

Symmetry Term at High Density from Heavy Ion Collisions

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for the ASY-EOS Collaboration

- Symmetry energy
- Reanalysis of the FOPI/LAND data
[P. Russotto et al., PLB 697 (2011) 471]
- ASY-EOS Experiment of May 2011 (GSI)

Symmetry Energy (nucleus)

$$\frac{E_B}{A} = \underbrace{-a_V}_{Volume} + \underbrace{a_S \frac{1}{A^{1/3}}}_{Surface} + \underbrace{a_C \frac{Z^2}{A^{4/3}}}_{Coulomb} + \underbrace{a_A \left(\frac{N-Z}{A} \right)^2}_{Symmetry} \pm \underbrace{a_P \frac{1}{A^{3/2}}}_{Pairing}$$

Binding Energy/Nucleon
(Bethe-Weizsäcker)
acc. better than 1%

$$a_V = 15.8$$

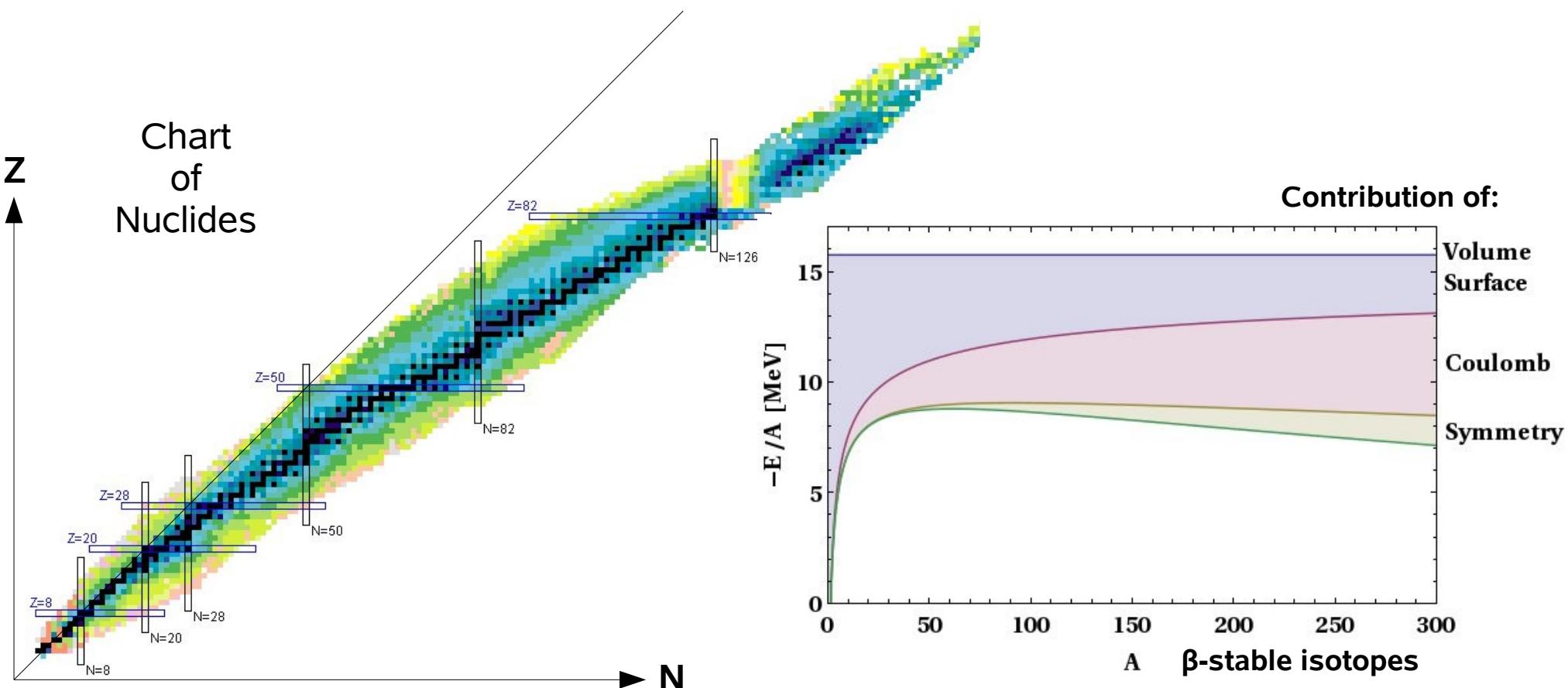
$$a_S = 18.0$$

$$a_C = 0.72$$

$$a_A = 23.5$$

$$a_P = 11.5$$

$$[MeV]$$



Symmetry Energy (nuclear matter)

$$\frac{E_B}{A} = \underbrace{-a_V}_{\text{Volume}} + \underbrace{a_S \frac{1}{A^{1/3}}}_{\text{Surface}} + \underbrace{a_C \frac{Z^2}{A^{4/3}}}_{\text{Coulomb}} + \underbrace{a_A \left(\frac{N-Z}{A} \right)^2}_{\text{Symmetry}} \pm \underbrace{a_P \frac{1}{A^{3/2}}}_{\text{Pairing}}$$

Energy per nucleon in nuclear matter (EoS):

$$E(\rho, \delta) = E(\rho, 0) + E_{sym}(\rho) \delta^2 + o(\delta^4)$$

dominant symmetric matter ($N=Z$) term:

$$E(\rho, 0) \approx -a_V + \frac{K}{18} \left(\frac{\rho - \rho_o}{\rho_o} \right)^2 + \dots$$

symmetry term:

$$E_{sym}(\rho) \approx a_A^V + \frac{L}{3} \left(\frac{\rho - \rho_o}{\rho_o} \right) + \frac{K_{sym}}{18} \left(\frac{\rho - \rho_o}{\rho_o} \right)^2 + \dots$$

$\rho_n, \rho_p \rightarrow$ neutron, proton densities

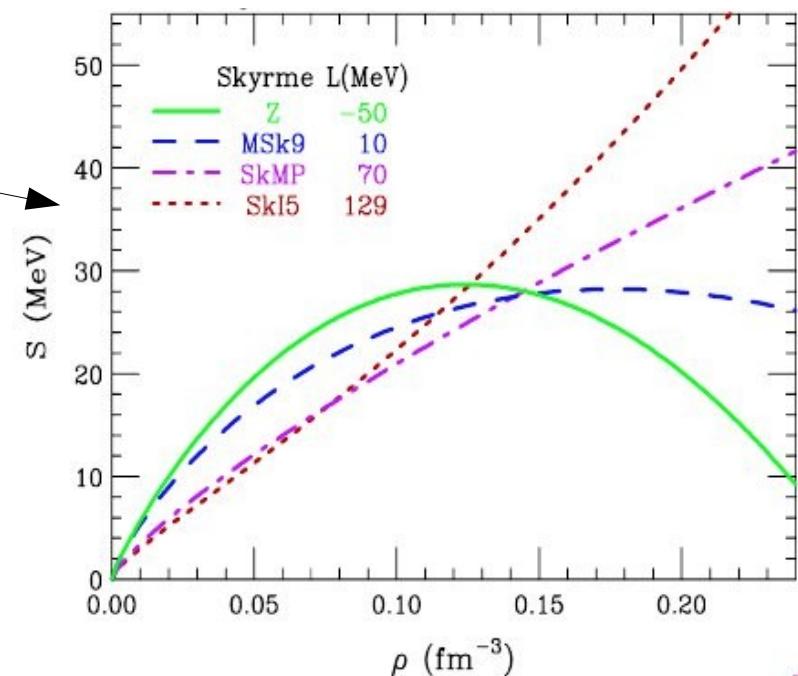
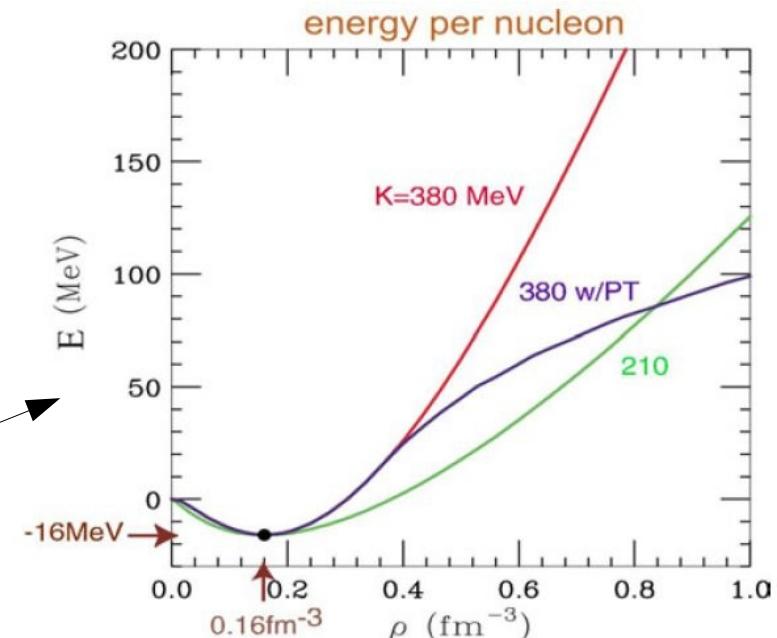
$\rho = \rho_n + \rho_p \rightarrow$ nucleon density

$\delta = \frac{\rho_n - \rho_p}{\rho} \rightarrow$ relative neutron excess

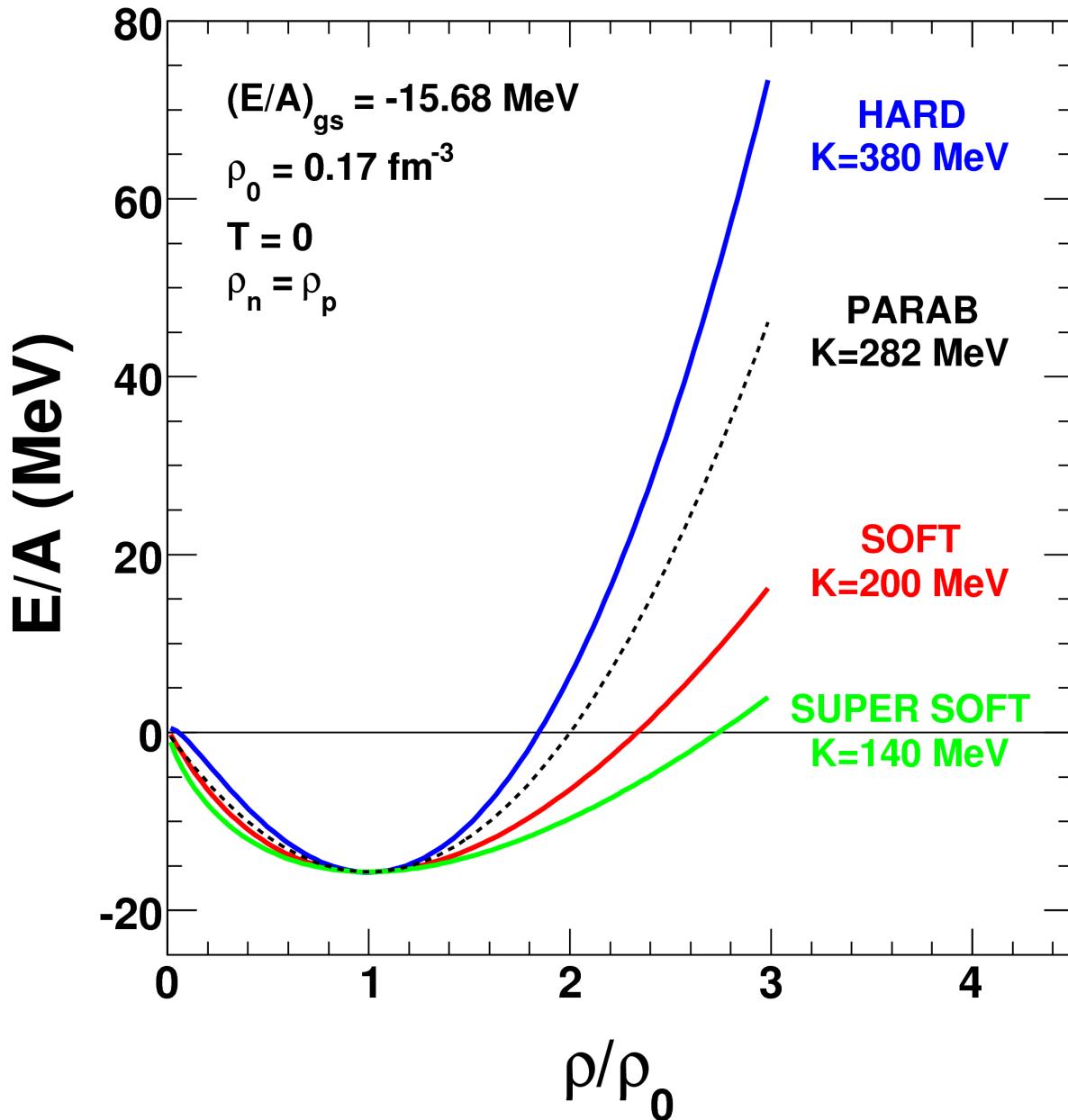
$L = 3 \rho_o \left. \frac{\partial E_{sym}}{\partial \rho} \right|_{\rho=\rho_o} \rightarrow \sim$ symmetry pressure

