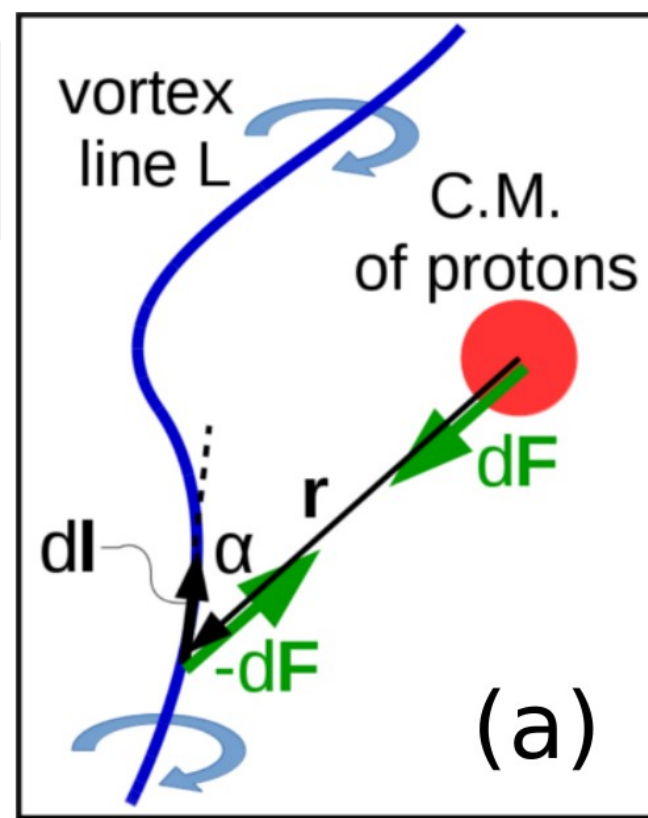


# Force per unit length

Inspired by the vortex filament model  
(describes vortex-vortex interaction)

$$d\mathbf{F} = f(r) \sin \alpha \hat{\mathbf{r}} dl \quad \hat{\mathbf{r}} = \mathbf{r}/r$$

$$\mathbf{F} = \int_L d\mathbf{F} \quad \leftarrow \text{Total force acting on nucleus}$$



# Force per unit length

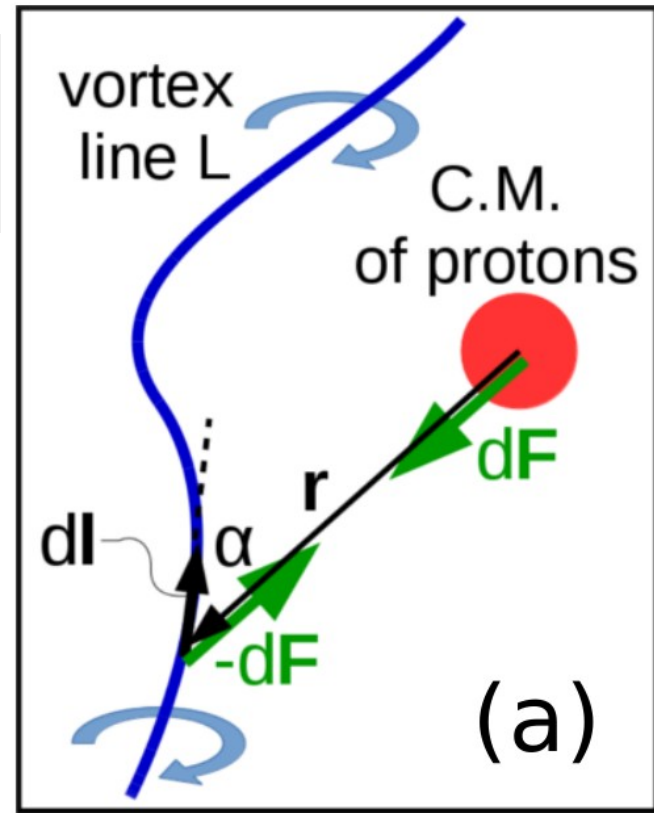
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$$d\mathbf{F} = f(r) \sin \alpha \hat{\mathbf{r}} dl \quad \hat{\mathbf{r}} = \mathbf{r} / r$$

$$\mathbf{F} = \int_L d\mathbf{F} \quad \leftarrow \text{Total force acting on nucleus}$$

Ansatz for f(r) - Pade approximant

$$f(r) = \frac{\sum_{k=0}^n a_k r^k}{1 + \sum_{k=1}^{n+3} b_k r^k}$$



(a)

We minimize chi2 with respect of {a<sub>k</sub>, b<sub>k</sub>}

$$\chi_w^2 = \sum_{i=1}^N w(|\mathbf{F}_i|) \left( \mathbf{F}_i - \mathbf{F}_i^{(f)}(\{a_k, b_k\}) \right)^2$$

“frames of movie”

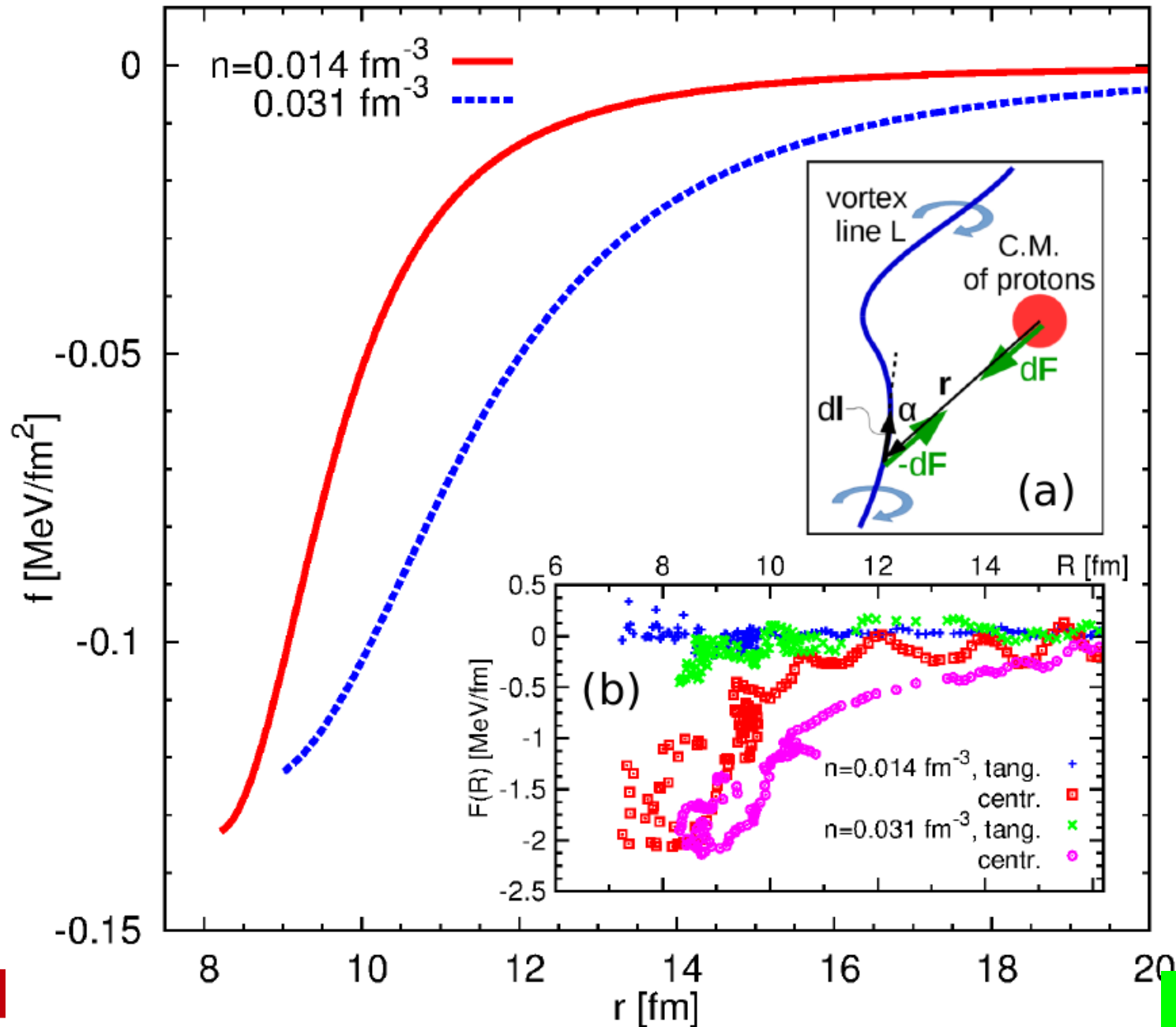
Measured force

Predicted for given set {a<sub>k</sub>, b<sub>k</sub>}

$$\mathbf{F}_i^{(f)} = \int_{L_i} f(r; \{a_k, b_k\}) \sin \alpha \mathbf{e}_r dl$$

$f(r \rightarrow \infty) \propto 1/r^3$   
from irrotational and incompressible hydrodynamics

