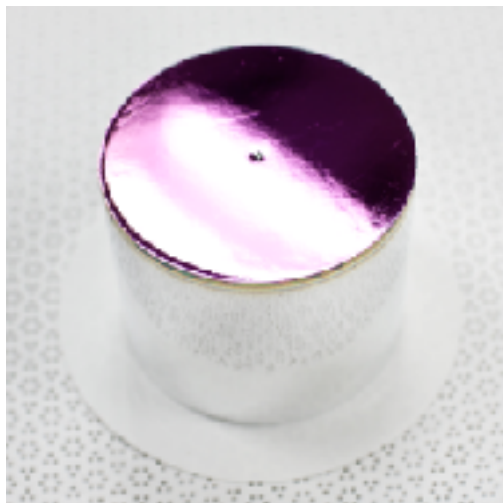


The Neutrino's Mysterious Mass



Jason Detwiler
REU Meeting
July 31, 2017

The Ubiquitous Neutrino

A night sky photograph showing the Milky Way galaxy arching across the frame, with snow-capped mountains visible in the foreground. The stars are sharp and numerous, and the overall color palette is dominated by deep blues and blacks.

1 neutrino for every ~ 3 photons

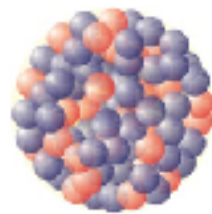
The Sun in Neutrinos



Credit: SuperKamiokande
Actual size of sun: ~one pixel

The Neutrino

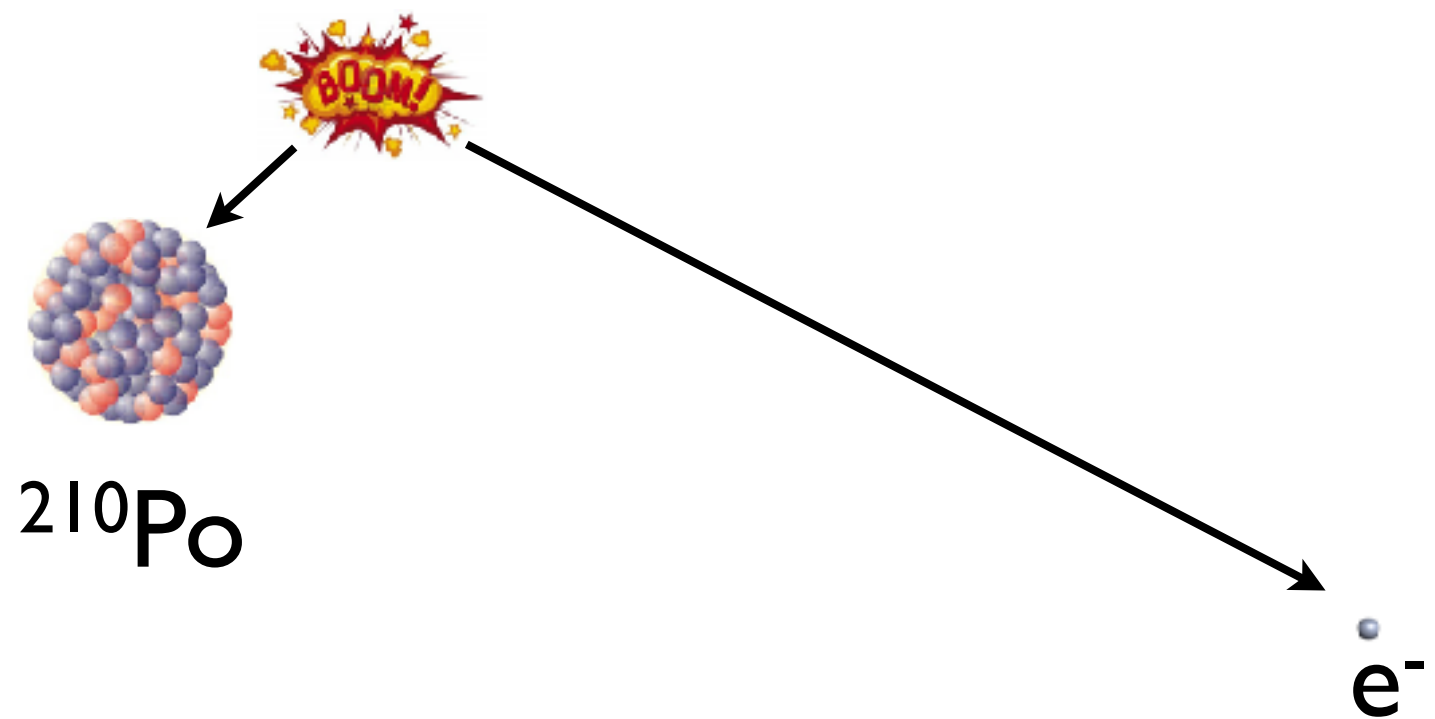
Meitner and
Hahn (1911):



^{210}Bi

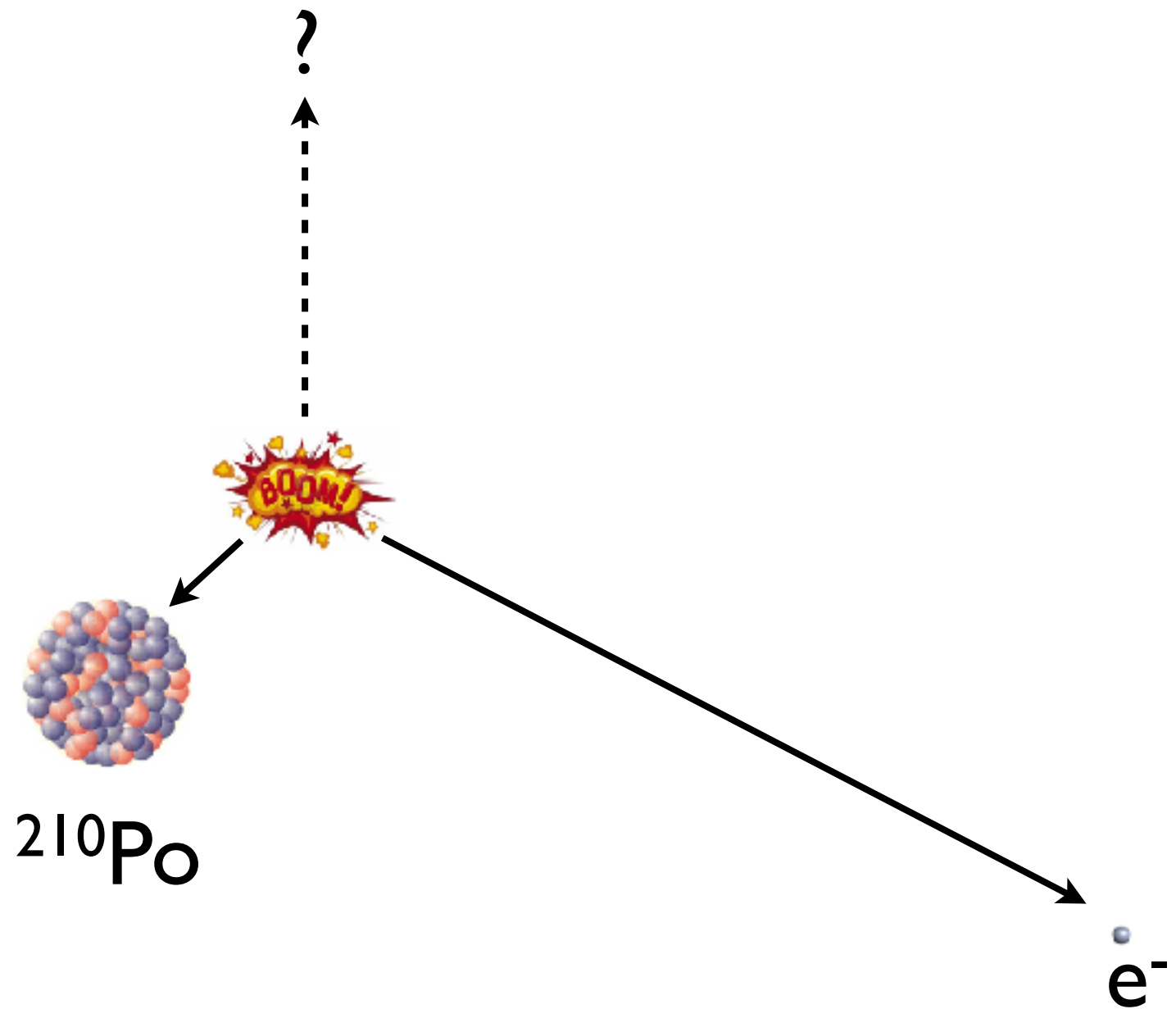
The Neutrino

Meitner and
Hahn (1911):



The Neutrino

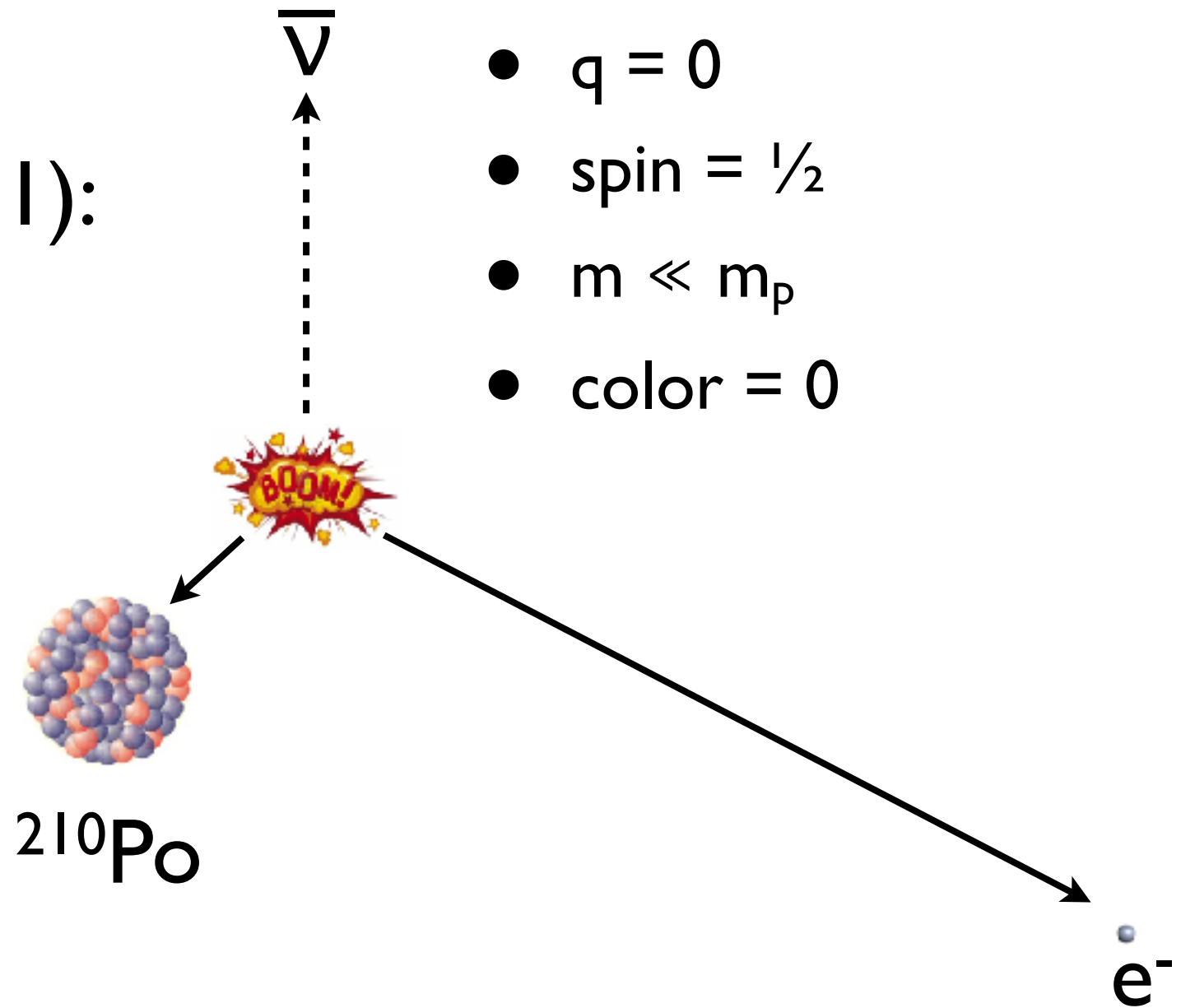
Meitner and
Hahn (1911):



The Neutrino

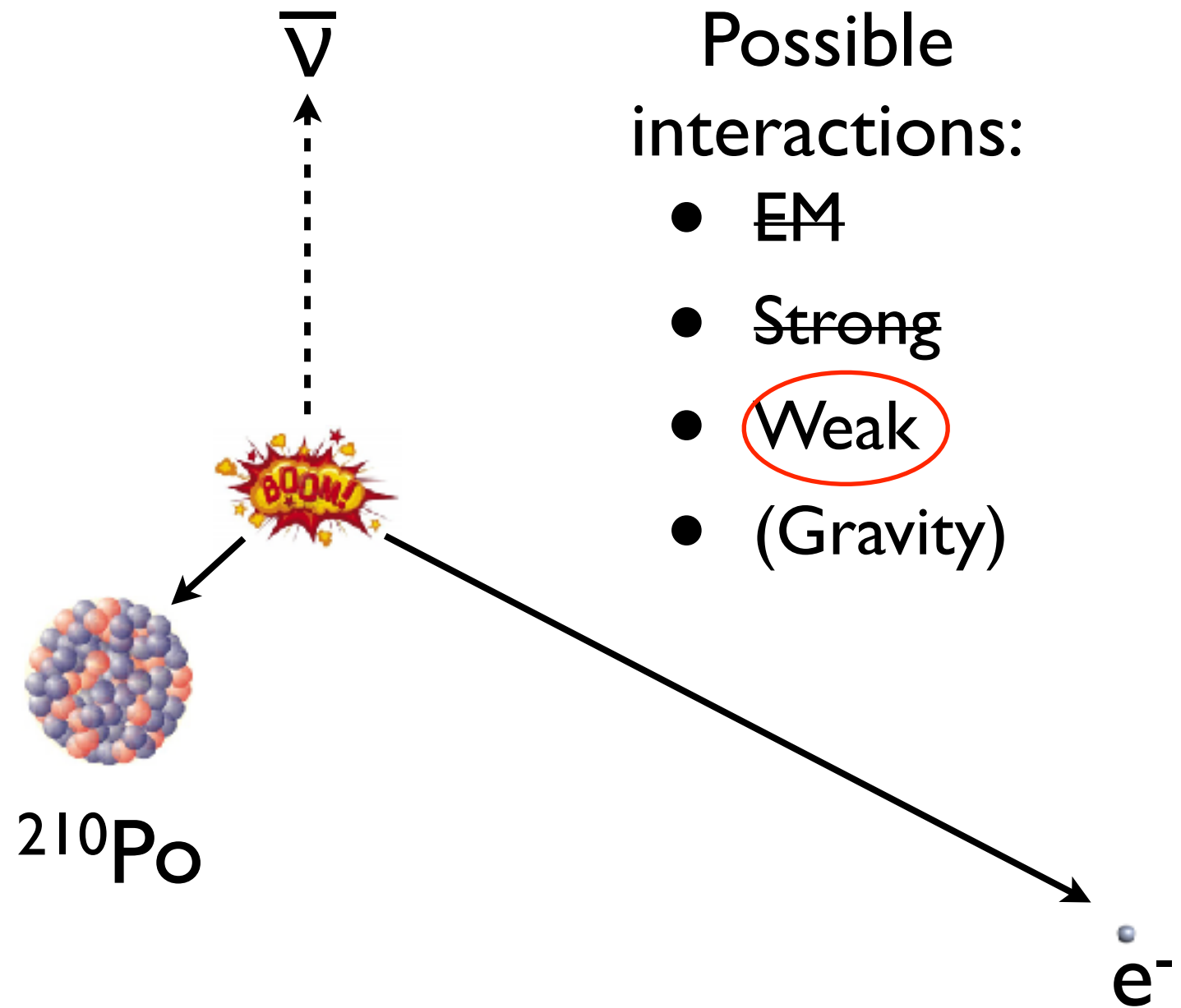
Wolfgang Pauli (1931):

- $q = 0$
- $\text{spin} = 1/2$
- $m \ll m_p$
- $\text{color} = 0$



The Neutrino

Enrico Fermi (1934):
“Little neutral one”

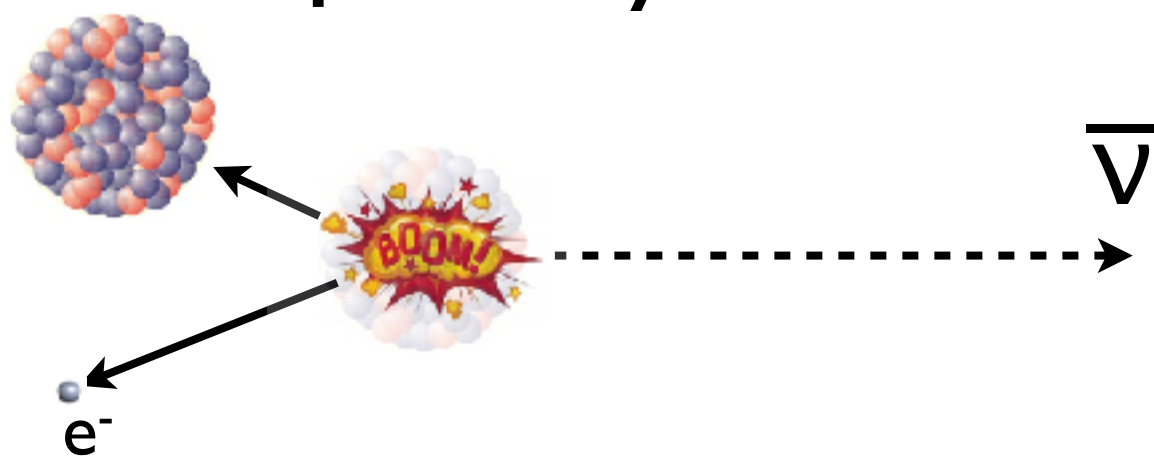


Possible interactions:

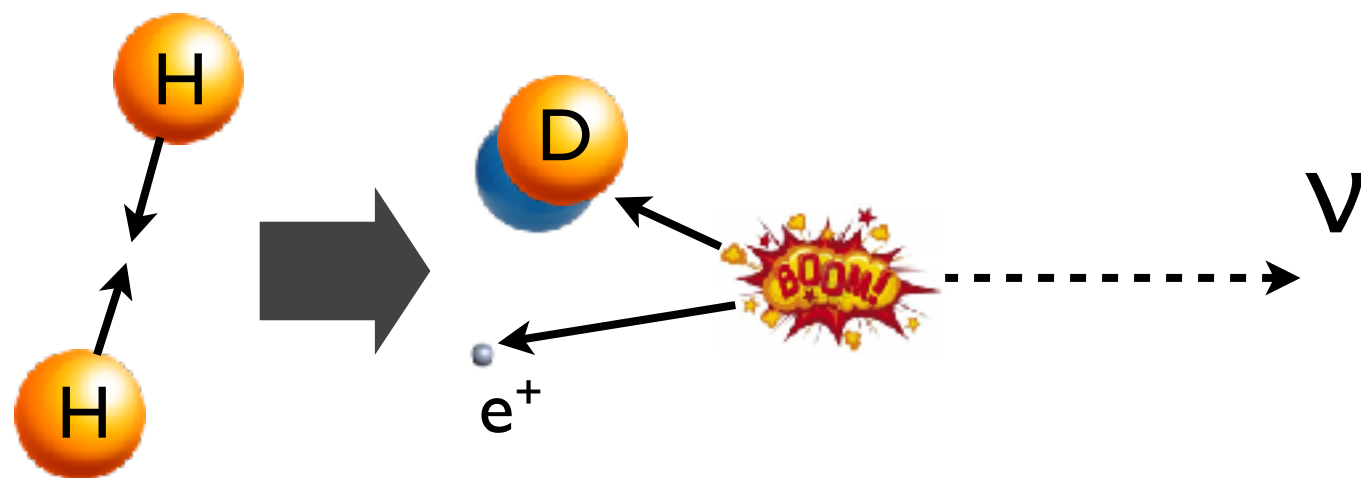
- EM
- Strong
- **Weak**
- (Gravity)

