

Toward a new effective interaction in energy density functional theory

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With M. Grasso, D. Lacroix, U. van Kolck

INT workshop

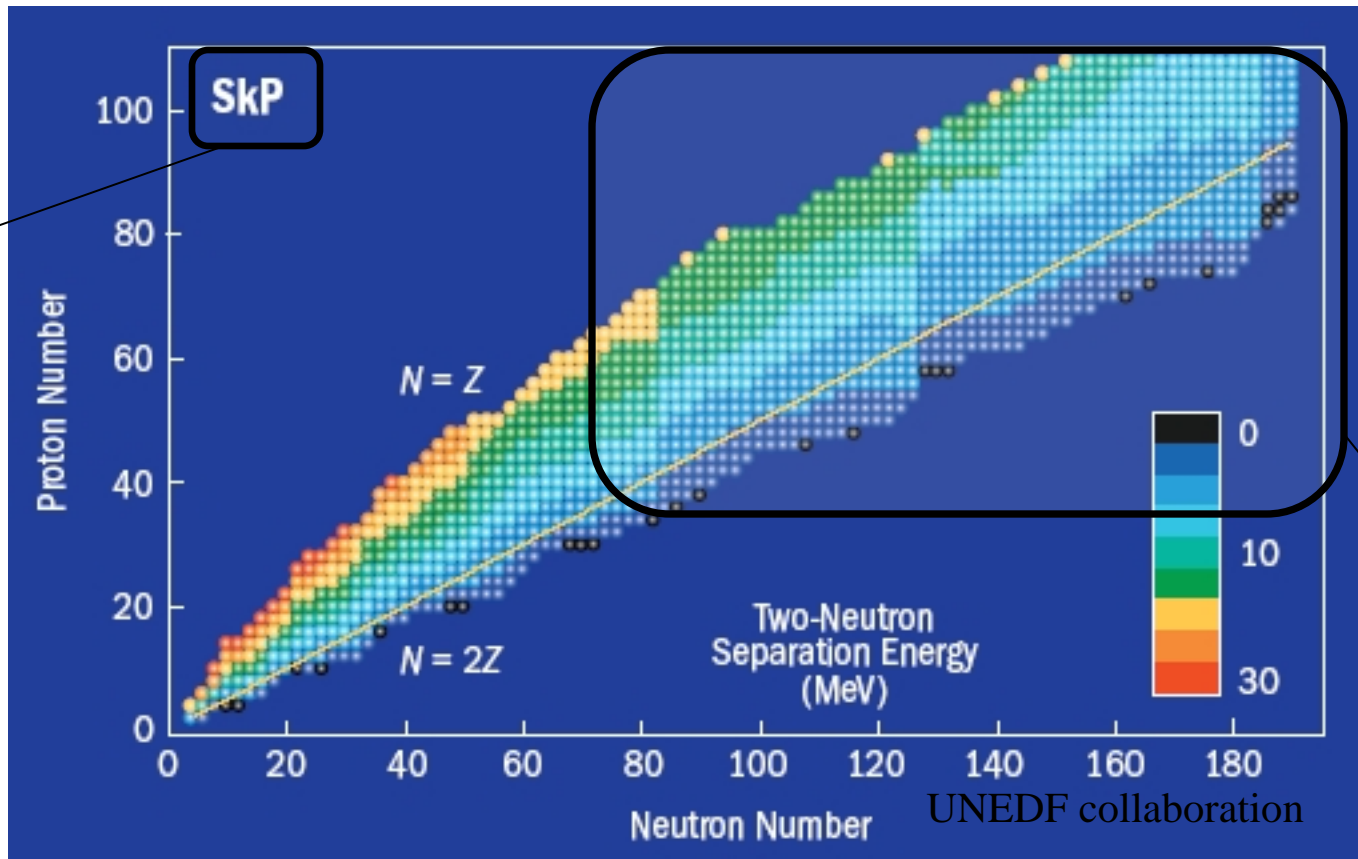
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Outline

- **Beyond mean field approach.** (C.J. Yang, M. Grasso, X. Roca-Maza, G. Colo, and K. Moghrabi, arXiv:1604.06278)
- **Renormalizability.** (C.J. Yang, M. Grasso, K. Moghrabi, U van Kolck, coming soon!)
- **Low density limit.** (C.J. Yang, M. Grasso, D Lacroix, arXiv:1604.06587)
- **Conclusion and future work.**

Motivation



Skyrme-type interaction works o.k. (able to do the fitting in EDF framework)

No way to get with ab-initio!

Need to think about other expansion (than on NN d.o.f.).

Present status of EDF

- Energy density functional (EDF) framework gives reasonable results at mean field, when sufficient amount of parameters (~ 10) are included.
- Include **more parameters won't necessarily help**.
 - Limited predictive power.
- Maybe the correct theory has a structure where **different terms appears at different order**.
 - Need to go beyond mean field to perform the test.

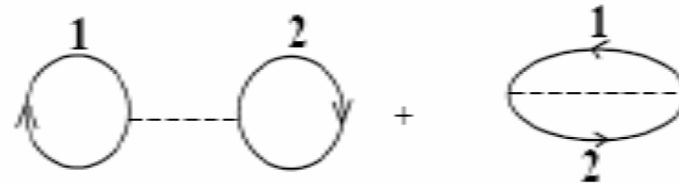
Interaction & mean field EoS

Interaction: Skyrme without spin-orbit

$$\begin{aligned}
 v = & \underbrace{t_0(1+x_0P_\sigma)}_{S\text{-wave } O(0)} + \frac{1}{2} \underbrace{t_1(1+x_1P_\sigma)(k'^2+k^2)}_{S\text{-wave } O(q^2)} + \underbrace{t_2(1+x_2P_\sigma)\mathbf{k}'\cdot\mathbf{k}}_{p\text{-wave } O(q^2)} \\
 & + \frac{1}{6} \underbrace{t_3(1+x_3P_\sigma)\rho^\alpha}_{s\text{-wave, higher body}}. \qquad P_\sigma = \frac{1}{2}(1+\sigma_1\cdot\sigma_2)
 \end{aligned}$$

No pion! Like pionless EFT, except for the density-dependent term.

$$EoS: \quad \frac{E}{A} \propto \frac{1}{\rho} \int_0^{k_{F1}} d^3\mathbf{k}_1 \int_0^{k_{F2}} d^3\mathbf{k}_2 v$$



$$\left(\mathbf{k} = \frac{\mathbf{k}_1 - \mathbf{k}_2}{2}, \mathbf{k}' = \frac{\mathbf{k}_1 - \mathbf{k}_2}{2} + \mathbf{q} \right)$$

2nd order correction (symmetric & neutron matter)

$$\frac{E}{A} = \underbrace{\frac{E^{(0)}}{A}}_{\text{mean field}} + \underbrace{\frac{E^{(2)}}{A}}_{2^{\text{nd}} \text{ order}} + \dots$$

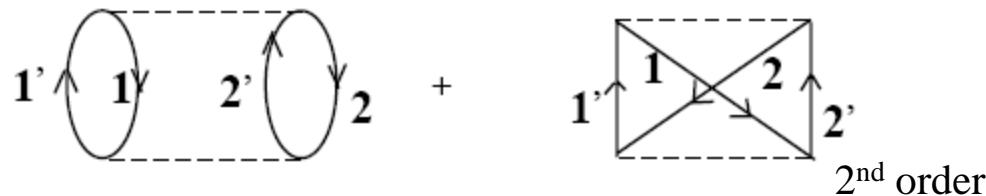
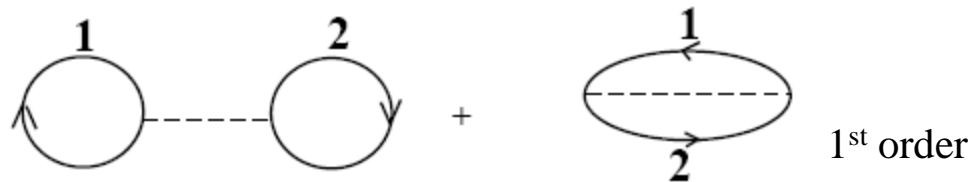
$$\frac{E_{\text{sym}}^{(2)}}{A} = \frac{-3m^*}{64\pi k_F^3 (2\pi)^6} \sum_{S,T} (2T+1)(2S+1) \int_{C_I} d^3\mathbf{k}_1 \int d^3\mathbf{k}_2 \int d^3\mathbf{q} [\mathbf{v}G\mathbf{v}]$$

$$G = \frac{1}{q^2 + \mathbf{q} \cdot (\mathbf{k}_1 - \mathbf{k}_2)}$$

Contour of integral (C_I):

$$|\mathbf{k}_{1,2}| \in [0, k_{F_{1,2}}]$$

$$|\mathbf{k}_1 + \mathbf{q}| > k_{F_1}, |\mathbf{k}_2 - \mathbf{q}| > k_{F_2}$$



Previous attempts

1. t_0 - t_3 model, done by: K. Moghrabi, M. Grasso, G. Colo, and N.V. Giai, Phys. Rev. Lett. 105, 262501 (2010).

