

### Why holography for HLBL? -

- Holographic principle: Conjectured duality between strongly coupled gauge theories in d dimensions and gravitational weakly coupled theories in (d + 1) dimensions. AdS/CFT one of the most powerful examples.
- Lagrangian formulation: consistent treatment of the different channels (no double counting).
- (leading) short distances: automatically implemented due to the conformal invariance of AdS.
- Large- $N_c$ : full realization of the large- $N_c$  limit ( $\infty$  resonances with short distances).
- number of parameters: minimal.

### Motivating questions .

- Which are the parameters that mostly affect the uncertainty on the HLBL?
- Given

$$K(Q_1^2,Q_2^2) \simeq 1 + \hat{\alpha} \ (Q_1^2 + Q_2^2) + \hat{\beta} \ Q_1^2 Q_2^2 + \hat{\gamma} \ (Q_1^4 + Q_2^4)$$

what is the impact of the low energy parameters  $\alpha$ ,  $\beta$  and  $\gamma$ ?  $\alpha$  determined by CLEO, but  $\gamma$  and  $\beta$  out of current experimental reach. Can one obtain reasonable predictions?

- Is vector meson dominance a good enough approximation?
- How accurate is the pion-pole approximation?

# The setting -

• AdS<sub>5</sub> space with metric

$$ds^{2} = g_{MN} dx^{M} dx^{N} = \frac{1}{y^{2}} (-dy^{2} + \eta_{\mu\nu} dx^{\mu} dx^{\nu})$$

with  $\eta_{\mu\nu}$  mostly negative.



• Intuitively, UV boundary captures pQCD (leading term trivial: conformal limit reproduced by AdS metric). IR deals with non-perturbative physics (OPE and non-OPE, confinement scale,  $S\chi$ SB, spectrum patterns)

# Holographic models

• There are different realizations of the holographic principle for QCD. Generically,

$$S_5 = S_{\rm YM} + S_X + S_{\rm CS}$$

where

$$S_{\rm YM} \left[ \mathcal{B}_L, \mathcal{B}_R \right] = \operatorname{tr} \int d^4x \int_0^{z_0} dz \ e^{-\Phi(z)} \frac{-1}{8g_5^2} w(z) \left[ \mathcal{F}_{(L)}^{MN} \mathcal{F}_{(L)MN} + \mathcal{F}_{(R)}^{MN} \mathcal{F}_{(R)MN} \right]$$
$$S_X \left[ X \right] = \operatorname{tr} \int d^4x \int_0^{z_0} dz \ e^{-\Phi(z)} w(z)^3 \left[ D^M X D_M X^{\dagger} + V(X^{\dagger} X) \right]$$
$$S_{\rm CS} \left[ \mathcal{B}_L, \mathcal{B}_R \right] = \frac{N_c}{24\pi^2} \int \operatorname{tr} \left[ \mathcal{B}_L \mathcal{F}_L^2 - \frac{i}{2} \mathcal{B}_L^3 \mathcal{F}_L - \frac{1}{10} \mathcal{B}_L^5 - (L \to R) \right]$$

- Notice that chiral symmetry is incorporated.
- The models differ mostly in  $S_X$  (spontaneous symmetry breaking) and  $(\Phi(z), z_0)$  (spectrum).
- Major drawback: plethora of models...
- Advantage: stability tests of phenomenological results.

# The $F_{\pi^0\gamma\gamma}$ form factor

• From holography  $F_{\pi^0\gamma\gamma}$  can be extracted from the Chern-Simons 5-dimensional term: [Grigoryan et al'08,09]

$$S_{\rm CS} \left[ \mathcal{B} \right] = \frac{N_c}{24\pi^2} \int \operatorname{tr} \left[ \mathcal{B}_L \mathcal{F}_L^2 - \mathcal{B}_R \mathcal{F}_R^2 \right] + \cdots$$
$$= \frac{N_c}{4\pi^2} \epsilon^{\mu\nu\rho\sigma} \int_0^{z_0} dz \left( \partial_z \beta \right) \int d^4 x \ \pi^a \left( \partial_\rho V_\mu^a \right) \left( \partial_\sigma \hat{V}_\nu \right)$$

