

# Transport properties of a non-relativistic dilute Yukawa liquid

Postnikov Sergey

Ohio University

supervised by Madappa Prakash

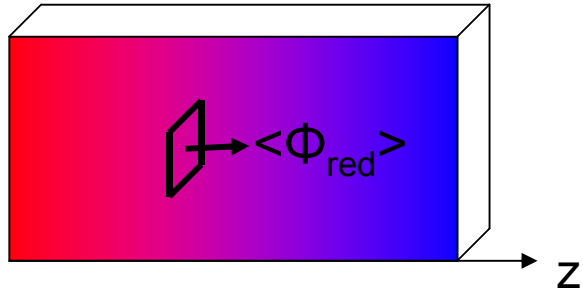
“Neutron star crust” workshop, INT, July 2007

# Content

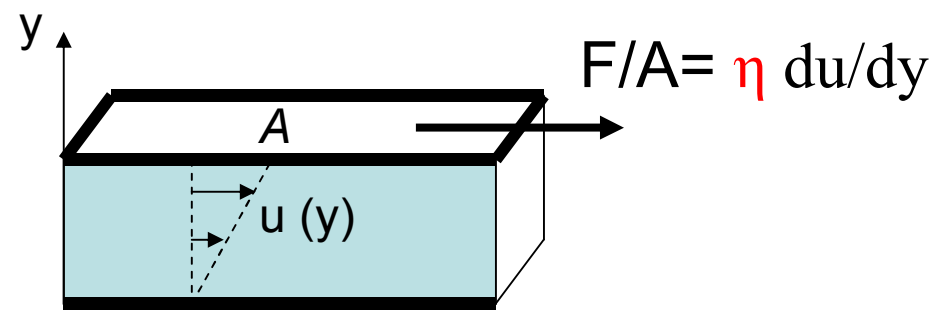
- Introduction to transport phenomena
- Boltzmann equation and the equation of change
- Enskog solution of Boltzmann equation for a dilute system close to equilibrium
- Screened Coulomb potential: preliminary results for Yukawa liquid of  $C^{12}$
- Relevance to Neutron Star physics and computer simulations (Horowitz & Caballero)

