

# The NEXT-100 Double Beta Decay Experiment - Progress and Perspectives for the Ton-scale

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# Dual-purpose Concept

## 1. $0\nu\beta\beta$ search:

NEXT-100 experiment at Canfranc is based on a high-pressure Xe gas TPC for better performance

- Spring-board for dual-purpose ton-scale system

## 2. WIMP search:

Novel approach for *directional* sensitivity in WIMP nuclear recoils exploits columnar recombination

- If successful, active mass  $\rightarrow$  ton-scale is possible

# Simultaneous searches?

- **Next generation projects will be expensive!**
  - A dual-purpose detector should be considered... *if it truly saves money and truly is dual-purpose*
- Xenon is an attractive choice for both searches
  - No long-lived isotopes
  - Relatively cheap, and easy to enrich
  - Can exchange: enriched  $\leftrightarrow$  depleted
  - Scales well as monolithic source = detector

# Xenon in Gas Phase?

- Gas phase offers attractive possibilities:
  - Normal energy partition fluctuations:  $F = 0.15$ 
    - Excellent correlation of ionization with deposited energy
      - » Remarkably good energy resolution ( $0\nu\text{-}\beta\beta$ )
    - Much better discrimination between electron/nuclear recoils
      - » Small S2/S1 fluctuations (WIMPs)
  - Visualization of event topology ( $0\nu\text{-}\beta\beta$  & WIMPs)
    - Must try to evade background dominant  $(Mt)^{1/4}$  regime
  - Nuclear recoil directional sensing possibility
    - *Optimal density? Maybe 10 bars! → x1000 advance?*

# Why Xenon Gas?

Energy resolution in Xenon depends very strongly on density

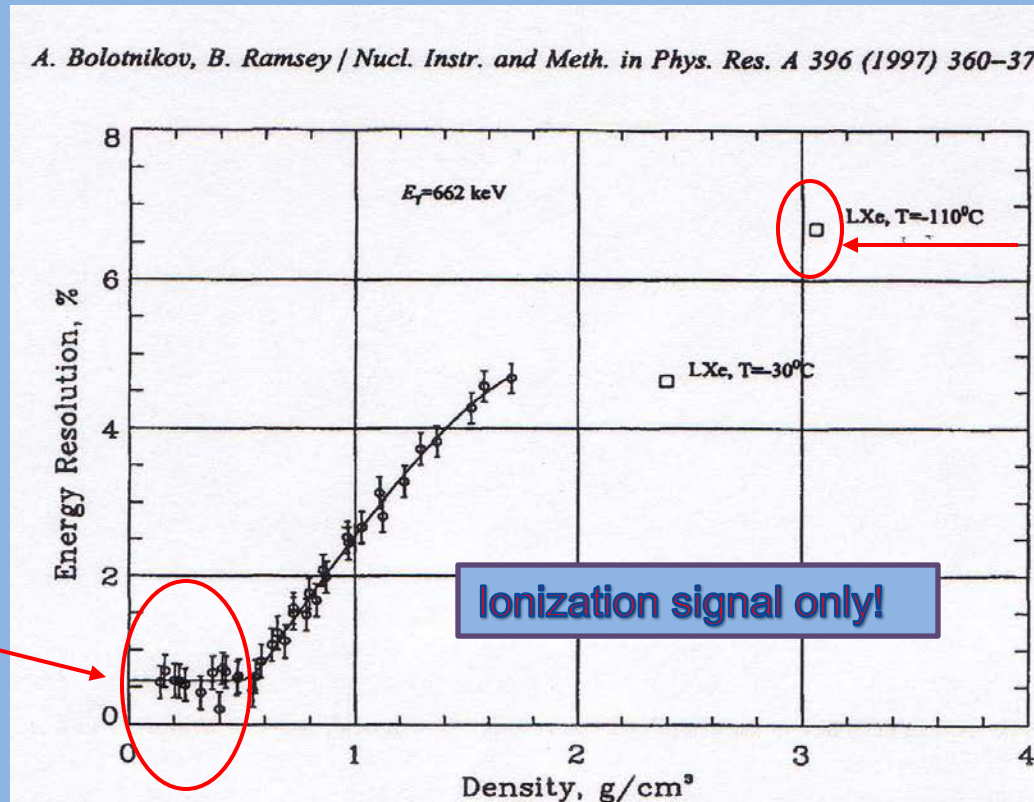
Here, the fluctuations are normal

Fano factor

$$F = 0.15$$

Unfolded resolution:

$$\delta E/E \sim 0.6\% \text{ FWHM}$$



Very large fluctuations between light/charge!

$$F \sim 20 !!$$

For  $\rho < 0.55 \text{ g/cm}^3$ , energy resolution from ionization is “intrinsic”

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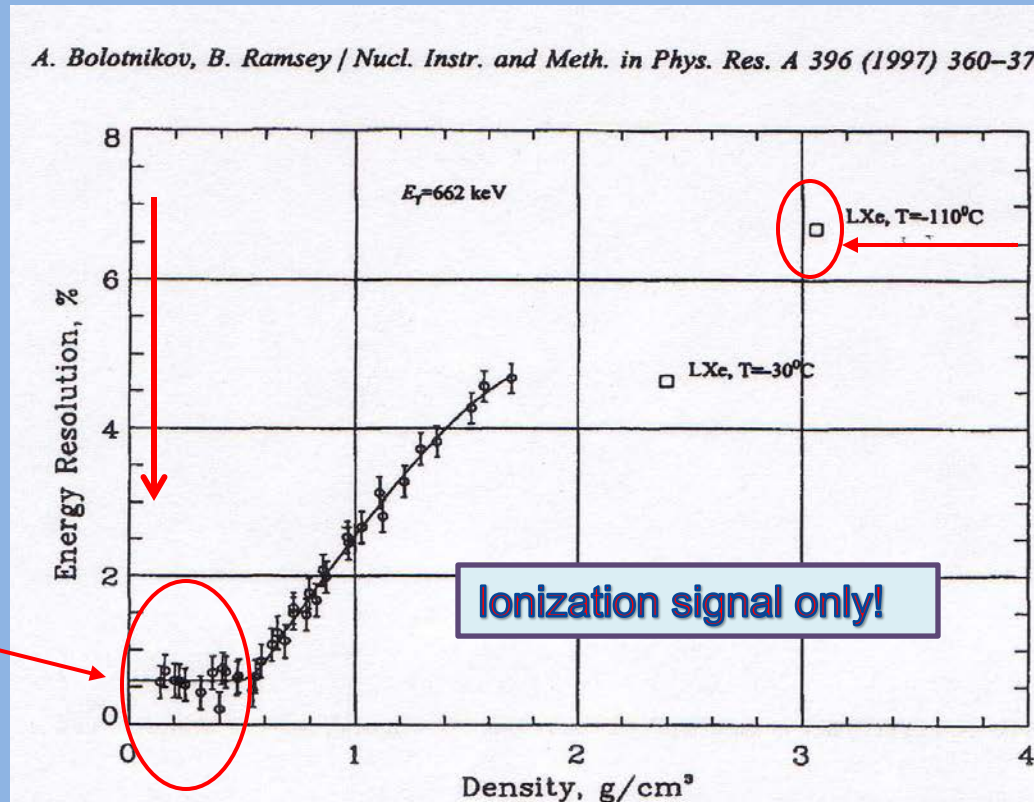
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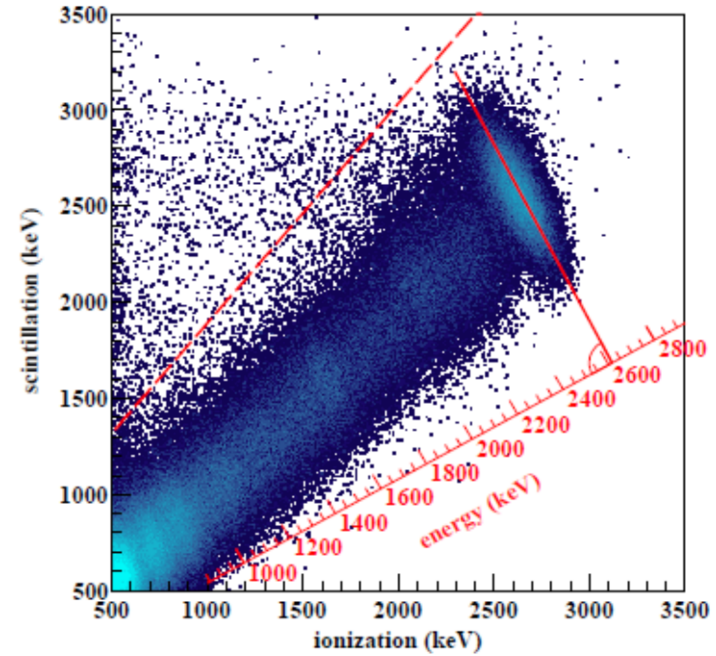
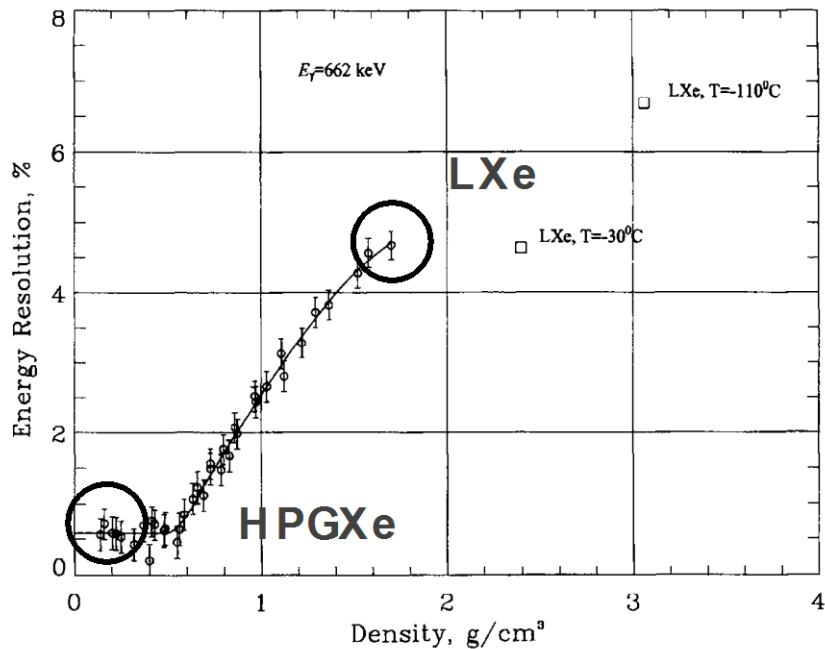
$$F \sim 20 !!$$

**WIMPs:**

**Large S2/S1 fluctuations**

For  $\rho < 0.55 \text{ g/cm}^3$ , energy resolution from ionization is "intrinsic"

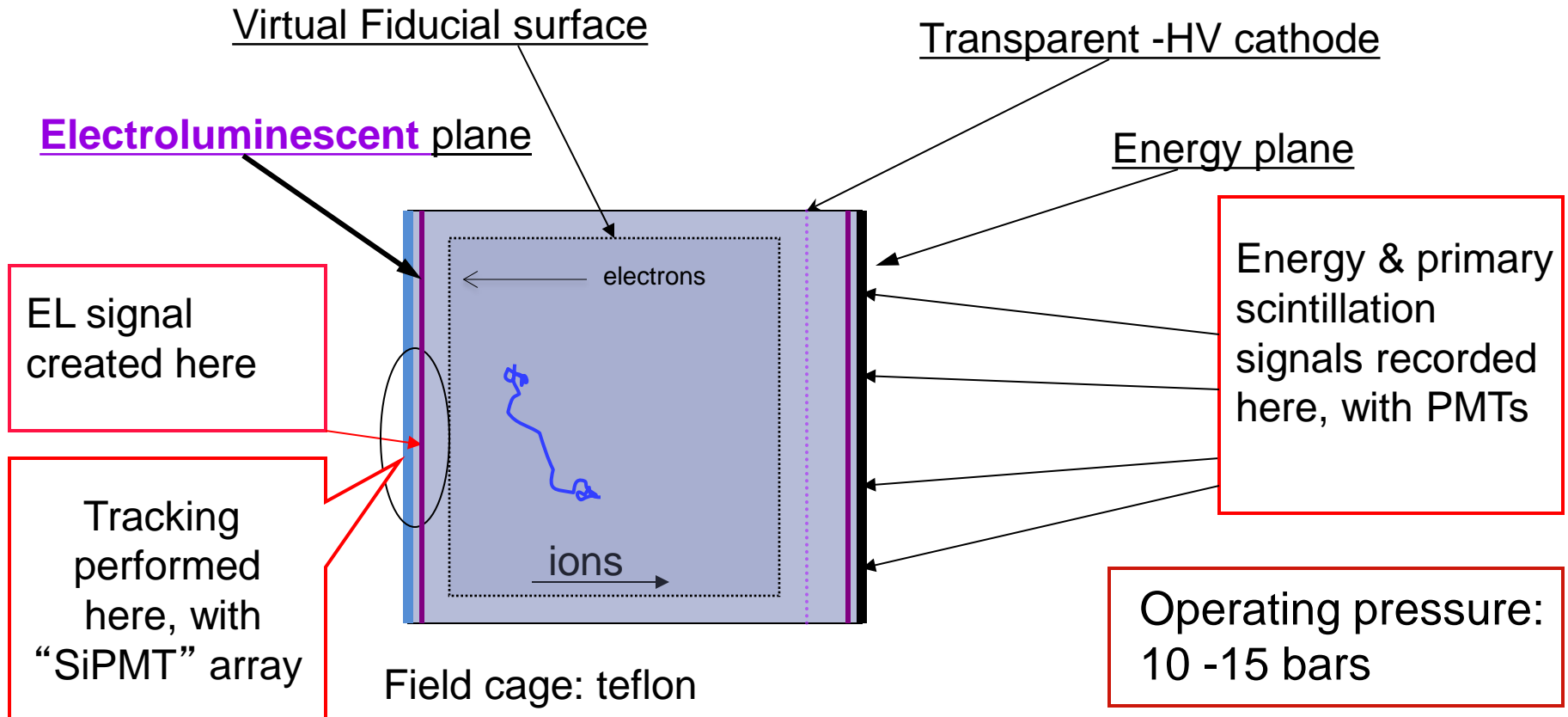
# LXe: Energy resolution



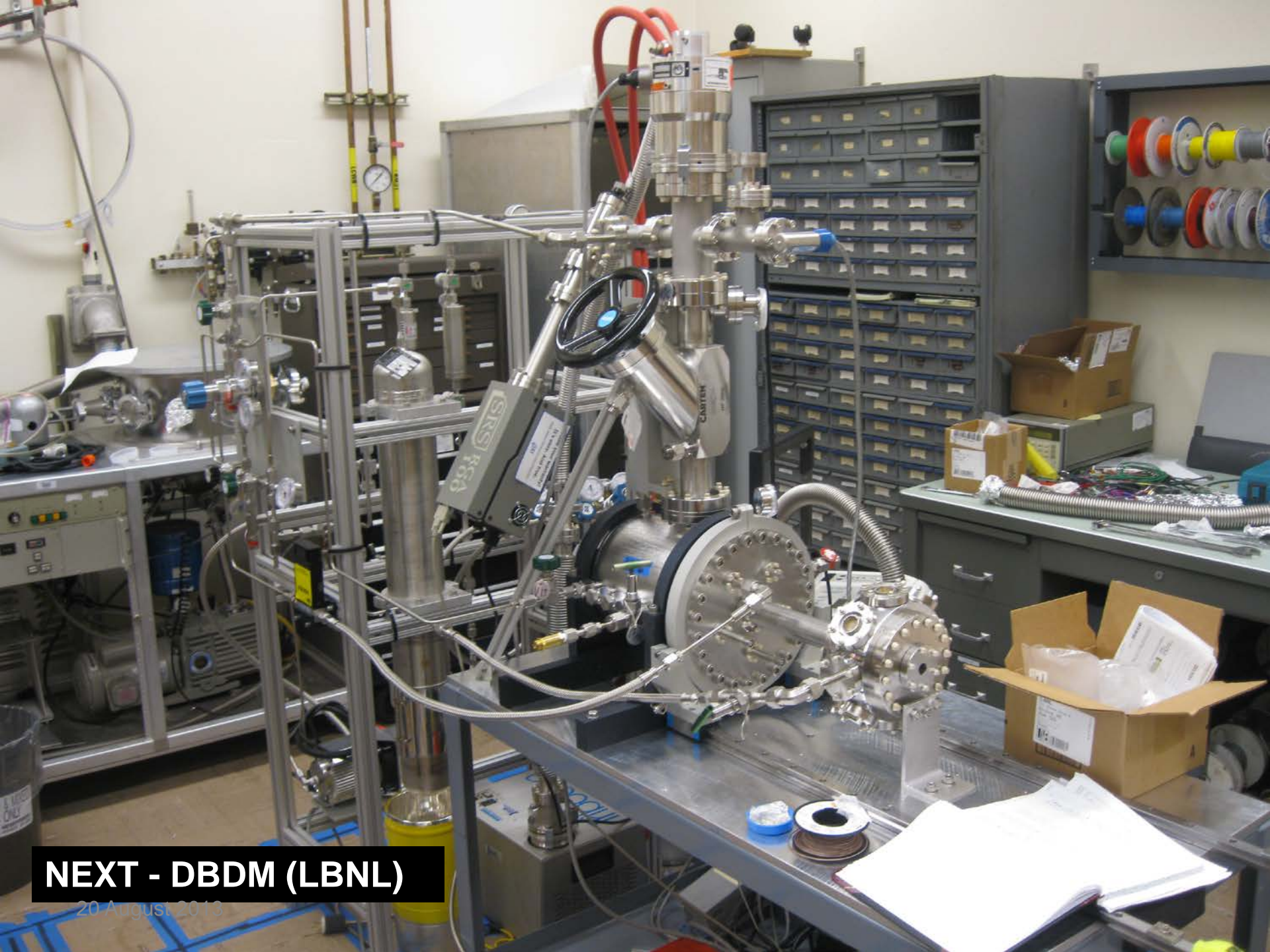
Energy resolution:  
Anomalous in LXe.  
Much worse than in  
HPXe.

