

Nuclear Aspects of Dark Matter Searches INT December 2014

Strongly-coupled composite dark matter and lattice field theory

Enrico Rinaldi

Nuclear and Chemical Sciences Division, Lawrence Livermore National Laboratory, Livermore, CA

Lattice Strong Dynamics collaboration

This research was performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under Contract DE-AC52-07NA27344 and supported by the LLNL LDRD "Illuminating the Dark Universe with PetaFlops Supercomputing" 13-ERD-023.

Computing support comes from the LLNL Institutional Computing Grand Challenge program. LLNL-PR

LLNL-PRES-665405



Lattice Strong Dynamics Collaboration

James Osborn



Rich Brower Michael Cheng Claudio Rebbi Oliver Witzel Evan Weinberg



Ethan Neil

Ethan Neil

Sergey Syritsyn













Evan Berkowitz Enrico Rinaldi Chris Schroeder Pavlos Vranas

Joe Kiskis



David Schaich



Tom Appelquist George Fleming Gennady Voronov

Mike Buchoff

Outline

- Features of strongly-coupled composite dark matter
- Searching for a class of models: guidelines
- Importance of lattice field theory simulations
- Polarizability of composite dark matter

Outline

- Features of strongly-coupled composite dark matter
- Searching for a class of models: guidelines
- Importance of lattice field theory simulations
- Polarizability of composite dark matter

in the end it's a nuclear physics problem!

- Dark matter is a composite object
- Composite object is electroweak neutral
- Constituents can have electroweak charges
- Dark matter is stable thanks to a global symmetry (like baryon number)

Dark matter is a composite object



Akin to a technibaryon

- Composite object is electroweak neutral
- Constituents can have electroweak charges
- Dark matter is stable thanks to a global symmetry (like baryon number)

- Dark matter is a composite object
 Composite object is electroweak neutral
- Constituents can have electroweak charges
- Dark matter is stable thanks to a global symmetry (like baryon number)



 Dark matter is stable thanks to a global symmetry (like baryon number)



What do we have in mind?

- In general we think about a new strongly-coupled gauge sector like QCD with a plethora of composite states in the spectrum
- Dark fermions have dark color and also have electroweak charges
- Depending on the model, dark fermions have electroweak breaking masses (chiral), electroweak preserving masses (vector) or a mixture
- A global symmetry of the theory naturally stabilizes the dark baryonic composite states (e.g. neutron)

What do we have in mind?

- In general we think about a new strongly-coupled gauge sector like QCD with a plethora of composite states in the spectrum
- Dark fermions have dark color and also have electroweak charges
- Depending on the model, dark fermions have electroweak breaking masses (chiral), electroweak preserving masses (vector) or a mixture
- A global symmetry of the theory naturally stabilizes the dark baryonic composite states (e.g. neutron)

we construct a minimal model with these features

What do we have in mind?

- In general we think about a new strongly-coupled gauge sector like QCD with a plethora of composite states in the spectrum
- Dark fermions have dark color and also have electroweak charges
- Depending on the model, dark fermions have electroweak breaking masses (chiral), electroweak preserving masses (vector) or a mixture
- A global symmetry of the theory naturally stabilizes the dark baryonic composite states (e.g. neutron)

today focus on the general features

"How dark is dark matter?"

Interactions of neutral object with photons

- dimension 5 ➡ magnetic dipole
- $\frac{(\bar{\chi}\sigma^{\mu\nu}\chi)F_{\mu\nu}}{\Lambda_{\rm dark}}$ $\frac{(\bar{\chi}\chi)v_{\mu}\partial_{\nu}F^{\mu\nu}}{\Lambda_{\rm dark}^{2}}$
- dimension 6 \blacktriangleright charge radius
- dimension 7 \blacktriangleright polarizability



 $(\bar{\chi}\chi)v_{\mu}\partial_{\nu}F^{\mu\nu}$

2 dark

"How dark is dark matter?"

