

TITAN

ISAC-TRIUMF

stephan ettenauer for the TITAN collaboration



Weakly Bound Systems in Atomic and Nuclear Physics, March 2010

### Experimental Program on Halo Nuclei

with non-accelerated Beams at TRIUMF

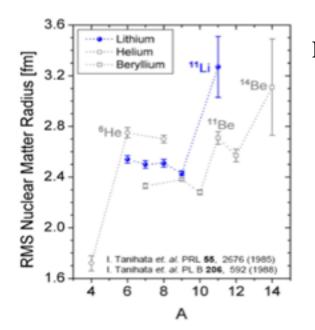


- · Overview: Experimental Probes on Halo
- · Production of Halo Nuclei
- non-accelerated Halos @ TRIUMF
  - Laser Spectroscopy
  - Mass Measurements in Penning Trap
- Conclusion & Outlook

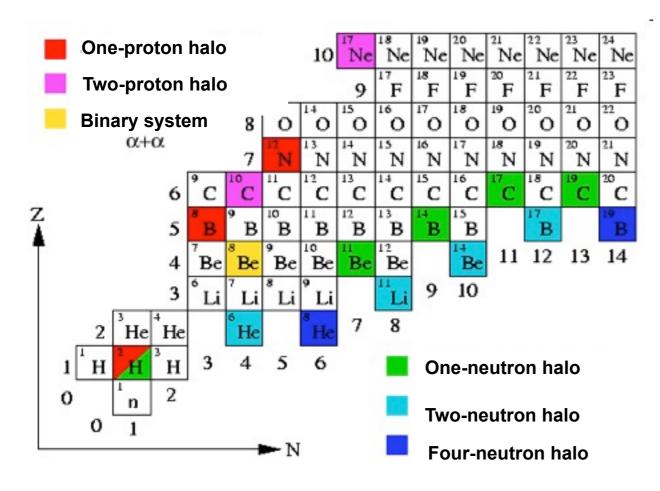


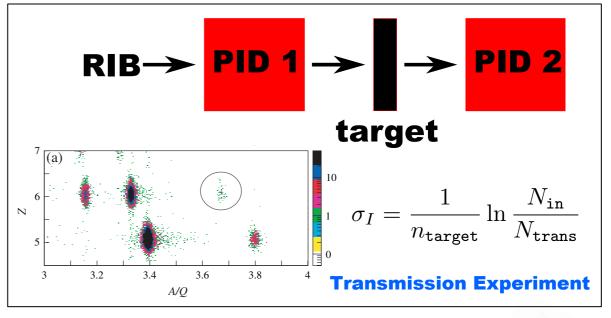


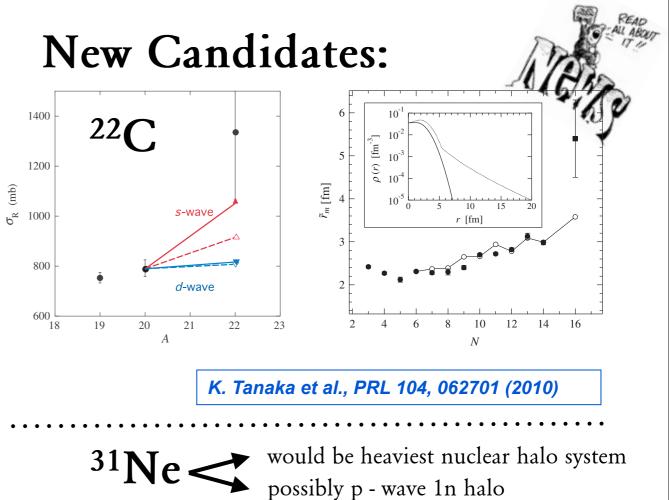
# Halo Nuclei



- In 1985 Tanihata et al.:
- interaction cross section measurements (transmission experiment)
- <sup>11</sup>Li much larger than expected from general rule of stables:  $R_N \sim r_0 A^{1/3}$
- extra neutrons (or protons) in classically forbidden region
- I. Tanihata et al., PRL 55, 2676 (1985)



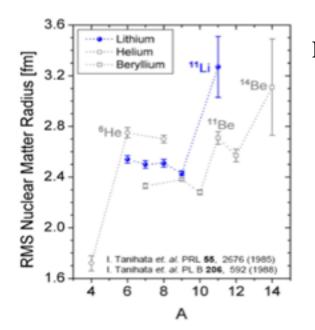




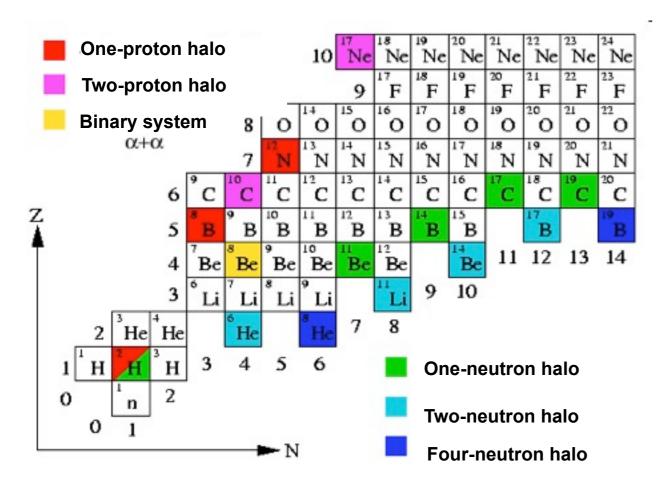
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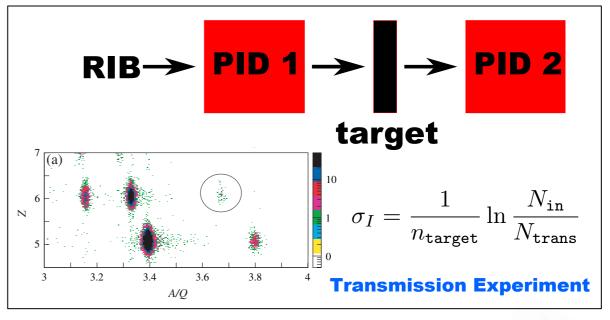