Energy resolution: 4%  
FWHM at Q, using anti-  
correlation between  
scintillation and  
ionization

# Asymmetric TPC with “Separated functions”

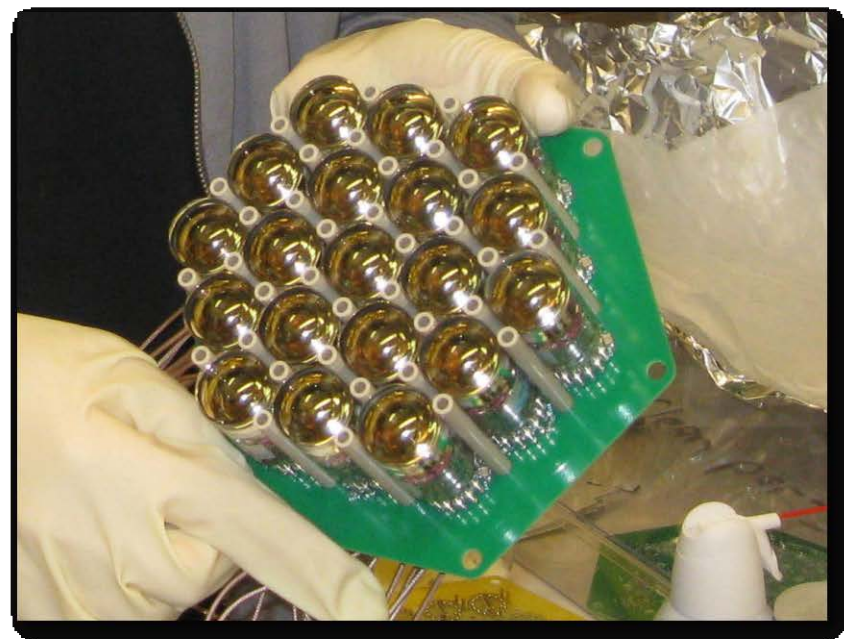
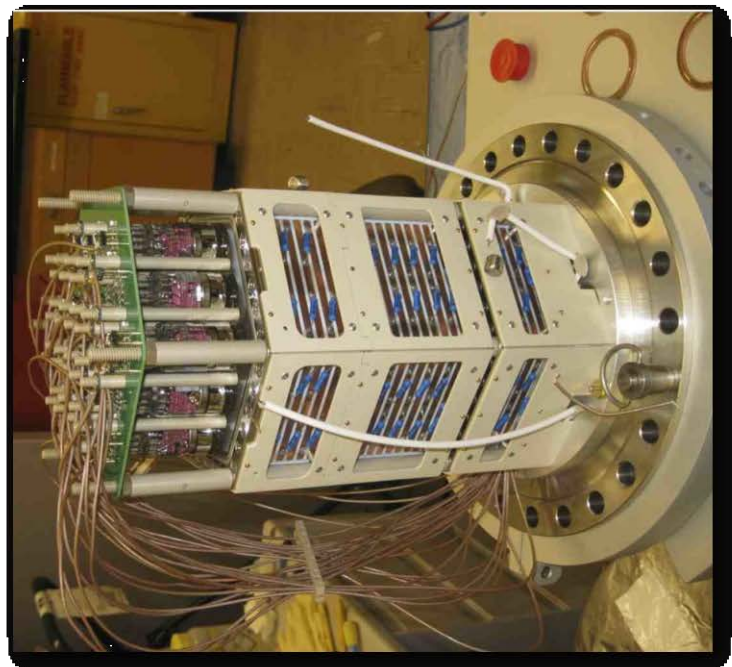
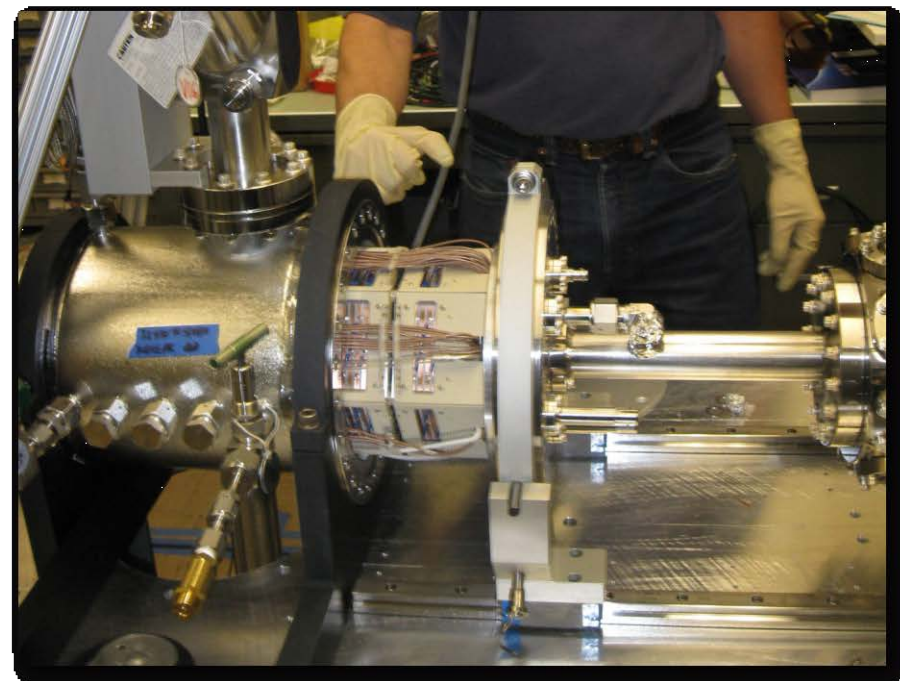
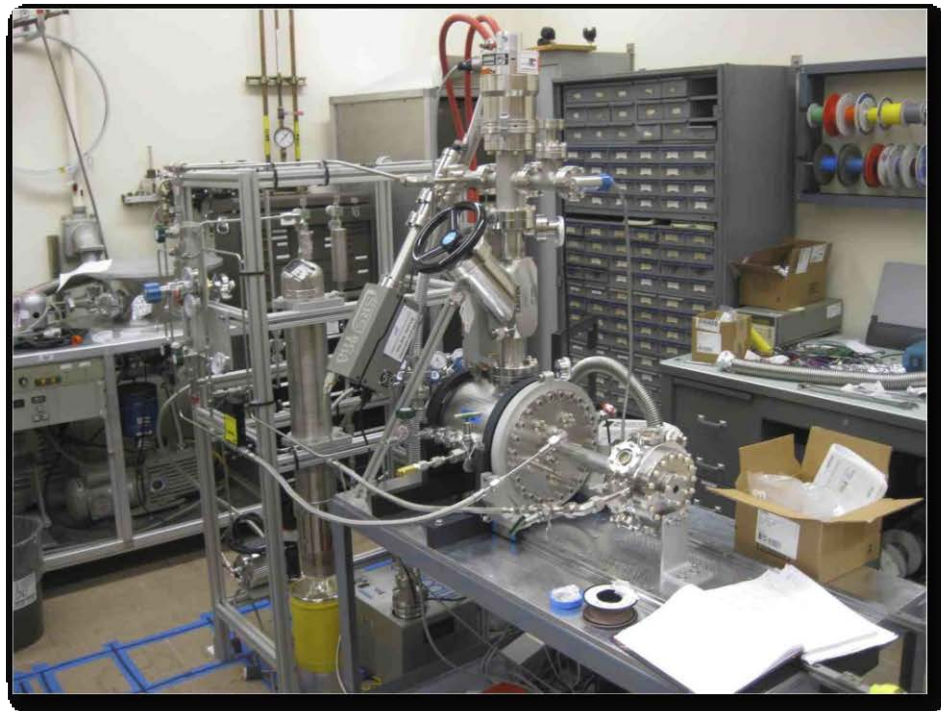




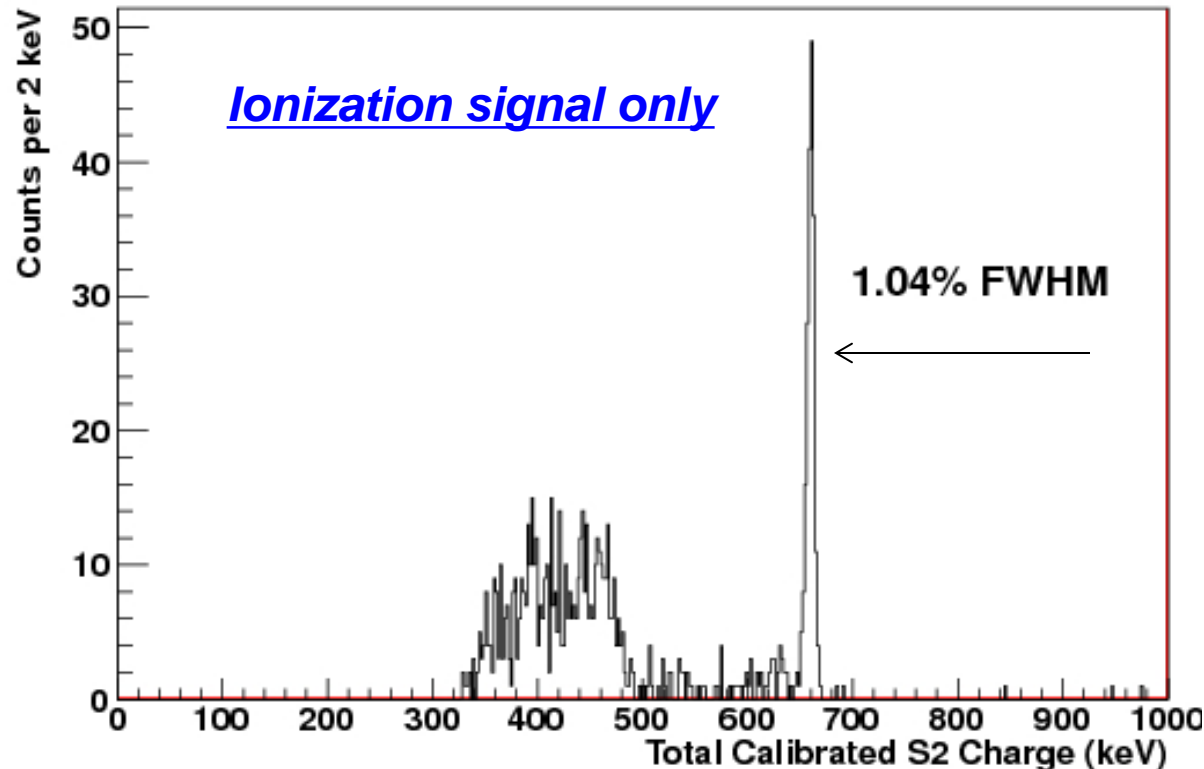


**NEXT - DBDM (LBNL)**

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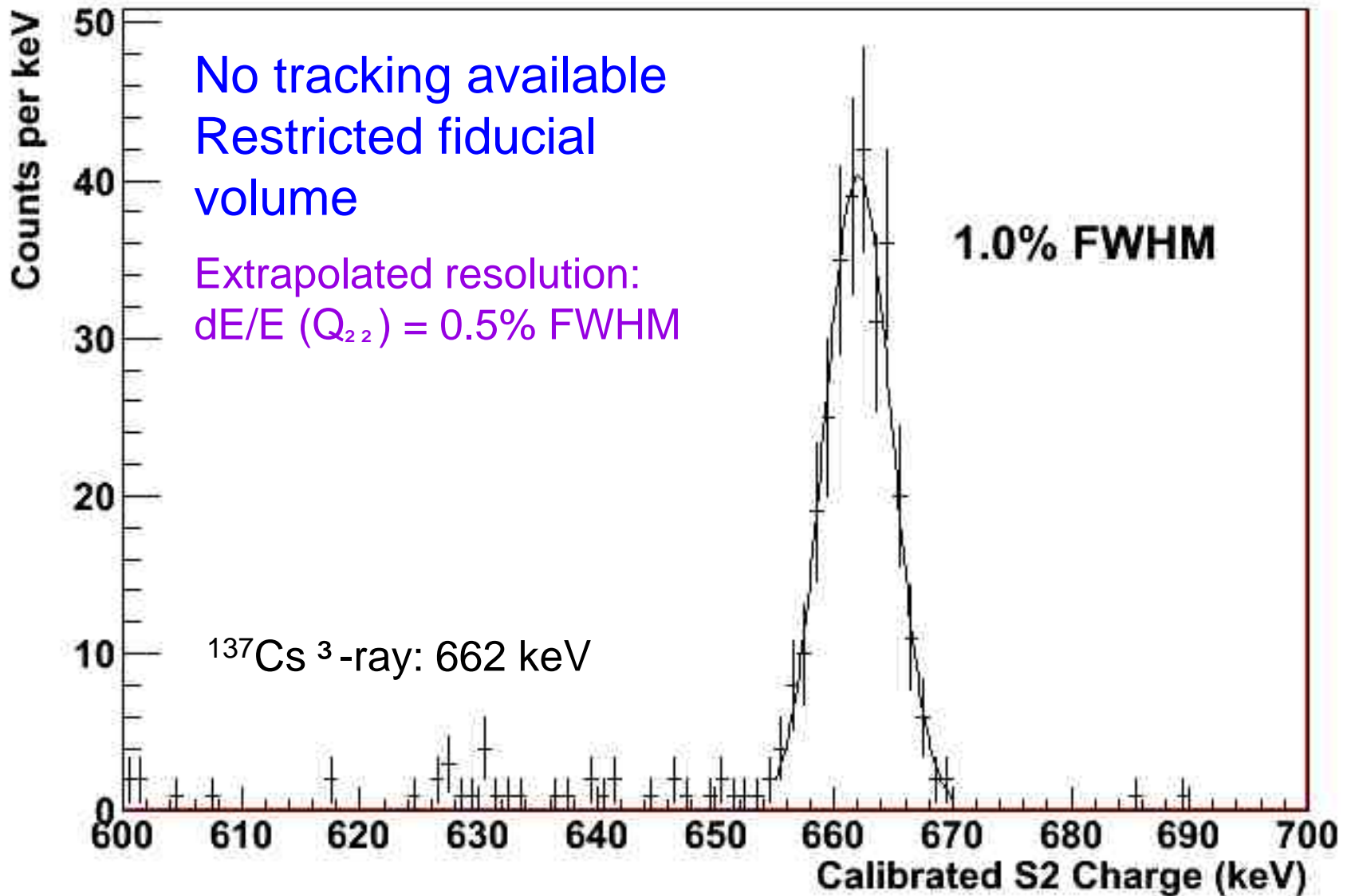
# World record: Energy resolution $\sigma E/E = 1\%$ FWHM for $^{137}\text{Cs}$ 662 keV $\gamma$ -rays in xenon!