• Applying the holographic recipe for correlators:

$$K(Q_1^2, Q_2^2) = -\int_0^{z_0} \mathcal{J}(Q_1, z) \mathcal{J}(Q_2, z) \,\partial_z \beta(z) \,dz$$

• For instance,

$$\mathcal{J}(Q, z) = Qz \left[ K_1(Qz) + I_1(Qz) \frac{K_0(Qz_0)}{I_0(Qz_0)} \right]$$
  
$$\beta(z) = 1 - \frac{z^2}{z_0^2}$$

# $\checkmark$ The $F_{\pi^0\gamma\gamma}$ form factor (high energies)

• Master formula for the different holographic models. It only depends on  $z_0$ , which for consistency (matching with the  $\rho(770)$  mass) has to be

$$z_0 = \frac{2}{g_5^2 f_\pi} = \frac{N_c}{6\pi^2 f_\pi}$$

• In any model with (asymptotic) AdS metric, it can be shown that (at least) the leading short distance constraints from QCD are satisfied. Not surprising: the asymptotic behavior of  $\mathcal{J}(Q, z)$  fixed by AdS, while the pion wave function is only relevant at the origin:



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# $\checkmark$ The $F_{\pi^0\gamma\gamma}$ form factor (low energies)

• Low energies will depend on the specific model. At small virtualities one can expand  $K(Q_1^2,Q_2^2)$  in the form

$$K(Q_1^2, Q_2^2) \simeq 1 + \hat{\alpha} (Q_1^2 + Q_2^2) + \hat{\beta} Q_1^2 Q_2^2 + \hat{\gamma} (Q_1^4 + Q_2^4)$$

- Experimentally,  $\hat{\alpha} = -1.76 \pm 0.22$ ,  $\hat{\gamma}$  waiting for better statistics while  $\hat{\beta}$  challenging (2 virtual photons needed).
- Predictions from holographic models:

Model	$\hat{lpha}$ (GeV $^{-2}$ )	$\hat{eta}$ (GeV $^{-4}$ )	$\hat{\gamma}$ (GeV $^{-4}$ )
HW1	-1.60	3.01	2.63
HW2 (AdS)	-1.81	3.65	3.06
HW2 (Flat)	-1.37	2.25	2.25
SS	-2.04	4.56	3.55
SW	-1.66	3.56	2.76

• Compliance with experiment used as a holographic model filter. From the "acceptable" models we can then estimate the quartic terms to be

$$\hat{\beta} = 3.33 \pm 0.32$$
  
 $\hat{\gamma} = 2.84 \pm 0.21$ 

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#### - Vector meson dominance revisited -

• Spectral decomposition:

$$\mathcal{J}(z,Q) = \sum_{n=1}^{\infty} \frac{f_n}{Q^2 + m_n^2} \psi_n(z); \qquad K(Q_1^2,Q_2^2) \equiv \sum_{k,l=1}^{\infty} \frac{B_{kl}}{(Q_1^2 + m_k^2)(Q_2^2 + m_l^2)}$$

• Graphically:



• Each resonance contribution can be evaluated

Model	$\hat{lpha}_n/\hat{lpha}$		$\hat{eta}_n/\hat{eta}$		$\hat{\gamma}_n/\hat{\gamma}$				
HW1	1.20	-0.18	-0.04	1.10	-0.06	0.01	1.20	-0.22	0.06
HW2 (AdS)	1.30	-0.37	0.06	1.10	-0.11	0.01	1.30	-0.37	0.08
HW2 (flat)	0.99	0.01	0.00	1.00	0.00	0.00	1.00	0.00	0.00
SS	1.70	-1.10	0.49	1.30	-0.34	0.07	1.60	-1.10	0.54
SW	0.75	0.14	0.05	0.87	0.09	0.02	0.88	0.09	0.02

• Lowest meson dominance qualitatively holds. Enough for the accuracy of HLBL?

