Nuclear Matrix Elements for Tensor Interactions that Violate Local Lorentz Invariance

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New Test of Local Lorentz Invariance Using a ²¹Ne-Rb-K Comagnetometer

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We develop a new comagnetometer using ²¹Ne atoms with nuclear spin I = 3/2 and Rb atoms polarized by spin exchange with K atoms to search for tensor interactions that violate local Lorentz invariance. We frequently reverse the orientation of the experiment and search for signals at the first and second harmonics of the sidereal frequency. We constrain 4 of the 5 spatial Lorentz-violating coefficients c_{jk}^n that parametrize anisotropy of the maximum attainable velocity of a neutron at a level of 10^{-29} , improving previous limits by 2 to 4 orders of magnitude and placing the most stringent constraint on deviations from local Lorentz invariance.







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The Michelson-Morley experiment and its successors have established that the speed of light is isotropic to a part in 10¹⁷ [1,2]. Similarly, possible anisotropy in the maximum attainable velocity (MAV) for a massive particle [3] has been constrained by Hughes and Drever NMR experiments [4,5] and their successors to a part in 10²⁷ [6]. These experiments form the basis for the principle of local Lorentz invariance (LLI). Together with the weak equivalence principle and the position invariance principle, they constitute the Einstein equivalence principle (EEP) that is the basis of general relativity [7].



Here we describe a new comagnetometer that is sensitive to anisotropy in neutron MAV at 10^{-29} level. The idea of the experiment is based on the K-³He comagnetometer, previously used to constrain Lorentz-violating vector spin interactions [18]. The ³He (I = 1/2) is replaced by ²¹Ne (I = 3/2) to allow measurements of tensor anisotropy. In addition, since the gyromagnetic ratio of ²¹Ne is about an order of magnitude smaller than that of ³He, the comagnetometer has an order of magnitude better energy resolution for the same level of magnetic field sensitivity. The electric quadrupole interactions of ²¹Ne cause several difficulties.



CPT-even Lorentz violation

$$\mathcal{L} = -\overline{\psi}(m + a_{\mu}\gamma^{\mu} + b_{\mu}\gamma_{5}\gamma^{\mu})\psi + a_{\mu\nu}\psi + c_{\mu\nu}\gamma^{\mu} + d_{\mu\nu}\gamma_{5}\gamma^{\mu})\overleftrightarrow{\partial}^{\nu}\psi \qquad a,b - CPT-odd c,d - CPT-even$$

• Maximum attainable particle velocity

Coleman and Glashow

$$V_{MAX} = c(1 - c_{00} - c_{0j} \hat{v}_j - c_{jk} \hat{v}_j \hat{v}_k)$$
 Jacobson

- Implications for ultra-high energy cosmic rays, Cherenkov radiation, etc
- Best limit $c_{00} \sim 10^{-23}$ from Auger ultra-high energy cosmic rays
- Many laboratory limits (optical cavities, cold atoms, etc)
- Motivation for Lorentz violation (without breaking CPT)
 - Doubly-special relativity
 - Horava-Lifshitz gravity

Something special needs to happen when particle momentum reaches Plank scale!



Search for CPT-even Lorentz violation with nuclear spin

- Need nuclei with orbital angular momentum and total spin >1/2
- Quadrupole energy shift proportional to the kinetic energy of the valence nucleon

$$E_Q \sim (c_{11} + c_{22} - 2c_{33}) \left\langle p_x^2 + p_y^2 - 2p_z^2 \right\rangle$$

- Previosly has been searched for in two experiments using ²⁰¹Hg and ²¹Ne with sensitivity of about 0.5 µHz
- Bounds on neutron c_n~10⁻²⁷ already most stringent bound on c coefficient!
 Sidereal Variation

Sidereal Variation Semi-sidereal Variation

$$\Delta E(t) = E_0 + E_{1X} \cos \Omega t + E_{1Y} \sin \Omega t + E_{2X} \cos 2\Omega t + E_{2Y} \sin 2\Omega t$$

$$C_{\mu\nu} = \begin{pmatrix} C_{TT} & C_{TX} & C_{TY} & C_{TZ} \\ C_{XT} & C_{XX} & C_{XY} & C_{XZ} \\ C_{YT} & C_{YX} & C_{YY} & C_{YZ} \\ C_{ZT} & C_{ZX} & C_{ZY} & C_{ZZ} \end{pmatrix} \xrightarrow{\bullet} \begin{cases} 2^{nd} \text{ Harmonic} \\ 3^{1st} \text{ Harmonic} \\ 3^{1st} \text{ Harmonic} \\ 3^{1st} \text{ Suppressed by } v_{\text{Earth}} \end{cases}$$



Comagnetometer

Magnetic field self-compensation







FIG. 1 (color online). The experimental apparatus is rotated around the local vertical. ²¹Ne spins are polarized down along $-\hat{z}$ and the probe beam is directed horizontally along $-\hat{x}$.



Lots of data!

