

Is the Neutrino its Own Antiparticle?

Jason Detwiler
CENPA

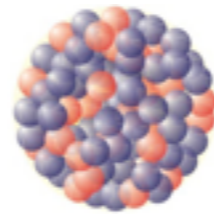
REU Summer Seminar Series
University of Washington, Seattle, WA
July 22, 2013

Outline

- What's a neutrino?
- The case for Majorana neutrinos
- Probing the nature of the neutrino with neutrinoless double-beta decay

What's a Neutrino?

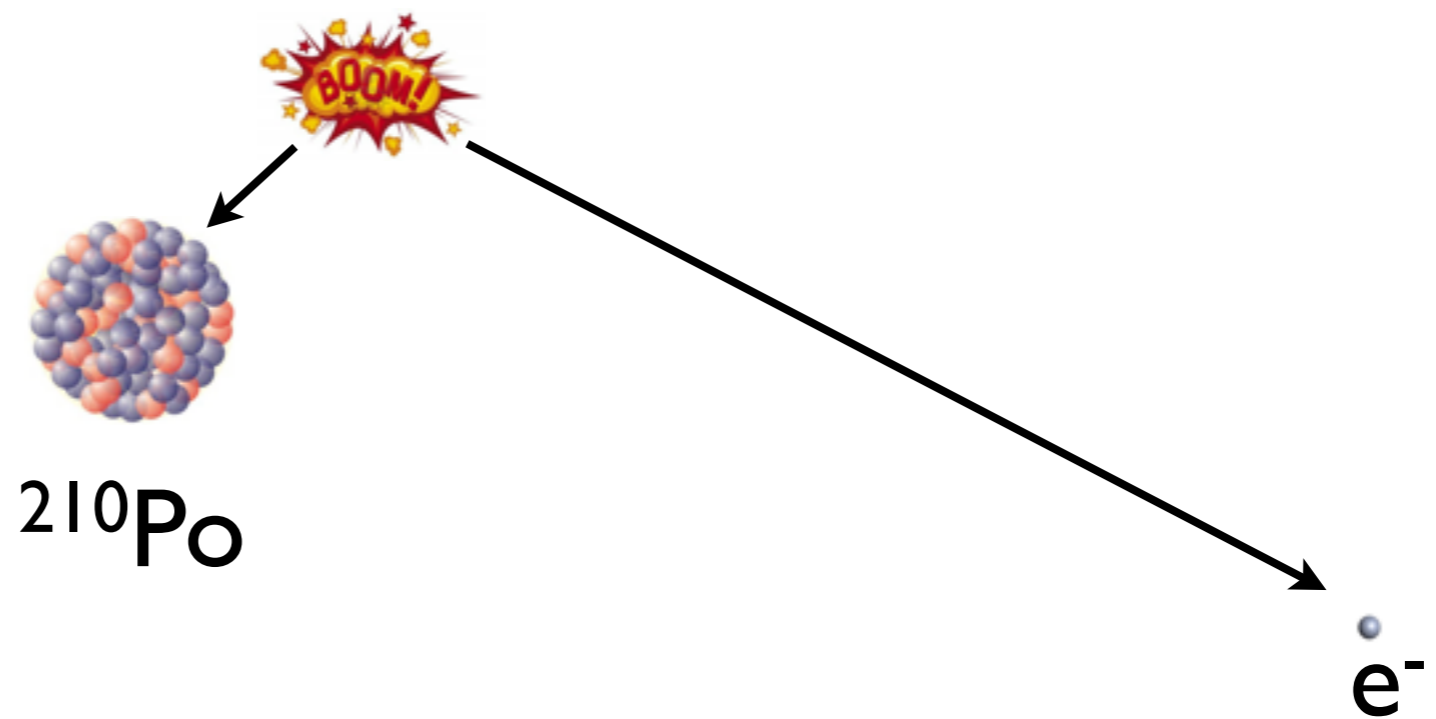
Meitner and
Hahn (1911):