Interactions of neutral object with photons

 $\frac{(\bar{\chi}\sigma^{\mu\nu}\chi)F_{\mu\nu}}{\Lambda_{\rm dark}}$

- dimension $5 \Rightarrow$ magnetic dipole
- dimension $6 \Rightarrow$ charge radius
- dimension 7 \blacktriangleright polarizability

our guy for today

 $(\bar{\chi}\chi)F_{\mu
u}F^{\mu
u}$

- remove magnetic dipole moment:
 - lightest stable baryon is a boson with S=0
- remove charge radius:
 - 2 flavors with degenerate masses

polarizability can not be removed

$$\mathcal{O}_F^{\chi} = \mathcal{C}_F^{\chi} \, \bar{\chi} \chi F^{\mu\nu} F_{\mu\nu}$$









Importance of lattice field theory techniques

- lattice simulations are naturally suited for models where dark fermion masses are comparable to the confinement scale
- controllable systematic errors and room for improvement
- Naive dimensional analysis and EFT approaches can miss important non-perturbative contributions
- NDA is not precise enough when confronting experimental results and might not work for certain situations

Importance of lattice field theory techniques

- lattice simulations are naturally suited for models where dark fermion masses are comparable to the confinement scale
- controllable systematic errors and room for improvement
- Naive dimensional analysis and EFT approaches can miss important non-perturbative contributions
- NDA is not precise enough when confronting experimental results and might not work for certain situations



[Detmold, Tiburzi & Walker-Loud, Phys. Rev. D79 (2009) 094505 and Phys. Rev. D81 (2010) 054502] [LSD collab., in preparation]

Polarizability of DM from lattice simulations

- Background field method: response of neutral baryon to external electric field ${\mathcal E}$
- Measure the shift of the baryon mass as a function of $\ensuremath{\mathcal{E}}$

$$E_{SU(3)} = M_{\chi} + \frac{1}{2} (\mathcal{C}_F^{\chi} + \frac{\mu^2}{4M_{\chi}^3}) \mathcal{E}^2 + \text{h.o.}$$
$$E_{SU(4)} = M_{\chi} + \frac{1}{2} (\mathcal{C}_F^{\chi}) \mathcal{E}^2 + \text{h.o.}$$

Precise lattice results



[Pospelov & Veldhuis, Phys. Lett. B480 (2000) 181] [Weiner & Yavin, Phys. Rev. D86 (2012) 075021] [Frandsen et al., JCAP 1210 (2012) 033] [Ovanesyan & Vecchi, arxiv:1410.0601]

Nuclear polarizability (Rayleigh scattering)

- several attempts to estimate this in the past, with increasing level of complexity in a perturbative setup
 multiple scales are probed by the momentum transfer in the virtual photons loop
 - $\begin{array}{cccc} \chi & \mathcal{O}\chi & \chi \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ \end{array}$
- mixing operators and threshold corrections appear at leading order and interference is possible
- nuclear matrix element has non-trivial excited state structure that requires non-perturbative treatment
- $\langle A | \bar{\chi} \chi F^{\mu\nu} F_{\mu\nu} | A \rangle$
- similar structure arising in double beta decay matrix elements

 $\mathcal{O}^{\chi} M_{\gamma} \mathcal{O}^{M}$ M CM MM $M \mathcal{O}^M M$ E \mathcal{F}_{gg}^{gg} qqΑ Ø A A $\mathcal{O}_G^{\chi} = \mathcal{C}_G^{\chi} \,\bar{\chi} \chi G^{\mu\nu} G_{\mu\nu} \qquad \mathcal{O}_F^{\chi} = \mathcal{C}_F^{\chi} \,\bar{\chi} \chi F^{\mu\nu} F_{\mu\nu}$ $\mathcal{O}_q^{\chi} = \mathcal{C}_q^{\chi} m_q \, \bar{\chi} \chi \bar{q} q$ m_Q m_q Q









Naive dimensional analysis of $\langle A | \bar{\chi} \chi F^{\mu\nu} F_{\mu\nu} | A \rangle$



- to asses the impact of uncertainties on the total cross section we start from naive dimensional analysis
- we allow a "magnitude" factor M_F^A to change from 1 to 25













Summary and future directions

- strongly-coupled composite dark matter is an interesting scenario that should not be overlooked
- within this space of theories it is not hard to find regions where all interactions with SM are suppressed up to dimension-7 operators
- dimension-7 EM polarizability can not be eliminated in any case
- lattice simulations can calculate the EM form factors of the composite object with controllable errors (using mature LQCD techniques)
- nuclear physics input is needed and nuclear matrix elements have the largest uncertainties that should be assessed

Questions ?