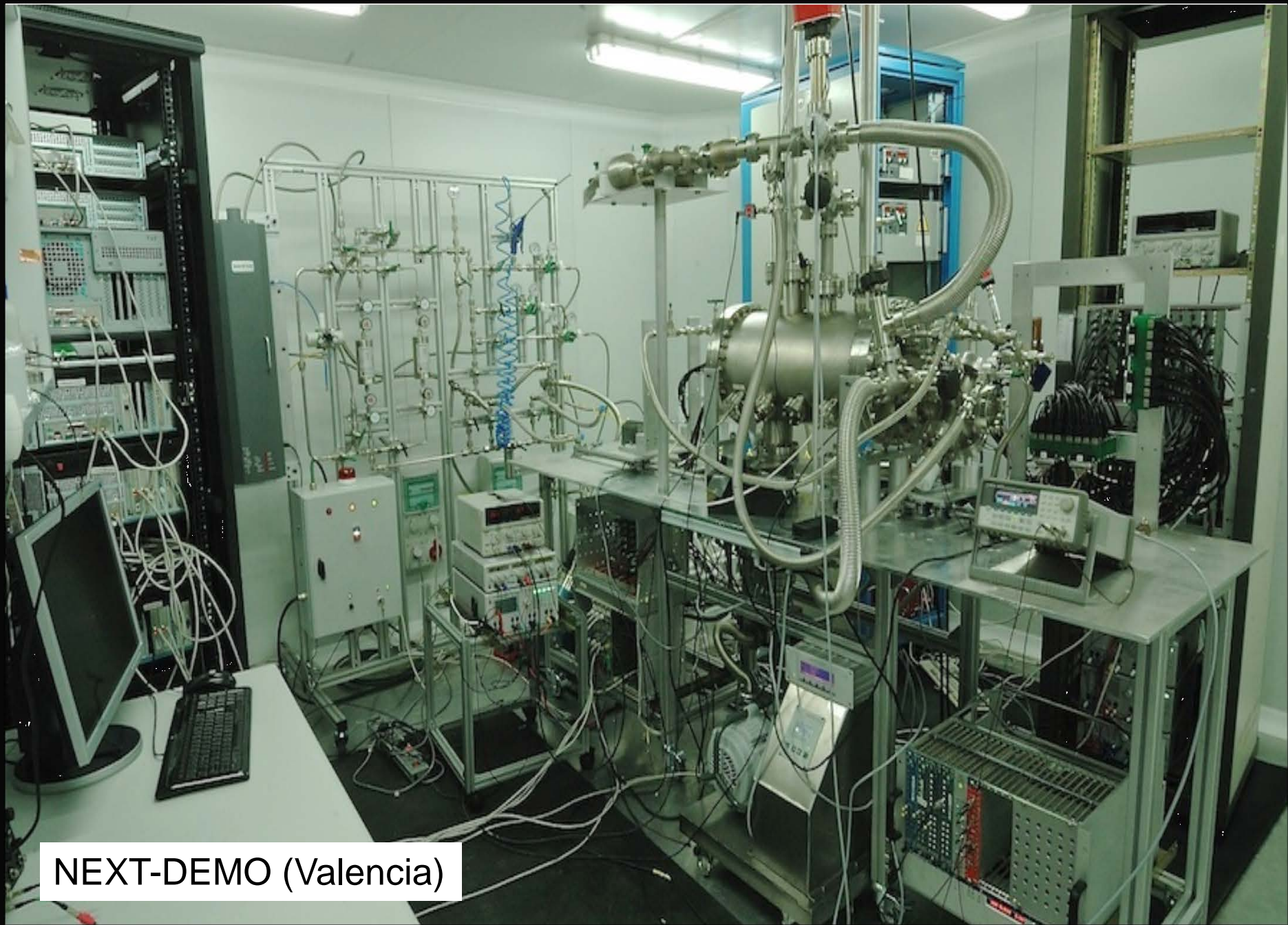


**Data** from  
LBNL-TAMU  
HPXe TPC

This result is  
important for  
both  $0\nu\beta\beta$  &  
WIMP searches

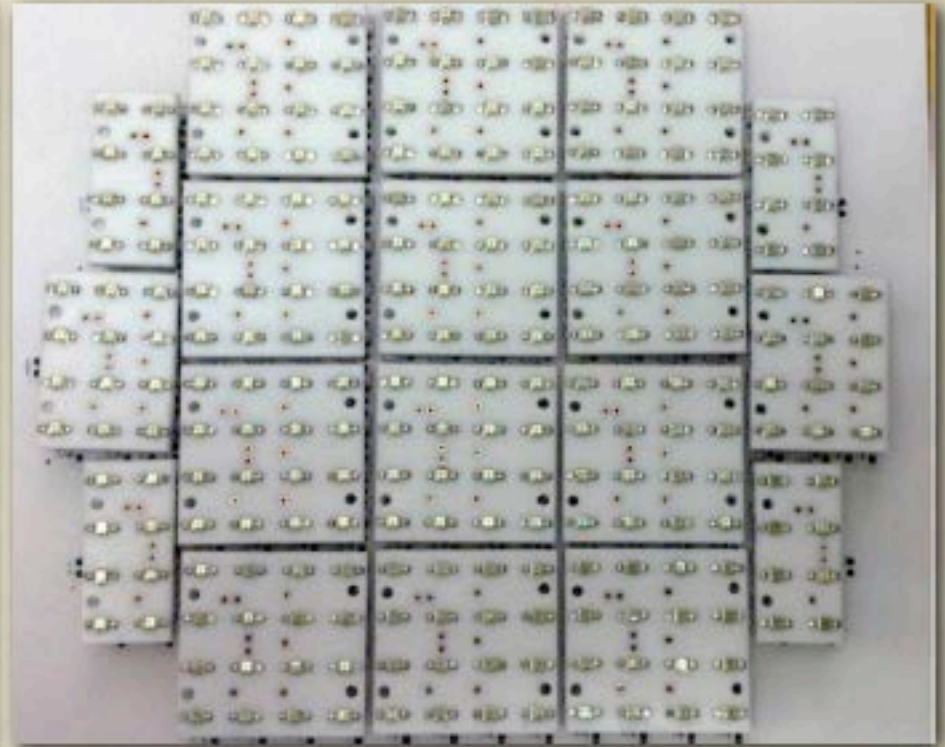
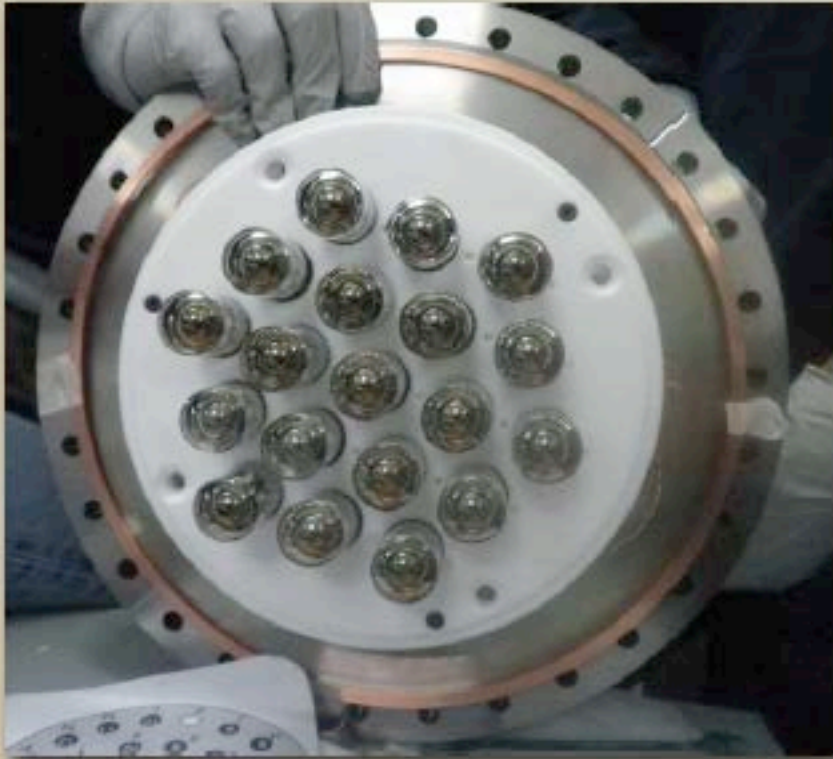
This result shows that fluctuations are “normal” in HPXe



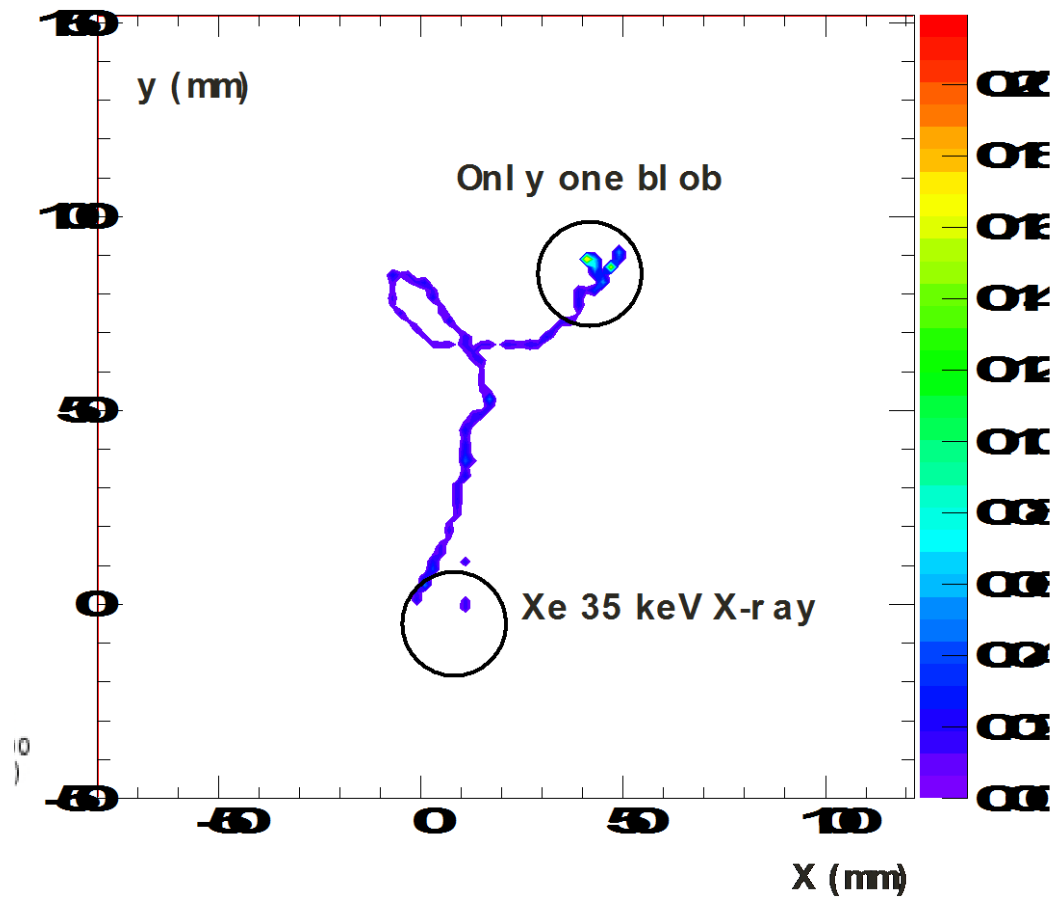
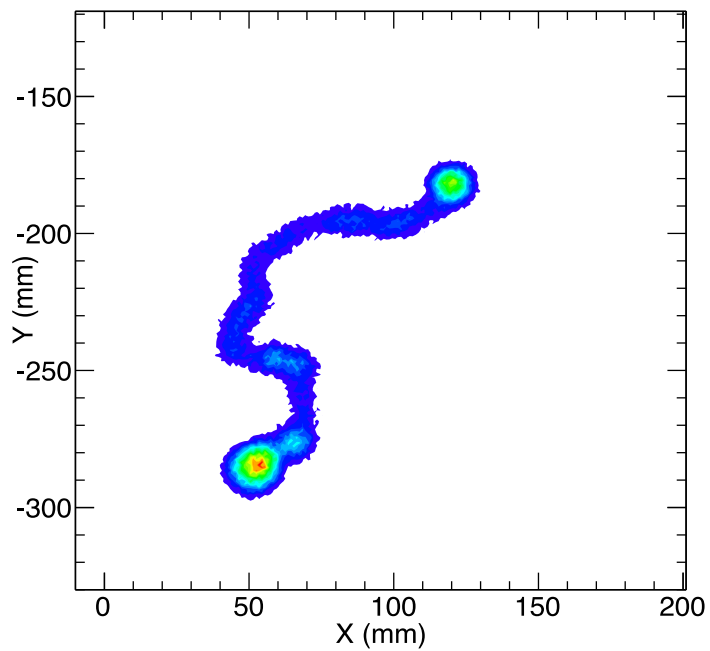


NEXT-DEMO (Valencia)

# Tracking: PMTs $\rightarrow$ SiPMs

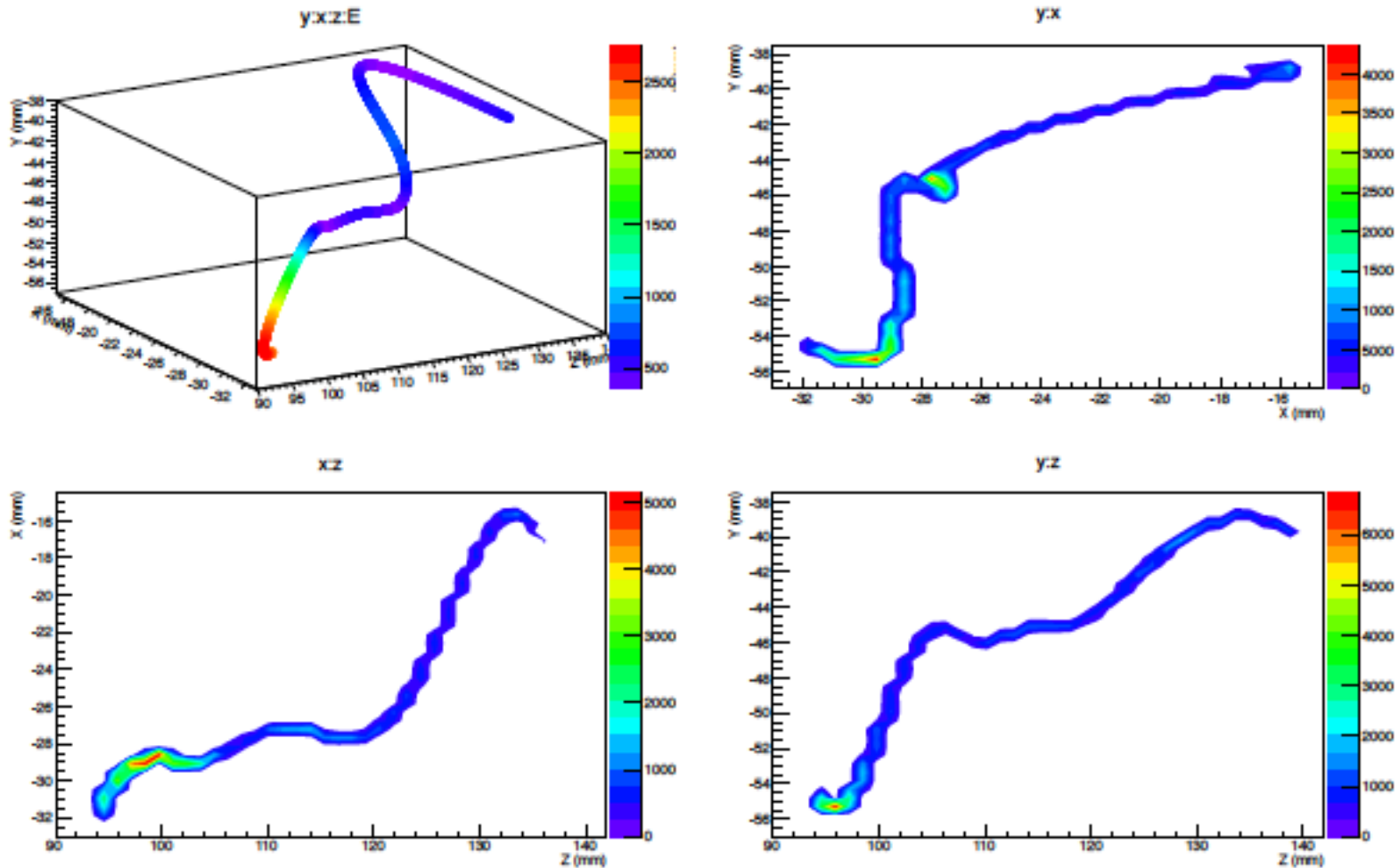


NEXT-DEMO (IFIC, Valencia)



Simulations!

# Real track from $^{137}\text{Cs}$ $\beta$ -ray – reconstructed with SiPMs



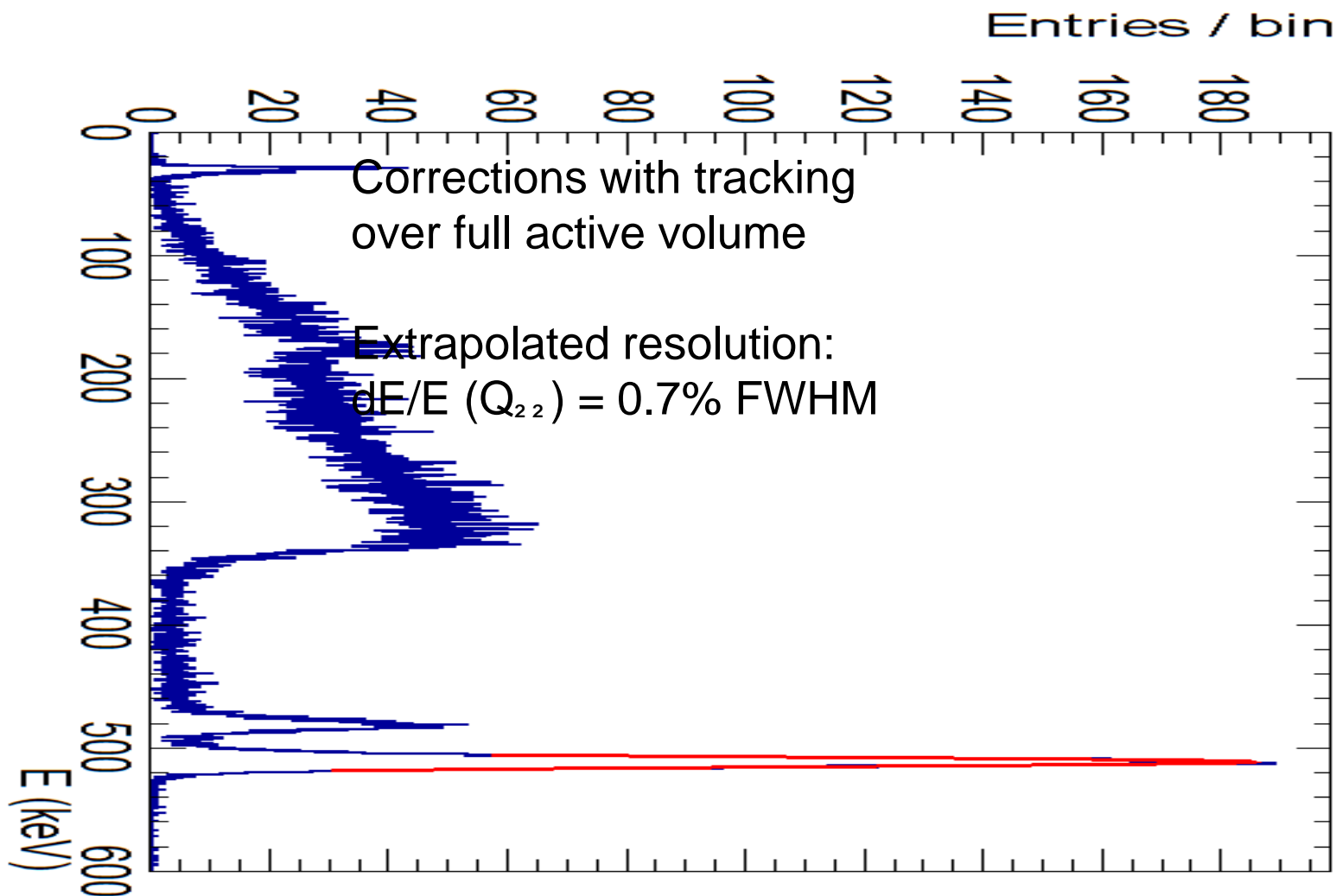
DATA!