Signal: Lock-In X Lock-In Y Background Signal Background Times

Feedback monitors:

Probe Oscillator Intensity Monitor Probe Wavelength Feedback Probe Wavelength Fabry-Perot Signal Probe Tapered Amplifier Current Probe Intensity Pump Intensity Feedback Pump Intensity Pump Wavelength Feedback Pump Wavelength Signal Pre-Oven Heater Power Oven Heater Power Shutter Duty Cycle Motion: Probe Horizontal Position Probe Vertical Position Probe Pre-Cell Horizontal Position Probe Pre-Cell Vertical Position Tube Position Z Table Position X1 Table Position X1 Table Position Z1 Table Position X2 Table Position Y2 Table Position Y3

Zeroing control:

Zero Pump Wavelength (Lz) Zero Pockel Cell (Lx) Pockel Cell Background Zero Bz Orthogonal Zero By Orthogonal Zero Bx Orthogonal Sensitivity Calibration Zero Mirror Linear Stage Polarization Pz

Environmental:

Shield #1 Temperature Shield #3 Temperature Shield #5 Temperature Internal G10 Tube Temperature Air Vent Temperature Probe Laser Temperature Cooling Shield Temperature Cooling Shield Temperature Inner Oven Temperature Input Air Temperature Outer Oven Temperature Cell Temperature Faraday Rotator Return Water Temperature Barometric Pressure Humidity

... and more to come!



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$$C_{\mu\nu} = \begin{pmatrix} C_{TT} & C_{TX} & C_{TY} & C_{TZ} \\ C_{XT} & C_{XX} & C_{XY} & C_{XZ} \\ C_{YT} & C_{YX} & C_{YY} & C_{YZ} \\ C_{ZT} & C_{ZX} & C_{ZY} & C_{ZZ} \end{pmatrix} \begin{bmatrix} e^{-1} e^{-$$



To convert our measurements from magnetic field units, we need an estimate of the nuclear operator $\langle I, I | \mathcal{P}_0^2 | I, I \rangle = \langle I, I | 2p_z^2 - p_x^2 - p_y^2 | I, I \rangle / \sqrt{6}$. Within the Schmidt model, ²¹Ne has a valence neutron in the $d_{3/2}$ state. This gives $\langle I, I | \mathcal{P}_0^2 | I, I \rangle = -\sqrt{2/3} \langle p^2 \rangle / 5 = -\sqrt{8/3} m E_k / 5$, and we take the kinetic energy of the valence nucleon $E_k \sim$ 5 MeV [6]. ²¹Ne is better described by a collective wave function within the *sd* shell model, and it should be possible to calculate the nuclear operator more precisely [29].



Matrix elements required for experiments involving nuclear spin coupling

 $\mu_{th} = 1.148 \; (d_{3/2} \; model)$

 $\mu_{th} = -0.750 \text{ (full sd model)}$

 $\mu_{exp} = -0.662$

 $(S_p, S_n) = (0.022, 0.292)$



Dimensionless harmonic-oscillator potential

$$H = (p^2 + r^2)/2$$

means that the form of the wf in momentum space are the same as the those in coordinate space. But there is a phase factor

$$\Psi(\vec{r}) = i^N \Psi(\vec{p})$$

where

$$N = n_x + n_y + n_z = 2n + \ell$$

The dimensions for $< r^2 >$ are $b^2 = \hbar/m\omega$ and those for $< p^2 >$ are m $\hbar\omega$.



Table 1. One-body transition density					
a ⁺	a	simple		sd-shell	
		proton	neutron	proton	neutron
d5	d5	0	0	-0.253	-0.237
d5	d3	0	0	-0.038	-0.133
d5	s1	0	0	-0.270	-0.345
d3	d5	0	0	0.038	0.134
d3	d3	0	1	-0.035	-0.044
d3	s1	0	0	0.104	0.073
s 1	d5	0	0	-0.270	-0.345
s1	d3	0	0	-0.104	-0.073

Table 1: One-body transition density



The sd-shell results for ²¹Ne are

$$(\langle 2r_z^2 - r_y^2 - r_x^2 \rangle_p, \langle 2r_z^2 - r_y^2 - r_x^2 \rangle_n) = (2.37, 2.81) b^2$$

So the momentum results are

$$\begin{aligned} (<2p_z^2 - p_y^2 - p_x^2 >_p, <2p_z^2 - p_y^2 - p_x^2 >_n) &= (2.37, 2.81) \text{ m } \hbar \omega \end{aligned}$$

The simple neutron $d_{3/2}$ model gives
 $(<2p_z^2 - p_y^2 - p_x^2 >_p, <2p_z^2 - p_y^2 - p_x^2 >_n) &= (0, -1.4) \text{ m } \hbar \omega \end{aligned}$
 $(1.4 = 7/5).$



With $b^2 = 3.40 \text{ fm}^2$ (from ground-state rms charge radius) the Q moments are

 $(Q_p, Q_n) = (5.7, 6.7) \text{ fm}^2$

The charge Q moment is $5.7 e^2 \text{ fm}^2$ compared to the experimental value of $10.3(8) e^2 \text{ fm}^2$.