Neutrinos and Antineutrinos

β^- decay

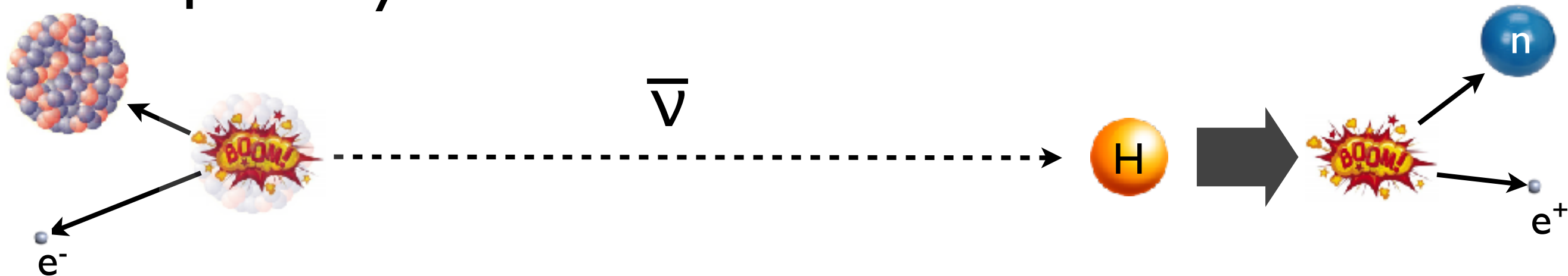


H fusion

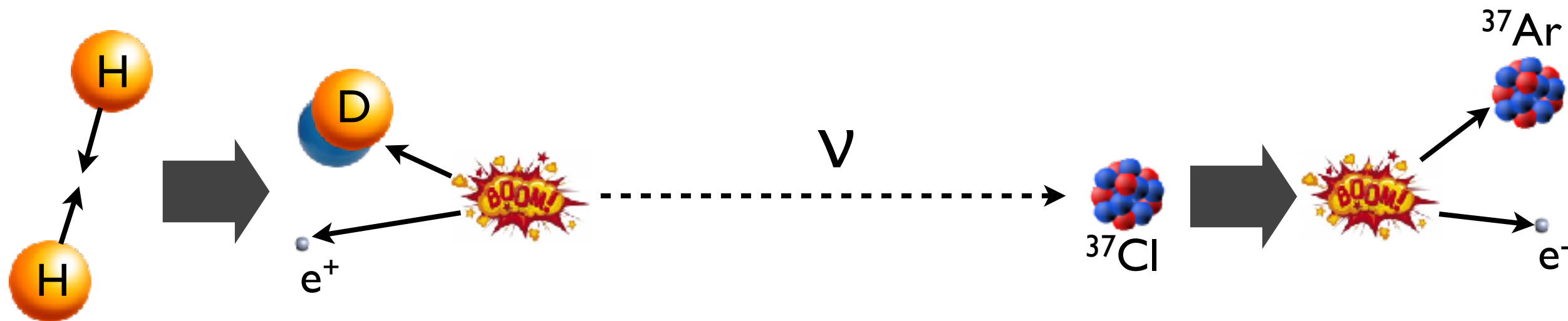


Neutrinos and Antineutrinos

β^- decay



H fusion



Neutrinos and Antineutrinos

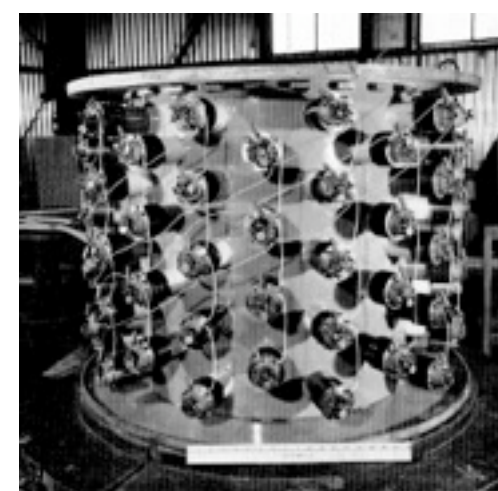
Nuclear Reactor



$\bar{\nu}$

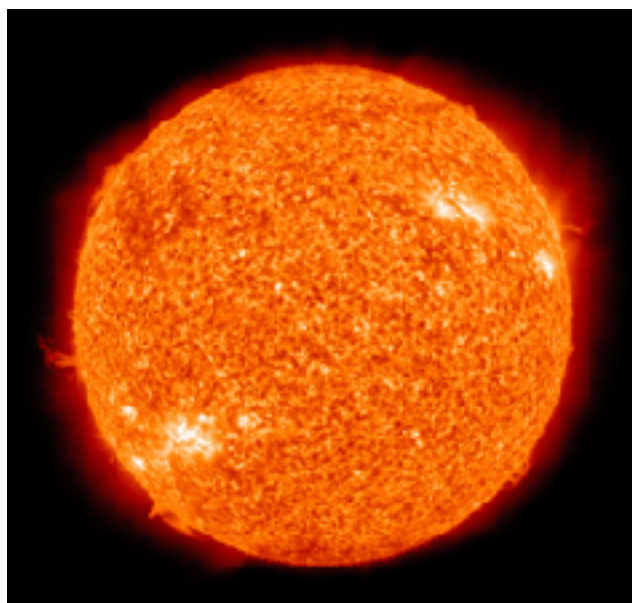


Scintillator (C_xH_y)



PMTs

The Sun



ν



Cleaning fluid (Cl)



+ Ar detector

Neutrinos and Antineutrinos

Nuclear Reactor

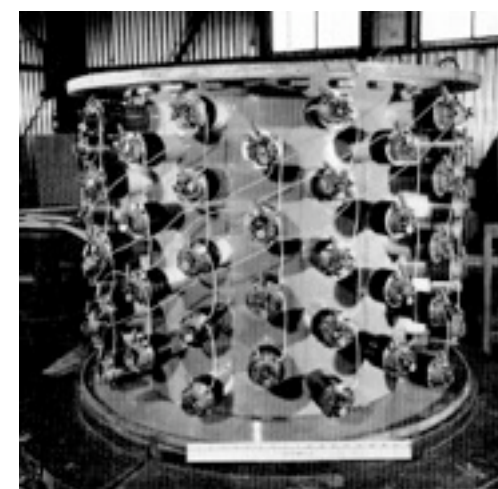


$\bar{\nu}$



Cowan and Reines (1956)

Scintillator (C_xH_y)



PMTs

The Sun



ν



Ray Davis Jr. (1964)

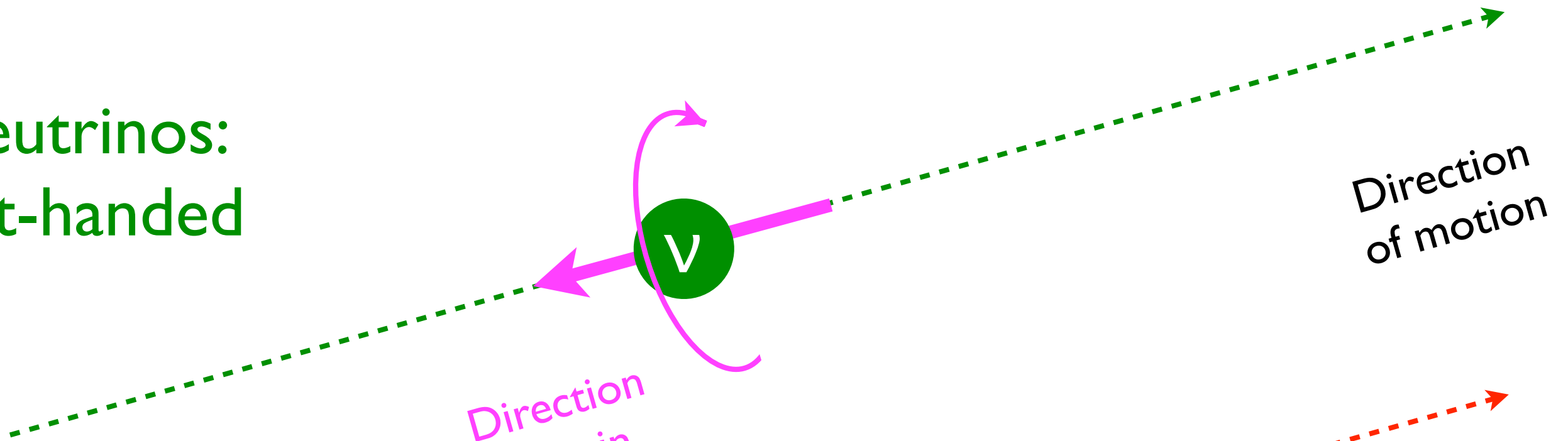
Cleaning fluid (Cl)



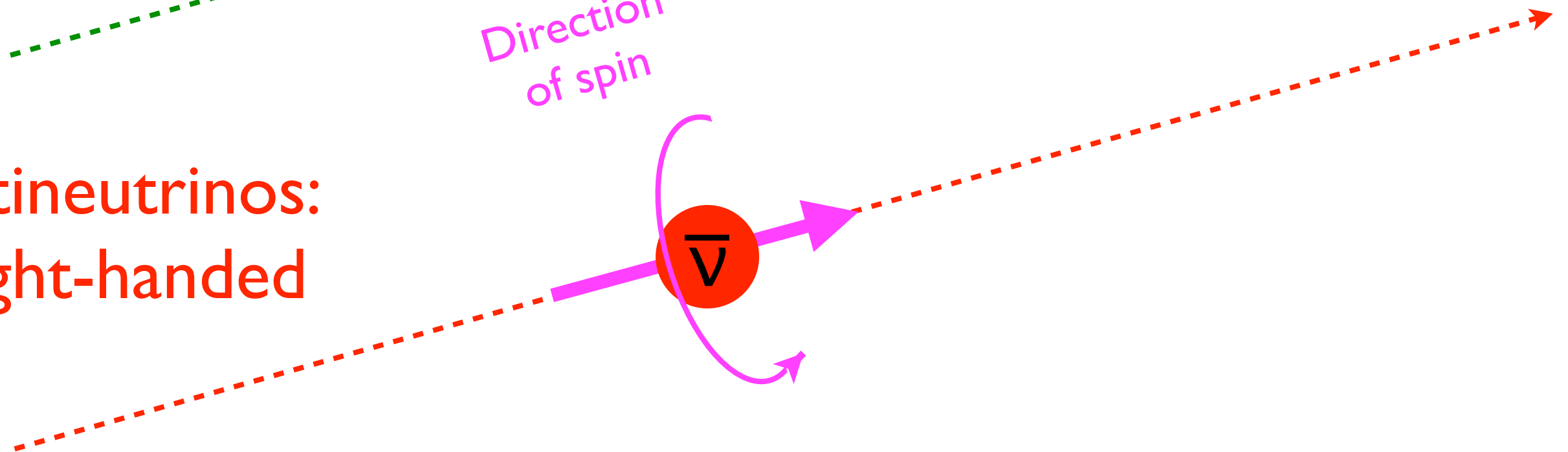
+ Ar detector

Neutrino Handedness

neutrinos:
left-handed



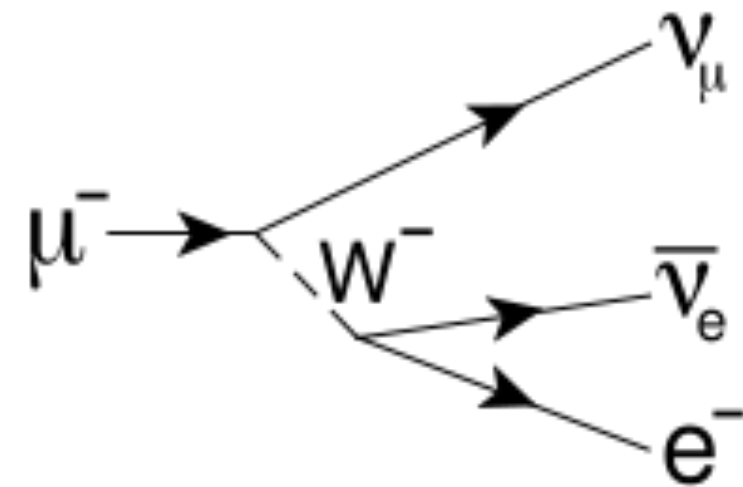
antineutrinos:
right-handed



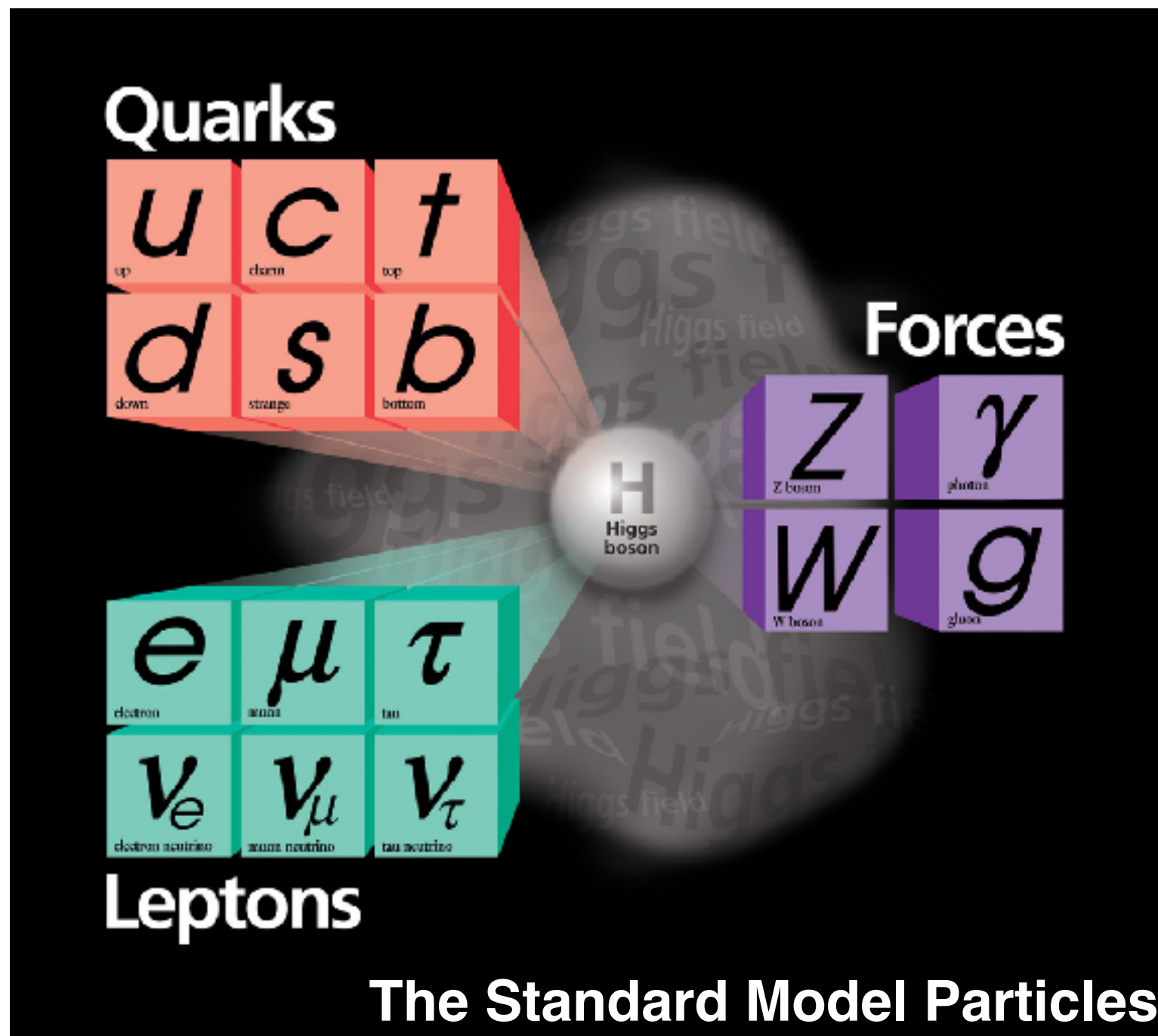
Neutrino Flavors



The Flavors of Neutrinos



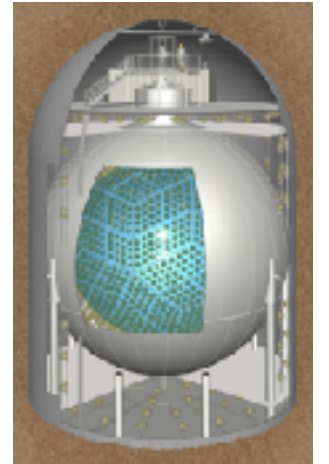
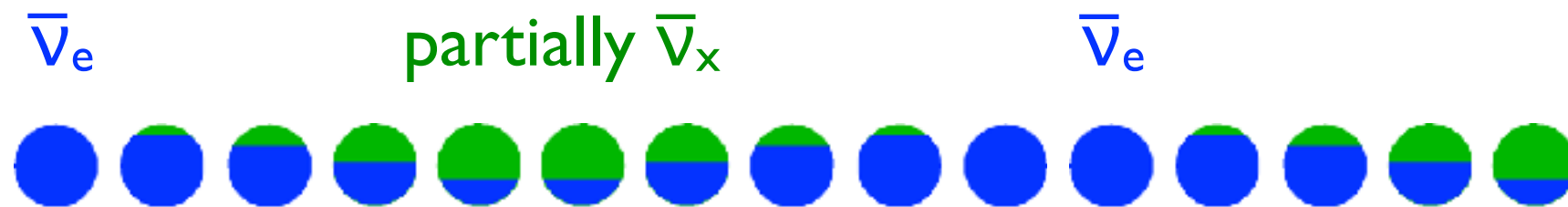
Standard Model Neutrinos



- $q = 0$
- color = 0
- spin = $1/2$
- 3 flavors (e, μ , τ)
- left-handed ν , right-handed $\bar{\nu}$
- $m_\nu < 2 \text{ eV}$
($m_e / 250000$)

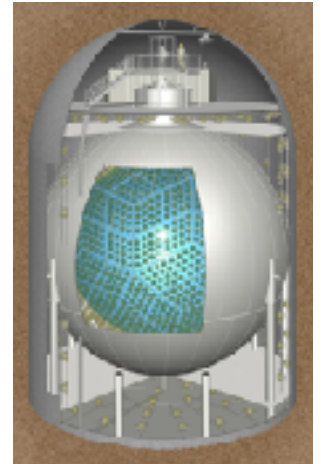
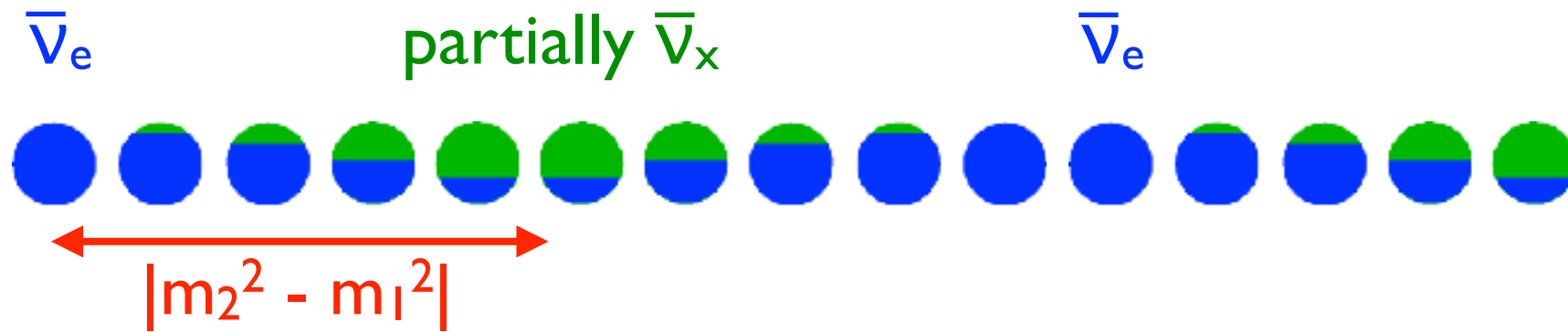
$m_\nu = 0?$

Neutrino Oscillation



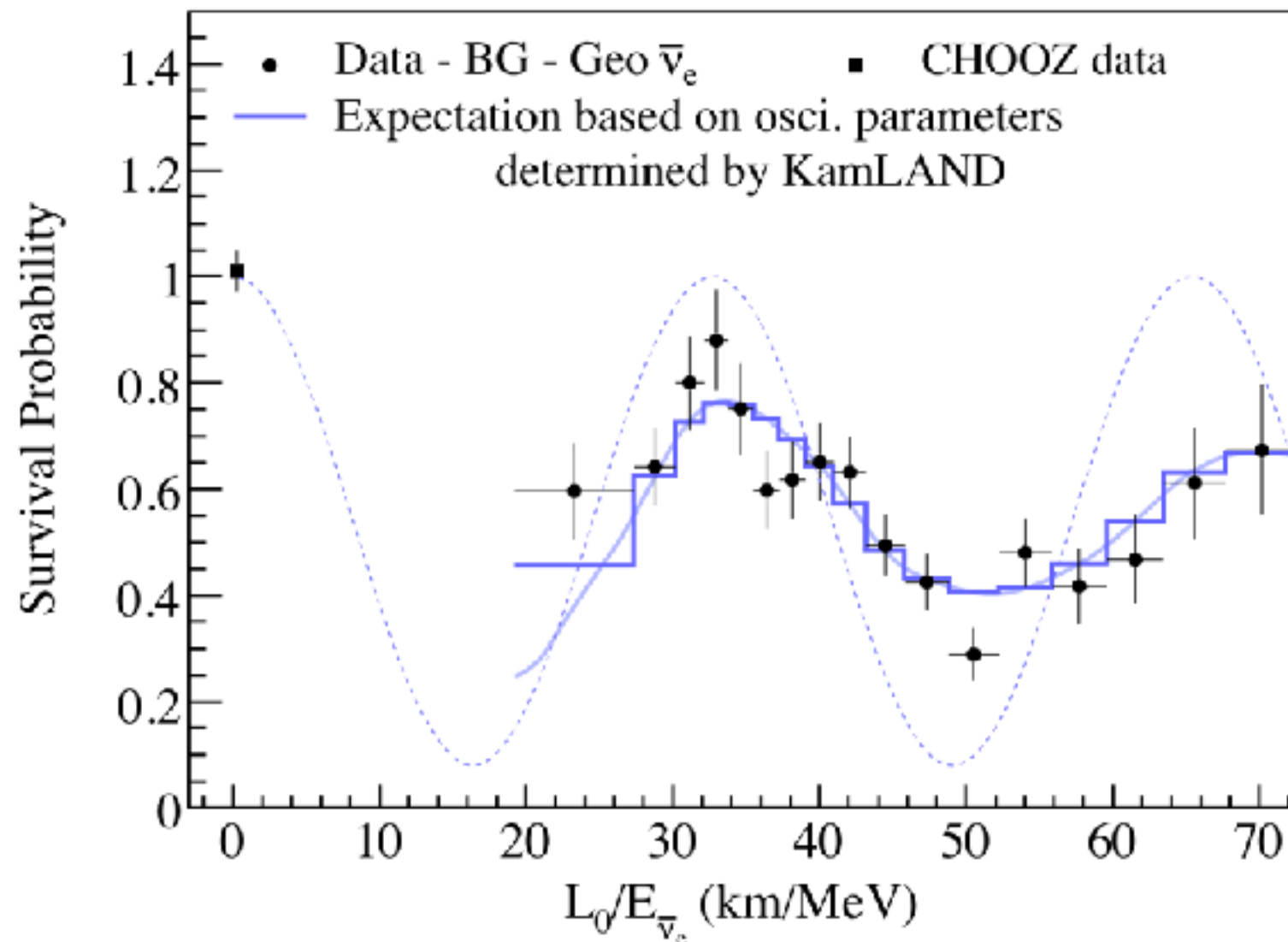
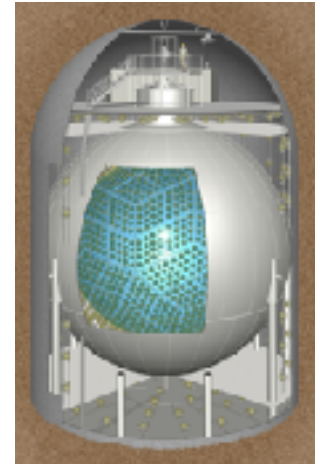
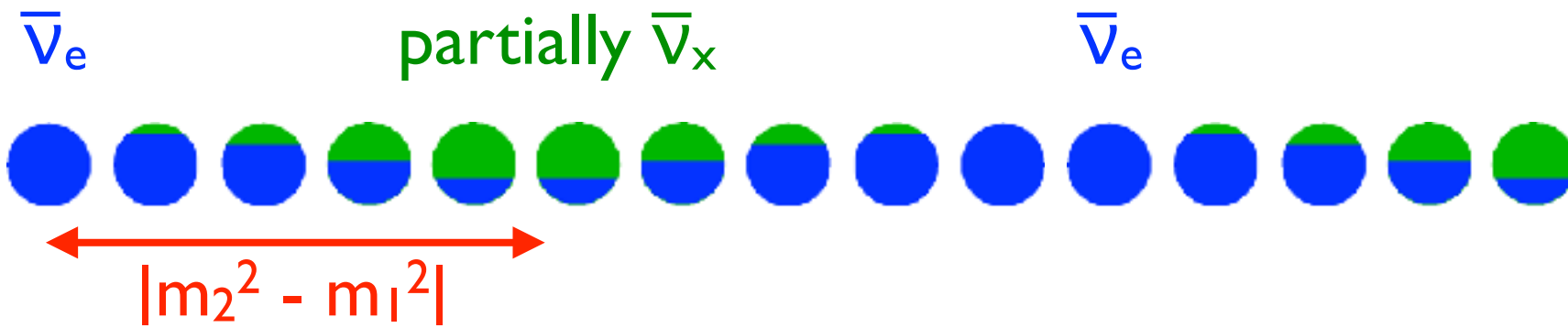
Neutrino Oscillation

