Interaction *vs* correlation effects in many-body systems

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<span id="page-0-0"></span>Workshop on Nuclear Structure and Dynamics at Short Distances INT, Seattle, February 13, 2013

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 $\star$  Defining correlations: quite an elusive issue

- $\triangleright$  correlations in the absence of interaction
- $\triangleright$  interaction without correlations
- $\star$  Theoretical description of correlations
	- $\triangleright$  particle and hole propagation in interacting many-body systems
- $\star$  Empirical evidence of nucleon-nucleon correlations
	- $\triangleright$  nucleon knockout processes
	- $\triangleright$  Final State Interactions (FSI) in inclusive processes
	- $\triangleright$  the EMC effect

 $\star$  Summary & Outlook

### Defining correlations

- $\star$  Consider a system of N interacting particle described by the wave function  $\Psi(x_1, \ldots, x_N)$ , with  $x_i \equiv (\mathbf{r}_i, \sigma_i)$
- $\star$  Probability of finding particles  $1, \ldots, n$  at positions  $\mathbf{r}_1, \ldots, \mathbf{r}_n$

$$
\rho^{(n)}(\mathbf{r}_1,\ldots,\mathbf{r}_n)=\frac{N!}{(N-n)!}\sum_{\sigma_1,\ldots,\sigma_N}\int d\mathbf{r}_{n+1}\ldots d\mathbf{r}_N\,|\Psi(x_1,\ldots,x_N)|^2
$$

 $\star$  Particles 1 and 2 are correlated if

$$
\rho^{(2)}(\mathbf{r}_1, \mathbf{r}_2) \neq \rho^{(1)}(\mathbf{r}_1)\rho^{(1)}(\mathbf{r}_2)
$$

 $\star$  The quantity

$$
g(\mathbf{r}_1, \mathbf{r}_2) = \frac{\rho^{(2)}(\mathbf{r}_1, \mathbf{r}_2)}{\rho^{(1)}(\mathbf{r}_1)\rho^{(1)}(\mathbf{r}_2)}
$$

provides a measure of correlations in coordinate space

# The archetype corelated system: the Van der Waals liquid



 $\star$  Equation of state at particle density ρ and temperature *<sup>T</sup>*

$$
P = \frac{\rho T}{1 - \rho b} - a \rho^2 ,
$$

- $\triangleright$  *b* ∝ *d*<sup>3</sup> is the "excluded" volume"
- . *<sup>a</sup>* <sup>∼</sup> integral of the attractive part of the interaction
- $\star$  The *full* Van der Waals potential provides a good description of atomic systems. Hovever, its use in perturbation theory involves non trivial problems.

### Correlations in the non interacting Fermi gas

 $\star$  Enter Pauli's principle. Consider the ground state of a translationally invariant fermion system at density  $\rho = N/V = k_F^3/3\pi^2$ 

$$
\Psi_0(x_1,\ldots,x_N)=\frac{1}{N!}\ \det\left[\phi_{\alpha_i}(x_i)\right]\ ,\ \ \phi_{\alpha_i}(x_i)=\frac{1}{V^{1/2}}\ e^{i\mathbf{k}\cdot\mathbf{r}_i}\ \chi_{\sigma_i}\ ,\ \ |\mathbf{k}_i|
$$

 $\triangleright$  Statistical correlations are described by the function  $(r = |{\bf r}_1 - {\bf r}_2|)$ 

$$
g_{FG}(r) = \frac{\rho^{(2)}(r)}{\rho^2} = 1 - \frac{1}{2}\ell^2(k_F r)
$$

$$
\ell(x) = 3\frac{\sin x - x\cos x}{x^3}
$$



- $\star$  Bottom line: correlations are best defined in coordinate space.
- $\star$  To see this, consider the non interacting Fermi gas again. The joint probability of finding two particles with momenta  $\mathbf{k}_1$  and  $\mathbf{k}_2$  is

$$
n_{FG}(\mathbf{k}_1, \mathbf{k}_2) = \theta(k_F - |\mathbf{k}_1|)\theta(k_F - |\mathbf{k}_2|) \left[1 - \frac{1}{N}\frac{\rho}{2} (2\pi)^3 \delta(\mathbf{k}_1 - \mathbf{k}_2)\right]
$$

? *In the absence of long range order*, a similar result holds true in interacting systems

$$
n(\mathbf{k}_1, \mathbf{k}_2) = n(\mathbf{k}_1)n(\mathbf{k}_2) [1 + O(1/N)]
$$

 $\star$  In momentum space, non trivial correlations effects on  $n(k_1, k_2)$  vanish in the  $N \to \infty$  limit. However, correlations strongly affect the behaviour of  $n(\mathbf{k})$  at  $|\mathbf{k}| > k_F$ .

