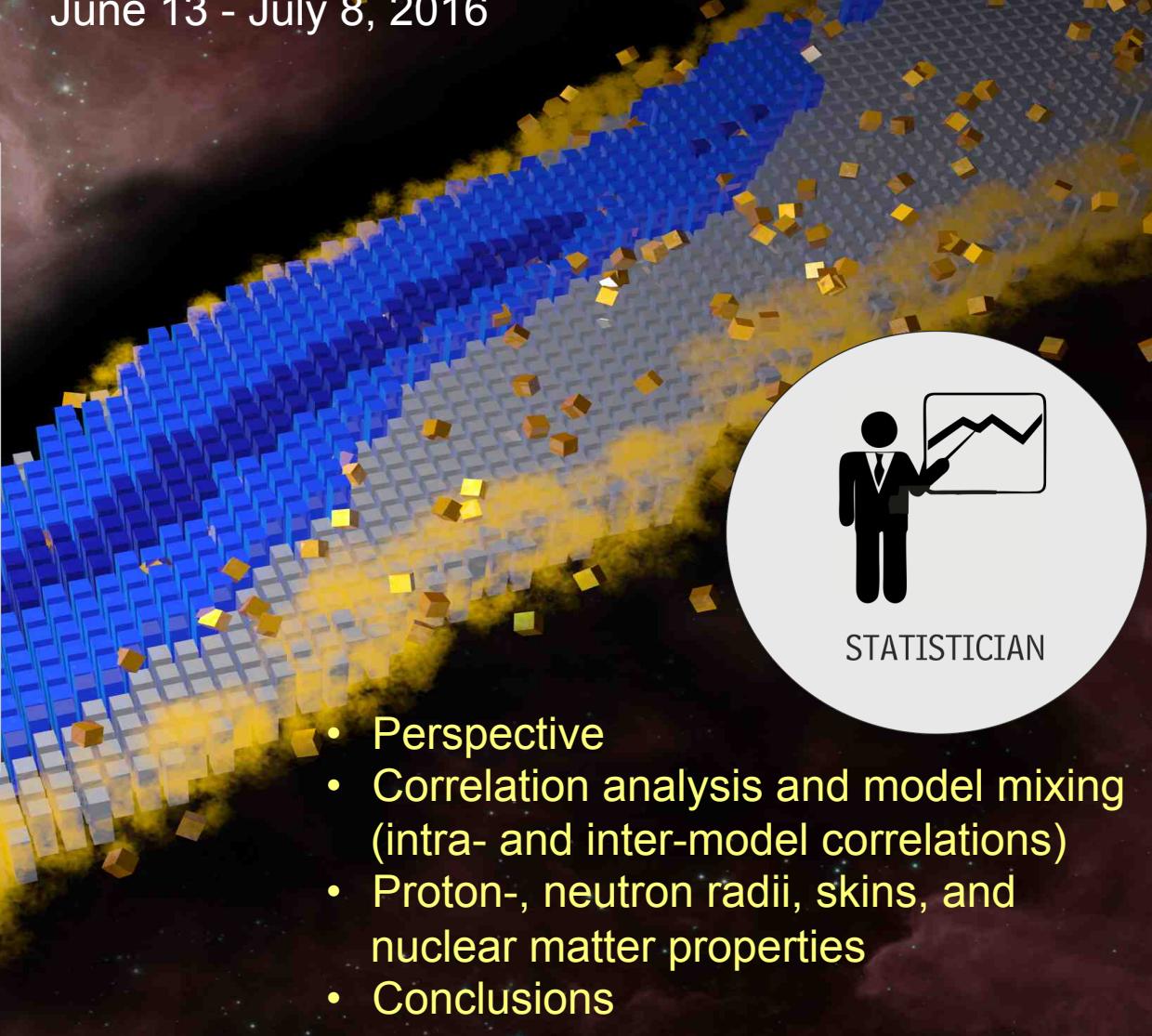
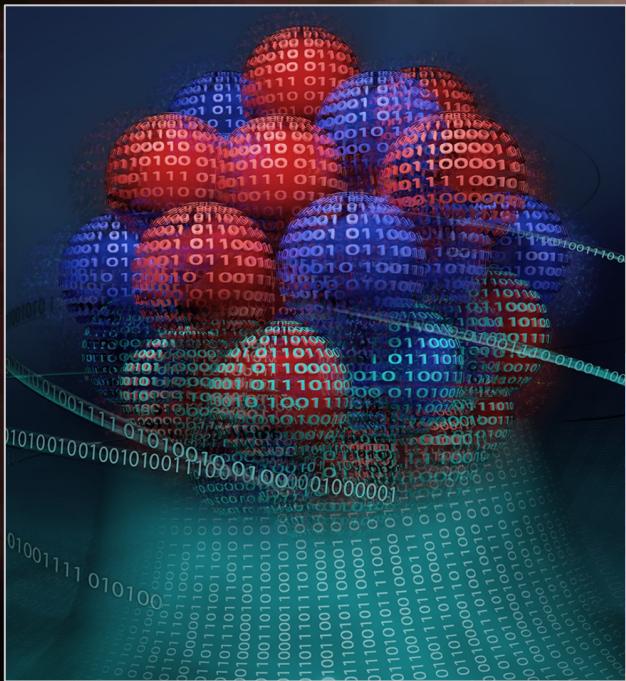


# Nuclear charge and neutron radii and nuclear matter: correlation analysis

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INT Program INT-16-2a: Bayesian Methods in Nuclear Physics

June 13 - July 8, 2016



# Classification of theories

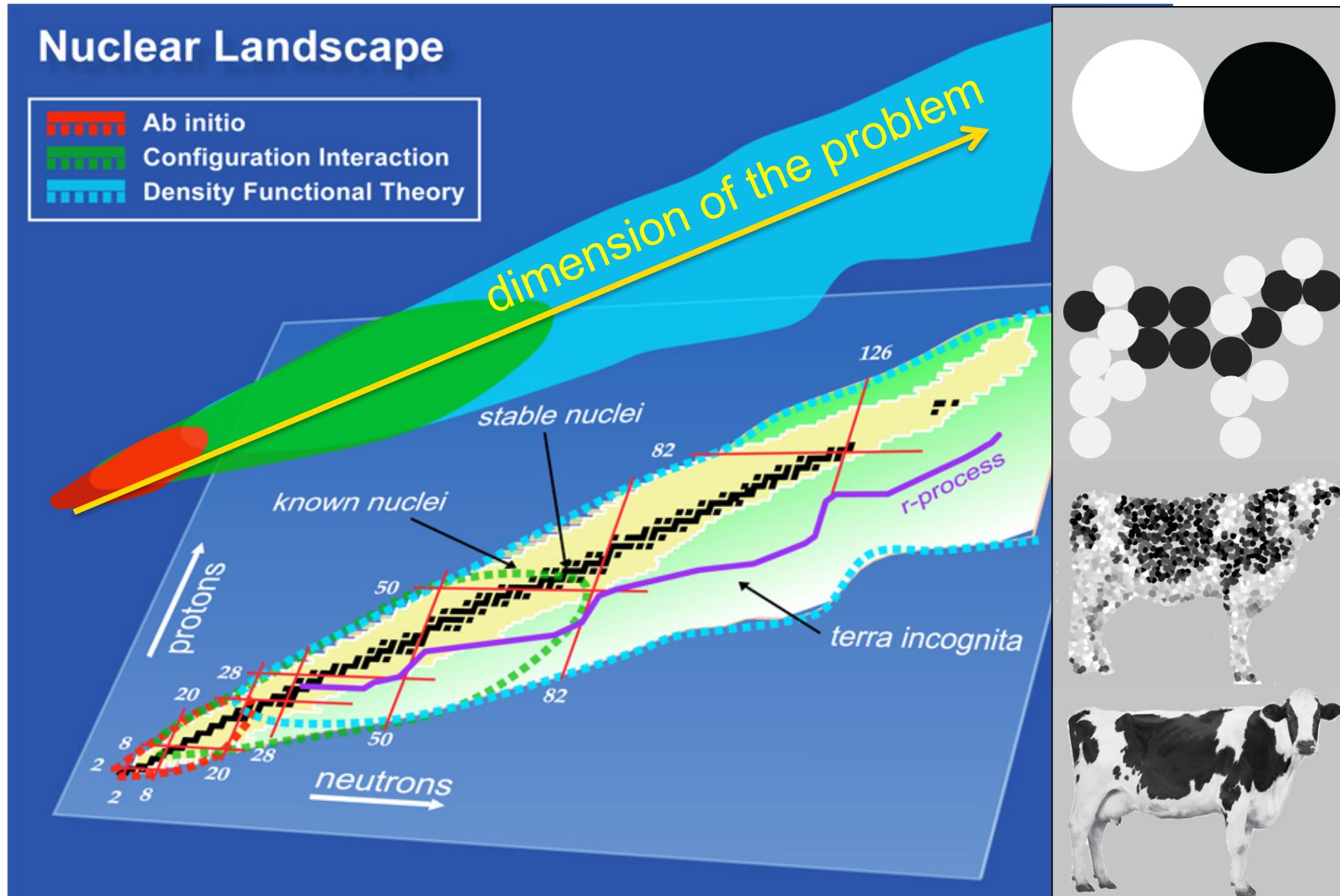
(according to Alexander I. Kitaigorodskii)

- A third rate theory explains after the event (postdictive, retrodictive)
- A second rate theory forbids
- A first rate theory predicts (predictive)

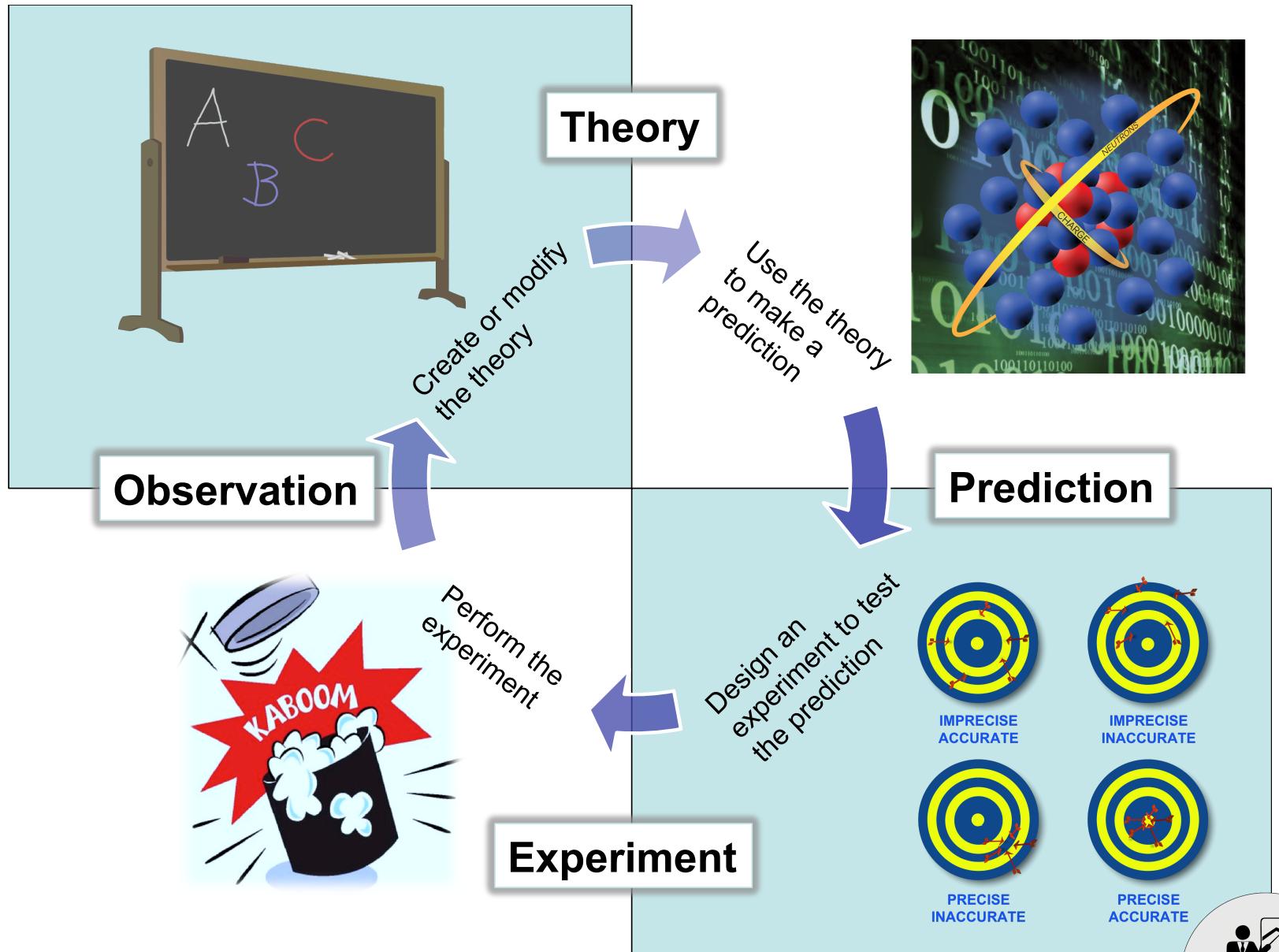
UQ is crucial to make  
this assessment



How to explain the nuclear landscape from the bottom up? **Theory roadmap**



The resolving power of a theoretical model should always be as low as reasonably possible for the question at hand



