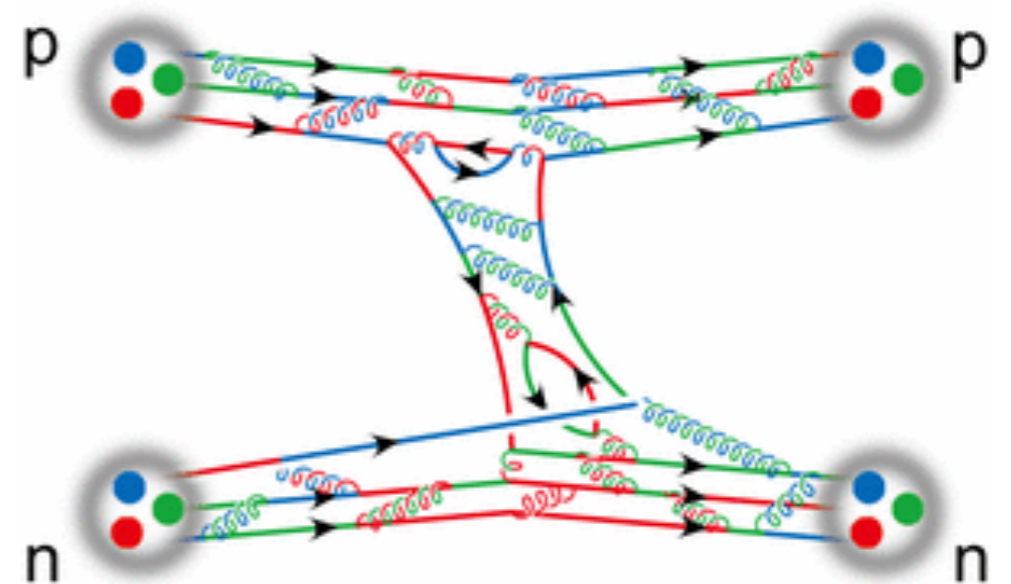


Overview of the HAL QCD potential method and recent results

Sinya Aoki

Yukawa Institute for Theoretical Physics
Kyoto University



INT Program INT-16-1

Nuclear Physics from Lattice QCD

March 21 - May 27, 2016, INT, University of Washington, Seattle, USA

For HAL QCD Collaboration

Sinya Aoki (YITP, Kyoto)

Takumi Doi (Riken)

Faisal Etminan (Birjand)

Sinya Gongyo (YITP, Kyoto->France)

Tetsuo Hatsuda (Riken)

Yoichi Ikeda (Riken->RCNP)

Takashi Inoue (Nihon)

Takumi Iritani (YITP, Kyoto->StonyBrook)

Noriyoshi Ishii (RCNP, Osaka)

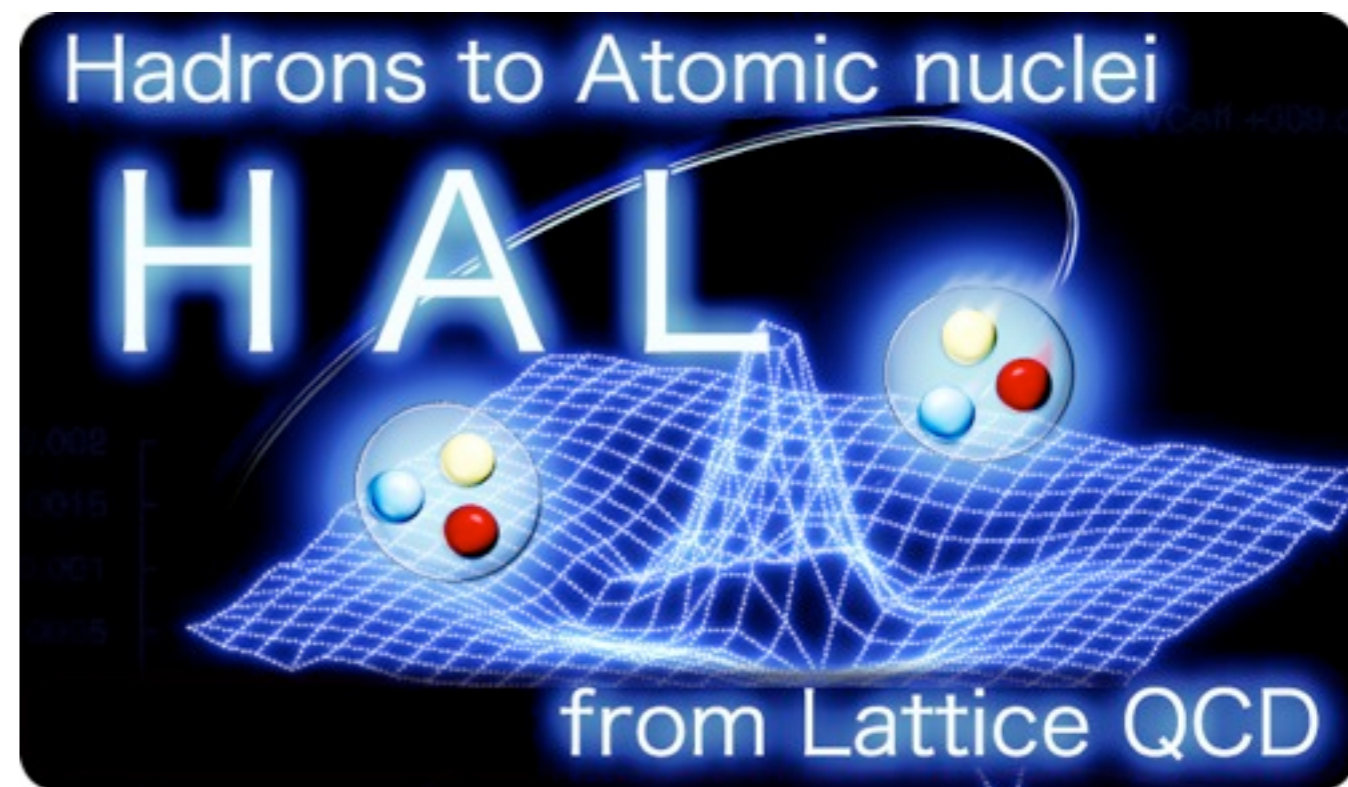
Daisuke Kawai* (Kyoto)

Takaya Miyamoto* (YITP, Kyoto)

Keiko Murano (RCNP, Osaka)

Hidekatsu Nemura (Tsukuba)

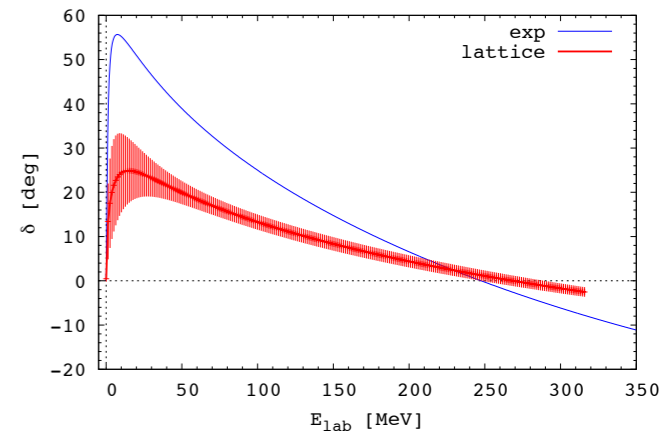
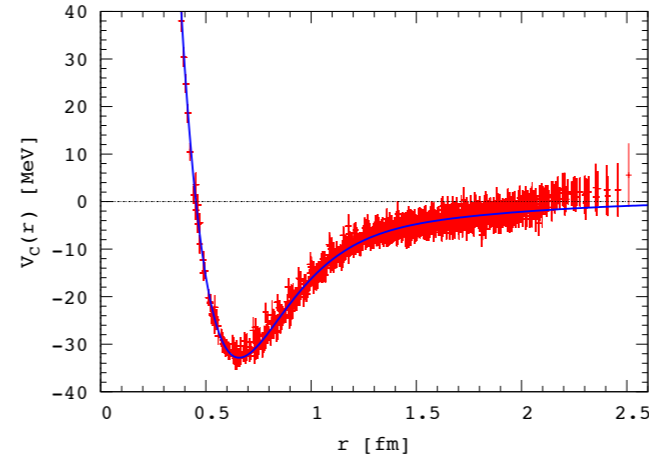
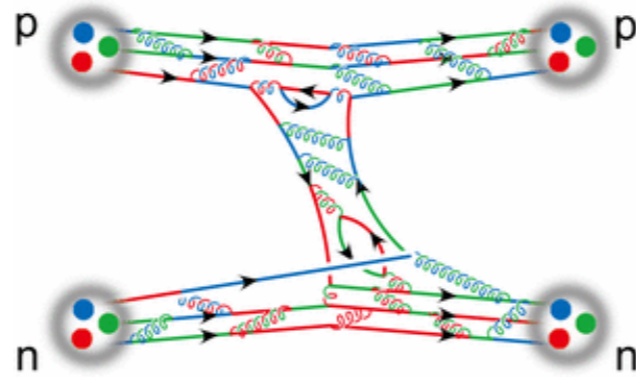
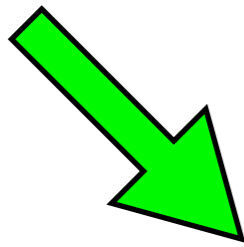
Kenji Sasaki (Tsukuba->YITP)



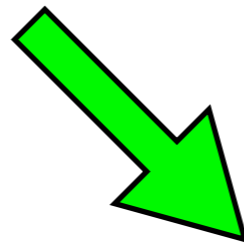
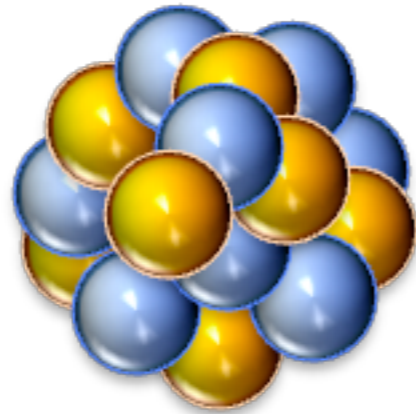
* PhD students

Our strategy to Nuclear/Astro physics from QCD

Potentials from lattice QCD

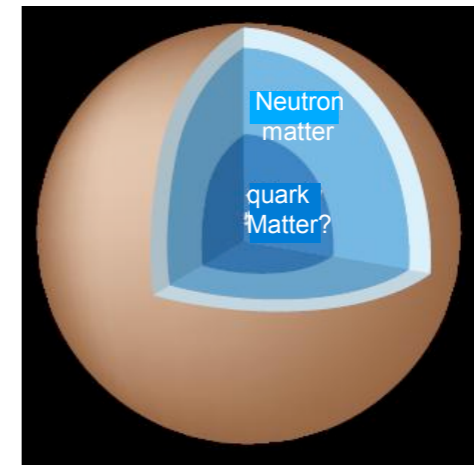
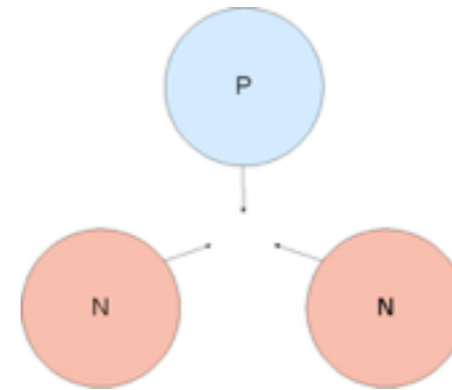


Nuclear Physics with these potentials



Neutron stars
Supernova explosion

parameters of EFT



1. HAL QCD method

--Overview--

HAL QCD strategy

$$NN \rightarrow NN$$

~~$$NN \rightarrow NN + \text{others}$$~~

Aoki, Hatsuda & Ishii, PTP123(2010)89.

Define a non-local but energy-independent potential below inelastic threshold in QCD

$$[\epsilon_k - H_0] \varphi_{\mathbf{k}}(\mathbf{x}) = \int d^3y \underline{U(\mathbf{x}, \mathbf{y})} \varphi_{\mathbf{k}}(\mathbf{y})$$

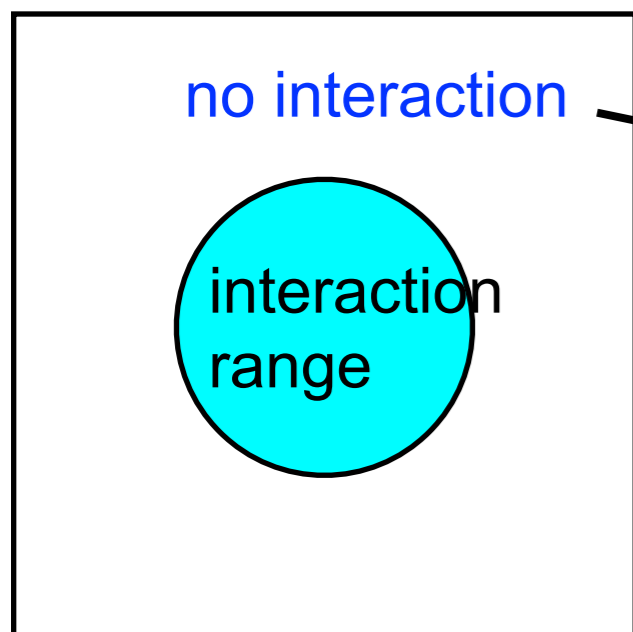
$$\epsilon_k = \frac{\mathbf{k}^2}{2\mu} \quad H_0 = \frac{-\nabla^2}{2\mu}$$

general $U_{\mathbf{k}}(\mathbf{x}, \mathbf{y}) \rightarrow U(\mathbf{x}, \mathbf{y}) \longleftrightarrow V_{\mathbf{k}}(\mathbf{x})$

from Nambu-Bethe-Salpeter (NBS) wave function

energy $W_k = 2\sqrt{\mathbf{k}^2 + m_N^2}$

$$\varphi_{\mathbf{k}}(\mathbf{r}) = \langle 0 | N(\mathbf{x} + \mathbf{r}, 0) N(\mathbf{x}, 0) | NN, W_k \rangle$$



$$r = |\mathbf{r}| \rightarrow \infty$$

$$\varphi_{\mathbf{k}}(\mathbf{r}) \simeq \sum_{l,m} C_l \frac{\sin(kr - l\pi/2 + \delta_l(k))}{kr} Y_{ml}(\Omega_{\mathbf{r}})$$

$$\delta_l(k)$$

scattering phase shift (phase of the S-matrix by unitarity) in QCD.

potential $U(x, y)$ is faithful to QCD phase shift $\delta_l(k)$.

A non-local but **energy-independent** potential exists.

Proof

$$U(\mathbf{x}, \mathbf{y}) = \sum_{\mathbf{k}, \mathbf{k}'}^{W_{\mathbf{k}}, W_{\mathbf{k}'} \leq W_{\text{th}}} [\epsilon_{\mathbf{k}} - H_0] \varphi_{\mathbf{k}}(\mathbf{x}) \eta_{\mathbf{k}, \mathbf{k}'}^{-1} \varphi_{\mathbf{k}'}^\dagger(\mathbf{y})$$

inner product

$\eta_{\mathbf{k}, \mathbf{k}'}^{-1}$: inverse of $\eta_{\mathbf{k}, \mathbf{k}'} = (\varphi_{\mathbf{k}}, \varphi_{\mathbf{k}'})$

For $\forall W_{\mathbf{p}} < W_{\text{th}} = 2m_N + m_\pi$ (threshold energy)

$$\int d^3y U(\mathbf{x}, \mathbf{y}) \phi_{\mathbf{p}}(\mathbf{y}) = \sum_{\mathbf{k}, \mathbf{k}'} [\epsilon_{\mathbf{k}} - H_0] \varphi_{\mathbf{k}}(x) \eta_{\mathbf{k}, \mathbf{k}'}^{-1} \eta_{\mathbf{k}', \mathbf{p}} = [\epsilon_{\mathbf{p}} - H_0] \varphi_{\mathbf{p}}(x)$$

Derivative (velocity) expansion

$$U(\mathbf{x}, \mathbf{y}) = V(\mathbf{x}, \nabla) \delta^3(\mathbf{x} - \mathbf{y})$$

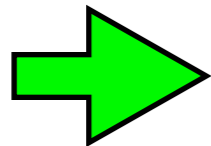
$$V(\mathbf{x}, \nabla) = \underbrace{V_0(r)}_{\text{LO}} + \underbrace{V_\sigma(r)}_{\text{LO}} (\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2) + \underbrace{V_T(r)}_{\text{LO}} S_{12} + \underbrace{V_{\text{LS}}(r)}_{\text{NLO}} \mathbf{L} \cdot \mathbf{S} + \underbrace{O(\nabla^2)}_{\text{NNLO}}$$

tensor operator $S_{12} = \frac{3}{r^2} (\boldsymbol{\sigma}_1 \cdot \mathbf{x})(\boldsymbol{\sigma}_2 \cdot \mathbf{x}) - (\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2)$

spins

At LO, for example, we simply have

$$V_{\text{LO}}(\mathbf{x}) = \frac{[\epsilon_k - H_0] \varphi_{\mathbf{k}}(\mathbf{x})}{\varphi_{\mathbf{k}}(\mathbf{x})}$$

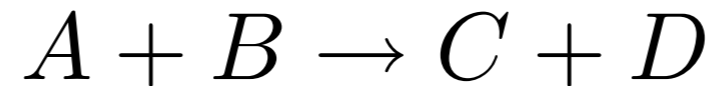


phase shifts and binding energy below inelastic threshold

Note truncation of the derivative expansion introduces some systematics, which fortunately can be estimated explicitly.

Advantages

- milder finite volume corrections
- extension to coupled channel problems is easy. **Inelastic scattering** can be treated.



- extensions to 3-body potential /particle production are possible with non-relativistic approximation, though numerically demanding.



- **ground state saturation** is not required to extract the potential. (See later.)

Disadvantages

- higher numerical cost than the standard method in both Flops and Bytes.

2. Previous results

Extraction of potentials

Standard method

NBS wave function

$$\varphi_{\mathbf{k}}(\mathbf{r}) = \langle 0 | N(\mathbf{x} + \mathbf{r}, 0) N(\mathbf{x}, 0) | NN, W_k \rangle$$



Potential

$$[\epsilon_k - H_0] \varphi_{\mathbf{k}}(\mathbf{x}) = \int d^3 y U(\mathbf{x}, \mathbf{y}) \varphi_{\mathbf{k}}(\mathbf{y})$$

4-pt Correlation function

source for NN

$$F(\mathbf{r}, t - t_0) = \langle 0 | T \{ N(\mathbf{x} + \mathbf{r}, t) N(\mathbf{x}, t) \} \underline{\overline{\mathcal{J}}(t_0)} | 0 \rangle$$

complete set for NN

$$\begin{aligned} F(\mathbf{r}, t - t_0) &= \langle 0 | T \{ N(\mathbf{x} + \mathbf{r}, t) N(\mathbf{x}, t) \} \sum_{n, s_1, s_2} \underline{|2N, W_n, s_1, s_2\rangle \langle 2N, W_n, s_1, s_2| \overline{\mathcal{J}}(t_0)} | 0 \rangle + \dots \\ &= \sum_{n, s_1, s_2} A_{n, s_1, s_2} \varphi^{W_n}(\mathbf{r}) e^{-W_n(t-t_0)}, \quad A_{n, s_1, s_2} = \langle 2N, W_n, s_1, s_2 | \overline{\mathcal{J}}(0) | 0 \rangle. \end{aligned}$$

ground state saturation at large t

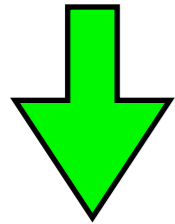
$$\lim_{(t-t_0) \rightarrow \infty} F(\mathbf{r}, t - t_0) = \underline{A_0 \varphi^{W_0}(\mathbf{r})} e^{-W_0(t-t_0)} + O(e^{-W_{n \neq 0}(t-t_0)})$$

NBS wave function

This is a standard method in lattice QCD and was employed for our first calculation.

normalized 4-pt function

$$R(\mathbf{r}, t) \equiv F(\mathbf{r}, t) / (e^{-m_N t})^2 = \sum_n A_n \varphi^{W_n}(\mathbf{r}) e^{-\Delta W_n t}$$



potential

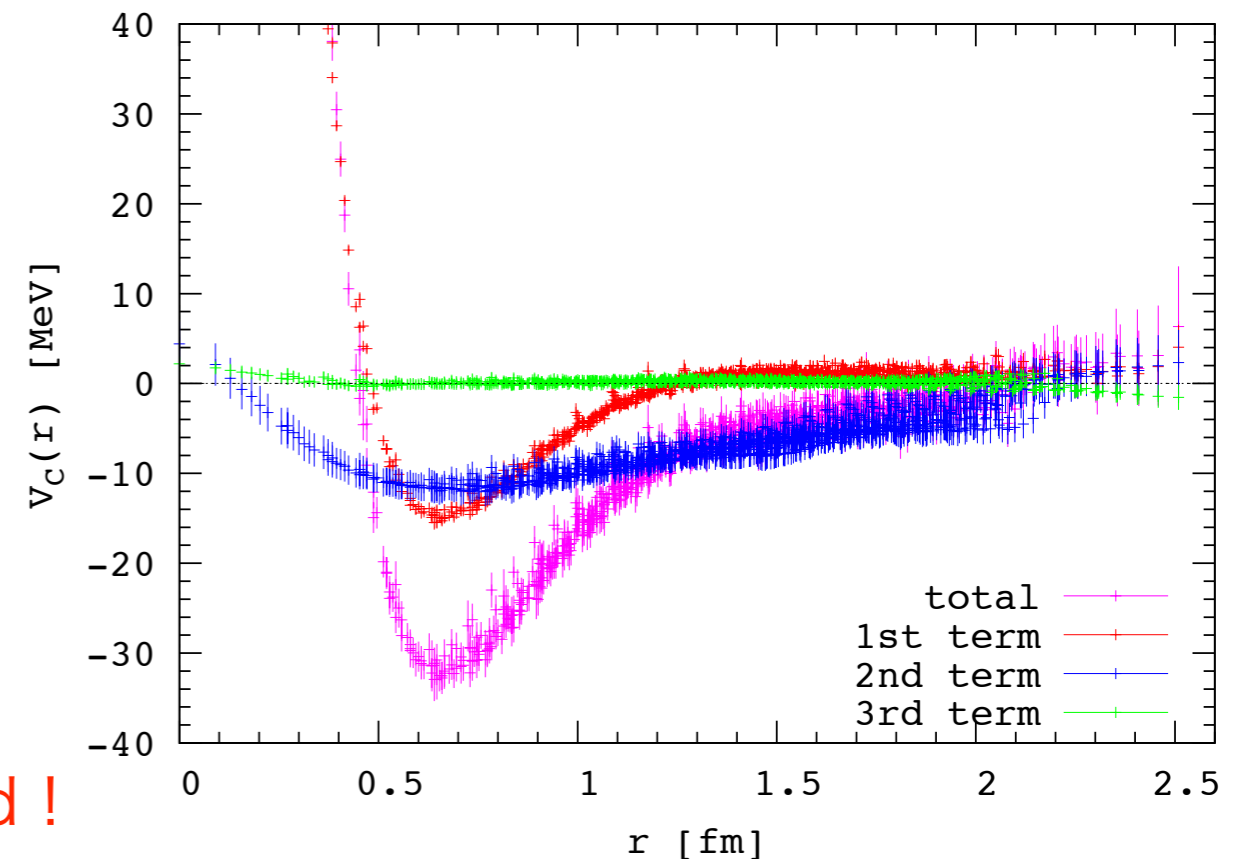
$$\Delta W_n = W_n - 2m_N = \frac{\mathbf{k}_n^2}{m_N} - \frac{(\Delta W_n)^2}{4m_N}$$

$$-\frac{\partial}{\partial t} R(\mathbf{r}, t) = \left\{ H_0 + U - \frac{1}{4m_N} \frac{\partial^2}{\partial t^2} \right\} R(\mathbf{r}, t)$$

Leading Order

$$\left\{ \underbrace{-H_0}_{1st} - \underbrace{\frac{\partial}{\partial t}}_{2nd} + \underbrace{\frac{1}{4m_N} \frac{\partial^2}{\partial t^2}}_{3rd} \right\} R(\mathbf{r}, t) = \int d^3 r' U(\mathbf{r}, \mathbf{r}') R(\mathbf{r}', t) = \underbrace{V_C(\mathbf{r})}_{total} R(\mathbf{r}, t) + \dots$$

3rd term (relativistic correction) is negligible.

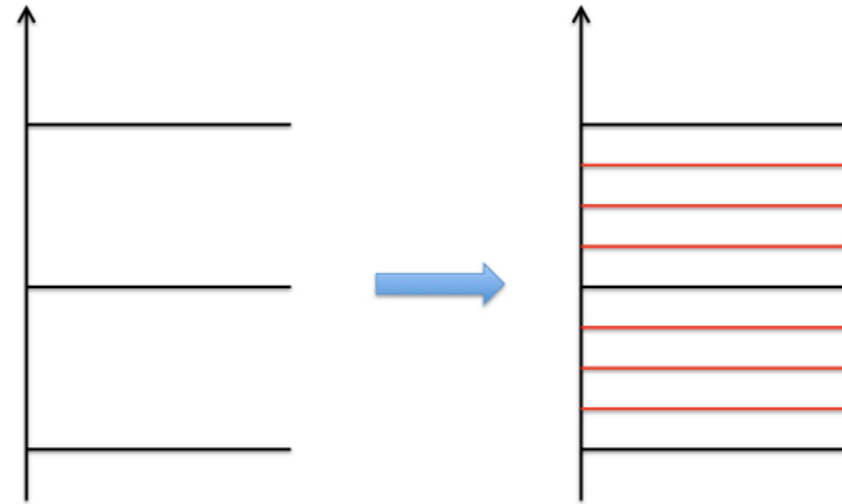


Ground state saturation is no more required!
(advantage over the finite volume method.)

