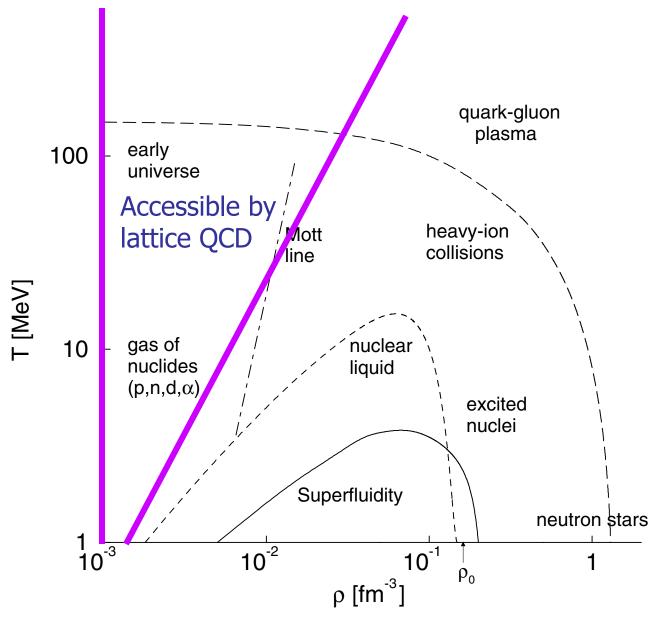
# Lattice Simulations of Cold Dilute Neutron Matter

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#### Outline

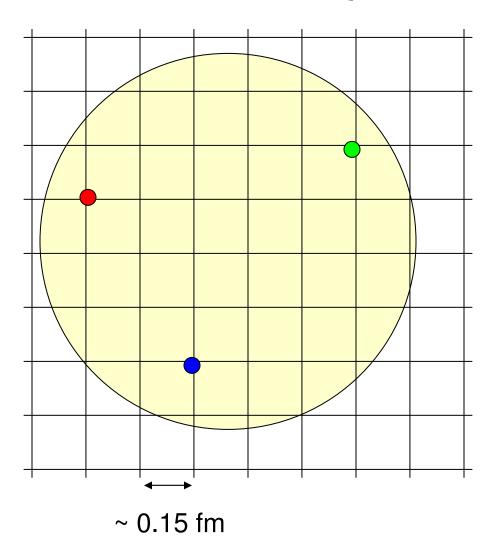
- 1. Lattice simulations with effective field theory
- 2. Neutron scattering and universality
- 3. A puzzle
- 4. High temperature/low density calculations
- 5. Virial coefficients
- 6. Unitary limit and scaling
- 7. Results in the unitary limit



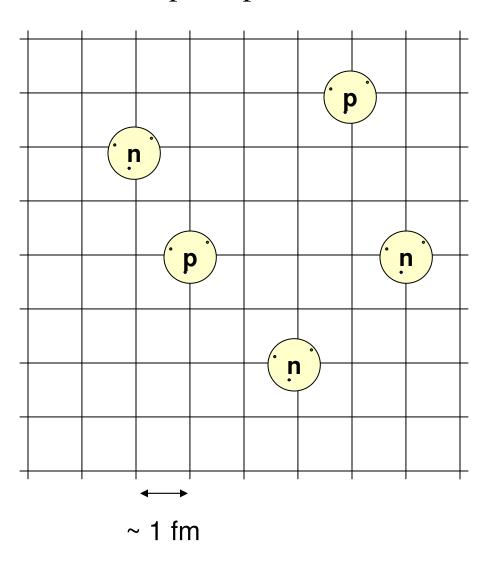
from Ropke and Schell, Prog. Part. Nucl. Phys. 42, 53 (1999)

# Why do nuclear lattice simulations?

#### Nucleon in lattice QCD



#### Nucleons as point particles on lattice



# Simulations with Effective Field Theory

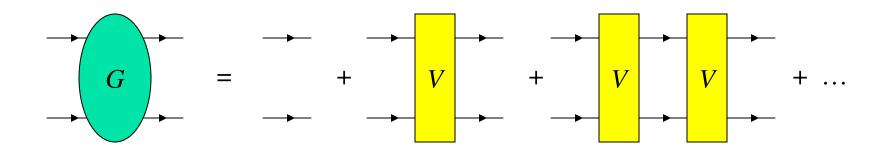
Non-perturbative lattice simulations of effective field theory of low energy pions and nucleons.

Non-perturbative effective field theory?... but isn't effective field theory based upon an expansion?