Today's posterior is tomorrow's prior



Consider a model described by coupling constants  $\boldsymbol{\theta} = \{\theta_1, \theta_2, \dots, \theta_k\}$ . Any predicted expectation value of an observable  $Y_i$  is a function of these parameters. Since the number of parameters is much smaller than the number of observables, there *must exist* correlations between computed quantities. Moreover, since the model space has been optimized to a limited set of observables, there may also exist correlations between model parameters.



$$\chi^2(\boldsymbol{\theta}) = \sum_i^{n_y} \left( \frac{Y_i(\boldsymbol{\theta}) - Y_i(\text{exp})}{\sigma_i} \right)^2$$

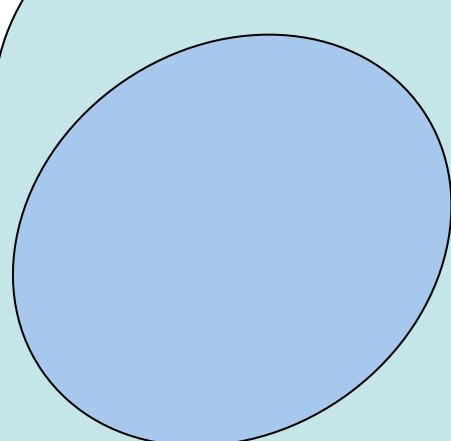
Objective  
function

Model predictions

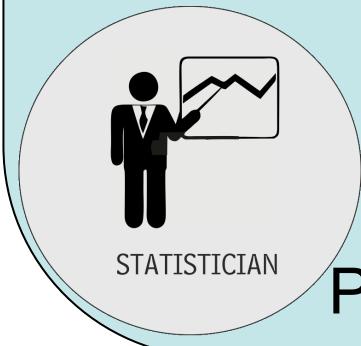
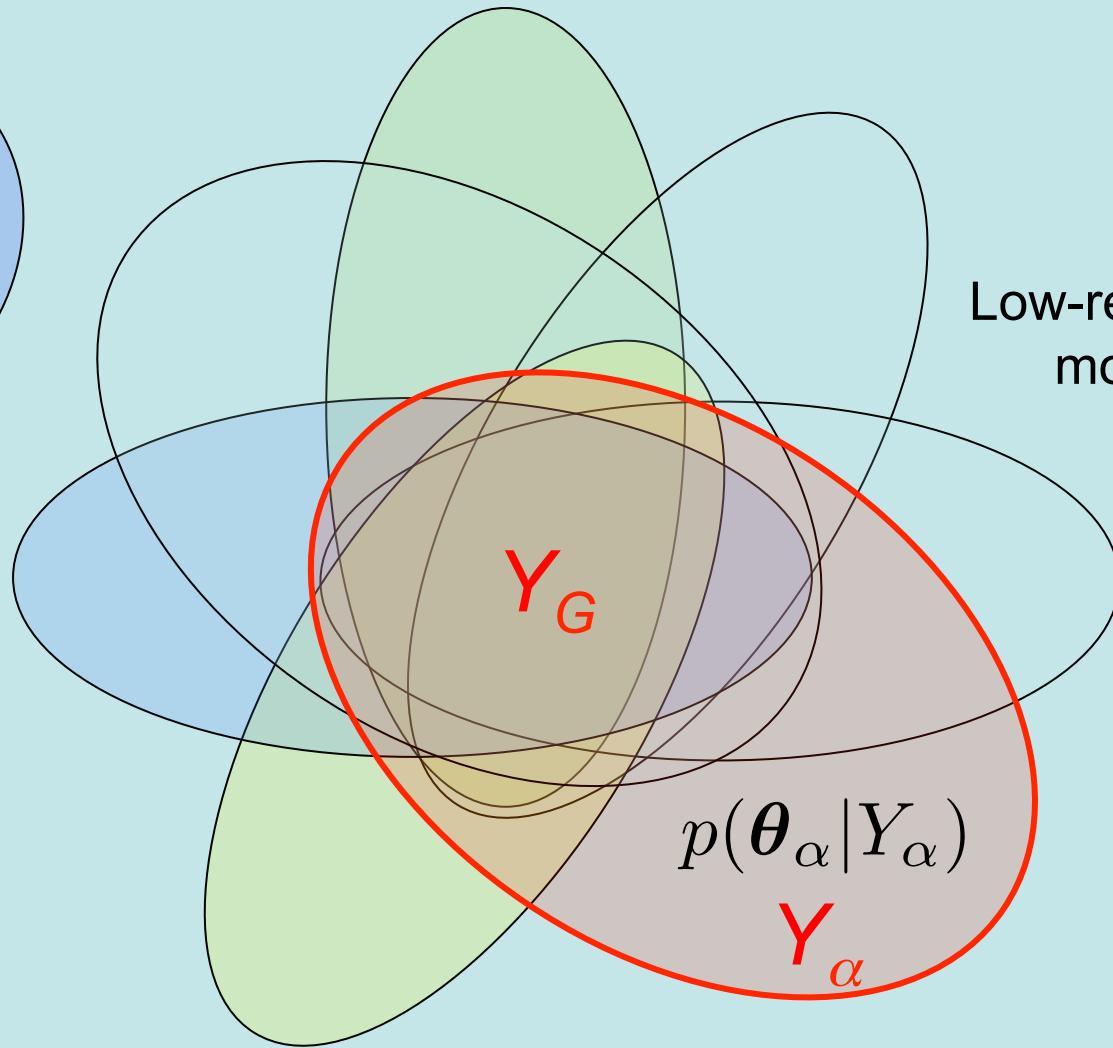
Expected uncertainties

fit-observables  
(may include pseudo-data)

# How to quantify inter-model correlations?



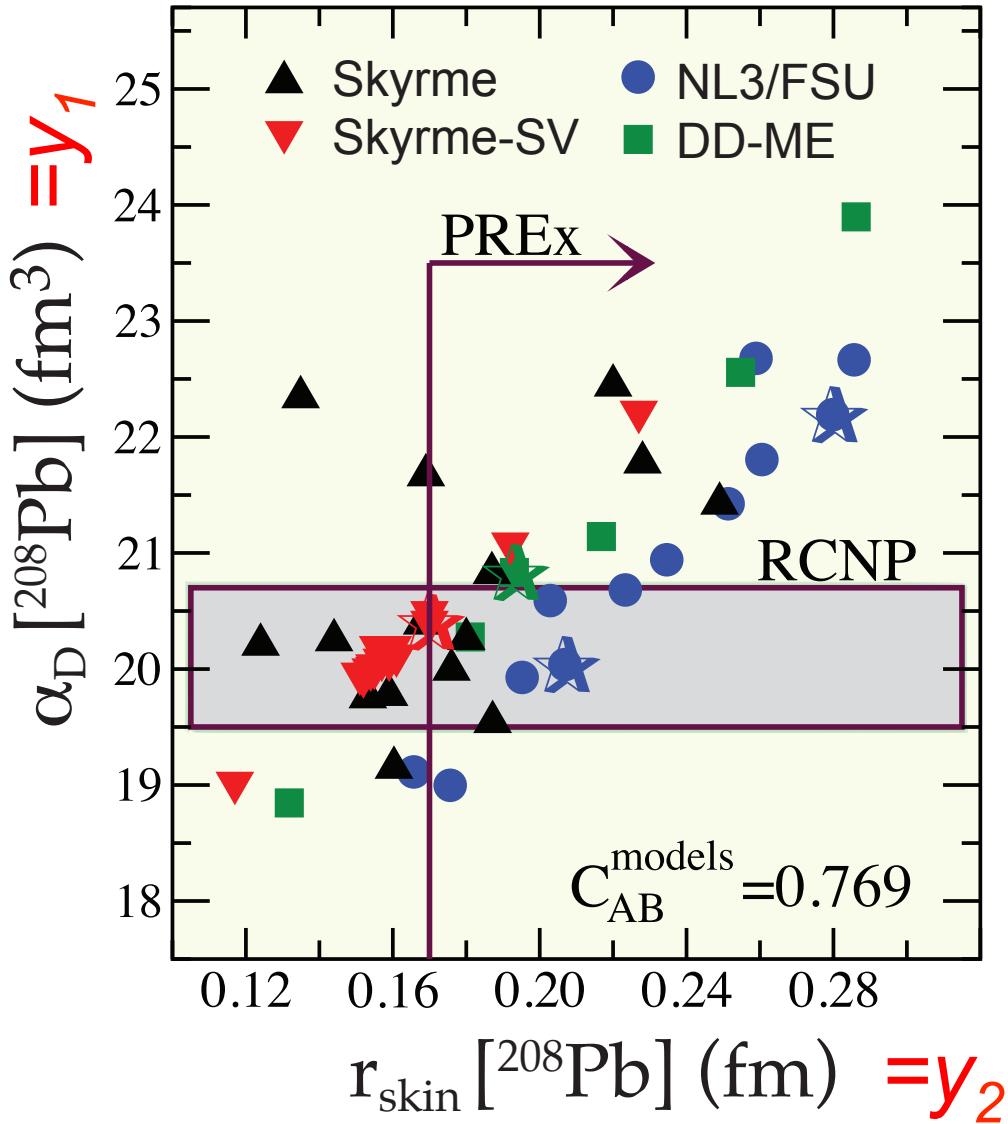
High-resolution  
models



Parameter estimation. The set of fit-observables

# Example of inter-model correlation analysis

J. Piekarewicz et al., Phys. Rev. C(1)  
85, 041302(R) (2012)



Model	$\alpha_D [{}^{208}\text{Pb}]$		
	$C_{AB}^{\text{model}}$	Slope	Intercept
Skyrme	0.9959	29.0847	15.5290
DD-ME	0.9939	31.9907	14.5206
NL3/FSU	0.9941	29.8864	13.9692