Remarks

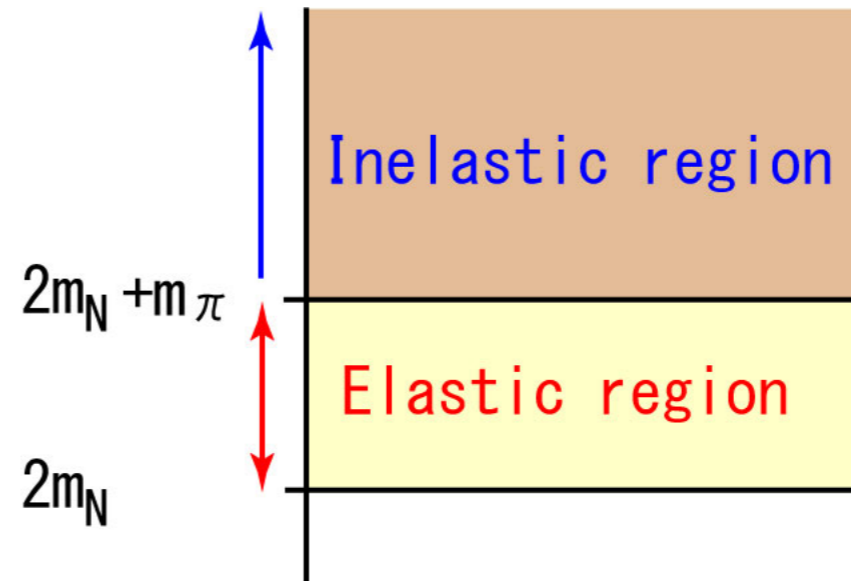
excited state contributions become bigger
in the larger volume

$$\Delta E \propto \frac{1}{L^2}$$



time-dependent HAL QCD method makes
this difficulty milder

$$\Delta E \simeq m_\pi$$



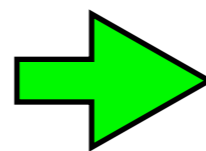
remaining t-dependence of the potential

1. Inelastic contributions (including excited states of one baryon)

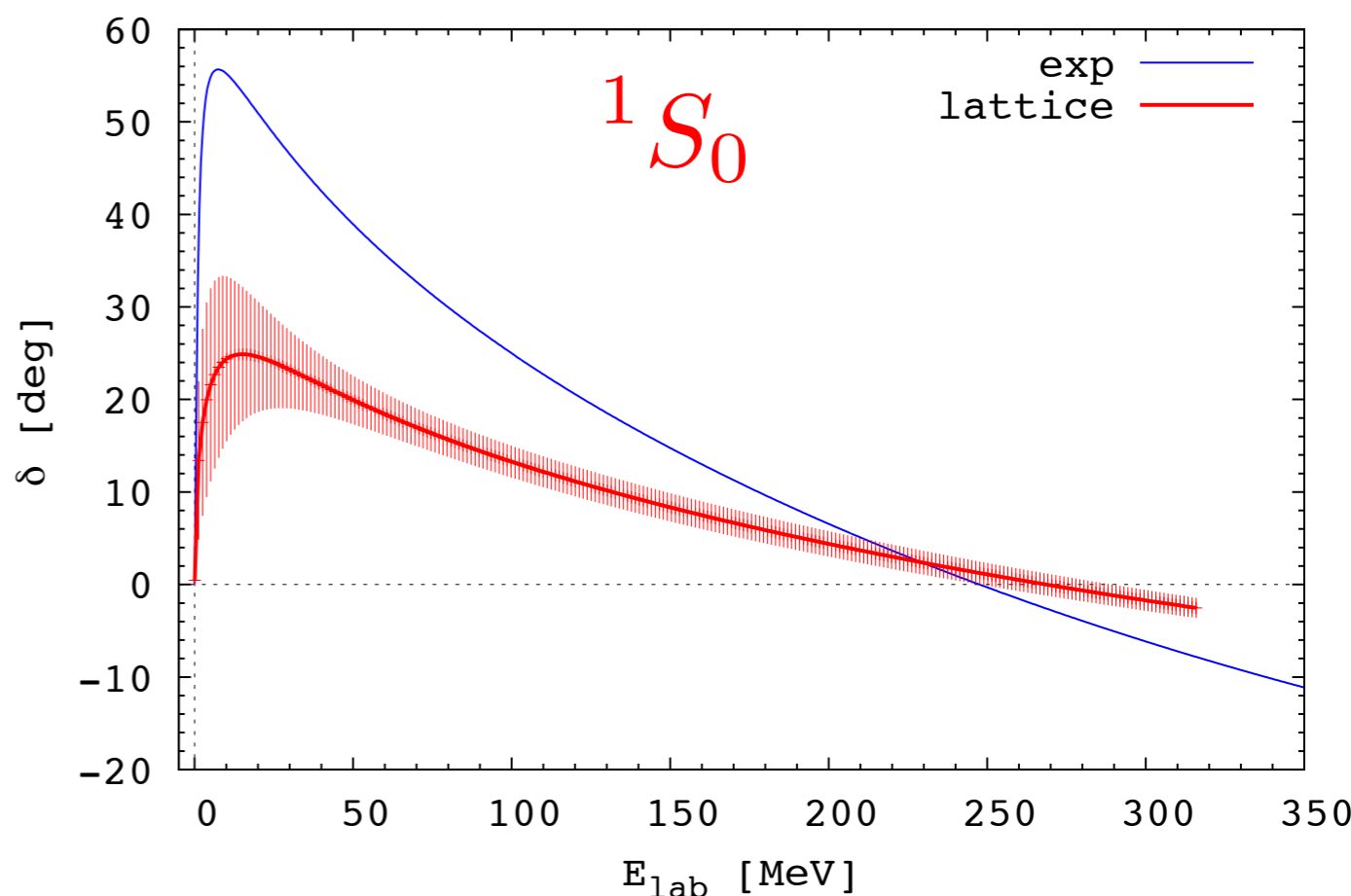
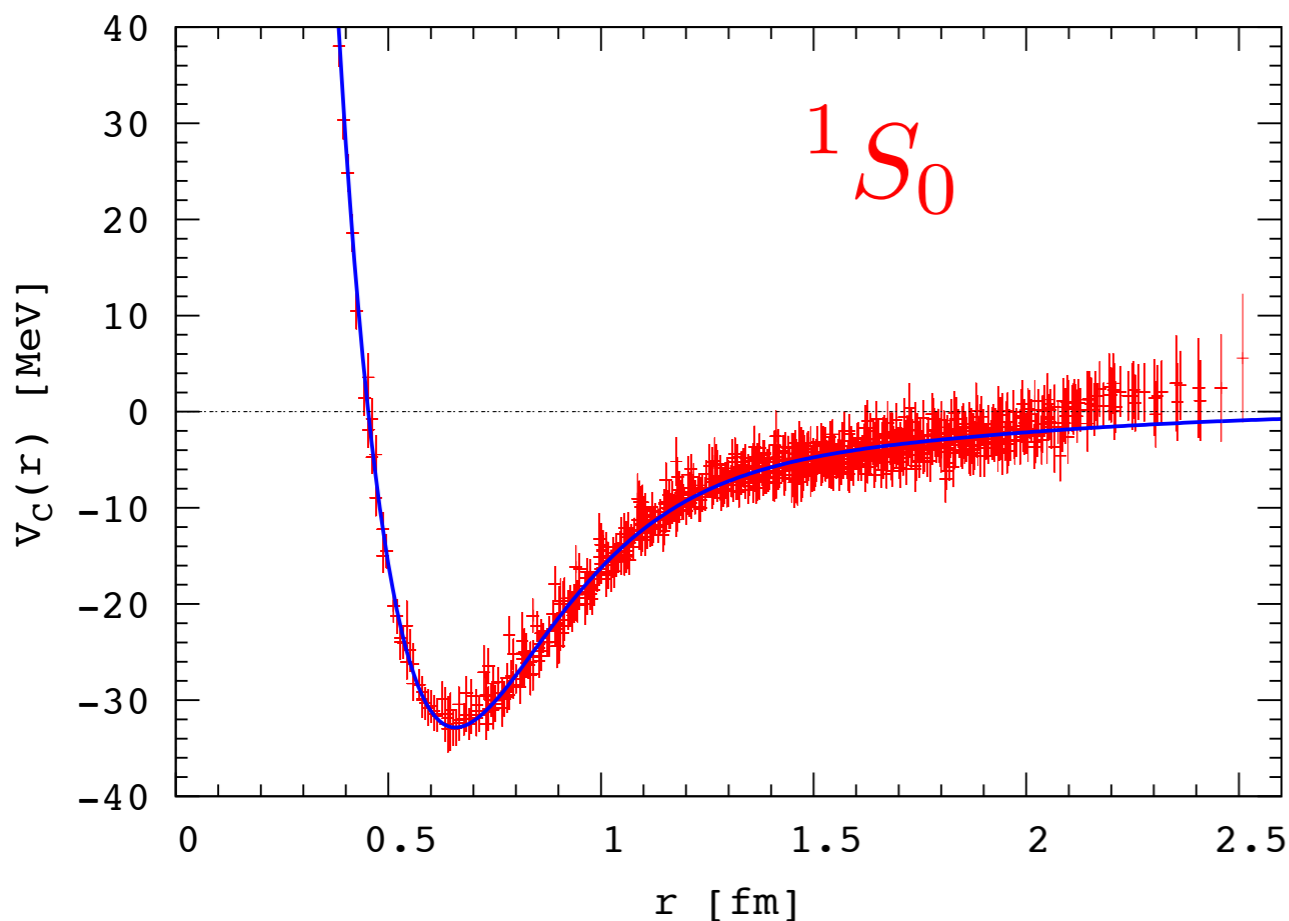
$$R(\mathbf{r}, t) = F(\mathbf{r}, t) / G_N(t)^2$$

2. Higher order terms in the derivative expansion

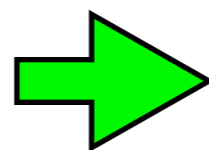
NN potential



phase shift

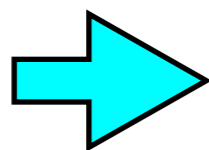


Qualitative features of NN potential are reproduced.



It has a reasonable shape. The strength is weaker due to the heavier quark mass.

Need calculations at physical quark mass on "K" computer.



Doi's Talk.

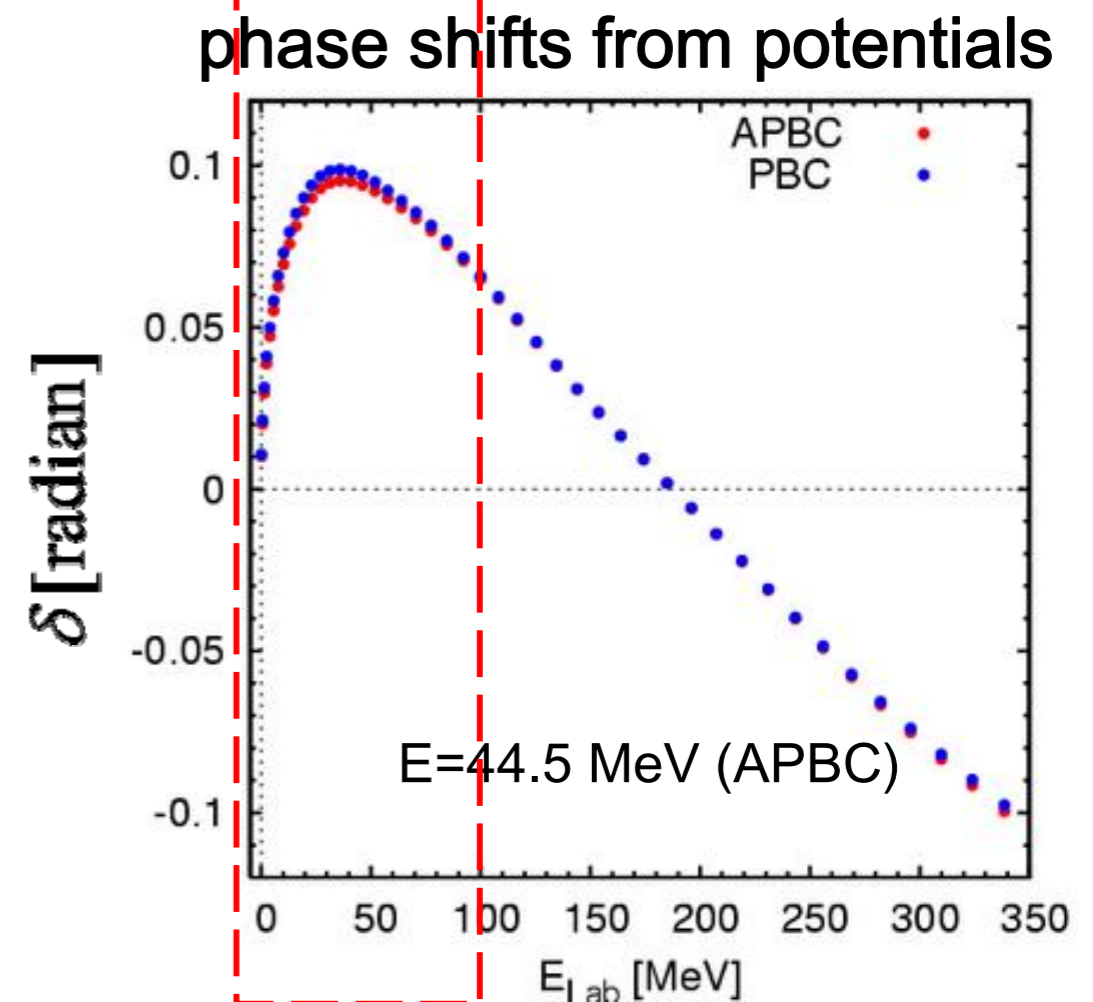
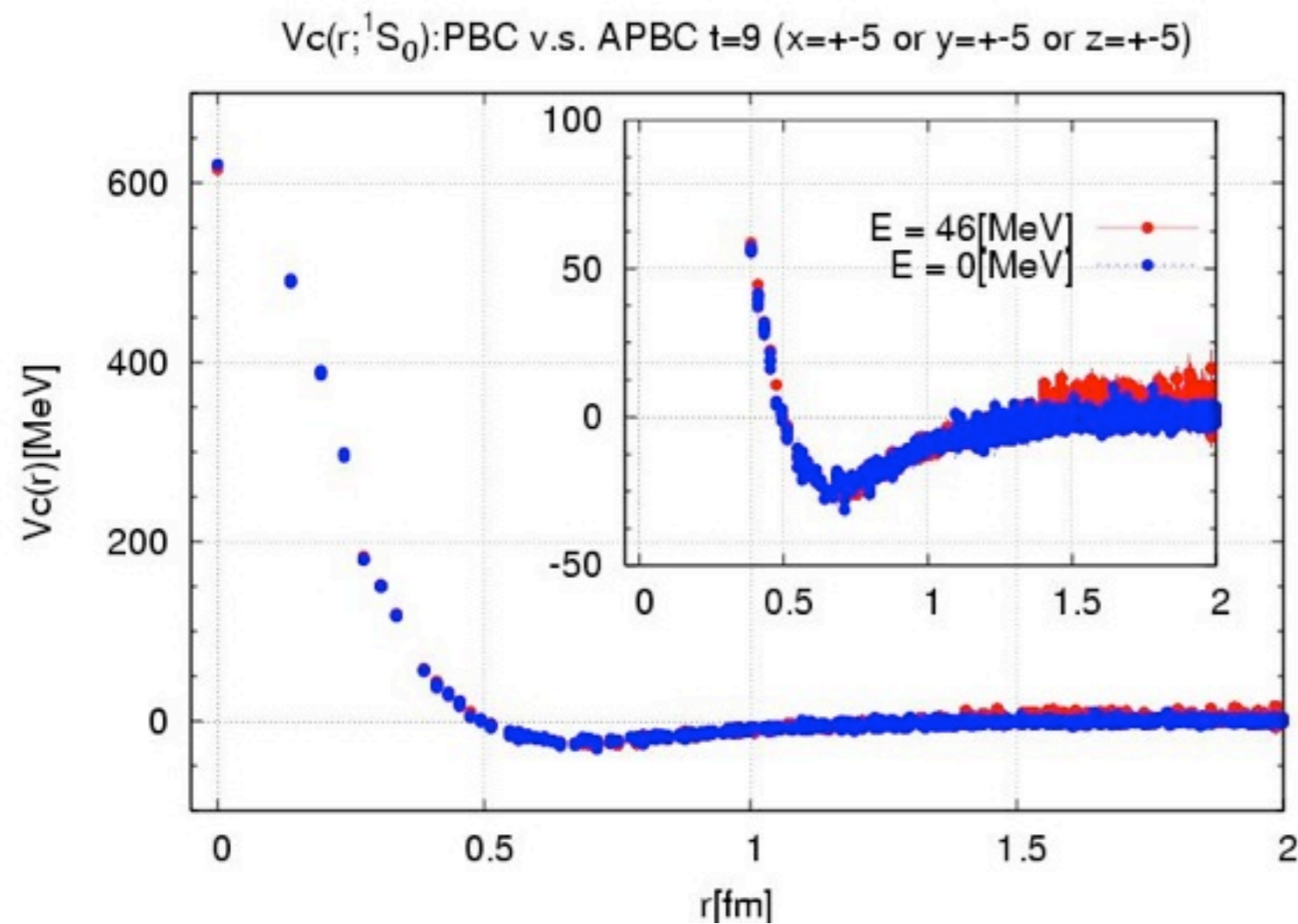


Convergence of velocity expansion

If the higher order terms are large, LO potentials determined from NBS wave functions at **different energy** become different.(cf. LOC of ChPT).

Numerical check in quenched QCD

K. Murano, N. Ishii, S. Aoki, T. Hatsuda [PTP 125 \(2011\)1225.](#)



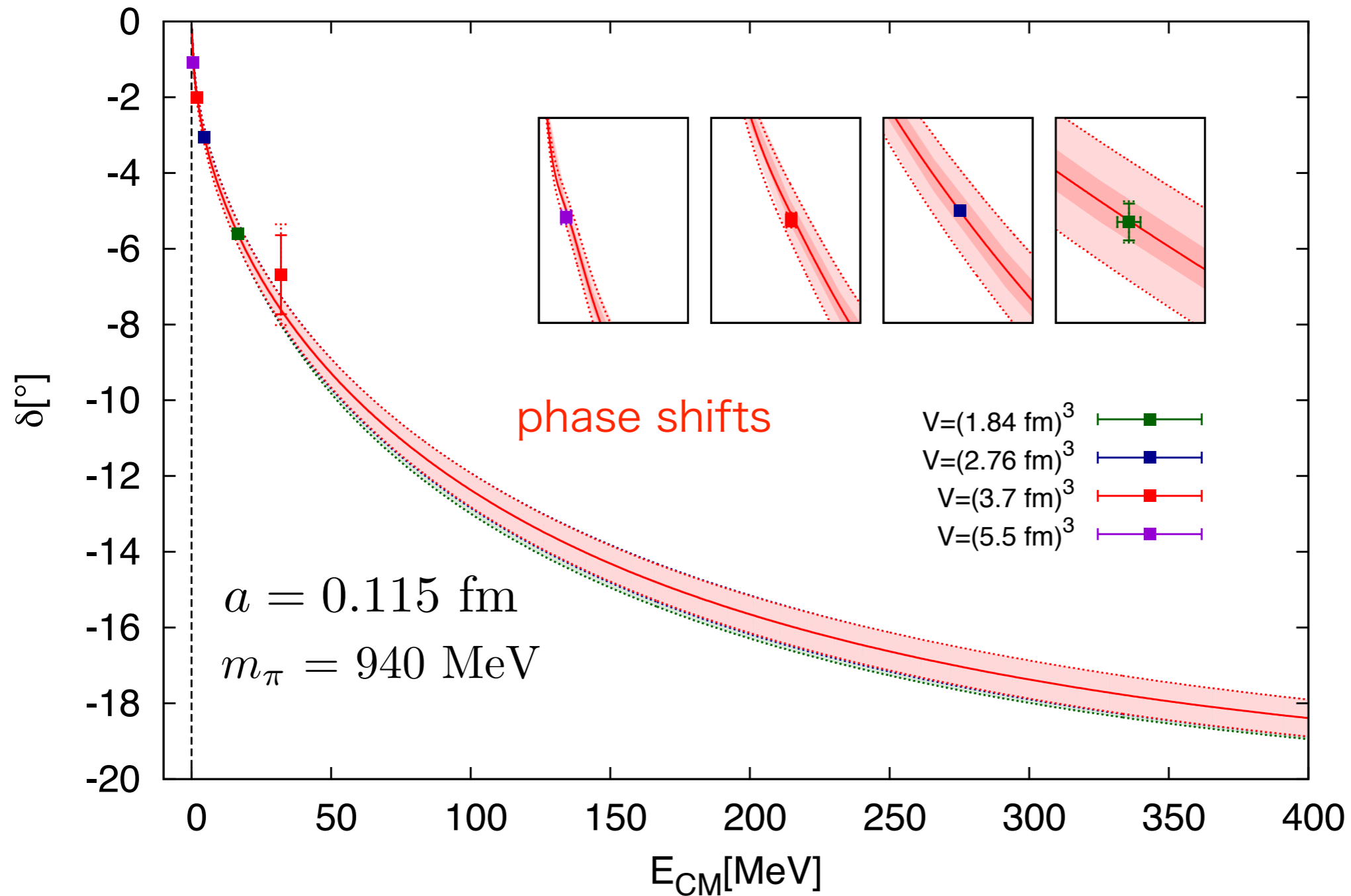
Higher order terms turn out to be very small at low energy in our scheme.

Recent comparison: [Iritani's talk](#)

Potential vs. Finite volume

Kurth, Ishii, Doi, Aoki & Hatsuda, JHEP 1312(2013)015

$I = 2$ $\pi\pi$ scattering in quenched QCD



both methods
agree very well.

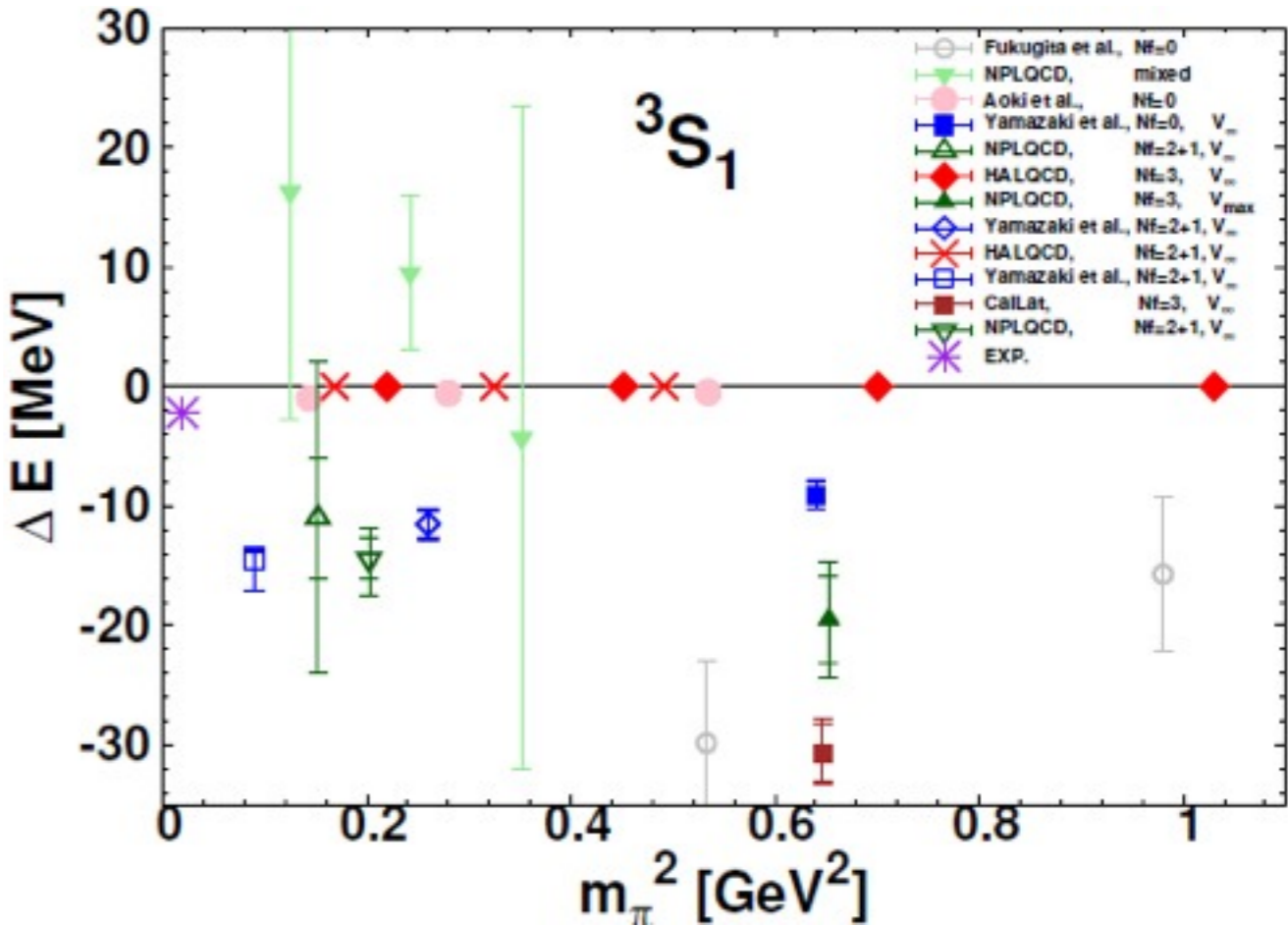
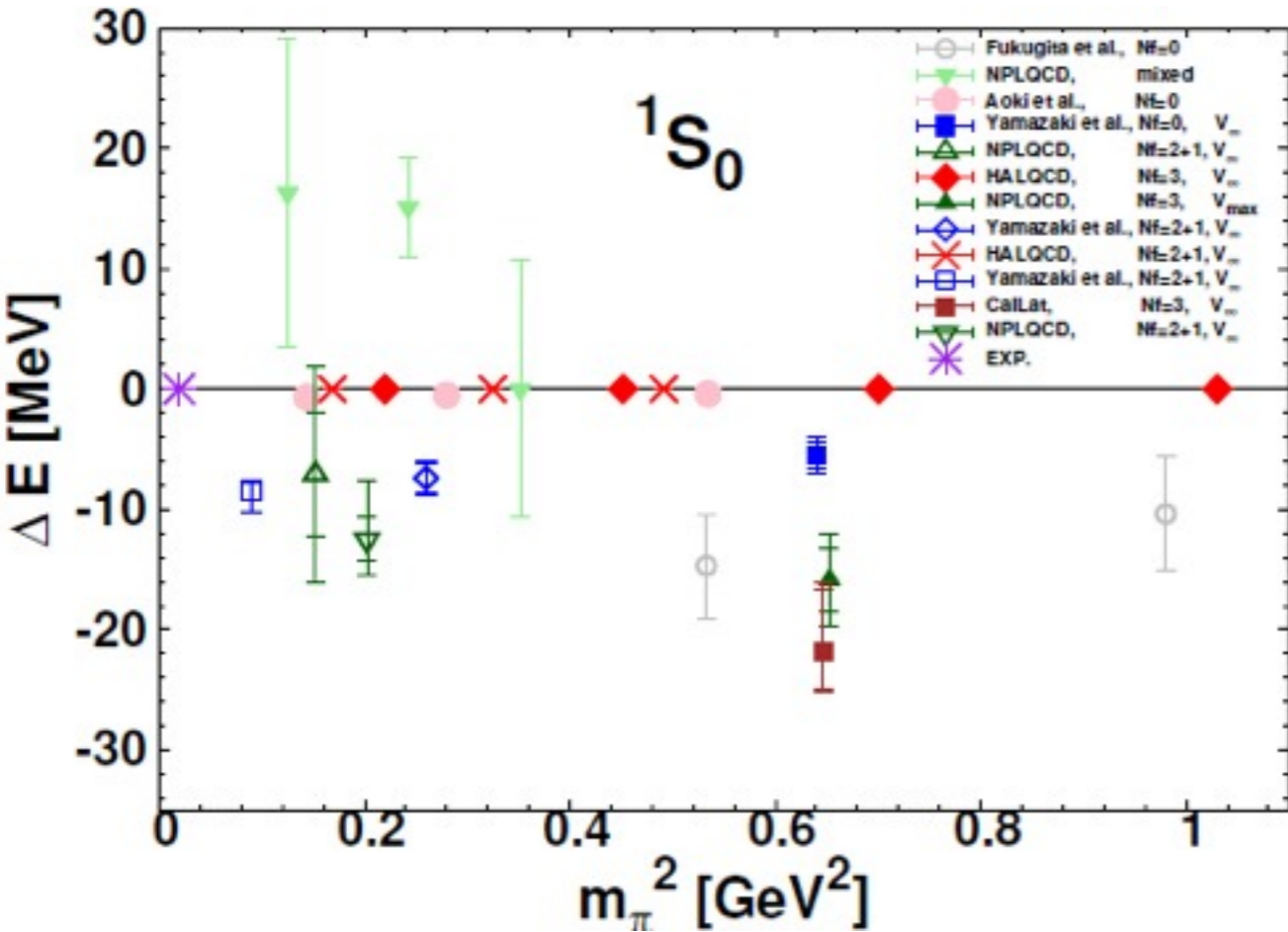
This establishes a validity of the potential method and shows a good convergence of the velocity expansion.