### Interaction without correlation: the mean field picture

 $\star$  Dynamical correlations are induced by two-body interactions described by the potential  $v_{ii}$  appearing in the N-particle hamiltonian

$$
H = \sum_{i=1}^{N} -\frac{\nabla_i^2}{2m} + \sum_{j>i=1}^{N} v_{ij},
$$

 $\star$  The mean field approximation is based on the replacements

$$
\sum_{j>i=1}^{N} v_{ij} \rightarrow \sum_{i=1}^{N} U_i \quad , \quad H \rightarrow \sum_{i=1}^{N} h_i = \sum_{i=1}^{N} \left( -\frac{\nabla_i^2}{2m} + U_i \right)
$$
   
implying

$$
H|\Psi_0\rangle = E_0|\Psi_0\rangle \rightarrow h_i|\phi_i\rangle = \epsilon_i|\phi_i\rangle
$$

 $\star$  Within the mean field approximation

$$
E_0 = \sum_{i \in \{F\}} \epsilon_i , \quad \Psi(x_1, \dots, x_N) = \det[\phi_i(x_i)]
$$

$$
\rho^{(2)}(\mathbf{r}_1, \mathbf{r}_2) = \sum_{i} \phi_i^{\dagger}(\mathbf{r}_1) \phi_j^{\dagger}(\mathbf{r}_2) \left[ \phi_i(\mathbf{r}_1) \phi_j(\mathbf{r}_2) - \phi_j(\mathbf{r}_1) \phi_i(\mathbf{r}_2) \right]
$$

- $\star$  The mean field approach provides a remarkably accurate description of a variety of properties of interacting many-body systems. However, one should keep in mind that
	- $\triangleright$  dynamical correlations are not taken into acount

*<sup>i</sup>*,*j*∈{*F*}

- $\triangleright$  including their effects as corrections to the mean field approximation may be highly misleading, as the definition of the mean field itself is model dependent
- $\star$  Theoretical studies aimed at pinning down the role of correlations should be carried out within *ab initio* many body approaches

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# Model independent determination of correlations

 $\star$  Definition of Green's function

 $iG(x - x') = \langle 0|T[\hat{\psi}(x)\hat{\psi}^{\dagger}(x')]|0\rangle$ 

After Fourier transformation ( $\eta = 0^+$ )

$$
G(\mathbf{k},E) = \sum_{n} \left\{ \frac{|\langle n_{(N+1)}(\mathbf{k})|a_{\mathbf{k}}^{\dagger}|0_{N}\rangle|^{2}}{E - (E_{n} - E_{0}) + i\eta} + \frac{|\langle n_{(N-1)}(-\mathbf{k})|a_{\mathbf{k}}|0_{N}\rangle|^{2}}{E + (E_{n} - E_{0}) - i\eta} \right\}
$$

$$
= G_{p}(\mathbf{k},E) + G_{h}(\mathbf{k},E) = \int dE' \left[ \frac{P_{p}(\mathbf{k},E')}{E - E' + i\eta} + \frac{P_{h}(\mathbf{k},E')}{E + E' - i\eta} \right]
$$

 $\star$  Spectral functions of hole and particle states

$$
P_h(\mathbf{k}, E) = \sum_n |\langle n_{(N-1)}(\mathbf{k})|a_{\mathbf{k}}|0_N\rangle|^2 \delta(E - E_n + E_0) = \frac{1}{\pi} \operatorname{Im} G_h(\mathbf{k}, E)
$$

$$
P_p(\mathbf{k}, E) = \sum_n |\langle n_{(N+1)}(\mathbf{k})|a_{\mathbf{k}}^\dagger|0_N\rangle|^2 \delta(E + E_n - E_0) = \frac{1}{\pi} \operatorname{Im} G_p(\mathbf{k}, E)
$$

