## Viscous Effects on the Mapping of the Initial to Final State in Heavy Ion Collisions F.Gardim, J.Noronha-Hostler, M.Luzum, FG, PRC91 (2015) 034902, F.Gardim, FG, M.Luzum, J.-Y.Ollitrault, PRC85 (2012) 024908.

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#### **Outline**



- [Setting up the estimator](#page-5-0)
- [Computing the flow coefficients with hydro](#page-8-0)
- [Results for the mapping](#page-13-0)

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#### Examples of Initial Conditions Energy Density profile



Schenke, Tribedy, Venugopalan,Phys.Rev.Lett. 108 (2012) 252301

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Drescher,Nara, Phys. Rev. C 75, 034905 (2007); ,76, 041903 (2007).



Drescher,Hladik,Ostapchenko,Pierog, Werner, Phys.Rept. [35](#page-3-0)[0,](#page-1-0) [93 \(](#page-2-0)[2](#page-3-0)[00](#page-1-0)[1\)](#page-2-0)<br> $\leftarrow$   $\Box$   $\rightarrow$   $\leftarrow$   $\overline{\Box}$   $\rightarrow$   $\leftarrow$   $\overline{\Xi}$   $\rightarrow$   $\leftarrow$   $\overline{\Xi}$  $2Q$ 

## Hydrodynamics prediction for flow harmonics

#### NeXSPheRIO (ideal hydro) fits the flow harmonics well



<span id="page-3-0"></span>So do models with shear viscosity and shear+bulk viscosities.

## **Objective**

#### Study the mapping between initial conditions and flow harmonics.

#### Many works on the subject:

- H. Petersen, G. Y. Qin, S. A. Bass and B. Muller, Phys.Rev. C82 , 041901 (2010)
- B. Schenke, S. Jeon and C. Gale, Phys. Rev. Lett.106, 042301 (2011)
- Z. Qiu and U. W. Heinz, Phys. Rev. C84, 024911 (2011)
- Z. Qiu, C. Shen and U. Heinz, Phys. Lett. B707, 151(2012)
- B. Alver and G. Roland, Phys. Rev. C81, 054905 (2010) [Erratum-ibid. C82 , 039903 (2010)]
- B. H. Alver, C. Gombeaud, M. Luzum and J. Y. Ollitrault, Phys. Rev. C82, 034913 (2010)
- M. Luzum and H. Petersen, J. Phys. G41, 063102 (2014)
- F. G. Gardim, F. Grassi, M. Luzum and J. -Y. Ollitrault, Phys. Rev. C85, 024908 (2012)
- H. Niemi, G. S. Denicol, H. Holopainen and P. Huovinen, Phys. Rev. C87, 054901 (2013)  $\bullet$  ...

#### <span id="page-4-0"></span>Here: systematic study.

## Characterizing the initial state

Use the cumulant expansion of Teaney & Yan PRC 83, 064904 (2011): ρ(**x**) ∼ energy or entropy density  $\longrightarrow$   $\rho(\mathbf{k}) = \int d^2x \rho(\mathbf{x})e^{i\mathbf{k}\cdot\mathbf{x}}$  $\longrightarrow W(\mathbf{k}) \equiv \ln \rho(\mathbf{k}) = \sum_{m=1}^{\infty} \sum_{m=1}^{\infty} W_{n,m} k^{m} e^{-in\phi_{k}}$ *n*=−∞ *m*≥|*n*|

Useful properties:

- *n* specifies the rotation symmetry and only *n* ≥ 0 matters (because *W*−*n*,*<sup>m</sup>* ∝ *W*<sup>∗</sup> *<sup>n</sup>*,*m*),
- *m* − *n* is even and *m* ≥ *n*,
- $\bullet$  Hydro sensitive only to large scale  $\Rightarrow$  truncate the sum  $m \leq m_{max}$ .

<span id="page-5-0"></span>Examples  $({} \}$ = weighted average over  $\rho(\mathbf{x})$ :  $W_{1,1} = 0$ ,  $W_{0,2} \propto \{r^2\}$ ,  $W_{2,2} \propto \{r^2 e^{i2\phi}\}$ ,  $W_{1,3} \propto \{r^3 e^{i\phi}\}$  etc. So:  $\epsilon_2 = |2W_{2,2}/W_{0,2}|$ ,  $\epsilon_1 \propto |W_{1,3}|$  etc. KO K KØ K K E K K E K V R K K K K K K K K K

## Find a best estimator *Vest*,*<sup>n</sup>*

#### **Step 1:** Find *Vest*,*<sup>n</sup>* computed from the intial state properties, estimator for

 $V_n = v_n e^{in\Psi_n}$ , complex flow coefficient, obtained from hydro.

Since  $W_{n\neq0,m}$  small, write (including only cumulants leading to same rotation properties as *Vn*):

 $V_{est,n} = \sum_{m=n}^{m_{max}} k_{n,m} W_{n,m}$  $+ \sum_{l=1}^{m} \sum_{m=1}^{m} \sum_{m=1}^{m} \sum_{m'=|n-l|}^{m} K_{l,m,m'} W_{l,m} W_{n-l,m'}$  $+O(W^3)$ .

