

# Magnetic Properties of Light Nuclei from Lattice QCD

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**INT-Program 16-1**  
***Nuclear Physics from Lattice QCD***

B C Tiburzi  
18 May 2016

My work funded by



Done in collaboration with ***Nuclear Physics Lattice QCD =***





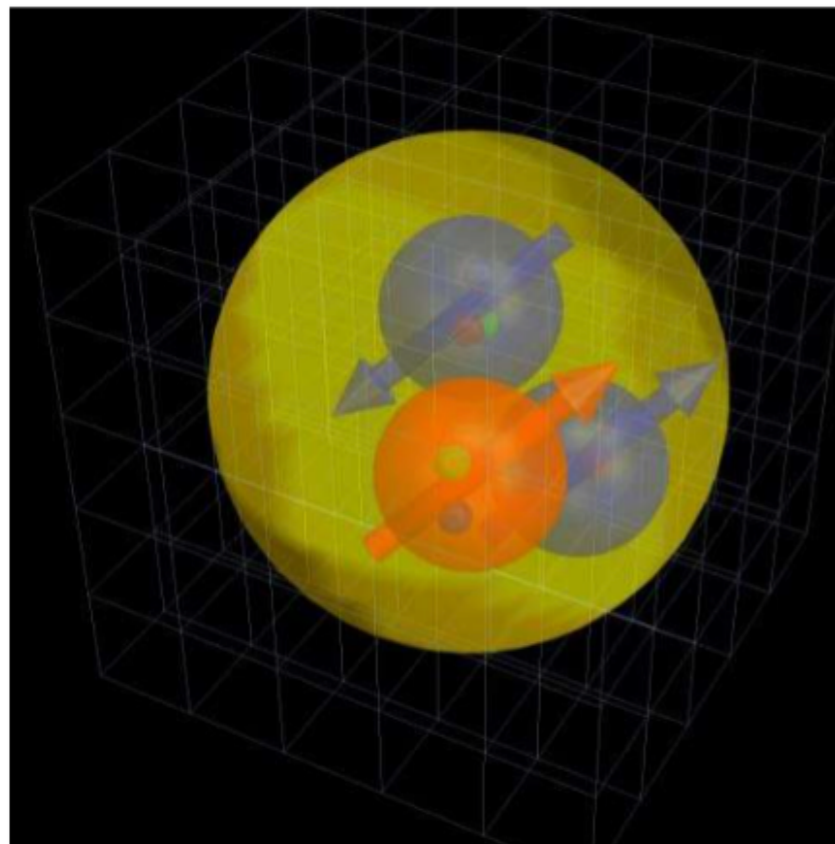
# Magnetic Moments of Light Nuclei

Home Physics General Physics February 2, 2015

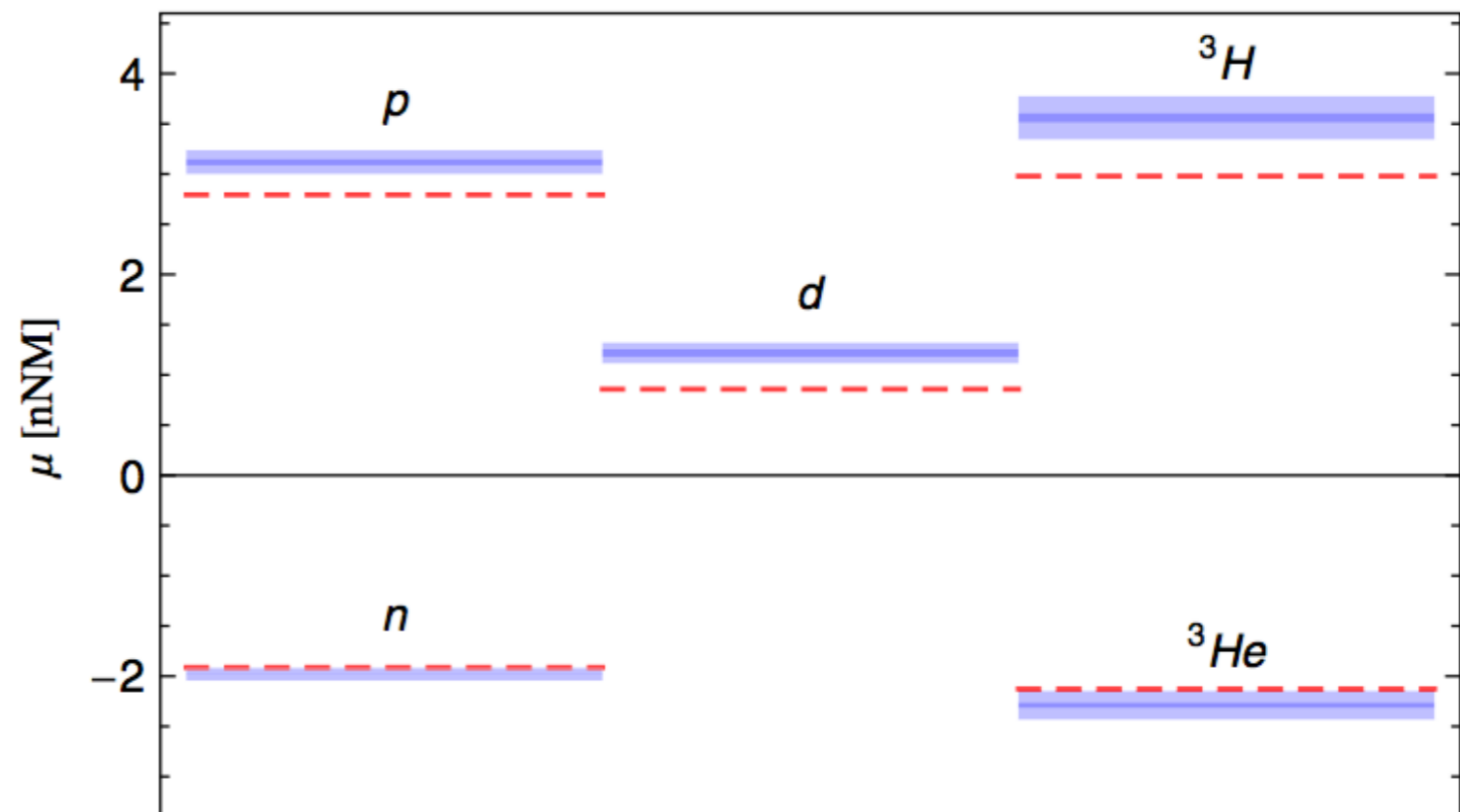


## Pinpointing the magnetic moments of nuclear matter

February 2, 2015 by Kathy Kincaide



Artist's impression of a triton, the atomic nucleus of a tritium atom. The image shows a red neutron with quarks inside; the arrows indicate the alignments of the spins. Credit: William Detmold, MIT



A team of nuclear physicists has made a key discovery in its quest to shed light on the structure and behavior of subatomic particles.

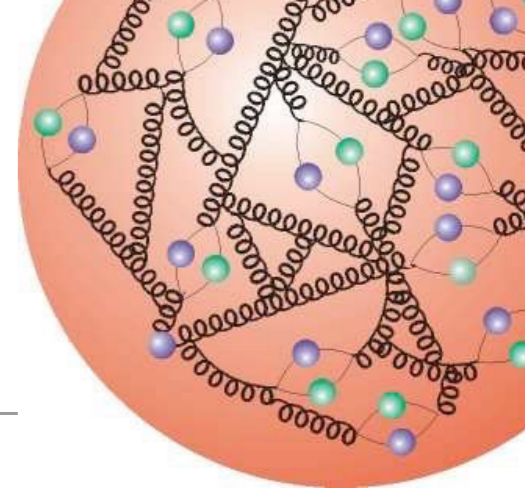
Beane, *et al.* (NPLQCD), Phys.Rev. Lett.113, 2014.

### First Computation:

$$m_u = m_d = (m_s)_{\text{phys}}$$

$$m_\pi \sim 800 \text{ MeV}$$

# Grand Overview



## Electroweak Interactions: Nucleons and Nuclei

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- *Lattice QCD* continues to sharpen our knowledge of The Standard Model (e.g. *CKM* extraction,  $K \rightarrow \pi \pi$ )
- Nucleons and light nuclei present ~~challenge~~ opportunity
- QCD relevant for high-precision low-energy experiments



Magnetic properties of light nuclei



# Quark Interactions to Nuclear Physics



Ken Wilson  
1936-2013

- **Textbook:** gauge theories defined in perturbation theory
- **QCD:** short distance perturbative, long distance non-perturbative

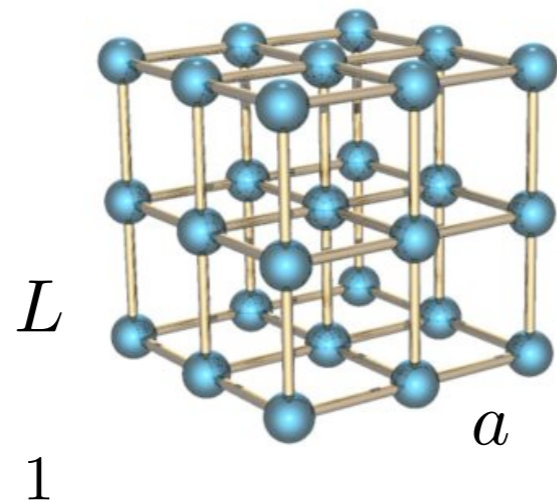
$$\bar{q} (\not{D} + m_q) q + \frac{1}{4} G_{\mu\nu} G_{\mu\nu} \quad \xrightarrow{\text{Many Technicalities}} \quad M_N \quad \delta_{NN}(k) \quad \epsilon_b(D)$$

Non-perturbative definition of asymptotically free gauge theories

Strong interaction observables

One step:  $\int [\mathcal{D}A_\mu] e^{-S_{\text{YM}}(A_\mu)} \approx \frac{1}{N_{\text{cfg}}} \sum_{\{A_\mu\}} e^{-S_{\text{YM}}(A_\mu)}$  stat. evaluation

Another step:



sys. approx.  $U_\mu(x) = e^{igaA_\mu(x)} \in SU(3)$

Quarks:

$$\frac{1}{\not{D}(U_\mu) + m_q}$$

Quark electroweak interactions

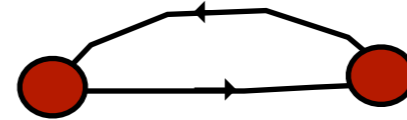
fortunately perturbative ...

$$J_\mu = \bar{q} \gamma_\mu q$$

# Particle Physics (B=0) vs. Nuclear Physics (B>0)

## Pion Correlation Function

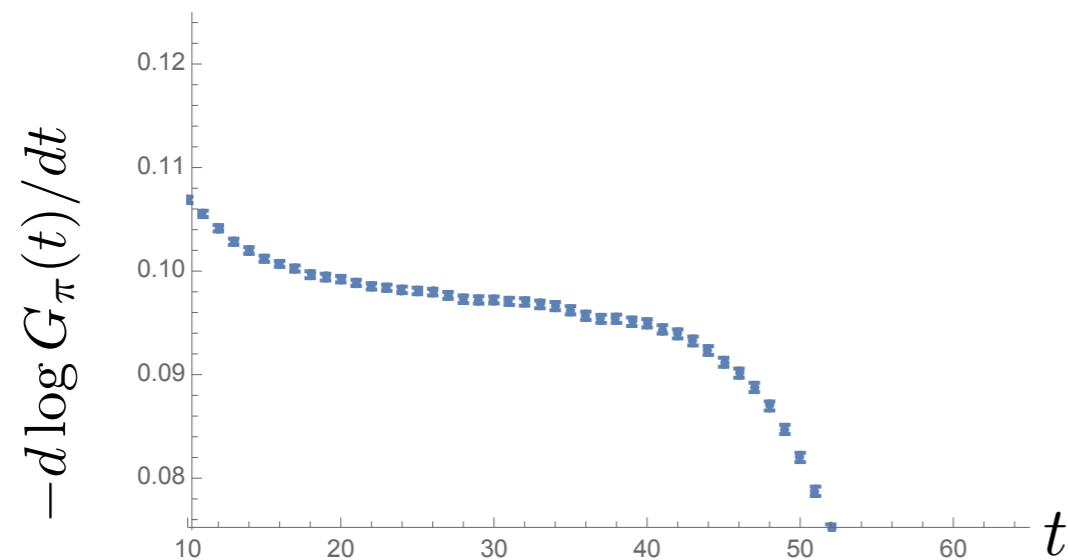
Signal  $\sum_{\{A_\mu\}} \langle q\bar{q}(t)q\bar{q}(0) \rangle \sim e^{-m_\pi t}$



Signal/Noise

Noise<sup>2</sup>  $\sum_{\{A_\mu\}} \langle q\bar{q}(t)q\bar{q}(t)q\bar{q}(0)q\bar{q}(0) \rangle \sim e^{-2m_\pi t}$

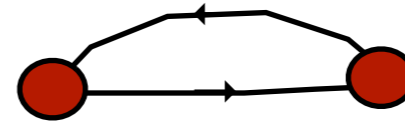
$\sim \text{const}$



# Particle Physics (B=0) vs. Nuclear Physics (B>0)

## Pion Correlation Function

Signal  $\sum_{\{A_\mu\}} \langle q\bar{q}(t)q\bar{q}(0) \rangle \sim e^{-m_\pi t}$



Signal/Noise

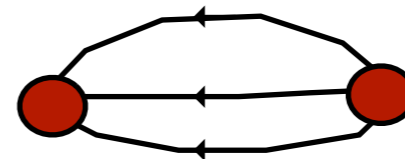
Noise<sup>2</sup>  $\sum_{\{A_\mu\}} \langle q\bar{q}(t)q\bar{q}(t)q\bar{q}(0)q\bar{q}(0) \rangle \sim e^{-2m_\pi t}$

