

Canonical-basis HFB method

— drip lines and beyond —

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Contents

1. A new method to obtain the solutions of the HFB equation directory in its canonical form
 - advantages and disadvantages
 - practical and conceptual significance
2. Extension of the code to the full Skyrme energy density functional
 - is under way. A new progress is an extension to $N \neq Z$ nuclei.
 - Spin-orbit and Coulomb forces are not considered.
3. Calculation of the isotope chain of ^{14}Si
 - reaching the neutron drip line
 - beyond the neutron drip line
4. Summary

To treat the neutron-rich side of the nuclear chart, one has to take into account

1. Pairing in the continuum

$$\epsilon_F > -\frac{1}{2}\hbar\omega_{osc} \text{ (at least half of the shell needed)} \rightarrow \frac{10^4}{2} \text{ nuclides}$$

2. Long tail for large r $\epsilon_F > -1 \text{ MeV}$ for halo

$$\rho(r) \sim \left(\frac{e^{-\kappa r}}{r}\right)^2, \quad \kappa = \frac{\sqrt{-2m\epsilon_F}}{\hbar}$$

pairing \rightarrow shorter tail

$$\epsilon_F \rightarrow \epsilon_F - E_{qp}^{(\min)} \text{ (deepened), } E_{qp}^{(\min)} \sim \Delta(\text{pairing gap})$$

3. Deformation all but near spherical magics

\Rightarrow Coordinate-space Hartree-Fock-Bogoliubov method

such as 3D Cartesian mesh
feasible for Skyrme force

not HF+BCS

\Rightarrow Now there is a new method to solve it : the Canonical-basis HFB method !

HFB in quasi-particle method

$$|\psi\rangle = \prod_{i=1}^{\# \text{basis}} b_i |0\rangle$$

$$b_i = \sum_s \int d^3r \left\{ \phi_i^*(\vec{r}, s) a(\vec{r}, s) + \psi_i(\vec{r}, s) a^\dagger(\vec{r}, s) \right\}$$

$$\begin{pmatrix} -h & \tilde{h} \\ \tilde{h} & h \end{pmatrix} \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix} = \epsilon_i \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix}$$

b_i^\dagger : quasi-particle states, including information on excitations.

HFB in canonical-basis method

HFB solutions can be expressed in the BCS form

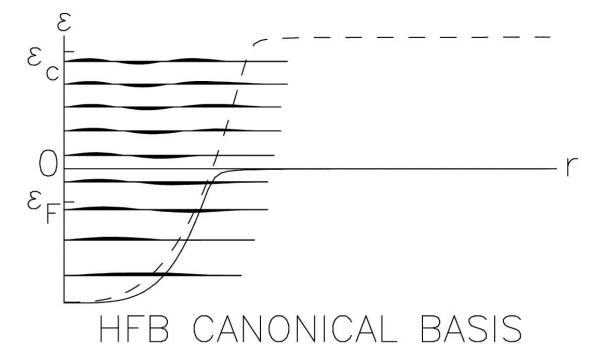
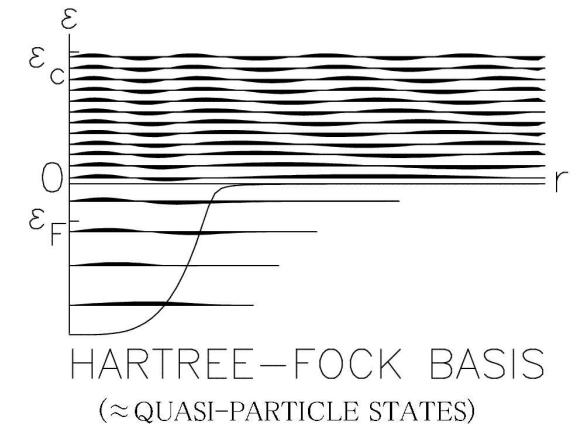
$$|\Psi\rangle = \prod_{i=1}^{i_{\max}} (u_i + v_i a_i^\dagger a_{\bar{i}}^\dagger) |0\rangle$$

$$a_i^\dagger = \sum_s \int d^3r \psi_i(\vec{r}, s) a^\dagger(\vec{r}, s) \quad : \text{HFB canonical basis}$$

$$\frac{\delta E}{\delta \psi_i^*} = \mathcal{H}_i \psi_i = \sum_j \lambda_{ij} \psi_j, \quad \mathcal{H}_i = v_i^2 h + u_i v_i \tilde{h}$$

Exact when $i_{\max} = \frac{1}{2} \# \text{basis}$ (Bloch-Messiah theorem)