$K = 9 \rho_o^2 \left. \frac{\partial^2 E}{\partial \rho^2} \right|_{\rho=\rho_o} \rightarrow$ compressibility

$a_V \approx 15.8 \text{ MeV}$ $a_A^V \approx 32.5^{(*)} \text{ MeV}$ $\delta^2(^{208}\text{Pb}) \approx 0.04$ $\delta^2(^{271}\text{Pb}) \approx 0.16$
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Symmetric matter ($\delta=0$)



Simple Skyrme parametrization

$$\frac{E}{A}(\rho) = \frac{3 p_F^2(\rho)}{10 m} + \frac{\alpha}{2} \frac{\rho}{\rho_0} + \frac{\beta}{1+\gamma} \left(\frac{\rho}{\rho_0} \right)^\gamma$$

Compressibility parameter

$$K \equiv 9 \rho^2 \left. \frac{\partial^2 E/A}{\partial \rho^2} \right|_{\rho=\rho_0}$$

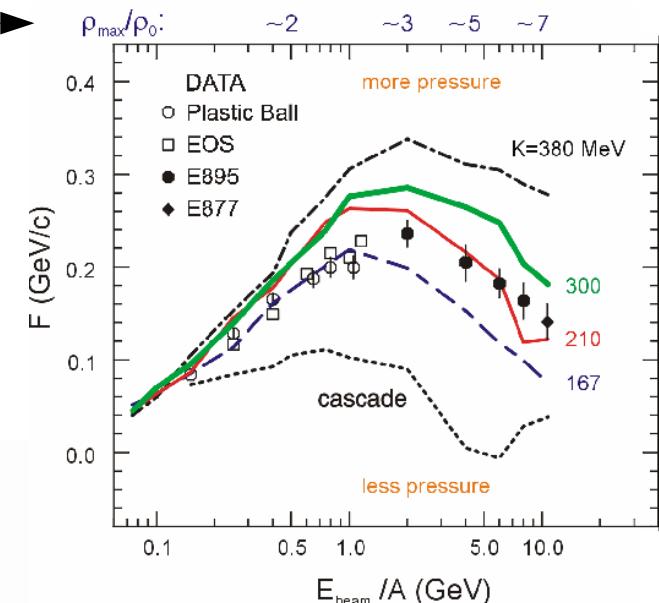
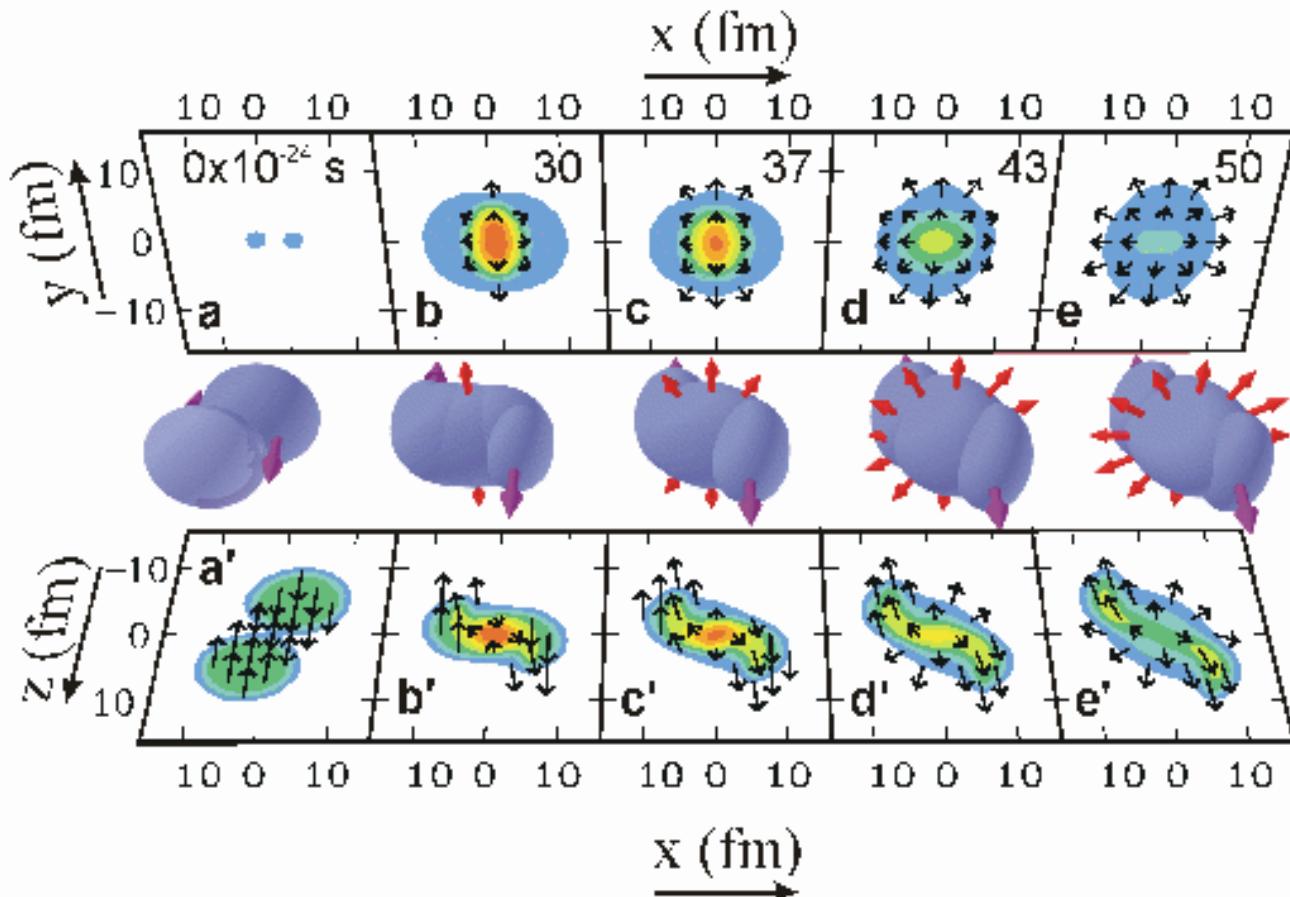
Flow \rightarrow a density sensitive observable

$$F = \left. \frac{d \langle p_x / A \rangle}{d(y/y_{cm})} \right|_{y/y_{cm}=1} \rightarrow \text{slope at midrapidity of the mean transverse in-plane momentum per nucleon}$$

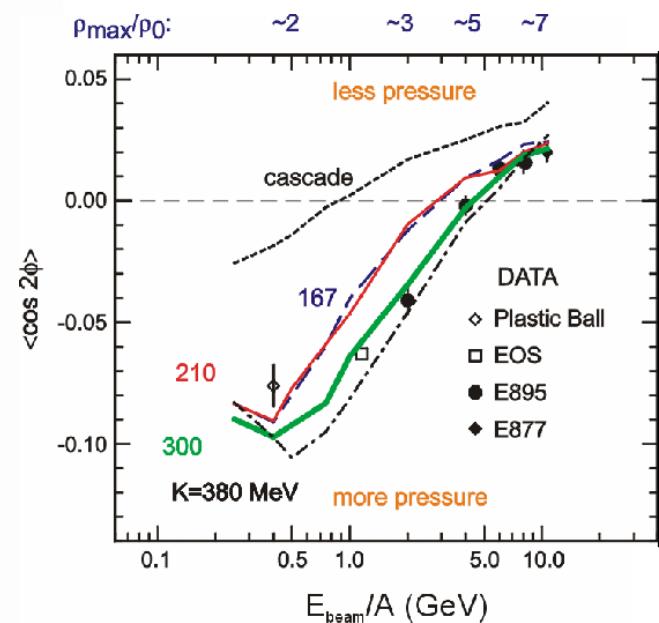
$$E \frac{d^3 N}{d p^3} = \frac{1}{2\pi} \frac{d^2 N}{p_T dp_T dy} \left(1 + \sum_{n=1}^{\infty} 2v_n \cos(n(\phi - \phi_{RP})) \right)$$

with: $v_1 = \langle \cos \Delta \phi \rangle, v_2 = \langle \cos 2 \Delta \phi \rangle, \dots, v_n = v_n(b, Z, A, y, p_T)$

Au+Au @ 2 AGeV, $b=6$ fm (BEM)



Elliptic flow



parameter dependence of elliptic flow

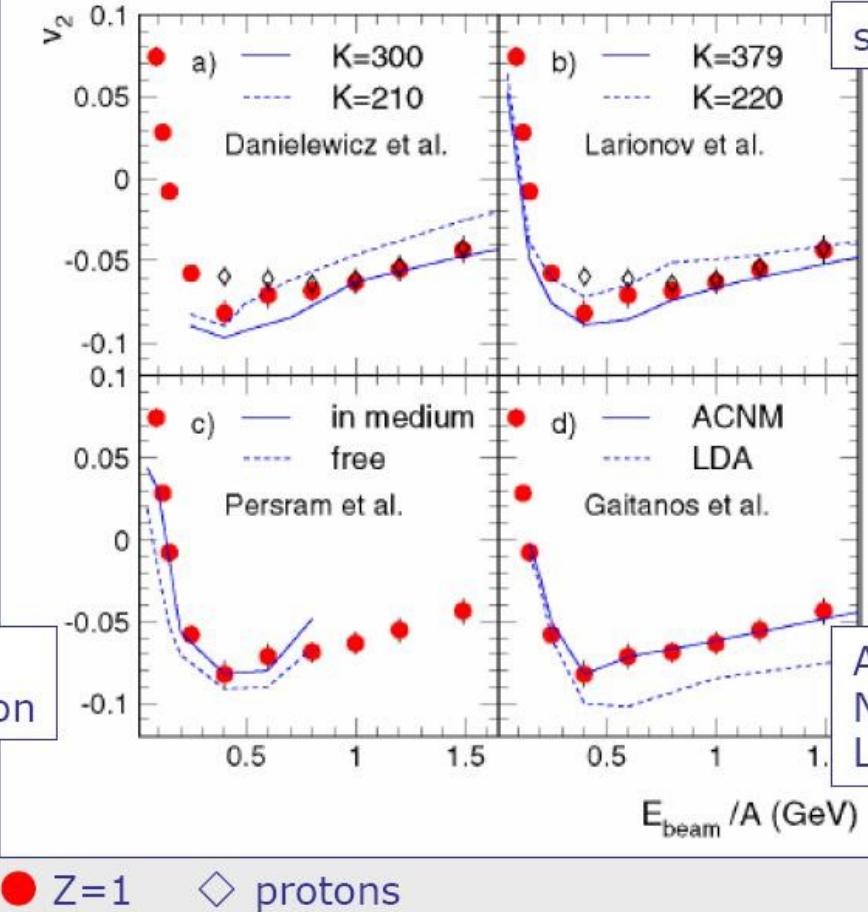
P. Danielewicz, R. Lacey,
W.G. Lynch,
Science 298 (2002) 1592

D. Persram, C. Gale,
PRC 65 (2002) 064611

in-medium
nucleon-nucleon Xsection

Andronic et al.