### Analytic structure of the Green's function

 $\star$  In interacting systems, the Green's function (e.g. for hole states) can be written in terms of the particle self energy  $\Sigma(k, E)$ 

$$
G_h(\mathbf{k}, E) = \frac{1}{E - |\mathbf{k}|^2 / 2m - \Sigma(\mathbf{k}, E)}
$$

- ? Landau's quasiparticle picture: isolate contributions of <sup>1</sup>*<sup>h</sup>* (*bound*) intermediate states, exhibiting poles at energies  $\epsilon_k$ , given by  $\epsilon_k = |\mathbf{k}|^2 / 2m + \text{Re }\Sigma(\mathbf{k}, \epsilon_k)$ , as Im  $\Sigma(\mathbf{k}, E) \to 0$  (Fermi surface)
- $\star$  The resulting expression is

<span id="page-9-0"></span>
$$
G_h(\mathbf{k}, E) = \frac{Z_k}{E - \epsilon_k - i Z_k \operatorname{Im} \Sigma(\mathbf{k}, e_k)} + G_h^B(\mathbf{k}, E)
$$

where  $Z_k = |\langle -\mathbf{k} | a_{\mathbf{k}} | 0 \rangle|^2$ , and  $G_h^B(\mathbf{k}, E)$  is a smooth contribution, arising<br>from  $2h - 1p$ ,  $3h - 2p$  (continuum) intermediate states from  $2h - 1p$ ,  $3h - 2p$ , . . . (*continuum*) intermediate states

# Correlated Basis Functions (CBF) approach

 $\star$  Correlated states obtained from Fermi gas states through the transformation

$$
|n\rangle = \frac{F}{\langle n_{FG} | F^{\dagger} F | n_{FG} \rangle} | n_{FG} \rangle \quad , \quad F = S \prod_{j>i} f_{ij}
$$

 $\star$  The two-nucleon correlation operator reflects the complexity of the nucleon-nucleon (NN) force [spin-isospin (ST) dependent, non central]

$$
f_{ij} = \sum_{TS} [f_{TS}(r_{ij}) + \delta_{S1} f_{Tt}(r_{ij}) S_{ij}] P_{TS}
$$

*P*<sub>*TS*</sub> : spin – isospin projectors,  $S_{ij} = \sigma_i^{\alpha} \sigma_j^{\beta} \left( 3r_{ij}^{\alpha} r_{ij}^{\beta} - \delta^{\alpha \beta} \right)$ 

 $\star$  Shapes of  $f_{TS}$ ,  $f_{tT}$  determined from minimization of ground state energy

 $\star$  Split the hamiltonian according to

 $H = H_0 + H_I$ 

 $\langle m|H_0|n\rangle = \delta_{mn}\langle m|H|n\rangle$ ,  $\langle m|H_I|n\rangle = (1 - \delta_{mn})\langle m|H|n\rangle$ 

 $\star$  If correlated states have large overlaps with the eigenstates of the hamiltonian, the matrix elements of  $H<sub>I</sub>$  are small and perturbation theory can be used to obtain, e.g., the ground state from

$$
\widetilde{|0\rangle} = \sum_{m} (-)^{m} \left( \frac{H_{I} - \Delta E_{0}}{H_{0} - E_{0}^{V}} \right)^{m} |0\rangle
$$
  

$$
\Delta E_{0} = E_{0}^{V} - E_{0} = \langle 0|H|0\rangle - E_{0}
$$

### Hole spectral function of nuclear matter from CBF

$$
P_h(\mathbf{k}, E) = \frac{1}{\pi} \frac{Z_k^2 \operatorname{Im} \Sigma(\mathbf{k}, \epsilon_k)}{[E - \mathbf{k}^2 / 2m - \operatorname{Re} \Sigma(\mathbf{k}, E)]^2 + [Z_k \operatorname{Im} \Sigma(\mathbf{k}, \epsilon_k)]^2} + P_h^B(\mathbf{k}, E)
$$



