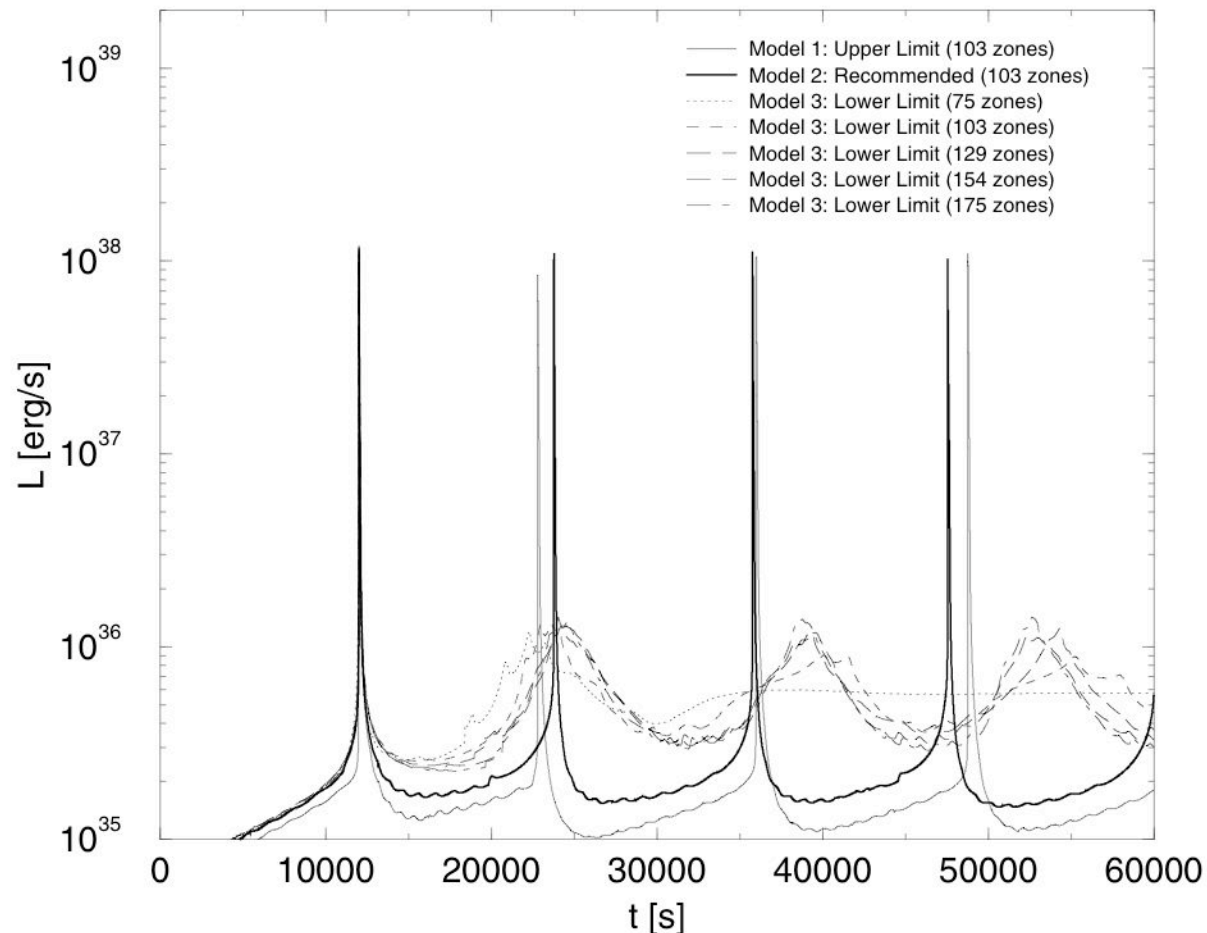


Experimental Efforts to Determine the $^{15}\text{O}(\alpha,\gamma)^{19}\text{Ne}$ Reaction Rate

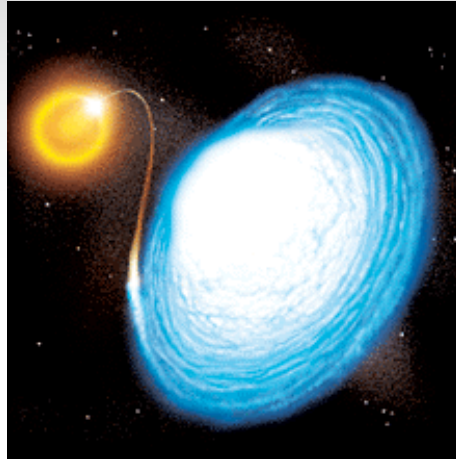
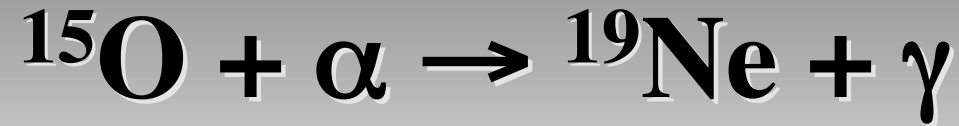
Barry Davids, TRIUMF