^{210}Bi

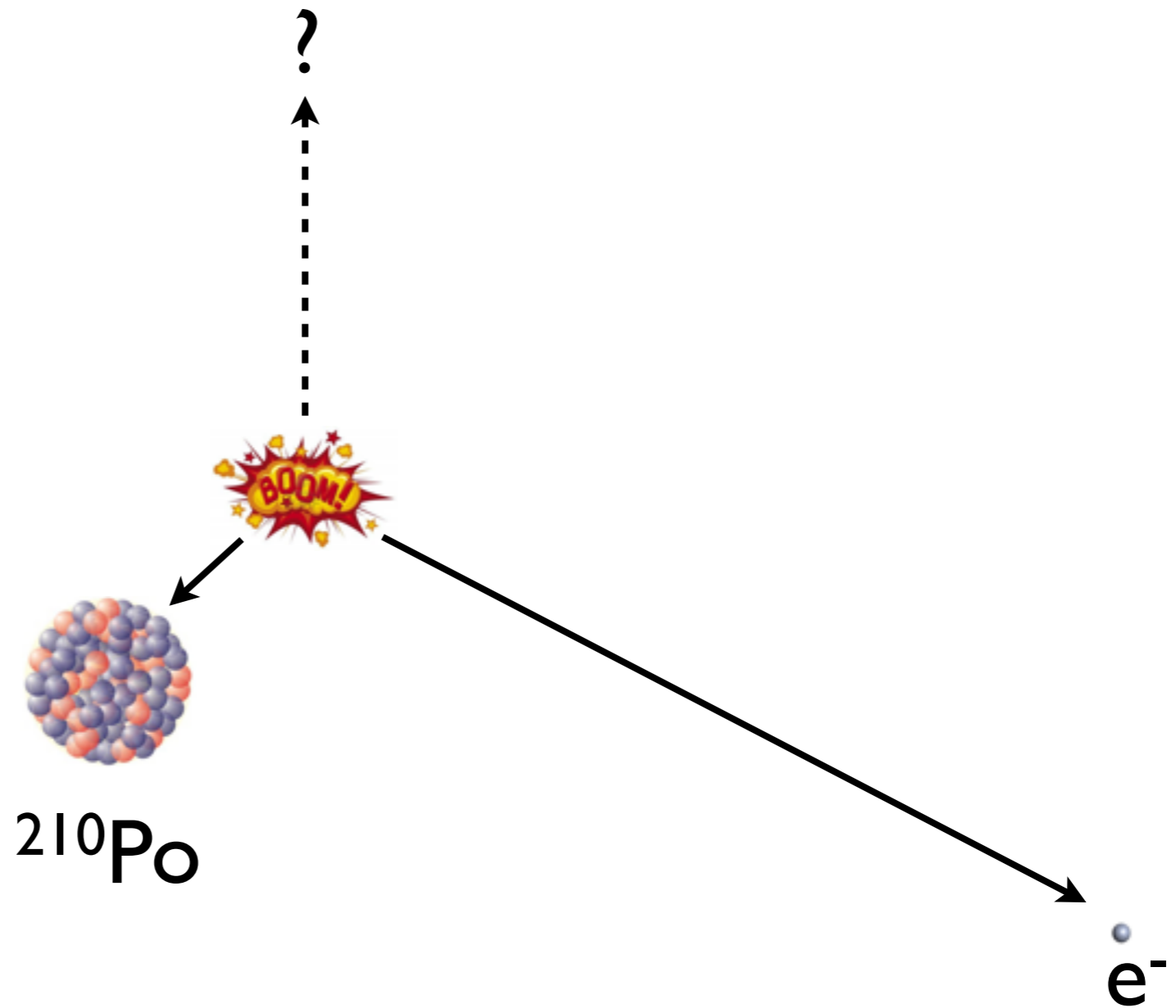
What's a Neutrino?