$$\hat{\theta}(M_\alpha) \equiv \theta(M_\alpha)_{\text{MLE}}$$

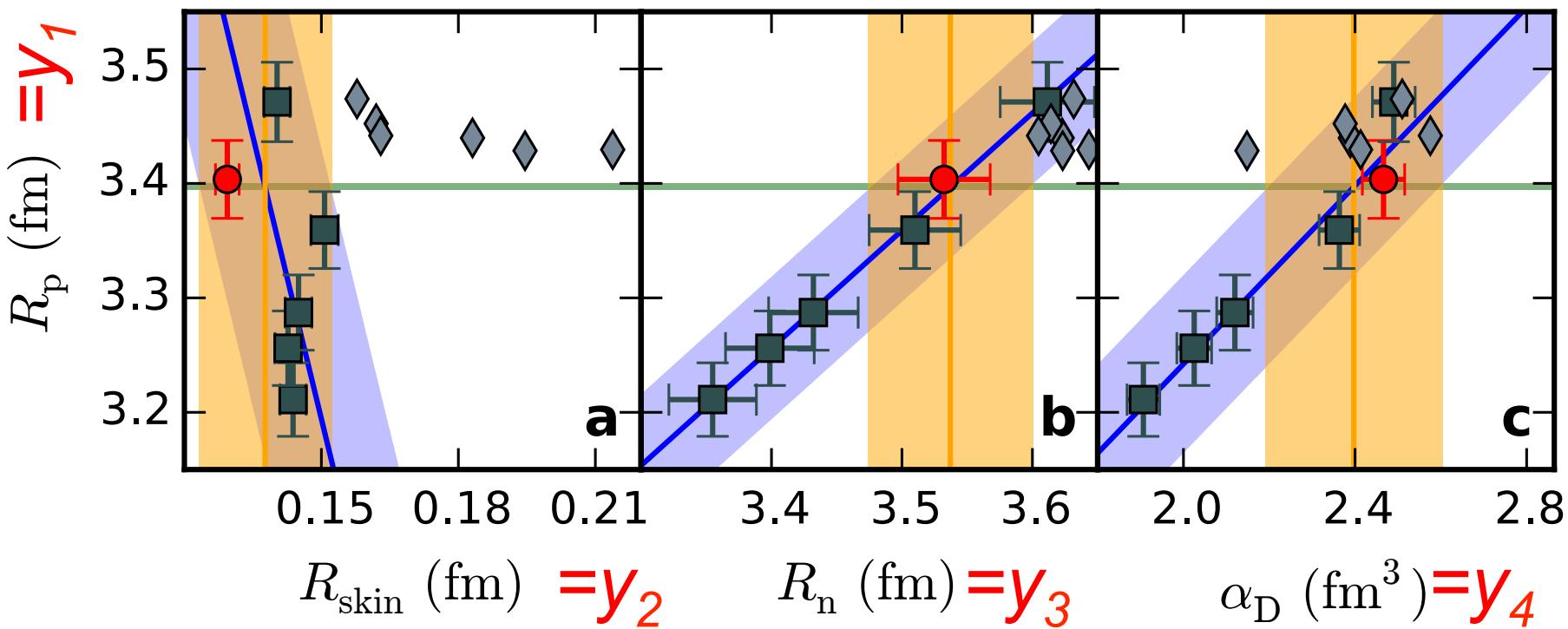
$$y_i[\hat{\theta}(M_\alpha)]$$

## Purpose:

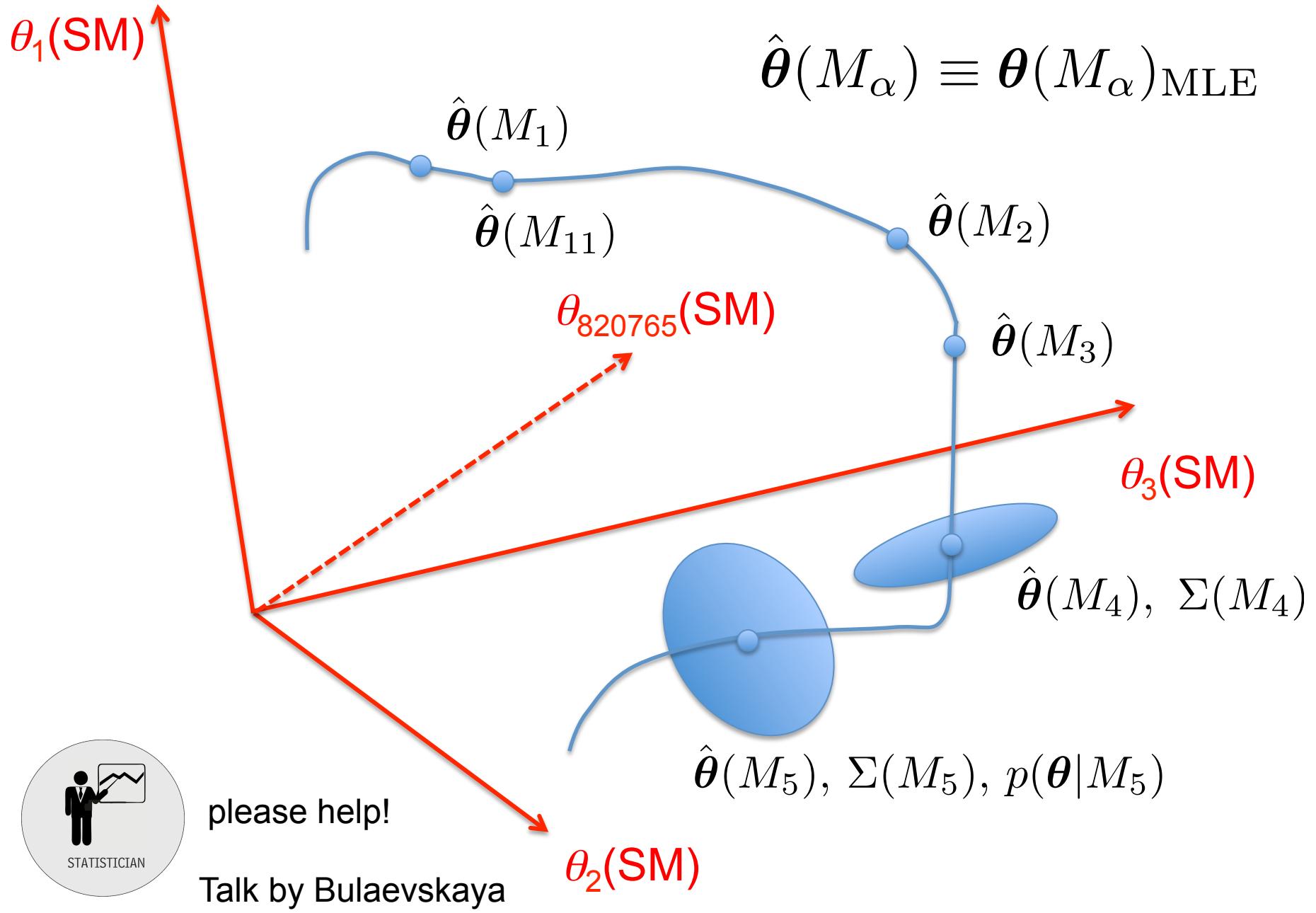
- Determine the new global relation/law
- Determine *unknown*  $y_2$  given measured  $y_1$
- Learn about constraints on models

## Example of inter-model correlation analysis (2)

G. Hagen et al., *Nature Physics* **12**, 186 (2016)



$$y_i[\hat{\theta}(M_\alpha)]$$



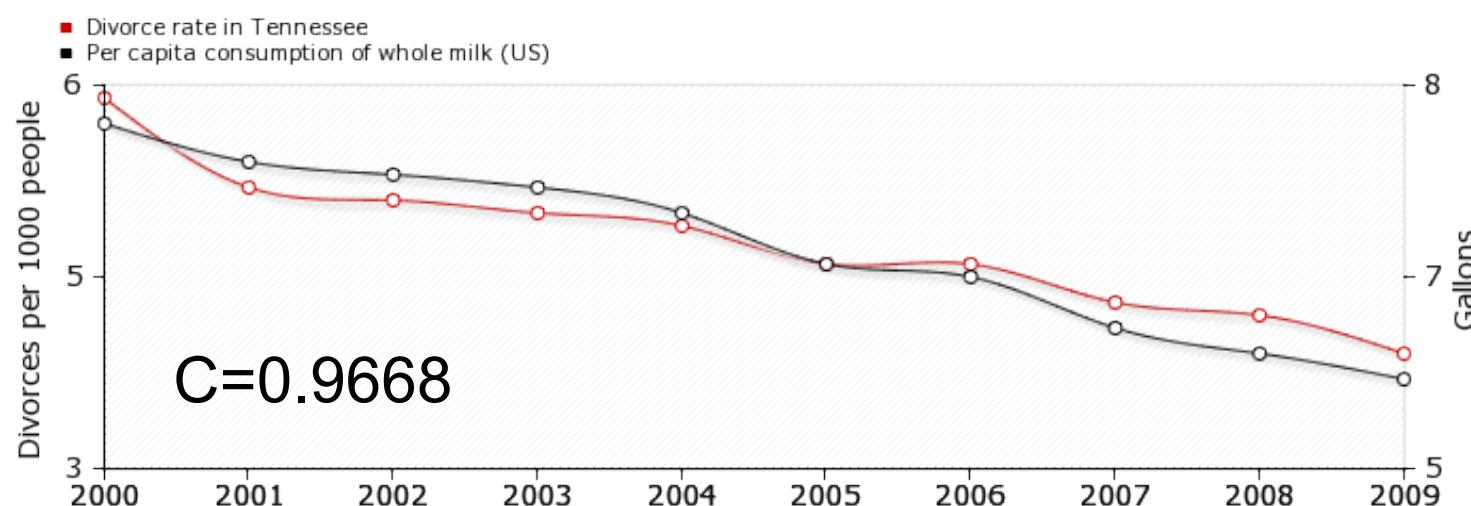
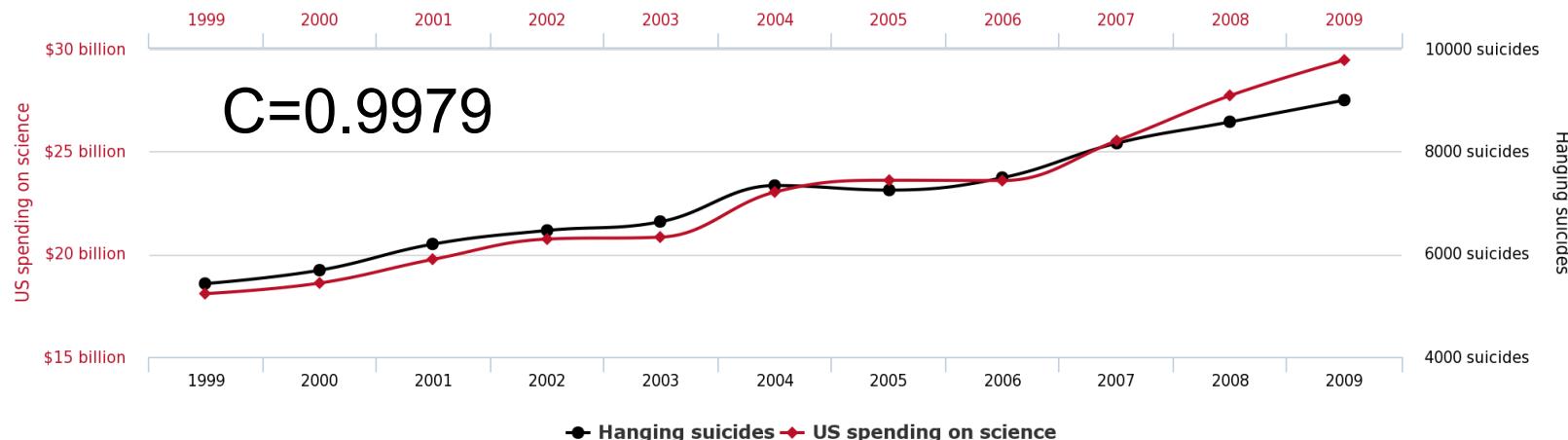
# Beware of spurious correlations!

