Dark Matter without Astrophysics: Halo-independent methods at Direct Detection Experiments

Patrick Fox

‡ Fermilab

Based on work with A. Anderson, G. Jung, Y. Kahn, G. Kribs, J Liu, M. McCullough, T. Tait, P Sorensen and N. "break dancer" Weiner

Astrophysics independence

Comparing two positive results: Drees and Shan; 0803.4477, 0809.1295

vmin space: PJF, Liu, Weiner; 1011.1915

Applications to CDMS, CoGeNT, COUPP, CRESST, DAMA, LUX, XENON, light DM, etc; 1103.3481, 1106.0743, 1106.6241, 1107.0715, 1107.0741, 1110.5338, 1304.6066, 1304.6183,....

Annual modulation: Herrero-Garcia et al.; 1112.1627, 1205.0134

Including detector effects: Gelmini and Gondolo; 1202.6359

Extension to iDM: Bozorgnia et al.; 1305.3575

Extension to general interactions and kinematics: Del Nobile et al.; 1306.7273

Unbinned analysis: PJF, Kahn, McCullough; 1403.6830

Wednesday, 10 December 14

Astrophysics independence

Comparing two positive results: Drees and Shan; 0803.4477, 0809.1295



$$\frac{dR}{dE_R} = \frac{N_T \rho_{\chi}}{m_{\chi}} \int_{v_{\min}}^{v_{\max}} d^3 \vec{v} f(\vec{v}(t)) \frac{d\sigma |\vec{v}|}{dE_R}$$

with,

$$\frac{d\sigma}{dE_R} = F_N^2(E_R)\frac{m_N}{\mu v^2}\bar{\sigma}(v, E_R) \qquad \bar{\sigma}_i(v, E_R) = \begin{cases} \sigma_{i0} \\ \sigma_{i0}F_{\chi_i}^2(E_R) \\ \sigma_{i0}(v)F_{\chi_i}^2(E_R) \\ \sigma_{i0}(v, E_R) \end{cases}$$

Usual case (SI elastic WIMP)

$$\sigma_{N} = \frac{(Zf_{p} + (A - Z)f_{n})^{2}}{f_{p}^{2}} \frac{\mu_{N\chi}^{2}}{\mu_{n\chi}^{2}} \sigma_{p} \qquad v_{min} = \sqrt{\frac{m_{N}E_{R}}{2\mu_{N\chi}^{2}}}$$

Theory

$$\frac{dR}{dE_R} = \frac{N_T \rho_{\chi}}{m_{\chi}} \int_{v_{\min}}^{v_{\max}} d^3 \vec{v} f(\vec{v}(t)) \frac{d\sigma |\vec{v}|}{dE_R}$$

Expt.

$$\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE'_R G(E_R, E'_R) \epsilon(E'_R) F^2(E'_R) g(v_{min}(E'_R))$$

Theory

$$\frac{dR}{dE_R} = \frac{N_T \rho_{\chi}}{m_{\chi}} \int_{v_{\min}}^{v_{\max}} d^3 \vec{v} f(\vec{v}(t)) \frac{d\sigma |\vec{v}|}{dE_R}$$
Expt.
resolution field efficiency

 $\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE_R' G(E_R, E_R') \epsilon(E_R') F^2(E_R') g(v_{min}(E_R'))$

There are known knowns. These are things we know that we know. There are known unknowns. That is to say, there are things that we know we don't know. But there are also unknown unknowns. There are things we don't know we don't know.



 $\frac{dR}{dE_R} = \frac{N_T \rho_{\chi}}{m_{\chi}} \int_{v_{\min}}^{v_{\max}} d^3 \vec{v} f(\vec{v}(t)) \frac{d\sigma |\vec{v}|}{dE_R}$

Expt.