$\sim \text{const}$

Baryons are statistically noisy

## Nucleon Correlation Function

Signal  $\sum_{\{A_\mu\}} \langle qqq(t)\bar{q}\bar{q}\bar{q}(0) \rangle \sim e^{-Mt}$

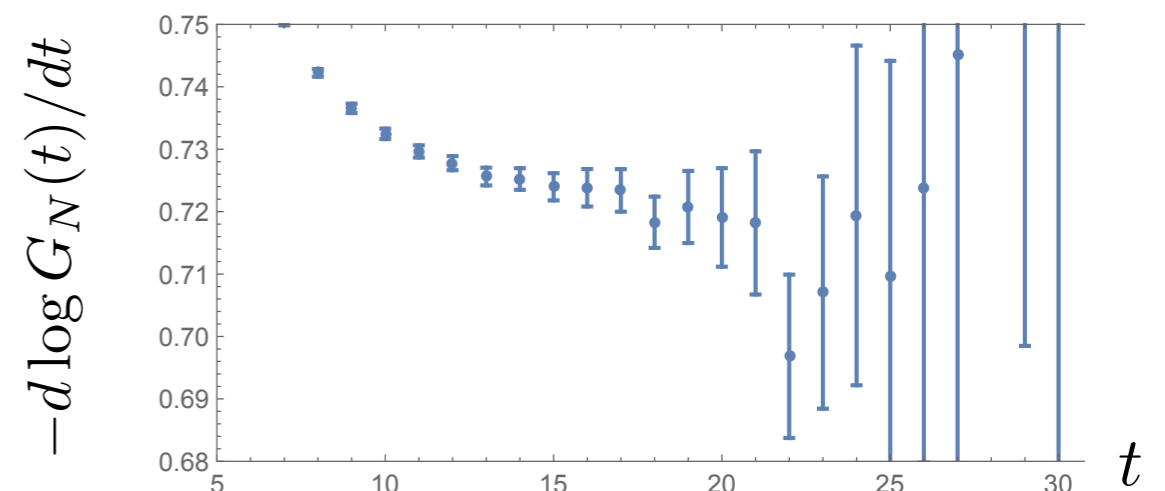
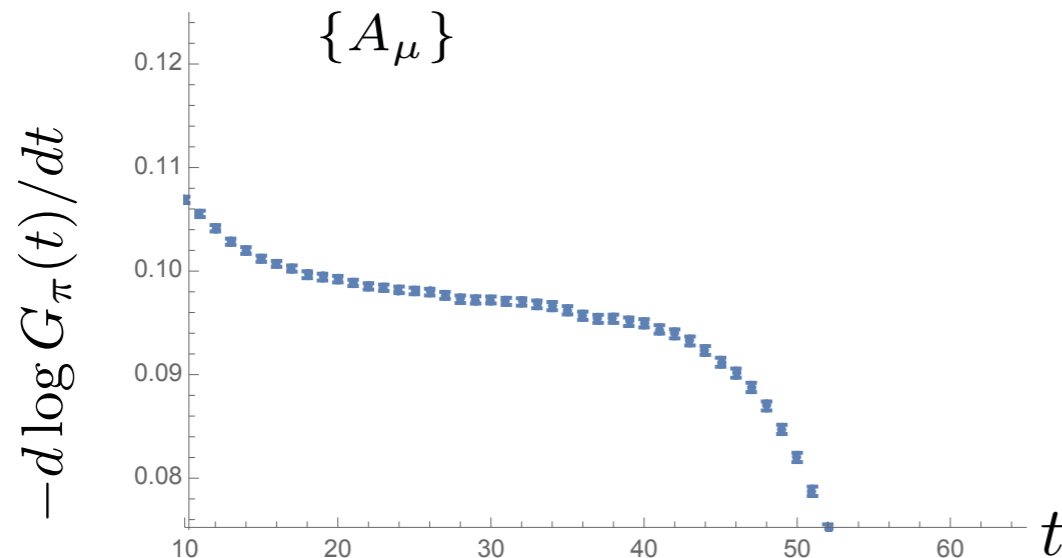


Scales exponentially with B  
in asymptotic time limit

Noise<sup>2</sup>  $\sum_{\{A_\mu\}} \langle qqq(t)\bar{q}\bar{q}\bar{q}(t)qqq(0)\bar{q}\bar{q}\bar{q}(0) \rangle \sim e^{-3m_\pi t}$

Signal/Noise

$\sim e^{-(M - \frac{3}{2}m_\pi)t}$

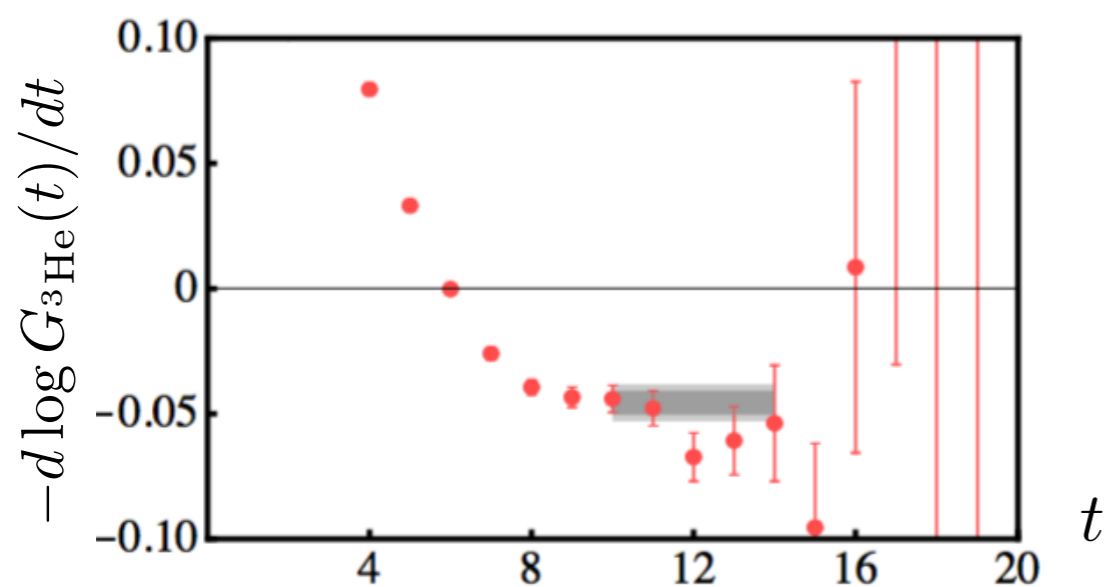




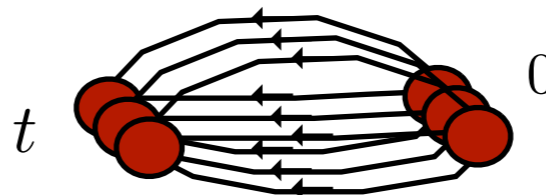
# Nuclear Physics @ $m_\pi = 800$ and 450 MeV

Beane, Chang, Cohen, Detmold, Lin, Luu, Orginos, Parreño, Savage, Walker-Loud **PRD87** (2013)  
 Orginos, Parreño, Savage, Beane, Chang, Detmold **PRD92** (2015)

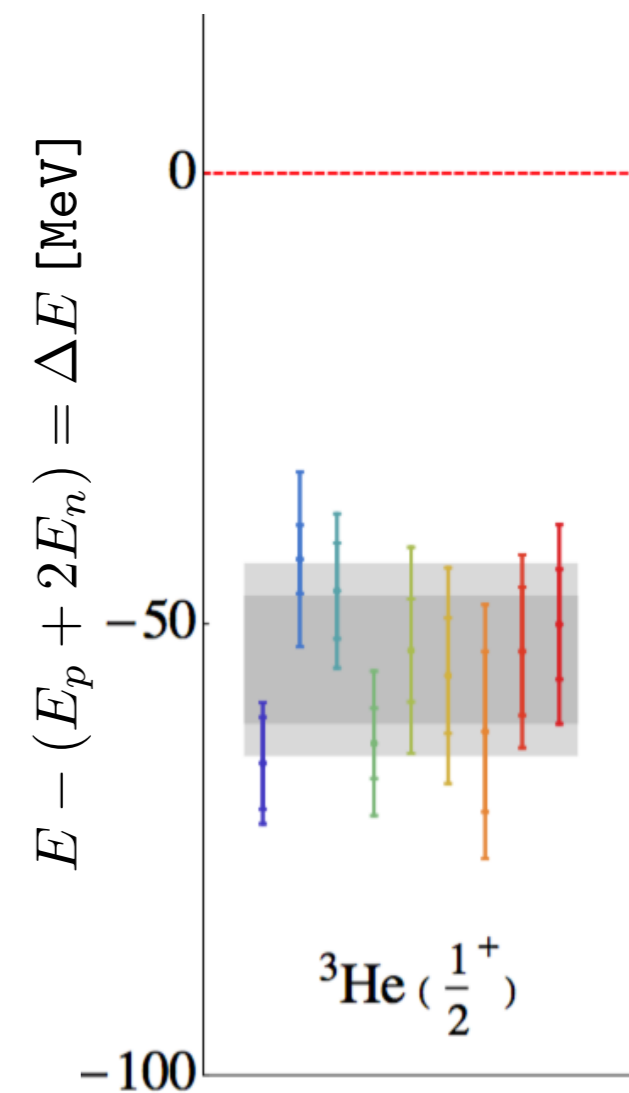
Spectrum



e.g.  $m_\pi = 800$   $G_{3\text{He}}(t)$



$$\sum_{\{A_\mu\}} \langle qqqqqqqqq(t) \overline{qqqqqqqqq}(0) \rangle$$

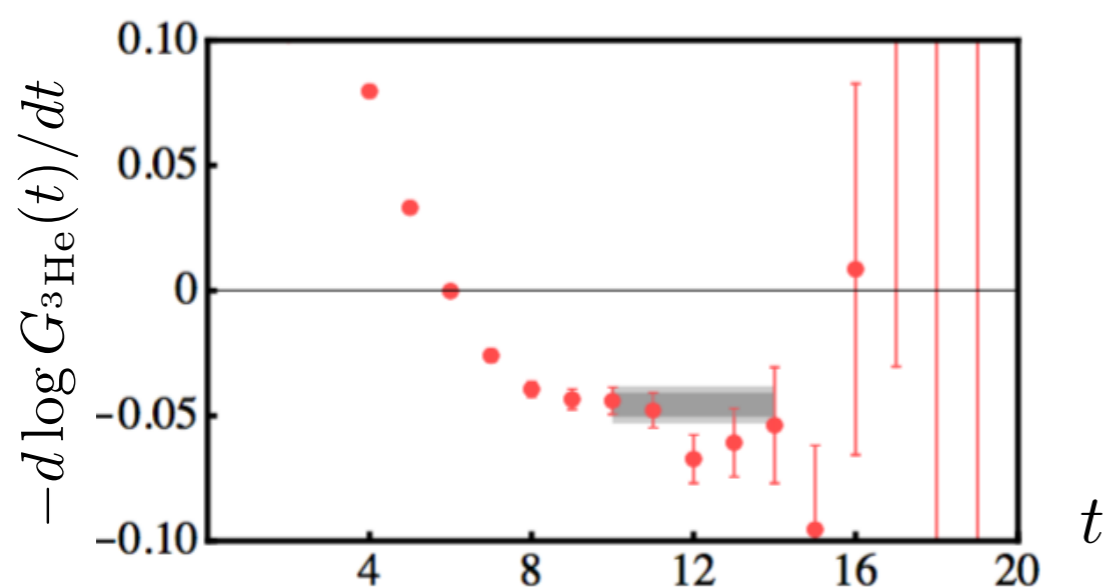




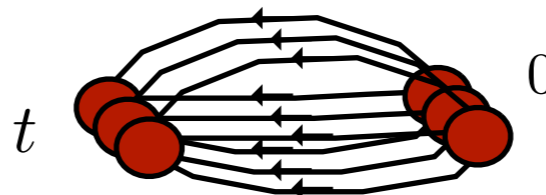
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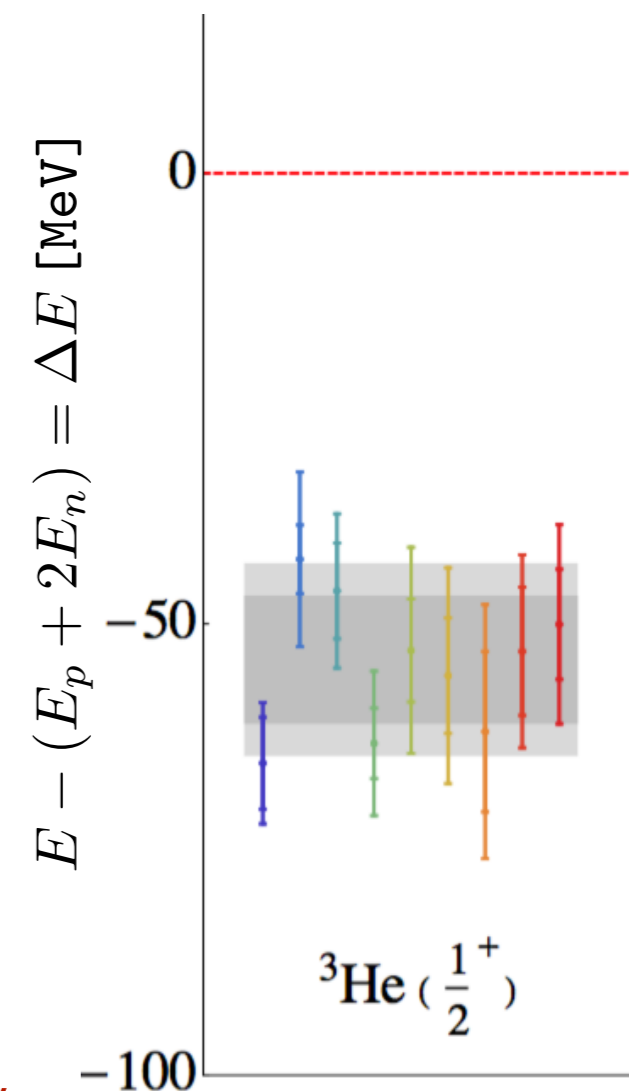
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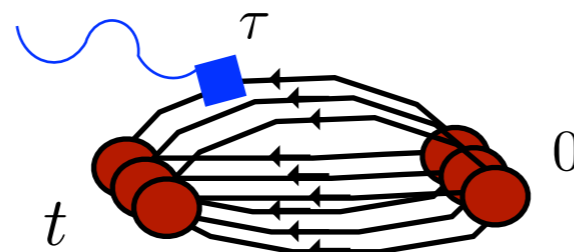