# Force per unit length



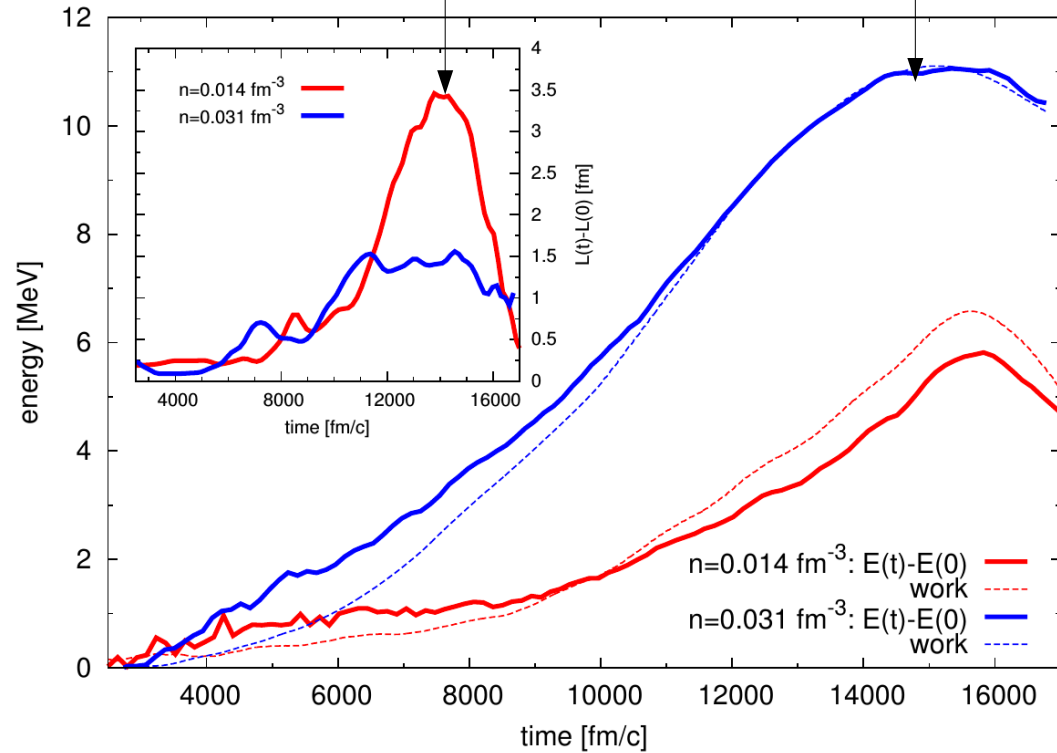
# Vortex tension

**Tension:** energy needed to increase vortex length by unit



Change of system energy

Vortex length change

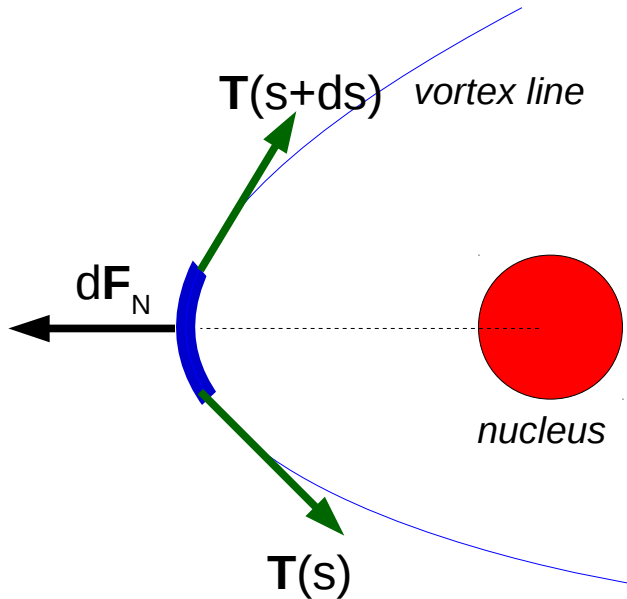


$$T \lesssim 1.4 \text{ MeV/fm} \quad \text{and} \quad 7.3 \text{ MeV/fm}$$

$$n=0.014 \text{ fm}^{-3} \qquad \qquad n=0.031 \text{ fm}^{-3}$$

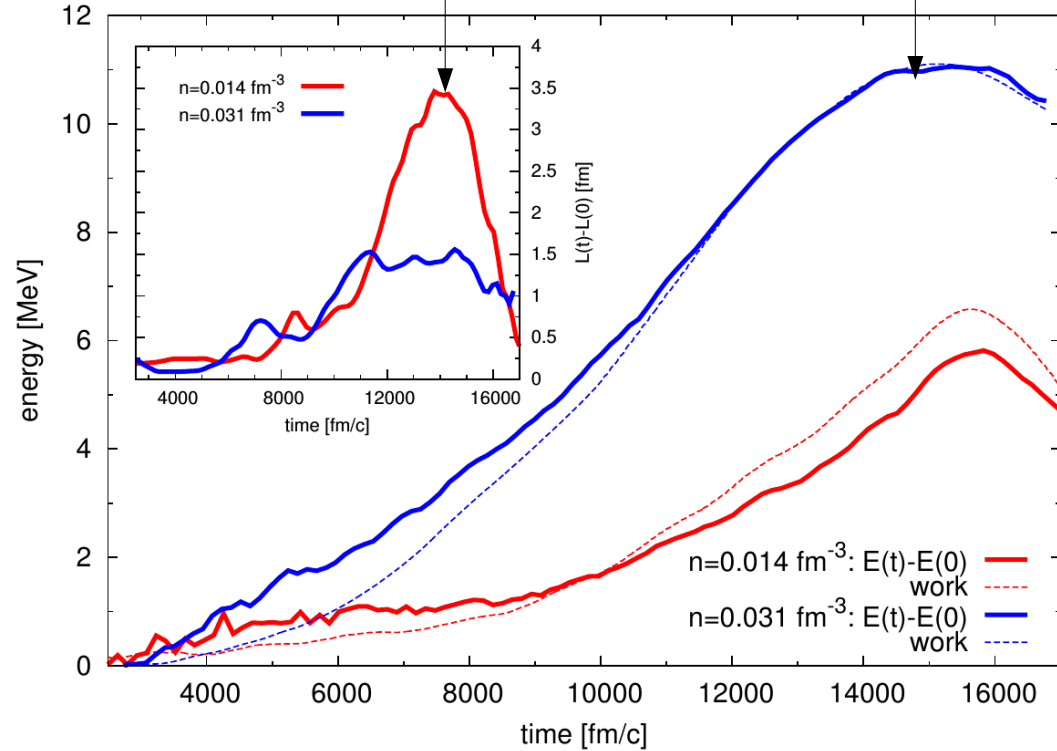
# Vortex tension

**Tension:** energy needed to increase vortex length by unit



Change of system energy

Vortex length change



$$dF_M = dF_N + T(s + ds) - T(s)$$

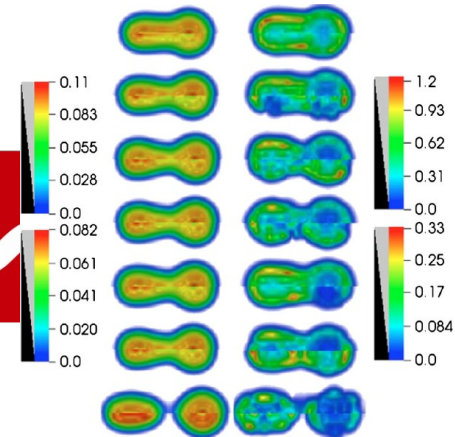
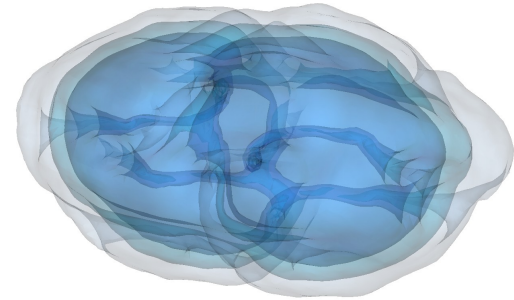
By inverting equations of motion for vortex elements we can gain detailed information about tension

# CONCLUSIONS:

- Using a **qualitatively new approach**, we follow the dynamics as superfluid **vortices** move in response to the presence of “**nuclei**”
- Nuclei repel vortices** in the neutron star crust, leading thus to **interstitial** vortex **pinning**
- The approach can provide **detailed** insight into vortex-nucleus interaction
- Time-dependent simulations can be used to extract **various** quantities.

- Dynamics in ultracold atoms
  - vortex dynamics* (*Phys. Rev. Lett.* 112, 025301 (2014))
  - quantum turbulence* (*Phys. Rev. A* 91, 031602(R) (2015))
  - shock waves* (*Phys. Rev. Lett.* 108, 150401 (2012))

- Dynamics of nuclear systems
  - fission* (*Phys. Rev. Lett.* 116, 122504 (2016))
  - relativistic coulomb excitation* (*Phys. Rev. Lett.* 114, 012701 (2015))
  - isovector giant dipole resonance* (*Phys. Rev. C* 84, 051309(R) (2011))

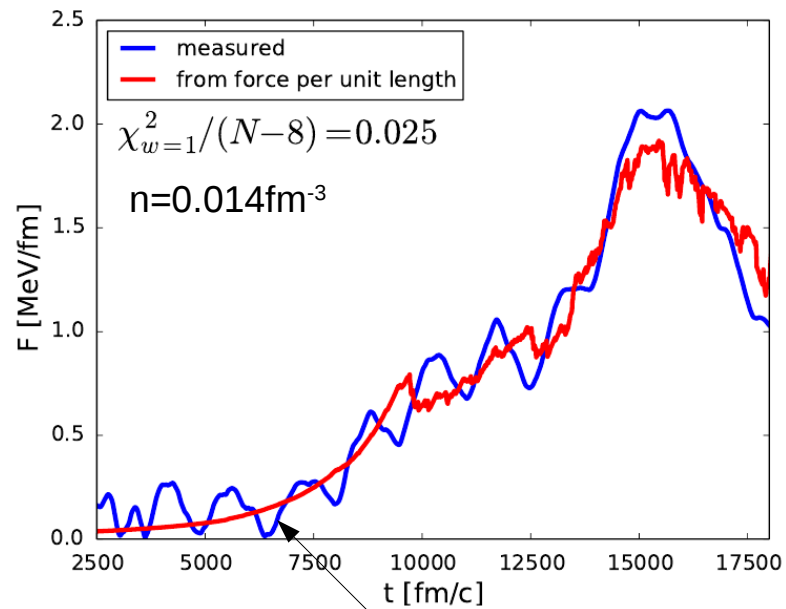
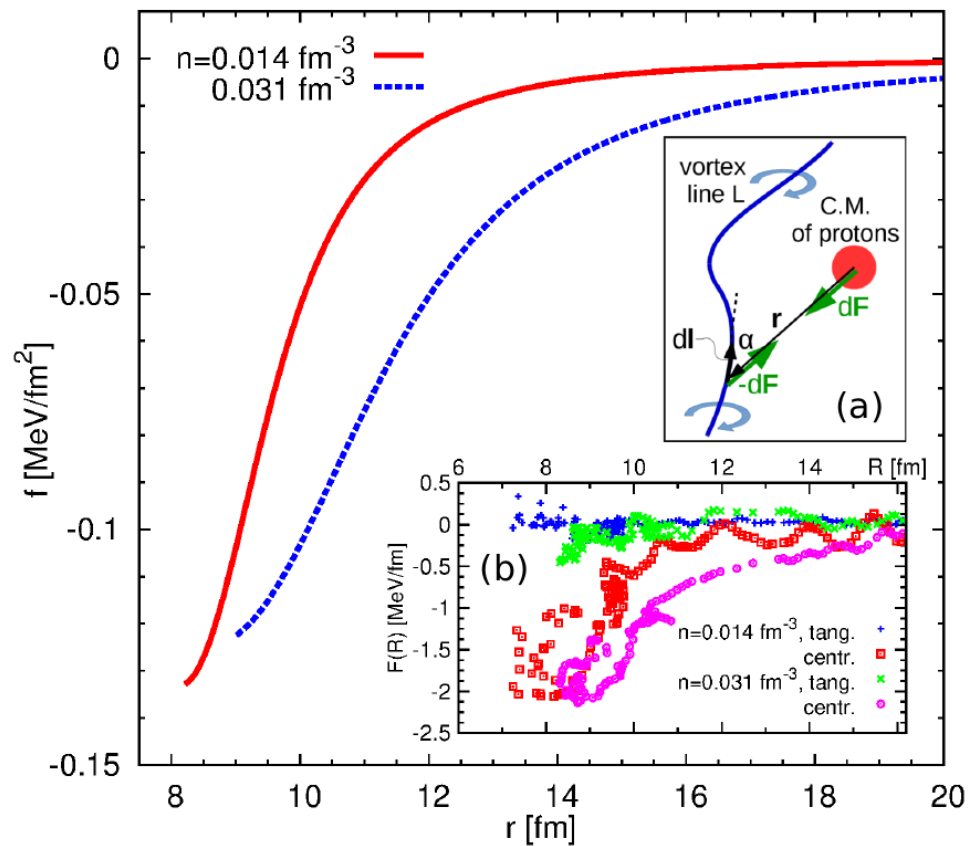


