

Electroweak properties of Weakly-Bound Light Nuclei

Doron Gazit

האוניברסיטה העברית בירושלים
The Hebrew University of Jerusalem



Weakly-Bound Systems in Atomic and Nuclear Physics – March 2010

INSTITUTE FOR NUCLEAR THEORY



Collaborators



Sonia Bacca

Winfried Leidemann, Giuseppina



Orlandini



Sofia Quaglioni, Petr Navratil



Achim Schwenk

האוניברסיטה העברית בירושלים
The Hebrew University of Jerusalem



Nir Barnea



Ho-Ung Yee

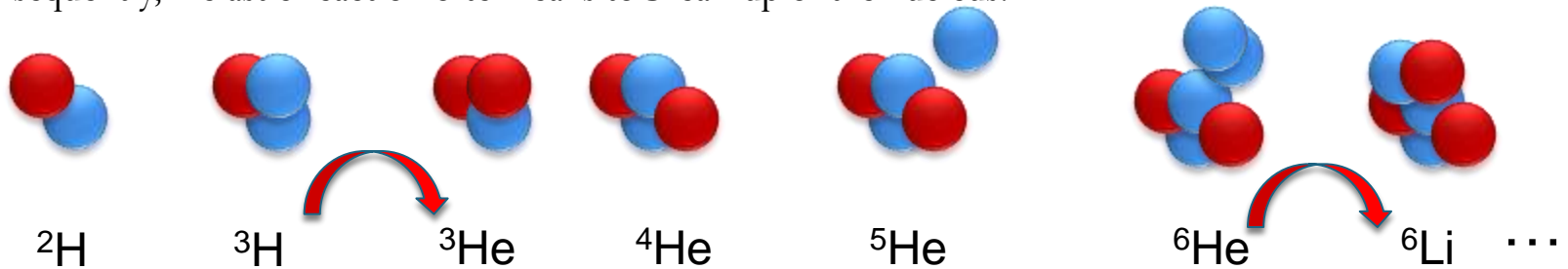


Introduction

- *Light nuclei are special:*

- Weakly-bound:

- Usually include few (if any) bound excited states.
- Consequently, inelastic reaction often leads to break-up of the nucleus.



B.E [MeV]	-2.22	-8.48	-7.72	-28.3	unbound	-29.3	-32.0
-----------	-------	-------	-------	-------	---------	-------	-------

B.E/mass	$1 \cdot 10^{-3}$	$3 \cdot 10^{-3}$	$3 \cdot 10^{-3}$	$7 \cdot 10^{-3}$	unbound	$5 \cdot 10^{-3}$	$5 \cdot 10^{-3}$
----------	-------------------	-------------------	-------------------	-------------------	---------	-------------------	-------------------





Introduction

- *Why looking at electro-weak properties?*

- One can immediately relate **electro-weak reaction** of a probe with nucleus to currents inside the nucleus.

$$\hat{H}_W \sim \int d^3\vec{x} \hat{j}_\mu^+(\vec{x}) \hat{J}^{\mu-}(\vec{x})$$

Lepton current

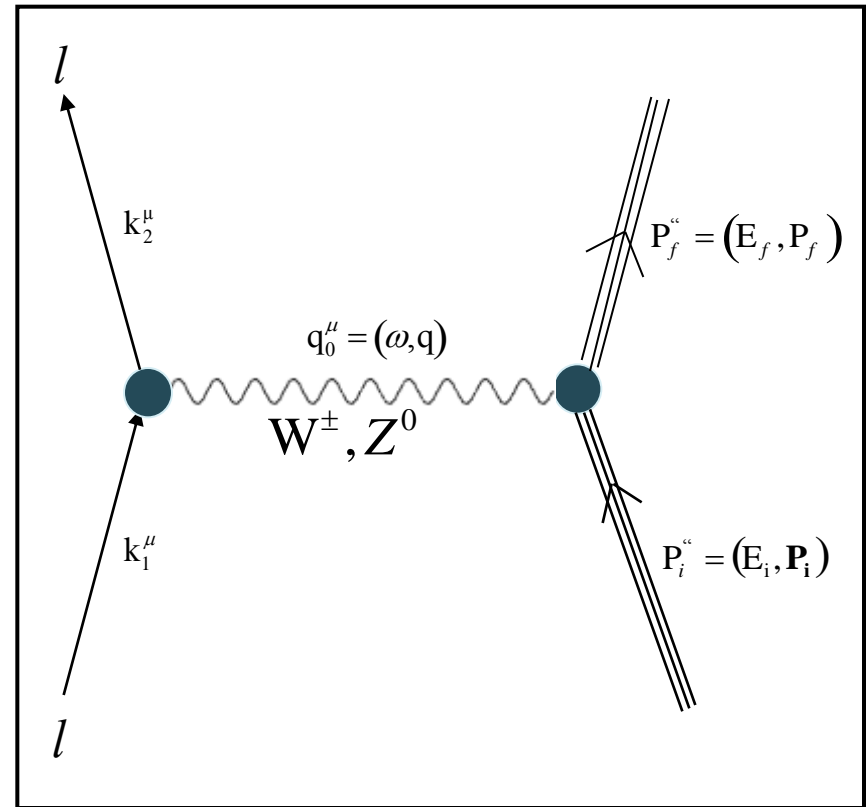
Nuclear current

$$\mathcal{O} \sim \langle \psi_i | \hat{J}^\mu | \psi_f \rangle$$

Scattering operator



Currents in the nucleus



$$W, Z \text{ propagator} = \frac{g_{\mu\nu} + \frac{q_\mu q_\nu}{M_W^2}}{q^2 + M_W^2} \xrightarrow{q \ll M_W} \frac{g_{\mu\nu}}{M_W^2}$$



Introduction

- *Why looking at electro-weak properties?*
 - One can immediately relate *electro-weak reaction* of a probe with nucleus to currents inside the nucleus.
 - The currents are *reflections* of the *symmetries* of the nuclear interaction.
 - Thus, one can *relate electro-weak properties and reaction rates* with non-trivial properties not only of the *target*, but also of the *fundamental theory* leading to its structure!
- In addition, electro-weak properties are important as a microscopic input for simulations of astrophysical phenomena.
 - Solar fusion.
 - Supernovae.
- These are often very challenging, or even impossible to measure, thus need accurate, parameter free predictions.