Potential vs. Direct

Reviewed in T.Doi, PoS LAT2012,009 (+ updates)

“di-neutron”

“deuteron”



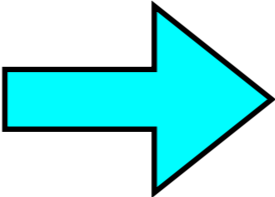
HAL (potential) method (HAL) :

unbound

Direct method (PACS-CS (Yamazaki et al.)/NPL/CalLat):

bound

Which is correct ?



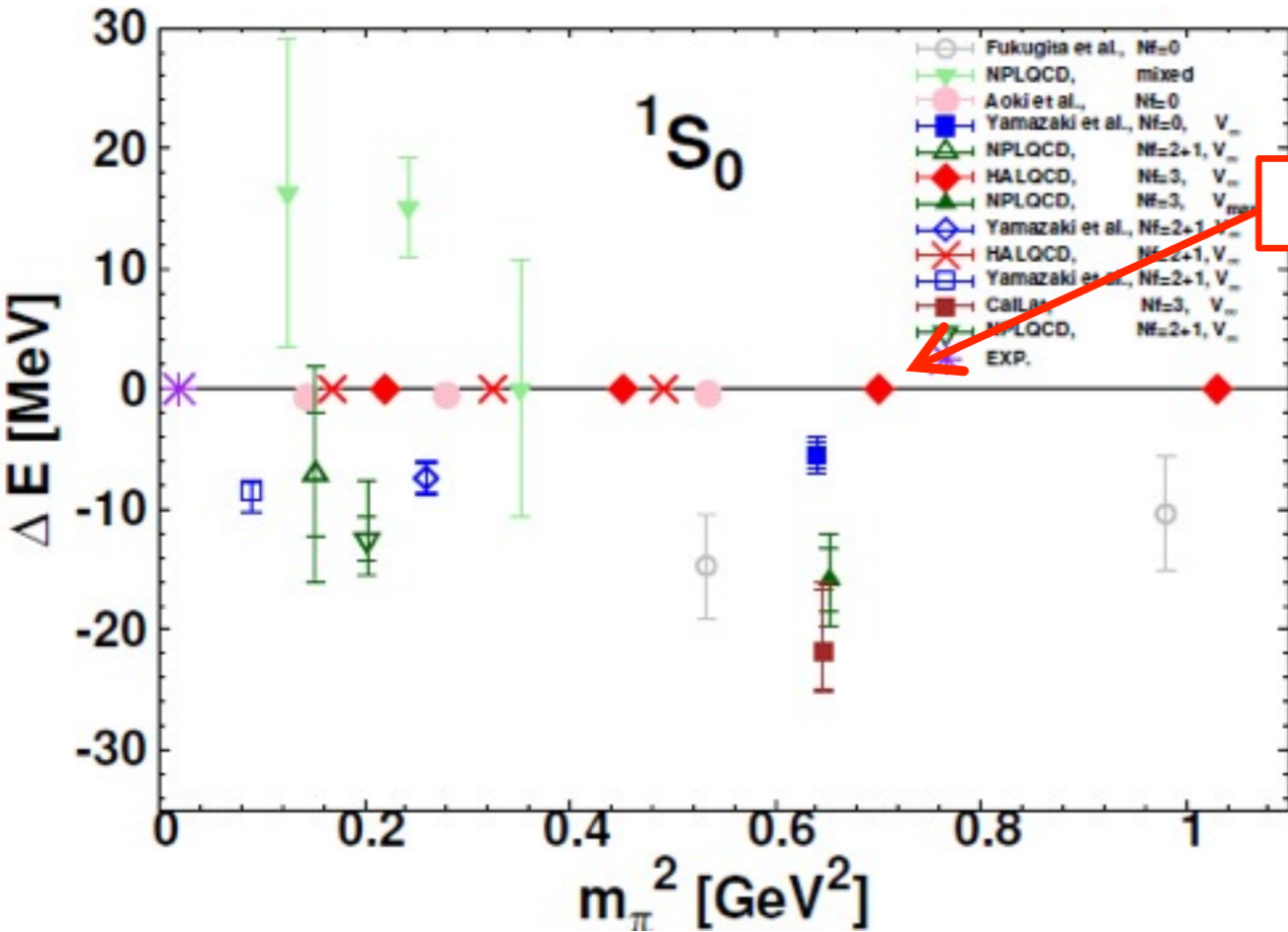
Iritani's talk

Potential vs. Direct

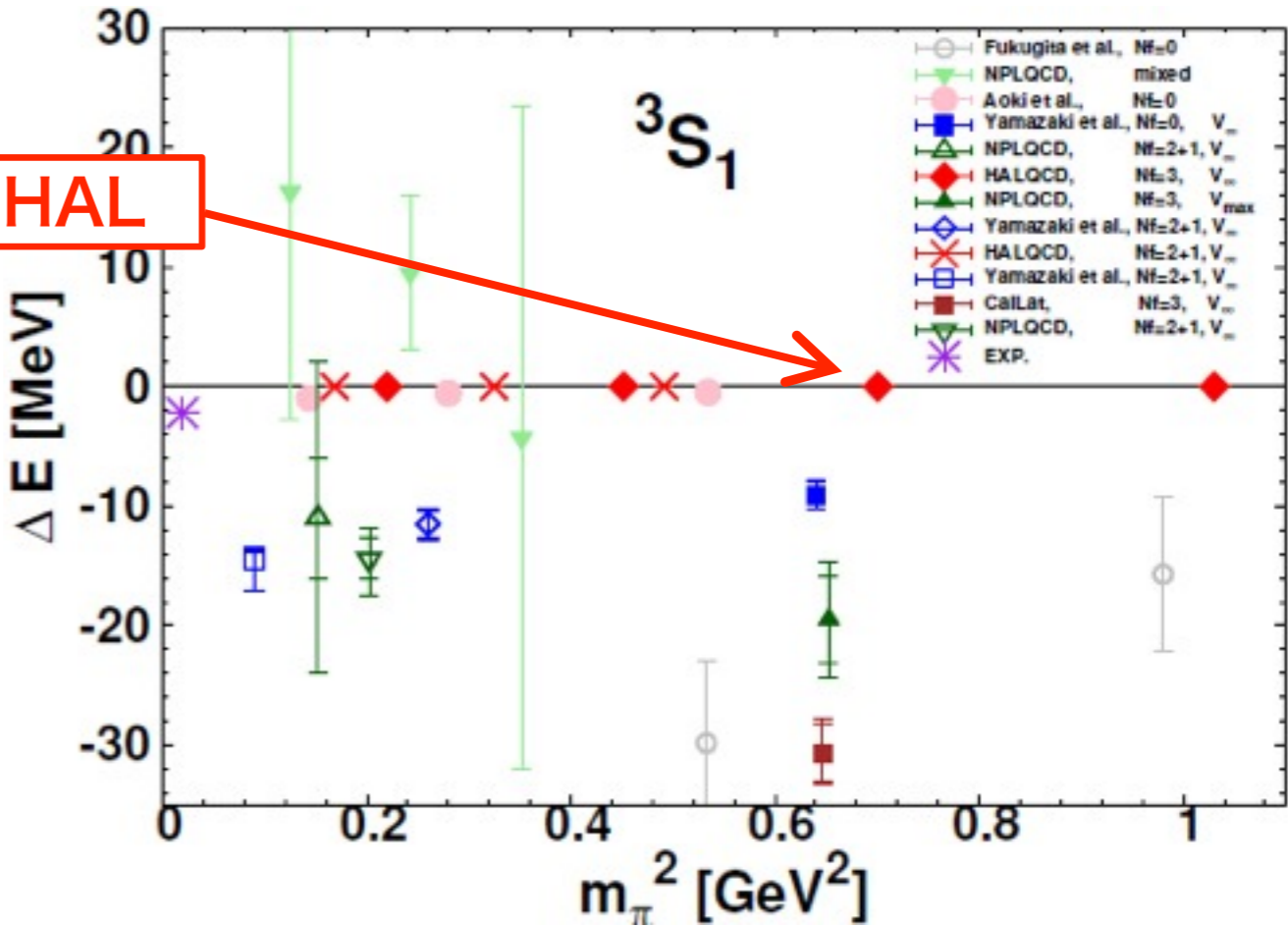
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HAL



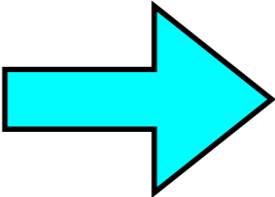
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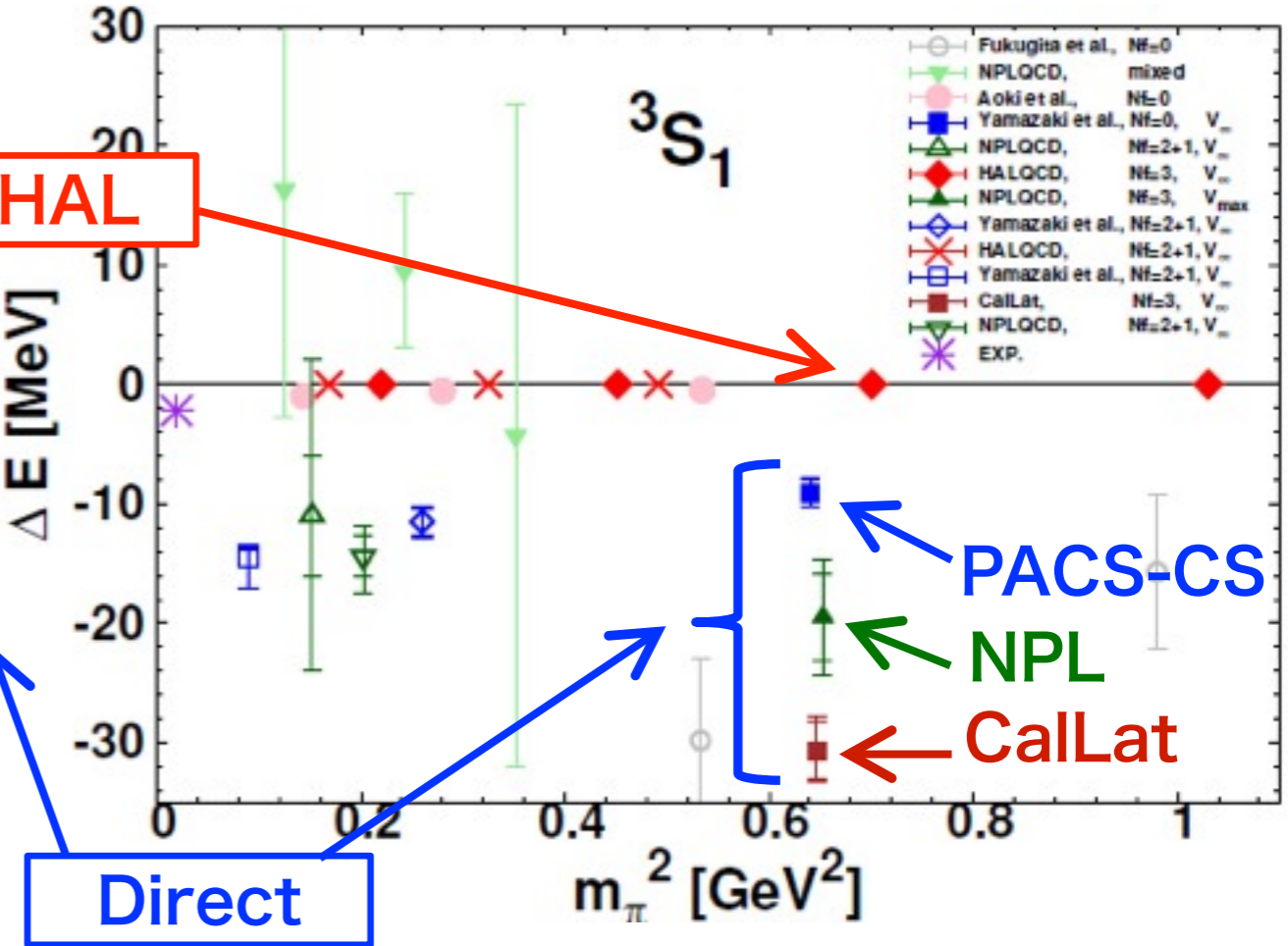
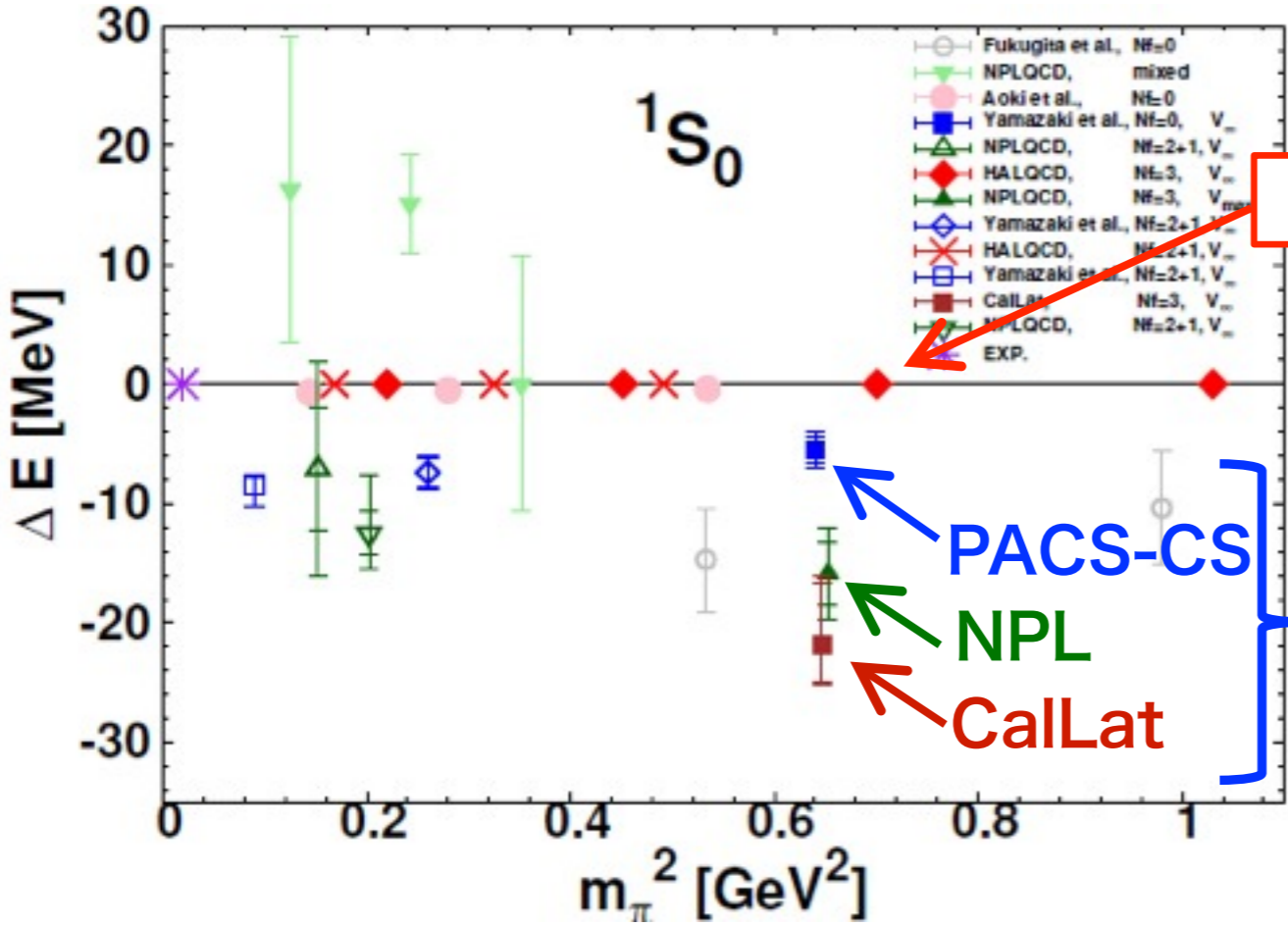
Iritani's talk

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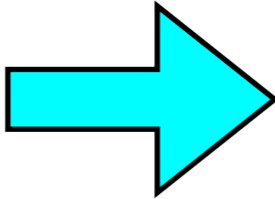
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unbound

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bound

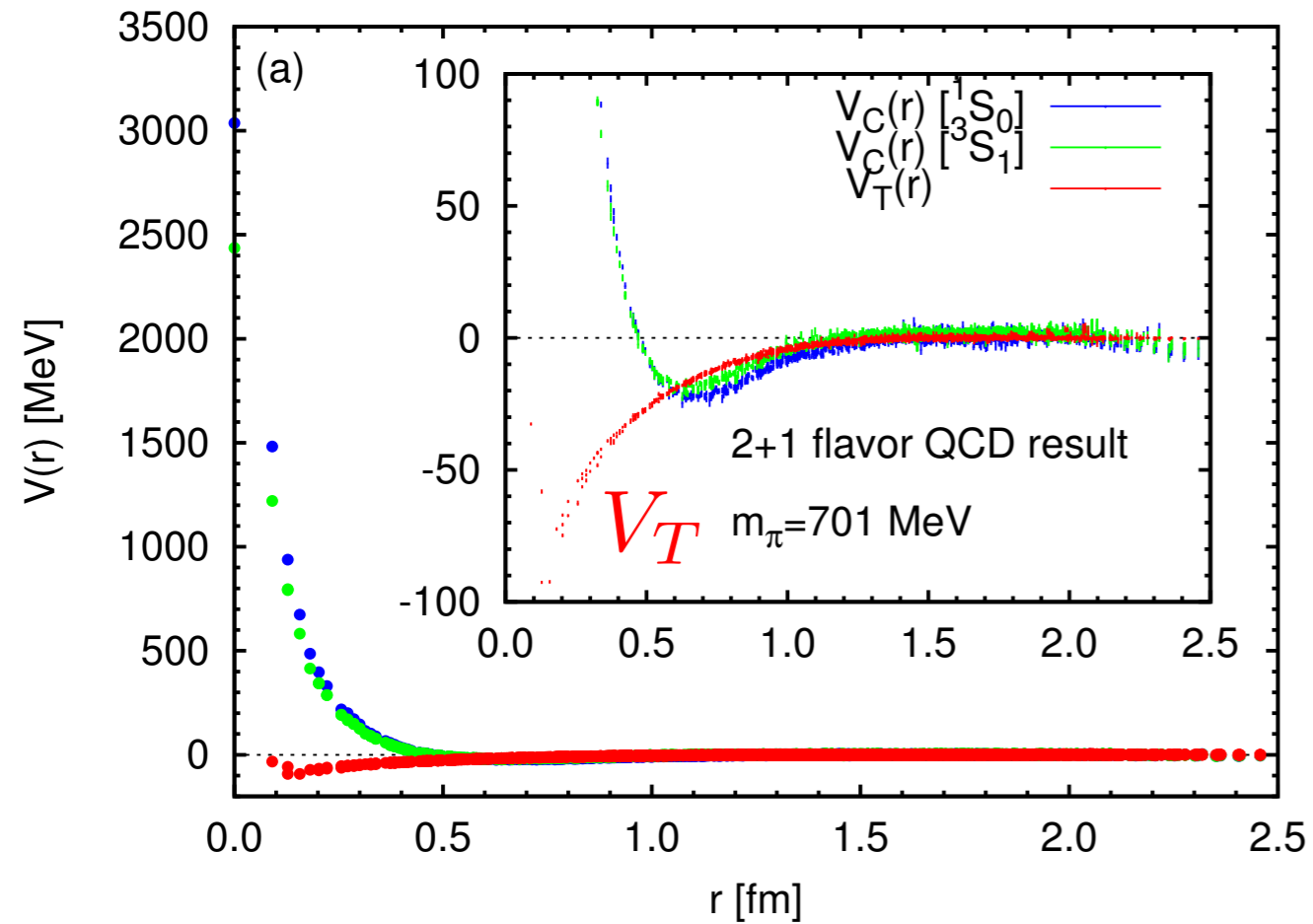
Which is correct ?



Iritani's talk

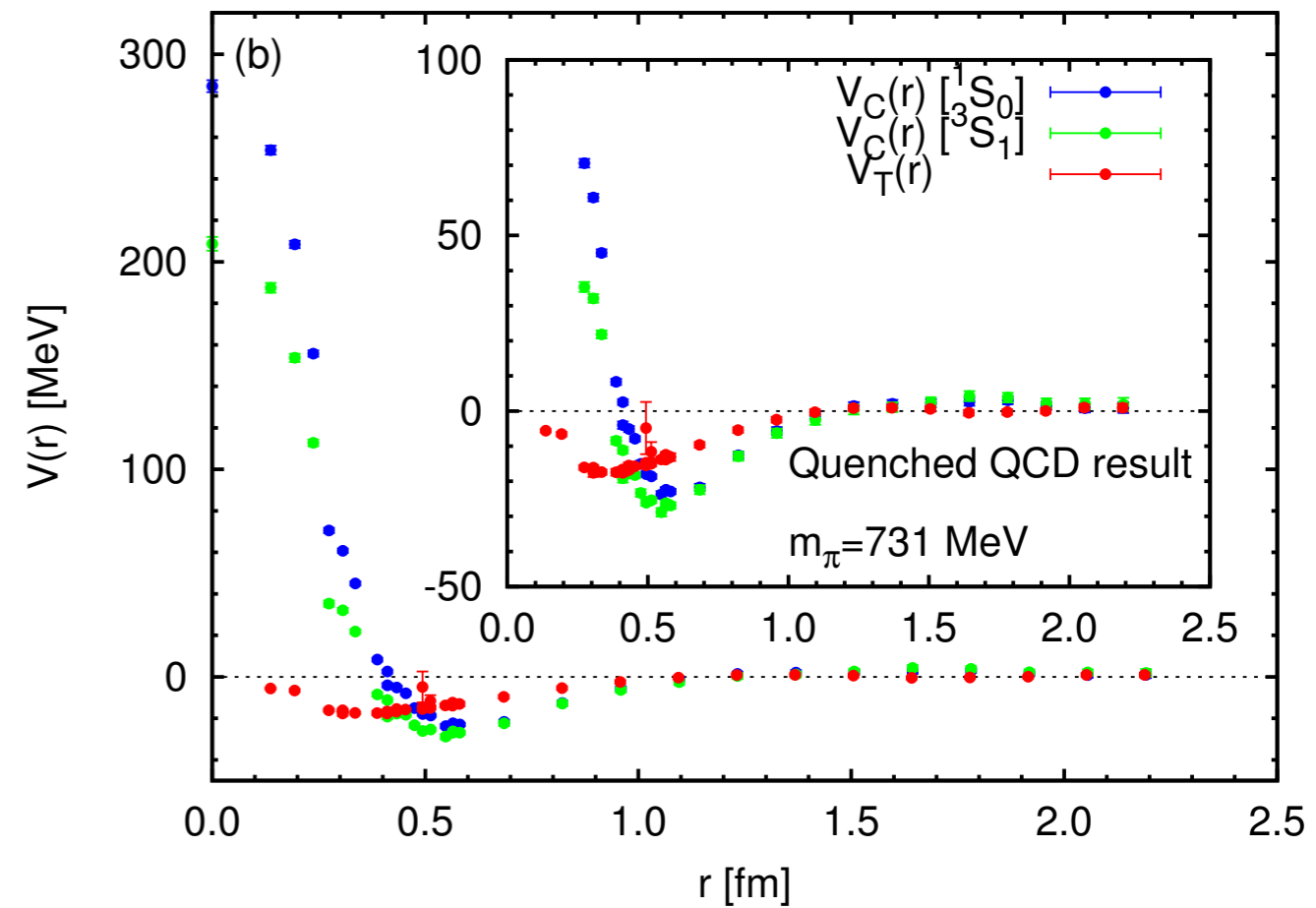
LO Tensor potential

full QCD



$a \simeq 0.091$ fm $L \simeq 2.9$ fm

quenched QCD



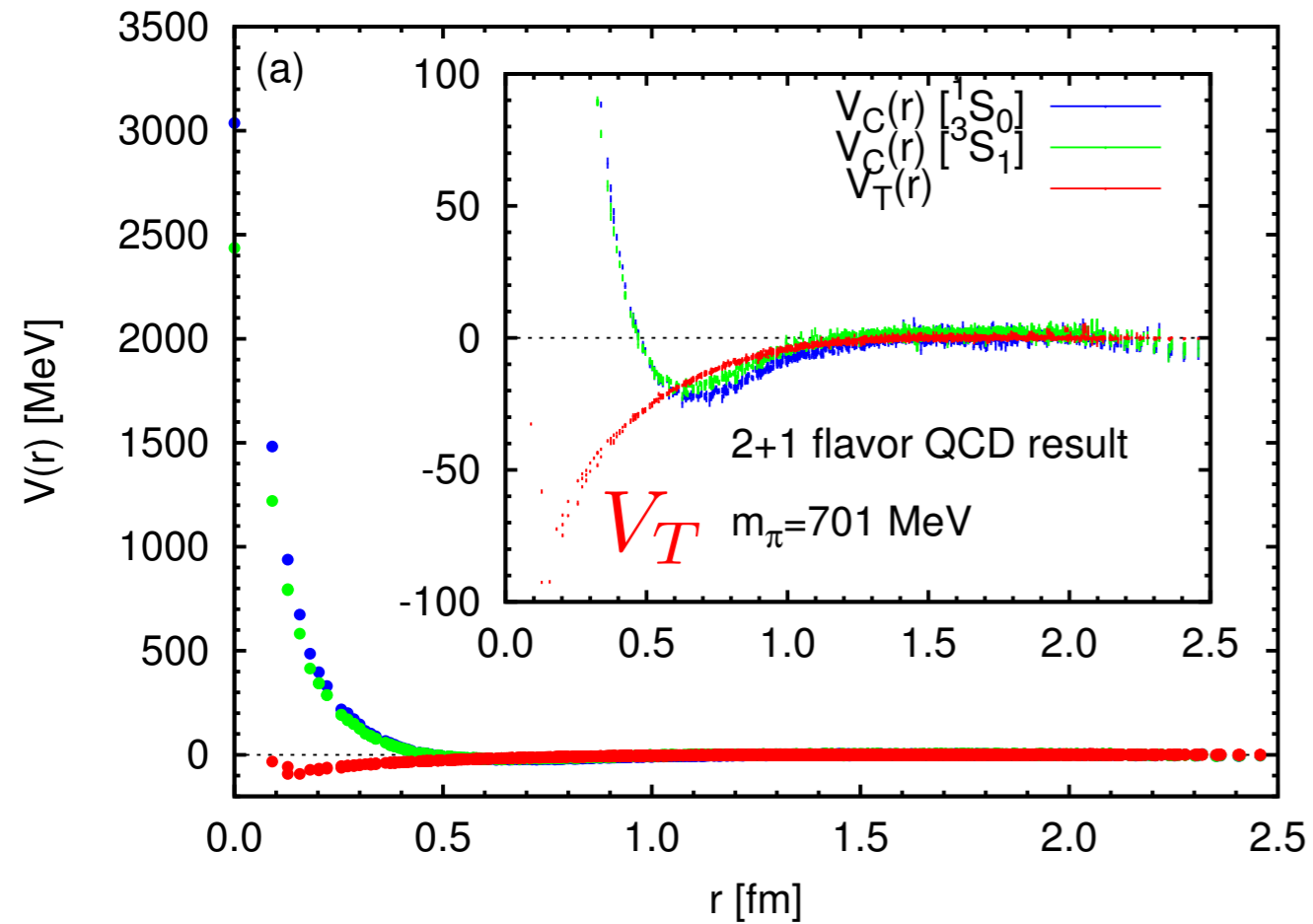
$a \simeq 0.137$ fm $L \simeq 4.4$ fm

- no repulsive core in the tensor potential.
- the tensor potential is enhanced in full QCD