<http://www.tylervigen.com/spurious-correlations>

## US spending on science, space, and technology

correlates with

## Suicides by hanging, strangulation and suffocation



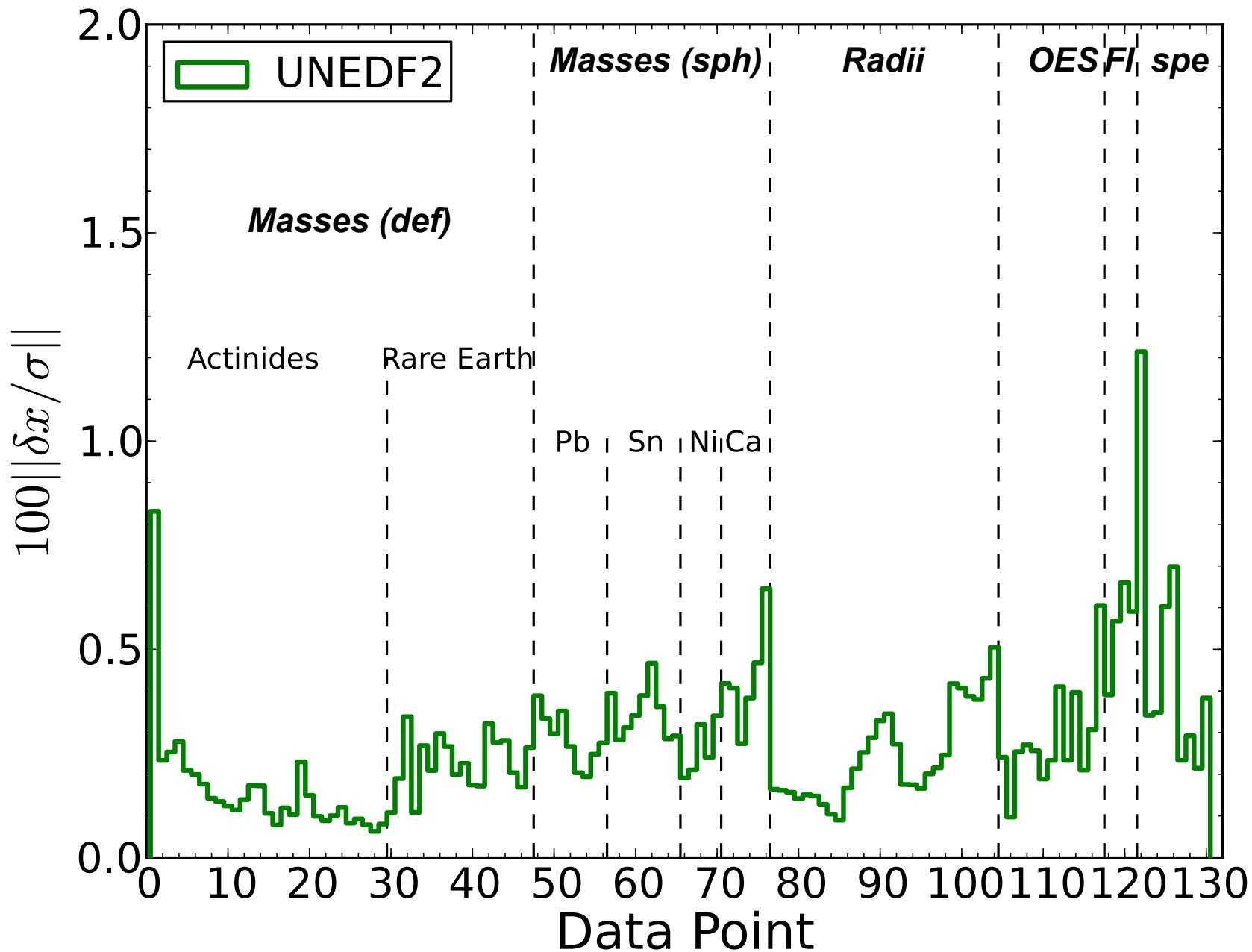
tylervigen.com

## Naïve nuclear theorist's approach to a systematic (model) error estimate:

- Take a set of *reasonable* models  $M_i$
- Make a prediction  $E(y; M_i)$
- Compute average and variation within this set
- Compute rms deviation from existing experimental data. If the number of fit-observables is large, statistical error is small and the error is predominantly systematic.

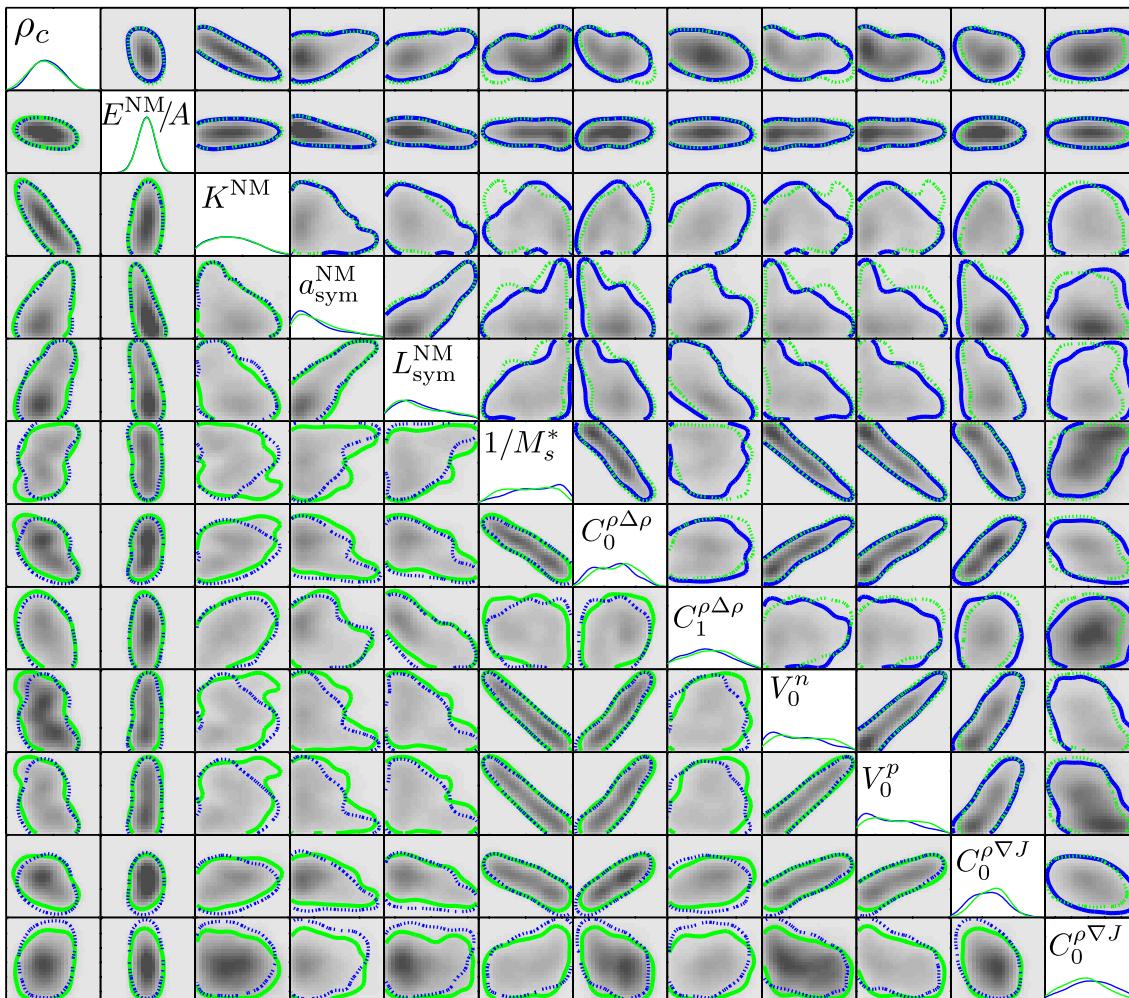
# UNEDF2 functional 12 parameters

Phys. Rev. C 89, 054314 (2014)



Uncertainty Quantification for Nuclear Density Functional Theory and Information Content of New Measurements, J. McDonnell et al., Phys. Rev. Lett. 114, 122501 (2015).

$$p(\theta_{\text{UNEDF1}} | Y_{\text{UNEDF1}})$$



UNEDF1<sub>CPT</sub>

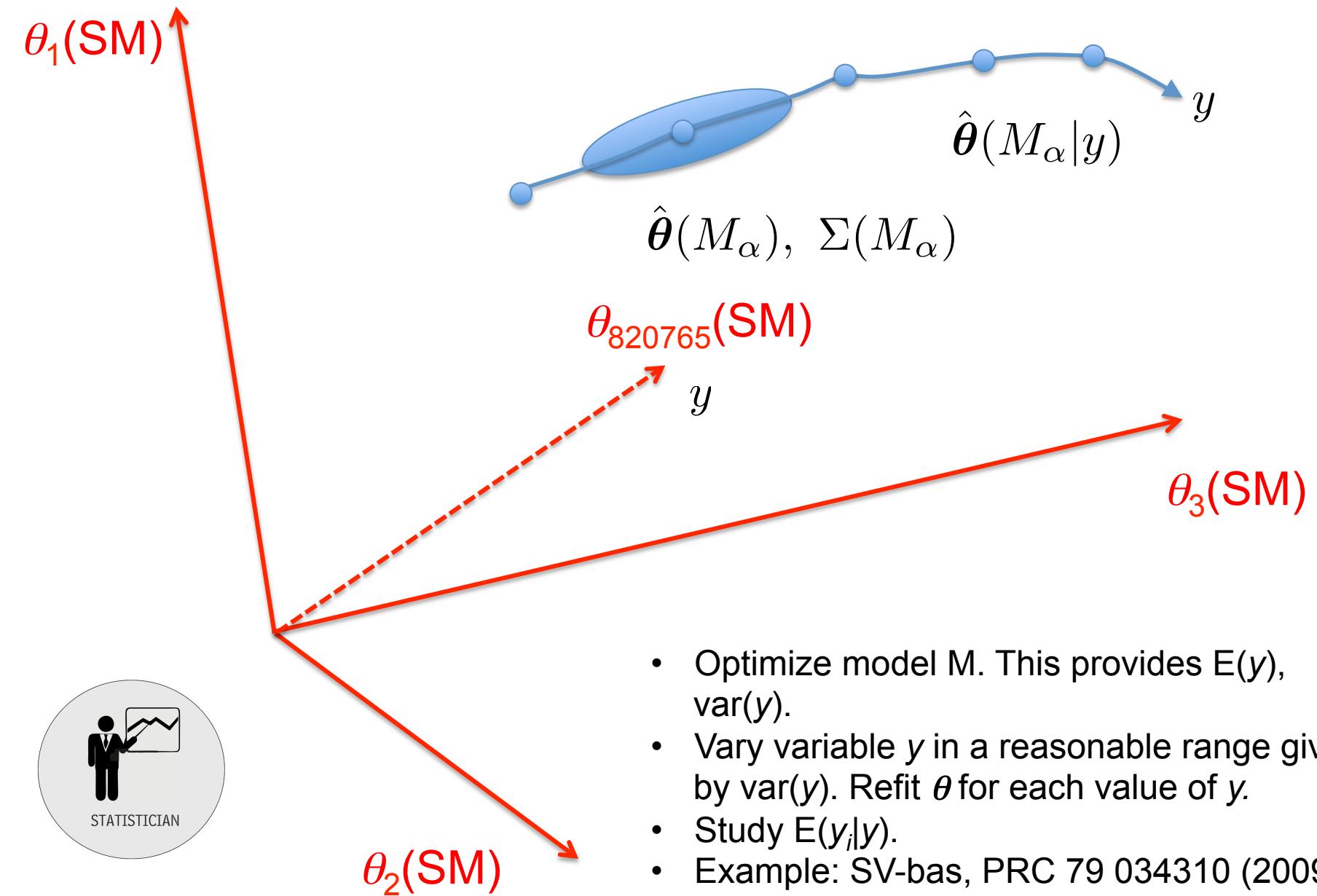
Pilot Study Applied to UNEDF1

- Massively Parallel Approach
- 130 data points (including deformed nuclei)
- Gaussian process response surface
- 200 Test Parameter Sets
- Latin hyper-rectangle

