Lattice Effective Field Theory for Nuclei from $A = 4$ to $A = 28$

Nuclear Lattice EFT Collaboration

Evgeny Epelbaum (Bochum) Hermann Krebs (Bochum) Timo A. Lähde (Jülich) Dean Lee (NC State) Ulf-G. Meißner (Bonn/Jülich) Gautam Rupak (MS State)

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HELMHOLTZ GEMEINSCHAFT

Outline

Brief introduction to Lattice EFT for nuclei

Carbon-12 and the Hoyle state

Production of Carbon-12 in red giant stars

Bounds on the anthropic scenario

Preliminary results up to $A = 28$

Low-energy nucleons

Chiral effective field theory on the lattice ...

Current status of lattice chiral EFT Improved NNLO interaction ...

Epelbaum, Hammer, Meißner, Rev. Mod. Phys. 81 (2009) 1773

$$
\mathcal{A}_{\text{LO}} = C_{S=0,I=1} f(\boldsymbol{q}) \left(\frac{1}{4} - \frac{1}{4} \boldsymbol{\sigma}_i \cdot \boldsymbol{\sigma}_j \right) \left(\frac{3}{4} + \frac{1}{4} \boldsymbol{\tau}_i \cdot \boldsymbol{\tau}_j \right) \n+ C_{S=1,I=0} f(\boldsymbol{q}) \left(\frac{3}{4} + \frac{1}{4} \boldsymbol{\sigma}_i \cdot \boldsymbol{\sigma}_j \right) \left(\frac{1}{4} - \frac{1}{4} \boldsymbol{\tau}_i \cdot \boldsymbol{\tau}_j \right) \n- \tilde{g}_{\pi N}^2 \boldsymbol{\tau}_i \cdot \boldsymbol{\tau}_j \frac{\boldsymbol{\sigma}_i \cdot \boldsymbol{q} \boldsymbol{\sigma}_j \cdot \boldsymbol{q}}{\boldsymbol{q}^2 + M_{\pi}^2},
$$

Smearing of LO contact interactions

$$
C_0 = \frac{3}{4} C_{S=0,I=1} + \frac{1}{4} C_{S=1,I=0}
$$

$$
C_I = \frac{1}{4} C_{S=0,I=1} - \frac{1}{4} C_{S=1,I=0}
$$

Fix constants from the two-nucleon sector Lattice EFT is predictive for A > 2 ...

$$
\cos \delta_L \cdot j_L(kR_{wall}) = \sin \delta_L \cdot y_L(kR_{wall}),
$$

$$
\delta_L = \tan^{-1} \left[\frac{j_L(kR_{wall})}{y_L(kR_{wall})} \right].
$$

Gaussian smearing of contact terms

Borasoy, Krebs, Lee, Meißner, Nucl. Phys. A768 (2006) 179; Eur. Phys. J. A31 (2007) 105; Lee, Prog. Part. Nucl. Phys. 63 (2009) 179

Euclidean time projection Ground state energy ...

$$
Z_A(t) = \langle \psi_A | \exp(-tH) | \psi_A \rangle
$$

(discretized)

Choice of trial wavefunction:

- Standing waves
- Alpha clusters
- Lattice Hamiltonian \vert Shell model wavefunctions

Decoupling of nucleon-nucleon interactions Hubbard-Stratonovich transformation ...

$$
\exp\left[-\frac{C}{2}(N^{\dagger}N)^{2}\right] = \sqrt{\frac{1}{2\pi}} \int_{-\infty}^{\infty} ds \exp\left[-\frac{1}{2}s^{2} + \sqrt{-C} s(N^{\dagger}N)\right]
$$

Auxiliary Field Quantum Monte Carlo (AFQMC) Discretized Euclidean time evolution ...

$$
\Big\| = M_{\text{LO}} \qquad \Big\| = M_{\text{approx}} \qquad \Big\| = O_{\text{observable}}
$$

sign problem

positive definite (not too severe) strength can be varied

Hybrid Monte Carlo sampling

$$
\Rightarrow Z_{n_t,\text{LO}} = \langle \psi_{\text{init}} | \boxed{\text{min}} \boxed{\text{min}} \boxed{\text{min}} \boxed{\text{min}} \boxed{\text{min}} \boxed{\text{min}} \ket{\psi_{\text{init}}}
$$

$$
e^{-E_{0,\text{LO}}a_t} = \lim_{n_t \to \infty} Z_{n_t+1,\text{LO}} / Z_{n_t,\text{LO}}
$$

$$
\langle O \rangle_{0,\text{LO}} = \lim_{n_t \to \infty} Z_{n_t,\text{LO}}^{\langle O \rangle} / Z_{n_t,\text{LO}}
$$

For a thorough review, see: Lee, Prog. Part. Nucl. Phys. 63 (2009) 179

AFQMC + Hybrid Monte Carlo

Substantial investment of supercomputing time ...