Nature and properties of these states?

$$\sum_{\{A_\mu\}} \langle qqqqqqqqq(t) [\bar{q}\gamma_\mu q](\tau) \overline{qqqqqqqqq}(0) \rangle \stackrel{t-\tau \gg 1, \tau \gg 1}{\sim} \langle {}^3\text{He} | J_\mu | {}^3\text{He} \rangle$$

Ground state requires

$$t - \tau, \tau \gtrsim 10$$



Current technology

$$t > 20$$

**Nuclear Matrix Elements from QCD?**

**... not yet attempted**

Better sources  
 Greater statistics

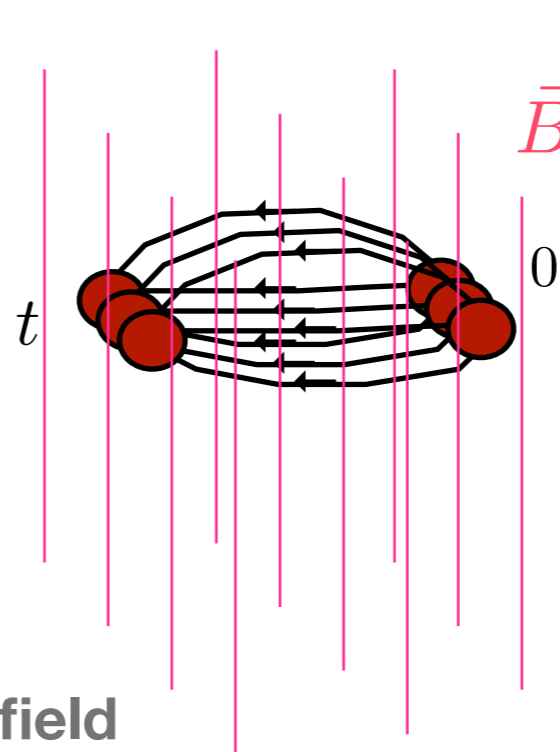
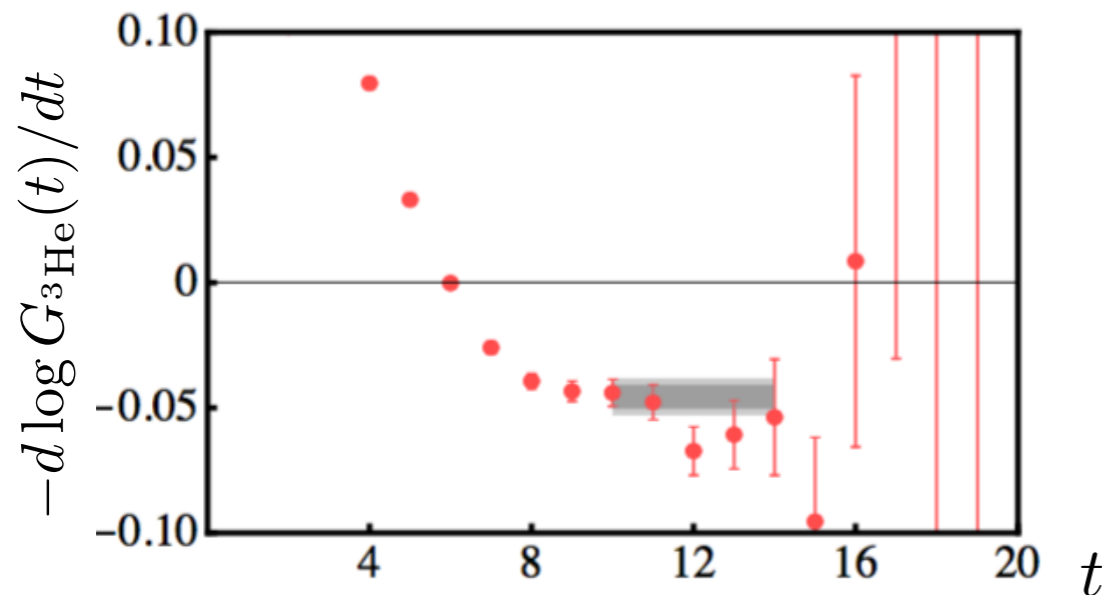




# Nuclear Properties @ $m_\pi = 800$ ?

Spectrum responds to external fields:

e.g. uniform magnetic fields



$$G_{\text{He}}^3(t)_{\vec{B}} = \sum_{\{A_\mu\}} \langle qqqqqqqqq(t) \overline{qqqqqqqqq}(0) \rangle_{\vec{B}}$$

Compute spectrum as a function of applied field

- I). In weak enough fields, can utilize same sources
- II). Need roughly same statistics for each field strength
- III). Requires fitting the field-strength dependence
- IV). Limited number of properties for a given type of field

Practical Solution:

*Lattice QCD + Classical Fields*

Beane, et al. **PRL:113** (2014)  
 Beane, et al. **PRL:115** (2015)  
 Chang, et al. **PRD:92** (2015)  
 Detmold, et al. **PRL:116** (2016)

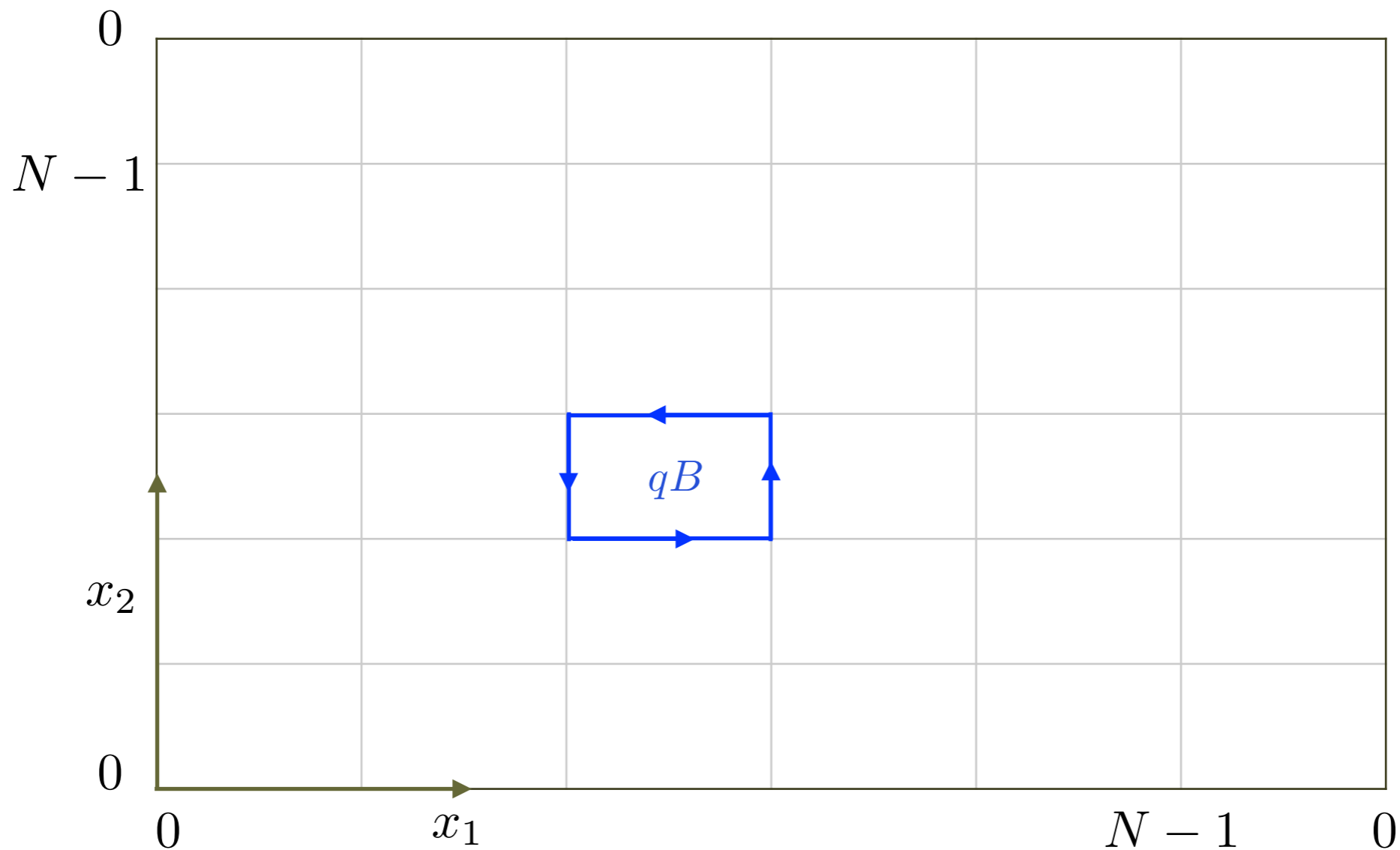
# Magnetic Field on a Periodic Lattice

Gauge links:

$$U_\mu(x) = e^{igG_\mu(x)} \in SU(3)$$

$$U_\mu^{\text{e.m.}}(x) = e^{iqA_\mu(x)} \in U(1)$$

Seek uniform **B-field**  $U_\mu(x) = e^{-iqx_2 B \delta_{\mu 1}}$



$$U_1(x)U_2(x + \hat{i})U_2^\dagger(x + \hat{i} + \hat{j})U_1^\dagger(x + \hat{j}) = e^{iqF_{12}} = e^{iqB}$$