One may neglect $v^2 \ll 1$ states to describe the ground state.
i.e., $i_{\max} = \mathcal{O}(A) \ll \# \text{basis}$ to a good approximation



h : Hartree-Fock Hamiltonian
 \tilde{h} : Pairing Hamiltonian

Nature of the positive energy canonical basis

Positive-energy canonical basis

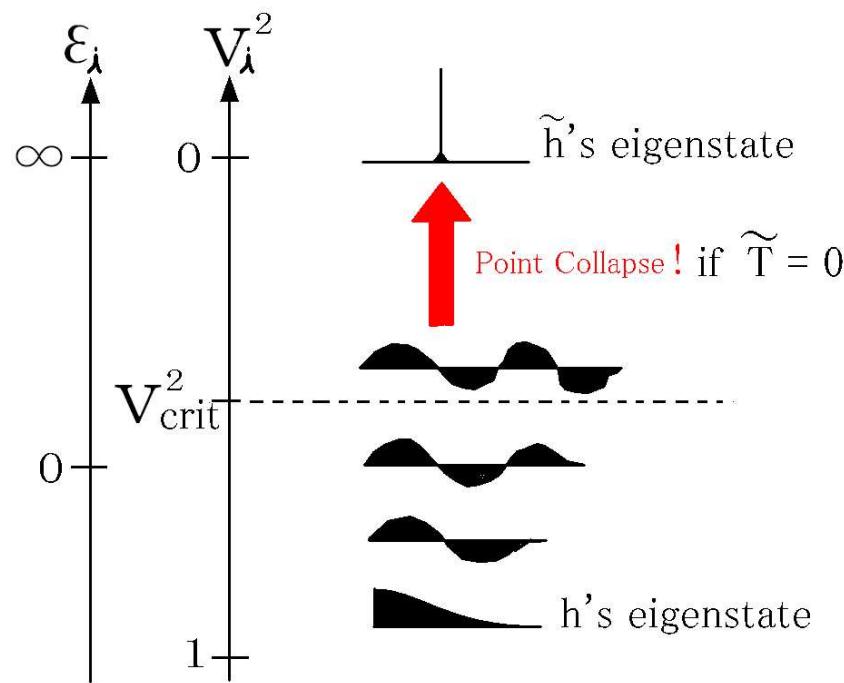
- are spatially localized by \tilde{h} .
- have a discrete spectrum.
- are determined by \tilde{h} , not by h .

Comparison of the two methods

method	#basis	orthogonality condition	pairing force
quasi particle	\propto box vol.	redundant	δ -func, + dens. dep.
canonical basis	\propto nucl. vol.	essential	+ mom. dep.

- Hamiltonian becomes **state-dependent**.
⇒ Orthogonality among the canonical basis states should be explicitly taken care of.
- Necessity of momentum-dependent or finite-range pairing interaction
⇐ Point collapse problem (see the next slide)

Point Collapse



How to avoid the collapse

- Sharp cutoff \rightarrow too few orbitals
- Smooth cutoff \rightarrow no q.p. Hamiltonian
- Density dependence
 \rightarrow cannot prevent the collapse
- Pairing density dependence \rightarrow OK
- Momentum dependence \rightarrow OK

HFB with the Skyrme force

Mean-field interaction

$$\begin{aligned}\hat{v} = & \textcolor{red}{t}_0(1 + x_0 P_\sigma)\delta + \frac{1}{2}\textcolor{red}{t}_1(1 + x_1 P_\sigma)(\vec{k}^2\delta + \delta\vec{k}^2) + t_2(1 + x_2 P_\sigma)\vec{k} \cdot \delta\vec{k} \\ & + \frac{1}{6}\rho^\alpha\textcolor{red}{t}_3(1 + x_3 P_\sigma)\delta + iW(\vec{\sigma}_1 + \vec{\sigma}_2) \cdot \vec{k} \times \delta\vec{k}\end{aligned}$$

Pairing interaction: different parameters assumed, only for (S=0, T=1) pairs.

$$\hat{v}_p = \textcolor{red}{v}_p \frac{1 - P_\sigma}{2} \left\{ \left[1 - \frac{\rho}{\rho_c} - \frac{\rho_n - \rho_p}{\rho'_c} T_z - \left(\frac{\tilde{\rho}}{\tilde{\rho}_c} \right)^2 \right] \delta - \frac{1}{2\textcolor{red}{k}_c^2} (\vec{k}^2\delta + \delta\vec{k}^2) \right\}$$

- v_p : overall strength, to be adjusted depending on the cutoff.
Cutoff can be controlled by the number of (explicitly considered) canonical basis with $v^2 > 0$.
- $\rho_c = 0.32 \text{ fm}^{-1}$, $\rho'_c = \tilde{\rho}_c = \infty$: density dependence, insufficient information
- $k_c = 2 \text{ fm}^{-1}$: momentum dependence (=finite range effect), prevents the point collapse