CPU time allocations:

- JUQUEEN (FZ Jülich), 30 Mcore-h (project) + > 100 Mcore-h (institutional)

- RWTH cluster (Aachen), 1.3 Mcore-h (project) + "free CPU time" (long queue)

Figure courtesy of Jülich Supercomputer Centre (JSC)

AFQMC results for 12C (ground state) Improved NNLO interaction ... $a = 1.97$ fm

Epelbaum, Krebs, D.L, Meißner, PRL 106 (2011) 192501 Epelbaum, Krebs, Lähde, D.L, Meißner, PRL 109 (2012) 252501

AFQMC - ground and Hoyle states of ¹² C Multiple trial wavefunctions ... $a = 1.97$ fm

Epelbaum, Krebs, Lähde, D.L, Meißner, PRL 109 252501 (2012)

Alpha cluster structure of 12C ...

Ground state

Hoyle state

24 rotational orientations

Production of 12C in red giant stars Resonant production via ⁸Be and Hoyle state ...

$$
r_{3\alpha} = 3^{\frac{3}{2}} N_\alpha^3 \left(\frac{2\pi\hbar^2}{M_\alpha k_\mathrm{B} T} \right)^3 \frac{\Gamma_\gamma}{\hbar} \, \exp\left(-\frac{\Delta E_{h+b}}{k_\mathrm{B} T}\right)
$$

Is the Universe fine-tuned?

Energy of Hoyle state in 12C relative to triple alpha $\Delta E_b + \Delta E_h = E_{12}^{\star} - 3E_4$

> Experiment: 379.47 ± 0.18 keV

What if the Hoyle state is moved? Calculations of stellar nucleosynthesis ...

Schlattl et al., Astrophys. Space Sci. 291, 27-56 (2004)

 $|\delta(\Delta E_{h+b})|$ < 100 keV

Anthropic bound on (ad hoc) variation of the Hoyle state

More fundamental description - Chiral EFT Sources of quark mass dependence ...

Figure courtesy of U.-G. Meißner

AFQMC calculation for 4He, 8Be and 12C ...

$$
E_{i} = E_{i}(\tilde{M}_{\pi}, m_{N}(M_{\pi}), \tilde{g}_{\pi N}(M_{\pi}), C_{0}(M_{\pi}), C_{I}(M_{\pi}))
$$
\nSmall shifts around

\n
$$
\frac{\partial E_{i}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\text{ph}}} = \frac{\partial E_{i}}{\partial \tilde{M}_{\pi}}\Big|_{M_{\pi}^{\text{ph}}} + x_{1} \frac{\partial E_{i}}{\partial m_{N}}\Big|_{m_{N}^{\text{ph}}} + x_{2} \frac{\partial E_{i}}{\partial \tilde{g}_{\pi N}}\Big|_{\tilde{g}_{\pi N}^{\text{ph}}}
$$
\n
$$
+ x_{3} \frac{\partial E_{i}}{\partial C_{0}}\Big|_{C_{0}^{\text{ph}}} + x_{4} \frac{\partial E_{i}}{\partial C_{I}}\Big|_{C_{I}^{\text{ph}}}
$$
\n
$$
x_{1} := \frac{\partial m_{N}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\text{ph}}}
$$
\n
$$
x_{2} := \frac{\partial \tilde{g}_{\pi N}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\text{ph}}}
$$
\n
$$
x_{3} := \frac{\partial C_{0}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\text{ph}}}, \quad x_{4} := \frac{\partial C_{I}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\text{ph}}}
$$
\n
$$
x_{4} := \frac{\partial C_{I}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\text{ph}}}
$$
\nHowever, the result of the physical point

\n
$$
x_{5} := \frac{\partial C_{0}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\text{ph}}}, \quad x_{6} := \frac{\partial C_{I}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\text{ph}}}
$$
\nThus-nucleon scattering