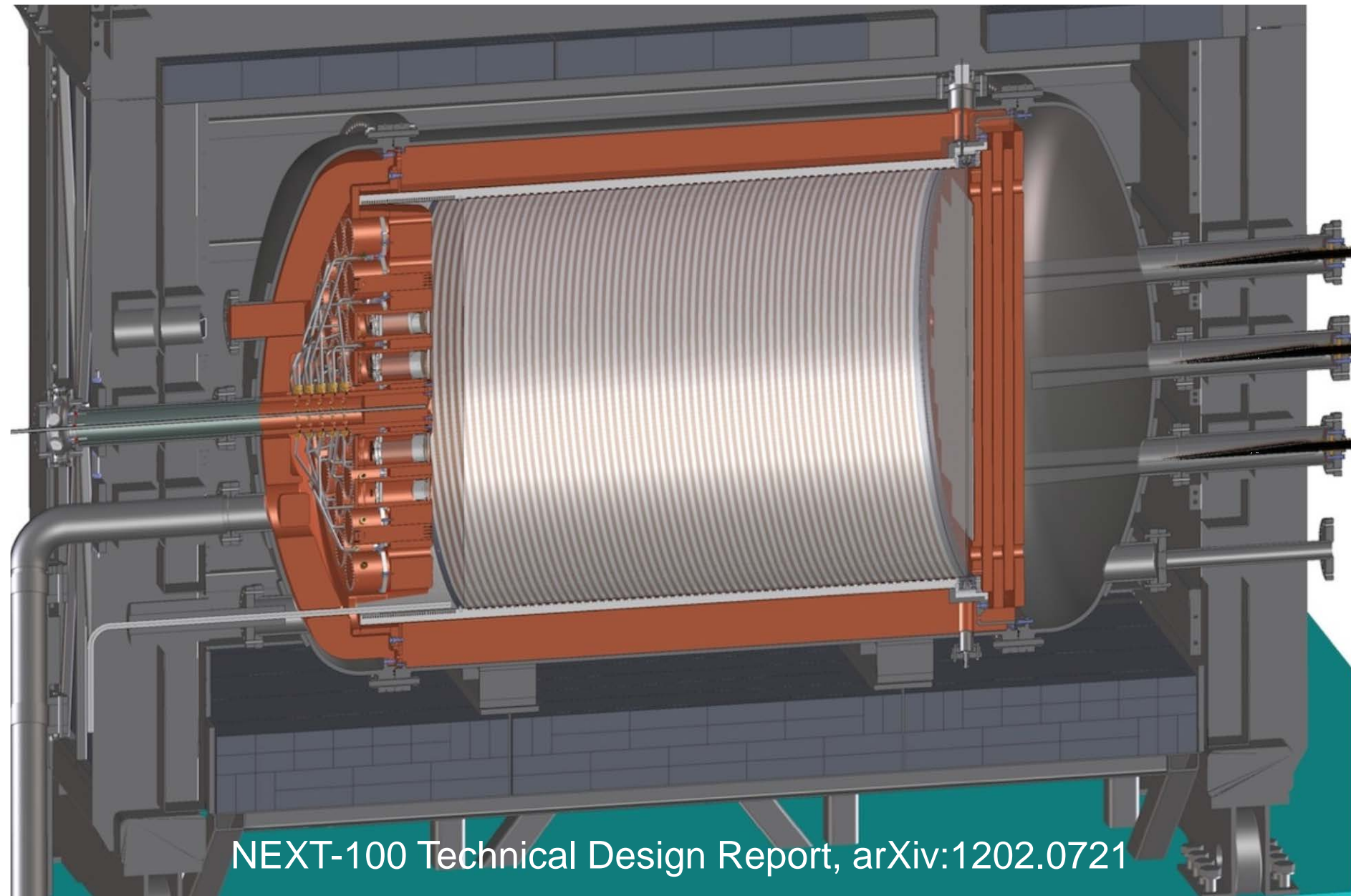
NEXT-DEMO

IFIC, Valencia



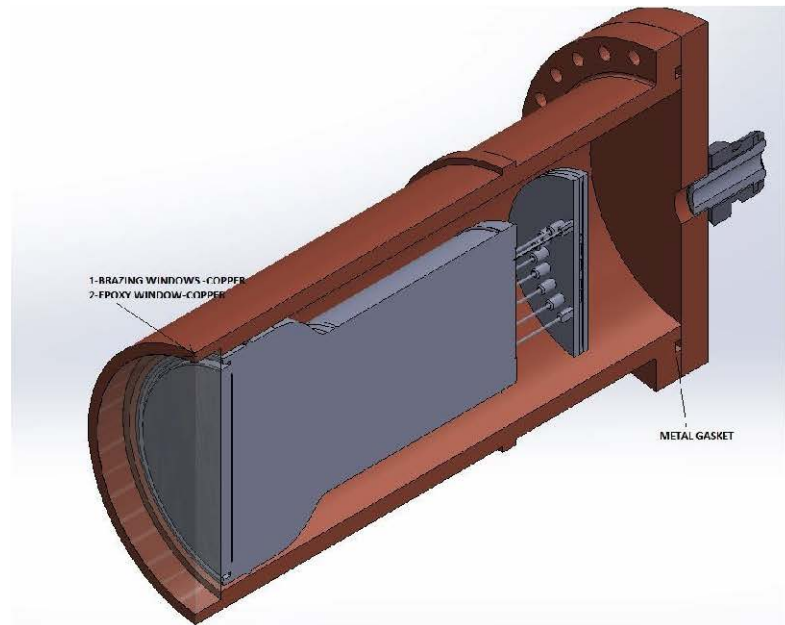
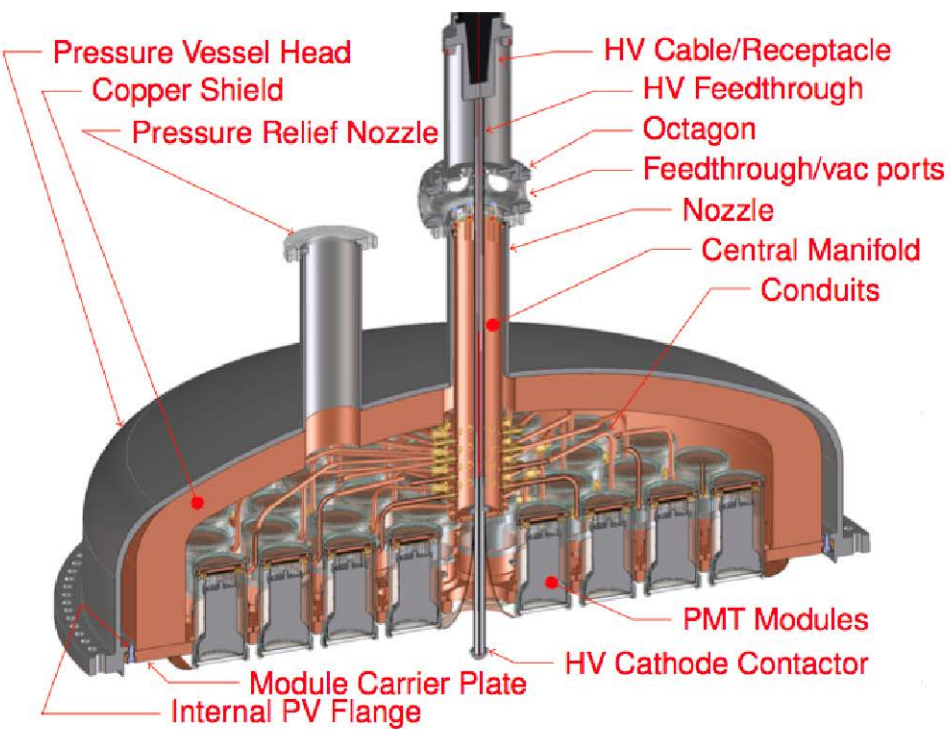
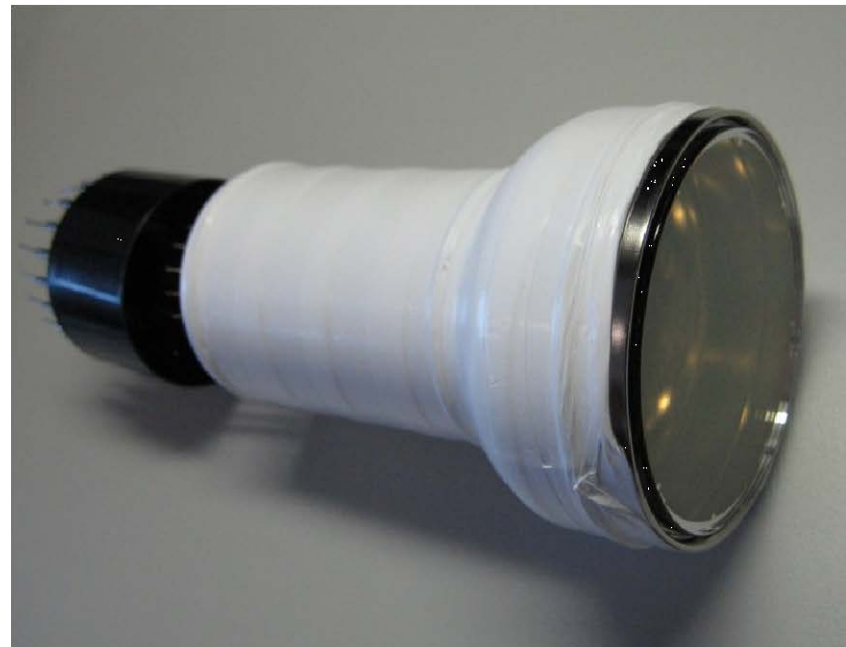
## Energy resolution @ 511 keV: NEXT-DEMO (Valencia)





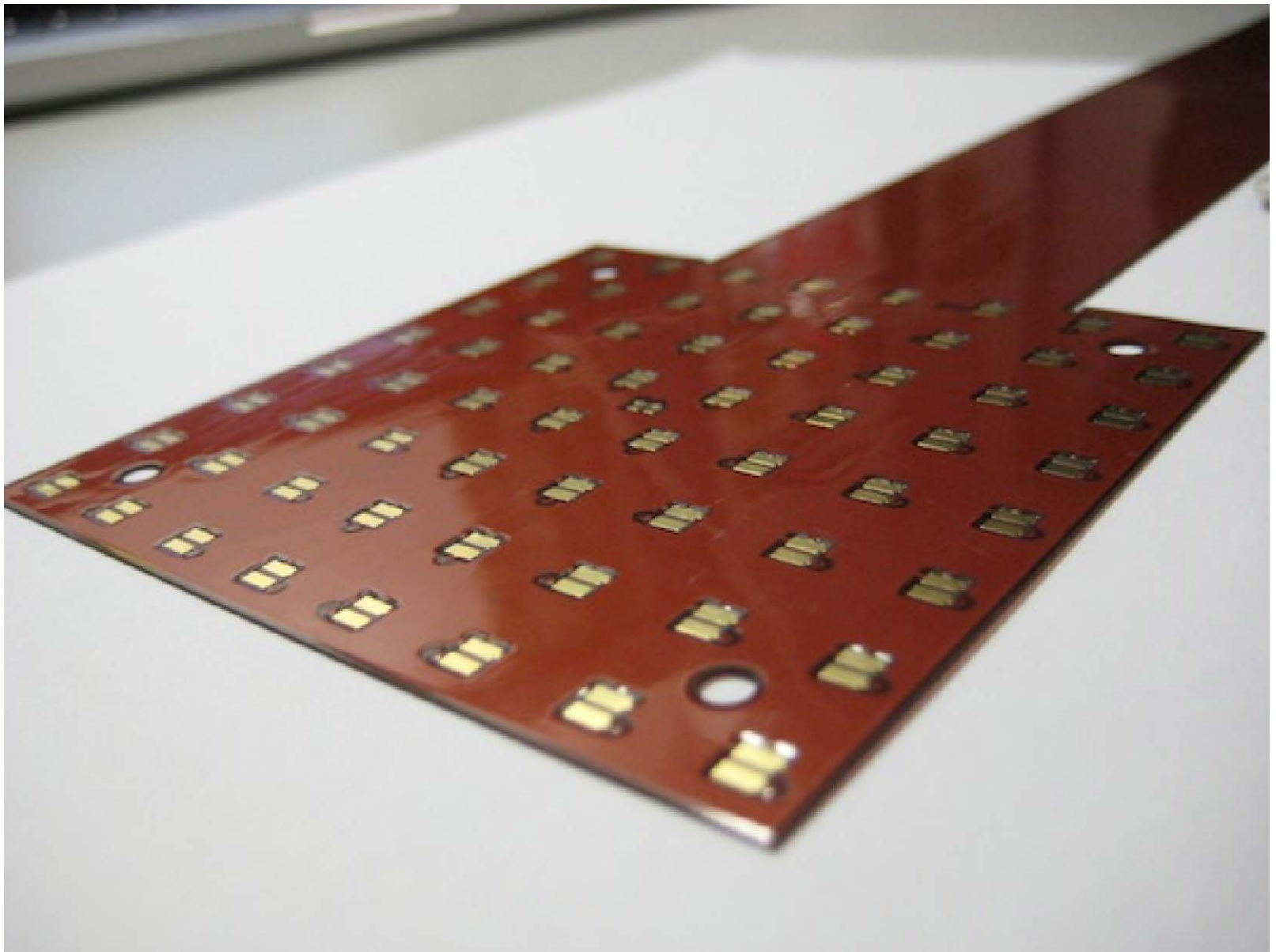
NEXT-100 Technical Design Report, arXiv:1202.0721

# Energy plane



NEXT-100 Pressure Vessel

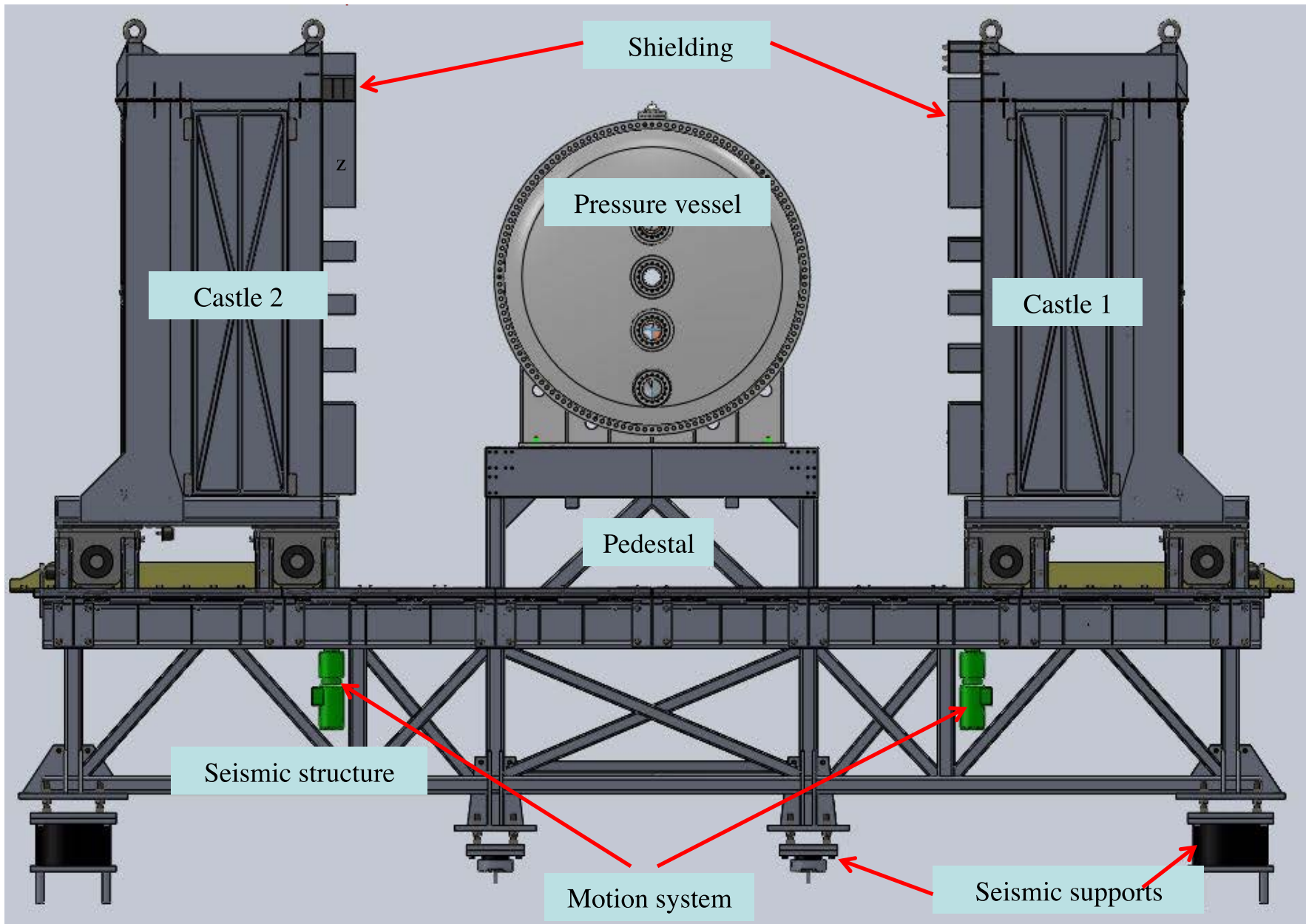




# NEXT at Laboratorio Subterranneo de Canfranc (LSC), Spain



The experiment will be located at Hall A at the LSC. Working platform and basic gas system already in place.



# Pure xenon gas: $0\nu\beta\beta$ :

- 1% FWHM energy resolution verified by NEXT-DBDM at 662 keV  $\rightarrow$  0.5% FWHM @ Q-value
  - Dangerous  $^{214}\text{Bi}$   $\gamma$ -ray at 2448 keV
- Track reconstruction with SiPMs verified; confirms x 30 – 50 background rejection
  - Background rate:  $4 \times 10^{-4}$  counts/keV/kg/year
- NEXT-100, with 100 kg enriched  $^{136}\text{Xe}$ , should “touch” inverted hierarchy if backgrounds are as low as measurements + simulations



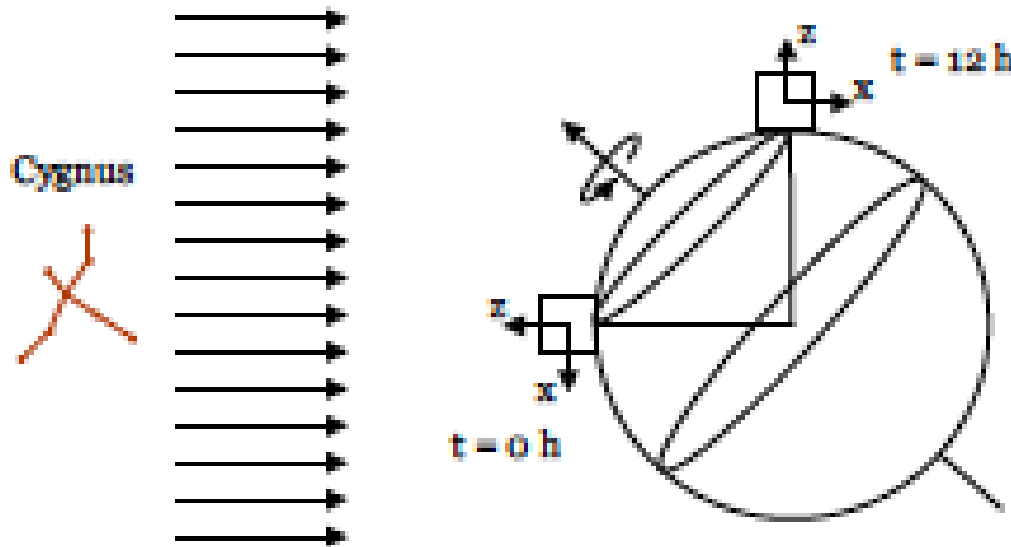
# Criteria for “Discovery”

- What criteria should be established and met for a claim of “discovery” of  $0\nu\text{-}\beta\beta$  decay?
  - Excellent energy resolution has been insufficient...
  - Is *event topology* an essential component?
- What criteria should be established and met for “discovery” of WIMP dark matter?
  - Backgrounds lurk everywhere at keV energies...
  - Is “recoil directionality” an essential criterion?