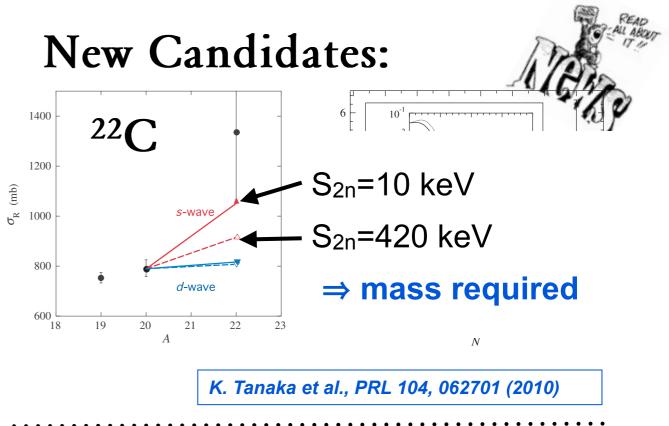
# Halo Nuclei



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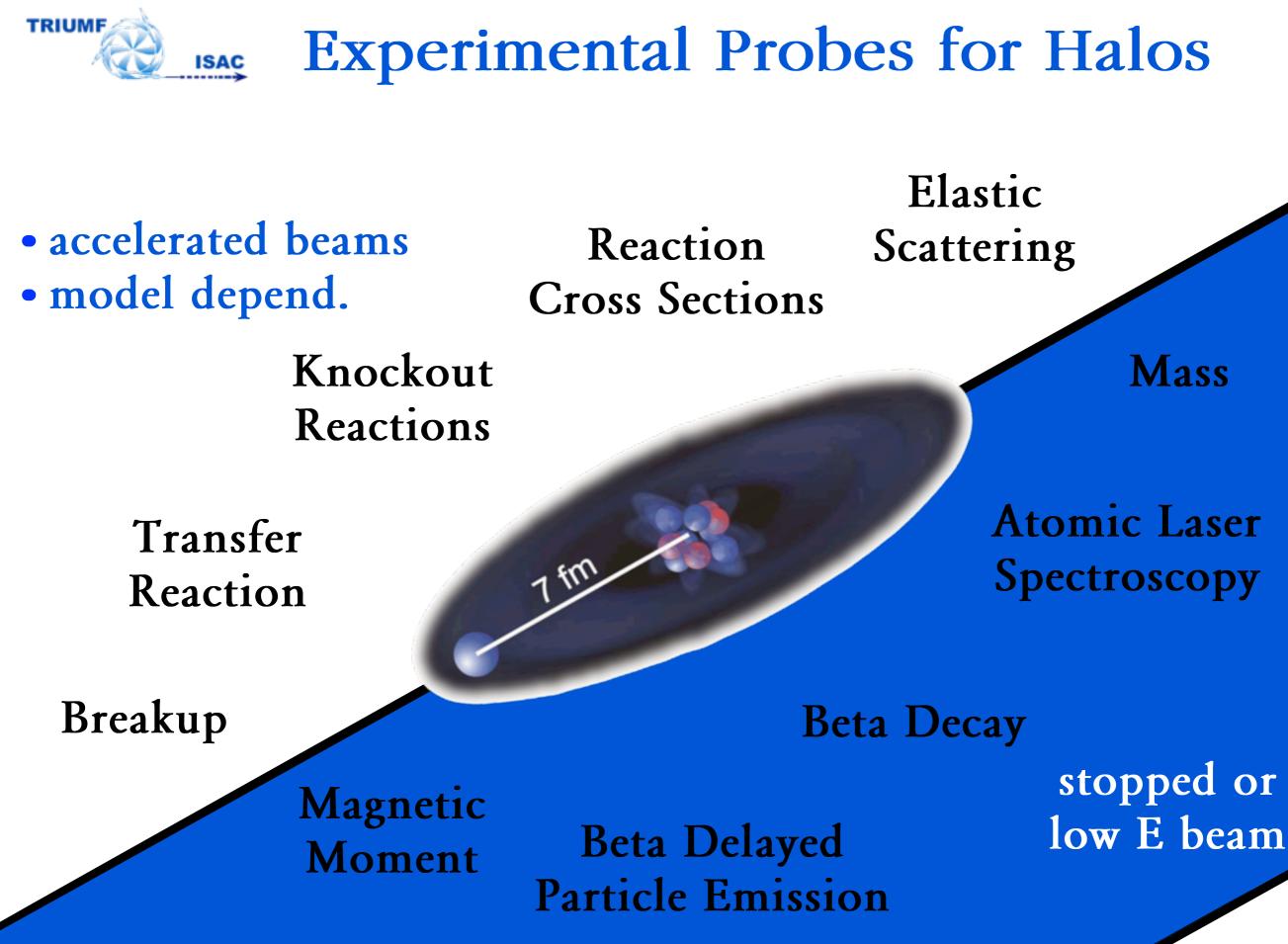


<sup>31</sup>Ne <

would be heaviest nuclear halo system
possibly p - wave 1n halo

3

T. Nakamura et al., PRL 103, 262501 (2009)





**Experimental Probes for Halos** 

Elastic

Scattering

**Beta Decay** 

accelerated beamsmodel depend.

