

(arXiv:1009.1161, in collaboration with G.

 $D'Amhrecio and L'annialio)$ 

### Why holography for HLBL? -

- Holographic principle: Conjectured duality between strongly coupled gauge theories in  $d$  dimensions and gravitational weakly coupled theories in  $(d + 1)$ dimensions.  ${\sf AdS}/\mathsf{CFT}$  one of the most powerful examples.
- Lagrangian formulation: consistent treatment of the different channels (no double counting).
- • (leading) short distances: automatically implemented due to the conformal invariance of AdS.
- Large- $N_c$ : full realization of the large- $N_c$  limit  $(\infty$  resonances with short distances) distances).
- number of parameters: minimal.

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### Motivating questions

- Which are the parameters that mostly affect the uncertainty on the HLBL?
- Given

$$
K(Q_1^2, Q_2^2) \simeq 1 + \hat{\alpha} (Q_1^2 + Q_2^2) + \hat{\beta} Q_1^2 Q_2^2 + \hat{\gamma} (Q_1^4 + Q_2^4)
$$

what is the impact of the low energy parameters  $\alpha$ ,  $\beta$  and  $\gamma$ ?  $\alpha$  determined by CLEO, but  $\gamma$  and  $\beta$  out of current experimental reach. Can one obtain reasonable predictions?

- Is vector meson dominance <sup>a</sup> good enoug<sup>h</sup> approximation?
- How accurate is the pion-pole approximation?

## The setting -

 $\bullet$  AdS $_5$  space with metric

$$
ds^2 = g_{MN}dx^M dx^N = \frac{1}{y^2}(-dy^2 + \eta_{\mu\nu}dx^{\mu}dx^{\nu})
$$

with  $\eta_{\mu\nu}$  mostly negative.



• Intuitively, UV boundary captures pQCD (leading term trivial: conformal limit reproduced by AdS metric). IR deals with non-perturbative physics (OPE andnon-OPE, confinement scale,  $S\chi SB$ , spectrum patterns)

## Holographic models

• There are different realizations of the holographic principle for QCD. Generically,

$$
S_5 = S_{\rm YM} + S_X + S_{\rm CS}
$$

where

$$
S_{\text{YM}} [\mathcal{B}_L, \mathcal{B}_R] = \text{tr} \int d^4x \int_0^{z_0} dz \ e^{-\Phi(z)} \frac{-1}{8g_5^2} w(z) \left[ \mathcal{F}_{(L)}^{MN} \mathcal{F}_{(L)MN} + \mathcal{F}_{(R)}^{MN} \mathcal{F}_{(R)MN} \right]
$$
  

$$
S_X [X] = \text{tr} \int d^4x \int_0^{z_0} dz \ e^{-\Phi(z)} w(z)^3 \left[ D^M X D_M X^{\dagger} + V(X^{\dagger} X) \right]
$$
  

$$
S_{\text{CS}} [\mathcal{B}_L, \mathcal{B}_R] = \frac{N_c}{24\pi^2} \int \text{tr} \left[ \mathcal{B}_L \mathcal{F}_L^2 - \frac{i}{2} \mathcal{B}_L^3 \mathcal{F}_L - \frac{1}{10} \mathcal{B}_L^5 - (L \to R) \right]
$$

- Notice that chiral symmetry is incorporated.
- The models differ mostly in  $S_X$  (spontaneous symmetry breaking) and  $(\Phi(z), z_0)$ (spectrum).
- Major drawback: plethora of models...
- Advantage: stability tests of phenomenological results.