# Diffusivity

$$\langle \Phi_{\text{red}} \rangle = D \, dn_{\text{red}} / dz$$

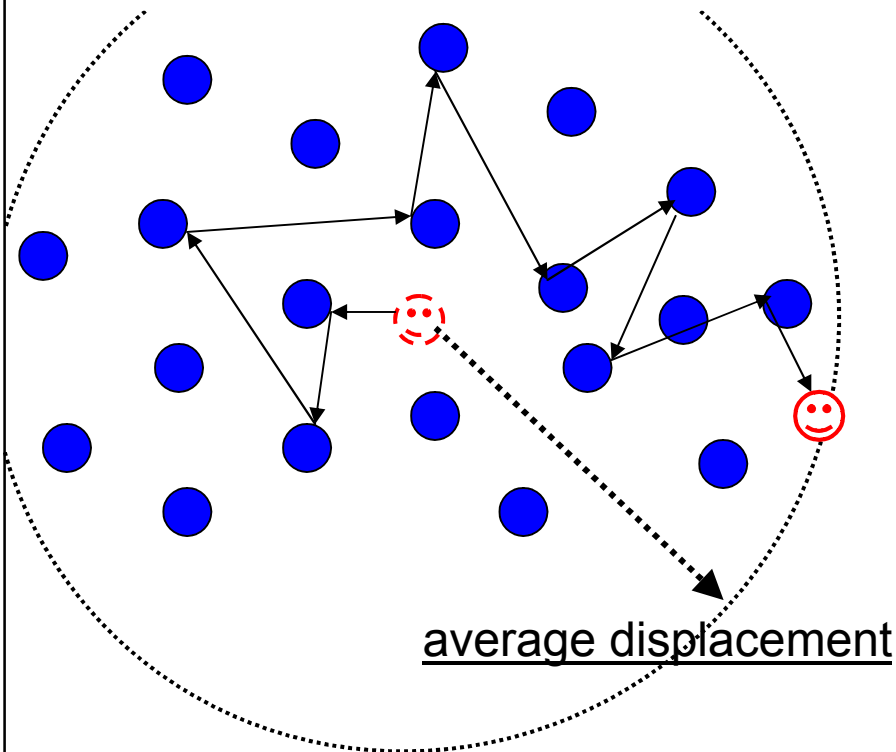


# Shear viscosity

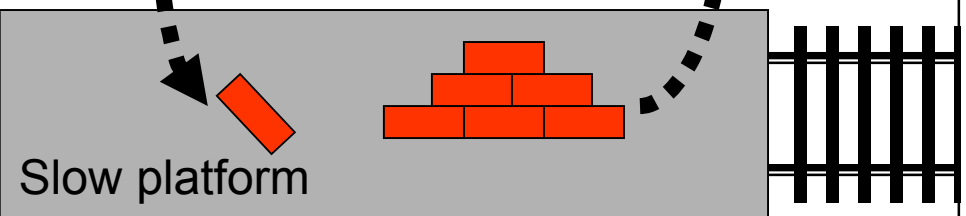
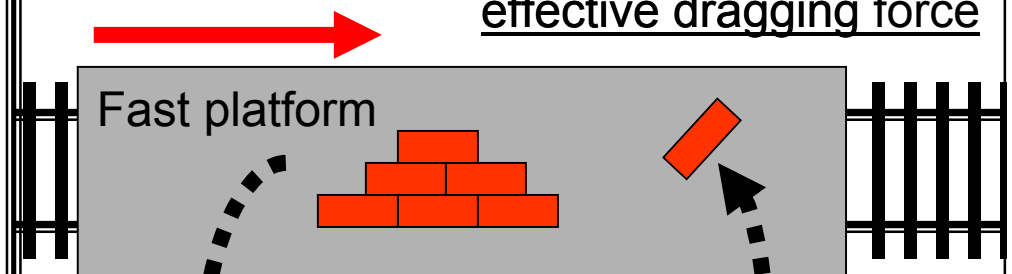