$$\bar{\nu}_e = U_{e1}\bar{\nu}_1 + U_{e2}\bar{\nu}_2 + U_{e3}\bar{\nu}_3$$



Neutrino Oscillation

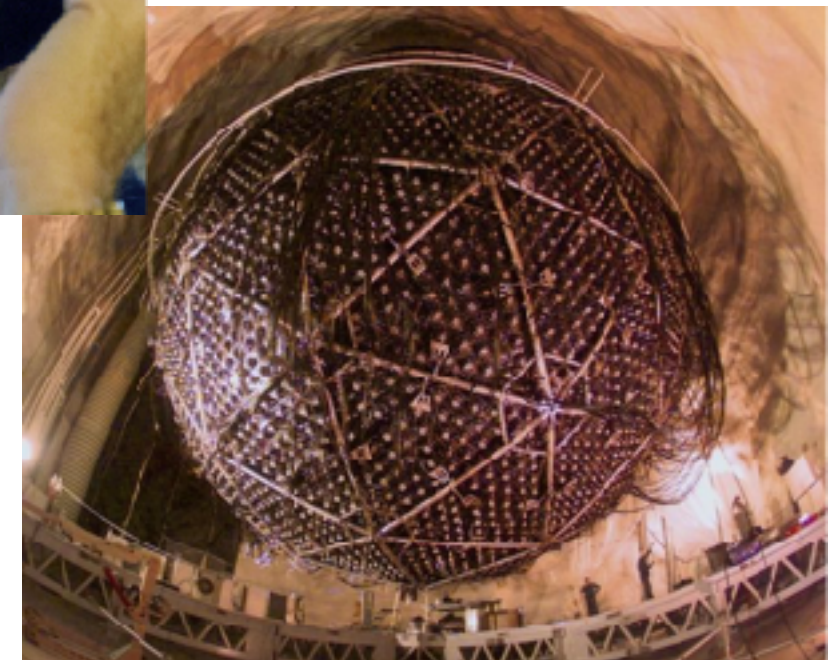
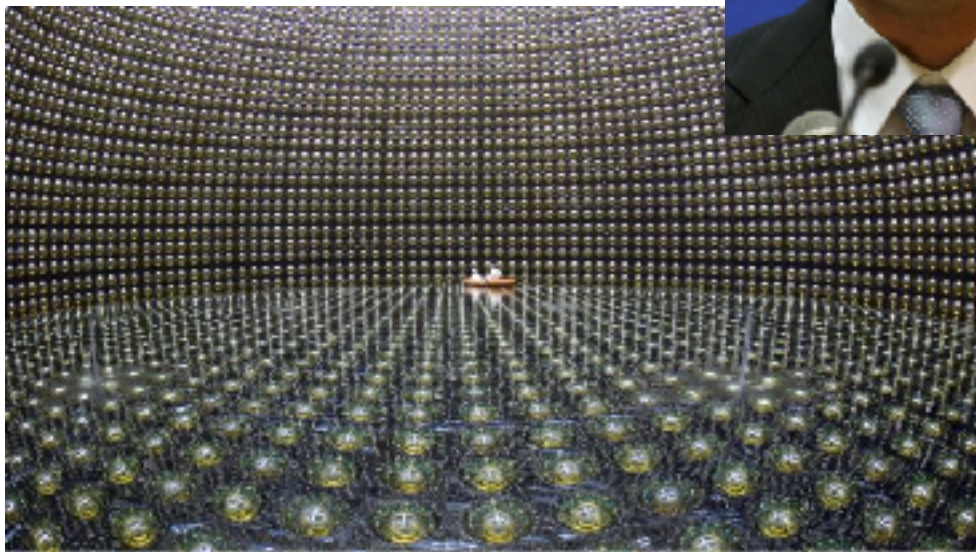
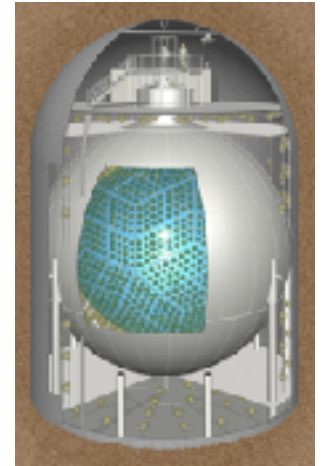
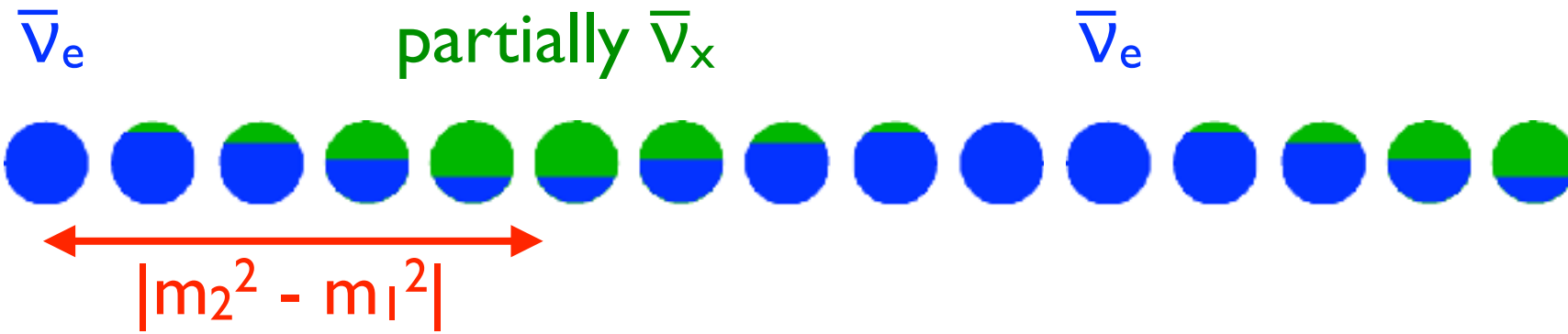
$$\bar{\nu}_e = U_{e1}\bar{\nu}_1 + U_{e2}\bar{\nu}_2 + U_{e3}\bar{\nu}_3$$



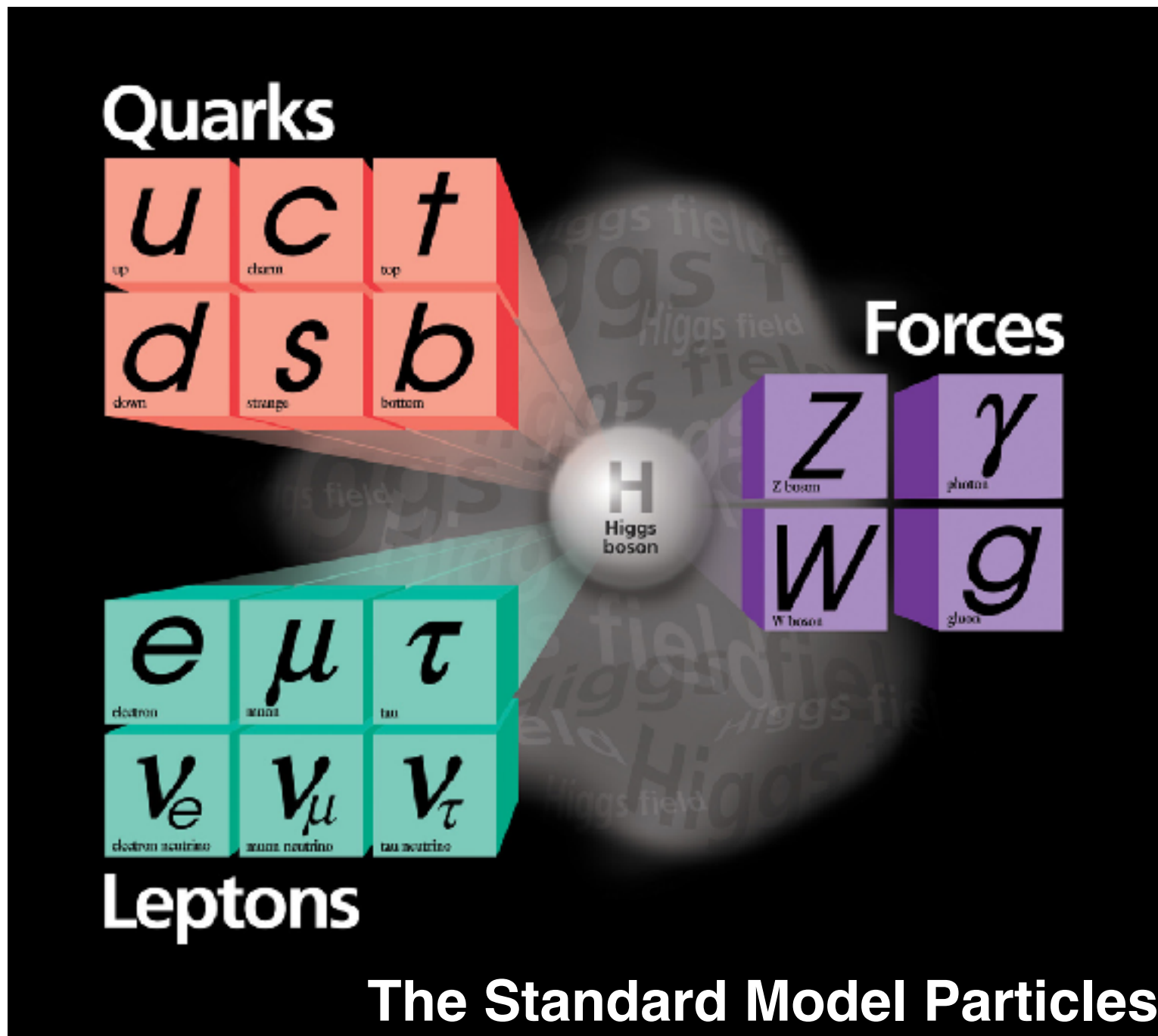


Neutrino Oscillation

$$\bar{\nu}_e = U_{e1}\bar{\nu}_1 + U_{e2}\bar{\nu}_2 + U_{e3}\bar{\nu}_3$$

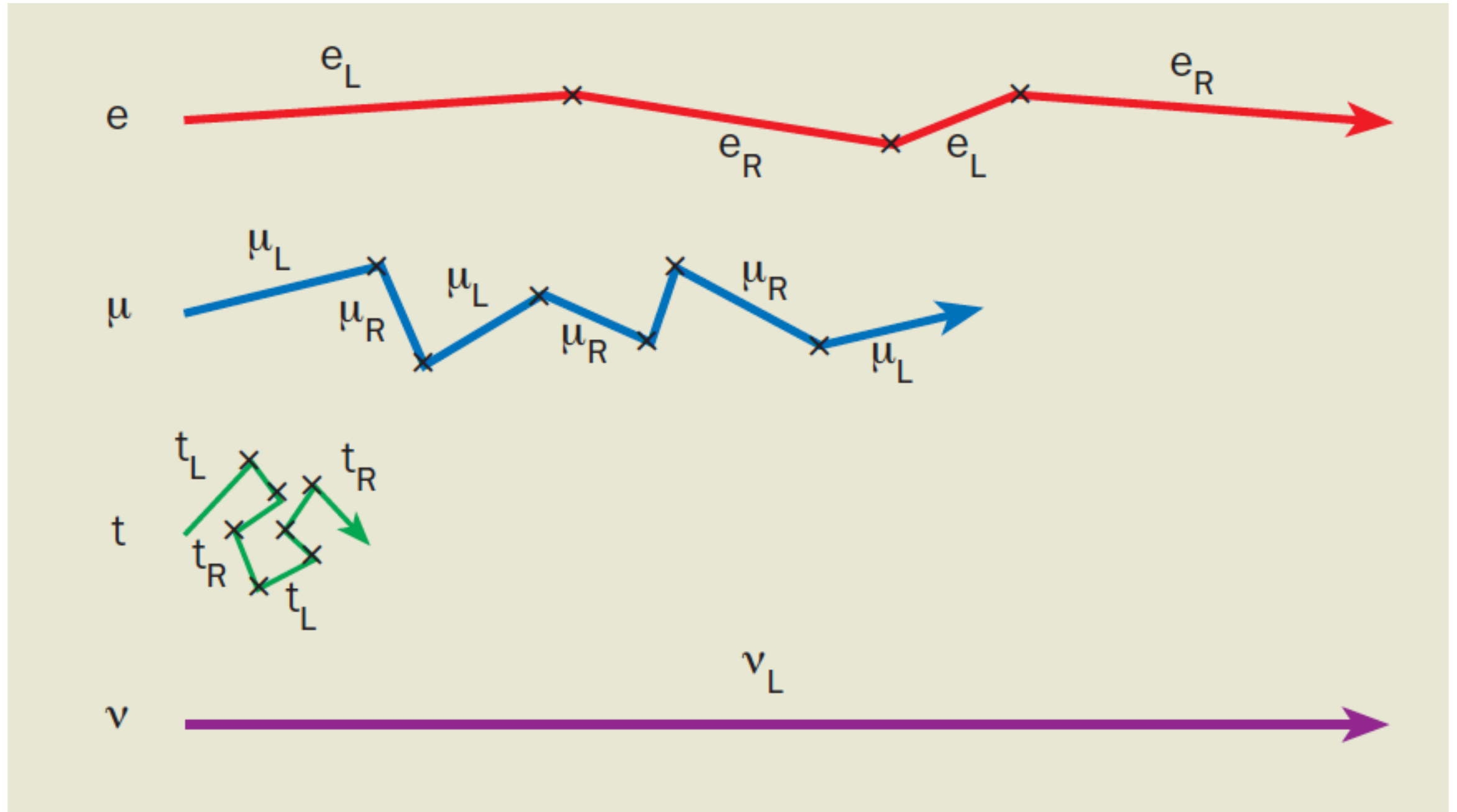


Standard Model Neutrinos

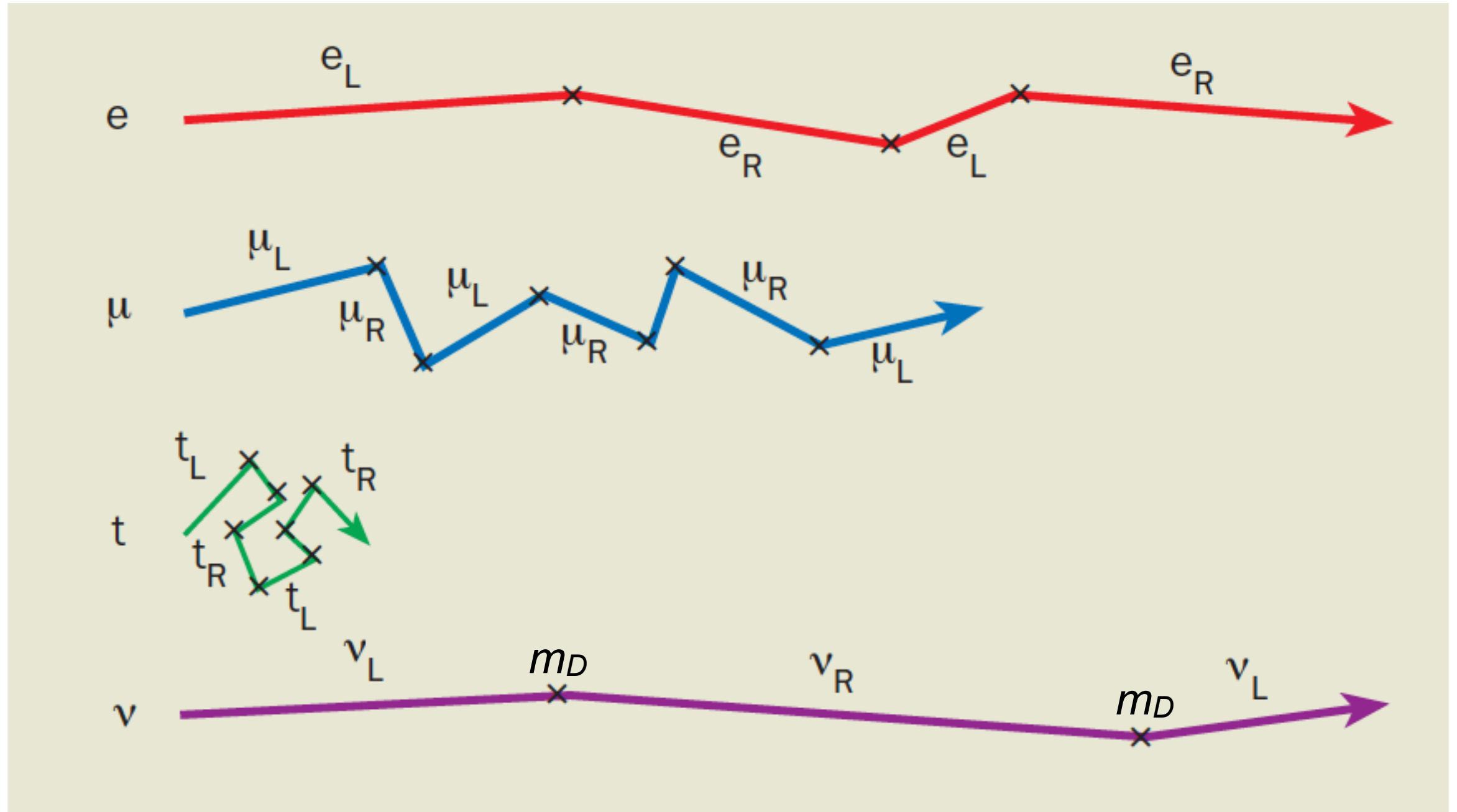


- $q = 0$
- $\text{color} = 0$
- $\text{spin} = 1/2$
- 3 flavors (e, μ, τ)
- left-handed ν , right-handed $\bar{\nu}$
- $0.02 \text{ eV} < m_\nu < 2 \text{ eV}$
- How light are they?
- Where is the right-handed ν ?