FOPPI Collaboration / Physics Letters B 612 (2005) 173–180

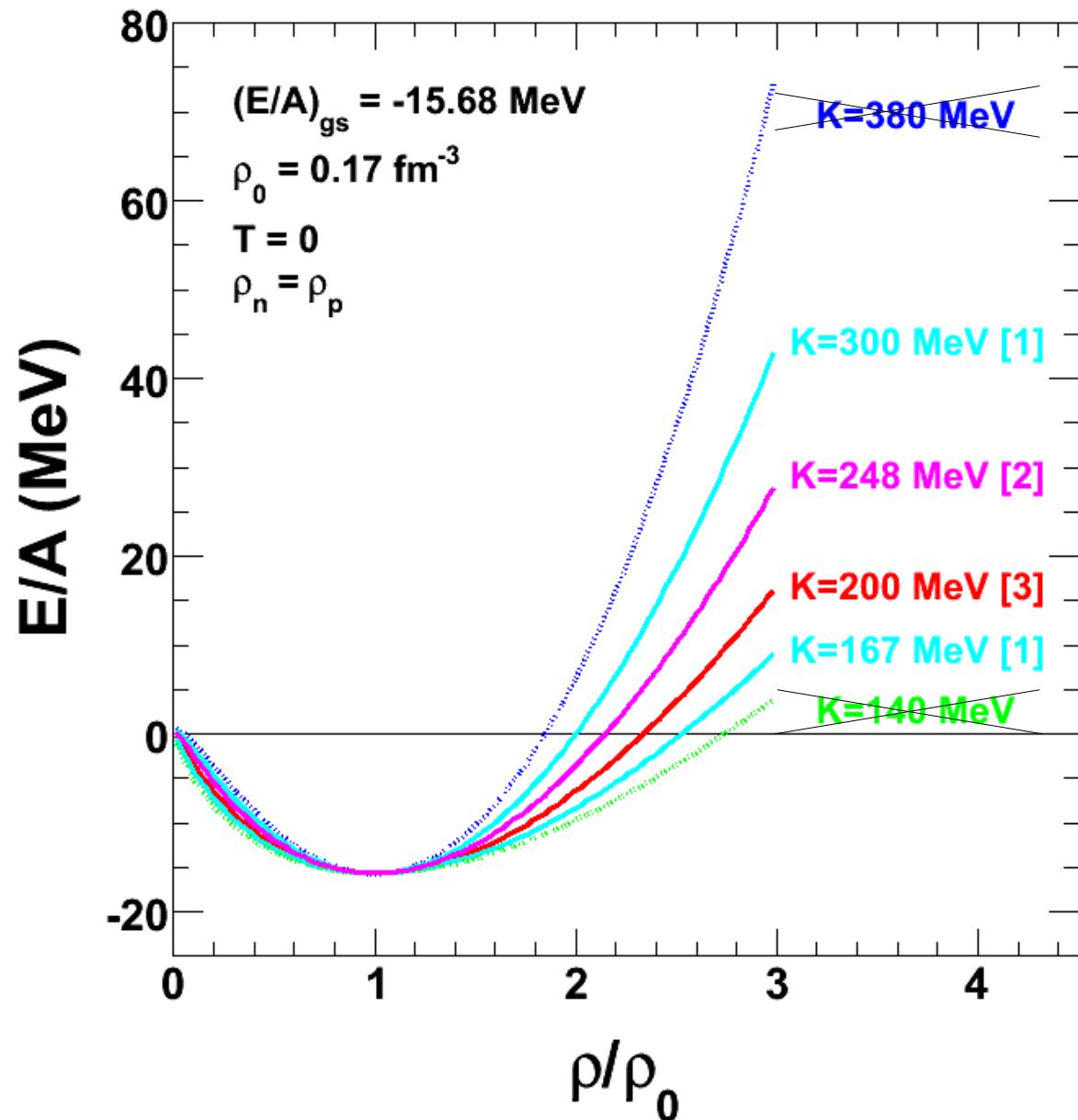


A.B. Larionov, W. Cassing,
C. Greiner, U. Mosel,
PRC 62 (2000) 064611

T. Gaitanos, C. Fuchs,
H.H. Wolter,
NPA 741 (2004) 287

ACNM: Asymmetric Colliding
Nuclear Matter Approx.
LDA: Local Density Approx.

Symmetric matter ($\delta=0$)



[1] Flow:
P. Danielewicz et al., Science 298 (02) 1592
 $K = 167\text{-}300$ MeV

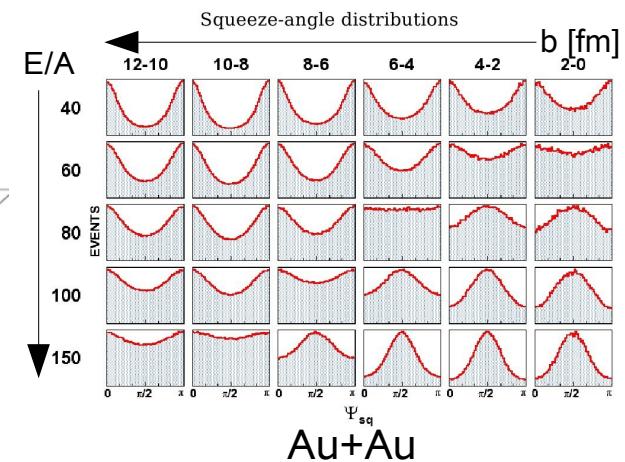
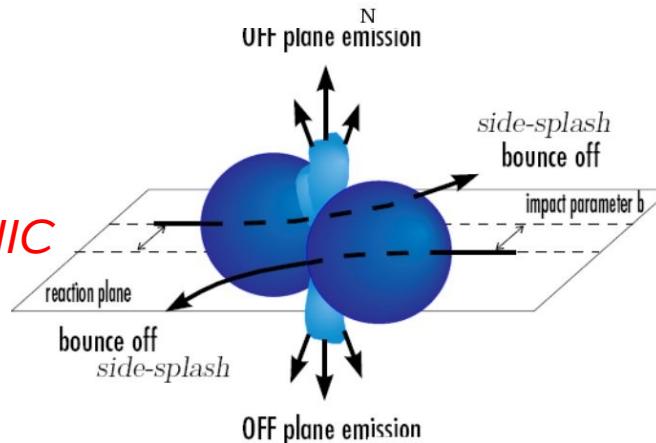
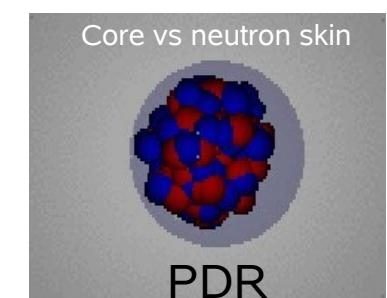
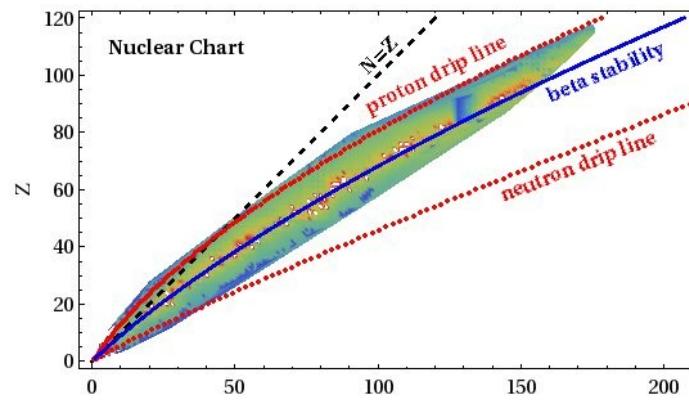
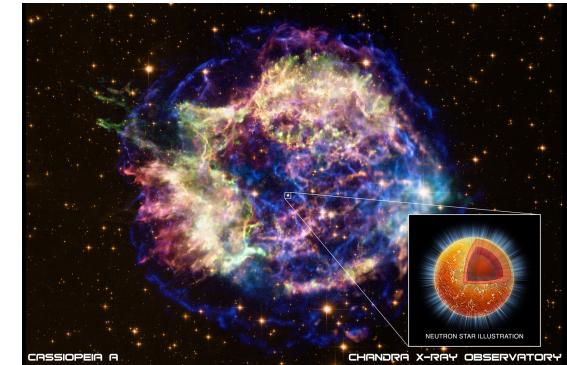
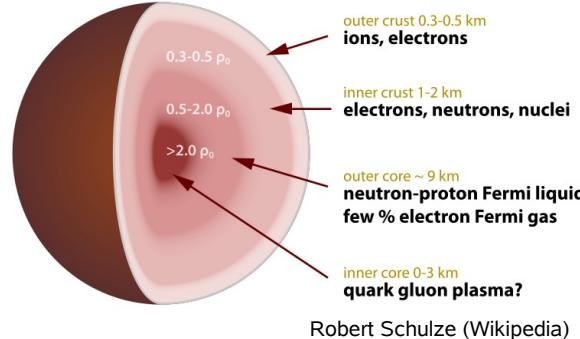
[2] ISGMR:
J. Piekarewicz, PRC 69 (04) 041301
RMF: $K=248$ MeV
G. Colò et al., PRC 70 (04) 024307
Skyrme HF: $K=230$ MeV

[3] Subthreshold K^+ :
C. Sturm et al., PRL 86 (01) 39
Ch. Hartnack et al., PRL 96 (06) 012302
 $K=200$ MeV

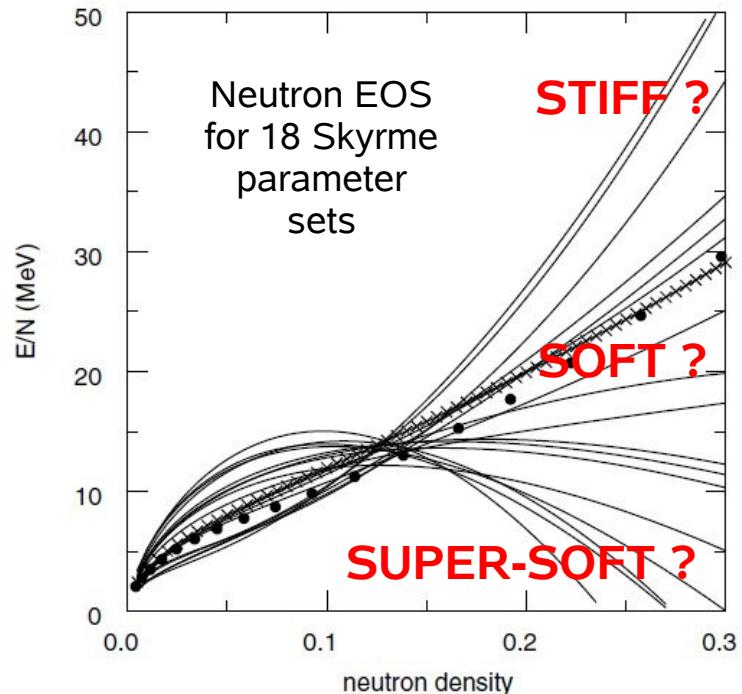
Model and data dependent
values of K
Still $\sim 20\text{-}30\%$ uncertainty

Symmetry energy affects:

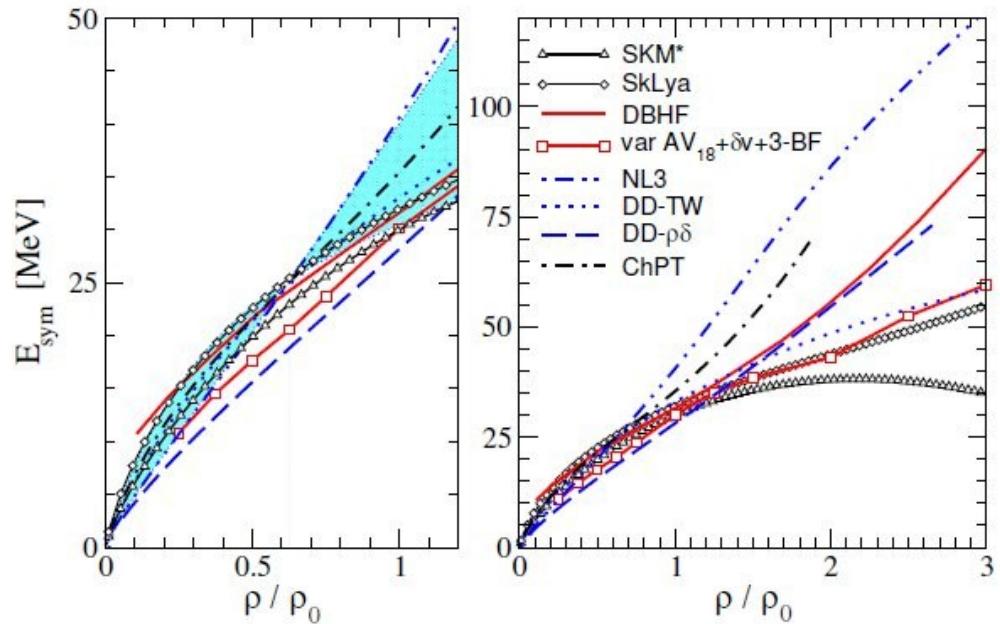
- Neutron star structure, composition, size, mass and cooling
- Supernova explosions
- Structure of the nucleus (masses, drip lines)
- Neutron skin thickness
- IvGDR
- Pygmy resonances
- Differences between IAS
- *Flow patterns in HIC*
- *n/p, t/³He, π/π⁺, K⁺/K⁰ ratios in HIC*



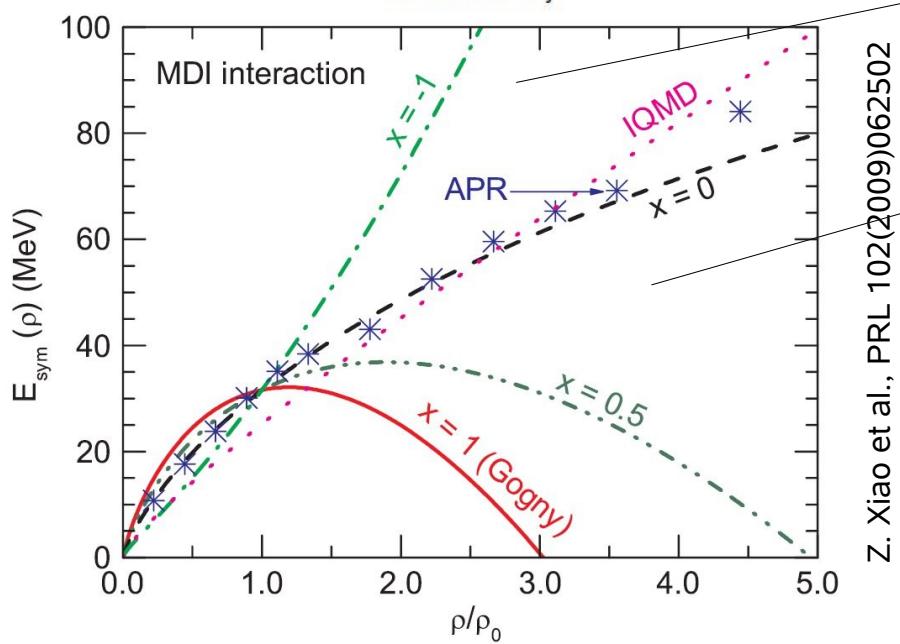
Symmetry term. Why so uncertain?