# Directional sensing for nuclear recoils?

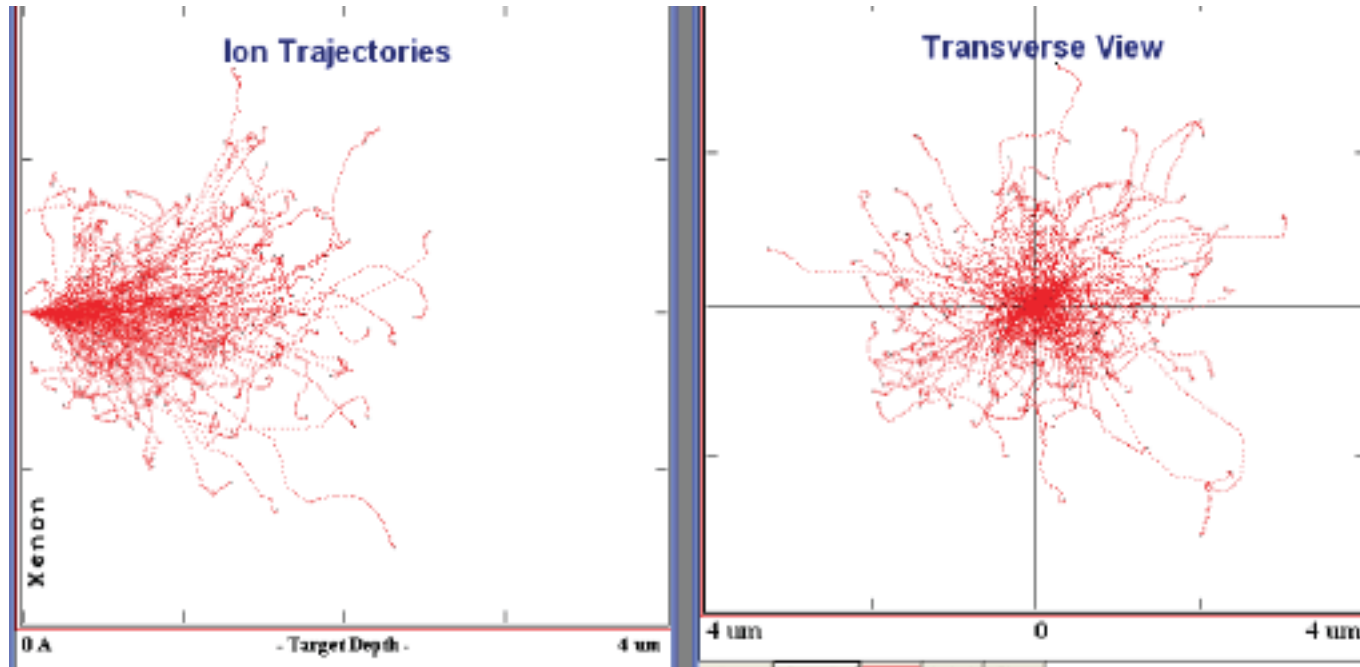
A sidereal variation of “WIMP wind from Cygnus”

WIMP  $\langle V \rangle$  comparable to earth’s velocity:  $\sim 230$  km/s



A substantial anisotropy in nuclear recoils should be observable

# Do nuclear recoils retain directionality?



**SRIM:** 200 Xenon 30 keV nuclear recoil events in HPXe Xenon – **unweighted** by energy loss

# Target Ionization

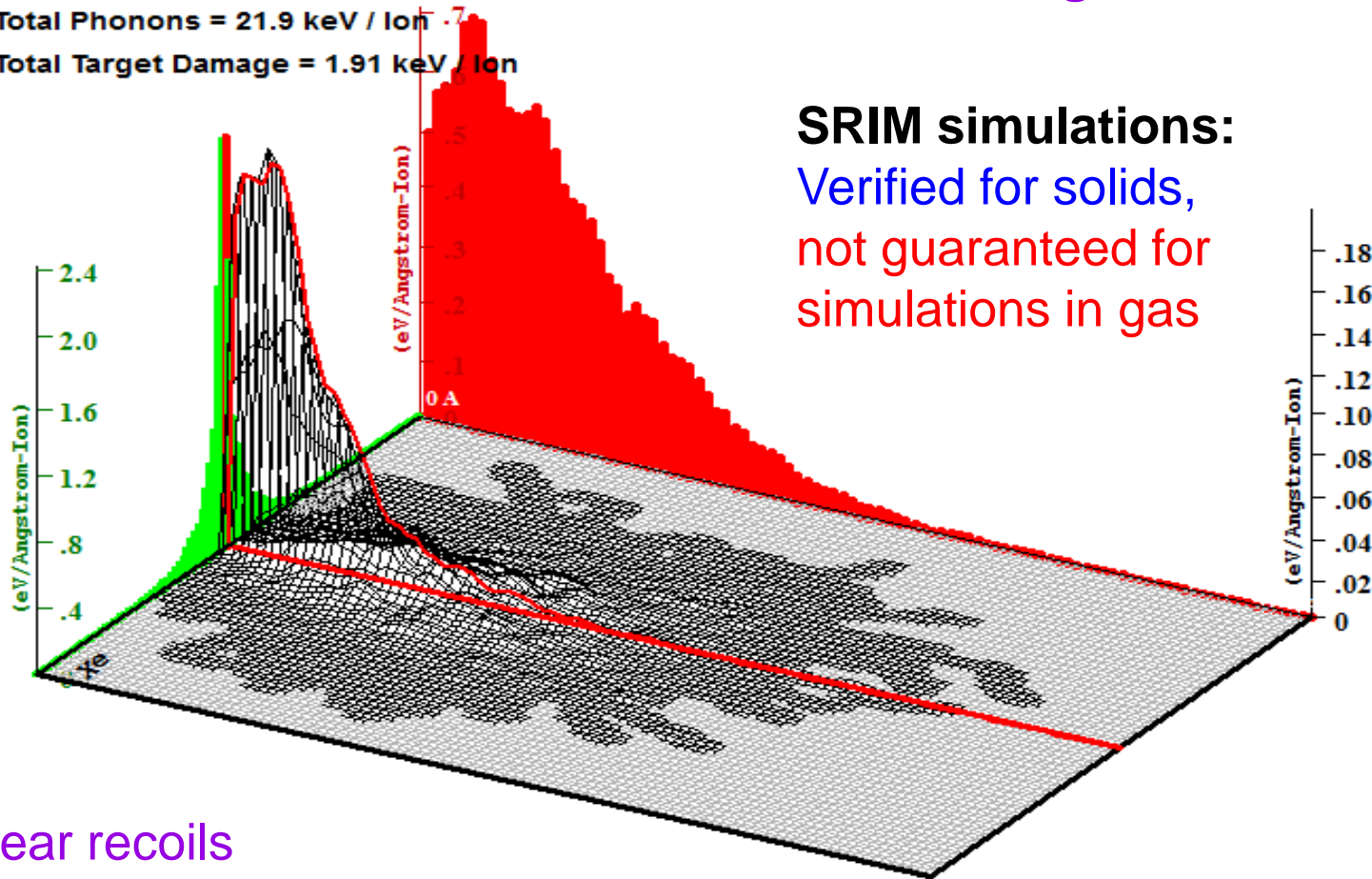
Total Ionization = 6.2 keV / Ion

Total Phonons = 21.9 keV / Ion

Total Target Damage = 1.91 keV / Ion

30 keV Xe ions in  
10 bars xenon gas

**SRIM simulations:**  
Verified for solids,  
not guaranteed for  
simulations in gas



Nuclear recoils  
are messier  
than  $\pm$ -particles

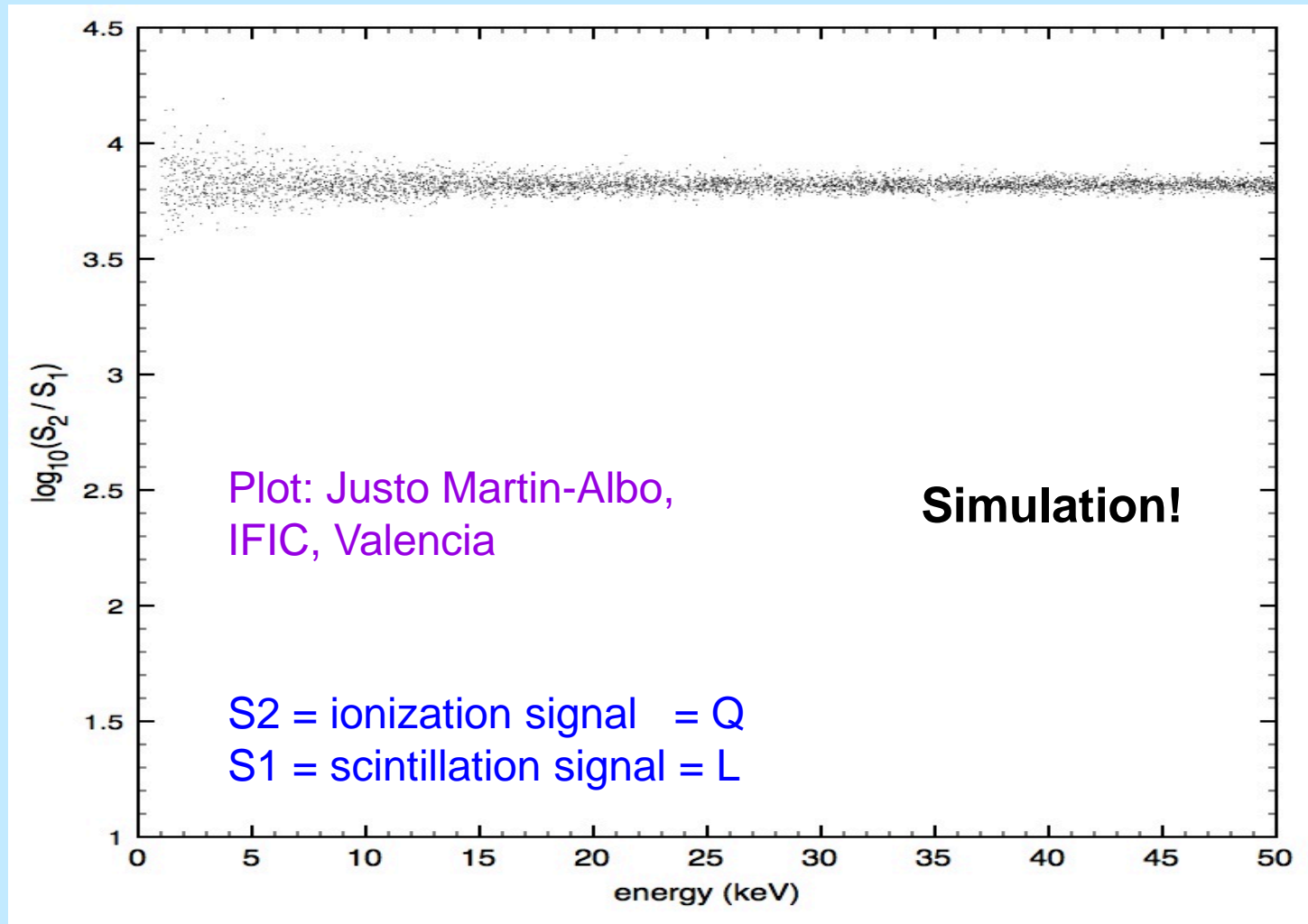
Plot Window goes from 0 A to 4 um; cell width = 400 A  
Press PAUSE TRIM to speed plots. Rotate plot with Mouse.

**Ion = Xe (30. keV)**

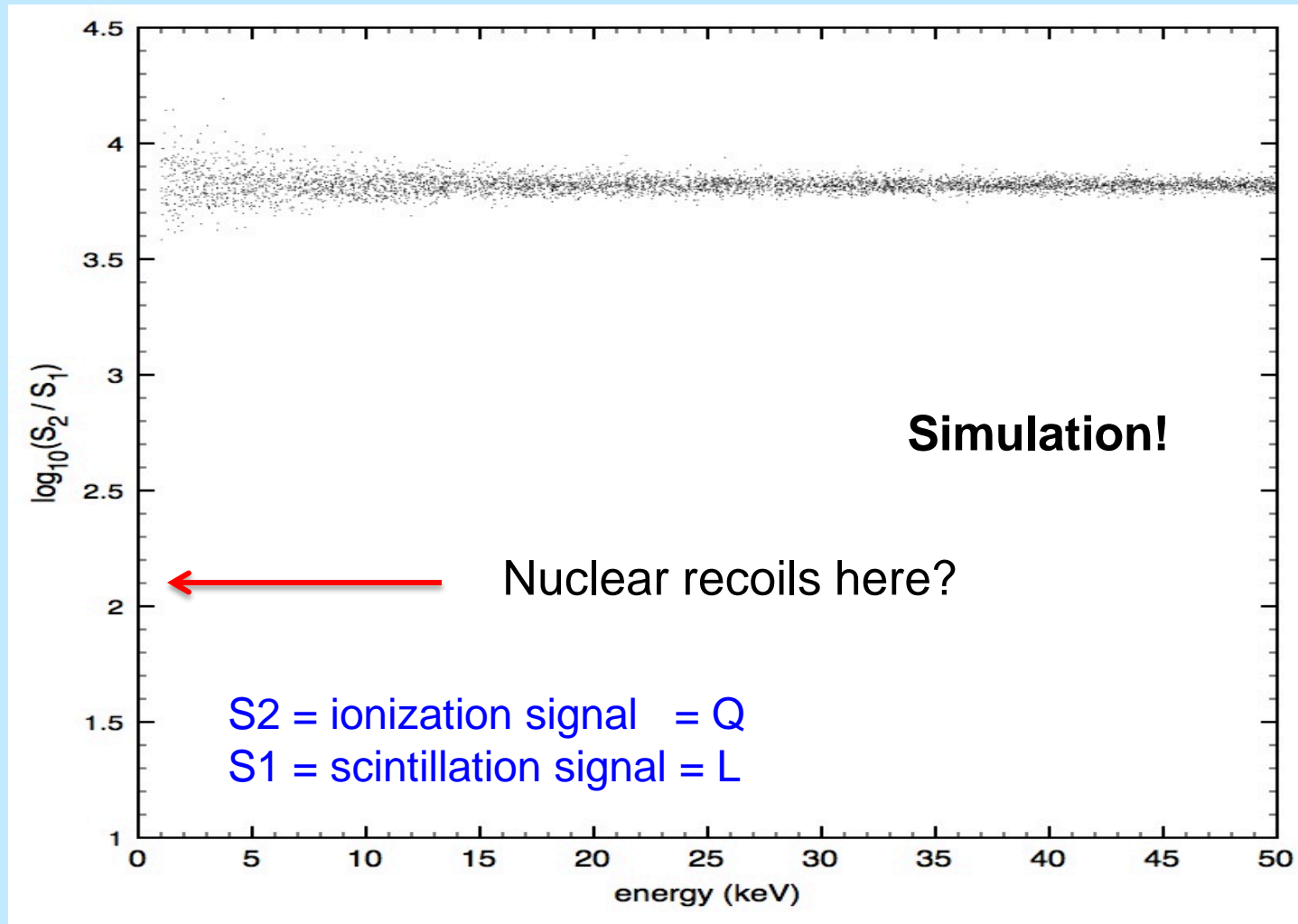
# Dark Matter Search with HPXe

- The superb energy resolution available, in principle, also helps the WIMP search
  - Intrinsic S2/S1 fluctuations are much smaller
- Gas phase permits molecular additives that offer remarkable performance opportunities
  - Beneficial for  $0\nu\text{-}\beta\beta$ , too, but no time to discuss

# Simulation: electron recoils in pure HPXe, $F = 0.15$ , 10% optical efficiency



# Simulation: electron recoils in pure HPXe, $F = 0.15$ , 10% optical efficiency



# Today's techniques

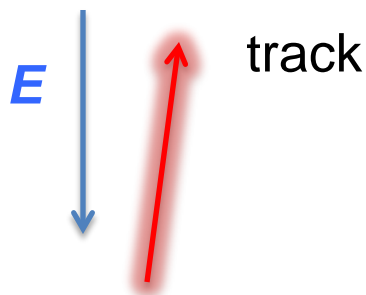
- All current approaches attempt to **visualize** the nuclear recoil track:
  - **Nuclear Emulsions** (expanded optical readout !! )
- **Low-pressure TPCs** (~50 g/detector)
  - **D<sup>3</sup>** (GEM + pixel ASIC)
  - **DMTPC** (CF<sub>4</sub> optical CCD)
  - **DRIFT** (CS<sub>2</sub><sup>-</sup>, CF<sub>4</sub>, MWPC )
  - **MIMAC** (CF<sub>4</sub> + μMegas )
  - **NEWAGE** ( Gem, μ-dots )



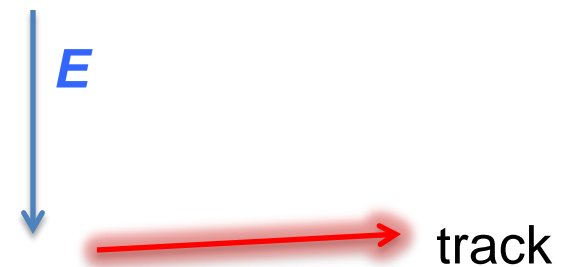
# Columnar Recombination: a nuisance?

- or a new way to “see” directionality?

- Columnar Recombination (CR) occurs when:
  - A drift electric field  $E$  exists;
  - Tracks are highly ionizing;
  - Tracks display an approximately linear character;
  - The angle between  $E$  and track is small:



**Substantial CR: more light, less charge**



**CR small: less light, more charge**





# Columnar recombination and Directionality sensing in nuclear recoils

- **Columnar recombination** (CR) can be quite sensitive to the **angle** between a highly ionizing track and an electric field  $E$ ;
- For a given event energy, more recombination would yield more scintillation, less ionization
- Therefore, a comparison – event by event – of **scintillation/ionization** is a measure of CR

## CR Exists!

Evidence for columnar recombination in  $\pm$ -particle tracks in dense xenon gas.

FWHM depends on E-field and density!

Bolotnikov & Ramsey  
NIM A 428 (1999)  
pp 391-402

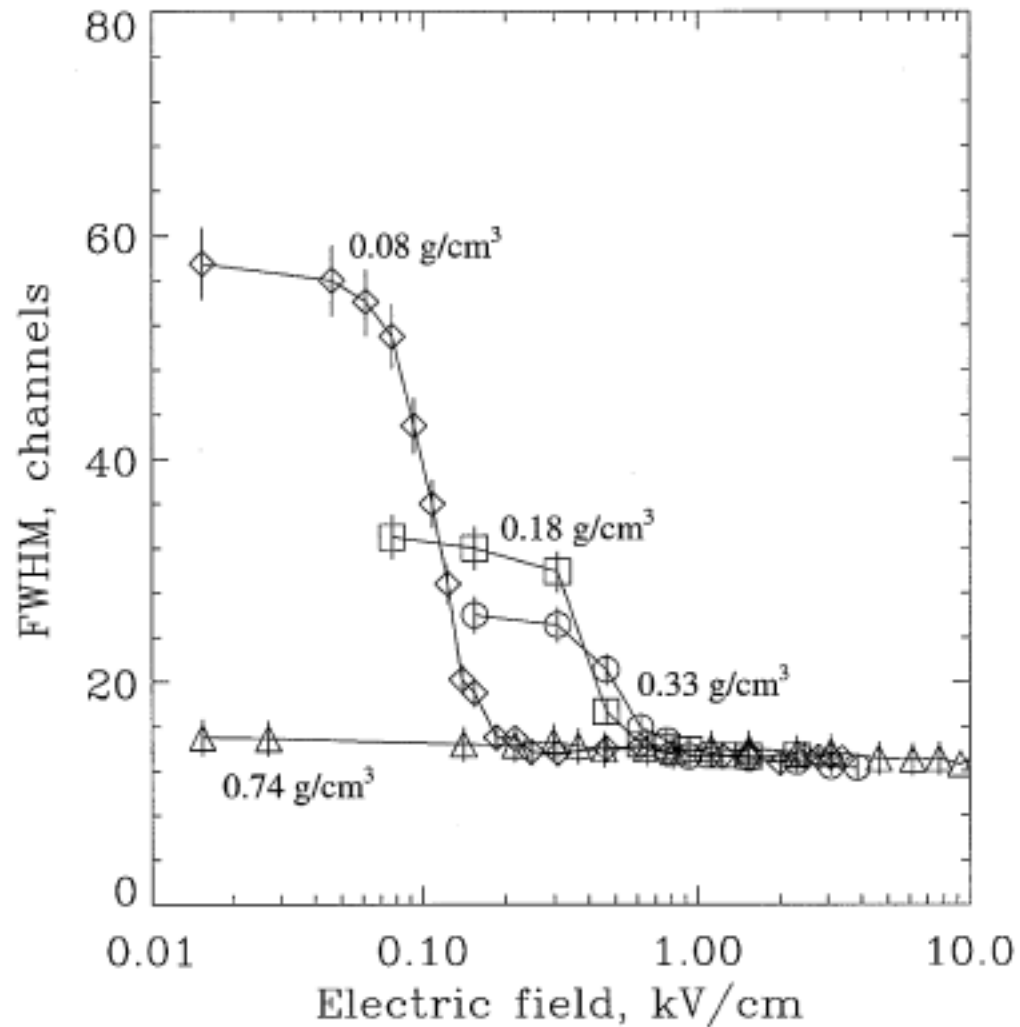


Fig. 5. FWHM of the peaks in pulse-height spectra of the amplitude of the light signals versus the electric field strength measured at  $0.08 \text{ g/cm}^3$  (diamonds),  $0.18 \text{ g/cm}^3$  (squares),  $0.33 \text{ g/cm}^3$  (circles), and  $0.74 \text{ g/cm}^3$  (triangles).