Introduction

- *Why looking at electro-weak properties?*
- *Why light Nuclei?*
 - Available methods for *solving exactly the Schrödinger equation for few body systems*, from nucleonic dof: no core shell model, expansions in Hyperspherical Harmonics, Green's function Monte Carlo,
 - *Chiral effective field theory (χ PT)*, enables a connection between the fundamental theory of QCD and the nuclear interaction.
- *Allow accurate, parameter free calculations, of weakly bound systems, and their properties*, from their nucleonic dof.



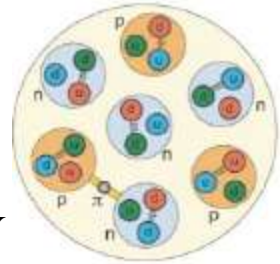
Chiral Effective Field Theory

- Symmetries are important **NOT** degrees of freedom.
- In QCD – an approximate chiral symmetry:

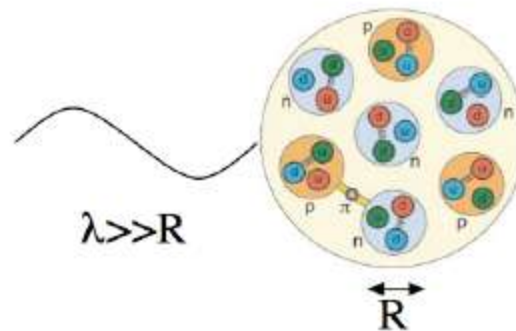
- The u and d are (almost) massless.

$$SU(2)_L \times SU(2)_R \cong SU(2)_V \times SU(2)_A \rightarrow SU(2)_V$$

- The $SU(2)_V$ symmetry is the isospin symmetry.
- However, no degenerate parity doublets are found in the spectrum.
- The axial symmetry is spontaneously broken.
- Chiral EFT is based on this observation.
- Identify Q – the momentum scale of the process.
- Choose Λ – the theory cutoff.
- In view of these-identify the effective degrees of freedom.



$$\lambda \sim \frac{1}{Q} \gg \frac{1}{\Lambda} \sim R$$





Chiral Effective Field Theory

- The pions are interpreted as the Goldstone bosons of the spontaneously broken $SU(2)_A$ symmetry.
- Their mass is a result of the explicit symmetry breaking due to the finite u and d masses. This introduces an additional scale.
 - If $Q \ll \Lambda \ll m_\pi$ then the effective theory is of point particles (pionless χ PT).
 - If $Q \sim m_\pi \ll \Lambda$ then the effective theory should consist of both pions and nucleons (pionfull χ PT), and even higher dof: Delta resonance, etc.
- Write a Lagrangian composed of ALL possible operators invariant under symmetries of the underlying theory.
- Find a *systematic* way to organize diagrams according to their contribution to the observable.



Weinberg's Power Counting Scheme

- Each Feynman diagram can be characterized by: $\left(\frac{Q}{\Lambda}\right)^{\nu}$
- Weinberg showed that ν is bound from below.
- In addition, expand in the inverse of the nucleon's mass (take $\Lambda \sim M_N$) \rightarrow **Heavy Baryon χ PT**.
- This power counting is based on an expansion around the RG fixed point $Q=0$.
- There are indications that this is correct only for a limited range of Λ , due to the existence of a non-trivial, unitary, fixed point.
- This is also evident in the abnormal size of the NN interaction induced by a pion exchange (however, expansions based on the latter seem to have convergence problems (KSW)).

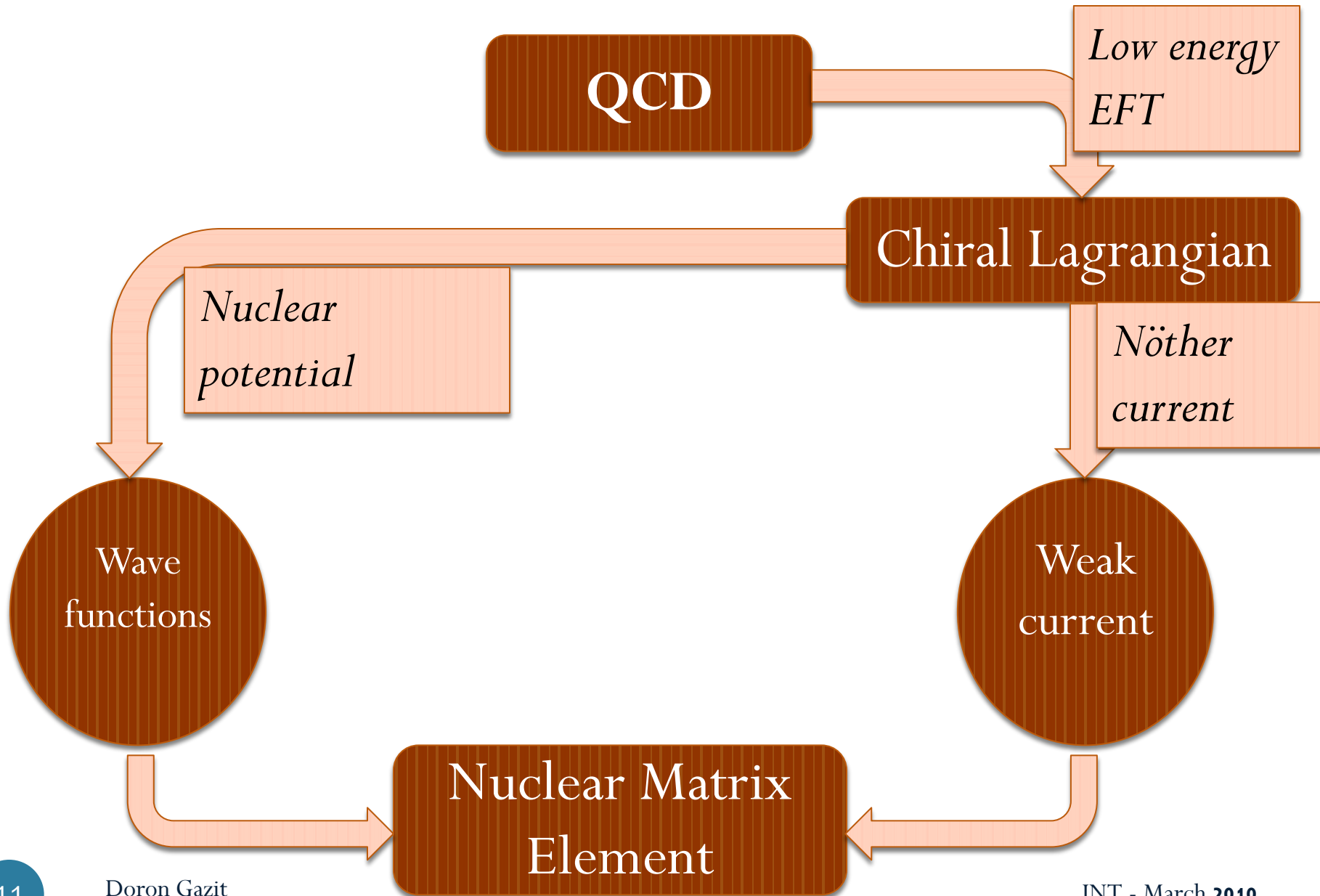
20 years of debate led by: Weinberg, Kaplan, Savage, Wise, van-Kolck, Nogga, Timmermans, Birse, Meissner, Epelbaum...