INT, 12 July 2007

Type I X-Ray Bursts: Dependence on $^{15}\text{O}(\alpha,\gamma)^{19}\text{Ne}$

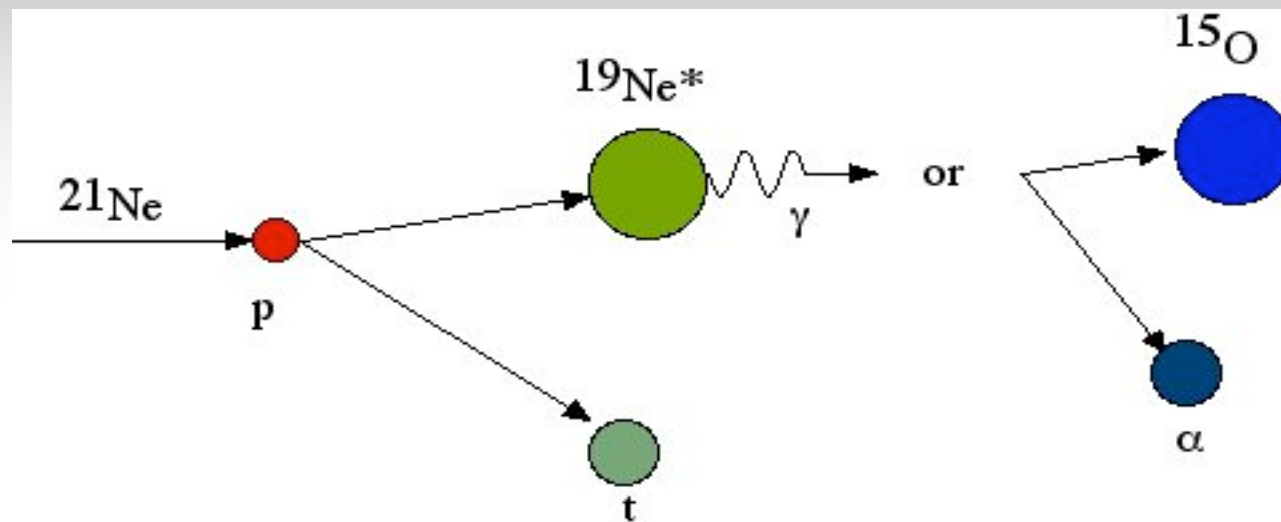


Fisker et al., ApJ 650, 332 (2006): reaction rate must be above certain value or accreted matter burns stably without bursting



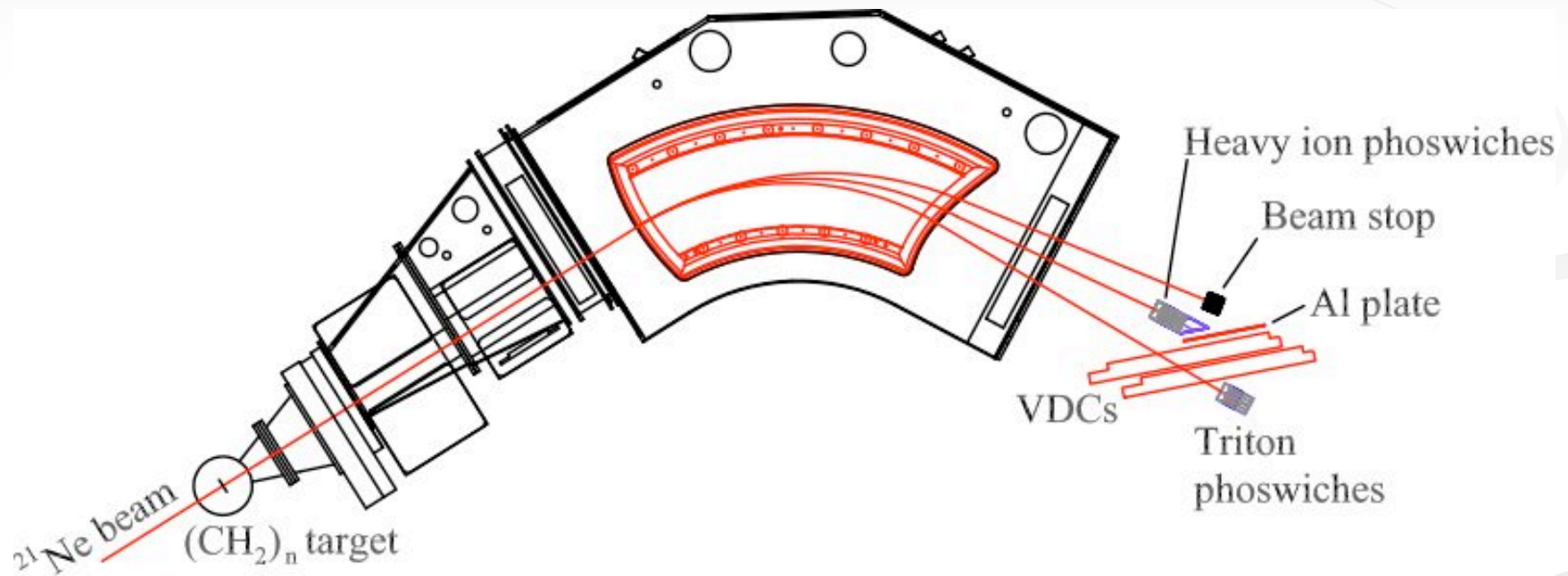
- No direct measurement is presently feasible (^{15}O is radioactive)
- Reaction proceeds resonantly at temperatures characteristic of x-ray bursts
- For narrow, isolated resonances, contributions add incoherently
- Contribution of each resonance to reaction rate proportional to its strength $\omega\gamma$
- Resonance strength $\omega\gamma \propto \Gamma_\alpha \Gamma_\gamma / \Gamma$
- $\Gamma_\alpha \Gamma_\gamma / \Gamma = B_\alpha (1 - B_\alpha) \hbar/\tau$, where B_α is the alpha-decay branching ratio and τ the mean lifetime of the state