# What is the optimum Xe density?

- Define (*electrostatic*) **Columnarity:  $C$**
- $C = \mathcal{R}/r_0$
- $\mathcal{R}$  = the nuclear recoil track *range*
- $r_0$  = Onsager radius  $r_0 = e^2/\epsilon \bar{E}$ , where  $\bar{E}$  is electron energy (usually taken as kT)
  - in xenon gas for  $\rho \approx 0.05 \text{ g/ cm}^3$ :
    - $R_0 \sim 70 \text{ nm}$
    - $\mathcal{R} \sim 2100 \text{ nm}$  for 30 keV nuclear recoil
    - $C \approx 30$  in this example
  - Hopeless for liquid density:  $C < 1$

# Columnarity is key

*We want  $C$  to be fairly large, i.e.  $C > 10$*

- This condition is probably met for  $KE \geq 20$  keV in xenon gas for  $\rho \approx 0.05$  g/cm<sup>3</sup>, or less
- Figure of Merit  $M = V_{\text{det}}/V_{\text{track}} = 10^{17}$  per m<sup>3</sup>
- CR  $M$  is better than low-density TPC by  $\times 10^9$

# Molecular gymnastics can help

- Primary excitations  $\sim$  ionization
  - Excitations carry no directional information!
    - Convert excitations to ionization by Penning effect
      - Use appropriate molecular additive – which one?
      - **Trimethylamine** (TMA) displays a strong Penning effect in Xe
  - Molecular additive:
    - will cool electrons – facilitates CR
    - Neutralizes xenon ions by charge exchange
    - Track “image” transformed to molecular ion image
    - Molecular ions recombine with electrons  $\rightarrow$  “light”



# How to maximize a CR signal...

- Large size of gas-phase TPC requires optical detection by wavelength-shifting plastic (WLS)
  - WLS: maximum efficiency at 300 nm
  - WLS: negligible efficiency at 173 nm VUV of xenon
- **Miracle needed**: Penning molecule must display efficient UV fluorescence at  $\sim 300$  nm
- **Providence**: **trimethylamine** is known to fluoresce very efficiently at  $\sim 300$  nm !!!

# Scenario

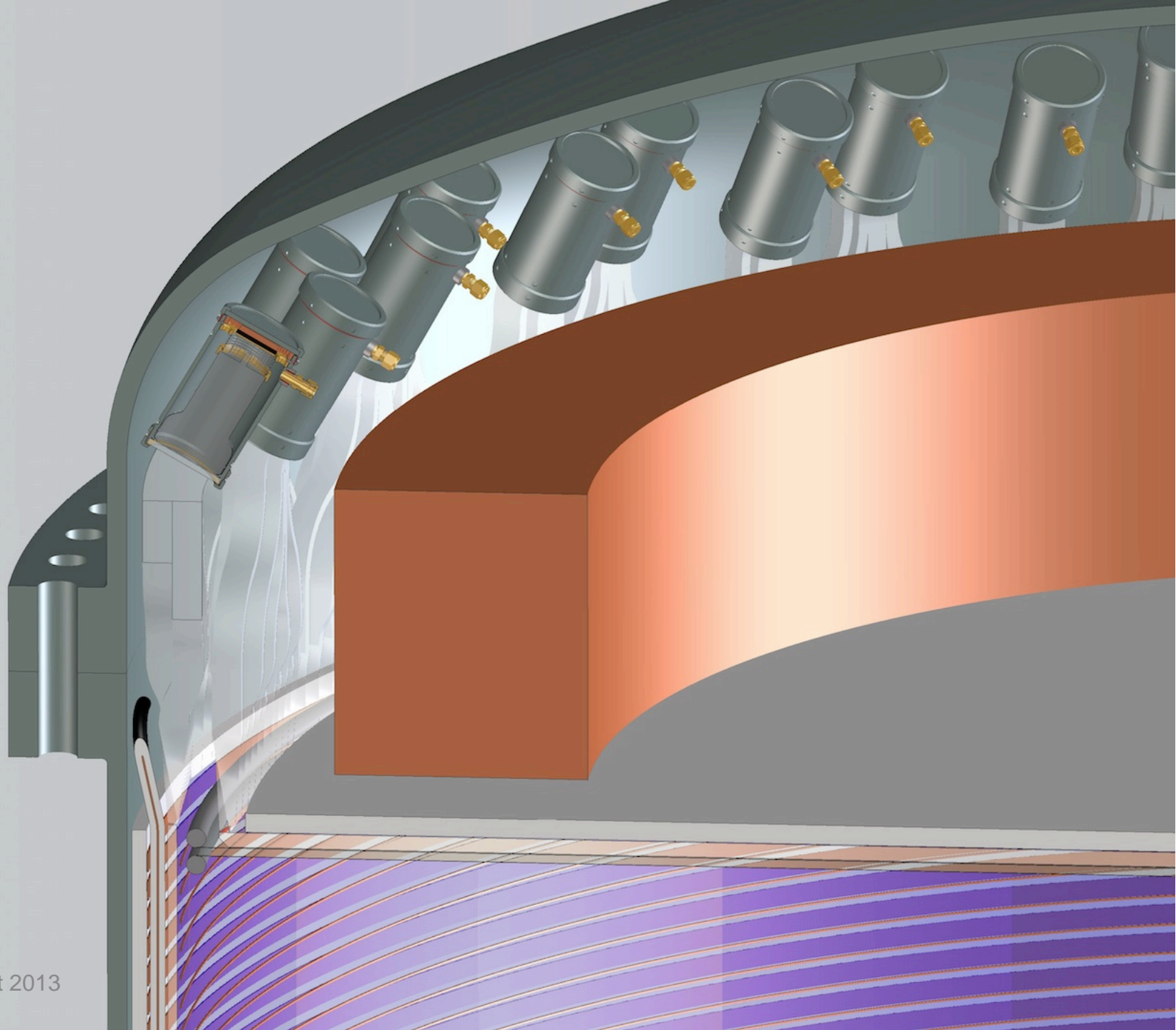
- “S1” signal may display strong *columnar recombination*
  - A substantial effect for nuclear recoils (~100x minimum ionizing)
  - A negligible effect for electron recoils
- This may provide a way to “see” WIMP directionality *without direct imaging* of nuclear recoil tracks
  - Density restriction on gas is moved to ~10 bars
  - X100 increase relative to low density TPC concept
  - Drift length restriction due to diffusion is removed
  - Simpler spatial detection requirements at anode plane
  - Larger monolithic detector possible, x10 – 100 volume
- *Several hundred kg active mass possible, if true!*



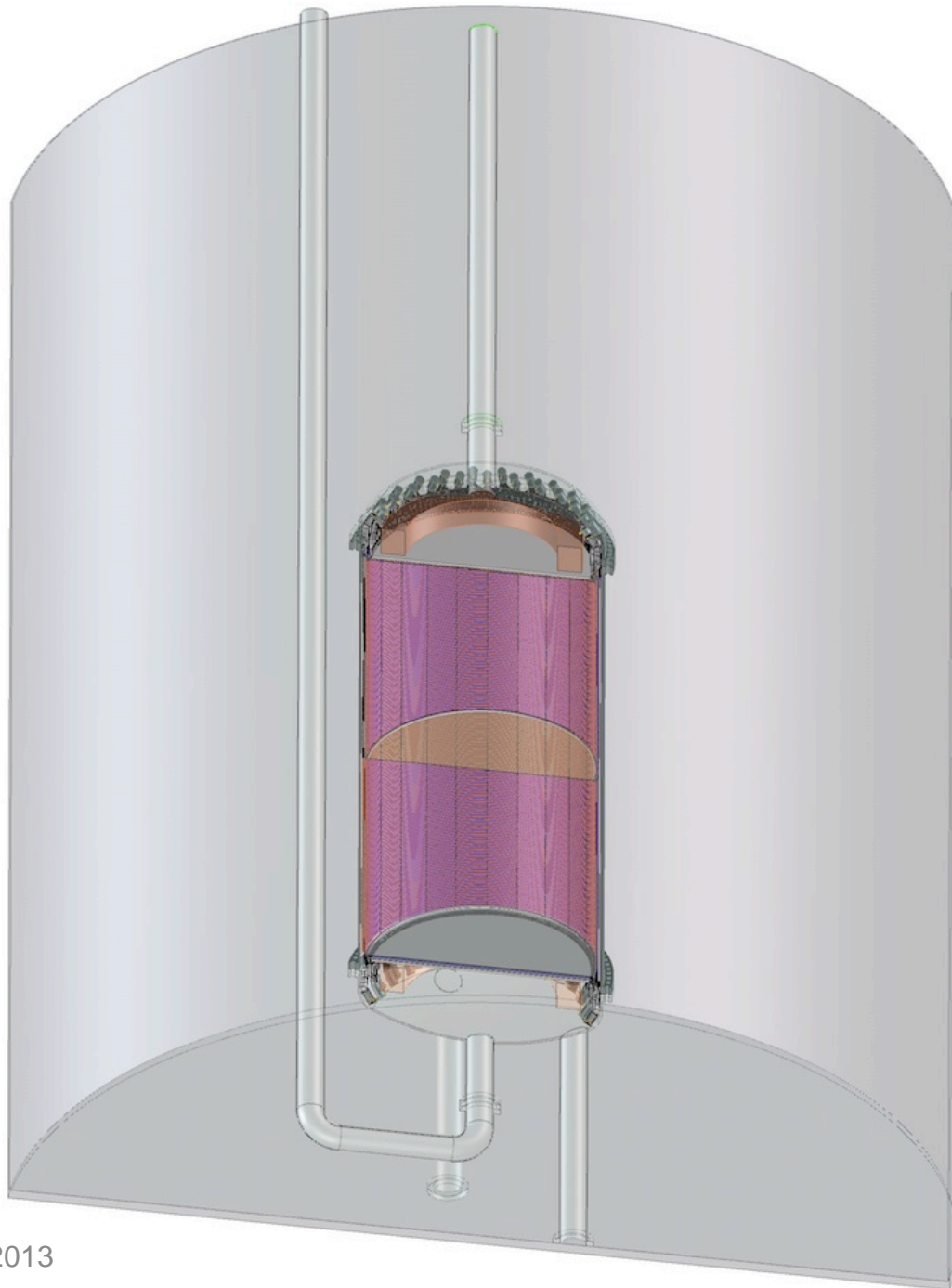
Size: 2.3 meter diameter  
2 x 3 meter drift length

With WLS, only a few  
dozen PMTs are needed

With E-field in opposing  
directions, a “head-tail”  
effect might show up

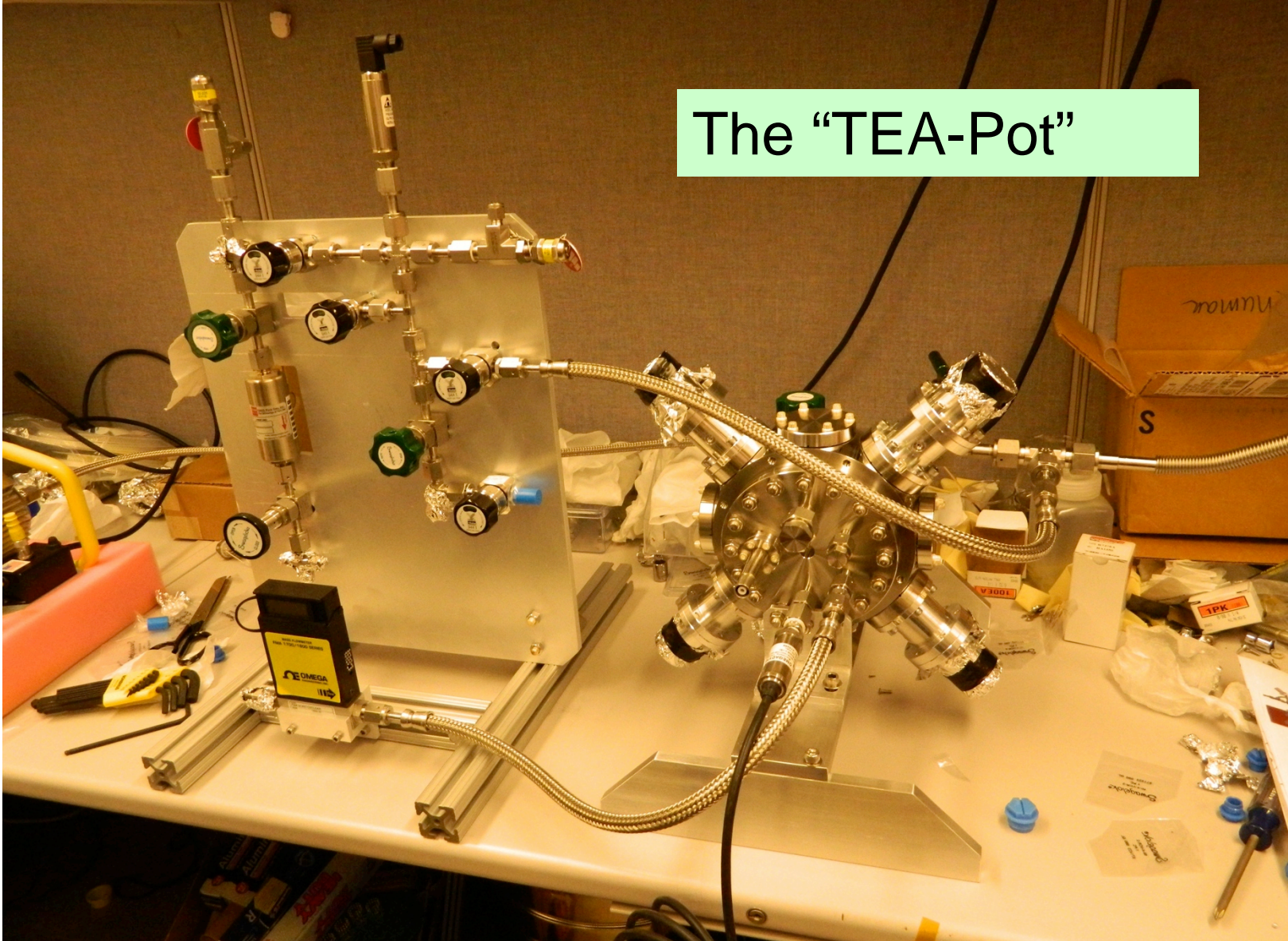


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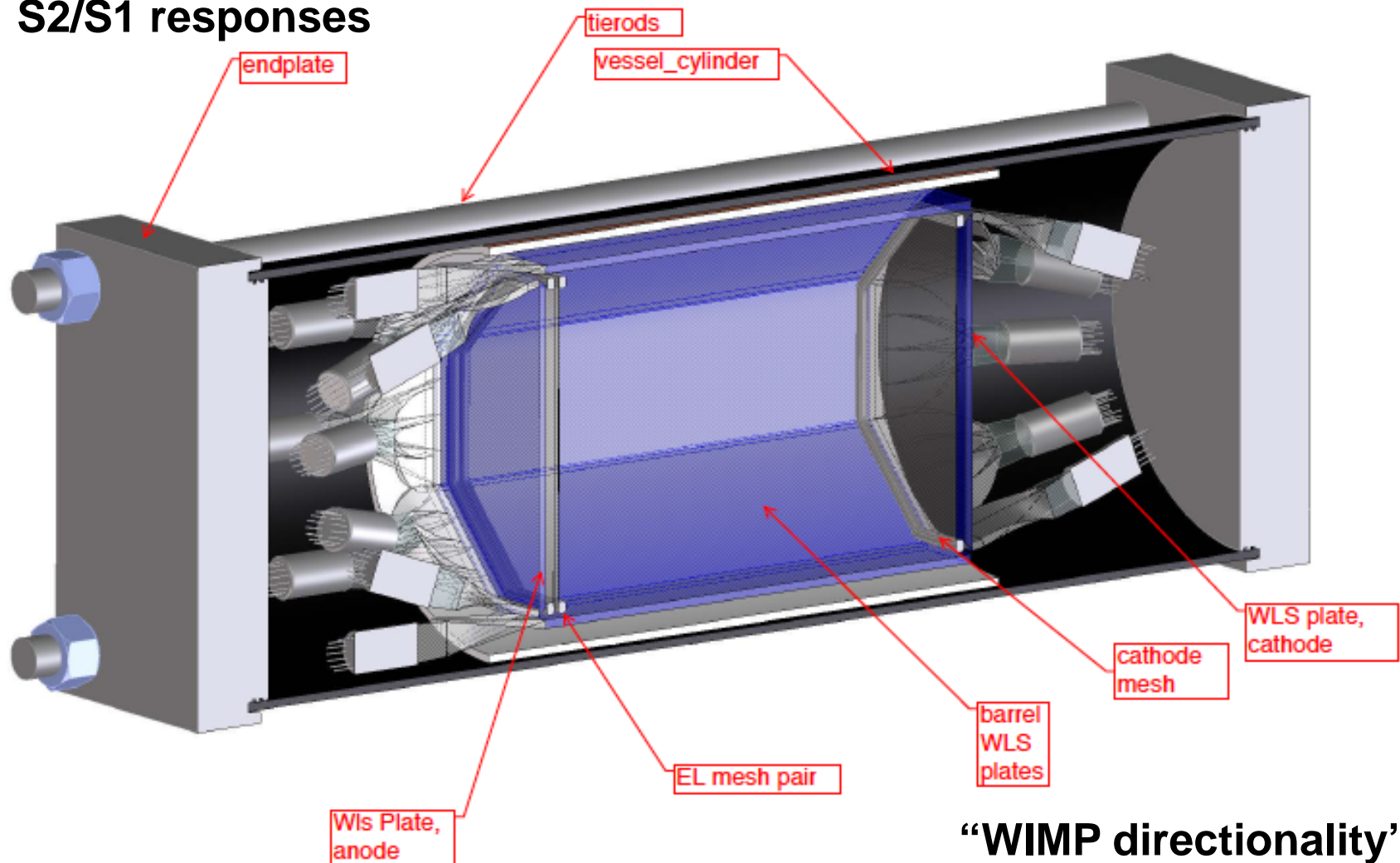
Large, but  
maybe not  
too large...