# Spectral function of infinite nuclear matter

 $\star$  Results obtained using CBF perturbation theory and the U14+TNI hamiltonian



 $\star$  The correlation contribution can be identified by its distinctive energy dependence

# Momentum distribution and spectroscopic factors

 $\star$  In analogy with the spectral function, the momentum distribution can be split into quasi particle (pole) and and correlation (continuum) contributions

$$
n(\mathbf{k}) = \int dE P(\mathbf{k}, E) = Z_k \theta(k_F - |\mathbf{k}|) + \int dE P_B(\mathbf{k}, E) = Z_k \theta(k_F - |\mathbf{k}|) + n_B(\mathbf{k})
$$



# Exploiting the (near) universality of correlations

 $\star$  Local density approximation

 $P(k, E) = P_{MF}(k, E) + P_{corr}(k, E)$ 

 $\Rightarrow$   $P_{MF}(\mathbf{k}, E) \rightarrow$  from  $(e, e'p)$  data

$$
P_{MF}(\mathbf{k}, E) = \sum_{n} Z_n |\phi_n(\mathbf{k})|^2 F_n(E - E_n)
$$

 $\triangleright$  *P*<sub>corr</sub>( $\mathbf{p}, E$ )  $\rightarrow$  from uniform nuclear matter calculations at different densities:

$$
P_{corr}(\mathbf{k}, E) = \int d^3r \, \rho_A(r) \, P_{corr}^{NM}(\mathbf{k}, E; \rho = \rho_A(r))
$$

- $\star$  Widely and successfully employed to analize (*e*, *e'*) data at beam<br>energies  $\approx 1 GeV$ energies ∼ 1*GeV*
- $\star$  Warnings: model dependence, chance of double counting

# Theory vs data ( $E_e = 1.3$  GeV,  $\theta_e = 37.5^\circ$ )

 $\star$  Note: calculations involve no adjustable parameters



 $\star$  The measured x-section can be described, except in the *dip* region, between the quasi elastic and ∆-production peaks, and the low enrgy loss tail, where FSI (not included) play a role  $\Omega$ 

### Correlation effects on the nuclear response

 $\star$  Consider scattering of a scalar probe, for simplicity

$$
\frac{d\sigma}{d\Omega d\omega} \propto S(\mathbf{q}, \omega) = \sum_{n} \langle 0 | \rho_{\mathbf{q}}^{\dagger} | n \rangle \langle n | \rho_{\mathbf{q}} | 0 \rangle \delta(E_0 + \omega - E_n)
$$

$$
\rho_{\mathbf{q}} = \sum_{\mathbf{k}} a_{\mathbf{k}+\mathbf{q}}^{\dagger} a_{\mathbf{k}} \quad , \quad H|0\rangle = E_0|0\rangle \quad , \quad H|n\rangle = E_n|n\rangle
$$

 $\star$  Rewrite the response in the form

$$
S(\mathbf{q}, \omega) = \sum_{n} \left| \sum_{k} \langle n | a_{\mathbf{k}+\mathbf{q}}^{\dagger} a_{\mathbf{k}} | 0 \rangle \right|^{2} \delta(\omega + E_{0} - E_{n})
$$
  

$$
= \int \frac{dt}{2\pi} e^{i(\omega + E_{0})t} \sum_{\mathbf{p}, \mathbf{k}} \langle 0 | a_{\mathbf{p}+\mathbf{q}} a_{\mathbf{p}}^{\dagger} e^{-iHt} a_{\mathbf{k}+\mathbf{q}}^{\dagger} a_{\mathbf{k}} | 0 \rangle
$$

 $\star$  *S*(q,  $\omega$ ) can be expressed in terms of interactions and Green functions describing nucleons in particle and hole states

# Effects of interactions on the nuclear response

- $\star$  In the absence of correlations, the only possible final states are one particle-one hole states
- $\star$  For example, according to the Fermi gas model