2. Full Skyrme (no spin-orbit):

K. Moghrabi, M. Grasso, G. Colo, X. Roco-Maza, Phys. Rev. C 85, 044323 (2012).

K. Moghrabi, M. Grasso, Phys. Rev. C 86, 044319 (2012).



Some mistakes

3. Full Skyrme (no density-dep.): N.Kaiser, J. Phys. G 42,095111(2015)

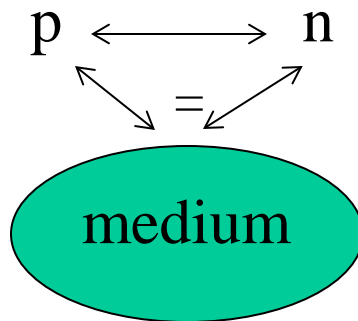
Details: partial-waves mixing

In general, need to sum over contribution from nn, pp and np.

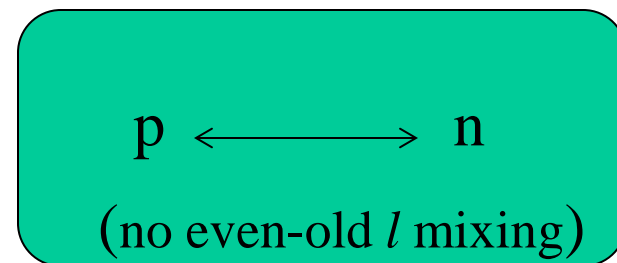
If $k_{F1} \neq k_{F2}$, G no longer symmetric under the contour C_I !

Thus, $\int_{C_I} d^3\mathbf{q} d^3\mathbf{k}_1 d^3\mathbf{k}_2 [v_l G v_{l'}] \neq 0. \Rightarrow$ In general any partial - waves can mix!

Symmetric or pure n: $k_{F1} = k_{F2}$



Reduce to



Through 2nd order effect, medium (G) act as 3rd particle, but does not destroy the original parity of the interaction.

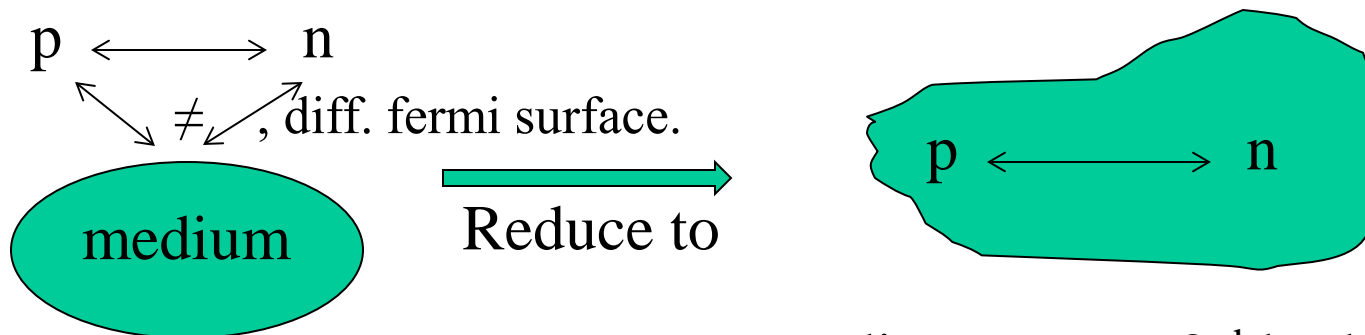
Details: partial-waves mixing

In asymmetric matter, need to sum over contribution from nn, pp and np.

For np part, $k_{F1} \neq k_{F2}$, G no longer symmetric under the contour C_I !

Thus, $\int_{C_I} d^3\mathbf{q} d^3\mathbf{k}_1 d^3\mathbf{k}_2 [v_l G v_l] \neq 0. \Rightarrow$ any partial - waves can mix!

Asymmetric case: $k_{F1} \neq k_{F2}$



medium acts as 3rd body (by G),
and *mixes even-odd partial-wave*.

Results for nuclear matter

In agreement with N.Kaiser, J. Phys. G 42,095111(2015)

$$\frac{\Delta E_{sym(l=0)}^{(2)}}{A} = -\frac{mk_F^4}{110880\hbar^2\pi^4} \left\{ \begin{array}{l} \left[\begin{array}{l} -6534 + 1188\ln[2] + 3564\lambda - 19602\lambda^3 - 5940\lambda^5 \\ + (1782 - 20790\lambda^4)\ln[\frac{\lambda-1}{\lambda+1}] \\ + (24948\lambda^5 - 5940\lambda^7)\ln[\frac{\lambda^2-1}{\lambda^2}] \end{array} \right] \tilde{T}_{03}^2 \\ - \left[\begin{array}{l} 14696 + 2112\ln[2] + 5280\lambda - 2860\lambda^3 \\ -48840\lambda^5 - 18480\lambda^7 + (2640 - 55440\lambda^6)\ln[\frac{\lambda-1}{\lambda+1}] \\ + (71280\lambda^7 - 18480\lambda^9)\ln[\frac{\lambda^2-1}{\lambda^2}] \end{array} \right] k_F^2 \tilde{T}_{03} \tilde{T}_1 \\ + \left[\begin{array}{l} -9886 + 1128\ln[2] + 2520\lambda + 147\lambda^3 - 3654\lambda^5 \\ -35280\lambda^7 - 15120\lambda^9 + (1260 - 41580\lambda^8)\ln[\frac{\lambda-1}{\lambda+1}] \\ + (55440\lambda^9 - 15120\lambda^{11})\ln[\frac{\lambda^2-1}{\lambda^2}] \end{array} \right] k_F^4 \tilde{T}_1^2 \end{array} \right\} \quad \text{Diverge as } \Lambda^5$$

$$\frac{\Delta E_{sym(l=1)}^{(2)}}{A} = -\frac{mk_F^8}{73920\hbar^2\pi^4} \left\{ \left[\begin{array}{l} -1033 + 156\ln[2] + 420\lambda + 140\lambda^3 - 840\lambda^5 \\ -5880\lambda^7 - 2520\lambda^9 + (-210 + 6930\lambda^8)\ln[\frac{\lambda-1}{\lambda+1}] \\ + (9240\lambda^9 - 2520\lambda^{11})\ln[\frac{\lambda^2-1}{\lambda^2}] \end{array} \right] \tilde{T}_2^2 \right\},$$