Reaction Cross Sections

Knockout Reactions

Transfer Reaction

Breakup

Magnetic Moment

Beta Delayed Particle Emission stopped or low E beam

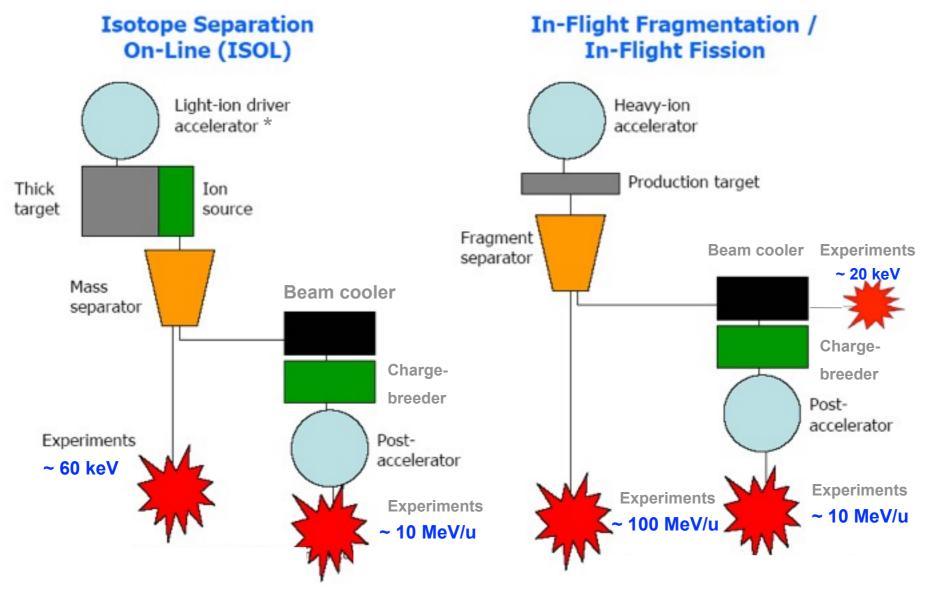
Mass

this talk

**Atomic Laser** 

Spectroscopy

### **Rare Isotope Production**



### ISOL (TRIUMF, ISOLDE@CERN):

TRIUMF

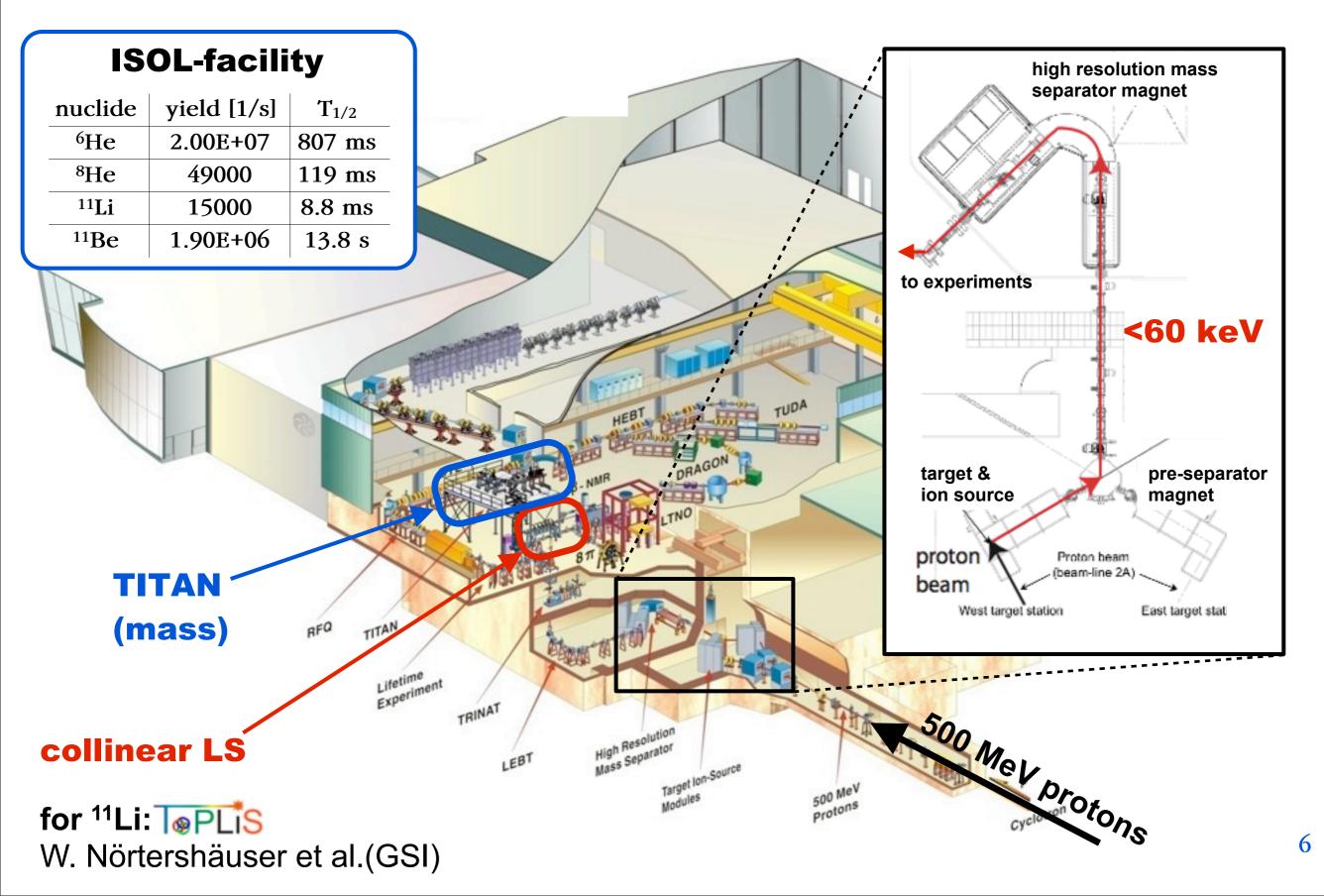
ISAC

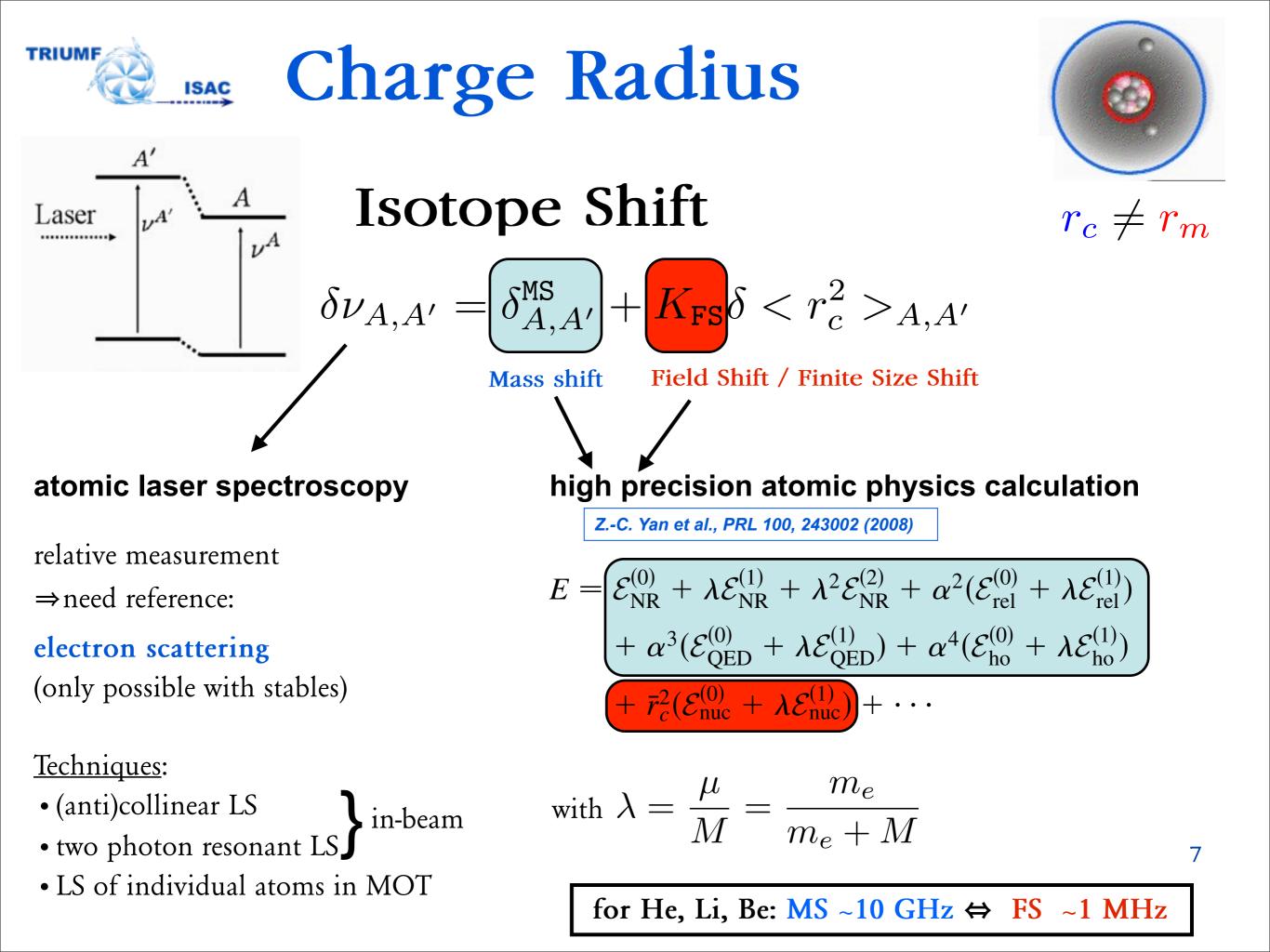
- Production: slow (~5 ms) BUT high intensity
- Low beam energy, ideal for decay and trap exp.
- Good beam quality (even cooled) & purity
- Post-acceleration for reaction studies
- BUT element selective ionization
  - $\Rightarrow$  some elements not possible!

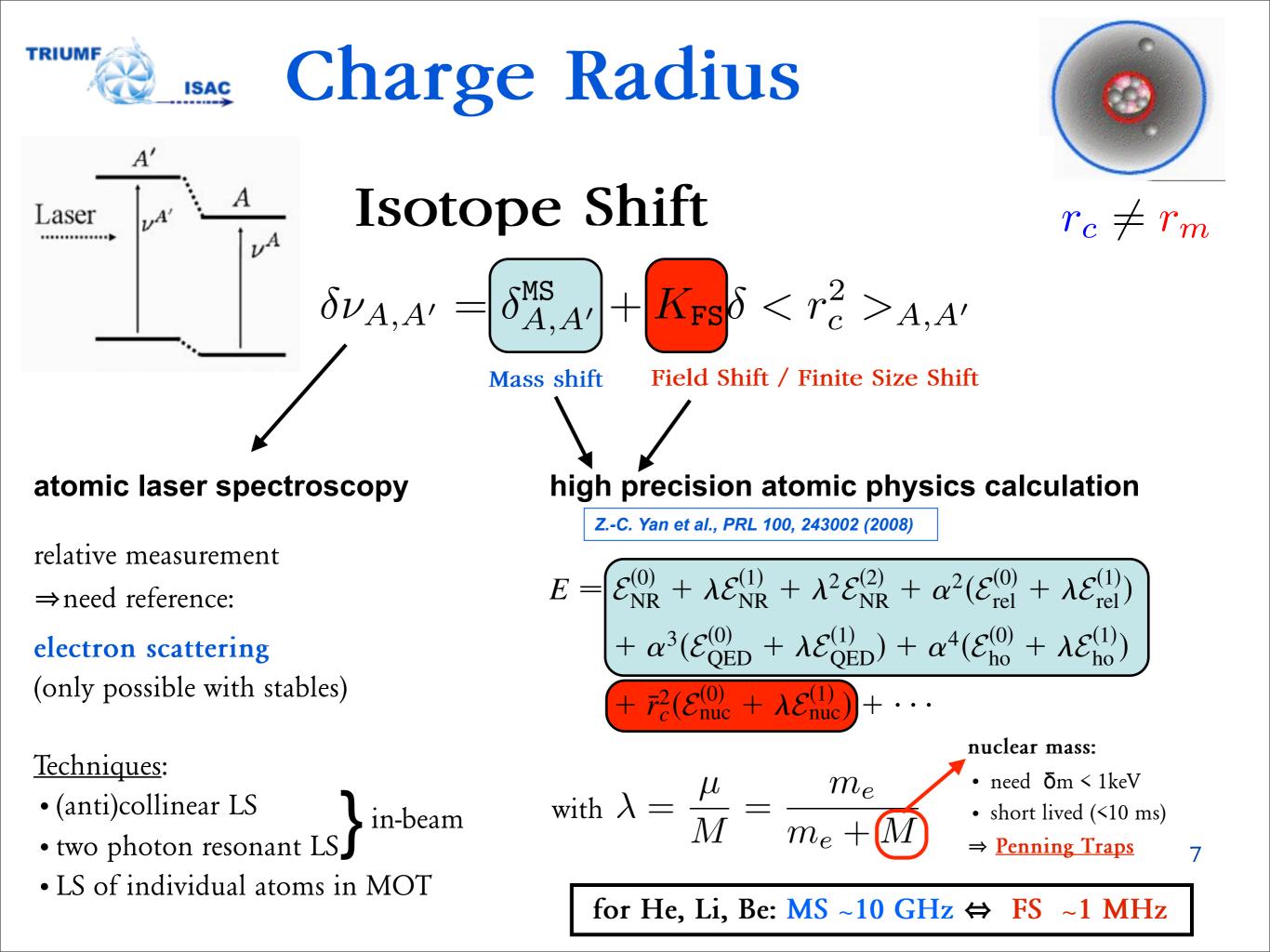
### In-Flight (MSU, GSI, RIKEN, GANIL):