UNEDF1

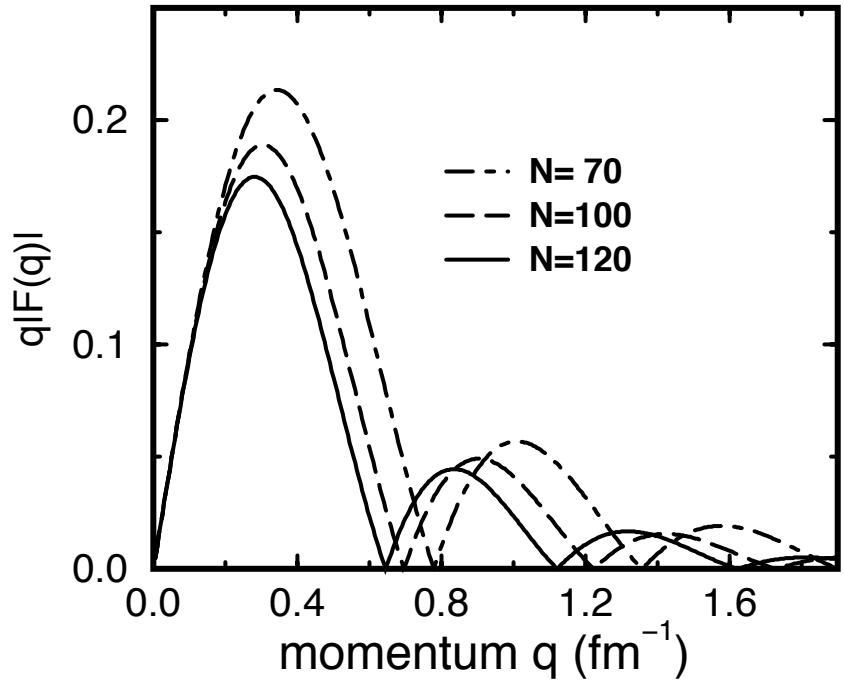
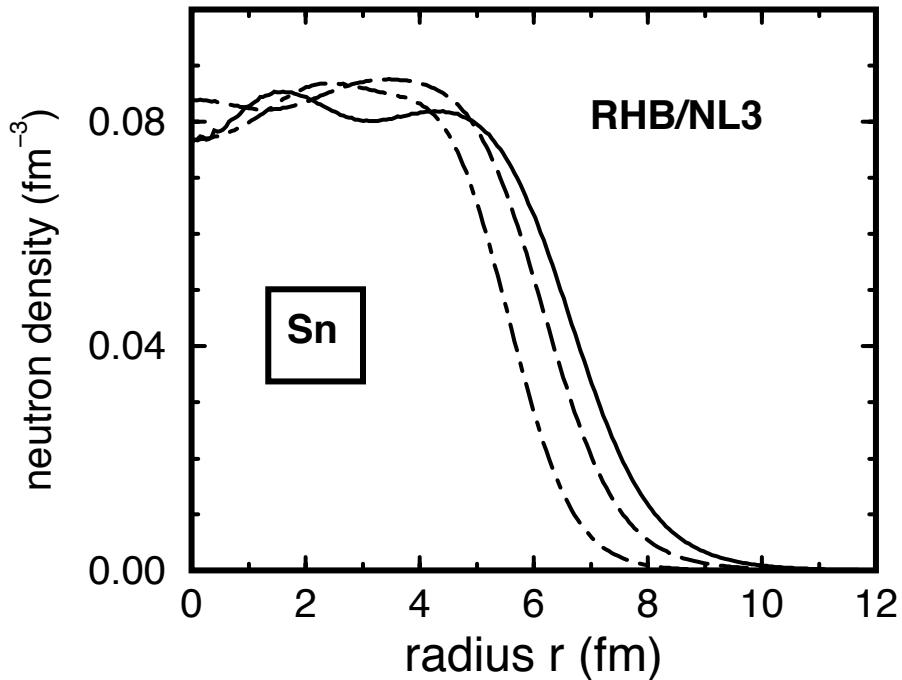
No improvement on model's predictability except for postdictions on additional data

# How to assess systematic trends?



# Radii in nuclear DFT

S. Mizutori et al., *Phys. Rev. C* **61**, 044326 (2000)

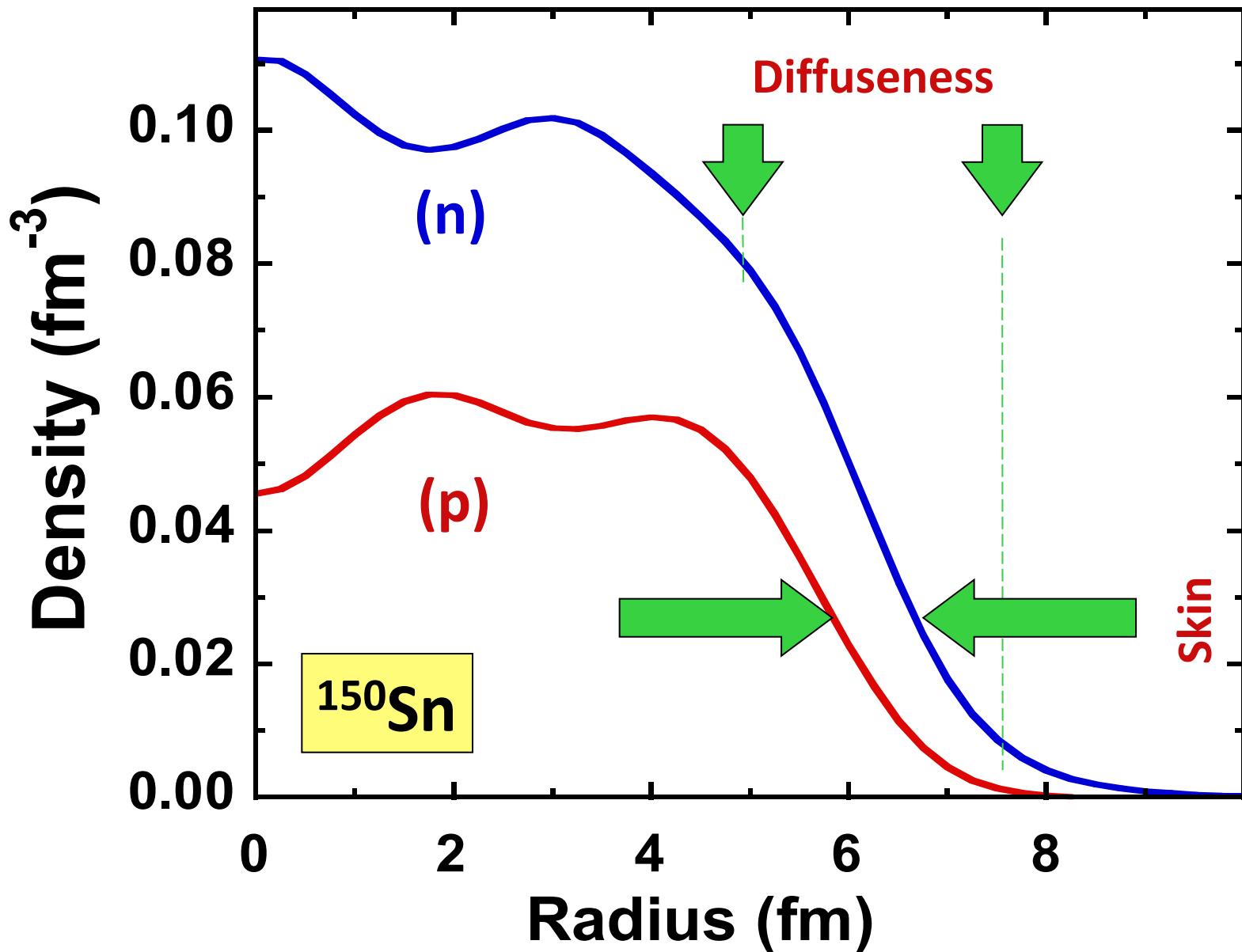


$$R_{\text{diff}} = 4.49341/q_1$$

first zero of  $F(q)$

$$R_{\text{diff}} \approx r_0 A^{1/3}$$

# Neutron & proton density distributions

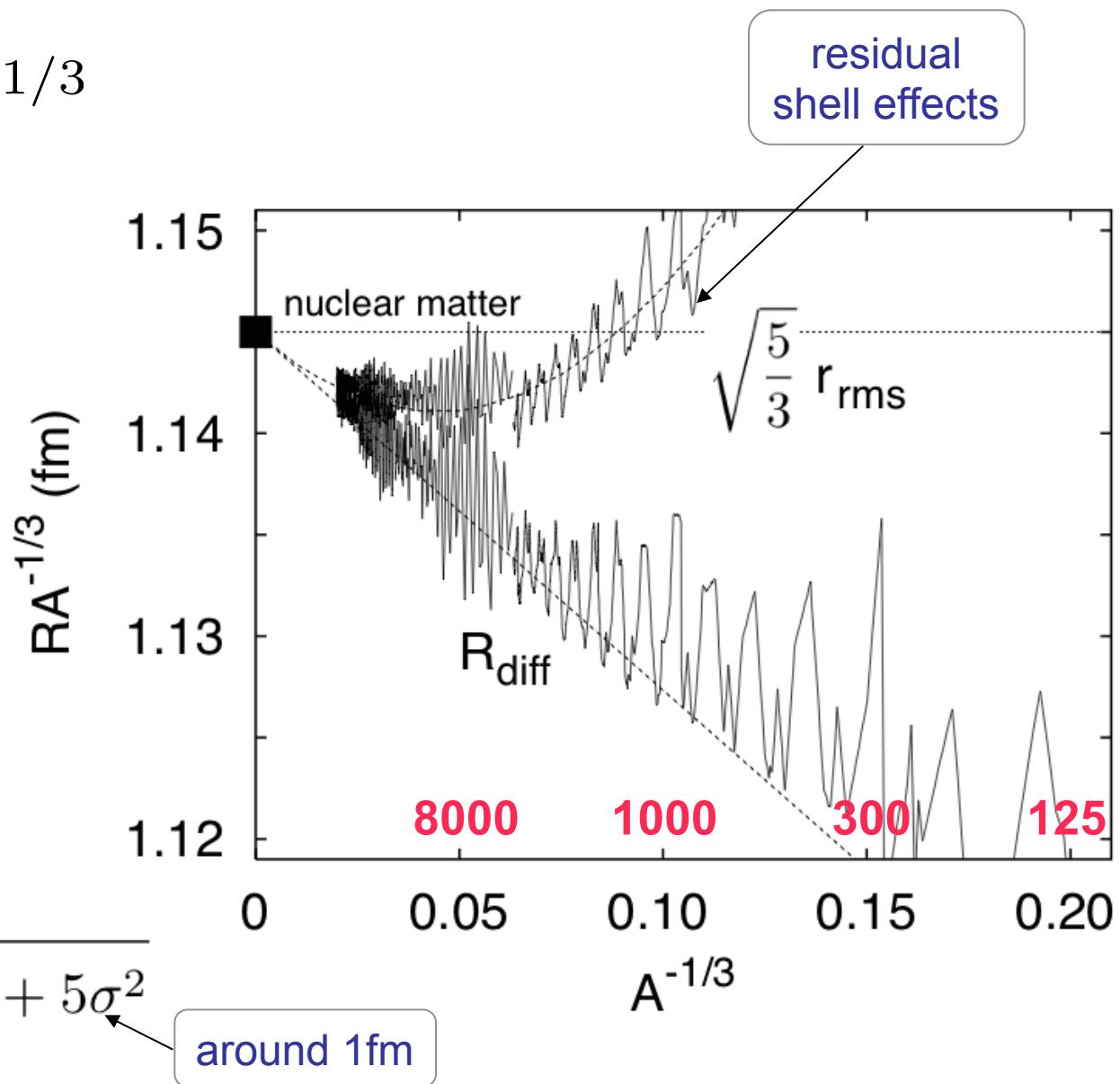


# Finite size effects and leptodermous expansion

Phys. Rev. C 73, 014309 (2006)