**Thank you**



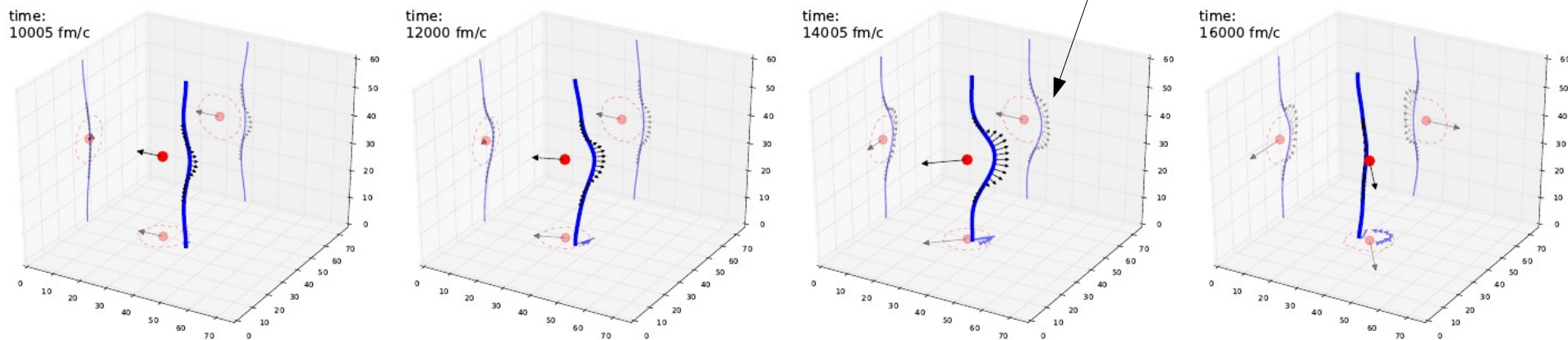


# Force per unit length

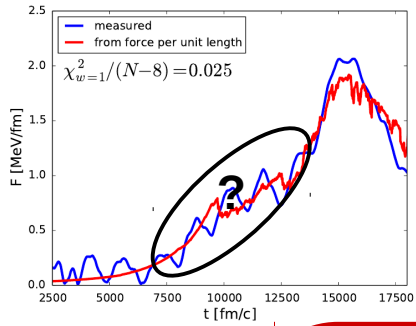


Fluctuations?

Only part close to nucleus surface contributes...



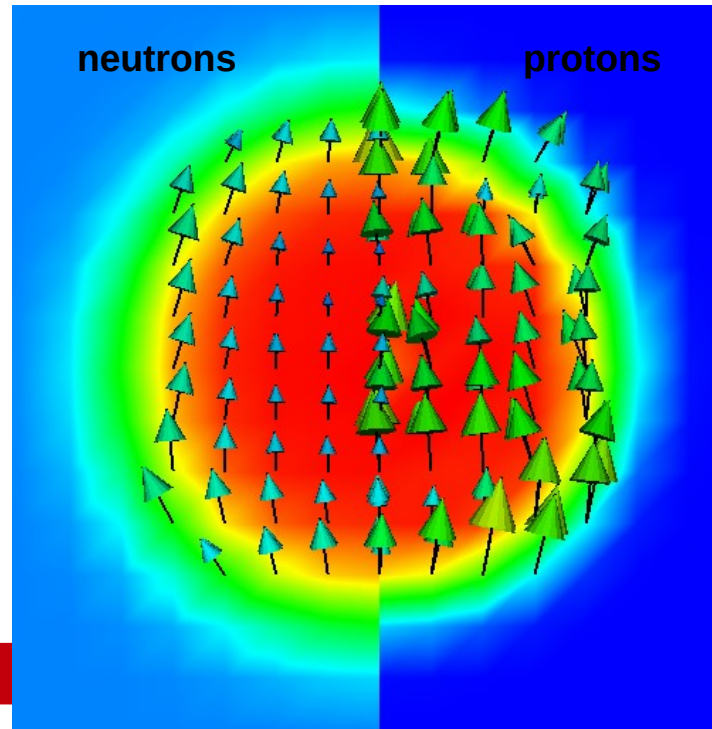
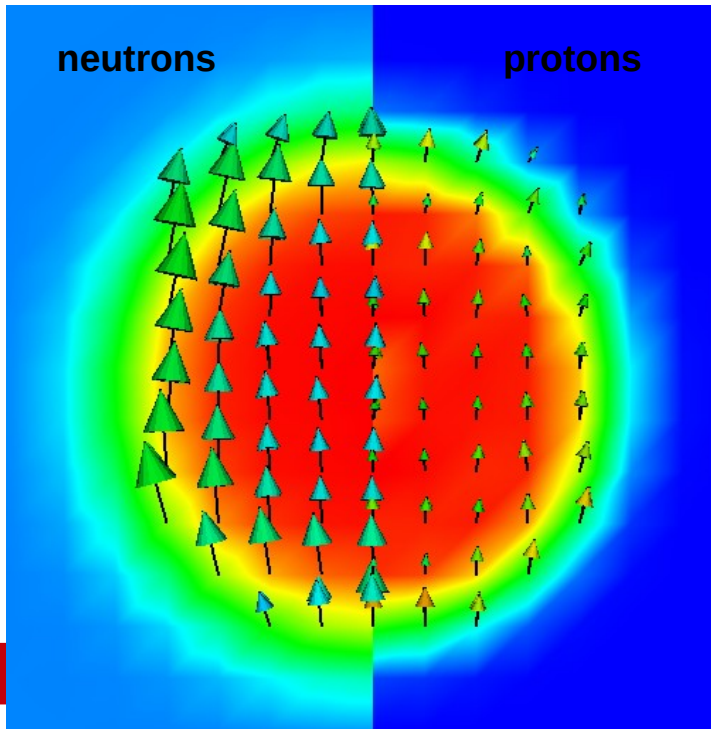
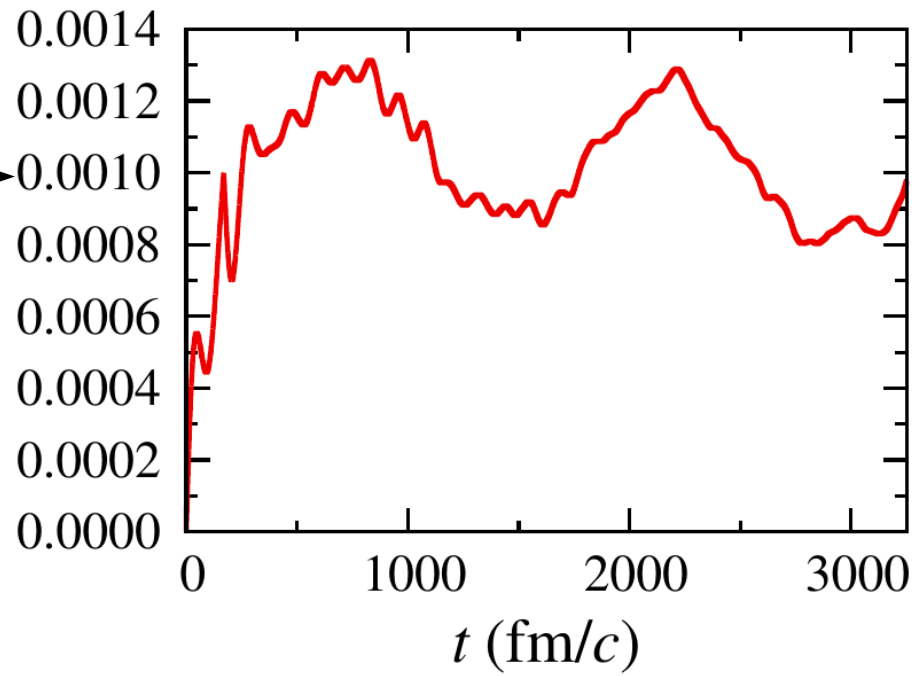
# Force fluctuations



We accelerated nucleus to  $v_0$  (by dragging protons) and let it go...

