# Theory and Calculation of X-ray spectra

J. J. Kas



# **Theoretical** Spectroscopy Calculations

• GOAL: Next Generation Theory for Next Generation X-ray Sources

## • TALK

I Introduction Current approximations II State-of-the-art Ab initio calculations III Next generation Multi-electron excitations & strong correlations

# What is x-ray absorption?



# What is x-ray absorption?

Cu Absorption (log-log)



# What is x-ray absorption?

Cu Absorption (log-log)



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# What is EXAFS?



# What is EXAFS? What is EXAFS?<br>• Quantum interference





# Qualitative Interpretation



MSE-333 5/8/2013

# Introduction

Promise of Next generation Light sources: NSLSII, LCLS …

Powerful probes of multiple length and time scales Partial solution to DOE 5 Grand Challenges, etc.

BUT



# I Current methods

JULY 2NDO

### **Reviews Modern Physics**

VOLUME 72 . NUMBER 3 Fublished by THE AMERICAN PHYSICAL SOCIETY through the AMERICAN INSTITUTE OF PHYSICS

# **J. J. Rehr & R.C. Albers Rev. Mod. Phys. 72, 621 (2000)**

## EXAFS & XANES

# Automated & Integrated Automated & Integrated<br>XAS codes<br>FEFF<br>> 20 yrs development

XAS codes

FEFF

# Real-space Green's Function (RSGF) Theory VIS – X-ray



Paradigm shift:

Ψ

$$
\mu(E) \sim -\tfrac{1}{\pi}\mathrm{Im} \bra{\mathsf{i}} \hat{\epsilon} \cdot \mathbf{r}^{\prime} \operatorname{G}(\mathbf{r}^{\prime}, \mathbf{r}, E) \, \hat{\epsilon} \cdot \mathbf{r} \ket{\mathsf{i}}
$$

EFFICIENT: No sums over final states !



- features, including absence of white line at  $L_2$ -edge.
- white line intensity.

## Spectra: RIXS, XES, Compton, …

PHYSICAL REVIEW B 83, 235114 (2011)

### Real-space Green's function approach to resonant inelastic x-ray scattering

J. J. Kas,<sup>1</sup> J. J. Rehr,<sup>1,\*</sup> J. A. Soininen,<sup>2</sup> and P. Glatzel<sup>3</sup>

<sup>1</sup>Department of Physics, Box 351560, University of Washington, Seattle, Washington 98195-1560, USA <sup>2</sup>Department of Physics, P.O. Box 64, University of Helsinki, FI-00014 Helsinki, Finland <sup>3</sup>European Synchrotron Radiation Facility, B.P. 220, F-38043 Grenoble, France





# II. State of the art theory **THEPTS<br>FEFF9**<br>JJR et al., Comptes Rendus<br>Physique 10, 548 (2009)

FEFF9

# Physique 10, 548 (2009)

State of the art theory GOAL: ab initio many-body effects no adjustable parameters

GW Self-energy Debye Waller factors RPA core-hole



Need: multiple codes: DFT, self-energy, Phonons No one group or code can do it all



\* Also used in BSE Improves on final state rule, Z+1, half hole

## Many-pole GW Self-energy

**Many-pole GW Self-energy<br>Energy dependent**  $\Sigma(E)=iGW$  **– replaces**  $V_{xc}$ **<br>MFP, energy shifts cf. Hedin-Lundqvist plasmon-pole Many-pole GW Self-energy<br>Energy dependent**  $\Sigma(E)=iGW$  **- replaces**  $V_{xc}$ **<br>MFP, energy shifts cf. Hedin-Lundqvist plasmon-pole<br>Improved XANES** 



### Ab initio XAS Debye Waller Factors  $\int e^{-2\sigma^2 k^2}$ 2

An Initio Determination of Extended X-Ray Absorption Fine Structure Debye-Waller Factors

> Fernando D. Vila, G. Shu, and John J. Rehr Department of Physics, University of Washington, Seattle, WA 98195



### Many Pole model

$$
\sigma^2 = \frac{\hbar}{\mu_i} \int_0^\infty \rho(\omega^2) \coth \frac{\beta \hbar \omega}{2} d\omega
$$