Questions ? YES.

- Q: How much does the nuclear matrix element influence the conclusions?
 - we tried to estimate this and it seems a O(25) change could affect results
- Q: What methods can be used to evaluate the nuclear matrix element?
 - can we learn something from double-beta decays? or electron-nuclei scattering?
- Q: Are there experimental limits that can bound the matrix element?
 - not sure at the moment.

Backup slides

S. Nussinov, Phys. Lett. B165 (1985) 55

S. Barr, R. S. Chivukula, and E. Farhi, Phys. Lett. B241 (1990) 387

D. B. Kaplan, Phys. Rev. Lett. 68 (1992) 741

Asymmetric dark matter

S. Nussinov, Phys. Lett. B165 (1985) 55 S. Barr, R. S. Chivukula, and E. Farhi, Phys. Lett. B241 (1990) 387 D. B. Kaplan, Phys. Rev. Lett. 68 (1992) 741

Asymmetric dark matter

 It is an observational fact that the number density for dark matter and baryonic matter are of the same order of magnitude

 $\Omega_{\rm DM} \approx 5 \ \Omega_{\rm B}$



Asymmetric dark matter

 It is an observational fact that the number density for dark matter and baryonic matter are of the same order of magnitude

$$\Omega_{\rm DM} \approx 5 \ \Omega_{\rm B}$$

• This can be explained in Technicolor theories where dark matter is a baryon of a new strongly-coupled sector which shares an asymmetry with standard baryonic matter

$$n_{\rm DM} - \bar{n}_{\rm DM} \approx n_{\rm B} - \bar{n}_{\rm B}$$

Dark Matter 26.8% Driversy Matter 88.3% EPlanck and ESA

S. Nussinov, Phys. Lett. B165 (1985) 55 S. Barr, R. S. Chivukula, and E. Farhi, Phys. Lett. B241 (1990) 387 D. B. Kaplan, Phys. Rev. Lett. 68 (1992) 741

Asymmetric dark matter

 It is an observational fact that the number density for dark matter and baryonic matter are of the same order of magnitude

$$\Omega_{\rm DM} \approx 5 \ \Omega_{\rm B}$$

S. Nussinov, Phys. Lett. B165 (1985) 55

D. B. Kaplan, Phys. Rev. Lett. 68 (1992) 741

S. Barr, R. S. Chivukula, and E. Farhi, Phys. Lett. B241 (1990) 387

 This can be explained in Technicolor theories where dark matter is a baryon of a new strongly-coupled sector which shares an asymmetry with standard baryonic matter

$$n_{\rm DM} - \bar{n}_{\rm DM} \approx n_{\rm B} - \bar{n}_{\rm B}$$





A composite dark matter model

- Let's focus on a SU(N) dark gauge sector with N=4
- Let dark fermions interact with the SM Higgs and obtain current/chiral masses
- Let's introduce vector-like masses for dark fermions that do not break EW symmetry

Field	$ \mathrm{SU}(N)_D $	$(\mathrm{SU}(2)_L, Y)$	Q
$F_1 = \begin{pmatrix} F_1^u \\ F_1^d \end{pmatrix}$	N	(2,0)	$\left(\begin{array}{c} +1/2\\ -1/2 \end{array}\right)$
$F_2 = \begin{pmatrix} F_2^u \\ F_2^d \end{pmatrix}$	$\overline{\mathbf{N}}$	(2 ,0)	$\begin{pmatrix} +1/2 \\ -1/2 \end{pmatrix}$
F_3^u	N	(1, +1/2)	+1/2
F_3^d	N	(1, -1/2)	-1/2
F_4^u	N	(1, +1/2)	+1/2
F_4^d	$\overline{\mathbf{N}}$	(1, -1/2)	-1/2

