

Three-cluster dynamics within an ab initio framework

*Universality in Few-Body Systems:
Theoretical Challenges and New Directions*

INT 14-1

S. Quaglioni

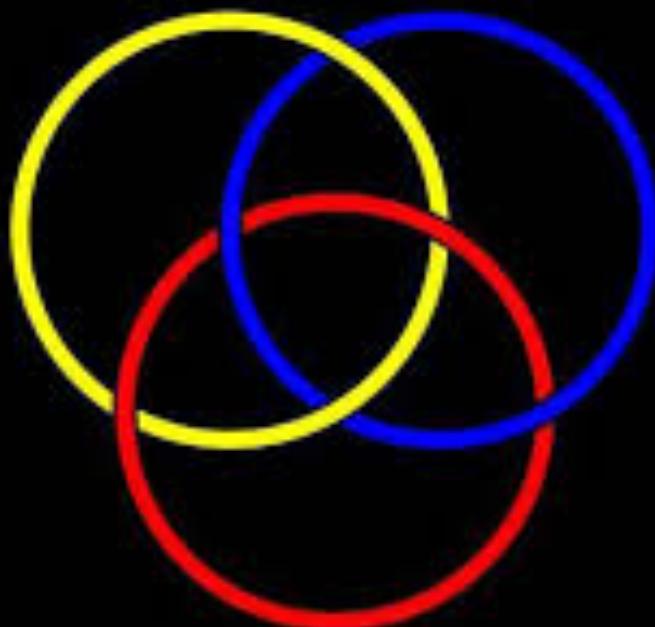


Collaborators:

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LLNL-PRES-652341

This work was performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under contract DE-AC52-07NA27344. Lawrence Livermore National Security, LLC



Outline

- Introduction
- Microscopic three-cluster problem
- Formalism for the $(A-2)+1+1$ mass partition
- Applications to ${}^6\text{He}$
- Conclusions
- Outlook



Our goal is to develop a fundamental theory for the description of thermonuclear reactions and exotic nuclei

Standard solar model

pp chain

$$p + p \rightarrow ^2H + e^+ + \nu_e$$

$$^2H + p \rightarrow ^3He + \gamma$$

$$^3He + ^4He \rightarrow ^7Be + \gamma$$

$$^7Be + e^- \rightarrow ^7Li + \nu_e$$

$$^7Li + p \rightarrow ^4He + ^4He$$

$$^8B \rightarrow ^8Be^* + e^+ + \nu_e$$

$$^8Be^* \rightarrow ^4He + ^4He$$

$$^3He + ^3He \rightarrow ^4He + 2p$$

Stellar Nucleosynthesis

$$^4He + ^4He + ^4He \rightarrow ^{12}C + \gamma$$

$$^{12}C + ^4He \rightarrow ^{16}O + \gamma$$

Fusion Energy Generation

$$d + ^3H \rightarrow ^4He + n$$

$$^3H + ^3H \rightarrow ^4He + 2n$$

$$d + ^3H \rightarrow ^4He + n + \gamma$$

What is the nature of the nuclear force?

One-Neutron Halo

Two-Neutron Halo

$1H$

$2D$

$6He$

$8He$

$$^{10}Ne$$

$$^{11}Li$$

$$^{12}Be$$

$$^{13}B$$

$$^{14}Be$$

$$^{15}C$$

$$^{16}O$$

$$^{17}N$$

$$^{18}O$$

$$^{19}C$$

$$^{*}\alpha + 4n^{**}$$

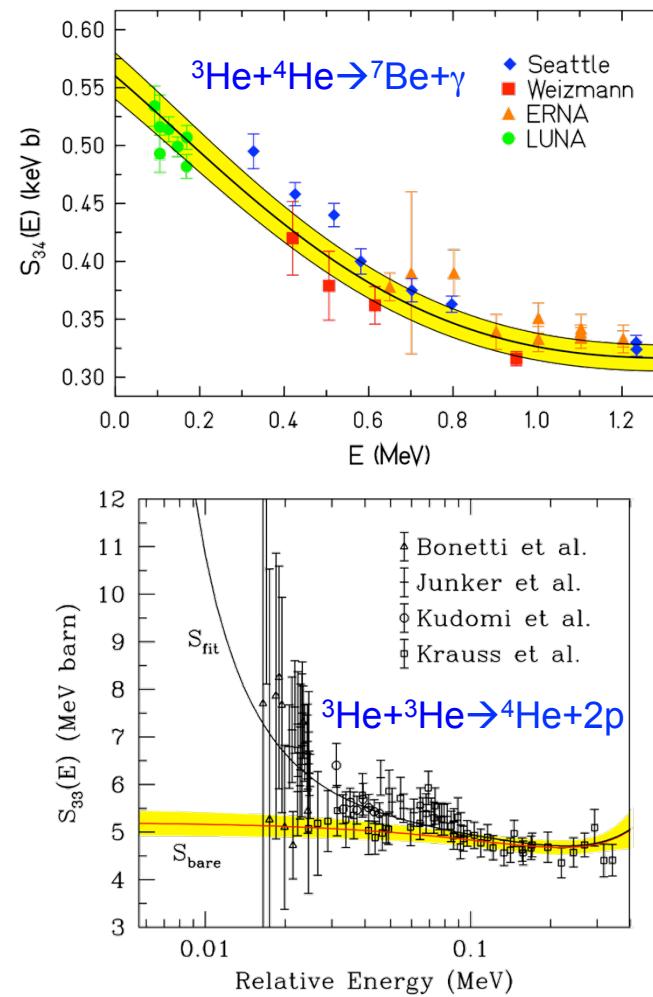
Theory needed because fusion reactions are difficult or impossible to measure at astrophysical energies

- The nuclear fusion process operates mainly by tunneling through the Coulomb barrier

- Extremely low rates

$$\sigma(E) = \frac{S(E)}{E} \exp\left(-\frac{2\pi Z_1 Z_2 e^2}{\hbar \sqrt{2mE}}\right)$$

- Projectiles and targets are not fully ionized
 - Electron screening can mask “bare” nuclear cross section

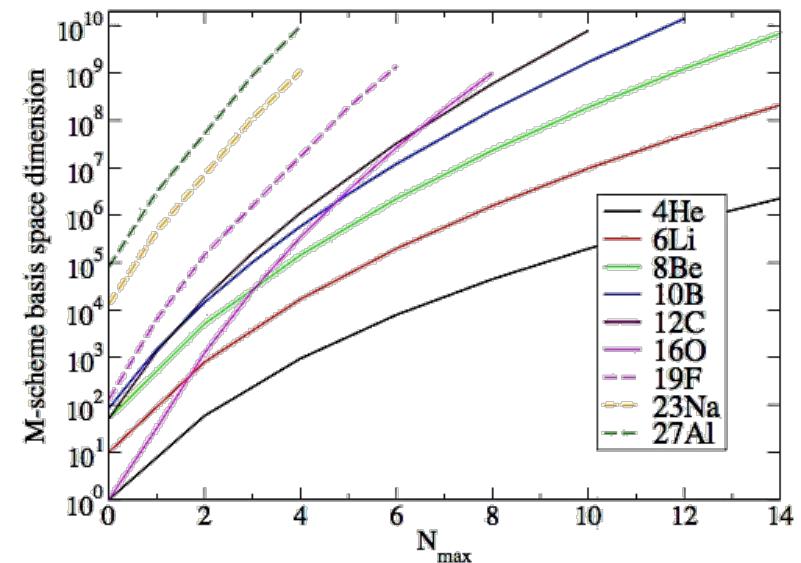
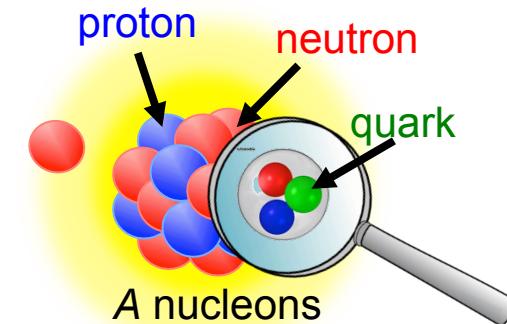


Developing such a fundamental theory is extremely complicated and a longstanding goal of nuclear theory

Ab initio many-body calculations:

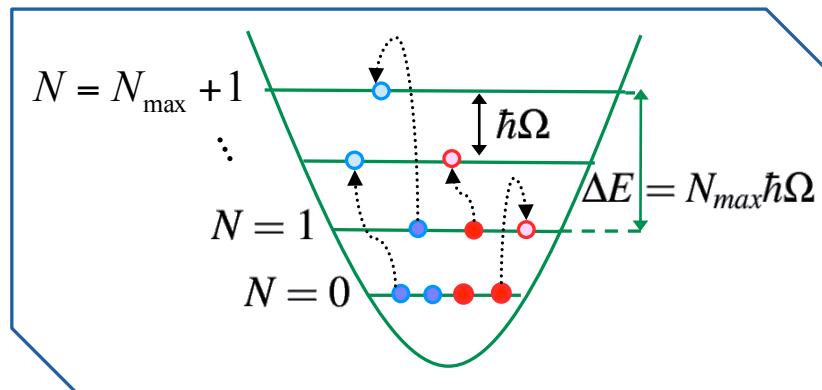
- A (all active) point-like nucleons
- Nuclear two- and three-body (NN+NNN) forces guided by Quantum Chromodynamics (QCD)
- Unitary transformation to soften bare Hamiltonian: e.g., Similarity Renormalization Group (SRG)

Efficient theoretical framework and High Performance Computing (HPC)