B. Alex Brown, PRL 85(2000)5296



C. Fuchs and H.H. Wolter, EPJA 30(2006)5



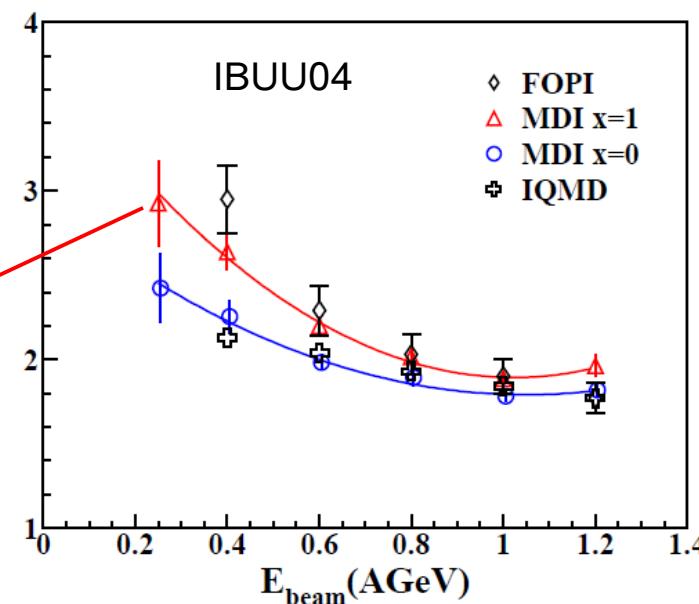
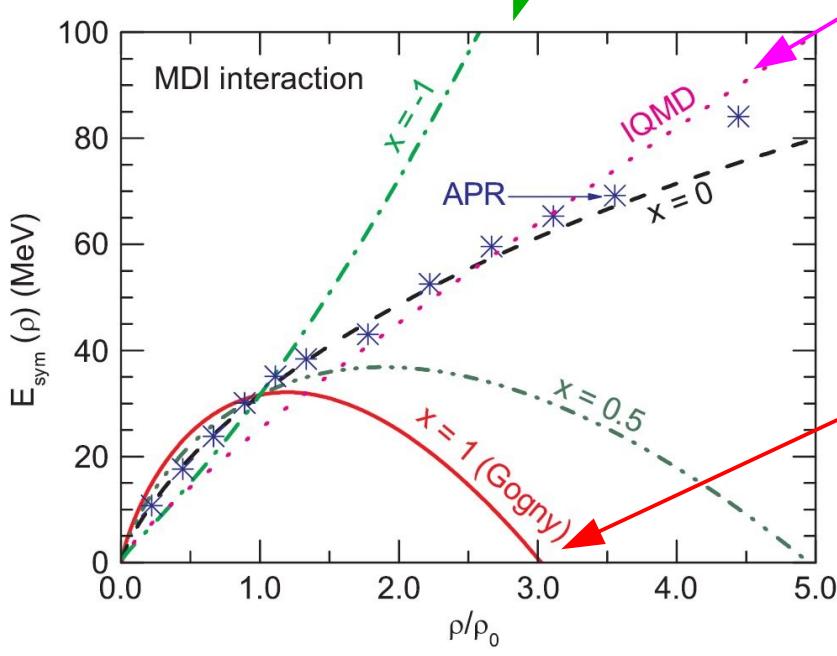
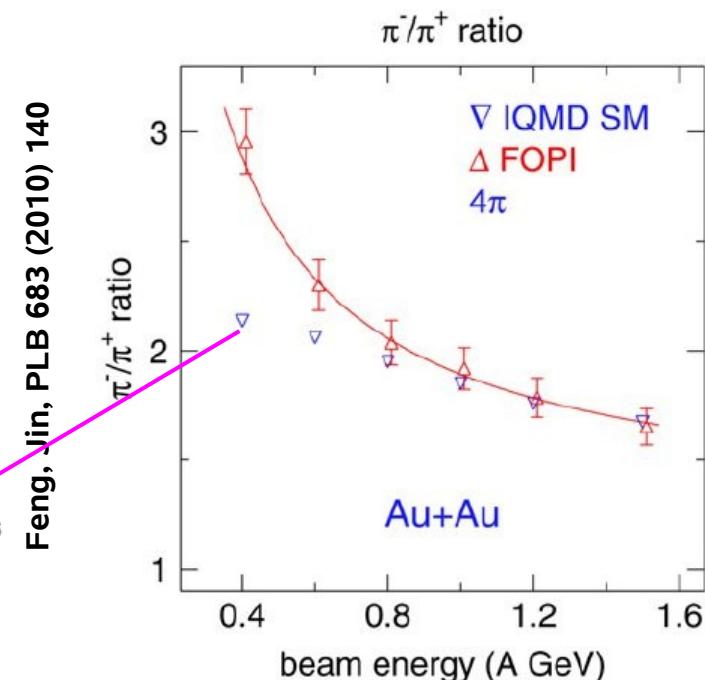
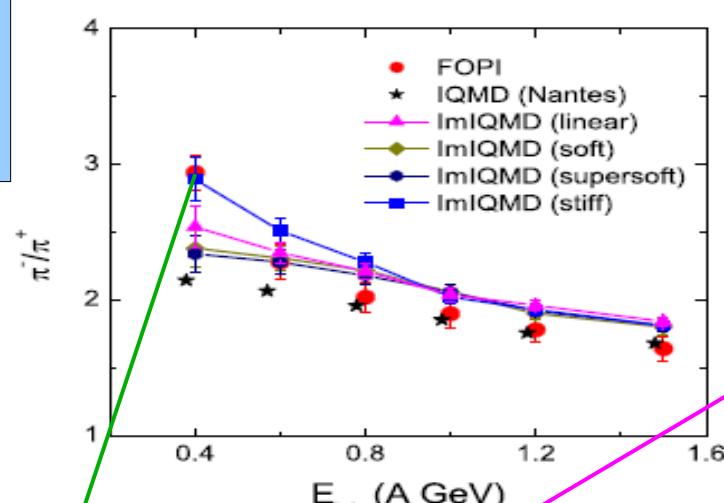
Thicker neutron skins,
larger proton fraction in NS

Thinner neutron skins,
smaller proton fraction in NS

Symmetry energy uncertain at high density
and modified by clustering at low density

Pion ratios (FOPI, Au+Au)

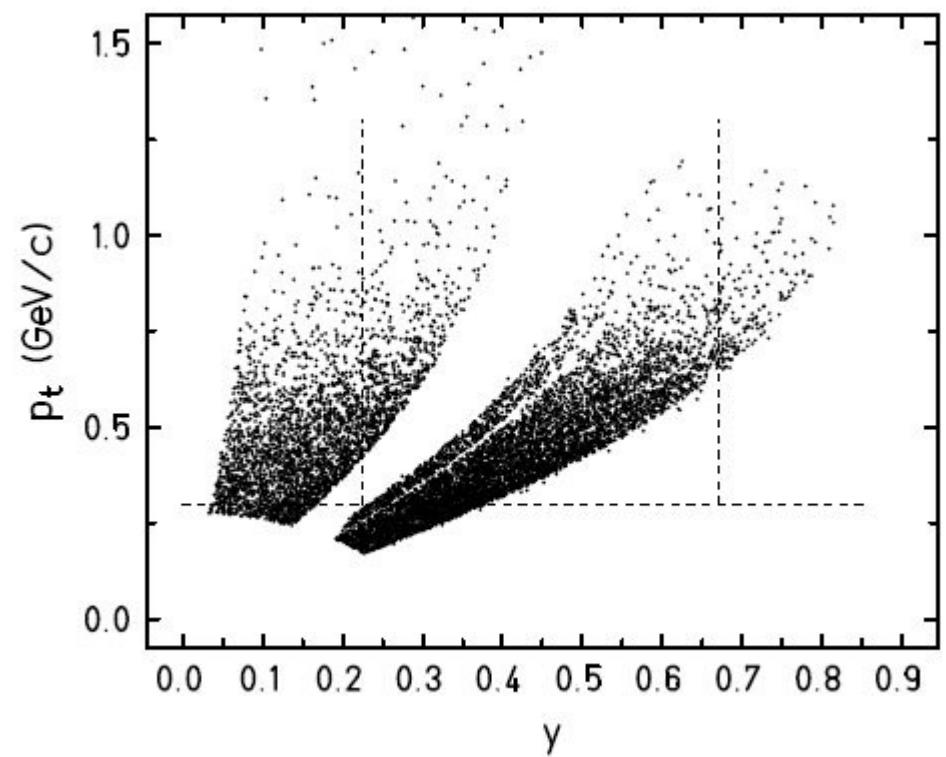
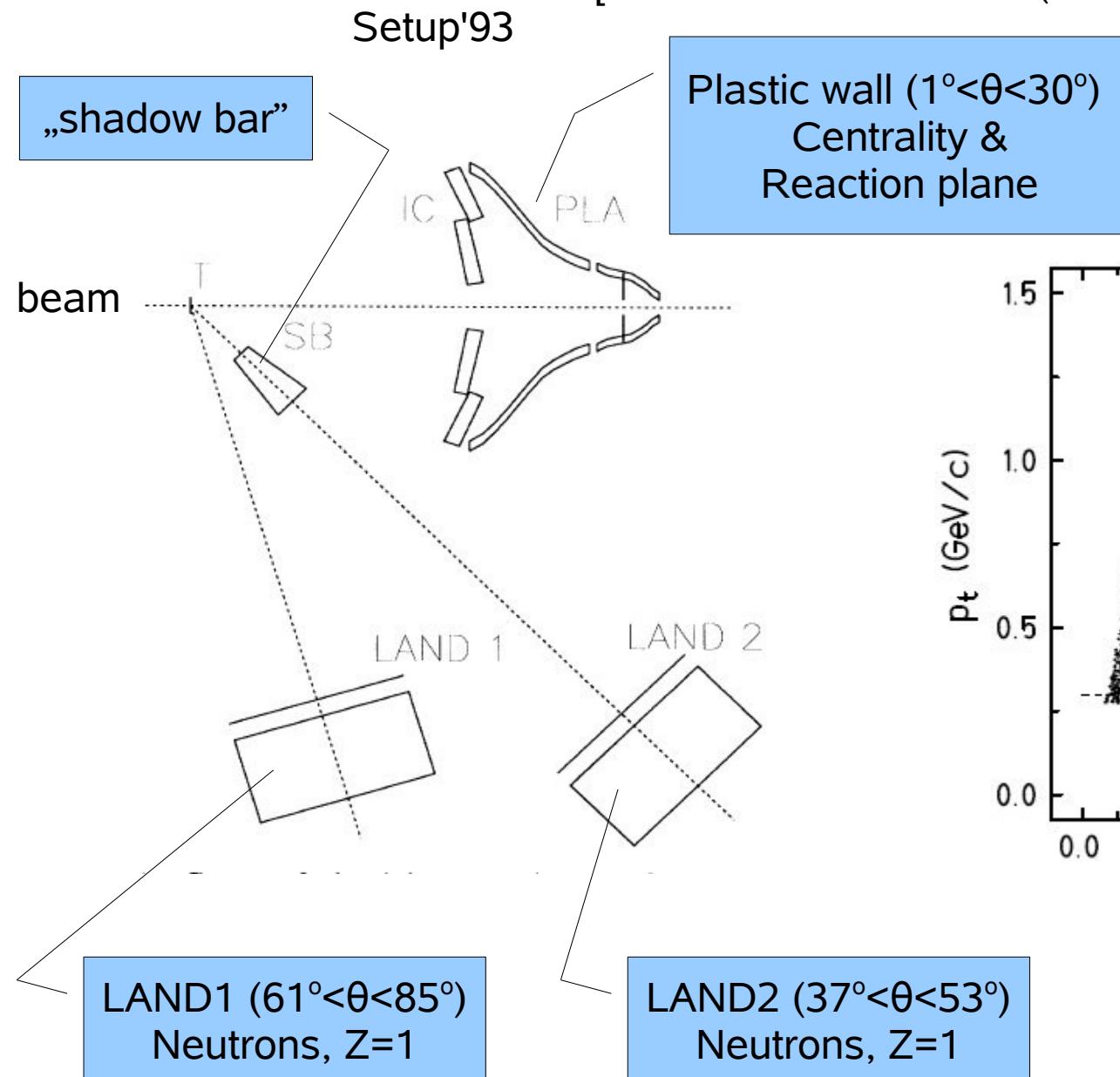
$$\frac{\pi^-}{\pi^+} = \frac{5n^2 + np}{5p^2 + np} \approx \left(\frac{n}{p}\right)^2$$



Z. Xiao et al., PRL 102 (2009) 062502

Reanalysis of FOPI/LAND Au+Au @ 400, 600, 800 AMeV

[Y. Leifels et al. PRL 71 (1993) 963]



Acceptance for neutrons
and cuts for analysis (dashed)