E2 matrix elements for the sd-shell require effective changes of $e_p = 1.0 + 0.37$ and $e_n = 0 + 0.45$ that give

 $Q_p = 5.7 + 5.7 \ge 0.37 + 6.7 \ge 0.45 = 10.8 \text{ fm}^2$

 $Q_n = 6.7 + 6.7 \ge 0.45 + 5.7 \ge 0.37 = 11.8 \text{ fm}^2$

This is attributed to the core-polarization corrections coming from mixing of configurations with $\Delta N = 2$.



















The i^N factor gives a reduction in the momentum matrix elements by factors of $e'_p = 1.0 - 0.37$ and $e'_n = -0.45$

$$(\langle 2p_z^2 - p_y^2 - p_x^2 \rangle_p, \langle 2p_z^2 - p_y^2 - p_x^2 \rangle_n) = (2.37, 2.81) \text{ m } \hbar\omega$$
to

$$(\langle 2p_z^2 - p_y^2 - p_x^2 \rangle_p, \langle 2p_z^2 - p_y^2 - p_x^2 \rangle_n) = (0.22, 0.67) \text{ m } \hbar\omega$$



Bohr and Mottelson: if the deformation of the density distribution is equal to that of the self-consistent potential then for E2 we have $e_p + e_n = 1 + 1 = 2$ and for the momentum matrix elements 1 - 1 = 0. That is $\langle p_z^2 \rangle = \langle p_x^2 \rangle = \langle p_y^2 \rangle$.



Other experiments of this type require these momentum matrix elements for 131 Xe (I=3/2) and 201 Hg (I=3/2).

We have a good Hamiltonian for the $(0g_{7/2}, 1d_{5/2}, 1d_{3/2}, 2s_{1/2}, 0h_{11/2})$ (jj55) model space.

Results for 131 Xe can be obtained with NuShellX M=3/2 dimension is 388,283,806 J=3/2 dimension is 14,021,322





*.int Hamiltonian files

NSCI

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$$< B_f, J \mid H_{pn} \mid B_i, J > = \sum_{pp'nn'\lambda} F_{\lambda}(pp'nn') \ \Gamma_{\lambda} \ \mathrm{RDM}(p_f, p_i, p, p', \lambda) \ \mathrm{RDM}(n_f, n_i, n, n', \lambda)$$

where, for example, p_f , stands for labels (J_{p_f}, α_{p_f}) ,

$$\Gamma_{\lambda} = \left\{ \begin{array}{ccc} J_{p_{f}} & J_{p_{i}} & \lambda \\ J_{n_{f}} & J_{n_{i}} & \lambda \\ J & J & 0 \end{array} \right\}$$

and RDM are the reduced density matricies:

$$RDM(p_f, p_i, p, p', \lambda) = < [(J_{p_f}, \alpha_{p_f}) || [a_p^+ \tilde{a}_{p'}]^{\lambda} || [(J_{p_i}, \alpha_{p_i}) >$$

and

$$RDM(n_f, n_i, n, n', \lambda) = < [(J_{n_f}, \alpha_{n_f}) || [a_n^+ \tilde{a}_{n'}]^{\lambda} || [(J_{n_i}, \alpha_{n_i}) >$$

The key is to optimize the sums in this equation for OpenMP and/or MPI







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jj55 model space



Search for Neutrinoless Double-Beta Decay in ¹³⁶Xe with EXO-200

Mihai Horoi (next week) – also requires addition of spin-orbit partners

¹³⁷Ba – first observation of double-gamma decay (D.J. Millener, R.J. Sutter, D.E. Alburger)





(72)







Result is $(Q_p, Q_n) = (-1.91, -7.83) e^2 \text{ fm}^2$. Compared to $Q_{exp} = -11.4(1) e^2 \text{ fm}^2$.

Now we need to add both $\Delta N = 0$ and $\Delta N = 2$ core-polarization corrections.

The results will be something like

 $Q_p = -1.91 \text{ (bare)} - (1.91 + 7.83) (0.2) (\Delta N = 0)$ - (1.91 + 7.83) (0.8) ($\Delta N = 2$) = -11.6



The results will be something like

$$Q_p = -1.91 \text{ (bare)} - (1.91 + 7.83) (0.2) (\Delta N = 0)$$

- (1.91 + 7.83) (0.8) ($\Delta N = 2$) = -11.6

and the momentum matrix elements will be proportional to

$$P_p = -1.91 \text{ (bare)} - (1.91 + 7.83) (0.2) (\Delta N = 0) + (1.91 + 7.83) (0.8) (\Delta N = 2) = +3.9$$
$$P_n = -7.83 \text{ (bare)} - (1.91 + 7.83) (0.2) (\Delta N = 0) + (1.91 + 7.83) (0.8) (\Delta N = 2) = -2.0$$



How small are these quadrupole momentum matrix elements in ab-initio and no-core approaches?

For example ⁹Be 3/2- ground state



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