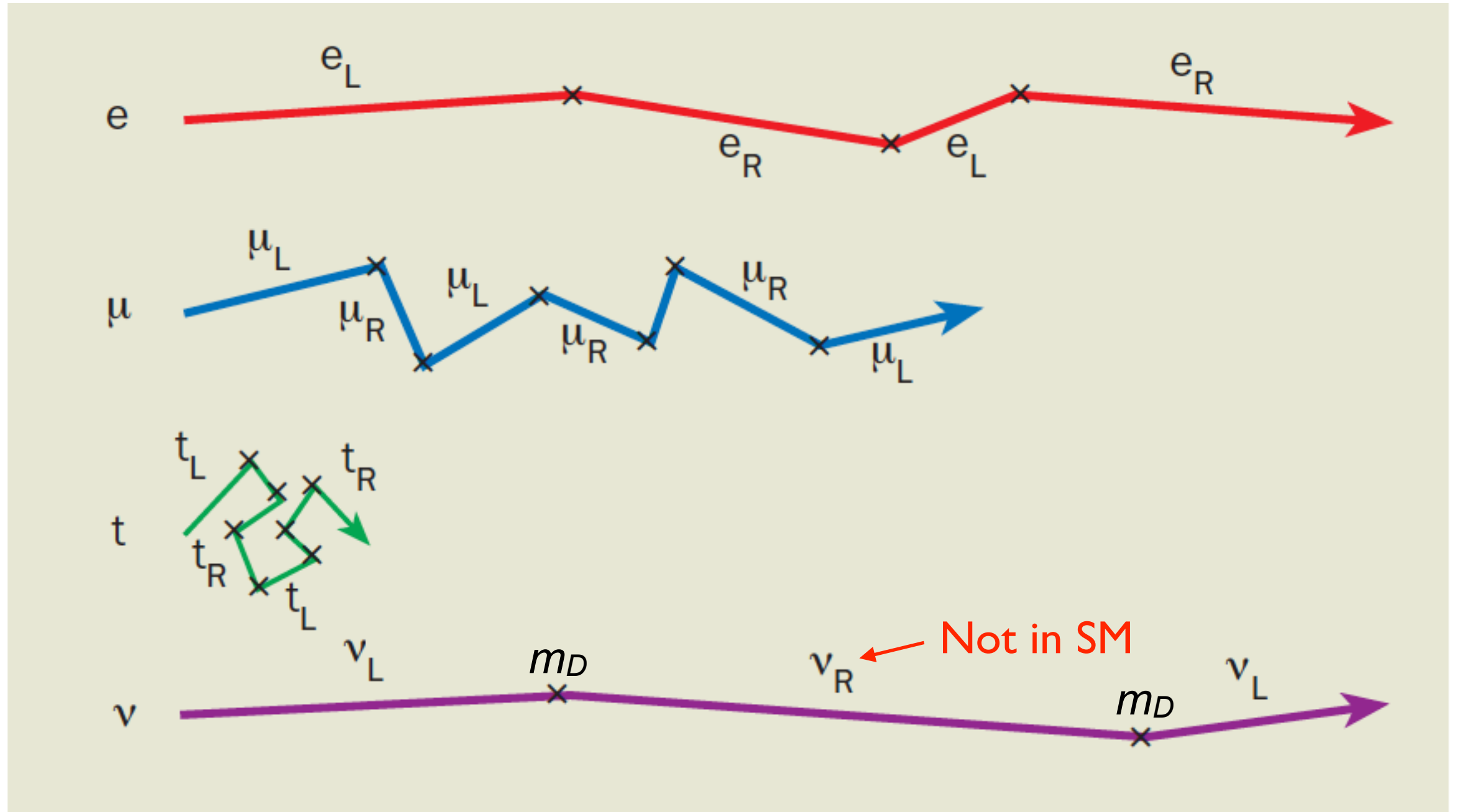
Incorporating ν Mass



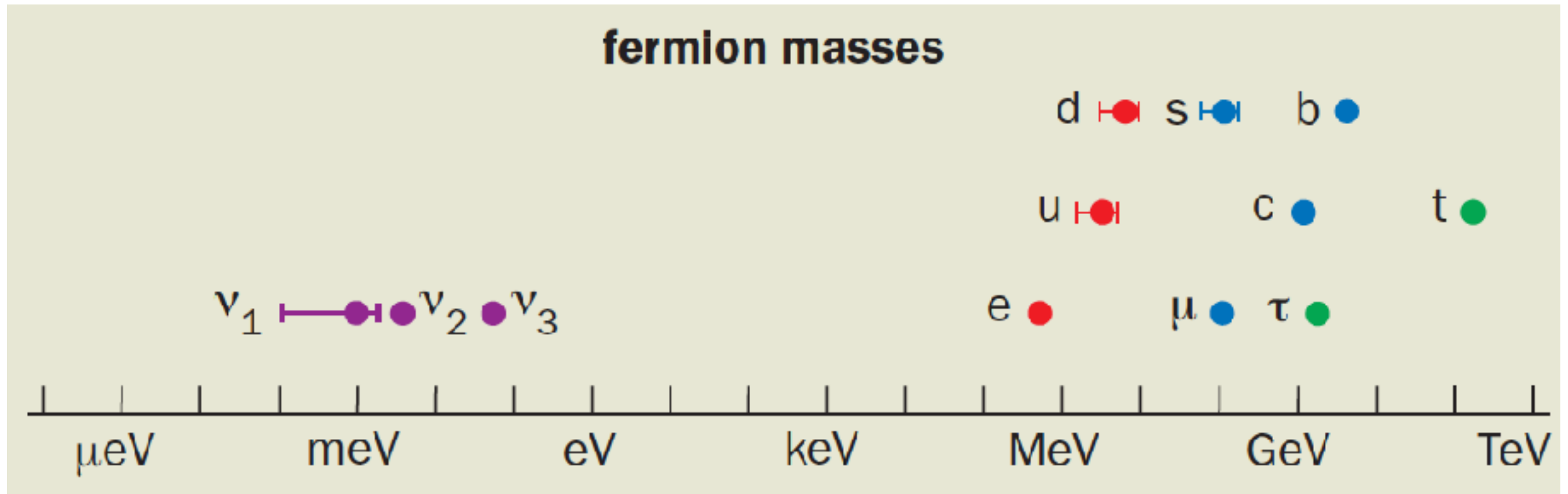
Incorporating ν Mass



Incorporating ν Mass



Incorporating ν Mass



The Majorana Equation



Schrodinger:
$$i \frac{\partial}{\partial t} \Psi + \frac{1}{2m} \nabla^2 \Psi = 0$$

The Majorana Equation



Schrodinger:
$$i \frac{\partial}{\partial t} \Psi + \frac{1}{2m} \nabla^2 \Psi = 0$$



Dirac:
$$-i \gamma^\mu \partial_\mu \psi + m \psi = 0$$

The Majorana Equation



Schrodinger:
$$i \frac{\partial}{\partial t} \Psi + \frac{1}{2m} \nabla^2 \Psi = 0$$

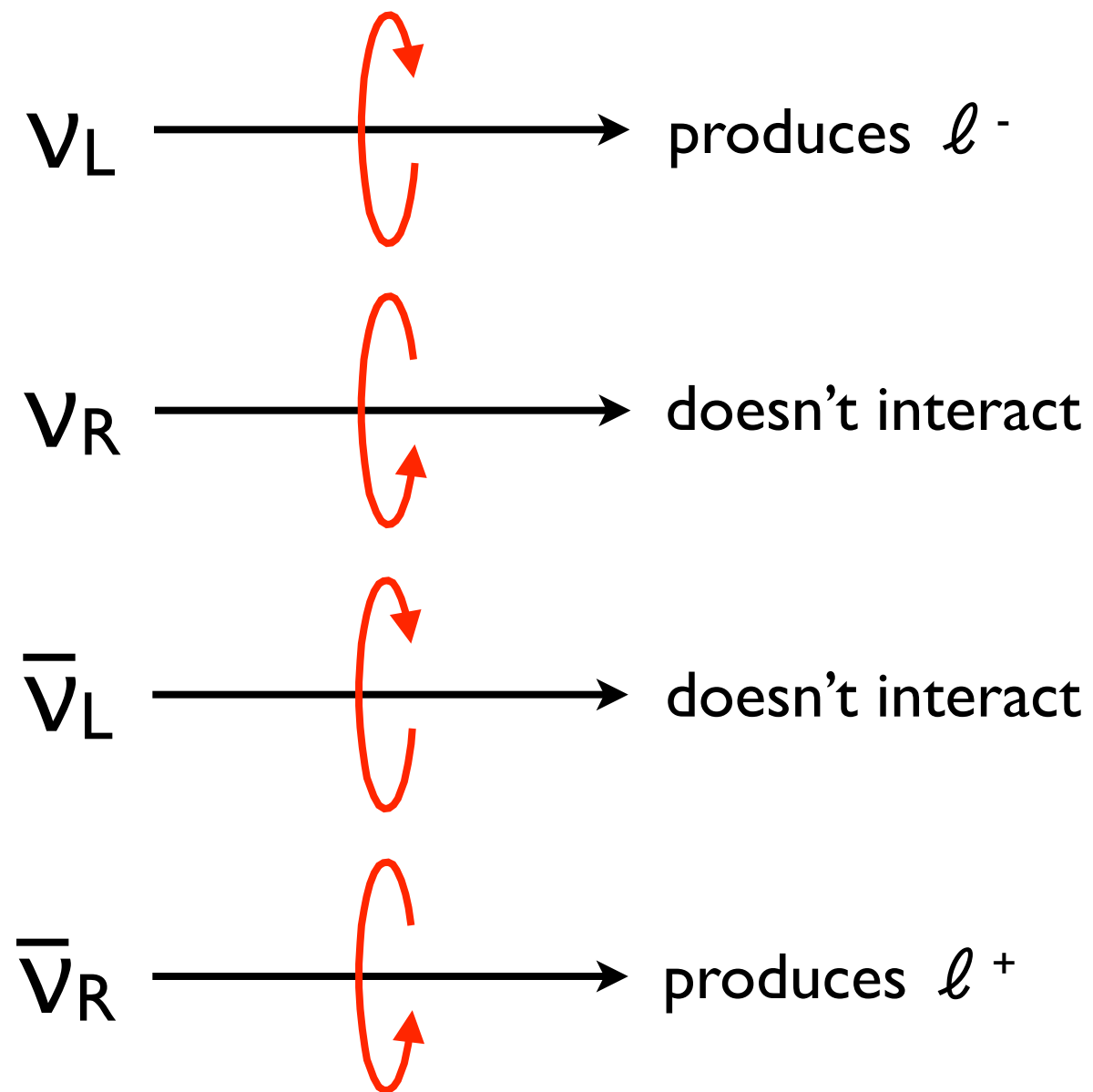


Dirac:
$$-i \gamma^\mu \partial_\mu \psi + m \psi = 0$$



Majorana:
$$\sigma_\pm^\mu \partial_\mu \chi \pm m \sigma_2 \chi^* = 0$$

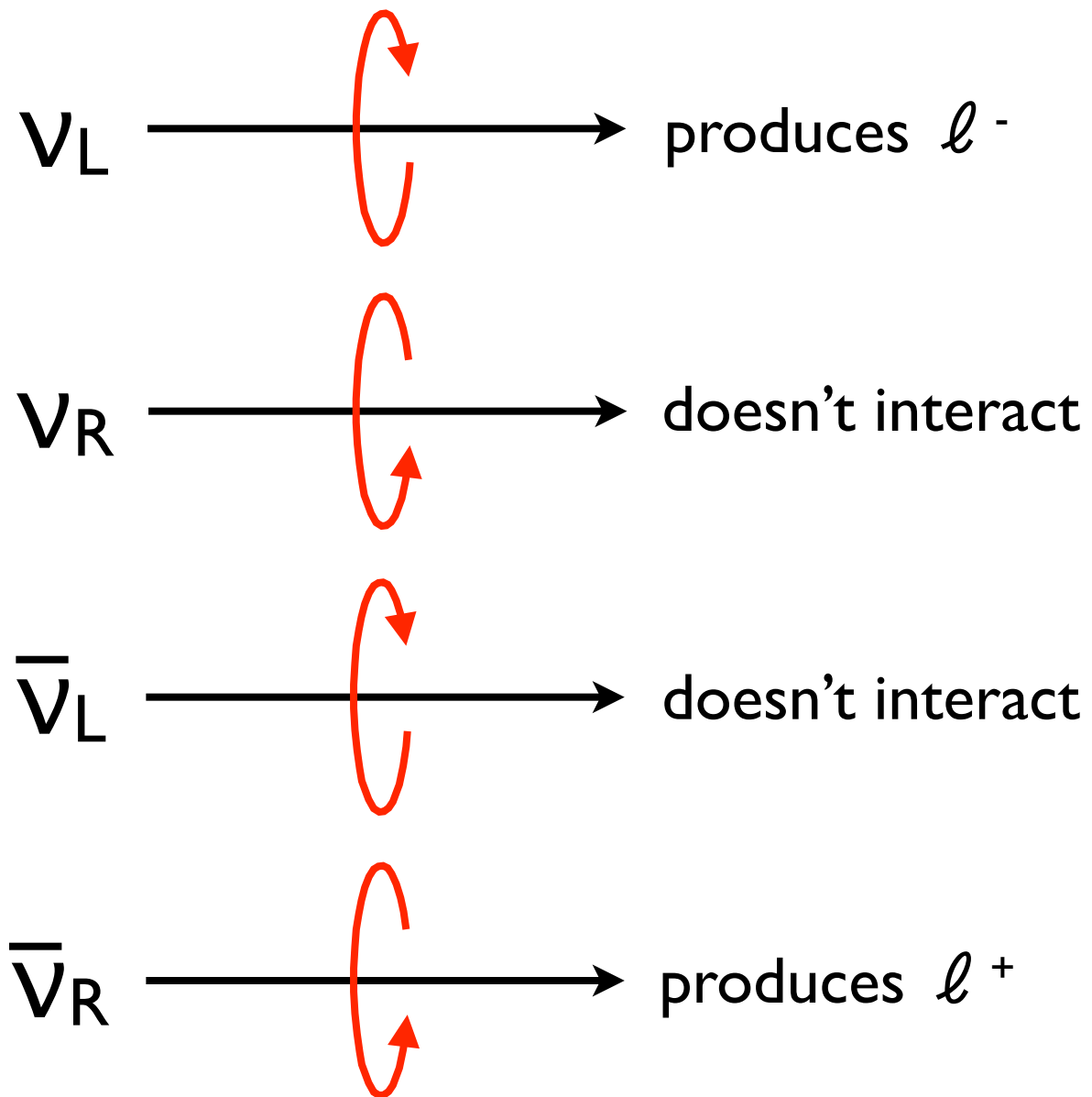
Dirac neutrinos



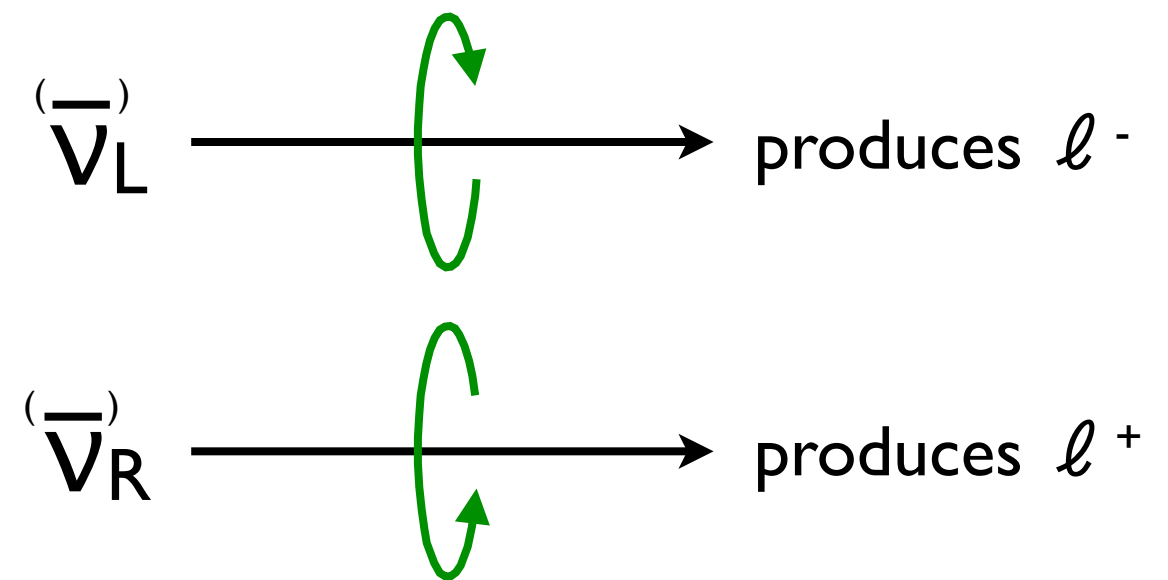
Majorana: $\nu = \bar{\nu}$



Dirac ($\nu \neq \bar{\nu}$)

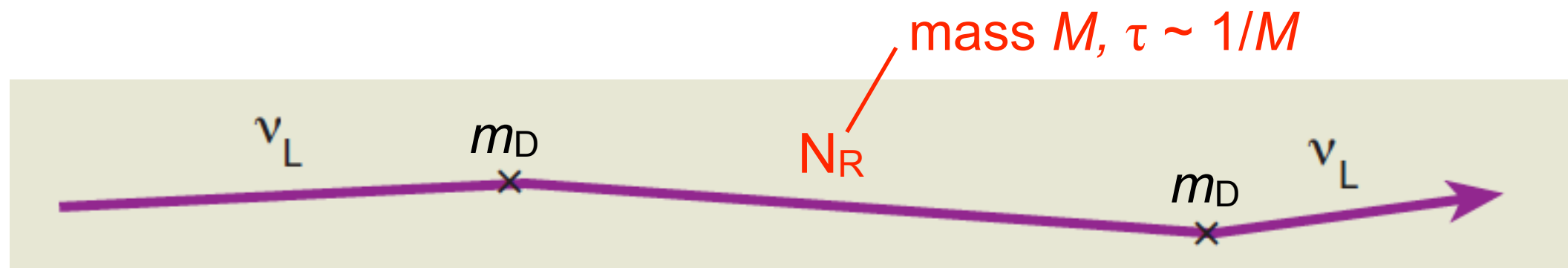


Majorana ($\nu = \bar{\nu}$)

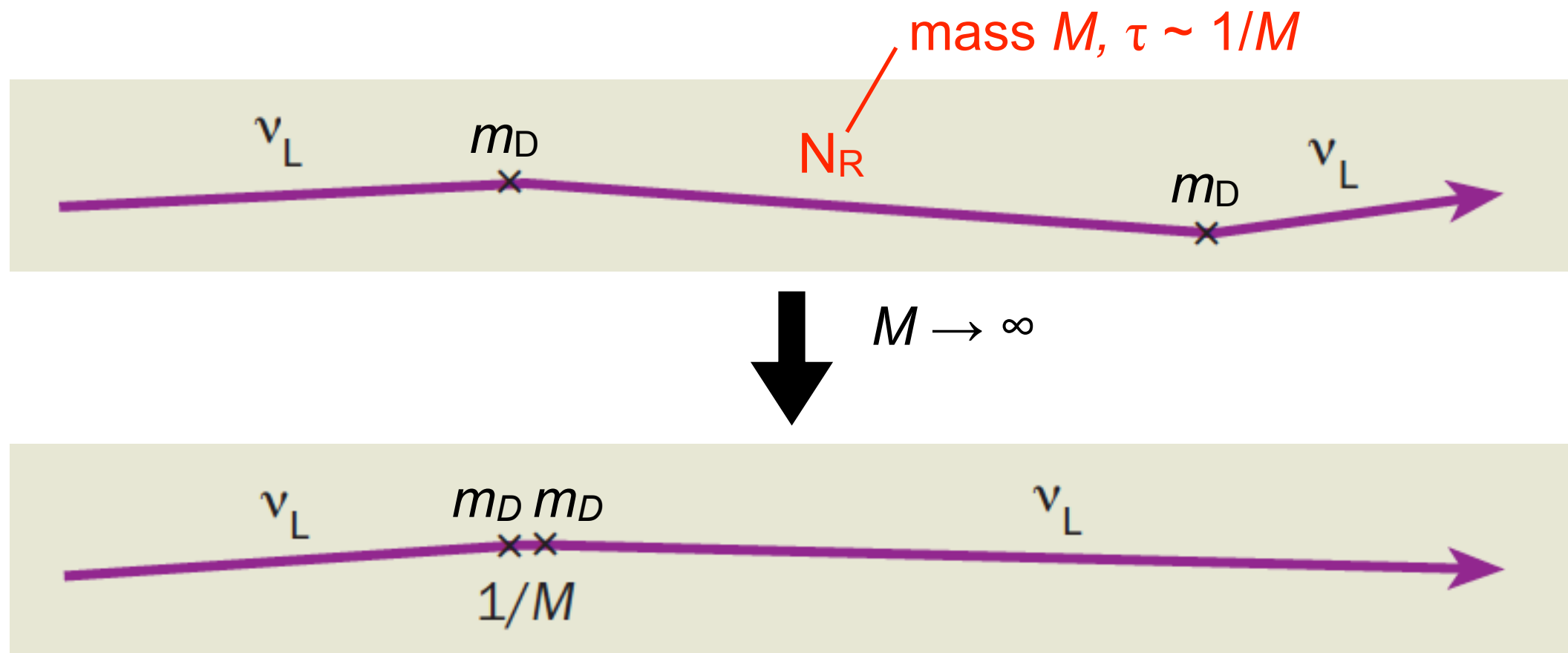


- Requires $Q = -Q = 0$
- Appear “naturally” in Grand Unification theories, explains why ν are so much lighter than e, μ, τ
- Implies Lepton number is not conserved