Theory

resolution — $\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n_\chi}^2} C_T^2(A, Z) \int dE_R' G(E_R, E_R') \epsilon(E_R') F^2(E_R') g(v_{min}(E_R'))$

Wednesday, 10 December 14

There are known knowns. These are things we know that we know. There are known unknowns. That is to say, there are things that we know we don't know. But there are also unknown unknowns. There are things we don't know we don't know.



 $\frac{dR}{dE_R} = \frac{N_T \rho_{\chi}}{m_{\chi}} \int_{v_{\min}}^{v_{\max}} d^3 \vec{v} f(\vec{v}(t)) \frac{d\sigma |\vec{v}|}{dE_R}$

Expt.

Theory

$\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE'_R G(E_R, E'_R) \epsilon(E'_R) F^2(E'_R) g(v_{min}(E'_R))$

"In theory there is no difference between theory and practice. But in practice there is.--Yogi Berra"

The uncertain world of direct detection

Nuclear form factors, SI/SD + >2 others [Atomic physics, detector response.
e.g. Leff, keVee → keVnr

[see talks from this workshop]

 Astrophysics How is DM distributed in our neighbourhood? So far we have only transited ~0.007pc, way below Nbody resolution **Baryons? SIDM?** Dark disc? [Randall et al.] **Debris flow?** [Lisanti et al.] Gravitational focusing?

Dependence on details of velocity distribution



Wednesday, 10 December 14

Dependence on details of velocity distribution



Wednesday, 10 December 14

Is a signal a measurement of particle physics or astrophysics?

The only way we have of probing our **local** DM distribution

 $f_1(v_{\min}(E_R)) = -\frac{4\mu^2 E_R^2}{m_N^2 E_R^2 - \mu^2 \delta^2} \frac{1}{\mathcal{N}\sigma_0(v_{\min}(E_R)) F_\chi^2(E_R)} \left(\frac{d\mathcal{R}}{dE_R} - \mathcal{R}\frac{1}{F_\chi^2(E_R)} \frac{dF_\chi^2(E_R)}{dE_R}\right)$

f-condition: $f(v) \ge 0$

(Deconvoluted) rate is a monotonically decreasing function, or there is non-standard particle physics e.g. inelastic or an increasing DM form factor

Is a signal a measurement of particle physics or astrophysics?

The only way we have of probing our **local** DM distribution

$$f_{1}(v_{\min}(E_{R})) = -\frac{4\mu^{2}E_{R}^{2}}{m_{N}^{2}E_{R}^{2} - \mu^{2}\delta^{2}} \frac{1}{\mathcal{N}\sigma_{0}(v_{\min}(E_{R}))F_{\chi}^{2}(E_{R})} \left(\frac{d\mathcal{R}}{dE_{R}} - \mathcal{R}\frac{1}{F_{\chi}^{2}(E_{R})}\frac{dF_{\chi}^{2}(E_{R})}{dE_{R}}\right)$$

$$f_{1}(v) = \int d\Omega f(\vec{v}).$$

f-condition: $f(v) \ge 0$

(Deconvoluted) rate is a monotonically decreasing function, or there is non-standard particle physics e.g. inelastic or an increasing DM form factor

Is a signal a measurement of particle physics or astrophysics?

The only way we have of probing our **local** DM distribution

$$f_{1}(v_{\min}(E_{R})) = -\frac{4\mu^{2}E_{R}^{2}}{m_{N}^{2}E_{R}^{2} - \mu^{2}\delta^{2}} \frac{1}{\mathcal{N}\sigma_{0}(v_{\min}(E_{R}))F_{\chi}^{2}(E_{R})} \left(\frac{d\mathcal{R}}{dE_{R}} - \mathcal{R}\frac{1}{F_{\chi}^{2}(E_{R})} \frac{dF_{\chi}^{2}(E_{R})}{dE_{R}} \right)$$

$$f_{1}(v) = \int d\Omega f(\vec{v}). \qquad \mathcal{R} \equiv \frac{1}{F_{N}^{2}(E_{R})} \frac{dR}{dE_{R}}$$

f-condition: $f(v) \ge 0$

(Deconvoluted) rate is a monotonically decreasing function, or there is non-standard particle physics e.g. inelastic or an increasing DM form factor

Neal is not a great break dancer...