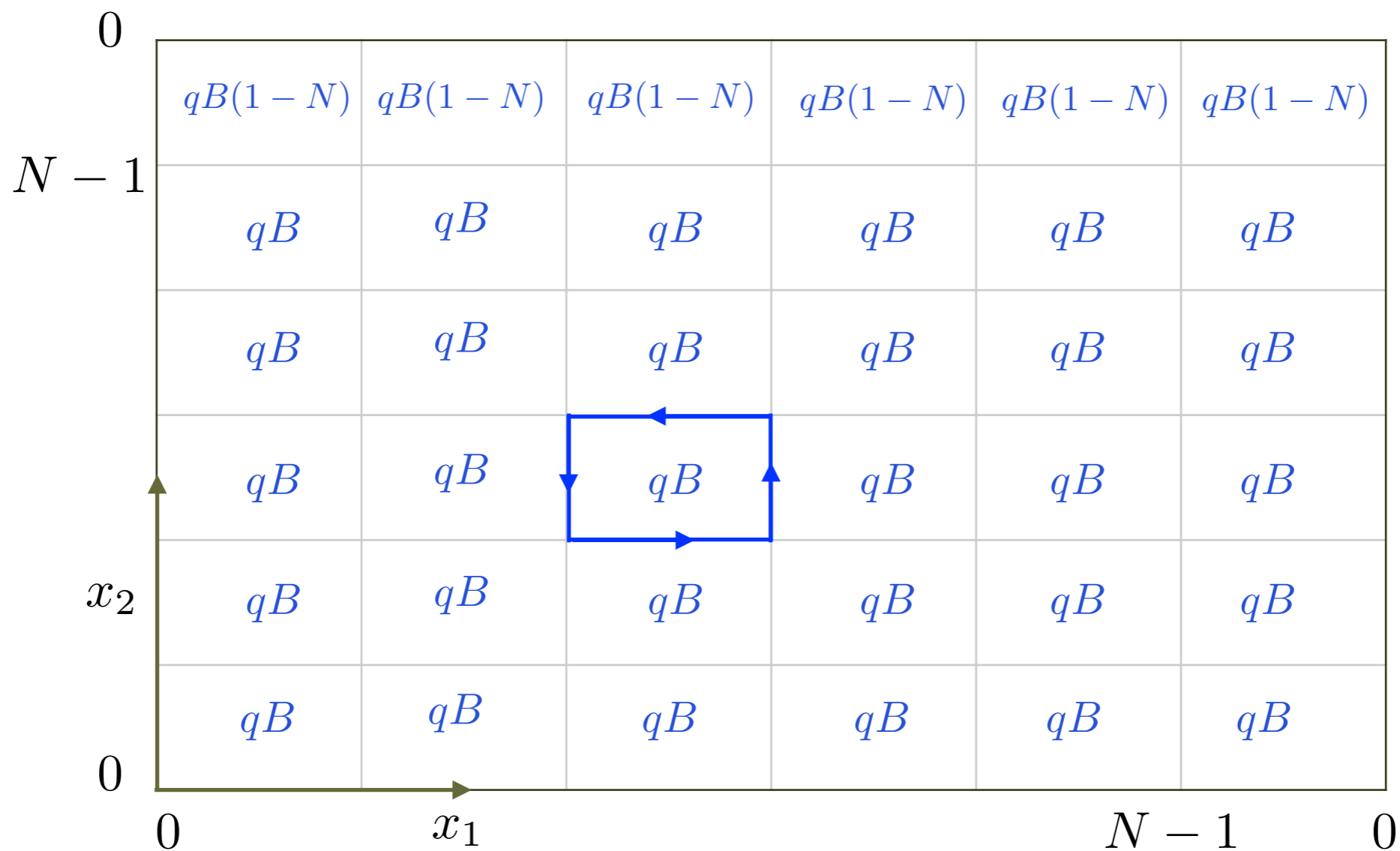
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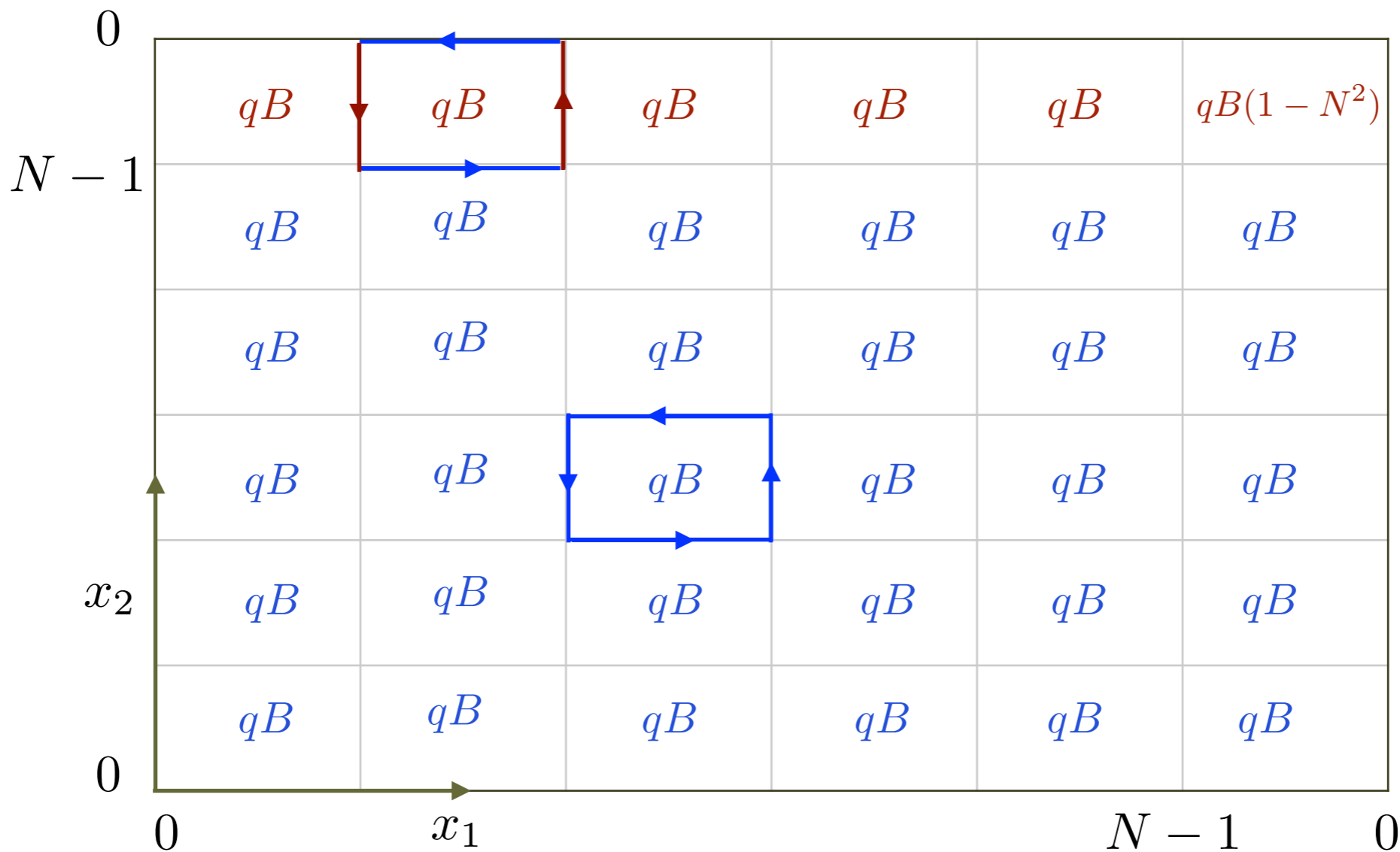
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Seek uniform B-field  $U_\mu(x) = e^{-iqx_2 B \delta_{\mu 1}} e^{+iqx_1 B N \delta_{\mu 2} \delta_{x_2, N-1}}$



Flux quantization

$$qB = \frac{2\pi}{N^2} n_\Phi$$

$$N = 32$$

We choose:

$$n_\Phi = +3, -6, +12$$

$$U_1(x)U_2(x + \hat{i})U_2^\dagger(x + \hat{i} + \hat{j})U_1^\dagger(x + \hat{j}) = e^{iqF_{12}} = e^{iqB}$$

# Magnetic Moments of Octet Baryons

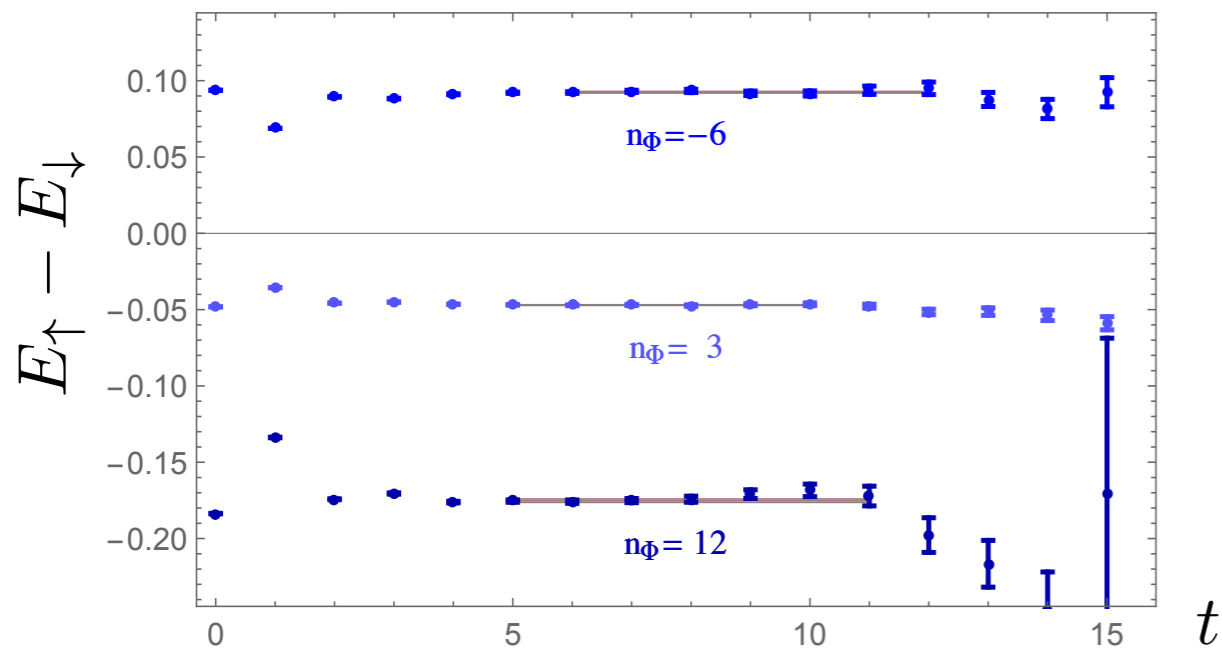


**Compute Zeeman Effect** using Lattice QCD + Uniform Magnetic fields  $E(B, J_z)$

Proton

$m_\pi \sim 800 \text{ MeV}$

$\gtrsim 10^{19} \text{ G}$



Units!

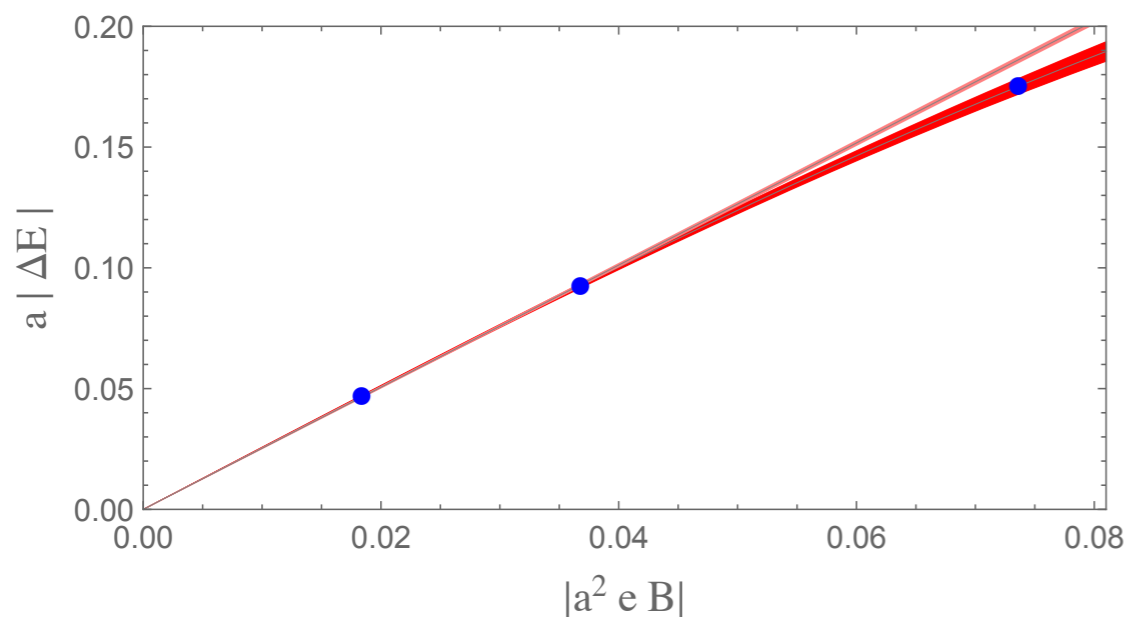
$$\mu_p = 2.560(09)(52) \text{ [LatM]}$$