$$r_s = \left( \frac{3}{4\pi\rho_0} \right)^{1/3}$$

Wigner-Seitz  
radius



$$r_{\text{rms}} = \sqrt{\frac{3}{5}} \sqrt{R_{\text{diff}}^2 + 5\sigma^2}$$

# Neutron-skin uncertainties of Skyrme EDF

M. Kortelainen et al., Phys. Rev. C 88, 031305 (2013)

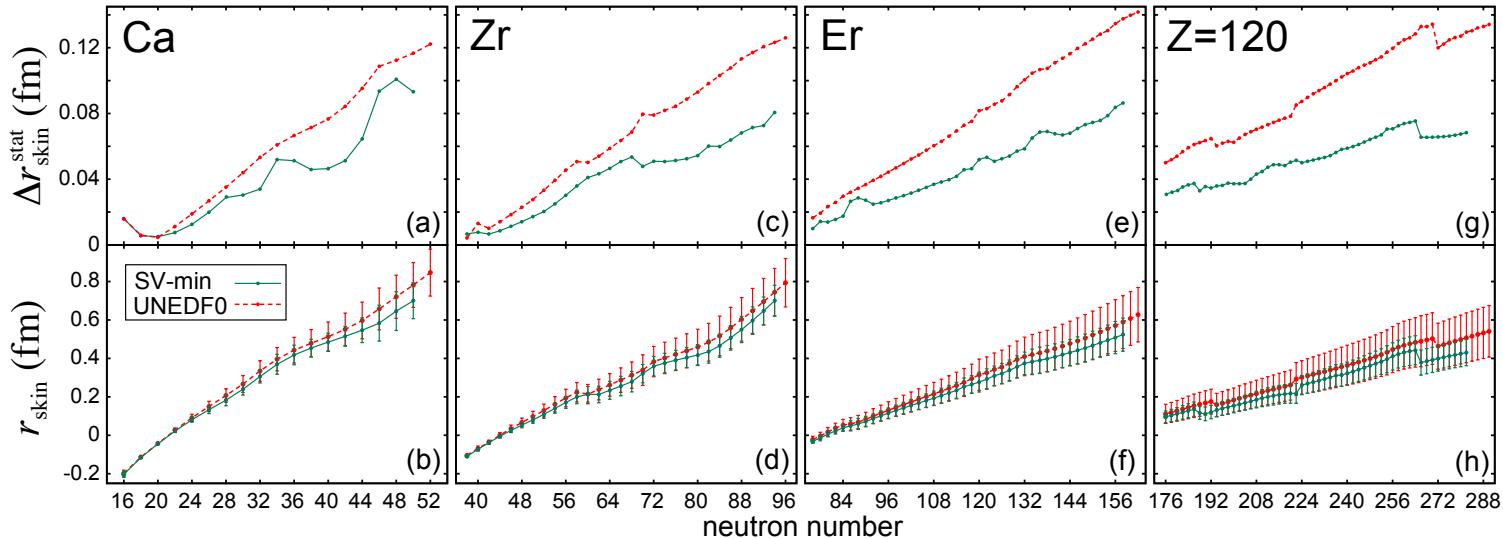
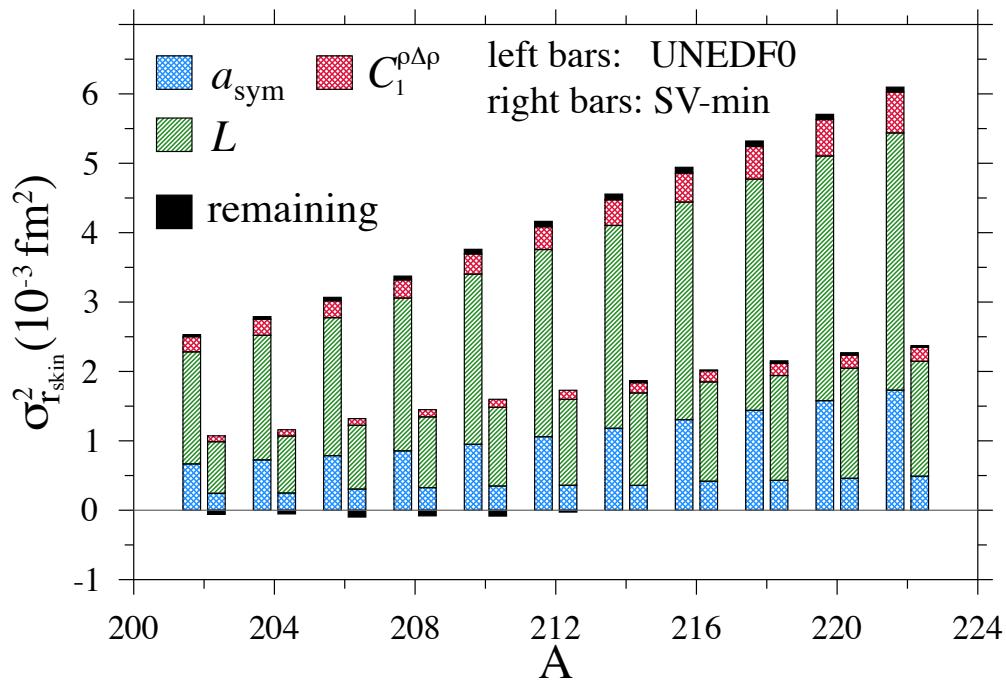


TABLE I. Theoretical uncertainties on  $r_{\text{skin}}$  in  $^{208}\text{Pb}$  and  $^{48}\text{Ca}$  (in fm). Shown are statistical errors of UNEDF0 and SV-min, systematic error  $\Delta r_{\text{skin}}^{\text{syst}}$ , the model-averaged deviation of Ref. [9], and errors of PREX [25] and planned PREX-II [29] and CREX [30] experiments.

nucleus	$\Delta r_{\text{skin}}^{\text{syst}}$ UNEDF0	$\Delta r_{\text{skin}}^{\text{syst}}$ SV-min	$\Delta r_{\text{skin}}^{\text{syst}}$ Ref. [9]	Experiment
$^{208}\text{Pb}$	0.058	0.037	0.013	0.022 0.18 [25], 0.06[29]
$^{48}\text{Ca}$	0.035	0.026	0.019	0.018 0.02 [30]



# Nuclear charge and neutron radii and nuclear matter: trend analysis in Skyrme-DFT approach



STATISTICIAN

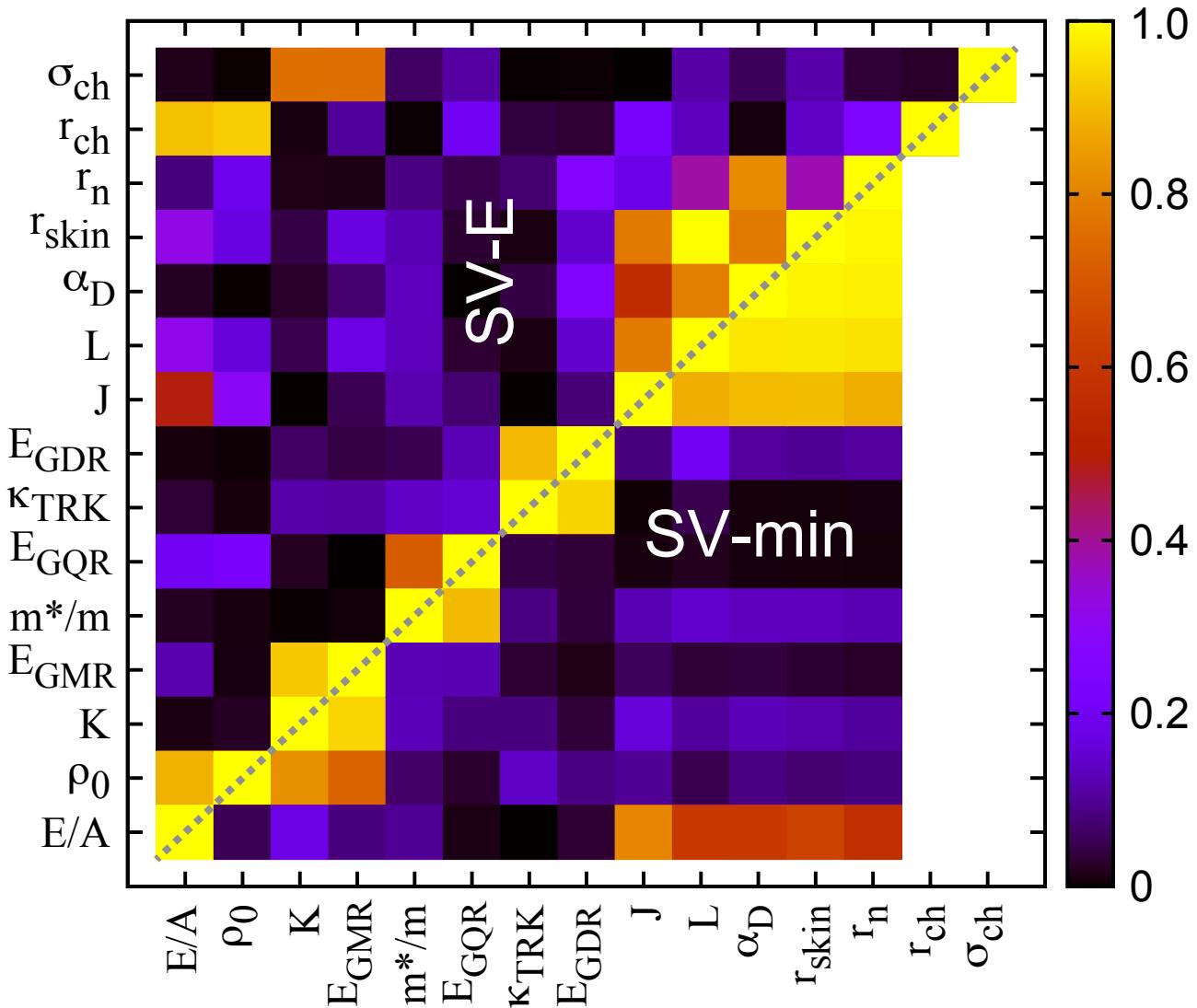
P.-G. Reinhard and WN, PRC 93, 051303 (R) (2016)