The big deal in χ PT

- A perturbation theory/ expansion in small parameter of the observable, gives control over the accuracy of the calculation.
- Varying the cutoff gives estimate of the theoretical error-bar.
- Allows connection between *a-priori* unrelated operators:
 - In particular the nuclear force and the electro-weak currents in the nucleus (that the $Su(2)\times Su(2)$ structure is a gauging of).
- When the low-energy constants are known: the calculations are predictions of QCD.

χ PT approach for low-energy EW nuclear reactions:





Forces in χ PT

- The leading order NNN forces are at N^2 LO.
- They include 2 new contact parameters.
- No new parameters at N^3 LO.

χ^2/datum for the reproduction of the
1999 np database

Bin (MeV)	# of data	N^3 LO	NNLO	NLO	AV18
0–100	1058	1.06	1.71	5.20	0.95
100–190	501	1.08	12.9	49.3	1.10
190–290	843	1.15	19.2	68.3	1.11
0–290	2402	1.10	10.1	36.2	1.04

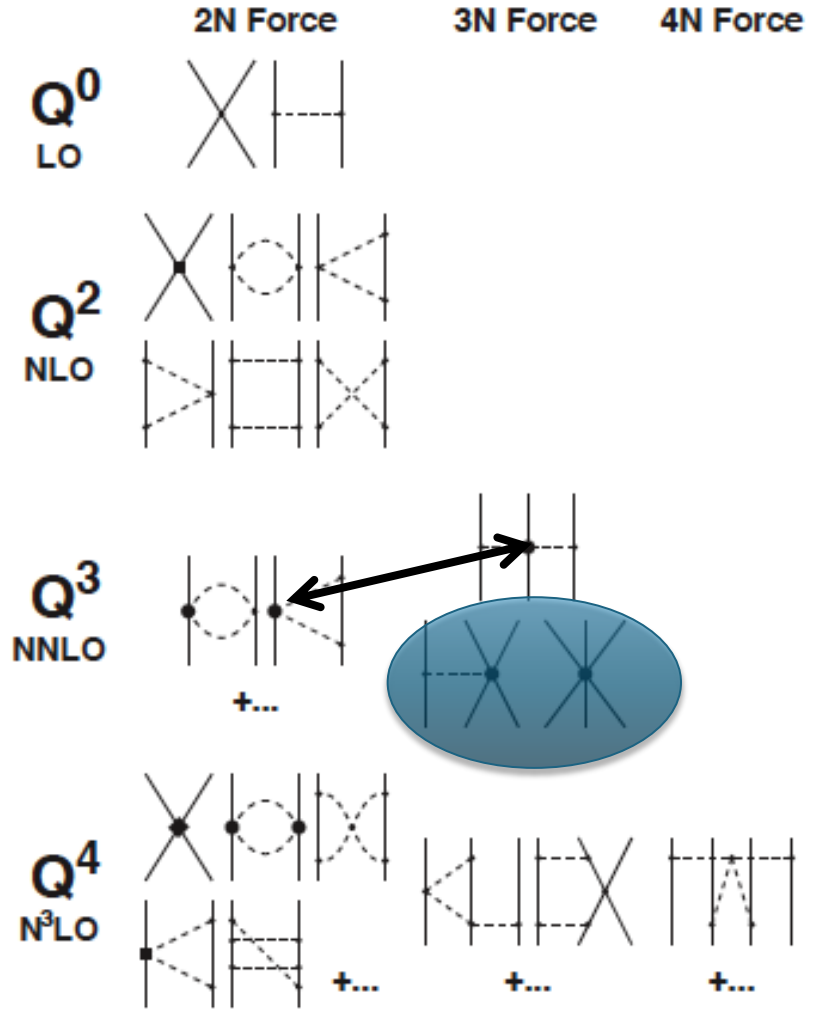




Photo-dissociation

- Current conservation leads to a connection between the spatial current and charge density.
- At low energy transfer the scattering operator is simply the dipole operator (Siegert theorem).
- This approximation is accurate to about 10% at 100 MeV.

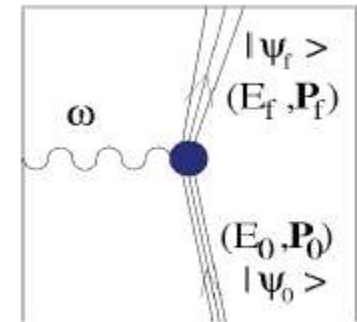
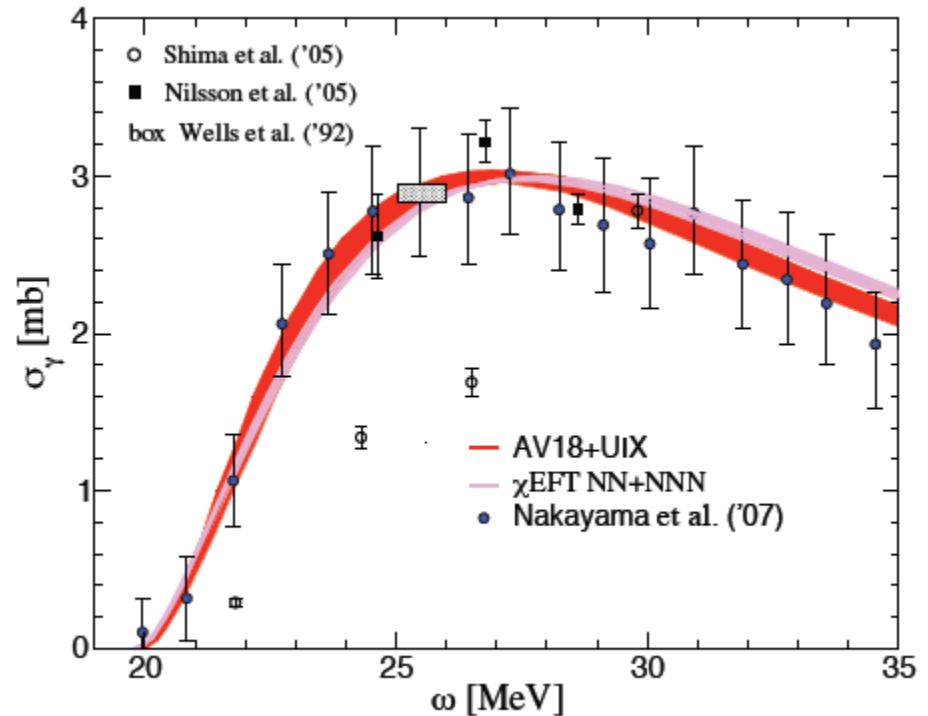