Experimental Technique: B_α



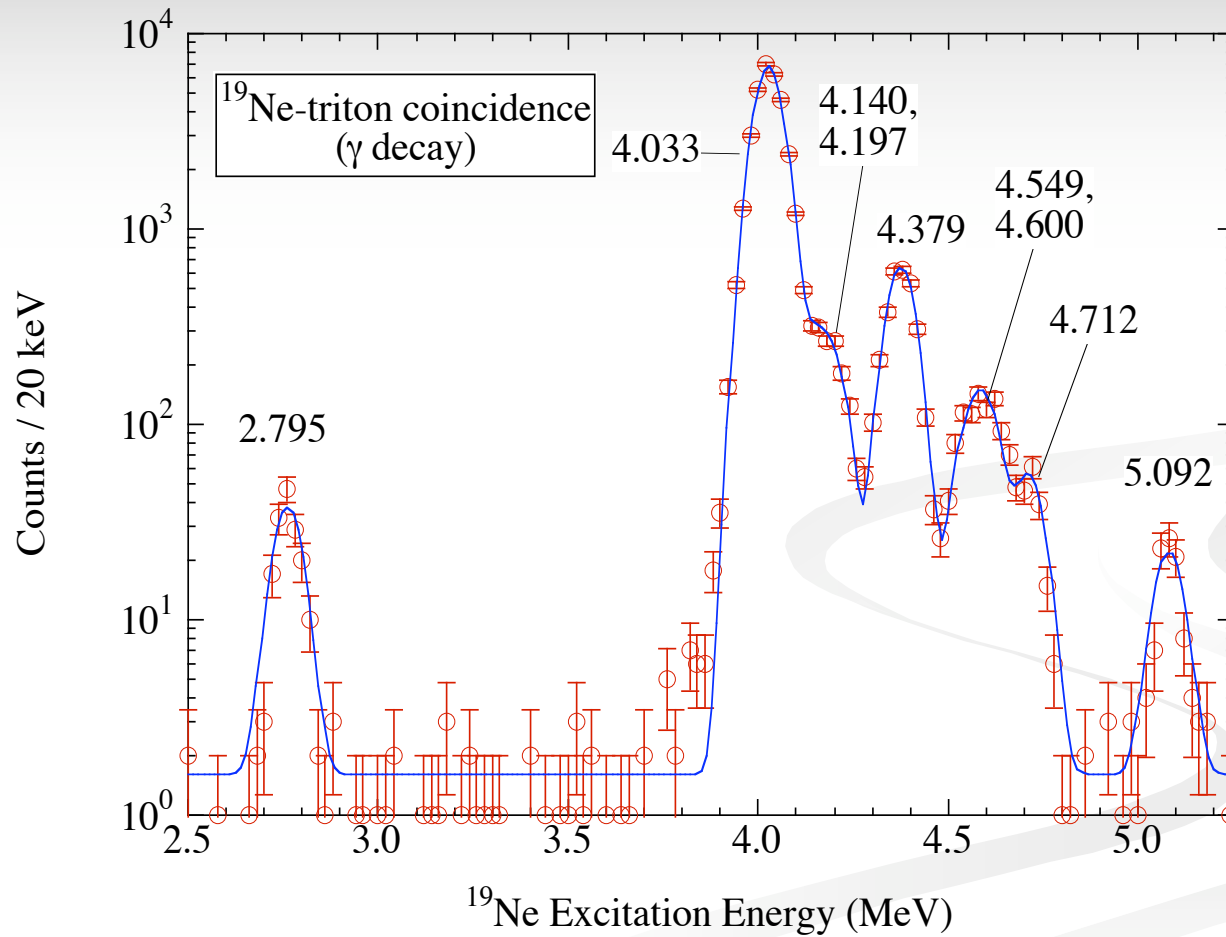
- Decay properties of excited states in ^{19}Ne determine reaction rate
- Populate states via transfer reaction, study decays
- States of interest decay exclusively by α or γ emission
- Populate states, count relative numbers of α & γ decays to obtain B_α
- Used $^{21}\text{Ne} + p \rightarrow ^{19}\text{Ne} + t$ reaction at KVI, Groningen, Netherlands