Reanalysis of FOPI/LAND Au+Au @ 400, 600, 800 AMeV

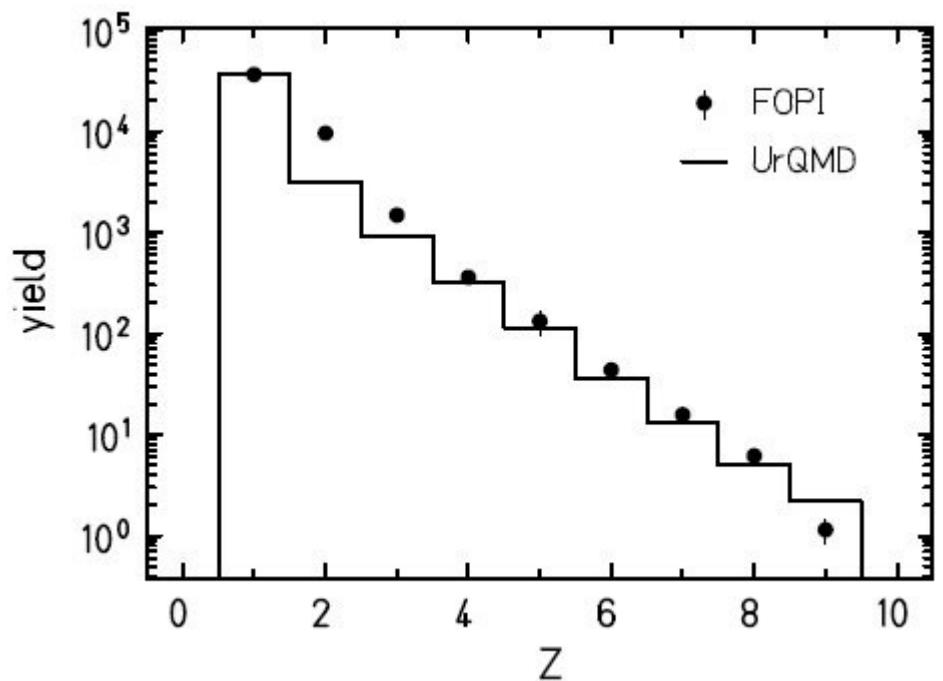
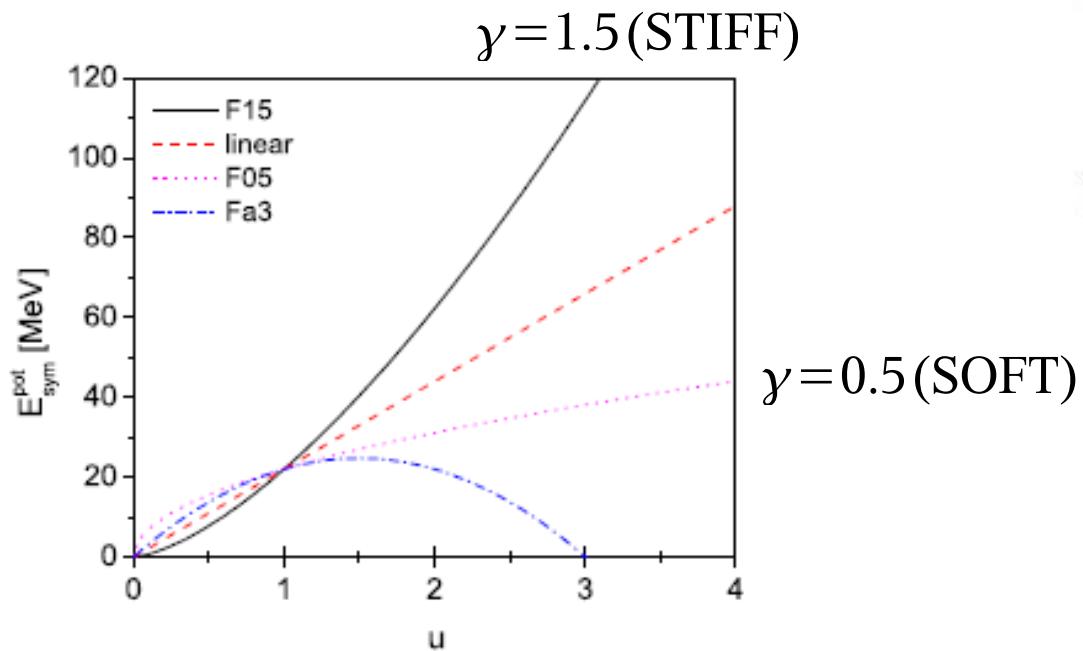
[Y. Leifels et al. PRL 71 (1993) 963] -> [P. Russotto et al. PLB 697 (2011) 471]

UrQMD, Q. Li, J.Phys. G 31(2005)1359

P. Russotto et al. PLB 697 (2011) 471
Data: W. Reisdorf, et al., NPA 612 (1997) 493
Central collisions, Au+Au @ 400 AMeV

„Fermi-gas“ parametrization of the symmetry term:

$$E_{sym} = E_{sym}^{pot} + E_{sym}^{kin} = 22 \text{ MeV} \left(\frac{\rho}{\rho_o} \right)^{\gamma} + 12 \text{ MeV} \left(\frac{\rho}{\rho_o} \right)^{2/3}$$

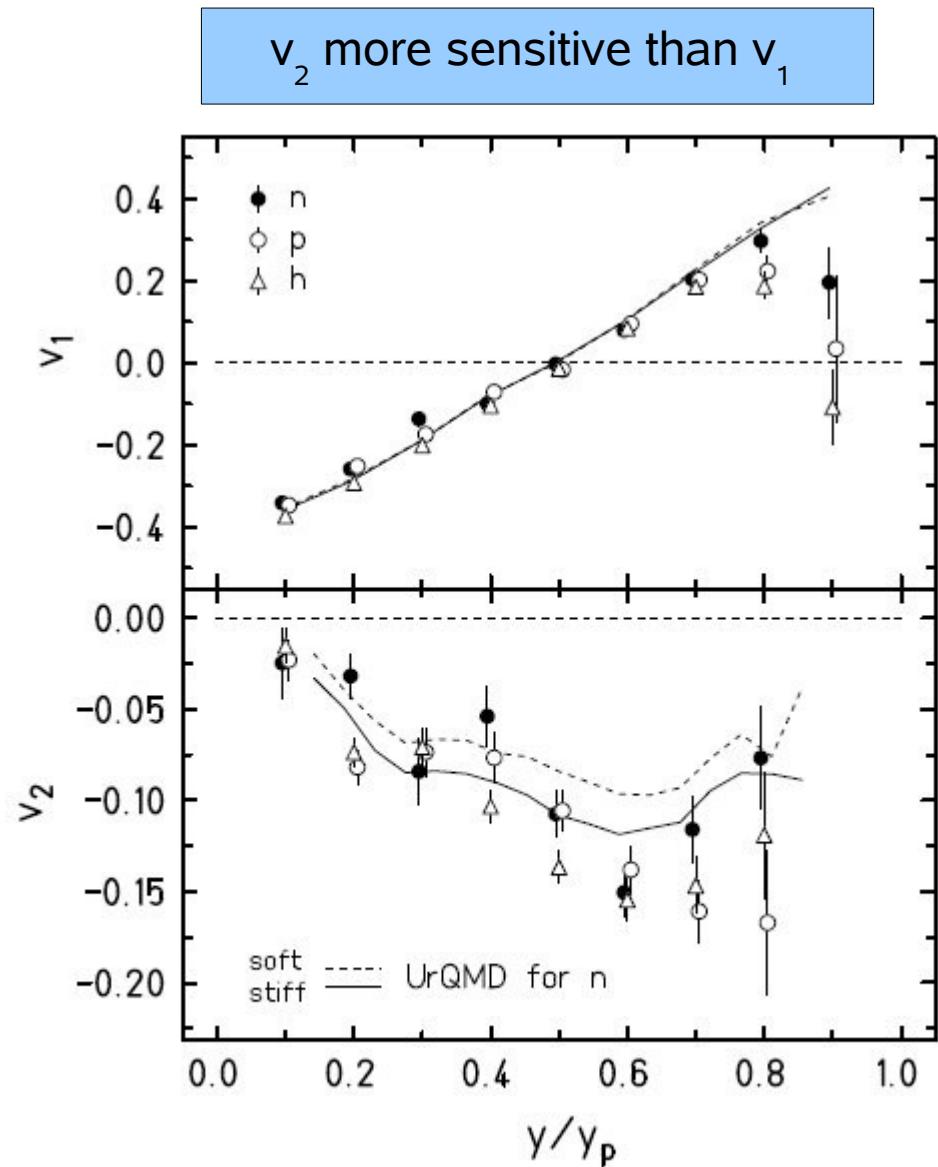
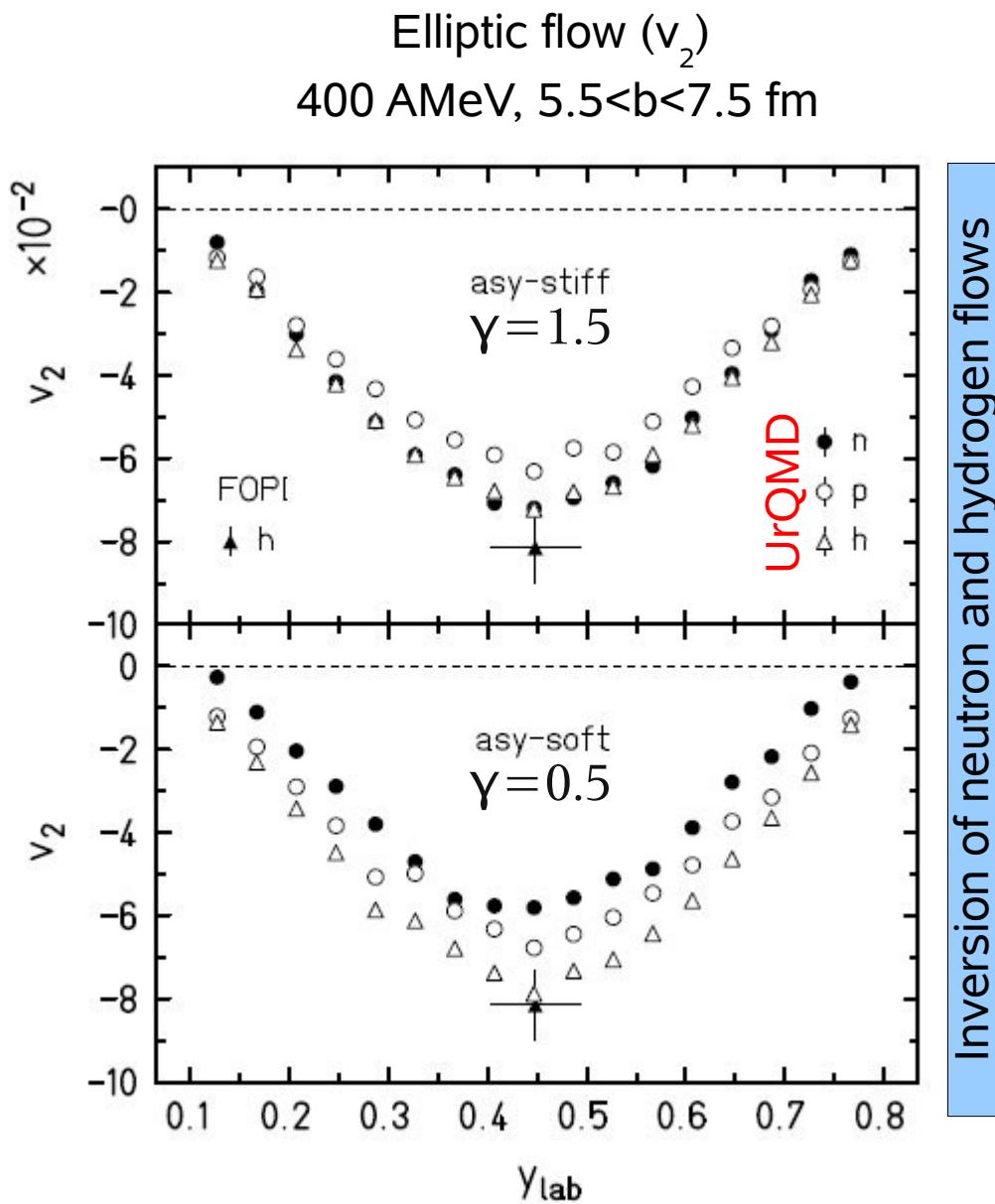


Fragment yields
(test of clustering procedure
with $\Delta r=3$ fm and $\Delta p=275$ MeV/c)
Normalization at $Z=1$

d,t, α underpredicted (x 2-3)

