Direct and indirect torsion-balance constraints on monopole-dipole forces from axion-like particles

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Outline:

 discuss the 2 most recent direct lab results: try to get sensitivity to high-mass "home made" ALPs challenging experiments with simple interpretation

 indirect lab results (see Raffelt's talk) get g_s(N) from ISL and WEP tests very sensitive experiments but some subtleties in interpretation get g_p(e) from spin-spin experiments 5 orders of magnitude less sensitive than astrophysical limits

• Is it worth pushing a 2nd generation of direct experiments?

innovative, high-tech Birmingham / Imperial College spin-bulk experiment G D Hammond et al, PRL 98, 081101 (2007) PRD 77, 036005 (2008)

the challenge: align a lot of electron spins in close proximity to a test body while minimizing magnetic effects on the unpolarized test body



top view of test body in a core



Mu-Metal core: r_{inner} = 4.5 mm, h=80 mm OFHC test body: r=0.5 mm, h=30 mm

Eöt-Wash spin-bulk experiment



"Improved Constraints on Axion-mediated Forces", S. Hoedl, F. Fleischer, E. G. Adelberger and B. R. Heckel, Physical Review Letters **106**, 041801 (2011).

Exchange of virtual axions mediates a monopole-dipole force

$$V(\hat{\sigma}, \mathbf{r}) = \frac{\hbar^2(\hat{\sigma} \cdot \hat{r})}{8\pi m_e} \left(\frac{g_s^a g_p^e}{\hbar c}\right) \left(\frac{1}{r\lambda_{\rm ALP}} + \frac{1}{r^2}\right) e^{-r/\lambda_{\rm ALP}}$$

where g_s^a is proportional to Θ_{QCD}





electrostatic shield

electrically isolated magnet poles used to lock pendulum when reversing B

Tb coating canceled Si diamagnetism to within 5% This experiment was very brave: we got sensitivity to higher axion masses by using an unshielded test body

our pendulums are normally shielded from electrical and magnetic fields and very little power is dissipated inside the instrument to avoid temperature variations.

In this case pendulum is subjected reversing kG fields and magnets windings dissipate ~30 W

pendulum is slightly paramagnetic (a weak-field seeker) and so is subject to a "magnetic torsional spring" that is much stronger than the fiber spring

B must be reversed very precisely and any stray B fields carefully minimized to keep the magnet spring constant

our experiment is most sensitive for axions (ALPs) at the high end of the allowed axion mass window

The pendulum: • hi-purity single-Xtal Si (vol mag susceptability -3.7x10-6) • coated with 30 nm of Tb (vol mag susceptability .112, Curie point 222K) • finish coat of Au

The magnet:
•core of magnet iron
•temp controlled by coolant in direct contact with iron ΔT=(-.32 +- .27)mK
•potted in heat conducting exoxy
•electrically insulated pole shims improved B field uniformity used to capture pendulum during B reversals
•surrounded by 2 layers of mu-metal shielding



a run consisted of 12 cycles of +/magnetic-field reversals

a scan consisted of a magnet degauss followed by runs taken with pendulum at 8 positions within the magnet gap

uncertainties were dominated by "degauss scatter" that is an order magnitude larger than scatter within a run



we suspect the effect arises from inhomogeneities in the magnetic iron



20 scans

we exploited the finite range of the monopole-dipole force by varying the distance between pendulum and pole

Constraints on couplings of axions (ALPs) in the allowed mass range



We improved direct constraints on ALPs with masses >1meV by more than a factor of 10¹⁰

The upgrade path for our experiment:

replace magnet core with laminated hi-mu material (Conetic AA)

- lower saturation B
- much smaller coercive force

replace pendulum

- Ge instead of Si
- reduce susceptability with conventional paramagnetic coating
 much larger "cutouts"

don't try to cancel magnetic spring



Now review some of the lab experiments that Raffelt uses for indirect constraints on the monopole-dipole force:

get g_s(N) from •ISL tests •WEP tests

assume deviations prop to $|g_s(N)|^2$

Eöt-Wash 42-hole inverse-square law pendulum



PhD project of Dan Kapner, PRL 98, 021101(2007)

95% confidence upper limits on ISL violation D J Kapner et al., PRL 98, 021101 (2007)



the Fourier-Bessel pendulum



pendulum & attracter are 50µm think W foils glued to glass plates



PhD project of Ted Cook



Preliminary 95% CL limits from Cook's expt.