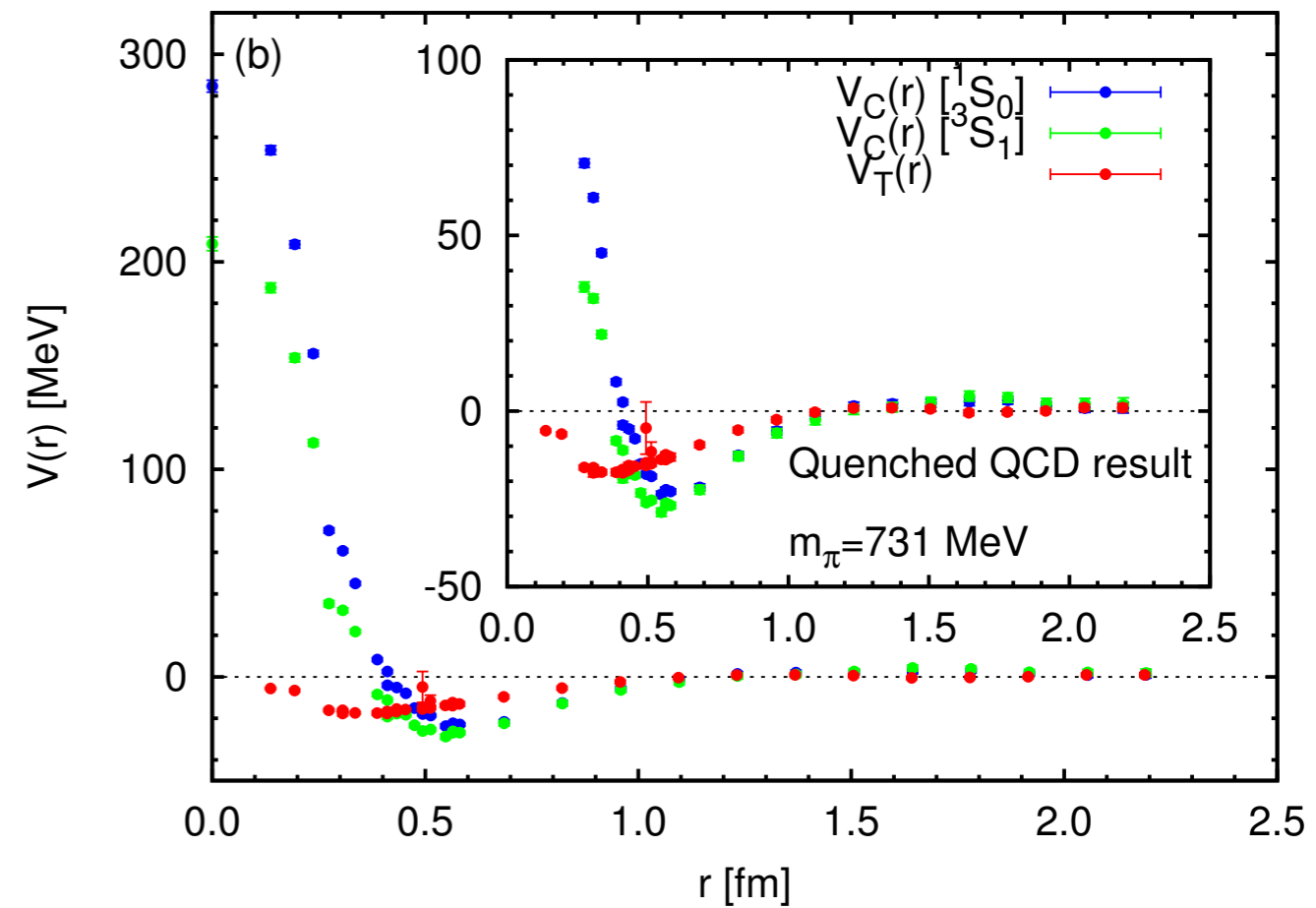
LO Tensor potential

full QCD

quenched QCD

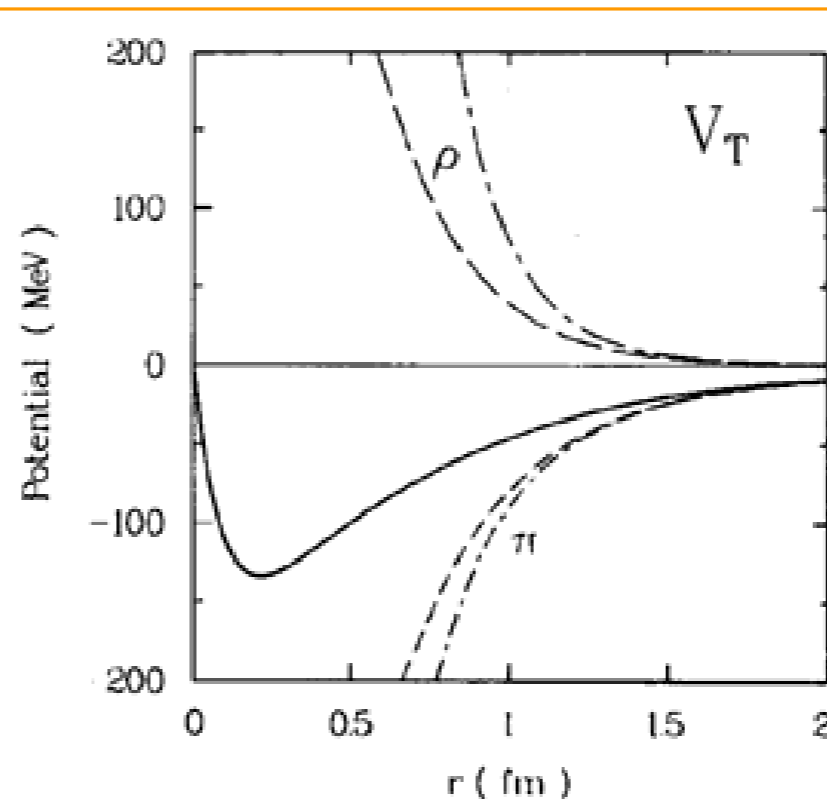


$a \simeq 0.091 \text{ fm}$ $L \simeq 2.9 \text{ fm}$



$a \simeq 0.137 \text{ fm}$ $L \simeq 4.4 \text{ fm}$

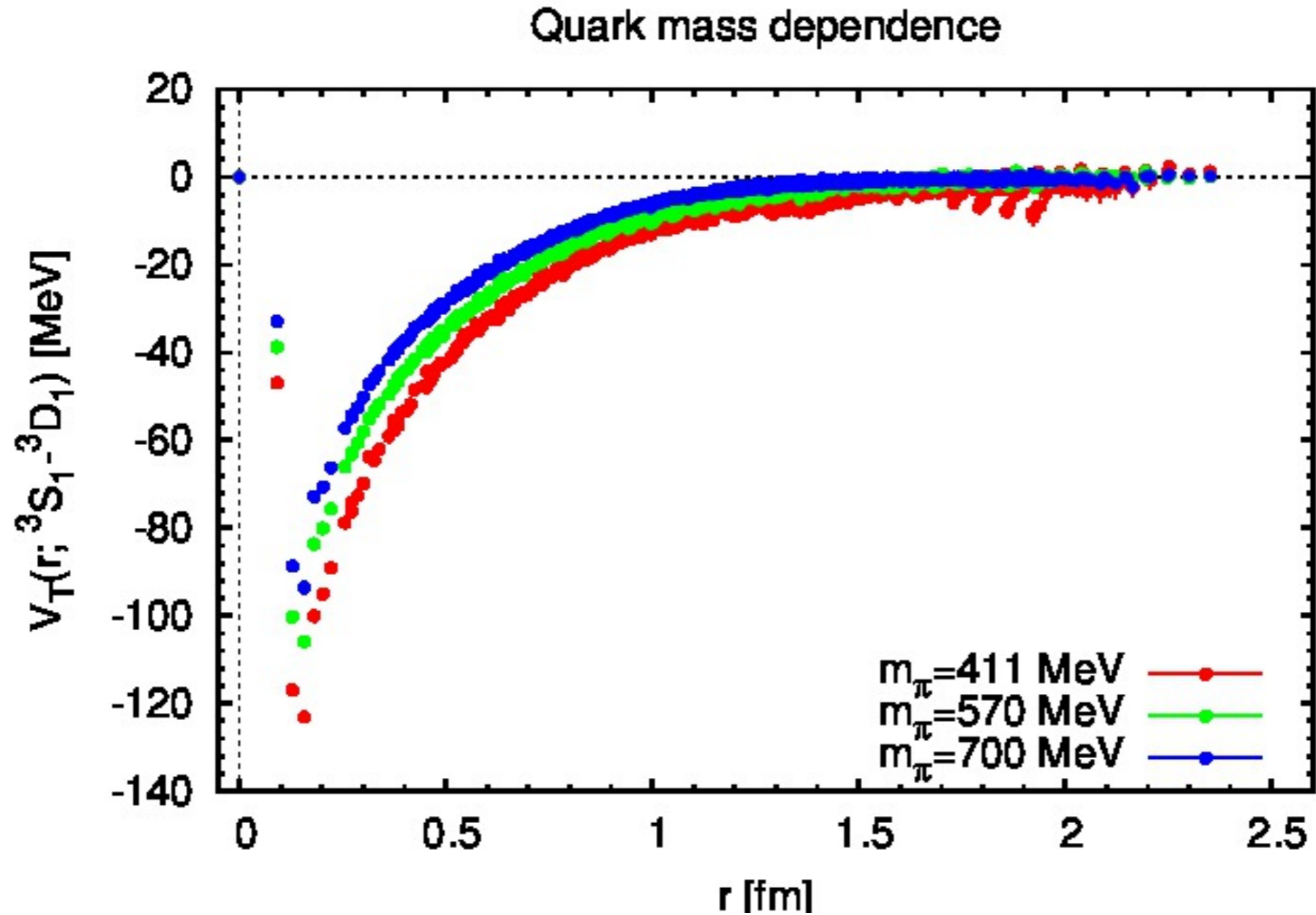
- no repulsive core in the tensor potential.
- the tensor potential is enhanced in full QCD



from
R.Machleidt,
Adv.Nucl.Phys.**19**

Fig. 3.7. The contributions from π and ρ (dashed) to the $T = 0$ tensor potential. The solid line is the full potential. The dash-dot lines are obtained when the cutoff is omitted.

Quark mass dependence (full QCD)

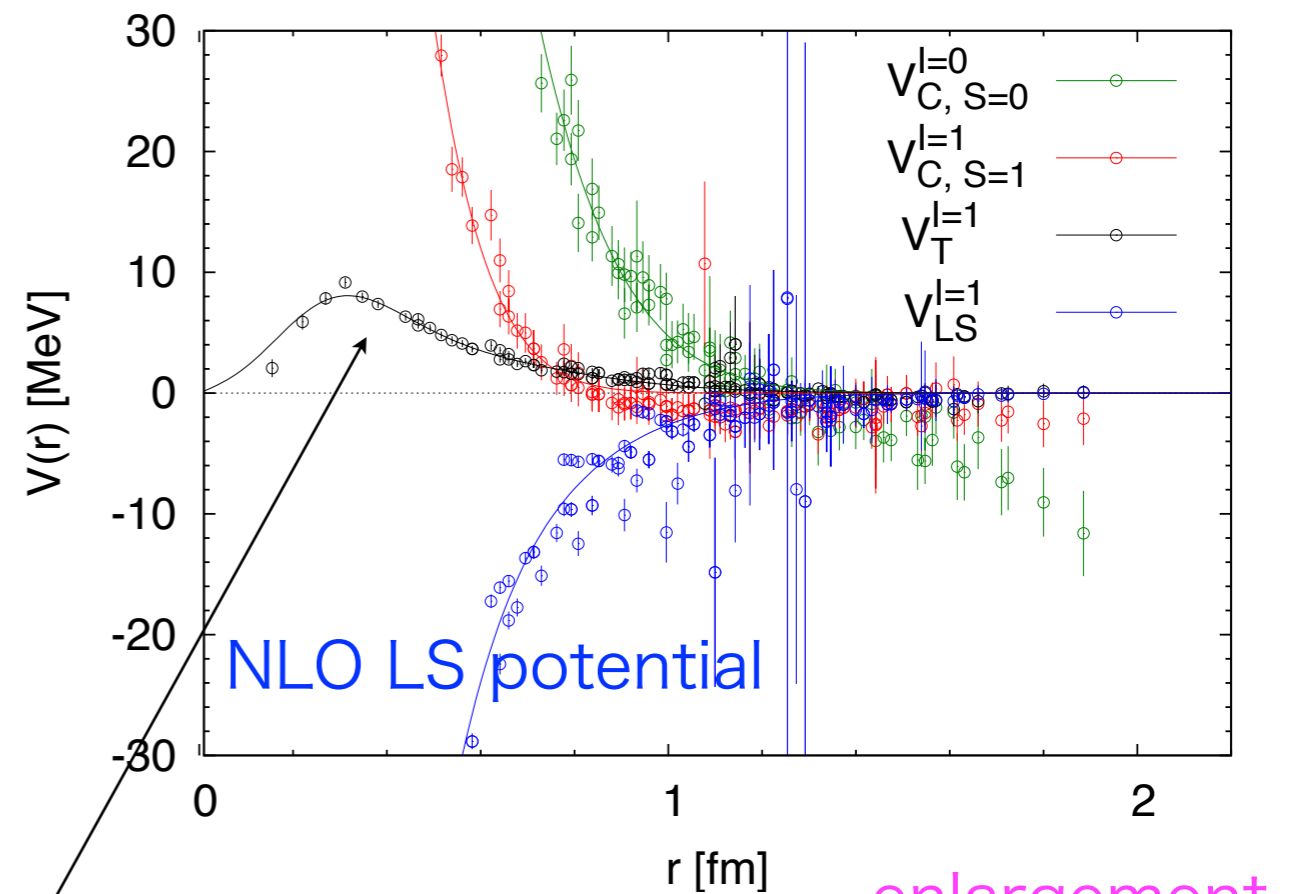
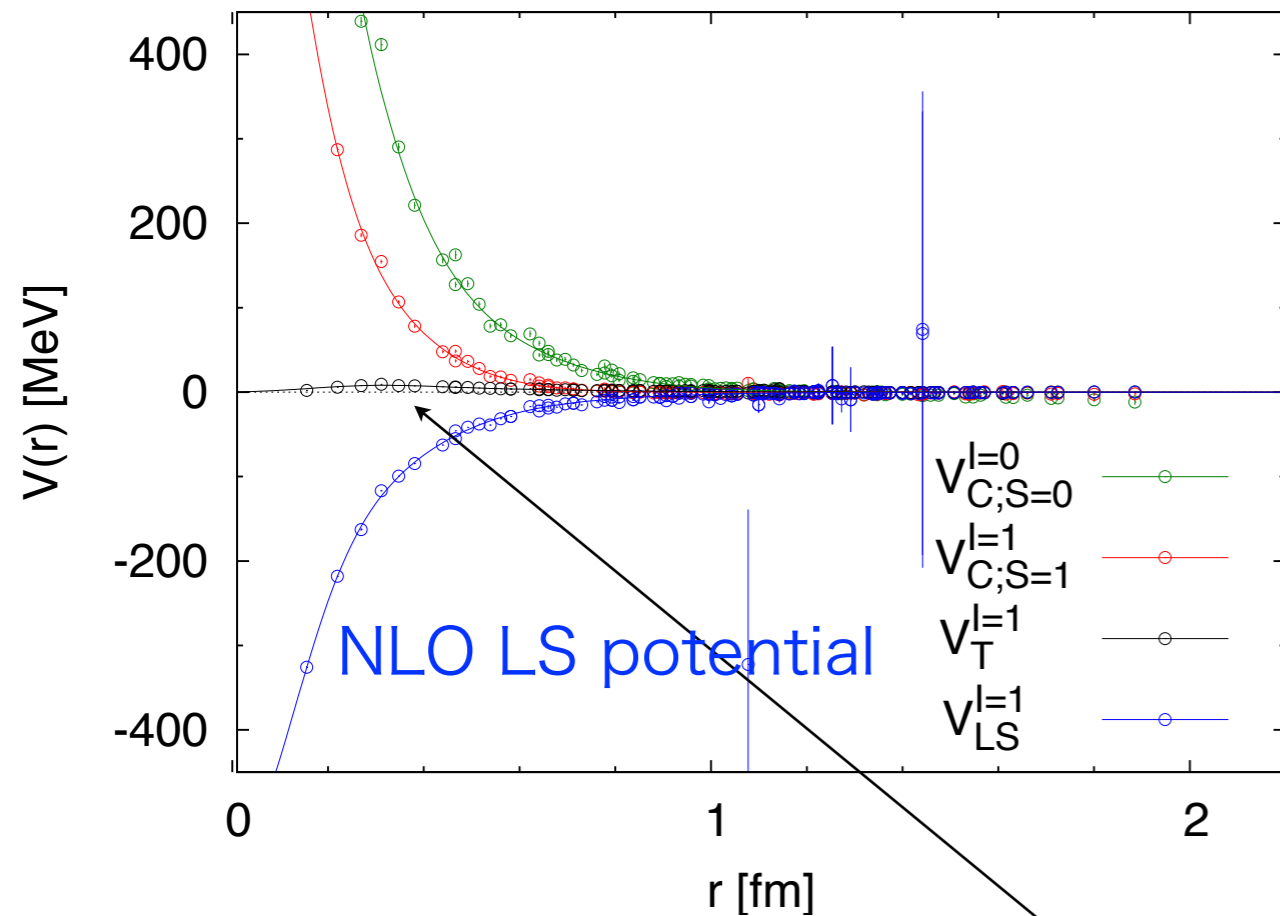


- the tensor potential increases as the pion mass decreases.
 - manifestation of one-pion-exchange ?

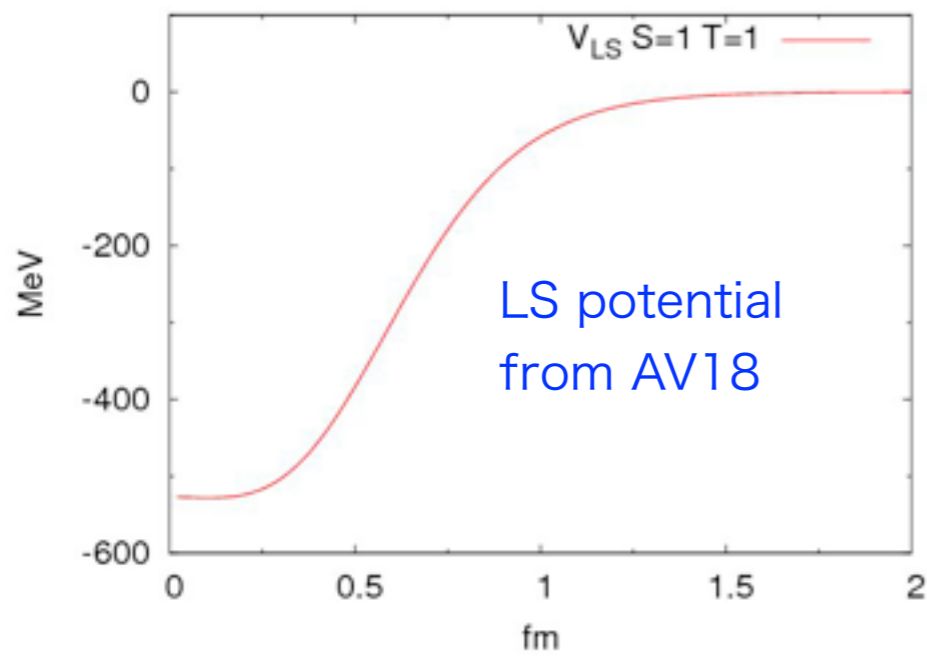
Negative parity potentials

$a = 0.16$ fm, $L = 2.5$ fm, $m_\pi = 1100$ MeV

Murano et al. (HAL QCD), PLB735(2014)19



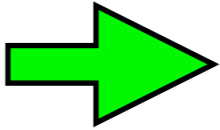
enlargement



tensor potential is very weak.

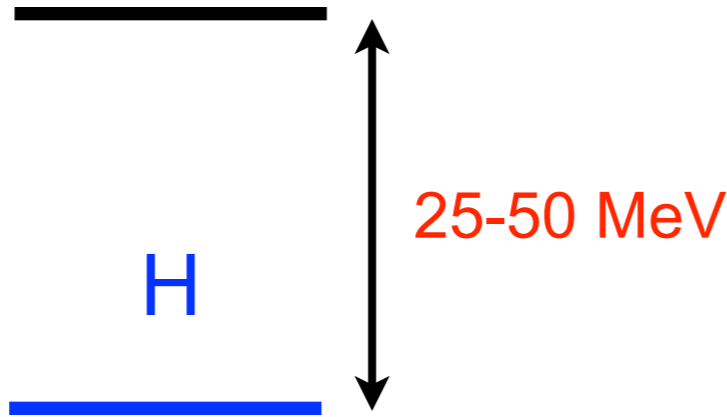
3. Recent results

H-dibaryon



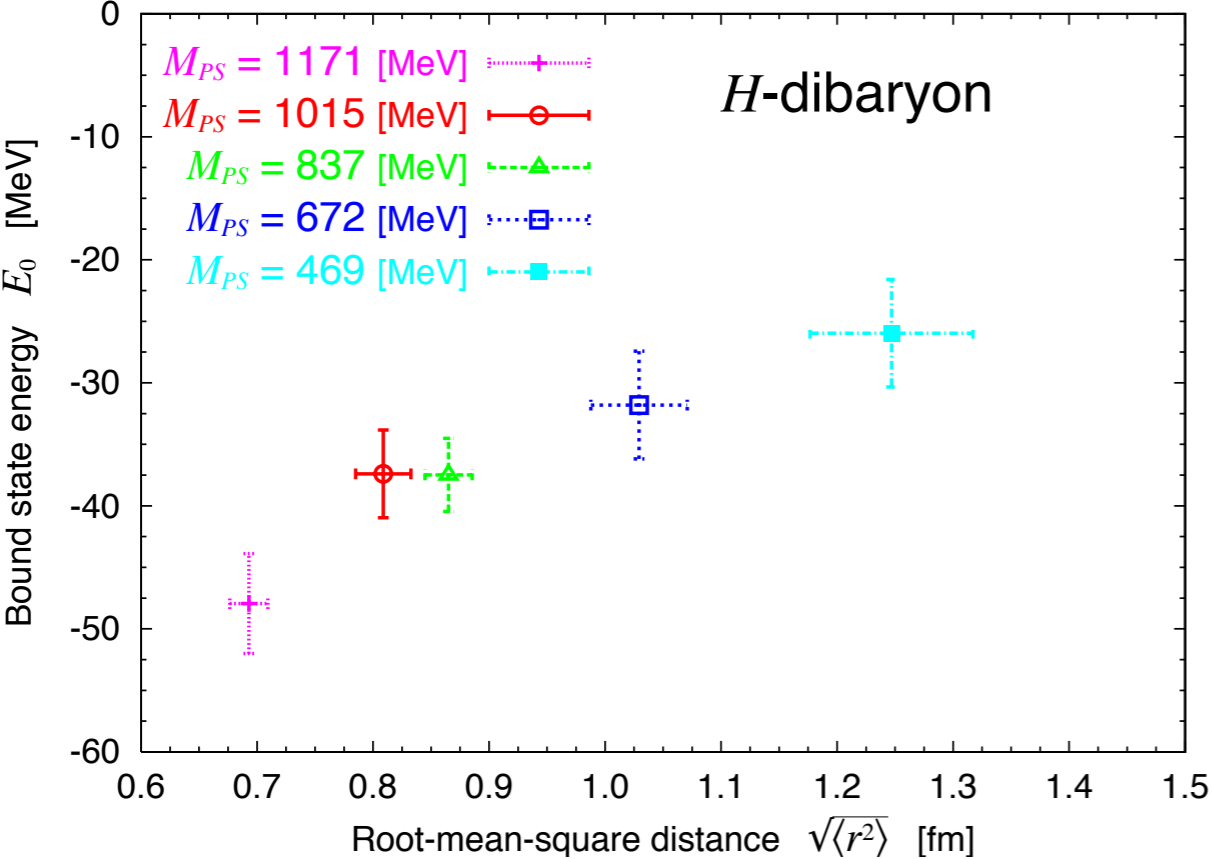
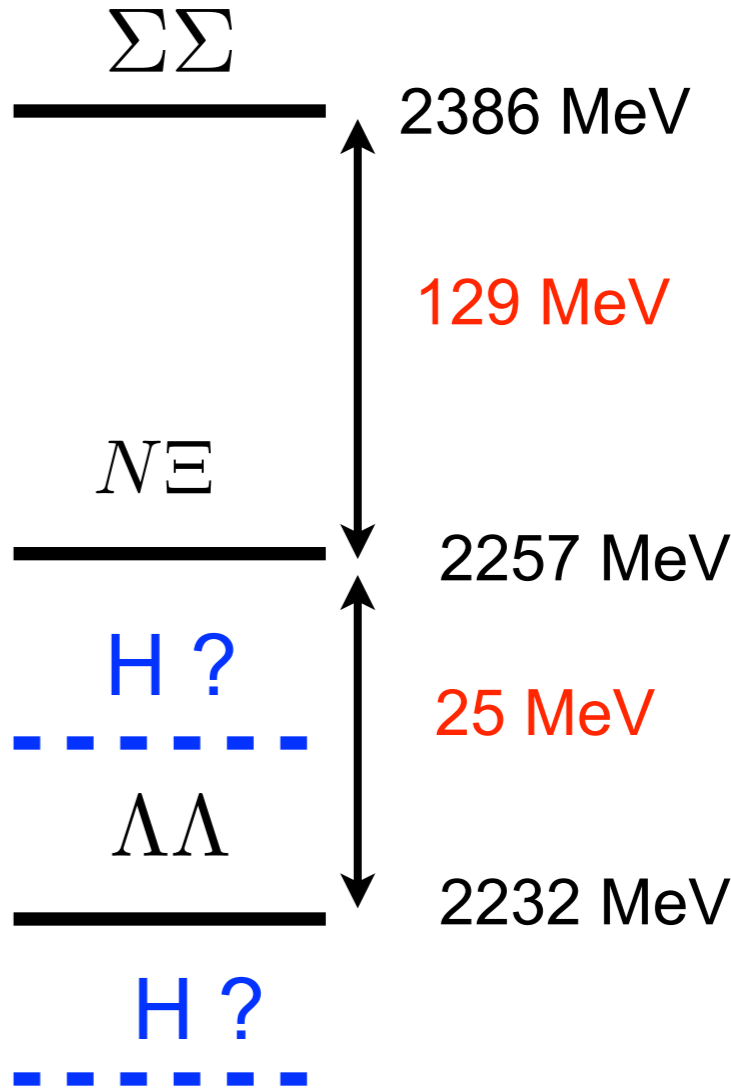
SU(3) limit

$\Lambda\Lambda - N\Xi - \Sigma\Sigma$



Real world

$m_u = m_d \neq m_s$



coupled channel analysis is needed.