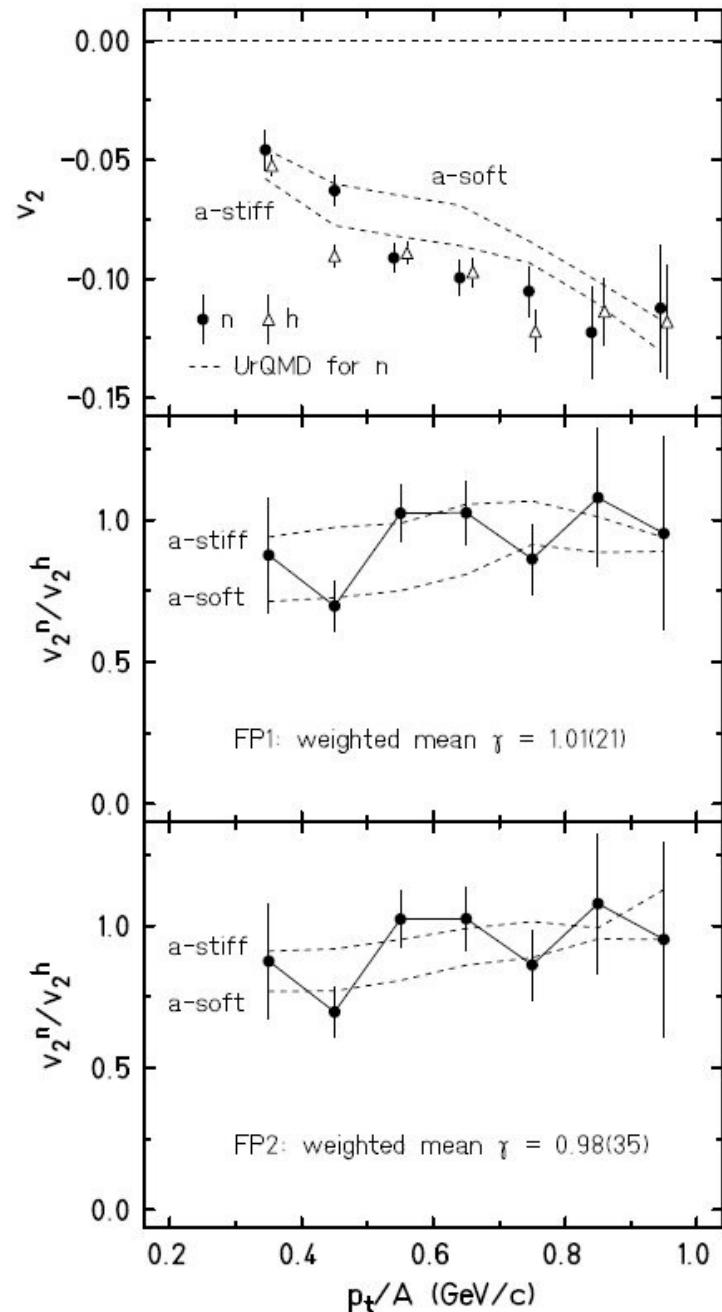
Reanalysis of FOPI/LAND Au+Au @ 400, 600, 800 AMeV

[P. Russotto et al. PLB 697 (2011) 471]



Reanalysis of FOPI/LAND Au+Au @ 400, 600, 800 AMeV

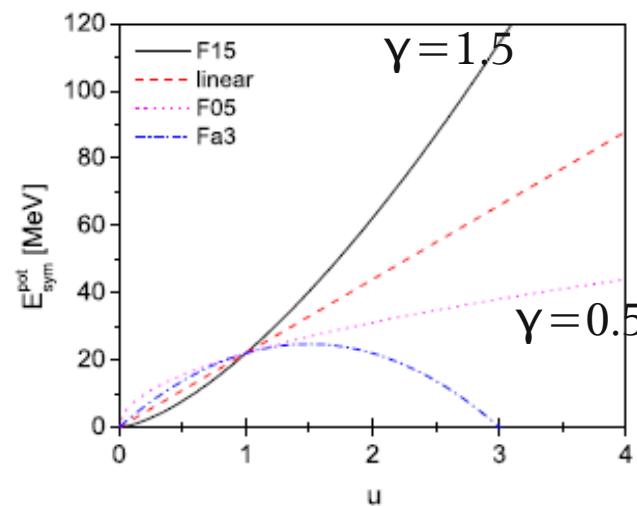
[P. Russotto et al. PLB 697 (2011) 471]



v_2^n/v_2^H ratio vs p_T sensitive to the symmetry term

neutron/hydrogen
 FP1: $\gamma = 1.01 \pm 0.21$
 FP2: $\gamma = 0.98 \pm 0.35$

neutron/proton
 FP1: $\gamma = 0.99 \pm 0.28$
 FP2: $\gamma = 0.85 \pm 0.47$
 adopted: $\gamma = 0.9 \pm 0.4$

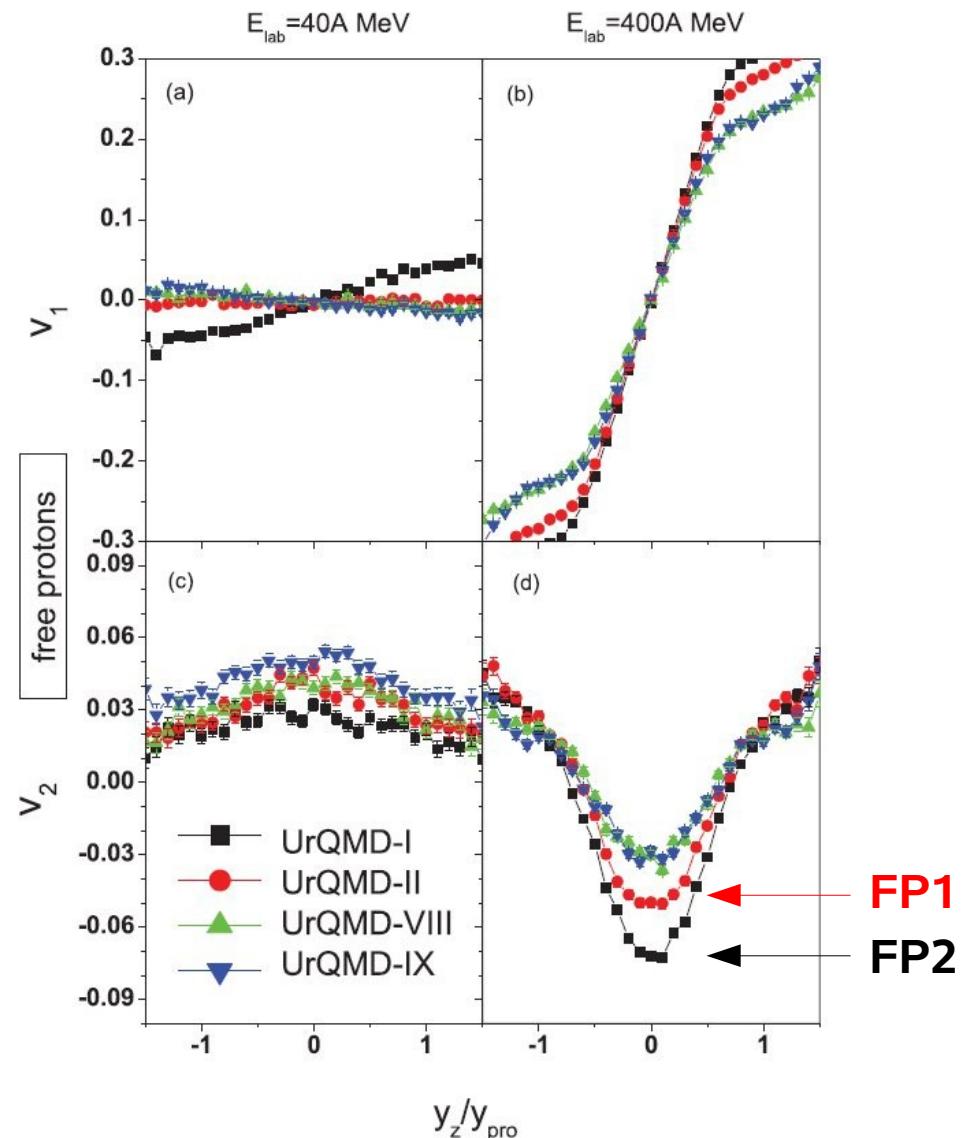
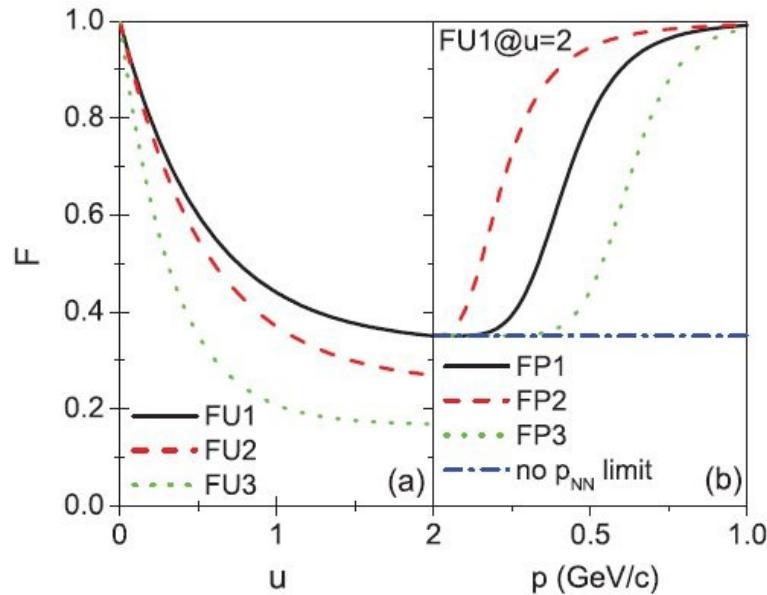


FP1, FP2: different parametrizations of the Momentum dependence of the elastic nucleon-nucleon cross section
 [Q. Li et al. PRC 83(2011)044617]

Medium correction factors to the elastic σ_{NN}

[UrQMD, Q. Li et al., PRC 83(2011)044616]

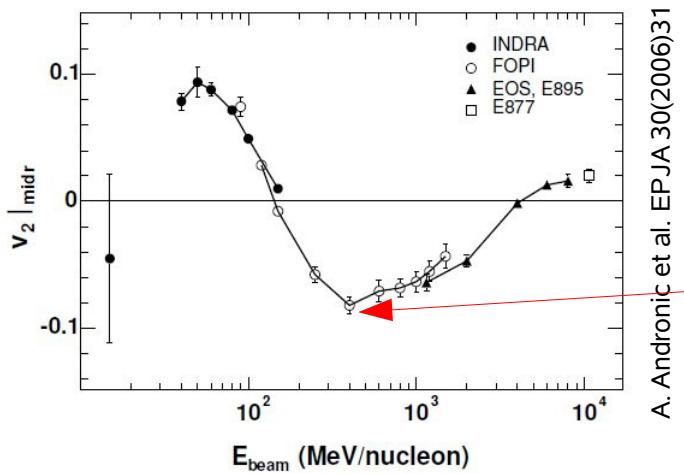
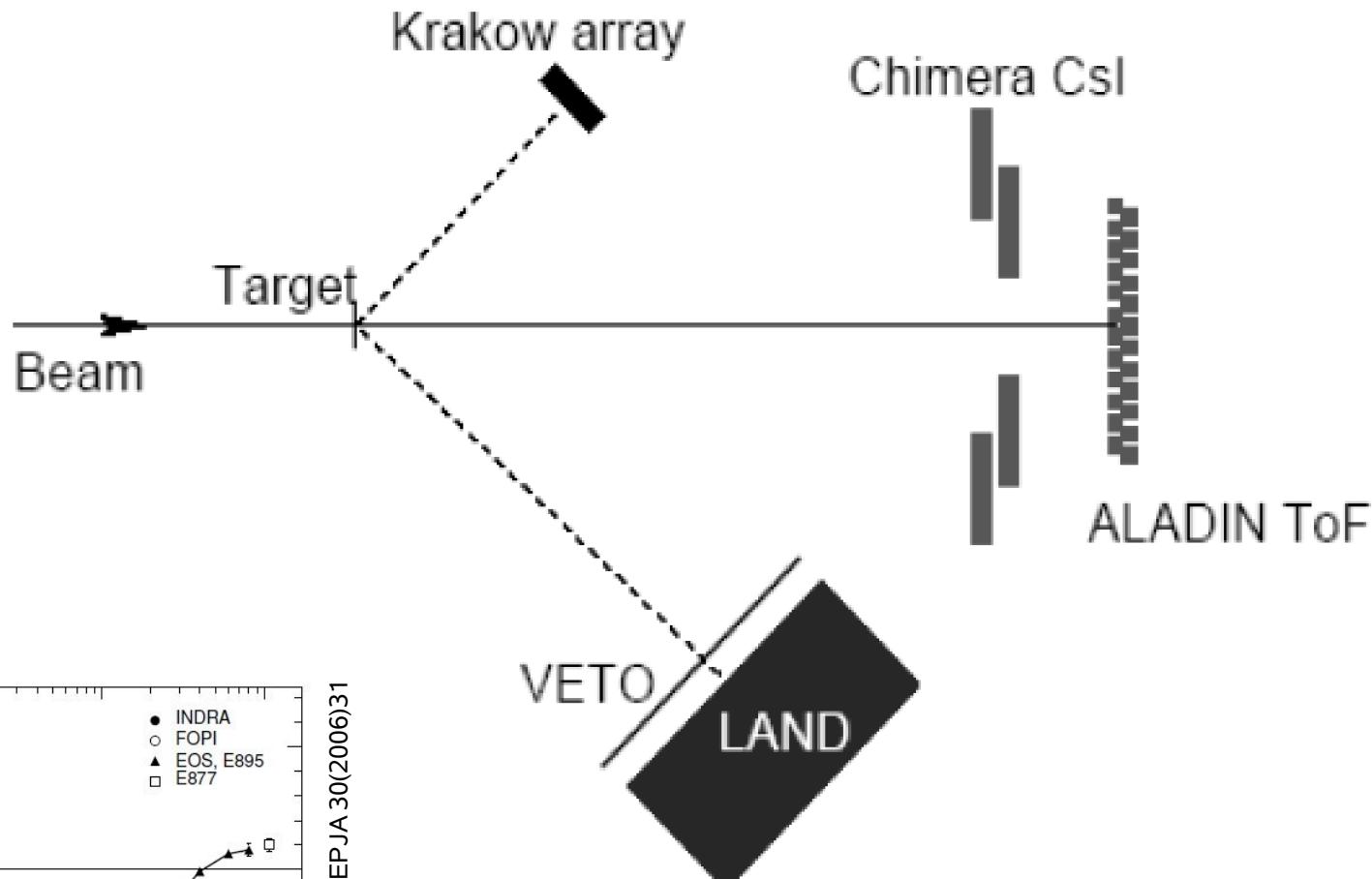
$$\sigma_{\text{tot}}^* = \sigma_{\text{in}} + \sigma_{\text{el}}^* = \sigma_{\text{in}}^{\text{free}} + F(\rho, p) \sigma_{\text{el}}^{\text{free}}$$