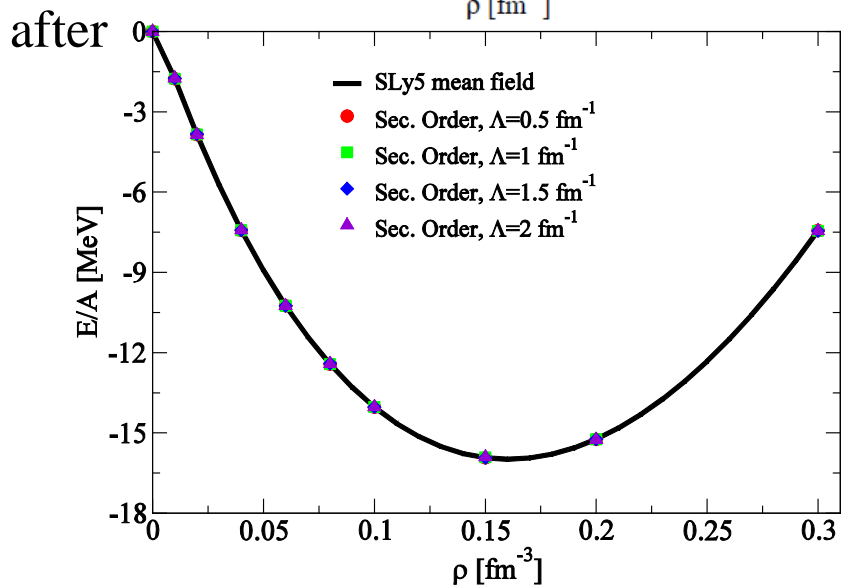
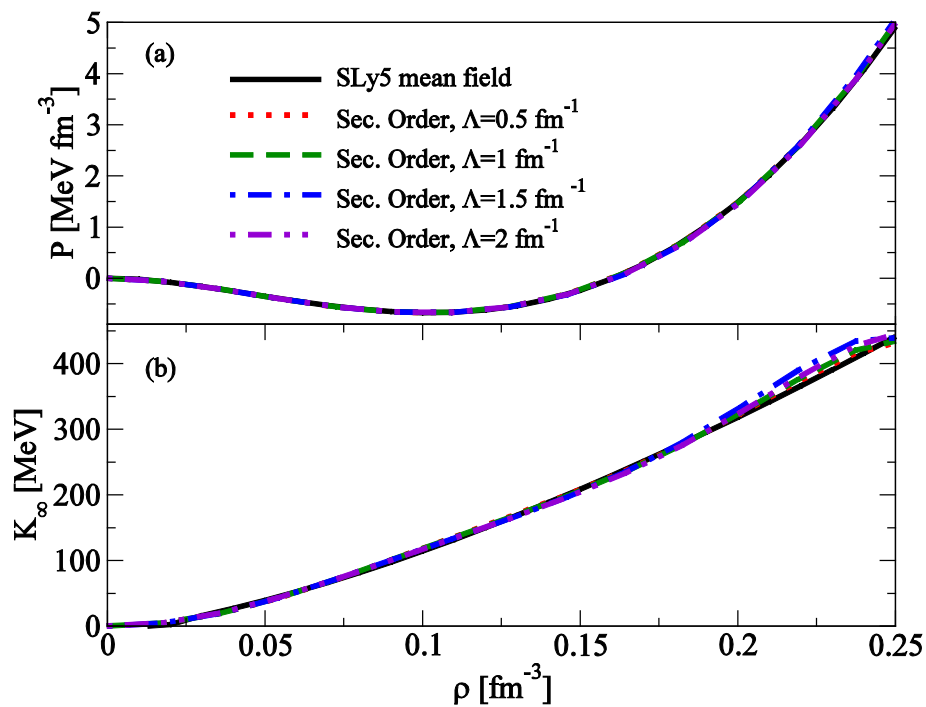
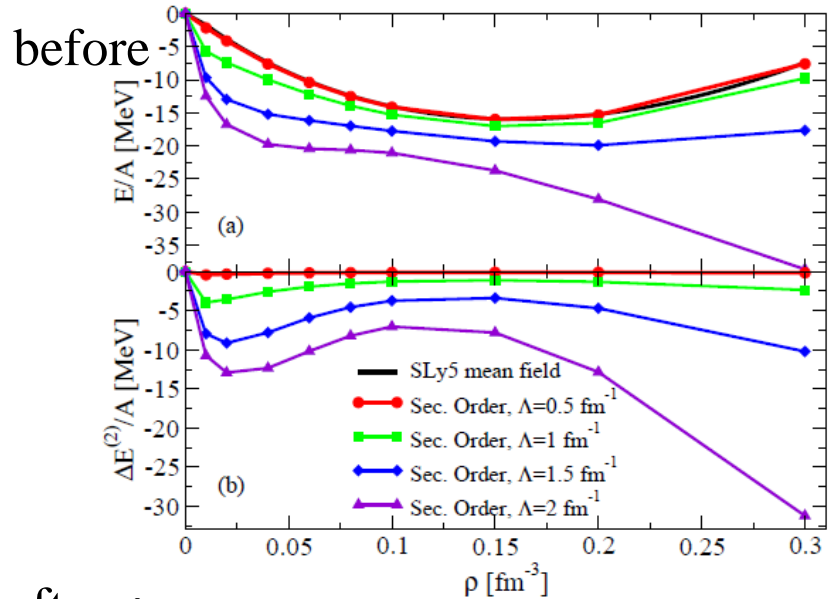
$$\frac{\Delta E_{neutr(l=0)}^{(2)}}{A} = -\frac{mk_{F_N}^4}{166320\hbar^2\pi^4} \left\{ \begin{array}{l} \left[\begin{array}{l} -6534 + 1188\ln[2] + 3564\lambda - 19602\lambda^3 - 5940\lambda^5 \\ + (1782 - 20790\lambda^4)\ln[\frac{\lambda-1}{\lambda+1}] \\ + (24948\lambda^5 - 5940\lambda^7)\ln[\frac{\lambda^2-1}{\lambda^2}] \end{array} \right] T_{03}^2 \\ - \left[\begin{array}{l} 14696 + 2112\ln[2] + 5280\lambda - 2860\lambda^3 \\ -48840\lambda^5 - 18480\lambda^7 + (2640 - 55440\lambda^6)\ln[\frac{\lambda-1}{\lambda+1}] \\ + (71280\lambda^7 - 18480\lambda^9)\ln[\frac{\lambda^2-1}{\lambda^2}] \end{array} \right] k_{F_N}^2 T_{03} T_1 \\ + \left[\begin{array}{l} -9886 + 1128\ln[2] + 2520\lambda + 147\lambda^3 - 3654\lambda^5 \\ -35280\lambda^7 - 15120\lambda^9 + (1260 - 41580\lambda^8)\ln[\frac{\lambda-1}{\lambda+1}] \\ + (55440\lambda^9 - 15120\lambda^{11})\ln[\frac{\lambda^2-1}{\lambda^2}] \end{array} \right] k_{F_N}^4 T_1^2 \end{array} \right\} \quad \text{Diverge as } \Lambda^5$$

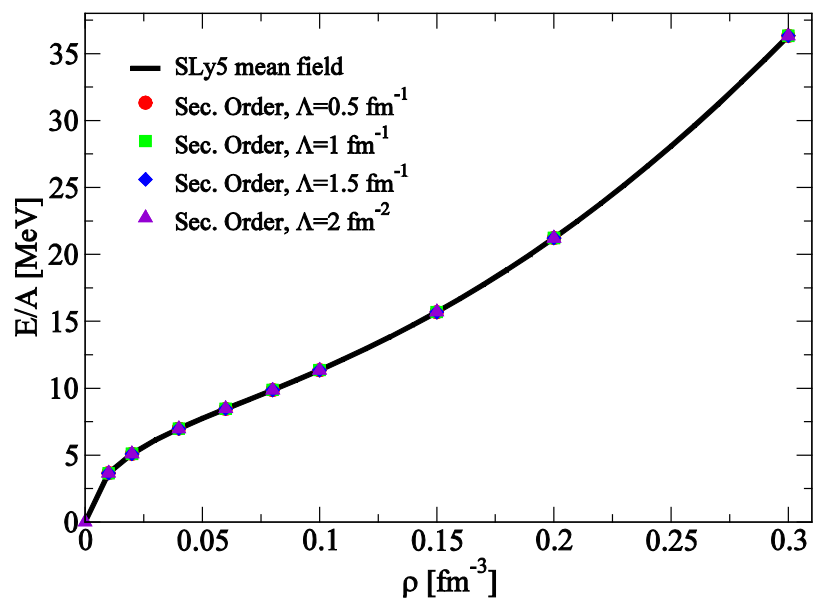
$$\frac{\Delta E_{neutr(l=1)}^{(2)}}{A} = -\frac{mk_{F_N}^8}{110880\hbar^2\pi^4} \left\{ \left[\begin{array}{l} -1033 + 156\ln[2] + 420\lambda + 140\lambda^3 - 840\lambda^5 \\ -5880\lambda^7 - 2520\lambda^9 + (-210 + 6930\lambda^8)\ln[\frac{\lambda-1}{\lambda+1}] \\ + (9240\lambda^9 - 2520\lambda^{11})\ln[\frac{\lambda^2-1}{\lambda^2}] \end{array} \right] T_2^2 \right\},$$