# The "TEA-Pot"



# OSPREY: “Opportunities for Superior Performance in Rare Event Yields”

## S2/S1 responses



“WIMP directionality”

# Perspective

- Is this a true story, or a fairy tale?
  - At least, serves as imagination stretcher...
- Plausible at each step, but **unknowns** exist:
  - Has Nature chosen WIMP mass: 50 – 350 GeV?
  - Penning efficiency of TMA?
  - Fluorescence efficiency of TMA in recombination?
  - Rate of ionic charge exchange?
  - Cooling rate of electrons after ionization?
  - Head-Tail sensitivity?
- Simulation and experimental effort starting...



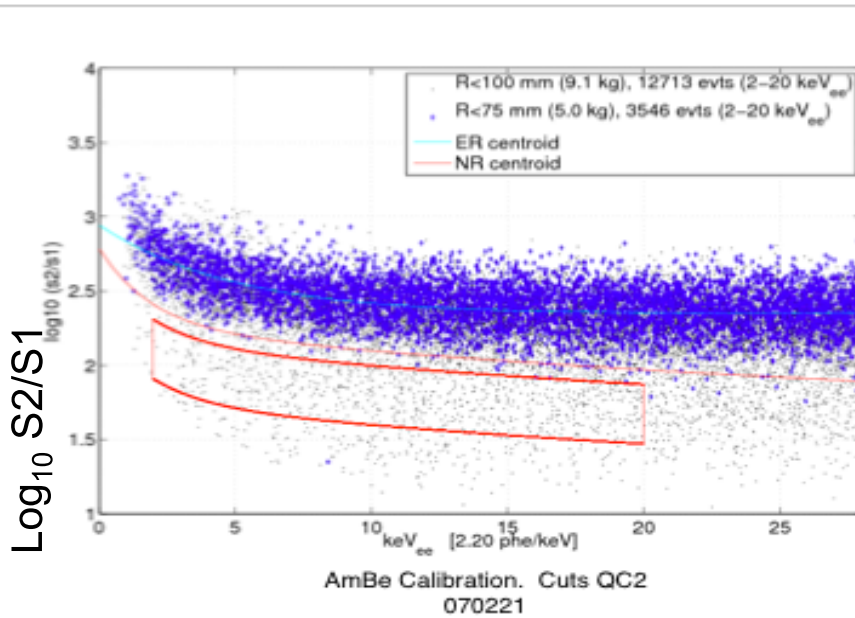
# Perspective

- **Gas phase** offers superb energy resolution, event visualization, and flexibility in operation
- **EL gain stage** is a key element for near-intrinsic energy resolution for  $0\nu\bar{\nu}\nu$  search and low energy signals
- Small energy partition fluctuations imply **superb S2/S1 discrimination** between electron and nuclear recoils
- **Directionality signal** for WIMP search at 100's of kg in monolithic TPC would exceed current reach by  $>1000$

**Thank you**

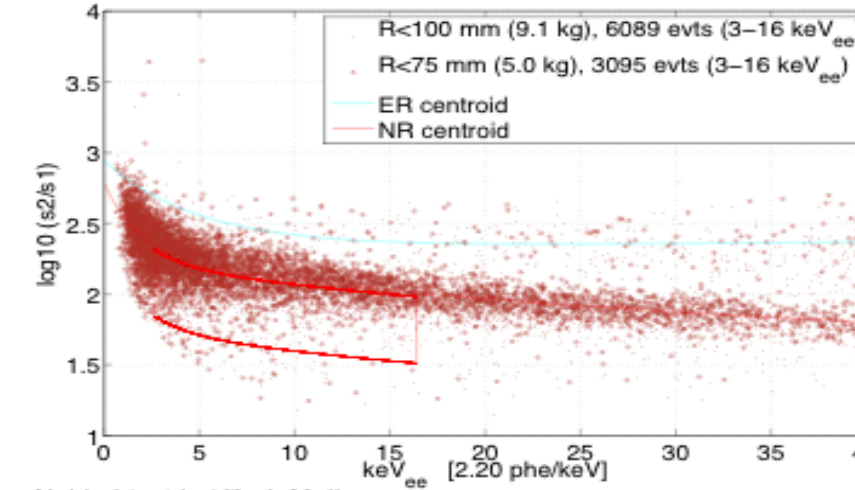
**S2 = Primary ionization signal**  
**S1 = Primary scintillation signal**

## Xenon10 – LXe data



Gamma events (e - R)

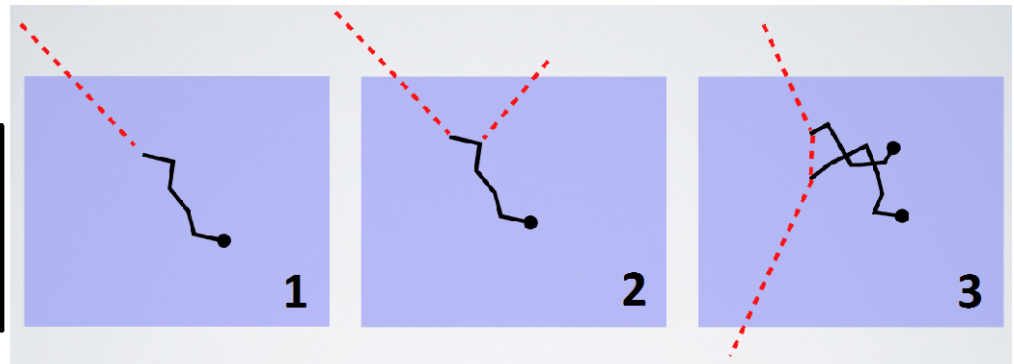
$\gamma$  events show large  $S_2/S_1$  fluctuations at all energies, not improving with energy



Neutron events (N - R)

# Radiopurity: BACKGROUND MODEL

Simulations made with **NEXUS**, a GEANT4 based software developed by NEXT



**Example 1:** Electron photo-produced by 2448 keV gamma from  $^{214}\text{Bi}$  decay

**Example 2:** Electron photo-produced by 2448 keV gamma from  $^{214}\text{Bi}$  decay that undergoes Bremsstrahlung

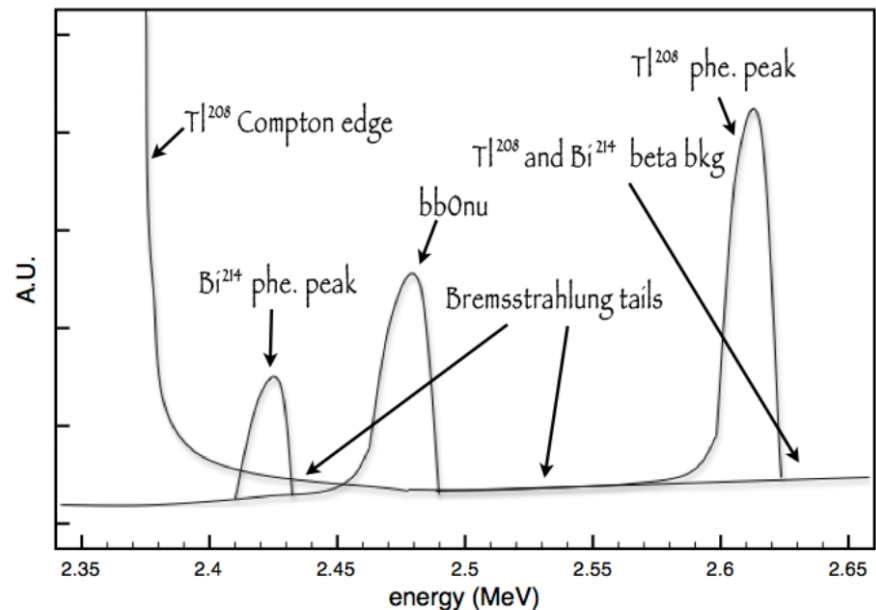
**Example 3:** Two electron Compton scattered from 2615 keV gamma from  $^{208}\text{Tl}$  decay

requirements from Background Model  
(counts/kg/keV/year)

$^{214}\text{Bi}$ : 0.18 - 0.40  $e^{-3}$

$^{208}\text{Tl}$ : 0.21 - 0.48  $e^{-3}$

Total: 0.38 - 0.88  $e^{-3}$



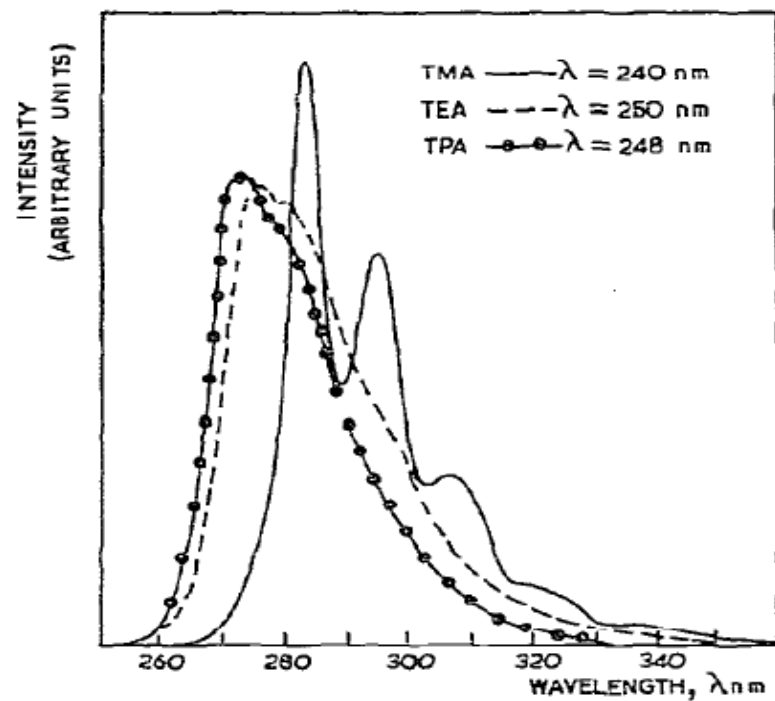
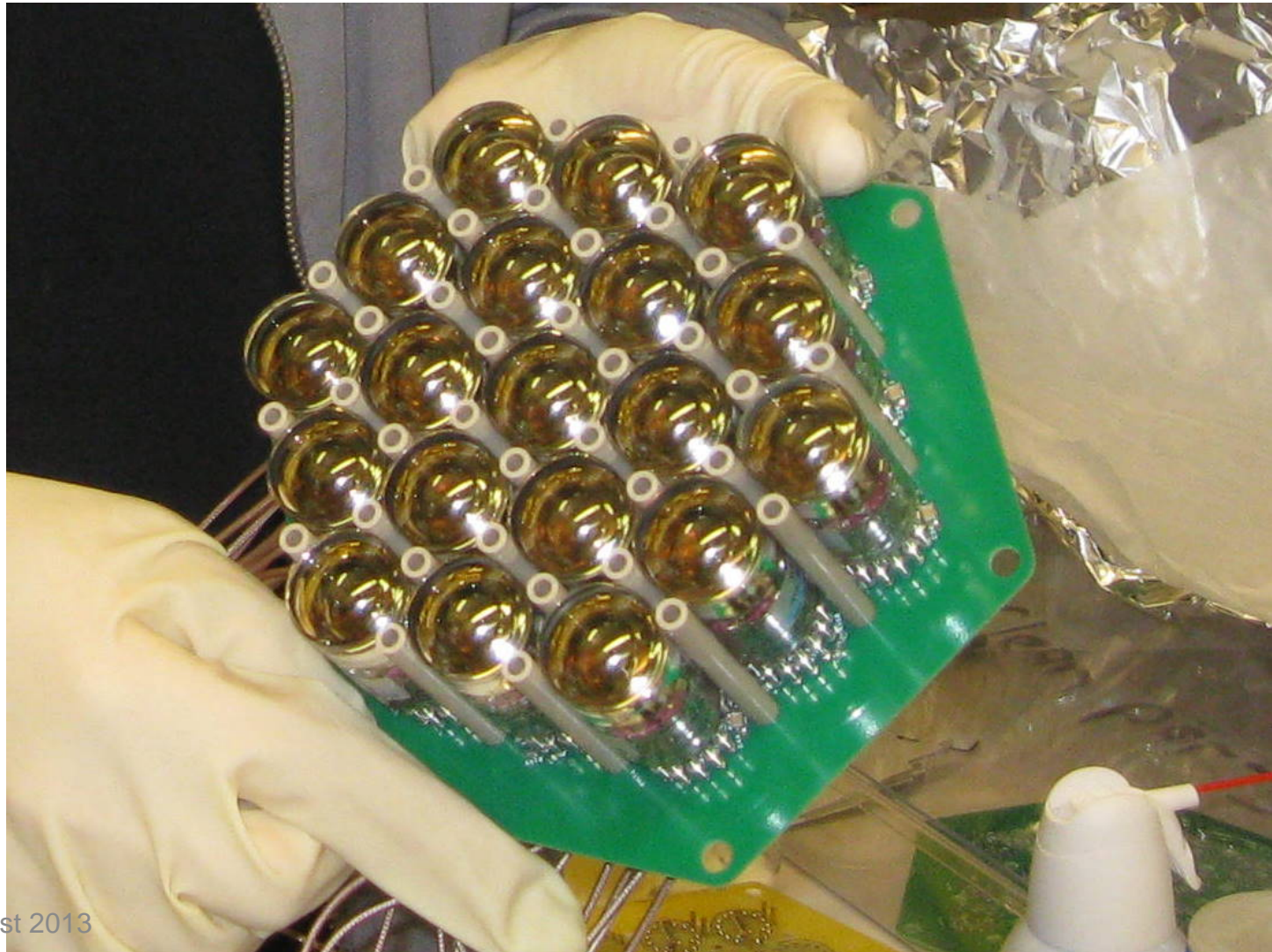


Fig. 4. Vapour-phase fluorescence spectra of TMA, TEA and TPA at excitation wavelengths indicated.



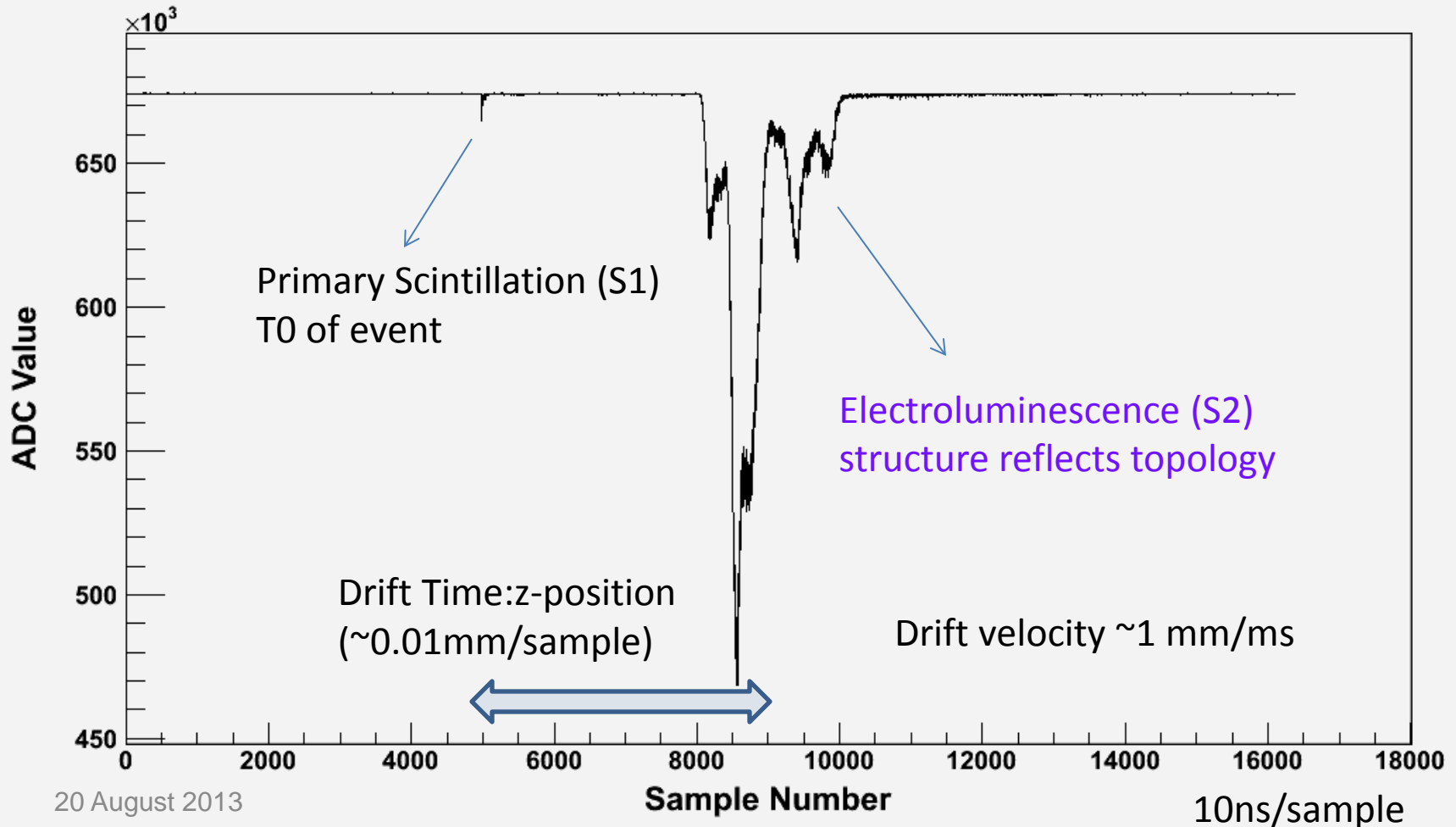
# PMT Array: inside the pressure vessel

Quartz window 2.54 cm diameter PMTs



A typical  $^{137}\text{Cs}$   $\gamma$  waveform (sum of 19 PMTs)  
~300,000 detected photoelectrons