## for phonons

$$
\rho(\omega^2) = \langle Q_i | \delta(\omega^2 - D) Q_i \rangle
$$
  
= {6 - step Lanczos recursion}



\*Phys. Rev. B 76, 014301 (2007)

### Example: XAFS Debye-Waller Factor of Ge



Expt: Dalba et al. (1999)

# Real-time Finite T DFT-MD / XAS **Real-time Finite T DFT-MD /**<br>Nanoscale Pt catalysts\*<br> $Pt_{10} / \gamma$ -Al<sub>2</sub>O<sub>3</sub> Expt. vs DFT<br>+ FEFF X

# + FEFF XAS Theory\*



\*F. Vila, J.J. Rehr, J. Kas, R.G. Nuzzo, A.I. Frenkel, Phys. Rev. B 78, 121404(R) (2008)



MANGANESE

**MUINOSHS** 

**IRON** 

**COBALT** 

**NICKEL** 

**COPPER** 

ZINC

21

III. Next generation XAS Calculations Next generation XAS Calculations<br>- beyond-quasiparticles

II. Next generation XAS Calcula<br>- beyond-quasiparticles<br>Excitonic effects BSE<br>Multi-electron excitations Multi-electron excitations Strong correlations



## Plus integrated Codes: DFT, GW/BSE, Real-time etc.

# $GW$  Bethe-Salpeter Equation calculations of core excitation spectra<br>the-Salpeter equation calculations of core excitation spectra<br>Department of Physics, University of Washington, Seattle, Washington 98195, USA

E. L. Shirley

with experiment and with other theoretical approaches.

BSE OCEAN\*



\*Obtaining Core Excitations  $\begin{array}{ll}\n\textbf{BSE} \\
\textbf{BSE} \\
\textbf{F}_{\text{other}} \\
\text{SINR} \\
\text{BINR} \\$ 

 $\begin{array}{lll} \mathcal{P}^{land 20899, \; USA} & \mathbf{BSE} \\ \mathcal{S} & \mathcal{S}^{total} & \mathcal{S}^{total} \\ \mathcal{S}^{total} & \mathcal{S}^{total} & \mathcal{S}^{total} & \mathcal{S}^{$ + PAW + MPSE

cf. EXC, EXC!TING, BerkeleyGW

### Many-body Amplitudes in XAS  $S_n^2$ 2

**Many-body Amplitudes in XAS**<br>• Many-body X\*S ≈ Convolution<br> $\mu(\omega) = \int_0^\infty d\omega' \tilde{A}(\omega, \omega') \mu_{qp}(\omega - \omega')$ **Many-body Amplitudes in XAS** S<br>
• Many-body X\*S ≈ Convolution<br>  $\mu(\omega) = \int_0^\infty d\omega' \tilde{A}(\omega, \omega') \mu_{qp}(\omega - \omega')$ <br>  $\equiv \langle \mu_{qp}(\omega) \rangle$  Na XPS<br>
• Source: multi-electron excitations • Many-body X\*S ≈ Convolution<br>  $\mu(\omega) = \int_0^\infty d\omega' \tilde{A}(\omega, \omega') \mu_{qp}(\omega - \omega')$ <br>  $\equiv \langle \mu_{qp}(\omega) \rangle$  Na XPS<br>
• Source: multi-electron excitations<br>
• Spectral function ~ XPS<br>  $A = - \text{Im } G$ Na XPS

- 
- $A = \operatorname{Im} G$



# Cumulant<br>  $\Sigma = i GWT$ <br>  $W = \epsilon^{-1}v$ Which Green's function ? GW vs Cumulant

GW Cumulant

J. Phys.: Condens. Matter 11 (1999) R489–R528<br>
relation effects in electron spectroscopies and the  $\Sigma = i\ GW\Gamma$ <br>
pproximation  $W = \epsilon^{-1}v$ <br>
Lars Hedin<br>  $G( w ) = G_0 + G_0 \Sigma G$ <br>  $G = \int_{\Gamma} + G_0 \Sigma G$ <br>  $\Gamma = 1$   $\Sigma^{GW} = iGW$   $\Gamma = 1$  $G(t) = G_0(t) e^{C(t)}$ <br> $C \sim \text{Im } \Sigma^{GW}$  $G$ <br>  $T=1$   $\Sigma^{GW}$  = iGW<br>  $T=1$   $\Sigma^{GW}$  = iGW<br>  $\Gamma^{SM}$  = iGW<br>  $\Gamma^{SM}$ 