ASY-EOS experimental setup

May 2011

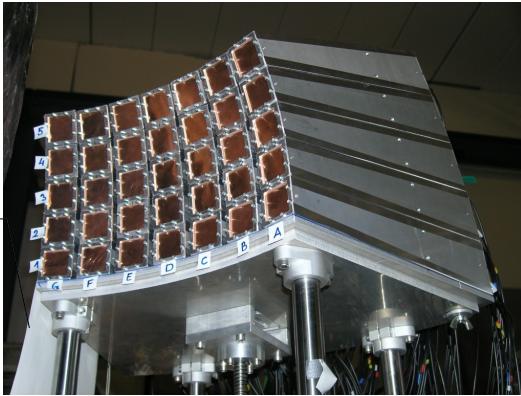
Setup from the proposal of 2009



$^{197}\text{Au} + ^{197}\text{Au}$	@ 400 AMeV	$\delta^2 = 0.039$
$^{96}\text{Zr} + ^{96}\text{Zr}$	@ 400 AMeV	$\delta^2 = 0.028$
$^{96}\text{Ru} + ^{96}\text{Ru}$	@ 400 AMeV	$\delta^2 = 0.007$

ASY-EOS

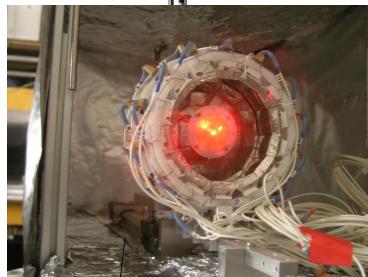
5x7 triple telescopes, $20^\circ < \theta < 60^\circ$
 Si-CsI-CsI
 Midrapidity pdt +
 Isotopes of $Z < 9$



Start + ROLU



Beam → Target



μ -Ball + Halo

4 rings, $\theta > 60^\circ$, CsI(Tl)
 Discriminate target vs air
 interactions, remove halo,
 possibly centrality + rpl



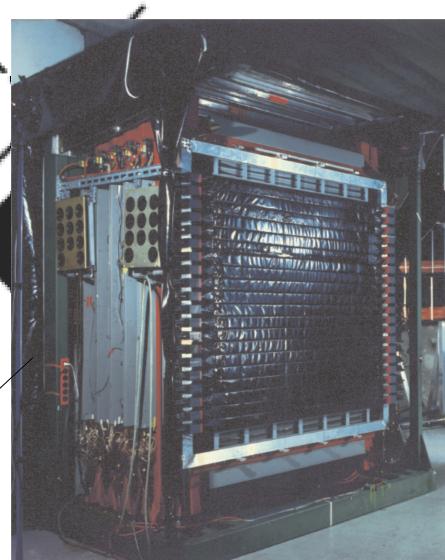
Chimera CsI



ALADIN ToF



SHADOW BAR

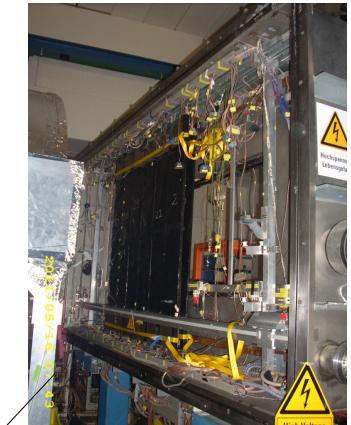


LAND+VETO

2x2x1 m³ plastic/Fe sandwich
 + plastic veto wall
 Midrapidity neutrons & $Z=1$

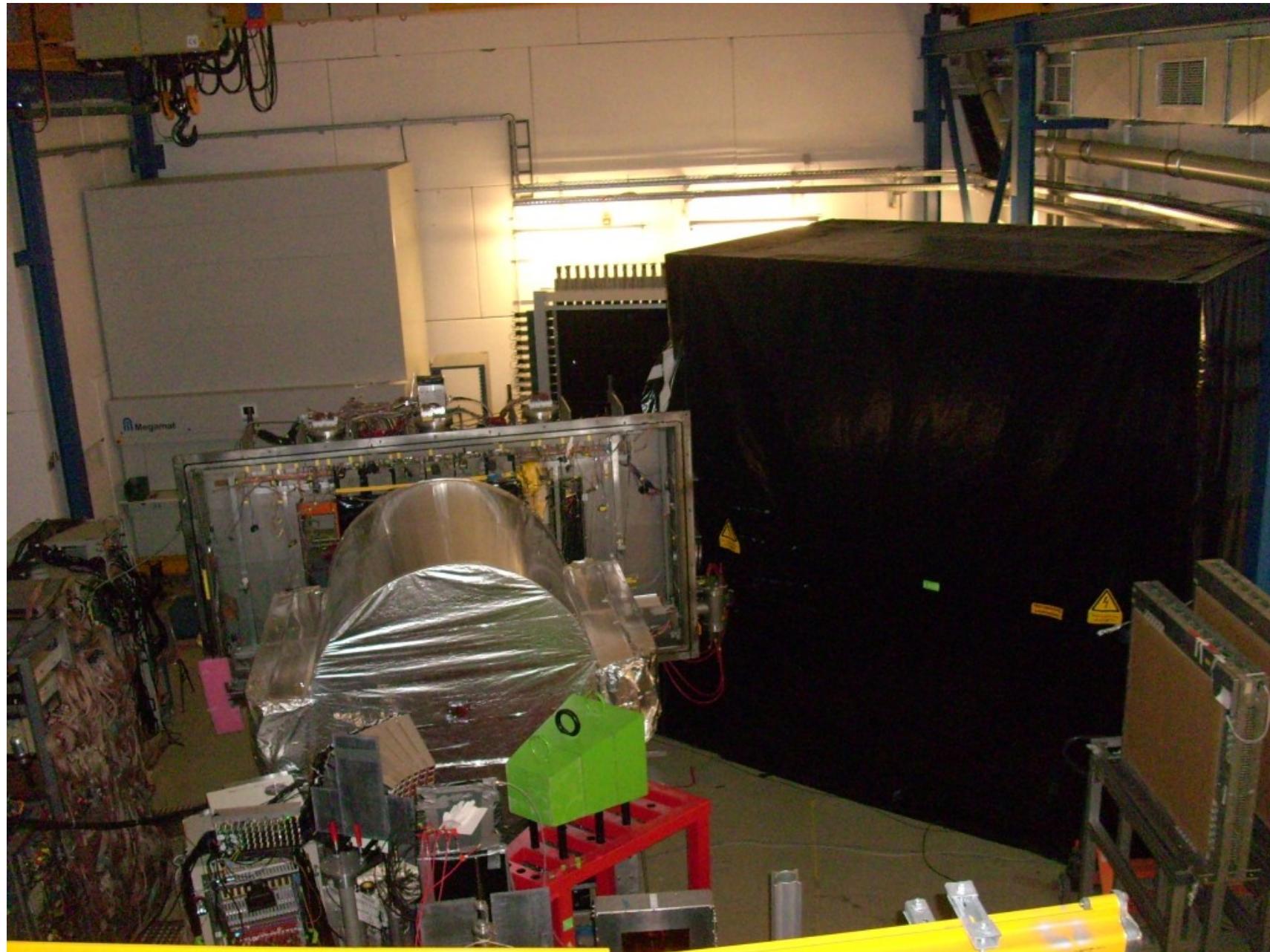
setup

4 rings, 352 CsI(Tl), $7^\circ < \theta < 20^\circ$
 Centrality
 &
 Reaction plane



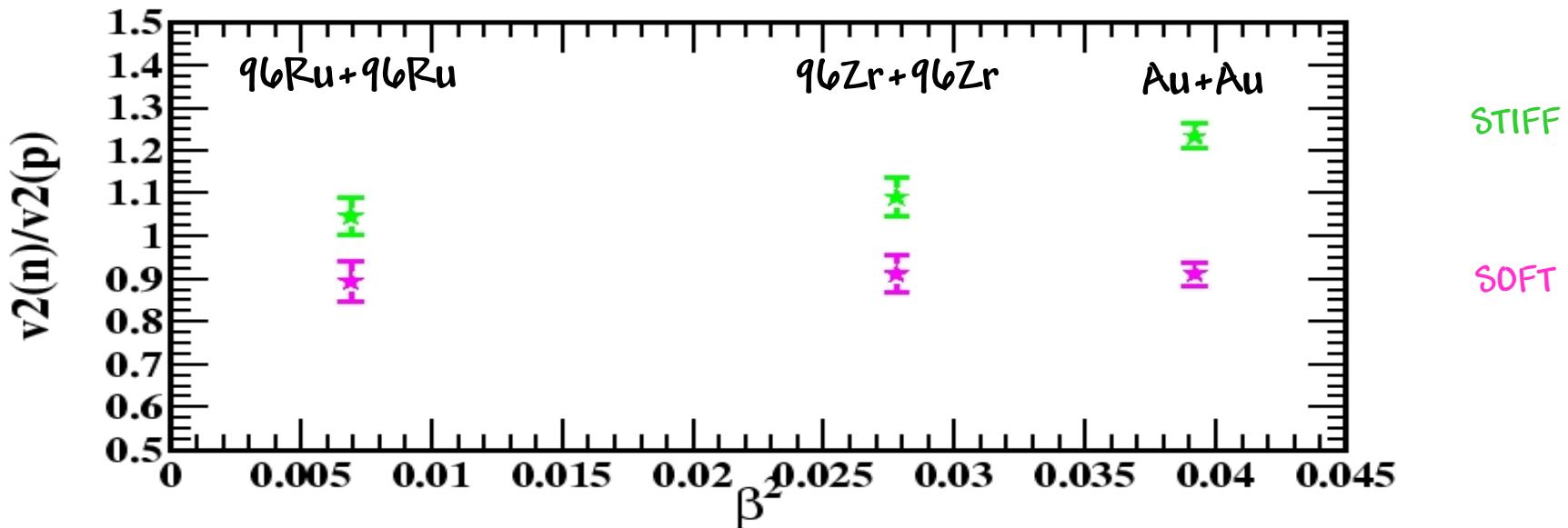
96 plastic bars
 x-y positions, centrality,
 reaction plane, trigger