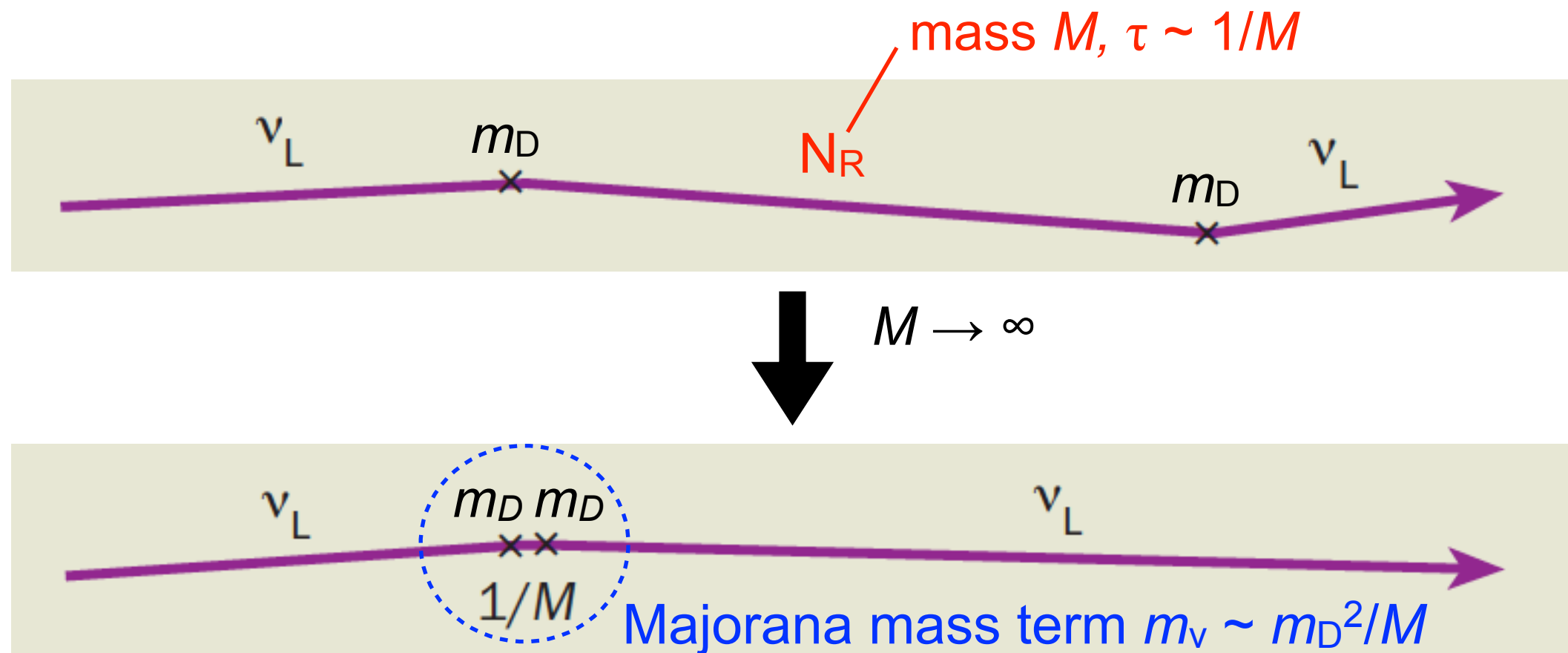
Seesaw Mechanism



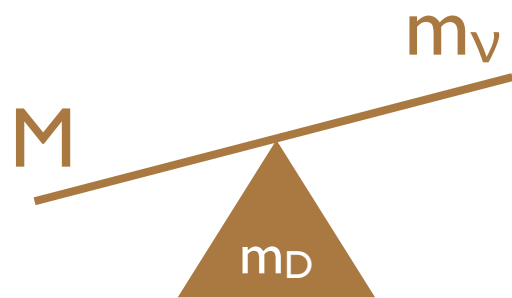
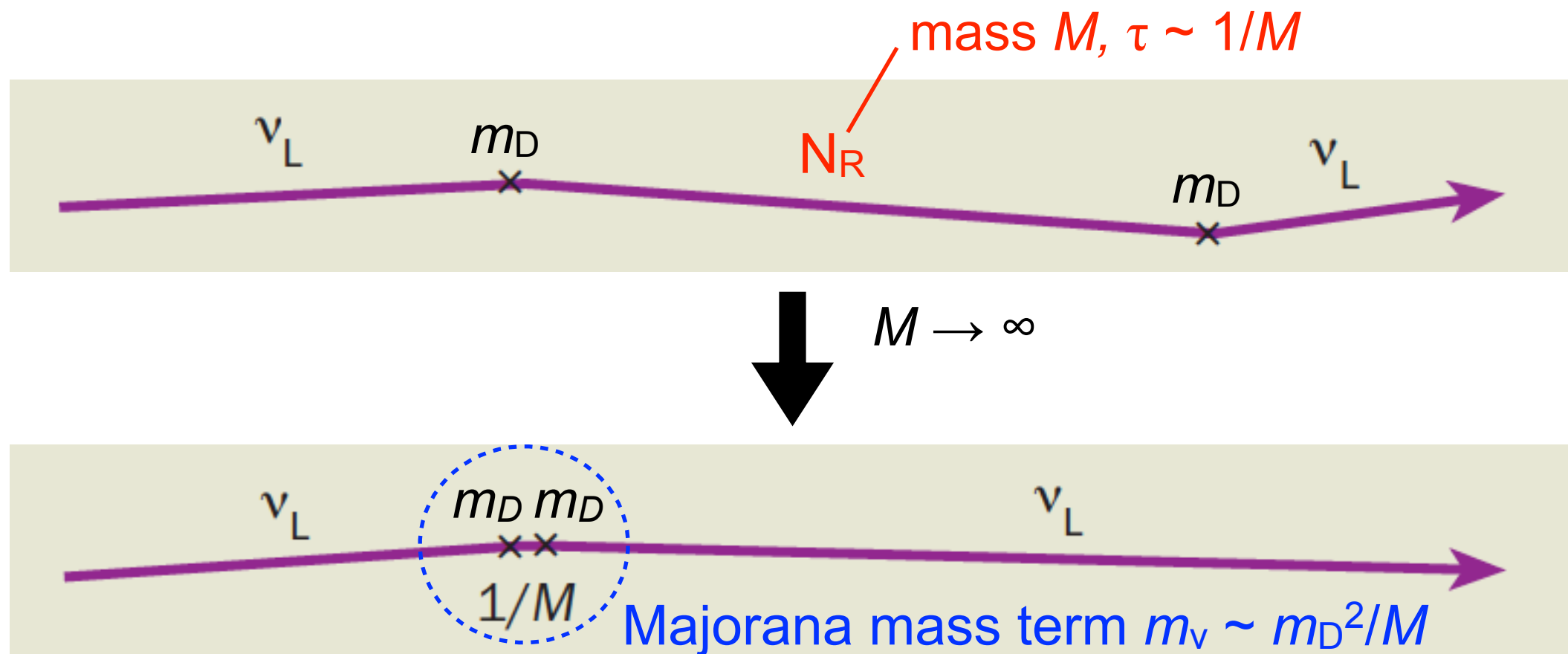
Seesaw Mechanism



Seesaw Mechanism

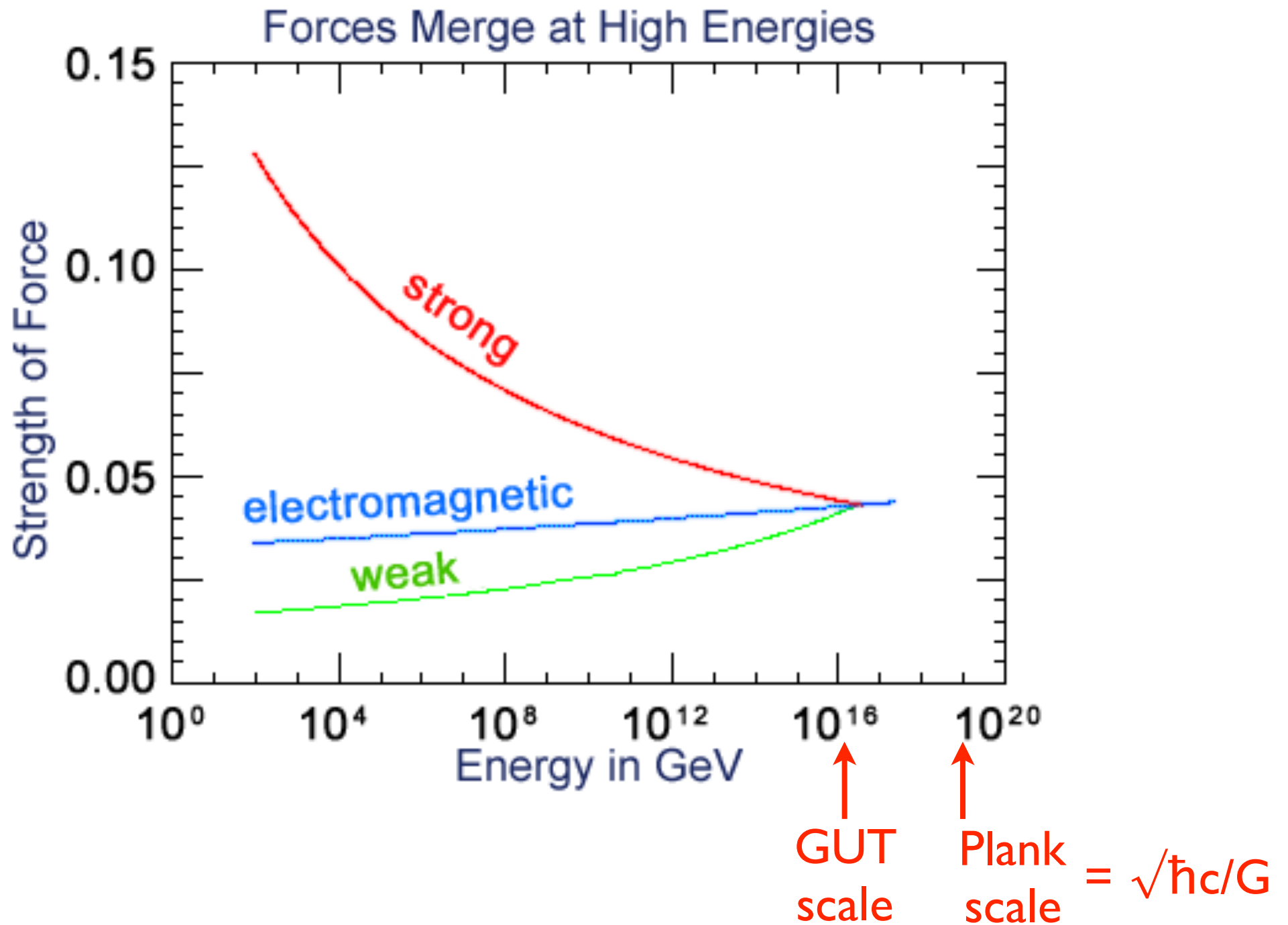


Seesaw Mechanism



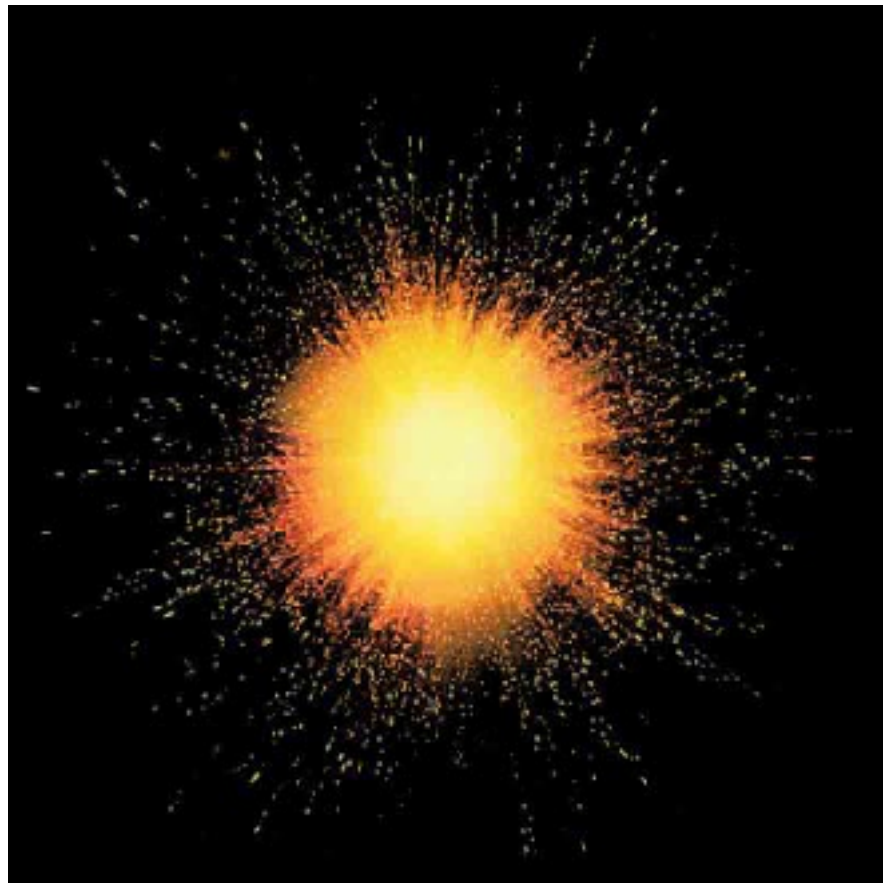
for $m_D \sim \text{GeV-TeV}$:
 $m_\nu \sim \text{meV-eV} \leftrightarrow M \sim 10^{16}-10^{19} \text{ GeV}$

Grand Unification



Matter-Antimatter Asymmetry

The Big Bang



matter + antimatter

The Universe Today



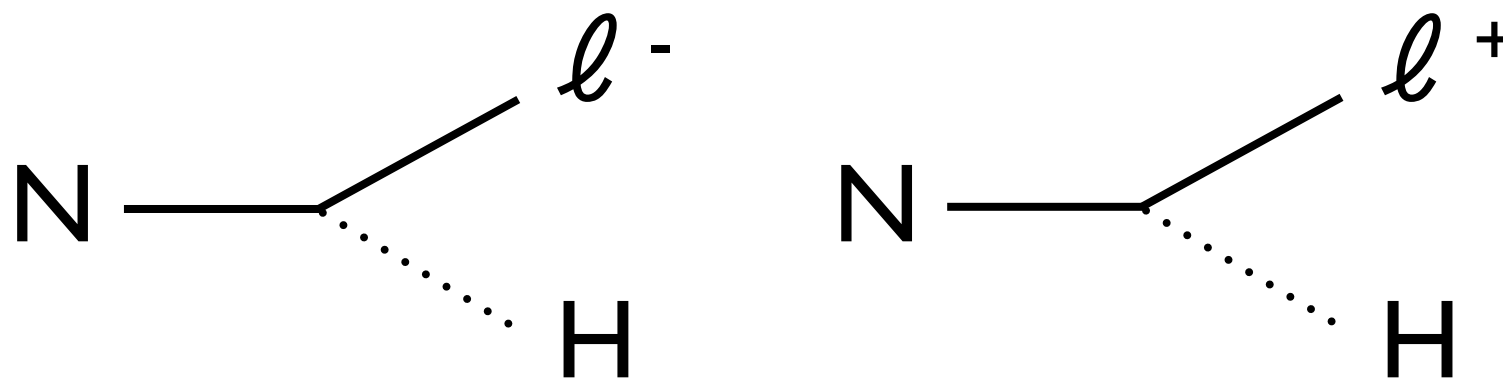
matter only

Sakharov Conditions

- Interactions out of thermal equilibrium
- Baryon number violation (baryogenesis)
- C (charge) and CP (charge-parity) violation

Leptogenesis

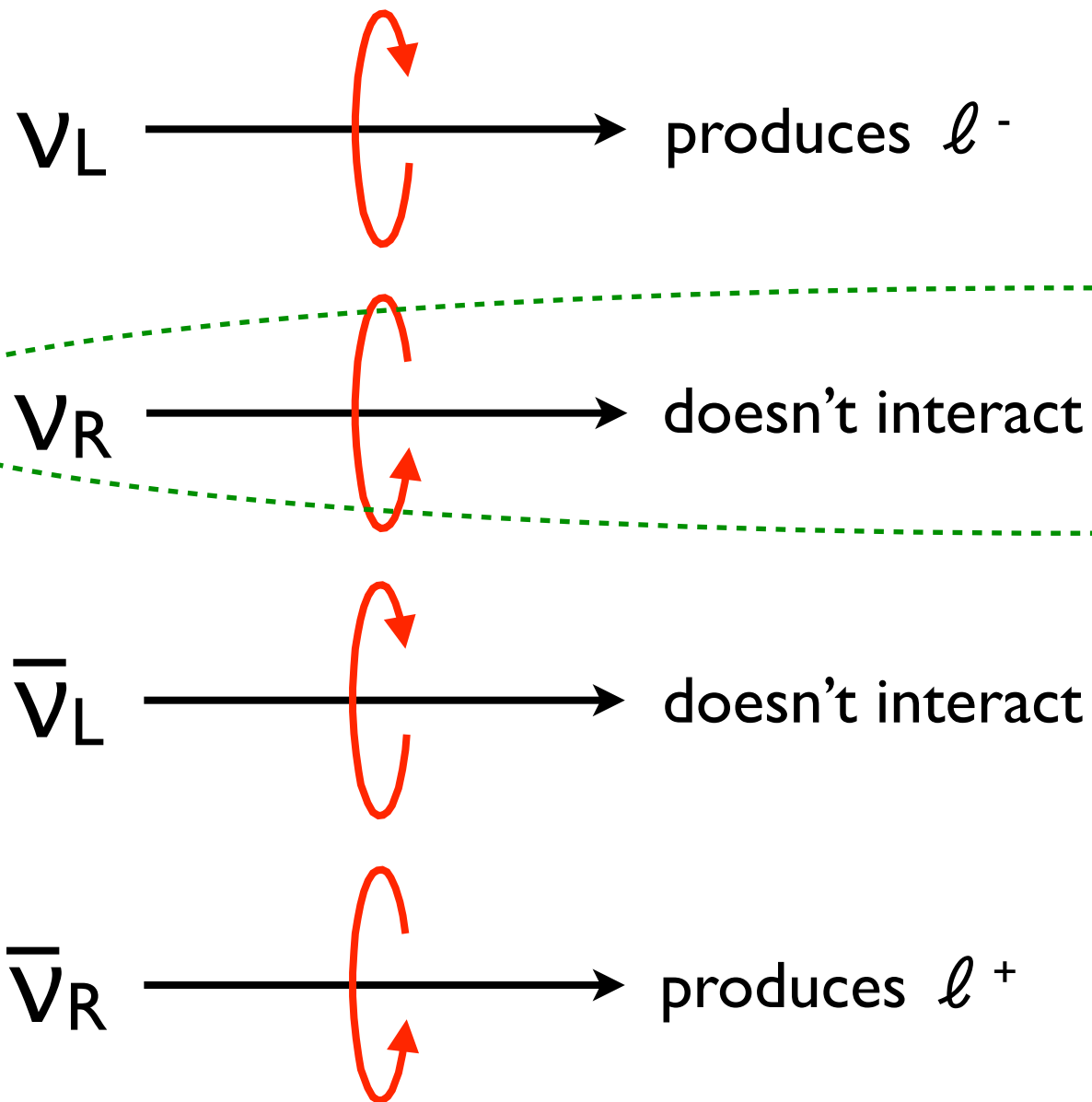
- Decay of heavy Majorana neutrino (N) into SM leptons (ℓ^\pm) and Higgs (H):



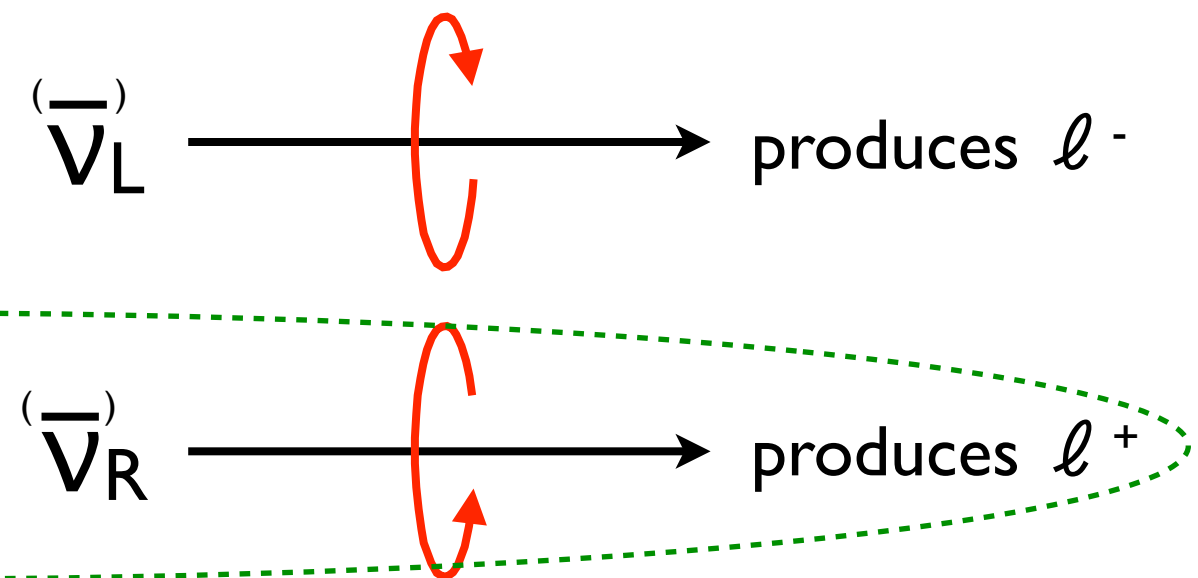
- CP violation in ν sector could give these different branching ratios
- SM processes could convert L to B: baryogenesis!
- Majorana neutrinos could be the reason we exist at all!