Similar ingredients; all many-body effects in  $\Sigma^{GW}$ 

# Test: Satellites in XPS of simple metals\*

**ES IN XPS of simple meta**<br>
Satellites in Na and Al Spectral Functions<br>
(*b Initio* Cumulant Expansion<br>
tiawan,<sup>1,2</sup> L. Hedin,<sup>1</sup> and K. Karlsson<sup>3</sup><br>
Phys Rev Lett 77, 2268 (1996)<br> **Muasi-particle peak agrees wit** 

Good news: GW quasi-particle peak agrees with XPS

**Bad news:** GW has one satellite  $\omega$  wrong energy  $\omega_{\text{max}}$  (b)

XPS has multiple satellites



# Test 2: Multiple Satellites in XPS of Si

PRL 107, 166401 (2011)

PHYSICAL REVIEW LETTERS

week ending 14 OCTOBER 2011

### **Valence Electron Photoemission Spectrum of Semiconductors: Ab Initio Description of Multiple Satellites**

Matteo Guzzo,<sup>1,2,\*</sup> Giovanna Lani,<sup>1,2</sup> Francesco Sottile,<sup>1,2</sup> Pina Romaniello,<sup>3,2</sup> Matteo Gatti,<sup>4,2</sup> Joshua J. Kas,<sup>5</sup> John J. Rehr,<sup>5,2</sup> Mathieu G. Silly,<sup>6</sup> Fausto Sirotti,<sup>6</sup> and Lucia Reining<sup>1,2,†</sup>



### XPS Intensity Ratios in Molecules

PRL 108, 193005 (2012)

PHYSICAL REVIEW LETTERS

week ending 11 MAY 2012

### **Nonstoichiometric Intensities in Core Photoelectron Spectroscopy**

J. Söderström, <sup>1</sup> N. Mårtensson, <sup>1</sup> O. Travnikova, <sup>2</sup> M. Patanen, <sup>2</sup> C. Miron, <sup>2</sup> L. J. Sæthre, <sup>3</sup> K. J. Børve, <sup>3</sup> J. J. Rehr,<sup>4</sup> J. J. Kas,<sup>4</sup> F. D. Vila,<sup>4</sup> T. D. Thomas,<sup>5</sup> and S. Svensson<sup>1</sup>







<XPS> = XAS

Intensity ratio  $\sim$  Z

including all losses explains nonstoichiometric ratios

# Question: Can the cumulant method work<br>for CT satellites in correlated systems ? for CT satellites in correlated systems ?

# Hedin's answer \* MAYBE

"Calculation similar to core case … but with more complicated fluctuation potentials … Iculation similar to core case ... but with more<br>complicated fluctuation potentials ...<br>ot question of principle, but of computational<br>\* L. Hedin, J. Phys.: Condens. Matter 11, R489 (1999)

… not question of principle, but of computational work..."

# Approach: Cumulant expansion in real-time<br><br>leal time cumulant approach for charge transfer satellites in x-ray photoemission

J. J. Kas,<sup>1</sup> F. D. Vila,<sup>1</sup> J. J. Rehr,<sup>1</sup> and S. Chambers<sup>2</sup>

### arXiv:1408.2508

<sup>1</sup>Dept. of Physics, Univ. of Washington, Seattle, WA 98195-1560 <sup>2</sup>Physical Sciences Division, Pacific Northwest National Laboratory, Richland, WA 99352 (Dated: August 7, 2014)

**Langreth** 
$$
C(t) = \sum_{\mathbf{q}, \mathbf{q'}} V_{\mathbf{q}}^* V_{\mathbf{q'}} \int d\omega S(\mathbf{q}, \mathbf{q'}, \omega) \frac{e^{i\omega t} - i\omega t - 1}{\omega^2} = \int d\omega \beta(\omega) \frac{e^{i\omega t} - i\omega t - 1}{\omega^2}
$$