# The  $F_{\pi^0\gamma\gamma}$  form factor

 $\bullet$  From holography  $F_{\pi^0\gamma\gamma}$  can be extracted from the Chern-Simons 5-dimensional term:[Grigoryan et al'08,09]

$$
S_{\text{CS}} [\mathcal{B}] = \frac{N_c}{24\pi^2} \int \text{tr} \left[ \mathcal{B}_L \mathcal{F}_L^2 - \mathcal{B}_R \mathcal{F}_R^2 \right] + \cdots
$$
  

$$
= \frac{N_c}{4\pi^2} \epsilon^{\mu\nu\rho\sigma} \int_0^{z_0} dz \, (\partial_z \beta) \int d^4 x \, \pi^a \left( \partial_\rho V_\mu^a \right) \left( \partial_\sigma \hat{V}_\nu \right)
$$

• Applying the holographic recipe for correlators:

$$
K(Q_1^2, Q_2^2) = -\int_0^{z_0} \mathcal{J}(Q_1, z) \mathcal{J}(Q_2, z) \partial_z \beta(z) dz
$$

• For instance,

$$
\mathcal{J}(Q, z) = Qz \left[ K_1(Qz) + I_1(Qz) \frac{K_0(Qz_0)}{I_0(Qz_0)} \right]
$$
  

$$
\beta(z) = 1 - \frac{z^2}{z_0^2}
$$

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# The  $F_{\pi^0\gamma\gamma}$  form factor (high energies)

 $\bullet$  $\bullet$  Master formula for the different holographic models. It only depends on  $z_0$ , which for consistency (matching with the  $\rho(770)$  mass) has to be

$$
z_0 = \frac{2}{g_5^2 f_\pi} = \frac{N_c}{6\pi^2 f_\pi}
$$

• In any model with (asymptotic) AdS metric, it can be shown that (at least) the leading short distance constraints from QCD are satisfied. Not surprising: theasymptotic behavior of  $\mathcal{J}(Q,z)$  fixed by AdS, while the pion wave function is only relevant at the origin:



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# The  $F_{\pi^0\gamma\gamma}$  form factor (low energies)

• Low energies will depend on the specific model. At small virtualities one canexpand  $K(Q_1^2,Q_2^2)$  in the form

$$
K(Q_1^2, Q_2^2) \simeq 1 + \hat{\alpha} (Q_1^2 + Q_2^2) + \hat{\beta} Q_1^2 Q_2^2 + \hat{\gamma} (Q_1^4 + Q_2^4)
$$

- Experimentally,  $\hat{\alpha} = -1.76 \pm 0.22$ ,  $\hat{\gamma}$  waiting for better statistics while  $\beta$ ˆchallenging (2 virtual photons needed).
- Predictions from holographic models:



• Compliance with experiment used as <sup>a</sup> holographic model filter. From the "acceptable" models we can then estimate the quartic terms to be

$$
\hat{\beta} = 3.33 \pm 0.32
$$
  

$$
\hat{\gamma} = 2.84 \pm 0.21
$$

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### Vector meson dominance revisited

• Spectral decomposition:

$$
\mathcal{J}(z,Q) = \sum_{n=1}^{\infty} \frac{f_n}{Q^2 + m_n^2} \psi_n(z); \qquad K(Q_1^2, Q_2^2) \equiv \sum_{k,l=1}^{\infty} \frac{B_{kl}}{(Q_1^2 + m_k^2)(Q_2^2 + m_l^2)}
$$

• Graphically:



• Each resonance contribution can be evaluated



• Lowest meson dominance qualitatively holds. Enough for the accuracy of HLBL?

## Estimation of the  $\pi^0$ -HLBL

• Strategy: encode the short and long distance information in an interpolator (DIP):

$$
K(q_1^2, q_2^2) = 1 + \lambda \left( \frac{q_1^2}{q_1^2 - m_V^2} + \frac{q_2^2}{q_2^2 - m_V^2} \right) + \eta \frac{q_1^2 q_2^2}{(q_1^2 - m_V^2)(q_2^2 - m_V^2)}
$$

- $\bullet\,$  Used before as low energy parameterization for kaon decays  $\,$  [D'Ambrosio et al'97]
- Consistent with the anomaly and Bose symmetry.
- $\bullet\,$  Parameters easily interpreted in terms of  $\hat{\alpha},\,\hat{\beta}$  and  $\hat{\gamma}.$
- Interesting feature: it can be cast in the genera<sup>l</sup> form

$$
F_{\pi^0 \gamma^* \gamma^*}(q_1^2, q_2^2) = -\frac{N_c}{12\pi^2 f_\pi} \left[ f(q_1^2) - \sum_i \frac{1}{q_2^2 - m_i^2} g_i(q_1^2) \right]
$$