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## - Estimation of the $\pi^0$ -HLBL

• Strategy: encode the short and long distance information in an interpolator (DIP):

$$K(q_1^2, q_2^2) = 1 + \lambda \left( \frac{q_1^2}{q_1^2 - m_V^2} + \frac{q_2^2}{q_2^2 - m_V^2} \right) + \eta \frac{q_1^2 q_2^2}{(q_1^2 - m_V^2)(q_2^2 - m_V^2)}$$

- Used before as low energy parameterization for kaon decays [D'Ambrosio et al'97]
- Consistent with the anomaly and Bose symmetry.
- Parameters easily interpreted in terms of  $\hat{\alpha}$ ,  $\hat{\beta}$  and  $\hat{\gamma}$ .
- Interesting feature: it can be cast in the general form

$$F_{\pi^0\gamma^*\gamma^*}(q_1^2, q_2^2) = -\frac{N_c}{12\pi^2 f_\pi} \Big[ f(q_1^2) - \sum_i \frac{1}{q_2^2 - m_i^2} g_i(q_1^2) \Big]$$

with

$$f(q^{2}) = 1 + \lambda + (\lambda + \eta) \frac{q^{2}}{q^{2} - m_{V}^{2}}$$
$$g(q^{2}) = -m_{V}^{2} \left[\lambda + \eta \frac{q^{2}}{q^{2} - m_{V}^{2}}\right]$$

which allows to easily perform the two-loop integrals as [Knecht et al'01]

$$a_{\mu}^{\pi^{0}} = \left(\frac{\alpha_{em}}{\pi}\right)^{3} \int_{0}^{\infty} dQ_{1} \int_{0}^{\infty} dQ_{2} \left[w_{1}G_{1} + w_{2}(m_{V})G_{2} + w_{3}(m_{V})G_{3} + w_{3}(m_{\pi})G_{4}\right]$$

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#### - Numerical analysis

• Our parameterization has 3 free parameters,  $\lambda$ ,  $\eta$  and  $m_V$ , to be fixed by

$$1 + 2\lambda + \eta = 0, \qquad \text{(OPE)}$$
  

$$\lambda + \eta = -\frac{4\pi^2 f_{\pi}^2}{3m_V^2}, \qquad \text{(OPE)}$$
  

$$\frac{\lambda}{m_V^2} = -1.76 \pm 0.22 \ (= \hat{\alpha}) \qquad (\text{exp.})$$

• This yields

$$\lambda = -0.73 \pm 0.05 ,$$
  

$$\eta = 0.46^{+0.10}_{-0.13} ,$$
  

$$m_V = (0.64^{+0.07}_{-0.06}) \text{ GeV} ,$$
  

$$\chi_0 = \frac{N_c}{4\pi^2 f_\pi^2} (1+\lambda) = (2.42 \pm 0.17) \text{ GeV}^{-2}$$

Several comments:

- $\eta = m_V^4 \hat{\beta}$ , fixed by short distance constraints, agrees very well with holographic predictions. Recall that  $\lambda = m_V^2 \hat{\alpha}$  also does.
- $m_V = 0.64^{+0.07}_{-0.06}$  GeV determined dynamically. In LMD models one assumes  $m_V \equiv m_{\rho}$ . In our case  $m_V$  comes from using the experimental slope of  $F_{\pi^0\gamma\gamma}$ . Effective resummation of  $\infty$  resonances.

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- DIP naturally gives  $\chi_0 \neq 0$ . Therefore, it provides one way to estimate the 'off-shellness' through a short distance constraint.
- Result for *HLBL*:

Model	$w_1G_1$	$w_2G_2$	$w_3G_3$	$w_4G_4$	$a_{\mu}$
LMD	+0.015	+0.042	+0.0016	-0.0002	$7.3 \cdot 10^{-10}$
$DIP_{\hat{lpha}}$	+0.018(3)	+0.034(4)	+0.0016	-0.0002	$6.7(3) \cdot 10^{-10}$
$DIP_{m_{ ho}}$	+0.015	+0.043	+0.0016	-0.0002	$7.35 \cdot 10^{-10}$

- Difference with LMD only due to the vector mass scale. With long distance constraints, no need to take  $m_V = m_{\rho}$ .
- Intriguing: parameterization dependence (DIP vs VMD) seems to be extremely small.
- Contribution dominated by  $w_1G_1, w_2G_2$ , peaked at (very) low energies.