We induce giant dipole oscillations

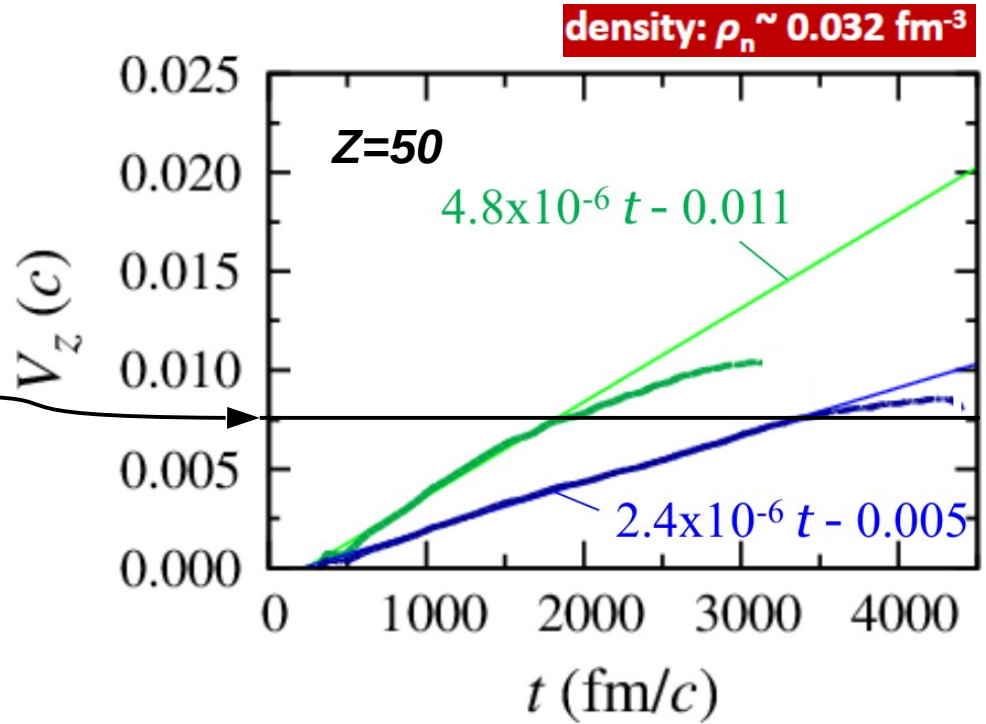
$v(c)$



$$F = M^* a$$

$$M^* \simeq 218 m_N$$

“critical” velocity



# Vortex model - idea

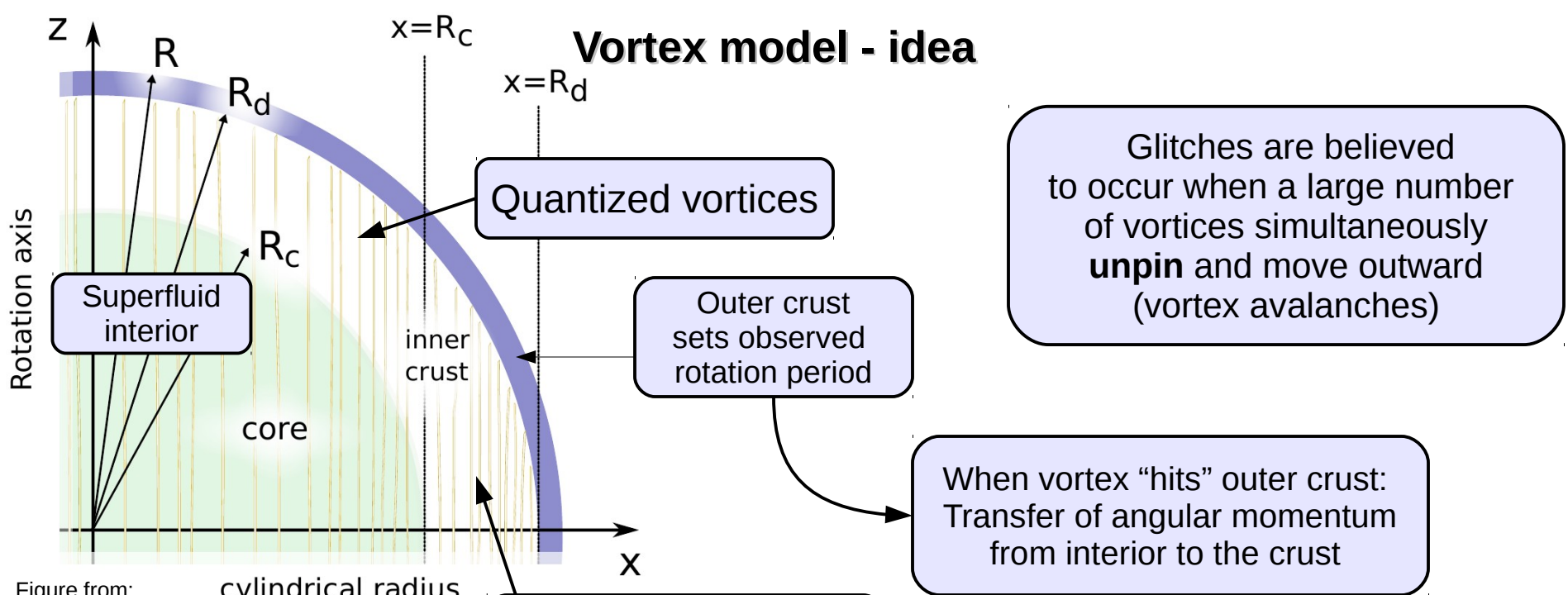


Figure from: M. Antonelli, P. Pizzochero, arXiv:1603.02838

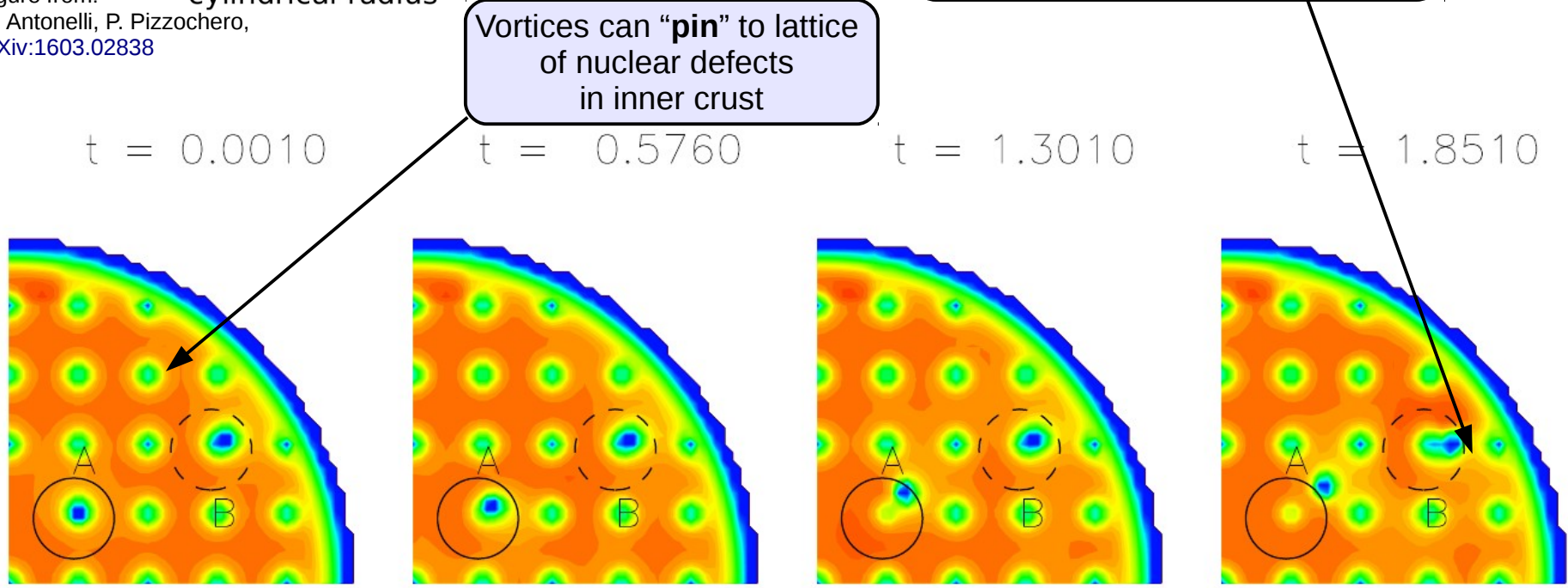


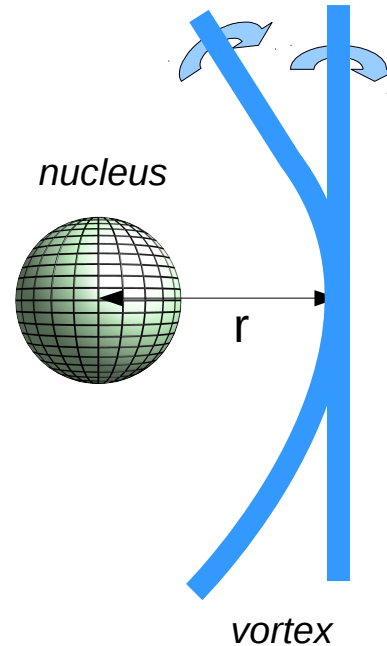
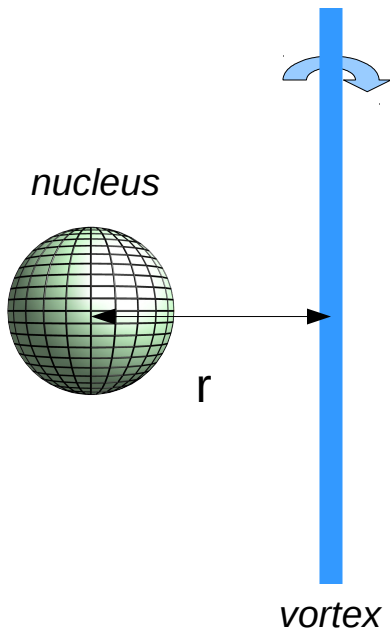
Figure from: L. Warszawski, A. Melatos, N. G. Berloff, Phys. Rev. B 85, 104503 (2012)

# Force per unit length

From pinning energy the average force can be deduced:  $F_{\text{pin}} \approx -\frac{\Delta E}{\Delta r}$

One should consider energy change as a function of separation  $r$ ...

... as well as function of deformation



Note:  
Vortex impurity  
is not a function  
of distance  $r$  only!!!

Very problematic within static calculations...

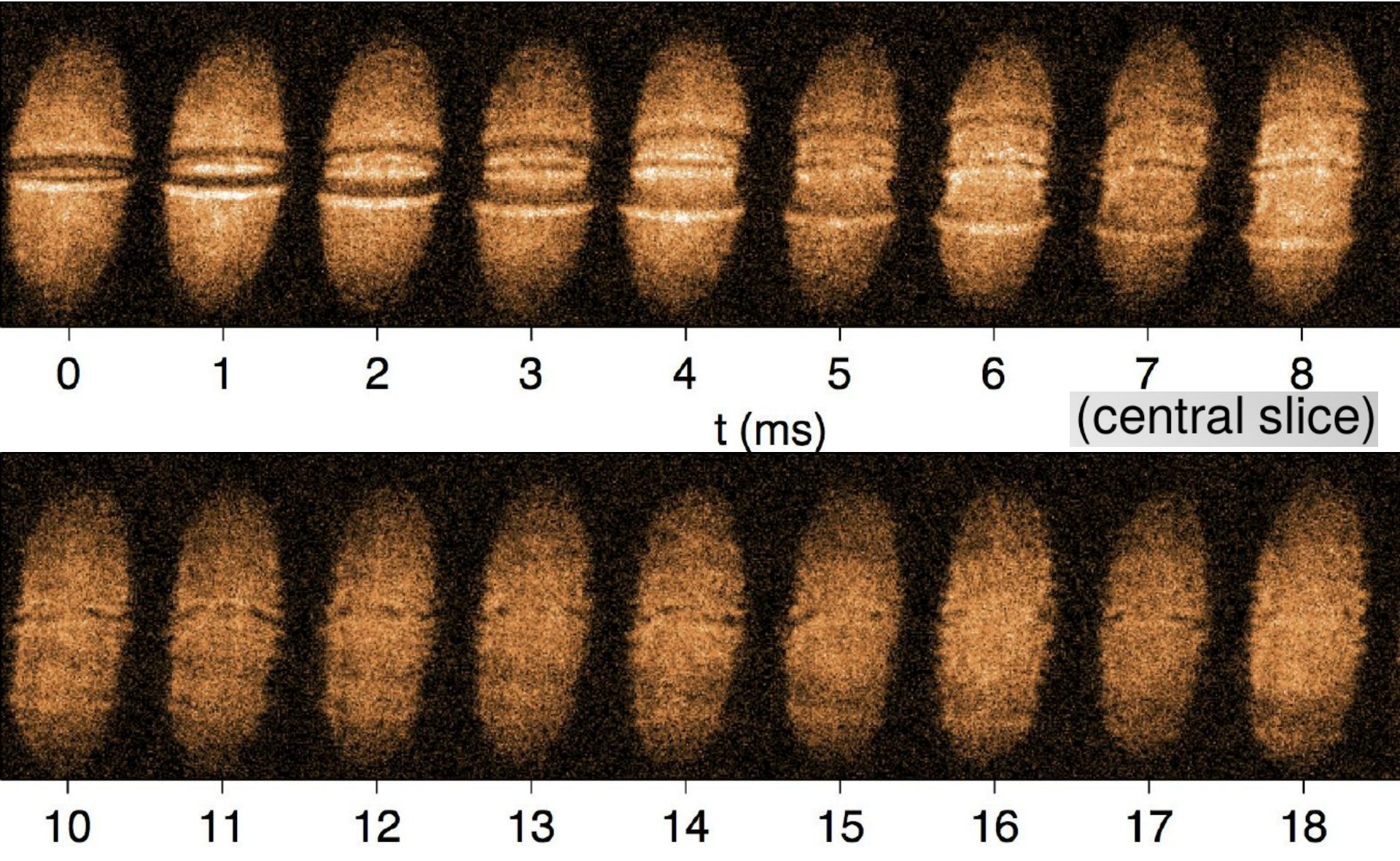


# Experimental results – Cascade of Solitary Waves

Figures taken from: M. Zwierlein talk, ([http://en.sif.it/activities/fermi\\_school/mmxiv](http://en.sif.it/activities/fermi_school/mmxiv))

School of Physics E. Fermi – Quantum Matter at Ultralow Temperatures Varenna, July 9th , 2014

See also: Mark J.H. Ku, et al., Phys. Rev. Lett. 116, 045304 (2016)