Renormalization I

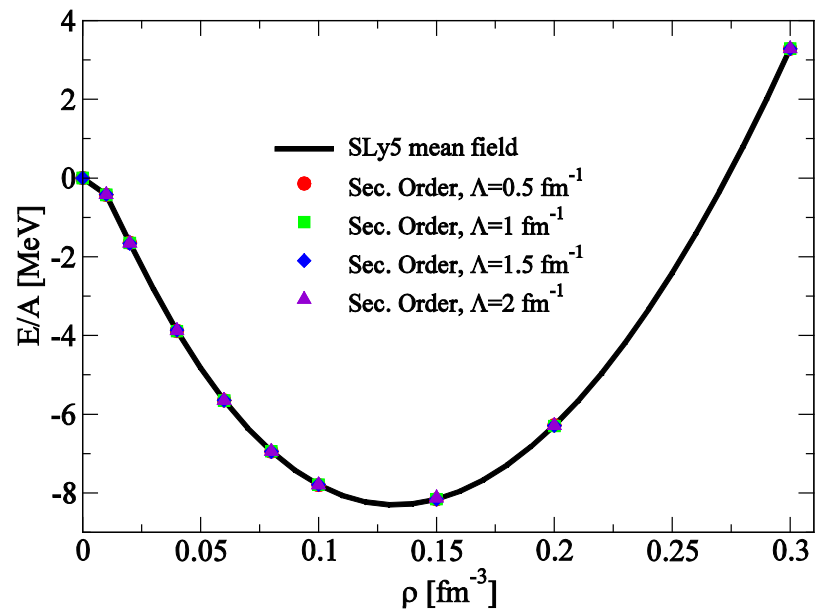
- Pick a Λ , and simply re-adjust the 9 (skyrme) parameters.

Symmetric matter



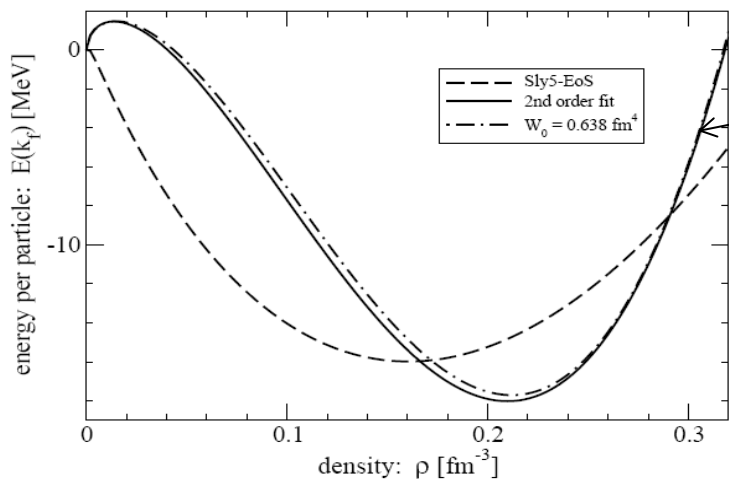


Pure neutron matter

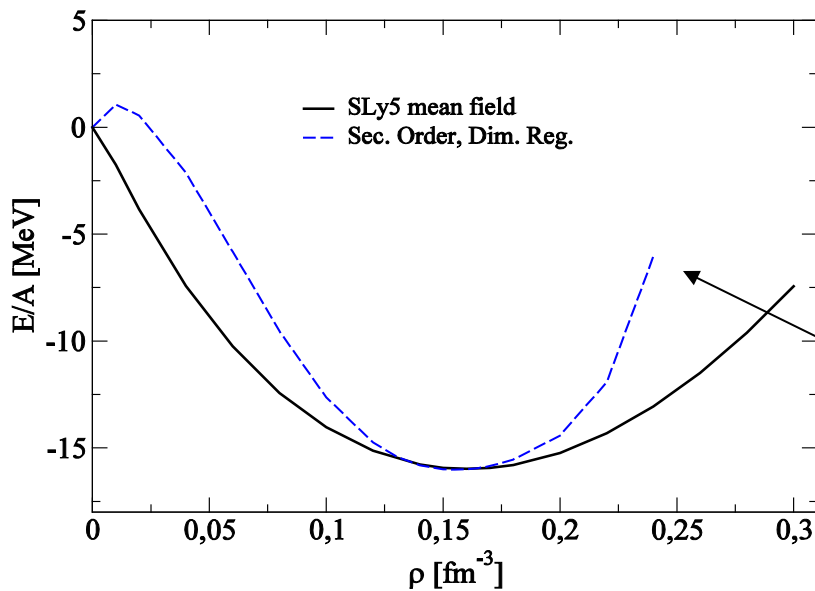
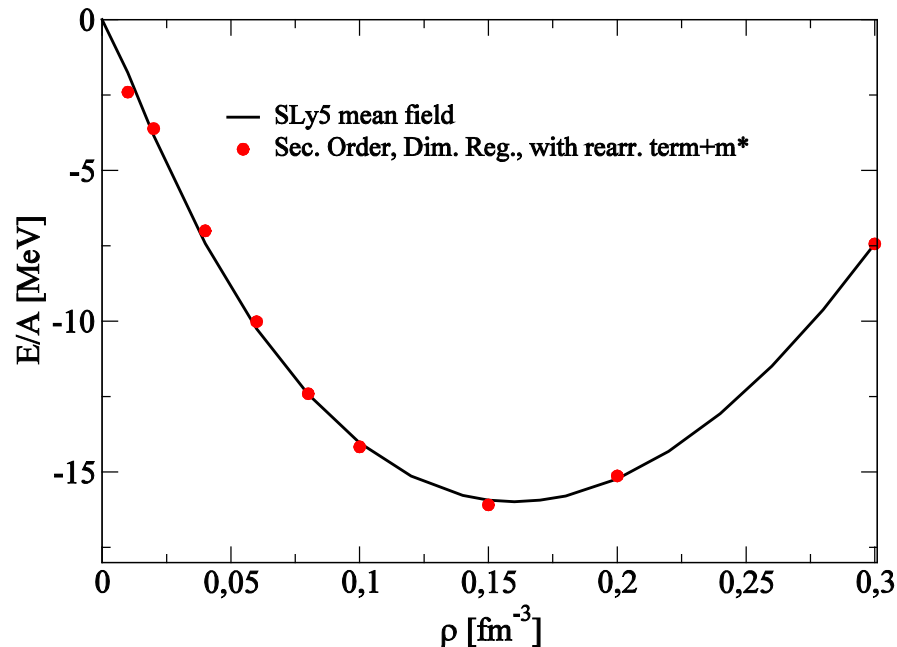


Asymmetric matter ($\delta=0.5$)

Keep only the finite part (Dimensional regularization)