Hamiltonian density for even-even nuclei

$$\begin{aligned}
E = \int \mathcal{H} d\vec{r}, \quad \mathcal{H} = & \frac{\hbar^2}{2m} (\tau_n + \tau_p) \\
& + C_1 (\rho_n^2 + \rho_p^2) + C'_1 \rho_n \rho_p \\
& + C_2 (\rho_n \tau_n + \rho_p \tau_p) + C'_2 (\rho_n \tau_p + \rho_p \tau_n) \\
& + C_3 (\rho_n \vec{\nabla}^2 \rho_n + \rho_p \vec{\nabla}^2 \rho_p) + C'_3 (\rho_n \vec{\nabla}^2 \rho_p + \rho_p \vec{\nabla}^2 \rho_n) \\
& + C_4 \rho^\alpha (\rho_n^2 + \rho_p^2) + C'_4 \rho^\alpha \rho_n \rho_p \\
& + C_5 (\rho_n \vec{\nabla} \cdot \vec{J}_n + \rho_p \vec{\nabla} \cdot \vec{J}_p) + C'_5 (\rho_n \vec{\nabla} \cdot \vec{J}_p + \rho_p \vec{\nabla} \cdot \vec{J}_n) \\
& + C_6 \tilde{\rho}_n^2 + C_7 \rho_n \tilde{\rho}_n^2 + C'_7 \rho_p \tilde{\rho}_n^2 + C_8 \tilde{\rho}_n^4 + C_9 \tilde{\rho}_n (\tilde{\tau}_n - \vec{\nabla}^2 \tilde{\rho}_n) \\
& + C_6 \tilde{\rho}_p^2 + C_7 \rho_p \tilde{\rho}_p^2 + C'_7 \rho_n \tilde{\rho}_p^2 + C_8 \tilde{\rho}_p^4 + C_9 \tilde{\rho}_p (\tilde{\tau}_p - \vec{\nabla}^2 \tilde{\rho}_p) \\
& + C_{10} V_C \rho_p + C_{11} \rho_p^{4/3}
\end{aligned}$$

where

$C_1 = \frac{1}{4} t_0 (1 - x_0)$ $C_2 = \frac{1}{8} (t_1(1 - x_1) + 3t_2(1 + x_2))$ $C_3 = \frac{3}{32} (t_1(x_1 - 1) + t_2(x_2 + 1))$ $C_4 = \frac{1}{24} t_3 (1 - x_3)$ $C_5 = -W$ $C_6 = \frac{1}{4} v_P$ $C_7 = -\frac{1}{4} v_P \left(\frac{1}{\rho_c} + \frac{1}{\rho'_c} \right)$ $C_8 = -\frac{1}{4} \frac{v_P}{\tilde{\rho}_c^2}$ $C_{10} = \frac{1}{2}$	$C'_1 = \frac{1}{2} t_0 (2 + x_0)$ $C'_2 = \frac{1}{8} (t_1(2 + x_1) + t_2(2 + x_2))$ $C'_3 = \frac{1}{32} (-3t_1(2 + x_1) + t_2(2 + x_2))$ $C'_4 = \frac{1}{12} t_3 (2 + x_3)$ $C'_5 = -\frac{1}{2} W$ $C'_7 = -\frac{1}{4} v_P \left(\frac{1}{\rho_c} - \frac{1}{\rho'_c} \right)$ $C_9 = -\frac{1}{4} \frac{v_P}{k_c^2}$ $C_{11} = -\frac{3}{4} e^2 \left(\frac{3}{\pi} \right)^{1/3}$
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Densities ($\mathbf{q} = \mathbf{n}$, \mathbf{p})

$$\begin{aligned}\tau_{\mathbf{q}}(\vec{r}) &= 2 \sum_{i>0,\sigma} v_{qi}^2 |\vec{\nabla} \psi_{qi}(\vec{r}, \sigma)|^2, & \tilde{\tau}_{\mathbf{q}}(\vec{r}) &= 2 \sum_{i>0,\sigma} u_{qi} v_{qi} |\vec{\nabla} \psi_{qi}(\vec{r}, \sigma)|^2, \\ \rho_{\mathbf{q}}(\vec{r}) &= 2 \sum_{i>0,\sigma} v_{qi}^2 |\psi_{qi}(\vec{r}, \sigma)|^2, & \tilde{\rho}_{\mathbf{q}}(\vec{r}) &= 2 \sum_{i>0,\sigma} u_{qi} v_{qi} |\psi_{qi}(\vec{r}, \sigma)|^2 \\ \vec{\nabla} \cdot \vec{J}_{\mathbf{q}} &= \frac{2}{i} \sum_{i>0,\sigma,\sigma'} v_{qi}^2 \vec{\nabla} \psi_{qi}^*(\vec{r}, \sigma) \times \vec{\nabla} \psi_{qi}(\vec{r}, \sigma') \cdot \langle \sigma | \vec{\sigma} | \sigma' \rangle\end{aligned}$$

