

Covariant density functional theory for excited states in nuclei

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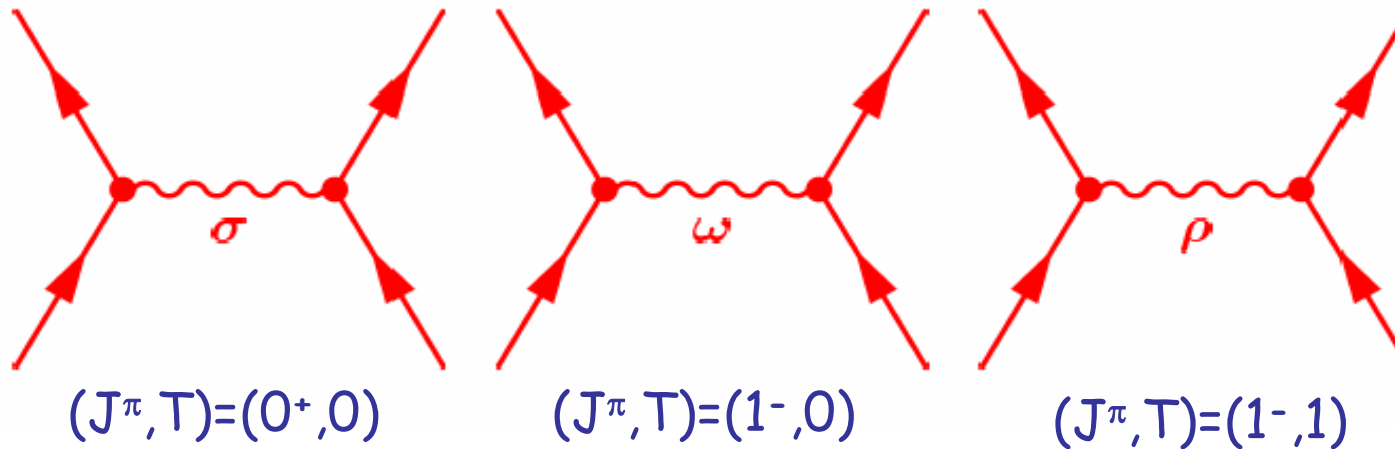
- rotational excitations (Cranked RHB)
- vibrational excitations (Rel. QRPA)

● Methods beyond mean field

- projected density functionals (PDFT)
- relativistic GCM
- particle vibrational coupling (PVC)
- decay width of Giant resonances

Covariant density functional theory

Nucleons are coupled by exchange of mesons through an effective Lagrangian (EFT)



$$S(\mathbf{r}) = g_\sigma \sigma(\mathbf{r})$$

Sigma-meson:
attractive scalar field

$$V(\mathbf{r}) = g_\omega \omega(\mathbf{r}) + g_\rho \vec{\tau} \vec{\rho}(\mathbf{r}) + eA(\mathbf{r})$$

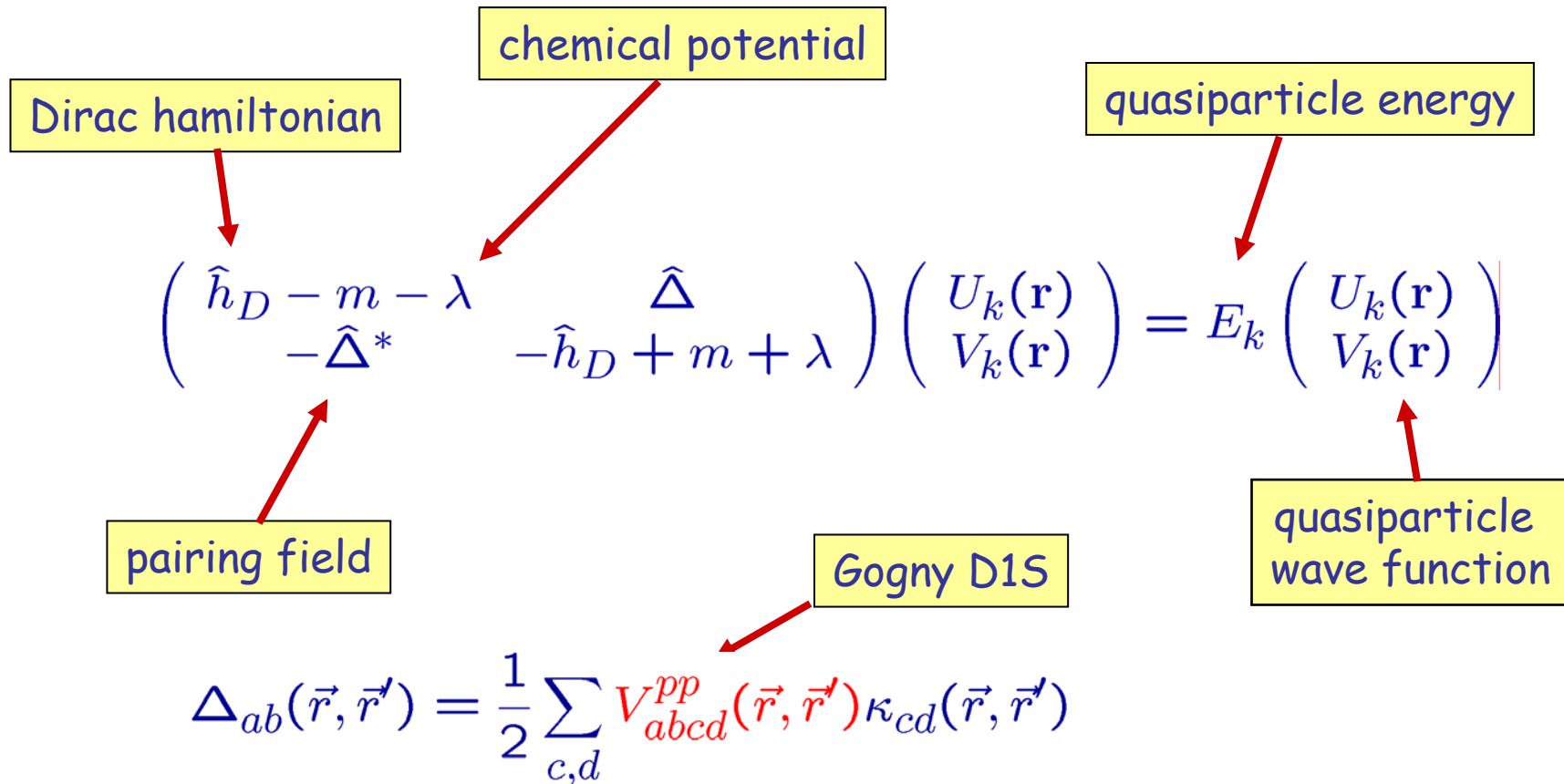
Omega-meson:
short-range repulsive

Rho-meson:
isovector field

Relativistic Hartree Bogoliubov (RHB)

Ground-state properties of weakly bound nuclei far from stability

→ Unified description of mean-field and pairing correlations



Effective density dependence:

non-linear potential:

NL1, NL3..

Boguta and Bodmer, NPA. 431, 3408 (1977)

$$\frac{1}{2}m_{\sigma}^2\sigma^2 \Rightarrow U(\sigma) = \frac{1}{2}m_{\sigma}^2\sigma^2 + \frac{1}{3}g_2\sigma^3 + \frac{1}{4}g_3\sigma^4$$

density dependent coupling constants:

R.Brockmann and H.Toki, PRL 68, 3408 (1992)

S.Typek and H.H.Wolter, NPA 656, 331 (1999)

$$g_o, g_{\omega}, g_{\rho} \Rightarrow g_o(\rho), g_{\omega}(\rho), g_{\rho}(\rho)$$

new

$$\mathbf{g} \rightarrow \mathbf{g}(\rho(\mathbf{r}))$$

DD-ME1, DD-ME2

Niksic et al, PRC 66, 024306 (2002)

Lalazissis et al Niksic, PRC 71, 024312 (2005)

Nuclei used in the fit for DD-ME2

Nucleus	B.E (MeV)	r_c (fm)	$r_n - r_p$ (fm)	dE	dr_c	dr_{np}
^{16}O	127.801 (127.619)	2.727 (2.730)	-0.03	0.1	-0.1	
^{40}Ca	342.741 (342.052)	3.464 (3.485)	-0.05	0.2	-0.6	
^{48}Ca	414.770 (415.991)	3.481 (3.484)	0.18	-0.3	-0.1	
^{72}Ni	612.655 (613.173)	3.914	0.28	-0.1		
^{90}Zr	783.155 (783.893)	4.275 (4.272)	0.07	-0.1	0.1	
^{116}Sn	986.928 (988.681)	4.615 (4.626)	0.12 (0.12)	-0.2	-0.2	3.8
^{124}Sn	1048.859 (1049.962)	4.671 (4.674)	0.21 (0.19)	-0.1	-0.1	10.7
^{132}Sn	1103.469 (1102.860)	4.718	0.26	0.1		
^{204}Pb	1608.506 (1607.520)	5.500 (5.486)	0.17	0.1	0.3	
^{208}Pb	1639.826 (1636.446)	5.518 (5.505)	0.19 (0.20)	0.2	0.2	-4.7
^{214}Pb	1661.182 (1663.298)	5.568 (5.562)	0.24	-0.1	0.1	
^{210}Po	1649.695 (1645.228)	5.552	0.17	0.3		

Nuclear matter:

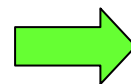
$E/A = -16$ MeV (5%), $\rho_0 = 1,53$ fm $^{-3}$ (10%)

$K = 250$ MeV (10%), $a_4 = 33$ MeV (10%)

How many parameters ?

7 parameters

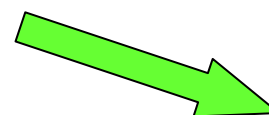
symmetric nuclear matter: $E/A, \rho_0$



$$\frac{g_\sigma}{m_\sigma}$$

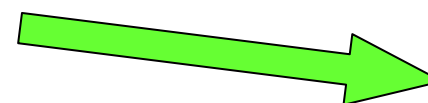
$$\frac{g_\omega}{m_\omega}$$

finite nuclei (N=Z): $E/A,$ radii
spinorbit o.k.



$$m_\sigma$$

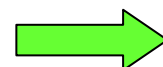
(N≠Z): Coulomb, symmetry energy: a_4



$$\frac{g_\rho}{m_\rho}$$

density dependence: T=0

$$K_\infty$$



$$g_2$$

$$g_3$$

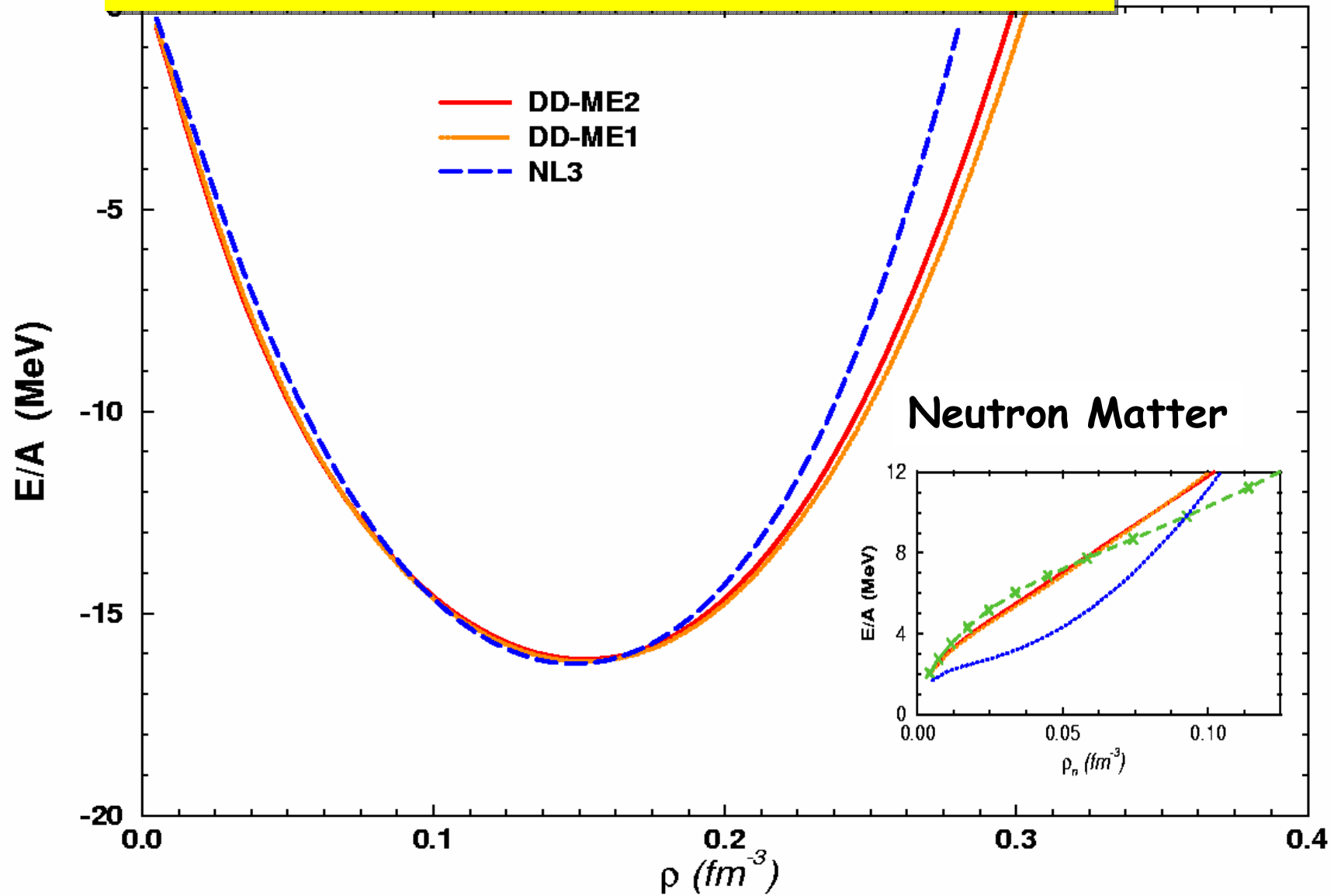
T=1

$$r_n - r_p$$

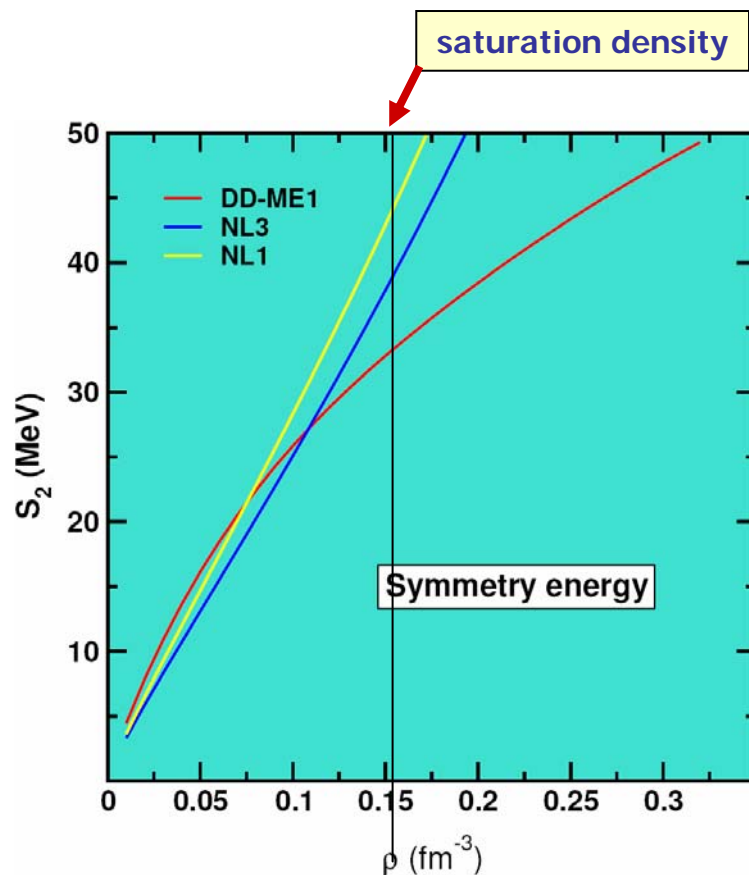


$$a_\rho$$

Nuclear matter equation of state



Symmetry energy



$$E(\rho, \alpha) = E(\rho, 0) + S_2(\rho)\alpha^2 + S_4(\rho)\alpha^4 + \dots$$

$$\alpha \equiv \frac{N-Z}{N+Z}$$

$$S_2(\rho) = a_4 + \frac{p_0}{\rho_{\text{sat}}^2} (\rho - \rho_{\text{sat}}) + \frac{\Delta K_0}{18\rho_{\text{sat}}^2} (\rho - \rho_{\text{sat}})^2 + \dots$$

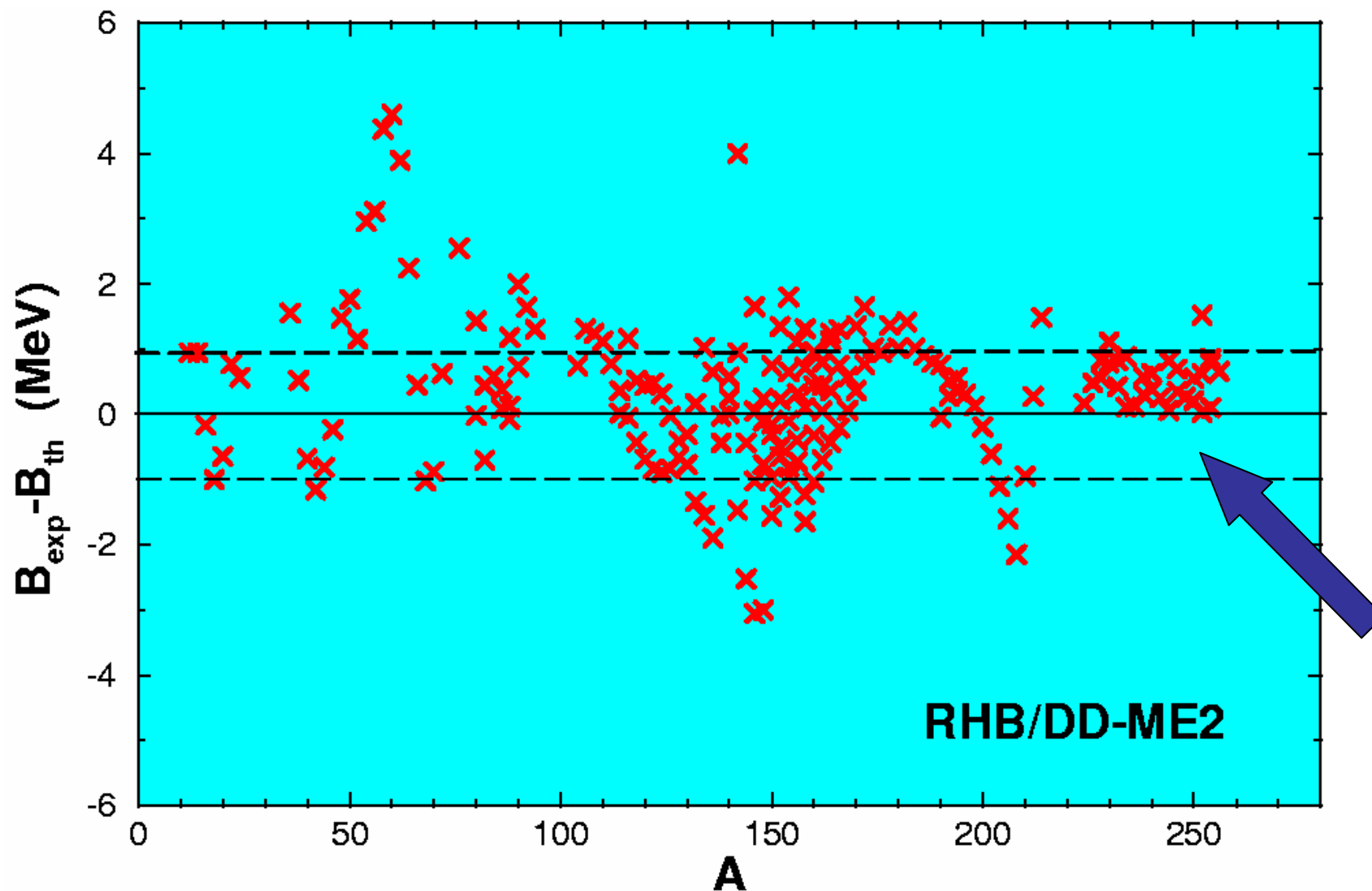
empirical values:

$$\begin{aligned} 30 \text{ MeV} &< a_4 < 34 \text{ MeV} \\ 2 \text{ MeV}/\text{fm}^3 &< p_0 < 4 \text{ MeV}/\text{fm}^3 \\ -200 \text{ MeV} &< \Delta K_0 < -50 \text{ MeV} \end{aligned}$$

	DD-ME1	NL3	NL1
a_4 (MeV)	33.1	37.9	43.7
p_0 (MeV/ fm^3)	3.26	5.92	7.0
ΔK_0 (MeV)	-128.5	52.1	67.3


Furnstahl, NPA 705 (2002) 85

rms-deviations: masses: $\Delta m = 900 \text{ keV}$
radii: $\Delta r = 0.015 \text{ fm}$



Excited States: Time dependence:

$$\delta \int dt \left\{ \langle \Phi(t) | i\partial_t | \Phi(t) \rangle - E[\hat{\rho}(t)] \right\} = 0$$



$$i\partial_t \hat{\rho} = [\hat{h}(\hat{\rho}) + \hat{f}, \hat{\rho}]$$

Rotational Motion:

$$\rho(t) = e^{-i\vec{\Omega}t \cdot \vec{j}} \rho_{\Omega} e^{i\vec{\Omega}t \cdot \vec{j}}$$

$$[h - \vec{\Omega} \cdot \vec{j}, \rho_{\Omega}] = 0$$


Vibrational Motion:


$$\hat{\rho}(t) = \hat{\rho}^{(0)} + \delta\hat{\rho}(t)$$

$$A, B \sim \delta^2 E / \delta\rho\delta\rho$$

ground-state density

$$\begin{pmatrix} A & B \\ -B^* & -A^* \end{pmatrix} \begin{pmatrix} X \\ Y \end{pmatrix} = \hbar\omega \begin{pmatrix} X \\ Y \end{pmatrix}$$

 $\delta\rho_{ph}$

 $\delta\rho_{hp}$

Cranked relativistic Hartree+Bogoliubov theory

CRHB equations for the fermions in the rotating frame

$$\begin{pmatrix} \hat{h}_D - \lambda_\tau - \Omega \hat{J}_x & \hat{\Delta} \\ -\hat{\Delta}^* & \hat{h}_D + \lambda_\tau + \Omega \hat{J}_x \end{pmatrix} \begin{pmatrix} U_k \\ V_k \end{pmatrix} = E_k \begin{pmatrix} U_k \\ V_k \end{pmatrix}$$