Experimental Setup at KVI's Big-Bite Spectrometer

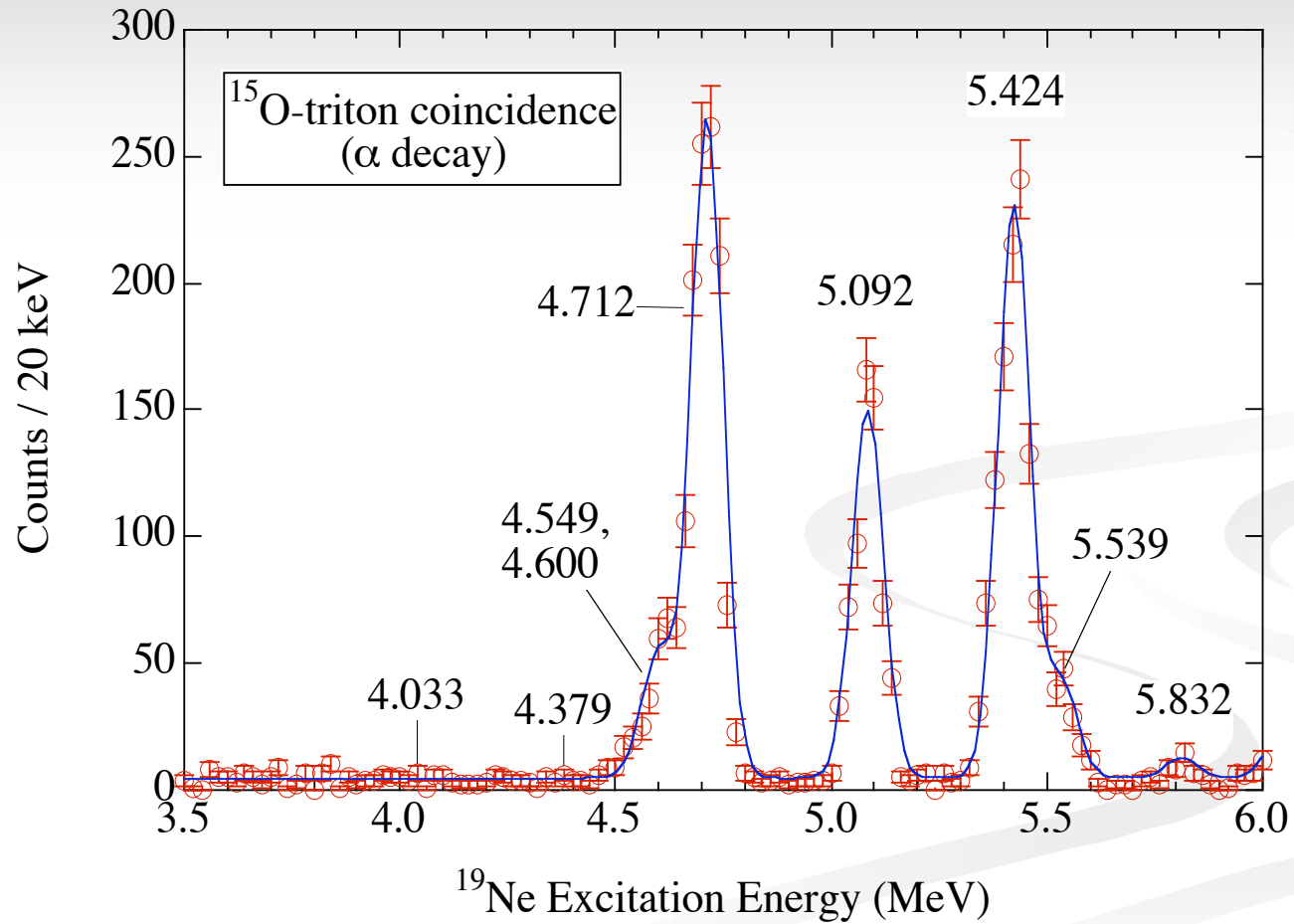


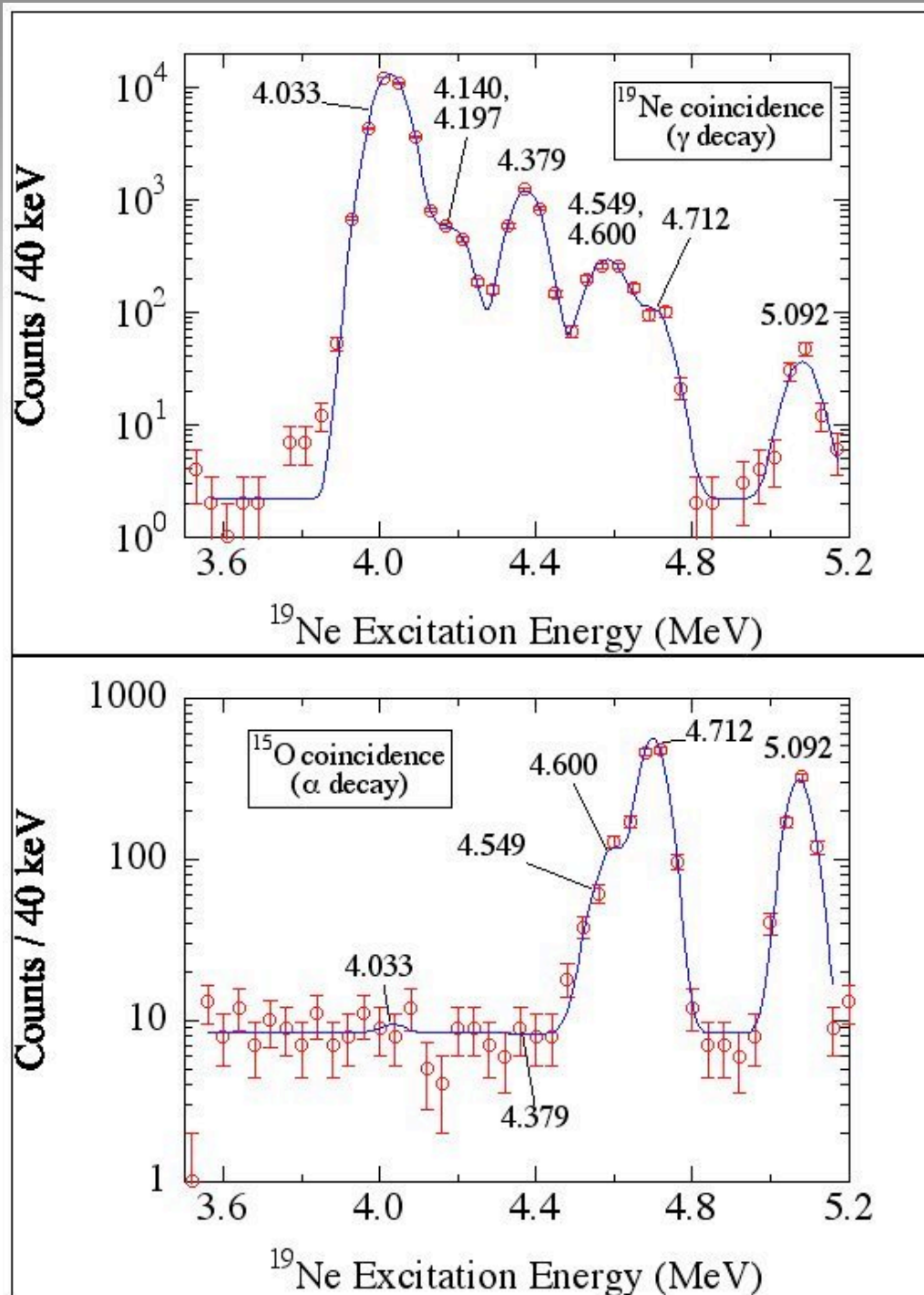
Triton determines ^{19}Ne excitation energy
Heavy ion in coincidence reveals decay mode