For pions the expansion is simple

$$G = G_0 + G_2 + \dots$$

For nucleons we must take care of infrared singularities [Weinberg, PLB 251 (1990) 288, NPB 363 (1991) 3]



We will iterate "everything"

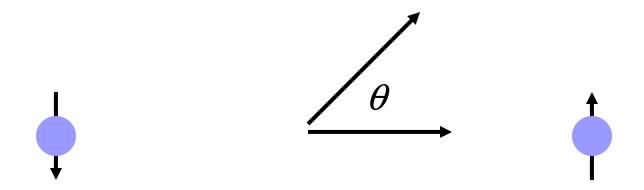
$$G = \frac{\int D\pi DND\bar{N} \ G(\pi, N, \bar{N}) \ e^{-S(\pi, N, \bar{N})}}{\int D\pi DND\bar{N} \ e^{-S(\pi, N, \bar{N})}}$$

$$= \frac{\int D\pi DND\bar{N} \ G(\pi, N, \bar{N}) \ e^{-S(\pi, N, \bar{N})}}{\int D\pi DND\bar{N} \ e^{-\sum_{i \leq k} S_i(\pi, N, \bar{N})}}$$

A complete summation of all diagrams involving interaction terms with order  $\leq k$ . [D.L., Borasoy, Schäfer, PRC70 (2004) 014007]

#### Pure neutron matter

[D.L. and Schäfer, PRC72 (2005) 024006]



Incoming and scattered wave

$$\psi(\mathbf{r}) \sim e^{i\mathbf{k}\cdot\mathbf{r}} + f(\mathbf{k}',\mathbf{k}) \frac{e^{i\mathbf{k}'\cdot\mathbf{r}}}{\mathbf{r}}$$

Partial wave decomposition

$$f(\mathbf{k}', \mathbf{k}) = \sum_{l=0}^{\infty} f_l(k) P_l(\cos \theta)$$

Phase shifts

$$f_l(k) = \frac{2l+1}{2ik} \left( e^{2i\delta_l(k)} - 1 \right)$$

S-wave scattering dominant at lowest energies

$$f_0(k) = \frac{1}{k \cot \delta_0(k) - ik}$$

S-wave scattering length

$$a_{scatt} = -\lim_{k \to 0} \frac{\delta_0(k)}{k}$$

Effective range expansion

$$k \cot \delta_0(k) \approx -a_{scatt}^{-1} + \frac{1}{2}k^2r_0$$

Neutron-neutron scattering length is -18 fm while the range is only 2.8 fm.

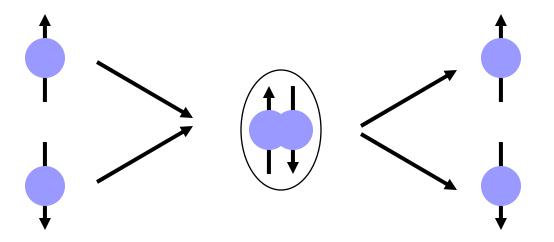
Theoretically interesting ... for dilute neutron matter one is close to *unitary regime* or *universal scaling* limit where magnitude of scattering length  $\rightarrow \infty$ , range  $\rightarrow 0$ .

In this limit, no dimensionful parameters as  $T \rightarrow 0$ , and so we expect the energy per particle and superfluid gap to satisfy

$$\frac{E}{A} = \xi \frac{3}{5} \frac{k_F^2}{2m}, \quad \Delta = \zeta \frac{k_F^2}{2m}$$

#### Feshbach resonance

Experiments done with cold Li and K atoms, which can form diatomic molecules, Li<sub>2</sub> and K<sub>2</sub>.



Tune energy of the diatomic molecule with external magnetic to produce a resonance near threshold. [O'Hara et. al., Science 298 (2002) 2179; Regal, Jin, PRL (2003) 230404; etc.]

For dilute neutron matter we have the effective Hamiltonian,

$$H = -\sum_{i} \int d^{3}x a_{i}^{\dagger} \frac{\nabla^{2}}{2m} a_{i} + C \int d^{3}x a_{\downarrow}^{\dagger} a_{\uparrow}^{\dagger} a_{\uparrow} a_{\downarrow}$$

On the lattice,

$$H - \mu N = \sum_{\vec{n}_s, i} \left[ (-\mu + \frac{3}{m}) a_i^{\dagger}(\vec{n}_s) a_i(\vec{n}_s) \right]$$

$$- \frac{1}{2m} \sum_{\vec{n}_s, l_s, i} \left[ a_i^{\dagger}(\vec{n}_s) a_i(\vec{n}_s + \hat{l}_s) + a_i^{\dagger}(\vec{n}_s) a_i(\vec{n}_s - \hat{l}_s) \right]$$

$$+ C \sum_{\vec{n}_s} a_{\downarrow}^{\dagger}(\vec{n}_s) a_{\uparrow}^{\dagger}(\vec{n}_s) a_{\uparrow}(\vec{n}_s) a_{\downarrow}(\vec{n}_s)$$

We use a Hubbard-Stratonovich transformation to rewrite the interaction as

$$\exp\left[-\frac{C}{2}(a_{\uparrow}^{\dagger}a_{\uparrow} + a_{\downarrow}^{\dagger}a_{\downarrow})^{2}\right]$$

$$= \sqrt{\frac{1}{2\pi}} \int_{-\infty}^{\infty} ds \exp\left[-\frac{1}{2}s^{2} + s\sqrt{-C}(a_{\uparrow}^{\dagger}a_{\uparrow} + a_{\downarrow}^{\dagger}a_{\downarrow})\right]$$

We then integrate out the neutron field. The resulting action has no signs or phases, as the determinant of the matrix is positive semi-definite.