Coriolis term

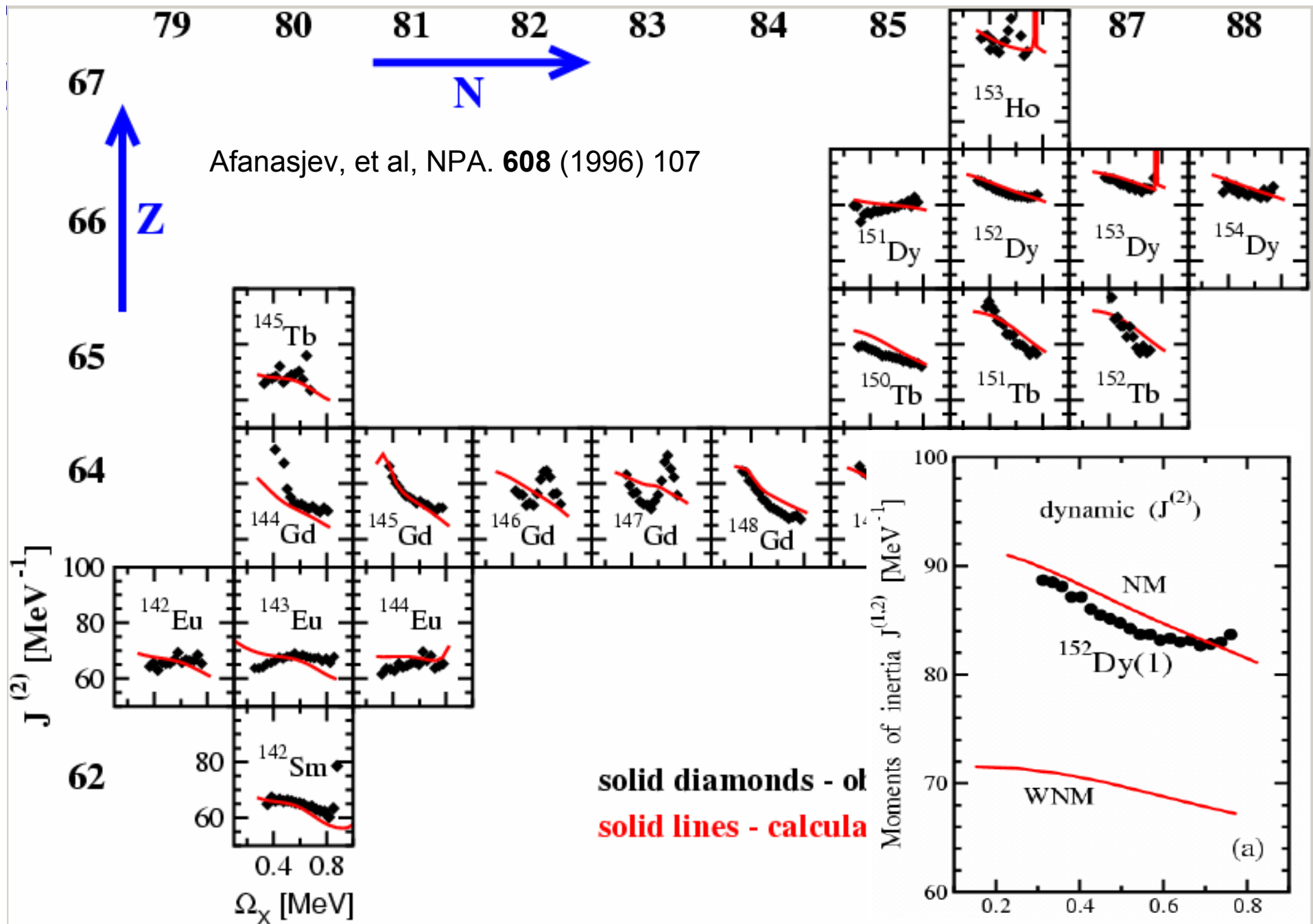
$$\hat{h}_D = \alpha(-i\vec{\nabla} - \vec{V}(\vec{r})) + V_0(\vec{r}) + \beta(m - S(\vec{r}))$$

Magnetic potential

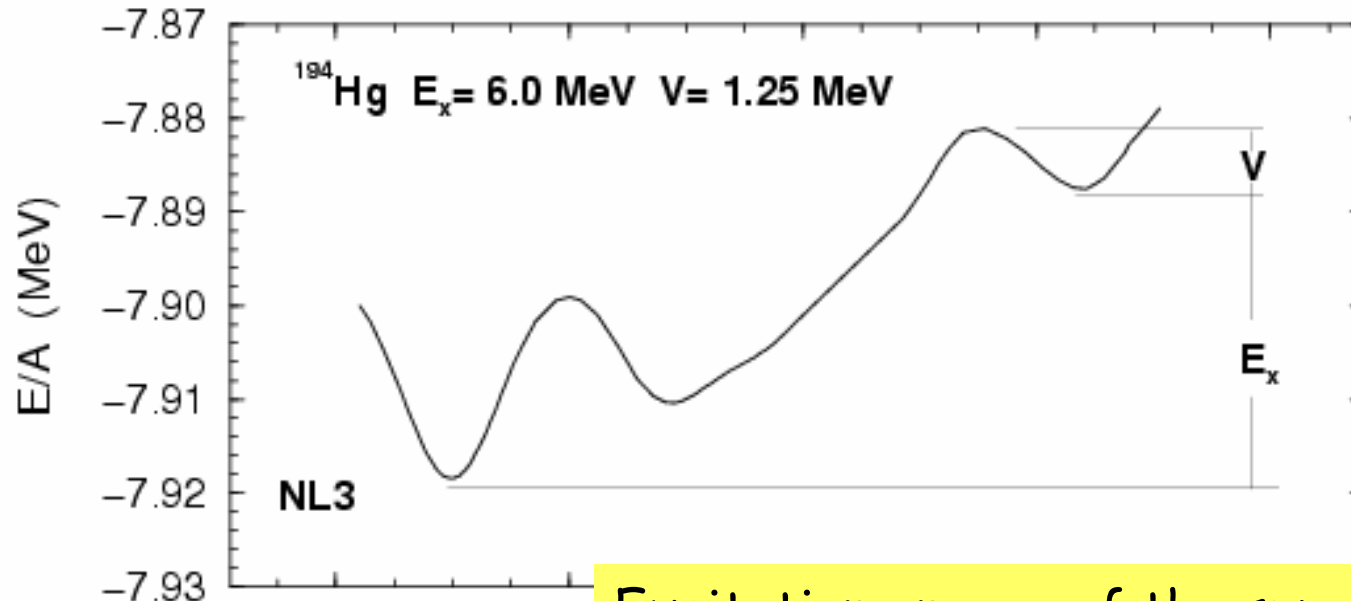
$$\vec{V}(\vec{r}) = g_\omega \vec{\omega}(\vec{r}) + g_\rho \tau_3 \vec{\rho}(\vec{r}) + e \frac{1 - \tau_3}{2} \vec{A}(\vec{r})$$

space-like components of vector mesons
behaves in Dirac equation like a magnetic field

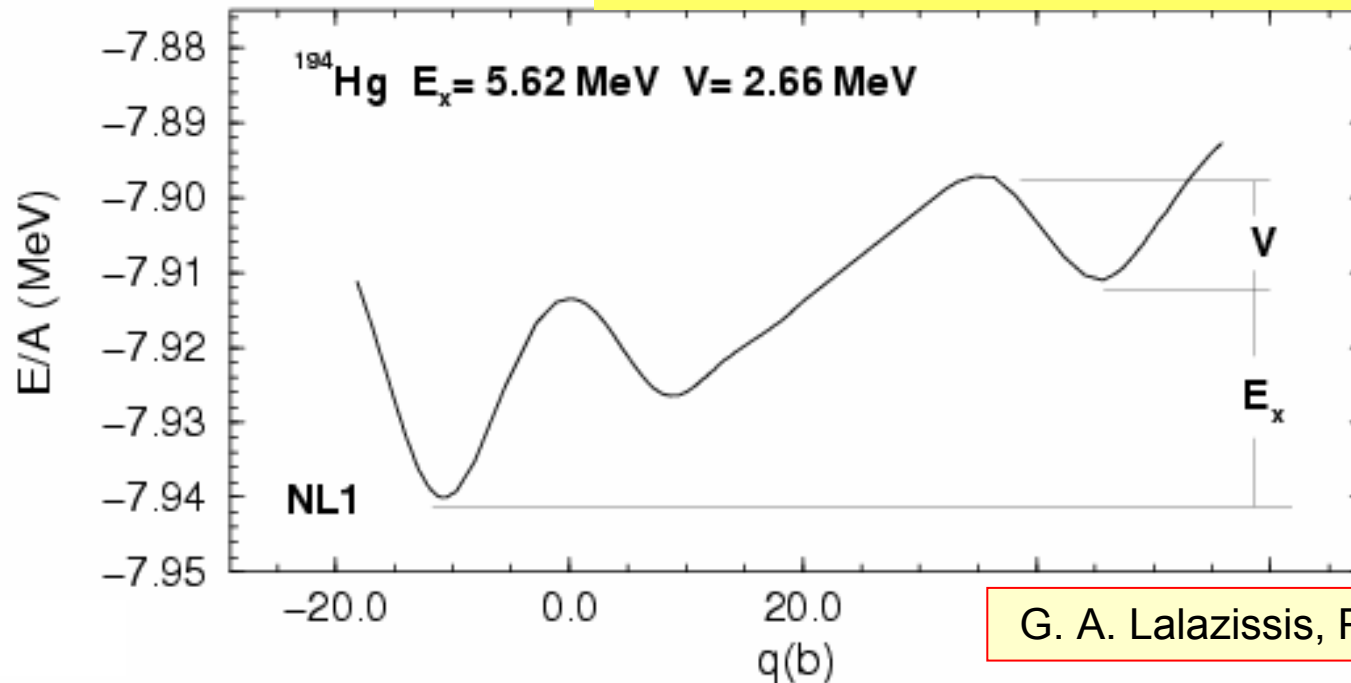
Nuclear magnetism



¹⁹⁴Hg



Excitation energy of the superdeformed minimum:

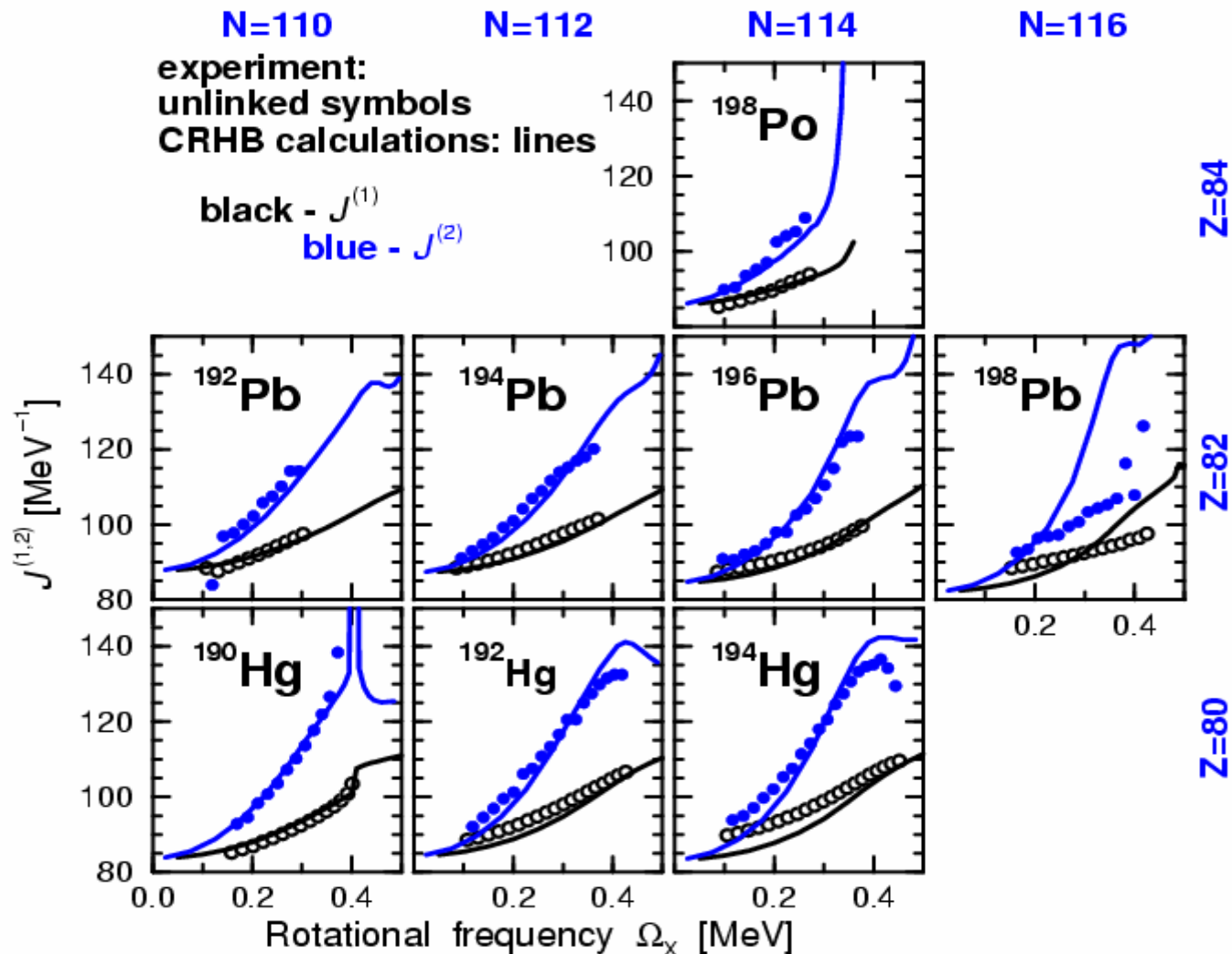


Exp:	$E_x = 6.02$
NL3:	$= 6.0$
NL1:	$= 5.6$
Gogny:	$= 6.9$
Skyrme:	$= 5.0$
WS:	$= 4.6$

G. A. Lalazissis, P. Ring, PLB 427 (1998) 225

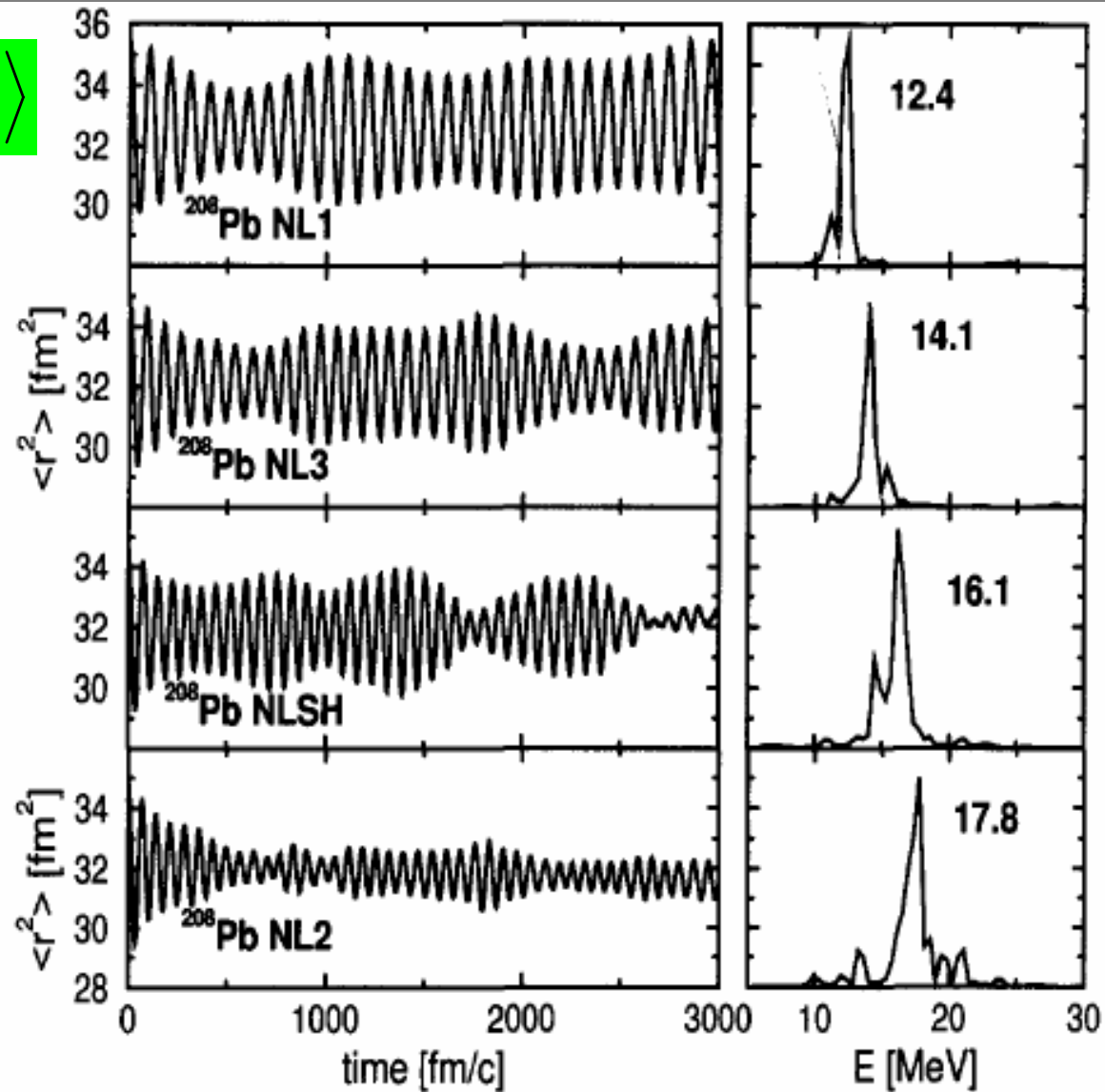
Cranked RHB

A.V.Afanasjev, P. Ring, J. König
 Phys. Rev. C60 (1999) 051303; Nucl. Phys. A 676 (2000) 196



Time-dependent RMF: breathing mode, ^{208}Pb :

$$\langle \Phi(t) | r^2 | \Phi(t) \rangle$$



$$K_\infty = 211$$

$$K_\infty = 271$$

$$K_\infty = 355$$

Pb: lowlying discrete spectrum

Calculated and experimental excitation energies, and $B(EL)$ values for the low-lying vibrational states in ^{208}Pb

L^π	E_{th}	E_{exp}	$B(EL)_{\text{th}}$	$B(EL)_{\text{exp}}$
3^-	2.76	2.61	499×10^3	$(540 \pm 30) \times 10^3$
5^-	3.26	3.71	201×10^6	330×10^6
2^+	4.99	4.07	2816	2965
4^+	4.95	4.32	998×10^4	1287×10^4

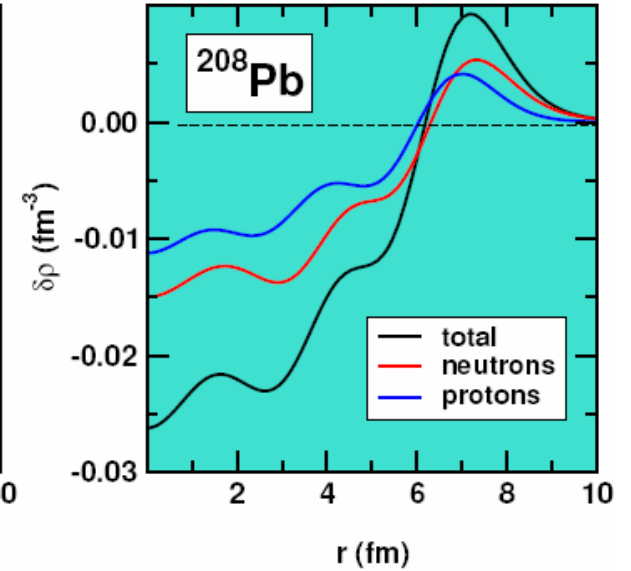
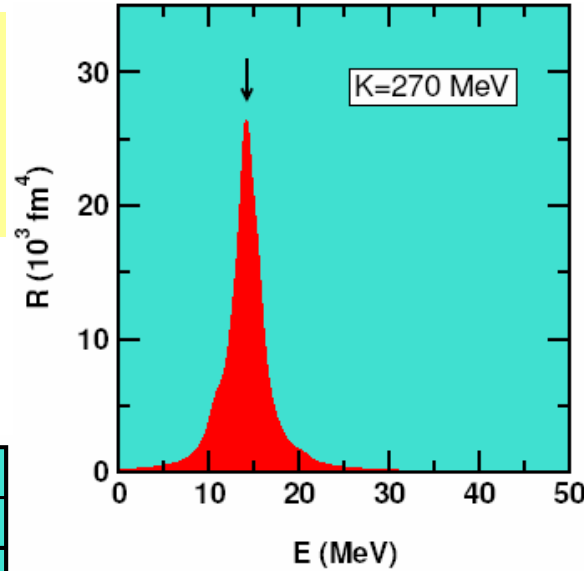
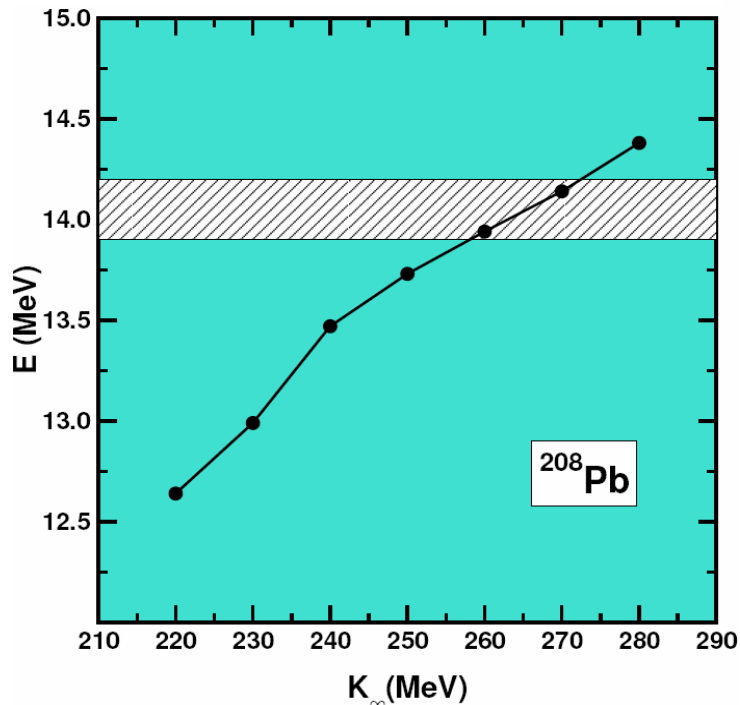
The calculated values correspond to NL3 parameterization, the data are from Ref. [29]. Energies are in MeV, $B(EL)$ values in $e^2 \text{fm}^{2L}$.