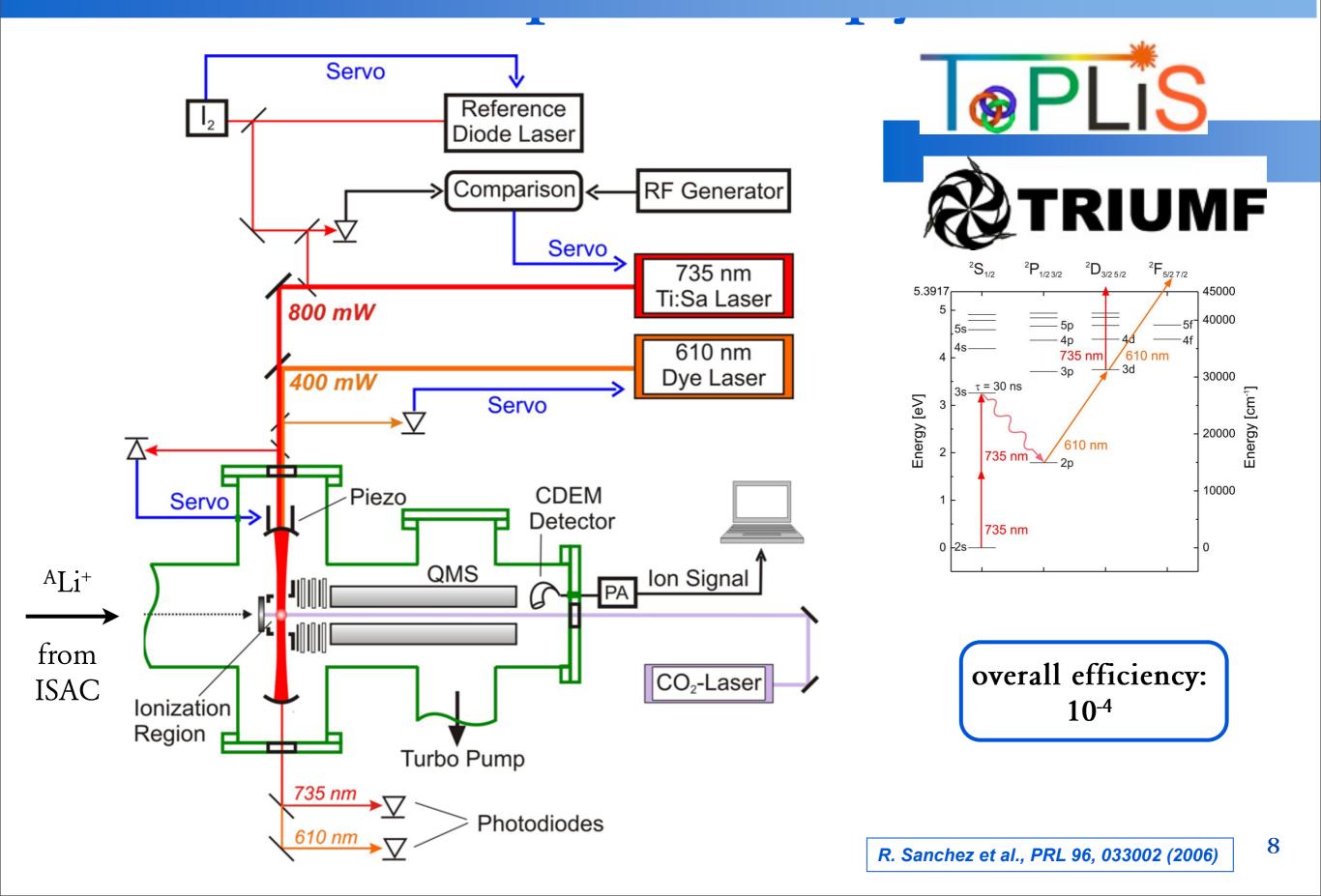
- Production: fast, no chemistry involved
- High beam energy, ideal for reaction exp.
- Life-time, masses, & basic discovery
- Low intensity, poor beam quality & purity