Effective masses and single-particle potentials

$$\begin{aligned}B_{\mathbf{n}} &= \frac{\hbar^2}{2m} + C_2 \rho_{\mathbf{n}} + C'_2 \rho_{\mathbf{p}} \\ \tilde{B}_{\mathbf{n}} &= C_9 \tilde{\rho}_{\mathbf{n}} \\ V_{\mathbf{n}} &= 2C_1 \rho_{\mathbf{n}} + C'_1 \rho_{\mathbf{p}} + C_2 \tau_{\mathbf{n}} + C'_2 \tau_{\mathbf{p}} + 2C_3 \vec{\nabla}^2 \rho_{\mathbf{n}} + 2C'_3 \vec{\nabla}^2 \rho_{\mathbf{p}} \\ &\quad + \rho^{\alpha-1} [(\alpha+2)C_4 \rho_{\mathbf{n}}^2 + (\alpha C_4 + C'_4) \rho_{\mathbf{p}}^2 + (2C_4 + (\alpha+1)C'_4) \rho_{\mathbf{n}} \rho_{\mathbf{p}}] \\ &\quad + C_5 \vec{\nabla} \cdot \vec{J}_{\mathbf{n}} + C'_5 \vec{\nabla} \cdot \vec{J}_{\mathbf{p}} + C_7 \tilde{\rho}_{\mathbf{n}}^2 + C'_7 \tilde{\rho}_{\mathbf{p}}^2 \\ \vec{W}_{\mathbf{n}} &= -C_5 \vec{\nabla} \rho_{\mathbf{n}} - C'_5 \vec{\nabla} \rho_{\mathbf{p}} \\ \tilde{V}_{\mathbf{n}} &= 2C_6 \tilde{\rho}_{\mathbf{n}} + 2C_7 \rho_{\mathbf{n}} \tilde{\rho}_{\mathbf{n}} + 2C'_7 \rho_{\mathbf{p}} \tilde{\rho}_{\mathbf{n}} + 4C_8 \tilde{\rho}_{\mathbf{n}}^3 + C_9 (\tilde{\tau}_{\mathbf{n}} - 2\vec{\nabla}^2 \tilde{\rho}_{\mathbf{n}})\end{aligned}$$

State dependent Hamiltonian

$$\begin{aligned}h_{\mathbf{q}} &= -\vec{\nabla} \cdot B_{\mathbf{q}} \vec{\nabla} + V_{\mathbf{q}} + i \vec{W}_{\mathbf{q}} \cdot \vec{\sigma} \times \vec{\nabla} && : \text{mean-field Hamiltonian} \\ \tilde{h}_{\mathbf{q}} &= -\vec{\nabla} \cdot \tilde{B}_{\mathbf{q}} \vec{\nabla} + \tilde{V}_{\mathbf{q}} && : \text{pairing Hamiltonian} \\ \mathcal{H}_{\mathbf{q}i} &= v_{\mathbf{q}i}^2 h_{\mathbf{q}} + u_{\mathbf{q}i} v_{\mathbf{q}i} \tilde{h}_{\mathbf{q}} && : \text{Hamiltonian of } i\text{th canonical orbital}\end{aligned}$$

Approach to the drip line (and beyond)

isotope chain:

$\text{Si}(Z=14)$, $6 \leq N \leq 32$ have bound states

Mean filed:

SIII force, ls and Coulomb excluded.

Pairing force:

$$k_c = 2 \text{ fm}^{-1}, \rho_c = 0.32 \text{ fm}^{-3}.$$

$v_p(\text{MeV fm}^3)$	-680	-780	-880
$\Delta_n(N = 28)(\text{MeV})$	1.4	2.0	3.0

Mainly, $v_p = -780$ is used.