Classic Newtonian laminar flow

## Random walk of drunk sailor



## effective dragging force

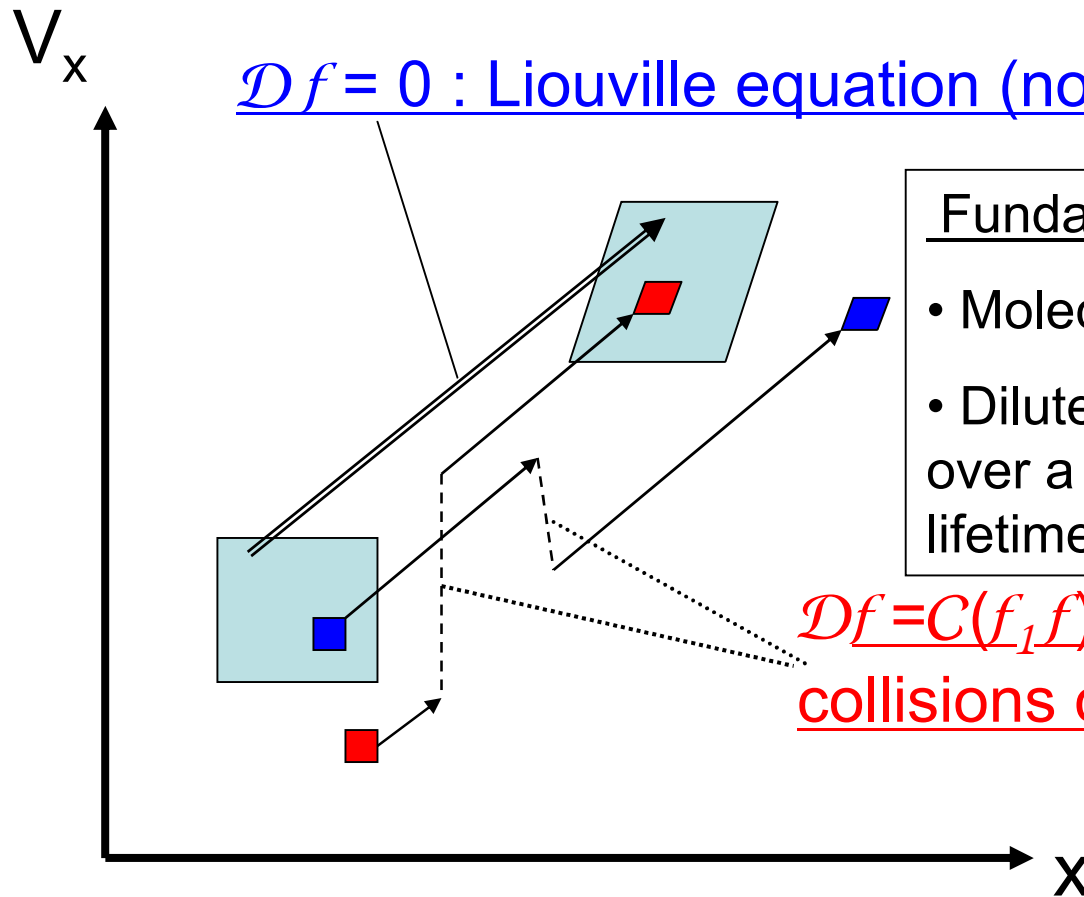


## effective pushing force

# Boltzman equation for the particle distribution function: $f$ (phase space)

*Equations of motion  $\rightarrow$  time evolution  $\rightarrow$  operator  $\mathcal{D}$*

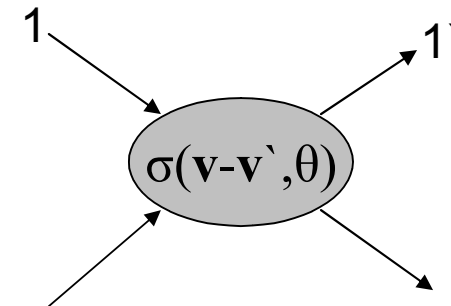
$\mathcal{D}f = 0$  : Liouville equation (no collisions)



Fundamental assumptions:

- Molecular chaos – no correlations
- Dilute system – encounters occur over a small fraction of molecular lifetime

$\mathcal{D}f = C(f_1 f)$ : Boltzman equation (binary collisions only)



# Equation of change for property $\phi$

$$\int \phi \mathcal{D}f \, dv = n \langle \Delta\phi \rangle$$

$$\langle \Delta\phi \rangle = n^{-1} \int \phi C(f_1 f) \, dv$$

**Table 14.2** Summary of phenomenological transport laws

Effect	Flux of particle property $\phi$	Gradient	Coefficient	Law	Name of law	Approximate expression for coefficient
Diffusion	Number	$\frac{dn}{dz}$	Diffusivity $D$	$\mathbf{J}_n = -D \text{grad } n$	Fick's law	$D = \frac{1}{3}\bar{c}l$
Viscosity	Transverse momentum	$M \frac{dv_x}{dz}$	Viscosity $\eta$	$\frac{F_x}{A} = J_p^x = -\eta \frac{dv_x}{dz}$	Newtonian viscosity	$\eta = \frac{1}{3}\rho\bar{c}l$
Thermal conductivity	Energy	$\frac{d\rho_u}{dz} = \hat{C}_V \frac{dT}{dz}$	Thermal conductivity $K$	$\mathbf{J}_u = -K \text{grad } \tau$	Fourier's law	$K = \frac{1}{3}\hat{C}_V\bar{c}l$
Electrical conductivity	Charge	$-\frac{d\phi}{dz} = E_z$	Conductivity $\sigma$	$\mathbf{J}_q = \sigma \mathbf{E}$	Ohm's law	$\sigma = \frac{nq^2l}{M\bar{c}}$

SYMBOLS:  $n$  = number of particles per unit volume  
 $\bar{c}$  = mean thermal speed =  $\langle |v| \rangle$   
 $l$  = mean free path  
 $\hat{C}_V$  = heat capacity per unit volume  
 $\rho_u$  = thermal energy per unit volume  
 $F_x/A$  = shear force per unit area