#### Refinements and sanity checks -

- We would like to use holographic input  $(\hat{eta}, \hat{\gamma})$  and check for stability.
- Simplest generalised version of the interpolator:

$$K(q_1, q_2) = 1 + \sum_{i}^{2} \lambda_i \left( \frac{q_1^2}{q_1^2 - m_i^2} + \frac{q_2^2}{q_2^2 - m_i^2} \right) + \sum_{i}^{2} \eta_i \frac{q_1^2 q_2^2}{(q_1^2 - m_i^2)(q_2^2 - m_i^2)}$$

• It still satisfies

$$F_{\pi^0\gamma^*\gamma^*}(q_1^2, q_2^2) = -\frac{N_c}{12\pi^2 f_\pi} \left[ f(q_1^2) - \sum_i \frac{1}{q_2^2 - m_i^2} g_i(q_1^2) \right]$$

• Two candidates:

$$K_{(1)}(q_1, q_2) = 1 + \lambda \left( \frac{q_1^2}{q_1^2 - m_1^2} + \frac{q_2^2}{q_2^2 - m_1^2} \right) + \sum_i^2 \eta_i \frac{q_1^2 q_2^2}{(q_1^2 - m_i^2)(q_2^2 - m_i^2)}$$
  

$$K_{(2)}(q_1, q_2) = 1 + \sum_i^2 \lambda_i \left( \frac{q_1^2}{q_1^2 - m_i^2} + \frac{q_2^2}{q_2^2 - m_i^2} \right) + \eta \frac{q_1^2 q_2^2}{(q_1^2 - m_1^2)(q_2^2 - m_1^2)}$$

#### - Constraints -

• The constraints we will impose now are the following:

(a) Long distance constraints:

$$\begin{aligned} \hat{\alpha} &= \sum_{i}^{2} \frac{\lambda_{i}}{m_{i}^{2}} &= -1.76 \pm 0.22 \text{ , (exp.)} \\ \hat{\beta} &= \sum_{i}^{2} \frac{\eta_{i}}{m_{i}^{4}} &= 3.33 \pm 0.32 \text{ , (holography)} \\ \hat{\gamma} &= -\sum_{i}^{2} \frac{\lambda_{i}}{m_{i}^{4}} &= 2.84 \pm 0.21 \text{ , (holography)} \end{aligned}$$

(b) Short distance constraints:

$$1 + 2\sum_{i}^{2} \lambda_{i} + \sum_{i}^{2} \eta_{i} = 0,$$
  
$$\sum_{i}^{2} m_{i}^{2} (\lambda_{i} + \eta_{i}) = -\frac{4\pi^{2} f_{\pi}^{2}}{3},$$
  
$$1 + \sum_{i}^{2} \lambda_{i} = -\frac{4\pi^{2} f_{\pi}^{2}}{3} \chi_{0}$$

• Slightly over-constrained.

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- Results

Model	$w_1G_1$	$w_2G_2$	$a_{\mu}$
$DIP_{\hat{lpha}}$	+0.018(3)	+0.034(4)	$6.7(3) \cdot 10^{-10}$
$DIP^{(1)}_{\hat{eta};m_2=m_ ho}$	+0.014	+0.037	$6.52(15)(10) \cdot 10^{-10}$
$\underline{DIP^{(2)}_{\hat{\beta};m_2=m_\rho}}$	+0.014	+0.037	$6.55(21)(6) \cdot 10^{-10}$
$DIP^{(1)}_{\hat{\beta}:0<\chi_0<8.9}$	[+0.003; +0.047]	[+0.043; +0.022]	$[5.9; 8.9] \cdot 10^{-10}$
$DIP_{\hat{\beta}; 0 < \chi_0 < 4.4}^{(2)}$	[+0.002; +0.027]	[+0.044; +0.025]	$[6.0; 6.7] \cdot 10^{-10}$

- $\hat{\alpha}$  drives the error over  $\hat{\beta}$  and  $\hat{\gamma}$ .
- If I don't look at  $\chi_0$ :

$$a_{\mu} = 6.54(25) \cdot 10^{-10}$$

compatible with previous estimates. The uncertainty is driven by the experimental situation on  $\hat{\alpha}$ , whereas (holographic) model dependences turn out to be negligible.