RT Cumulant: RT-TDDFT

$$
S(\boldsymbol{q},\boldsymbol{q'},\omega)=\int\frac{dt}{2\pi}e^{i\omega t}\langle\rho_{\boldsymbol{q}}(t)\rho_{\boldsymbol{q'}}(0)\rangle
$$



## XPS & Real-space Interpretation



Interpretation: satellites due to local ligand-metal charge fluctuations at frequency  $\sim \omega_{ct}$  in response to transient core-hole potential JJ Kas, FD Vila, JJR & S.Chambers arXiv:1408.2508

# Conclusions and Outlook

EXAFS – well understood with current methods RSGF + DFT/MD **Conclusions and Outlook**<br>EXAFS - well understood with current methods<br>RSGF + DFT/MD<br>XANES – semi-quantitative state of the art theory<br>GW, GW/BSE, DW factors, ... GW, GW/BSE, DW factors, … XAS, XPS – Multi-electron excitations & correlations understood with current methods<br>RSGF + DFT/MD<br>ni-quantitative state of the art theory<br>GW, GW/BSE, DW factors, ...<br>Multi-electron excitations & correlations<br>via quasi-boson/cumulant theory<br>n eslaulations nessible in exhitre

Full spectrum calculations possible in arbitrary materials with NG theory and codes.

### Rehr group & collaborators



# Thanks for Listening!

# Many Success Stories

# XAS of doped SiC

M Ohkubo et al., Sci. Rep., 2, 831 (2012)





"FEFF8.4 played an important role in analysis of SCIENTIFIC<br>
REPARTS<br>
We Are also spectra of N-doped SiC.<br>
The also spectra of N-doped SiC. I appreciate your<br>
States in the spectra of N-doped SiC. I appreciate your<br>
States spectra of N-doped SiC. I appreciate your<br>
State effort for developing this fantastic software …" S<br>
Fray absorption near edge spectroscopy<br>
ith a superconducting detector for<br>
trogen dopants in SiC<br>
wed an important role in analysis of<br>
yed an important role in analysis of<br>
yed an important role in analysis of<br>
loping

# Cumulant Expansion & Spectral Function

$$
G_k(t) = e^{i\epsilon_k^0 t} e^{C(t)} \qquad \qquad C(t) = \int d\omega' \beta(\omega') \frac{e^{i\omega' t} - i\omega' t - 1}{\omega'^2}
$$

### Spectral Function

**Spectral Function**  
\n
$$
A_k(\omega) = \int \frac{dt}{2\pi} e^{i(\omega - \epsilon_k)t} \exp \left\{ \int d\omega' \beta(\omega') \frac{e^{i\omega' t} - i\omega' t - 1}{\omega'^2} \right\}
$$
\nSee L. Hedin, J. Phys.: Condens. Matter **11**, R489 (1999)

\ncf. D.C. Langreth *Phys. Rev.* B **1**, 471 (1970) (Linked cluster theorem)

\*see L. Hedin, J. Phys.: Condens. Matter 11, R489 (1999)

# Other properties

Kernel:  $\beta(\omega) = |\text{Im }\Sigma_k(\omega)| \sim \text{Im } \varepsilon^{-1} \sim \text{loss function}$ Peaks in  $A(\omega) \sim$  Peaks in loss function  $Z = e^{-\overline{n}}$   $\overline{n} = \sum_{q} g_q^2 / \omega_q^2 = \int \beta(\omega) / \omega^2 d\omega$ .

 $\overline{n} = 0.201 r_s^{3/4}$  mean number of shake-up bosons Dimensionless measure of correlation strength  $n \sim 0.3$  to 0.4 even in s-p systems!