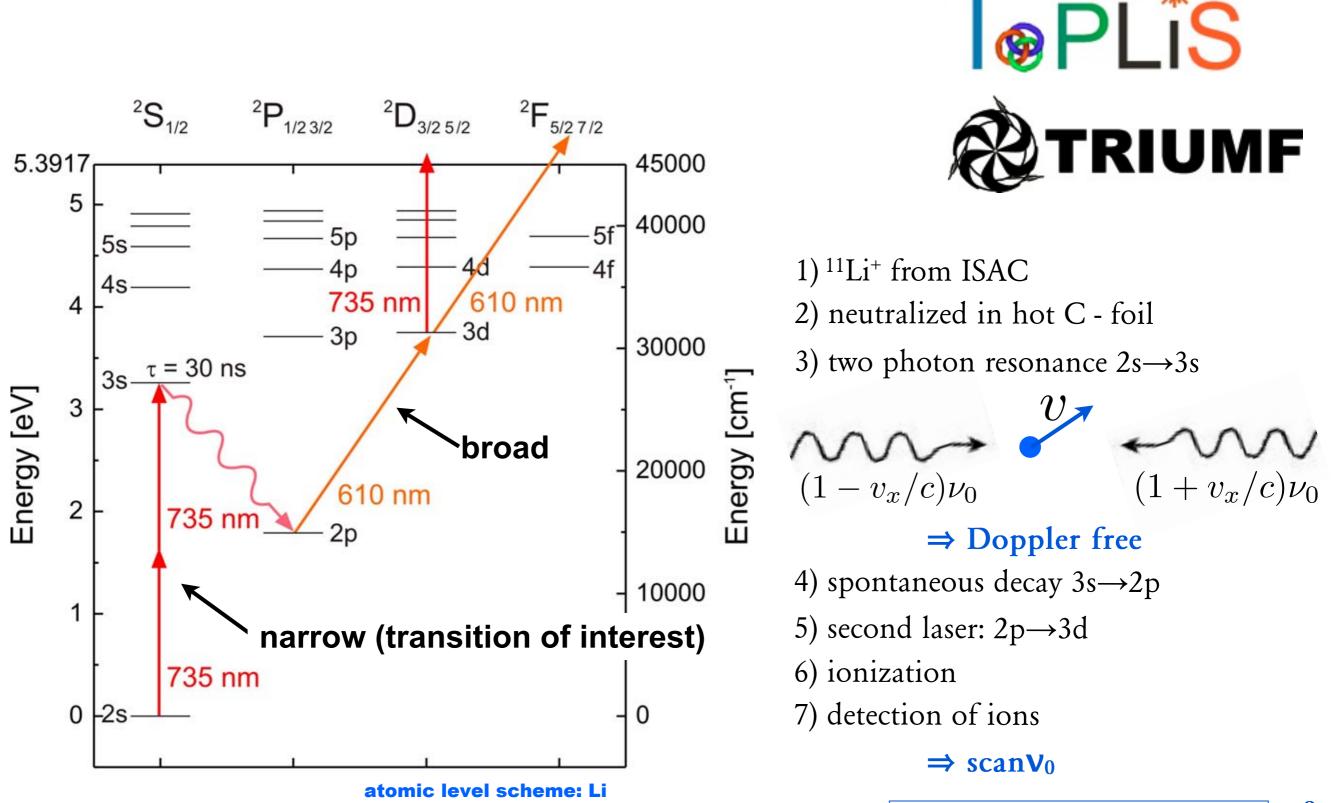












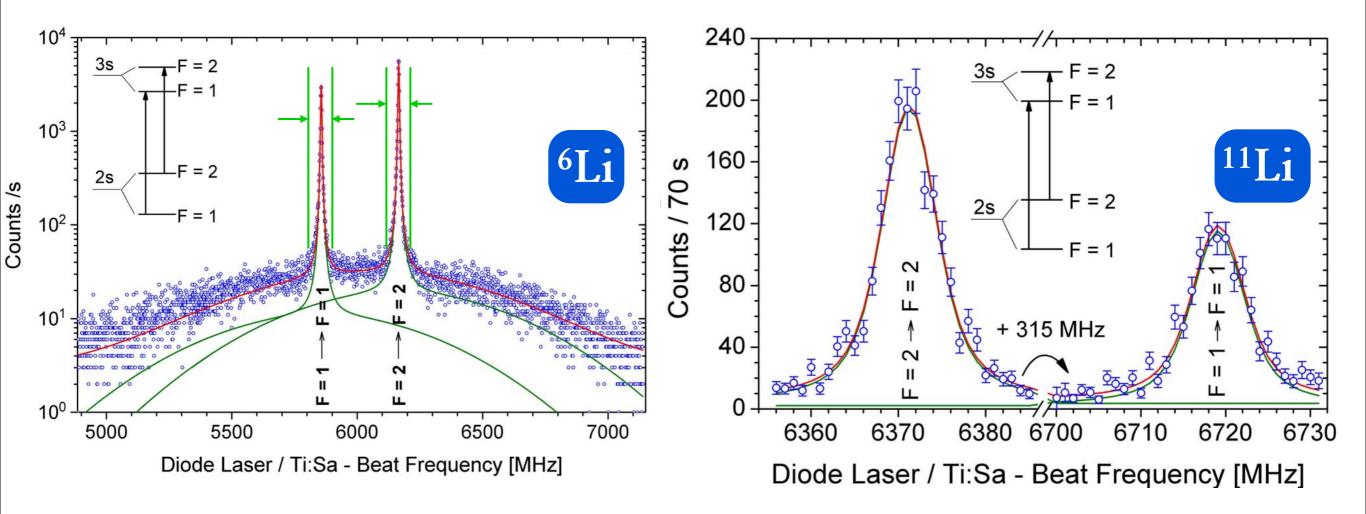
R. Sanchez et al., PRL 96, 033002 (2006)

9











### isotope shifts <sup>7</sup>Li-<sup>A</sup>Li:

- 2s→3s
- reference  $r_c(^7Li) = 2.39(3)$  fm

#### At. Data Nucl. Data Tables 14, 479 (1974)

Isc	otope	Isotope Shift, kHz
<sup>6</sup> Li	TRIUMF	-11 453 984(20)
	GSI	-11453950(130)
	avg	-11 453 983(20)
<sup>8</sup> Li	TRIUMF	8635781(46)
	GSI	8635790(150)
	avg	8635782(44)
<sup>9</sup> Li	TRIUMF	15 333 279(40)
	GSI	15 333 140(180)
	avg	15 333 272(39)
<sup>11</sup> Li	TRIUMF	25 101 226(125) <sup>a</sup>

R. Sanchez et al., PRL 96, 033002 (2006)



 $\delta \nu_{A,A'} = \delta^{\mathrm{MS}}_{A,A'} + K_{\mathrm{FS}} \delta < r_c^2 >_{A,A'}$ 

#### mass shifts

Isotopes	$2^{2}P_{1/2} - 2^{2}S$	$2^{2}P_{3/2} - 2^{2}S$	$3^{2}S - 2^{2}S$
<sup>7</sup> Li - <sup>6</sup> Li	-10532.111(6)	-10532.506(6)	-11452.821(2)
$^{7}$ Li $ ^{8}$ Li	7940.627(5)	7940.925(5)	8634.989(2)
<sup>7</sup> Li – <sup>9</sup> Li	14098.840(14)	14 099.369(14)	15331.799(13)
$^{7}$ Li $- {}^{11}$ Li <sup>a</sup>	23 082.642(24)	23 083.493(24)	25 101.470(22)
$^{9}$ Be $ ^{7}$ Be	-49 225.765(19)	-49 231.814(19)	-48514.03(2)
${}^{9}\text{Be} - {}^{10}\text{Be}$	17 310.44(6)	17 312.57(6)	17060.56(6)
$^{9}$ Be - $^{11}$ Be	31 560.01(6)	31 563.89(6)	31 104.60(6)

Z.-C. Yan et al., PRL 100, 243002 (2008)

*M. Puchalski et al., PRL 97,133001 (2006)* 

 $r_{c}$  (<sup>11</sup>Li) = 2.423(17)(30) fm

reference r<sub>c</sub>



### isotope shifts 7Li-ALi:

- 2s→3s
- reference  $r_c(^7Li) = 2.39(3)$  fm

#### At. Data Nucl. Data Tables 14, 479 (1974)

**r**<sub>c</sub> (

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_	Isotopes	$2^{2}P_{1/2} - 2^{2}S$	$2^2 P_{3/2} - 2^2 S$	$3^{2}S - 2^{2}S$
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(	$^{9}Be - {}^{10}Be$	mass: MI	STRAL (2005)	17 060.56(6)
_	$^{9}$ Be $ ^{11}$ Be	31 560.01(6)	31 563.89(6)	31 104.60(6)
_		!r	need mass !	243002 (2008)
			M. Pucha ki et al., PRL s	97,133001 (2006)
Li) = 2.423	3(17)(30)	fm	mass: AME	<b>'03</b>
reference	e r <sub>c</sub>	r <sub>c</sub>	$(^{11}\text{Li}) = 2.465(19)$	9)(30) fm
Li) = 2.423 reference		fm	M. Puchatki et al., PRL s mass: AME	97,133001 (200 '03



TITAN ISAC-TRIUMP

### TITAN

Cooler

Trap

Penning

RFQ

Buncher & Cooler

**ISAC Beamline** 

### masses of halos:

• reflect binding energy

ISAC beam: A<sup>+</sup>

- separation energy: S<sub>n</sub>, Sp
- input to extract physical quantities from exp. (e.g.  $r_c$ ) Measurement Penning trap



EBIT

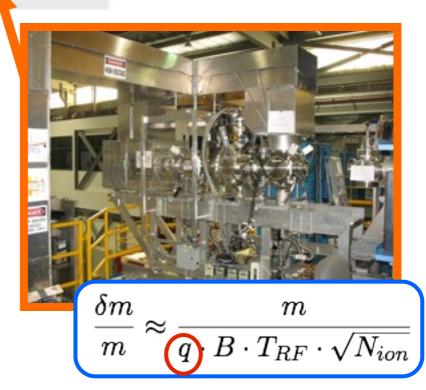


### Penning traps:

- highest precision
- previously shortest <sup>74</sup>Rb with T<sub>1/2</sub>=65 ms ISOLTRAP @ CERN

A. Kellerbauer et al., PRL 93, 072502 (2004)

```
• but <sup>11</sup>Li T_{1/2} = 8.8 ms
```





### **Measurement Principle**

В

confinement:

- strong axial, hom. B-field (3.7 T)
- electrostatic quadrupolar field
- 3 eigenmotions

 $v_{+} >> v_{z} >> v_{-}$ 

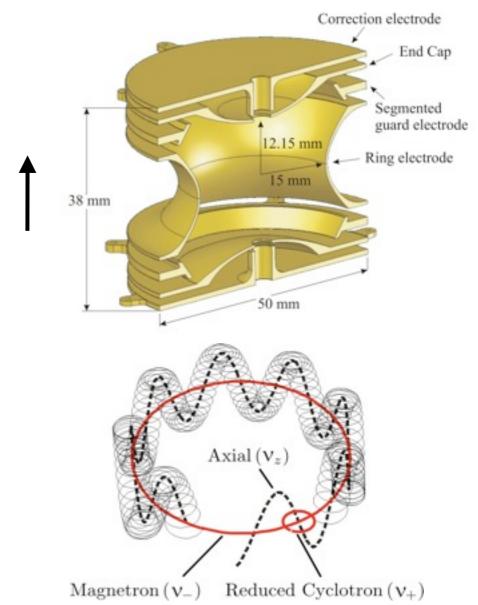
cyclotron frequency

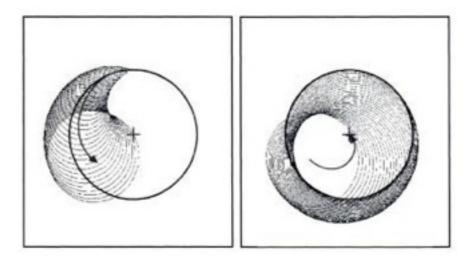
$$\mathbf{v}_c = \mathbf{v}_+ + \mathbf{v}_- = \frac{1}{2\pi} \frac{q}{m} B$$

quadrupolor rf- field (ring electrode) leads to conversion:

magnetron ↔ reduced cyclotron radial energy:

$$E_r(t) \propto \omega_+^2 \rho_+(t)^2 + \omega_-^2 \rho_-(t)^2 \approx \omega_+^2 \rho_+(t)^2$$





