Precision Determination of

$\alpha_{s}(m_{z})$ from Thrust Data

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Frontiers in QCD, INT, 6th October 2011

The Strong Coupling Constant α_s

α_s is a key parameter for the analysis of all collider experiments QCD: $\alpha_s(M_z)=0.1184(07)$

S. Bethke, Eur. Phys. J. C64 (2009) 689-703





Thrust:
$$T = \max_{i} \left[\frac{\sum_{j} |\vec{p}_{j}(t)|}{\sum_{j} |\vec{p}_{j}|} \right]$$
 use $\tau = 1 - T$







Thrust:
$$T \equiv \max_{i} \left[\frac{\sum_{j} \left| \vec{p}_{j} \cdot \vec{t} \right|}{\sum_{j} \left| \vec{p}_{j} \right|} \right]$$

use τ = 1 - T







 $\frac{1}{\sigma_0} \frac{d\sigma}{d\tau} = \delta(\tau) + \alpha_s(c_1 \delta(\tau) + R_1(\tau > 0)) + \dots \implies \text{Observable is very sensitive to } \alpha_s !$



- Results
- Theory
- Numerical Analysis
- Moment Fits
- Summary



Results

 $\alpha_{\rm s}({\rm m_Z}) = 0.1135 \pm (0.0002)_{\rm expt} \pm (0.0005)_{\rm hadr} \pm (0.0009)_{\rm pert}$

with
$$\frac{\chi^2}{dof} = 0.91$$





Fit to Distribution:



 $\Omega_1(\mu_{\Delta}, R_{\Delta}) = 0.388 \pm 0.061$

 $\alpha_{\rm s}({\rm m_Z}) = 0.1135 \pm (0.0002)_{\rm expt} \pm (0.0005)_{\rm hadr} \pm (0.0009)_{\rm pert}$

with
$$\frac{\chi^2}{dof} = 0.91$$





Theory



- hard scale (cm energy)
- jet scale (invariant mass of all energetic particles in one hemisphere)
- soft scale (uniform soft radiation)



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Soft Collinear Effective Theory (SCET)

Bauer, Fleming, Luke, Pirjol, Rothstein, Stewart

 $\Lambda_{
m QCD}$

 Describes light-like particles (collinear) interacting with a low energetic background (soft)

- Expansion in $\lambda \approx \sqrt{\frac{\Lambda_{QCD}}{Q}}$
- Power counting:

soft: collinear:

 $p_{\mu} = (p_{+}, p_{-}, p_{\perp}) \propto Q(\lambda^{2}, \lambda^{2}, \lambda^{2})$ $p_{\mu} = (p_+, p_-, p_+) \propto Q(\lambda^2, 1, \lambda)$

$$\frac{1}{\sigma_0}\frac{d\sigma}{d\tau} = \delta(\tau) + \frac{\alpha_s R_1(\tau)}{\alpha_s R_1(\tau)} + \frac{\alpha_s^2 R_2(\tau)}{\alpha_s R_3(\tau)} + \dots$$

$$\frac{1}{\sigma_0}\frac{d\sigma}{d\tau} = \delta(\tau) + \alpha_s R_1(\tau) + \alpha_s^2 R_2(\tau) + \alpha_s^3 R_3(\tau) + \dots$$

Problems if: $\alpha_s \gtrsim 1$ coefficients not of O(1)







$$\frac{d\sigma}{d\tau} = \int dk \left(\frac{d\hat{\sigma}_s}{d\tau} + \frac{d\hat{\sigma}_{ns}}{d\tau} + \frac{\Delta d\hat{\sigma}_b}{d\tau} \right) \left(\tau - \frac{k}{Q} \right) S_{\tau}^{\text{mod}} \left(k - 2\overline{\Delta} \right) + O\left(\sigma_0 \frac{\alpha_s \Lambda_{QCD}}{Q} \right)$$

Methode to simultaneously describe all three regions and multiple Q: profile functions

Singular partonic:

NNLO matching, including full 3-loop hard function

N³LL resummation of large logs

Nonsingular partonic:

fixed order thrust distribution (subleading orders in SCET)

Nonperturbative soft function:

nonperturabtive effects treated within field theory Operator Product Expansion in tail $\rightarrow \Omega_1$

Interface between power and perturbative corrections:

renormalon subtraction

b-mass effects (~2% effect)

QED effects (~2% effect)

axial anomaly at $O(\alpha_s^2)$ (~1% effect)

Factorization Formula for all Thrust

peak: sum large logs, nonperturbative soft fct.