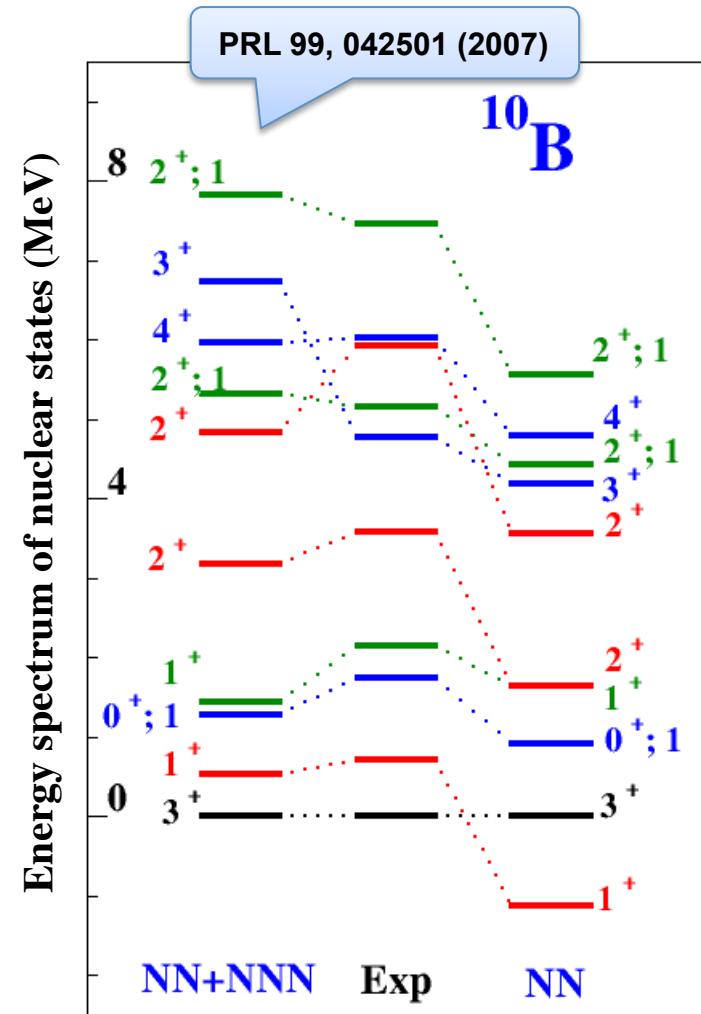


Our starting point is a method to describe static properties of light nuclei from first principles

- *Ab initio* no-core shell model (NCSM) approach

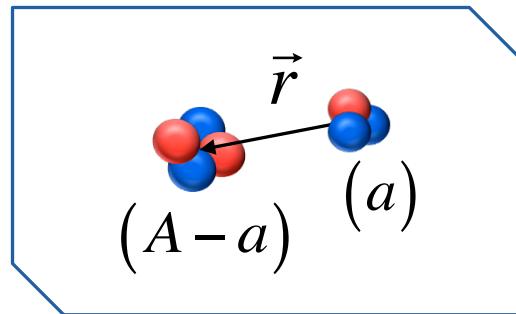


Helped to point out
the fundamental importance
of three-nucleon (NNN) forces
in structure calculations.



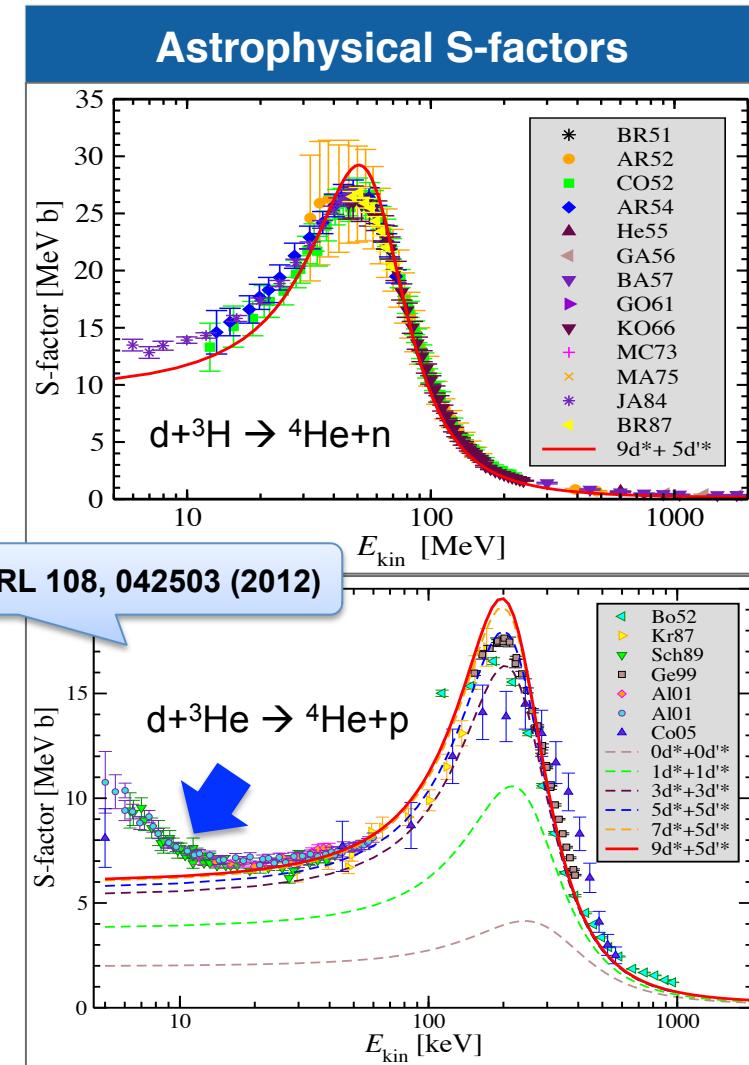
We extended this approach by adding the dynamics between nuclei with the resonating-group method (RGM)

- NCSM/RGM approach



- Ab initio* NCSM wave functions of the nuclei
- NN interactions

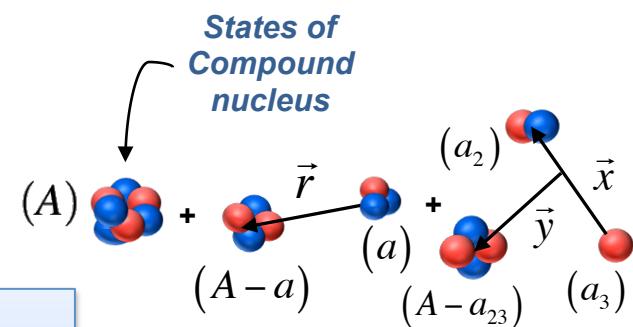
Pioneered *ab initio* calculations
of light-nuclei fusion reactions



We are now working to complete this picture

- Extended NCSM/RGM to include:

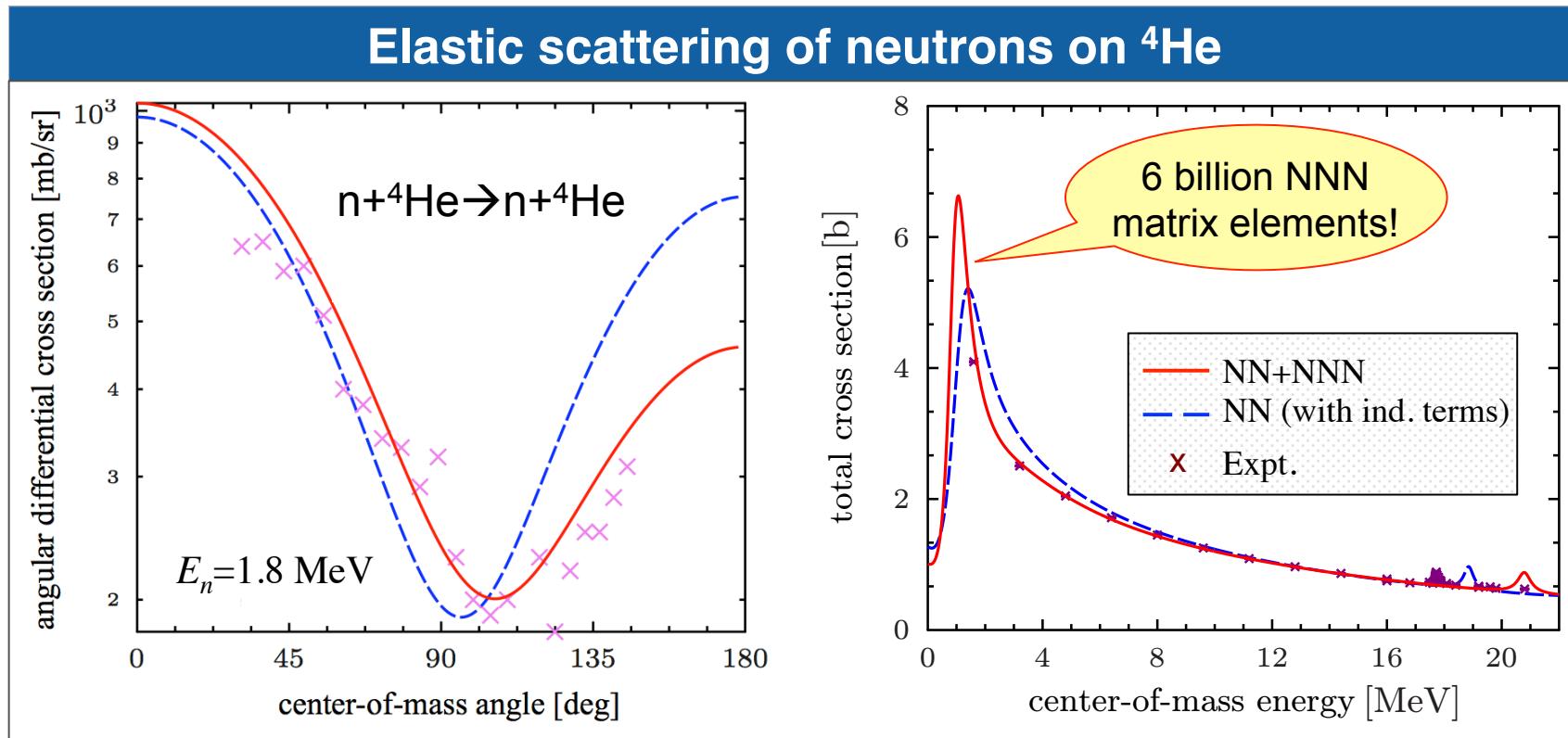
- 1) NNN force in reactions
- 2) States of the compound nucleus
- 3) Three-cluster states in the continuum