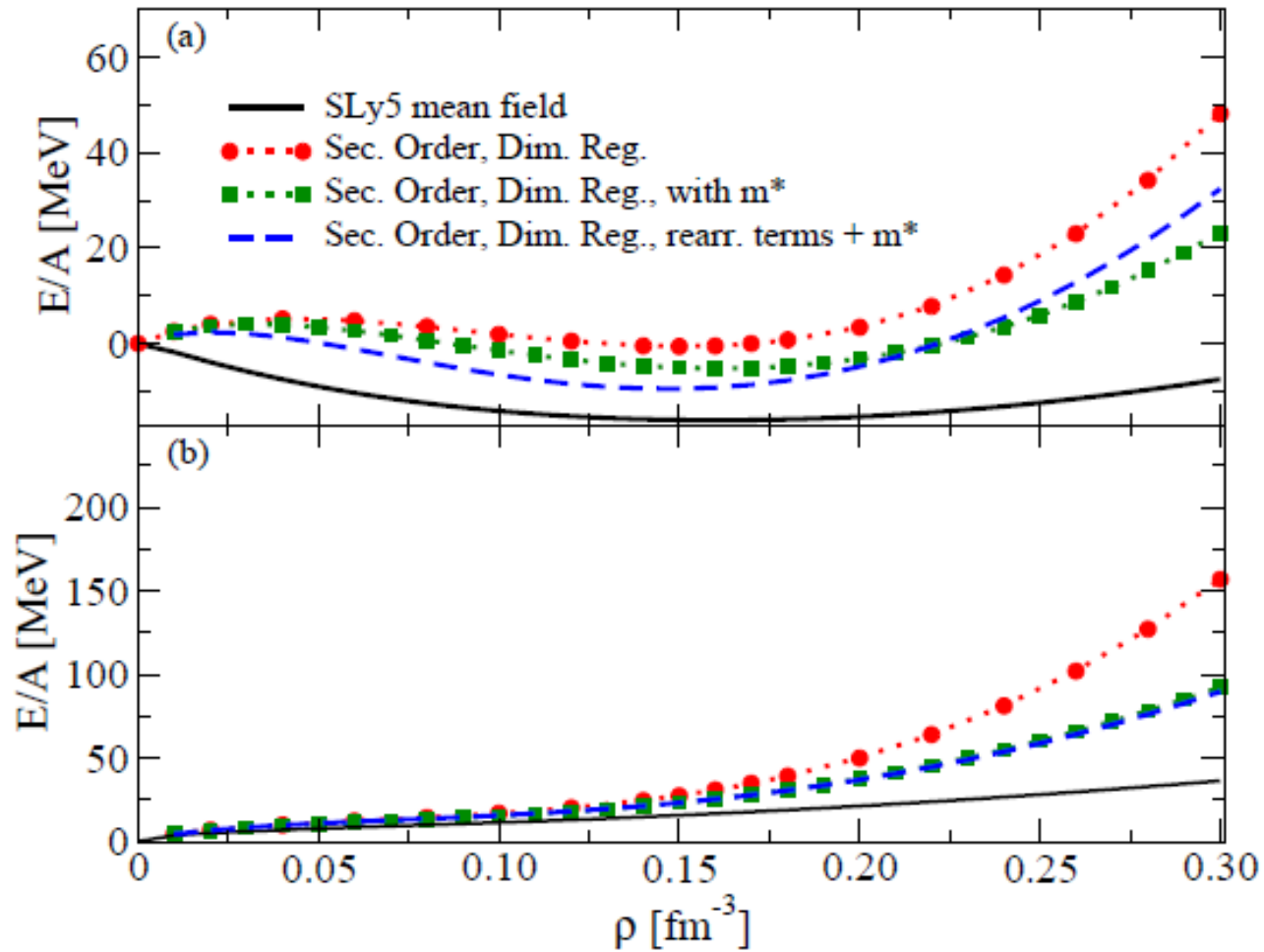


Without density-dep. (t_3) term, Kaiser 2015

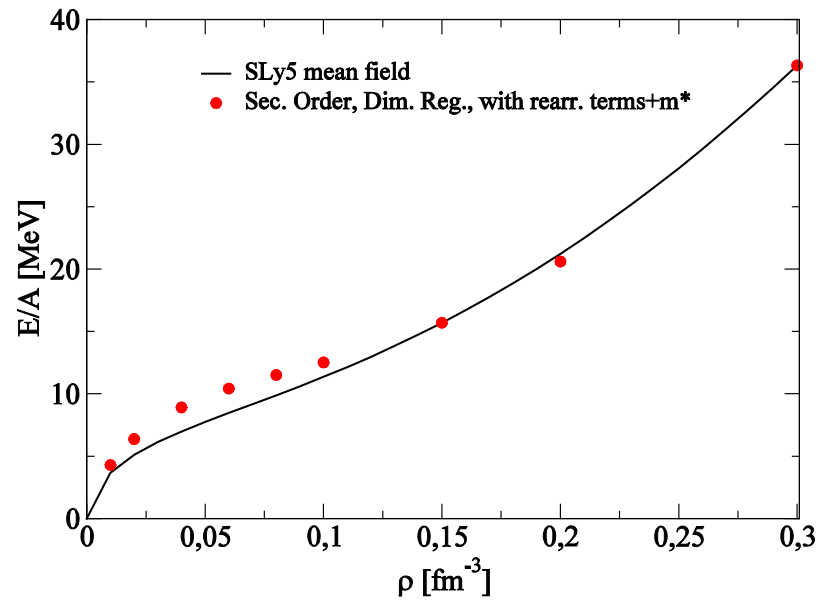
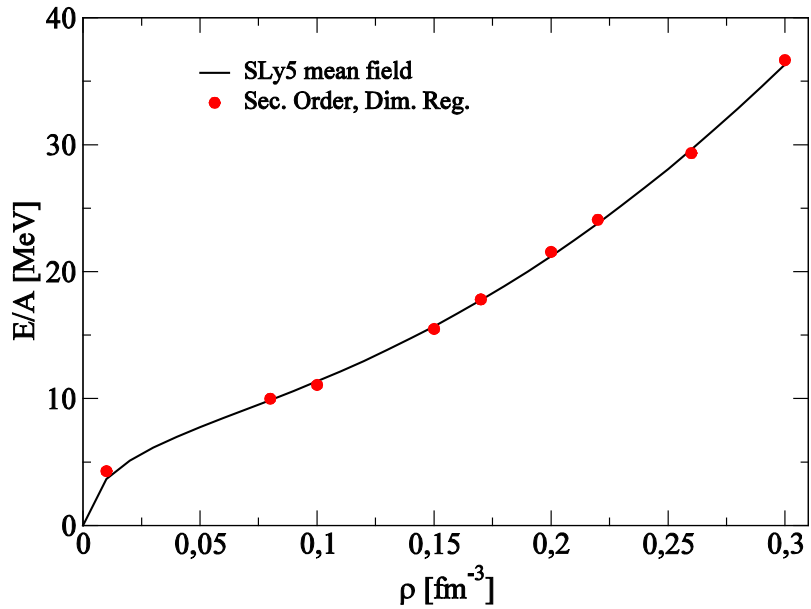


With density-dep., without rearrangement term

Before renormalization



Pure neutron matter (DR)



PART II:
RENORMALIZABILITY

- When $\Lambda \rightarrow \infty$, how the 2nd order terms behaves?

$$\frac{\Delta E_f^{(2)}(k_F)}{A} = \frac{3m}{2\pi^4 \hbar^2} k_F^4 [A_0 + A_1 T_3 k_F^{3\alpha} + A_2 T_3^2 k_F^{6\alpha} + A_3 k_F^2 + A_4 T_3 k_F^{2+3\alpha} + A_5 k_F^4], \quad \text{Converge terms}$$

$$\frac{\Delta E_a^{(2)}(k_F, \lambda)}{A} = -\frac{m}{8\pi^4 \hbar^2} \lambda k_F^3 [B_0(\lambda) + B_1(\lambda) T_3 k_F^{3\alpha} + B_2(\lambda) k_F^2], \quad \text{Diverge, } k_F\text{-dep appears in MF}$$

$$\frac{\Delta E_d^{(2)}(k_F, \lambda)}{A} = -\frac{m}{8\pi^4 \hbar^2} \lambda k_F^3 [C_0 T_3^2 k_F^{6\alpha} + C_1 T_3 k_F^{2+3\alpha} + C_2 k_F^4], \quad \text{Diverge, } k_F\text{-dep } \textit{not} \text{ in MF}$$

- Idea: Absorb the Λ -divergence in 2nd order into mean field terms with the same k_F -dependence.

$$\frac{\Delta E_f^{(2)}(k_F)}{A} = \frac{3m}{2\pi^4 \hbar^2} k_F^4 [A_0 + A_1 T_3 k_F^{3\alpha} + A_2 T_3^2 k_F^{6\alpha} + A_3 k_F^2 + A_4 T_3 k_F^{2+3\alpha} + A_5 k_F^4], \quad \text{converge}$$

$$\frac{\Delta E_a^{(2)}(k_F, \lambda)}{A} = -\frac{m}{8\pi^4 \hbar^2} \lambda k_F^3 [B_0(\lambda) + B_1(\lambda) T_3 k_F^{3\alpha} + B_2(\lambda) k_F^2], \quad \text{Diverge, } k_F\text{-dep appears in MF}$$

~~$$\frac{\Delta E_j^{(2)}(k_F, \lambda)}{A} = \frac{m}{8\pi^4 \hbar^2} \lambda k_F^3 [C_0 T_3^2 k_F^{6\alpha} + C_1 T_3 k_F^{2+3\alpha} + C_2 k_F^4], \quad \text{Diverge, } k_F\text{-dep } \textit{not} \text{ in MF}$$~~



eliminate by setting $\alpha=1/3$ and $t_1=t_2=0$, or setting $t_1=t_2=t_3=0$.