$$\text{[LatM]} = \frac{e a}{2}$$

$$a = 0.145(2) \text{ fm}$$

$$\mu_p = 1.770(06)(36)(19) \text{ [NM]}$$

$$\text{[NM]} = \frac{e}{2M_N}$$



# Magnetic Moments of Octet Baryons

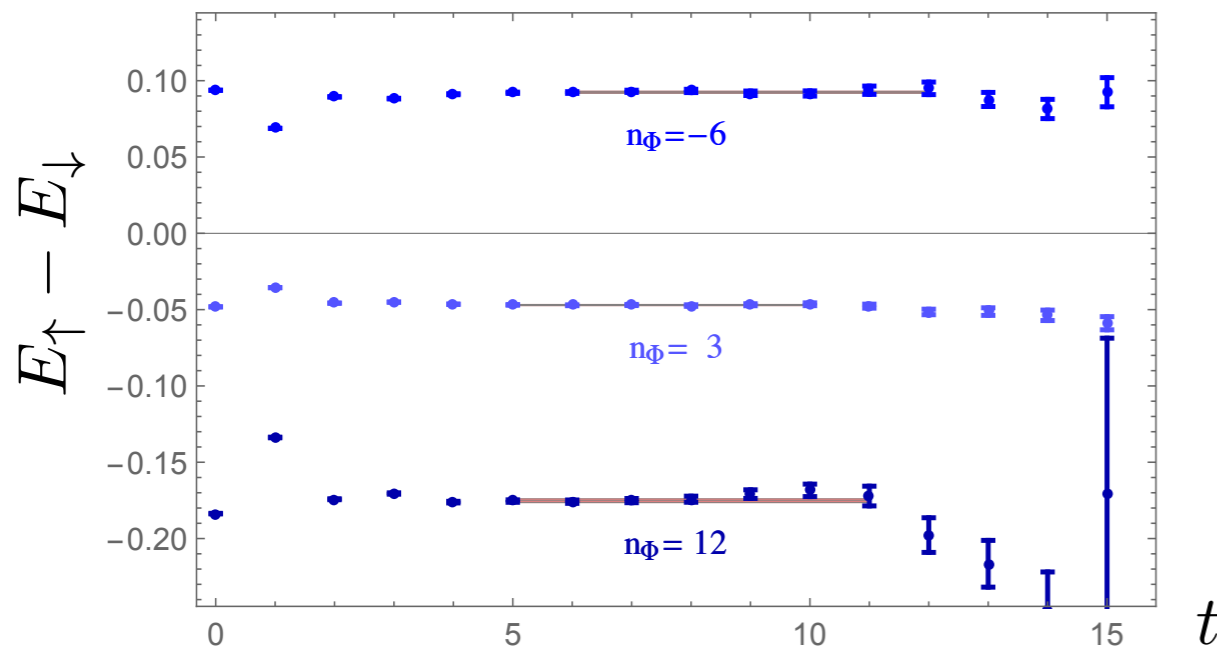


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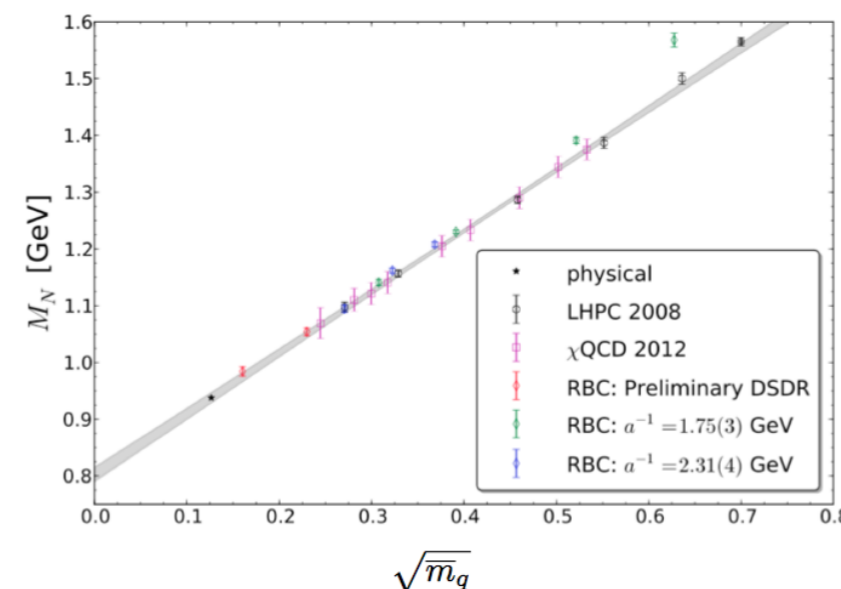
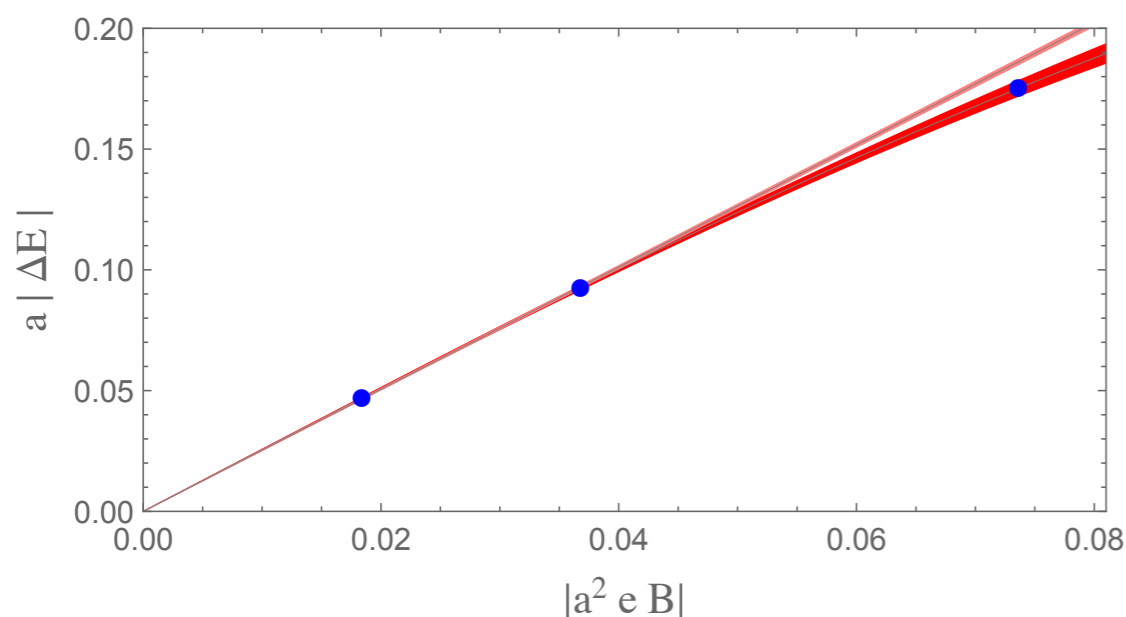


$$\mu_p = 1.770(06)(36)(19) \text{ [NM]}$$

$$\text{[NM]} = \frac{e}{2M_N}$$

**Ruler Mass Rule** (Walker-Loud, LHPC)

$$M_N(m_\pi) = 800 \text{ MeV} + m_\pi \sim 1,600 \text{ MeV}$$



$$\text{[nNM]} = \frac{e}{2M_N(m_\pi)}$$

# Magnetic Moments of Octet Baryons

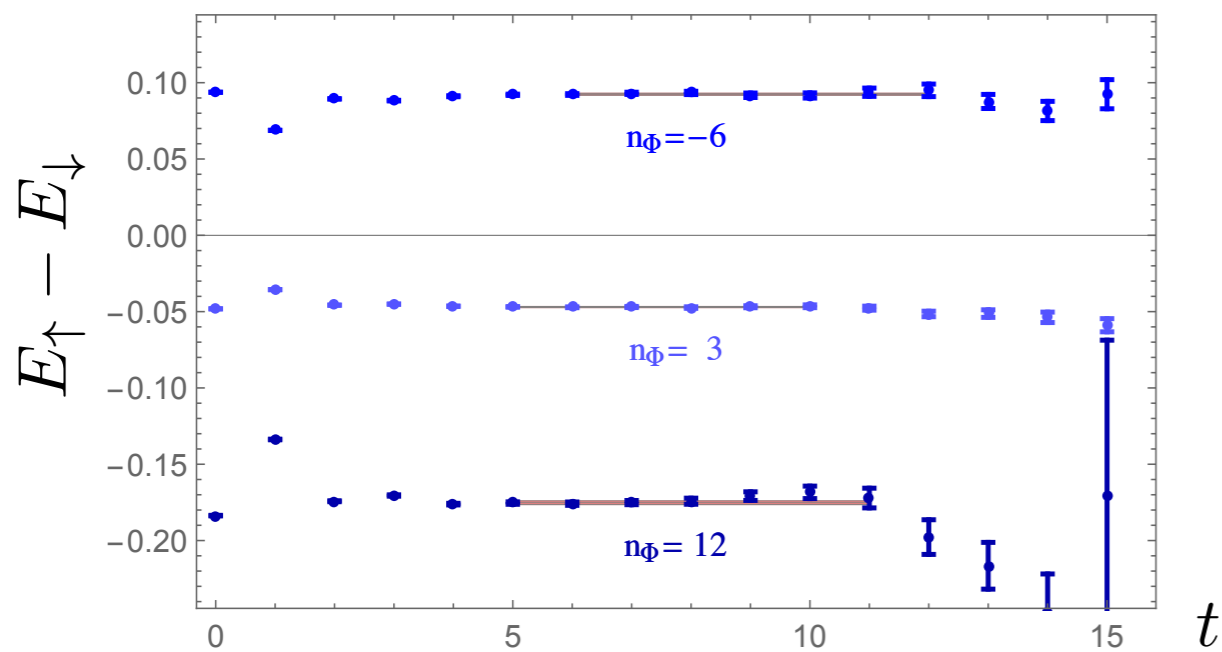


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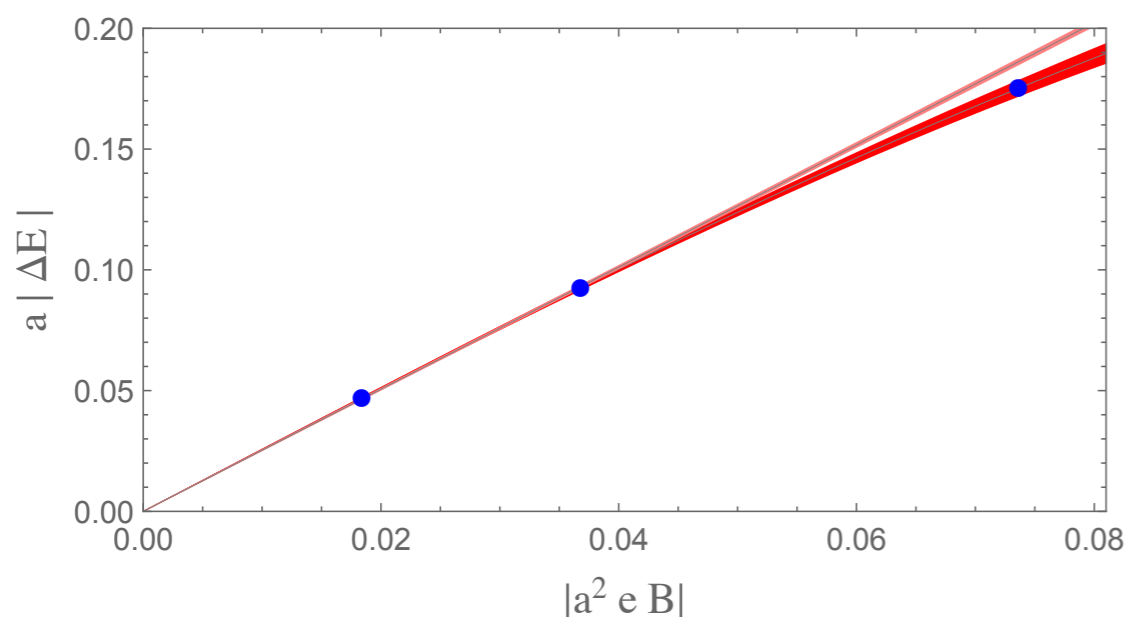
$$\mu_p = 1.770(06)(36)(19) \text{ [NM]}$$

$$\text{[NM]} = \frac{e}{2M_N}$$

Natural nucleon magnetons

$$\text{[nNM]} = \frac{e}{2M_N(m_\pi)}$$

$$\mu_p = 3.087(10)(62) \text{ [nNM]}$$



Dirac part is short-distance & guaranteed to  $\mathcal{O}(a^2 \Lambda_{\text{QCD}}^2)$

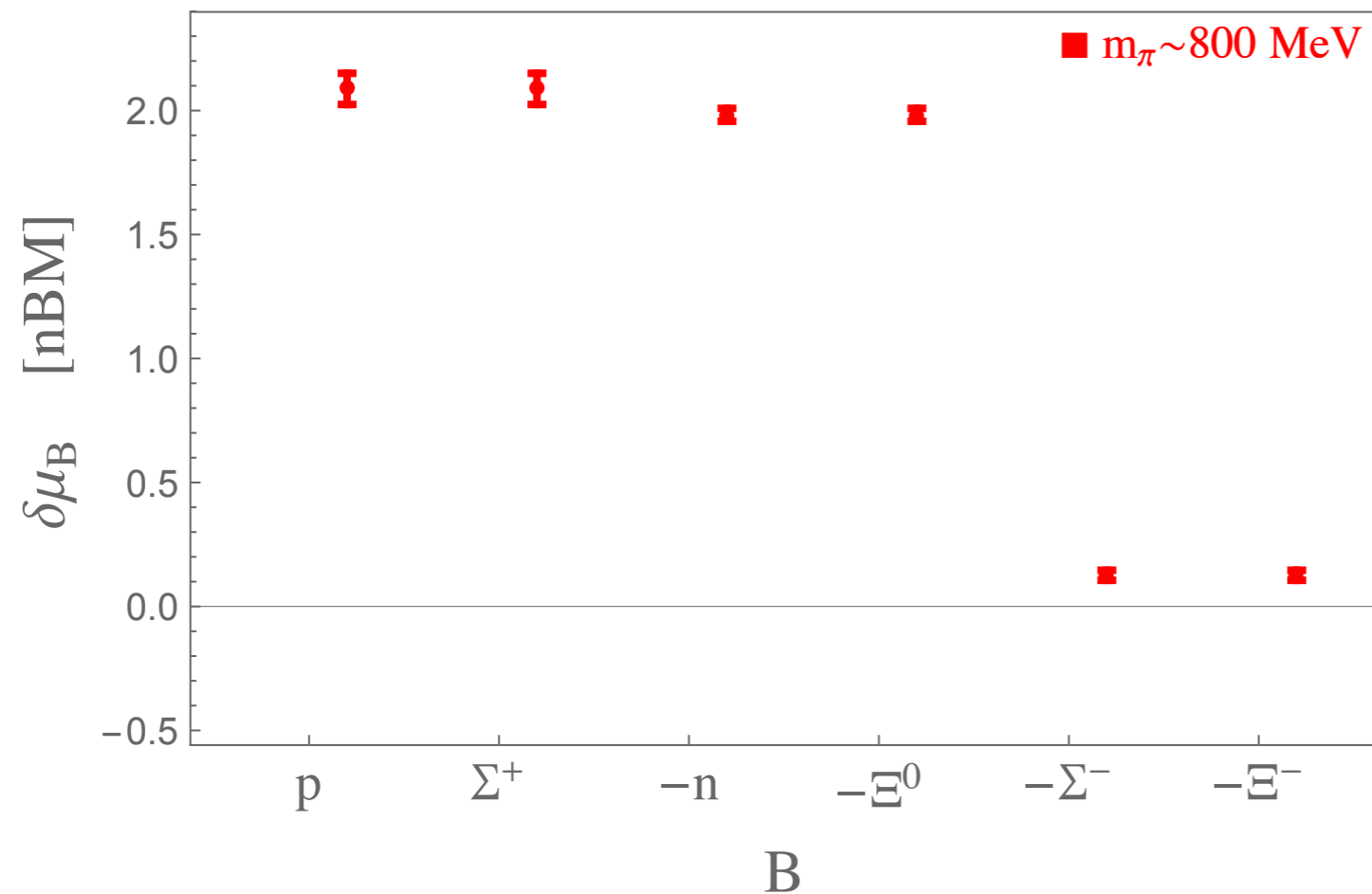
$$\delta\mu_p = 2.087(10)(62) \text{ [nNM]}$$