- tail:sum large logs,
series of nonperturbative
power corrections $Q >> Q\sqrt{\tau} >> Q\tau >> \Lambda_{QCD}$
- **far tail:** fixed order perturbation theory, power corrections



Profile Functions:



Scales must equal in the far tail: turns of resummation

Far Tail



$$\frac{d\sigma}{d\tau} = \int dk \left(\frac{d\hat{\sigma}_s}{d\tau} + \frac{d\hat{\sigma}_{ns}}{d\tau} + \frac{\Delta d\hat{\sigma}_b}{d\tau} \right) \left(\tau - \frac{k}{Q} \right) S_{\tau}^{\text{mod}} \left(k - 2\overline{\Delta} \right) + O\left(\sigma_0 \frac{\alpha_s \Lambda_{QCD}}{Q} \right)$$

Methode to simultaneously describe all three regions and multiple Q:profile functionsMoch, Value

Singular partonic:

Baikov et al.

NNLO matching, including full 3-loop hard function N³LL resummation of large logs

Nonsingular partonic:

Becher, Schwartz

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Moch, Vermaseren, Vogt Becher, Neubert Schwartz; Fleming et al. Becher, Schwartz; Hoang, Kluth Dasgupta, Salam

Summing large Logarithms



$$\frac{d\sigma}{d\tau} = \int dk \left(\frac{d\hat{\sigma}_s}{d\tau} + \left(\frac{d\hat{\sigma}_{ns}}{d\tau} + \frac{\Delta d\hat{\sigma}_b}{d\tau} \right) \left(\tau - \frac{k}{Q} \right) S_{\tau}^{\text{mod}} \left(k - 2\overline{\Delta} \right) + O\left(\sigma_0 \frac{\alpha_s \Lambda_{QCD}}{Q} \right) \right)$$

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N³LL resummation of large logs

Nonsingular partonic:

Ellis et al., Catani et al., Gehrmann et al., Weinzierl

fixed order thrust distribution (subleading orders in SCET)

Nonperturbative soft function: nonperturbative effects treated within field theory Operator Product Expansion in tail $\Rightarrow \Omega_1$ Interface between power and perturbative corrections: renormalon subtraction b-mass effects (~2% effect) QED effects (~2% effect) $d\hat{\sigma}_{ns} = \frac{d\sigma}{d\tau}\Big|_{fixed-order} - \frac{d\hat{\sigma}_s}{d\tau}\Big|_{no-resum.}$ axial anomaly at O(α_s^2) (~1% effect)

$$\frac{d\sigma}{d\tau} = \int dk \left(\frac{d\hat{\sigma}_s}{d\tau} + \frac{d\hat{\sigma}_{ns}}{d\tau} + \frac{\Delta d\hat{\sigma}_b}{d\tau} \right) \left(\tau - \frac{k}{Q} \right) S_{\tau}^{\text{mod}} \left(k - 2\overline{\Delta} \right) + O\left(\sigma_0 \frac{\alpha_s \Lambda_{QCD}}{Q} \right)$$

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Hoang, Stewart Ligeti, Stewart, Tackmann

Lee, Sterman Korchemsky, Sterman

Nonperturbative Corrections

Soft function from SCET factorization:

$$S_{\tau}(k,\mu) = \frac{1}{N_{c}} \left\langle 0 \left| Tr \overline{Y}_{\overline{n}} Y_{n} \delta(k-i\partial_{\tau}) Y_{n}^{+} \overline{Y}_{\overline{n}}^{+} \right| 0 \right\rangle$$
$$i\partial_{\tau} \equiv \theta(i\overline{n} \cdot \partial - in \cdot \partial) in \cdot \partial + \theta(in \cdot \partial - i\overline{n} \cdot \partial) i\overline{n} \cdot \partial$$