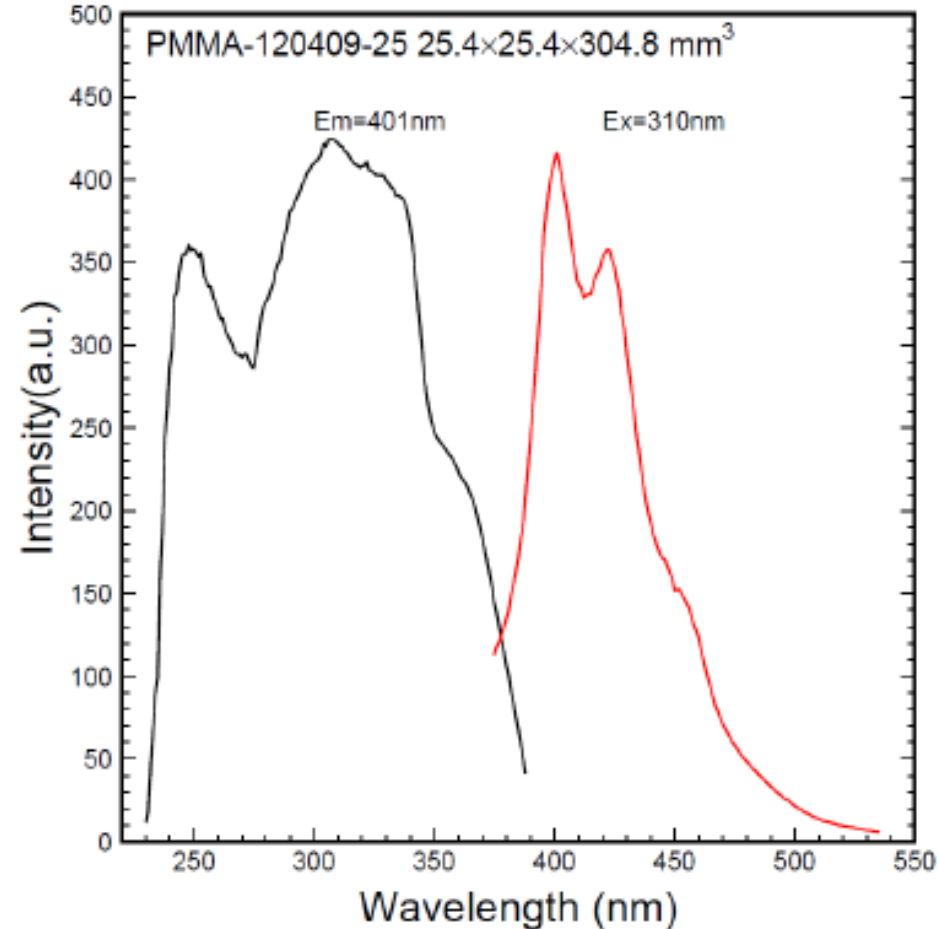
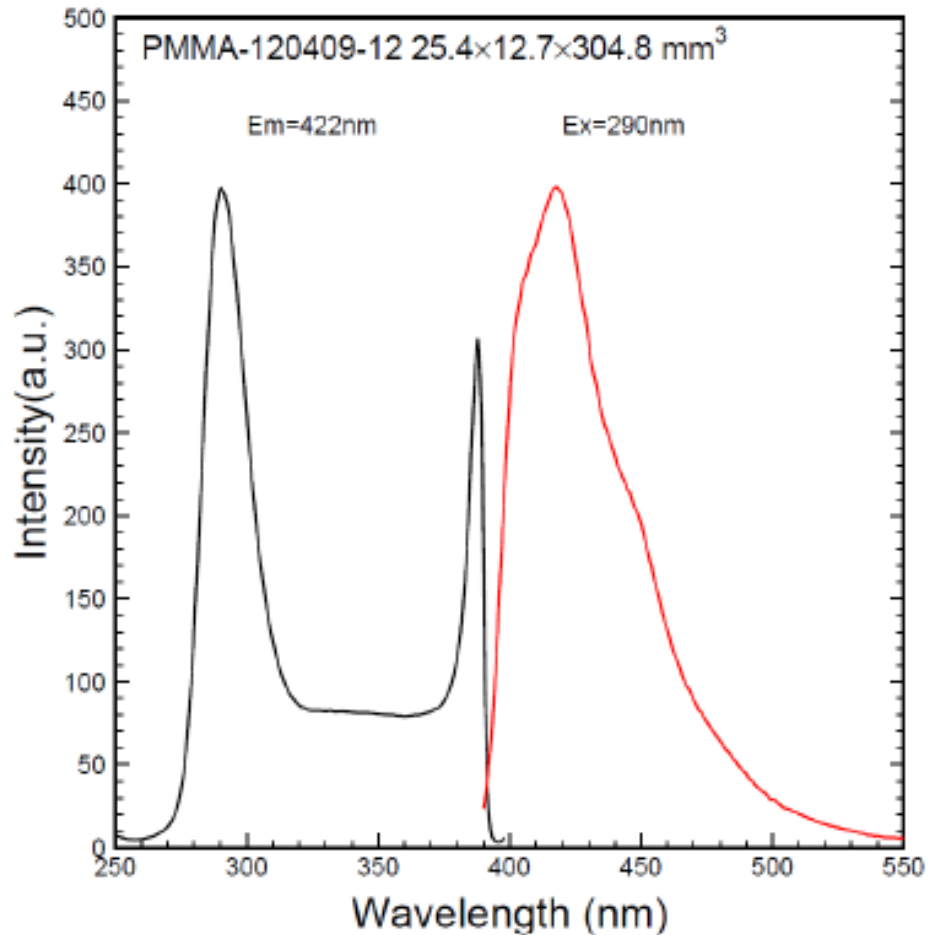
Waveform: Event 488, Channel 22





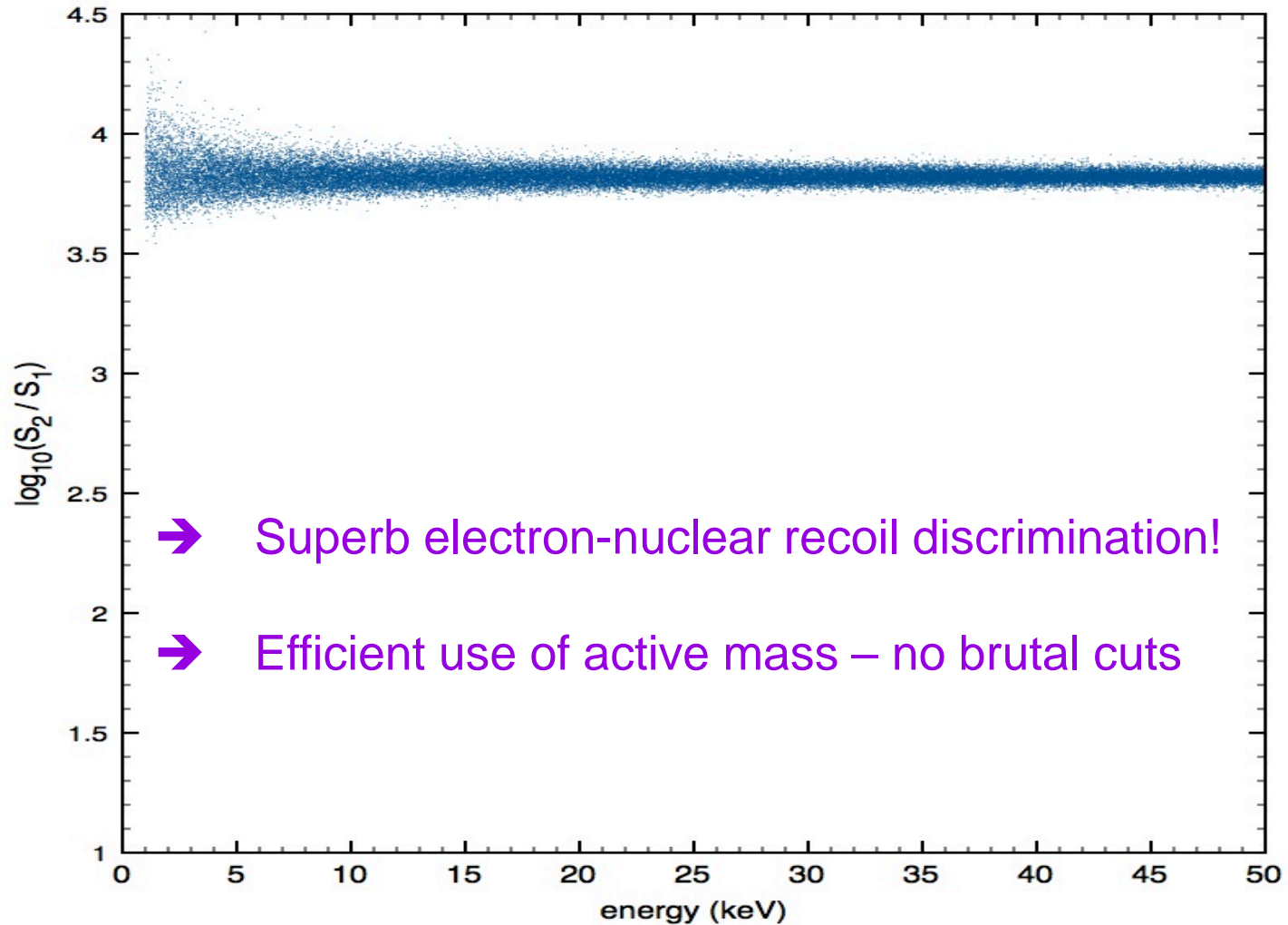
# Photo-Luminescence of PMMA

Different WLS nature observed for two PMMA Samples



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## Gaussian behavior persists at x10 number of events

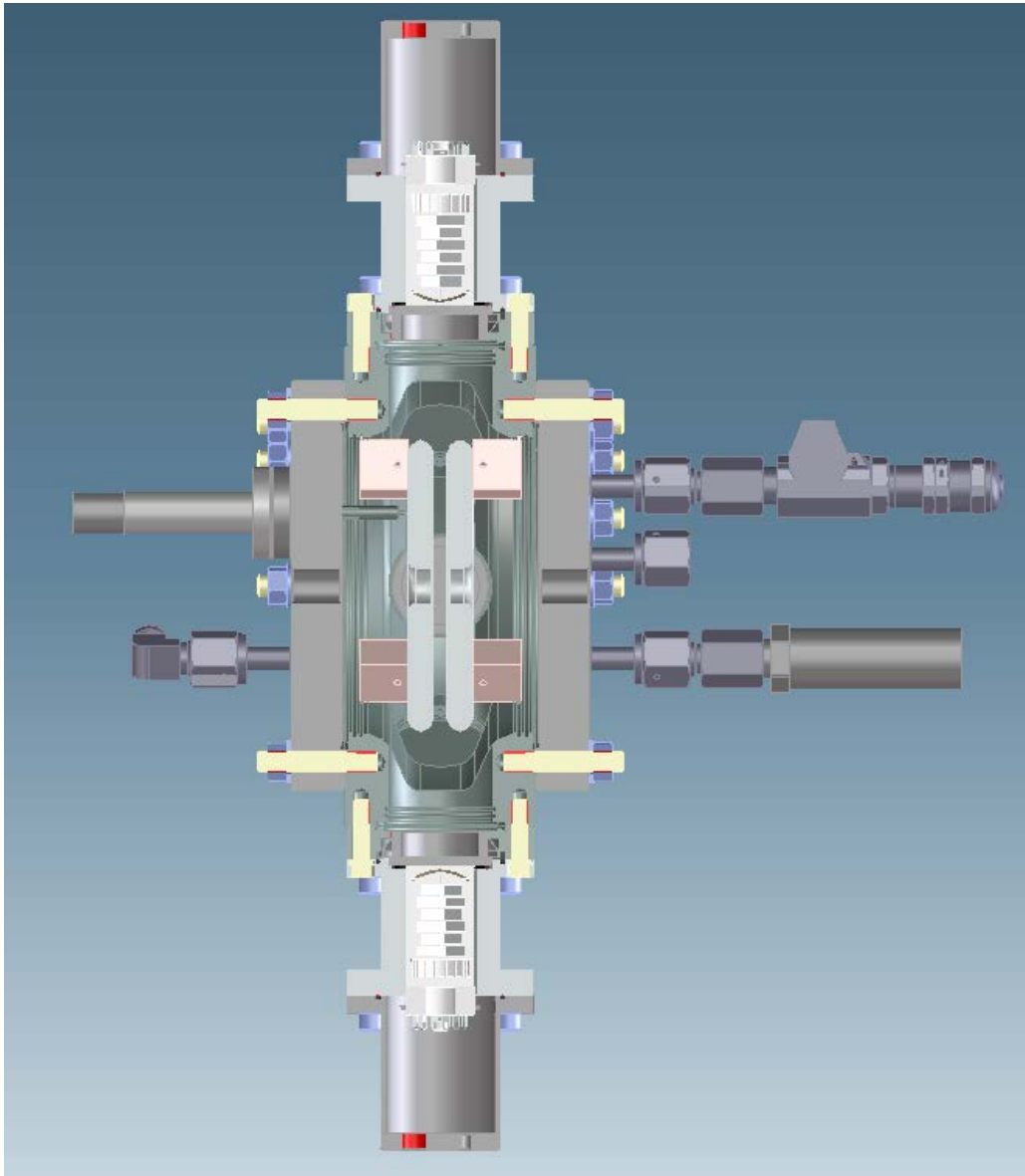


## the “TEA-pot”

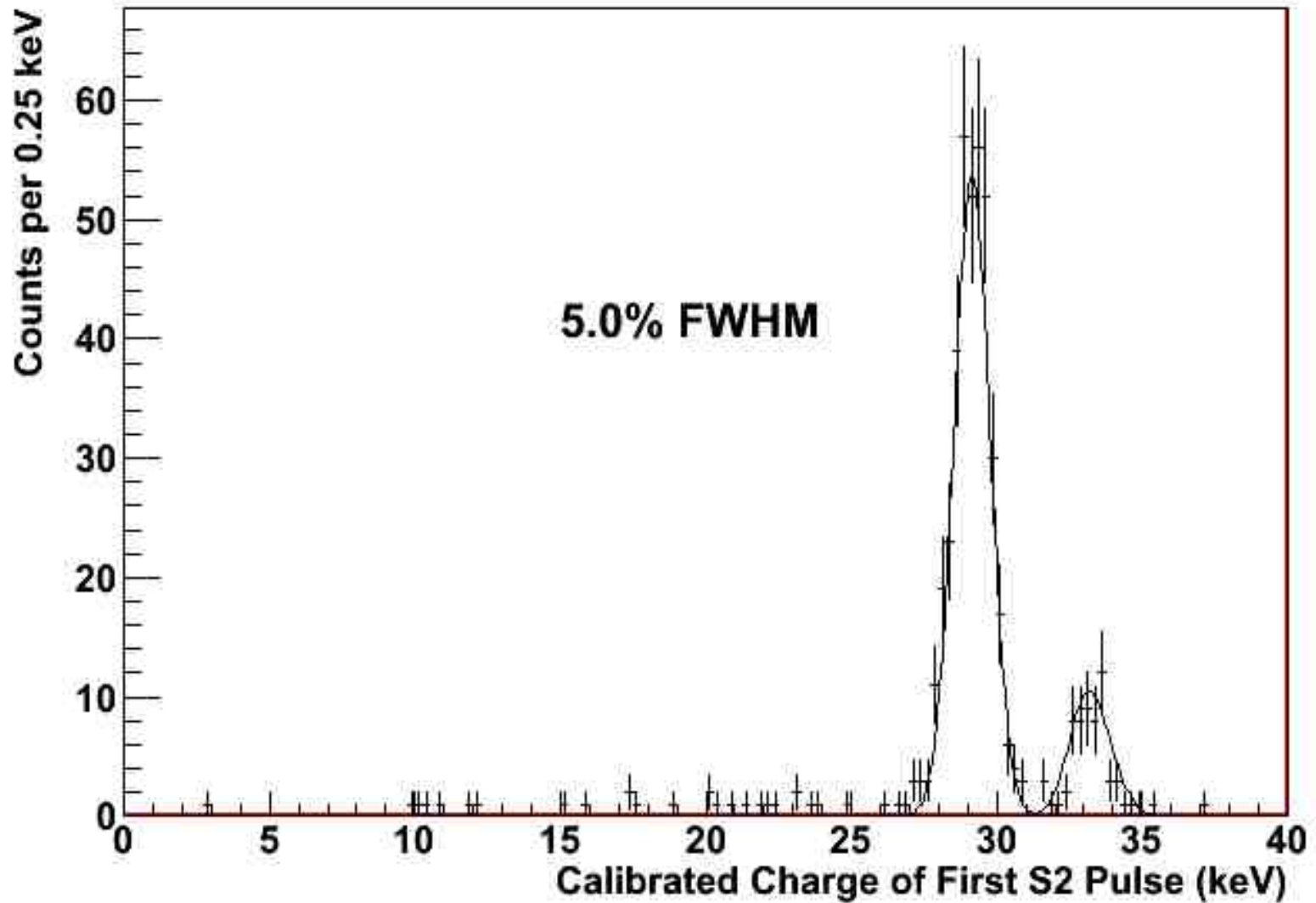
### Basic responses measurements:

A parallel-plate ionization chamber with optical sensing, using 4 PMTs that look at the gap from the sides

We will measure both light and charge as functions of density, electric field, and fraction of TMA/TEA,



The x-ray peaks around ~30 keV



# Energy resolution at $Q_{\beta\beta} = 2457$ keV

$$\delta E/E = 2.35 \cdot (F \cdot W/Q)^{1/2}$$

- $F \equiv$  Fano factor (HPXe) :  $F = 0.15$
- $w \equiv$  Average energy per ion pair:  $w \sim 25$  eV
- $Q \equiv$  Energy deposited from  $^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$ :  
 $N = Q/w \sim 100,000$  primary electrons  
 $\sigma_N = (F \cdot N)^{1/2} \sim 124$  electrons rms!

**$\delta E/E = 0.28\%$  FWHM      intrinsic HPXe**

Scaling our result:  $\delta E/E \sim 0.5\%$  FWHM @  $Q_{2,2}$

# NEXT Collaboration



~~CIEMAT (Madrid) • U. Girona • IFAE (Barcelona) •~~  
IFIC (Valencia) • U. Santiago • U.P.Valencia • U. Zaragoza



LBL • Texas A&M • ISU • UNM



U. Aveiro • U. Coimbra



CEA (Saclay)



JINR (Dubna)



UAN (Bogota)

## Spain provides:

Most of the collaborators

Most secured funding

Host Laboratory - LSC

## Key contributions from international groups

Engineering and integration

TPC expertise

high-pressure gas detectors

Xenon supply & enrichment