Neal is not a great break dancer...



... but his son is!

- •Cannot determine *anything* about DM from a single positive result
- •Assuming an escape velocity can place a *lower* bound on DM mass, but no upper bound
- •Luckily we have other complementary searches, and multiple direct detection expts.

Direct Detection without bias

[PF, Liu, Weiner]

$$\frac{dR}{dE_R} = \frac{N_T M_T \rho}{2m_\chi \mu^2} \int_{v_{min}}^{v_{max}} d^3 \vec{v} \frac{f(\vec{v}, \vec{v_E})}{v} \sigma(E_R)$$

$$\frac{dR}{dE_R} = \frac{N_T M_T F_N^2(E_r)}{2\mu^2} \frac{\rho\sigma}{m_\chi} g(v)$$

$$v_{min} = \sqrt{\frac{M_T E_R}{2\mu^2}}$$

Recoil energy uniquely determines minimum DM velocity

Direct Detection without bias

[PF, Liu, Weiner]



$$v_{min} = \sqrt{\frac{M_T E_R}{2\mu^2}}$$

Recoil energy uniquely determines **minimum** DM velocity

Direct Detection without bias

[PF, Liu, Weiner]



$$v_{min} = \sqrt{\frac{M_T E_R}{2\mu^2}}$$

Recoil energy uniquely determines **minimum** DM velocity

Comparing experiments-vmin space

 $N_T = \kappa N_A m_p / M_T$

Solve for g(v)

$$g(v_{min}) = \frac{2m_{\chi}\mu^2}{N_A \kappa m_p \rho \sigma(E_R)} \frac{dR_1}{dE_1}$$
$$\frac{dR_1}{dE_1} \iff g(v_{min}) \iff \frac{dR_2}{dE_2}$$
$$v_{min} = \sqrt{\frac{M_T E_R}{2\mu^2}}$$

The master formula (SI):

 $C_T^{(i)} = \kappa^{(i)} \left(f_p Z^{(i)} + f_n \left(A^{(i)} - Z^{(i)} \right) \right)^2$

$$\frac{dR_2}{dE_R} \left(E_2 \right) = \frac{C_T^{(2)}}{C_T^{(1)}} \frac{F_2^2(E_2)}{F_1^2 \left(\frac{\mu_1^2 M_T^{(2)}}{\mu_2^2 M_T^{(1)}} E_2 \right)} \frac{dR_1}{dE_R} \left(\frac{\mu_1^2 M_T^{(2)}}{\mu_2^2 M_T^{(1)}} E_2 \right)$$

Using vmin space



Using vmin space

Experiment I \longleftrightarrow Experiment 2 $[E_{low}^{(1)}, E_{low}^{(1)}] \iff [v_{min}^{low}, v_{min}^{high}] \iff [E_{low}^{(2)}, E_{high}^{(2)}]$



Using vmin space





A more direct comparison of data than sigma-m plots Easy to derive from data

For eDM (and single target expts.) need only show for one mass

Ultimately allows for measurements of g(v)Consistency of g(v) determines allowed DM params.



A more direct comparison of data than sigma-m plots Easy to derive from data

For eDM (and single target expts.) need only show for

one mass Ultimately allows for measurements of g(v) Consistency of g(v) determines allowed DM params.