S. Aoki, *et al.*, Proc. Jpn. Acad. Ser. B, 87 (2011) 509.

Results from HAL QCD Collaboration

Sasaki for HAL QCD Collaboration

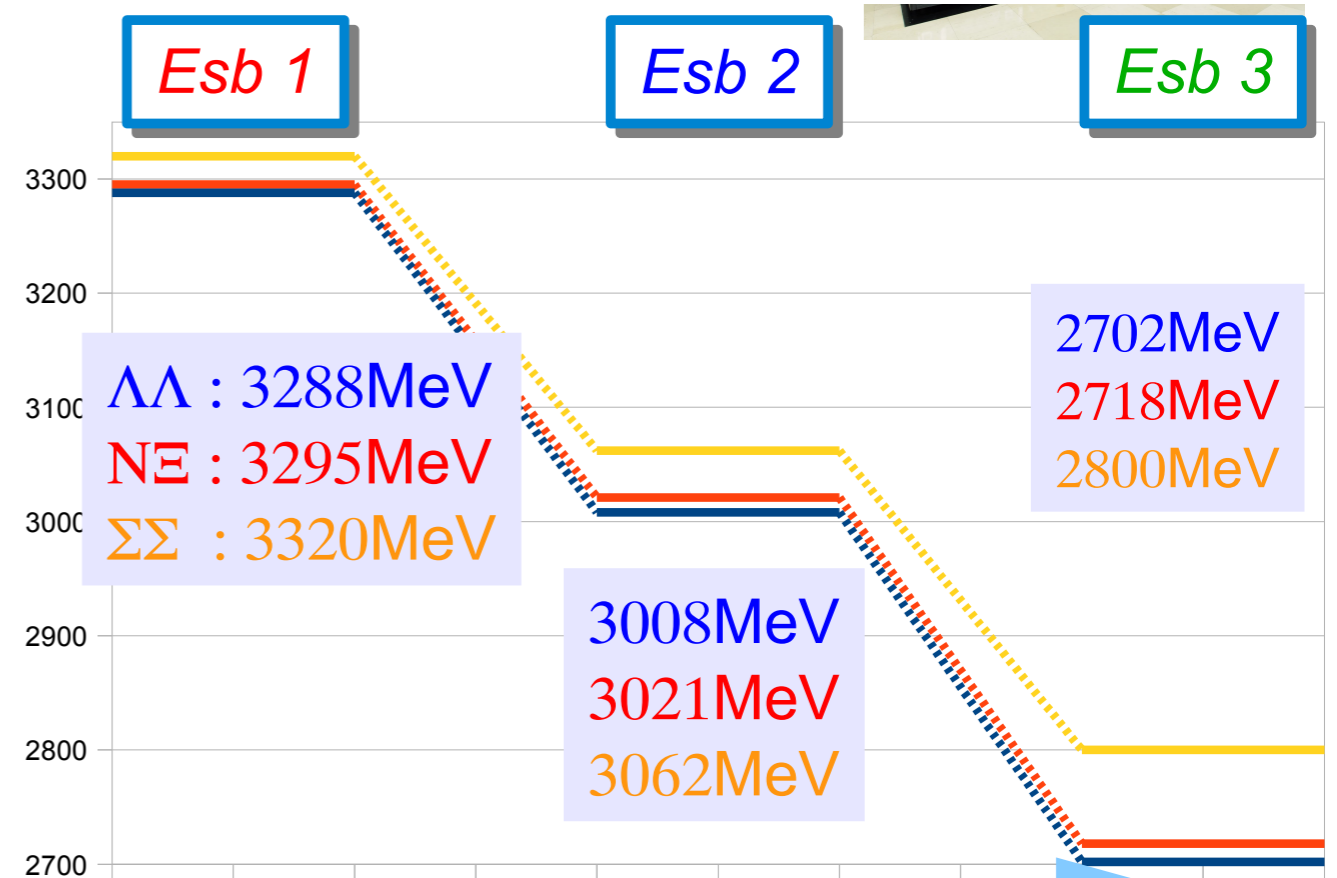
$N_f = 2 + 1$ full QCD with $L = 2.9$ fm

Gauge ensembles

In unit of MeV	<i>Esb 1</i>	<i>Esb 2</i>	<i>Esb 3</i>
π	701 ± 1	570 ± 2	411 ± 2
K	789 ± 1	713 ± 2	635 ± 2
m_π / m_K	0.89	0.80	0.65
N	1585 ± 5	1411 ± 12	1215 ± 12
Λ	1644 ± 5	1504 ± 10	1351 ± 8
Σ	1660 ± 4	1531 ± 11	1400 ± 10
Ξ	1710 ± 5	1610 ± 9	1503 ± 7

u,d quark masses lighter

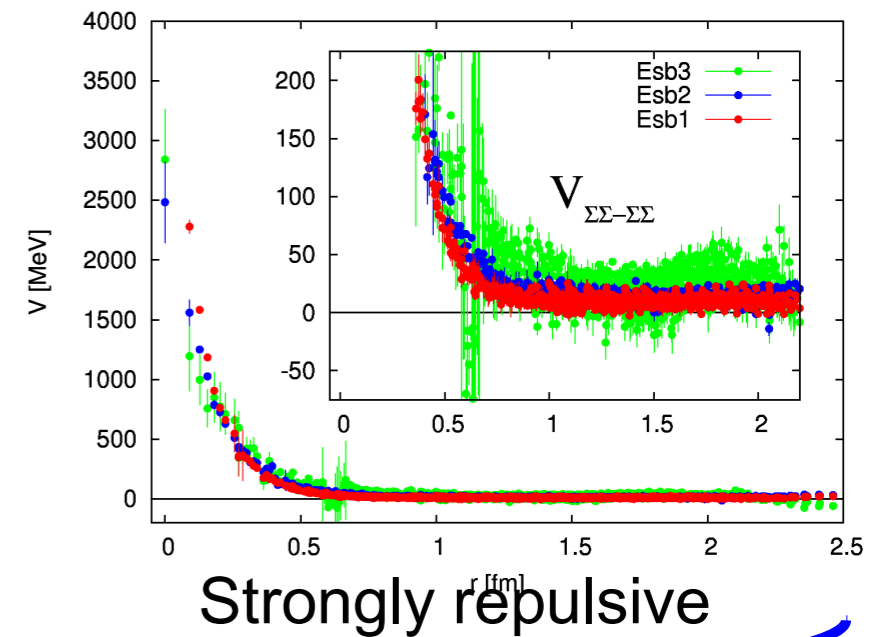
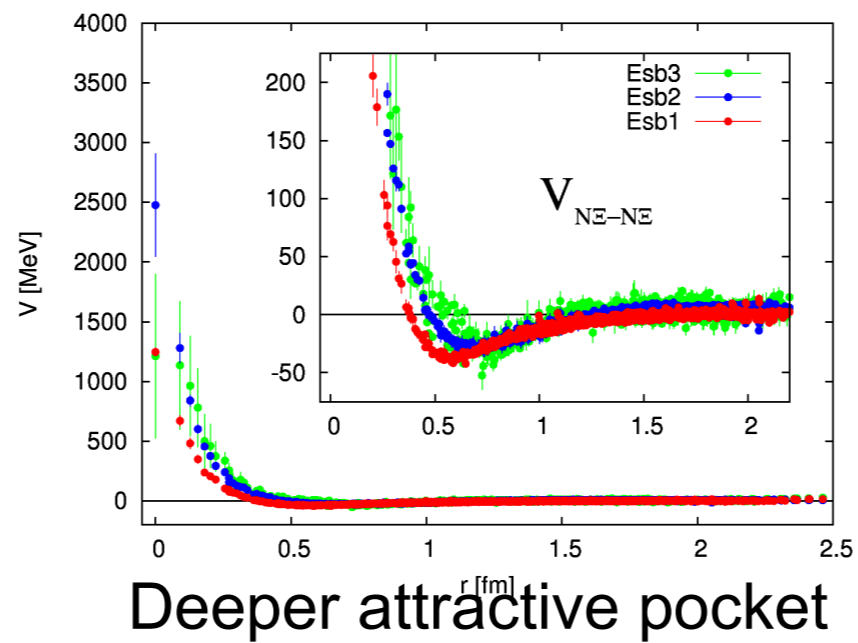
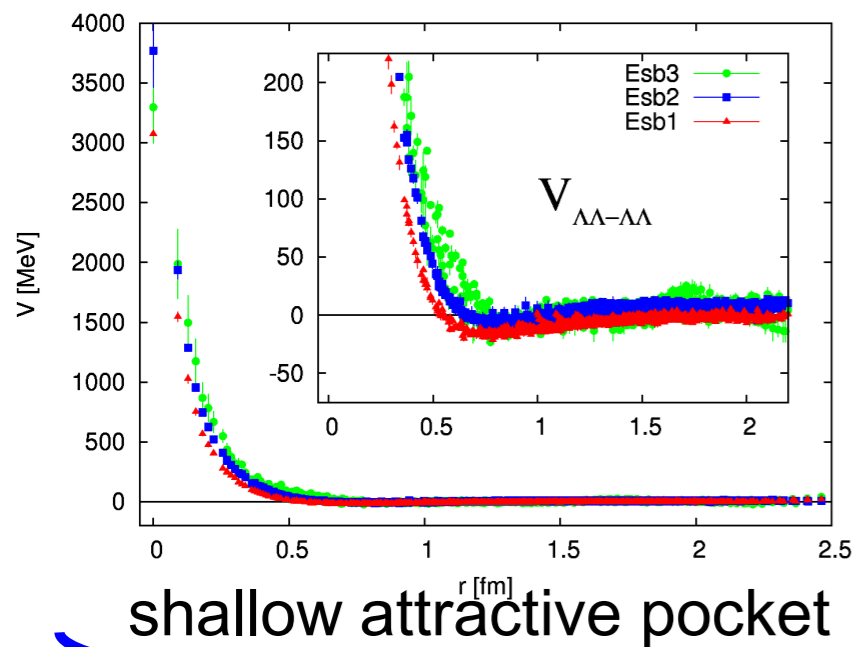
thresholds



SU(3) breaking effects becomes larger

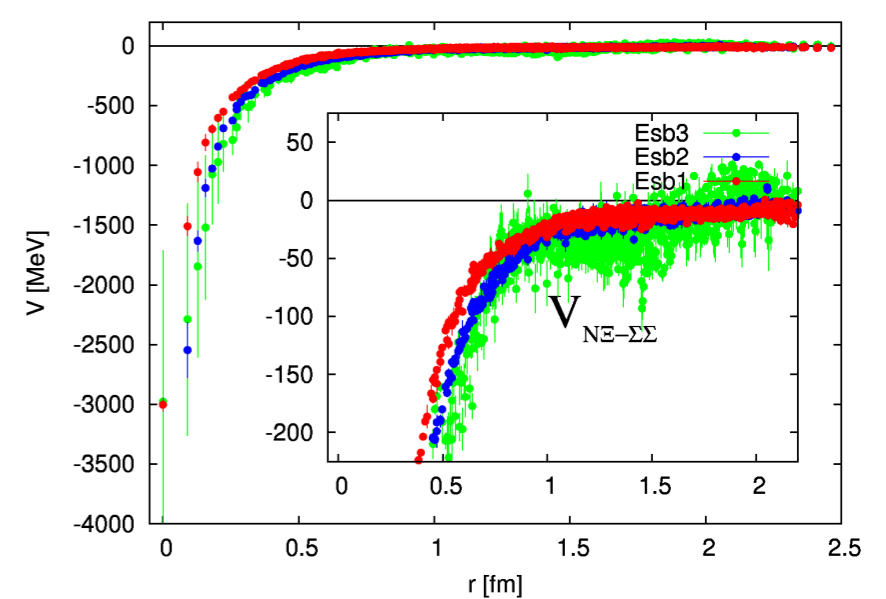
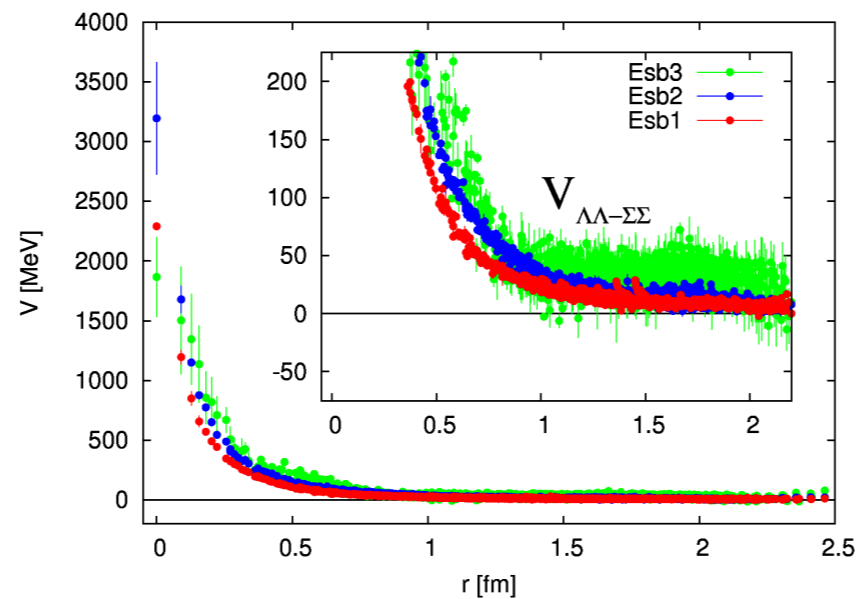
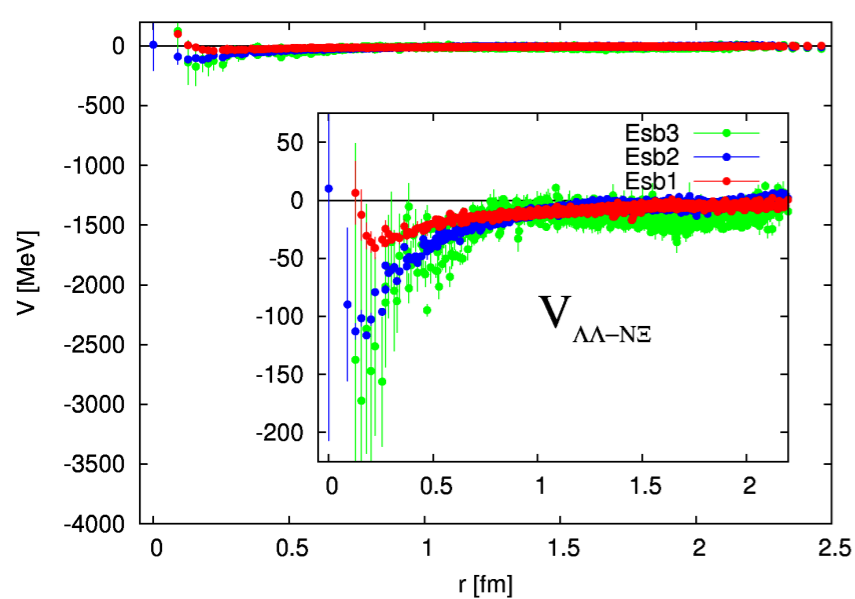
coupled channel 3x3 potentials

Diagonal elements



All channels have repulsive core

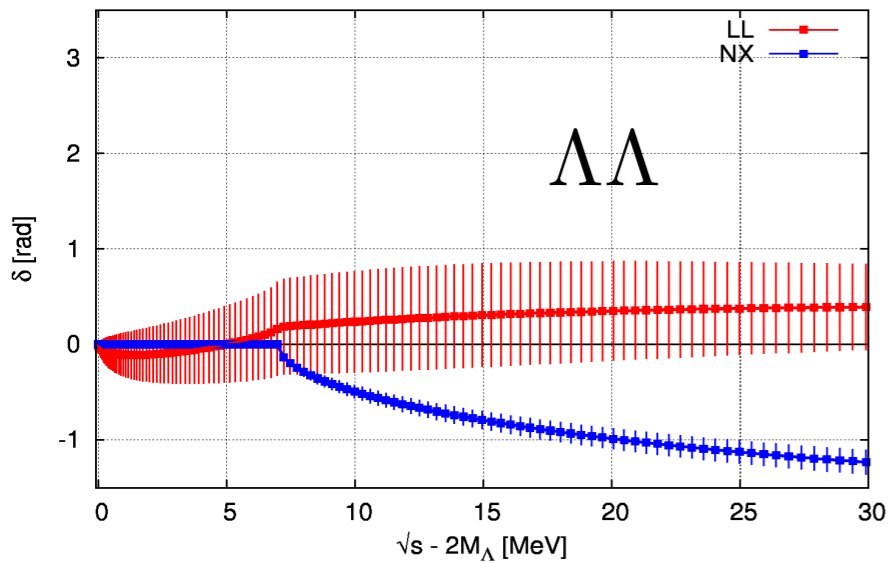
Off-diagonal elements



$\Lambda\Lambda$ and $N\Xi$ phase shift

Preliminary !

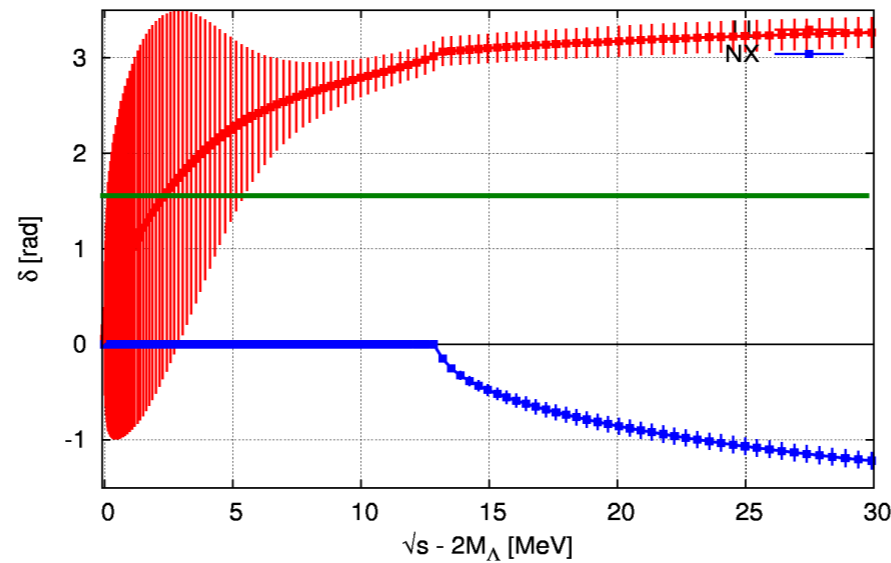
$m\pi = 700 \text{ MeV}$



$N\Xi$

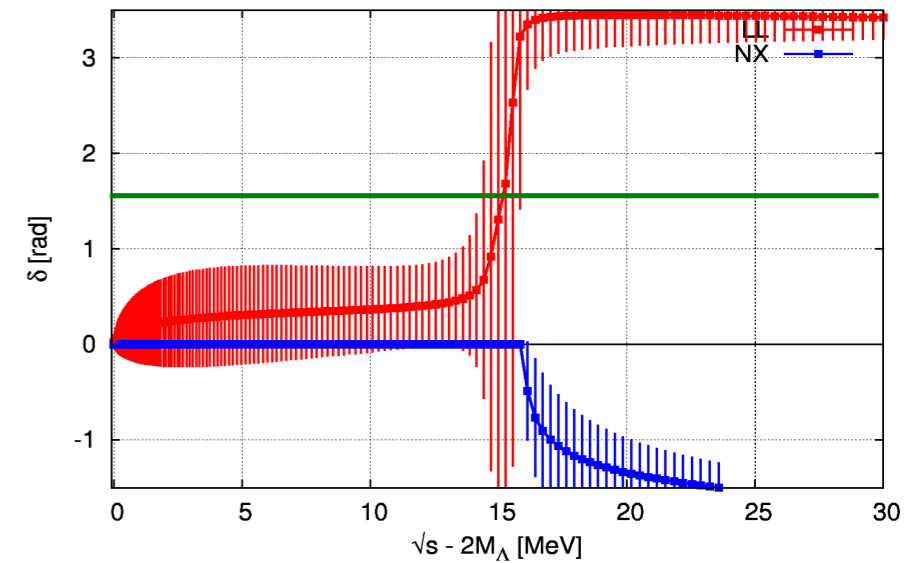
Bound H-dibaryon
coupled to $N\Xi$

$m\pi = 570 \text{ MeV}$



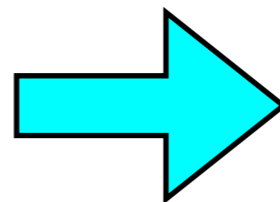
H as resonance
near $\Lambda\Lambda$ threshold
(H as bound $N\Xi$)

$m\pi = 410 \text{ MeV}$



H as resonance near
 $N\Xi$ threshold
(H as bound $N\Xi$)

This suggests that H-dibaryon becomes **resonance** at physical point.
Below or above $N\Xi$? Need simulation at physical point.