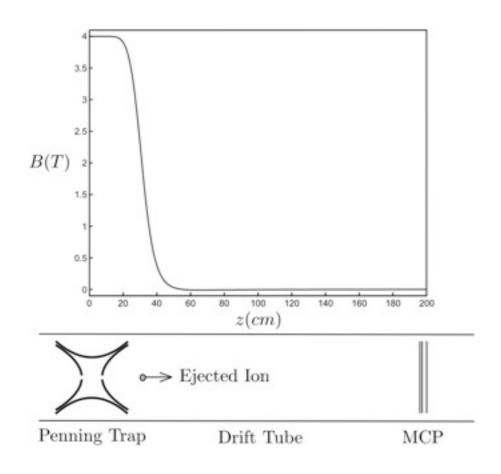


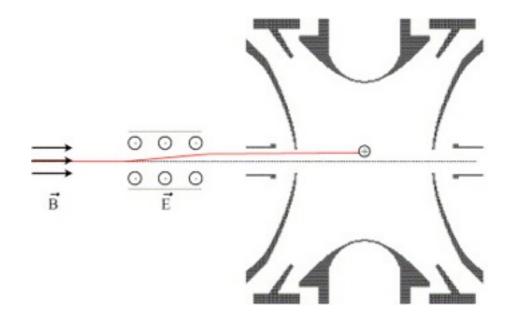
- initial magnetron preparation
  - dipolar RF excitation ~ 10 ms
  - Lorentz steerer
- · quadrupolor rf- field
- · extraction: through B-field  $E_r$  to  $E_l$
- $\cdot$  E<sub>1</sub> measured by TOF
- · minimum at  $v_c$

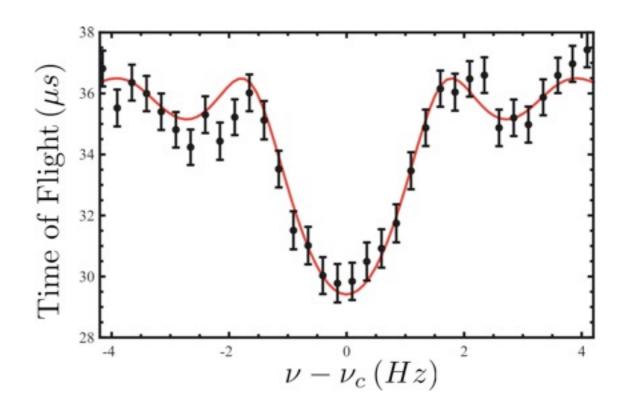
TRIUMI

•

· comparison to well known isotope







# Precise & Accurate

line width (FWHM):

ISAC

 $\Delta\nu\approx 1/T_{rf}$ 

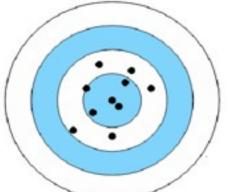
 $\Rightarrow$  resolution:

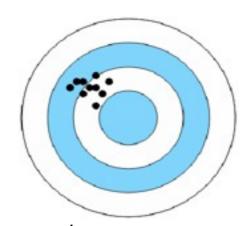
TRIUMF

$$R = \frac{m}{\Delta m} = \frac{\nu_c}{\Delta \nu_c} \approx \nu_c T_{rf}$$
$$\approx \frac{q B T_{rf}}{2\pi m}$$

 $\Rightarrow$  even for  $T_{rf} \sim 10ms$ 

$$(\delta m/m)_{\rm stat} < 10^{-7}$$





accurate, but not precise

precise, but not accurate

• exact theoretical description

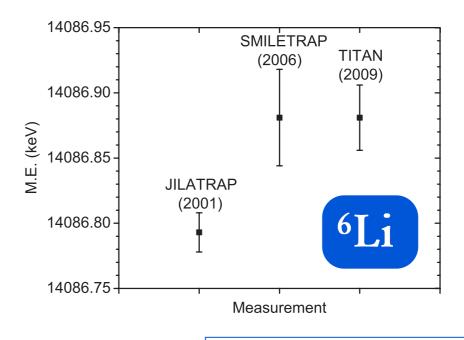
L.S. Brown and G. Gabrielse, Rev. Mod. Phys. 58, 233 (1986) G. Bollen et al., J. Appl. Phys. 88, 4355 (1990) M. König et al., Int. J. Mass Spect. 142, 95 (1995) M. Kretzschmarr, Int. J. Mass Spect. 246, 122 (2007)

• even for non-ideal traps

G. Bollen et al., J. Appl. Phys. 88, 4355 (1990) G. Gabrielse, PRL 102, 172501 (2009)

- off-line tests with stables
- $\Rightarrow$  control over systematics

<u>for TITAN:</u> < 5 ppb possible

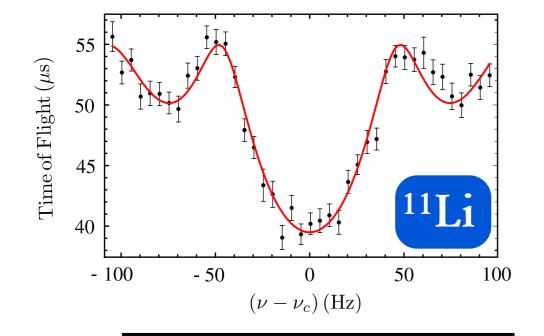


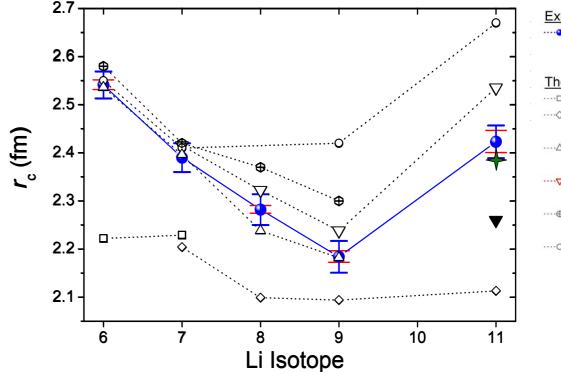
M. Brodeur et al, PRC 80, 044318 (2009) 15

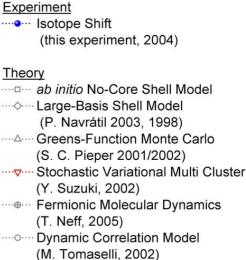


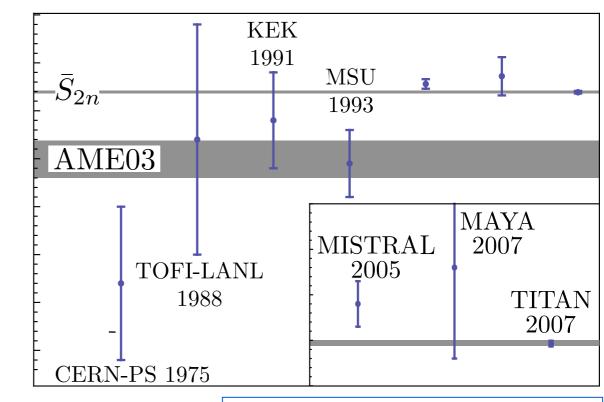
### Mass of the

 $S_{2n}$  (keV)









 $r_c$  (<sup>11</sup>Li) = 2.427(16)(30) fm

eliminates mass as source of uncertainty!