This talk

1) Importance of the NNN force in reactions

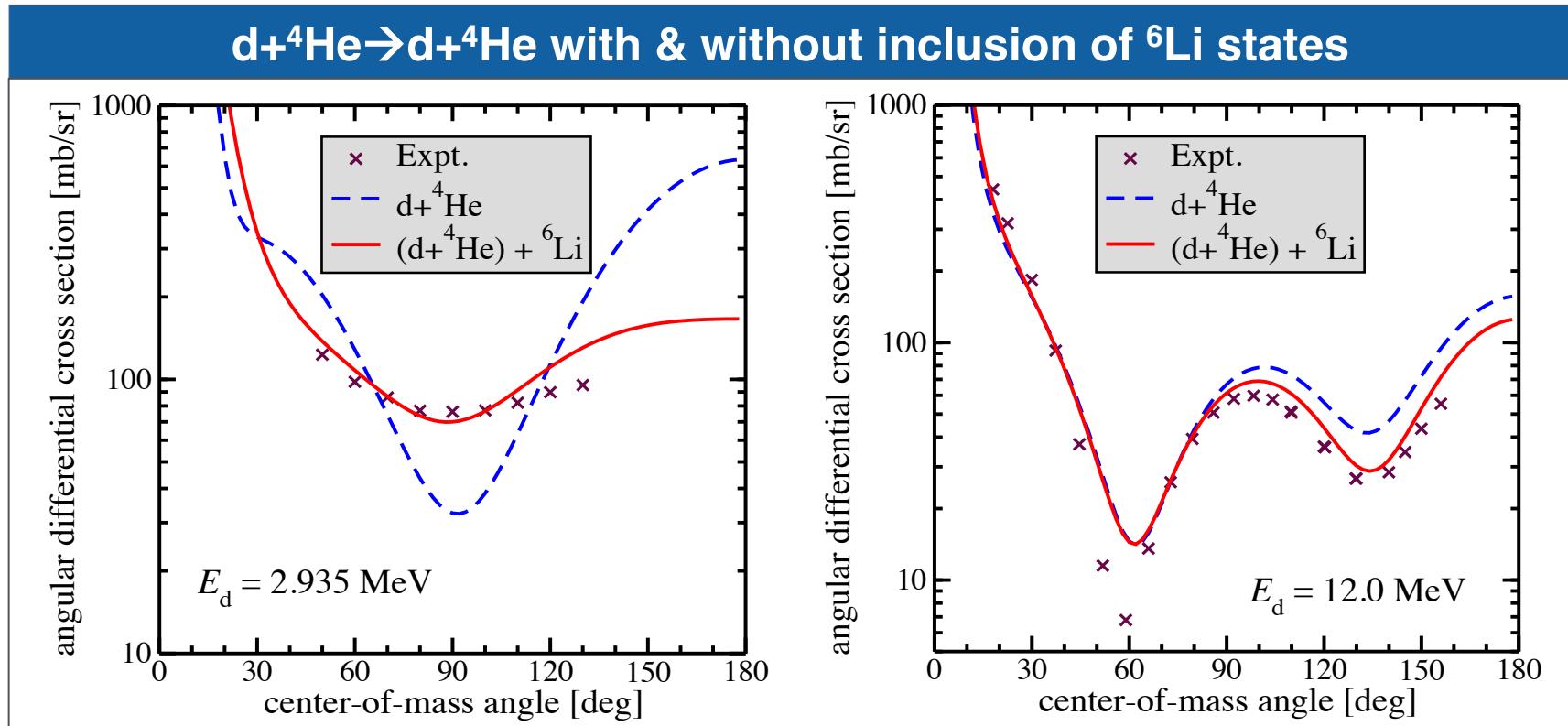
G. Hupin, J. Langhammer, P. Navratil, S. Quaglioni, A. Calci and R. Roth, Phys. Rev. C 88, 054622 (2013)



This work sets the stage for a truly accurate prediction of the $d + {}^3\text{H} \rightarrow {}^4\text{He} + n$ fusion from QCD-based NN+NNN forces

2) Importance of states of the compound system

G. Hupin, S. Quaglioni, and P. Navratil, in progress



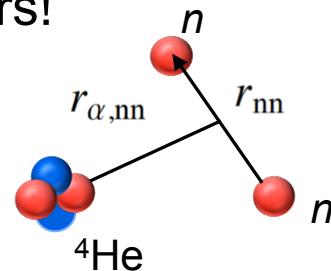
Six-body correlations important also for binding energy (~1 MeV)

3) We want to describe also systems for which the lowest threshold for particle decay is of the 3-body nature

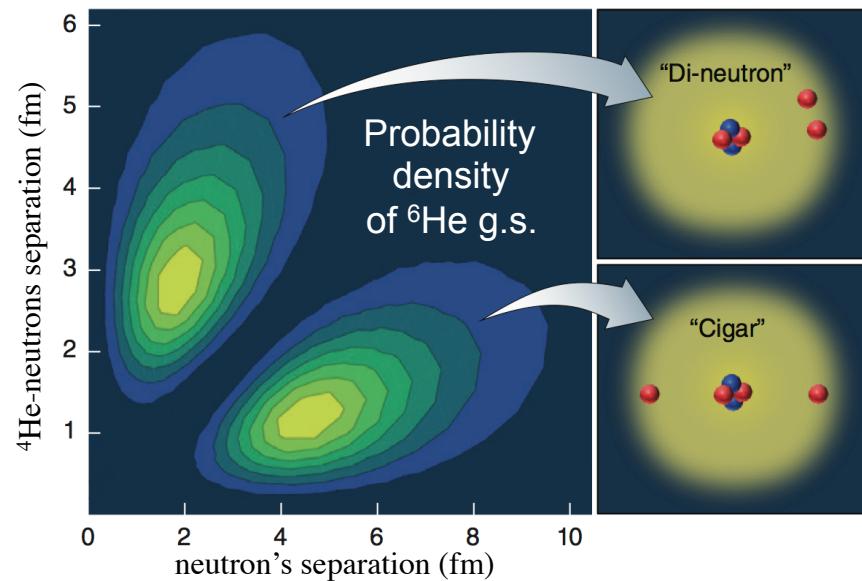
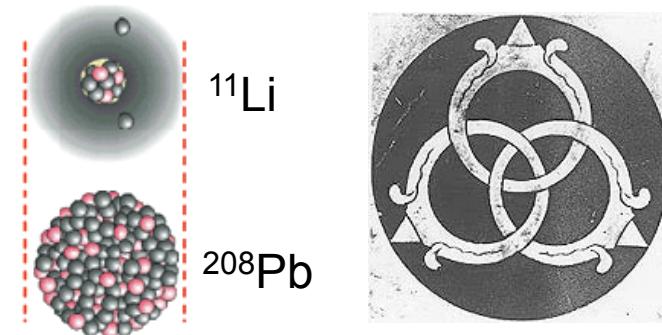
- Exotic nuclei, (Borromean halos, dripline nuclei)

- ${}^6\text{He} (= {}^4\text{He} + n + n)$
- ${}^6\text{Be} (= \alpha + p + p)$
- ${}^{11}\text{Li} (= {}^9\text{Li} + n + n)$
- ${}^{14}\text{Be} (= {}^{12}\text{Be} + n + n)$
- ...

- Constituents do not bind in pairs!



S. Quaglioni, C. Romero-Redondo,
P. Navrátil, Phys. Rev. C 88, 034320 (2013)



Microscopic three-cluster problem

- Starts from:

$$\Psi^{(A)} = \sum_v \iint d\vec{x} d\vec{y} G_v(\vec{x}, \vec{y}) \hat{A}_v |\Phi_{v\vec{x}\vec{y}}\rangle$$

3-body channels

$$\psi_{\alpha_1}^{(A-a_{23})} \psi_{\alpha_2}^{(a_2)} \psi_{\alpha_3}^{(a_3)} \delta(\vec{x} - \vec{\eta}_{23}) \delta(\vec{y} - \vec{\eta}_{1,23})$$

- Projects $(H - E)\Psi^{(A)} = 0$ onto the channel basis:

$$\sum_v \iint d\vec{x} d\vec{y} \left[H_{v'v}(\vec{x}', \vec{y}', \vec{x}, \vec{y}) - E N_{v'v}(\vec{x}', \vec{y}', \vec{x}, \vec{y}) \right] G_v(\vec{x}, \vec{y}) = 0$$

$$\langle \Phi_{v'\vec{x}'\vec{y}'} | \hat{A}_{v'} H \hat{A}_v | \Phi_{v\vec{x}\vec{y}} \rangle$$

Hamiltonian kernel

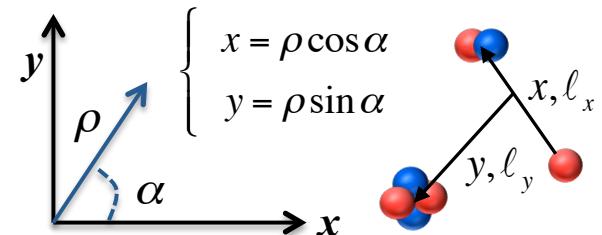
$$\langle \Phi_{v'\vec{x}'\vec{y}'} | \hat{A}_{v'} \hat{A}_v | \Phi_{v\vec{x}\vec{y}} \rangle$$