- Idea: Absorb the Λ -divergence in 2nd order into mean field terms with the same k_F -dependence.

$$\frac{\Delta E_f^{(2)}(k_F)}{A} = \frac{3m}{2\pi^4\hbar^2} k_F^4 [A_0 + A_1 T_3 k_F^{3\alpha} + A_2 T_3^2 k_F^{6\alpha} + A_3 k_F^2 + A_4 T_3 k_F^{2+3\alpha} + A_5 k_F^4], \quad \text{converge}$$

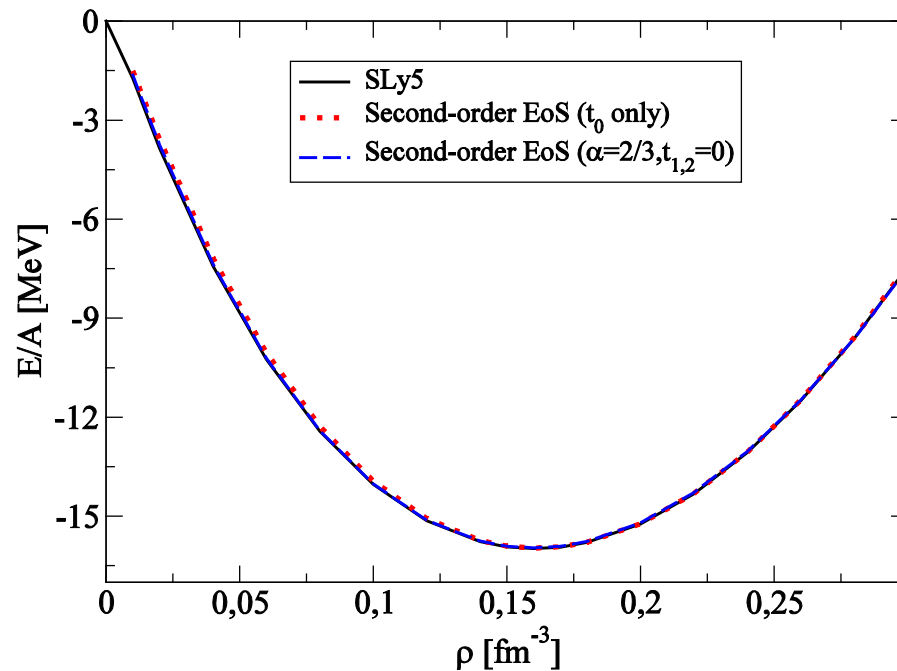
$$\frac{\Delta E_a^{(2)}(k_F, \lambda)}{A} = -\frac{m}{8\pi^4\hbar^2} \lambda k_F^3 [B_0(\lambda) + B_1(\lambda) T_3 k_F^{3\alpha} + B_2(\lambda) k_F^2], \quad \text{Diverge, } k_F\text{-dep appears in MF}$$



- Treatment 1: Absorb divergence into redefinition of parameters.
 Treatment 2: Add counter terms correspond to each divergence.

Results

- Treatment 1 doesn't work (cannot obtain reasonable fit).
- Treatment 2 works.



Lesson

- The leading order quite possible just contains only t_0 - t_3 terms.
- However, the regulator dependence tells us the power counting cannot be established in this way.

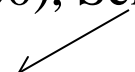
Part III: Matching the low density limit

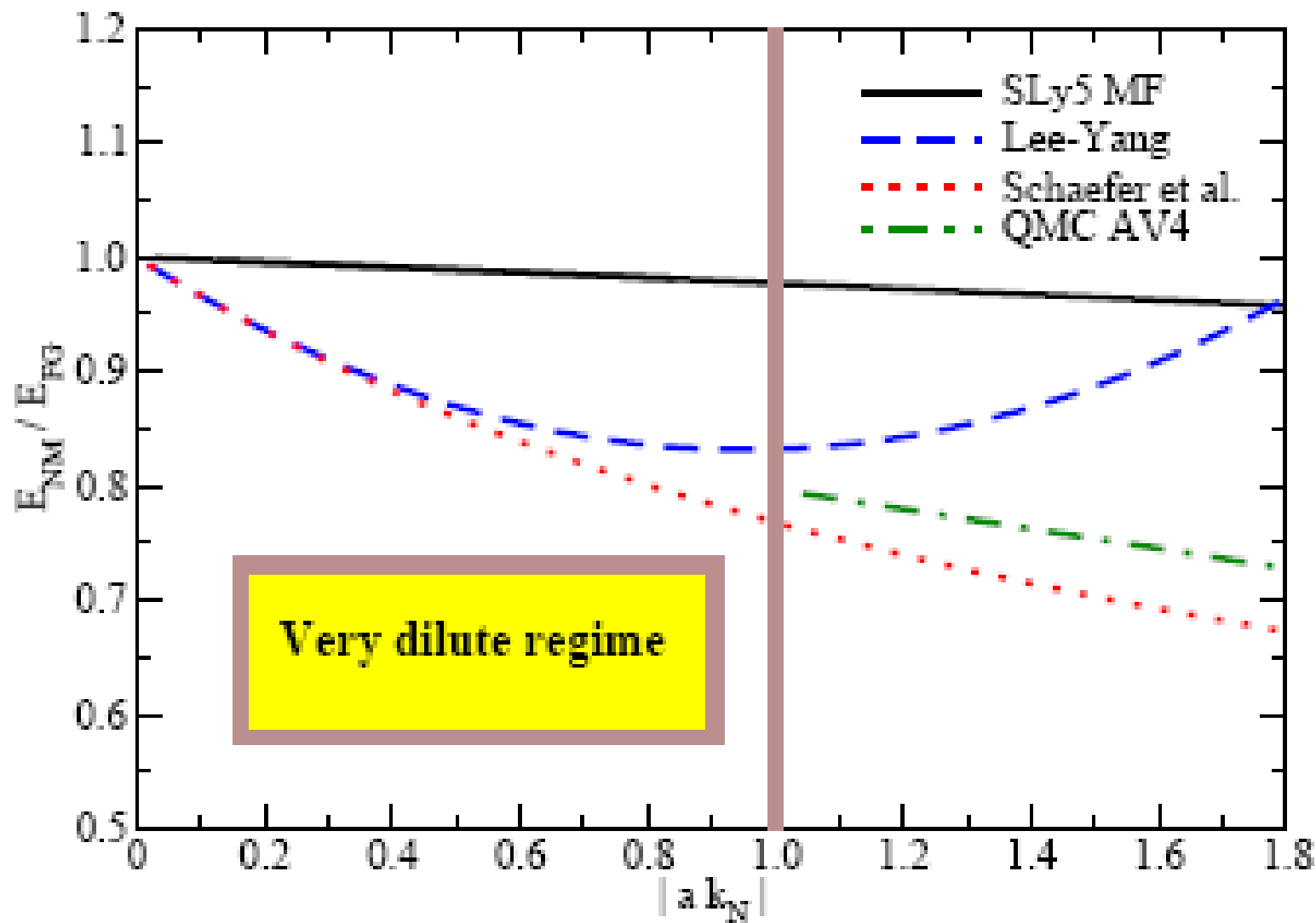
Neutron matter at very low- ρ

Lee & Yang formula (1957) describes the dilute system.

$$\frac{E_{NM}}{A} = \frac{\hbar^2 k_N^2}{2m} \left[\underbrace{\frac{3}{5}}_{K.E.} + \underbrace{\frac{2}{3\pi}(k_N a)}_{\text{fixed to } t_0 \text{ term}} + \underbrace{\frac{4}{35}(11-2\ln 2)(k_N a)^2}_{\text{automatically recover in 2}^{\text{nd}} \text{ of } t_0} + \underbrace{O(k_N^3)}_{\text{higher order}} \right]$$