$$\delta\mu_p^{\text{exp}} = 1.7929... \text{ [NM]}$$



# Magnetic Moments of Octet Baryons

**Compute Zeeman Effect** using Lattice QCD + Uniform Magnetic fields



Natural baryon magnetons

$$[\text{nBM}] = \frac{e}{2M_B(m_\pi)}$$

Anomalous magnetic moments

$$\delta\mu_B [\text{nBM}] = \mu_B [\text{nBM}] - Q_B$$

U-spin

$$U(3)_F \xrightarrow{Q} U(1)_U \times U(1)_{D+S} \times SU(2)_{U\text{-spin}}$$

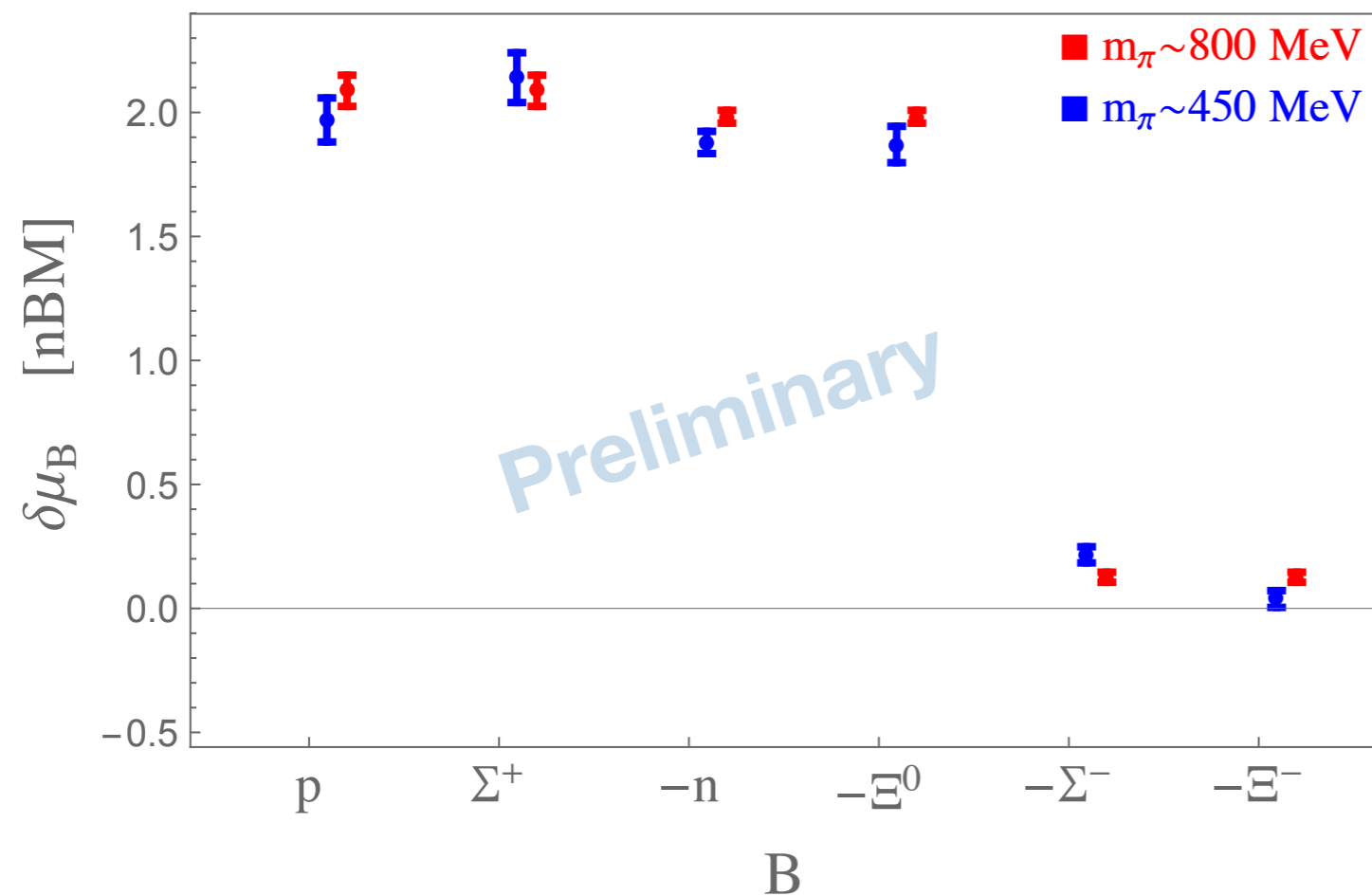
$$\begin{pmatrix} d \\ s \end{pmatrix} \xrightarrow{SU(2)} U \begin{pmatrix} d \\ s \end{pmatrix}$$





# Magnetic Moments of Octet Baryons

Compute **Zeeman Effect** using Lattice QCD + Uniform Magnetic fields



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$$[\text{nBM}] = \frac{e}{2M_B(m_\pi)}$$

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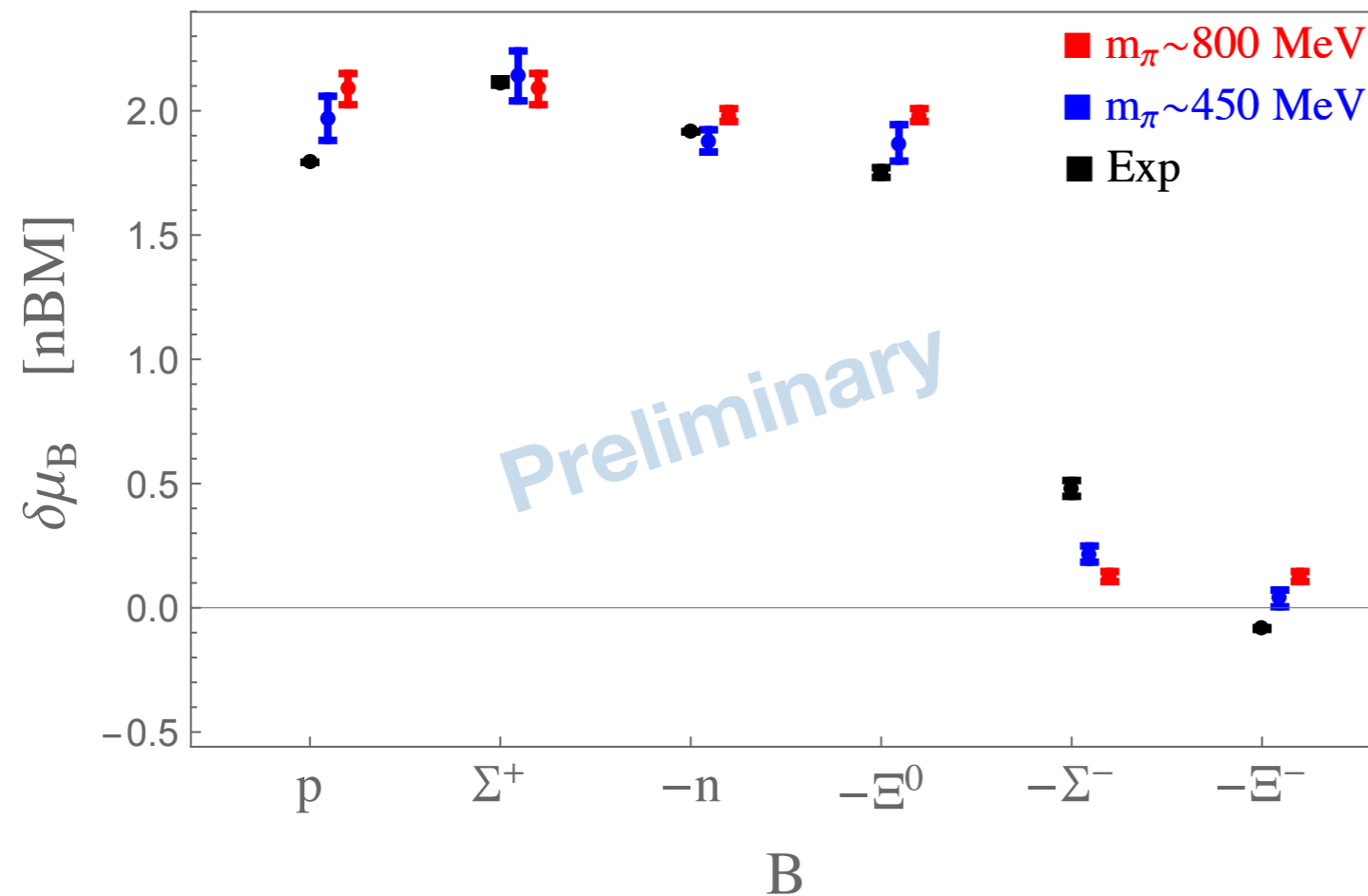
$$U(2)_I \times U(1)_S \xrightarrow{Q} U(1)_B \times U(1)_{I_3} \times U(1)_S \quad m_\pi \sim 450 \text{ MeV}$$

[Actually more complicated, our sea quarks are neutral]



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Compute **Zeeman Effect** using Lattice QCD + Uniform Magnetic fields



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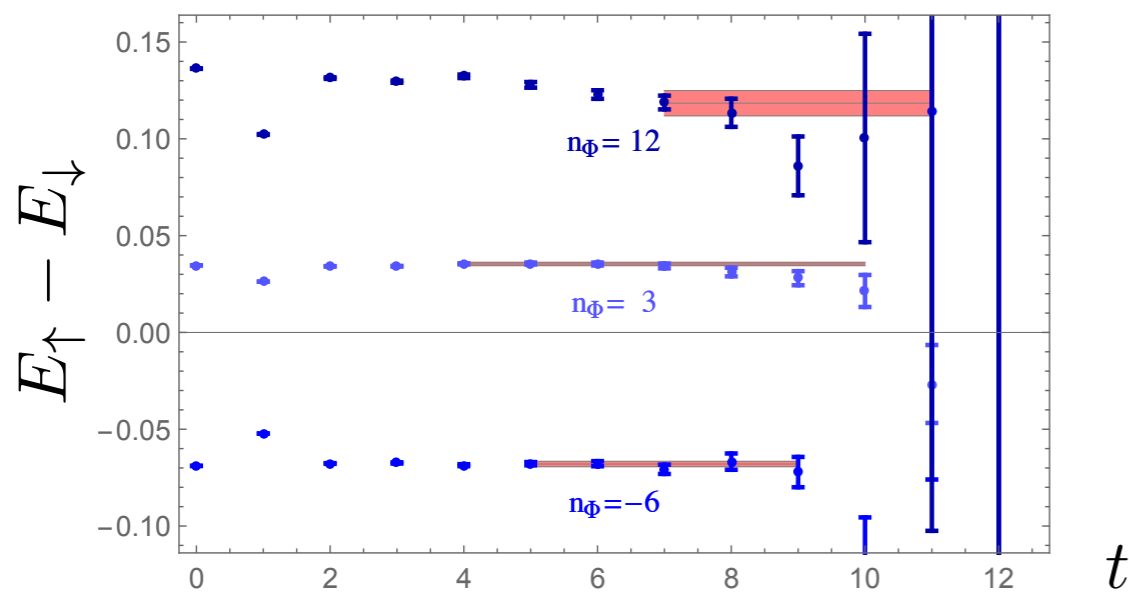
[Actually more complicated, our sea quarks are neutral]

# Magnetic Moments of Light Nuclei



Compute **Zeeman Effect** using Lattice QCD + Uniform Magnetic fields

${}^3\text{He}$



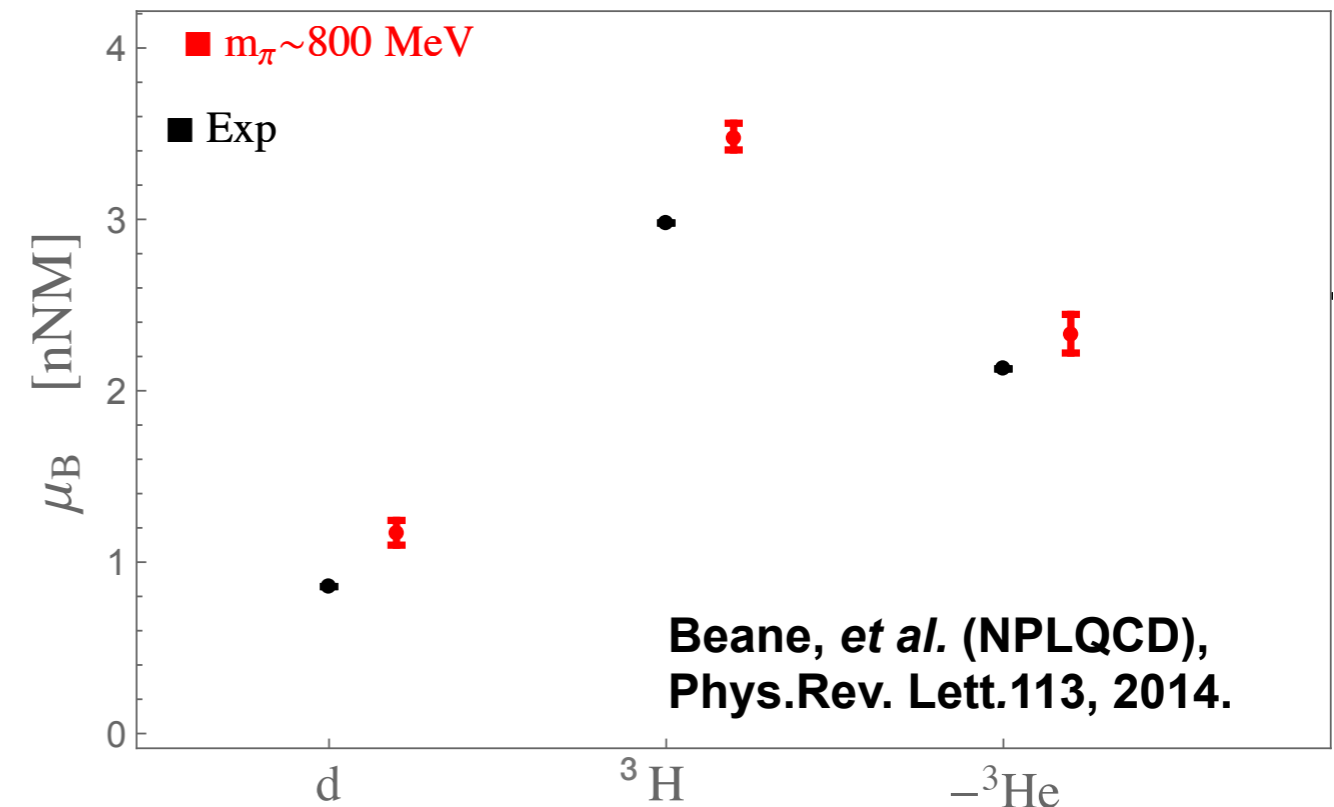
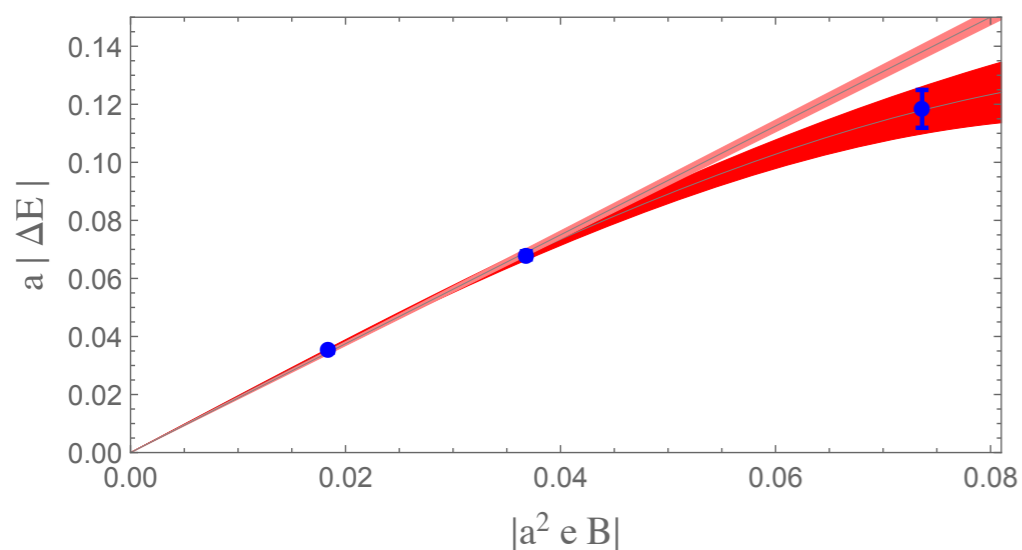
$t$