It is common to use the eccentricities to build estimators but they are not cumulants.  $\varepsilon_{n,m} e^{j\Phi_{n,m}} \equiv -\frac{\{r^m e^{in\phi}\}}{\{r^m\}}$ {*r <sup>m</sup>*} (We can revert to them in the end.)

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#### **Step 2:** Make the best estimator

For simplicity, suppose *Vest*,*<sup>n</sup>* has only one term and re-write

$$
V_{est,n} = k_{n,m} W_{n,m} = k V_{est,n}
$$

with *k* a real scaling coefficient and V*est*,*<sup>n</sup>* unscaled.

Define the error vector  $V_n \equiv V_{est,n} + \mathcal{E}_n = k V_{est,n} + \mathcal{E}_n$ Minimize  $\langle | \mathcal{E}_n|^2 \rangle$  in a given centrality bin

⇒  $\sqrt{ }$  $\left\vert \right\vert$  $\mathcal{L}$  $k = \frac{\text{Re}\langle V_n V_{est,n}^* \rangle}{\sqrt{N_n^2 + N_n^2}}$  $\langle$   $|\mathcal{V}_{est,n}|^2$  $\langle |\mathcal{E}_n|^2 \rangle = \langle |\mathcal{V}_n|^2 \rangle - \langle |\mathcal{V}_{est,n}|^2 \rangle$ . From this last equation, we define the quality of the estimator:

$$
Q_n^2 \equiv \frac{\langle |V_{est,n}|^2 \rangle}{\langle |V_n|^2 \rangle} = 1 - \frac{\langle |\mathcal{E}_n|^2 \rangle}{\langle |V_n|^2 \rangle}
$$

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<span id="page-7-0"></span>The closer *Q<sup>n</sup>* to 1, the better the estimator.

## Effect of viscosity on fluid evolution

Based on J.Noronha-Hostler et al. PRC88(2013)044916, PRC90(2014)034907 Equations of Motion Conservation of Energy and Momentum

$$
D_\mu T^{\mu\nu}=0,
$$

The energy-moment tensor contains a bulk viscous pressure Π and shear stress tensor  $\pi^{\mu\nu}$ 

$$
T^{\mu\nu} = \left(\epsilon + p + \Pi\right)u^{\mu}u^{\nu} - \left(p + \Pi\right)g^{\mu\nu} + \pi^{\mu\nu}
$$

<span id="page-8-0"></span>Using memory function method and minimal IS description  $\tau$ π $u^{\mu}D_{\mu}$ Π + Π =  $-(\zeta + \tau$ πΠ)  $D_{\mu}u^{\mu}$  $\tau_\pi\Delta^{\mu\nu\lambda\rho}\bm u^\alpha\bm D_\alpha\pi_{\lambda\rho}+\pi^{\mu\nu}=\eta\sigma^{\mu\nu}-\tau_\pi\pi^{\mu\nu}\bm D_\alpha\bm u^\alpha$  (standard notations)  $\rightarrow$  4 transport coefficients  $\zeta$ ,  $\eta$ ,  $\tau_{\Pi}$ ,  $\tau_{\pi}$ , which depend on T.  $\pi^{\mu\nu}_{\mathsf{Navier}-\mathsf{Stokes}} = \eta \sigma^{\mu\nu}$ : prevents deformations of fluid cell. Π*Navier*−*Stokes* = −ζ*D*µ*u* µ : negative pressur[e,](#page-7-0) [sl](#page-9-0)[o](#page-7-0)[wi](#page-8-0)[n](#page-9-0)[g](#page-7-0) [e](#page-8-0)[x](#page-13-0)[p](#page-7-0)[a](#page-8-0)[n](#page-12-0)[s](#page-13-0)[io](#page-0-0)[n](#page-18-0)

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## Effect of viscosity on fluid evolution

#### Initial Conditions:

- MC-Glauber: energy density  $= cn_{coll}(\vec{r})$  (*c* adjusted to get midrapidity multiplicity)

<span id="page-9-0"></span> $-\tau_0 = 1$  *fm* (tested)



# Effect of viscosity on fluid evolution



#### Shear+Bulk after  $\tau = 5.6$ *fm*







Shear after  $\tau = 5.6$ *fm* 



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• Viscosity attenuates other forces  $\rightarrow$  smearing of granularity. • Shear dominates, bulk barely affects expansion (ζ/*s* << η/*s*)

## Effect of viscosity on particle emission

Compute observables with Cooper-Frye formula: Particle spectra:  $E \frac{d^3 N}{dp^3} = \int_{f.o.} f(x, p)p^{\mu} d\sigma_{\mu}$  $f = f_{eq} + \delta f_{shear} + \delta f_{bulk}$ δ*fshear*

Common ansatz:  $\delta f_{shear} \sim \pi_{\mu\nu} p^{\mu} p^{\nu} / [(\epsilon + p) T^2]$ . Navier-Stokes limit, δ*fshear* ∝ (η/*s*)*p* 2

 $\rightarrow$  stronger effect for larger  $\eta$ /*s* and *p*.