Mass [u]

11.043 798(21)

11.043 715 7(54)

11.043 723 61 (69)

#### two neutron separation energy:

Reference

MISTRAL 2005

**TITAN 2007** 

**AME'03** 

 $S_{2n} = -M(A,Z) + M(A-2,Z) + 2n$ 

- asymptotic waveform for Borromean system
- soft electric-dipole excitation

T. Nakamura et al., PRL 96, 252502 (2006)

• models of <sup>11</sup>Li: adjust <sup>9</sup>Li-n interaction



### Other Halos: Laser Spectroscopy

### <u><sup>6</sup>He and <sup>8</sup>He</u>

- Argonne Lab / GANIL
- LS in MOT

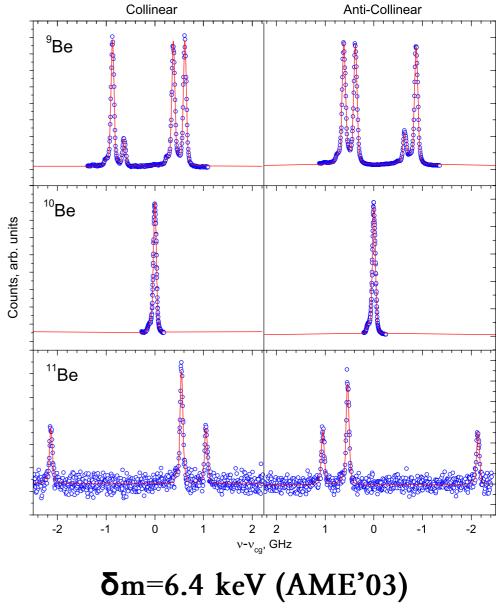
		all in I	MHz
<sup>6</sup> He		<sup>8</sup> He	
Value	Error	Value	Error
	0.008		0.032
	0.002		0.012
	0.002		0.024
			0.015
	0.030		0.045
	0.015		0.074
7			
0.110	0.000	0.165	0.000
-0.014	0.003	-0.002	0.001
-1.478	0.035	-0.918	0.097
	Value 0.110 -0.014	Value         Error           0.008         0.002           0.002         0.002           0.002         0.002           0.030         0.015           0.110         0.000           -0.014         0.003	<sup>6</sup> He <sup>8</sup> H Value Error Value 0.008 0.002 0.002 0.002 0.002 0.015 0.110 0.000 0.165 −0.014 0.003 −0.002