# IV. Problems with Cumulant Expansion\*<br>Electron spectral function

# BAD: Pathologies at Fermi energy\* 1 FOW

Time-ordered GF (TO) lacks satellites on both sides of QP peak near Fermi energy

Source: recoil approximation: neglect of "exchange diagrams"

### Electron spectral function



# Debye-Waller factors  $\sigma^2$  in Zr tungstate in Zr tungstate





# FIX: Retarded Cumulant expansion\*<br>
PHYSICAL REVIEW B 00, 005100 (2014)<br>
Cumulant expansion of the retarded one-electron Green function

J. J. Kas, 1,\* J. J. Rehr, 1,2,+ and L. Reining<sup>3,2,‡</sup> <sup>1</sup>Department of Physics, University of Washington, Seattle, Washington 98195, USA <sup>2</sup>European Theoretical Spectroscopy Facility (ETSF) <sup>3</sup>Laboratoire des Solides Irradiés, École Polytechnique, CNRS, CEA-DSM, F-91128 Palaiseau, France (Received 31 January 2014; revised manuscript received 28 July 2014; published xxxxxx)

## Retarded formalism Retarded formalism

$$
G_k^R(t) = -i\theta(t)e^{-i\epsilon_k^{HF}t}e^{\tilde{C}_k^R(t)},
$$
  
\n
$$
\tilde{C}_k^R(t) = \int d\omega \frac{\beta_k(\omega)}{\omega^2} (e^{-i\omega t} + i\omega t - 1),
$$
  
\n
$$
\beta_k(\omega) = \frac{1}{\pi} |\text{Im}\Sigma_k^R(\omega + \epsilon_k)|,
$$



\*Phys. Rev. B. (in press, Aug. 2014); arXiv:1402.0022

# Retarded Cumulant (GC) gives<br> **GOOD** quasi-particle properties GOOD quasi-particle properties



## Many-body Amplitudes in X-ray Spectra

**Many-body Amplitudes in X-ray Spe**<br>• Many-body X\*S ≈ Convolution<br> $\mu(\omega) = \int_0^\infty d\omega' \tilde{A}(\omega, \omega') \mu_{qp}(\omega - \omega')$ Many-body Amplitudes in X-ray Spectra<br>• Many-body X\*S ≈ Convolution<br> $\mu(\omega) = \int_0^\infty d\omega' \tilde{A}(\omega, \omega') \mu_{qp}(\omega - \omega')$ <br> $\equiv \langle \mu_{qp}(\omega) \rangle$ • Explains crossover: adiabatic<br>to sudden transition =  $\int_0^\infty d\omega' \tilde{A}(\omega, \omega') \mu_{qp}(\omega -$ <br>
=  $\langle \mu_{qp}(\omega) \rangle$ <br>
crossover: **adiabatic**<br>
to sudden transition<br>  $|z| = |g_q^{ext}|^2 + |g_q^{intrim}|^2 \cdot 2 g_q^{ext} g_q^{intrim}$  $\langle \tilde{A}(\omega,\omega')\mu_{qp}(\omega-\omega')\rangle$ <br>adiabatic<br>sudden transition<br> $|g_q^{intrim}|^2$ , 2  $g_q^{ext}g_q^{intrim}$ 

to sudden transition

 $|g_q|^{2} = |g_q^{ext}|^2 + |g_q^{intrim}|^2$ .  $2 g_q^{ext} g_q^{intrim}$ 

Interference reduces loss!

# Strong correlation effects: 1. Hubbard corrections **Strong correlation eff**<br> **1. Hubbard correction**<br>
PHYSICAL REVIEW B 85, 165123 (2012<br>
Hubbard model corrections in real-space x-ray s<br>
Towfiq Ahmed, J. J. Kas, and J. J. Reh<br>
CRPA calculated U







# FIX: Next generation approach Quasi-Boson Approximation

IDEA: Neutral Excitations - plasmons, phonons,<br>electron-hole pairs, magnons are **bosons**  $\gamma$ **FIX: Next generation approach<br>
Quasi-Boson Approximation<br>
IDEA: Neutral Excitations - plasmons, phonons,<br>
electron-hole pairs, magnons are bosons**  $\gamma$ **FIX: Next generation approach**<br>Quasi-Boson Approximation<br>IDEA: Neutral Excitations - plasmons, phonons,<br>electron-hole pairs, magnons are **bosons γ**<br>Many-body Model:  $|N\rangle = |e^r, h, \gamma \rangle$ 