Z.Y. Ma, A. Wandelt et al., NPA 694 (2001) 249

Isoscalar Giant Monopole Resonance: IS-GMR

The ISGMR represents the essential source of experimental information on the nuclear incompressibility

$$K_0 = p_f^2 \left. \frac{d^2 E/A}{dp_f^2} \right|_{p_{f0}}$$



constraining the nuclear matter compressibility

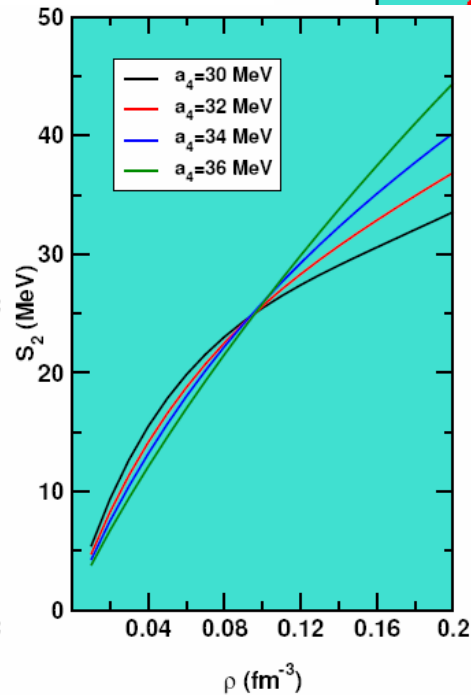
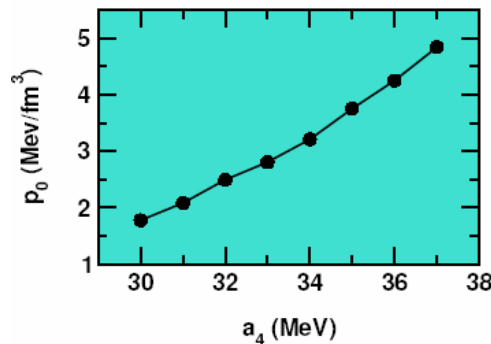
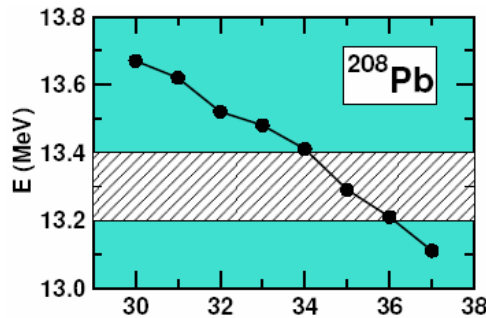
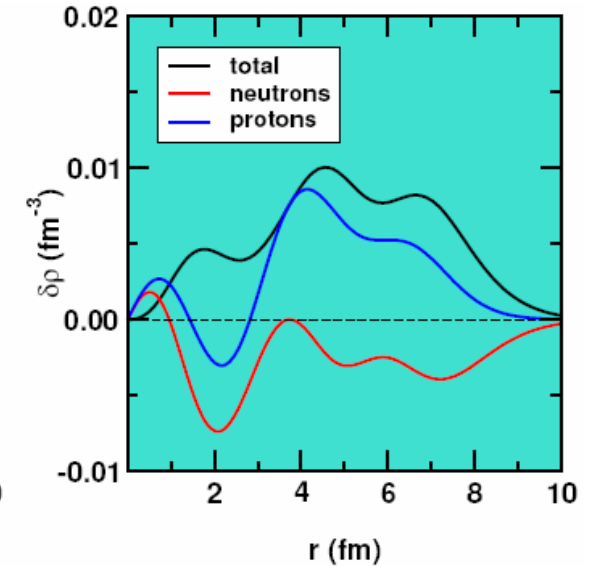
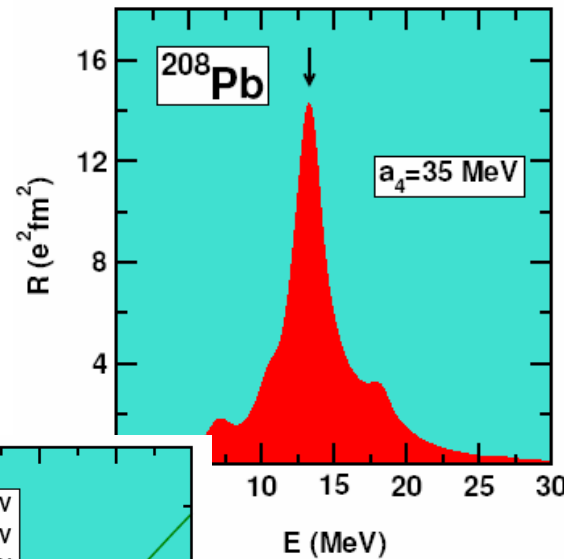


RMF models reproduce the experimental data only if

$$250 \text{ MeV} \leq K_0 \leq 270 \text{ MeV}$$

Isovector Giant Dipole Resonance: IV-GDR

the IVGDR represents one of the sources of experimental informations on the nuclear matter symmetry energy



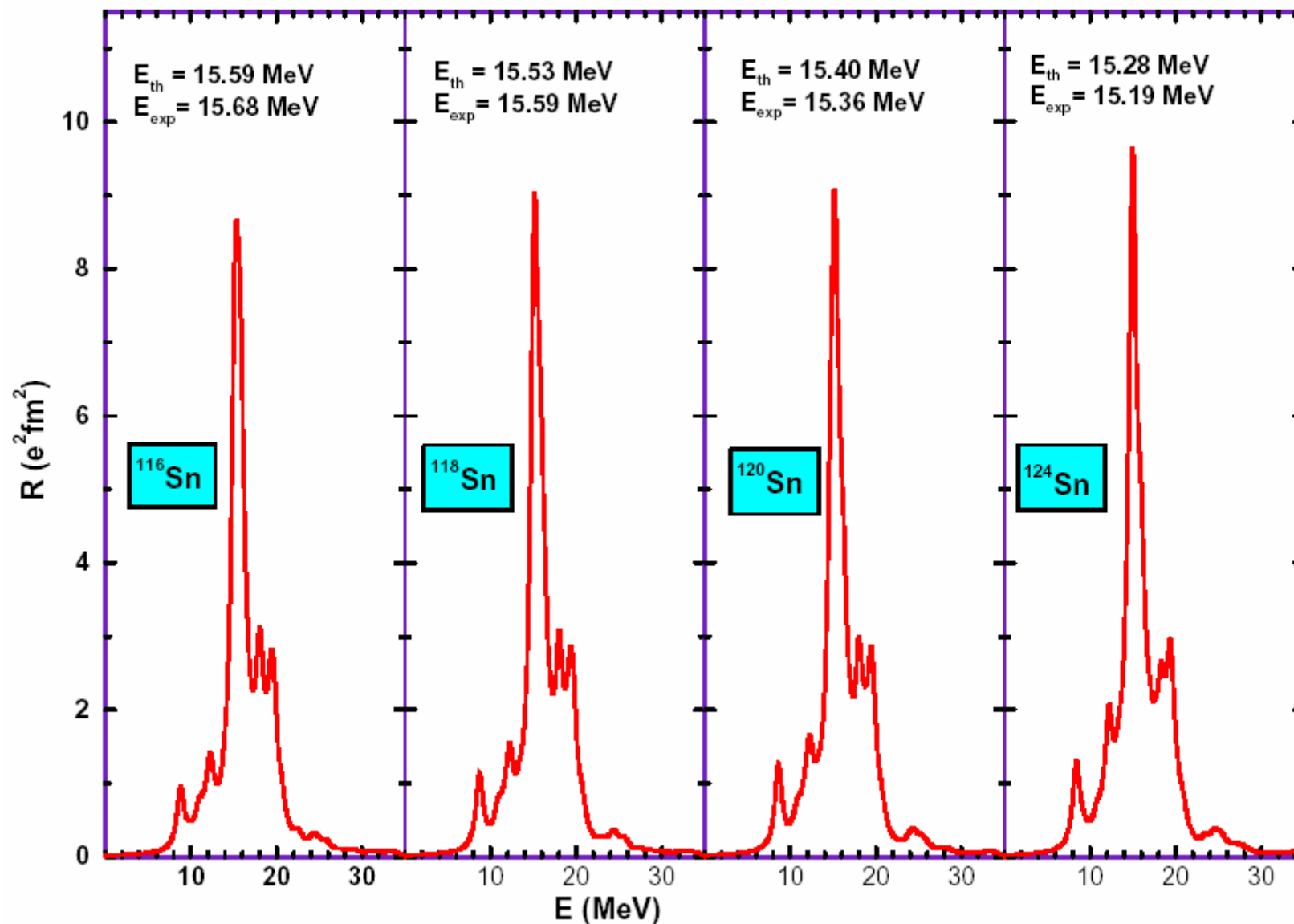
constraining the nuclear matter symmetry energy

the position of IVGDR is reproduced if

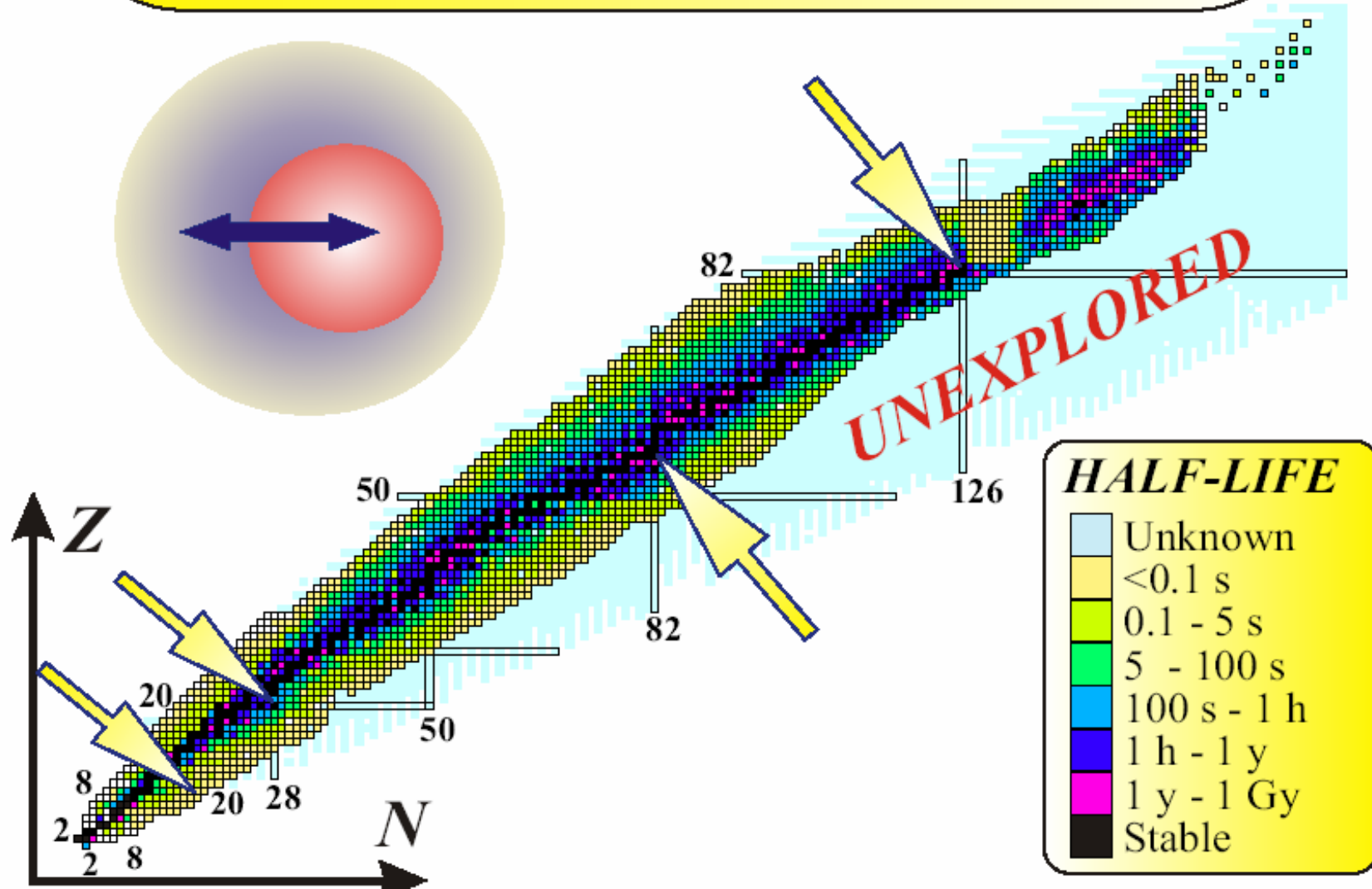
$$34 \text{ MeV} \leq a_4 \leq 36 \text{ MeV}$$

IV-GDR in Sn-isotopes

DD-ME2



Experimental indications of the soft dipole mode



Photoneutron Cross Sections for Unstable Neutron-Rich Oxygen Isotopes

A. Leistenschneider, T. Aumann, K. Boretzky, D. Cortina, J. Cub, U. Datta Pramanik, W. Dostal, Th. W. Elze, H. Emling, H. Geissel, A. Grünschloß, M. Hellstr, R. Holzmann, S. Ilievski, N. Iwasa, M. Kaspar, A. Kleinböhl, J. V. Kratz, R. Kulesa, Y. Leifels, E. Lubkiewicz, G. Münzenberg, P. Reiter, M. Rejmund, C. Scheidenberger, C. Schlegel, H. Simon, J. Stroth, K. Sümmerer, E. Wajda, W. Wälus, and S. Wan
Institut für Kernphysik, Johann Wolfgang Goethe-Universität, D-60486 Frankfurt, Germany
Gesellschaft für Schwerionenforschung (GSI), D-64291 Darmstadt, Germany
Institut für Kernchemie, Johannes Gutenberg-Universität, D-55099 Mainz, Germany
Institut für Kernphysik, Technische Universität, D-64289 Darmstadt, Germany
Instytut Fizyki, Uniwersytet Jagielloński, PL-30-059 Kraków, Poland
Sektion Physik, Ludwig-Maximilians-Universität, D-85748 Garching, Germany
 (Received 19 December 2000)

The dipole response of stable and unstable neutron-rich oxygen nuclei of masses $A=17$ to $A=22$ has been investigated experimentally utilizing electromagnetic excitation in heavy-ion collisions at beam energies about 600 MeV/nucleon. A kinematically complete measurement of the neutron decay channel in inelastic scattering of the secondary beam projectiles from a Pb target was performed. Differential electromagnetic excitation cross sections $d\sigma/dE$ were derived up to 30 MeV excitation energy. In contrast to stable nuclei, the deduced dipole strength distribution appears to be strongly fragmented and systematically exhibits a considerable fraction of low-lying strength.

DOI: 10.1103/PhysRevLett.86.2560

The study of the response of clear or electromagnetic field is the properties of the nuclear excitation energies above the par response of stable nuclei is distributions of various multipolarities, the giant resonance strength stable to exotic weakly bound n-to-proton ratios is presently unknown. For neutron-rich nuclei, pronounced effects, in particular strength towards lower excitation giant resonance region. The properties depend strongly on the effective interactions. In turn, measurements of response of exotic nuclei can depend on the isospin dependent nucleon-nucleon interaction [7]. Systematic experimental information on response of exotic nuclei, however, is scarce. For some light halo nuclei, low-lying strength is observed in electromagnetic excitation [8–11]. For the one-neutron halo nucleus ^{22}O [11], the observed dipole strength was interpreted as a threshold effect, involving nonvalence neutron into the continuum. For ^{16}O and ^{22}O , a coherent dipole resonance against the core was observed. The appearance of a collective dipole resonance (GDR) [19].

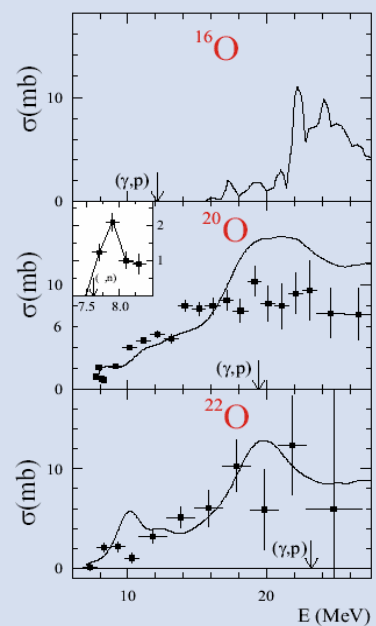


FIG. 2. Photoneutron cross sections for ^{16}O (upper panel) and for the unstable isotopes $^{20,22}\text{O}$ (lower panels) as extracted from the measured electromagnetic excitation cross section (symbols). The inset displays the cross section for near the neutron threshold on an expanded energy scale. The thresholds for decay channels involving protons (which were not observed in the present experiment) are indicated by arrows.

5442

0031-9007

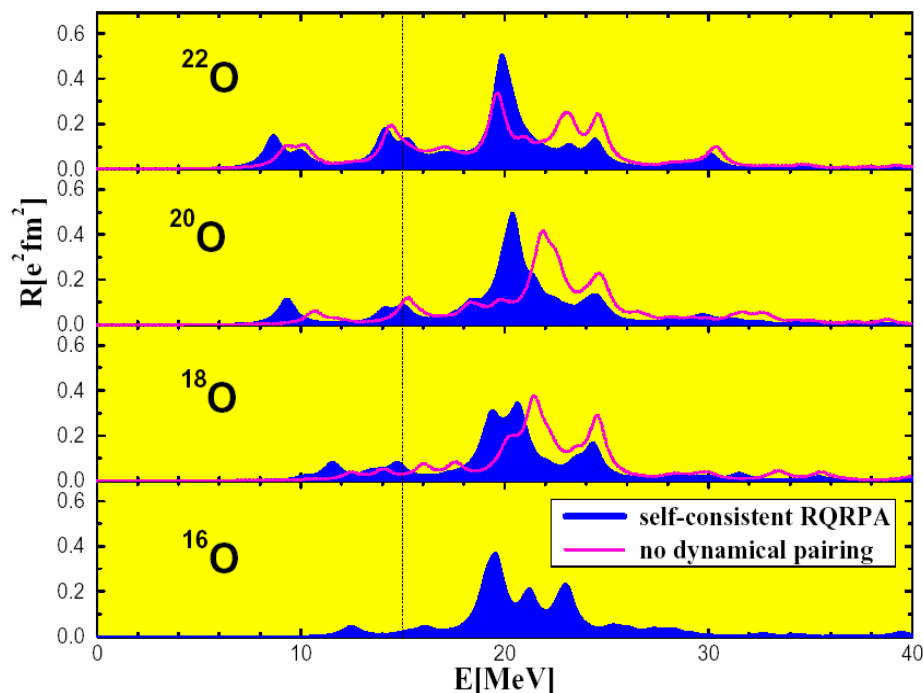
25.60.-t, 27.20.+n

giant resonance, may arise from neutrons vibrate against the core. It is interesting to note that a systematic fragmentation of the dipole strength in neutron-rich nuclei is observed. From physical aspects, e.g., calculations in the ^{16}O -process of the [21]. The giant resonance and lower lying resonances and lower lying strength investigated systematically for all neutron-rich oxygen isotopes. For strongly bound doubly magic nuclei, one may expect a dipole strength from the inert ^{16}O core. For ^{20}O , the first neutron is 7–8 MeV for ^{20}O , and about 4 MeV for the ^{22}O and about 16 MeV for ^{22}O . Thus the dipole strength might be good candidates for investigation. We use the electromagnetic excitation cross section for high targets. Similar to ^{16}O , it is mostly sensitive to electric dipole (E1) contributions. For $^{20,22}\text{O}$, the weighted sum rule for E1 strength is arbitrarily at an excitation energy of 10 MeV. The electromagnetic excitation cross section (mb), respectively (calculated as 10 mb), was demonstrated for a Pb target. It was demonstrated that the dipole strength distribution is qualitatively from a measurement of the electromagnetic dissociation cross section parameters by applying the sum rule [24]. The high secondary beam energy (600 MeV/nucleon) allows for the

Physical Society

Evolution of IV dipole strength in Oxygen isotopes

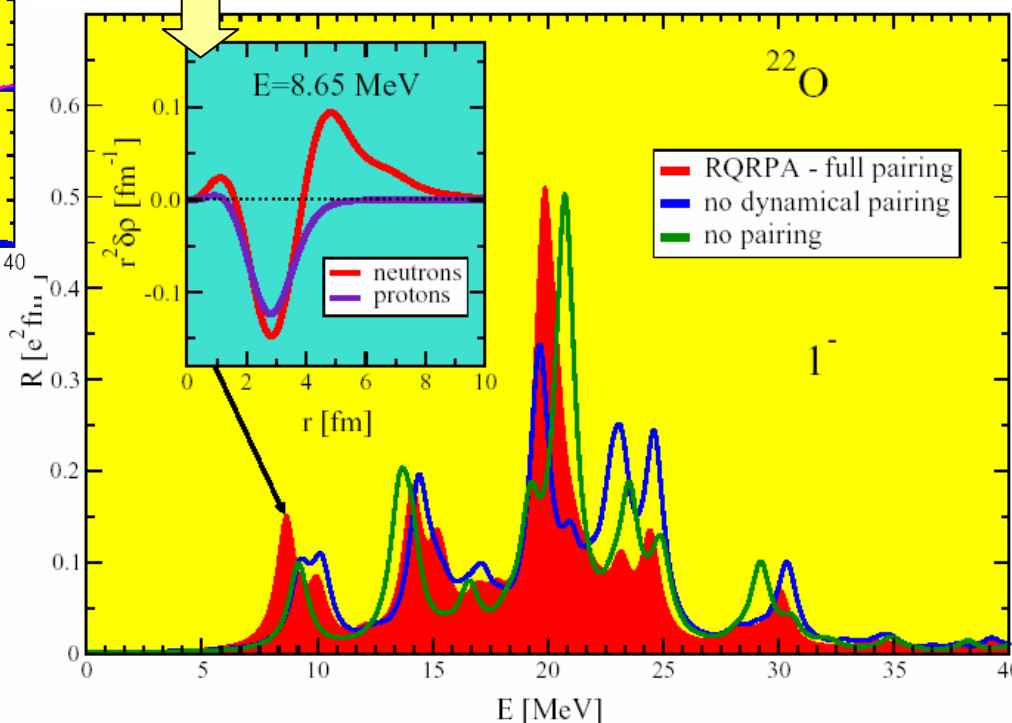
RHB + RQRPA calculations with the NL3 relativistic mean-field plus D15 Gogny pairing interaction.