The matrix has the structure

$$M(s)=M_{\uparrow}(s)\oplus M_{\downarrow}(s)$$
  $M_{\uparrow}(s)=M_{\downarrow}(s)$   $\det M(s)=\left(\det M_{\uparrow}(s)
ight)^2\geq 0$ 

So we can use standard pseudofermion methods with Hybrid Monte Carlo

$$\det M(s) = \left(\det M_{\uparrow}(s)\right)^{2}$$

$$\propto \int d\phi^{*} d\phi \, \exp\left[\left(M_{\uparrow}^{-1}(s)\phi\right)^{\dagger} \left(M_{\uparrow}^{-1}(s)\phi\right)\right]$$

Most computationally intensive step is conjugate gradient inversion

$$\left[M_{\uparrow}^{\dagger}(s)M_{\uparrow}(s)\right]v=b$$

Can be accelerated by diagonal preconditioning

$$\left[D^{-1}(s)M_{\uparrow}^{\dagger}(s)M_{\uparrow}(s)D^{-1}(s)\right]D(s)v = D^{-1}(s)b$$

where

$$D(s) = \operatorname{diag}\left[M_{\uparrow}^{\dagger}(s)M_{\uparrow}(s)\right]$$

# Operator coefficient on the lattice

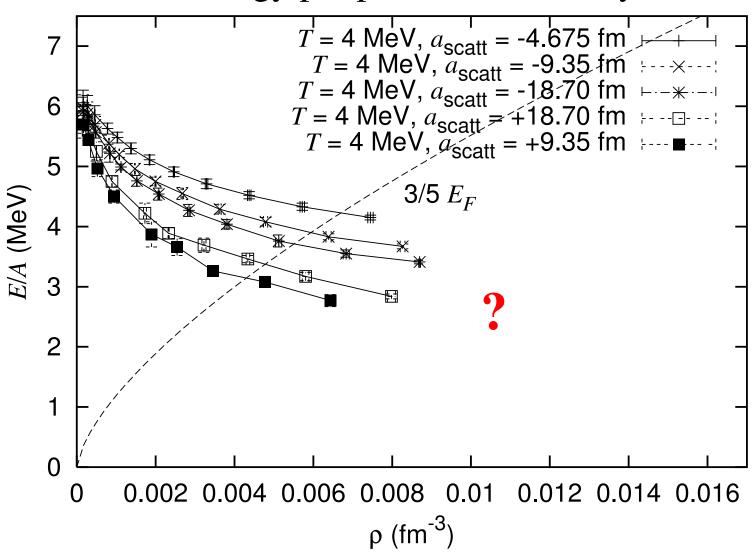
We use Lüscher's formula to set the operator coefficient *C* to give the physical s-wave scattering length for two-particle scattering.

$$E_0 = \frac{4\pi a_{scatt}}{mL^3} [1 - c_1 \frac{a_{scatt}}{L} + c_2 \frac{a_{scatt}^2}{L^2} + \cdots]$$