$\phi$  = electrostatic potential  
 $\mathbf{E}$  = electric field intensity  
 $q$  = electric charge  
 $M$  = mass of particle  
 $\rho$  = mass per unit volume  
 $\mathbf{p}$  = momentum

*"Thermal Physics" Ch. Kittel & H. Kroemer*

# Enskog approximate solution of the Boltzmann equation

System assumed to be only slightly disturbed from the Maxwell equilibrium state  $f(0)$ :

$$f = f(0) + f(1) + f(2) + \dots$$

Boltzmann equation:  $F[f] = 0$

$$F[f] = F(0)[f(0)] + F(1)[f(0), f(1)] + F(2)[f(0), f(1), f(2)] + \dots$$

$$F(0)[f(0)] = 0 \rightarrow \text{Maxwell distribution}$$

$$F(1)[f(0), f(1)] = 0 \rightarrow \text{first approximation}$$

$$F(2)[f(0), f(1), f(2)] = 0 \rightarrow \text{second approximation}$$

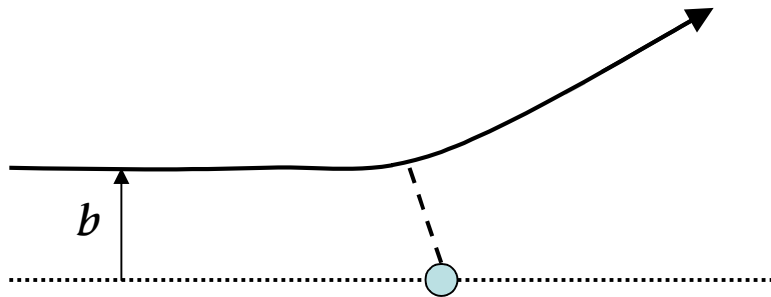
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# Classical vs quantum transport cross-section

The *transport cross-section of order n* is given by the integral

$$\phi^{(n)} = 2\pi \int_{-1}^{+1} d \cos \theta (1 - \cos^n \theta) \frac{d\sigma(k, \theta)}{d\Omega} \Big|_{c.m.},$$

## Classical case



$$\sigma(k, \theta) = \pi b^2$$

## Quantum case

the orthogonality of the Legendre polynomials  $P_l$

$$q^{(1)} \equiv \frac{\phi^{(1)}}{4\pi a^2} = \frac{2}{x^2} \sum_l' (2l + 1) \sin^2(\delta_l),$$

$$q^{(2)} \equiv \frac{\phi^{(2)}}{4\pi a^2} = \frac{2}{x^2} \sum_l' \frac{(l + 1)(l + 2)}{(2l + 3)} \sin^2(\delta_{l+2} - \delta_l),$$

- $a$  – characteristic length for a potential
- $x = ka$  – dimensionless momentum variable
- $l$  - quantum number for angular momentum
- $\delta_l$  – partial wave phase shifts

# Statistics and Transport Integrals

$$q_{(s)}^{(n)} = \frac{s+1}{2s+1} q_{Bose}^{(n)} + \frac{s}{2s+1} q_{Fermi}^{(n)}, \quad \text{for integer } s,$$

$$q_{(s)}^{(n)} = \frac{s+1}{2s+1} q_{Fermi}^{(n)} + \frac{s}{2s+1} q_{Bose}^{(n)}, \quad \text{for half-integer } s.$$

$$\omega^{(n,t)}(T) \equiv \int_0^\infty d\gamma e^{-\gamma^2} \gamma^{2t+3} q^{(n)}(x),$$

$$\gamma = \frac{\hbar k}{\sqrt{2\mu k_B T}} = \frac{x}{\sqrt{2\pi}} \left( \frac{\lambda(T)}{a} \right).$$