A composite dark matter model



A composite dark matter model



Higgs exchange cross section

- Need to non-perturbatively evaluate the σ-term of the dark baryon (scalar nuclear form factor)
- Effective Higgs coupling non-trivial with mixed chiral and vector-like masses
- Model-dependent answer for the cross-section in this channels
- A non-negligible vector mass is needed to evade direct detection bounds

$$m_{f}(h) = m + \frac{yh}{\sqrt{2}}$$

$$\alpha \equiv \frac{v}{m_{f}} \frac{\partial m_{f}(h)}{\partial h} \Big|_{h=v} = \frac{yv}{\sqrt{2}m + yv} \leq 1$$

$$\int_{1\times10^{-43}}^{5\times10^{-43}} M_{PS}/M_{V} = 0.77$$

Higgs exchange cross section

- Need to non-perturbatively evaluate the σ-term of the dark baryon (scalar nuclear form factor)
- Effective Higgs coupling non-trivial with mixed chiral and vector-like masses
- Model-dependent answer for the cross-section in this channels
- A non-negligible vector mass is needed to evade direct detection bounds

$$m_{f}(h) = m + \frac{yh}{\sqrt{2}}$$

$$\alpha \equiv \frac{v}{m_{f}} \frac{\partial m_{f}(h)}{\partial h} \Big|_{h=v} = \frac{yv}{\sqrt{2}m + yv} \leq 1$$

$$s \approx 10^{-43} \frac{M_{PS}/M_{V} = 0.77}{1 \times 10^{-43}} \frac{m_{PS}}{5 \times 10^{-44}} \frac{m_{PS}}{5 \times 10^{-44}} \frac{m_{PS}}{5 \times 10^{-46}} \frac{m_{PS}}{5 \times 10^{-46}}$$

Magnetic moment and charge radius of DM

- Need non-perturbative calculation of form-factors for DM composite object
- Negligible dependence on constituent mass and number of flavors
- Magnetic moment dominates for masses > 25GeV





$$\mathcal{O}_q^{\chi} = \mathcal{C}_q^{\chi} \, m_q \, \bar{\chi} \chi \bar{q} q$$

 $\left| \sigma \simeq \frac{\mu_{n\chi}^2}{\pi A^2} \left\langle \left| \sum_{q} \mathcal{C}_q^{\chi} f_q^A + \mathcal{C}_G^{\chi} f_G^A + \mathcal{C}_F^{\chi} f_F^A \right|^2 \right\rangle \right|$

$$\mathcal{O}_{q}^{\chi} = \mathcal{C}_{q}^{\chi} m_{q} \, \bar{\chi} \chi \bar{q} q$$
$$\mathcal{O}_{Q}^{\chi} = \mathcal{C}_{G}^{\chi} \, \bar{\chi} \chi G^{\mu\nu} G_{\mu\nu}$$

$$\sigma \simeq \frac{\mu_{n\chi}^2}{\pi A^2} \left\langle \left| \sum_q \mathcal{C}_q^{\chi} f_q^A + \mathcal{C}_G^{\chi} f_G^A + \mathcal{C}_F^{\chi} f_F^A \right|^2 \right\rangle$$

$$\mathcal{O}_{q}^{\chi} = \mathcal{C}_{q}^{\chi} m_{q} \, \bar{\chi} \chi \bar{q} q$$
$$\mathcal{O}_{F}^{\chi} = \mathcal{C}_{F}^{\chi} \, \bar{\chi} \chi F^{\mu\nu} F_{\mu\nu}$$
$$\mathcal{O}_{G}^{\chi} = \mathcal{C}_{G}^{\chi} \, \bar{\chi} \chi G^{\mu\nu} G_{\mu\nu}$$

$$\sigma \simeq \frac{\mu_{n\chi}^2}{\pi A^2} \left\langle \left| \sum_q \mathcal{C}_q^{\chi} f_q^A + \mathcal{C}_G^{\chi} f_G^A + \mathcal{C}_F^{\chi} f_F^A \right|^2 \right\rangle$$