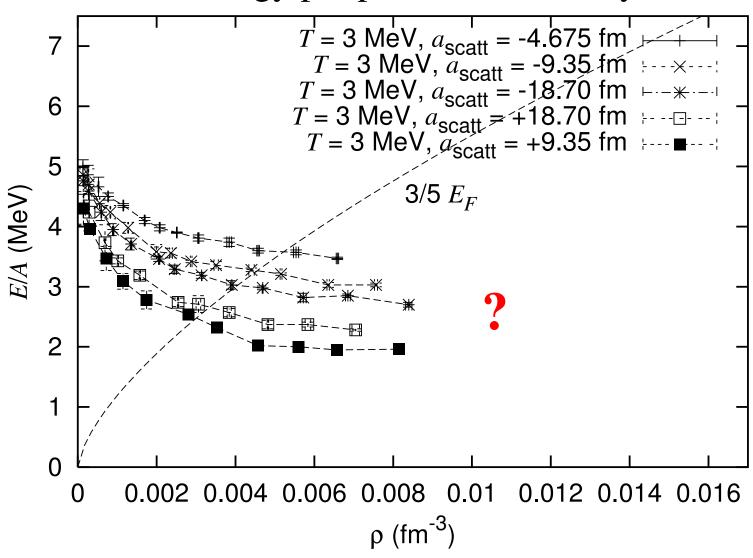
We sum the full set of bubble diagrams

$$\sum_{n=0}^{\infty} \cdots$$

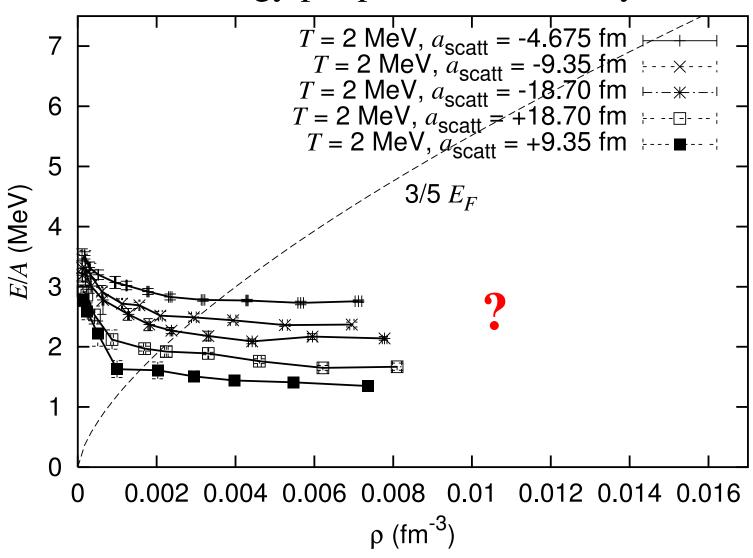
# Energy per particle vs. density



# Energy per particle vs. density



## Energy per particle vs. density



# Virial expansion

Expansion in fugacity,

$$z = e^{\beta \mu}$$

Thermal wavelength

$$\lambda_T = \sqrt{\frac{2\pi}{mk_B T}} = \left(\frac{2\pi}{m}\right)^{1/2} \beta^{1/2}$$

$$\frac{\ln Z_G}{V} = \frac{P}{k_B T} = \frac{2}{\lambda_T^3} \left[ z + b_2(T)z^2 + \cdots \right]$$

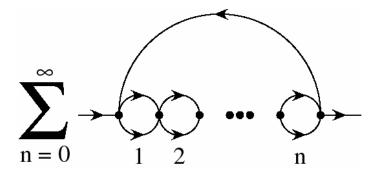
In the unitary regime (zero range and infinite scattering length)

$$b_2(T) \to \frac{3\sqrt{2}}{8} \approx 0.530$$

Second virial coefficient determined by two-particle interactions

# High temperature/low density

At low densities we can compute the self-energy by summing bubble chain diagrams



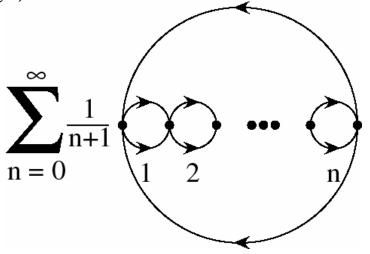
At T = 0, other diagrams are suppressed by factors of  $k_F |a_{nn}|$ . For T > 0, the thermal wavelength

$$\lambda_T = \sqrt{\frac{2\pi}{mT}}$$

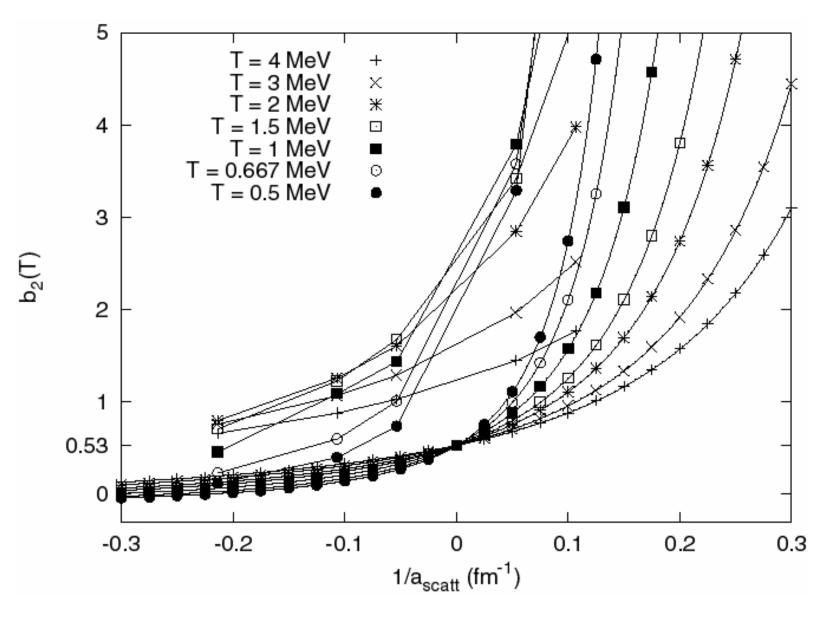
replaces the scattering length when it is the smaller of the two... comparable to expanding in fugacity