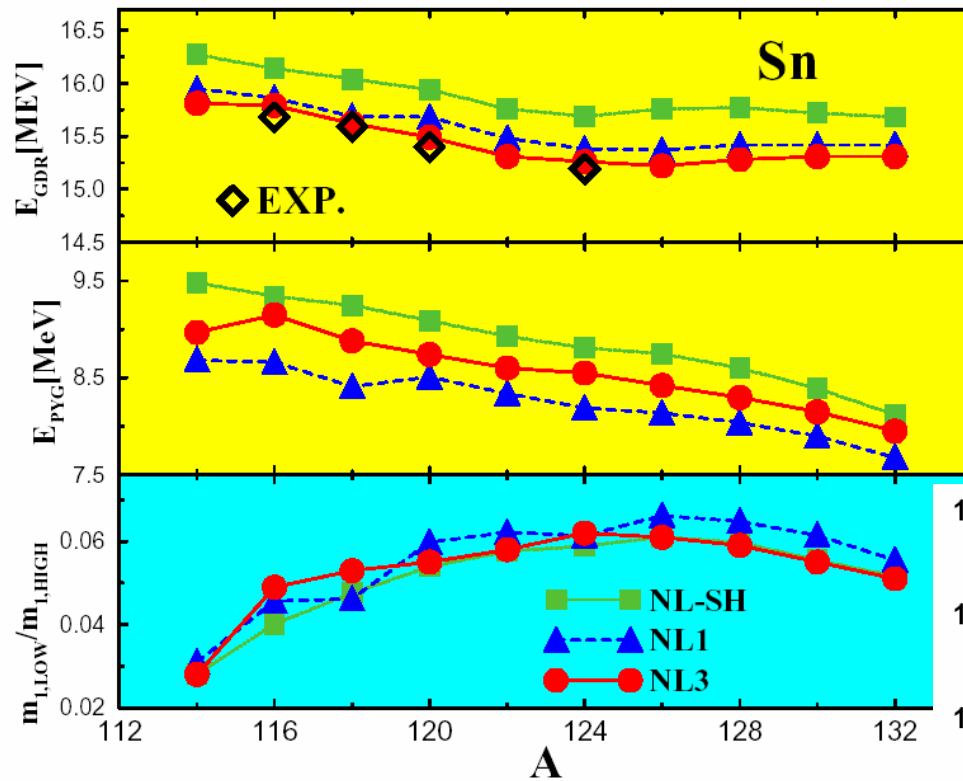


What is the structure of low-lying strength below 15 MeV ?

Effect of pairing correlations on the dipole strength distribution

Transition densities





Mass dependence of GDR and Pygmy dipole states in Sn isotopes. Evolution of the low-lying strength.

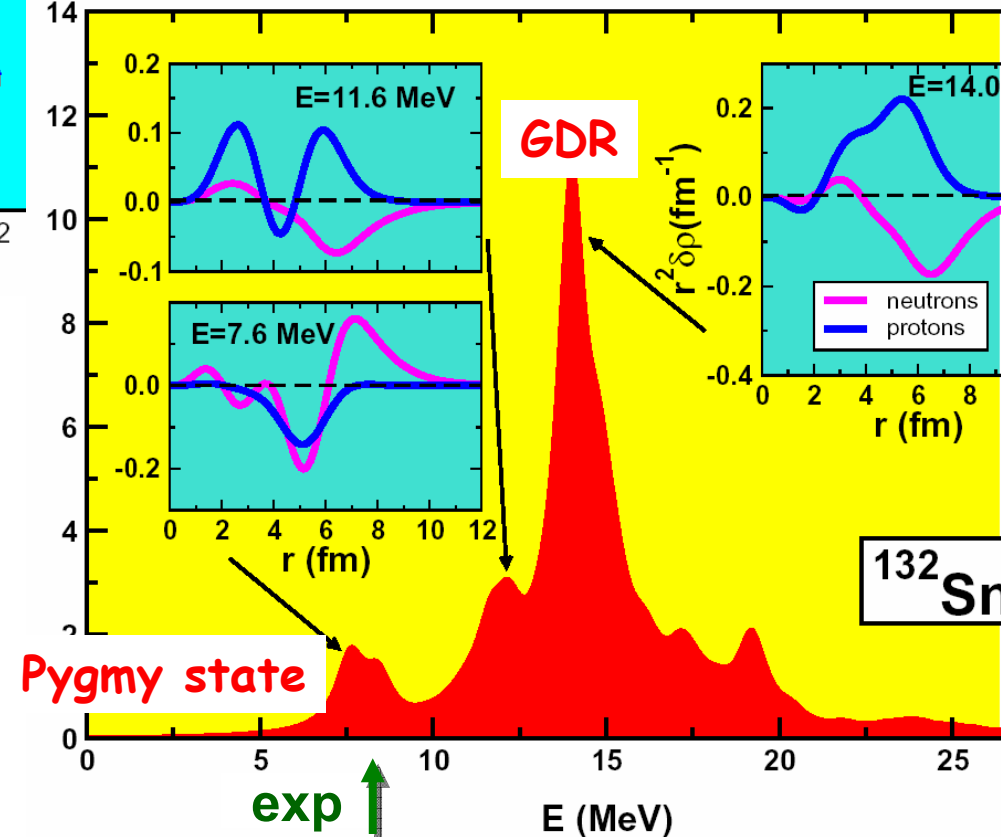
Isvector dipole strength in ^{132}Sn .

Nucl. Phys. A692, 496 (2001)

Distribution of the **neutron particle-hole configurations** for the peak at 7.6 MeV (1.4% of the EWSR)

^{132}Sn at 7.6 MeV

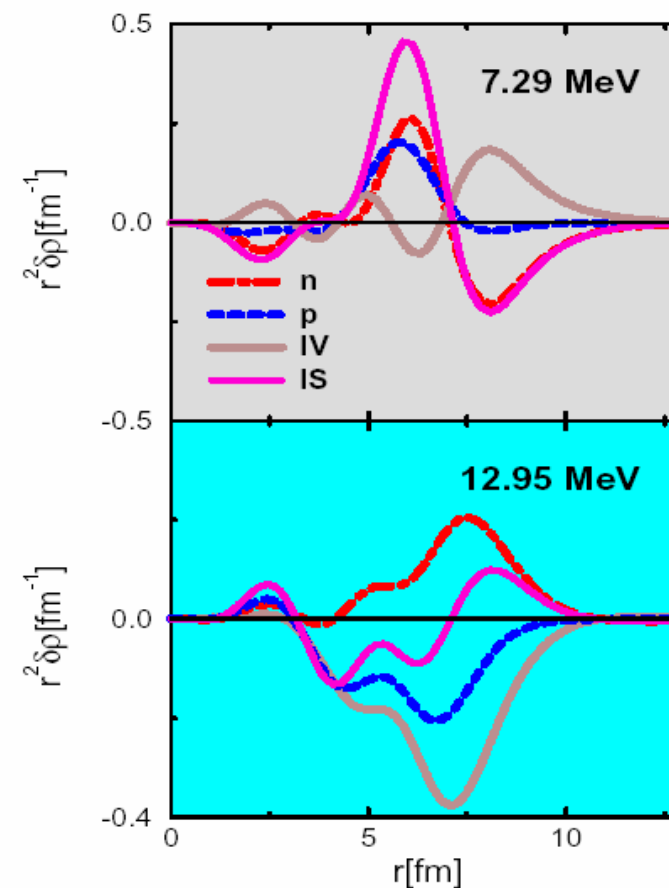
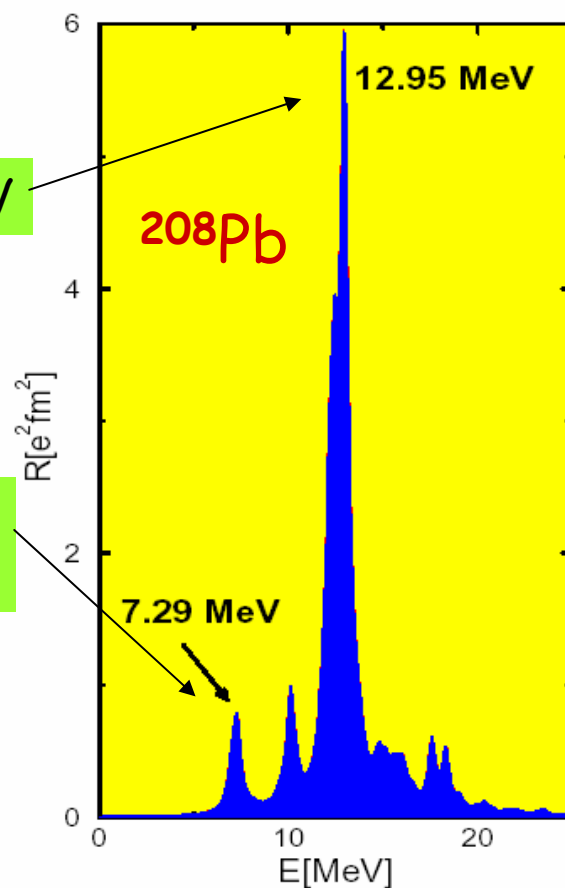
- 28.2% $2d_{3/2} \rightarrow 2f_{5/2}$
- 21.9% $2d_{5/2} \rightarrow 2f_{7/2}$
- 19.7% $2d_{3/2} \rightarrow 3p_{1/2}$
- 10.5% $1h_{11/2} \rightarrow 1i_{13/2}$
- 3.5% $2d_{5/2} \rightarrow 3p_{3/2}$
- 1.9% $1g_{7/2} \rightarrow 2f_{5/2}$
- 1.5% $1g_{7/2} \rightarrow 1h_{9/2}$
- 0.6% $1g_{7/2} \rightarrow 2f_{7/2}$
- 0.6% $2d_{3/2} \rightarrow 3p_{3/2}$



IV Dipole Strength for ^{208}Pb and transition densities
for the peaks at 7.29 MeV and 12.95 MeV *PRC 63, 047301 (2001)*

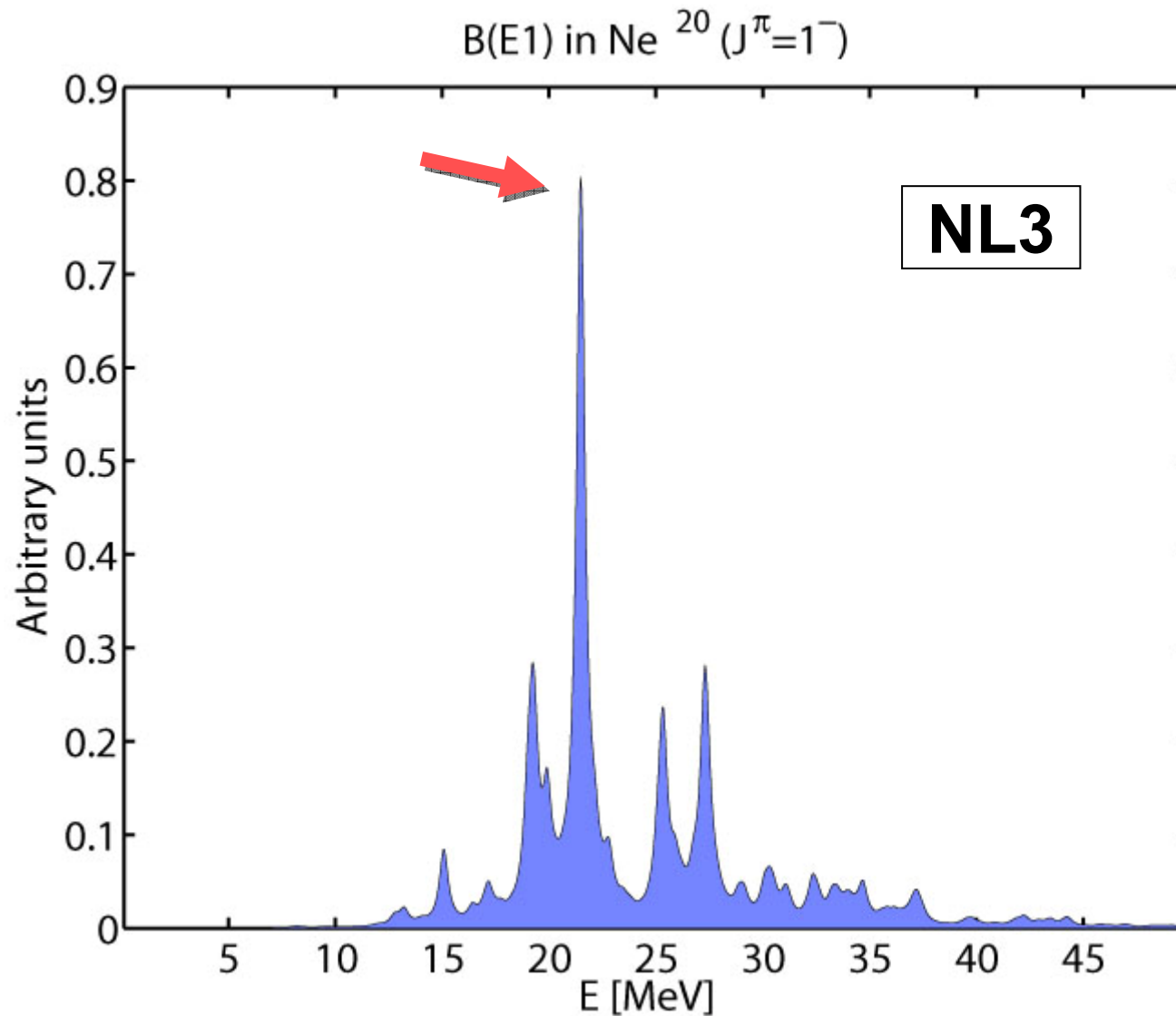
Exp GDR at 13.3 MeV

Exp PYGMY centroid
at 7.37 MeV

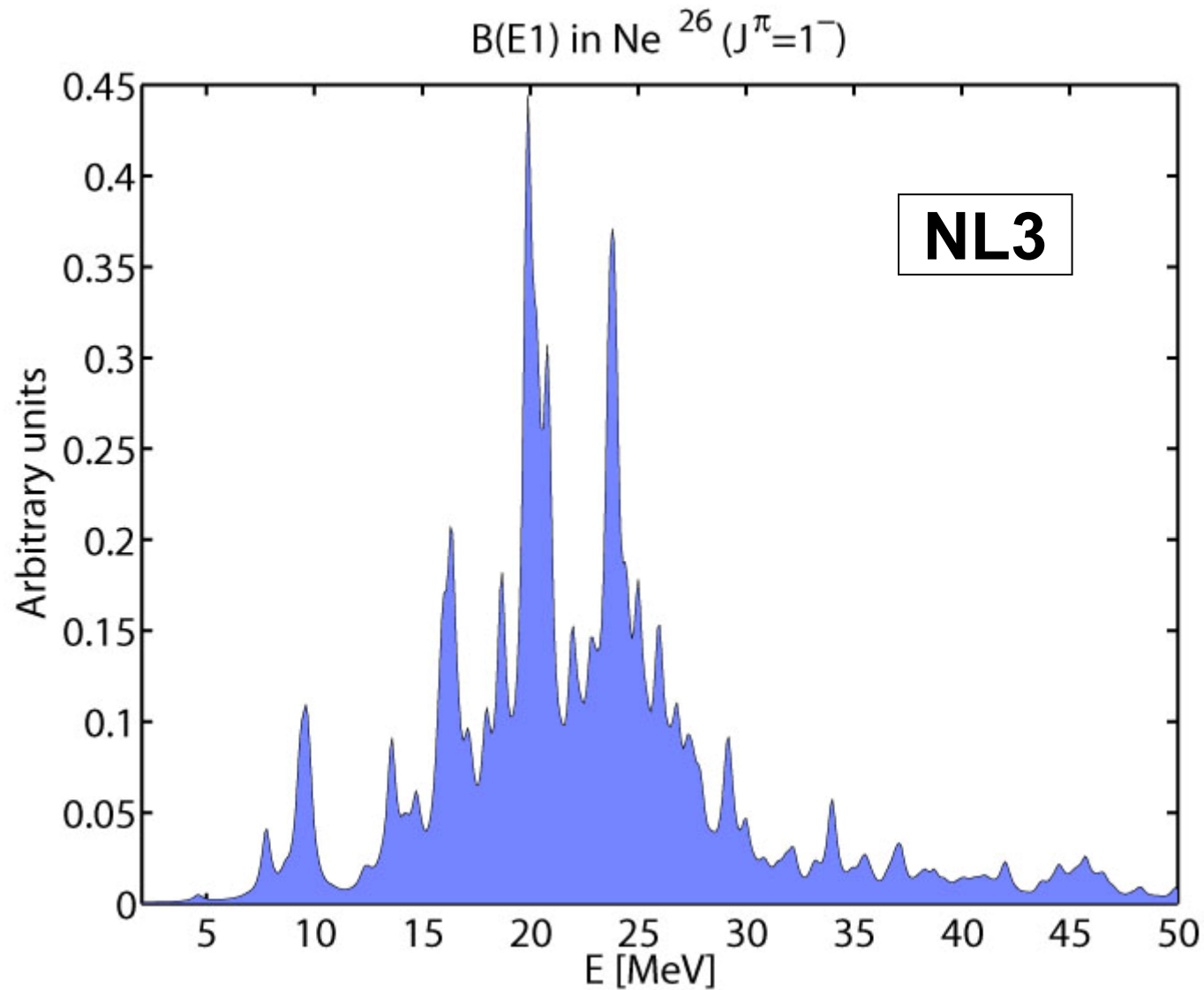


In heavier nuclei low-lying dipole states appear that are characterized by a more distributed structure of the RQRPA amplitude.