Photo-dissociation of ^4He

- Sensitivity to NN force model.
- Sensitivity vanishes when adding the 3NF.
- The theoretical prediction is much more accurate than the experimental measurement, and actually “chooses” the “correct” measurement.



AV18, UIX [D.Gazit, S.B. et al. PRL 96 112301 \(2006\)](#)

Chiral EFT [S.Quaglioni and P.Navratil PLB 652 \(2007\)](#)



The structure of ${}^4\text{He}$

- ${}^4\text{He}$ is a spherical nucleus ($J=0$), but what is its symmetry in the body frame?
- The unretarded dipole approximation to the photodissociation cross-section can be related to the mean inter-nucleon distances:

$$\begin{aligned}\Sigma_{BSR} &= \int_{\omega_h}^{\infty} \omega^{-1} \sigma_{\gamma}^{E1UR} d\omega = \frac{3}{4\pi^2\alpha} \langle 0 | \hat{D} \cdot \hat{D} | 0 \rangle \\ &= \frac{3}{4\pi^2\alpha} \left(Z^2 \langle r_p^2 \rangle - \frac{Z(Z-1)}{2} \langle r_{pp}^2 \rangle \right) = \\ &= \frac{3}{4\pi^2\alpha} \left(N^2 \langle r_n^2 \rangle - \frac{N(N-1)}{2} \langle r_{nn}^2 \rangle \right) = \\ &= \frac{3}{4\pi^2\alpha} \frac{NZ}{2} (\langle r_{np}^2 \rangle - \langle r_n^2 \rangle - \langle r_p^2 \rangle) =\end{aligned}$$

- In ${}^4\text{He}$ this is enough to reconstruct the structure:

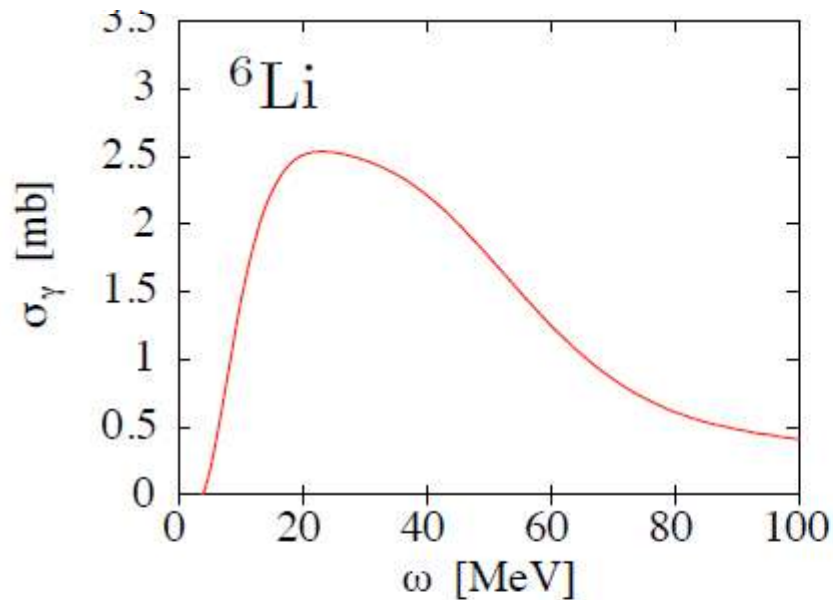
- Thus, ${}^4\text{He}$ has a slightly deformed internal tetrahedral symmetry...

$$\frac{\langle r_{pp}^2 \rangle}{\langle r_p^2 \rangle} = \frac{\langle r_{nn}^2 \rangle}{\langle r_n^2 \rangle} = 2.78 \text{ fm}^2 \frac{\langle r_{np}^2 \rangle}{\langle r_n^2 \rangle} = 2.62 \text{ fm}^2$$