Norm or Overlap kernel



This can be turned into a set of coupled-channels Schrödinger equations for the hyperradial motion

- Hyperspherical Harmonic (HH) functions form a natural basis:



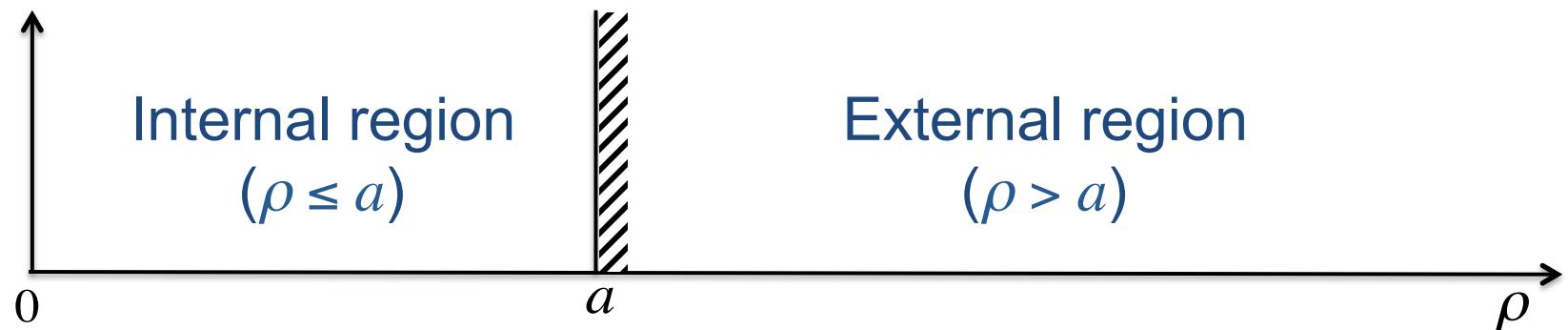
$$|\Phi_{v\vec{x}\vec{y}}\rangle = \sum_K \phi_K^{*\ell_x, \ell_y}(\alpha) |\Phi_{vK\rho}\rangle \quad \text{and} \quad \propto Y_L^{K\ell_x \ell_y}(\alpha_\eta, \hat{\eta}_{1,23}, \hat{\eta}_{23}) \frac{\delta(\rho - \rho_\eta)}{\rho^{5/2} \rho_\eta^{5/2}}$$

- Then, with orthogonalization and projection over $\phi_{K'}^{\ell'_x, \ell'_y}(\alpha')$:

$$\sum_{vK} \int d\rho \rho^5 [N^{-1/2} H N^{-1/2}]_{vv}(\rho', \rho) \frac{u_{Kv}(\rho)}{\rho^{5/2}} = E \frac{u_{K'v'}(\rho')}{\rho'^{5/2}}$$

$$[N^{1/2} G]_v(x, y) = \rho^{-5/2} \sum_K u_{vK}(\rho) \phi_K^{\ell_x \ell_y}(\alpha)$$

These equations can be solved using R-matrix theory



Expansion on a basis

$$u_{Kv}(\rho) = \sum_n c_n^{Kv} f_n(\rho)$$

Bound state asymptotic behavior

$$u_{Kv}(\rho) = C_{Kv} \sqrt{k\rho} K_{K+2}(k\rho)$$

Scattering state asymptotic behavior

$$u_{Kv}(\rho) = A_{Kv} \left[H_K^-(k\rho) \delta_{vv'} \delta_{KK'} - S_{vK,v'K'} H_K^+(k\rho) \right]$$



$^4\text{He} + \text{n} + \text{n}$ within the NCSM/RGM

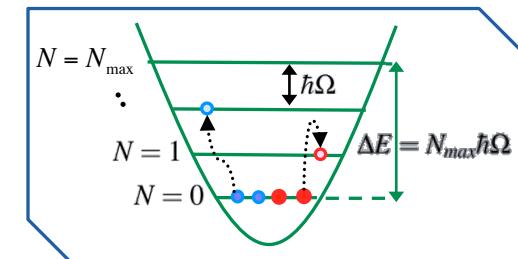
S. Quaglioni, C. Romero-Redondo, P. Navratil, Phys. Rev. C 88, 034320 (2013)

- Accurate soft NN interaction: SRG-evolved chiral N³LO potential with $\Lambda=1.5$ fm⁻¹
 - Fits NN data with high accuracy
 - But: misses both **chiral initial** and **SRG-induced** NNN force
 - Fortuitously**: two effects mostly compensate each other for very light systems

- ^4He ab initio wave function obtained within the NCSM

$$H^{(A-2)} \psi_{\beta_1}^{(A-2)}(\vec{r}_1, \vec{r}_2, \dots, \vec{r}_{A-2}) = E_{\beta_1}^{(A-2)} \psi_{\beta_1}^{(A-2)}(\vec{r}_1, \vec{r}_2, \dots, \vec{r}_{A-2})$$

- Large expansions in A -body **harmonic oscillator (HO)** basis
- Preserves: 1) Pauli principle, and 2) translational invariance
- Can include NNN interactions
- ^4He binding energy close to experiment: 28.22 MeV (expt.: 28.3 MeV)



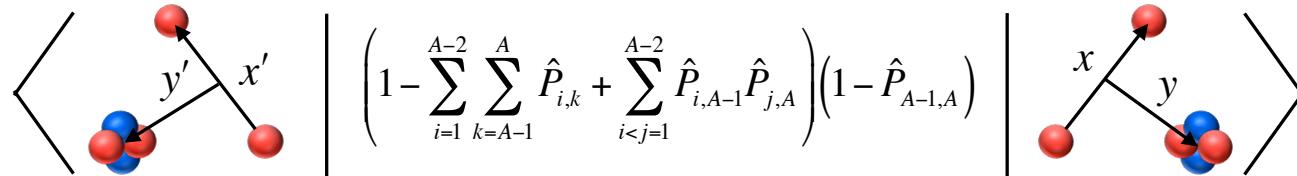
- Fully antisymmetric channel states:

$$\hat{A}_v = \sqrt{\frac{(A-2)!2!}{A!}} \left[1 - \sum_{i=1}^{A-2} \sum_{k=A-1}^A \hat{P}_{i,k} + \sum_{i < j=1}^{A-2} \hat{P}_{i,A-1} \hat{P}_{j,A} \right] \frac{1 - \hat{P}_{A-1,A}}{\sqrt{2}}$$

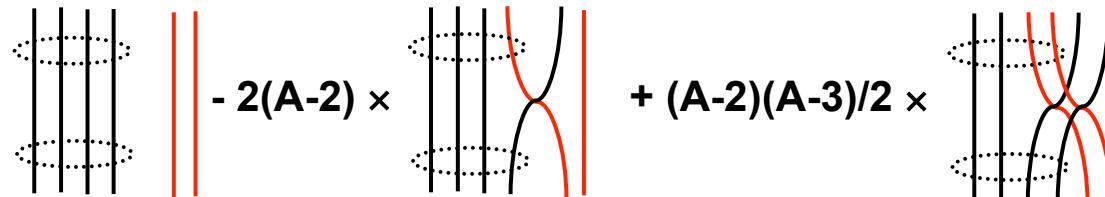


The formalism is general for (A-2)+1+1 mass partitions

Norm or overlap kernel (Pauli principle)



$$N_{\nu\nu}(x',y',x,y) = \frac{1}{2} \left[1 - (-1)^{\ell'_x + S'_{23} + T'_{23}} \right] \left[1 - (-1)^{\ell_x + S_{23} + T_{23}} \right] \times \left\{ \delta_{\nu\nu} \frac{\delta(x' - x)}{x'x} \frac{\delta(y' - y)}{y'y} \right. \\ - 2(A-2) \sum_{n'_x n'_y} \sum_{n_x n_y} R_{n'_x \ell'_x}(x') R_{n'_y \ell'_y}(y') \langle \Phi_{\nu'_3 n'_x n'_y} | P_{A-2,A} | \Phi_{\nu_3 n_x n_y} \rangle R_{n_x \ell_x}(x) R_{n_y \ell_y}(y) \\ \left. + \frac{(A-2)(A-3)}{2} \sum_{n'_x n'_y} \sum_{n_x n_y} R_{n'_x \ell'_x}(x') R_{n'_y \ell'_y}(y') \langle \Phi_{\nu'_3 n'_x n'_y} | P_{A-3,A-1} P_{A-2,A} | \Phi_{\nu_3 n_x n_y} \rangle R_{n_x \ell_x}(x) R_{n_y \ell_y}(y) \right\}$$