#### δ*fbulk*

Using method of moments as in Denicol, Niemi NPA904-905 (2013) 369c

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$$
\delta f^{(\pi)}_{\textit{bulk}} = f_{\textit{eq}} \times \Pi \times [B_0^{(\pi)} + D_0^{(\pi)} u.p + E_0^{(\pi)} (u.p)^2]
$$

 $B_0^{(\pi)} = -65.85$  fm<sup>4</sup>,  $D_0^{(\pi)} = 171,27$  fm<sup>4</sup> / GeV,  $E_0^{(\pi)} = -63.05$  fm<sup>4</sup> / GeV<sup>2</sup>

### Effect of viscosity on particle emission



• sensitive to value of viscosities

<span id="page-12-0"></span>• not to choice of  $\delta f_{bulk}$  (though the  $V_n(p_T)$ 's are).

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#### $v_2$  from  $\varepsilon_2$

$$
V_2 \sim V_{est,2} = k_{2,2} W_{2,2} + O(W^3) \sim k \, \epsilon_2 e^{j2\Phi_2}
$$
 if  $m_{max} = 2$ .



- **•**  $Q_2 \sim 1$ : *v*<sub>2</sub>  $\propto$  ε<sub>2</sub> and  $\Psi_2 \sim \Phi_2$
- Shear viscosity improves the estimator
- <span id="page-13-0"></span>NeXSPheRIO "∼ worst-case scenario": ideal + initial tranverse flow and rapidity-dependent fluctuations (not treated in this approach). $\left\{ \begin{array}{ccc} 1 & 0 & 0 \\ 0 & 1 & 0 \end{array} \right.$

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## $v_3$  from  $\varepsilon_3$

## $V_{est,3} = k_{3,3}W_{3,3} + k_{3,5}W_{3,5} + K_{132}W_{1,3}W_{2,2} + O(W^2).$



- $\bullet$   $\varepsilon_3$  and  $v_3$  are strong correlated
- Shear viscosity improves estimator
- $\bullet$   $\varepsilon_1 \varepsilon_2$  and  $\varepsilon_{3,5}$  help for peripheral collisions
- NeXSPheRIO again below.

<span id="page-14-0"></span>Obs.: To get unscaled cumulants, we shoud use  $W^{3/2}_{0,2}$  $\int_{0,2}^{3/2}$  but it makes little difference to u[s](#page-13-0)e  $\{r^3\}$  and revert to eccen[tric](#page-13-0)i[tie](#page-15-0)s[.](#page-14-0)

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#### $v_4$  from  $\varepsilon_4$ ,  $\varepsilon_2^2$  $\frac{2}{2}$  and  $\Phi_2 - \Phi_4$

 $V_{est,4} = k_{4,4}W_{4,4} + k_{2,2,2}W_{2,2}W_{2,2} + O(W^2)$ : both terms important



- Well predicted with two terms (better if there is viscosity).
- $v_4^2 = |V_4|^2 = k^2 \epsilon_4^2 + k'^2 \epsilon_2^4 + 2kk' \epsilon_4 \epsilon_2^2 \cos 4(\Phi_2 \Phi_4)$

<span id="page-15-0"></span>• Small improvement with more terms (5):  $V_{est,4} = k_{4,4}W_{4,4} + k_{2,2,2}W_{2,2}W_{2,2} + k_{2,4,2}W_{2,4}W_{2,2} +$  $K_{2,4,4}$   $W_{2,4}$   $W_{2,4}$  +  $K_{1,3,3}$   $W_{1,3}$   $W_{3,3}$  +  $O(W^2)$ 

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- Two terms OK for central collisions, not peripheral.
- Improvement with more terms (7)

 $\varepsilon_5 + \varepsilon_3 \varepsilon_2 + \varepsilon_{3,5} \varepsilon_2 + \varepsilon_3 \varepsilon_{2,4} + \varepsilon_{3,5} \varepsilon_{2,4} + \varepsilon_1 \varepsilon_4 + \varepsilon_{1,5} \varepsilon_4$ 

 $\left\{ \begin{array}{ccc} 1 & 0 & 0 \\ 0 & 1 & 0 \end{array} \right.$ 

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#### *v*1



- $v_1$  predicted by  $\varepsilon_1$  for central collisions.
- Both higher order terms (sensitive to smaller scale) and non-linear terms necessary for non-central collisions.
- **•** Large effect of viscosities.
- $v_1(p_T)$  changes sign. Smaller  $p_T$  range improves estimator

## Conclusion

- η/*s* and ζ/*s* improve the mapping of the initial state onto the final flow harmonics.
	- Correlation between  $v_n$  and  $\epsilon_n$  larger than for ideal hydro.
- NeXSPheRIO results generally below v-USPhydro. Central collisions more sensitive to this. (NeXSPheRIO has non-zero initial transverse flow and rapidity-dependant fluctuations.)
- <span id="page-18-0"></span>More higher order eccentricities and non-linear terms are necessary for peripheral collisions (what about pA?)

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