Speed distribution is positive semidefinite $f(v) \geq 0$



Integral monotonically decreases



"Least" monotonic function is a step function $\Theta(v_1 - v_{\min})$







Bounding g(v)

$$\frac{dR}{dE_R} = \frac{N_T M_T \rho}{2m_\chi \mu^2} \sigma(E_R) g_1 \Theta(v_1 - v_{min}(E_R))$$

Use standard statistical techniques (max gap, Pmax, etc) to bound $\rho\sigma g_1/m_{\chi}$

Most conservative possible limits

Measuring g(v)

$$g(v_{min}) = \frac{2m_{\chi}\mu^2}{N_A \kappa m_p \rho \sigma(E_R)} \frac{dR_1}{dE_1}$$

Convert binned recoil events into measurements in vmin space

Some potential problems:

Loss of information How to treat resolution, efficiencies [see eg Gelmini and Gondolo] Particularly a problem early on, with low statistics





[PF, Kahn, McCullough] see also [Feldstein, Kahlhoefer]

$$\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE'_R G(E_R, E'_R) \epsilon(E'_R) F^2(E'_R) g(v_{min}(E'_R))$$

$$L/2 = N_E - \sum_{i=1}^{N_O} \log \left. \frac{dR_T}{dE_R} \right|_{E_R = E_i}$$

Taking the moth

ut of the bin

[PF, Kahn, McCullough] see also [Feldstein, Kahlhoefer]

Address
Extended
$$\mathcal{L} = \frac{e^{-N_E}}{N_O!} \prod_{i=1}^{N_O} \frac{dR_T}{dE_R} \Big|_{E_R = E_i}$$

$$L = -2\log(\mathcal{L})$$

$$\frac{dR}{dE_R} = \frac{N_{\text{ext}} - n^{\text{ren}} R}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE_R' G(E_R, E_R') \epsilon(E_R') F^2(E_R') g(v_{min}(E_R'))$$

$$L/2 = N_E - \sum_{i=1}^{N_O} \log \left. \frac{dR_T}{dE_R} \right|_{E_R = E_i}$$

[PF, Kahn, McCullough] see also [Feldstein, Kahlhoefer]

$$\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE'_R G(E_R, E'_R) \epsilon(E'_R) F^2(E'_R) g(v_{min}(E'_R))$$

$$L/2 = N_E - \sum_{i=1}^{N_O} \log \left. \frac{dR_T}{dE_R} \right|_{E_R = E_i}$$

[PF, Kahn, McCullough] see also [Feldstein, Kahlhoefer]

$$\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE'_R G(E_R, E'_R) \epsilon(E'_R) F^2(E'_R) g(v_{min}(E'_R))$$

$$L/2 = N_E - \sum_{i=1}^{N_O} \log \left. \frac{dR_T}{dE_R} \right|_{E_R = E_i}$$

[PF, Kahn, McCullough] see also [Feldstein, Kahlhoefer]





Minimize L whilst preserving monotonicity



Best fit is given by a series of $N_{\cal O}$ step functions

For perfect energy resolution step positions given by energies of events $v_{min}(E_i)$

Minimize L over step heights

Confidence region for parameters

$$\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE'_R G(E_R, E'_R) \epsilon(E'_R) F^2(E'_R) g(v_{min}(E'_R))$$

In reality energy smearing means each step contributes at all energies What is the form of g(v) that minimizes LL?

Still series of N_O steps but now positions shift to the right Minimize over $2N_O$ parameters

$$\frac{dR}{dE_R} = \frac{N_A \rho_\chi \sigma_n m_n}{2m_\chi \mu_{n\chi}^2} C_T^2(A, Z) \int dE'_R G(E_R, E'_R) \epsilon(E'_R) F^2(E'_R) g(v_{min}(E'_R))$$

In reality energy smearing means each step contributes at all energies What is the form of g(v) that minimizes LL?

Still series of N_O steps but now positions shift to the right Minimize over $2N_O$ parameters

KKT - "Lagrange multipliers"

General resolution function (think Gaussian)

 $G(E_R, E'_R)$ to have the following properties:

- (i) $\int G(E_R, E'_R) dE'_R = 1$ for any E_R .
- (ii) As a function of E'_R for fixed E_R , $G(E_R, E'_R)$ has a single local maximum at $E'_R = E_R$ and no other local extrema.
- (iii) For $E_R \neq E'_R$, either $G(E_R, E'_R) = 0$ or $\partial G(E_R, E'_R) / \partial E'_R \neq 0$.