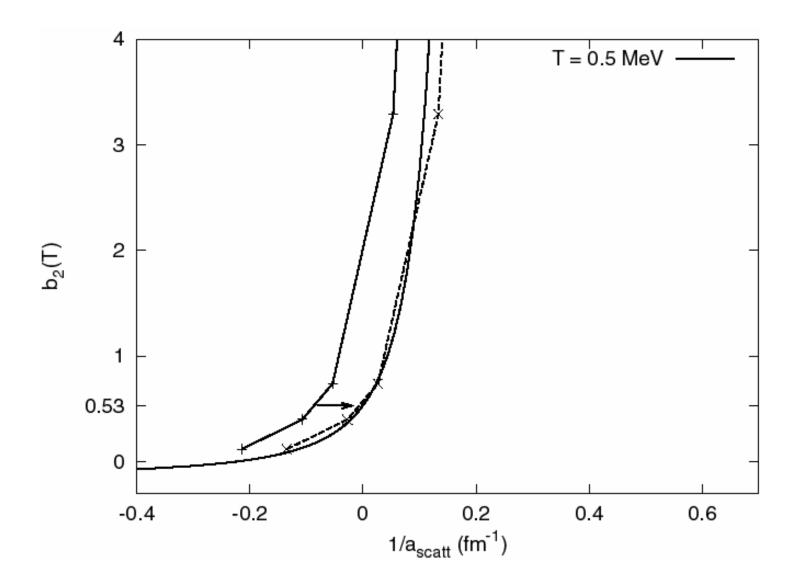
$$z = e^{\beta \mu}$$

Similarly we can compute the logarithm of the partition function (sum of connected diagrams with no external legs).



#### Lattice vs. continuum virial coefficients





# Fixing the problem...

Tune the operator coefficient to give the correct second virial coefficient at the given simulation temperature [D.L., Schäfer, nucl-th/0509017]

#### or

Use an improved lattice action [R. Thomson, D.L., work in progress]

# Scaling limit (scattering length $\rightarrow \pm \infty$ )

Hamiltonian lattice (temporal spacing = 0)

$$H - \mu N = \sum_{\vec{n}_s, i} \left[ (-\mu + \frac{3}{m}) a_i^{\dagger}(\vec{n}_s) a_i(\vec{n}_s) \right]$$

$$- \frac{1}{2m} \sum_{\vec{n}_s, l_s, i} \left[ a_i^{\dagger}(\vec{n}_s) a_i(\vec{n}_s + \hat{l}_s) + a_i^{\dagger}(\vec{n}_s) a_i(\vec{n}_s - \hat{l}_s) \right]$$

$$- \frac{\eta}{m} \sum_{\vec{n}_s} a_{\downarrow}^{\dagger}(\vec{n}_s) a_{\uparrow}^{\dagger}(\vec{n}_s) a_{\uparrow}(\vec{n}_s) a_{\downarrow}(\vec{n}_s)$$

$$\eta \simeq 3.96$$

Three-dimensional attractive Hubbard model with

$$U = -7.92t$$

We note that

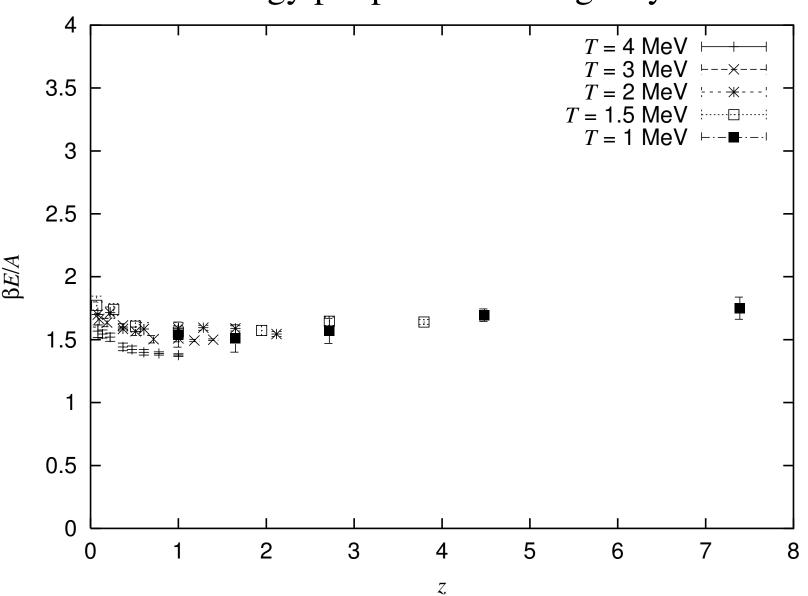
$$Z_G = Tr \left[ -\beta \left( H - \mu N \right) \right]$$

is a function of only the two dimensionless quantities

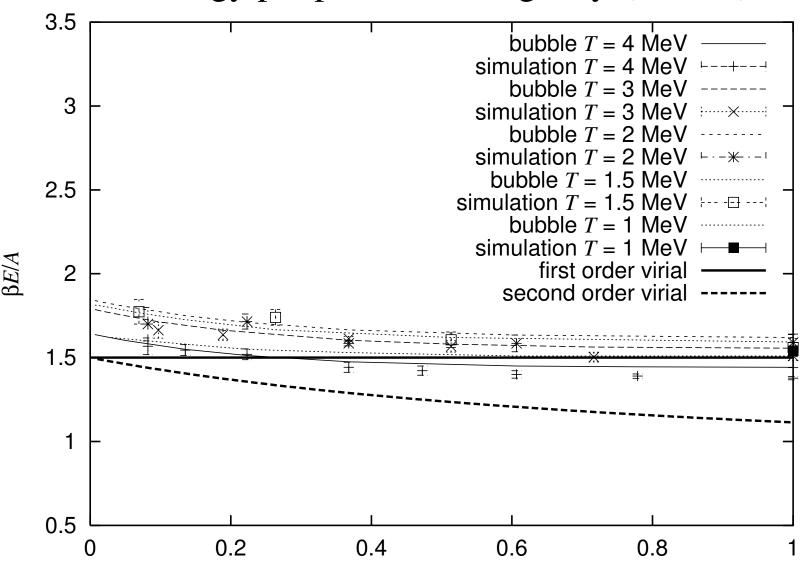
$$\frac{\beta}{2m} = \frac{1}{2m^{phys}T^{phys}a^2}$$
$$\beta\mu = \frac{\mu^{phys}}{T^{phys}}$$