Factorization in perturbative and non-perturbative part:

$$S_{\tau}(k,\mu) = \int dk' S_{\tau}^{part}(k-k',\mu) S_{\tau}^{mod}(k')$$
Hoang, Stewart
Ligeti, Stewart, Tackmann
complete basis
of functions

OPE:
$$S_{\tau}(k,\mu) = S_{\tau}^{part}(k,\mu) - 2\overline{\Omega}_{1} \frac{dS_{\tau}^{part}}{dk}(k,\mu) + \dots$$

Leading effect: shift in the thrust distribution $\tau \rightarrow \tau - 2\Lambda/Q$

$$\frac{1}{\sigma}\frac{d\sigma}{d\tau} = h\left(\tau - 2\frac{\Lambda}{Q}\right)$$

Manohar, Wise; Webber; Dokshitzer, Weber; Akhoury, Zakharov; Nason, Seymour; Korchemsky, Sterman; Movilla Fernandez, Bethke, Biebel, Kluth

Leading effect: shift in the thrust distribution $\tau \rightarrow \tau - 2 \Lambda / Q$

 $\frac{1}{\sigma} \frac{d\sigma}{d\tau} = h\left(\tau - 2\frac{\Lambda}{Q}\right) \approx h(\tau) - 2\frac{\Lambda}{Q}h'(\tau) \qquad \Lambda/Q \ll 1$

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-25

0.10

0.20

0.25

0.30

au

0.15

Leading effect: shift in the thrust distribution $\tau \rightarrow \tau - 2 \Lambda / Q$

 $\frac{1}{\sigma} \frac{d\sigma}{d\tau} = h \left(\tau - 2 \frac{\Lambda}{Q} \right) \approx h(\tau) - 2 \frac{\Lambda}{Q} h'(\tau) \qquad \Lambda / Q \ll 1$ $\frac{\delta \alpha_{S}}{\alpha_{S}} \approx \frac{h(\tau - 2\frac{\Lambda}{Q}) - h(\tau)}{h(\tau)} \approx -2\frac{\Lambda}{Q}\frac{h'(\tau)}{h(\tau)}$ *h* proportional to $\alpha_s \Rightarrow$ $\frac{h'(\tau)}{h(\tau)} \stackrel{0}{_{-5}} = \underbrace{\begin{array}{c} \bullet \quad \text{DELPHI} \\ \bullet \quad \text{ALEPH} \\ \bullet \quad \text{OPAL} \end{array}}_{OPAL}$ $Q = m_Z$ $\Lambda \approx 0.3 \text{ GeV}$ and -10 $h'/h \approx -14\pm 4$ in tail region(see figure) -15 $\Rightarrow \delta \alpha \sqrt{\alpha_s} \approx -(9\pm 3)\%$ -20 \Rightarrow **NOT** negligible for 2% analysis -250.20 0.100.150.250.30 au

Non-perturbative effects

Use tail fit results to predict peak



Non-perturbative effects

Use tail fit results to predict peak

 \Rightarrow much better prediction for peak



Non-perturbative effects



αS from full analysis is approximately
 9% smaller than αS without model,
 as predicted.

$$\frac{d\sigma}{d\tau} = \int dk \left(\frac{d\hat{\sigma}_s}{d\tau} + \frac{d\hat{\sigma}_{ns}}{d\tau} + \frac{\Delta d\hat{\sigma}_b}{d\tau} \right) \left(\tau - \frac{k}{Q} \right) S_{\tau}^{\text{mod}} \left(k - 2\overline{\Delta} \right) + O\left(\sigma_0 \frac{\alpha_s \Lambda_{QCD}}{Q} \right)$$

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Hoang, Stewart Hoang, Jain, Scimemi, Stewart

Renormalon Substraction

 $\overline{\text{MS}}$ perturbative series includes fluctuations with arbitrarily small momenta \rightarrow large unphysical corrections