Parameterization of the short-range terms

Lüscher formula **...**

$$
p \cot\delta = \frac{1}{\pi L} S(\eta) \approx -\frac{1}{a}, \qquad \eta:= m_N E\left(\frac{L}{2\pi}\right)^2
$$

$$
\bar{A}=\frac{\partial a^{-1}}{\partial M_\pi}=-\frac{1}{\pi L}S'(\eta)\frac{\partial \eta}{\partial M_\pi}
$$

$$
\zeta_i := \frac{m_N L}{4\pi^3} S'(\eta_i)
$$

$$
q_i := \frac{\partial E_i}{\partial C_0} \bigg|_{C^{\rm ph}_0}
$$

4He

$$
\frac{\partial E_4}{\partial m_\pi}\bigg|_{m_\pi^{\rm phys}} = -0.339(5)\left.\frac{\partial a_s^{-1}}{\partial m_\pi}\right|_{m_\pi^{\rm phys}} - 0.697(4)\left.\frac{\partial a_t^{-1}}{\partial m_\pi}\right|_{m_\pi^{\rm phys}} + 0.0380(14)_{-0.006}^{+0.008}
$$

8Be

$$
\frac{\partial E_8}{\partial m_\pi}\bigg|_{m_\pi^{\rm phys}}=-0.794(32)\left.\frac{\partial a_s^{-1}}{\partial m_\pi}\right|_{m_\pi^{\rm phys}}-1.584(23)\left.\frac{\partial a_t^{-1}}{\partial m_\pi}\right|_{m_\pi^{\rm phys}}+0.089(9)^{+0.017}_{-0.011}
$$

12C (ground)

$$
\frac{\partial E_{12}}{\partial m_{\pi}}\bigg|_{m_{\pi}^{\rm phys}}=-1.52(3)\left.\frac{\partial a_{s}^{-1}}{\partial m_{\pi}}\right|_{m_{\pi}^{\rm phys}}-2.88(2)\left.\frac{\partial a_{t}^{-1}}{\partial m_{\pi}}\right|_{m_{\pi}^{\rm phys}}+0.159(7)_{-0.018}^{+0.023}
$$

12C (Hoyle)

$$
\frac{\partial E_{12}^{\star}}{\partial m_{\pi}}\bigg|_{m_{\pi}^{\text{phys}}} = -1.588(11)\left.\frac{\partial a_s^{-1}}{\partial m_{\pi}}\right|_{m_{\pi}^{\text{phys}}} - 3.025(8)\left.\frac{\partial a_t^{-1}}{\partial m_{\pi}}\right|_{m_{\pi}^{\text{phys}}} + 0.178(4)_{-0.021}^{+0.026}
$$

Epelbaum, Krebs, Lähde, D.L, Meißner, PRL 110 (2013) 112502; ibid., arXiv:1303.4856 Berengut et al., Phys. Rev. D 87 (2013) 085018

$$
\left.\frac{\partial \Delta E_{h+b}}{\partial m_\pi}\right|_{m_\pi^{\text{phys}}}=-0.572(19)\left.\frac{\partial a_s^{-1}}{\partial m_\pi}\right|_{m_\pi^{\text{phys}}}-0.933(15)\left.\frac{\partial a_t^{-1}}{\partial m_\pi}\right|_{m_\pi^{\text{phys}}}+0.064(6)^{+0.010}_{-0.009}
$$

--> Viability of carbon-oxygen based life: $|\delta (\Delta E_{h+b})| < 100 \,\, \mathrm{keV}$

$$
\left[0.572(19) \bar{A}_s + 0.933(15) \bar{A}_t - 0.064(6) \right] \times \left(\frac{\delta m_q}{m_q} \right) \bigg| < 0.15\%
$$

Epelbaum, Krebs, Lähde, D.L, Meißner, PRL 110 (2013) 112502; ibid., arXiv:1303.4856 Berengut et al., Phys. Rev. D 87 (2013) 085018

Current theoretical knowledge of the quark mass dependence of the S-wave scattering lengths ...

Berengut et al., Phys. Rev. D 87 (2013) 085018

The "end of the world" plot :)

Epelbaum, Krebs, Lähde, Lee, Meißner, Phys. Rev. Lett. 110 (2013) 112502; arXiv:1303.4856

Upcoming results

Spectra of Oxygen-16 and Neon-20

Extension of nuclear lattice EFT up to $A = 28$

First AFQMC results for non-alpha-cluster nuclei

Extension of chiral NN interaction to N3LO

Effects of lattice spacing and finite volume

Preliminary results for binding energies up to A = 28 ...

(includes contact 4N correction)

Preliminary results for binding energies up to A = 28 ...

Improved NNLO interaction

 + smeared 4N correction