The dependence on the lattice spacing must drop out, so observables at different temperatures can be rescaled to a single universal function

### Energy per particle vs. fugacity

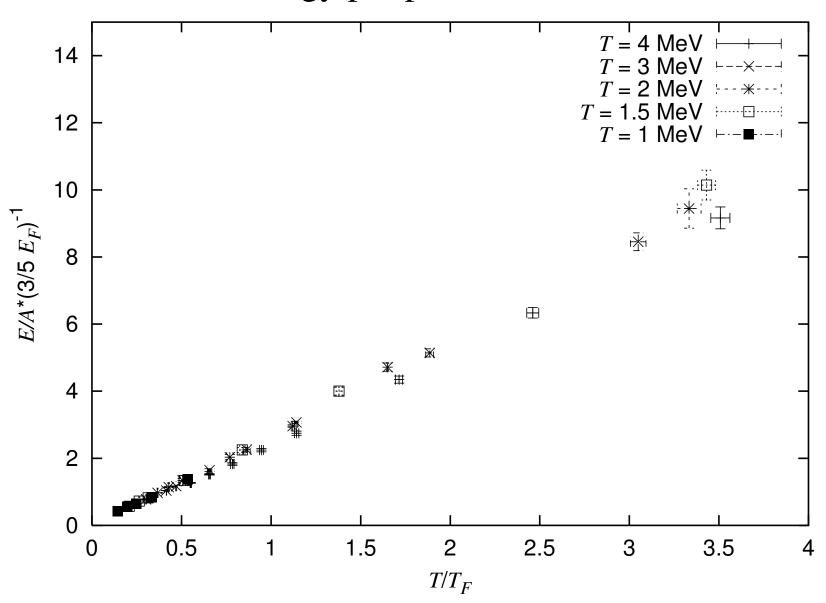


#### Energy per particle vs. fugacity (small z)

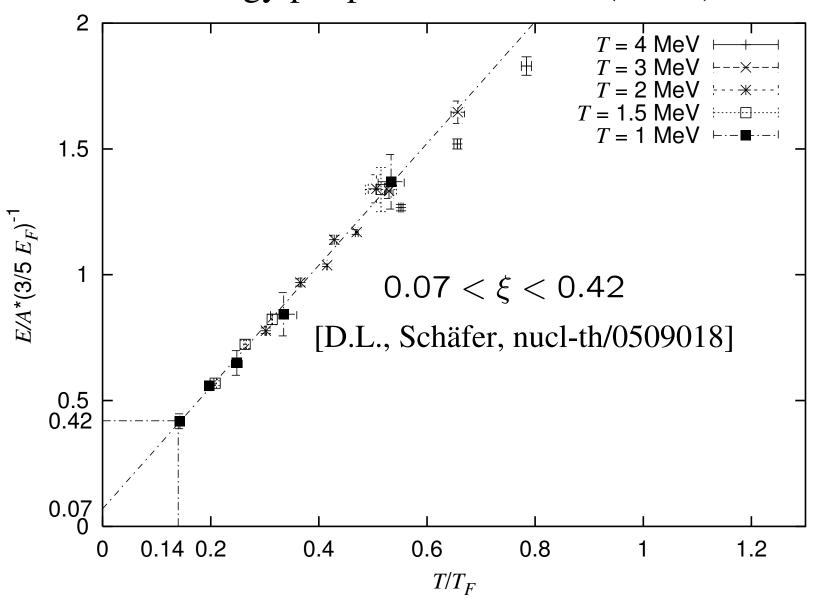


Z

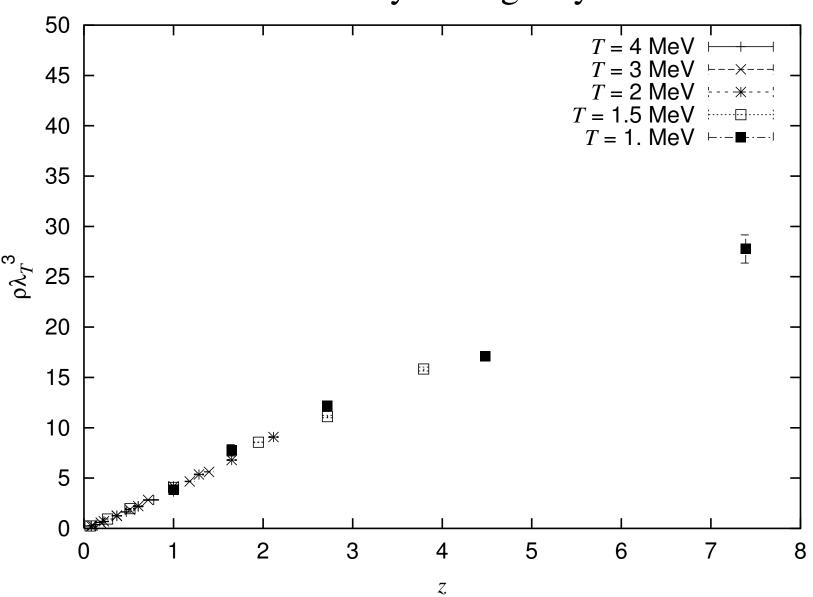
Energy per particle vs.  $T/E_F$ 



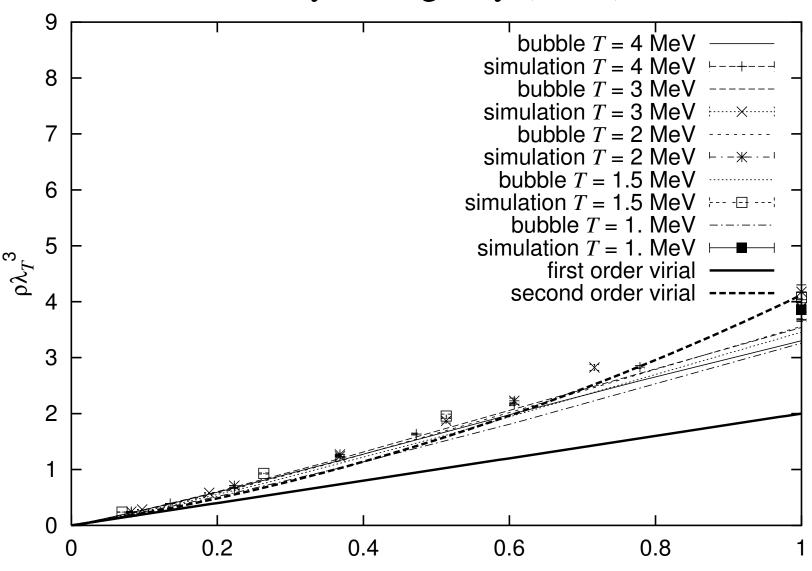
#### Energy per particle vs. $T/E_F$ (low T)