$\lambda(T)$  – thermal De-Broglie wavelength



# Diffusion and shear viscosity

$$\tilde{\mathcal{D}} = \frac{3h}{32\sqrt{2}\pi} \frac{1}{ma^3n},$$

$$\frac{[\mathcal{D}]_1}{\tilde{\mathcal{D}}} = \left( \frac{a}{\lambda(T)} \right) \frac{1}{\omega^{(1,1)}(T)},$$

$$\frac{[\mathcal{D}]_2}{[\mathcal{D}]_1} = 1 + \frac{(5\omega^{(1,1)}(T) - 2\omega^{(1,2)}(T))^2}{\omega^{(1,1)}(T) (30\omega^{(1,1)}(T) + 4\omega^{(1,3)}(T) + 8\omega^{(2,2)}(T)) - 4(\omega^{(1,2)}(T))^2},$$

$$\tilde{\eta} = \frac{5h}{32\sqrt{2}\pi} \frac{1}{a^3},$$

$$\frac{[\eta]_1}{\tilde{\eta}} = \left( \frac{a}{\lambda(T)} \right) \frac{1}{\omega^{(2,2)}(T)},$$

$$\frac{[\eta]_2}{[\eta]_1} = 1 + \frac{3(7\omega^{(2,2)}(T) - 2\omega^{(2,3)}(T))^2}{2 \left( \omega^{(2,2)}(T) (77\omega^{(2,2)}(T) + 6\omega^{(2,4)}(T)) - 6(\omega^{(2,3)}(T))^2 \right)},$$

# Thermal conductivity

$$[\lambda_T]_1 = \frac{5}{2} [\eta]_1 \left( \frac{c_v(n, T)}{m} \right),$$

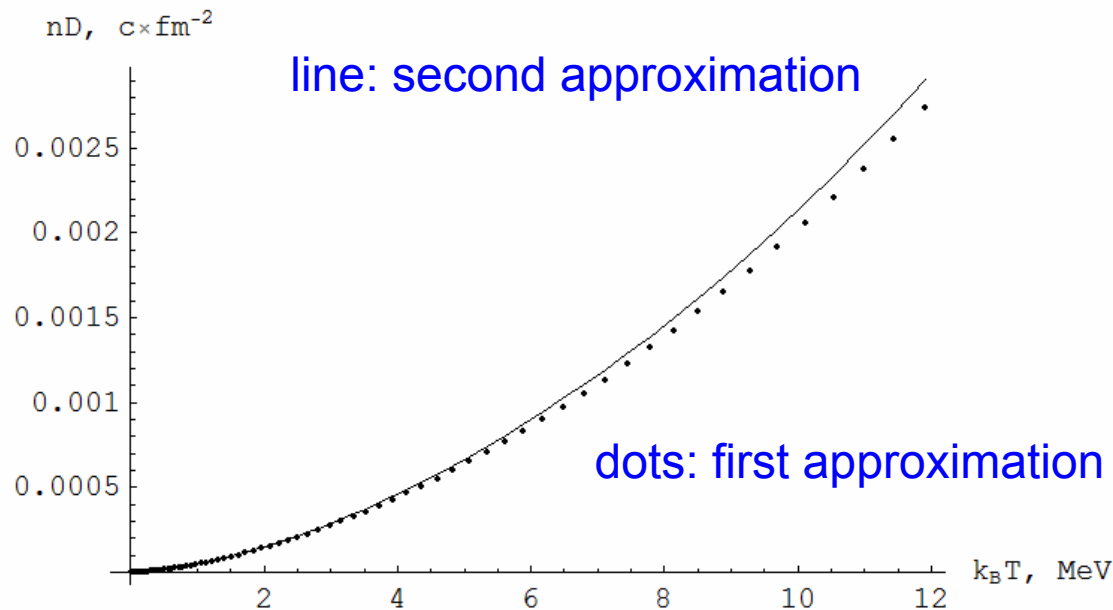
$$\tilde{\lambda}_T = \frac{25h}{64\sqrt{2}\pi} \frac{\tilde{c}_v(n)}{a^3 m},$$

$$\frac{[\lambda_T]_1}{\tilde{\lambda}_T} = \frac{c_v(n, T)}{\tilde{c}_v(n)} \left( \frac{a}{\lambda(T)} \right) \frac{1}{\omega^{(2,2)}(T)},$$

$$\frac{[\lambda_T]_2}{[\lambda_T]_1} = 1 + \frac{(7\omega^{(2,2)}(T) - 2\omega^{(2,3)}(T))^2}{4 \left( \omega^{(2,2)}(T) (7\omega^{(2,2)}(T) + \omega^{(2,4)}(T)) - (\omega^{(2,3)}(T))^2 \right)},$$