- The 2nd order EoS automatically recover the $(11-2\ln 2)(k_N a)^3$ term.
- If take physical value of $a = -18.9$ fm, then *impossible* to fit pure neutron matter EoS outside region $k_N a \ll 1$ (adding t_1, t_2, t_3 terms doesn't help).
- $(k_N a)$ needs to be re-summed. (Steele (2000), Schafer (2005), Kaiser (2011))

$$\frac{E_{NM}}{N} = \frac{\hbar^2 k_N^2}{2m} \left[\frac{3}{5} + \frac{2}{3\pi} \frac{k_N a}{1 - 6k_N a(11 - 2\ln 2)/(35\pi)} \right]$$




This work: Resumed-inspired functional

$$V = \frac{B_\beta}{1 - R_\beta \rho^{1/3} + \underbrace{C_\beta \rho^{2/3}}_{\text{higher order in L\&Y to be resumed}^*}} + \underbrace{D_\beta \rho^{2/3}}_{\text{velocity-dep term}^*} + \underbrace{F_\beta \rho^\alpha}_{3^+ \text{-body}}$$

B_β, R_β are fixed to reproduce first two term in Lee & Yang.

$$\Rightarrow B_\beta = 2\pi \frac{\hbar^2}{m} \frac{v-1}{v} a_\beta, \quad R_\beta = \frac{6}{35\pi} \left(\frac{6\pi^2}{v} \right)^{1/3} (11 - 2\ln 2) a_\beta.$$

(degeneracy: $\nu = 2(4)$ for $\beta = \begin{matrix} \text{0} \\ \text{sym} \end{matrix}$ ($\begin{matrix} \text{1} \\ \text{sym} \end{matrix}$))

$$a_0 = -18.9 \text{ fm}, \quad \underbrace{a_1 = -20 \text{ fm}}_{\text{avg. of } a_{nn}, a_{pp}, a_{np} \text{ in } ^1S_0}.$$

$$\frac{E}{A} = KE_\beta + \frac{B_\beta \rho}{1 - R_\beta \rho^{1/3} + C_\beta \rho^{2/3}} + D_\beta \rho^{5/3} + F_\beta \rho^{\alpha+1}$$

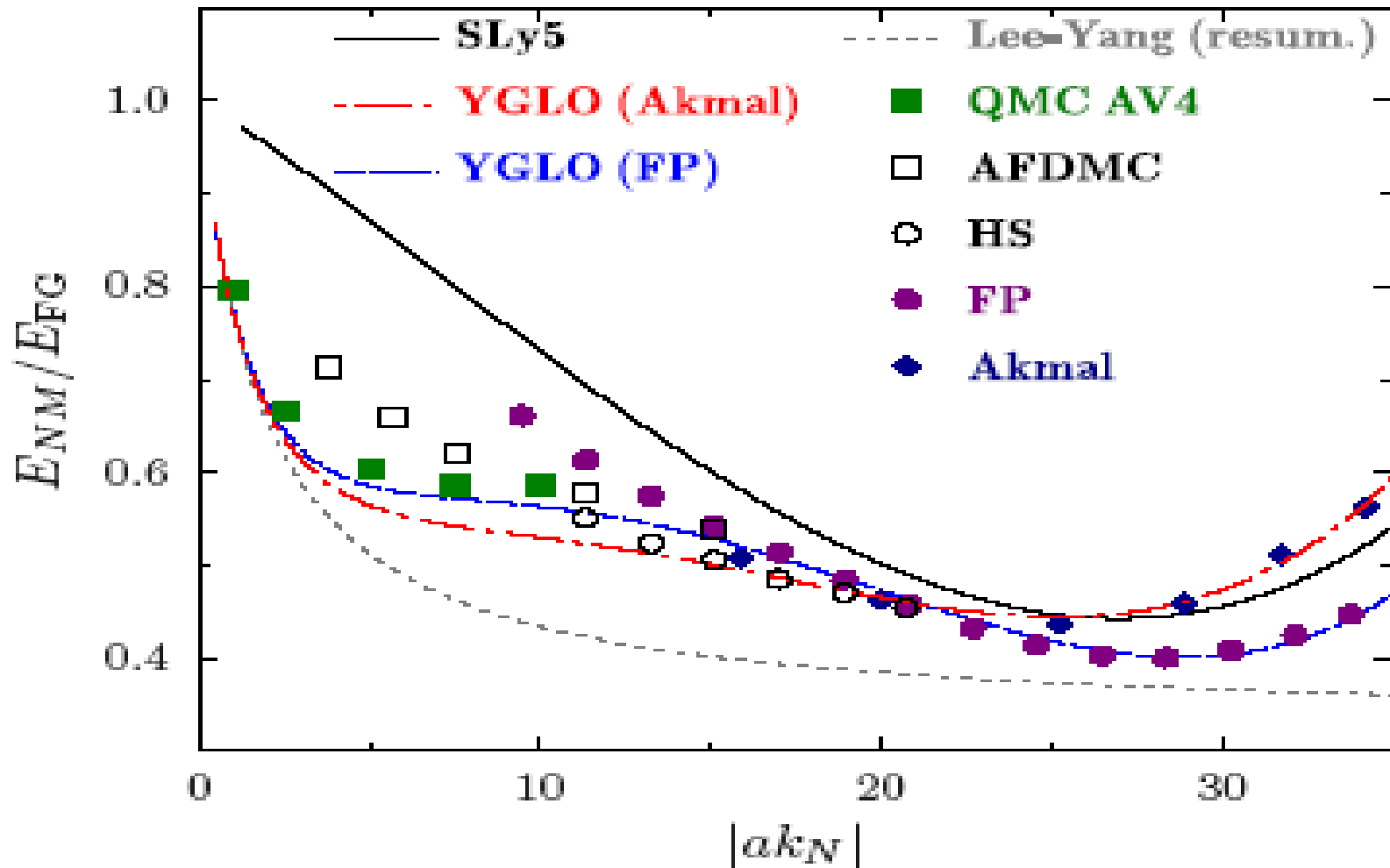
Results

Able to describe pure neutron matter from very-dilute-limit to twice saturation density.