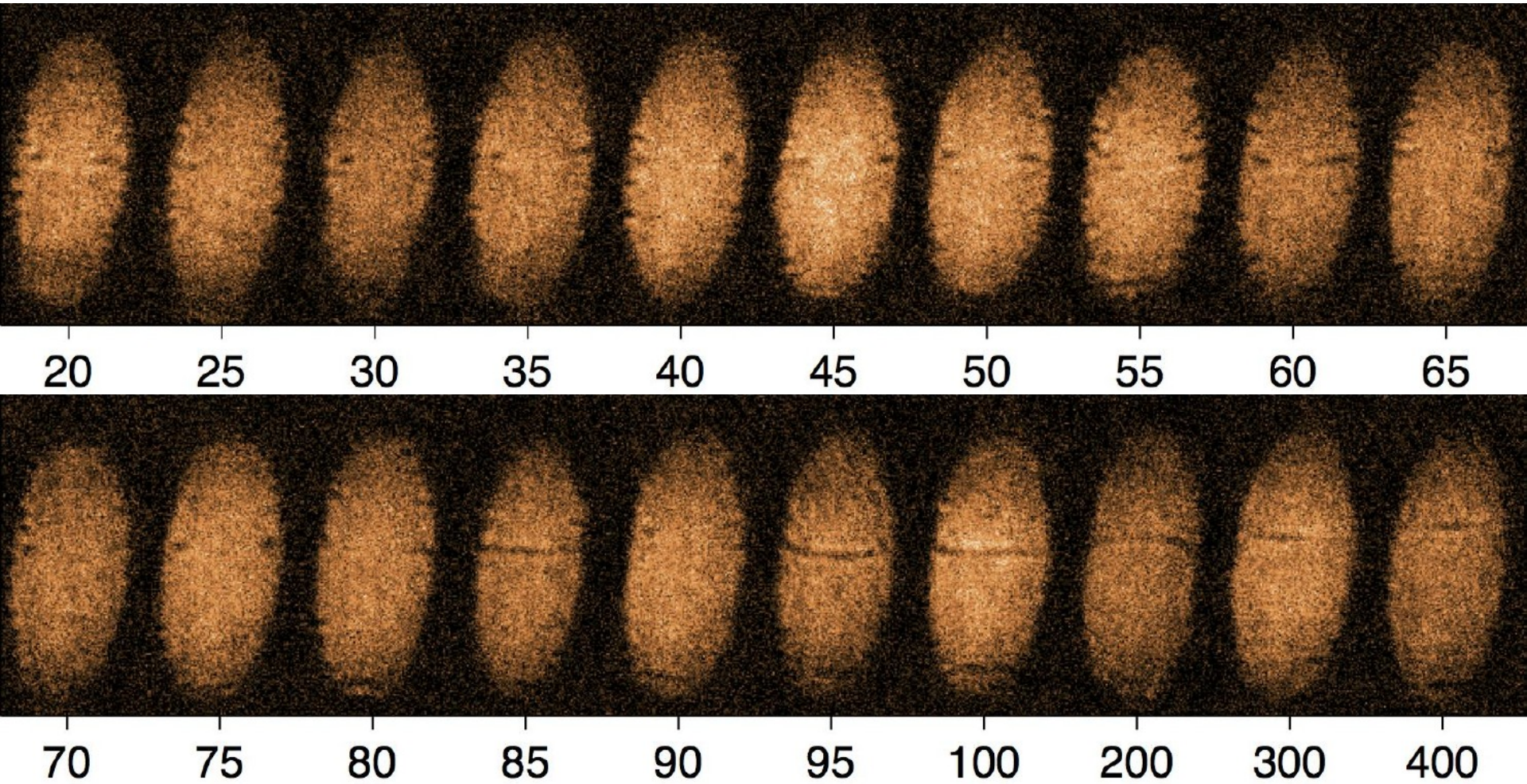




# Experimental results – Cascade of Solitary Waves

Figures taken from: M. Zwierlein talk, ([http://en.sif.it/activities/fermi\\_school/mmxiv](http://en.sif.it/activities/fermi_school/mmxiv))

School of Physics E. Fermi – Quantum Matter at Ultralow Temperatures Varenna, July 9th , 2014

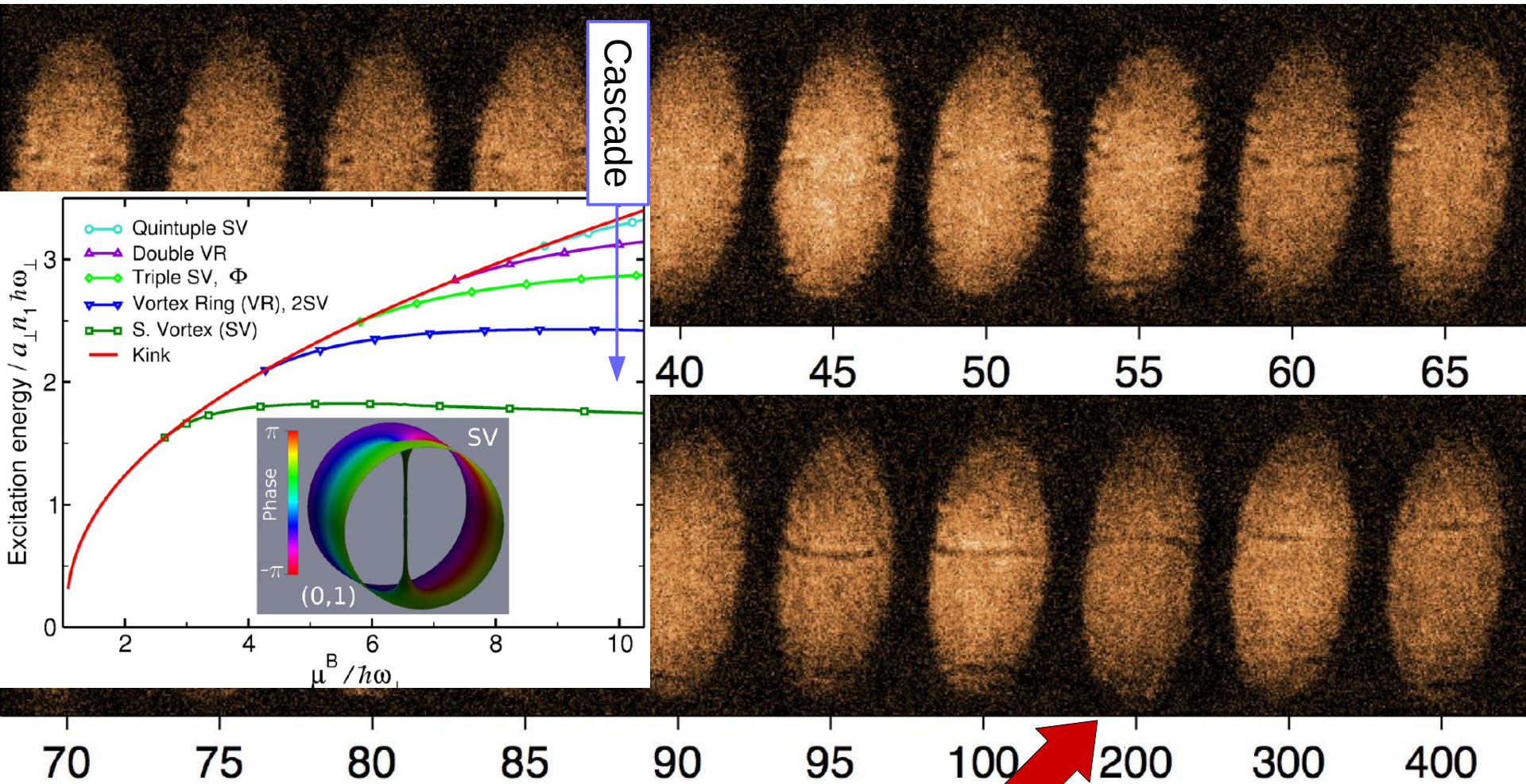




# Experimental results – Cascade of Solitary Waves

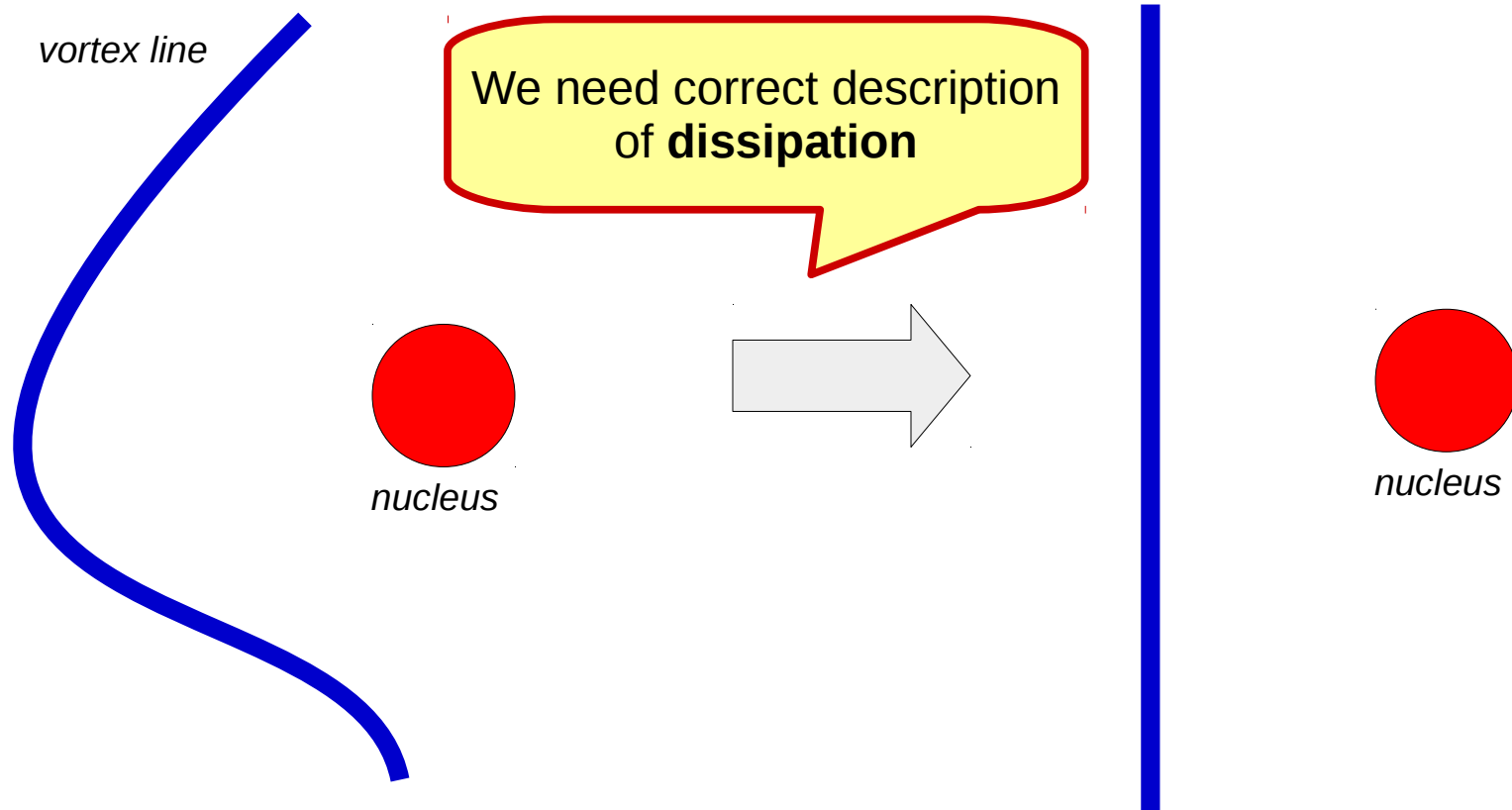
Figures taken from: M. Zwierlein talk, ([http://en.sif.it/activities/fermi\\_school/mmxiv](http://en.sif.it/activities/fermi_school/mmxiv))

School of Physics E. Fermi – Quantum Matter at Ultralow Temperatures Varenna, July 9th , 2014



Challenge for theory to describe all stages of the cascade!  
Dissipation is crucial...