10⁸ EXCLUDED Stanford REGION 10⁶ Colorado **10**⁴ 10 x Preliminary' improvement Analysis based on <u></u>² 10² for masses 1/3 of data of 10 meV 10⁰ Irvine 10^{-2} dark energy LEöt-wash 2004 scale Eöt-wash 2006 10^{-4} 10⁻⁵ 10⁻⁶ 10⁻² 10^{-4} 10^{-3} λ [m]





Limiting factor in Fourier-Bessel ISL test:

excess electrostatic noise for separations less than 100 microns

Eöt-Wash WEP test with a rotating torsion balance S. Schlamminger et al., PRL 100, 041101 (2008)



torsion pendulum of the recent Eöt-Wash EP test S. Schlamminger et al., PRL 100, 041101 (2008)



gravity-gradient compensation



gravity-gradiometer pendulums



q₄₁ configuration on a table

q₂₁ configuration installed

1σ statistical + systematic uncertainties from our Equivalence Principle experiment with beryllium and aluminum test bodies; beryllium and titanium data are similar

Source	Δa (cm/s²)	∆a/a _{source}
Earth	$(-1.2 \pm 2.2) \times 10^{-13}$	(-0.7± 1.3) x 10 ⁻¹³
Sun	$(-3.1 \pm 2.4) \times 10^{-13}$	$(-5.2 \pm 4.0) \times 10^{-13}$
Milky Way	$(-1.2 \pm 2.6) \times 10^{-13}$	$(-6.5 \pm 8.6) \times 10^{-6}$
CMB	$(-3.0 \pm 2.4) \times 10^{-13}$	(-3.4 ± 2.7)×10 ⁻⁴

PhD project of Todd Wagner

95% CL constraints on coupling to vector charges



Note that gap for ranges between 10⁴ m and 10⁶ m has been filled in by improved geophysical models T. A Wagner el al., CQG in press Limiting factors in our WEP tests:

thermal noise in suspension fiber replace W fibers with fused silica time-varying gravity gradients measure gradients continuously



The Eöt-Wash spin-spin experiments





BR Heckel, EGA and WA Terrano, to be published

$$V_1 = \frac{g_A^2}{4\pi r} (\hat{\sigma}_1 \cdot \hat{\sigma}_2) e^{-r/\lambda} \text{ and}$$
$$V_2 = -\frac{g_A g_V \hbar}{4\pi m_e c r^2} (\hat{\sigma}_1 \times \hat{\sigma}_2 \cdot \hat{r}) \left(1 + \frac{r}{\lambda}\right) e^{-r/\lambda}$$

$$V_{3} = - \frac{g_{P}^{2}\hbar^{2}}{16\pi m_{e}c^{2}r^{3}} \left[\left(\hat{\sigma}_{1} \cdot \hat{\sigma}_{2} \right) \left(1 + \frac{r}{\lambda} \right) - \left(\hat{\sigma}_{1} \cdot \hat{r} \right) \left(\hat{\sigma}_{2} \cdot \hat{r} \right) \left(3 + \frac{3r}{\lambda} + \frac{r^{2}}{\lambda^{2}} \right) \right] e^{-r/\lambda} ;$$

preliminary 2σ constraint on V3

5 orders of magnitude weaker than Raffelt astrophysical bound obtain high sensitivity for V1, V2 and V3 using horiz and vert source configs



The big question:

Is it worth pursuing 2nd generation monopole-dipole experiments?

recent review articles:

Torsion balance experiments: a low-energy frontier of particle physics EGA et al, PPNP 62, 102 (2009)

Torsion-balance tests of the Weak Equiv. Principle T. A. Wagner et al, CQG (in press)