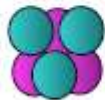


Photodissociation of 6 body nuclei

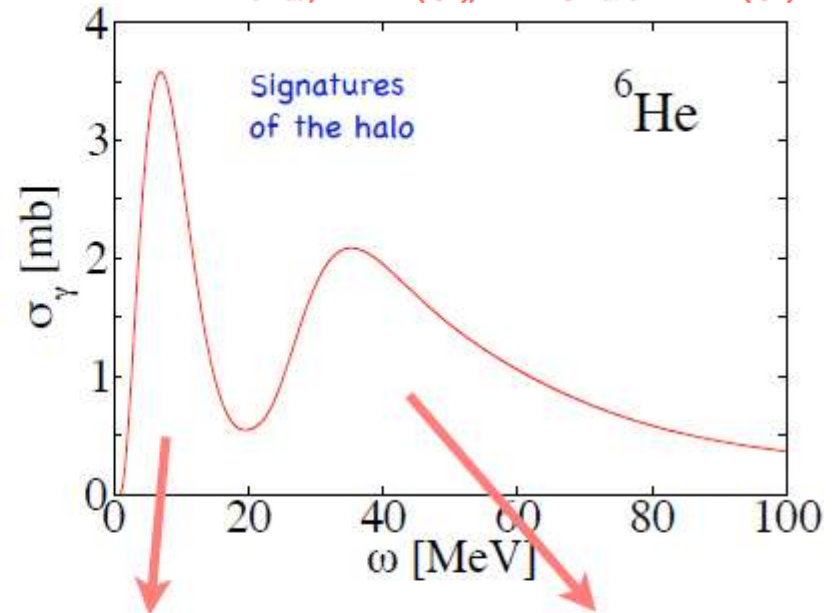
From S. Bacca



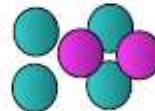
Giant Dipole Resonance



protons \longleftrightarrow neutrons

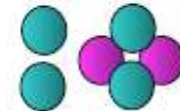


Soft-dipole Mode



neutron halo \longleftrightarrow α -core

Giant Dipole Mode



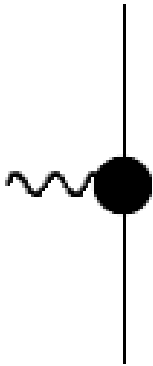
neutrons \longleftrightarrow protons



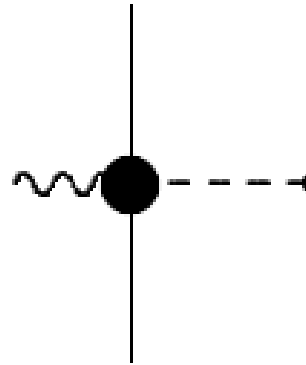
χ PT axial weak currents to fourth order

$$\hat{d}_R \equiv \frac{M_N}{\Lambda_\chi g_A} c_D + \frac{1}{3} M_N (c_3 + 2c_4) + \frac{1}{6}$$

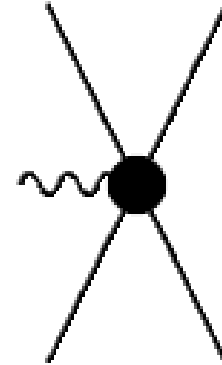
Single nucleon current



1 pion exchange

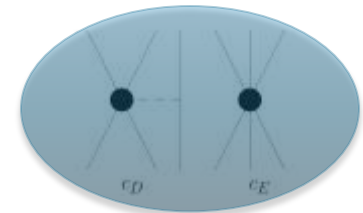


Contact term



Nucleon-pion interaction,
NO new parameters

Contact term



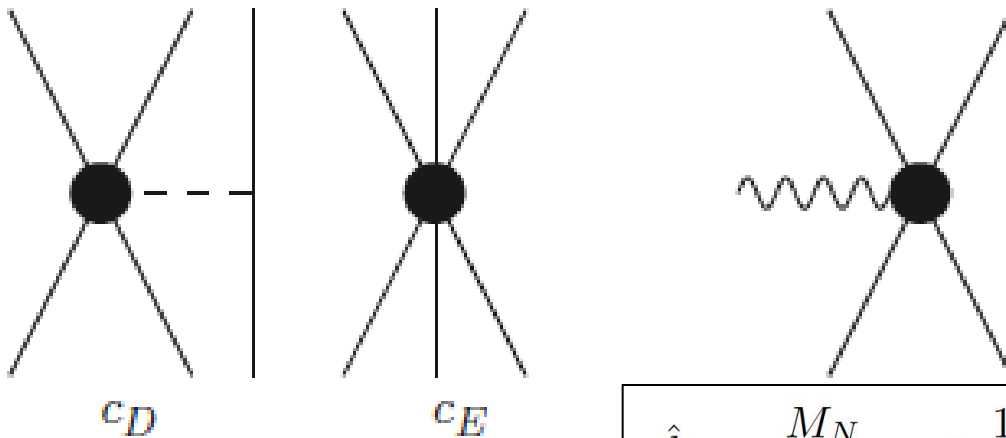


Axial MEC – remarks

- The MEC include “ $O\pi EC$ ” and contact topologies.
- MEC involve only TWO nucleons.
- Thus, in principle c_D can be calibrated using two-body weak processes.
- So – *three nucleon force constrained at the two nucleon level!*
- The most attractive process –
 - Muon capture on deuteron – known only at the 5% level. An experiment at PSI “MuD (MuSun)” aims to measure this process to 1%.
- However, many 3 nucleon processes are measured very well.



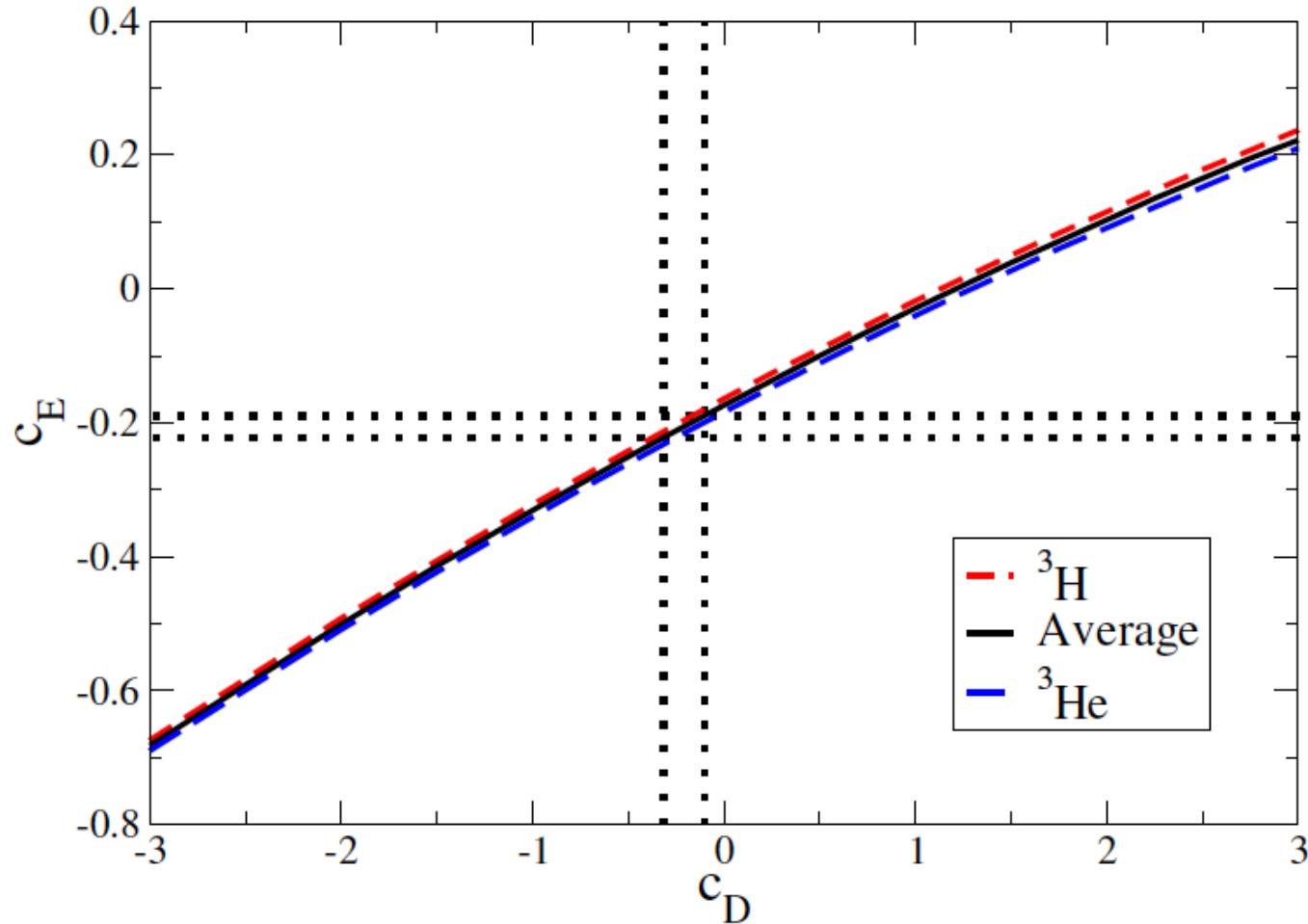
A calculation of ${}^3\text{H}$ β decay using consistent χPT interaction and currents