Both S_{τ}^{part} and $\overline{\Omega}_{1}$ suffer from renormalon

Introduce gap parameter Δ : $S_{\tau}^{\text{mod}}(k) \rightarrow S_{\tau}^{\text{mod}}(k-2\Delta)$ Hoang, Stewart

and
$$\Delta = \overline{\Delta}(R, \mu_s) + \delta(R, \mu_s)$$

 \rightarrow renormalon free soft function:

$$S_{\tau}(k,\mu) = \int dk' \left(e^{-2\delta \frac{\partial}{\partial k}} S_{\tau}^{part}(k-k',\mu) \right) S_{\tau}^{mod}(k'-2\overline{\Delta})$$

renormalon substracted

renormalon NOT substracted



MS perturbative series includes fluctuations with arbitrarily small momenta

→ large unphysical corrections

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Kniehl, Kuhn Hagiwara, Kuruma, Yamada



 $\alpha_s(m_Z)$ from global thrust fits

Comparison with similar analysis

	sum logs	power corrections	data	$\alpha_{S}(M_{Z})$
Dissertori et al.	no	Monte Carlo (MC)	ALEPH	0.1240(34)
Dissertori et al.	NLL	Monte Carlo	ALEPH	0.1224(39)
Becher, Schwartz	N3LL	uncertainty from MC	ALEPH, OPAL	0.1172(21)
Davison, Webber	NLL	effective coupling model	Most of data	0.1164(28)
Bethke et al.	NLL	Monte Carlo	JADE	0.1172(51)

Becher, Schwartz:

- no nonperturbative soft function
- different profile function
- different way of calculating binned cross section



Numerical Analysis









Perturbative Uncertainty

12 theory parameters:

- 6 parameters for the variation of the renormalization scales
- 3 parameters related to the statistical uncertainties of numerical fixed order calculations
- 3 parameters for Padé approximants of unknown constants
- 1. Flat random scan over theory parameters
- 2. Fitting for each parameter set
- Range of best fits → perturbative uncertainty





same analysis for other event shape variables: Heavy Jet Mass AHMS+Schwartz

> fits to **moment data** AFHMS



Moment Fits

*n*th moment:
$$M_n = \int_0^{\tau_{max}} d\tau \tau^n \int_0^{Q\tau} dp \frac{d\hat{\sigma}}{d\tau} (\tau - \frac{p}{Q}) S_{\tau}^{mod}(p) + O\left(\frac{\alpha_s \Lambda_{QCD}}{Q}\right)$$

purely perturbative
$$M_n = \sum_{k=0}^n \binom{n}{k} \hat{M}_k \left(\frac{2}{Q}\right)^{n-k} \Omega_{n-k} + O\left(\frac{\alpha_s \Lambda_{QCD}}{Q}\right)$$

where $\Omega_i = \int dp \left(\frac{p}{2}\right)^i S_{\tau}^{\text{mod}}(p)$ is defined as for the distribution

Higher sensitivity for higher moments with primed moments:

$$M'_{n} = \hat{M}'_{n} + \left(\frac{2}{Q}\right)^{n} \Omega'_{n} + O\left(\frac{\alpha_{S}\Lambda_{QCD}}{Q}\right)$$

appear to also be 1/Qⁿ!!!!

$$\Omega'_{1} = \Omega_{1}$$

$$\Omega'_{2} = \Omega_{2} - \Omega_{1}^{2}$$

$$\Omega'_{3} = \Omega_{3} - 3\Omega_{2}\Omega_{1} + 2\Omega_{1}^{3}$$

Global fits to M₁ data



Global fits to M₁ data





Comparison to Gehrmann



 \rightarrow fits to the higher moments are all sensitive to the same M_1 with dispersive model for power corrections:

$$P = \frac{4C_F}{\pi^2} \cdot M \cdot \left\{ \alpha_0 - \left(\alpha_S(\mu_R) + \frac{\beta_0}{\pi} \alpha_S^2(\mu_R) \left(\ln \frac{\mu_R}{\mu_I} + 1 + \frac{K}{2\beta_0} \right) + O(\alpha_S^3) \right\} \cdot \frac{\mu_I}{Q} \right\}$$

Summary

• $\alpha_s = 0.1135 \pm (0.0002)_{expt} \pm (0.0005)_{hadr} \pm (0.0009)_{pert}$



- SCET can improve convergence and therefore precision
- SCET allows to include non-perturbative effects in a systematic manner
- Profile functions allow to combine several kinematic regions



Backup Slides

Why do we need a global fit?