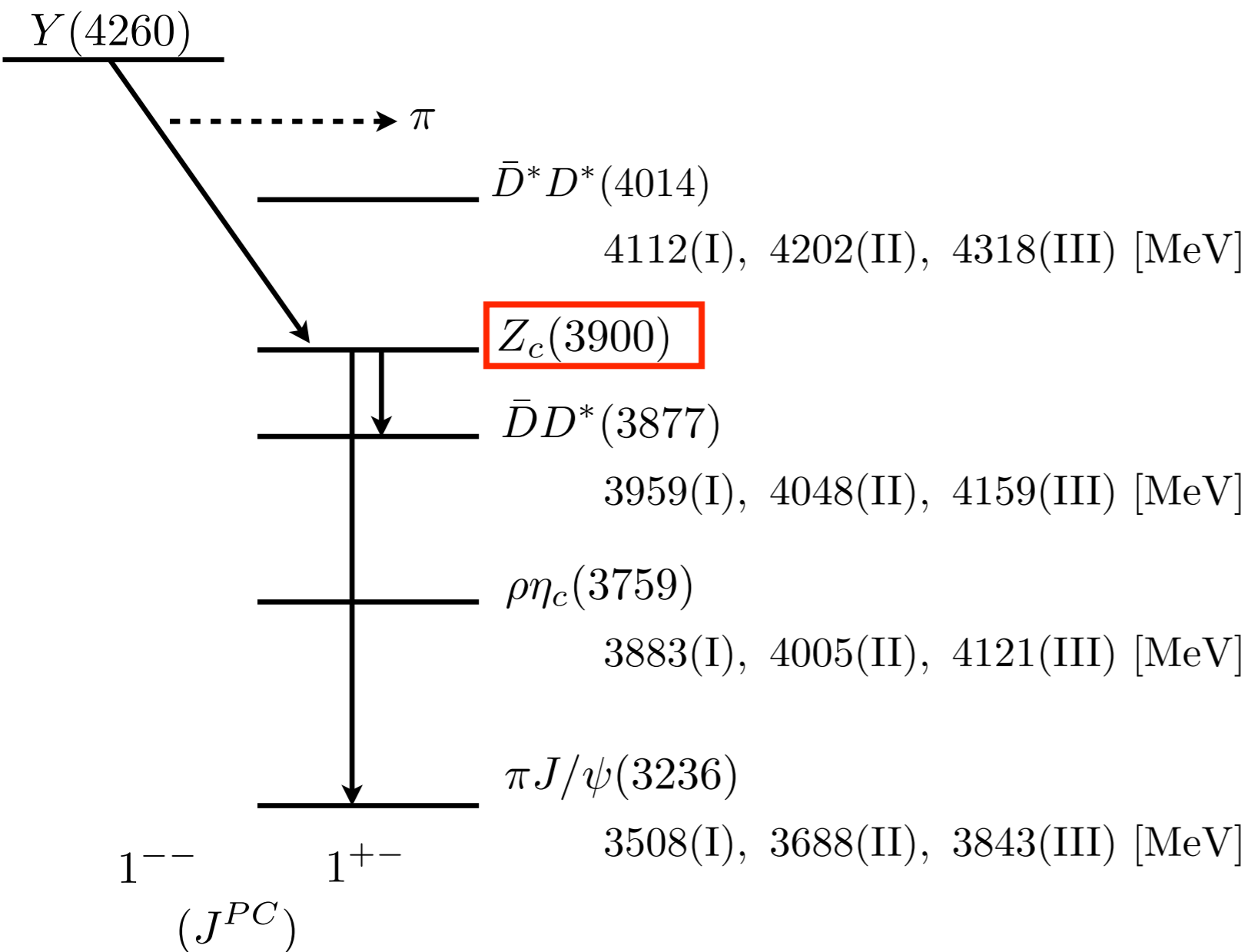


Doi's talk

$Z_c(3900)$

A tetraquark candidate

Y. Ikeda, *et al.*, arXiv:1602.03465[hep-lat]



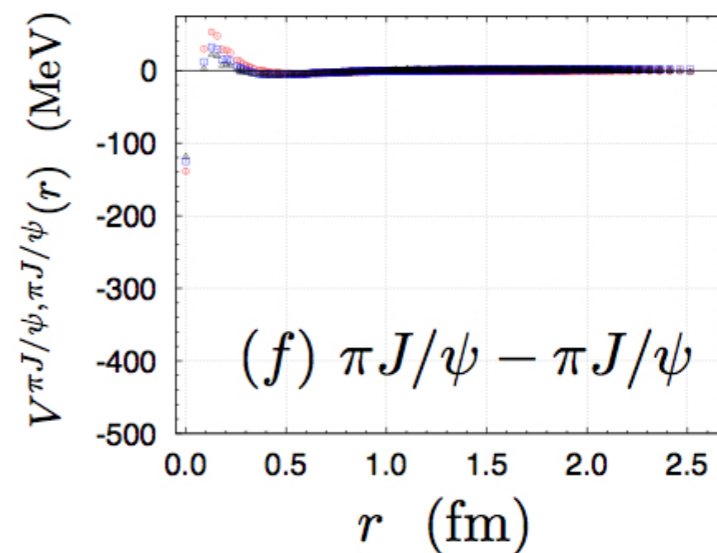
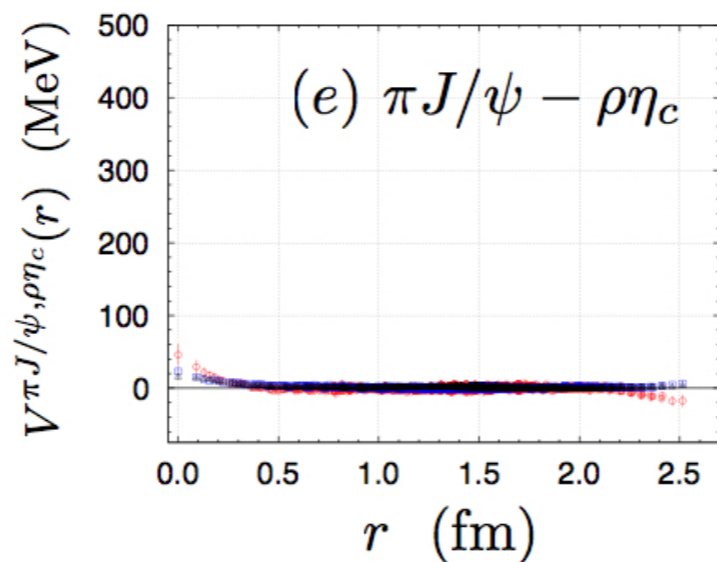
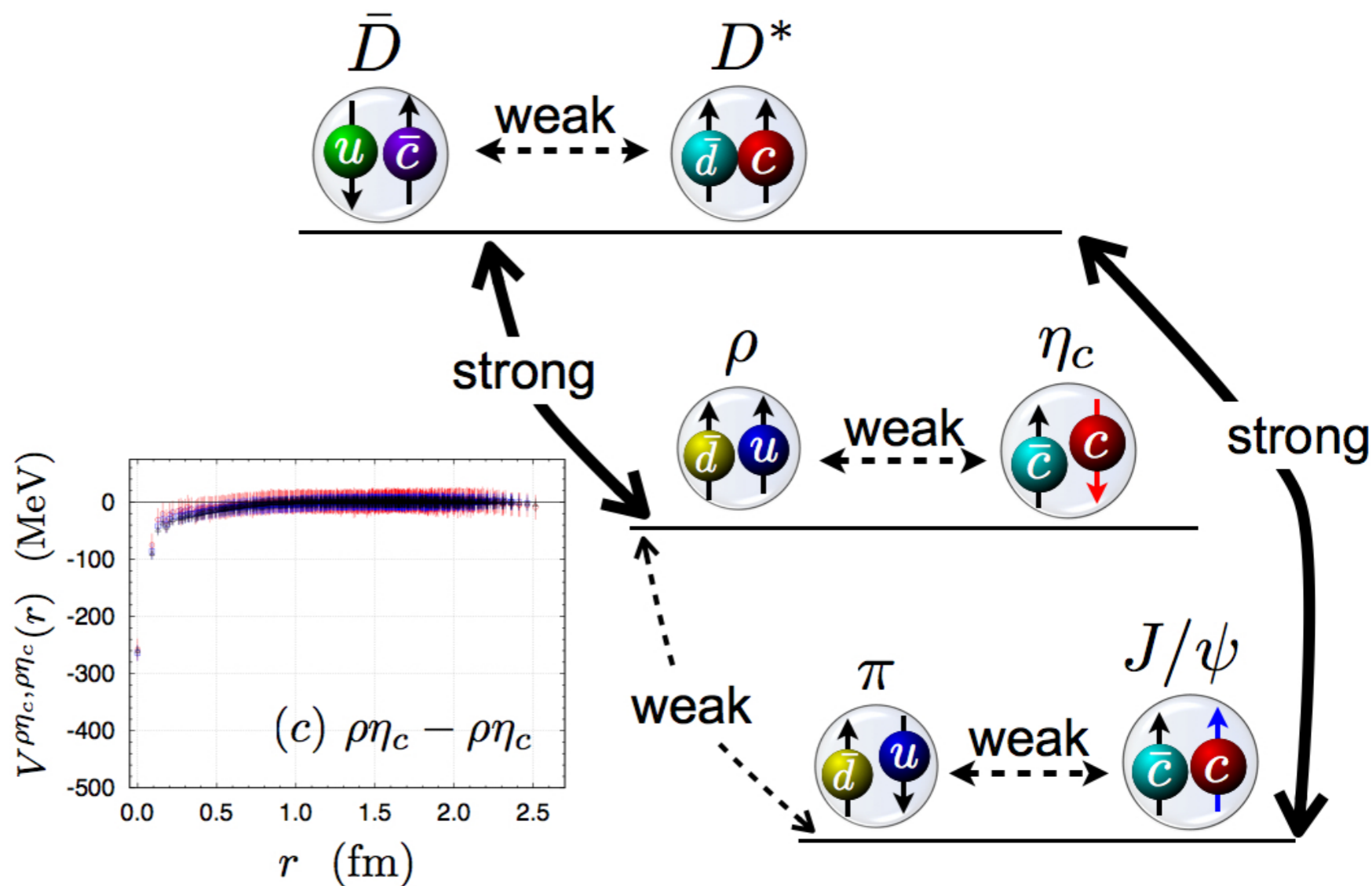
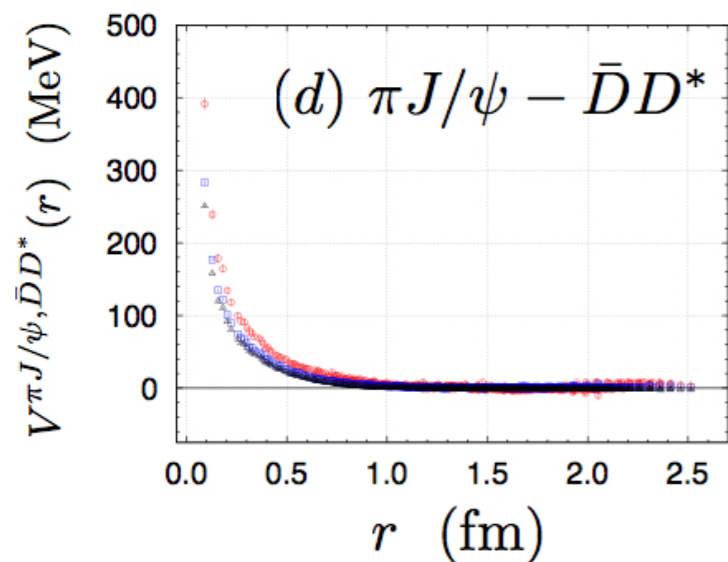
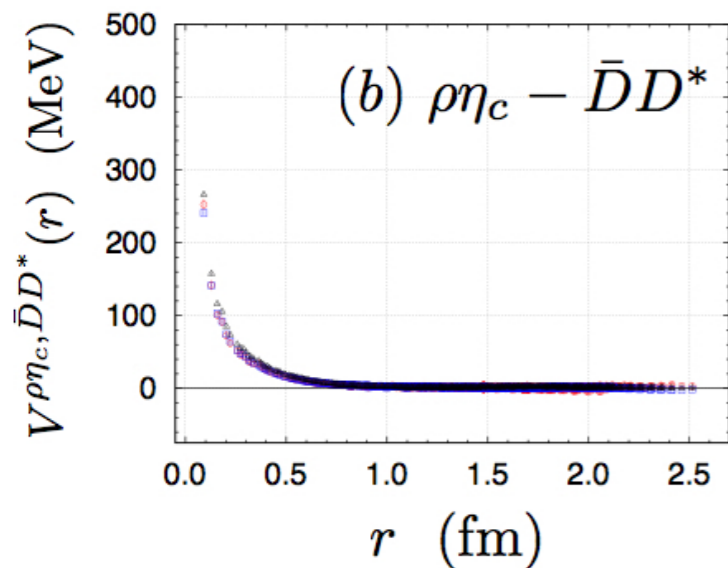
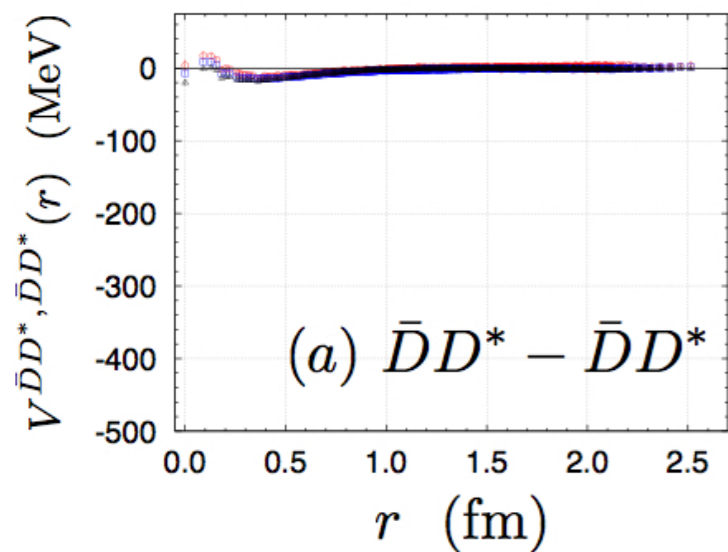
**coupled channel analysis
is needed.**

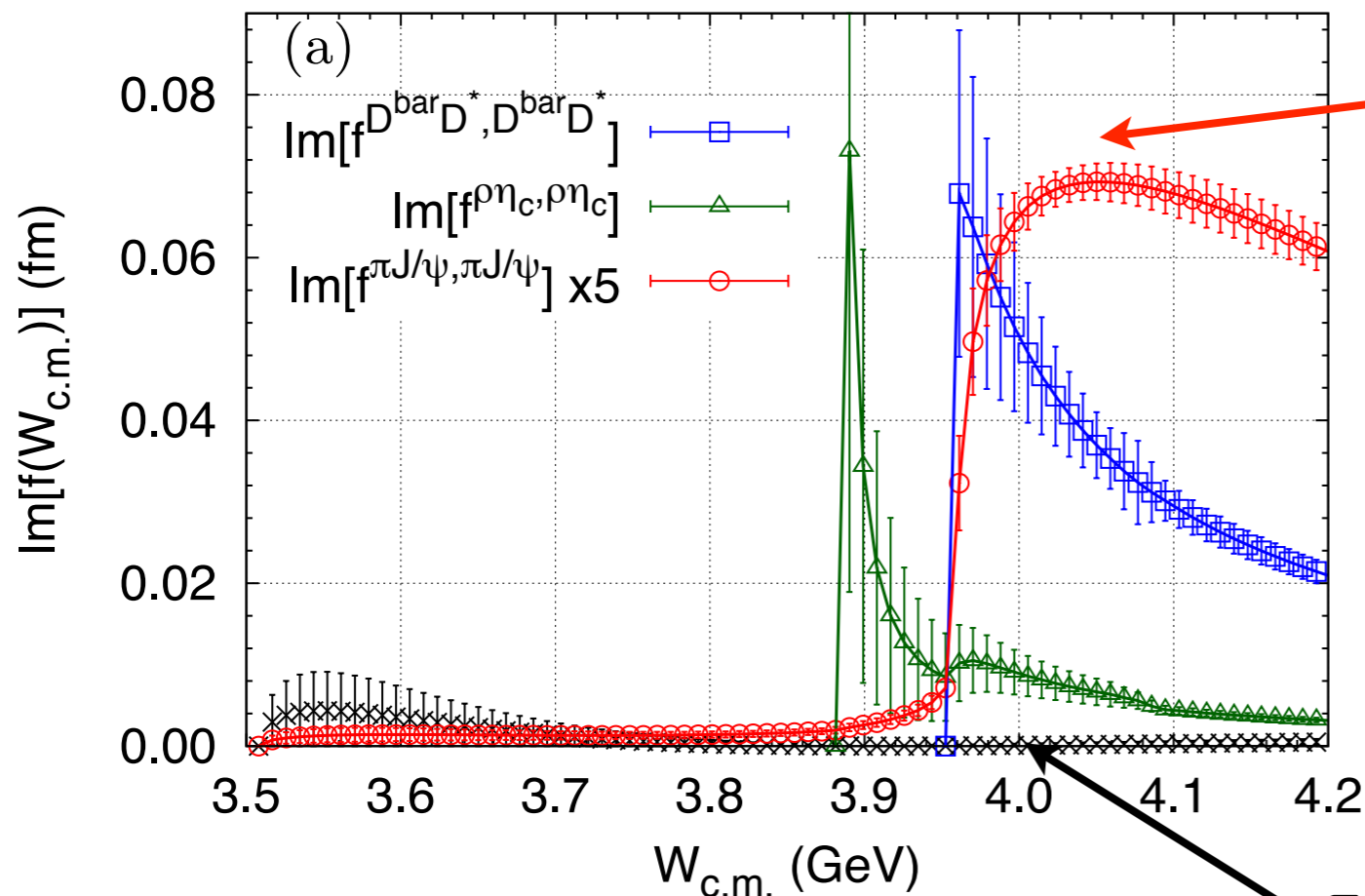
$m_\pi \simeq 410$ (Case I), 570 (Case II), 700 (Case III) MeV

$L \simeq 2.9$ fm, $a \simeq 0.09$ fm

coupled channel 3x3 potentials

Case I

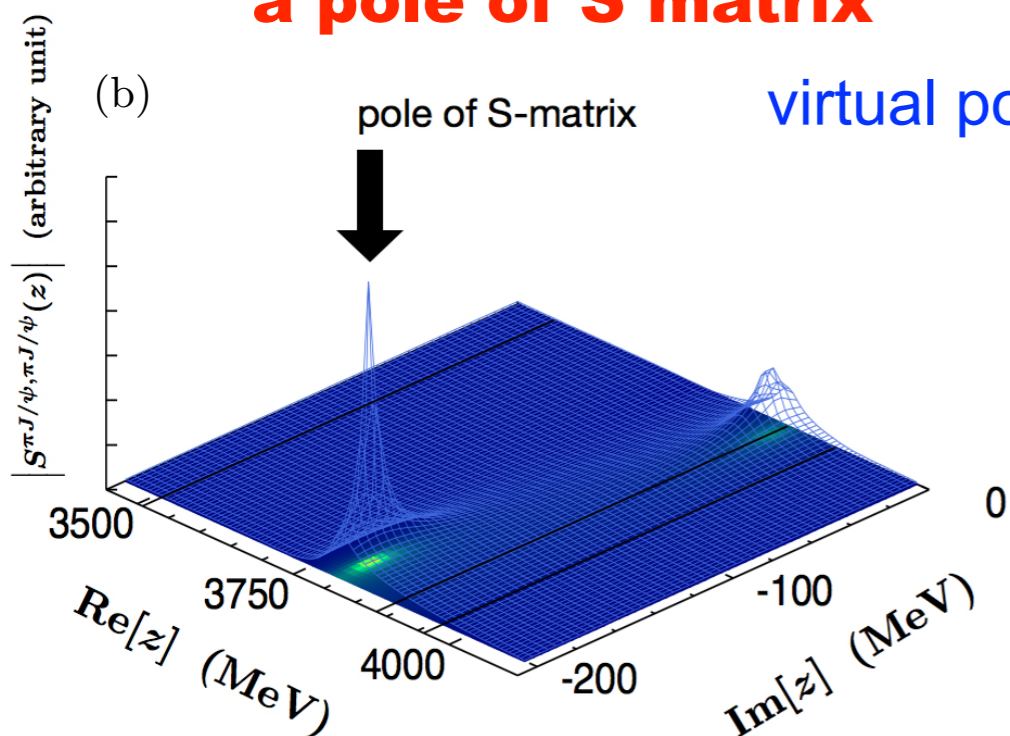




the peak structure in the $\pi J/\psi$ spectrum ($Z_c(3900)$) seems a "threshold cusp".

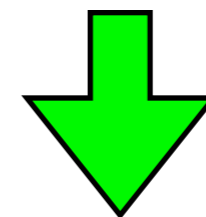
$\pi J/\psi$ mass spectrum w/o off-diagonal parts

a pole of S matrix



virtual pole in the second Riemann sheets of all channels

$$z_{\text{pole}} - (m_D + m_{D^*}) = -167(94) - 183(46)i$$



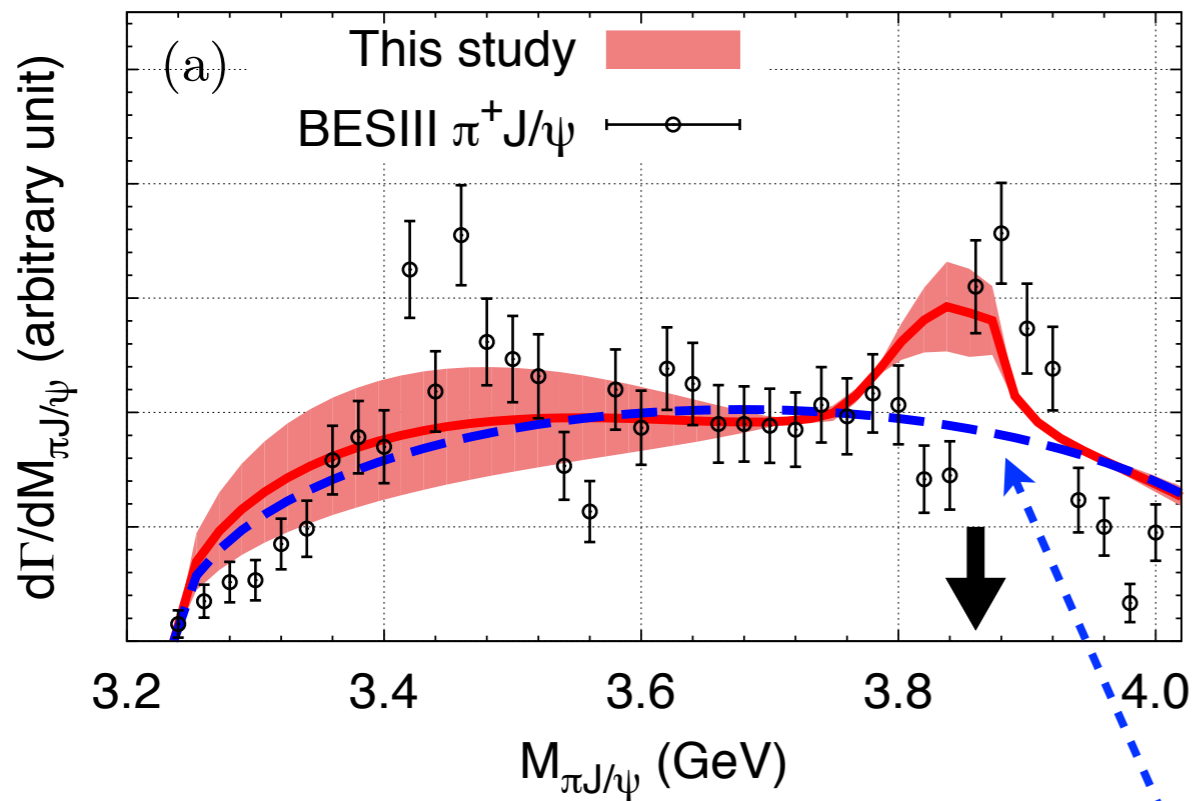
its contribution to the two-body amplitudes is highly suppressed.

model $Y(4260) \rightarrow \pi + \pi J/\psi, \pi + \bar{D}D^*$

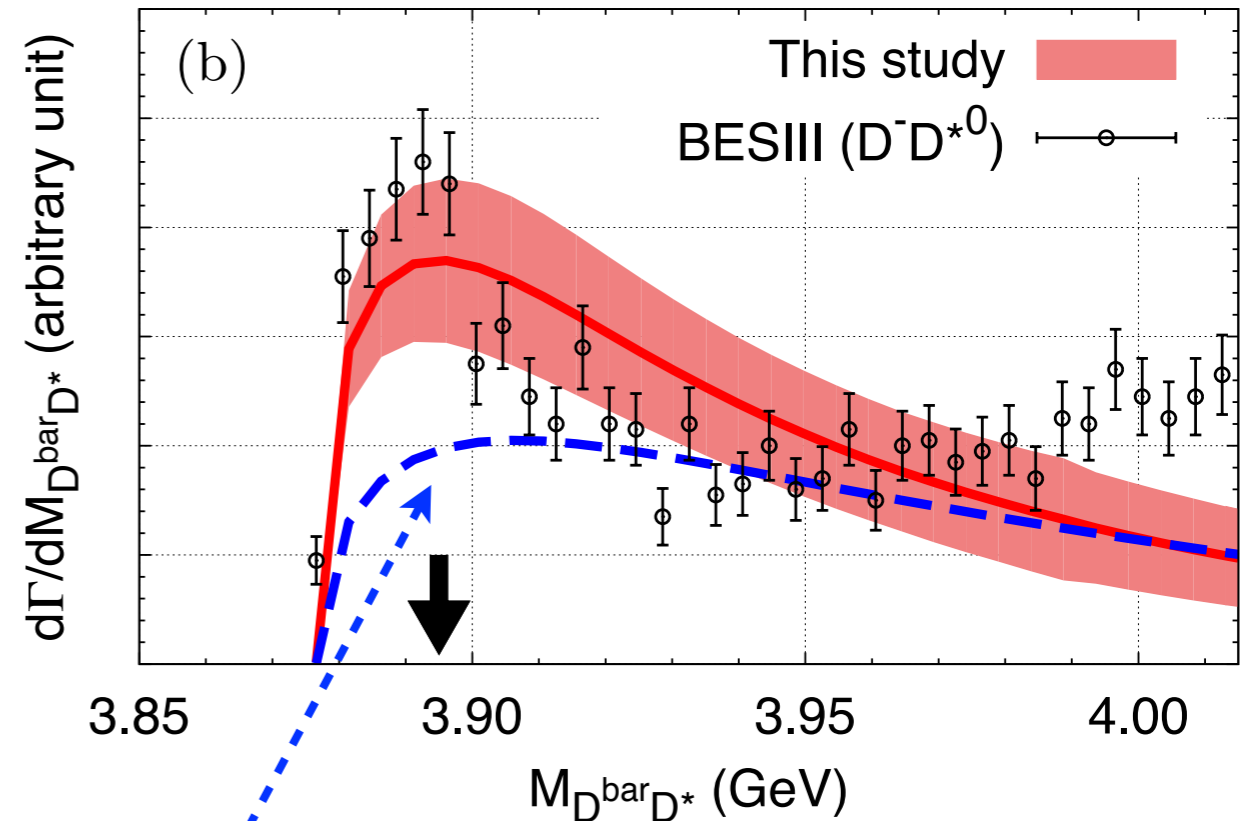
$$T^{Y \rightarrow \pi + \beta}(\vec{p}, \vec{q}_\beta; W_3) = \sum_{\alpha = \pi J/\psi, \bar{D}D^*} C^{Y \rightarrow \pi + \alpha} \leftarrow \text{fit parameters}$$

$$\times \left[\delta_{\alpha\beta} + \int d\vec{q}_\alpha \frac{T^{\alpha\beta}(\vec{q}_\alpha, \vec{q}_\beta, \vec{p}; W_3)}{W_3 - E_\pi(\vec{p}) - E_\alpha(\vec{p}, \vec{q}_\alpha) + i\epsilon} \right] \leftarrow \text{lattice data}$$