Testing $\nu = \bar{\nu}$

Dirac ν

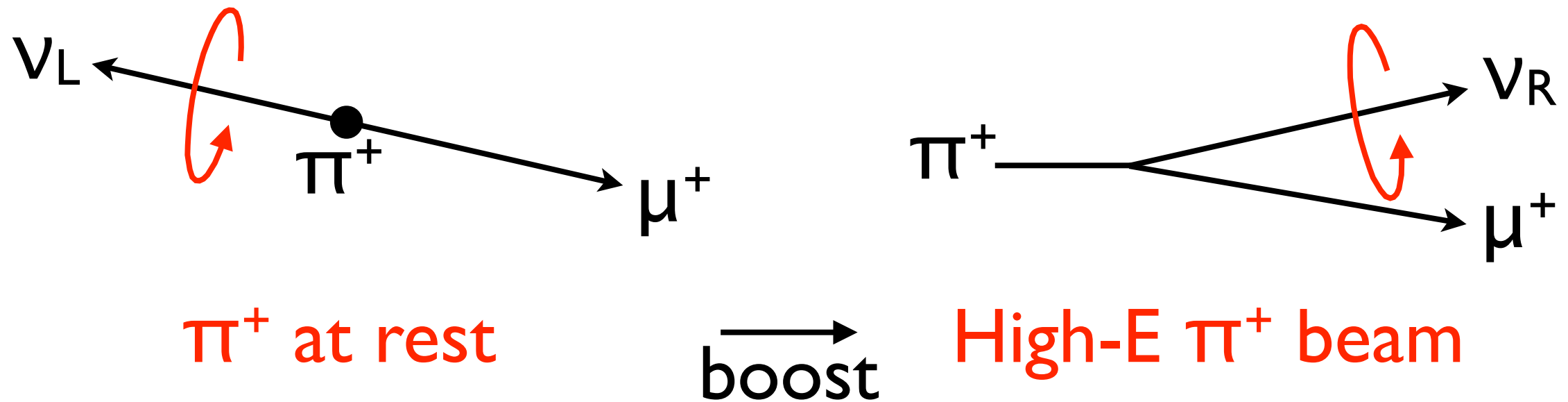


Majorana ν



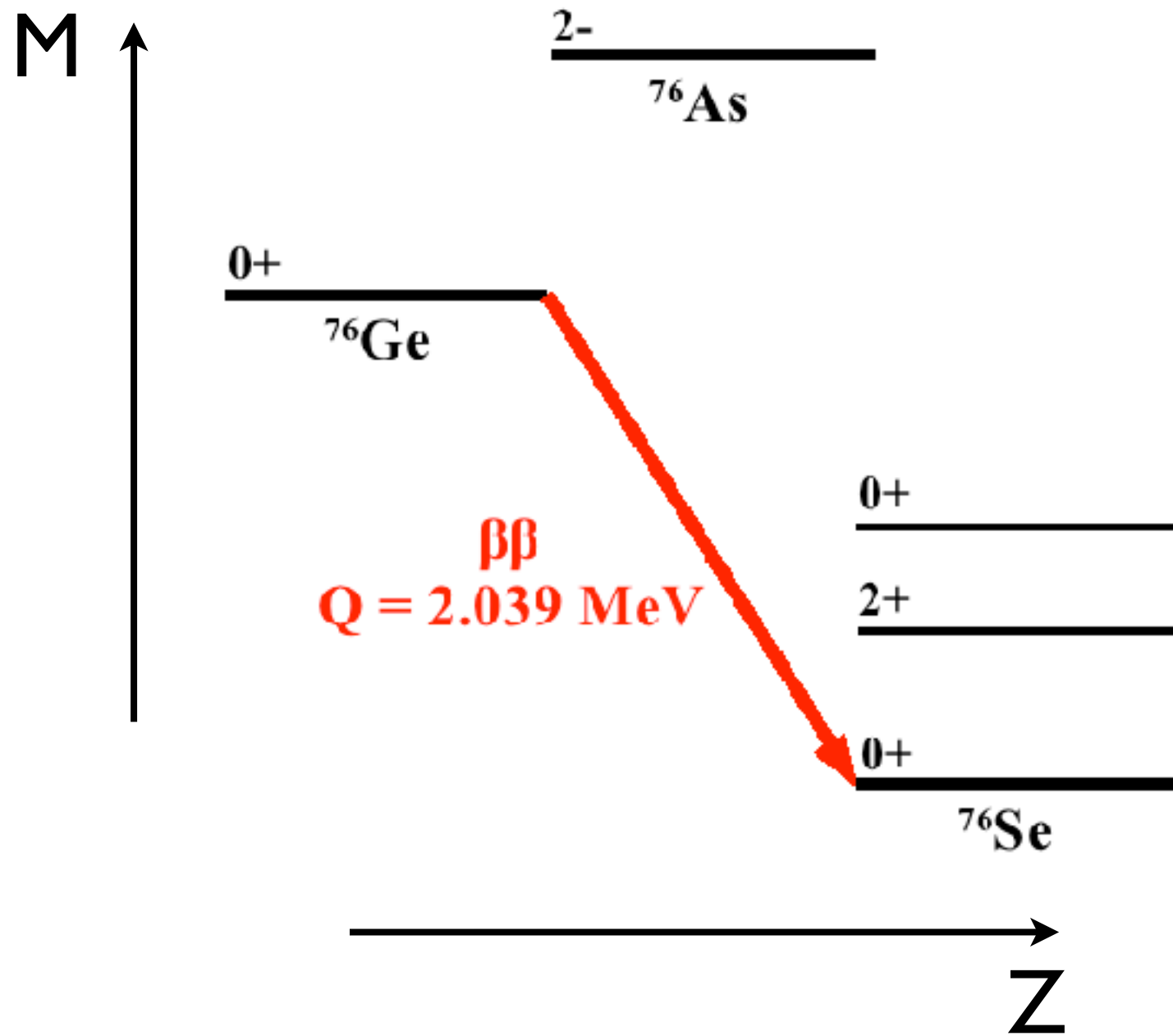
Why not generate ν_R in a beam and see if it produces ℓ^+ ?

Testing $\nu = \bar{\nu}$

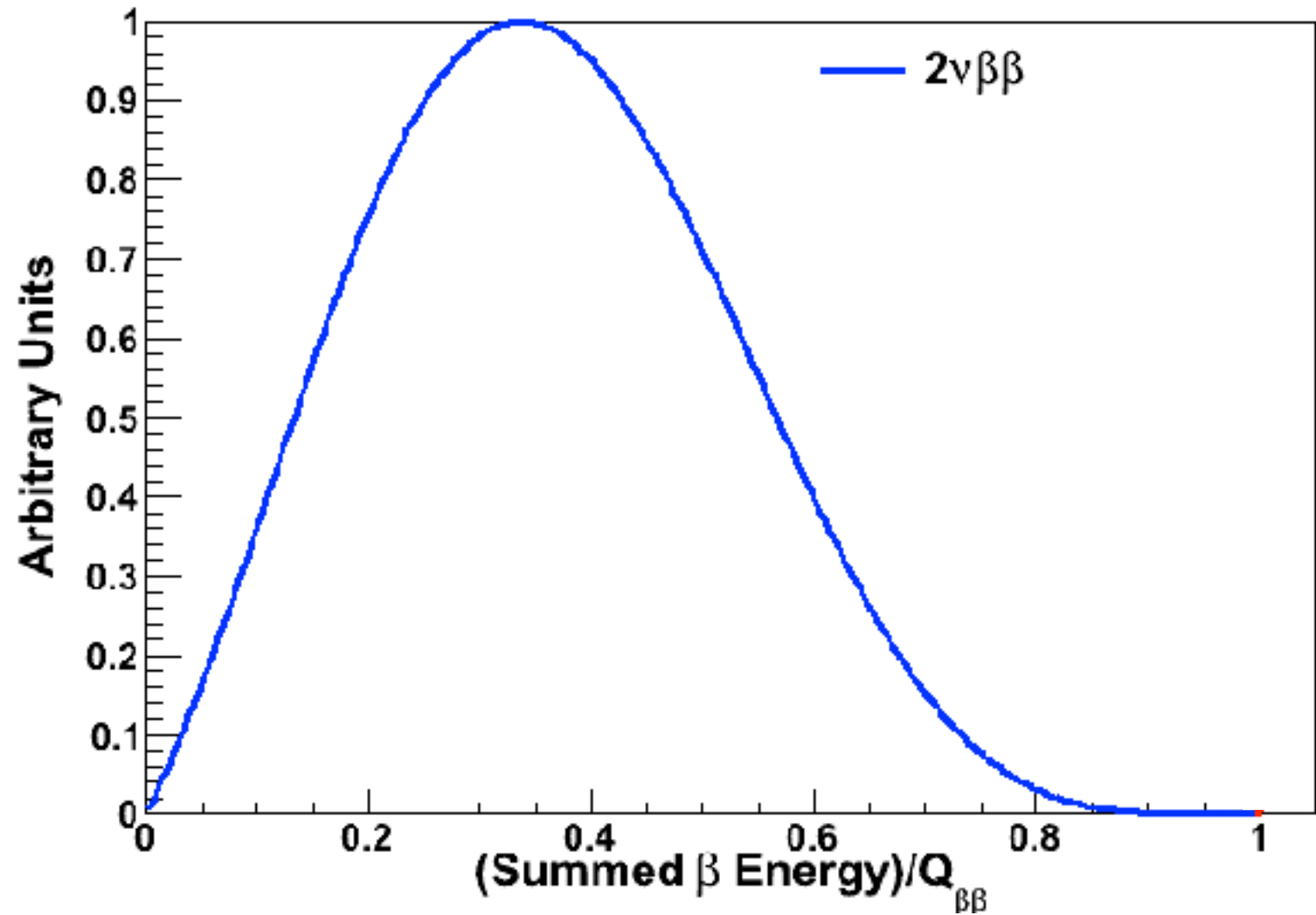
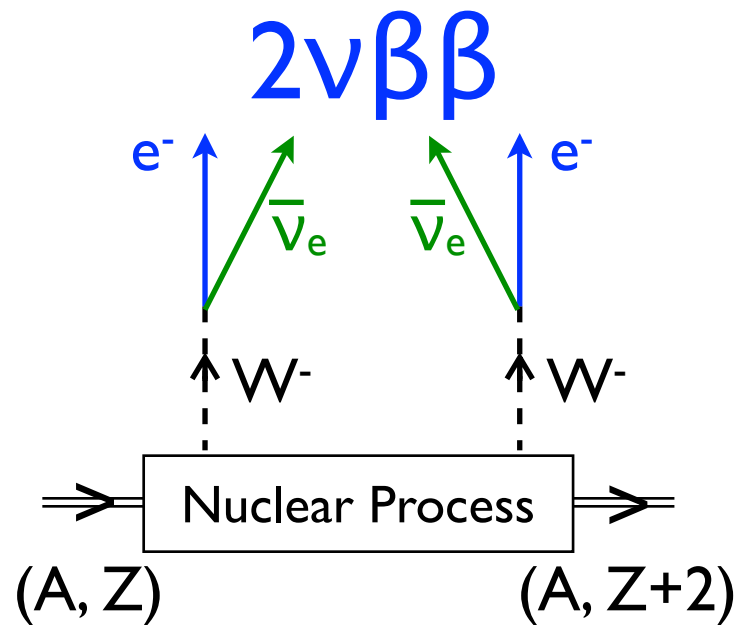


- Boost so that π^+ beam faster than ν_L from decay at rest: requires $E_\pi > 4 \text{ PeV}$ (n.b. LHC = 14 TeV)
- Fraction of decays with helicity flipped: $< 10^{-15}$
- “Since *L*-violation comes only from Majorana ν masses, any attempt to observe it will be at the mercy of the ν masses.”
- B. Kayser

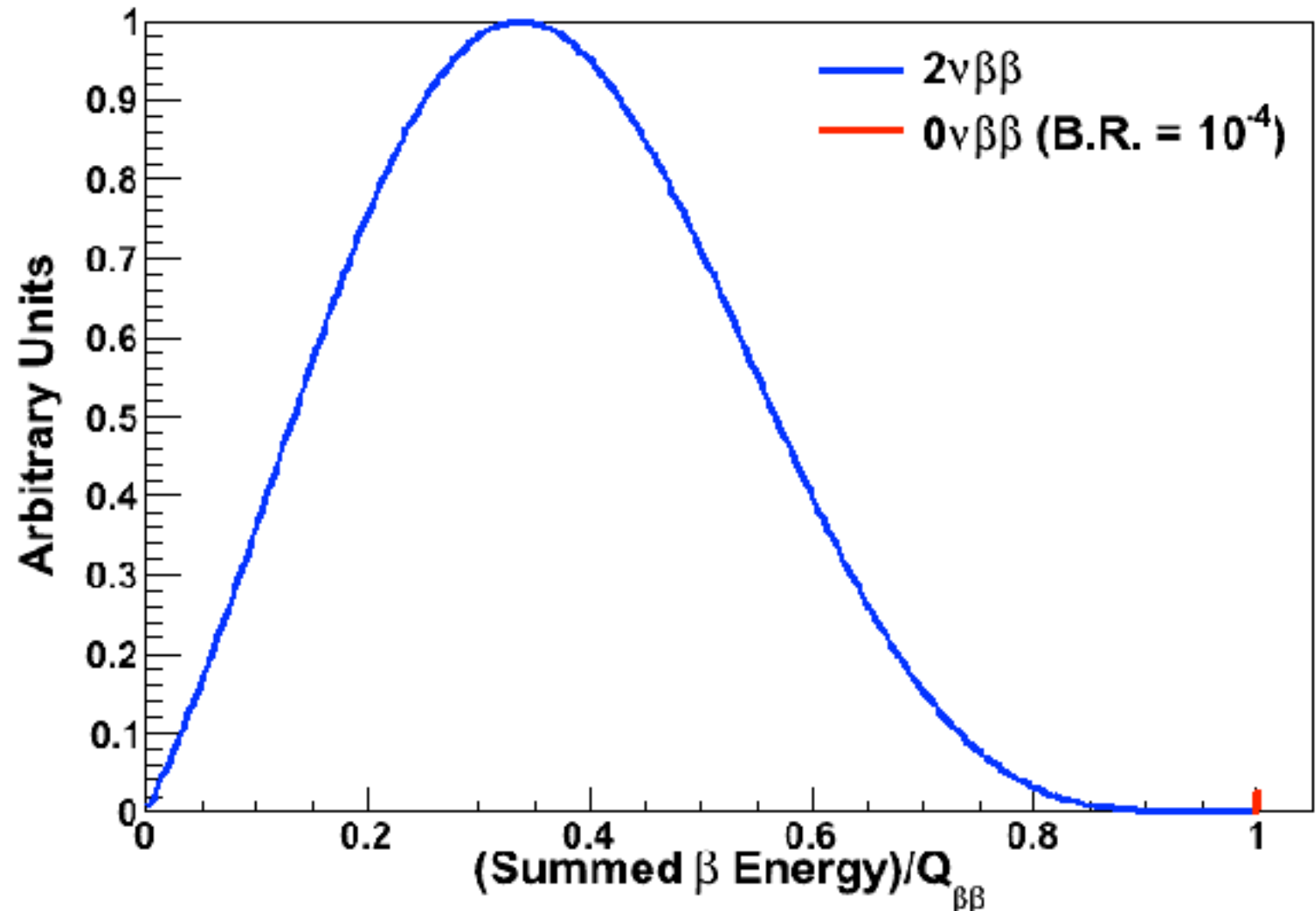
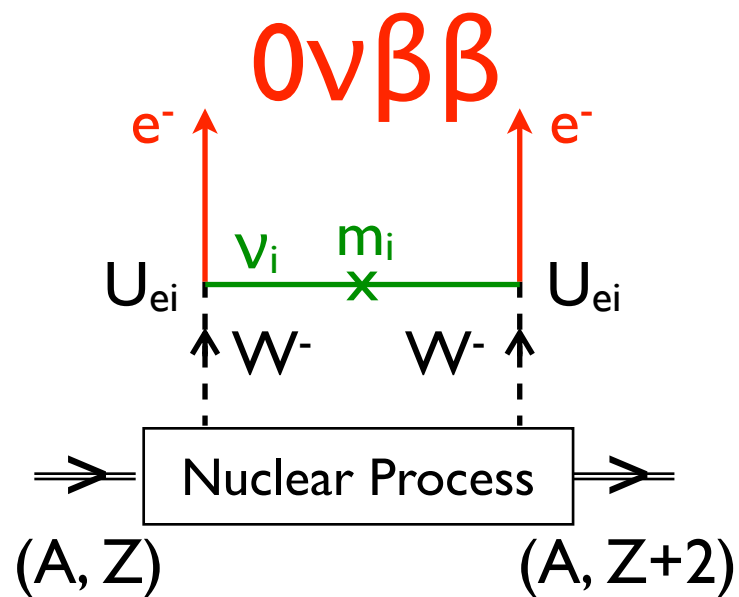
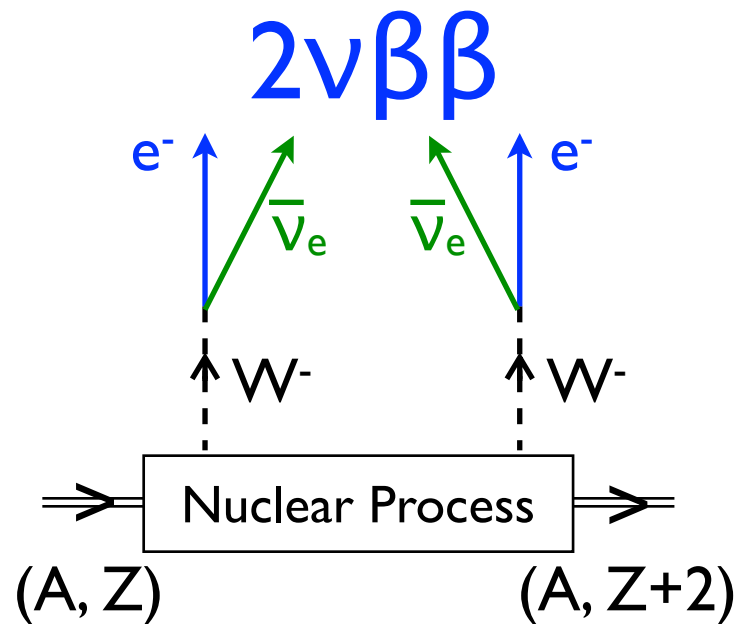
Double-Beta Decay



Double-Beta Decay



$0\nu\beta\beta$ Decay

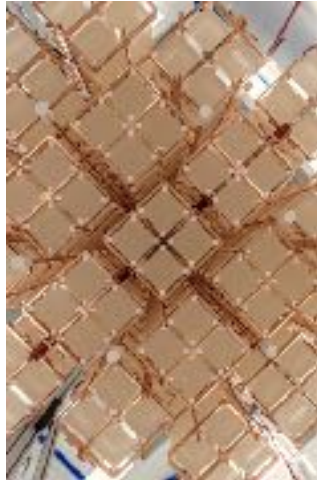


$$T_{1/2} = 1/G |M|^2 m_{\beta\beta}^2$$

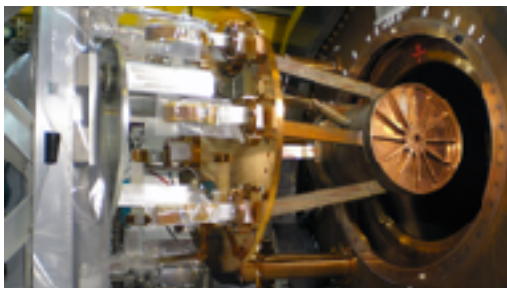
$$m_{\beta\beta} \equiv \left| \sum m_i U_{ei}^2 \right|$$

$0\nu\beta\beta$ Experiments

CUORE



EXO-200



NEMO3



Collaboration	Isotope	Technique	mass ($0\nu\beta\beta$ isotope)	Status
AMoRE	Mo-100	CaMoO4 bolometers (+ scint.)	5	Construction
CANDLES	Ca-48	305 kg CaF2 crystals - liq. scint	0.3 kg	Operating
CARVEL	Ca-48	⁴⁸ CaWO4 crystal scint.	16 kg	R&D
GERDA I	Ge-76	Ge diodes in LAr	15 kg	Operating
GERDA II	Ge-76	Point contact Ge in LAr	20 kg	Construction
MAJORANA DEMONSTRATOR	Ge-76	Point contact Ge in Lead	26 kg	Construction
1TGe (GERDA & MAJORANA)	Ge-76	Best of GERDA + MJD	~tonne	R&D
NEMO3	Mo-100 Se-82	Foils with tracking	6.9 kg 0.9 kg	Complete
SuperNEMO Demonstrator	Se-82	Foils with tracking	7 kg	Construction
SuperNEMO	Se-82	Foils with tracking	100 kg	R&D
MOON	Mo-100	Mo sheets	200 kg	R&D
CAMEO	Cd-116	CdWO4 crystals	21 kg	R&D
COBRA	Cd-116, Te-130	CdZnTe detectors	10 kg	Operating / Construction
CUORICINO	Te-130	TeO2 Bolometer	11 kg	Complete
CUORE-0	Te-130	TeO2 Bolometer	11 kg	Complete
CUORE	Te-130	TeO2 Bolometer	206 kg	Operating
SNO+	Te-130	0.3% ^{nat} Te in liquid scint.	800 kg	Construction
KamLAND-ZEN	Xe-136	2.7% in liquid scint.	370 kg	Operating
KamLAND2-ZEN	Xe-136	2.7% in liquid scint.	~tonne	R&D
NEXT-100	Xe-136	High pressure Xe TPC	10 kg	Construction
EXO-200	Xe-136	Xe liquid TPC	160 kg	Operating
nEXO	Xe-136	Xe liquid TPC	5 tonnes	R&D
DCBA	Nd-150	Nd foils & tracking chambers	30 kg	R&D