E1-strength distribution in deformed ^{20}Ne

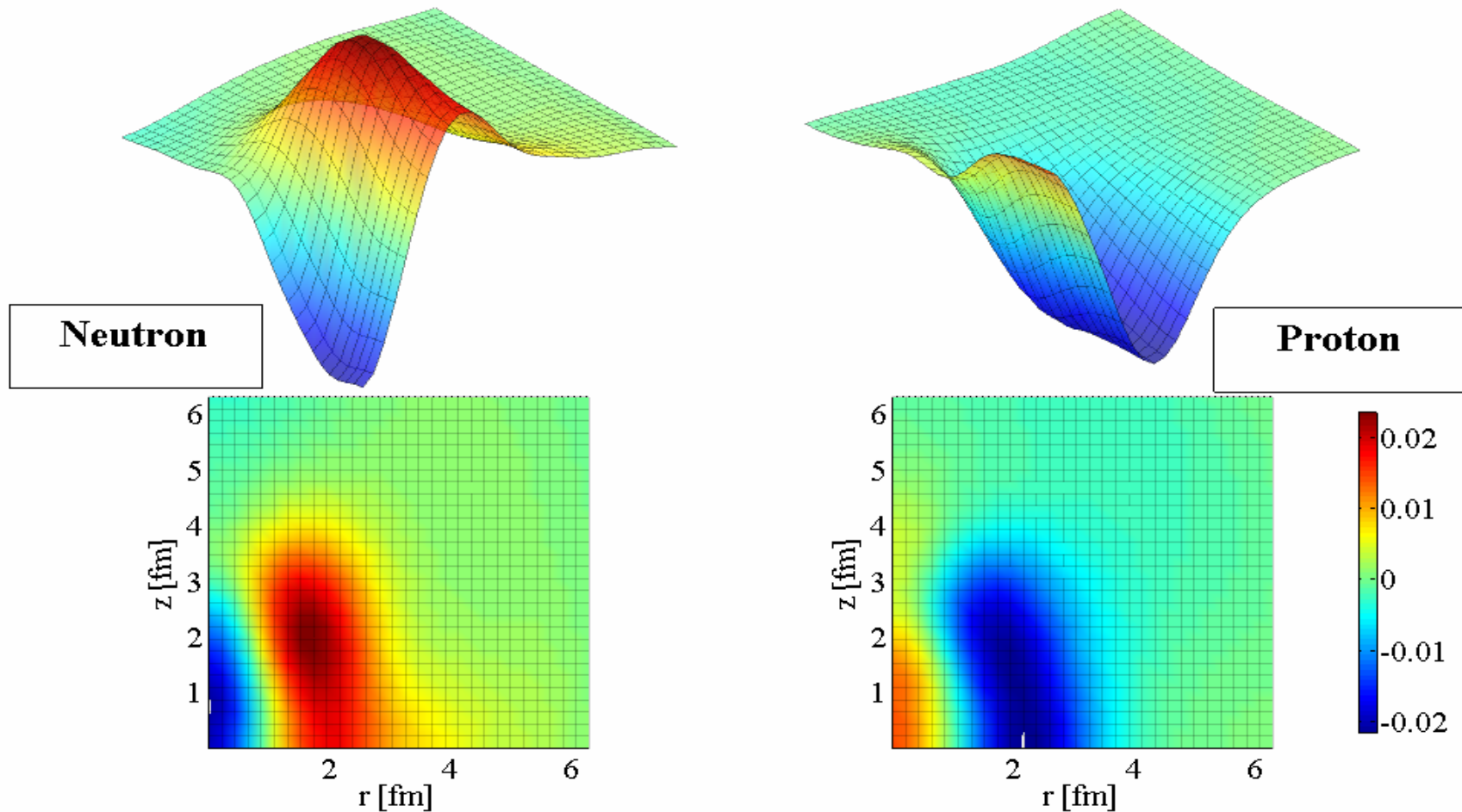


IV-E1-Resonances in deformed ^{26}Ne

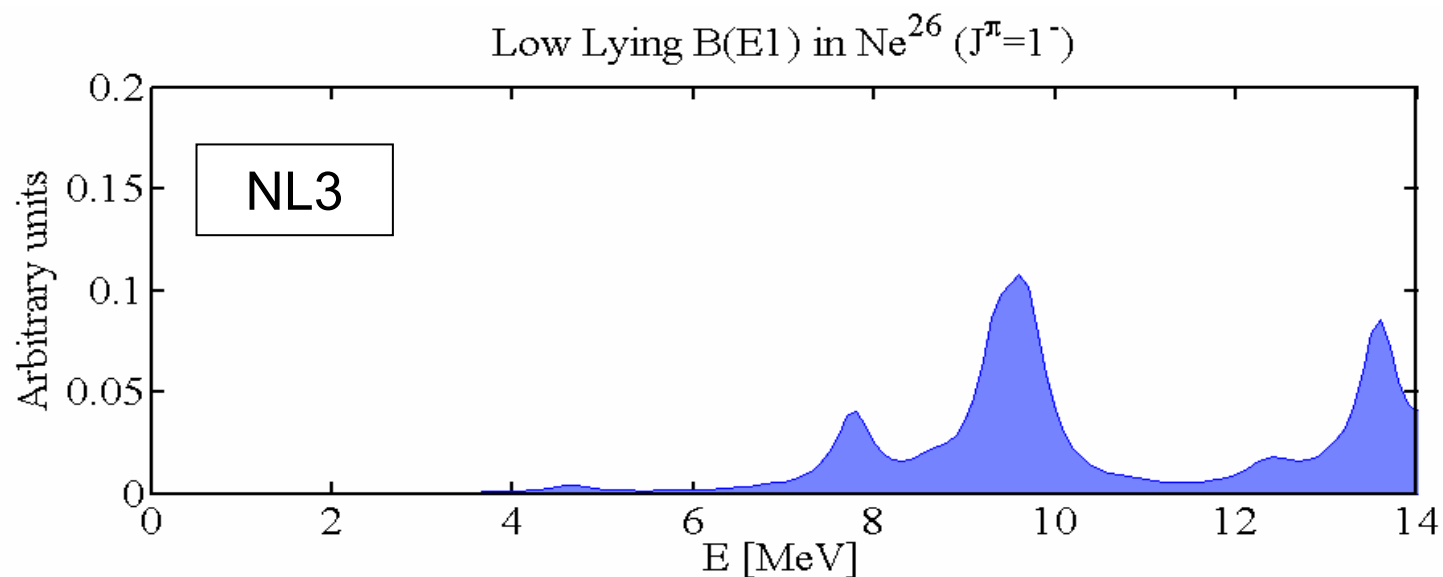
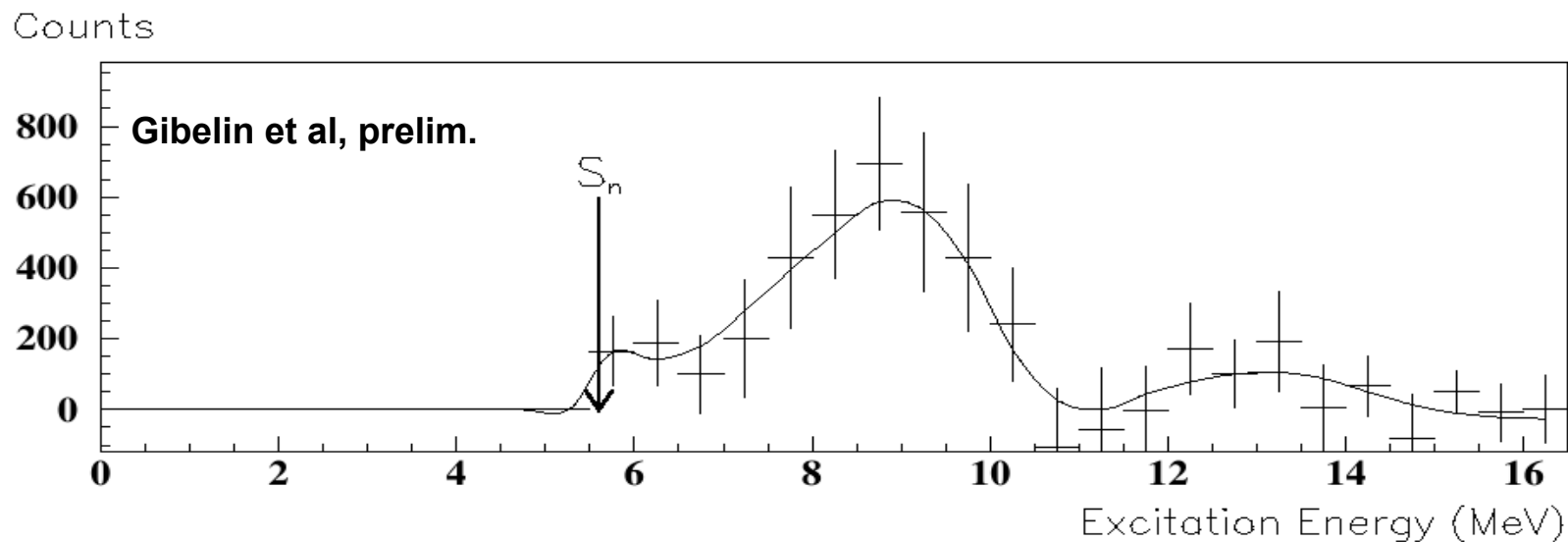


Transition density of the upper E1-peak

E1 Ne^{20} Transition Density, Peak at 21.4 MeV



Pygmy-Resonance in deformed ^{26}Ne



Isoscalar dipole compression - toroidal modes

Isoscalar GMR in spherical nuclei \rightarrow nuclear matter compression modulus K_{nm} .

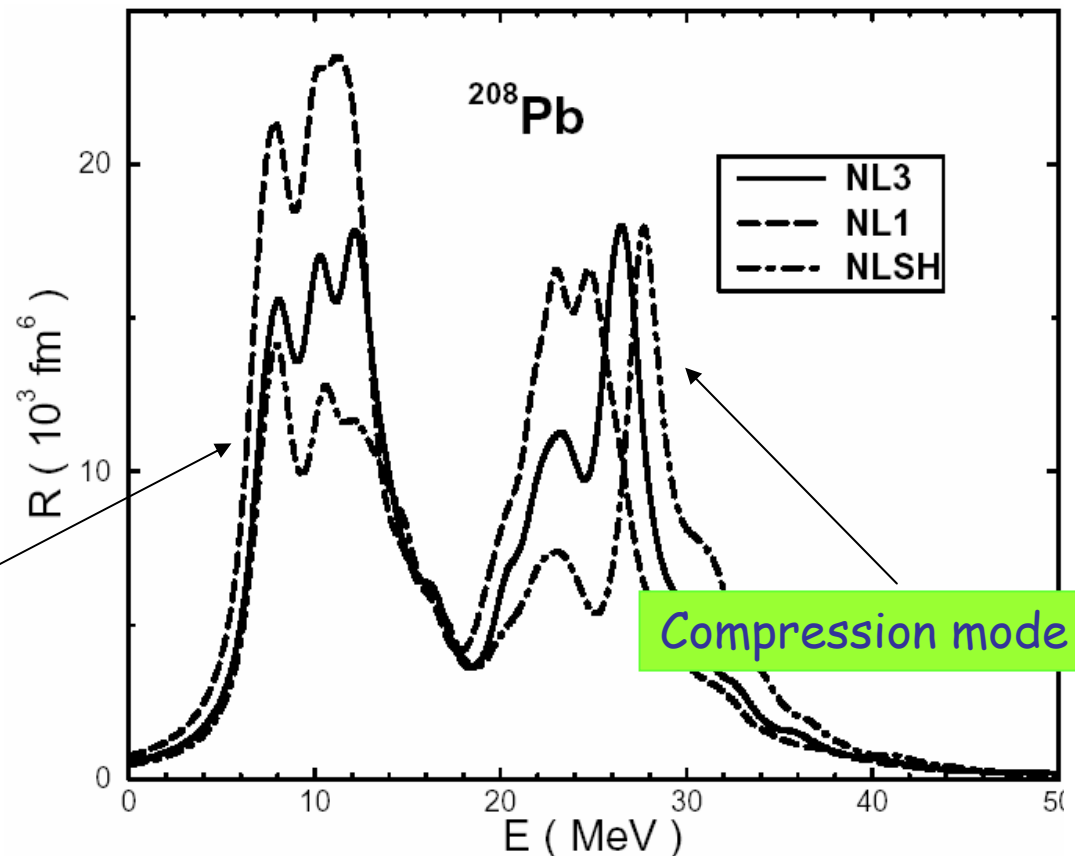
Giant isoscalar dipole oscillations \rightarrow additional information on the nuclear incompressibility.

$$\hat{Q}_{1\mu}^{T=0} = \sum_{i=1}^A \gamma_0 (r^3 - \eta r) Y_{1\mu}(\theta_i, \varphi_i)$$

ISGDR strength distributions
Effective interactions
with different K_{nm} .

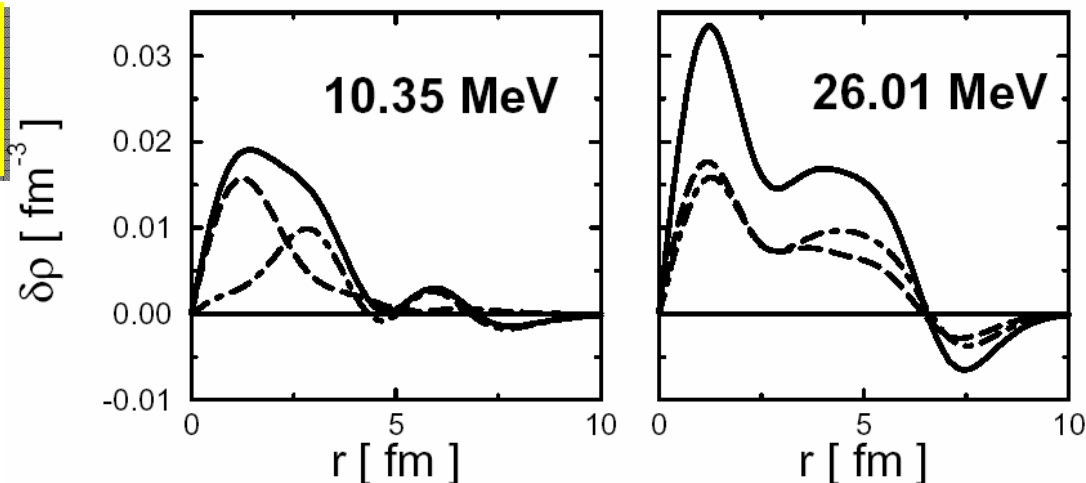
The low-energy strength
does not depend on K_{nm} !

Phys. Lett. B487, 334 (2000)



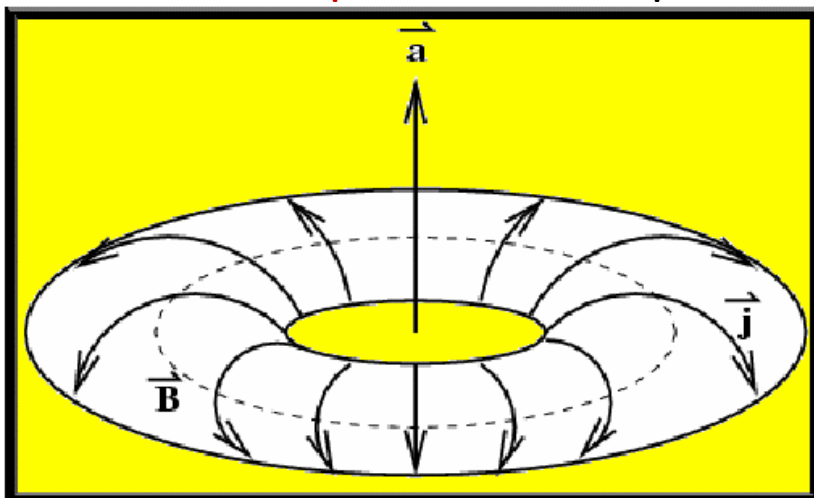
Toroidal motion:

ISGDR transition densities for ^{208}Pb (NL3 interaction)



multipole expansion of a four-current distribution:
 charge moments
 magnetic moments
 electric transverse moments \rightarrow toroidal moments

toroidal dipole moment: poloidal currents on a torus

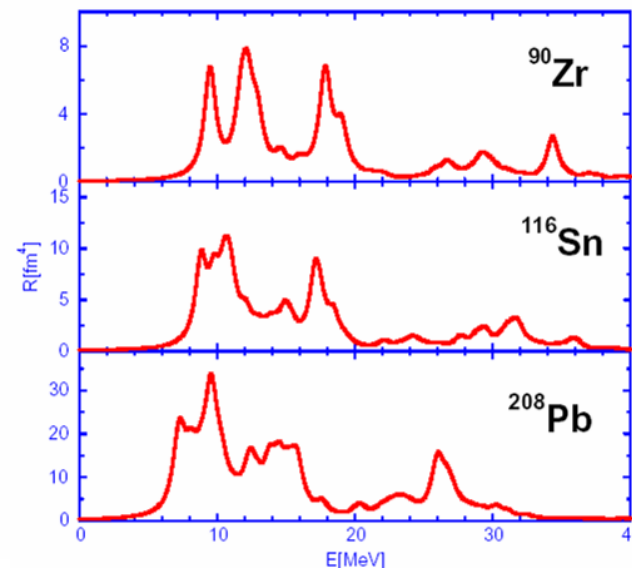
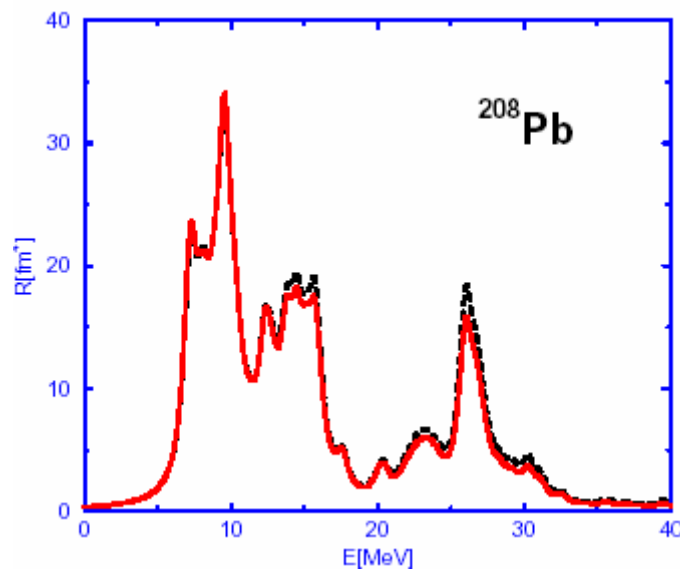


isoscalar toroidal dipole operator:

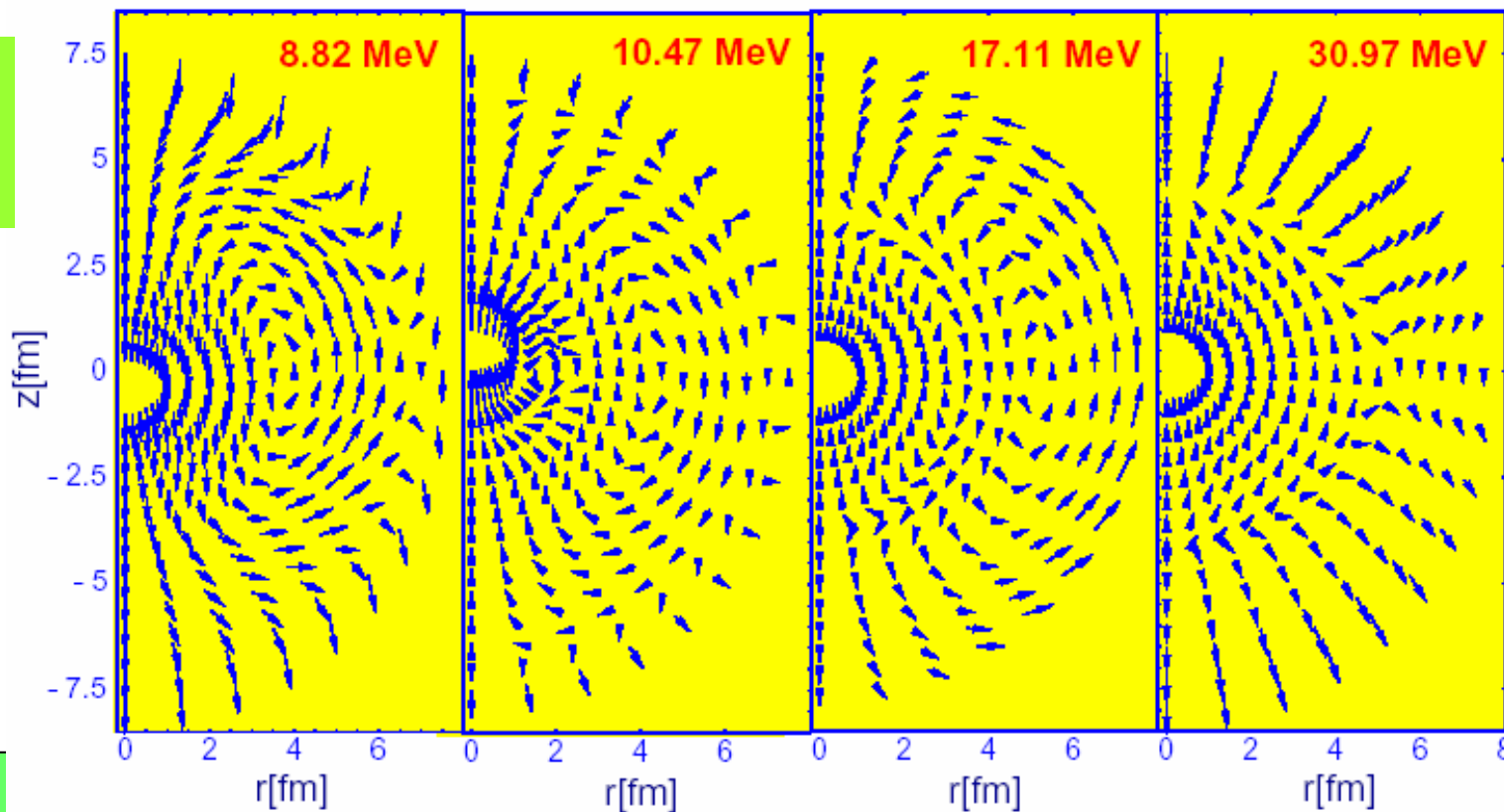
$$\hat{T}_{1\mu}^{T=0} \sim \int [r^2 \left(\vec{Y}_{10\mu}^* + \frac{\sqrt{2}}{5} \vec{Y}_{12\mu}^* \right) - \langle r^2 \rangle_0 \vec{Y}_{10\mu}^*] \cdot \vec{J}(\vec{r}) d^3r$$

Toroidal dipole strength distributions.

Vretenar, Paar, Niksic, Ring,
Phys. Rev. C65, 021301 (2002)



Velocity distributions in ^{116}Sn

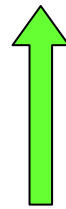


Spin-Isospin Resonances: IAR - GTR

Z, N

$Z+1, N-1$

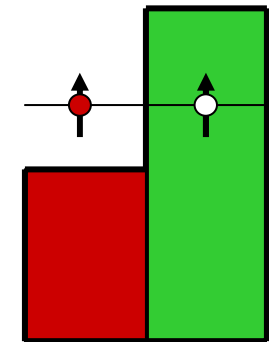
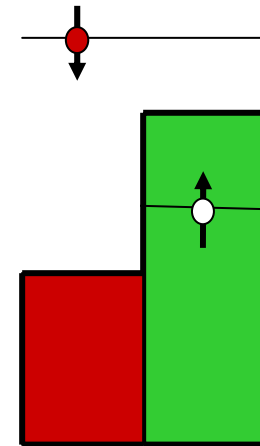
$$|\mathbf{GTR}\rangle = S_- T_- |\mathbf{Z}, N\rangle$$



spin flip σ

$$|\mathbf{Z}, N\rangle \longrightarrow |\mathbf{IAR}\rangle = T_- |\mathbf{Z}, N\rangle$$

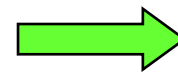
isospin flip τ



$$E_{\mathbf{GTR}} - E_{\mathbf{IAR}} \sim \Delta(l \cdot s) \sim \frac{dV}{dr} \sim \text{neutron skin} = r_n - r_p$$

Spin-Isospin Resonances: IAR and GTR

charge-exchange excitations



proton-neutron
relativistic QRPA

π and ρ -meson exchange
generate the spin-isospin
dependent interaction terms

$$\mathcal{L}_{\pi N} = -\frac{f_{\pi}}{m_{\pi}} \bar{\psi} \gamma_5 \gamma_{\mu} \partial^{\mu} \vec{\pi} \vec{\tau} \psi$$

the Landau-Migdal zero-range
force in the spin-isospin channel

$$V(1, 2) = g'_0 \left(\frac{f_{\pi}}{m_{\pi}} \right)^2 \vec{\tau}_1 \cdot \vec{\tau}_2 \Sigma_1 \cdot \Sigma_2 \delta(\mathbf{r}_1 - \mathbf{r}_2)$$