Gamma Decay



Alpha Decay





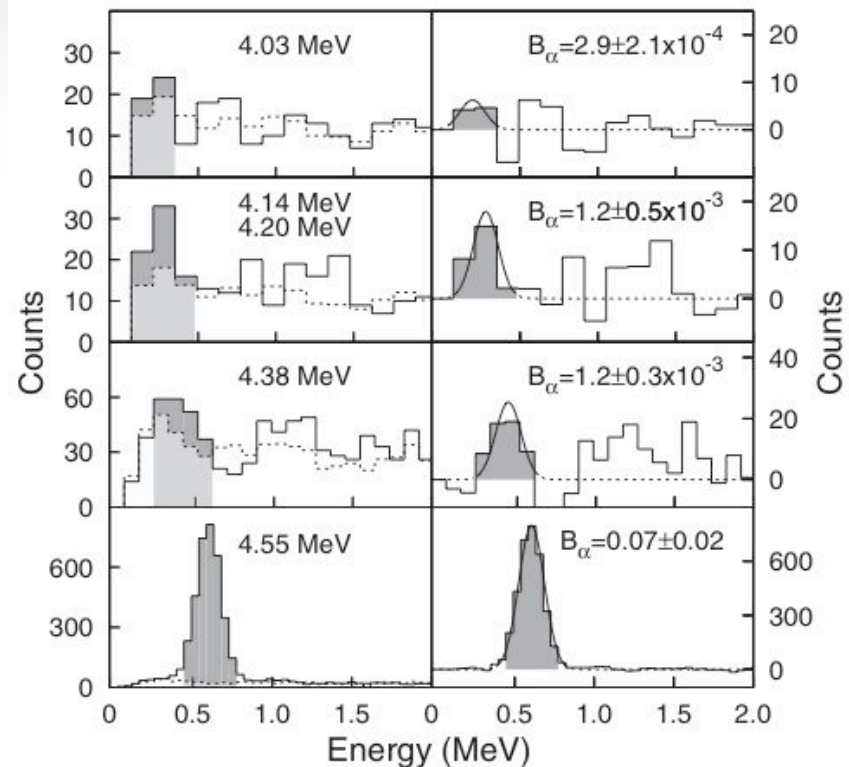
Branching Ratios

Excitation Energy (MeV)	Spin & Parity	B_{α} (this work)	B_{α} (Magnus <i>et al.</i>)	B_{α} (Laird <i>et al.</i>)
4.033	$3/2^+$	$< 4.3 \times 10^{-4}$		
4.379	$7/2^+$	$< 3.9 \times 10^{-3}$	0.044 ± 0.032	
4.549	$(3/2)^-$	0.16 ± 0.04	0.07 ± 0.03	
4.600	$(5/2^+)$	0.32 ± 0.04	0.25 ± 0.04	0.32 ± 0.03
4.712	$(5/2^-)$	0.85 ± 0.04	0.82 ± 0.15	
5.092	$5/2^+$	0.90 ± 0.06	0.90 ± 0.09	

Davids et al., Phys Rev C 67, 065808 (2003)

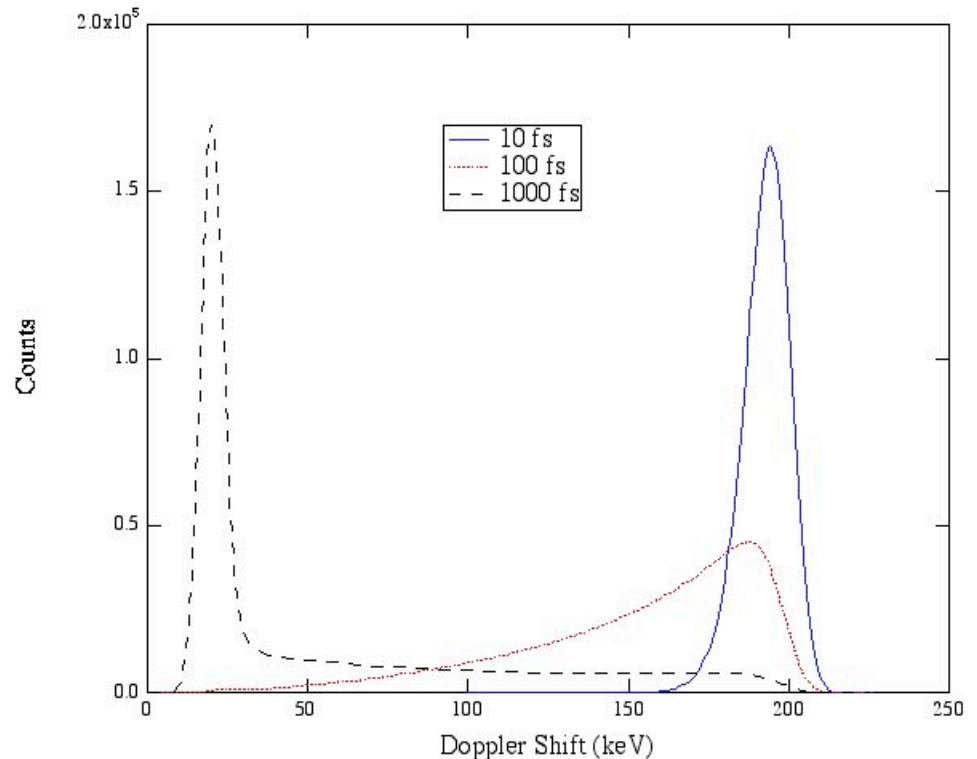
Notre Dame B_α Measurement?