Complete

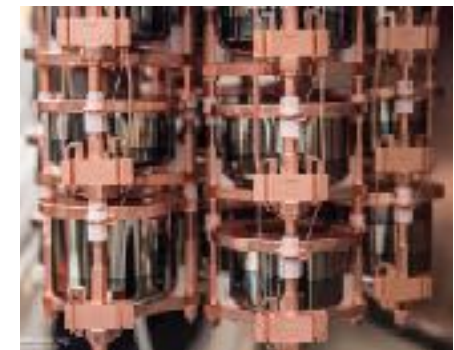
Construction

Operating

GERDA



MAJORANA

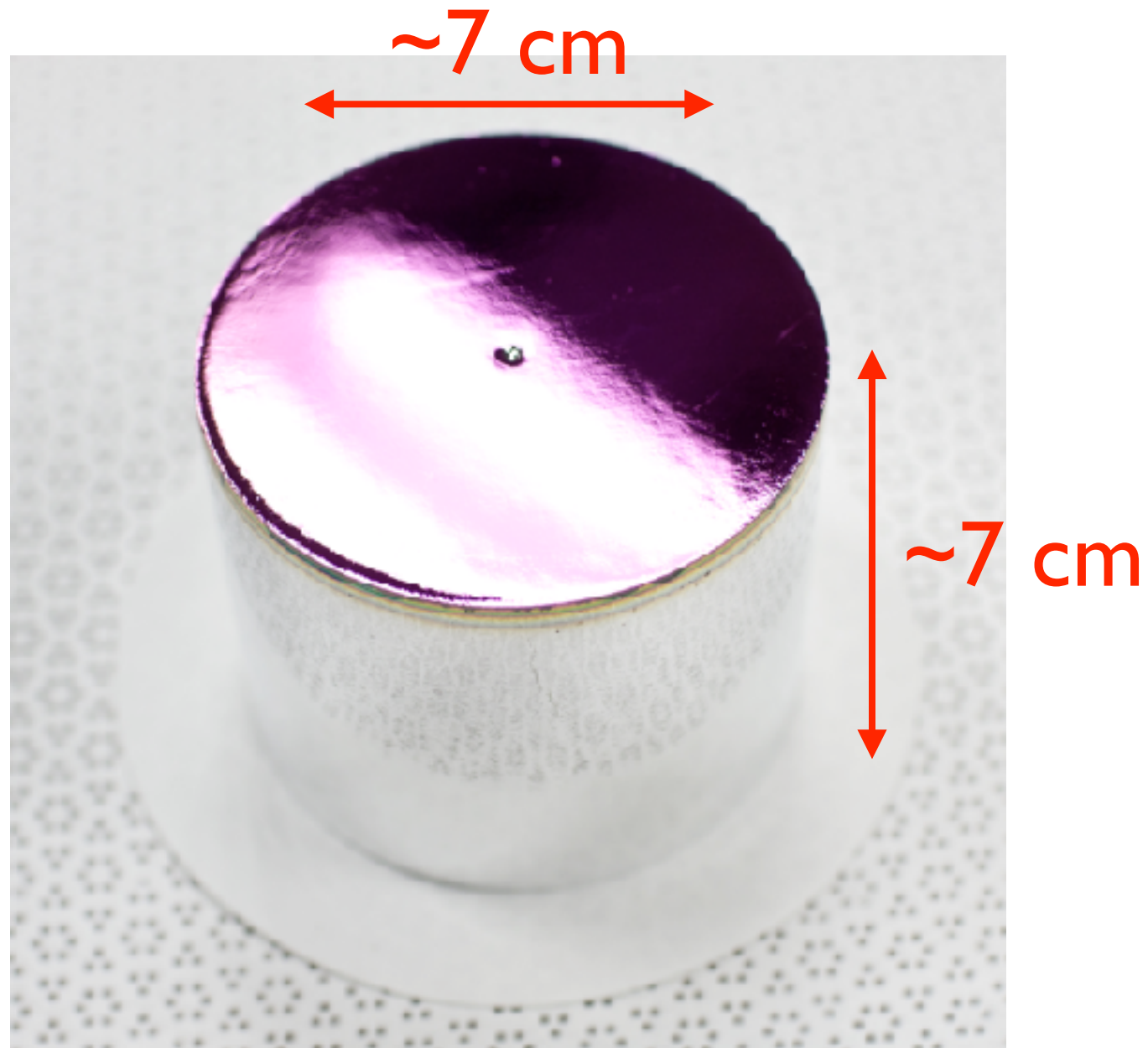


CANDLES

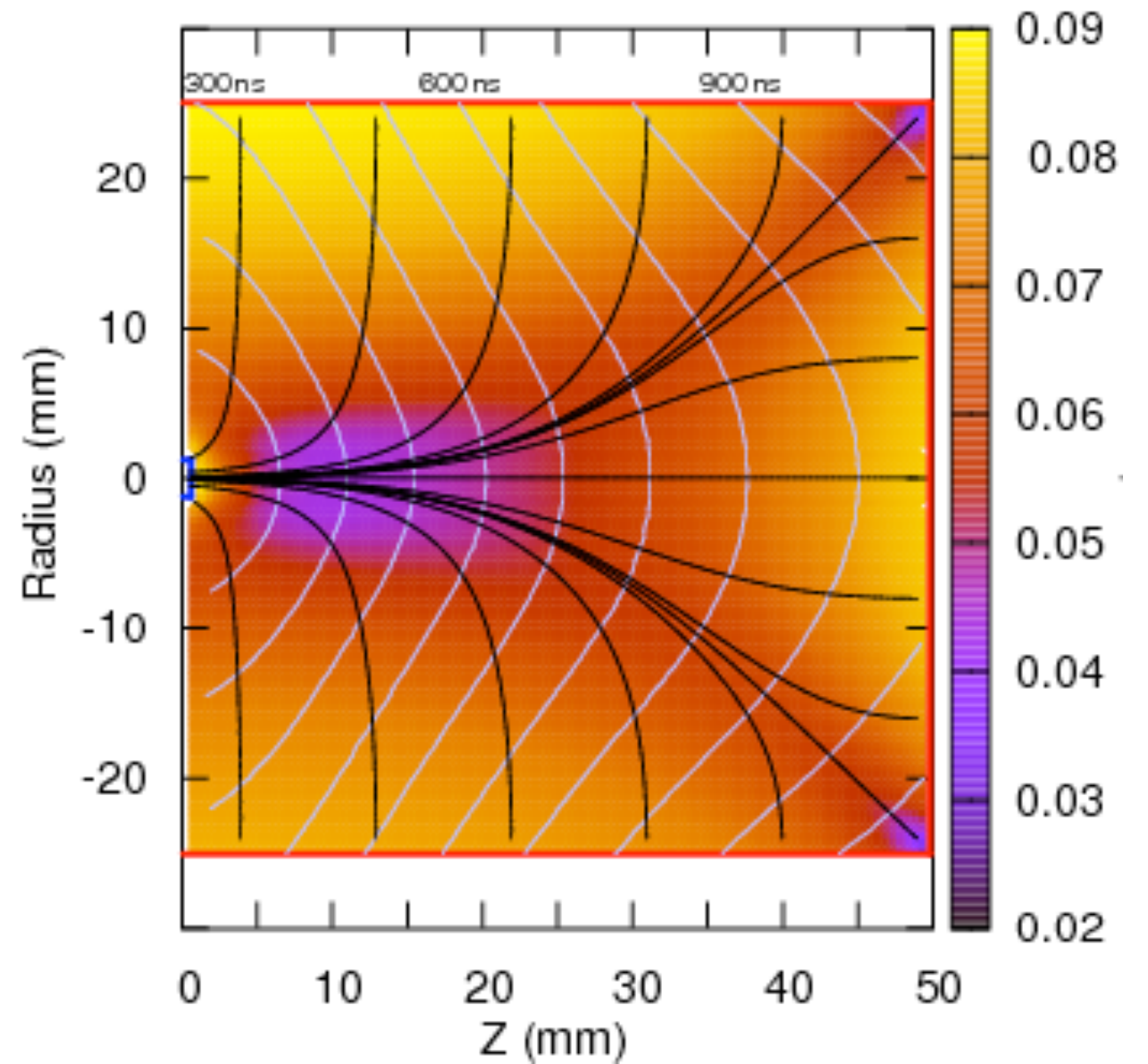
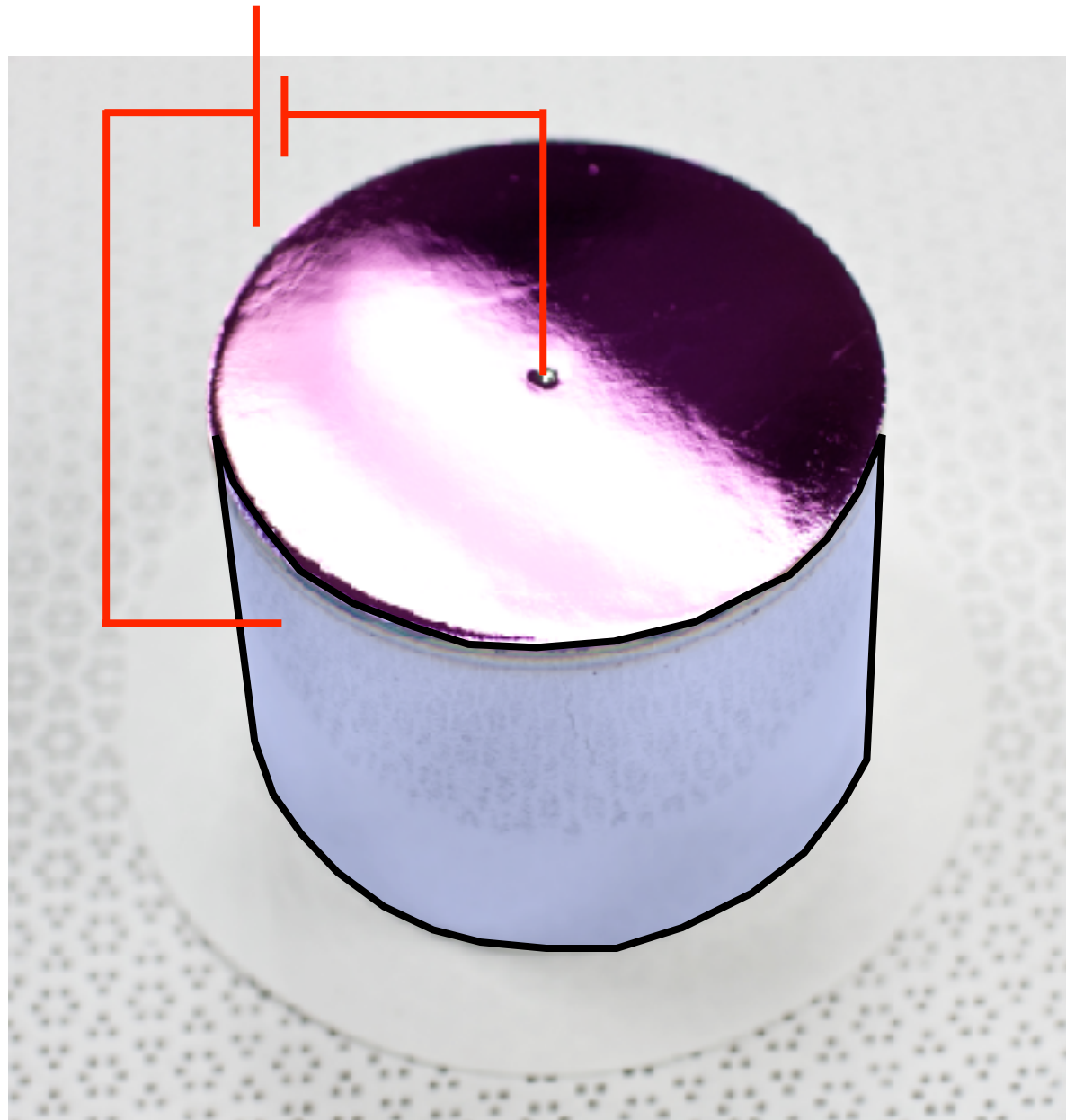


From J. F. Wilkerson

Germanium Detectors

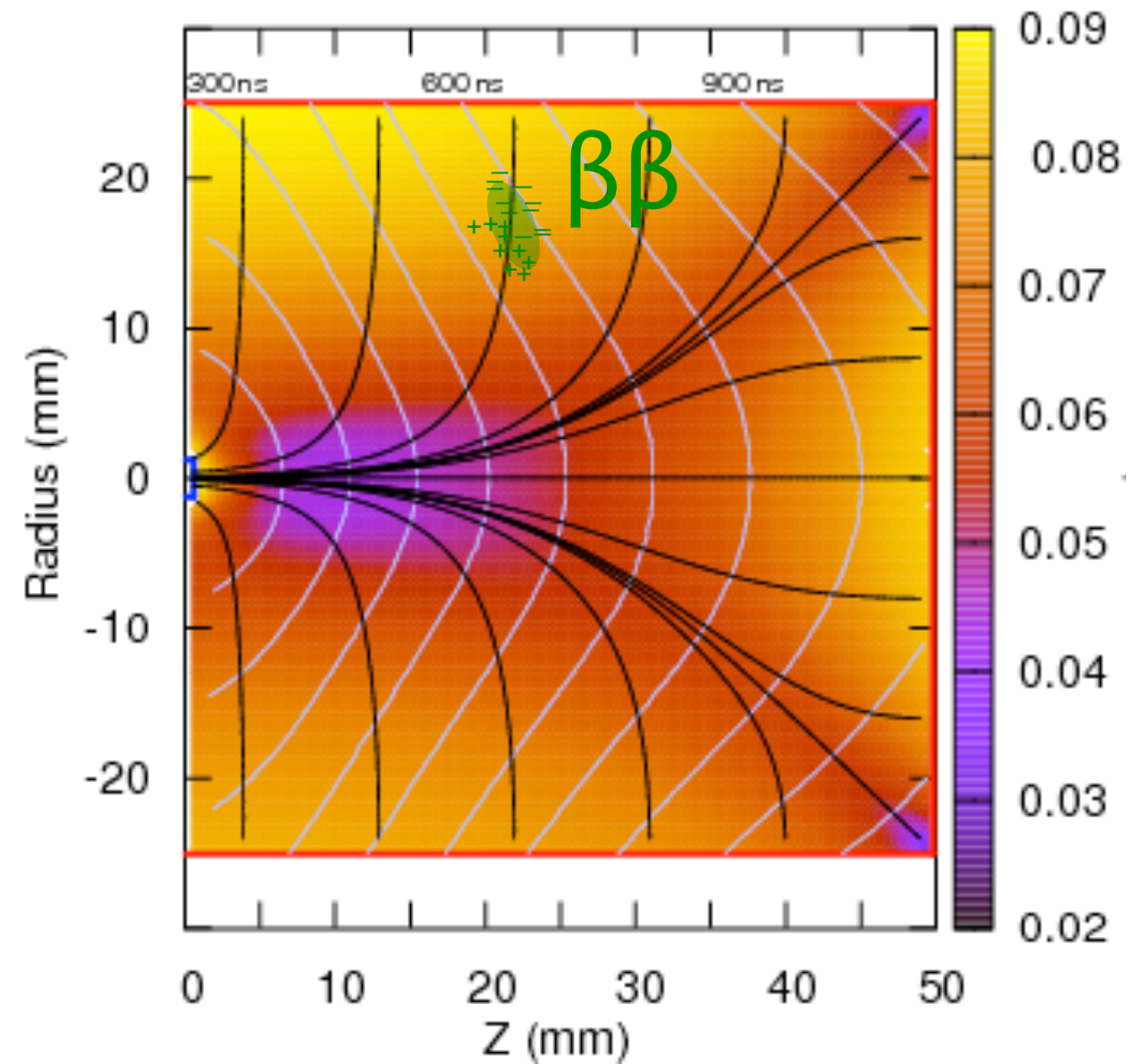
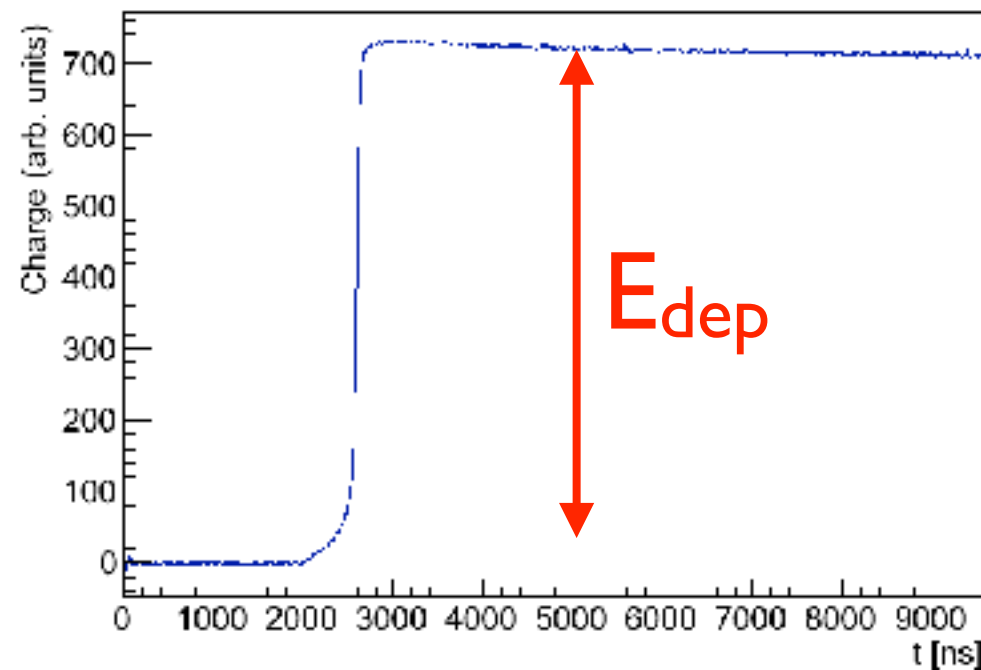
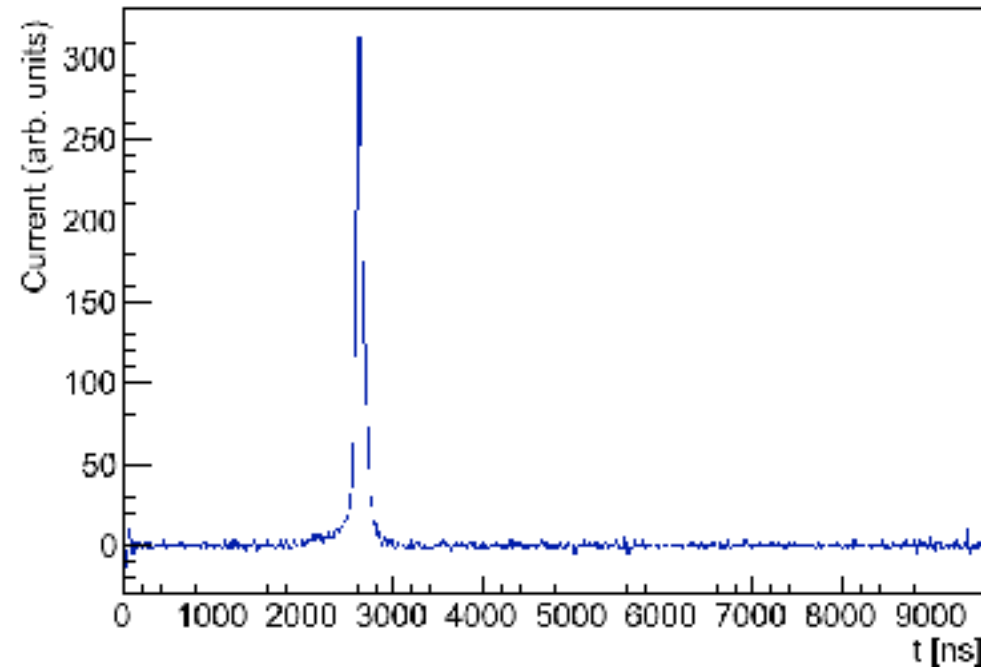


Germanium Detectors



Hole v_{drift} (mm/ns) w/ paths, isochrones

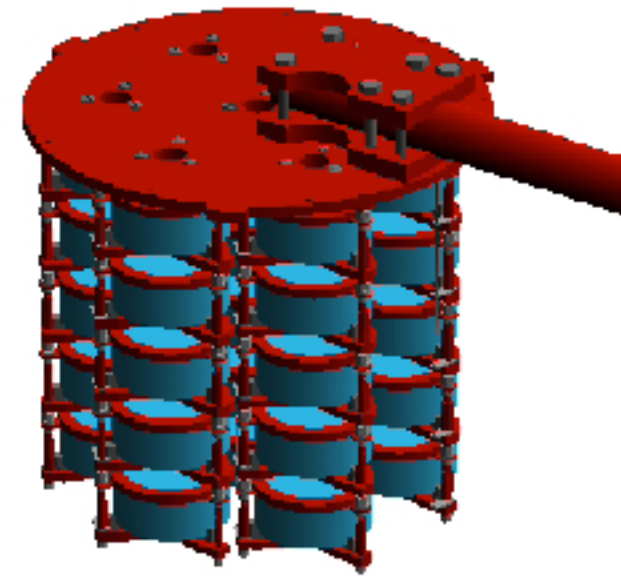
Germanium Detectors



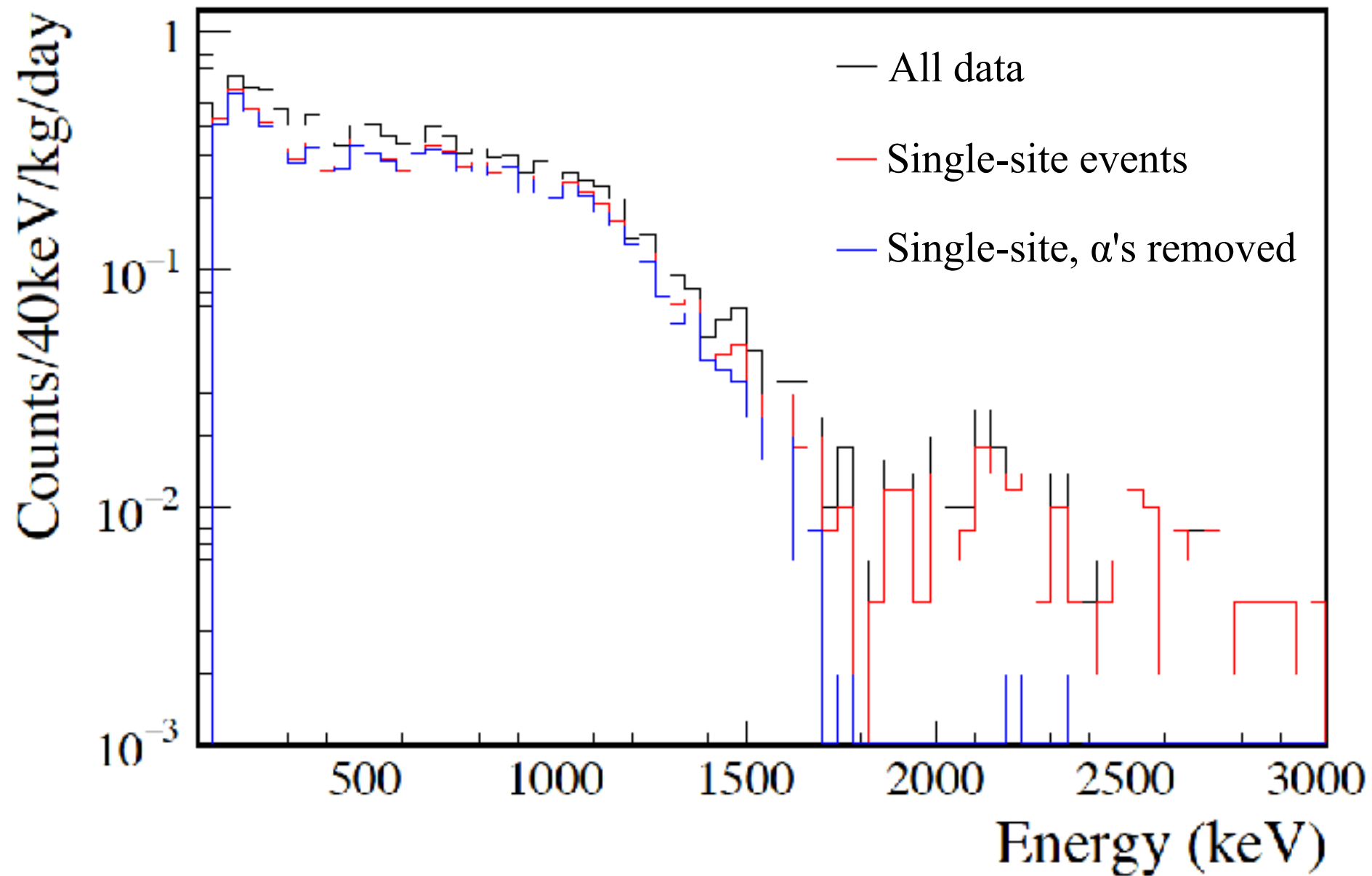
Hole v_{drift} (mm/ns) w/ paths, isochrones