• Notice stability: with the simplest interpolator,  $a_{\mu} = 6.7(3) \cdot 10^{-10}$ .

### - Recent determinations -

• Most recent estimates for the  $\pi^0$  and pseudoscalar exchange:

Model for $\mathcal{F}_{P^{(*)}\gamma^*\gamma^*}$	$a_{\mu}^{\text{LbyL};\pi^{0}} 10^{-11}$	$a_{\mu}^{\mathrm{LbyL;PS}} 10^{-11}$
modified ENJL [BPP]	59(9)	85(13)
VMD / HLS [HKS, HK]	57(4)	83(6)
nonlocal $\chi$ QM [DB]	65(2)	-
AdS/QCD [HoK]	69	107
LMD [KN]	73	-
LMD+V [KN1]	58(10)	83(12)
LMD+V [KN2]	63(10)	88(12)
LMD+V MV	77(7)	114(10)
LMD+V [N]	72(12)	99(16)
DIP	65(2)	-

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## - The status of $\chi_0$

• Showing up in the (off-shell) short distance constraint:

$$\lim_{Q^2 \to \infty} F_{\pi^0 \gamma \gamma}(Q^2, Q^2, 0) = \frac{f_\pi}{3} \chi_0 + \cdots$$

• Determinations range from  $0 \sim \chi_0 \sim 9 \text{ GeV}^{-2}$ .

Model	$\hat{lpha}$	$\hat{eta}$	$\hat{\gamma}$	$\chi_0$
$DIP_{\hat{lpha}}$	$-1.76^{*}$	2.67	4.25	2.42
$DIP_{m_{ ho}}$	-1.35	1.73	2.25	1.66
$DIP_{\hat{eta}}^{(1)}$	$-1.76^{*}$	$3.33^{*}$	3.78	1.61
$DIP_{\hat{\beta}}^{(2)}$	$-1.76^{*}$	3.33*	3.88	1.69
$DIP^{(1)}_{\hat{\beta},\chi_0}$	$-1.76^{*}$	$3.33^{*}$	$[3.10; -5\cdot 10^5]$	$[0; 8.9]^*$
$DIP_{\hat{eta},\chi_0}^{(2)^{0}}$	$-1.76^{*}$	$3.33^{*}$	$\left[ 3.19; -3.18 \right]$	$[0; 4.4]^*$

• Our analysis favors lower values of  $\chi_0$ . If so, mild effect and pion-pole approximation very successful. However,  $\chi_0$  9 GeV<sup>-2</sup> might amount to 10 - 15% systematic deviations in Hlbl.

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### First results and Outlook

- Holographic models of QCD aim at providing a full description of hadronic processes. Extremely useful for the HLBL (all-channel, all-state approach) where so far each contribution is evaluated in a fragmentary way (interpolators for different channels). In contrast, with holographic models one starts from one Lagrangian (predictability, unitarity, Green functions highly constrained).
- This is not (yet) an holographic determination of HLBL, rather a 'classic' approach with new ingredients: DIP interpolator, low energy parameters determined from holography. Full calculation underway.
- We provided: (a) predicted values for β̂ and γ̂ from holographic models (consistent with short distance QCD constraints and the experimental slope of the pion form factor), soon to be checked at KLOE-2; (b) assessed the quantitative validity of LMD in the pion-exchange diagram; (c) estimated the impact of 'pion off-shellness' through χ<sub>0</sub>.
- No free lunch: holographic models are not free from assumptions, but they open interesting avenues to study the hadronic contributions to  $(g-2)_{\mu}$ .