Many-body Model:  $|N\rangle = |e^{\tau}, h, \gamma \rangle$ 

- 
- 
- 
- 

 $V^n \to -\text{Im }\varepsilon^{-1}(\omega_n, q_n)$  "fluctuation potentials"

• Excitations:  $H_v = \sum_n \omega_n a_n^{\dagger} a_n$ <br>
• Electrons:  $h' = \sum_k \epsilon_k c_k^{\dagger} c_k$ <br>
• e-boson coupling  $V_{pv} = \sum_{nkk'} [V_{kk'}^n a_n^{\dagger} + (V_{kk'}^n)^* a_n] c_k^{\dagger} c_{k'}$ <br>
• Core-hole-boson coupling:  $V_{vc} = -\sum_n V_{bb}^n (a_n^{\dagger} + a_n)$ <br>  $V^n \rightarrow -Im \ \varepsilon^{-1} (\omega_m$ 

# GW++: Effective Green's Function **GW++: Effective Green's Function**<br>- **perturbation theory\***<br> $g_{\text{eff}}(\omega) = e^{-a} \left[ g'(\omega) + \sum_n \left( \frac{V_{bb}^n}{\omega_n} \right)^2 g'(\omega - \omega_n) - 2 \sum_n \frac{V_{bb}^n}{\omega_n} g'(\omega - \omega_n) V^n \right]$



presence of core-hole (~ final state rule! )

$$
g'\left(\omega\right) \equiv \left[\omega - h' - \Sigma\left(\omega\right) + i\gamma\right]^{-1}
$$

\*L. Campbell, L. Hedin, J. J. Rehr, and W. Bardyszewski, Phys. Rev. B 65, 064107 (2002)

# 2. Multiplets a la GW/BSE

Phys. Rev. B 86, 195135 (2012)

Phys. Rev. B 86, 195135 (2012)<br>
alculations of transition metal L-shell spectra<br>
J. Vinson and J. J. Rehr<br>
regt. of Physics,  $\frac{\text{Uncoal}}{\text{D}}$  (Dated: huly 3, 2012)<br>
the Euclide Salpher organic (Dated: huly 3, 2012)<br>
alcul

### GW/BSE + Spin & SO + MPSE + Slater F, G







### 3. Charge Transfer Satellites

PHYSICAL REVIEW B

VOLUME 60, NUMBER 11

15 SEPTEMBER 1999-I

### Transition from the adiabatic to the sudden limit in core-level photoemission: A model study of a localized system

J. D. Lee and O. Gunnarsson Max-Planck Institut fur Festkörperforschung, Heisenbergstrasse 1, D-70569 Stuttgart, Germany

L. Hedin Max-Planck Institut für Festkörperforschung, Heisenbergstrasse 1, D-70569 Stuttgart, Germany



Approaches: Anderson impurity, DMFT, multiplets, …. BUT still no satisfactory 1<sup>st</sup> principles theory ...

# Simple picture: 2-state model of LGH\*



## Goal: X-ray Spectroscopy Beamline



# EU Model



A new facility for research!

# European **Theoretical Spectroscopy**

### Facility www.etsf.eu

### **European Theoretical Spectroscopy Facility**

### An innovative project...

The Nanoquanta Network of Excellence, partially financed by the Furragism Union teader the Stillh Fra mework Programme, is building up a new type of facility called European Theoretical Spectroscopy Facility (ETSF). On the lines of the synchrotron facilities successfully shared by researchers across Europe, the FTSF will offer nears the opportunity to hearbit, menthe latest developments in the field of electronic exited states in matter.