The MAJORANA DEMONSTRATOR

- Goal: x100 reduction in background vs. previous efforts using clean materials, hit patterns, pulse-shapes
- Located at the 4850' level of Sanford Underground Laboratory in SD
- Modules:
 - Prototype: 3 strings $^{\text{nat}}\text{Ge}$ (Jun 2014 - Jun 2015)
 - Module 1: ~ 20 kg $^{\text{enr}}\text{Ge}$ (May 2015 - present; low-BG from Jan 2016)
 - Module 2: ~ 10 kg $^{\text{enr}}\text{Ge}$ + ~ 10 kg $^{\text{nat}}\text{Ge}$ (Summer 2016 - present)

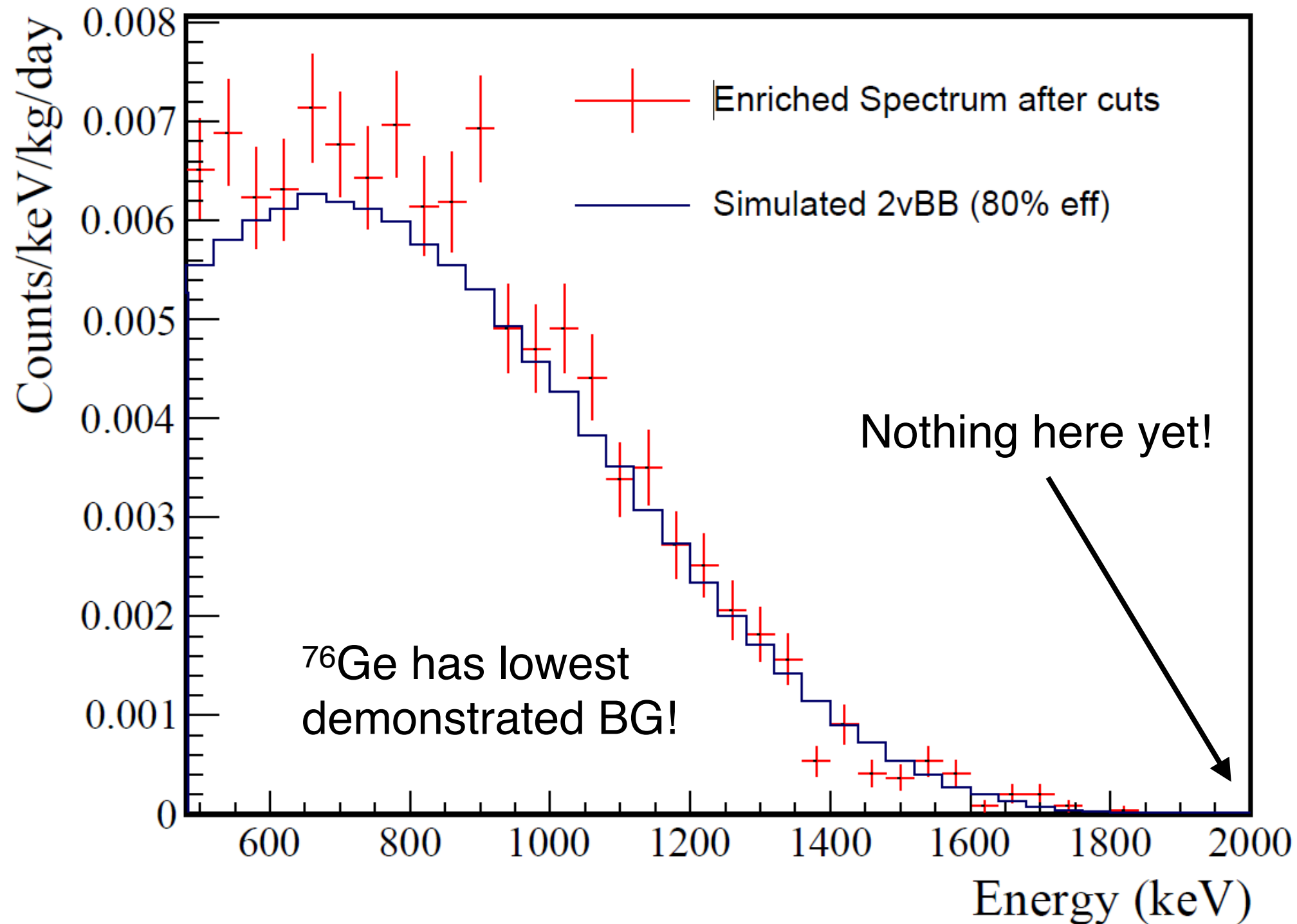


Preliminary Results



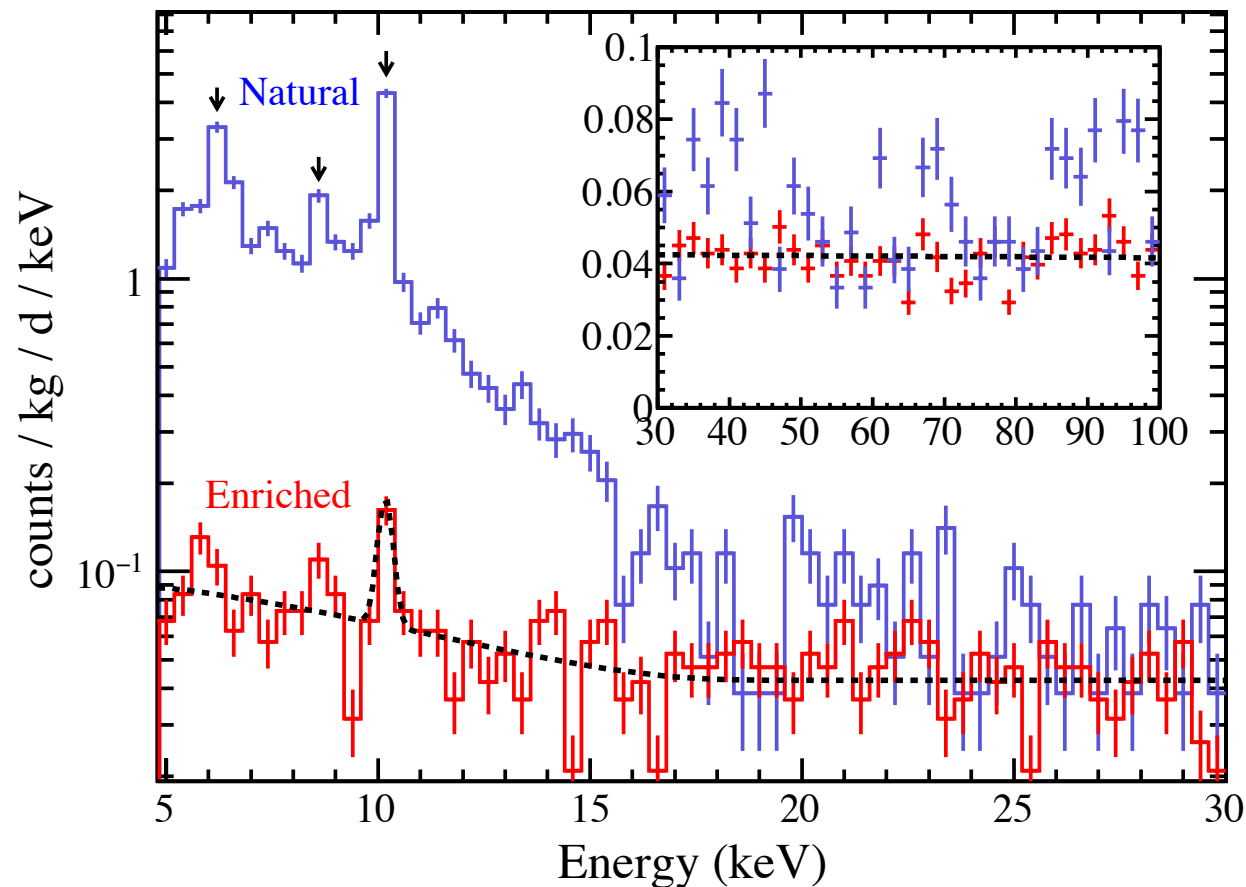
Campaign at UW to fully characterize α interactions

Preliminary Results

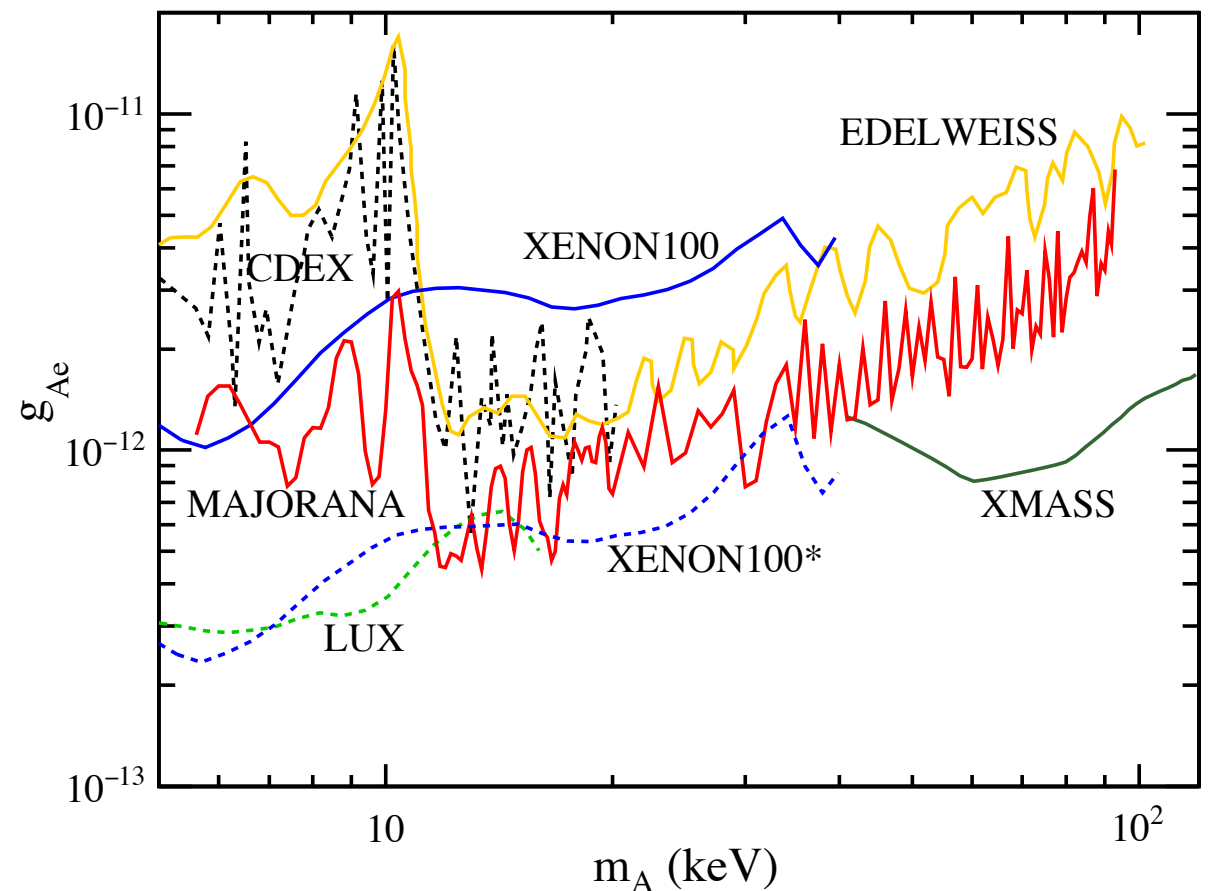


Low-Energy Physics

Data below 100 keV

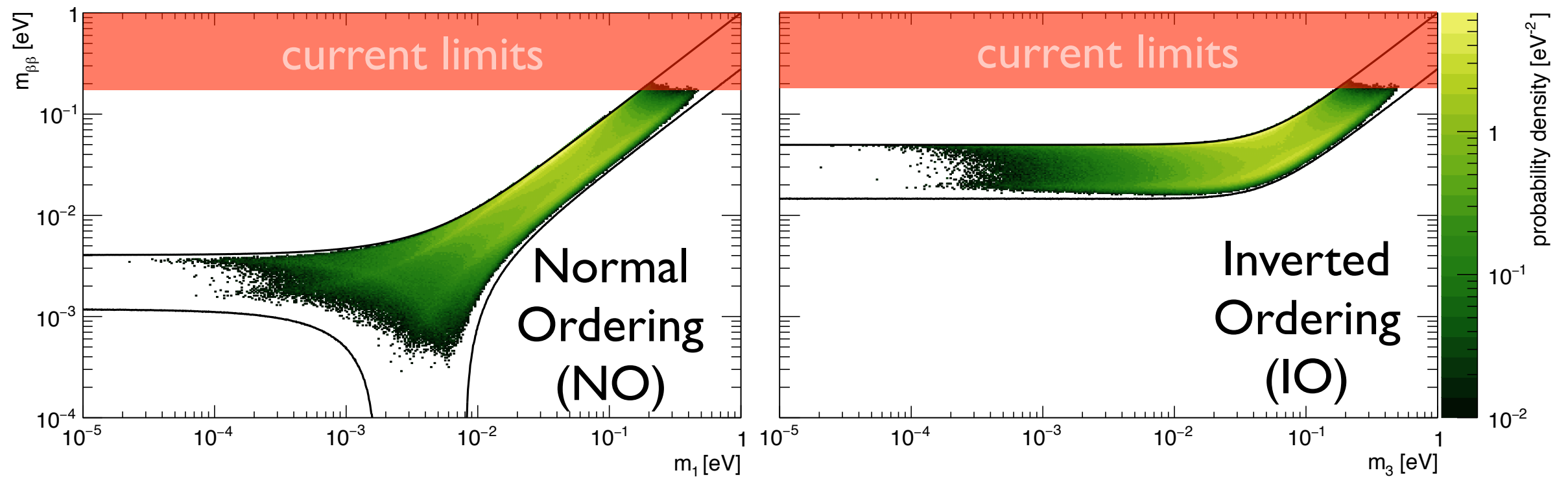


Bosonic Dark Matter Limits



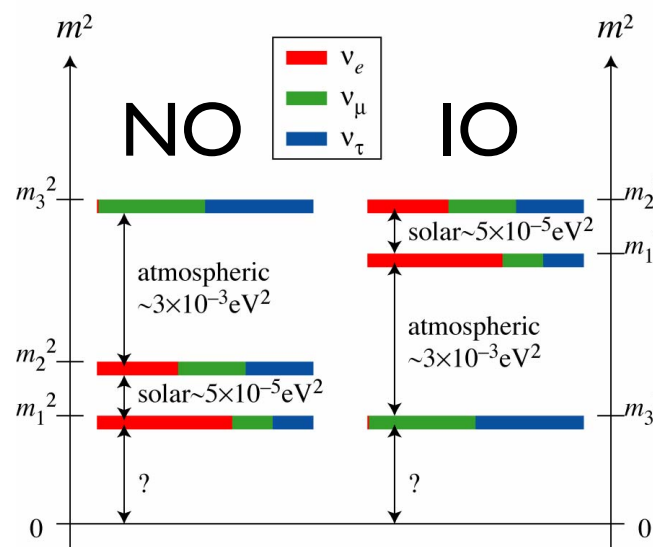
Also places limits on solar axions, Pauli Exclusion Principle violation, electron decay...

Next-Generation Goals



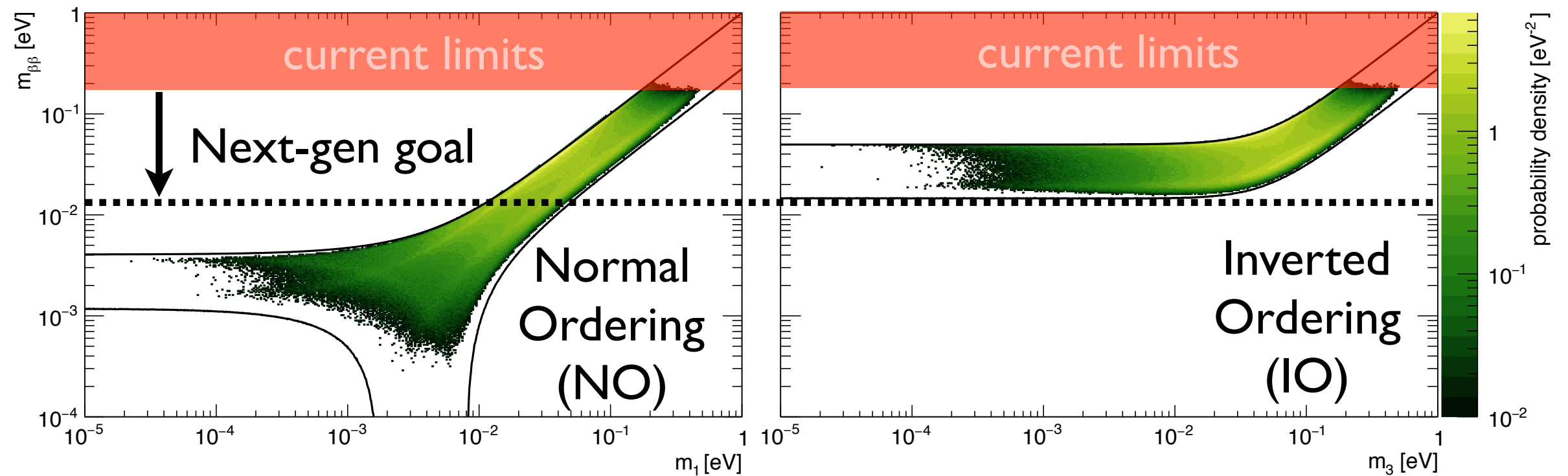
$$T_{1/2} = 1/G |M|^2 m_{\beta\beta}^2$$

$$m_{\beta\beta} \equiv \left| \sum m_i U_{ei}^2 \right|$$



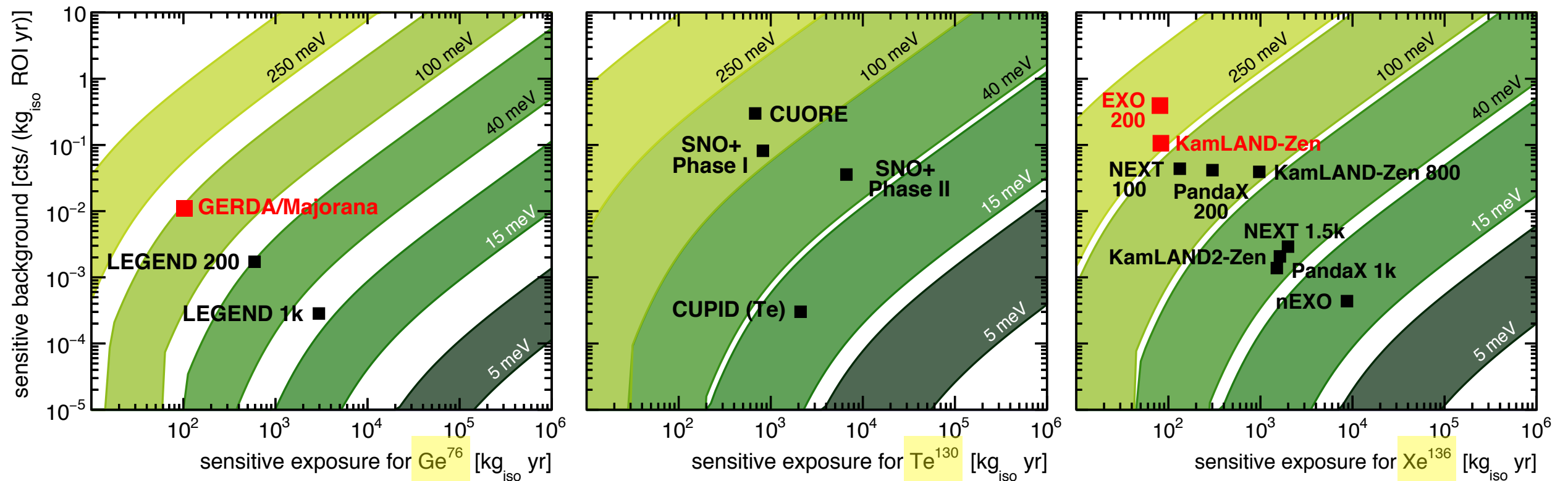
Parameter space weighted assuming non-vanishing lightest ν mass

Next-Generation Goals



Cost: $O(\$10^8)$
(per experiment)

Next-Gen Experiments



- Red dots: published limits. Black dots: 3σ discovery sensitivities with 5 yrs live time
- Discovery sensitivity after 10 yr is $\sim\sqrt{2}$ higher for all experiments
- Bands represent variation in nuclear matrix element calcs

Summary

- Neutrinos continue to give us new mysteries to solve
- Tests of Majorana neutrinos may give us insights into Grand Unification and why the Universe exists
- Next generation experiments will cover much of the remaining parameter space