- Recent PRL from Tan et al. claims to detect α decay from states below $E_x = 4.5$ MeV
- Data do not warrant claim
- Background poorly understood and modeled
- Statistical analysis flawed
- Reported branching ratios of 4.03, 4.14/4.20, and 4.38 MeV states are unreliable
- KVI data have highest sensitivity

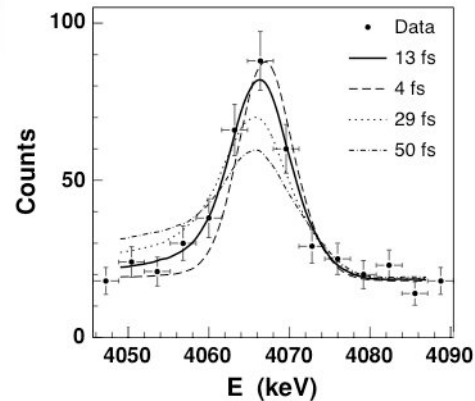


Lifetime Measurements

- Lifetimes measured via Doppler shift of emitted γ rays
- Fast decay \Rightarrow large Doppler shift, slow decay \Rightarrow small Doppler shift
- Shapes of detected γ ray lines yield lifetime; sensitive to fs lifetimes



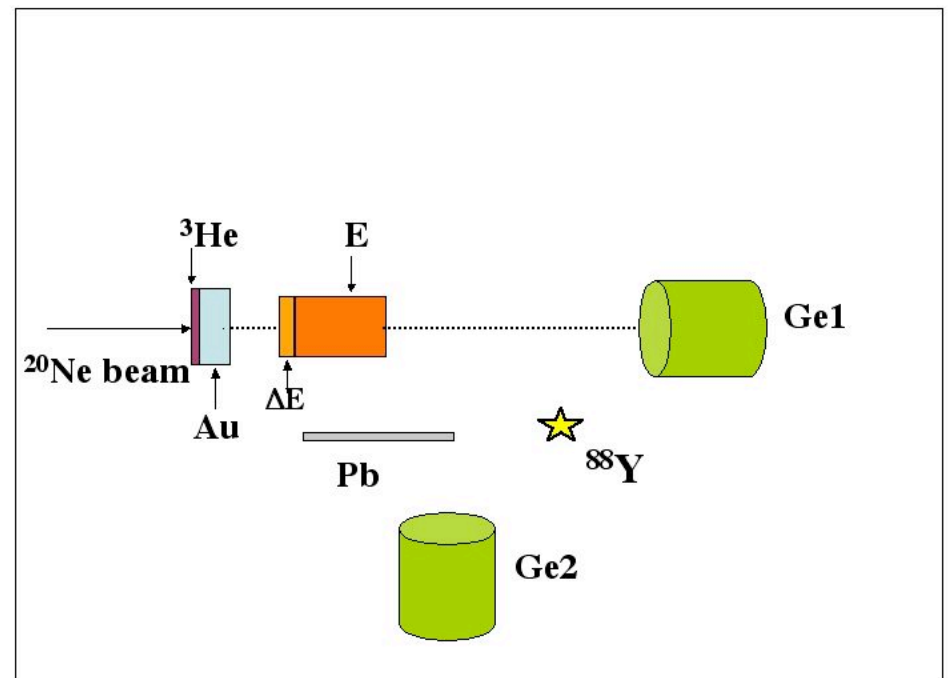
Notre Dame Data:

$$^{17}\text{O} + ^3\text{He} \rightarrow ^{19}\text{Ne} + \text{n}$$


Measured ten lifetimes, precisely determined transition energies
Lifetime of 4.03 MeV state = 13 (+ 9, - 6) fs