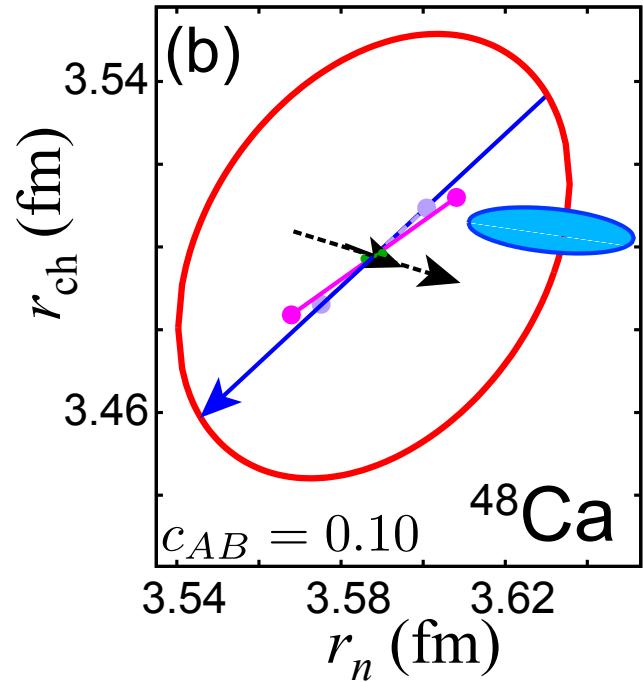
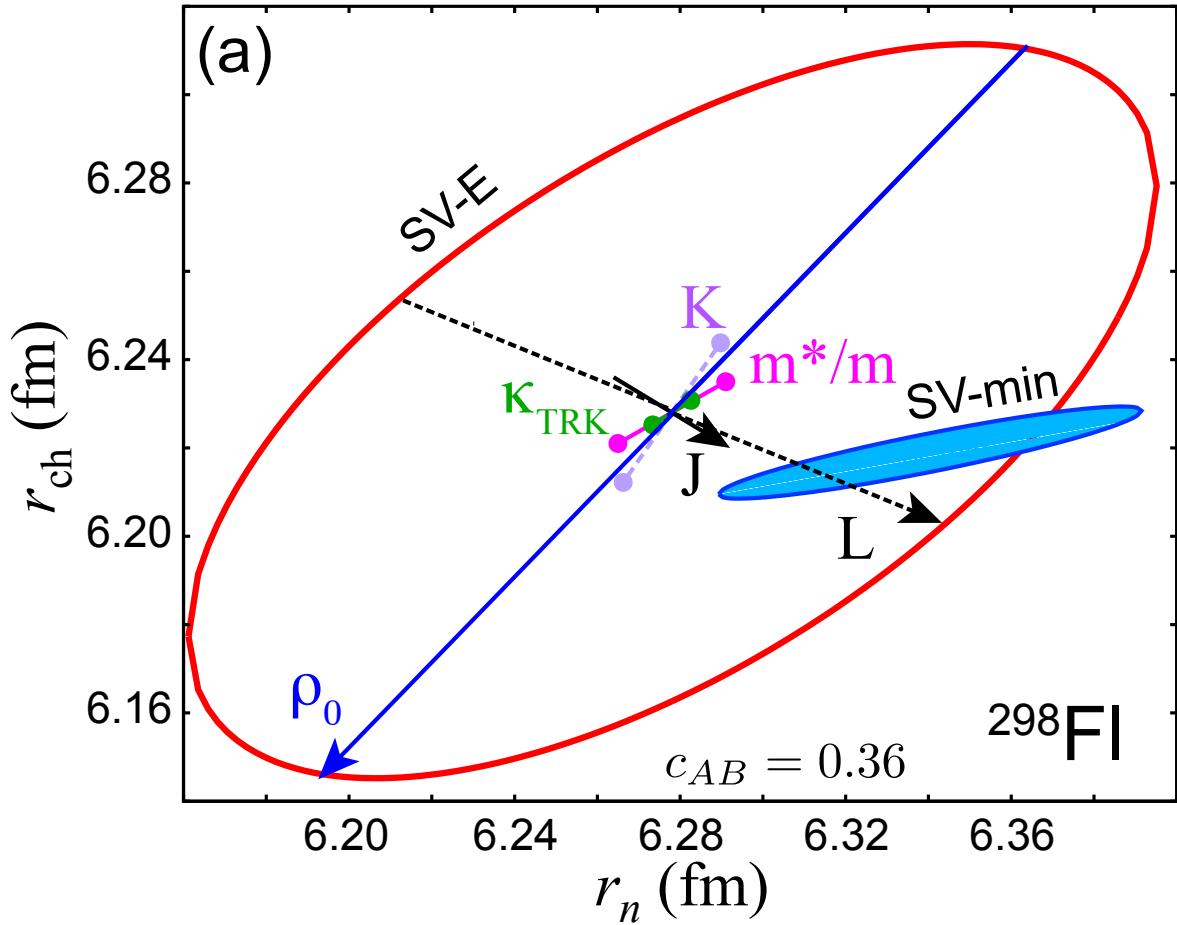
14-parameter model, optimized to 2 different sets of fit-observables

	SV-min	(Y=E, R)	SV-E	(Y=E)	
$\rho_0$ (MeV)	0.161085	$\pm$ 0.0011	0.154181	$\pm$ 0.0076	stiff
$E/A$ (MeV)	-15.9099	$\pm$ 0.04	-15.8120	$\pm$ 0.17	stiff
$K$ (MeV)	221.752	$\pm$ 8.1	273.733	$\pm$ 31.3	
$m^*/m$	0.951806	$\pm$ 0.067	1.07038	$\pm$ 0.103	
$J$ (MeV)	30.6570	$\pm$ 1.9	27.2333	$\pm$ 2.4	
$L$ (MeV)	44.8138	$\pm$ 25.7	2.92329	$\pm$ 62.9	sloppy
$\kappa_{\text{TRK}}$	0.076522	$\pm$ 0.1919	0.192	$\pm$ 0.349	
$C_0^{\Delta\rho}$ (MeV fm <sup>5</sup> )	107.657	$\pm$ 6.6	85.39992	$\pm$ 10.7	
$C_1^{\Delta\rho}$ (MeV fm <sup>5</sup> )	-141.506	$\pm$ 162	-80.90533	$\pm$ 391	sloppy
$C_0^{\nabla J}$ (MeV fm <sup>4</sup> )	-101.582	$\pm$ 5.5	-96.3170	$\pm$ 11.7	
$C_1^{\nabla J}$ (MeV fm <sup>4</sup> )	-22.9681	$\pm$ 16.2	-21.5881	$\pm$ 18.2	sloppy
$V_{\text{pair,p}}$ (MeV fm <sup>3</sup> )	601.160	$\pm$ 190	613.231	$\pm$ 209	
$V_{\text{pair,n}}$ (MeV fm <sup>3</sup> )	567.190	$\pm$ 154	568.739	$\pm$ 173	
$\rho_{0,\text{pair}}$ (fm <sup>-3</sup> )	0.211591	$\pm$ 0.052	0.202513	$\pm$ 0.046	

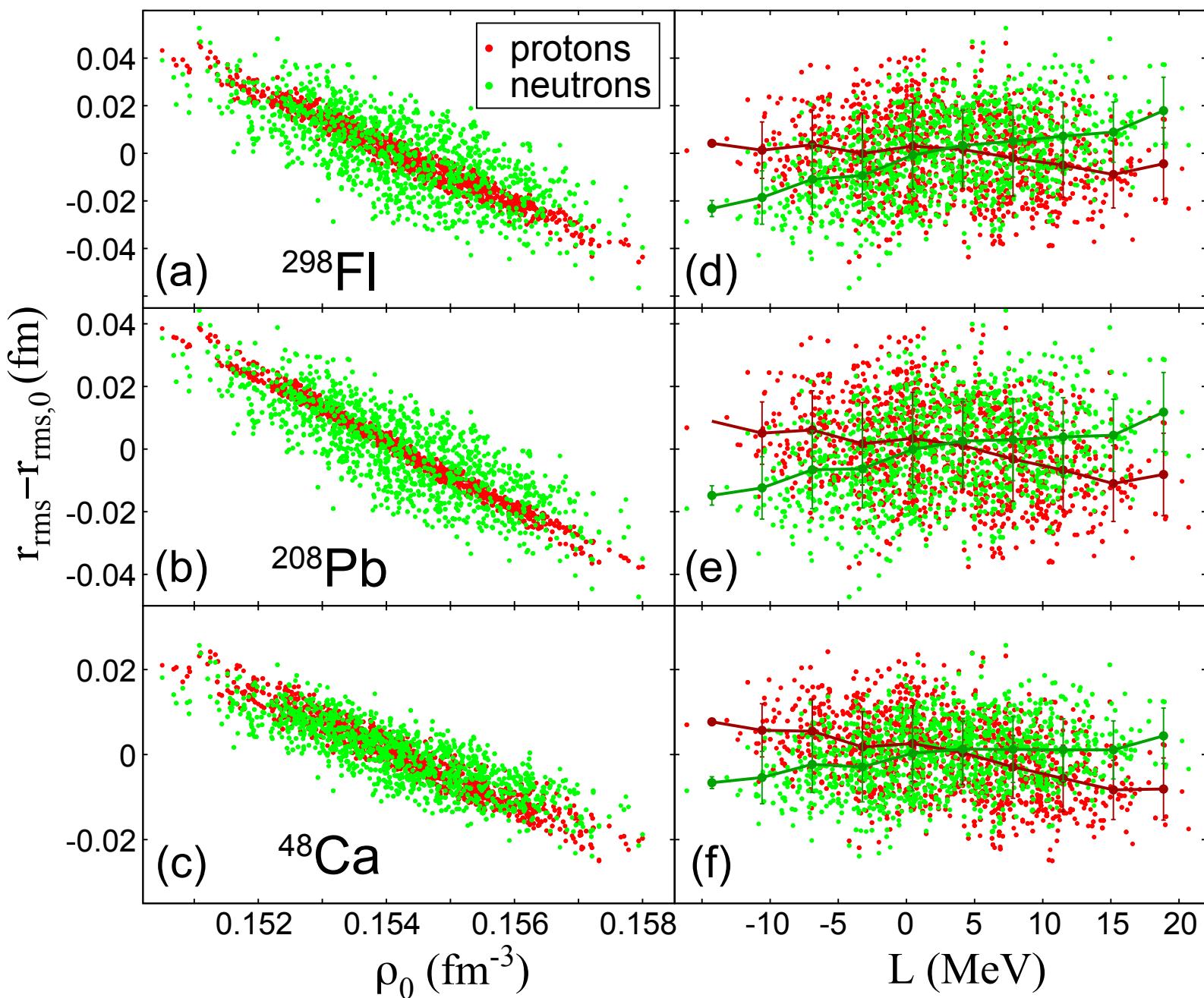
# Nuclear charge and neutron radii, and nuclear matter: intra-model trend analysis

P.-G. Reinhard and WN, PRC (R) (2016)