$$\hat{d}_R \equiv \frac{M_N}{\Lambda_\chi g_A} c_D + \frac{1}{3} M_N (c_3 + 2c_4) + \frac{1}{6}$$

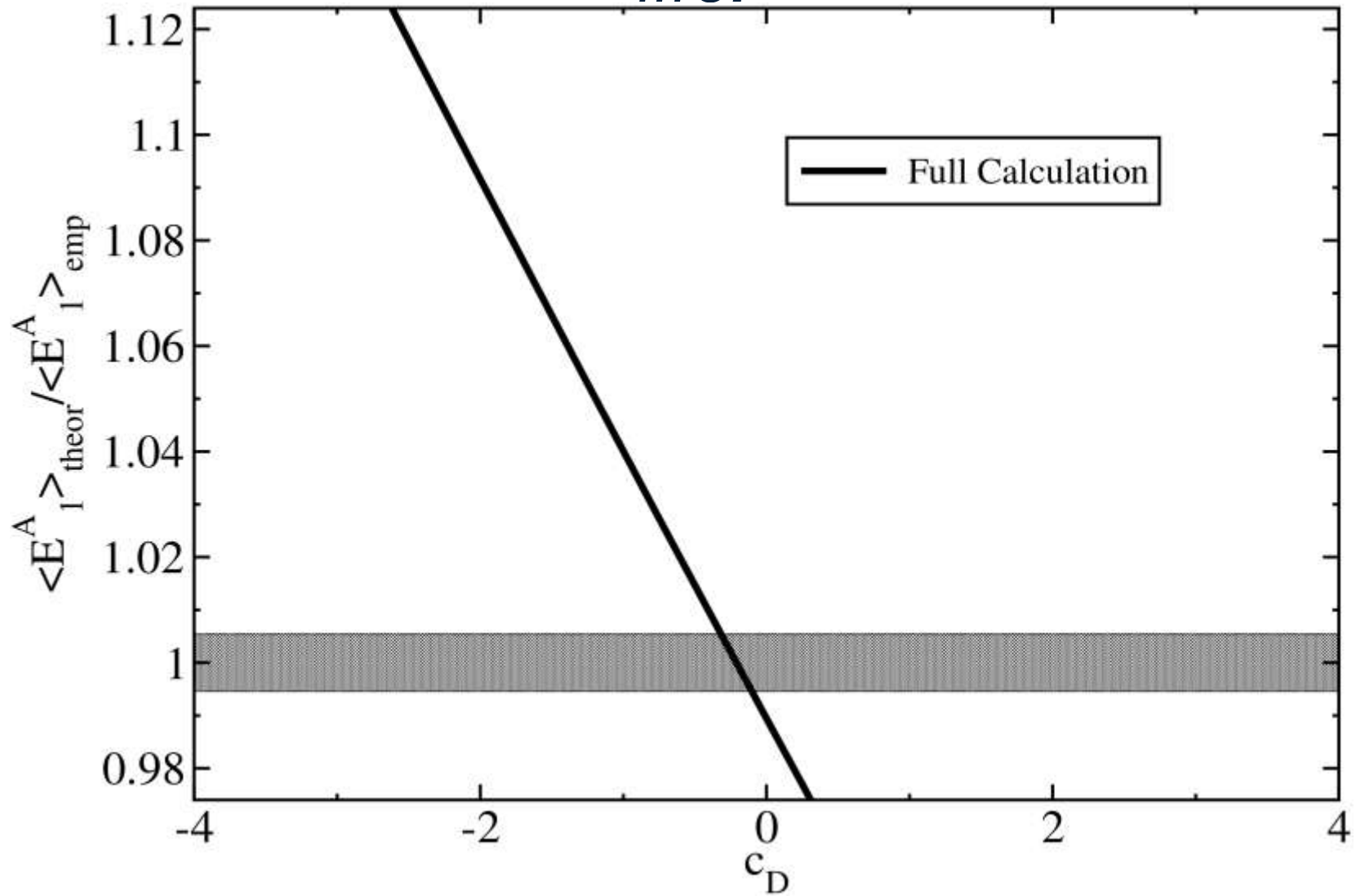
DG, Quaglioni, Navratil, Phys. Rev. Lett. **103**, 102502 (2009).