#### Density vs. fugacity

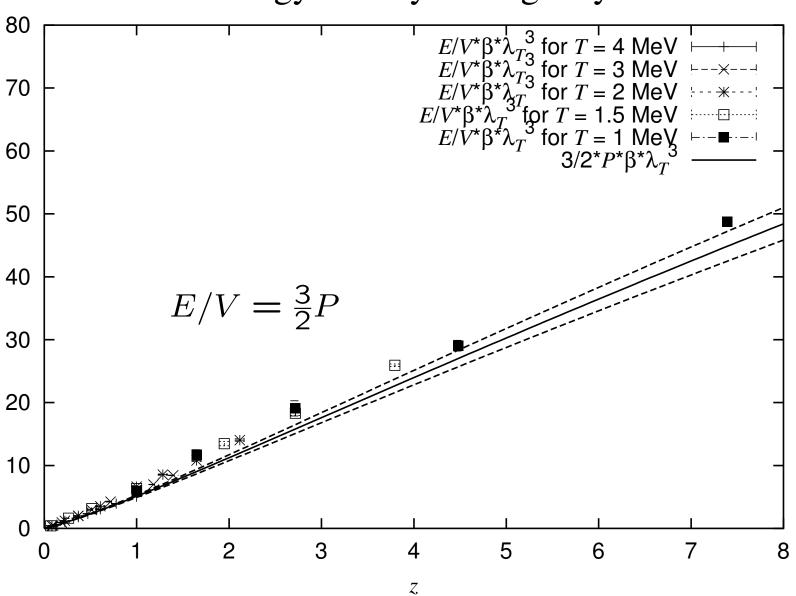


#### Density vs. fugacity (low z)



Z

#### Energy density vs. fugacity

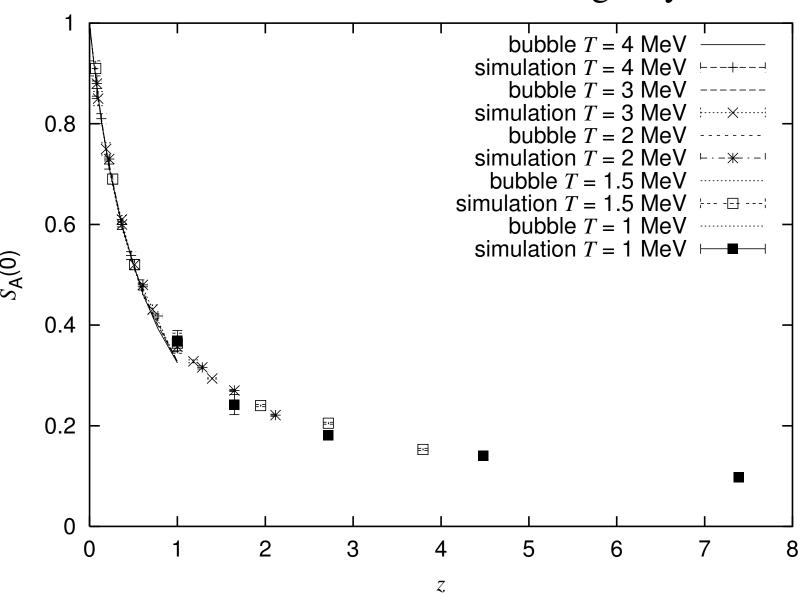


# Axial structure factor (or spin susceptibility)

$$S_A(0) = \frac{1}{\langle \hat{N} \rangle} \langle (\hat{N}_{\uparrow} - \hat{N}_{\downarrow}) (\hat{N}_{\uparrow} - \hat{N}_{\downarrow}) \rangle$$

measures spin correlation ... decreases if neutrons form spin zero pairs

#### Axial structure factor vs. fugacity

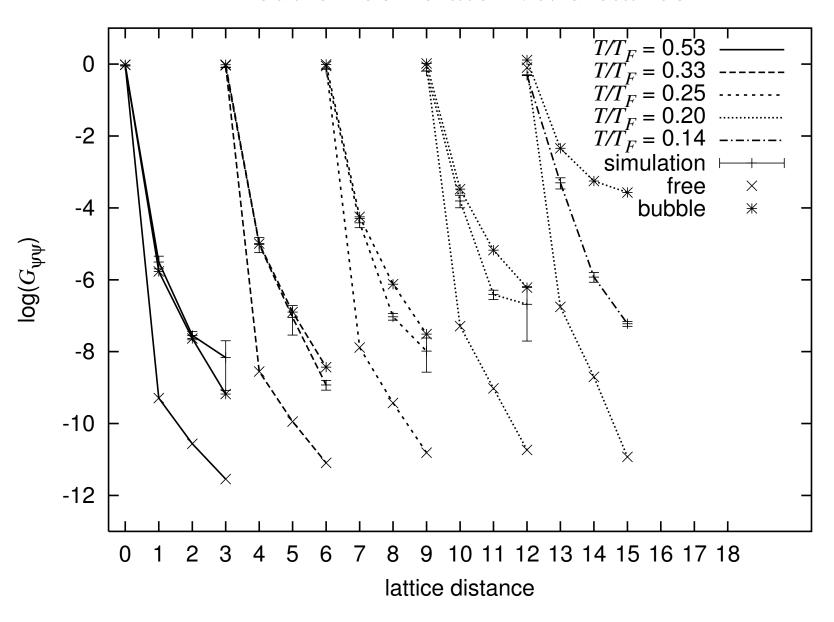


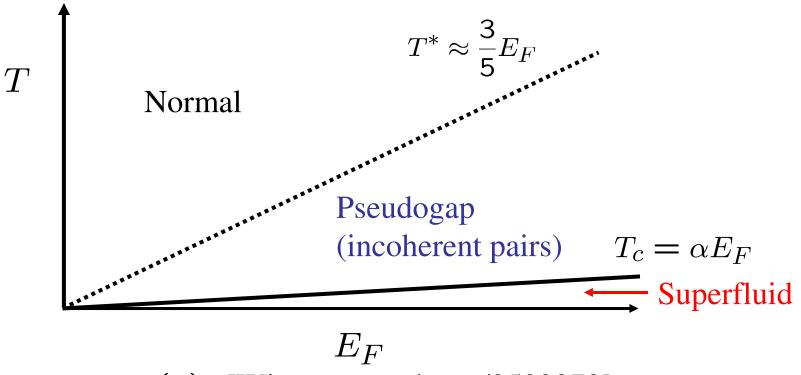
#### Dineutron correlator

$$\left\langle a_{\uparrow}(\vec{n}_s)a_{\downarrow}(\vec{n}_s)a_{\downarrow}^{\dagger}(0)a_{\uparrow}^{\dagger}(0)\right\rangle$$

Look for long range order and Bose condensate of dineutron pairs

#### Dineutron correlator vs. distance





$$\alpha = 0.035(4)$$
 [Wingate, cond-mat/0502372]

$$\alpha = 0.152(7)$$
 [Prokof'ev, Svistunov, preliminary]

$$\alpha = 0.22(3)$$
 [Bulgac, Drut, Magierski, cond-mat/0505374]

$$\alpha$$
 < 0.14 [D.L., Schäfer, nucl-th/0509018]

# Road map

- 1. Larger simulations of cold dilute neutron matter
- 2. Improved actions
- 3. Few body systems and three-body forces
- 4. Asymmetric nuclear matter without pions
- 5. Nuclear matter with pions