# Dissipation



adiabatic approximation  $\Rightarrow$  Memory effects are usually neglected

Result: dissipation effects are not correctly taken into account except for one-body dissipation

*...but our system is superfluid...*

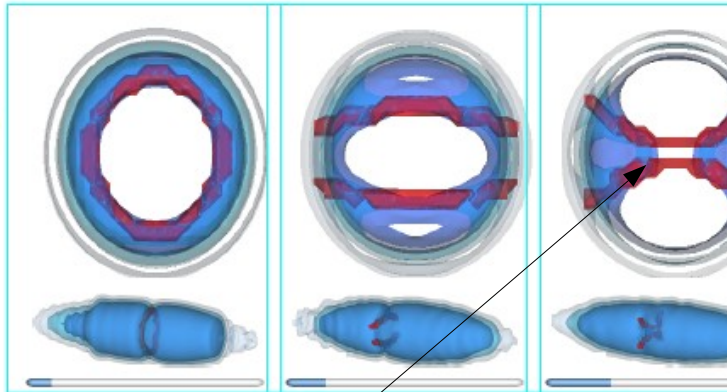
*... and one-body dissipation sufficient to describe dynamics of vortex in ultra-cold atoms...*

# What do fully 3D simulations reveal?

Phys. Rev. A 91, 031602 (2015)

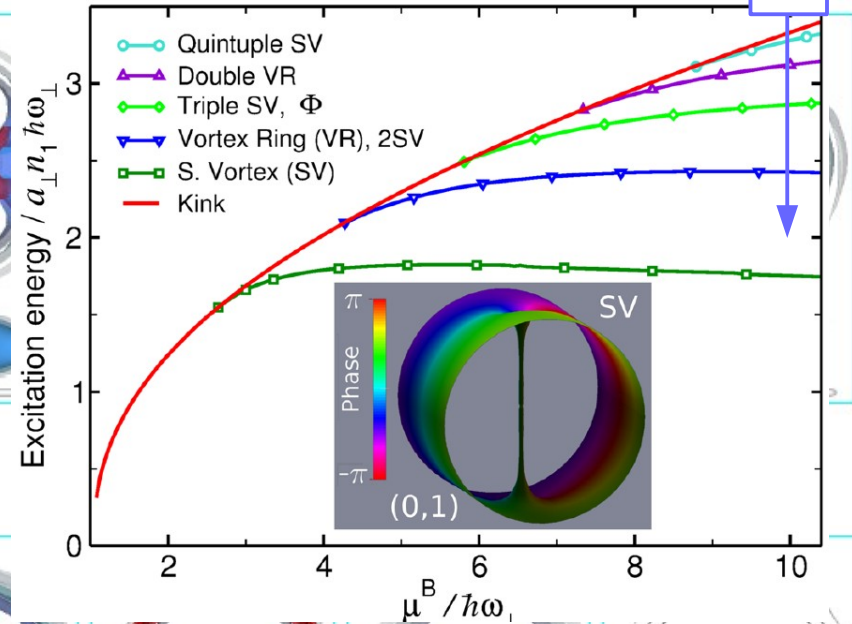
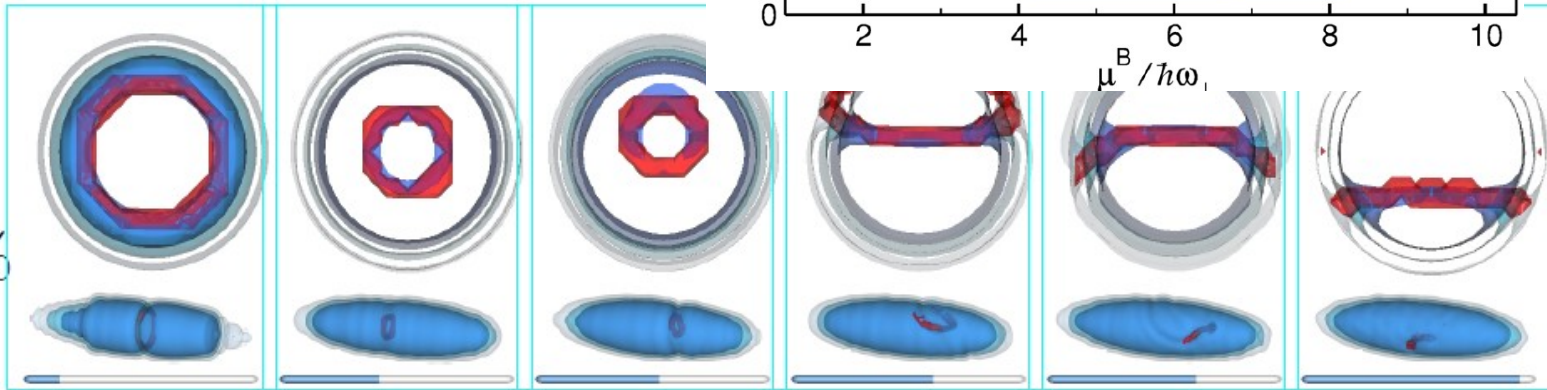
Cascade

$\delta \approx 10\%$   
 $\frac{C}{C_0} \approx 3\%$



**Crossing and reconnection!**

$\delta \approx 0.5\%$   
 $\frac{C}{C_0} \approx 0.15\%$





# Density Functional Theory – Idea

It can be shown that **instead of wave function** one may use a **density** distribution: *contains vastly more information than the one needed*

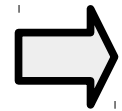
$$\rho_1(\mathbf{x}_1, \mathbf{x}'_1) = N \int \Psi(\mathbf{x}_1, \dots, \mathbf{x}_N) \Psi(\mathbf{x}'_1, \dots, \mathbf{x}_N)^* d\mathbf{x}_2 \cdots d\mathbf{x}_N$$

*reduced object - sufficient to extract one body observables*

**Theorem (Hohenberg & Kohn):**

**The energy of the nondegenerate ground state of the Fermi system is uniquely determined by its density distribution.**

$$\Psi \xleftrightarrow{1-1} \rho$$



$$E[\rho] = \langle \Psi[\rho] | \hat{H} | \Psi[\rho] \rangle$$

**The energy is a functional of the density**

In general:

$$\langle O \rangle = \langle \Psi[\rho] | O | \Psi[\rho] \rangle = O[\rho]$$

“Universal” part

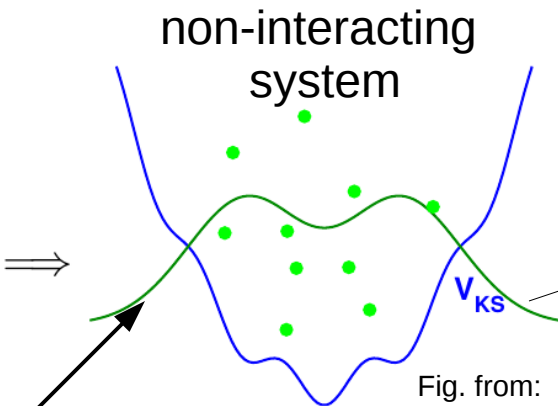
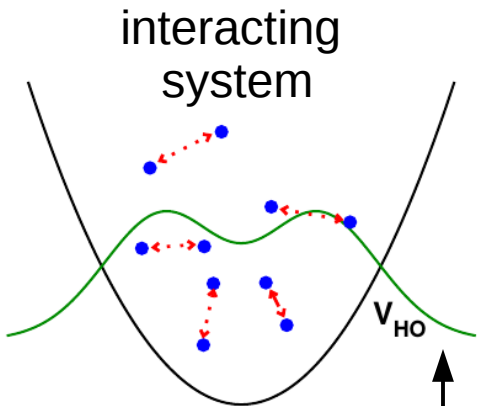
External field

$$E[\rho] = F[\rho] + \int d\mathbf{x} v_{\text{ext}}(\mathbf{x})\rho(\mathbf{x})$$

Formally rigorous way of approaching any **interacting problem** by **mapping** it exactly to a much easier-to-solve **noninteracting system**.

Kohn-Sham method:

easy, if Energy Density Functional (EDF) is known...



$$v_{KS} = \frac{\delta F}{\delta \rho} + v_{ext}$$

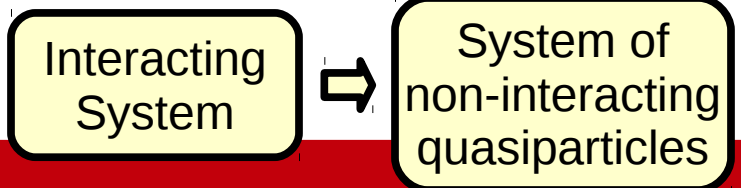
Both systems described by the same density

Fig. from: Prog.Part.Nucl.Phys.64:120-168,2010

$$[-\nabla^2/2m + v_{KS}(\mathbf{x})]\phi_i(\mathbf{x}) = \epsilon_i\phi_i(\mathbf{x})$$

$$\rho(\mathbf{x}) = \sum_i n_i |\phi_i(\mathbf{x})|^2 \quad n_i = \theta(\epsilon_F - \epsilon_i)$$

More general:



Note: There are easy and difficult observables in DFT. In general: **easy** observables are **one-body** observables. They are easily extracted and reliable.