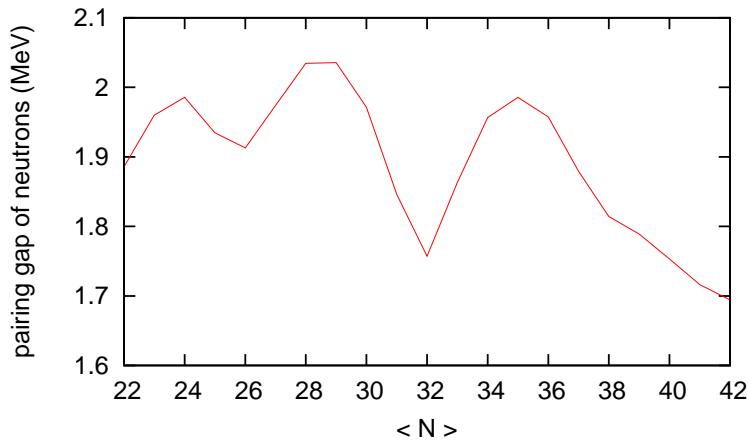
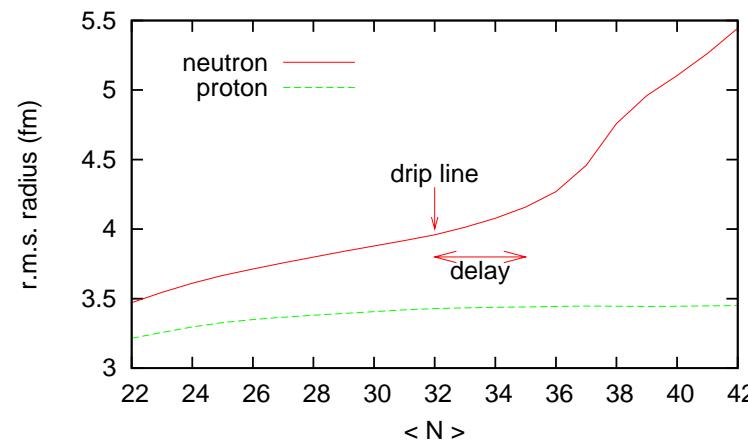
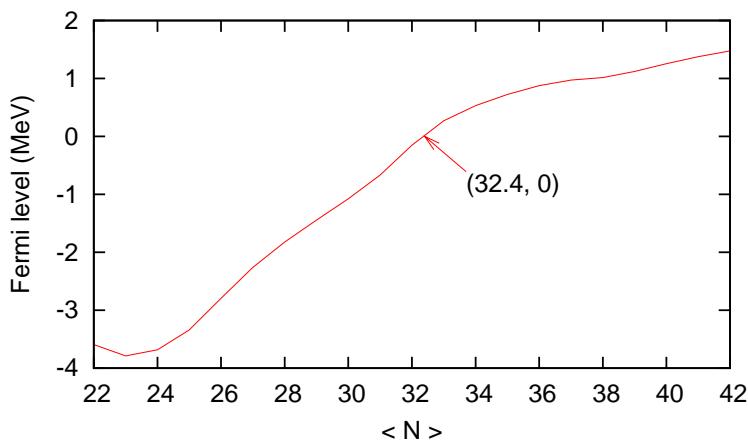
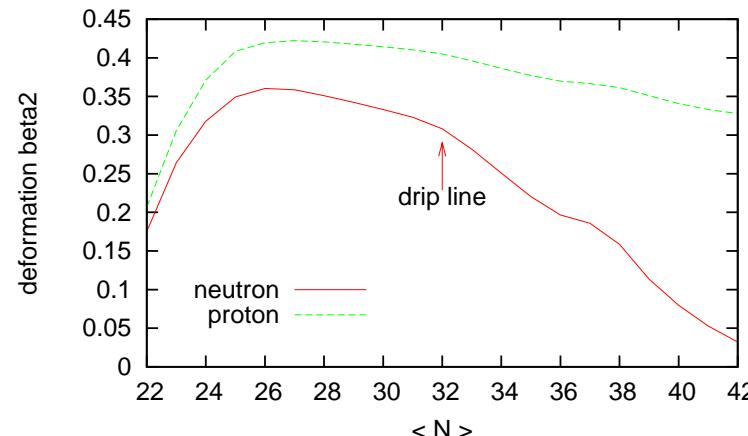
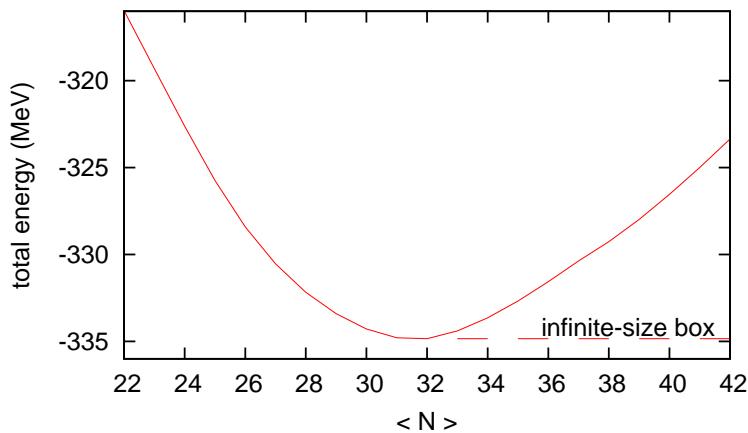
Box:

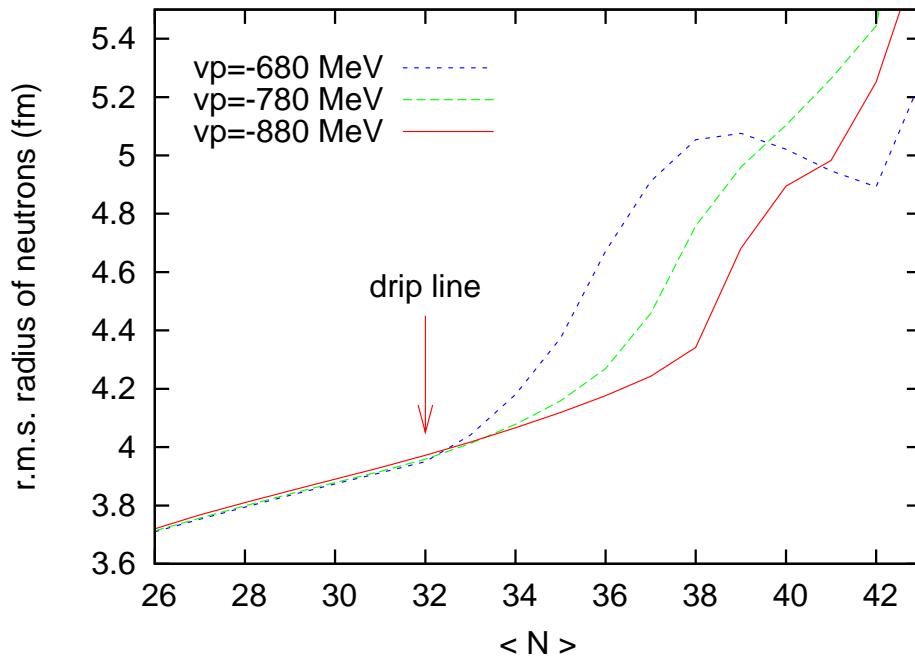
$$L = 32 \text{ fm}, \Delta x = 0.8 \text{ fm}$$