with

$$
f(q^2) = 1 + \lambda + (\lambda + \eta) \frac{q^2}{q^2 - m_V^2}
$$

$$
g(q^2) = -m_V^2 \left[ \lambda + \eta \frac{q^2}{q^2 - m_V^2} \right]
$$

which allows to easily perform the two-loop integrals as

[Knecht et al'01]

$$
a_{\mu}^{\pi^0} = \left(\frac{\alpha_{em}}{\pi}\right)^3 \int_0^{\infty} dQ_1 \int_0^{\infty} dQ_2 \Big[ w_1 G_1 + w_2(m_V) G_2 + w_3(m_V) G_3 + w_3(m_{\pi}) G_4 \Big]
$$

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### Numerical analysis

 $\bullet\,$  Our parameterization has 3 free parameters,  $\lambda,\,\eta$  and  $m_V$ , to be fixed by

$$
1 + 2\lambda + \eta = 0, \qquad \text{(OPE)}
$$
  
\n
$$
\lambda + \eta = -\frac{4\pi^2 f_\pi^2}{3m_V^2}, \qquad \text{(OPE)}
$$
  
\n
$$
\frac{\lambda}{m_V^2} = -1.76 \pm 0.22 \ (\text{= } \hat{\alpha}) \quad \text{(exp.)}
$$

• This <sup>y</sup>ields

$$
\lambda = -0.73 \pm 0.05 ,
$$
  
\n
$$
\eta = 0.46^{+0.10}_{-0.13} ,
$$
  
\n
$$
m_V = (0.64^{+0.07}_{-0.06}) \text{ GeV} ,
$$
  
\n
$$
\chi_0 = \frac{N_c}{4\pi^2 f_\pi^2} (1 + \lambda) = (2.42 \pm 0.17) \text{ GeV}^{-2}
$$

Several comments:

- $\eta = m_V^4 \hat{\beta}$ , fixed by short distance constraints, agrees very well with holographic predictions. Recall that  $\lambda = m_V^2 \hat{\alpha}$  also does.
- $m_V = 0.64^{+0.07}_{-0.06}$  GeV determined dynamically. In LMD models one assumes  $m_V \equiv m_\rho.$  In our case  $m_V$  comes from using the experimental slope of  $F_{\pi^0\gamma\gamma}.$ <br>Effective resummation of  $\infty$  resonances. Effective resummation of  $\infty$  resonances.
- $\bullet$  DIP naturally gives  $\chi_0\neq 0$ . Therefore, it provides one way to estimate the 'off-shellness' through <sup>a</sup> short distance constraint.
- Result for  $HLBL$ :



- Difference with LMD only due to the vector mass scale. With long distance constraints, no need to take  $m_V = m_\rho.$
- Intriguing: parameterization dependence (DIP vs VMD) seems to be extremely small.
- $\bullet\,$  Contribution dominated by  $w_1G_1, w_2G_2$ , peaked at (very) low energies.

### Refinements and sanity checks

- $\bullet\,$  We would like to use holographic input  $(\hat\beta,\,\hat\gamma)$  and check for stability.
- Simplest generalised version of the interpolator:

$$
K(q_1, q_2) = 1 + \sum_{i}^{2} \lambda_i \left( \frac{q_1^2}{q_1^2 - m_i^2} + \frac{q_2^2}{q_2^2 - m_i^2} \right) + \sum_{i}^{2} \eta_i \frac{q_1^2 q_2^2}{(q_1^2 - m_i^2)(q_2^2 - m_i^2)}
$$

• It still satisfies

$$
F_{\pi^0 \gamma^* \gamma^*} (q_1^2, q_2^2) = -\frac{N_c}{12\pi^2 f_\pi} \left[ f(q_1^2) - \sum_i \frac{1}{q_2^2 - m_i^2} g_i(q_1^2) \right]
$$

• Two candidates:

$$
K_{(1)}(q_1, q_2) = 1 + \lambda \left( \frac{q_1^2}{q_1^2 - m_1^2} + \frac{q_2^2}{q_2^2 - m_1^2} \right) + \sum_i^2 \eta_i \frac{q_1^2 q_2^2}{(q_1^2 - m_i^2)(q_2^2 - m_i^2)}
$$
  