$Y(4260) \rightarrow \pi + \pi J/\psi$



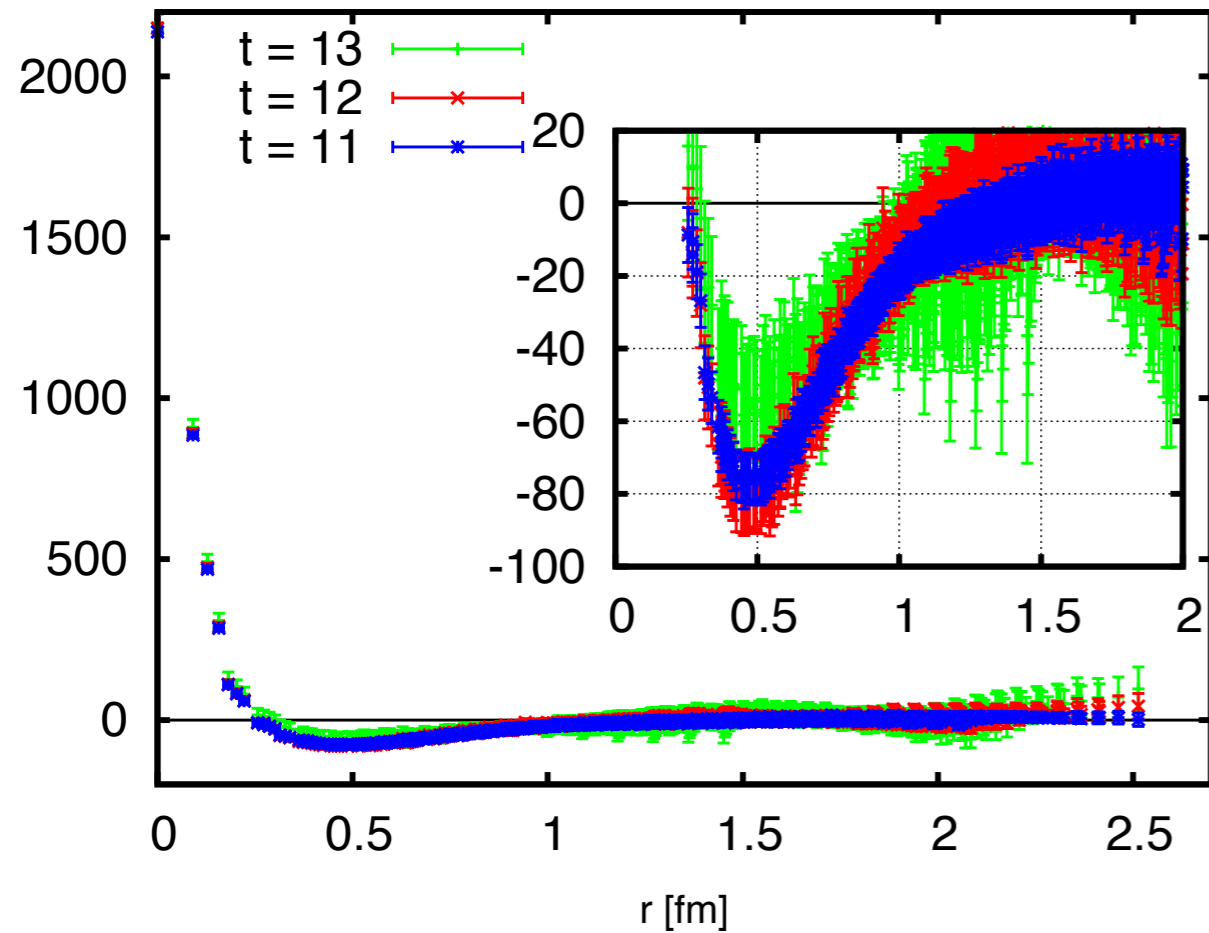
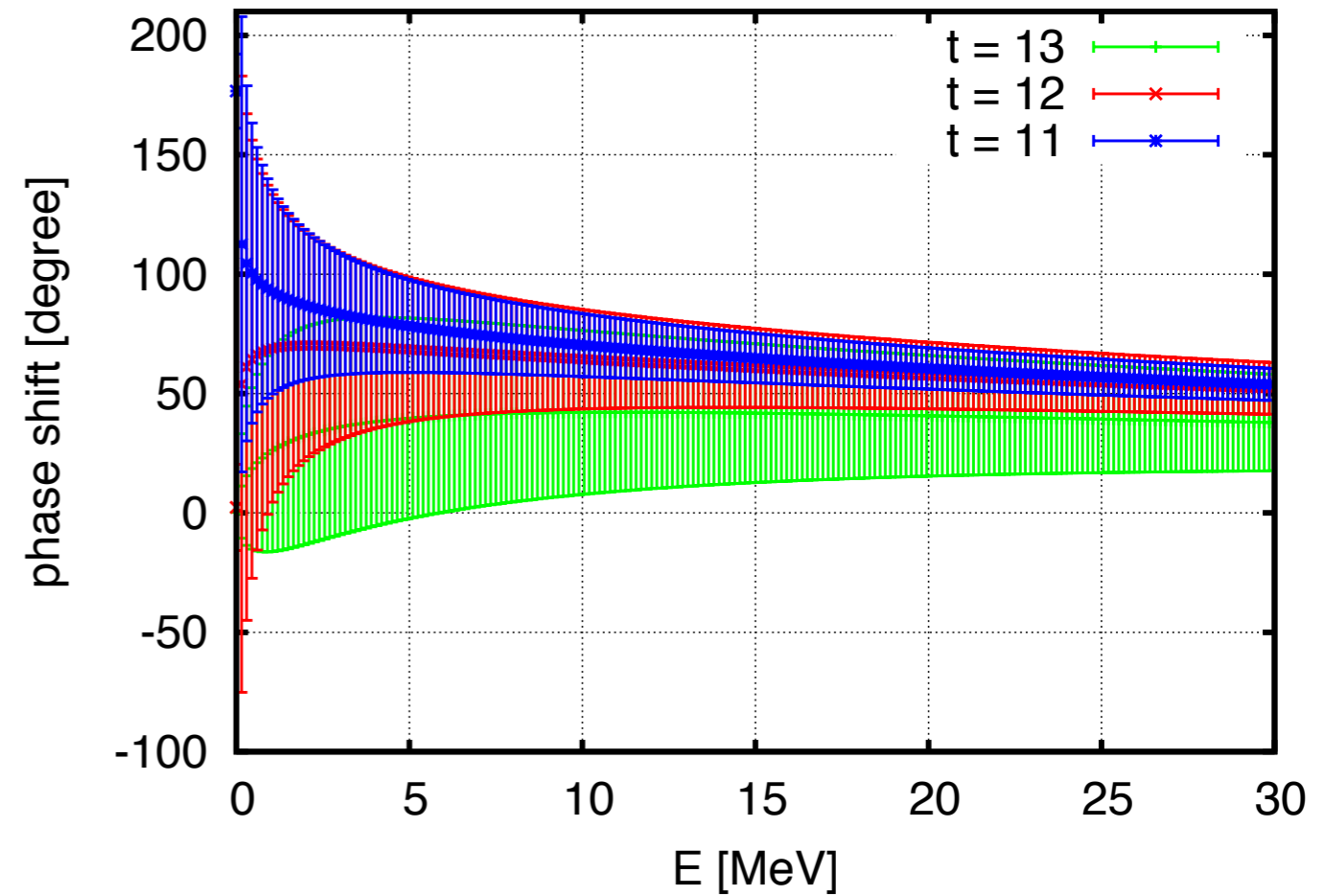
$Y(4260) \rightarrow \pi + \bar{D}D^*$



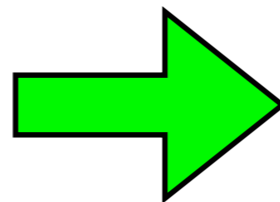
off-diagonals switched-off

4. Summary

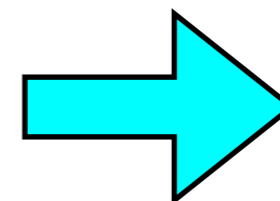
- HAL QCD strategy is a very powerful one for “nuclear physics from lattice QCD”.
- work well also for multi-channel scatterings
- More results

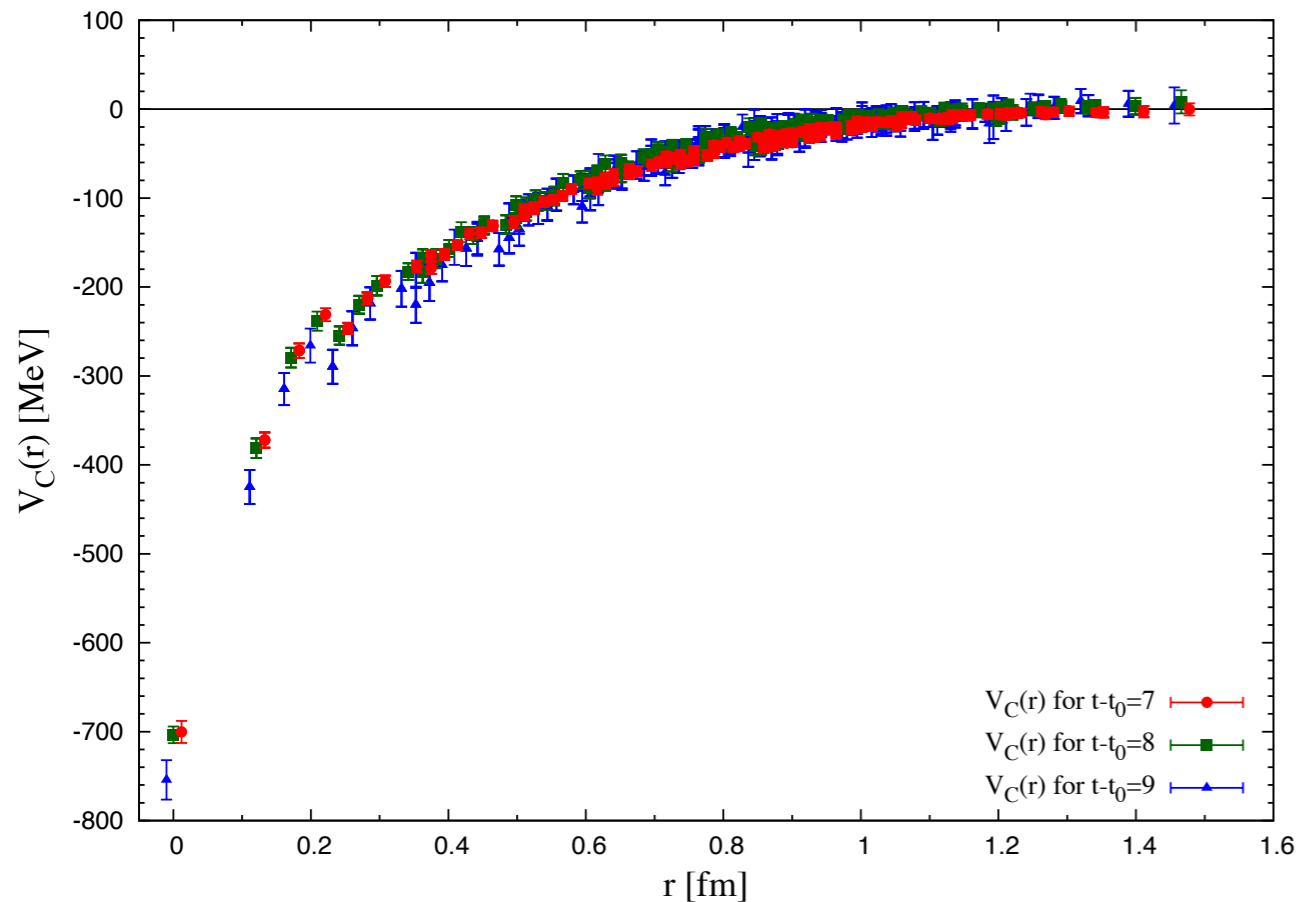
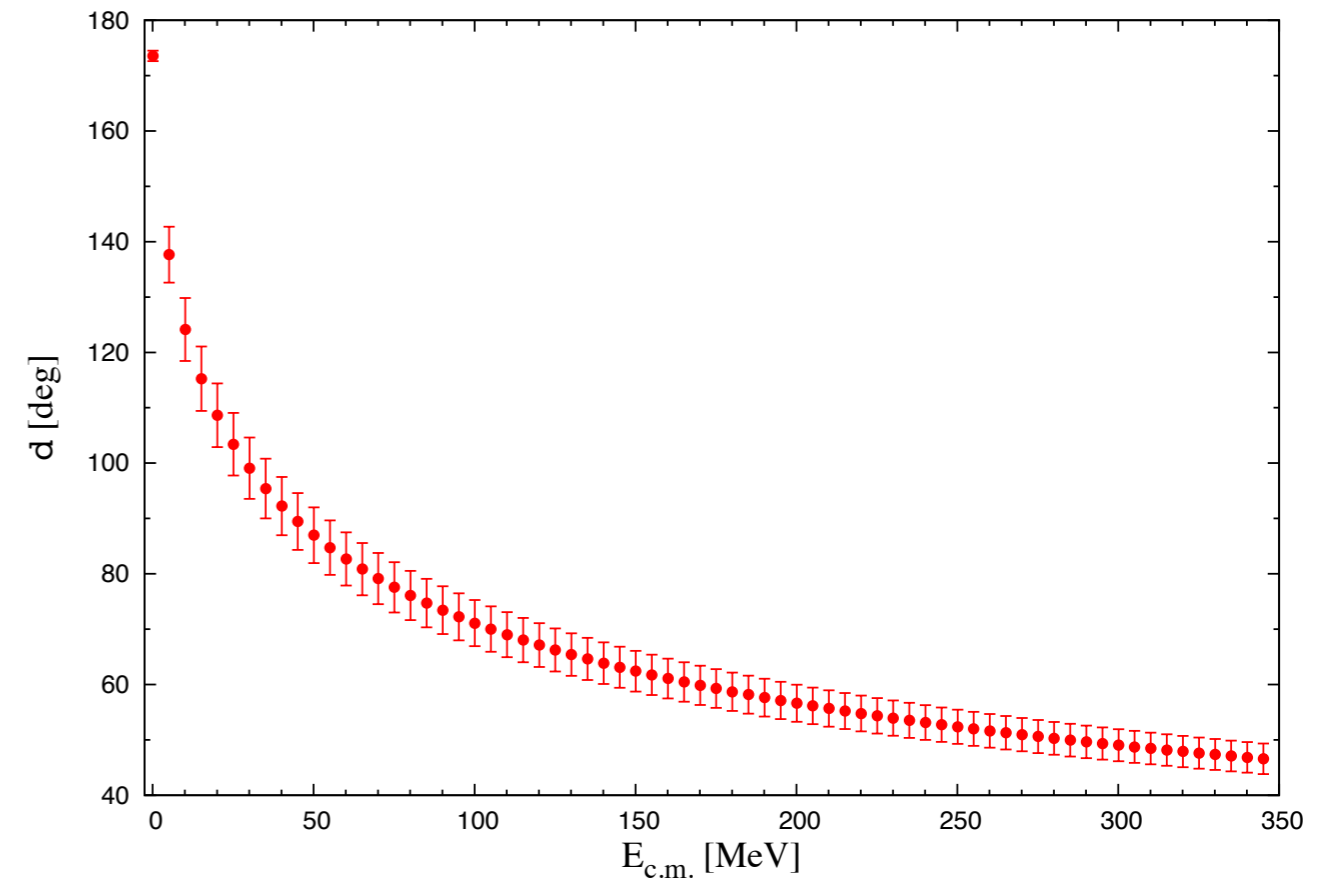
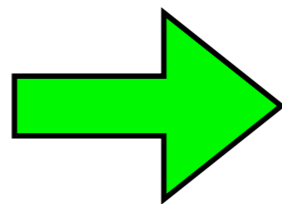
$\Omega\Omega$ M. Yamada, *et al.*, PTEP2015(2015)071B01 $m_\pi \simeq 700$ MeV, $L \simeq 2.9$ fm, $a \simeq 0.09$ fm**potential****phase shift**

NN-like structure, strong attraction

**about to be bound**
“Unitary region”

physical point ?

**Doi's talk**

$N\Omega$ F. Etminan, *et al.*, NPA928(2014)89 $m_\pi \simeq 875$ MeV, $L \simeq 1.9$ fm, $a \simeq 0.12$ fm**potential****phase shift**attractive potential
without repulsive core

a bound state

$$\begin{aligned}
 B_{N\Omega} &= 18.9(5.0)_{(-1.8)}^{(+12.1)} \text{ MeV}, \\
 a_{N\Omega} &= -1.28(0.13)_{(-0.15)}^{(+0.14)} \text{ fm}, \\
 (r_e)_{N\Omega} &= 0.499(0.026)_{(-0.048)}^{(+0.029)} \text{ fm}.
 \end{aligned}$$

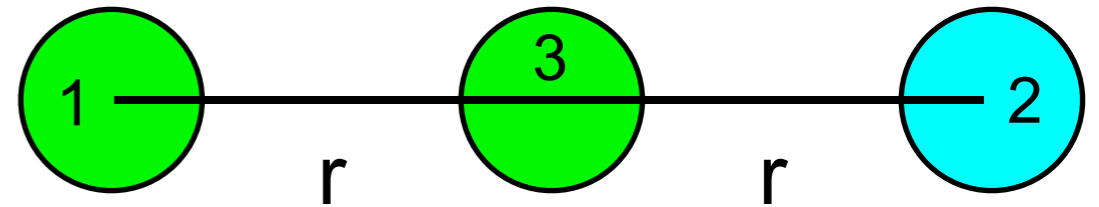
Back-up:
Some applications to
nuclear physics

Three nuclear force (3NF)

$$V_{3N}(\mathbf{x}_1 - \mathbf{x}_3, \mathbf{x}_2 - \mathbf{x}_3) = \sum_{i < j} V_{2N}(\mathbf{x}_i - \mathbf{x}_j) + V_{3NF}(\mathbf{x}_1 - \mathbf{x}_3, \mathbf{x}_2 - \mathbf{x}_3)$$

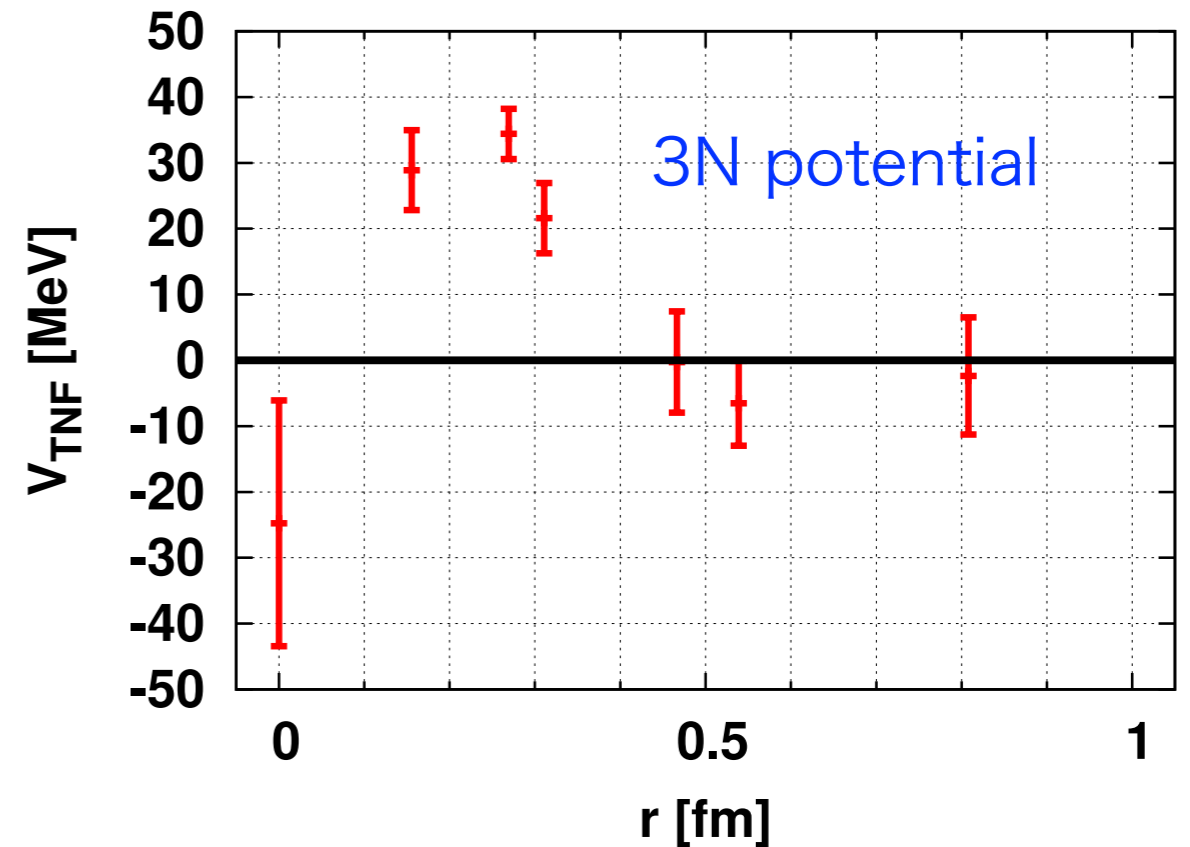
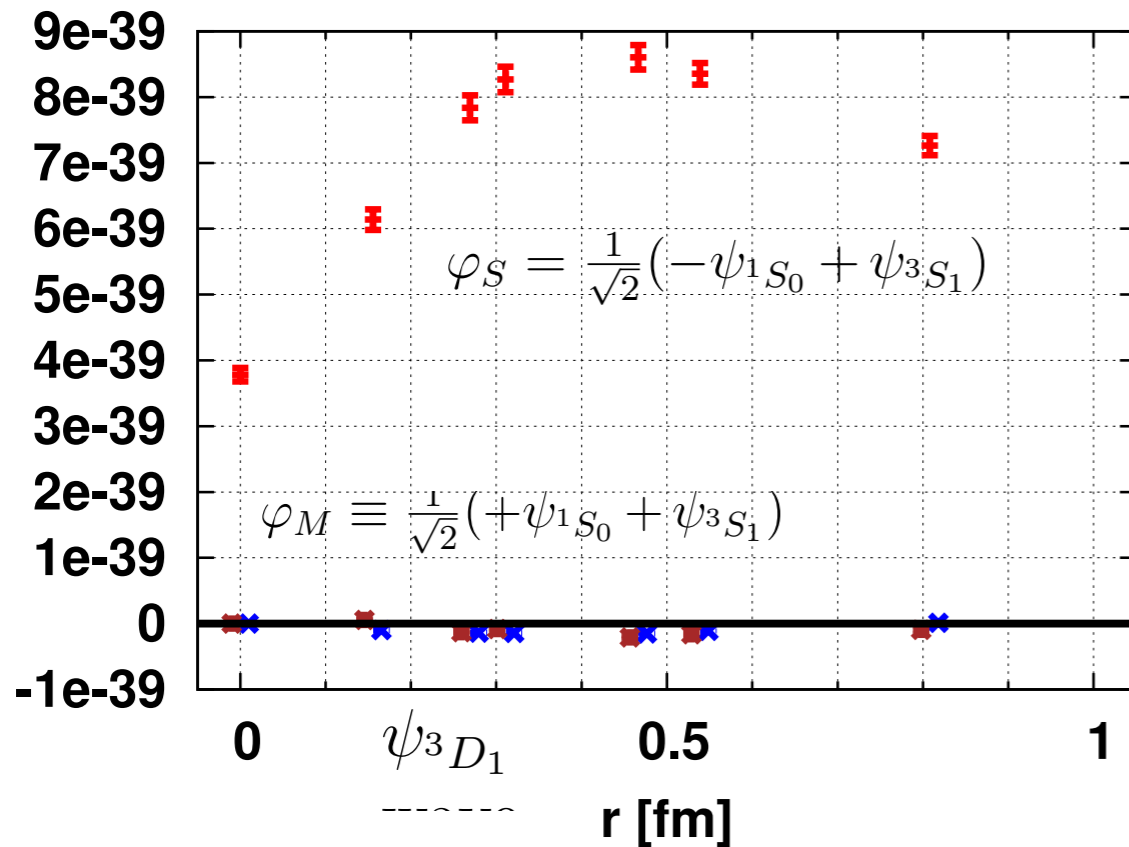
2-body potential
3-body potential

(1,2) pair $^1S_0, ^3S_1, ^3D_1$



Linear setup

wave functions



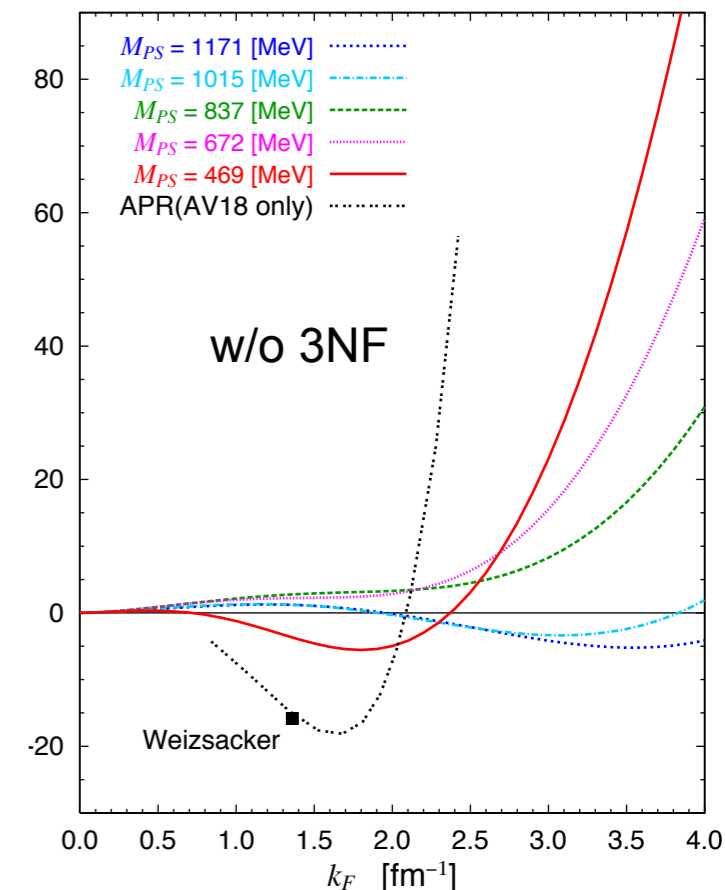
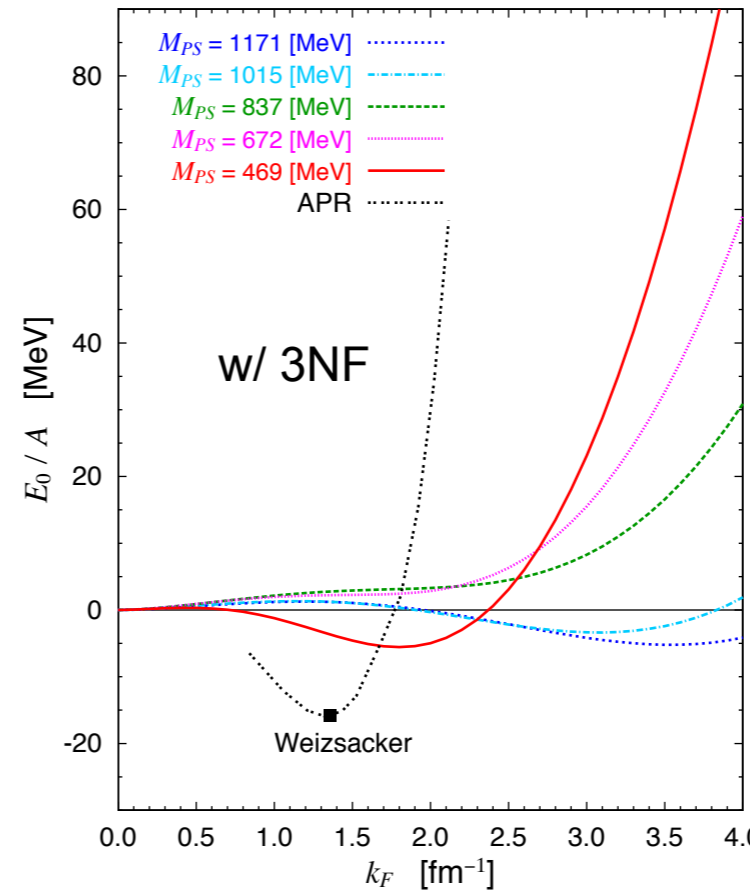
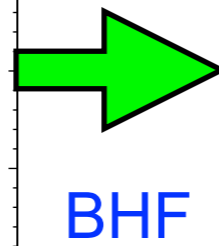
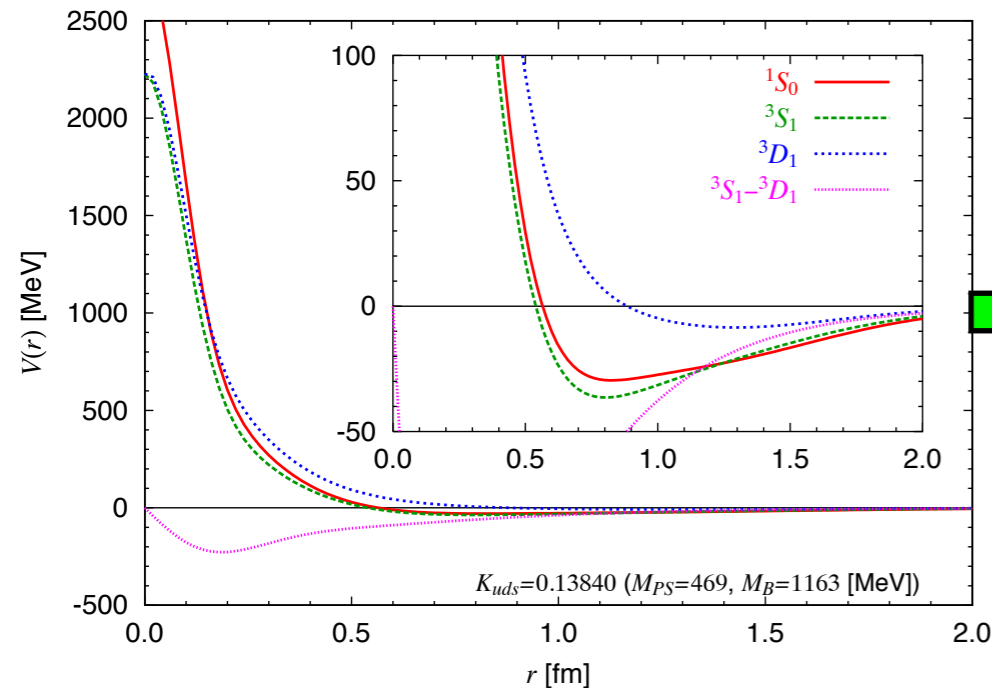
scalar/isoscalar 3NF is seen at short distance.