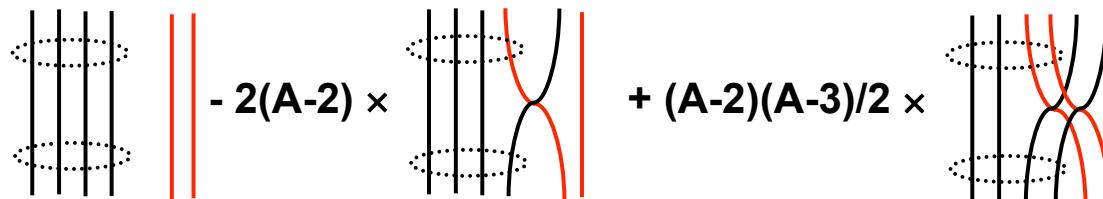
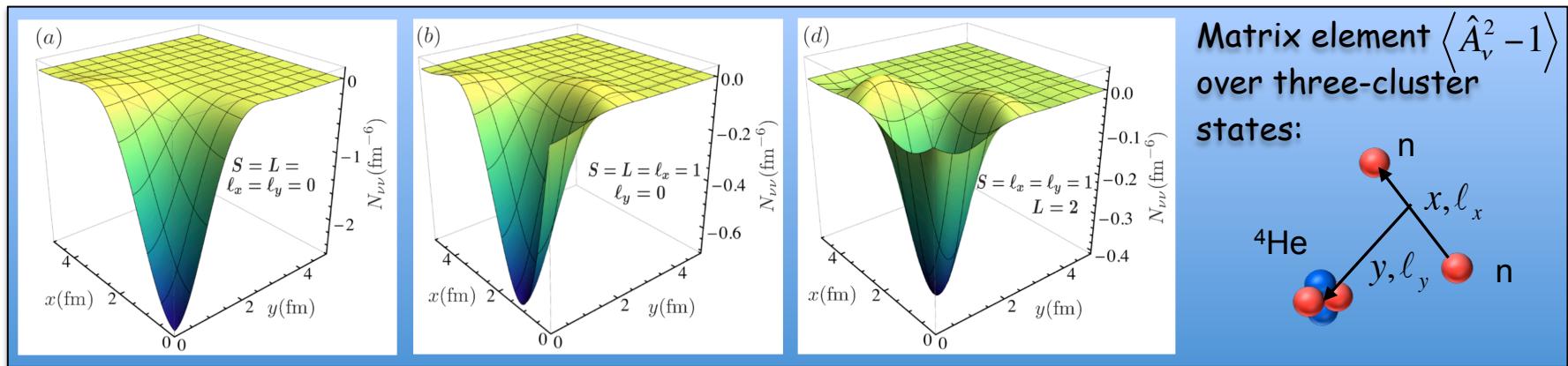
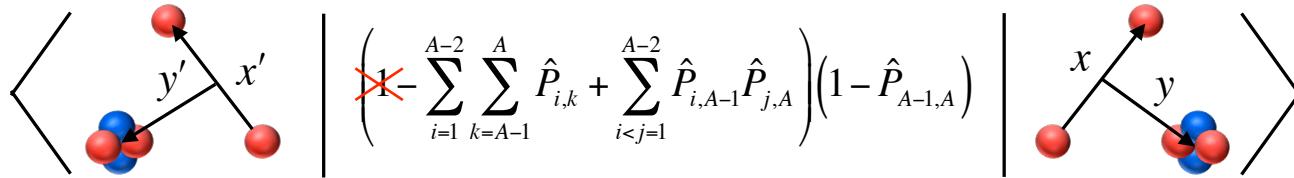


$${}_{SD} \langle \psi_{\mu_1}^{(A-2)} | a^+ a | \psi_{\nu_1}^{(A-2)} \rangle {}_{SD}$$

$${}_{SD} \langle \psi_{\mu_1}^{(A-2)} | a^+ a^+ a a | \psi_{\nu_1}^{(A-2)} \rangle {}_{SD}$$

The formalism is general for (A-2)+1+1 mass partitions

Norm or overlap kernel (Pauli principle)

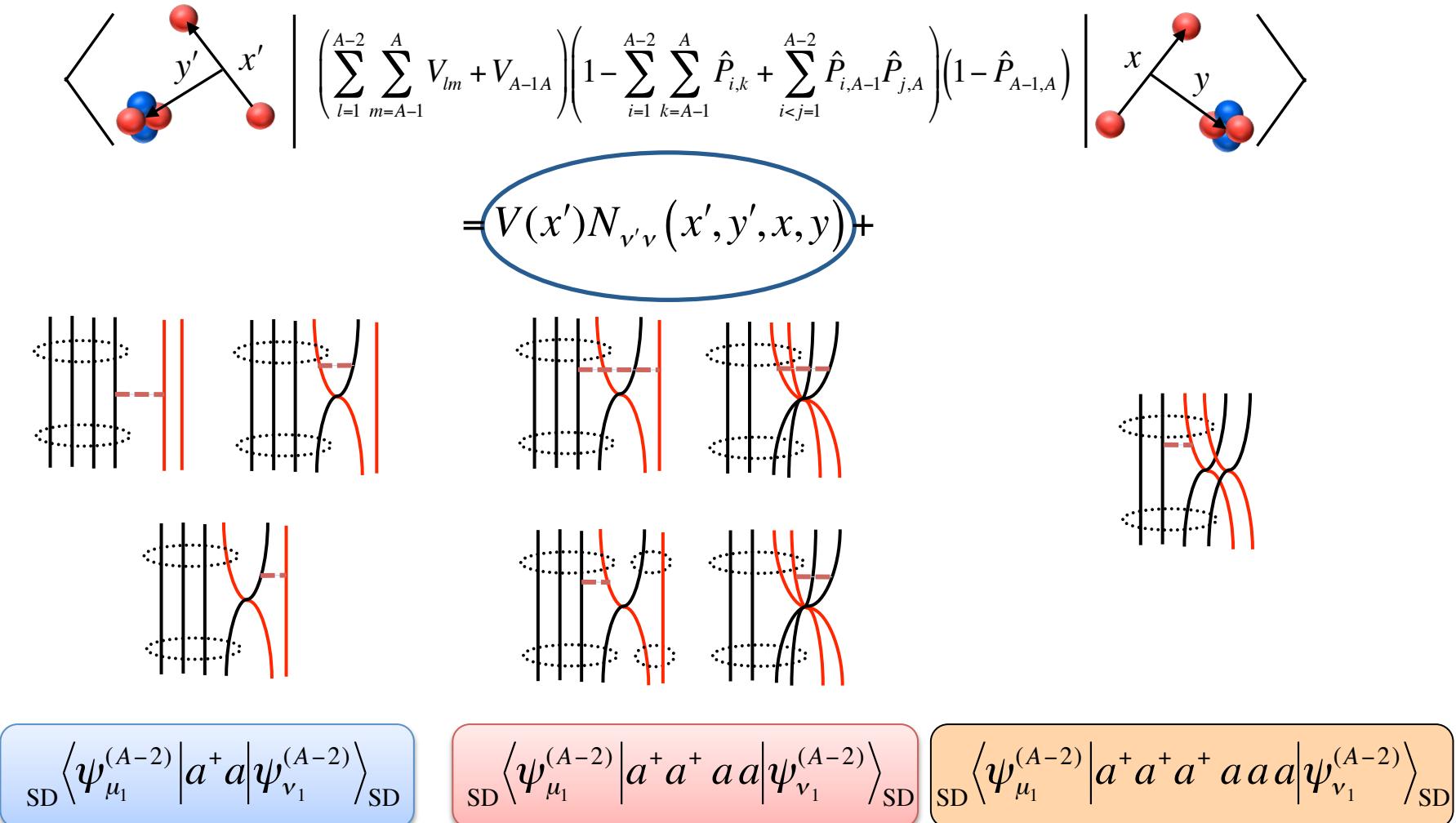


$${}_{\text{SD}} \left\langle \psi_{\mu_1}^{(A-2)} \left| a^+ a \right| \psi_{\nu_1}^{(A-2)} \right\rangle_{\text{SD}}$$

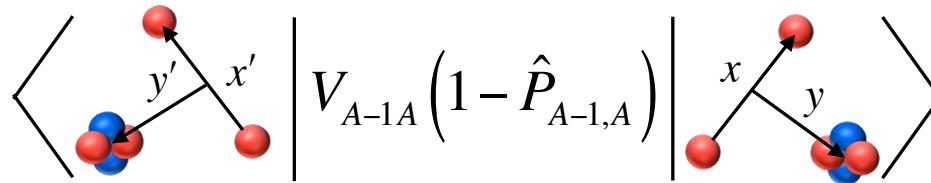
$${}_{\text{SD}} \left\langle \psi_{\mu_1}^{(A-2)} \left| a^+ a^+ a a \right| \psi_{\nu_1}^{(A-2)} \right\rangle_{\text{SD}}$$

The formalism is general for (A-2)+1+1 mass partitions

Hamiltonian kernel (nucleon-nucleon-target potentials)



Part of the interaction kernel is localized only in x, x'



$$\propto \sum_{n'_x n_x} R_{n'_x \ell'_x}(x') R_{n_x L_x}(x) \langle n'_x \ell'_x s_{23} J_{23} T_{23} | V | n_x L_x s_{23} J_{23} T_{23} \rangle$$

$$\times \left(1 - (-1)^{\ell_x + s_{23} + T_{23}} \right) \delta_{\tilde{\gamma}' \gamma} \frac{\delta(y' - y)}{y' y}$$

**Extended-size
HO expansion**
 $N_{\text{ext}} \gg N_{\text{max}}$

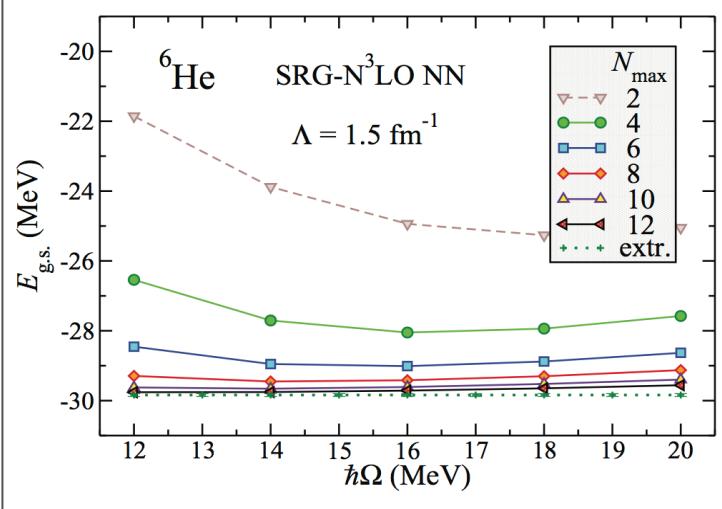
$$\approx \sum_{n_y} R_{n_y \ell_y}(y') R_{n_y \ell_y}(y)$$