\n
$$
K_{(2)}(q_1, q_2) = 1 + \sum_i^2 \lambda_i \left( \frac{q_1^2}{q_1^2 - m_i^2} + \frac{q_2^2}{q_2^2 - m_i^2} \right) + \eta \frac{q_1^2 q_2^2}{(q_1^2 - m_1^2)(q_2^2 - m_1^2)}
$$

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#### Constraints -

• The constraints we will impose now are the following:

(a) Long distance constraints:

$$
\hat{\alpha} = \sum_{i}^{2} \frac{\lambda_i}{m_i^2} = -1.76 \pm 0.22 , \text{ (exp.)}
$$
  

$$
\hat{\beta} = \sum_{i}^{2} \frac{\eta_i}{m_i^4} = 3.33 \pm 0.32 , \text{ (holography)}
$$
  

$$
\hat{\gamma} = -\sum_{i}^{2} \frac{\lambda_i}{m_i^4} = 2.84 \pm 0.21 , \text{ (holography)}
$$

(b) Short distance constraints:

$$
1 + 2\sum_{i}^{2} \lambda_{i} + \sum_{i}^{2} \eta_{i} = 0,
$$
  

$$
\sum_{i}^{2} m_{i}^{2}(\lambda_{i} + \eta_{i}) = -\frac{4\pi^{2}f_{\pi}^{2}}{3},
$$
  

$$
1 + \sum_{i}^{2} \lambda_{i} = -\frac{4\pi^{2}f_{\pi}^{2}}{3}\chi_{0}.
$$

• Slightly over-constrained.

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**Results** 



- $\bullet \;\; \hat{\alpha}$  drives the error over  $\hat{\beta}$  and  $\hat{\gamma}.$
- If I don't look at  $\chi_0$ :

$$
a_{\mu} = 6.54(25) \cdot 10^{-10}
$$

compatible with previous estimates. The uncertainty is driven by the experimental situation on  $\hat{\alpha}$ , whereas (holographic) model dependences turn out to be negligible.

 $\bullet\,$  Notice stability: with the simplest interpolator,  $a_{\mu}=6.7(3)\cdot 10^{-10}.$ 

### Recent determinations

 $\bullet$  Most recent estimates for the  $\pi^0$  and pseudoscalar exchange:



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## The status of  $\chi_0$

• Showing up in the (off-shell) short distance constraint:

$$
\lim_{Q^2 \to \infty} F_{\pi^0 \gamma \gamma}(Q^2, Q^2, 0) = \frac{f_\pi}{3} \chi_0 + \cdots
$$

• Determinations range from  $0 \sim \chi_0 \sim 9$  GeV $^{-2}$ .



 $\bullet\,$  Our analysis favors lower values of  $\chi_0.$  If so, mild effect and pion-pole approximation very successful. However,  $\chi_0$  9 GeV $^{-2}$  might amount to  $10-15\%$ systematic deviations in Hlbl.

### First results and Outlook

- Holographic models of QCD aim at providing <sup>a</sup> full description of hadronic processes. Extremely useful for the HLBL (all-channel, all-state approach) where so far each contribution is evaluated in <sup>a</sup> fragmentary way (interpolators for differentchannels). In contrast, with holographic models one starts from one Lagrangian(predictability, unitarity, Green functions highly constrained).
- This is not (yet) an holographic determination of HLBL, rather <sup>a</sup> 'classic' approachwith new ingredients: DIP interpolator, low energy parameters determined fromholography. Full calculation underway.
- $\bullet\,$  We provided: (a) predicted values for  $\hat{\beta}$  and  $\hat{\gamma}$  from holographic models (consistent with short distance QCD constraints and the experimental slope of the pion form factor), soon to be checked at KLOE-2; (b) assessed the quantitative validity of LMD in the pion-exchange diagram; (c) estimated the impact of 'pion off-shellness' through  $\chi_0.$
- No free lunch: holographic models are not free from assumptions, but they openinteresting avenues to study the hadronic contributions to  $(g\,$  $-2)$  $\mu$  .