Step 1: use the trinuclei binding energies to find a c_D - c_E relation



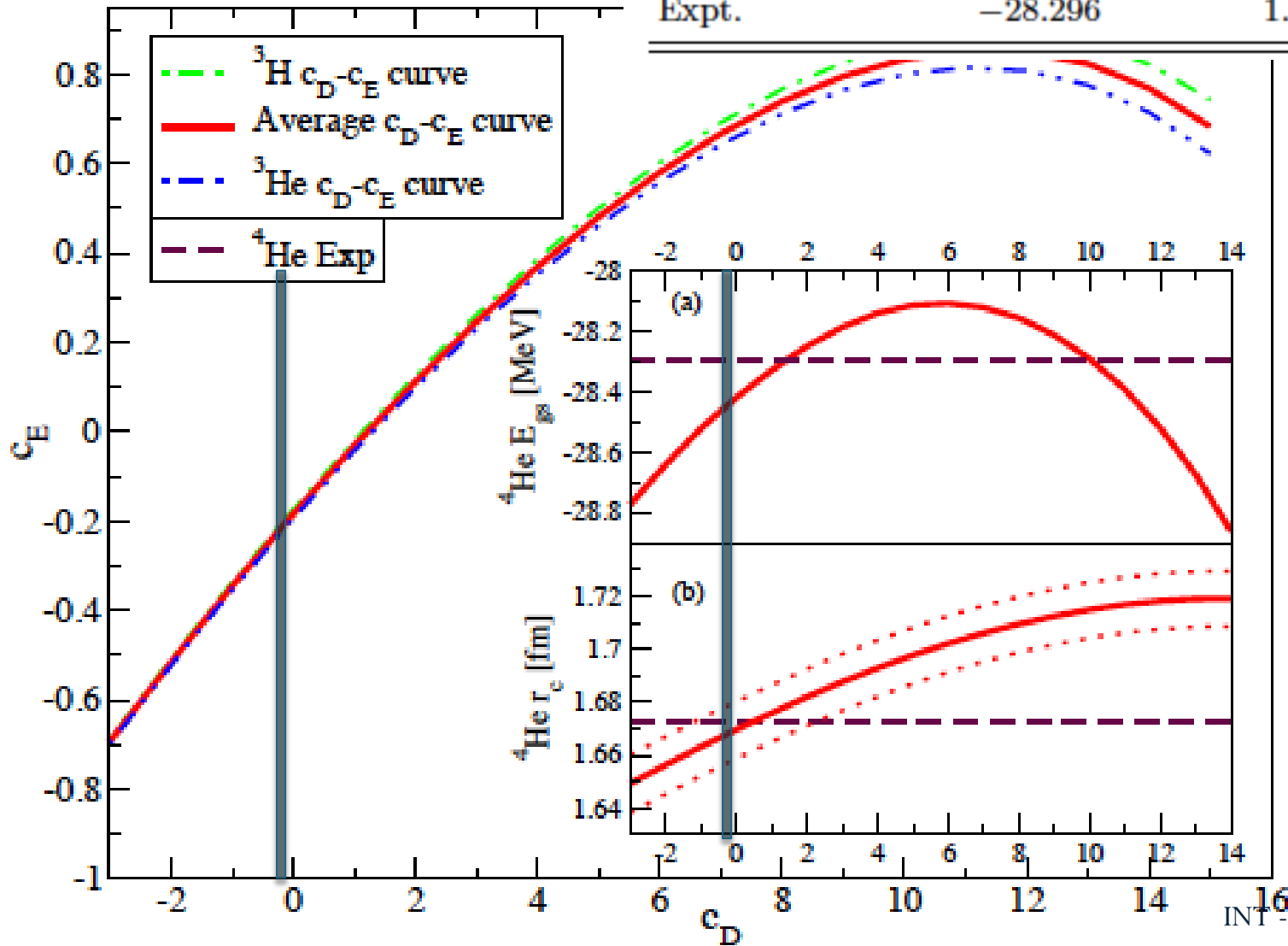
Navratil *et al.*, Phys. Rev. Lett. 99, 042501 (2007).

Step 2: calibrate c_D according to the triton half life.

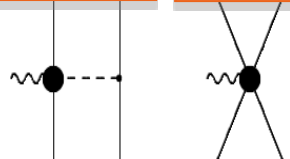


A prediction of ^4He

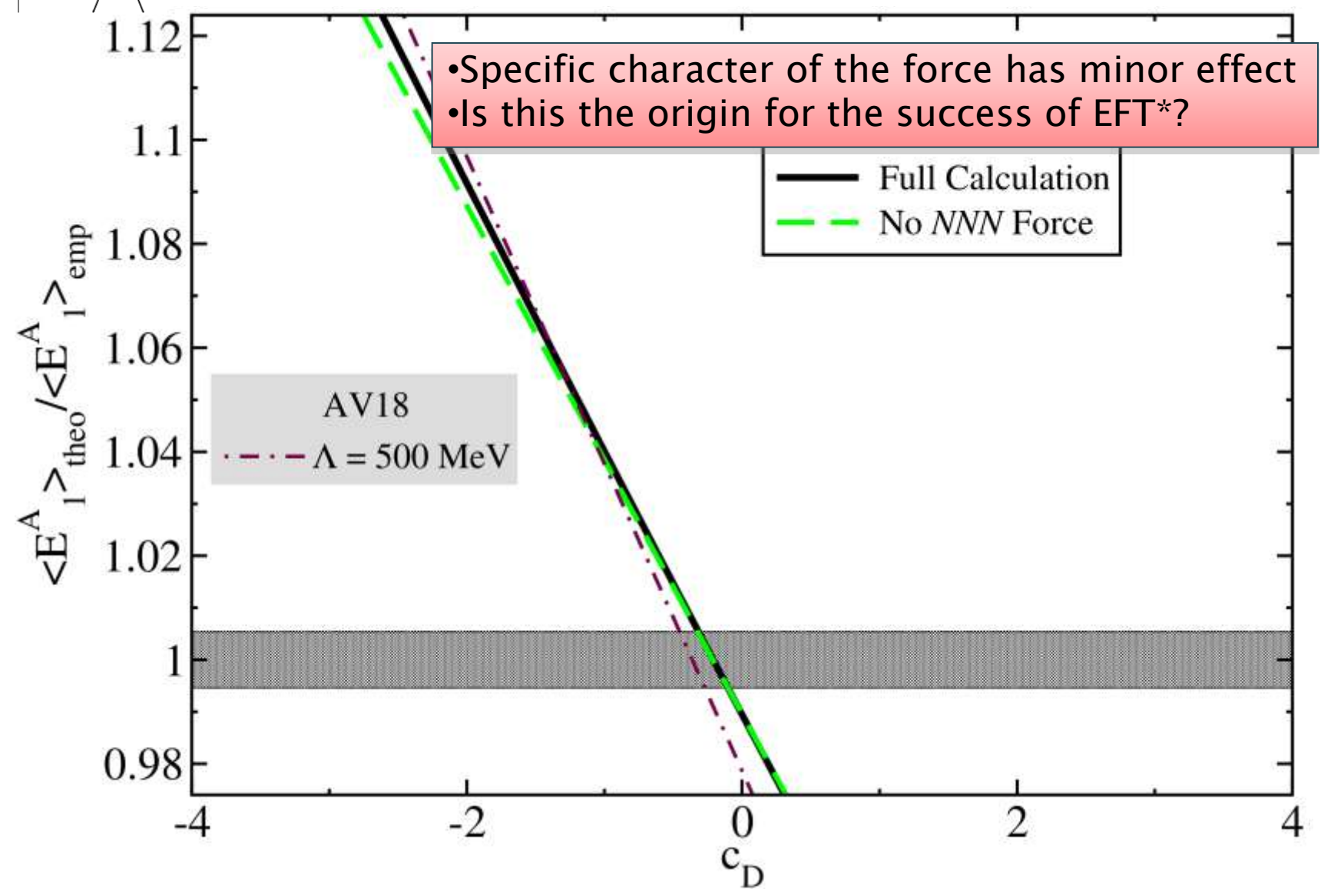
^4He		
	$E_{g.s.}$	$\langle r_p^2 \rangle^{1/2}$
NN	-25.39(1)	1.515(2)
$NN+NNN$	-28.50(2)	1.461(2)
Expt.	-28.296	1.467(13) [24]