Results for ${}^6\text{He}$ ground state

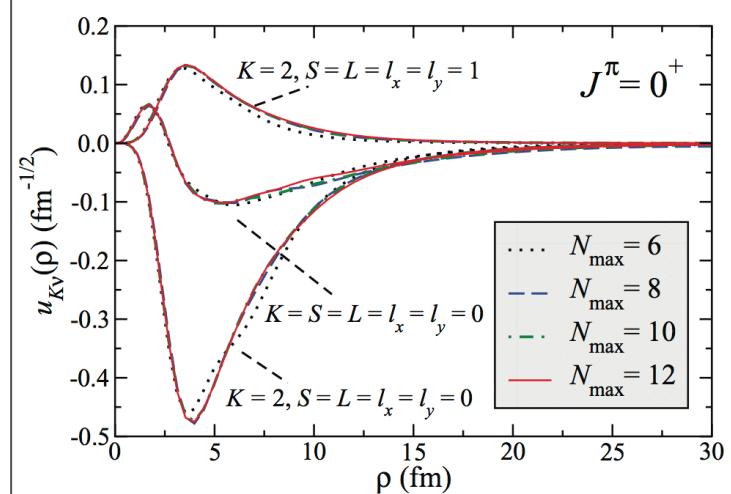
6-body diagonalization vs ${}^4\text{He}(\text{g.s.})+\text{n+n}$ calculation

$$\chi_\nu(x, y) = \frac{1}{\rho^{5/2}} \sum_K u_{\nu K}(\rho) \phi_K^{\ell_x \ell_y}(\alpha)$$

NCSM 6-body diagonalization



NCSM/RGM ${}^4\text{He}(\text{g.s.})+\text{n+n}$

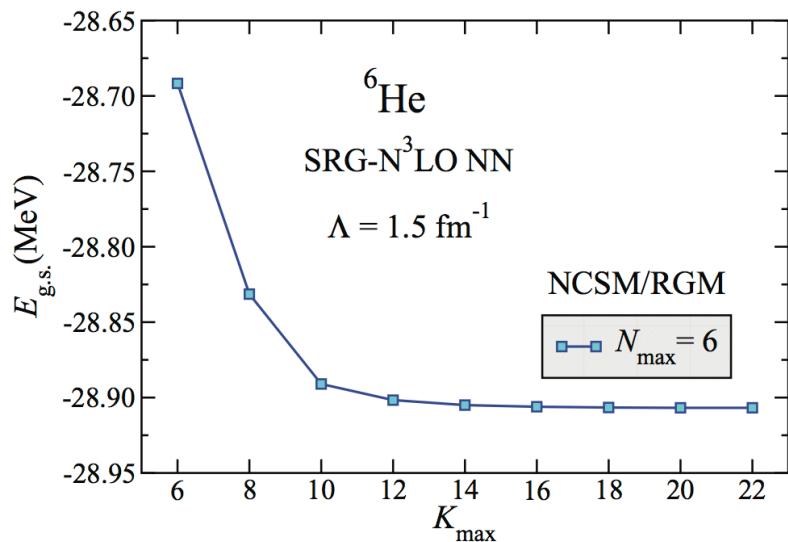


$N_{\text{tot}} = N_0 + N_{\text{max}}$	${}^4\text{He}$ NCSM	${}^6\text{He}$ NCSM/RGM	${}^6\text{He}$ NCSM
6	-27.984	-28.907	-27.705
8	-28.173	-28.616	-28.952
10	-28.215	-28.696	-29.452
12	-28.224	-28.697	-29.658
Extrapolation	-28.230(5)	-28.70(3)	-29.84(4)
Experimental	-28.296		-29.268

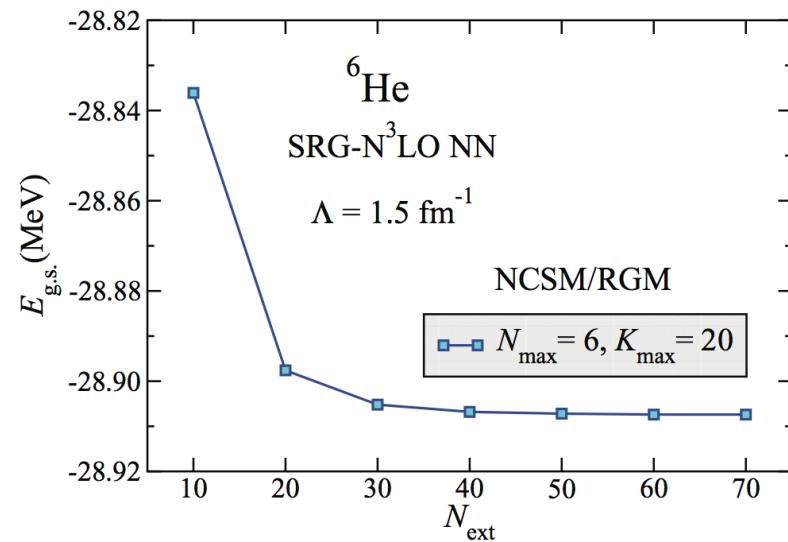
- Differences between NCSM 6-body and NCSM/RGM ${}^4\text{He}(\text{g.s.})+\text{n+n}$ results due to core polarization
- Contrary to NCSM, NCSM/RGM wave function has appropriate asymptotic behavior

Other convergence tests

- HH expansion



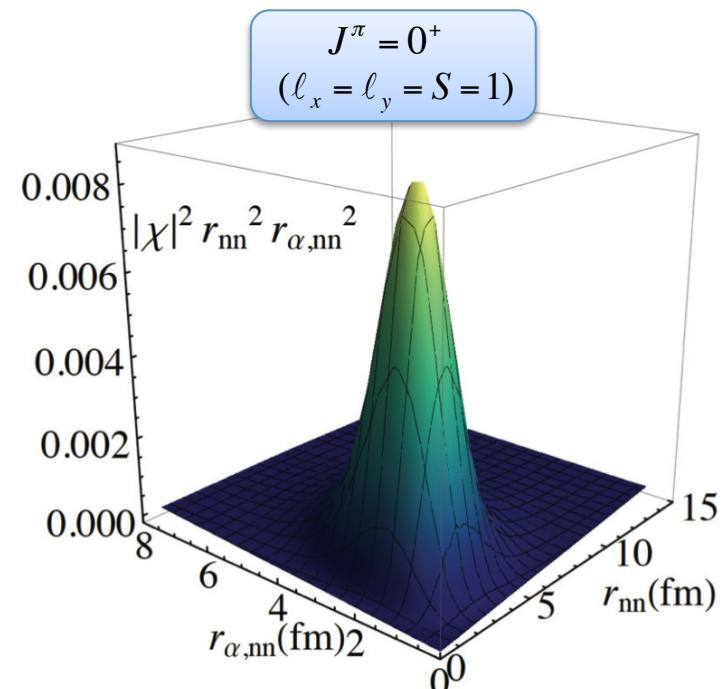
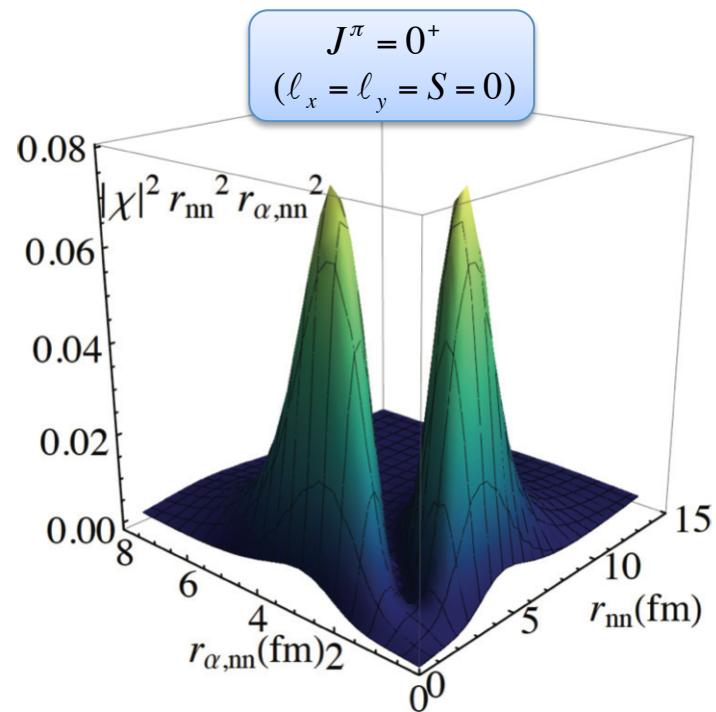
- Extended-size HO expansion