### To burst the secrets of matter

What happens in matter when it is exposed to ra-What does the response of the parte-<br>the radiation – that is, the spectra we<br>sire – testh us? That is what the resear-<br>ers of the FTSF want to know



In matter, the electrons occupy energy<br>Toyou, When absorbing redistion, there can be promoted us a higher lead, depending on the characteristics of the<br>system. These transitions must then be calculated taking into account the fact<br>that the electrons interior



Nevertheless much computer power is necesary to potein a reliable exeribility

I real matter. To do that, researchers develop software which arensates the coordistant into the language of com-**Euters** 

### 1-The core of the ETSF

A network of prominent European Condensed Matter theory groups, which will take responsibility for management of the ETSF and reach out

They developed, thanks to a 15 years collaboration, many of the theoretical approaches and computational solutions which are commonly used in the framework of calculations of electronic excitations from first principals.

### The ETSF will bridge the gap between fundamental research and technological applications

Theoretical spectroscopy and fight against cancer, an exal

- Theory and sof nly to know the properties Those it works and fromof marrer birt. fort to consome more ations.
- ford to consider important means time.<br>  $B_2$  -anary let the Green in the second Freezan (G+1), which gives the Aequates Victorial ethical information conservations, the<br>
gives the Aequates Victorial ethical consideration state)

A better understanding bill the fit creatent (GFR porphyrins, etc.) behaviour phobles to improve its use as a pip piptal matrix. The fight against the fit is exceeding to the fight against the form power of the state to d ones around them.

### 2-Associate theory groups

A broad and federal community of research groups workting on similar topics

They fully hencfit from the stientific exchange. enabled and stimulated by the ETSF and will be able o advertise on the ETSF website and call for proposals just like members of the Core.

### 3-Users

A large and varied group of researchers from the public or private sector

- Demand for collaboration using the developed sufficiare and support for these calculations is growing rapidly amongst theoreticians and experimentalists, some working at large institutions such as synchrotrons, others in individual laboratories and others in private companies. They benefit four the FTSF in different
- ways, depending on the proximity of the user to the field:
- · download software and use it with the help of publications, manuals and tutorials provided by the ETSF;
- · participate in training events;
- require ongoing consultancy for a project: demand pieces of theory or software that do nut yet exist but are sufficiently relevant tu br devised by a member of the ETSF, in collaboration with the proposer if appropriate; propose a relevant project that requires fulltime work of a member of the ETSF.
- Users of the ETSF may already propose pilot projects on different levels, according to their expertise. The first open call for user projects will start in 2005.



reland manuals and to organize training cycnts

### European Theoretical Spectroscopy Facility



### About the ETSF

### $\overline{\phantom{a}}$  Beamlines

- · Energy Loss Spectroscopy
- · Optics
- ▶ Photo-emission Spectroscopy
- · Quantum Transport
- 
- · Time-resolved Spectroscopy
- · Vibrational Spectroscopy
- o X-Rays Spectroscopy
- ▶ Services
- $\triangleright$  Resources
- · Funding
- Press
- · Impressum

### Home

### **Beamlines**

The ETSF is divided into 7 beamlines, each of which is concerned with a specific scientific topic. A beamline coordinator is responsible for the contact with the users of each line. He/She also serves as the contact person for users who want to submit a proposal to the ETSF.

Further details are available on the beamlines' description.

### **Beamlines and Coordinators**

### **Optics**

Dr. Olivia Pulci University of Rome Tor Vergata, Rome, Italy Olivia.Pulci@roma2.infn.it

### **Energy Loss Spectroscopy**

Dr. Francesco Sottile

Ecole Polytechnique, Palaiseau, France francesco.sottile@polytechnique.edu

### **Quantum Transport**

### **Dr. Peter Bokes**

Slovak University of Technology, Bratislava, Slovakia peter.bokes@stuba.sk

### **Time-resolved Spectroscopy**

### Dr. Alberto Castro

Instituto de Biocomputación y Física de Sistemas Complejos acastro@bifi.es

### **Photo-emission Spectroscopy**

Dr. Claudio Verdozzi Lund University, Lund, Sweden Claudio.Verdozzi@teorfys.lu.se



Networked "theoretical beamlines"

### **Vibrational Spectroscopy**

### Prof. Gian-Marco Rignanese

Université Catholique de Louvain, Louvain-la-Neuve, Belgium gian-marco.rignanese@uclouvain.be

### **X-Rays Spectroscopy**

### **Prof. John Rehr** University of Washington, Seattle, USA jjr@phys.washington.edu