



Johannes Hecker Denschlag Seattle, April 15, 2014







# ... another good thing from Ulm

(besides Einstein)







#### Trapped lons and Ultracold neutral Atoms



Good compatibility of traps!



#### Long range atom – ion interaction





#### An ion in a cloud of atoms, naive picture



- Thermalization of ion within a few collisions, sympathetic cooling
- No further dynamics afterwards....





# 0

#### Can atoms sympathetically cool hot ions?



Displaced ion



High potential energy





## The role of excess micromotion



# Elastic two-body atom-ion collisions

Thermal atom cloud T ~ 100nK





## Reactions

#### How do we detect a reaction?

• Ion turns dark, changes mass



- Release of energy
  - ion orbits outside of atoms
  - change in atom loss





## Ion-induced atom loss Rb<sup>+</sup> in cold Rb cloud



Interaction time  $\tau$  [s]

#### **Atom number distributions**



A. Härter et al. PRL 2012





#### **Atom-atom-ion three-body recombination**



#### **Measurement of the reaction energy**



## Data well described by three-body recombination dynamics



quadratic density dependence  $\rightarrow$  atom-atom-ion three-body coefficient

A. Härter et al. PRL 2012

## **Energy dependence**



Similar as for Ba\* + 2 Rb!

Chris Green and Jesus Perez-Rios can theoretically confirm this dependence! Calculations by Chris Greene and Jesus Perez-Rios (14. 5. 2014)

- Classical trajectory
- Monte-Carlo
- Heuristic argument for  $E^{-4/7} = E^{-5.7}$ dependence





#### **Reaction: Ba+ with Rb cloud**





Typical reaction: charge exchange



#### **Decay of Ba+ in atomic clouds of various densities**



#### Charge exchange in three-body process!





VS



Langevin collision: The rate  $\Gamma$ should be energy independent!

#### Energy dependence of three-body charge exchange



# Investigate reaction product states



Question: What quantum state is the molecule in?



#### **Three-body recombination**



#### Three-body recombination away from Feshbach resonance



# The set-up

 $4 \times 10^{4}$  <sup>87</sup>Rb atoms in an optical dipole trap at 1064nm; ~1 $\mu$ K temperature; density ~ 10<sup>13</sup> cm<sup>-3</sup>;

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> Stateselectively ionize molecule!

electron -

# Rb<sub>2</sub><sup>+</sup> ion is trapped!



### **Detecting dark ions**



#### **Use resonance enhanced multi-photon ionization!**







#### Plenty of resonances!



- very dense and fairly irregular spectrum (> 100 lines)
- distribution over many initial states
- selection rules: each level only gives rise to two or three lines

A. Härter, A. Krükow, M. Deiß, B. Drews, E. Tiemann, and J. Hecker Denschlag, Nature Physics (2013)



#### Study the line shape!



#### First assignment of rotational line spectrum





#### First assignment of rotational line spectrum





- Understand reaction pathways in all details
- Test theoretical models/ predictions for three-body recombination



A. Härter, A. Krükow, M. Deiß, B. Drews, E. Tiemann, and J. Hecker Denschlag, Nature Physics (2013)

# **Collisions of ultracold Rb<sub>2</sub> molecules**



Björn Drews Markus Deiss

Krzysztof Jachymski Tommaso Calarco Zbigniew Idziaszek

# **Collisions of ultracold Rb<sub>2</sub> molecules**



 $a^{3}\Sigma_{u}$ 

- -Vibrational ground state
- Molecule rotation R =0 or R = 2
- Precisely defined quantum state: R, I, F, J, m<sub>F</sub>, ...
- Quasi 1D trap ground state in transverse direction (~100 E<sub>r</sub>)
- Longitudinal energy
  ~ 100nK k<sub>B</sub>

# **Decay of Rb<sub>2</sub> molecules**



Data compatible with universal collisions?

$$a_{3D} = \bar{a}(1-i)$$
  
 $\bar{a} = 2\pi/\Gamma(1/4)^2 R_6$   
 $R_6 = (2\mu C_6/\hbar)^{1/4} \approx 270a_6$ 



## Four stories



4) Cold collisions of Rb<sub>2</sub> triplet molecules  $2 \operatorname{Rb}_2 \rightarrow \operatorname{loss}$ 