QED and b-mass effects



Cut on Dataset Ω_2 effects au_{min} increase $\frac{5}{O}$, 0.38 #bins au_{\min}, au_{\max} 0.09_0.33 394 0.80 $2\Omega_1$ $\left[\frac{5}{O}, 0.33\right]^{516}$ (GeV) $\frac{6}{Q_1}$ 0.38 decrease of $\Delta \tau$ 0.70 $\left[\frac{8}{O}, 0.33\right]^{442}$ \overline{O} vary, 0.33 \rightarrow increase of statistical $\left[\frac{6}{Q}, 0.33\right]_{\text{strict}}^{487}$ uncertainty 0.60 $\frac{8}{Q}$, 0.33 $\frac{427}{\text{strlet}}$ 0.50 $\frac{8}{Q}$, 0.25 strict 39% CL 68% CL au_{max} 0.40 0.112 0.113 0.114 0.115 0.116 0.111 $\alpha_s(m_Z)$ missing $\alpha_s \Lambda_{QCD}/Q$ effects become important

Theory Parameters

	range of values	default value	parameter
	1.5 to 2.5 GeV	$2{ m GeV}$	μ_0
profile function	2 to 8	5	n_1
(variation of renormalization scales)	0.20 to 0.30	0.25	t_2
	-1,0,1	0	e_J
	0.5 to 2.0	1	e_H
	-1,0,1	0	n_s
	-36.6 to -41.6	-39.1	s_2
Padé approximats	-1553.06 to $+4569.18$	1553.06	Γ_3^{cusp}
	-3000 to $+3000$	0	j_3
	-500 to $+500$	0	s_3
non-singular stat. uncertainty	-1,0,1	0	ϵ_2
	-1,0,1	0	ϵ_3

Include b-mass effects in Factorization Thm: (~ 2% effect)



- at this order it effects only the jet function and $_{ au}$ limits
- use SCET massive fact. thm

Fleming, Hoang, Mantry, Stewart

charm quarks are much smaller effect

Include QED effects in Factorization Thm: (~2% effect)

- count $\ lpha \sim lpha_s^2$, include only final state radiation
- include $\mathcal{O}(\alpha_s^2 \alpha)$ corrections to QCD β -function
- include one-loop QED corrections to H_Q, J_τ, S_τ

Include axial anomaly contribution $(\sim 1\% \text{ effect})$ • affects H_Q^{ua}, H_Q^{da} at $\mathcal{O}(\alpha_s^2)$ $H_Q^a = H_Q^v + H^{\text{singlet}} f^{da}(\tau, 1) = f^v(\tau, 1) + \frac{\alpha_s^2}{4\pi^2} f_{\text{singlet}} \left(\tau, \frac{Q^2}{4m_t^2}\right)$

due to large top-bottom mass splitting



Kniehl, Kuhn Hagiwara, Kuruma, Yamada

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Calculating binned cross-sections



How should we calculate the bins?

At this level of precission, it is important to calculate theory results for the entire bin (even though they are quite fine)

Difference of cumulants $\Sigma(\tau_2, \mu_i(\tau_2)) - \Sigma(\tau_1, \mu_i(\tau_1))$

[all classic resummation analyses] [Becher & Schwartz] [Chien & Schwartz]

Integrating the differential distribution

$$\int_{\tau_1}^{\tau_2} \mathrm{d}\tau' \frac{\mathrm{d}\sigma}{\mathrm{d}\tau'}(\tau',\mu_i(\tau'))$$

[AFHMS] [AHMS+Schwartz]

The two procedures look the same, but actually they are not



Enhanced uncertainty from the peak !