<span id="page-18-0"></span>
$$
M_n = \langle n | \sum_{\mathbf{k}} a_{\mathbf{k}+\mathbf{q}}^{\dagger} a_{\mathbf{k}} | 0 \rangle \to M_k = 1 \times \theta(k_F - |\mathbf{k}|) \theta(|\mathbf{k} + \mathbf{q}| - k_F)
$$
  

$$
S(\mathbf{q}, \omega) = \sum_{\mathbf{k}} |M_k|^2 \delta(\omega + e_0(\mathbf{k}) - e_0(\mathbf{k} + \mathbf{q})) \quad , \quad e_0(\mathbf{k}) = \frac{\mathbf{k}^2}{2m}
$$

 $\star$  Inclusion of interactions, through the replacement of Fermi gas states with CBF states, leads to a quenching of the transition matrix elements  $M_k$  and to a modification of the single particle spectrum  $e_0(\mathbf{k})$ 

### Correlations & interaction effects

- $\star$  Isospin symmetric nuclear natter at equilibrium density
- $\triangleright$  Correlations

 $M_{ph}$  < 1



 $\triangleright$  Mean field

 $m \rightarrow m^{\star} =$ 

<span id="page-19-0"></span> $\sqrt{1}$ *k*  $\left(\frac{de}{dk}\right)^{-1}$ 

### Correlation & interaction effects on the response

 $\star$  (A), (B), (C) → |q| = 0.3, 1.8, 3.0 fm<sup>-1</sup><br>2.5



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## Empirical evidence of correlation effects

Energy dependence of the spectroscopic strengths of shell model states<br>of  $\frac{208 \, Pb}{\text{m}}$  measured in high resolution (e e'n) experiments at NIKHEFof <sup>208</sup>*Pb*, measured in high resolution (*e*, *e'p*) experiments at NIKHEF-K



 $\star$  Theory: CBF nuclear matter results corrected for surface effects

# Measured correlation strength

- $\star$  The correlation strength in the 2p2h sector has been measured by the JLAB E97-006 Collaboration using a carbon target
- $\star$  Strong energy-momentum correlation:  $E \sim E_{thr} + \frac{A-2}{A-1}$  $\frac{A-2}{A-1}$   $\frac{\mathbf{k}^2}{2n}$ 2*m*



 $\star$  Measured correlation strength 0.61  $\pm$  0.06, to be compared with the theoretical predictions 0.64 (CBF) and 0.56 (G-Matrix)

### FSI in the impulse approximation regime

At momentum transfer  $|q|^{-1}$  >>  $2\pi/d$ , *d* being the average interparticle<br>separation distance separation distance

$$
S(\mathbf{q}, \omega) = \int d^3k dE P_h(\mathbf{k}, E) P_p(\mathbf{k} + \mathbf{q}, \omega - E)
$$

- $\triangleright$  *P<sub>h</sub>*  $\rightarrow$  many-body theory
- $\rightarrow$  *P<sub>p</sub>*  $\rightarrow$  many-body theory + eikonal approximation (OB, arXiv:1301.3357)
- $\star$  The struck particle travels along a straight trajectory with constant speed v. Its propagation is described by the Green's function  $(p = |k + q|)$

$$
G(\mathbf{r}_{\perp}, z) = -\frac{i}{\mathrm{v}} \delta(\mathbf{r}_{\perp}) \theta(z) \exp\left[i p z - \frac{i}{\mathrm{v}} \int_0^z d\zeta \ V(\zeta)\right]
$$

with

$$
V(\zeta) = \langle 0 | \sum_{j=2}^{N} \Gamma_{\mathbf{p}}(\mathbf{r}_{1j} + \hat{\mathbf{z}} \zeta) | 0 \rangle
$$

### Correlation effects in FSI

 $\star$  The interaction is described by the Fourier transform of the scattering amplitude

$$
\Gamma_{\mathbf{p}}(\mathbf{r}) = -\frac{2\pi}{m} \int \frac{d^3k}{(2\pi)^3} e^{-i\mathbf{k}\cdot\mathbf{r}} f_{\mathbf{p}}(\mathbf{k}) .
$$

with

$$
f_p(\mathbf{k}) = \frac{p}{4\pi} \sigma_p(\alpha_p + i) e^{-\beta_p \mathbf{k}^2}
$$

 $\star$  FSI are driven by the quantity

<span id="page-24-0"></span>
$$
V(\zeta) = \int d^3r \, g(r) \, \Gamma_p(\mathbf{r} + \hat{\mathbf{z}} \zeta)
$$

 $\star$  Under the assumptions underlying the eikonal approximation, correlations in coordinate space strongly affect the energy dependence of the spectral function.