In order to calculate thermal conductivity one needs to invoke the virial expansion or other means to find specific heat capacity  $c_v$

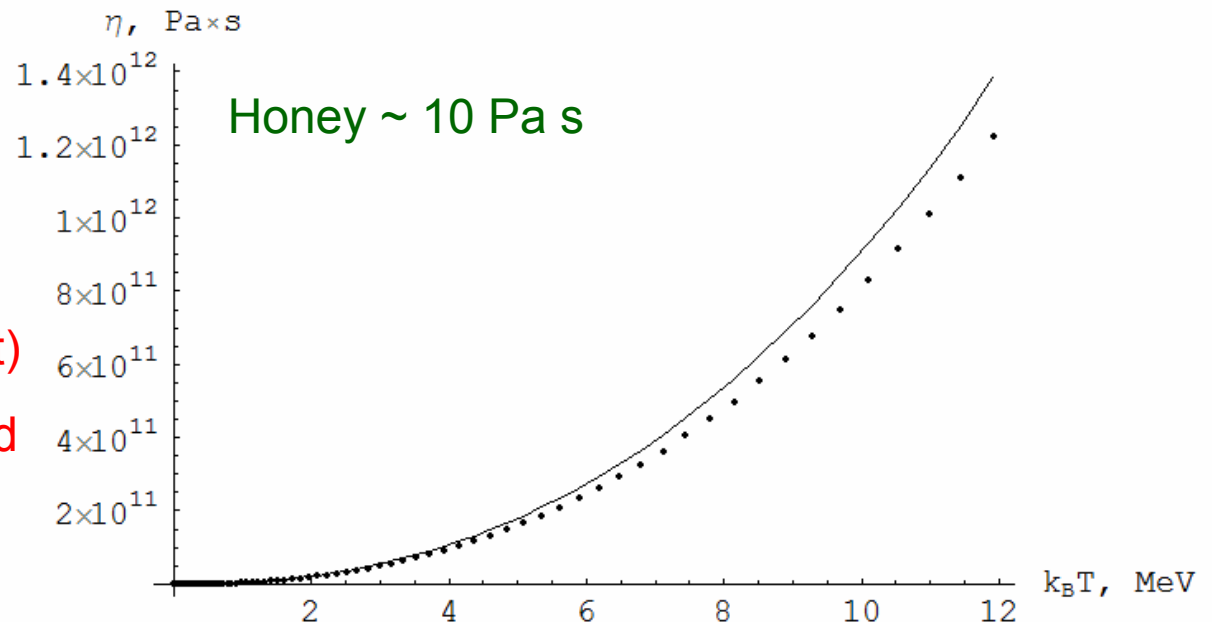
# Results for $C^{12}$ with $V(R) \rightarrow \exp[-R/a]/R$