ASY-EOS experimental setup



URQMD simulations: @ 400 AMeV

v_2 for $|(\gamma/\gamma_p)_{c.m.}| < 0.1$



$$\beta = (N-Z)/(N+Z)$$

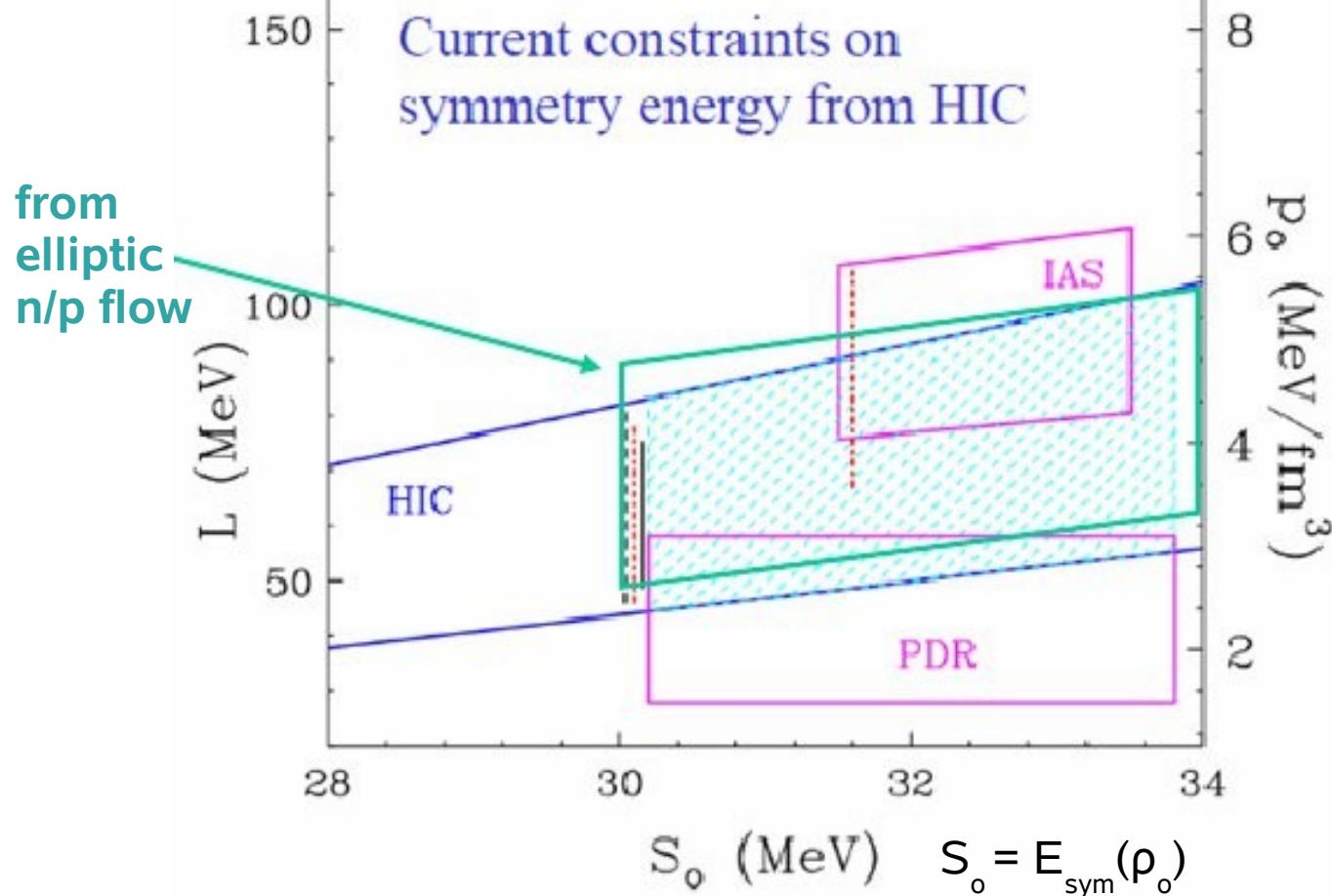
$\text{Au} + \text{Au}$ $b = 5.5 - 7.5$ fm

$^{96}\text{Zr} + ^{96}\text{Zr}$ $b = 4 - 6$ fm

$^{96}\text{Ru} + ^{96}\text{Ru}$ $b = 4 - 6$ fm

summary

$$L = 3 \rho_o \left. \frac{\partial E_{sym}}{\partial \rho} \right|_{\rho=\rho_o}$$



IAS -
isobaric analog states
Danielewicz/Lee 2008

HIC -
heavy-ion collisions
isospin diffusion, n/p ratios
Tsang et al., 2009

PDR -
pygmy dipole resonance
Klimkiewicz et al. 2007

$$P_o = (L/3) \rho_o$$

symmetry pressure

Difficulties in measuring the $E_{\text{sym}}(\rho)$

Experiment

- Mixture of density, temperature and time dependent processes
- Detection of neutrons and protons simultaneously
- Tiny effects – high precision and statistics needed
- Observables minimizing the influence of the isoscalar part
- Correlations of many observables needed
- Exotic beams (larger δ) would help

Model

- In-medium cross sections (ρ and p dependent)
- Realistic inelastic cross sections, particle production (π, K)
- Momentum dependence of the mean-field
- Control the competition between the mean-field and collisions
- Realistic description of cluster formation (at least $t/{}^3He$)
- Ability to describe „hot” and „cold” observables

The ASY-EOS Collaboration

Co-Spokespersons: R.C. Lemmon¹ and P. Russotto²

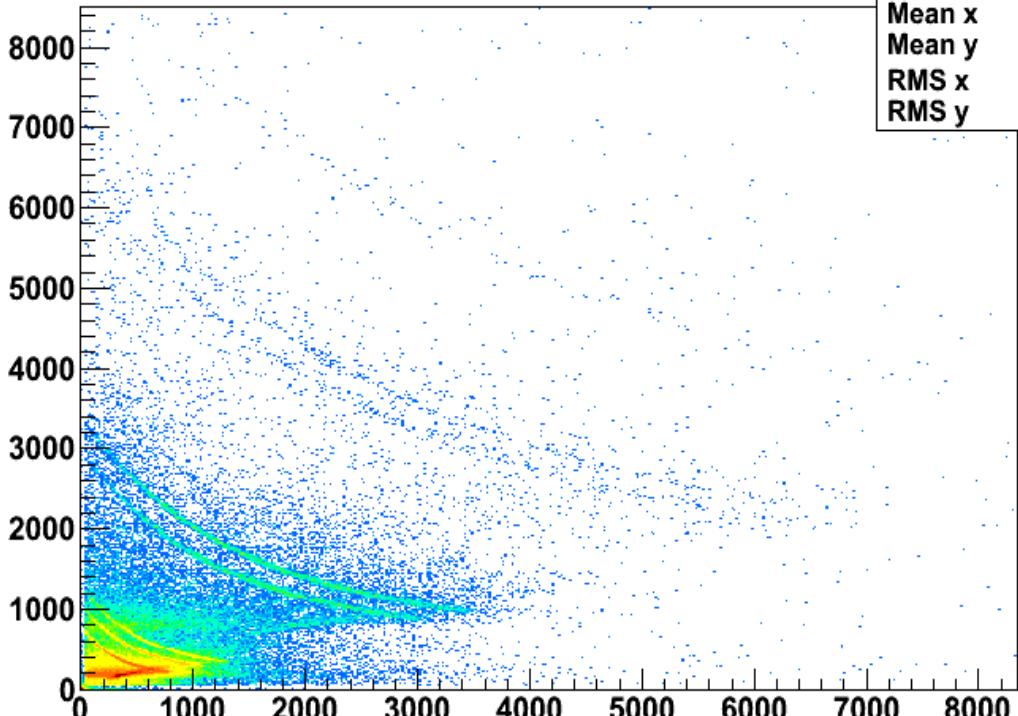
Collaboration

F. Amorini², A. Anzalone¹⁷, T. Aumann³, V. Avdeichikov¹², V. Baran²³, Z. Basrak⁴, J. Benlliure¹³, I. Berceanu¹¹, A. Bickley¹⁴, E. Bonnet⁶, K. Boretzky³, R. Bougault³⁰, J. Brzychczyk⁸, B. Bubak²², G. Cardella⁷, S. Cavallaro², J. Cederkall¹², M. Chartier⁵, M.B. Chatterjee¹⁶, A. Chbihi⁶, M. Colonna¹⁷, D. Cozma¹¹, B. Czech¹⁰, E. De Filippo⁷, K. Fissum¹², D. Di Julio¹², M. Di Toro², M. Famiano²⁷, J.D. Frankland⁶, E. Galichet¹⁸, I. Gasparic⁴, E. Geraci¹⁵, V. Giordano², P. Golubev¹², L. Grassi¹⁵, A. Grzeszczuk²², P. Guazzoni²¹, M. Heil³, J. Helgesson³¹, L. Isaksson¹², B. Jacobsson¹², A. Kelic³, M. Kis⁴, S. Kowalski²², E. La Guidara²⁰, G. Lanzalone²⁹, N. Le Neindre³⁰, Y. Leifels³, Q. Li⁹, I. Lombardo², O. Lopez³⁰, J. Lukasik¹⁰, W. Lynch¹⁴, P. Napolitani³⁰, N.G. Nicolis²⁴, A. Pagano⁷, M. Papa⁷, M. Parlog³⁰, P. Pawlowski¹⁰, M. Petrovici¹¹, S. Pirrone⁷, G. Politi¹⁵, A. Pop¹¹, F. Porto², R. Reifarth³, W. Reisdorf³, E. Rosato¹⁹, M.V. Ricciardi³, F. Rizzo², W.U. Schroder²⁸, H. Simon³, K. Siwek-Wilczynska²⁶, I. Skwira-Chalot²⁶, I. Skwirczynska¹⁰, W. Trautmann³, M.B. Tsang¹⁴, G. Verde⁷, E. Vient³⁰, M. Vigilante¹⁹, J.P. Wieleczko⁶, J. Wilczynski²⁵, P.Z. Wu⁵, L. Zetta²¹, W. Zipper²²

Symmetry energy from elliptic flow in $^{197}\text{Au} + ^{197}\text{Au}$ [PLB 697 (2011) 471]

P. Russotto, P.Z. Wu, M. Zoric, M. Chartier, Y. Leifels,
R.C. Lemmon, Q. Li, J. Łukasik, A. Pagano, P. Pawłowski,
W. Trautmann

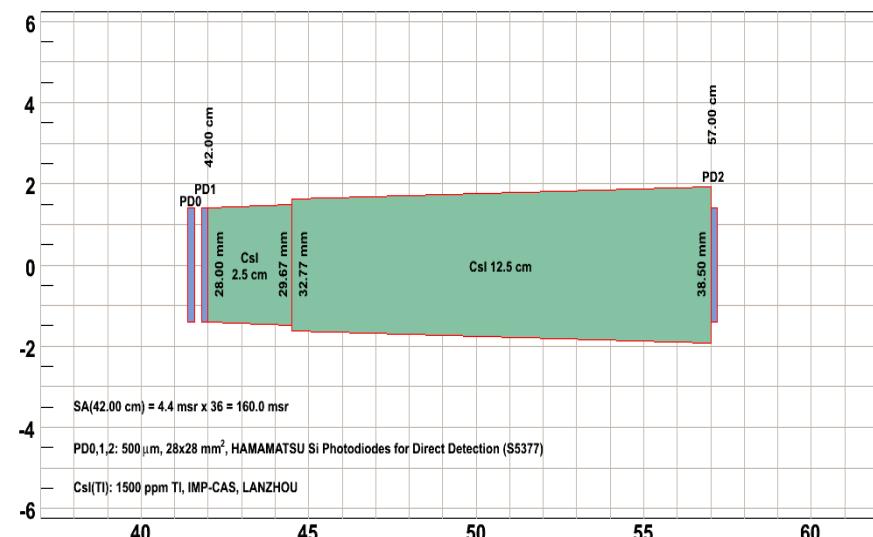
amp12_00 "A11" vs "A12"



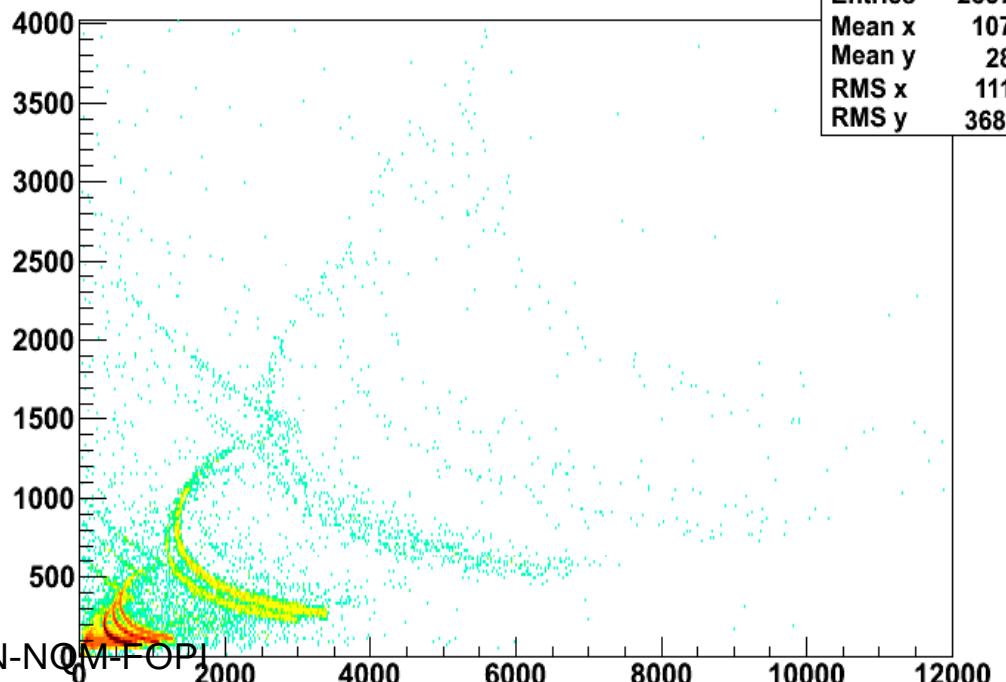
amp12_00
Entries 168692
Mean x 616.1
Mean y 470.5
RMS x 645
RMS y

TRIPLE TELESCOPE

Krakow array



amp01_00 "A10" vs "A11"



amp01_00

Differential flow

(minimizes the influence of the isoscalar part of the EOS)

Bao-An Li, PRL 85 (2000) 4221

$$F_{n-p}^x(y) \equiv \frac{1}{N(y)} \sum_{i=1}^{N(y)} p_i^x(y) \tau_i = \frac{N_n(y)}{N(y)} \langle p_n^x(y) \rangle - \frac{N_p(y)}{N(y)} \langle p_p^x(y) \rangle$$

where $N_n(y)$ is the total number of free nucleons at rapidity y , $p_i^x(y)$ is the transverse in-plane momentum of particle i and $\tau_i = 1$ (-1) for neutrons (protons).