GAMOW-TELLER RESONANCE:

$$S=1 \quad T=1 \quad J^{\pi} = 1^{+}$$

$$(g'_0=0.55)$$

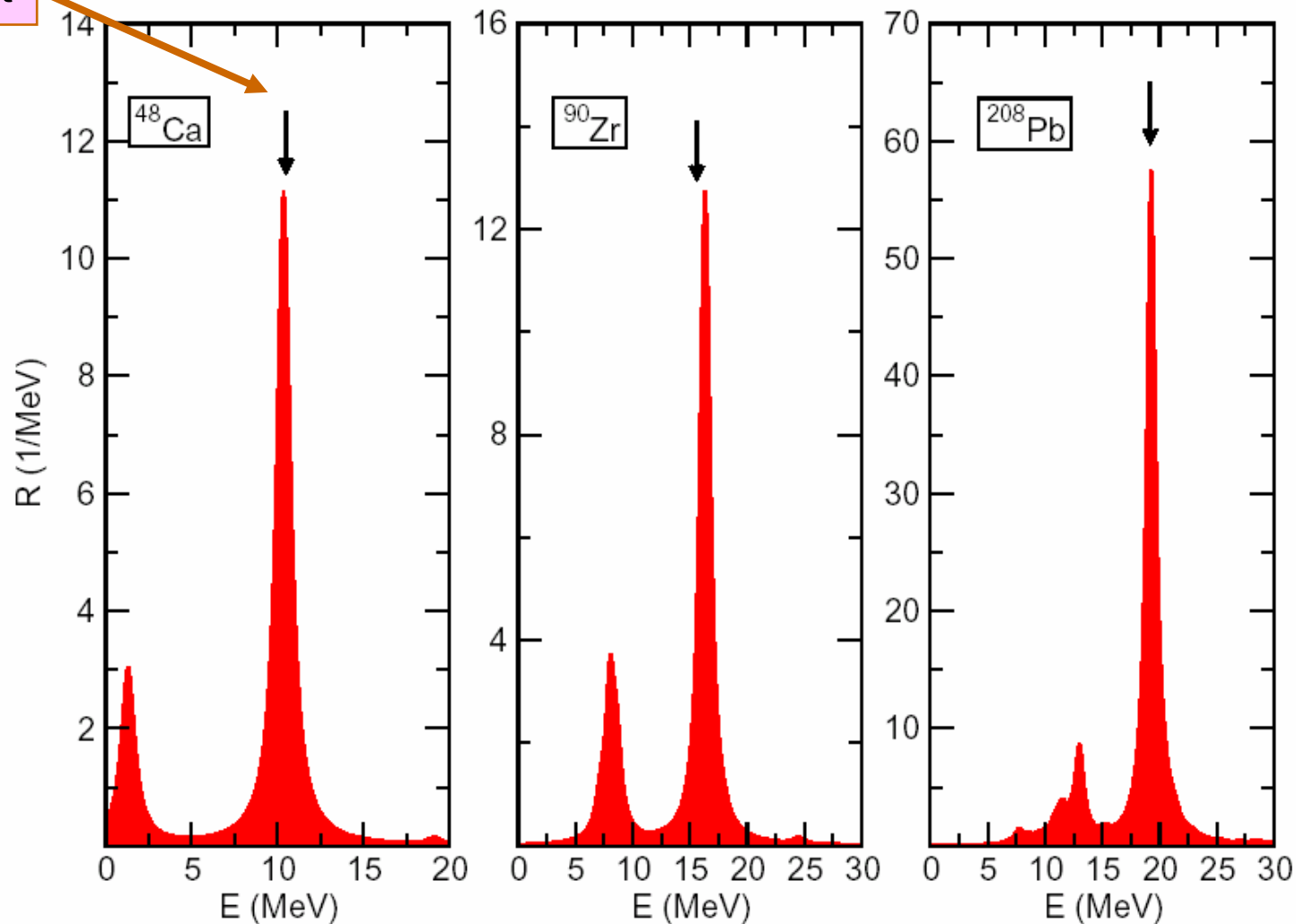
ISOBARIC ANALOG STATE:

$$S=0 \quad T=1 \quad J^{\pi} = 0^{+}$$

GT-Resonances

experiment

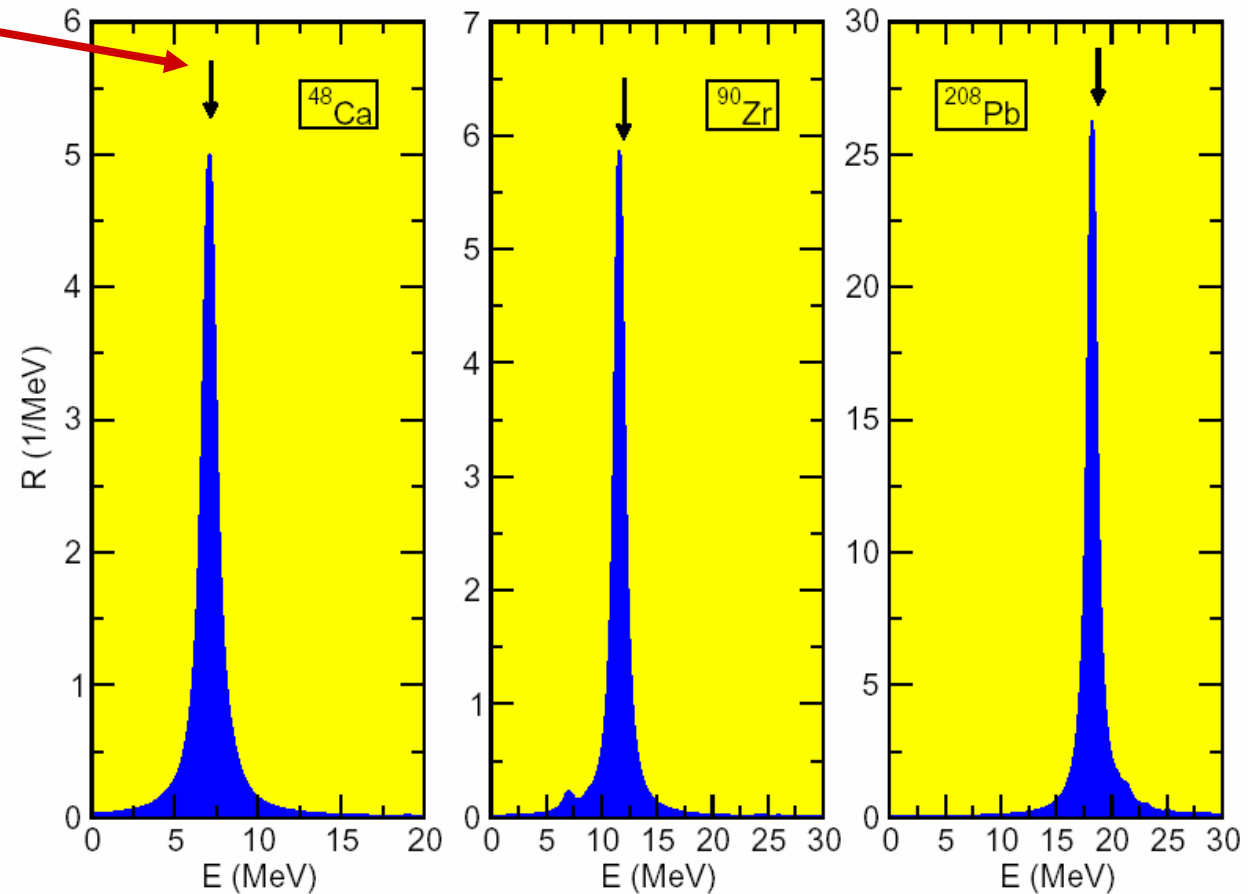
N. Paar, T. Niksic, D. Vretenar, P. Ring, PR C69, 054303 (2004)



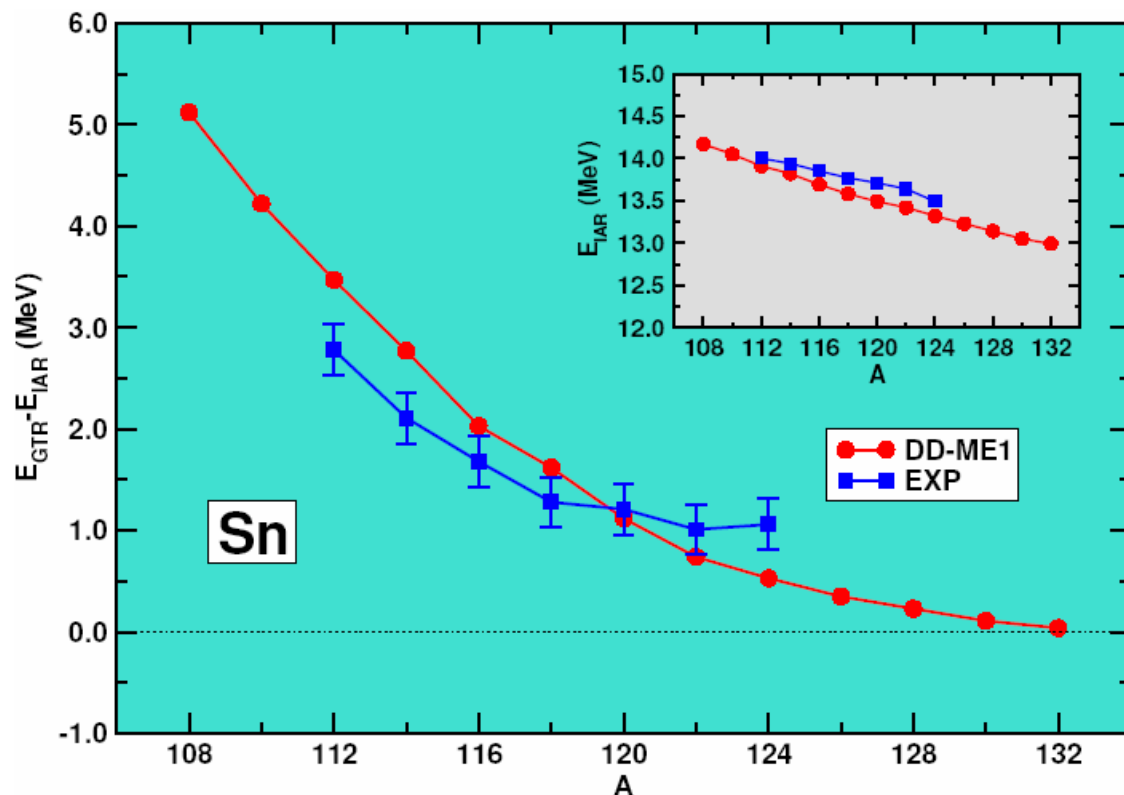
Isobaric Analog Resonance: IAR

N. Paar, T. Nikšić, D. Vretenar, P. Ring, PR C69, 054303 (2004)

experiment



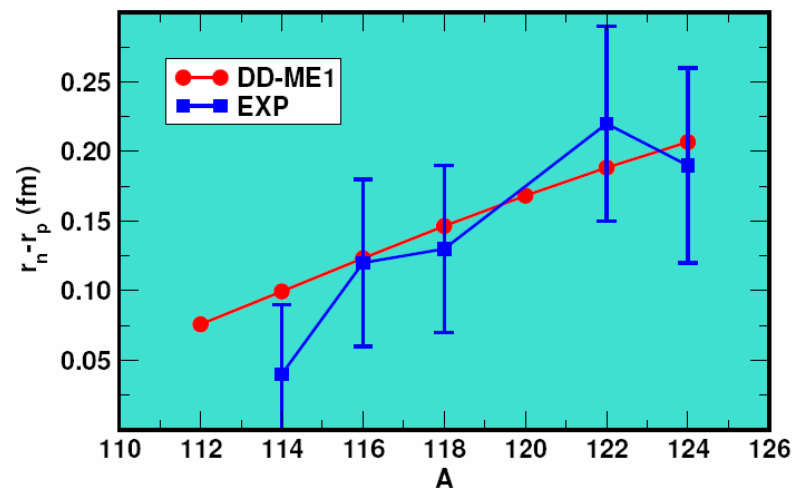
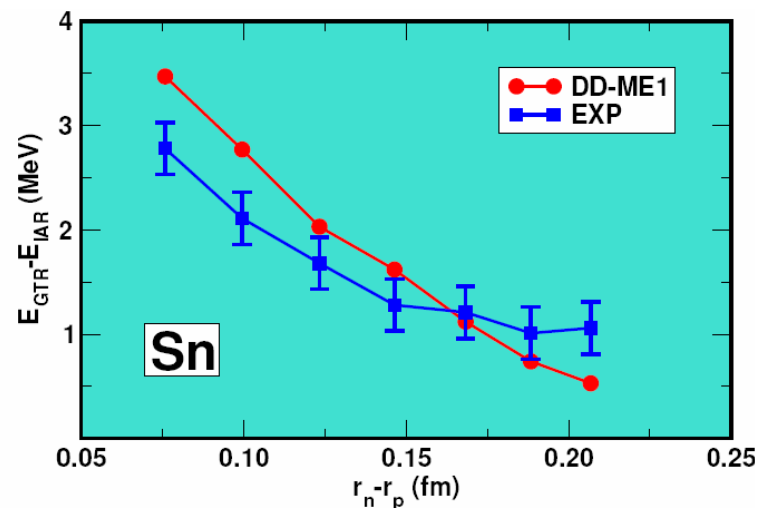
Neutron skin and IAR/GTR



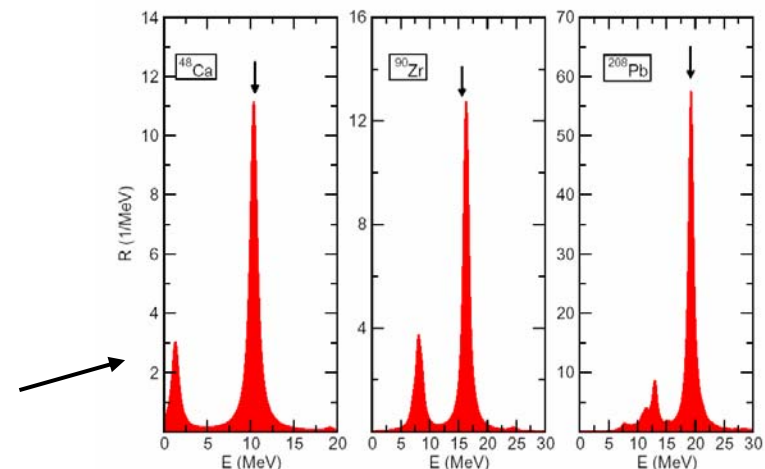
The isotopic dependence of the energy spacings between the GTR and IAS



direct information on the evolution of the neutron skin along the Sn isotopic chain



allowed β -decay :



* Important points:

- the tail of the GT-strength distribution at low energies
- the position of specific single particle levels (i.e. effective mass)
- effective pairing force in the T=1 and T=0 channel.
- in simple QRPA the lifetimes are too big

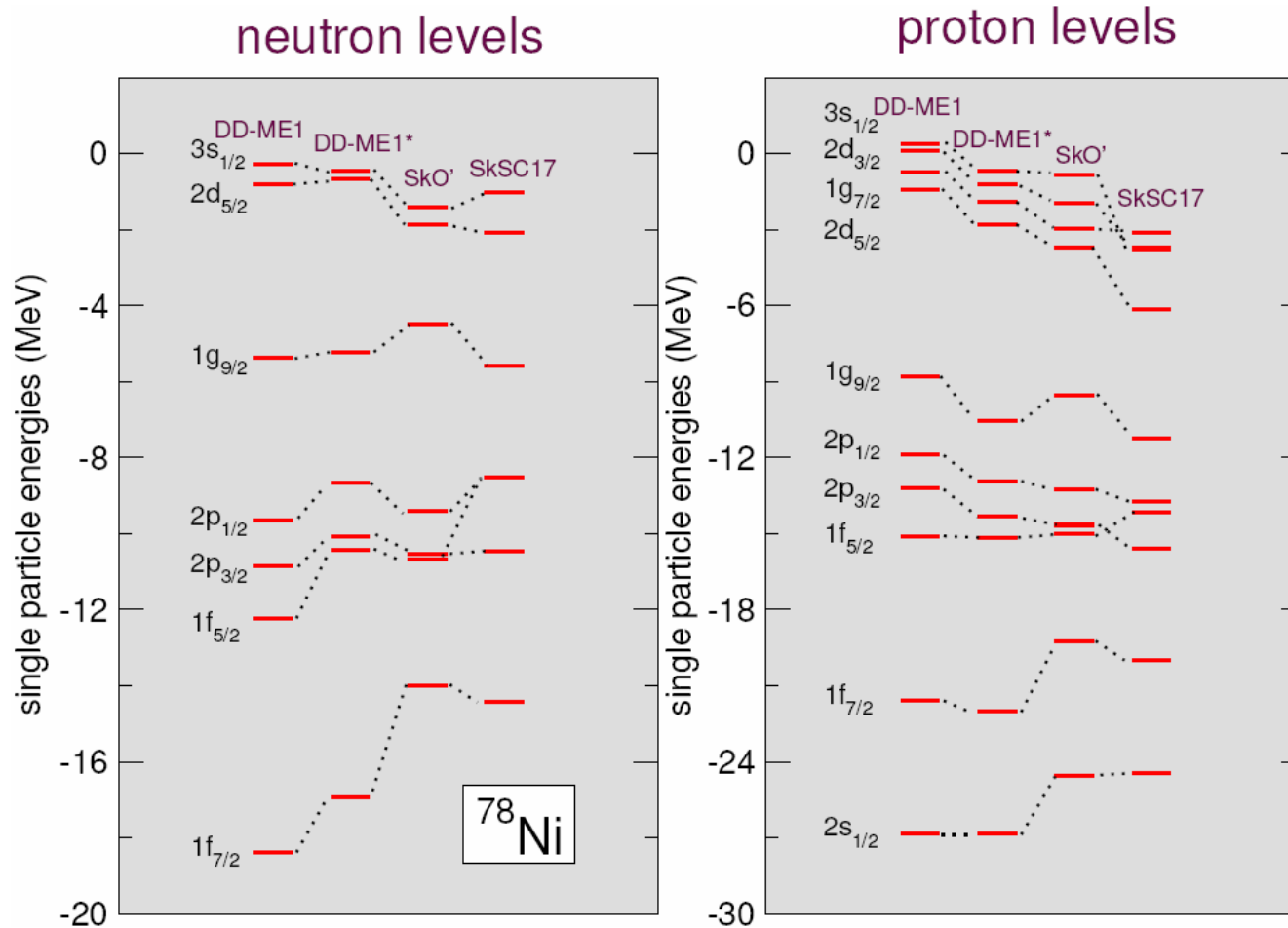
* Possible methods to improve the results:

- coupling to surface vibrations (difficult and beyond mean field)
- use of a tensor coupling in the ω -channel (one phenom. param.)
- T=0 pairing force with Gaussian character (one phen. parameter)

enhanced value of the
effective mass



increased density of states
around the Fermi surface



T. Niksic et al, PRC 71, 014308 (2005)

The nucleon effective mass m^* :

m^* represents a measure of the **density of states** around the Fermi surface

nonrelativistic mean-field models

effective mass: $m^*/m = 0.8 \pm 0.1$

relativistic mean-field models

Dirac mass: $m_D = m + S(r)$

effective mass: $m^* = m - V(r)$

conventional RMF models

spin-orbit splittings + nuclear matter binding

$$0.55m \leq m_D \leq 0.60m$$

$$0.64m \leq m^* \leq 0.67m$$

small density of states
-> **overestimated** β -decay lifetimes


Reduction of the spin-orbit in neutron-rich nuclei

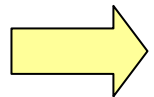
Lalazissis, Vretenar, Poeschl, Ring, Phys. Lett. B418, 7 (1998)

The **spin-orbit potential** originates from the addition of two large fields: the field of the vector mesons (short range repulsion), and the scalar field of the sigma meson (intermediate attraction).

$$V_{s.o.} \approx \frac{1}{r} \frac{\partial}{\partial r} V_{ls}(r)$$

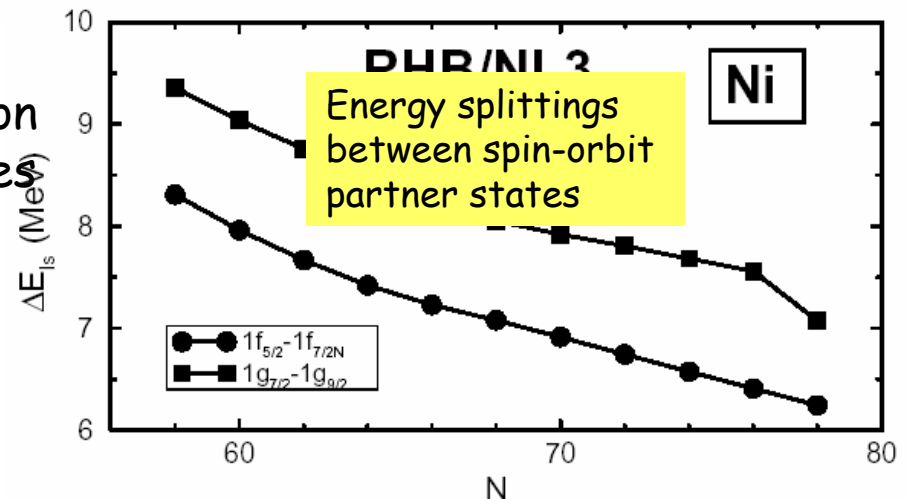
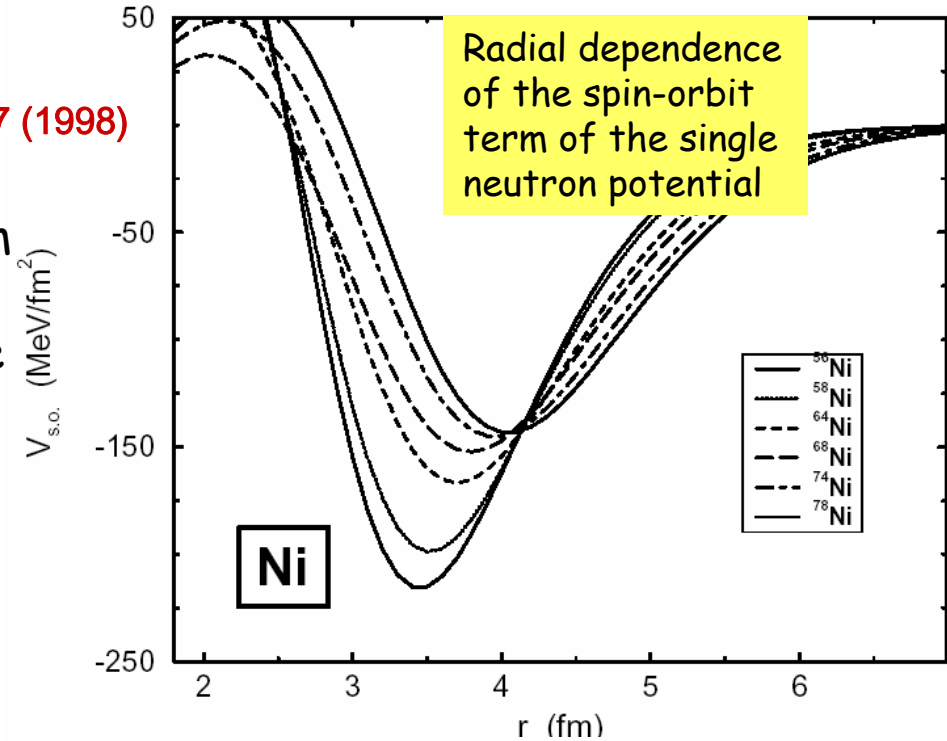
$$V_{ls} = \frac{m}{m_{eff}} (V - S)$$

 weakening of the effective single-neutron spin-orbit potential in neutron-rich isotopes

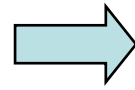


reduced energy spacings between spin-orbit partners

$$\Delta E_{ls} = E_{n,l,j=l-1/2} - E_{n,l,j=l+1/2}$$

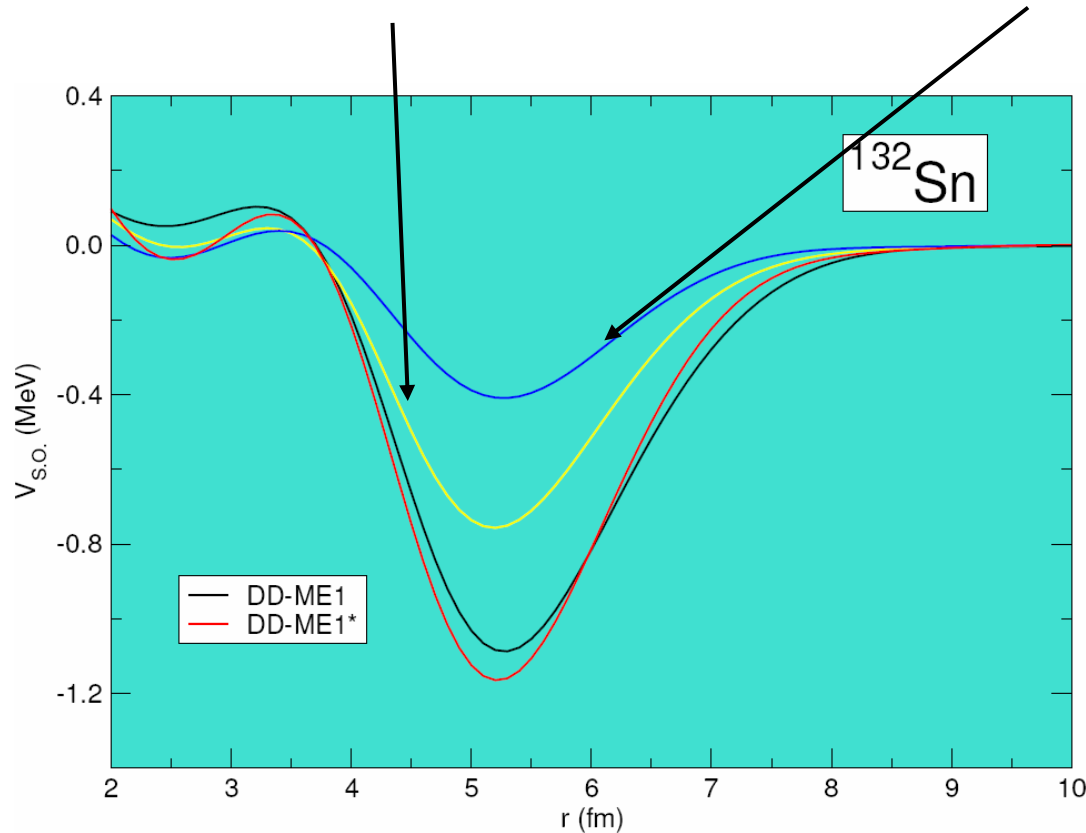


tensor omega-nucleon coupling
enhances the spin-orbit interaction



scalar and vector self-energies
can be reduced

$$V_{so} = \left[\frac{1}{4M^2} \frac{1}{r} \frac{d}{dr} (V - S) + \frac{f_V}{2MM} \frac{1}{r} \frac{d\omega}{dr} \right] \mathbf{l} \cdot \mathbf{s}$$