$$m_{\pi} \sim 800 \text{ MeV}$$

$$[\text{nNM}] = \frac{e}{2M_N(m_{\pi})}$$

$$\mu_{{}^3\text{He}} = -2.33(03)(11) [\text{nNM}]$$

$$\mu_{{}^3\text{He}}^{\text{exp}} = -2.1276... [\text{NM}]$$

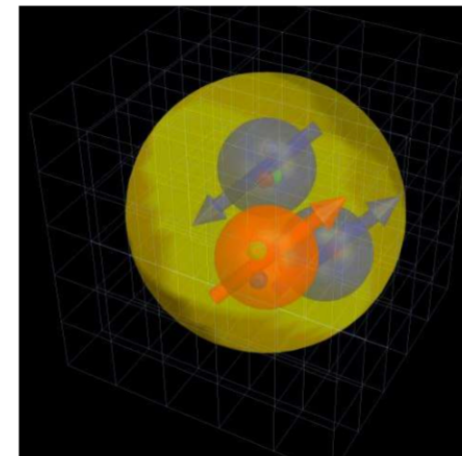
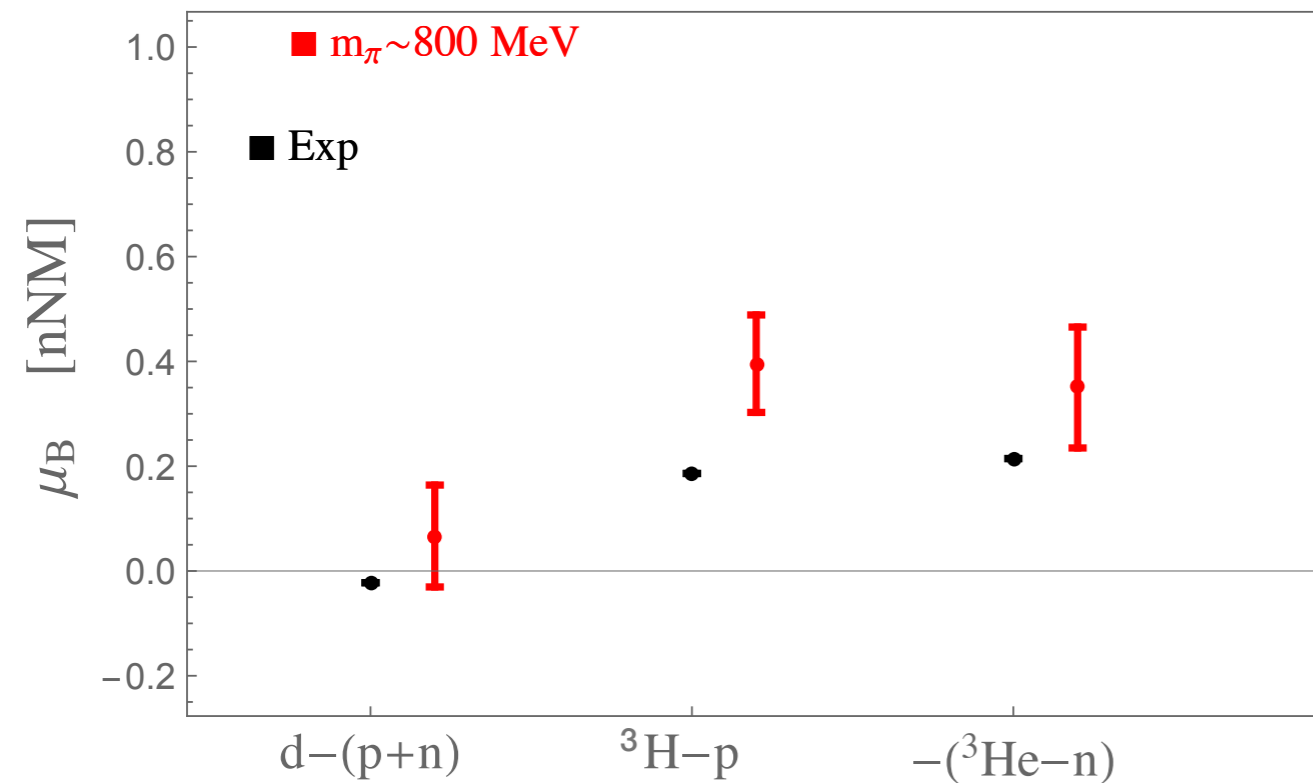


Beane, *et al.* (NPLQCD),  
Phys.Rev. Lett.113, 2014.

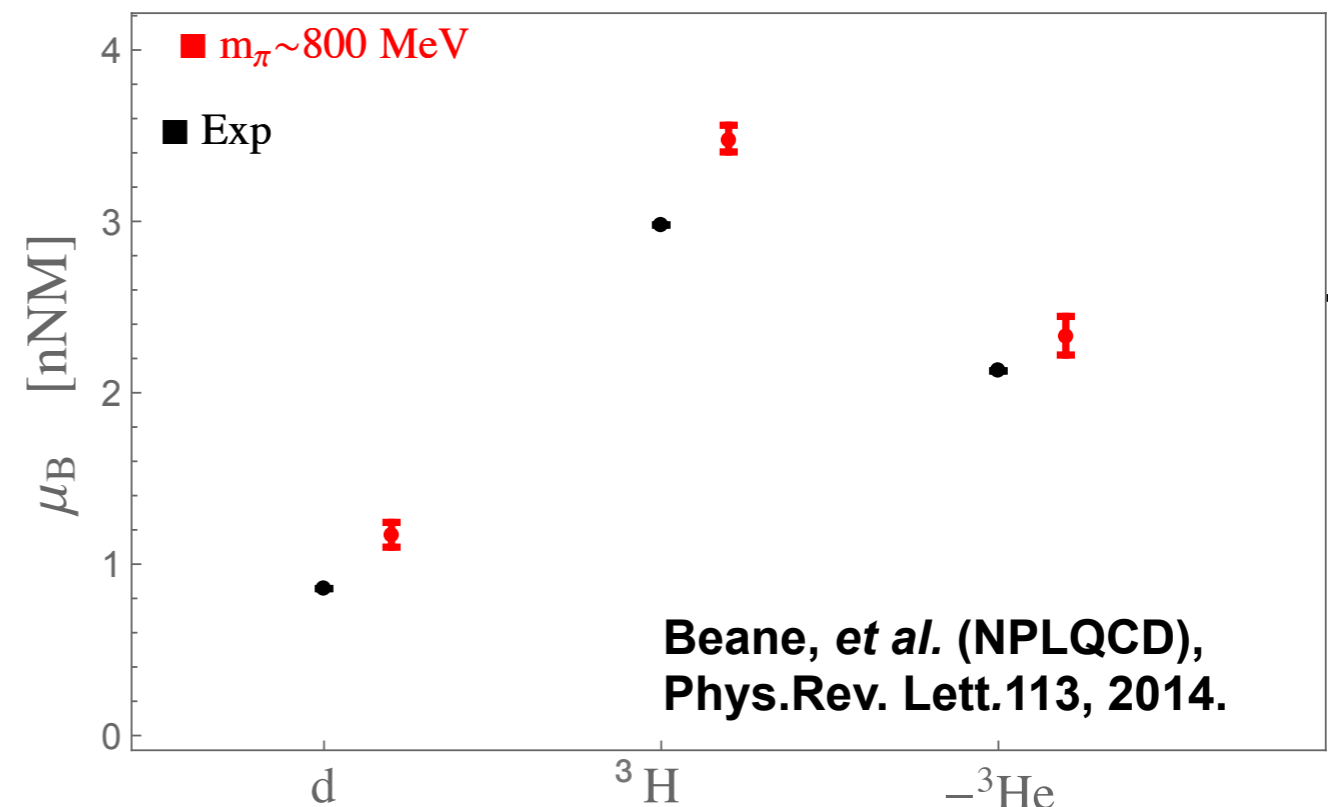


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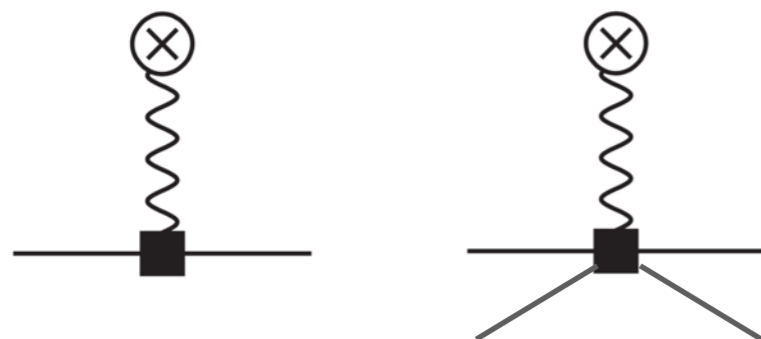
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$$[\text{nNM}] = \frac{e}{2M_N(m_\pi)}$$



Beane, *et al.* (NPLQCD),  
Phys.Rev. Lett.113, 2014.