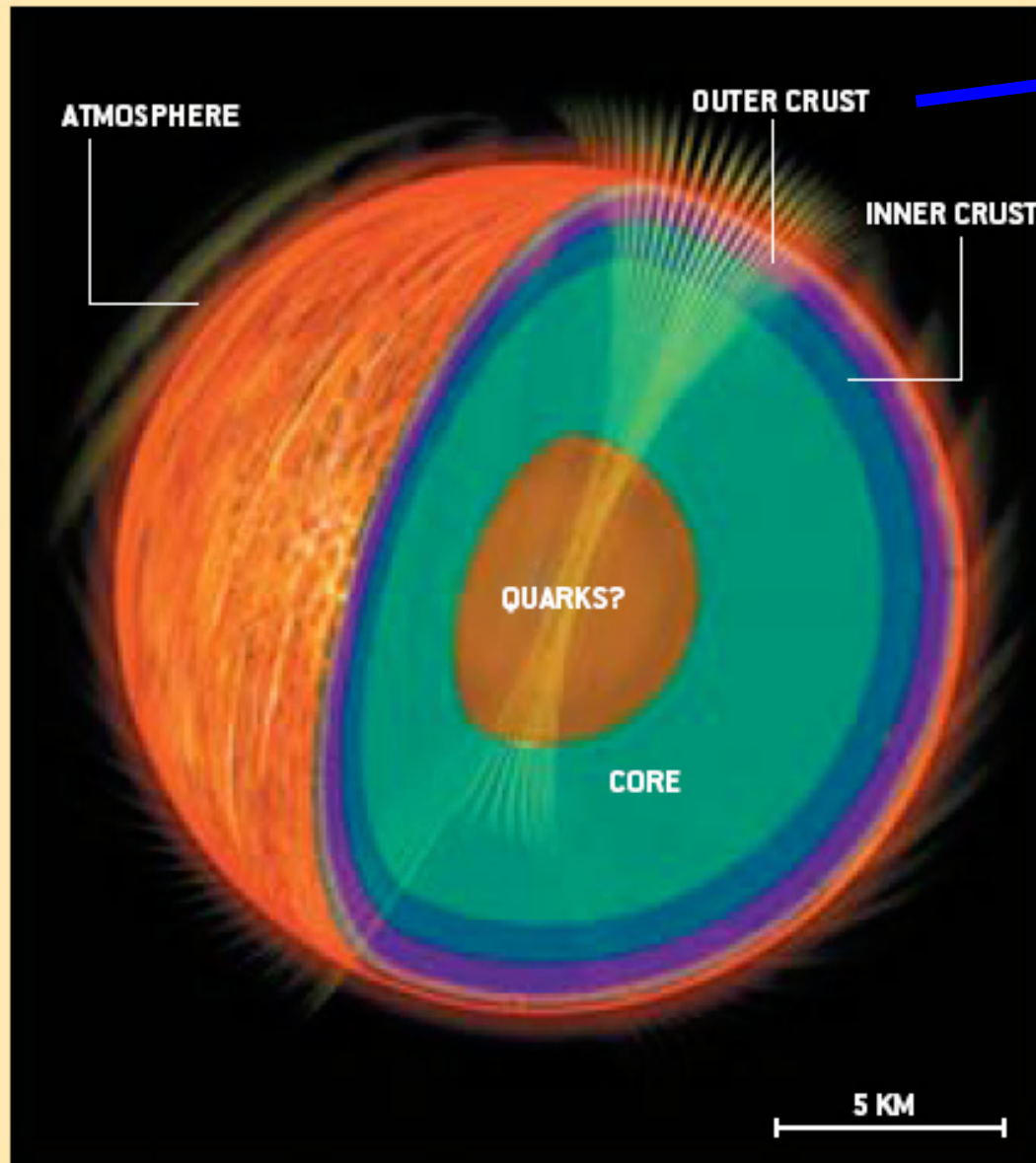


- “Low” temperatures
- Dilution parameter:  $n a^3 < 1$
- Density independence
- Large number of partial waves

## Ongoing work:

- Compare classical with quantum results
- Mixtures (multiple component)
- Study effects of magnetic field
- + the virial gives  $\eta/s$  (T)





STRUCTURE OF A NEUTRON STAR can be inferred from theories of nuclear matter. Starquakes can occur in the crust, a lattice of atomic nuclei and electrons. The core consists mainly of neutrons and perhaps quarks. An atmosphere of hot plasma might extend a grand total of a few centimeters.

- Diffusion:
  - element segregation
- Viscosity:
  - rotation
  - seismic modes
- Thermal conductivity
  - cooling
  - temperature map

Classical molecular-dynamical simulations (Horowitz & Caballero) are to be compared with the Enskog approach for a dilute system with quantum and classical collision cross-sections, allowing a check of the low density regime of the system.

...to be continued

*köszönöm ! תודה dĕkuji*

*mahalo* 고맙습니다

*thank you*

*merci* 谢谢 *danke*

Спасибо

Ευχαριστώ شڪرا

どうもありがとう *gracias*