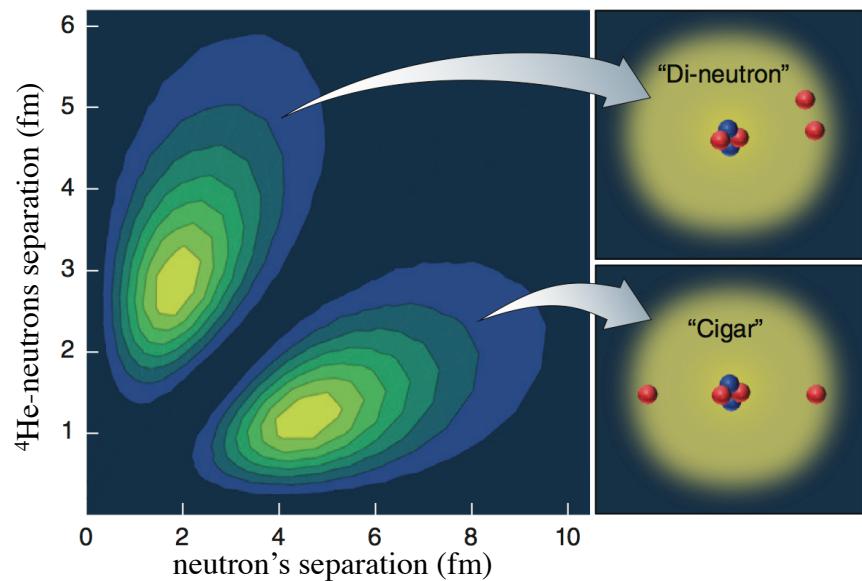
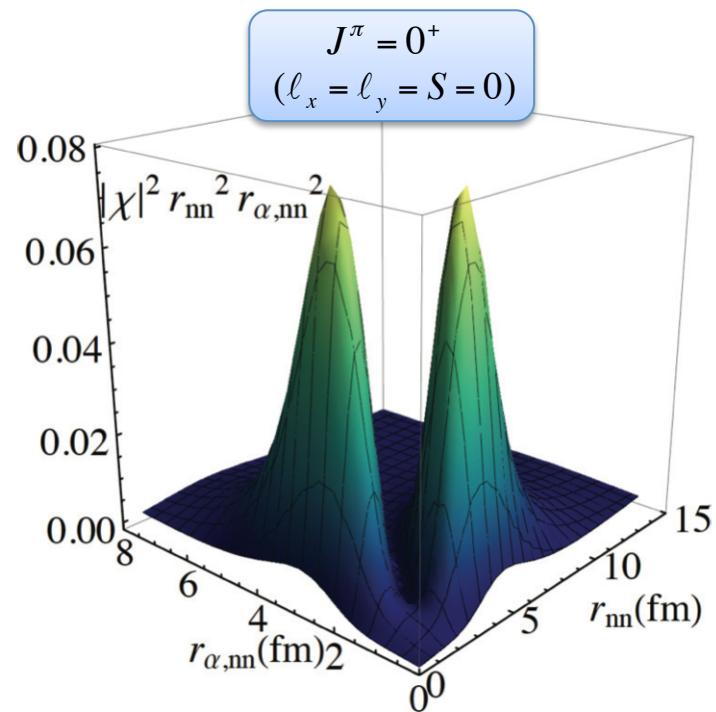
$$\chi_\nu(x, y) = \frac{1}{\rho^{5/2}} \sum_K^{K_{\text{max}}} u_{\nu K}(\rho) \phi_K^{\ell_x \ell_y}(\alpha)$$

$$\left\langle V_{A-1A} \left(1 - \hat{P}_{A-1,A} \right) \right\rangle \propto \sum_{n_y} R_{n_y \ell_y}(y') R_{n_y \ell_y}(y)$$

Probability density of ${}^6\text{He}$ ground state

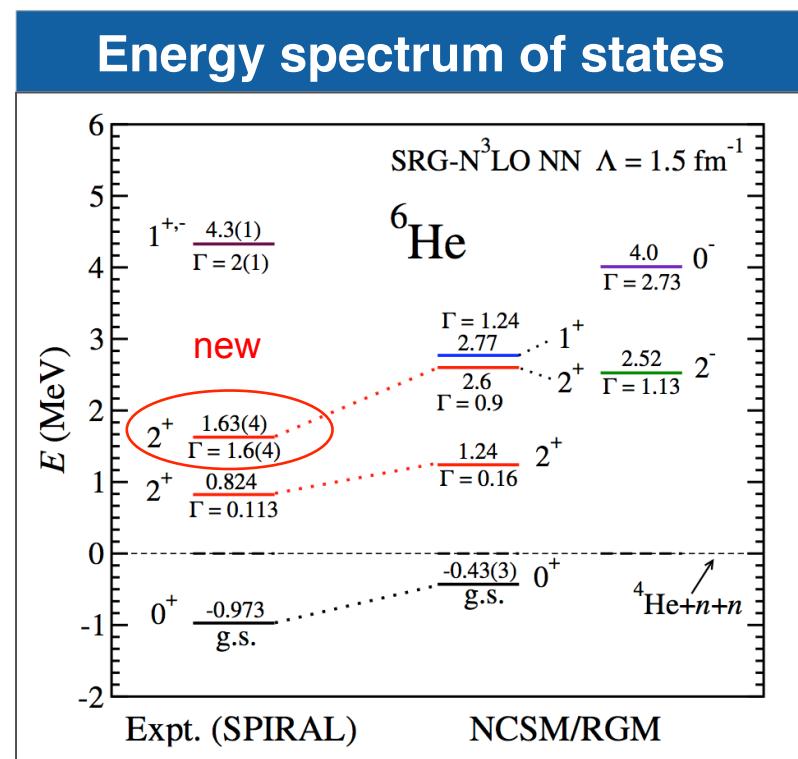
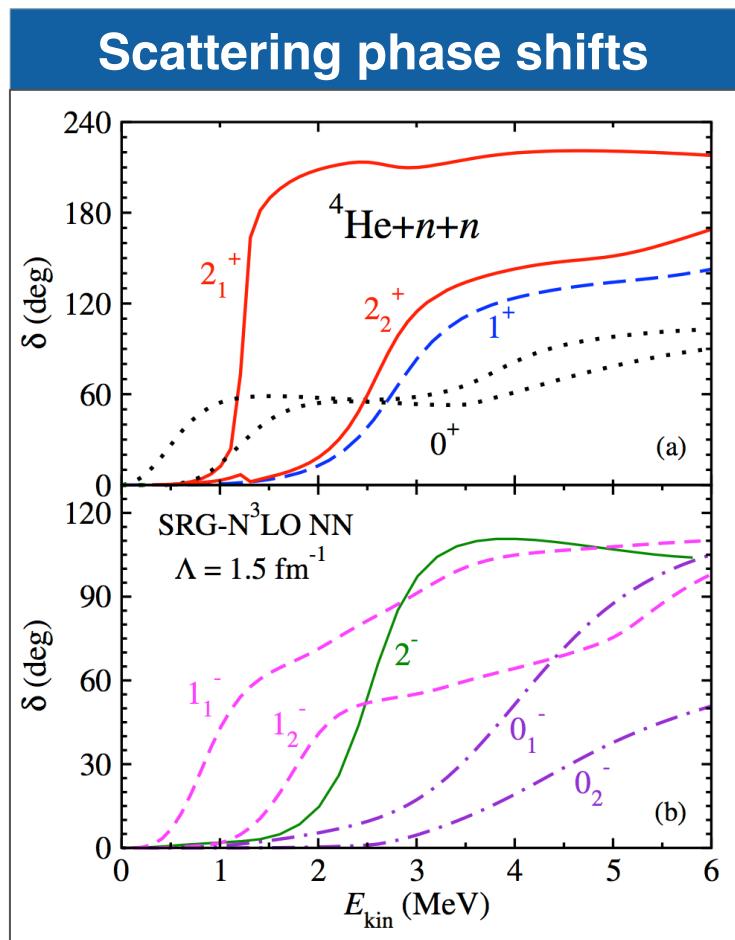


Probability density of ${}^6\text{He}$ ground state



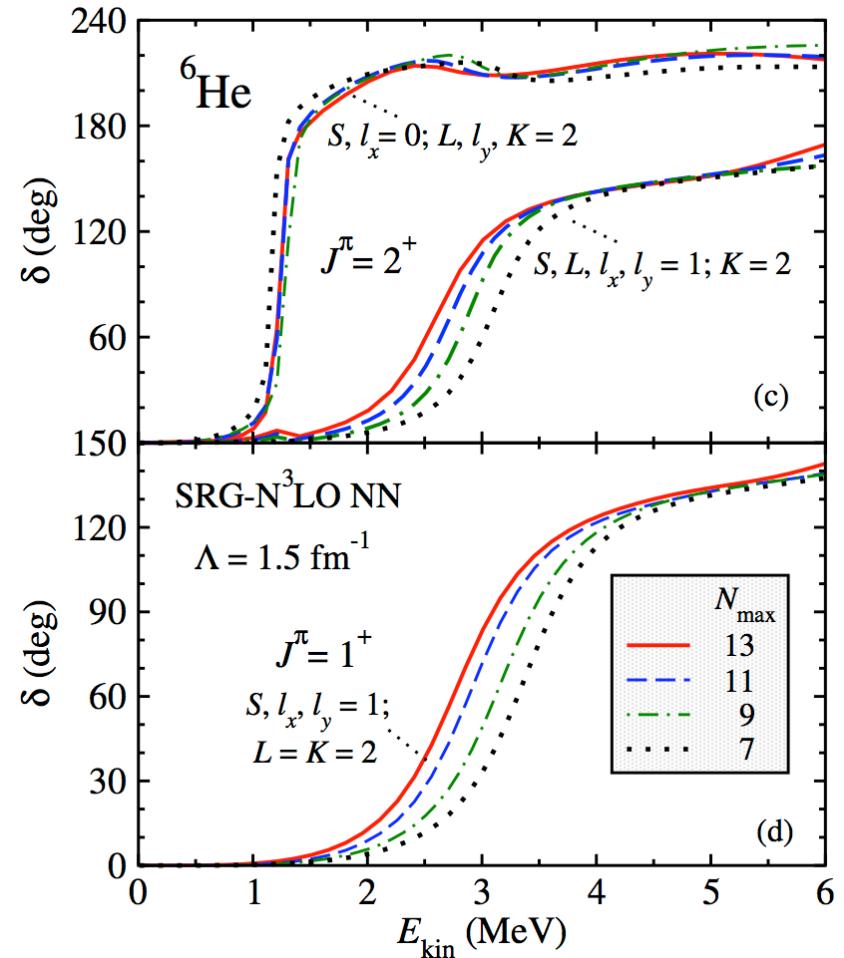
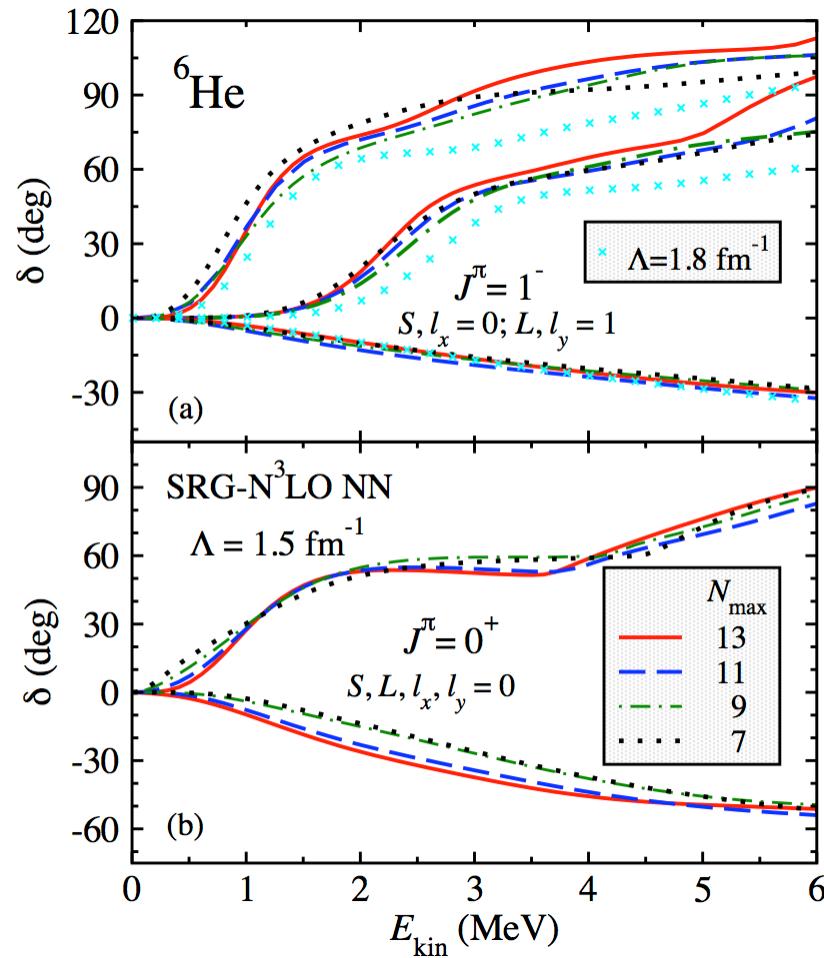
Results for ${}^4\text{He}(\text{g.s.}) + \text{n} + \text{n}$ continuum