### mass: dominating uncertainty

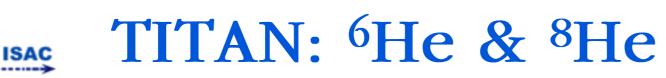
#### \_\_\_\_\_

<u>811Be:</u>

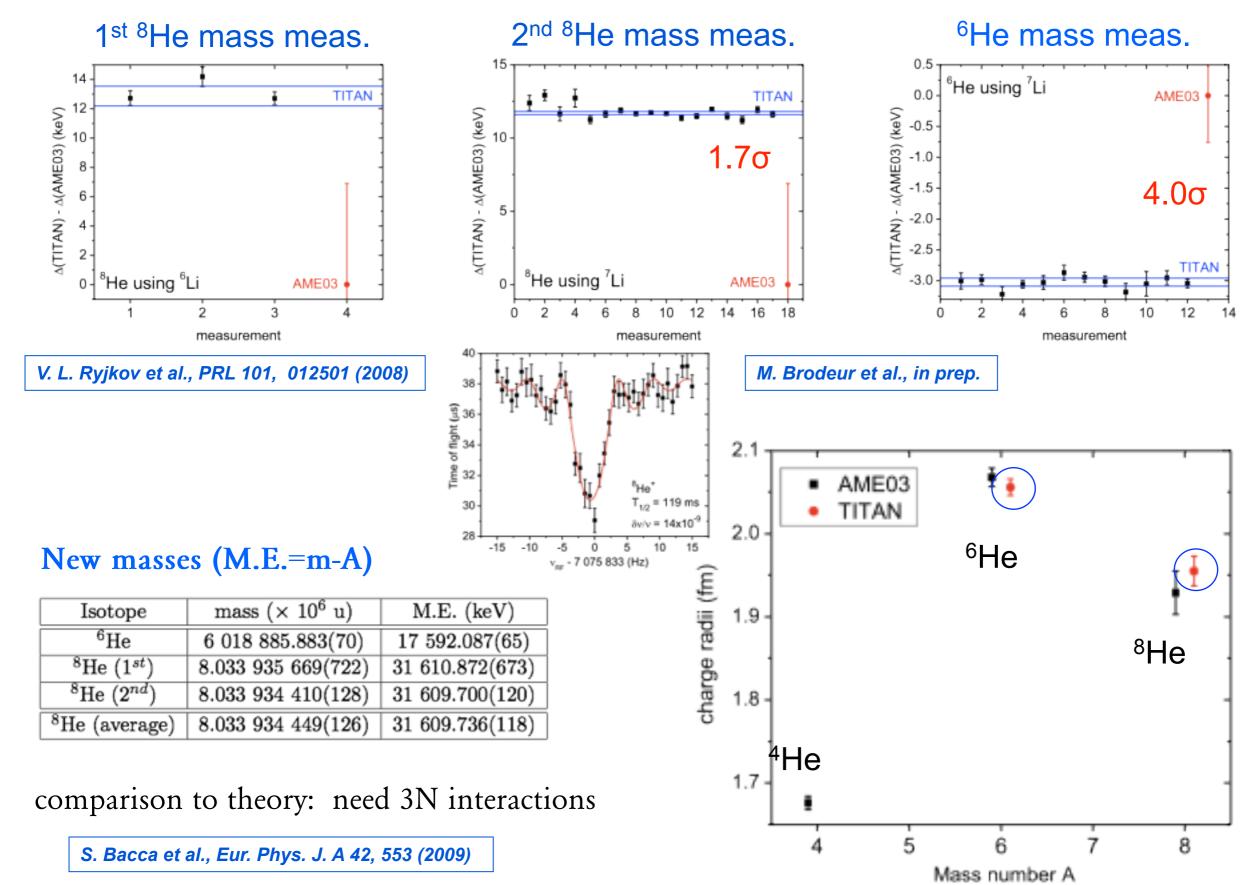
- GSI
- collinear LS

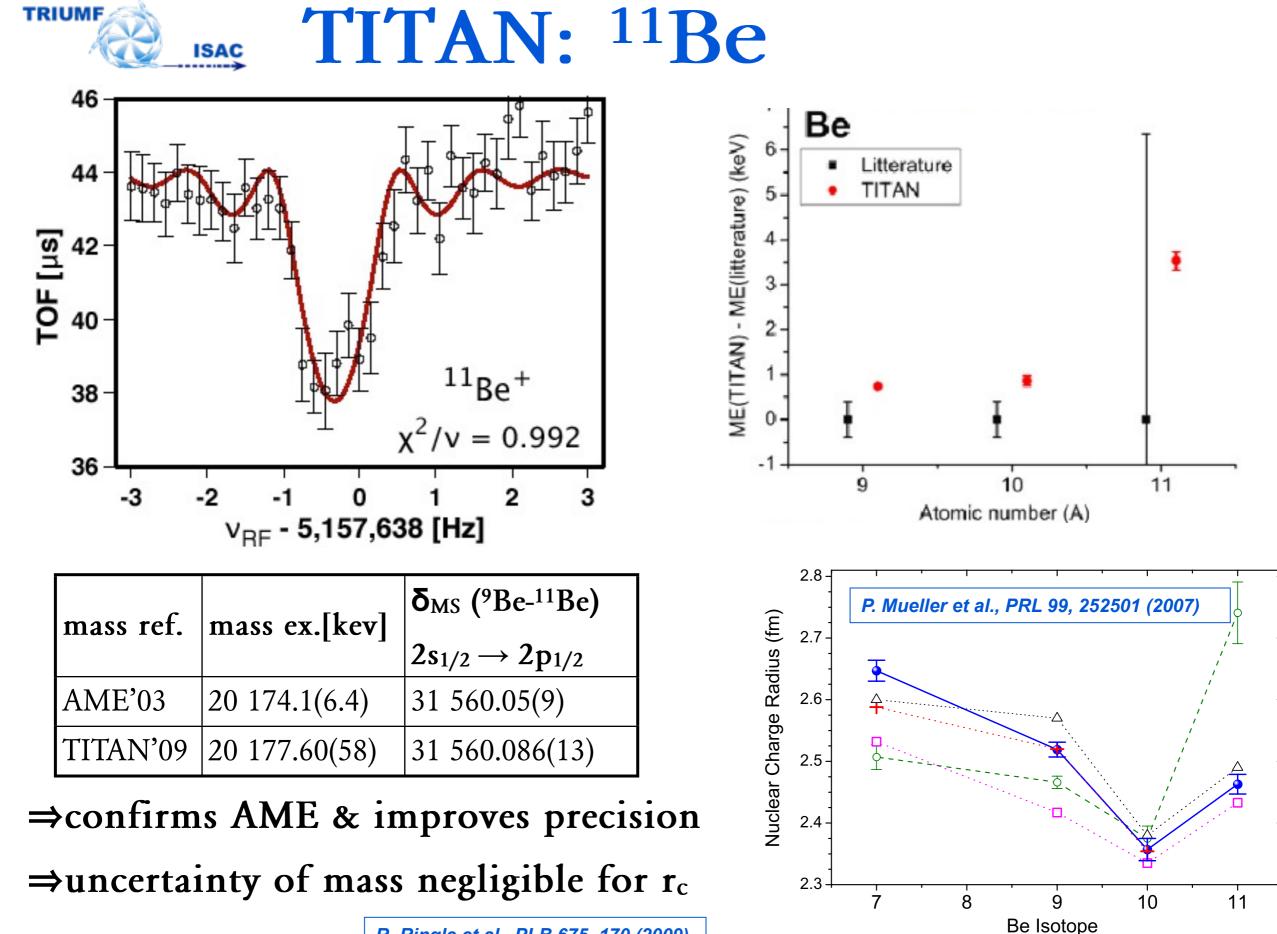


W. Nörtershäuseret al., PRL 102, 062503 (2009)



TRIUMF



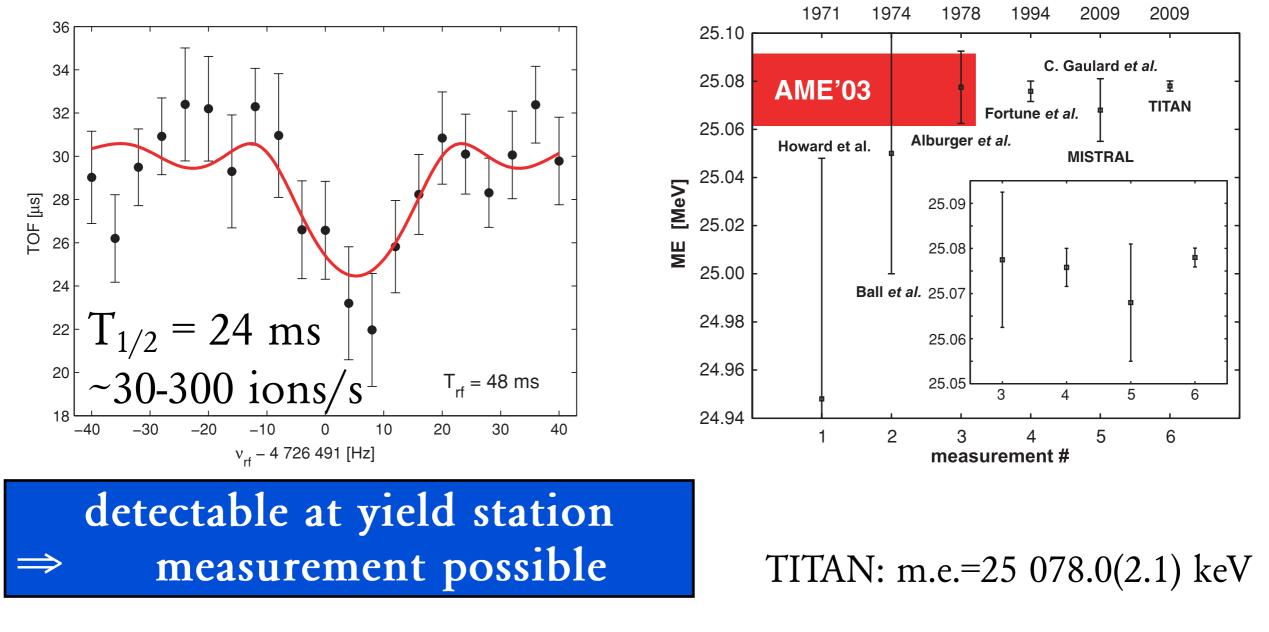


*R. Ringle et al., PLB 675, 170 (2009)* 



### calculation & measurement of $r_c$ in the near future

 $\rightarrow$  see talk of Thomas Neff





### Conclusions

- Interplay of various experimental approaches allow to identify & probe nuclear halos
- Combination of high precision
  - laser spectroscopy
  - mass measurements

• ⇒ charge radius

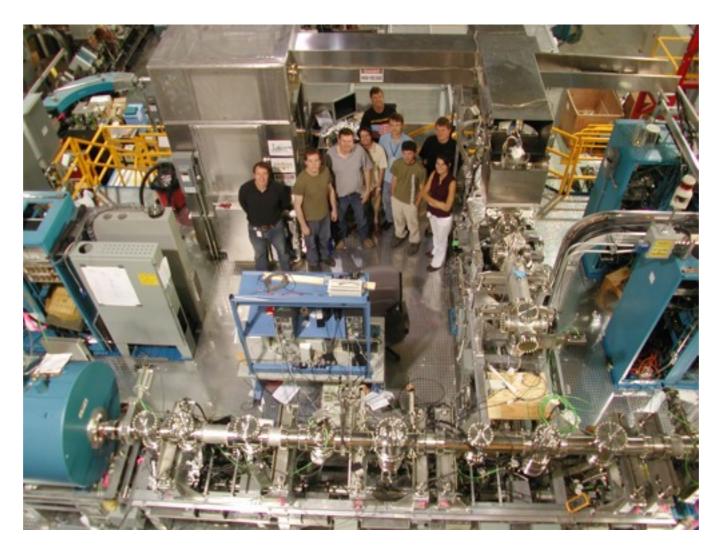
- atomic physics calculation
- benchmark theoretical models (mass, matter/charge radius, ..)

# Outlook (TRIUMF)

- later this year: electric quadrupole moment of <sup>11</sup>Li
- TITAN: masses
  - to investigate established halos  ${}^{14}Be(2n)$ ,  ${}^{19}C(1n)$ ,  ${}^{17}Ne(1p)$
  - needed to decide if halo structure in <sup>22</sup>C and <sup>31</sup>Ne



### **TITAN collaboration**



M. Brodeur, T. Brunner, S. Ettenauer, A. Gallant, V. Simon, M. Smith,A. Lapierre, R. Ringle, V. Ryjkov, M. Simon,M. Good, P. Delheij, D. Lunney, and J. Dilling for the TITAN collaboration





# Backup Slides

### <sup>11</sup>Be: Comparison to Models

TRIUMI

ISAC

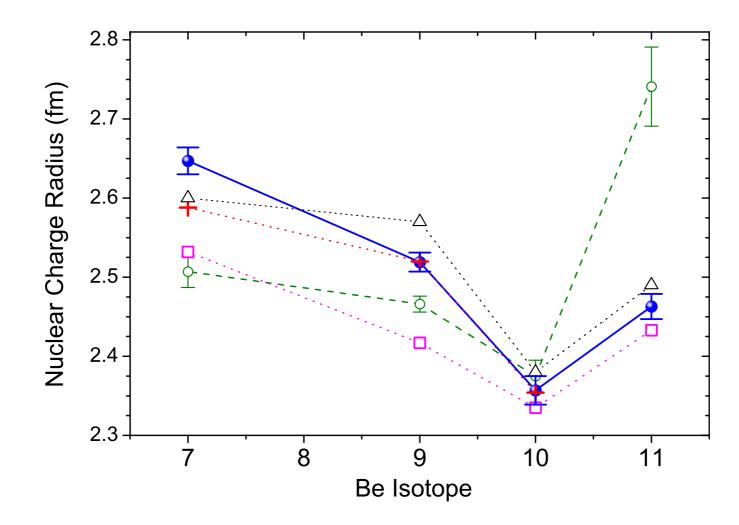


FIG. 3 (color online). Experimental charge radii of beryllium isotopes from isotope-shift measurements ( $\bullet$ ) compared with values from interaction cross-section measurements ( $\bigcirc$ ) and theoretical predictions: Greens-function Monte Carlo calculations (+) [2,24], fermionic molecular dynamics ( $\triangle$ ) [25], *ab initio* no-core shell model ( $\Box$ ) [13,26,27].