Minimize,

$$L[\tilde{g}] = \int dE'_R K(E'_R) \tilde{g}(E'_R) - \sum_{i=1}^{N_O} \log\left(\mu_i + \int dE'_R G(E_i, E'_R) K(E'_R) \tilde{g}(E'_R)\right)$$

subject to inequality constraint: $d\tilde{g}/dE_R' \leq 0$

KKT - "Lagrange multipliers"

General resolution function (think Gaussian)

 $G(E_R, E'_R)$ to have the following properties:

- (i) $\int G(E_R, E'_R) dE'_R = 1$ for any E_R .
- (ii) As a function of E'_R for fixed E_R , $G(E_R, E'_R)$ has a single local maximum at $E'_R = E_R$ and no other local extrema.
- (iii) For $E_R \neq E'_R$, either $G(E_R, E'_R) = 0$ or $\partial G(E_R, E'_R) / \partial E'_R \neq 0$.

Minimize, $I[\tilde{g}] = \int dE'_R K(E'_R) \tilde{g}(E'_R) - \sum_{i=1}^{N_O} \log \left(\mu_i + \int dE'_R G(E_i, E'_R) K(E'_R) \tilde{g}(E'_R) \right)$

subject to inequality constraint: $d\tilde{g}/dE_R' \leq 0$

KKT - "Lagrange multipliers"

General resolution function (think Gaussian)

 $G(E_R, E'_R)$ to have the following properties: (i) $\int G(E_R, E'_R) dE'_R = 1$ for any E_R . **normalised** (ii) As a function of E'_R for fixed E_R , $G(E_R, E'_R)$ has a single local maximum at $E'_R = E_R$ and no other local extrema.

(iii) For $E_R \neq E'_R$, either $G(E_R, E'_R) = 0$ or $\partial G(E_R, E'_R) / \partial E'_R \neq 0$.

not flat anywhere, unless O

Minimize,

$$L[\tilde{g}] = \int dE'_R K(E'_R) \tilde{g}(E'_R) - \sum_{i=1}^{N_O} \log\left(\mu_i + \int dE'_R G(E_i, E'_R) K(E'_R) \tilde{g}(E'_R)\right)$$

subject to inequality constraint: $d\tilde{g}/dE_R' \leq 0$

Karush-Kuhn-Tucker

To enforce monotonicity

$$\begin{split} L[\tilde{g}] &\to L[\tilde{g}] + \int dE'_R \, \frac{d\tilde{g}}{dE'_R} q(E'_R) \\ & \frac{\delta L}{\delta \tilde{g}} - \frac{dq}{dE'_R} = 0 \ , \\ & \frac{d\tilde{g}}{dE'_R} \leq 0 \ , \\ & q(E'_R) \geq 0 \ , \\ & \int dE'_R \, \frac{d\tilde{g}}{dE'_R} q(E'_R) = 0 \ . \end{split}$$

Karush-Kuhn-Tucker

To enforce monotonicity

$$\begin{split} L[\tilde{g}] \rightarrow L[\tilde{g}] + \int dE'_R \, \frac{d\tilde{g}}{dE'_R} q(E'_R) \\ & \frac{\delta L}{\delta \tilde{g}} - \frac{dq}{dE'_R} = 0 , \\ & \frac{d\tilde{g}}{dE'_R} \leq 0 , \\ & q(E'_R) \geq 0 , \\ \int dE'_R \, \frac{d\tilde{g}}{dE'_R} q(E'_R) = 0 . \end{split}$$