# Energy density functional

How to derive EDF?

$$E[\rho] = T[\rho] + V[\rho] + v_{ext}\rho$$

We build models of EDF (typically here we introduce approximation)

$$E[\rho_i, \tau_i, \dots] = \int d\mathbf{x} \mathcal{E}(\rho_i(\mathbf{x}), \tau_i(\mathbf{x}), \dots)$$

Example:

Local Density Approximation  
(only dependence on diagonal parts of densities  $\rho(\mathbf{x}) \equiv \rho(\mathbf{x}, \mathbf{x})$ )

$$[-\nabla^2/2m + v_{KS}(\mathbf{x})]\phi_i(\mathbf{x}) = \varepsilon_i\phi_i(\mathbf{x})$$

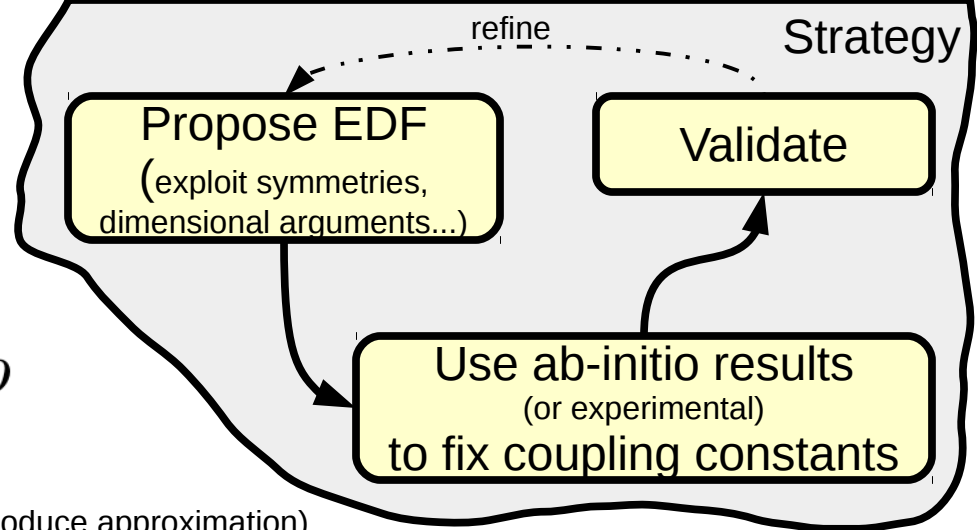
$$\rho(\mathbf{x}) = \sum_{i=1}^A |\phi_i(\mathbf{x})|^2$$

$$\tau(\mathbf{x}) = \sum_{i=1}^A |\nabla\phi_i(\mathbf{x})|^2$$

From orbitals  
we build other densities  
(correlated with  $\rho$ )

$$T_s[\rho] = -\frac{1}{2m} \sum_{i=1}^A \int d\mathbf{x} \phi_i^\dagger(\mathbf{x}) \nabla^2 \phi_i(\mathbf{x})$$

$$= \int d\mathbf{x} \frac{1}{2m} \tau(\mathbf{x})$$



# Ultracold atoms are superfluid!

Normal system

$$\begin{cases} [-\nabla^2/2m + v_{KS}(\mathbf{x})]\phi_i(\mathbf{x}) = \varepsilon_i\phi_i(\mathbf{x}) \\ v_{KS} = \frac{\delta F}{\delta \rho} + v_{ext} \end{cases} \quad \text{pairing (anomalous) density}$$

↓

$$F[\rho, \tau, \dots] \rightarrow F[\rho, \tau, \nu, \dots]$$

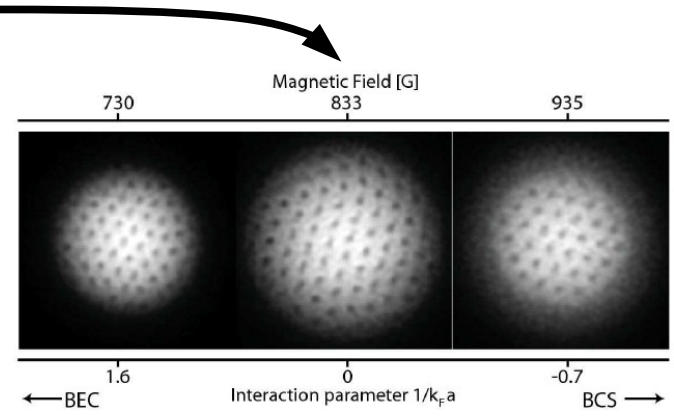


FIG. 36 Vortex lattice in a rotating gas of  ${}^6\text{Li}$  precisely at the Feshbach resonance and on the BEC and BCS side. Reprinted with permission from Zwierlein *et al.* (2005).

$$\begin{cases} [h(\mathbf{r}) - \mu]u_k(\mathbf{r}) + \Delta(\mathbf{r})v_k(\mathbf{r}) = E_k u_k(\mathbf{r}), \\ \Delta^*(\mathbf{r})u_k(\mathbf{r}) - [h(\mathbf{r}) - \mu]v_k(\mathbf{r}) = E_k v_k(\mathbf{r}), \end{cases}$$

$$h(\mathbf{r}) = -\nabla^2/2m + v_{KS}(\mathbf{x}) \quad \text{BCS-like}$$

$$\Delta = -\frac{\delta F}{\delta \nu^*}$$

$$n(\mathbf{r}) = 2 \sum_k |v_k(\mathbf{r})|^2, \quad \tau(\mathbf{r}) = 2 \sum_k |\nabla v_k(\mathbf{r})|^2,$$

$$\nu(\mathbf{r}) = \sum_k v_k^*(\mathbf{r})u_k(\mathbf{r}),$$

Note: diagonal part of pairing density is divergent  
Regularization required!

We use prescription given in:  
Bulgac, Yu, Phys. Rev. Lett. 88 (2002) 042504  
Bulgac, Phys. Rev. C65 (2002) 051305

# Extension to time-dependent case

$$i\hbar \frac{\partial}{\partial t} \psi = \hat{H} \psi, \quad \psi_0 = \psi(t_0) \quad \text{and} \quad \frac{\partial \rho}{\partial t} + \nabla \cdot \vec{j} = 0$$

Runge & Gross theorem

$$\left. \begin{array}{l} \rho(\vec{r}, t) \\ \psi(\dots, t_0) \end{array} \right\} \leftrightarrow e^{i\alpha(t)} \psi(\dots, t)$$

Up to arbitrary function  $\alpha(t)$

$$\vec{j}(\mathbf{r}) = -\frac{iN}{2} \int d\mathbf{r}_2 \cdots d\mathbf{r}_N \Psi(\mathbf{r}, \dots, \mathbf{r}_N)^* \nabla \Psi(\mathbf{r}, \dots, \mathbf{r}_N) - \Psi(\mathbf{r}_1, \dots, \mathbf{r}_N) \nabla \Psi(\mathbf{r}_1, \dots, \mathbf{r}_N)^*$$

... and consequently density functional exists...  
 NOTE: in general this functional:  
 - is initial state dependent  
 - at time t depends on densities in previous times (memory effect)

van Leeuwen's Theorem:

Very little is known about the memory terms, but in principle it can be long ranged (see eg. Dobson, Brunner, Gross, Phys. Rev. Lett. 79 (1997) 1905)

Time evolution of interacting system

Memory effects are usually neglected = adiabatic approximation  
 [Result: dissipation effects are not correctly taken into account except for one-body dissipation]



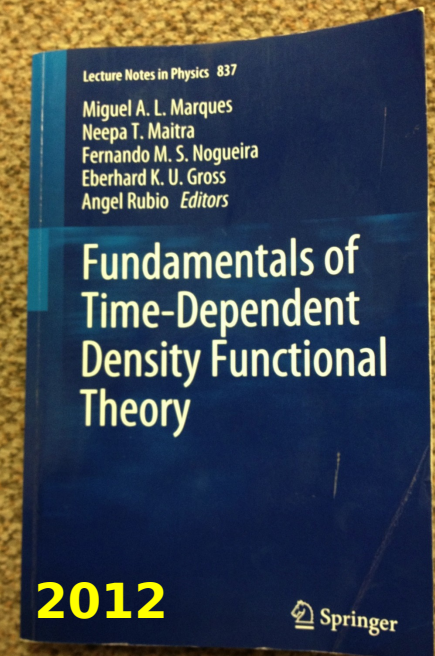
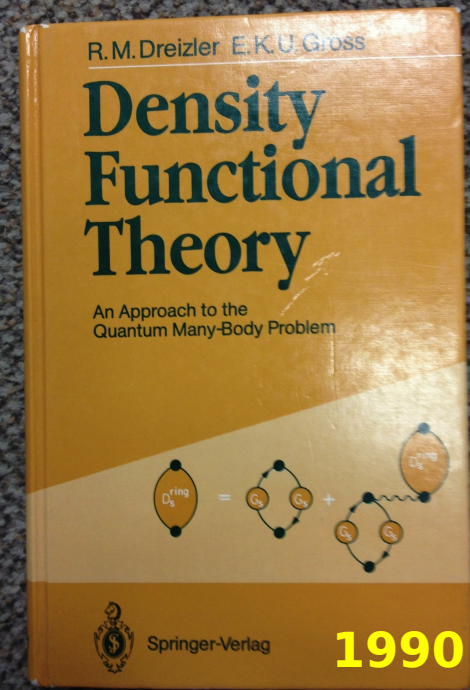
Time evolution of non-interacting system

*...but our system is superfluid...*

$$i\hbar \frac{\partial \phi_i(\mathbf{x}, t)}{\partial t} = \left[ -\frac{\nabla^2}{2m} + v_{KS}(\mathbf{x}, t) \right] \phi_i(\mathbf{x}, t)$$

# DFT: workhorse for electronic structure simulations

- ◆ The Hohenberg-Kohn theorem assures that the theory can reproduce exactly the ground state energy if the “exact” Energy Density Functional (EDF) is provided
- ◆ Often called as *ab initio* method
- ◆ Extension to Time-Dependent DFT is straightforward
- ◆ Can be extended to superfluid systems... (numerical cost increases dramatically)
- ◆ **Very successful** – DFT industry (commercial codes for quantum chemistry and solid-state physics)



# Solving time-dependent problem...

The real-time dynamics is given by equations, which are formally equivalent to the Time-Dependent HFB (TDHFB) or Time-Dependent Bogolubov-de Gennes (TDBdG) equations:

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_{n,a}(\mathbf{r}, t) \\ u_{n,b}(\mathbf{r}, t) \\ v_{n,a}(\mathbf{r}, t) \\ v_{n,b}(\mathbf{r}, t) \end{pmatrix} = \begin{pmatrix} h_a(\mathbf{r}, t) & 0 & 0 & \Delta(\mathbf{r}, t) \\ 0 & h_b(\mathbf{r}, t) & -\Delta(\mathbf{r}, t) & 0 \\ 0 & -\Delta^*(\mathbf{r}, t) & -h_a^*(\mathbf{r}, t) & 0 \\ \Delta^*(\mathbf{r}, t) & 0 & 0 & -h_b^*(\mathbf{r}, t) \end{pmatrix} \begin{pmatrix} u_{n,a}(\mathbf{r}, t) \\ u_{n,b}(\mathbf{r}, t) \\ v_{n,a}(\mathbf{r}, t) \\ v_{n,b}(\mathbf{r}, t) \end{pmatrix}$$

The mean field potentials  $h_i(\mathbf{r}, t)$  are derived from the EDF as functional derivative  $h_i = \frac{\delta E}{\delta n_i}$  and they explicitly depends on local densities  $n(\mathbf{r})$ ,  $\tau(\mathbf{r})$ , etc. The set of 4-component "wave functions" is in turn related to the densities and the pairing field  $\Delta(\mathbf{r})$

$$n_i(\mathbf{r}) = \sum_{E_n < E_c} |v_{n,i}(\mathbf{r})|^2,$$

$$\tau_i(\mathbf{r}) = \sum_{E_n < E_c} |\nabla v_{n,i}(\mathbf{r})|^2,$$

$$v(\mathbf{r}) = \sum_{E_n < E_c} u_{n,a}(\mathbf{r}) v_{n,b}^*(\mathbf{r}),$$

$$\Delta(\mathbf{r}) = -g_{\text{eff}}(\mathbf{r})v(\mathbf{r}).$$

**nonlinear  
coupled 3D  
Partial  
Differential  
Equations**

**Supercomputing**

We simulate fermionic systems consisting of  $10^3 - 10^4$  particles (**cold atoms**, neutron stars)