Equation of State of nuclear matter

Inoue et al. (HAL QCD Coll.), PRL111(2013)112503

NN potentials $m_\pi = 470$ MeV

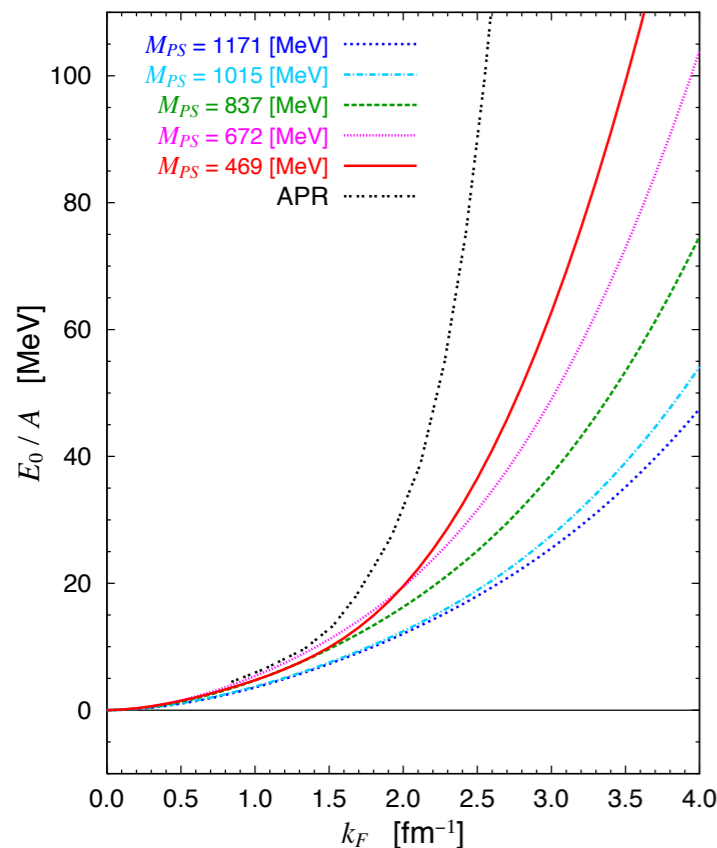
Energy density of Nuclear matter



Fermi momentum

A. Akmal, V.R. Pandharipande, G.G. Ravenhall, Phys. Rev. C58 1804 (1998)

Neutron matter

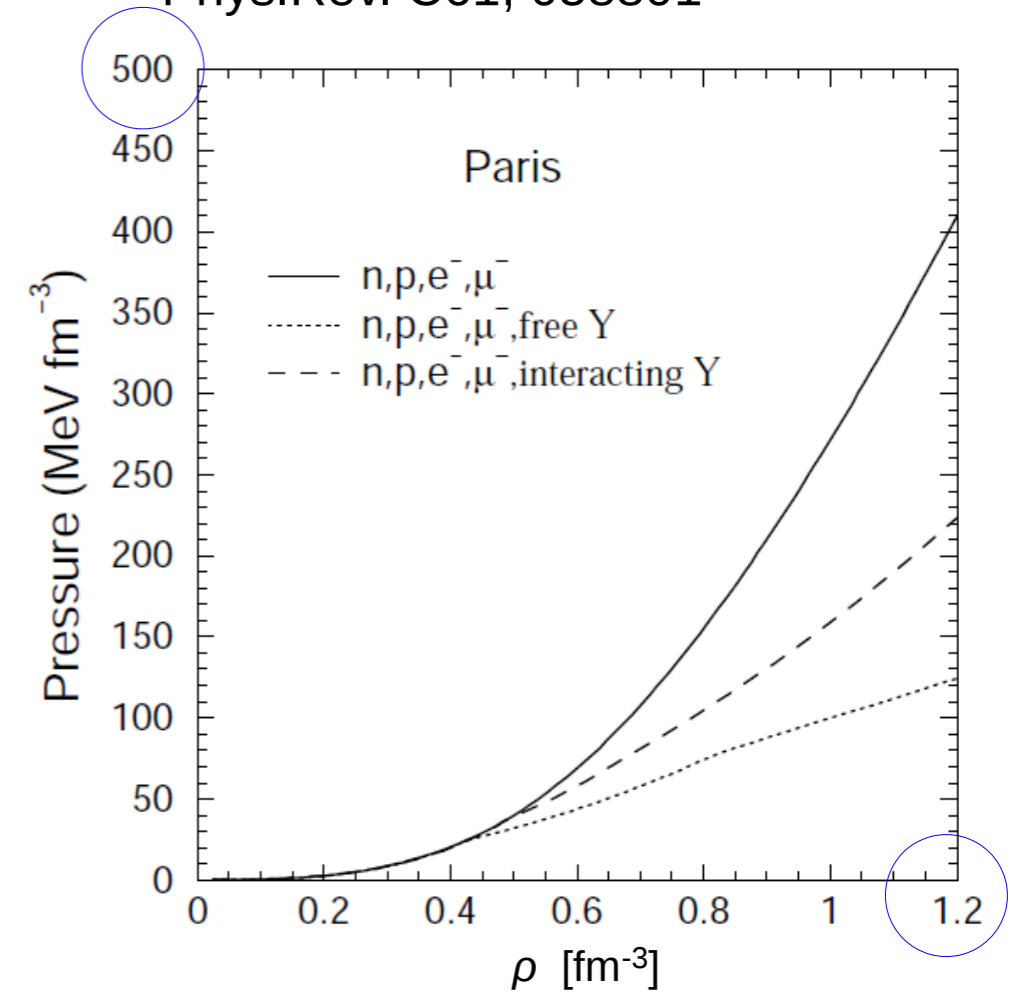
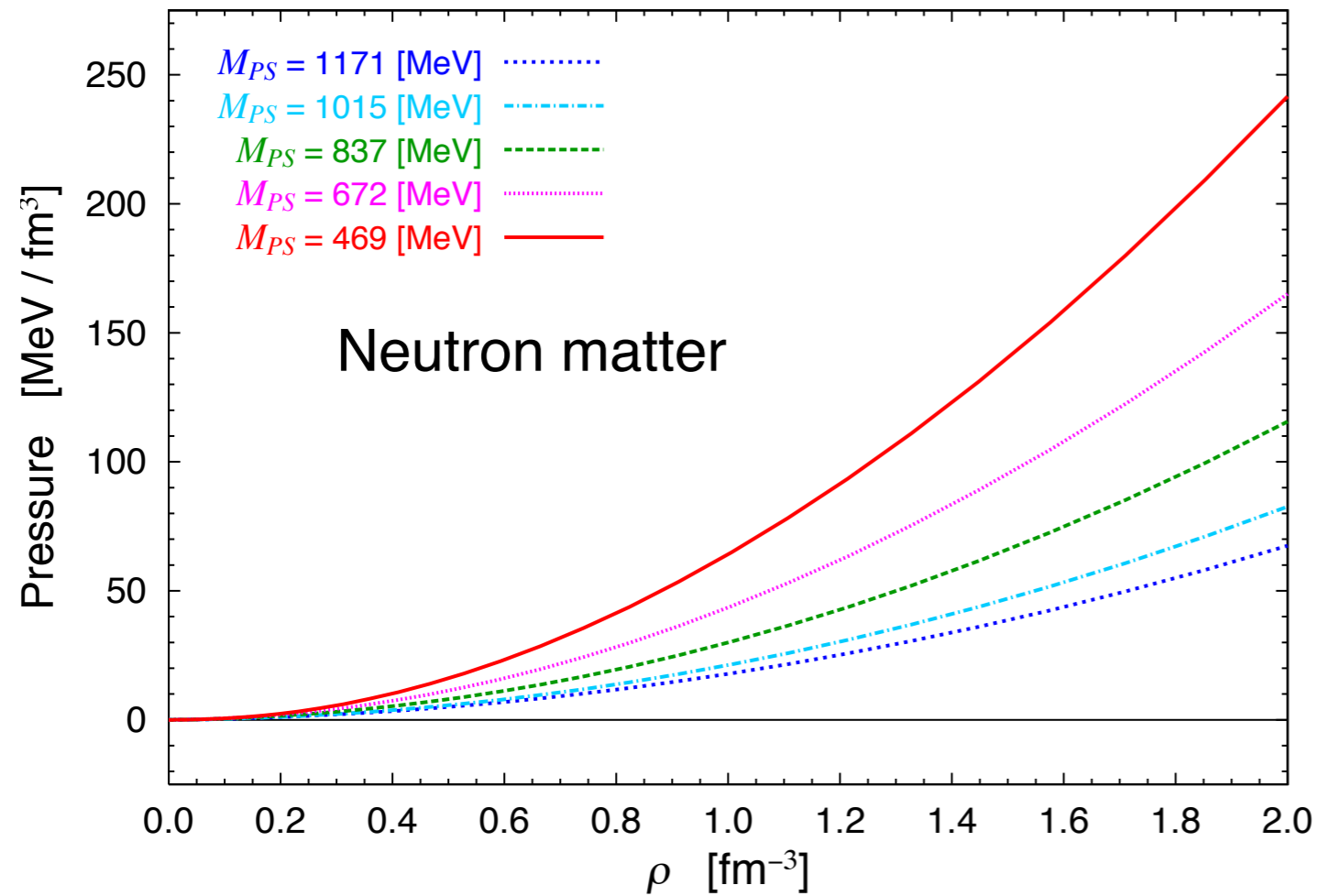


Nuclear matter shows the saturation at the lightest pion mass, but the saturation point deviates from the empirical one obtained by Weizsacker mass formula.

No saturation for Neutron matter.

Pressure of Neutron matter

M. Baldo, F. Burgio, H.-J.Schulze,
Phys.Rev. C61, 058801

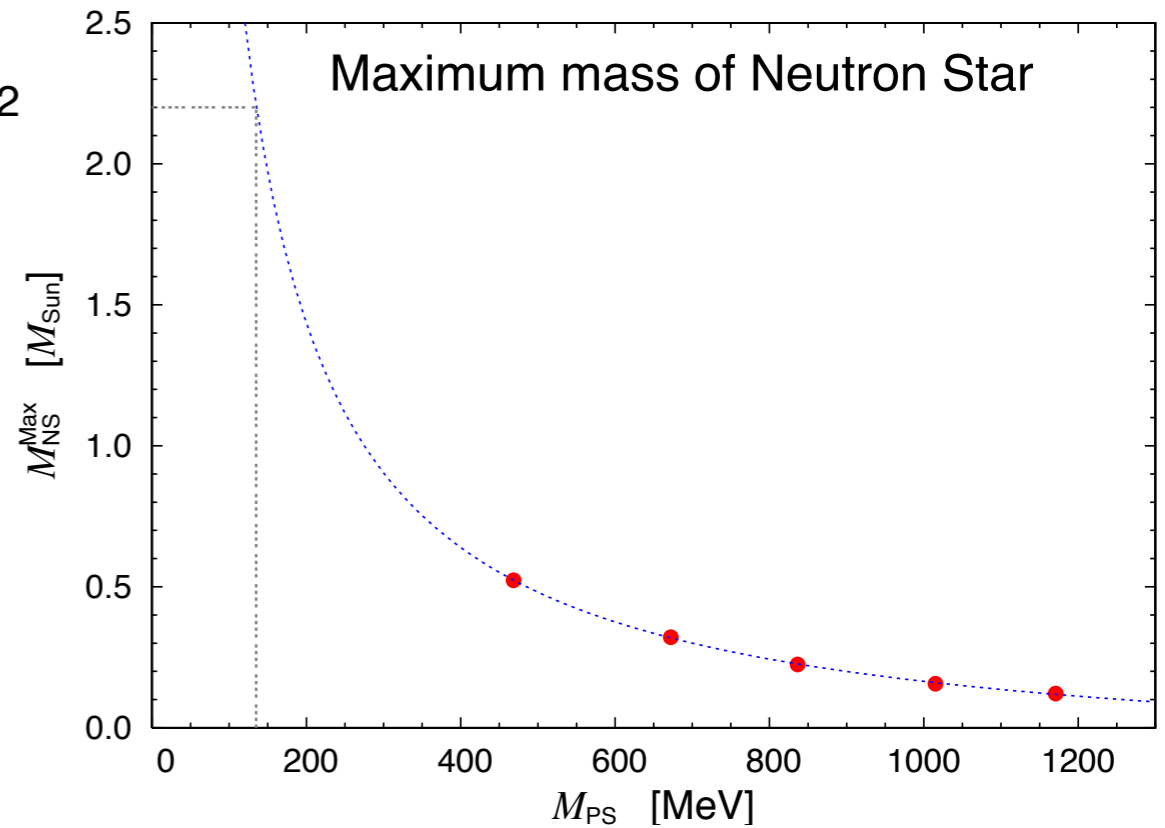
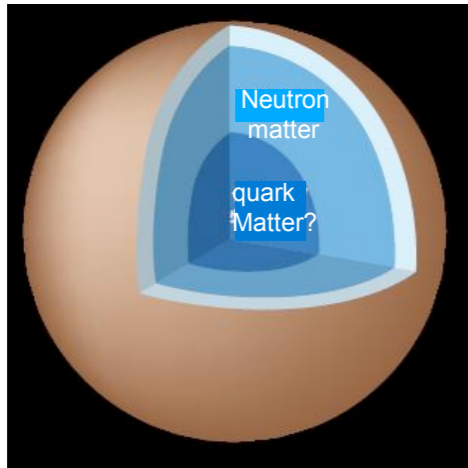
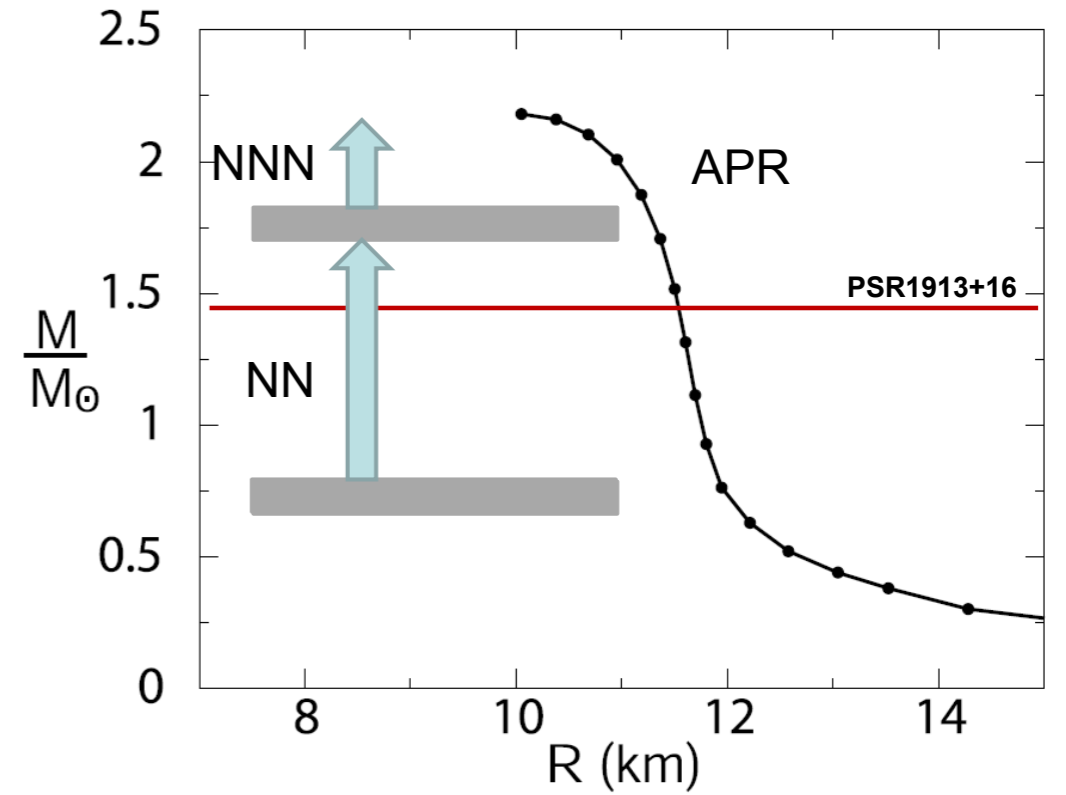
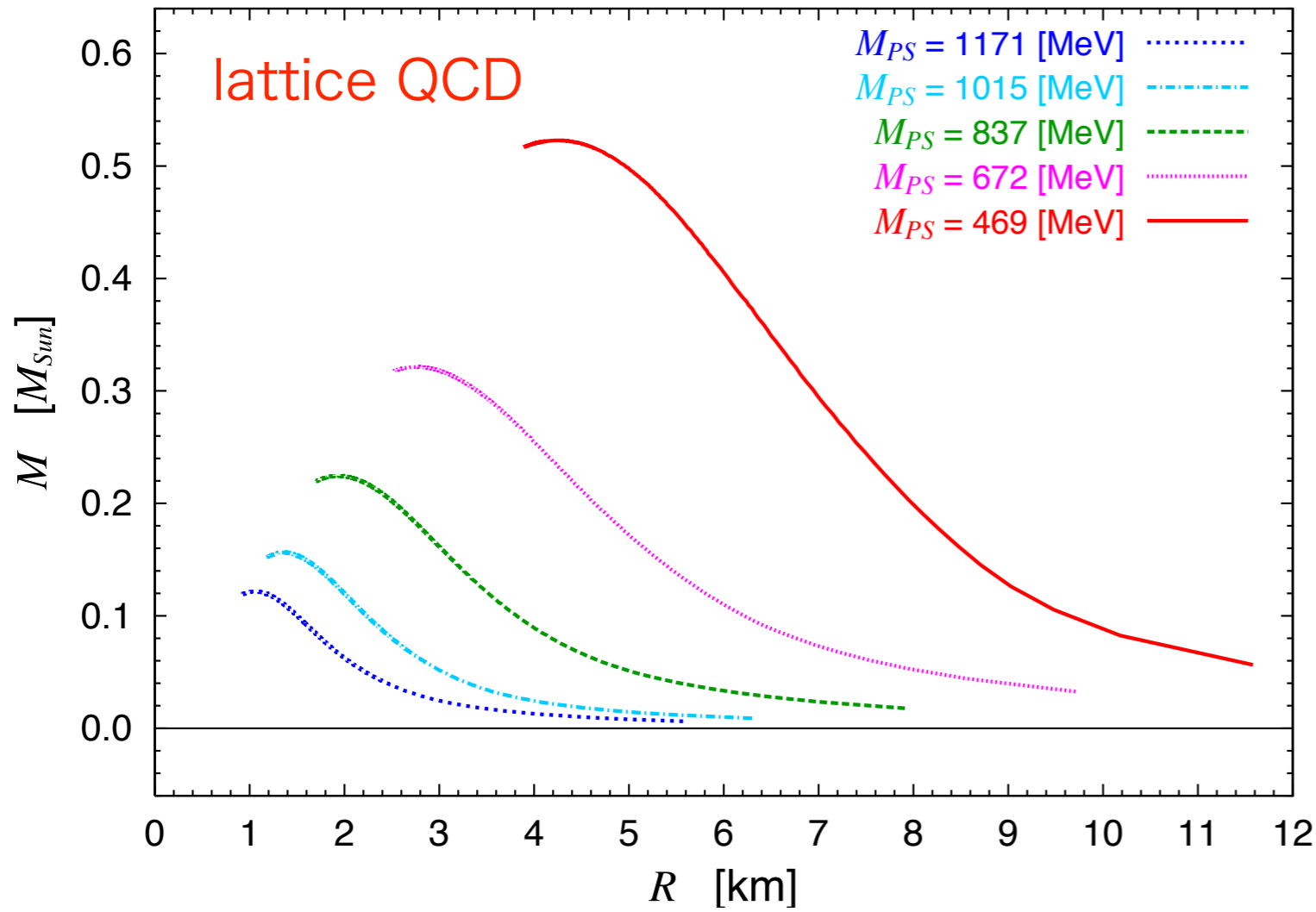


pressure $P = \rho^2 \frac{d(E_0/A)}{d\rho} = \frac{\gamma k_F^4}{18\pi^2} \frac{d(E_0/A)}{dk_F}$

density $\rho = \frac{\gamma k_F^3}{6\pi^2}$

Our Neutron matter becomes harder as the pion mass decreases, but it is still softer than phenomenological models.

Neutron star M-R relation



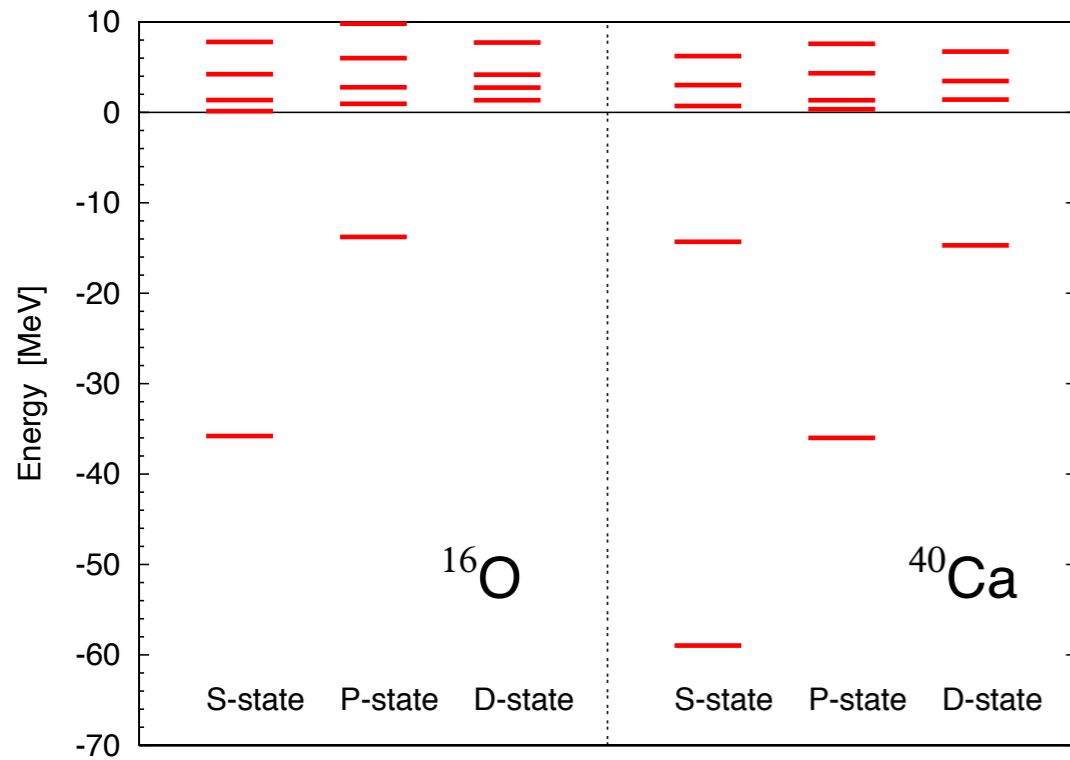
Maximum mass of Neutron vs. pion mass

Medium-Heavy Nuclei

Inoue et al. (HAL QCD Coll.), arXiv:1408.4829[hep-lat]

$$m_\pi = 470 \text{ MeV}$$

single particle level



	Single particle level				Total energy		Radius
	1S	1P	2S	1D	E_0	E_0/A	$\sqrt{\langle r^2 \rangle}$
^{16}O	-35.8	-13.8			-34.7	-2.17	2.35
^{40}Ca	-59.0	-36.0	-14.7	-14.3	-112.7	-2.82	2.78

$^{16}\text{O}(\text{exp})$	-127 MeV	2.73 fm
$^{40}\text{Ca}(\text{exp})$	-342 MeV	3.48 fm

E_0/A vs $A^{-1/3}$ from lattice QCD

Nucleon number distribution