# First “Nuclear Reaction” from QCD

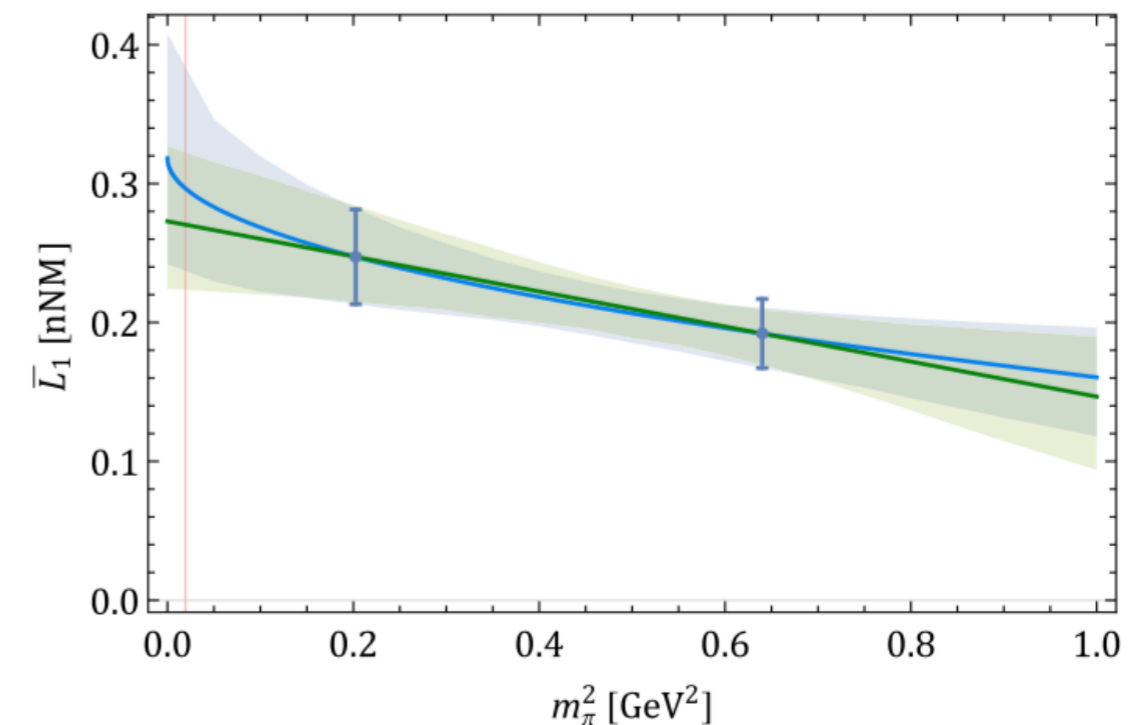
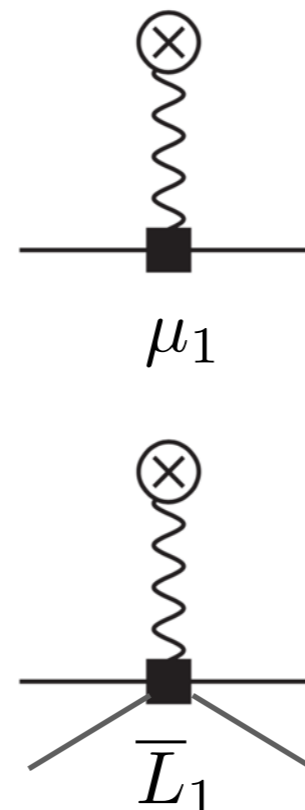
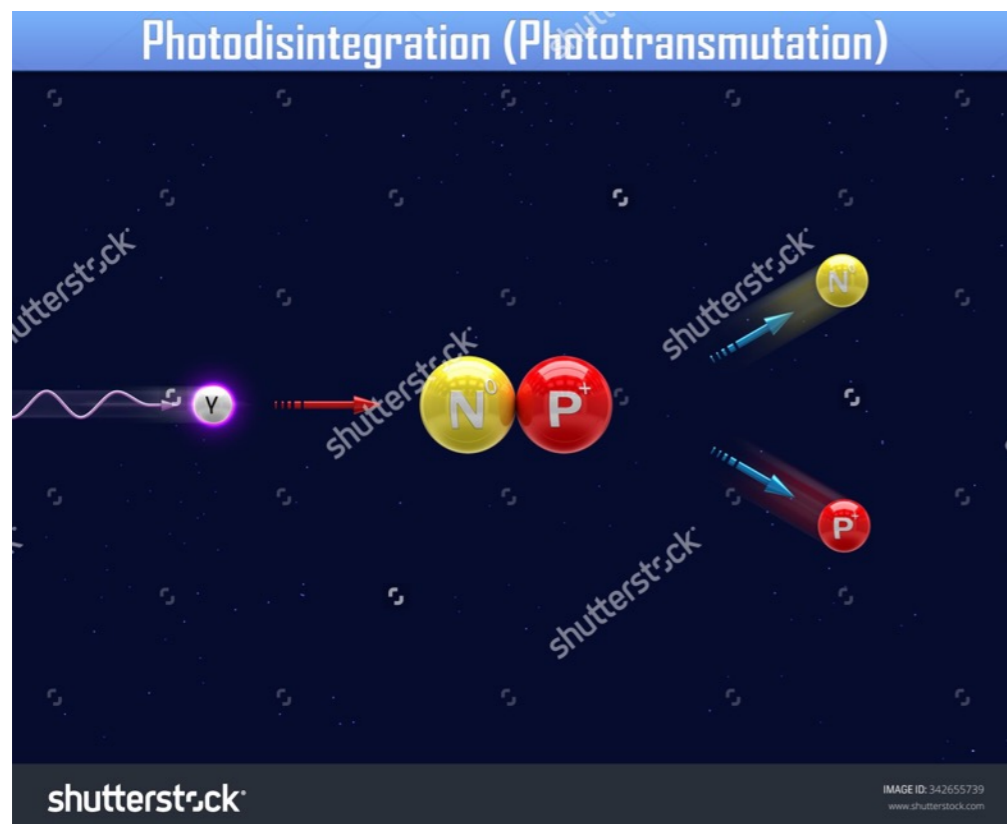
Dominant M1 transition @ Low Energy



Magnetically Coupled Channels

$$|\Delta I| = |\Delta J| = 1 \quad I_3 = j_z = 0$$

$$\mathbf{C}(t; \mathbf{B}) = \begin{pmatrix} C_{3S_1, 3S_1}(t; \mathbf{B}) & C_{3S_1, 1S_0}(t; \mathbf{B}) \\ C_{1S_0, 3S_1}(t; \mathbf{B}) & C_{1S_0, 1S_0}(t; \mathbf{B}) \end{pmatrix}$$



Two-body contribution isolated & compares favorably with EFT( $\pi$ ) phenomenology

# Extreme Magnetic Environments

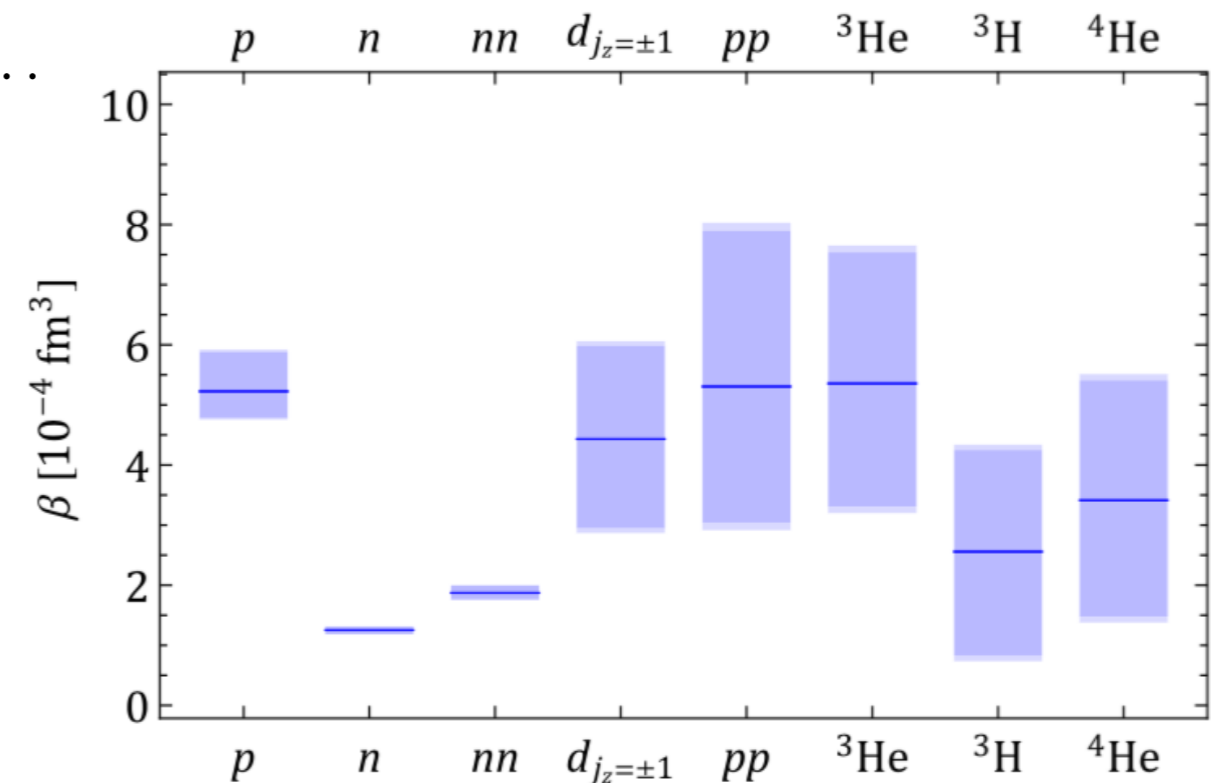
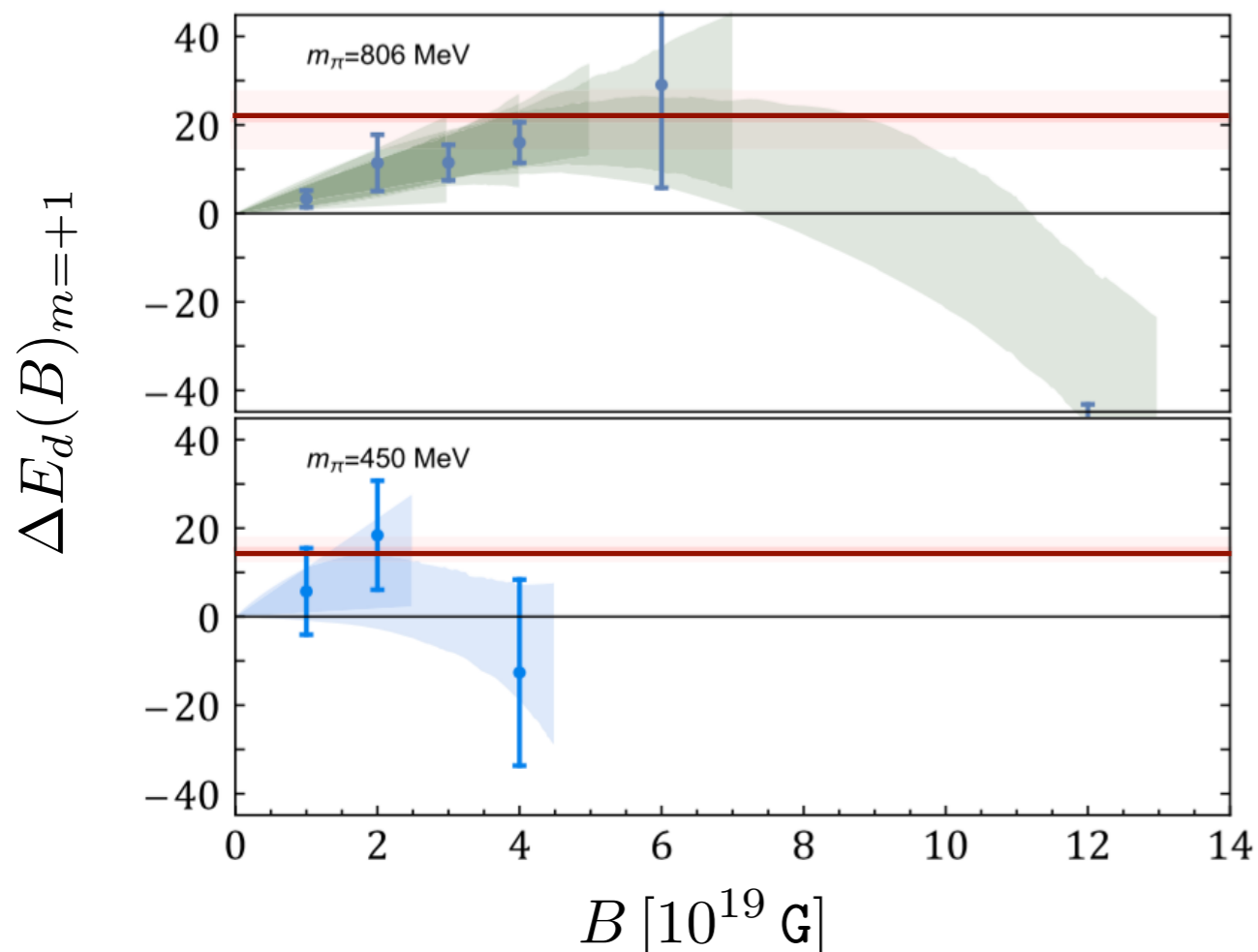


- Beyond Linear: Magnetic Polarizabilities

Chang, *et al.* (NPLQCD),  
Phys.Rev. D92, 2015.

$$E_{\uparrow+\downarrow}(B) = \sqrt{M^2 + (2n+1)|QB|} - \frac{1}{2}\beta_M B^2 + \dots$$

Landau identification dominates uncertainties



- Unitary NN Interactions?

$$a_{NN}(B) \rightarrow \infty?$$

Detmold, *et al.* (NPLQCD),  
Phys.Rev. Lett.116, 2016.

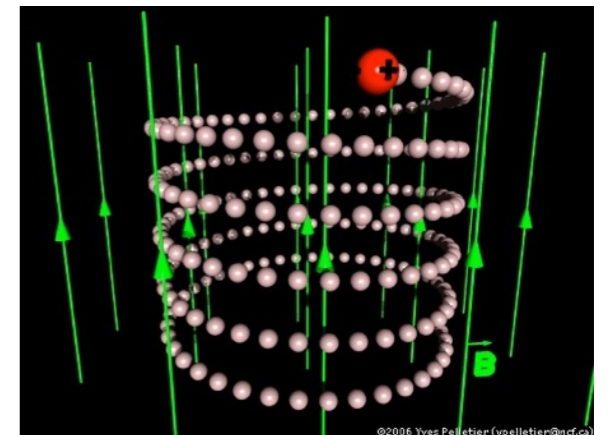
# Future Directions



+

- **Magnetic Structure of Nuclei**

Move beyond exploratory studies: remove systematics, lower pion mass, better treat Landau levels, sea quarks, ...



- **Electric Structure of Nuclei**

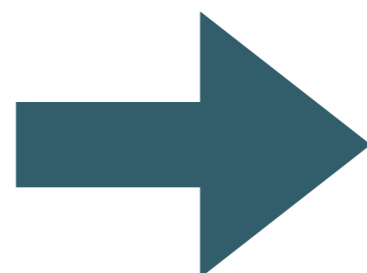
Electric polarizabilities?

EDMs of light nuclei from  $\theta$ -term?, BSM sources?

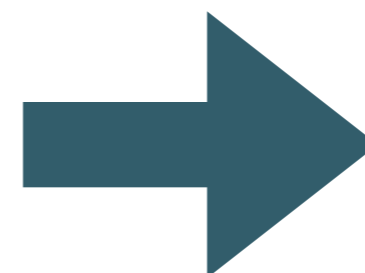
- **Nuclei in other classical fields...**

Gravitational?, Weak?

**Nuclear Physics  
from QCD**



**EW Reactions**



**BSM Physics**