Very dilute limit

$\rho(\text{fm}^{-3})$

0.005 0.0462 0.135

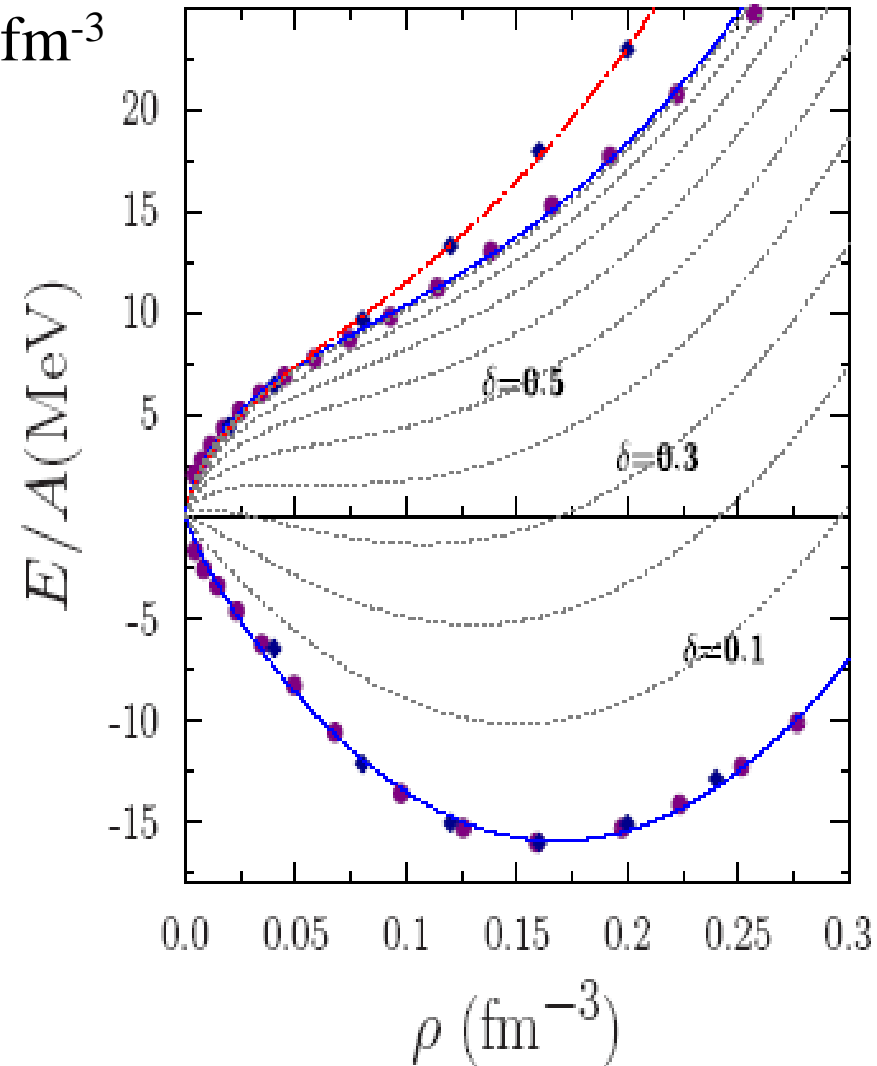


FP: B. Friedman and V. Pandharipande, Nucl. Phys. A361,502 (1981).

Akmal: A. Akmal, V. R. Pandharipande, and D. G. Ravenhall, Phys. Rev. C 58, 1804 (1998).

HS: K. Hebeler and A. Schwenk Phys. Rev. C **82**, 014314(2010).

Up to $\rho=0.3 \text{ fm}^{-3}$



Able to describe both sym and pure neutron matter EoS up to $2\rho_0$ very well with only 4 free parameters each.

Asymmetric case

Parabolic approximation

$$\frac{E_\delta}{A}(\rho) = \frac{E_{sym}}{A}(\rho) + S(\rho)\delta^2,$$

$$(\delta = (\rho_N - \rho_p) / (\rho_N + \rho_p))$$

$$L = 3\rho_0 (dS / d\rho)_{\rho=\rho_0}$$

Before: Lots of models fail

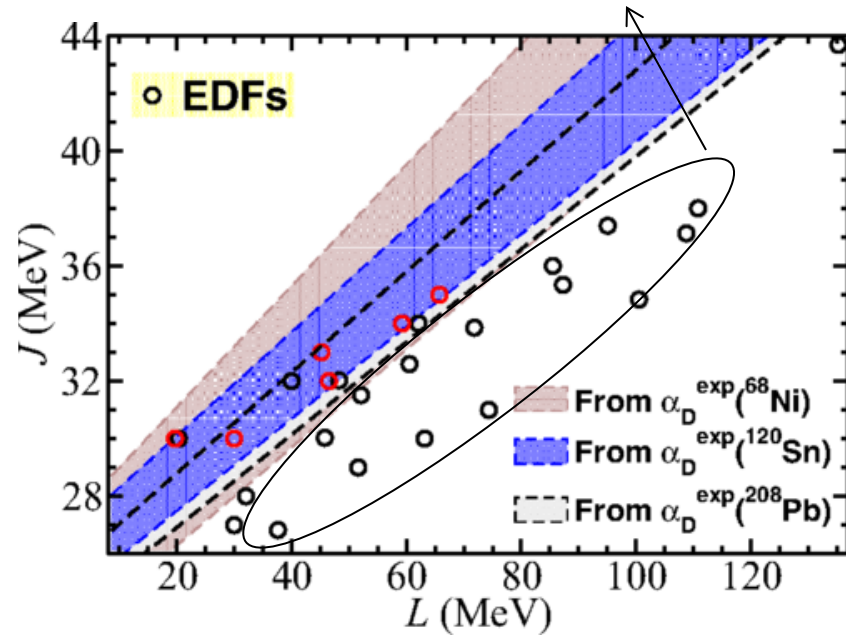


FIG. 4: Symmetry energy at saturation density as a function of its slope L . The black lines delimit the phenomenological area constrained by the experimental determination of the electric dipole polarizability in ^{208}Pb . The blue dotted lines delimit the area constrained by the same measurement in ^{68}Ni , and the red dashed lines refer to the measurement done in ^{120}Sn . The yellow area is the overlap. Inset: density dependence of the Symmetry energy for the two YGLO parametrizations of this work.

X. Roca-Maza, et al., (2015).

Asymmetric case

Our result (prediction)

Satisfies the experimental constraint.

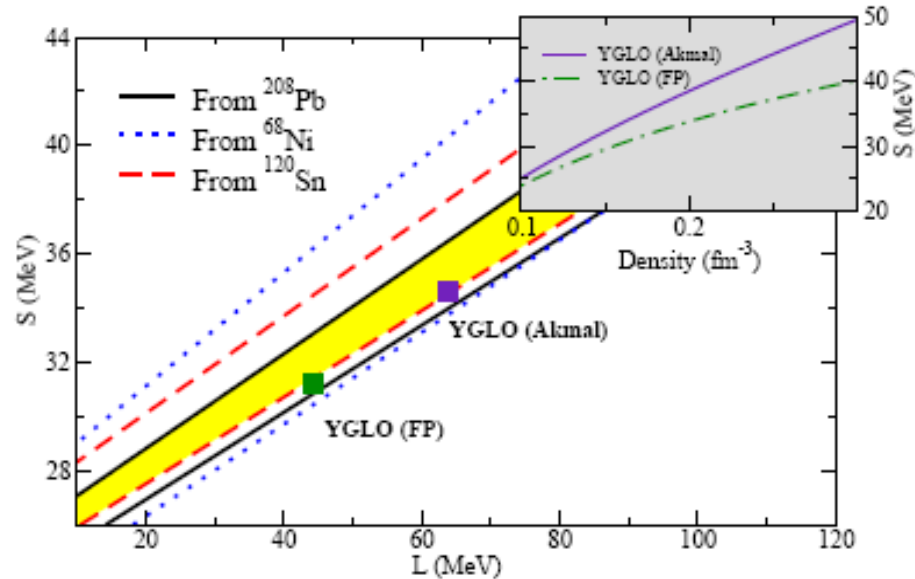


FIG. 4: Symmetry energy at saturation density as a function of its slope L . The black lines delimit the phenomenological area constrained by the experimental determination of the electric dipole polarizability in ^{208}Pb . The blue dotted lines delimit the area constrained by the same measurement in ^{68}Ni , and the red dashed lines refer to the measurement done in ^{120}Sn . The yellow area is the overlap. Inset: density dependence of the Symmetry energy for the two YGLO parametrizations of this work.

Summary and conclusion

- 2nd order contribution is considered, and with the rearrangement term included, results can be refitted to standard data very well in both cutoff- and dimensional-regularization.
- Renormalizability suggests that the leading order(MF) likely to be a t_0 - t_3 model (to avoid a rapid grows of counter terms).
- Resuming the t_0 part of interaction is necessary to describe low ρ limit correctly.

Future prospects

Try to bridge EFT ideas/techniques to mean field (and beyond) within EDF framework.

Mean field with potential models (effective interaction).
(e.g., Skyrme-type)

2nd order corrections

Add new effective interactions?

What is the proper form of it?

Higher order corrections

Is the improvement systematic?

Renormalization-group
analysis

+

power counting check

Goal:

Systematic treatment of the
interactions.

Thank you!

Simplification for symmetric and pure neutron matter

$$\text{Mean field: } E \propto \int_0^{k_{F1}} d^3\mathbf{k}_1 \int_0^{k_{F2}} d^3\mathbf{k}_2 v(\mathbf{k}', \mathbf{k}).$$

$$2^{\text{nd}} \text{ order correction: } E^{(2)} \propto \int_0^{k_{F1}} d^3\mathbf{k}_1 \int_0^{k_{F2}} d^3\mathbf{k}_2 \int_0^\Lambda v G v.$$

$$\text{Define } \mathbf{k} = \frac{\mathbf{k}_1 - \mathbf{k}_2}{2}, \mathbf{k}' = \frac{\mathbf{k}_1 - \mathbf{k}_2}{2} + \mathbf{q}$$

$$G_{\text{sym,neut}} \rightarrow \frac{-m^*}{k'^2 - k^2}.$$