Either saturate constraint (usual Lagrange multiplier) or q=0



Assume g(E) is <u>not flat</u> over some range. Then,

$$\sum_{i=1}^{N_O} \frac{G(E_i, E'_R)}{\gamma_i} = 1,$$

 $\frac{\delta L}{\delta \tilde{g}} - \frac{dq}{dE'_R} = 0 \; ,$

 $\int dE'_R \frac{d\tilde{g}}{dE'_R} q(E'_R) = 0 \; .$

 $\frac{d\tilde{g}}{dE'_{R}} \le 0 \ ,$

 $q(E'_R) \ge 0 \; ,$

with

$$\gamma_i = \mu_i + \int dE_R'' G(E_i, E_R'') K(E_R'') \tilde{g}(E_R'')$$

g(v) must be <u>flat</u> except at individual points i.e. steps



Assume g(E) is <u>not flat</u> over some range. Then,

depends on E' $\sum_{i=1}^{N_O} \frac{G(E_i, E'_R)}{\gamma_i} = 1, \quad \text{const.}$

 $\frac{\delta L}{\delta \tilde{g}} - \frac{dq}{dE'_R} = 0 \; ,$

 $\int dE'_R \, \frac{d\tilde{g}}{dE'_R} q(E'_R) = 0 \; .$

 $\frac{d\tilde{g}}{dE'_{\mathcal{P}}} \le 0 \ ,$

 $q(E'_R) \ge 0 \; ,$

with

$$\gamma_i = \mu_i + \int dE_R'' G(E_i, E_R'') K(E_R'') \tilde{g}(E_R'')$$

g(v) must be <u>flat</u> except at individual points i.e. steps

ΚΚΤ



$$\gamma_i = \mu_i + \int dE_R'' G(E_i, E_R'') K(E_R'') \tilde{g}(E_R'')$$

g(v) must be <u>flat</u> except at individual points i.e. steps



Can determine the positions of the steps

The the positions of the steps

$$\frac{dq}{dE'_R} = K(E'_R) \left(1 - \sum_{i=1}^{N_O} \frac{G(E_i, E'_R)}{\gamma_i} \right) \checkmark \int dE'_R \frac{d\tilde{g}}{dE'_R} q(E'_R) = 0.$$

At step energies (\tilde{E}_i) positivity of q requires

$$\frac{dq}{dE'_R}\Big|_{\widetilde{E}_j} = 0, \quad \frac{d^2q}{dE'^2_R}\Big|_{\widetilde{E}_j} \ge 0$$

Thus,

$$-\sum_{i=1}^{N_O} \frac{1}{\gamma_i} \left. \frac{\partial G(E_i, E'_R)}{\partial E'_R} \right|_{E'_R = \widetilde{E}_j} \gtrsim 0$$

Steps shift to the right



Varying the DM mass



$$v'_{min}(E_R) = \frac{\mu_{N\chi}}{\mu_{N\chi'}} v_{min}(E_R)$$



One plot is sufficient to determine results for all masses

Unbinned application

[PF, Kahn, McCullough]



Unbinned application

[PF, Kahn, McCullough]





F

0



Extension to inelastic DM

[see also Bozorgnia et al; del Nobile et al]

$$v_{\min} = \left| \delta + \frac{m_{\mathrm{N}} E_{\mathrm{R}}}{\mu} \right| \frac{1}{\sqrt{2 E_{\mathrm{R}} m_{\mathrm{N}}}}$$



Wednesday, 10 December 14

Extension to inelastic DM

[see also Bozorgnia et al; del Nobile et al]

$$v_{\min} = \left| \delta + \frac{m_{\mathrm{N}} E_{\mathrm{R}}}{\mu} \right| \frac{1}{\sqrt{2 E_{\mathrm{R}} m_{\mathrm{N}}}}$$



Conclusions

- Should analyse data independent of astro uncertainties
 With multiple experiments should compare g(v), tests consistency
- •Presenting experimental results in g(v) very useful
 - •One plot contains all information, for all masses
- •Find region of consistent parameter space
- Unbinned approach using likelihood techniques
 Maximal use of information
 - Independent of astrophysics, expts. agree/disagree?
- •Application: CDMS-Si is at odds with LUX
- •Ultimately may be able to extract f(v) by differentiating deconvoluted rate