The number of canonical basis states:

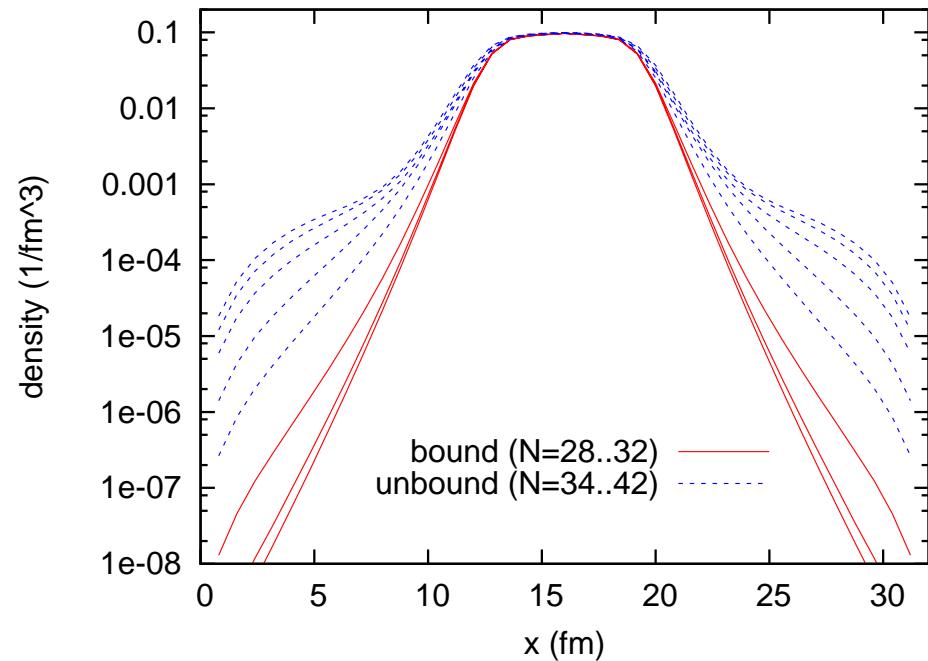
neutrons : 60×2 (2 for spin \uparrow and \downarrow)

