PION-MASS DEPENDENCE OF LIGHT NUCLEI.

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Proper references in arXiv:1509.07697 [nucl-th]







It we would like to simulate a universe, would it include a burning sun with its beauty reflected in the images of bubble chambers?



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THEORETICAL PATCHWORK.



2 NUCLEONS $({}^{1}S_{0})$.



2 NUCLEONS (${}^{3}S_{1}$).



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SCATTERING 2 NUCLEONS (I).



SCATTERING 2 NUCLEONS (II).

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3 NUCLEONS.



Nuclear $A \leq 4$ spectrum.



SCALES.



NLO $EFT(\pi)$ with single nucleons.

Regularization & Projection.



3-Nucleon scattering $EFT(\mathcal{F})$.



3-Nucleon correlation $EFT(\vec{x})$.



3-Nucleon correlation $EFT(\vec{x})$.



3-Nucleon correlation $EFT(\vec{x})$.



4-Nucleon correlation $EFT(\vec{\pi})$.



 B_{α} [MeV]

WHAT'S NEXT?

Nuclear characteristics beyond "something is bound":

- i, E&M properties. (NPLQCD, Barnea et al.)
- ii, Next order in the EFT (St) expansion. (B.Bazah's talk?)
- iii,³² a-wav

- Intuitive improvements: in <u>AB(2)</u> <u>amr</u> <u>mr=806MeV</u> ii, axial background. iii, SU(3) vs. SU(2) @ fixed mr.
- Few-body & EFT problems: i, role/effect of reemingly equivalent regulators in f-b calculations. ii, correlations for A24