 $\star$  Consider the simple case  $\alpha_p = \beta_p = 0$ , i.e.

Im 
$$
\Gamma_p(\mathbf{r}) = -\frac{1}{2}\rho v \sigma_p \delta(\mathbf{r})
$$

The corresponding eikonal phase is



 $\star$  After Fourier transformation, the *z*-dependence of *W* leads to a specific energy dependence of the eikonal spectral fun[cti](#page-24-0)[on](#page-26-0)  $\Omega$ 

# Preliminary results

 $\star$  Isospin symmetric nuclear matter at equilibrium density



- <span id="page-26-0"></span> $\star$  Main elements of the calculation<br> $\star$  medium modified nucleon-nuc
	- . medium modified nucleon-nucleon cross setctions
	- $\triangleright$  nucleon radial distribution function,  $g(r)$

### Nuclear binding, correlations and the EMC effect

 $\star$  The analysis of the dependence of the slope of the EMC ratio on the average nucleon removal energy, defined as

$$
\langle E \rangle = \int d^3k dE P(\mathbf{k}, E)
$$

requires a level of accuracy not yet achieved for nuclei with  $A > 3$ 

- $\star$  Green's Function Monte Carlo (GFMC) calculations provide the ground state energies,  $E_0$  and the expectation values of the kinetic energy operator,  $\langle T \rangle$ , of nuclei with  $A \leq 12$ , obtained from state-of-the-art nuclear hamiltonian
- $\star$  The corresponding average removal energies can be calculated using the GFMC results and the Koltun sum rule, stating that (up to a small correction arising from the three-body potential)

$$
\frac{E_0}{A} = \frac{1}{2} \left[ \frac{A-2}{A-1} \langle T \rangle - \langle E \rangle \right]
$$

 $\star$  The slope is analyzed in terms of the variable

 $\tilde{y} = v - |\mathbf{q}|$ 

that can be interpreted as the longitudinal momentum of the struck particle in the target rest frame. Note that  $\tilde{y}$  is trivially related to Nachtmann's variable through  $\tilde{y} = -\xi/m$ .



OB & I. Sick arXiv:1207.4595<br>
That Benhar (INFN, Roma)

- $\star$  The data shows an excellent correlation with  $\langle E \rangle$
- $\star$  The analysis includes the ratio obtained from the extrapolated nuclear matter data. The corresponding removal energy is obtained from the values of  $E_0$  and  $\langle T \rangle$  resulting from the CBF calculation of Akmal & Pandharipande
- $\star$  The values of  $\langle E \rangle$  employed in the analysis are significantly larger than those used in similar studies. For example, in Carbon the removal energy extracted from  $(e, e'p)$  data, corresponding to the shell model states, is<br> $\approx 25 \text{ MeV}$  to be compared to the GEMC result  $\approx 52 \text{ MeV}$  $\sim$  25 MeV, to be compared to the GFMC result ~ 52 MeV
- $\star$  The large values of  $\langle E \rangle$  are to be ascribed to strong nucleon-nucleon correlations, leading to the excitation of nucleons to states of high removal energy *and* high momentum

- $\star$  It is long known that correlation effects in nuclei are large. Back in 1952 AD, Blatt & Weiskopf pointed out that:
	- $\triangleright$  "The limitation of any independent particle model lies in its inability to encompass the correlation between the positions and spins of the various particles in the system"
- $\star$  While being best defined in coordinate space, correlations manifest themselves in a distinctive energy dependence of the Green's functions.
- $\star$  Pinning down pure correlation effetcs in a model independent fashion requires the calculation of the Green's function within ab initio many-body approaches.
- $\star$  There is ample empirical evidence of important correlation effects from electron-nucleus scattering data. However, the definition of correlation observables remains somewhat elusive.

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