OPEC Contact



A closer look into the weak axial correlations in 3H





EFT* approach for low-energy nuclear reactions:

Phenomenological
Hamiltonian

QCD

Low energy
EFT

Chiral Lagrangian

Solution of Schrödinger equation

*Nöther
current*

Wave
functions

Weak
current

*Nuclear Matrix
Element*



Applications of the EFT* approach

- $p+p$ fusion in the sun.
- ${}^3\text{He}+p$ fusion in the sun.

T.-S. Park *et al*, **Phys. Rev. C** **67**, 055206 (2003).
R. Schiavilla *et al*, **Phys. Rev. C** **58**, 1263 (1998).
M. Butler, J.-W. Chen, **Phys. Lett. B** **520**, 87 (2001).
L. Marcucci *et al*, **Phys. Rev. C** **66**, 054003 (2002)
Solar Fusion II, **Rev. Mod. Phys.** [to be published].

- The weak structure of the nucleon from μ capture on ${}^3\text{He}$.

DG, **Phys. Lett. B** **666**, 471 (2008).

- Neutrino scattering on light nuclei in core-collapse supernovae.

DG, Barnea **Phys. Rev. C** **70**, 048801 (2004); **Phys. Rev. Lett.** **75**, 192501 (2007); **Nucl. Phys. A** **790**, 356 (2007); **Few Body Syst.** (2008). DG, PhD. thesis, **arXiv: 0807.0216** (2007).
O'Connor, DG, Horowitz, Schwenk, Barnea, **Phys. Rev. C** **75**, 055803 (2007).

- β decay of ${}^6\text{He}$ and the suppression of the axial constant in nuclear matter.

Vaintraub, Barnea, DG, **Phys. Rev. C**, **79** 065501 (2009).

Go to summary



Few Solar Fusion open problems

T.-S. Park *et al*, **Phys. Rev. C** **67**, 055206 (2003).

R. Schiavilla *et al*, **Phys. Rev. C** **58**, 1263 (1998).

M. Butler, J.-W. Chen, **Phys. Lett. B** **520**, 87 (2001).

L. Marcucci *et al*, **Phys. Rev. C** **66**, 054003 (2002)

Solar Fusion II, **Rev. Mod. Phys.** [to be published].



pp fusion

- The weak fusion process $p+p \rightarrow d+\nu+e^+$, is sun's clock – it's the process determining the sun's evolution rate.
- An open field for state of the art/benchmark calculations with immense prospects.

$$4.01(1 \pm 0.007) \times 10^{-25} \text{ MeV b} \quad \text{potential models}$$

$$4.01(1 \pm 0.007) \times 10^{-25} \text{ MeV b} \quad \text{EFT}^*$$

$$3.99(1 \pm 0.030) \times 10^{-25} \text{ MeV b} \quad \text{pionless EFT (23)}$$

- Needed:
 - 3 nucleon currents in pionless EFT.
 - Consistent calculations in pionfull EFT.



hep process

- The weak fusion $p + {}^3\text{He} \rightarrow {}^4\text{He} + \nu + e^+$ is the source of the most energetic neutrinos.
- The single nucleon current is highly suppressed.
- The EFT* calculation depends strongly on the EFT cutoff.

	$E = 0 \text{ keV}$		$E = 5 \text{ keV}$		$E = 10 \text{ keV}$	
	${}^3\text{S}_1$	S+P	${}^3\text{S}_1$	S+P	${}^3\text{S}_1$	S+P
One-body	26.4	29.0	25.9	28.7	26.2	29.2
Full	6.38	9.64	6.20	9.70	6.36	10.1

$\Lambda \text{ (MeV)}$	500	600	800
$\bar{L}_1(q; A): 1\text{B}$	-0.081	-0.081	-0.081
$\bar{L}_1(q; A): 2\text{B (no contact term)}$	0.093	0.122	0.166
$\bar{L}_1(q; A): 2\text{B (with contact term)}$	-0.044	-0.070	-0.107
$\bar{L}_1(q; A): 2\text{B-total}$	0.049	0.052	0.059
S_{hep}	9.95	9.37	7.32

$$S_{\text{hep}}(0) = 8.3 \pm 1.3 \text{ keV} \cdot \text{b}$$



Summary and outlook

- Electro-weak reactions with light nuclei can be used to:
 - Constrain the Nuclear interaction.
 - Give information regarding the structure of the nucleus.
 - Extract microscopic information about the fundamental theory and its symmetries.
 - Predict, to a percentage level accuracy, reaction rates for astrophysical phenomena.
- Halo helium isotopes (${}^6\text{He}$ and ${}^8\text{He}$) are a great challenge for ab-initio calculation.
- Many challenges in the astrophysical sector. (pp fusion in a consistent manner as an important benchmark).