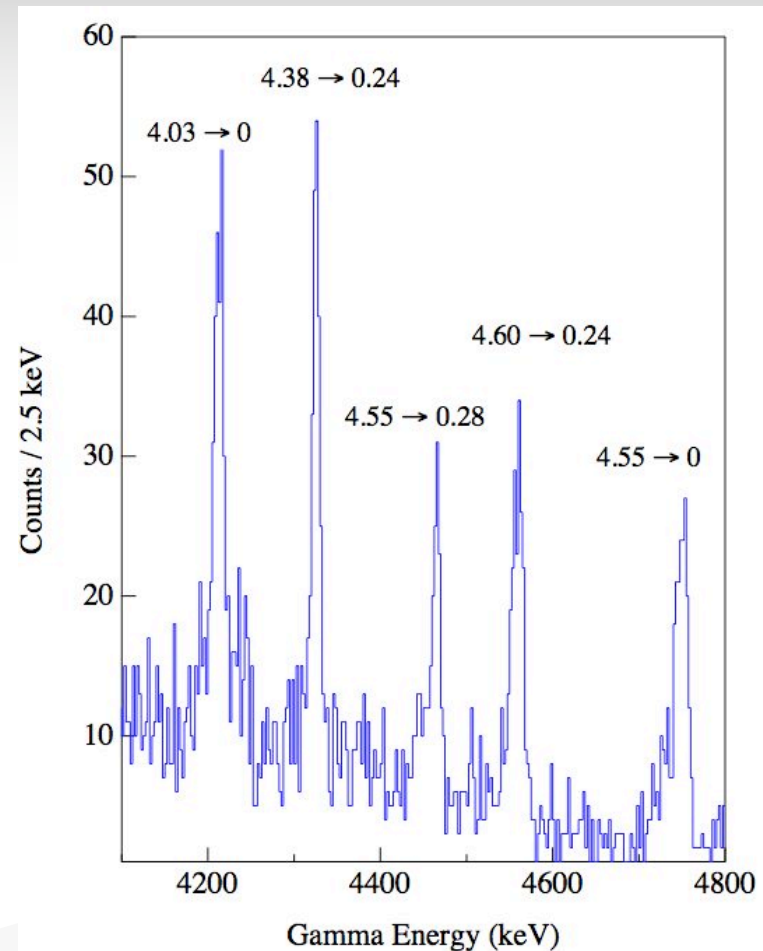
TRIUMF Measurement

- ^3He -implanted gold foil target
- $^{20}\text{Ne} + ^3\text{He} \rightarrow ^{19}\text{Ne} + \alpha$
- α particles detected in Si Δ E-E telescope
- ^{19}Ne emits a γ ray after slowing down in gold foil
- Detect Doppler-shifted de-excitation γ rays from ^{19}Ne states



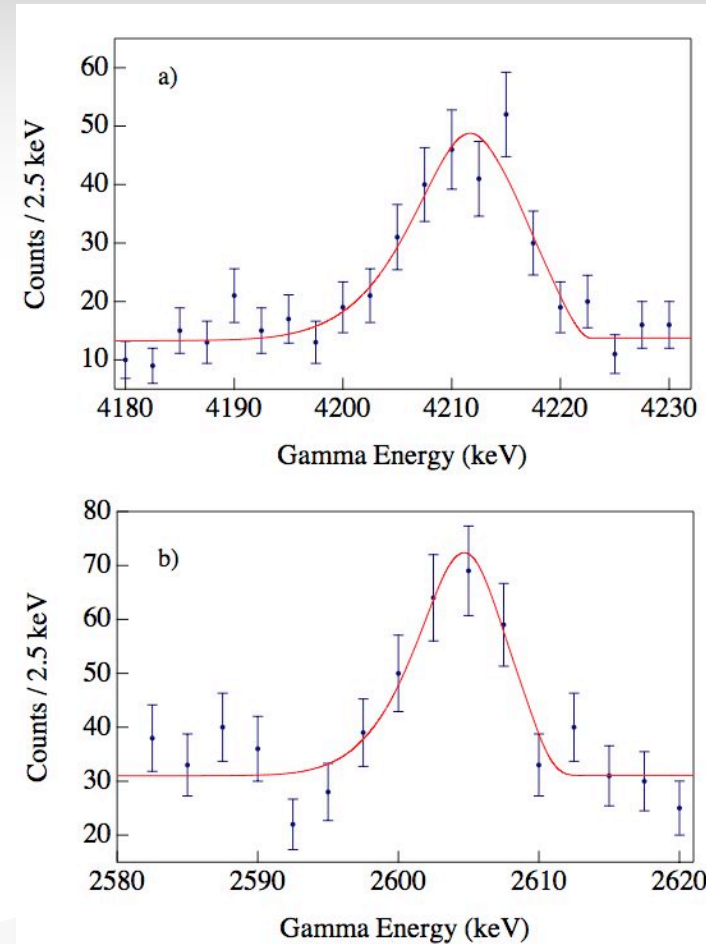
TRIUMF Data

- Measured lifetimes of 6 states above α threshold
- Two states observed via two transitions, other states only seen in one transition
- In general, lifetimes are consistent with and more precise than Notre Dame measurements

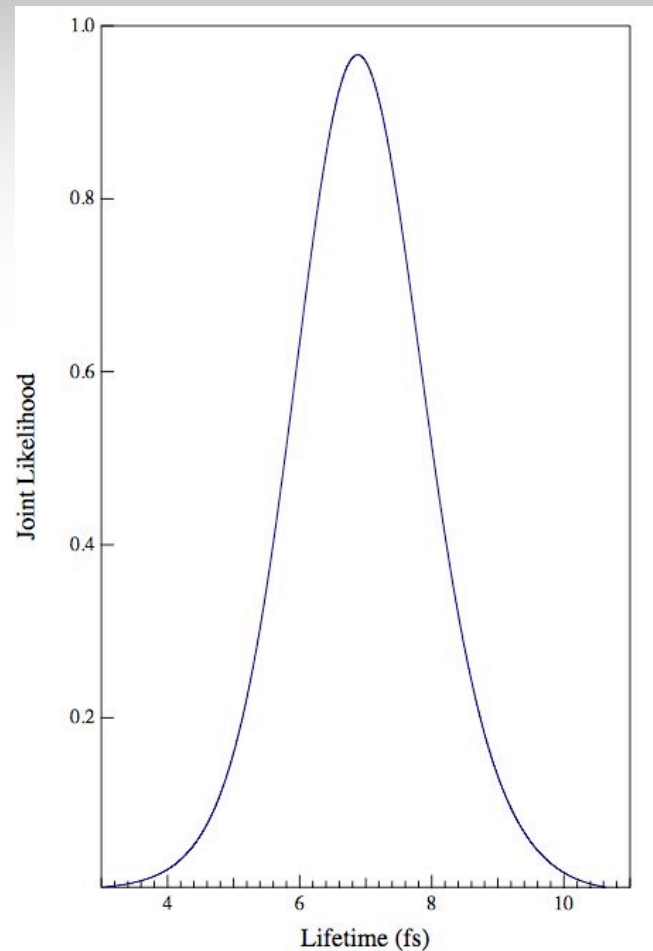


Transitions of 4.03 MeV State

- Two transitions observed, direct to ground state and also to 1536 keV state
- Two inferred lifetimes are mutually consistent