C. Romero-Redondo, S. Quaglioni, and P. Navratil, in progress

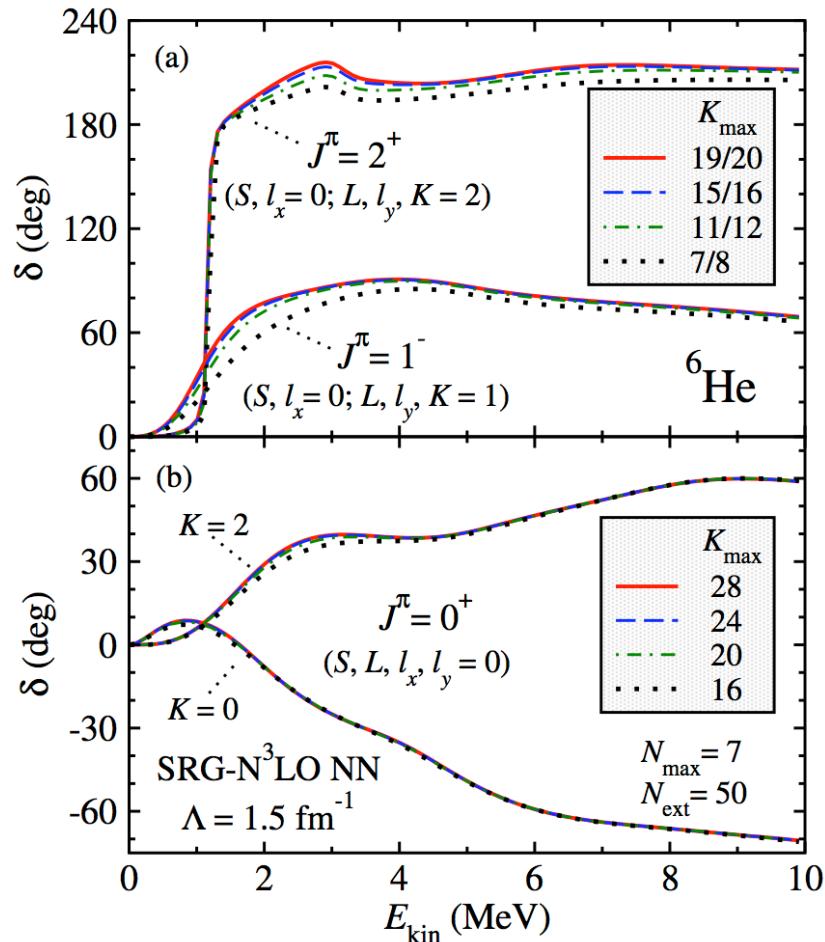


$$\Gamma = \frac{2}{\left. \frac{d\delta(E)}{dE} \right|_{E=E_R}}$$

Convergence with respect to HO model space size



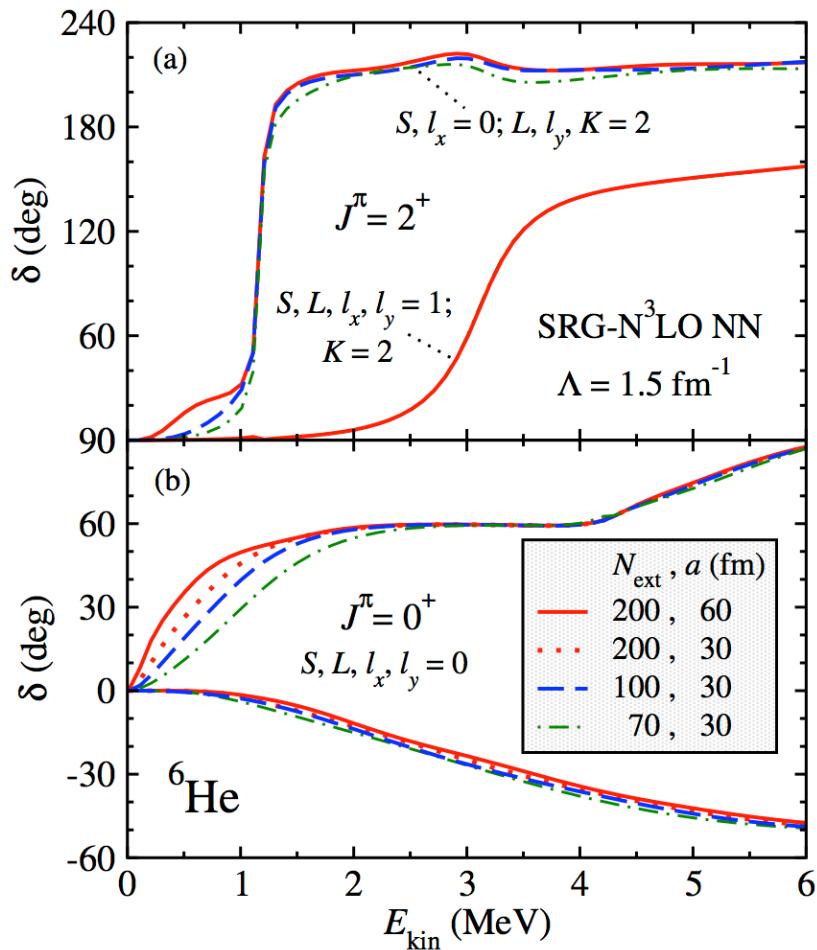
Other convergence tests



- HH expansion

$$\chi_\nu(x, y) = \frac{1}{\rho^{5/2}} \sum_K^{K_{\max}} u_{\nu K}(\rho) \phi_K^{\ell_x \ell_y}(\alpha)$$

Other convergence tests



- Extended-size HO expansion

$$\left\langle V_{A-1A} \left(1 - \hat{P}_{A-1,A} \right) \right\rangle \propto \sum_{n_y}^{N_{\text{ext}}} R_{n_y \ell_y}(y') R_{n_y \ell_y}(y)$$

- Sizable effects only when neutrons are in 1S_0 partial wave (strong attraction)

Conclusions

- We are building an efficient ab initio theory including the continuum
 - NCSM eigenstates → short- to medium-range A-body structure
 - NCSM/RGM cluster states → scattering physics of the system
- We map the many-body problem into a few-cluster problem
 - The Pauli exclusion principle is treated exactly
 - Inter-cluster interactions arise from underlying nuclear Hamiltonian
- First ab initio description of three-cluster dynamics
 - ${}^4\text{He} + \text{n} + \text{n}$ bound and continuum states
 - Good qualitative description of the low-lying spectrum of ${}^6\text{He}$



Outlook

- For a complete picture we need to:
 - Run calculations with NNN forces (codes are ready)
 - Introduce core excitations by coupling to NCSM A-body eigenstates
- Future applications of three-cluster formalism
 - Calculations of radii, electric dipole transitions
 - Other systems: ${}^5\text{H}$ (${}^3\text{H}+\text{n}+\text{n}$), ${}^{11}\text{Li}$ ($={}^9\text{Li}+\text{n}+\text{n}$), ${}^{12}\text{Be}$ ($={}^{10}\text{Be}+\text{n}+\text{n}$)
- Ultimate goal: binary & ternary light-nucleus fusion reactions
 - Transfer reactions: ${}^3\text{H}({}^3\text{H}, 2\text{n}) {}^4\text{He}$