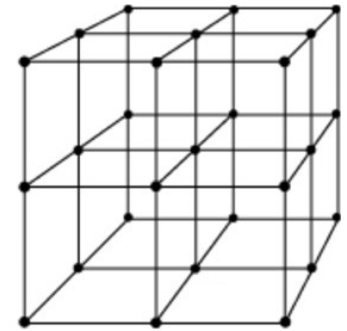
... also nuclear reactions (spin-orbit term required)



# Solving...

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_{n,a}(\mathbf{r}, t) \\ u_{n,b}(\mathbf{r}, t) \\ v_{n,a}(\mathbf{r}, t) \\ v_{n,b}(\mathbf{r}, t) \end{pmatrix} = \begin{pmatrix} h_a(\mathbf{r}, t) & 0 & 0 & \Delta(\mathbf{r}, t) \\ 0 & h_b(\mathbf{r}, t) & -\Delta(\mathbf{r}, t) & 0 \\ 0 & -\Delta^*(\mathbf{r}, t) & -h_a^*(\mathbf{r}, t) & 0 \\ \Delta^*(\mathbf{r}, t) & 0 & 0 & -h_b^*(\mathbf{r}, t) \end{pmatrix} \begin{pmatrix} u_{n,a}(\mathbf{r}, t) \\ u_{n,b}(\mathbf{r}, t) \\ v_{n,a}(\mathbf{r}, t) \\ v_{n,b}(\mathbf{r}, t) \end{pmatrix}$$

- ♦ The system is placed on a large 3D spatial lattice of size  $N_x \times N_y \times N_z$



- ★ Discrete Variable Representation (DVR) - solid framework (see for example: Bulgac, Forbes, Phys. Rev. C 87, 051301(R) (2013))
- ★ Errors are well controlled – exponential convergence
- ★ No symmetry restrictions
- ♦ Number of PDEs is of the order of the number of spatial lattice points
  - ★ Typically (without spin-orbit term):  $10^5 - 10^6$



# Solving...

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_{n,a}(\mathbf{r}, t) \\ u_{n,b}(\mathbf{r}, t) \\ v_{n,a}(\mathbf{r}, t) \\ v_{n,b}(\mathbf{r}, t) \end{pmatrix} = \begin{pmatrix} h_a(\mathbf{r}, t) & 0 & 0 & \Delta(\mathbf{r}, t) \\ 0 & h_b(\mathbf{r}, t) & -\Delta(\mathbf{r}, t) & 0 \\ 0 & -\Delta^*(\mathbf{r}, t) & -h_a^*(\mathbf{r}, t) & 0 \\ \Delta^*(\mathbf{r}, t) & 0 & 0 & -h_b^*(\mathbf{r}, t) \end{pmatrix} \begin{pmatrix} u_{n,a}(\mathbf{r}, t) \\ u_{n,b}(\mathbf{r}, t) \\ v_{n,a}(\mathbf{r}, t) \\ v_{n,b}(\mathbf{r}, t) \end{pmatrix}$$

- Derivatives are computed with FFT

- insures machine accuracy
- very fast

It sets scaling  
(N-number of lattice points)

$$O(\underbrace{N N}_{\text{Number of wave-functions}} \underbrace{\log N}_{\text{FFT}}) \Rightarrow \#O(N^2) \text{ for large lattice}$$

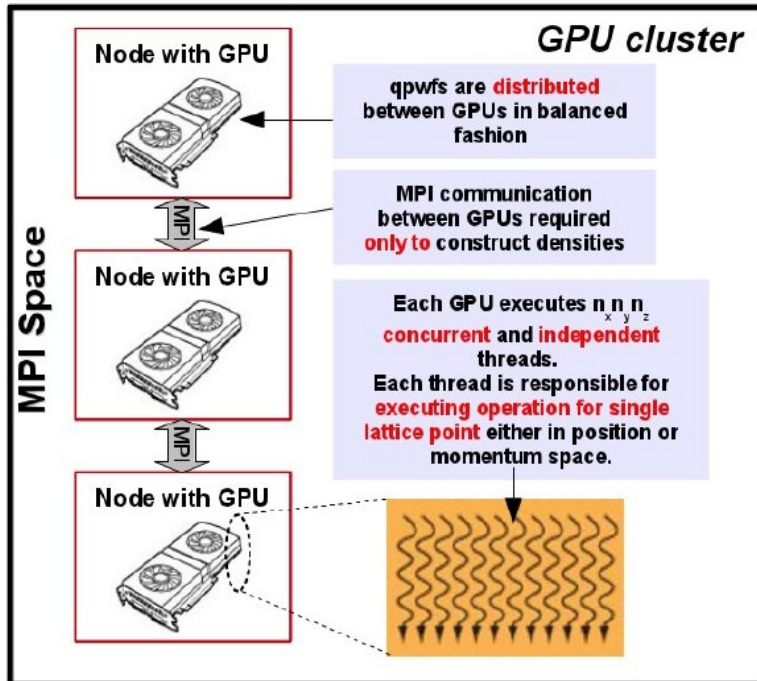
- Integration methods:

- Adams-Bashforth-Milne fifth order predictor-corrector-modifier integrator – very accurate but memory intensive
- Split-operator method that respects time-reversal invariance (third order) – very fast, but can work with simple EDF

If non local densities  
 $N^3!!!$   
(beyond our reach)

The spirit of SLDA is to exploit only local densities...

- ◆ Suitable for efficient parallelization (MPI)
- ◆ Excellent candidate for utilization multithreading computing units like GPUs



Lattice	# of GPUs	# of qpwfs	time per step [s]
24x24x96	64	24425	0.33
24x24x96	128	24425	0.17
24x24x96	256	24425	0.09
24x24x96	512	24425	0.06
32x32x128	256	57849	0.32
32x32x128	313	57849	0.25
48x48x128	512	129881	0.95
48x48x128	1024	129881	0.50
48x48x128	2048	129881	0.27
48x48x128	4096	129881	0.16
48x48x128	8192	129881	0.10

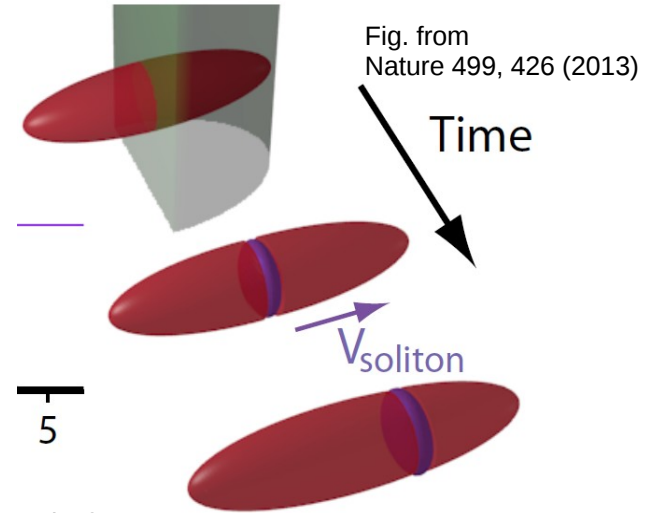
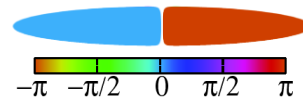
Lattice  $64^3$ , 137,062 (2-component) wave functions, ABM  
 CPU version running on  $16 \times 4096 = 65,536$  cores  
 GPU version running on 4096 GPUs

**15 times  
Speed-up!!!**

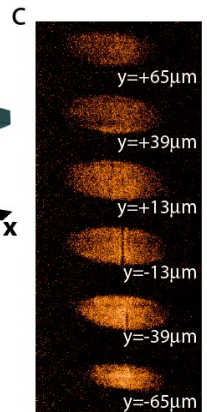
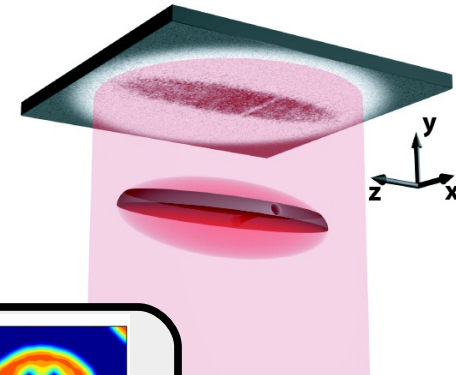
# Validation against dynamical properties of the system

Recent MIT experiments: Nature 499, 426 (2013), PRL 113, 065301(2014)

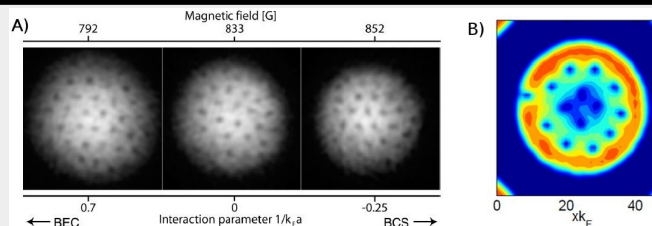
- ♦  $^6\text{Li}$  atoms near a Feshbach resonance ( $N \approx 10^6$ ) cooled in harmonic trap
- ♦ Step potential used to imprint a soliton (evolve to  $\pi$  phase shift)
- ♦ Let system evolve...
- ♦ Take picture (subtle imaging with tomography)



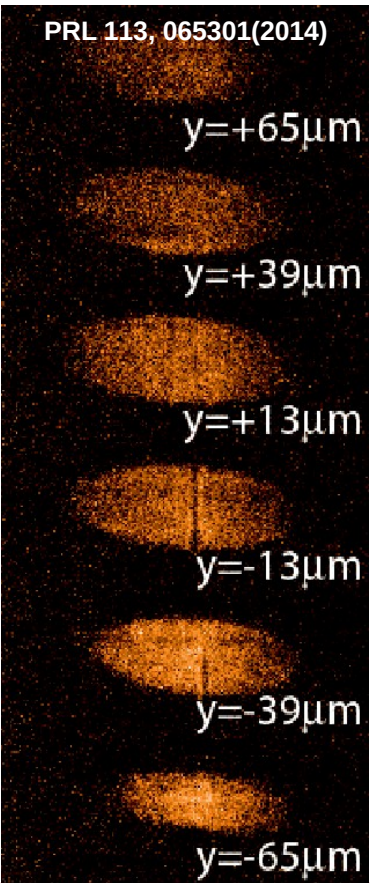
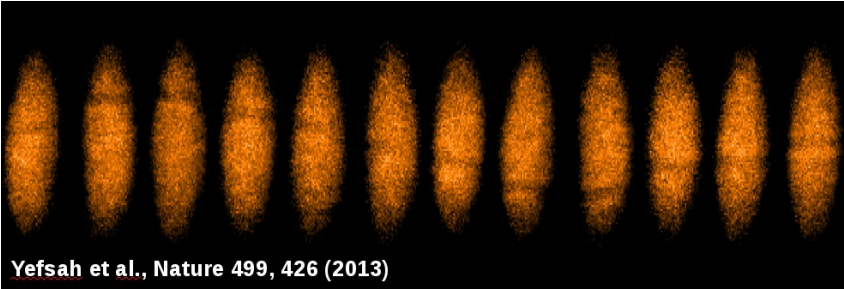
b Fig. from PRL 113, 065301(2014)



NOTE:  
We did tests against  
other dynamical properties



# Experimental results



## RESULTS:

- ◆ In the final state: Observe an oscillating **vortex line** with long period
- ◆ Inertial mass 200 times larger than the free fermion mass
- ◆ Precessional motion
- ◆ ...