protons : $21 \times 2 \rightarrow$ mostly in the normal state



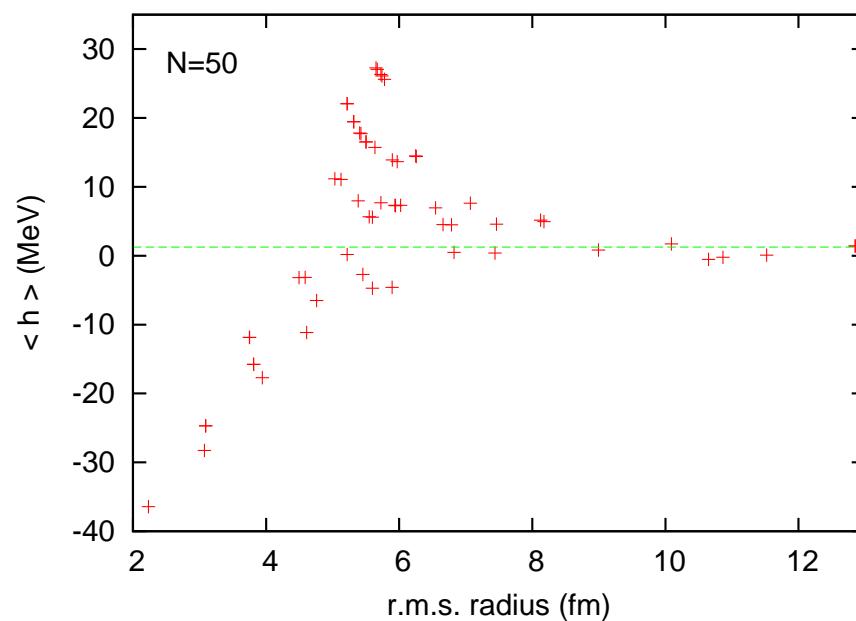
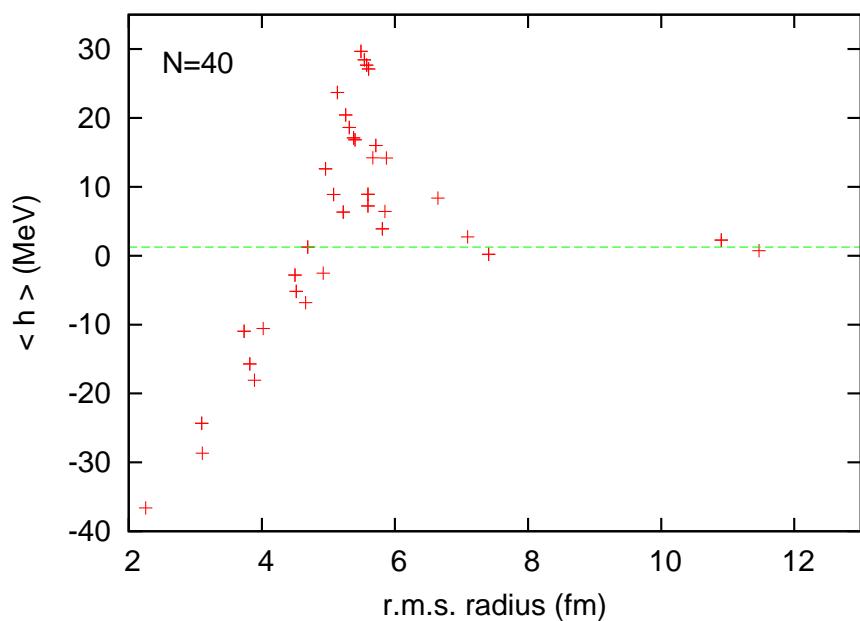
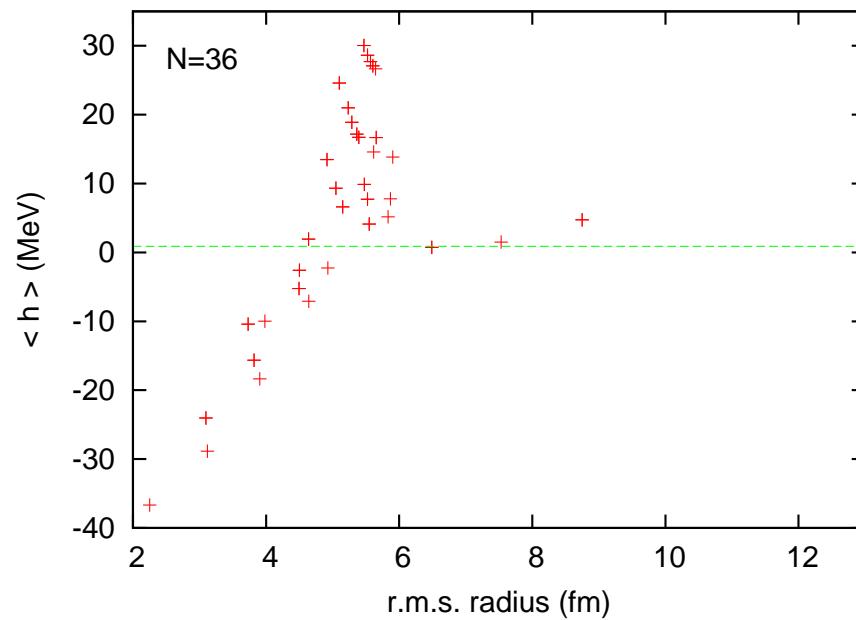
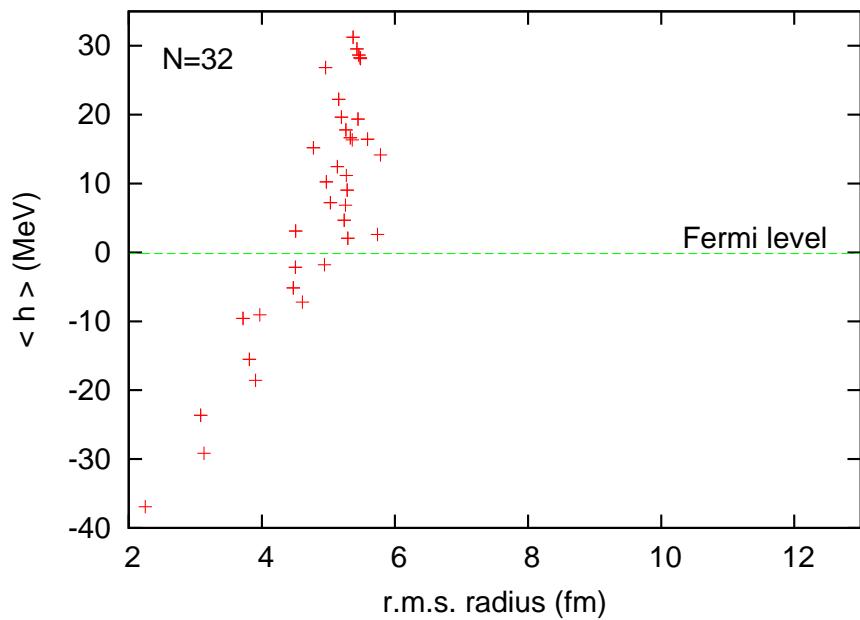