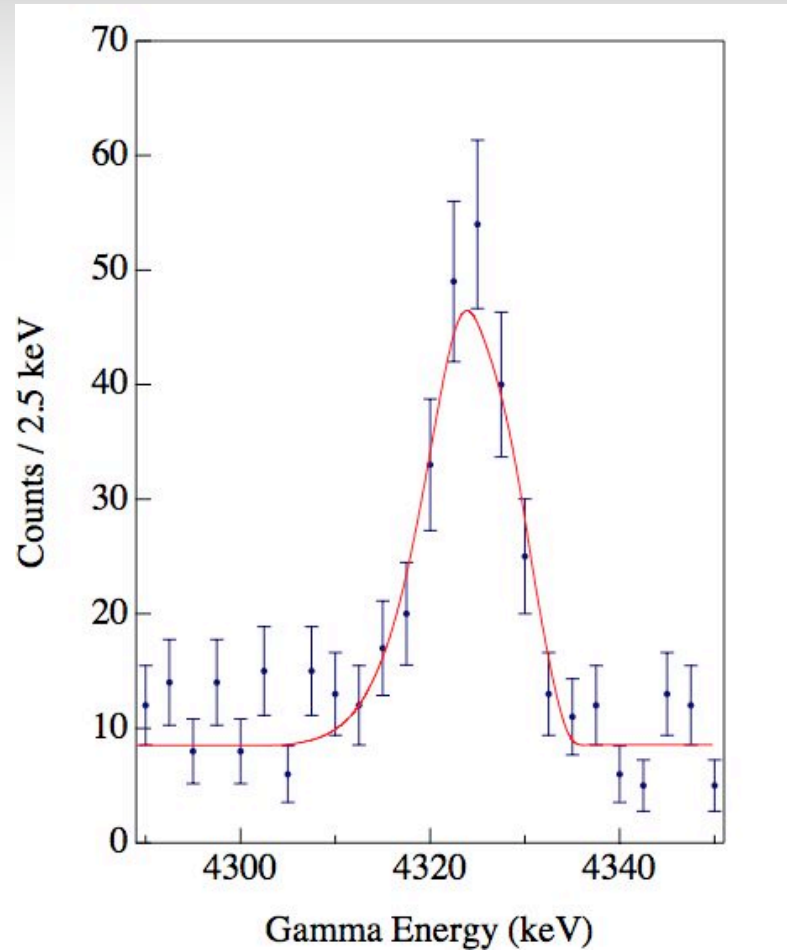


Lifetime of 4.03 MeV State



Joint likelihood analysis of two transitions yields
 $\tau = 6.9 \pm 1.5$ (statistical) ± 0.5 (systematic) fs

Lifetime of 4.38 MeV State



$$\tau = 2.9 \pm 1.4 \text{ (statistical)} \pm 0.4 \text{ (systematic) fs}$$

Comparison of Notre Dame and TRIUMF Lifetime Measurements

Level Energy (keV)	E_γ (keV) Ref. [9, 27]	Lifetime (fs) Ref. [9]	Lifetime (fs) Ref. [15]	Lifetime (fs) This work
1536	1297.8(4)	16 ± 4		$19.1_{-0.6}^{+0.7} \pm 0.9$
4034	2498.5(9)			$6.6_{-2.1}^{+2.4} \pm 0.5$
	4034.5(8)	13_{-6}^{+9}	11_{-3}^{+4}	$7.1_{-1.9}^{+1.9} \pm 0.6$
	Combined			$6.9_{-1.5}^{+1.5} \pm 0.5$
4144	2635.9(7)	18_{-3}^{+2}		$14.0_{-4.0}^{+3.5} \pm 1.2$
4200	2692.7(11)	43_{-9}^{+12}		$38_{-10}^{+20} \pm 2$
4378	4139.5(6)	5_{-2}^{+3}		$2.9_{-1.4}^{+1.4} \pm 0.4$
4548	4272.6(10)			$14.9_{-3.3}^{+4.3} \pm 1.9$
	4547.7(10)	15_{-5}^{+11}		$19.9_{-3.6}^{+3.0} \pm 1.9$
	Combined			$18.4_{-3.2}^{+3.3} \pm 1.9$
4602	4363.8(8)	7_{-4}^{+5}		$7.6_{-2.0}^{+2.0} \pm 0.7$

$^{15}\text{O}(\alpha,\gamma)^{19}\text{Ne}$: Status

- Lifetimes are now well measured
- B_α of important low-lying states @ 4.03 and 4.38 MeV still only constrained from above by upper limits
- Combining best experimental measurements, upper limit on reaction rate still much larger (~ 100 times) than theoretical lower limit
- New theoretical upper limit from Cooper and Narayan 2006: find that bursts would be observed from rapidly accreting neutron stars from which only stable burning is seen, if reaction rate were more than $\sim 1/25$ of experimental upper limit
- Theoretical bounds on rate are presently tightest (if they are to be believed)
- More sensitive B_α measurements required to make further progress

Credits

- Jacob Fisker, Lawrence Livermore National Lab (x-ray burst calculations)
- Wanpeng Tan, University of Notre Dame (lifetime measurements)
- Mythili Subramanian, TRIUMF and University of British Columbia (lifetime measurements)