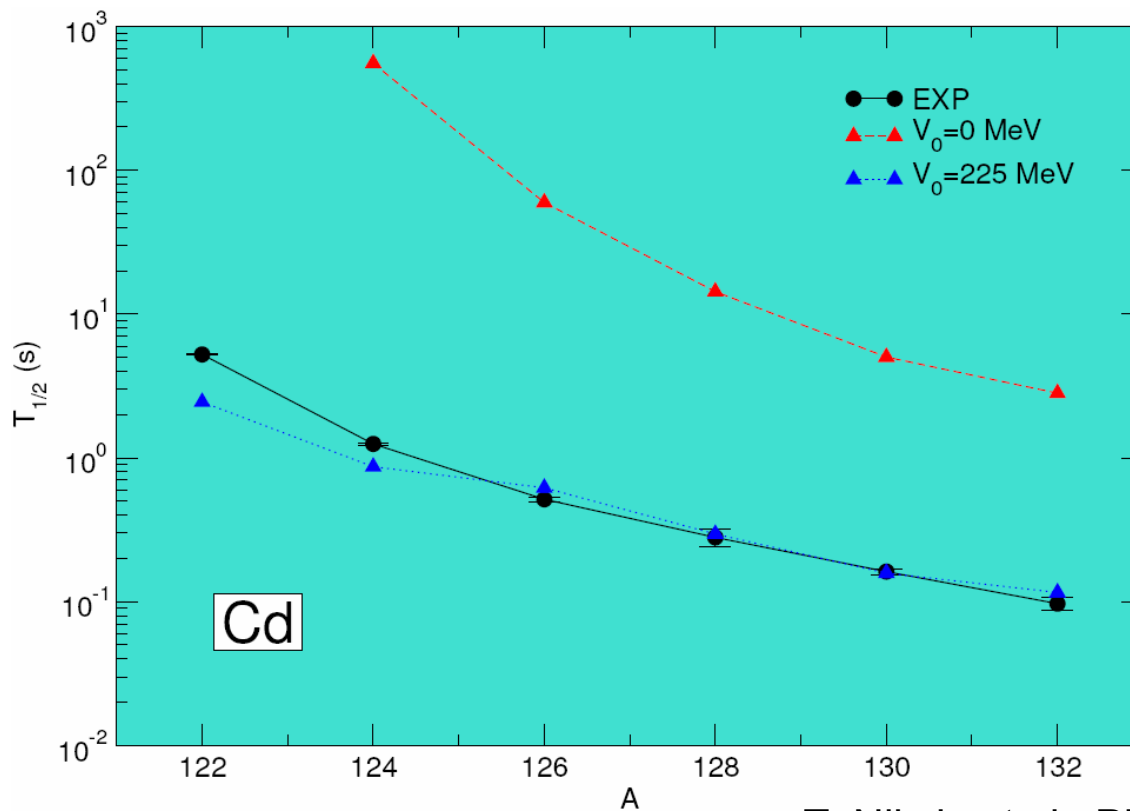


	DD-ME1	DD-ME1*
m_D/m	0.58	0.67
m^*/m	0.66	0.76

T. Niksic et al, PRC 71, 014308 (2005)

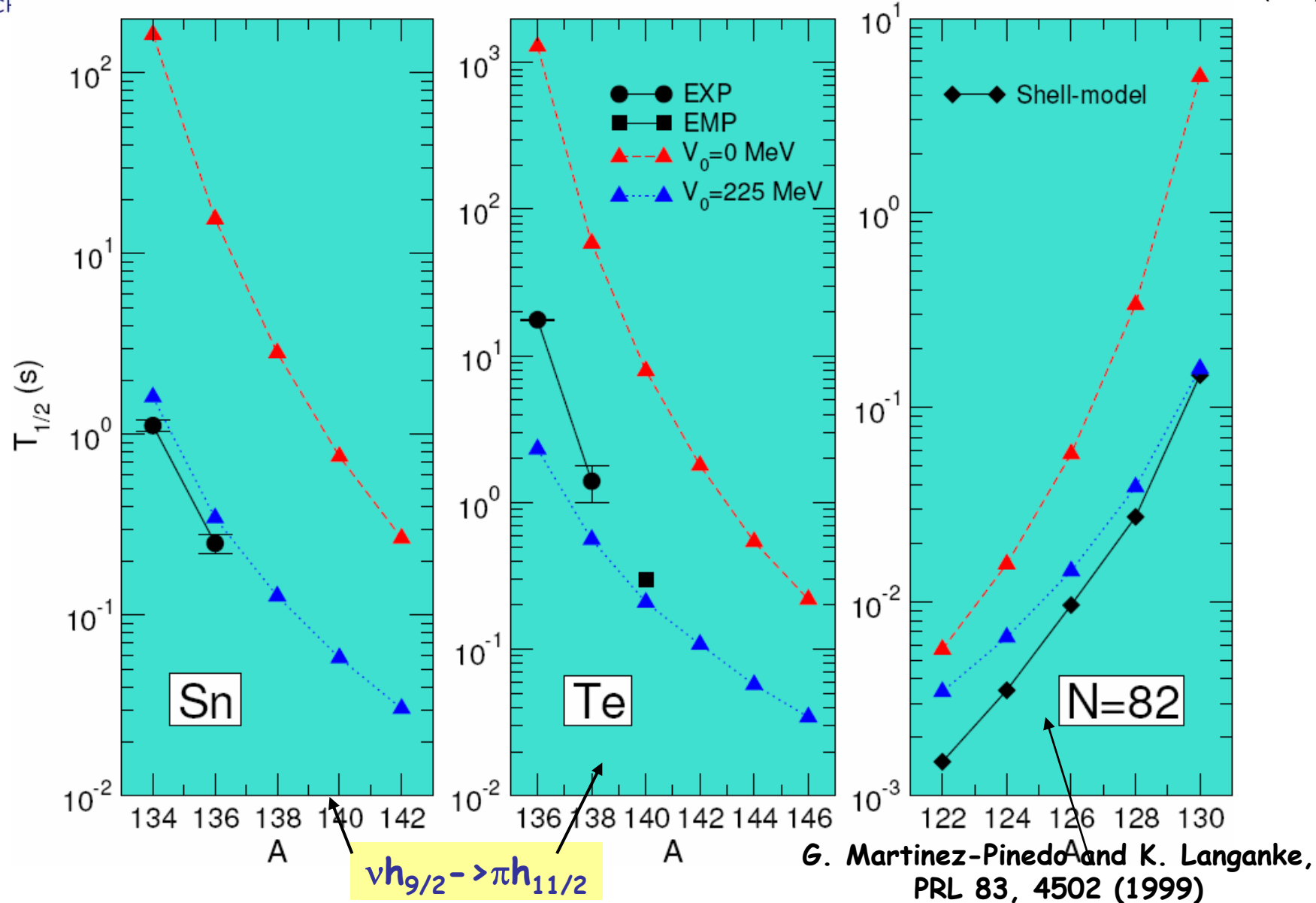
N \approx 82 region:

Cadmium isotopes: $\pi 1g_{9/2}$ level is partially empty
 \Rightarrow T=0 pairing has large influence on the $\nu 1g_{7/2} \rightarrow \pi 1g_{9/2}$ transition
 which dominates the β -decay process



T. Niksic et al, PRC 71, 014308 (2005)

An **increase** of the T=0 pairing **partially compenstates** for the fact that the density of states is still **rather low** Niksic et al, PRC 71, 014308 ('05)



Correlations beyond mean field

- Conservation of symmetries by **projection before variation**
- Motion with large amplitude by **Generator Coordinates**
- Coupling to **collective vibrations**
 - shifts of single particle energies
 - decay width of giant resonances

Projected Density Functionals

$$|\Psi^N\rangle = \hat{P}^N |\Phi\rangle = \delta(\hat{N} - N) |\Phi\rangle = \int \frac{d\varphi}{2\pi} e^{i\varphi(\hat{N} - N)} |\Phi\rangle$$

projected density functional:

$$E^N[\hat{\rho}, \hat{\kappa}] = \frac{\langle \Phi | \hat{H} \hat{P}^N | \Phi \rangle}{\langle \Phi | \hat{P}^N | \Phi \rangle}$$

analytic expressions

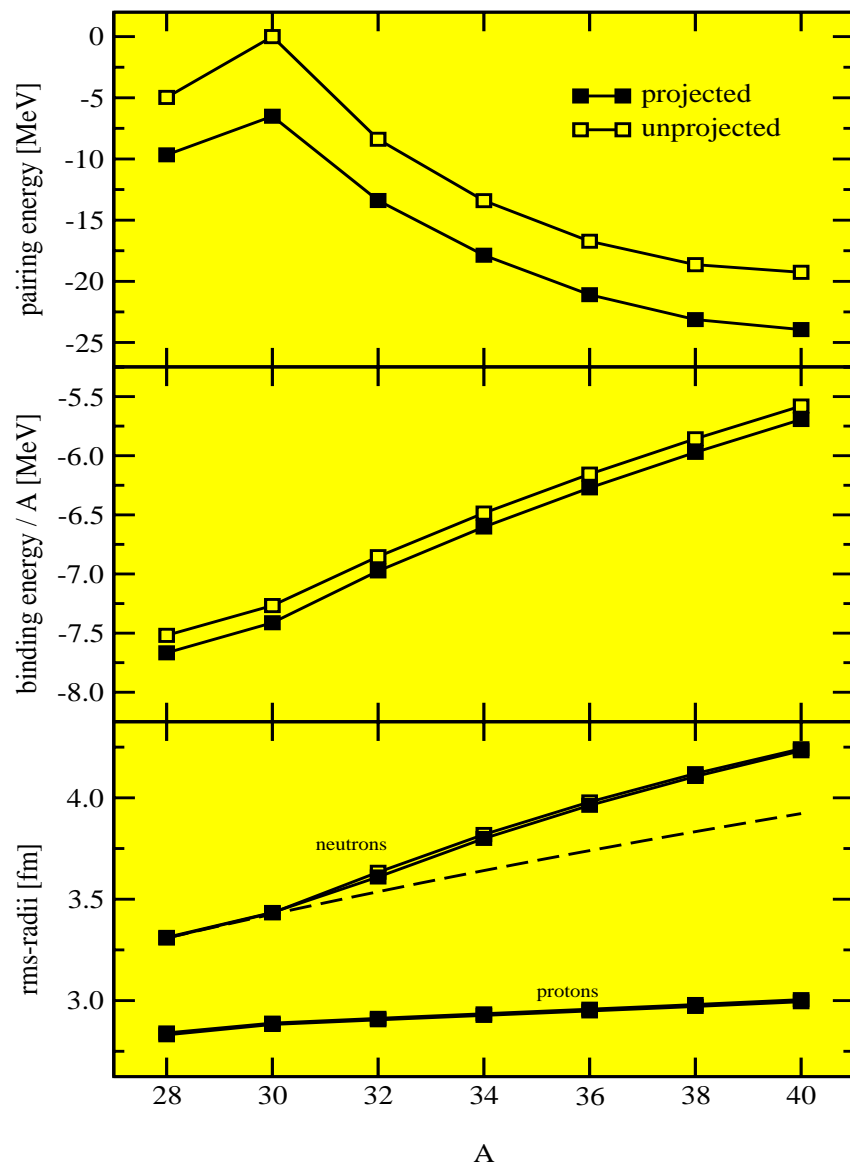
projected HFB-equations (variation after projection):

$$\begin{pmatrix} \hat{h}^N & \hat{\Delta}^N \\ -\Delta^{N*} & -\hat{h}^{N*} \end{pmatrix} \begin{pmatrix} U_k(\mathbf{r}) \\ V_k(\mathbf{r}) \end{pmatrix} = \begin{pmatrix} U_k(\mathbf{r}) \\ V_k(\mathbf{r}) \end{pmatrix} E_k$$

J.Sheikh and P. Ring NPA 665 (2000) 71

$$\hat{h}^N = \frac{\delta E^N}{\delta \hat{\rho}}$$

$$\hat{\Delta}^N = \frac{\delta E^N}{\delta \hat{\kappa}}$$



pairing energies

binding energies

rms-radii

L. Lopes, PhD Thesis, TUM, 2002

Generator Coordinate Method (GCM)

$$\langle \delta\Phi | \hat{H} - q\hat{Q} | \Phi \rangle = 0$$

Constraint Hartree Fock produces wave functions depending on a generator coordinate q

$$|q\rangle = |\Phi(q)\rangle$$

GCM wave function is a superposition of Slaterdeterminants

$$|\Psi\rangle = \int dq f(q) |q\rangle$$

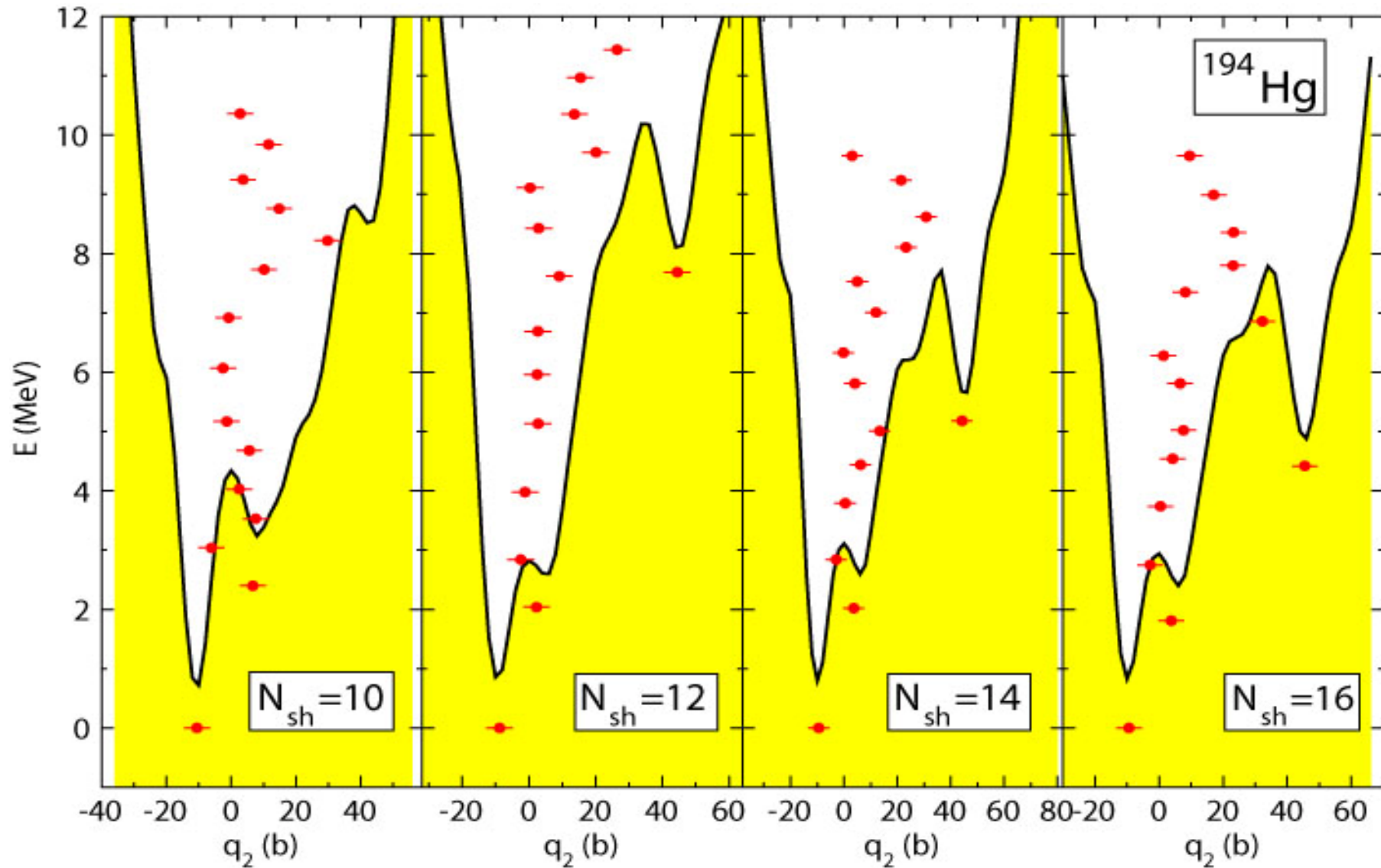
Hill-Wheeler equation:

$$\int dq [\langle q | H | q' \rangle - E \langle q | q' \rangle] f(q') = 0$$

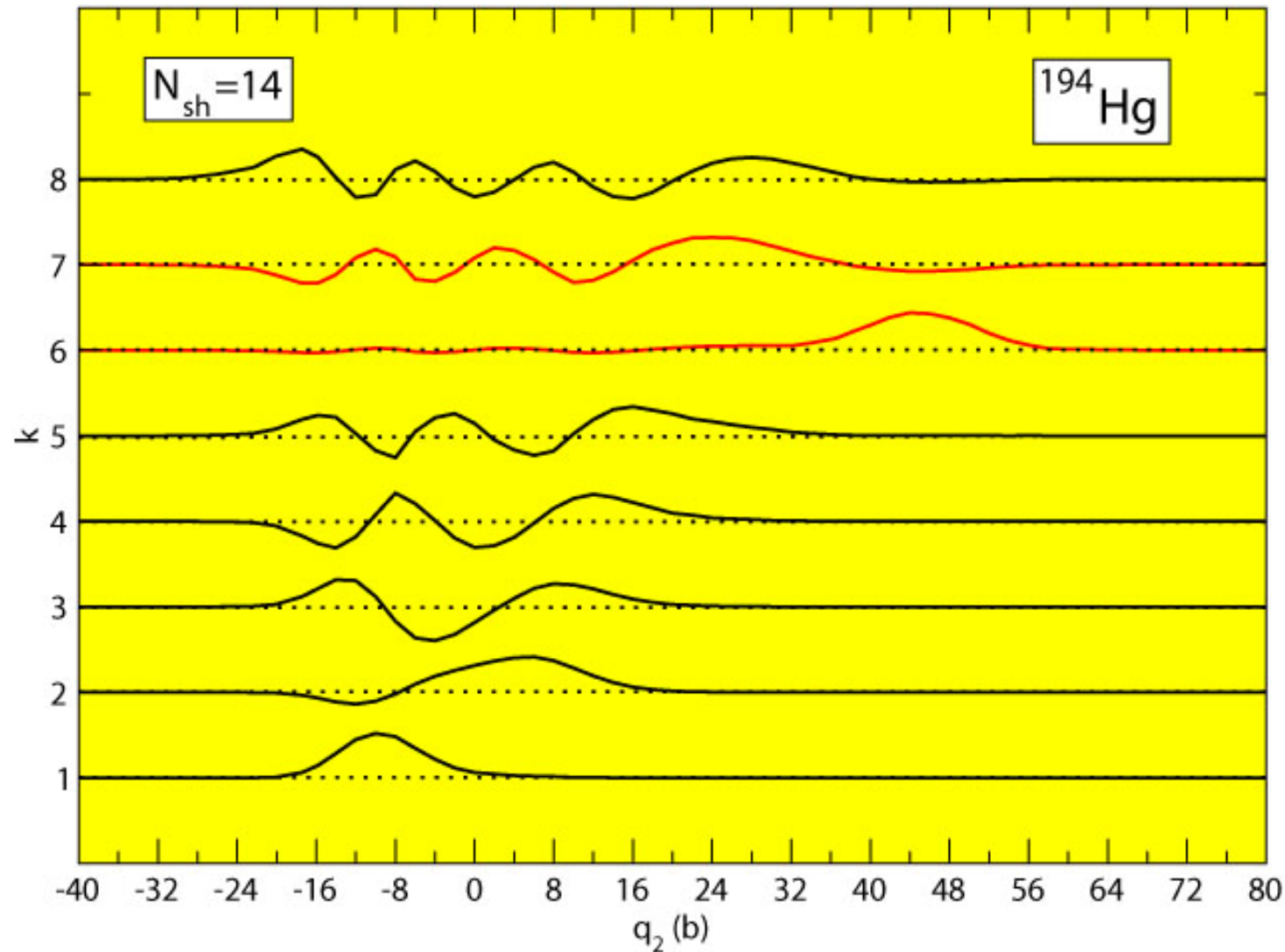
with projection:

$$|\Psi\rangle = \int dq f(q) \hat{P}^N \hat{P}^I |q\rangle$$

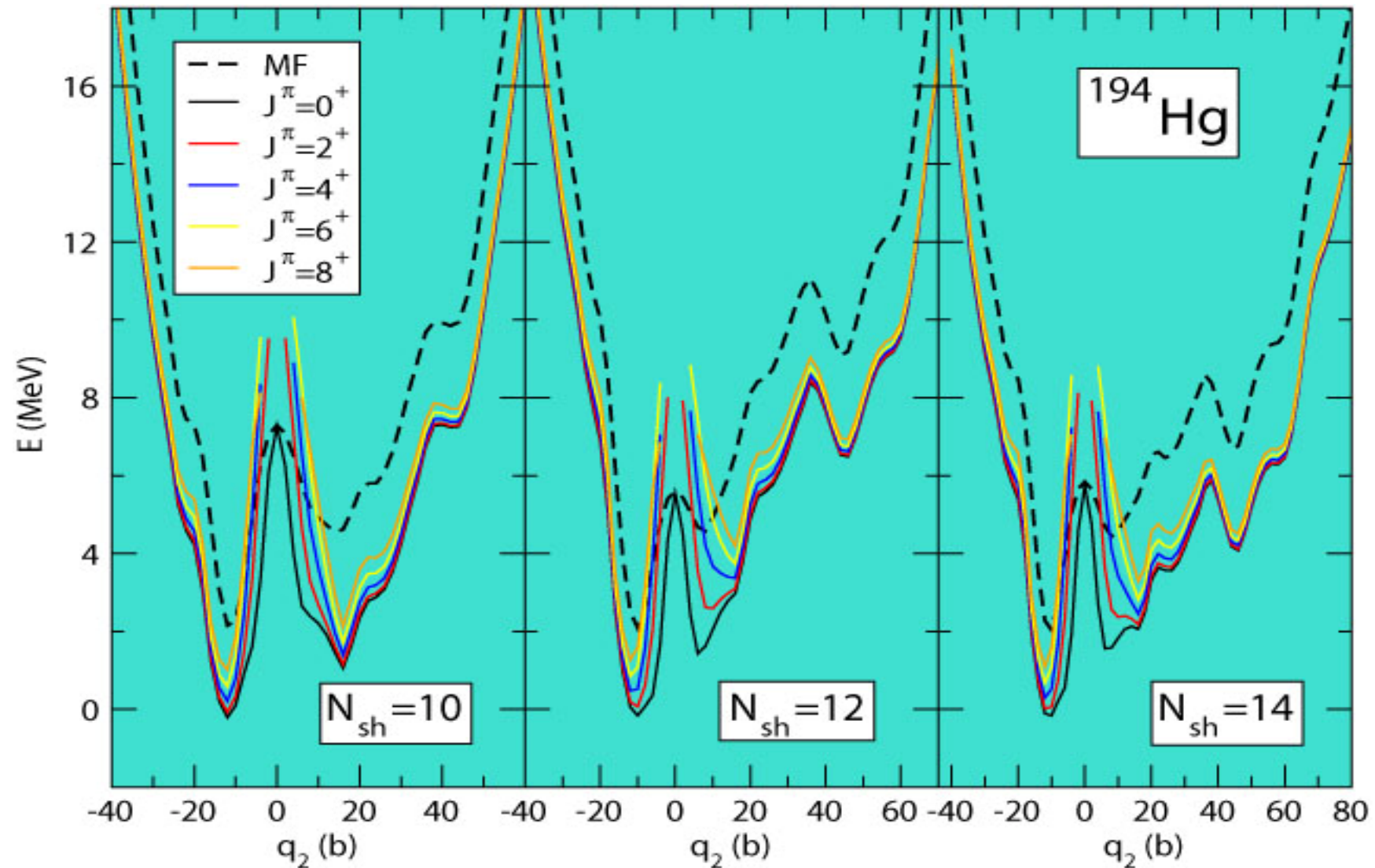
GCM without projection:



GCM-wave functions of the lowest states



Ang. momentum projected energy surfaces:



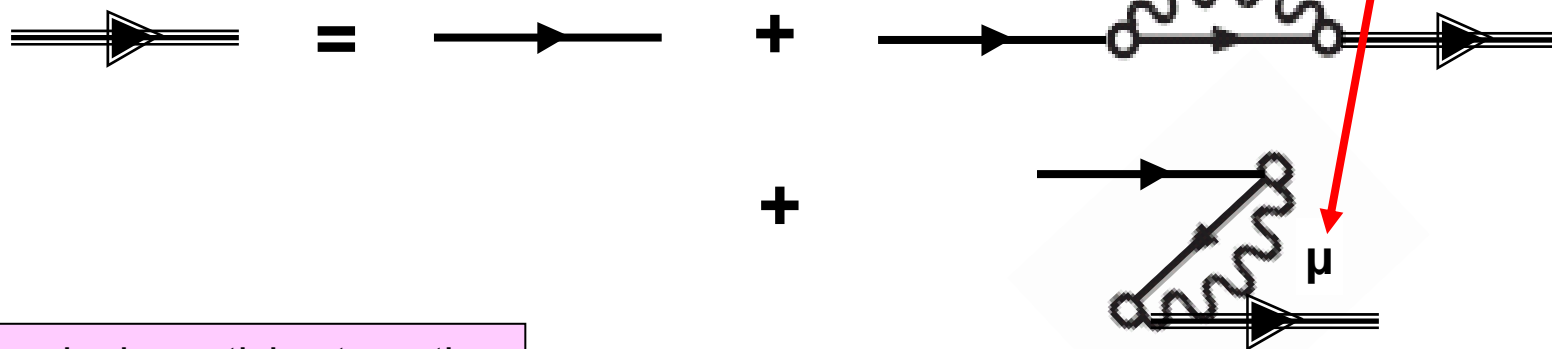
Vibrational Couplings: energy dependent self-energy:

$$\Sigma = S + V + \Sigma(\omega)$$

mean field

pole part

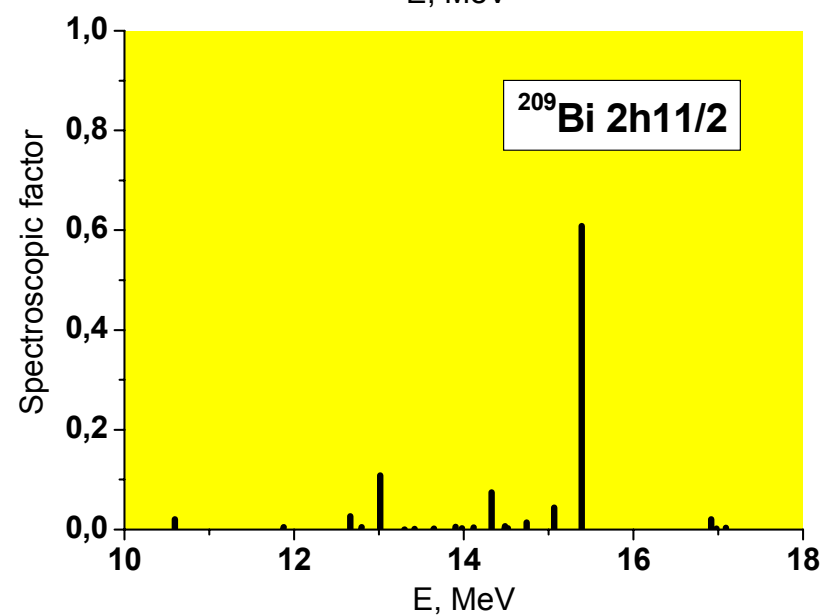
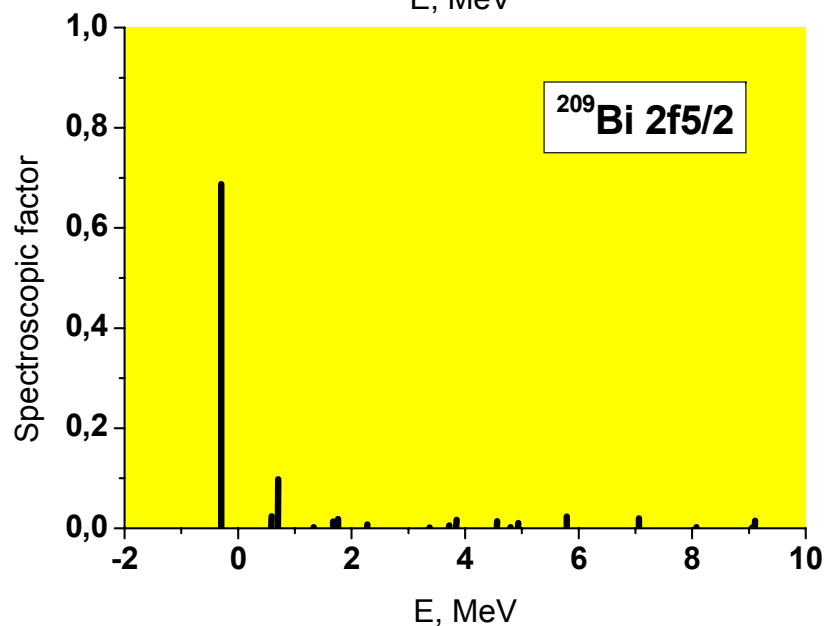
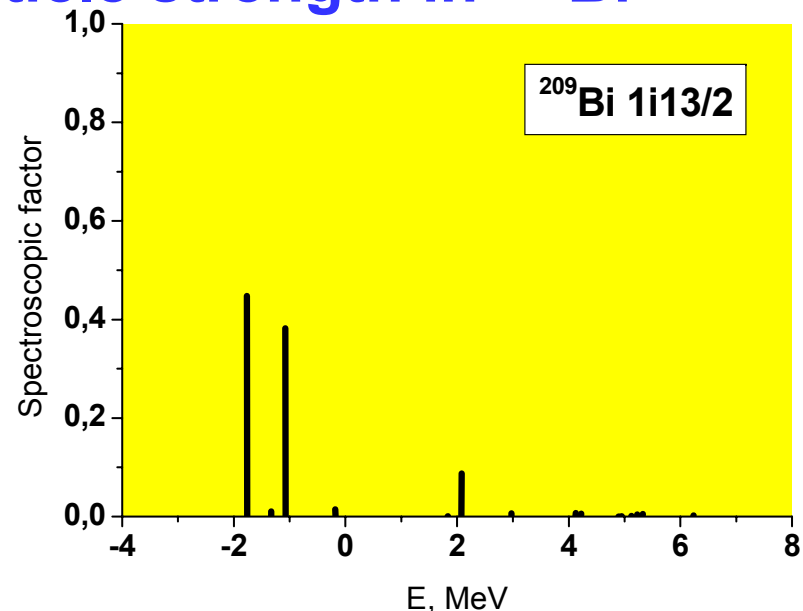
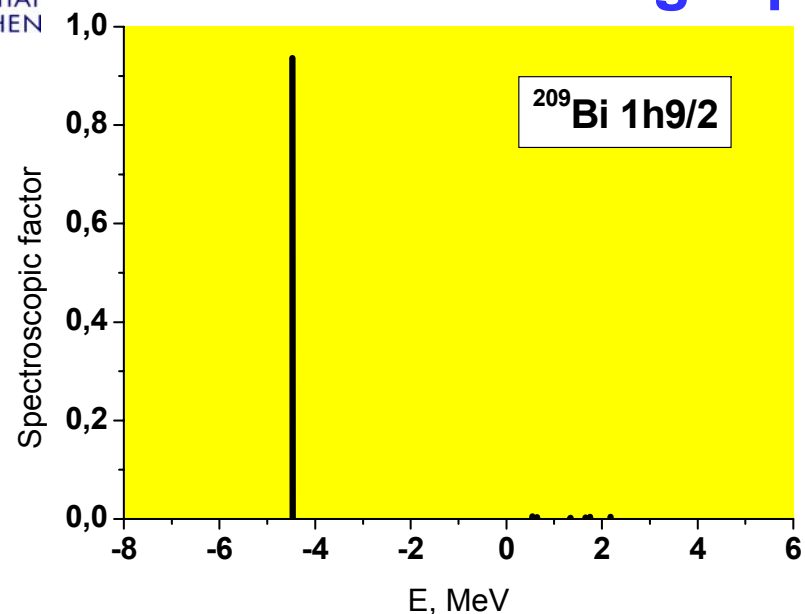
RPA-modes

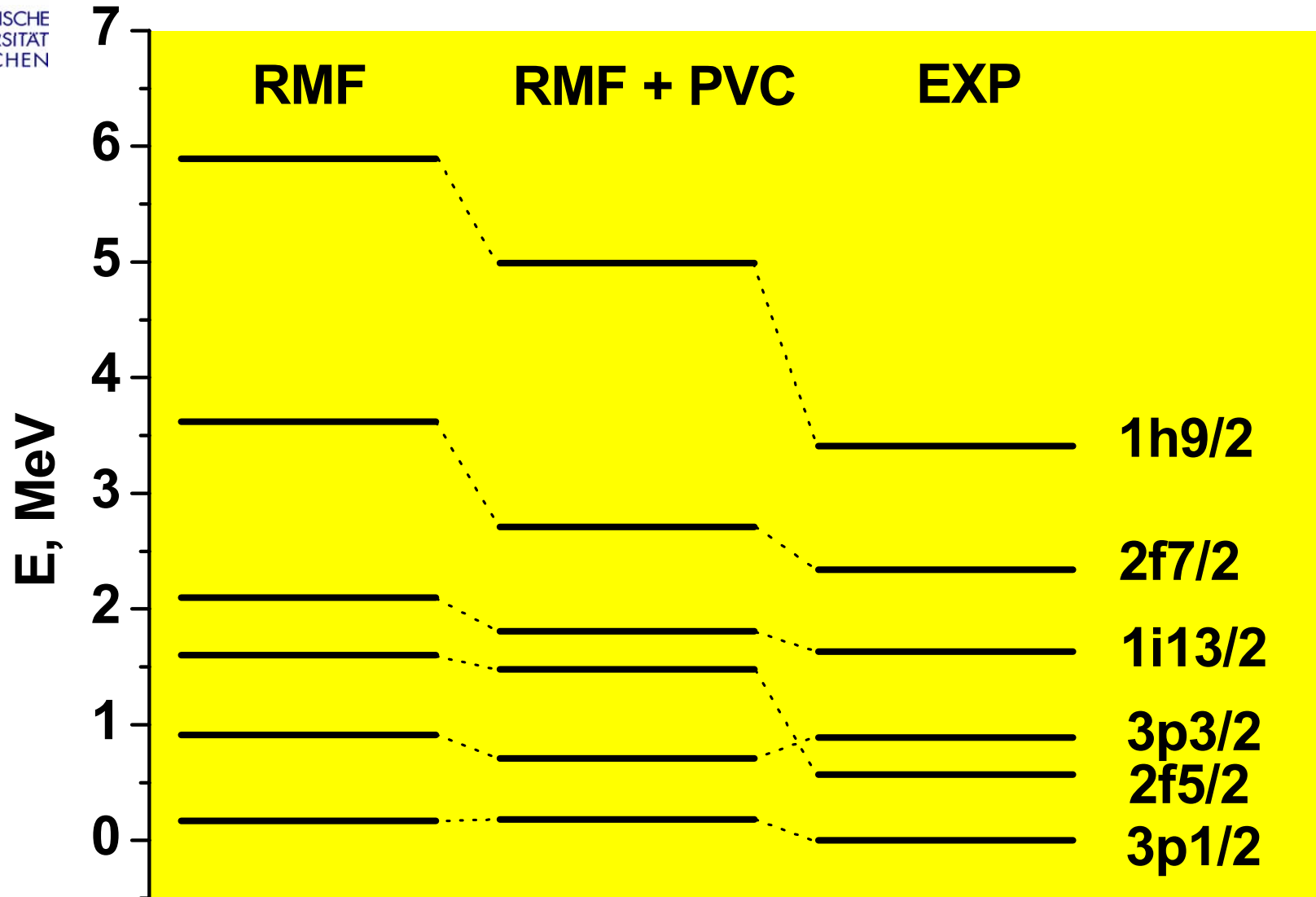


single particle strength:

$$z_\nu = \left[1 - \frac{d\Sigma_{\nu\nu}}{d\omega} \Big|_{\omega=\epsilon_\nu} \right]^{-1}$$

Distribution of single-particle strength in ^{209}Bi

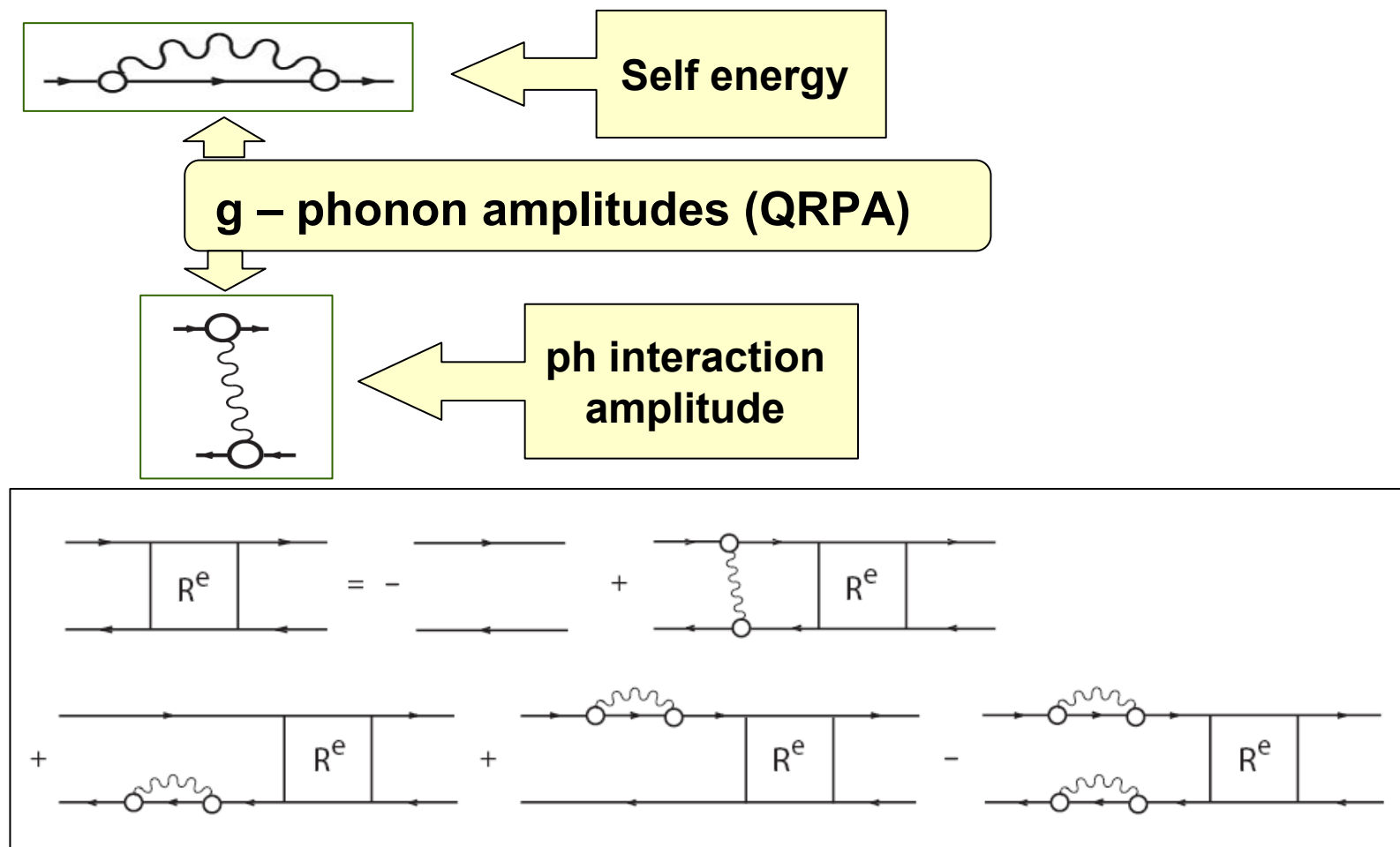




Level scheme for ^{207}Pb

Contributions of complex configurations

The full response contains energy dependent parts coming from vibrational couplings.

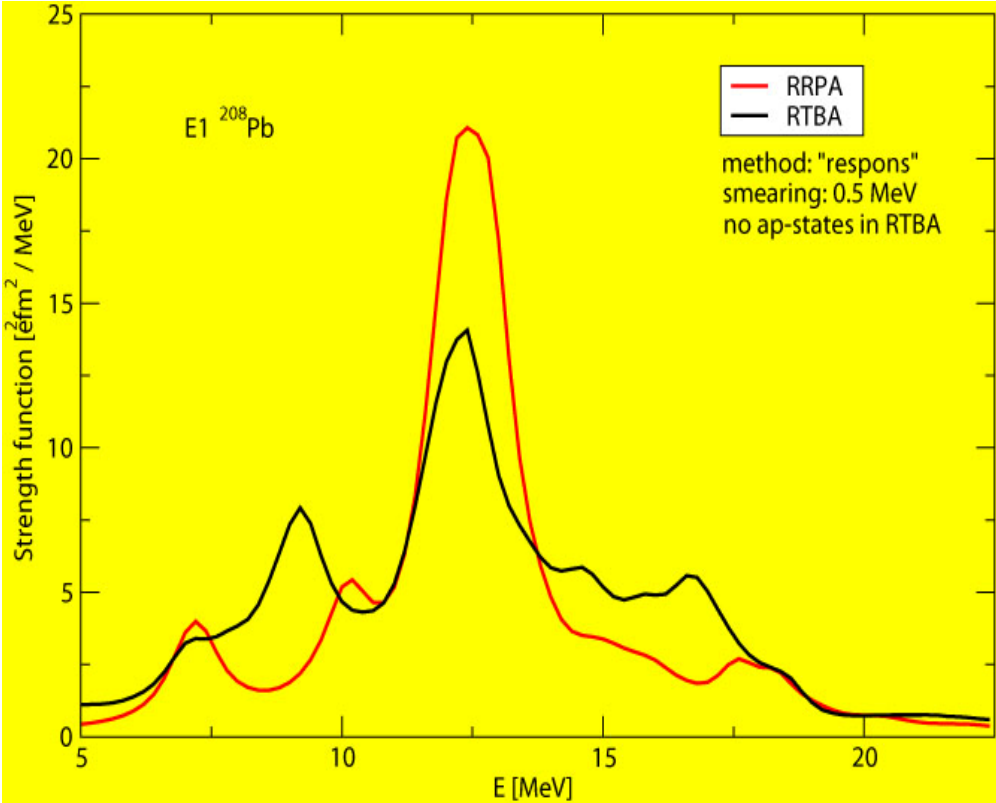


Decay-width of the Giant Resonances

$$S(E) = -\frac{1}{\pi} \text{Im} \Pi(E + i \Delta)$$

**E1 photoabsorption
cross section**

$$\sigma_{E1}(E) = \frac{16\pi^3 e^2}{9\hbar c} E S_{E1}(E)$$



Conclusions

On the way to a universal covariant density functional adjusted to ground state properties of finite nuclei.
≈7 parameters necessary for high precision

Time-dependent mean field theory provides a parameter-free theory for excited states

- rotational spectra (cranked RHB-theory)
- vibrational excitations (rel. QRPA)

Method beyond mean field:

- Projected functionals (PDFT)
- Generator Coordinate Method (GCM)
- Particle-Vibrational Coupling (PVC)

Open Problems:

Fock terms and tensor forces:

- **why is the first order pion-exchange quenched?**

Vacuum polarization:

- **renormalization in finite systems**

Simpler parametrizations:

- **point coupling**
- **simple pairing**

**Do we have to change the functional,
if we go beyond mean field?**

Open Problems:

Fock terms and tensor forces:

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- **simple pairing**

**Do we have to change the functional,
if we go beyond mean field?**

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