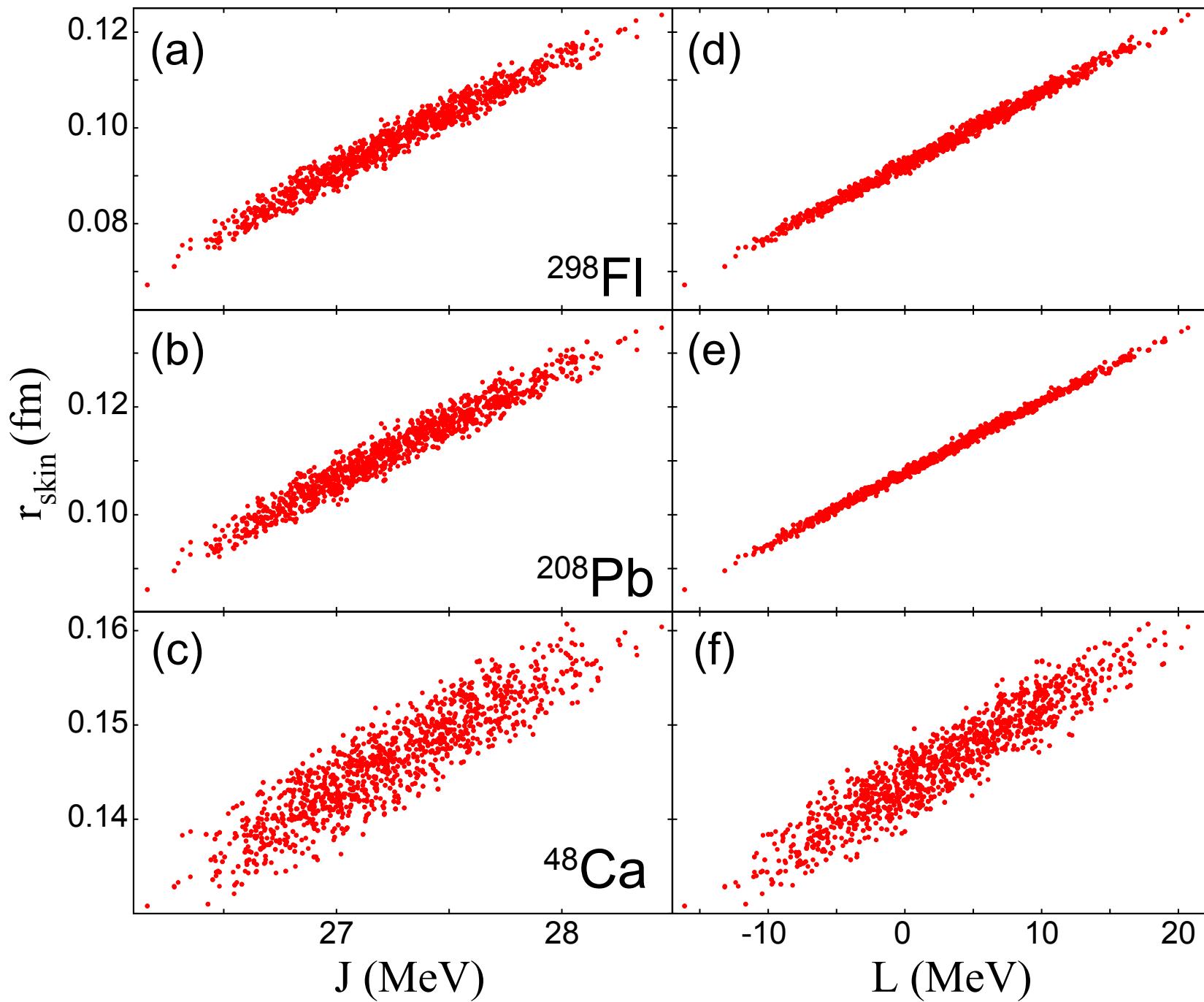


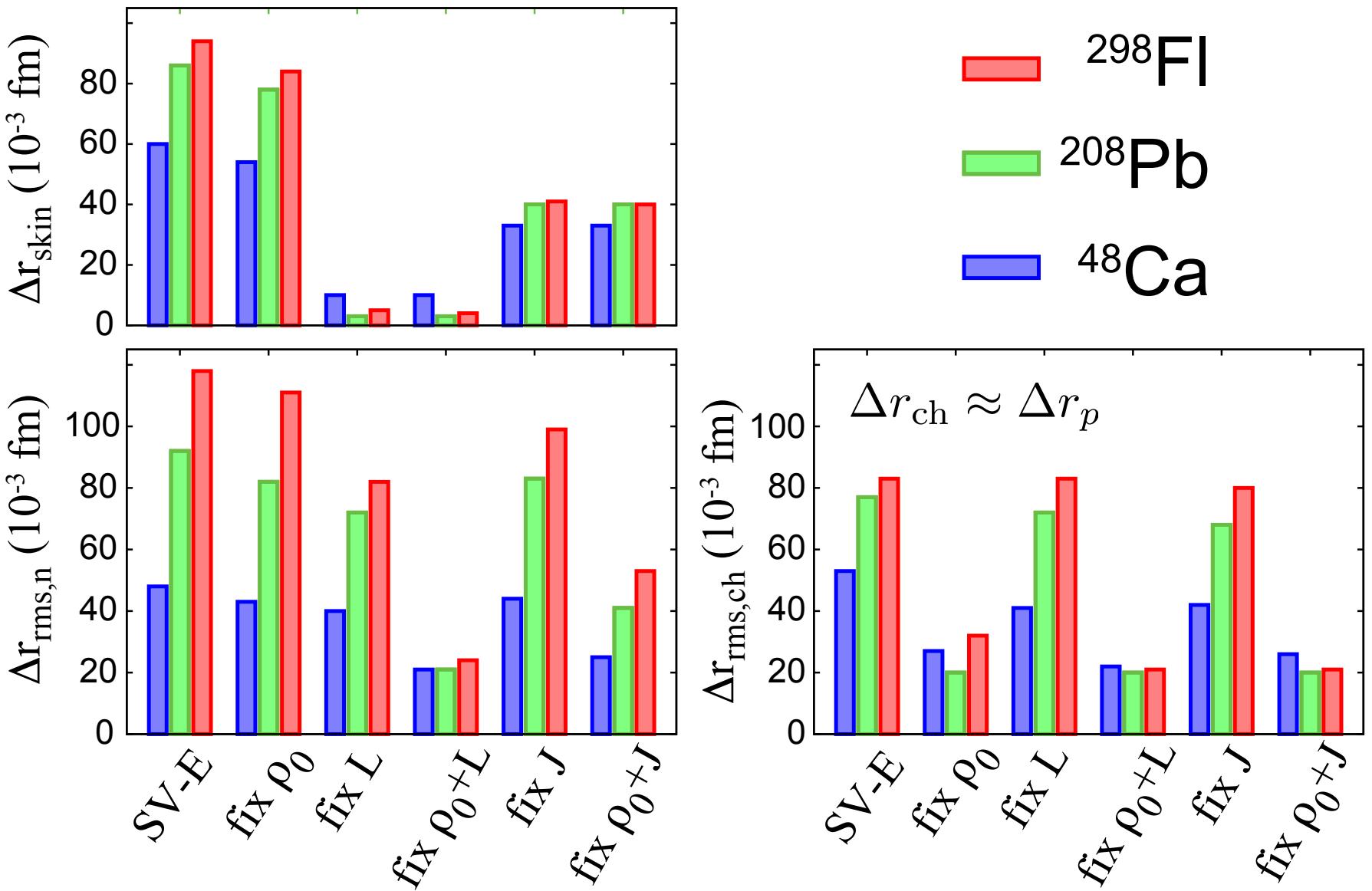


$$c_{AB} = \frac{|\overline{\Delta A} \overline{\Delta B}|}{\sqrt{\overline{\Delta A}^2 \overline{\Delta B}^2}}$$



2000 Gaussian samples of  $L(\theta)$





$$\Delta r_n = \Delta r_p + \Delta r_{\text{skin}}$$

- We explored various trends of charge and neutron radii with nuclear matter properties.
- There exist, at least within the Skyrme-DFT theory, only two strong correlations:
  - one-to-one relation between charge radii in finite nuclei and  $\rho_0$ :  $r_p \leftrightarrow \rho_0$
  - one-to-one relation between neutron skins in finite nuclei and  $L$ :  $r_{skin} \leftrightarrow L$
- By including charge radii in a set of fit-observables, as done for the majority of realistic Skyrme EDFs, one practically fixes the saturation density.
- The relation  $r_n \leftrightarrow \rho_0$  is much weaker than that for  $r_p$ , so by constraining the saturation density alone does not help significantly reducing the uncertainty on neutron (and mass) radii. However:

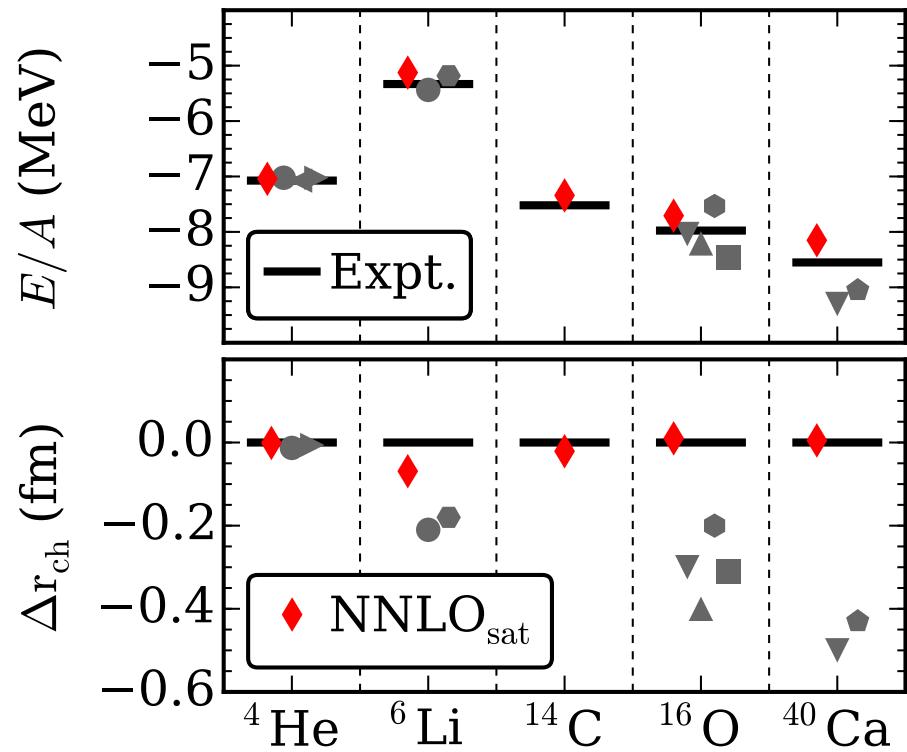
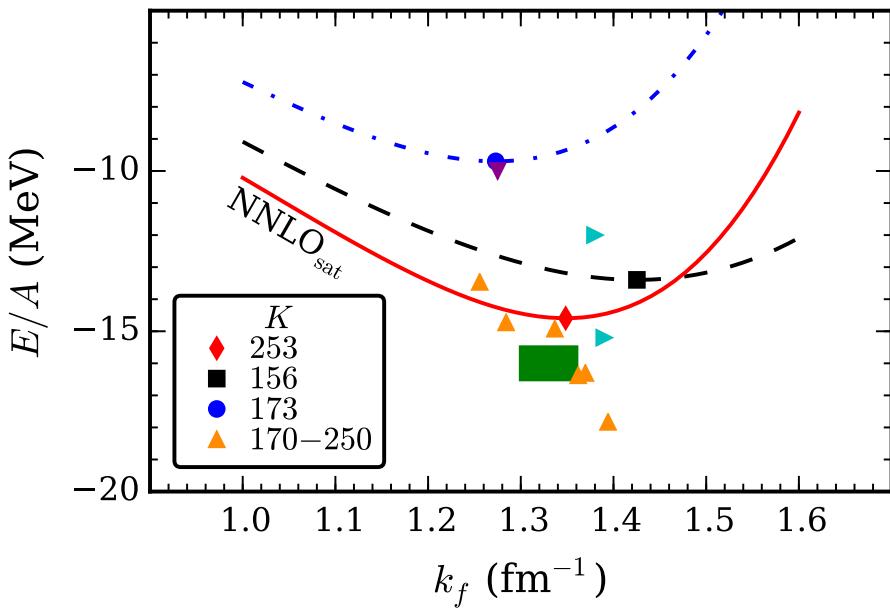
$$r_n = r_p + r_{skin}$$

- The  $r_n \leftrightarrow r_p$  relation is fairly complex: various trends are possible when moving along a trajectory in a parameter space.

# $N2LO_{sat}$ describes low-energy NN and Nuclei

A. Ekström et al. Phys. Rev. C 91, 051301(R) (2015)

- Order-by-order optimization
- Constrained by data on few-body systems and light nuclei



Coupled Cluster informing DFT  
and  
DFT informing Coupled Cluster