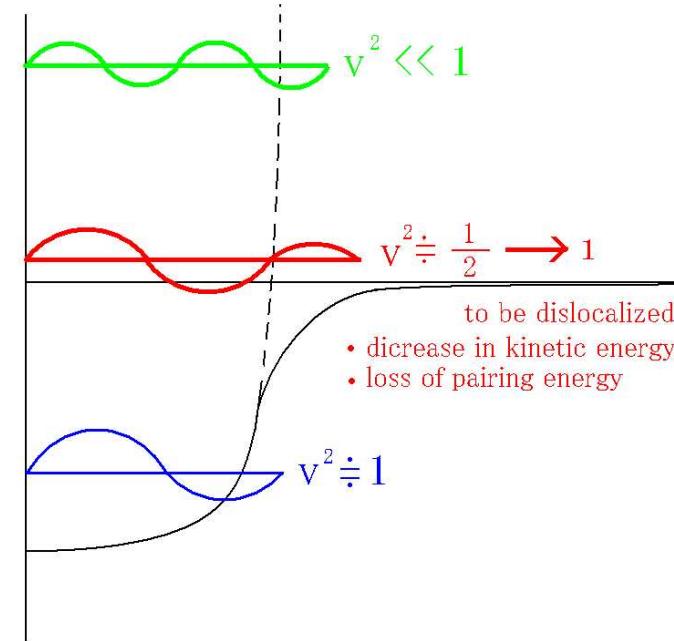
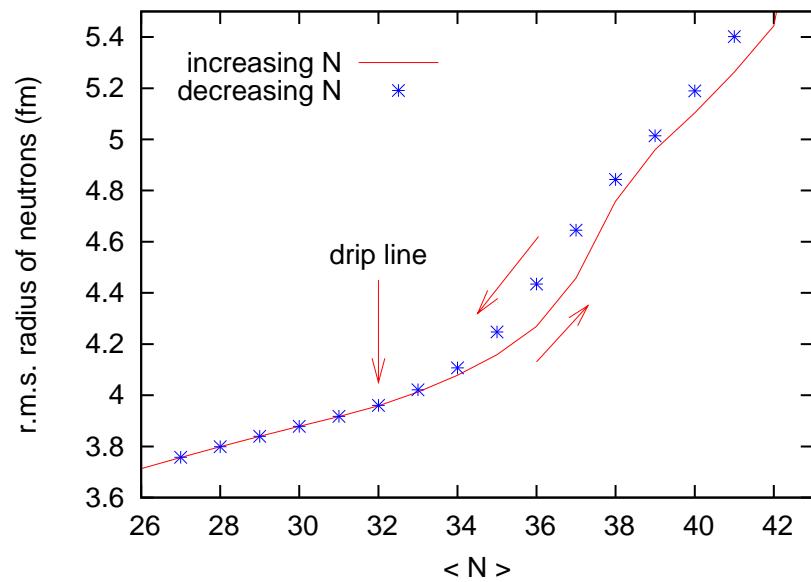
Stronger pairings lead to longer delays of dislocalization.



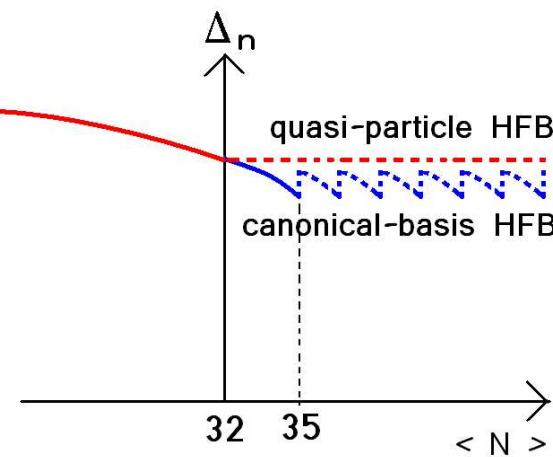
$$v_p = -780 \text{ MeV fm}^3$$

Canonical basis states of neutrons





- global minimum
- local minimum
- localized state
- states with dilute gas
of dripped neutrons



Summary

1. Canonical-basis HFB method is presented.

It is a much more economical way to express HFB ground states of neutron rich nuclei than expressing them as quasiparticle vacua.

Positive energy canonical-basis states are obtained as the bound states of the pairing Hamiltonian.

Consequently, they are guaranteed to be spatially localized and form a discrete spectrum.

2. The code has been extended to treat $N \neq Z$ systems.

Spin-orbit and Coulomb forces are not considered.

3. No difficulties are encountered in reaching the neutron drip line.

Approximate localized solutions for nearly bound systems can be obtained. It is probably not only for the finite-sized box but for the method to obtain solutions.

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