Meitner and
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What's a Neutrino?

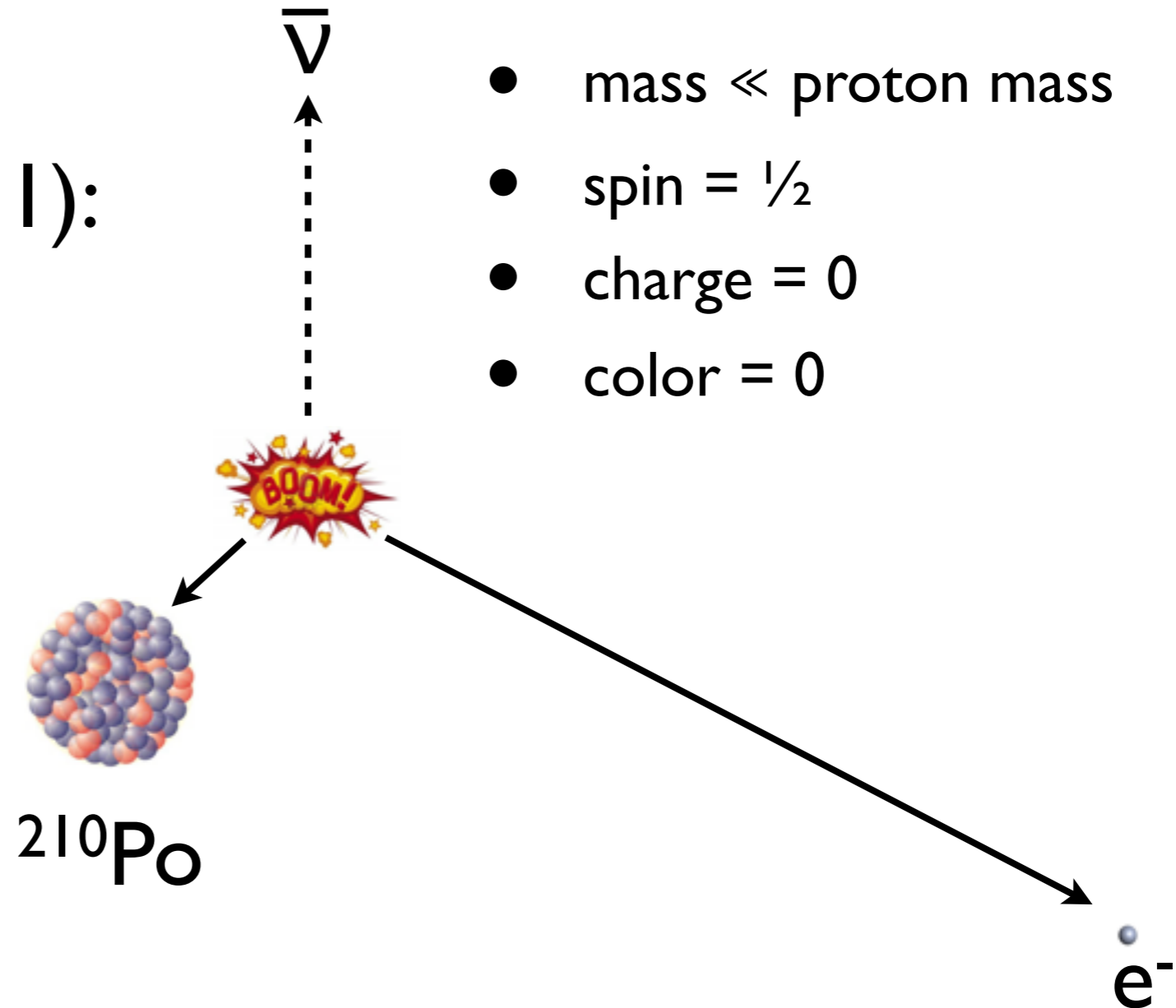
Meitner and
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What's a Neutrino?

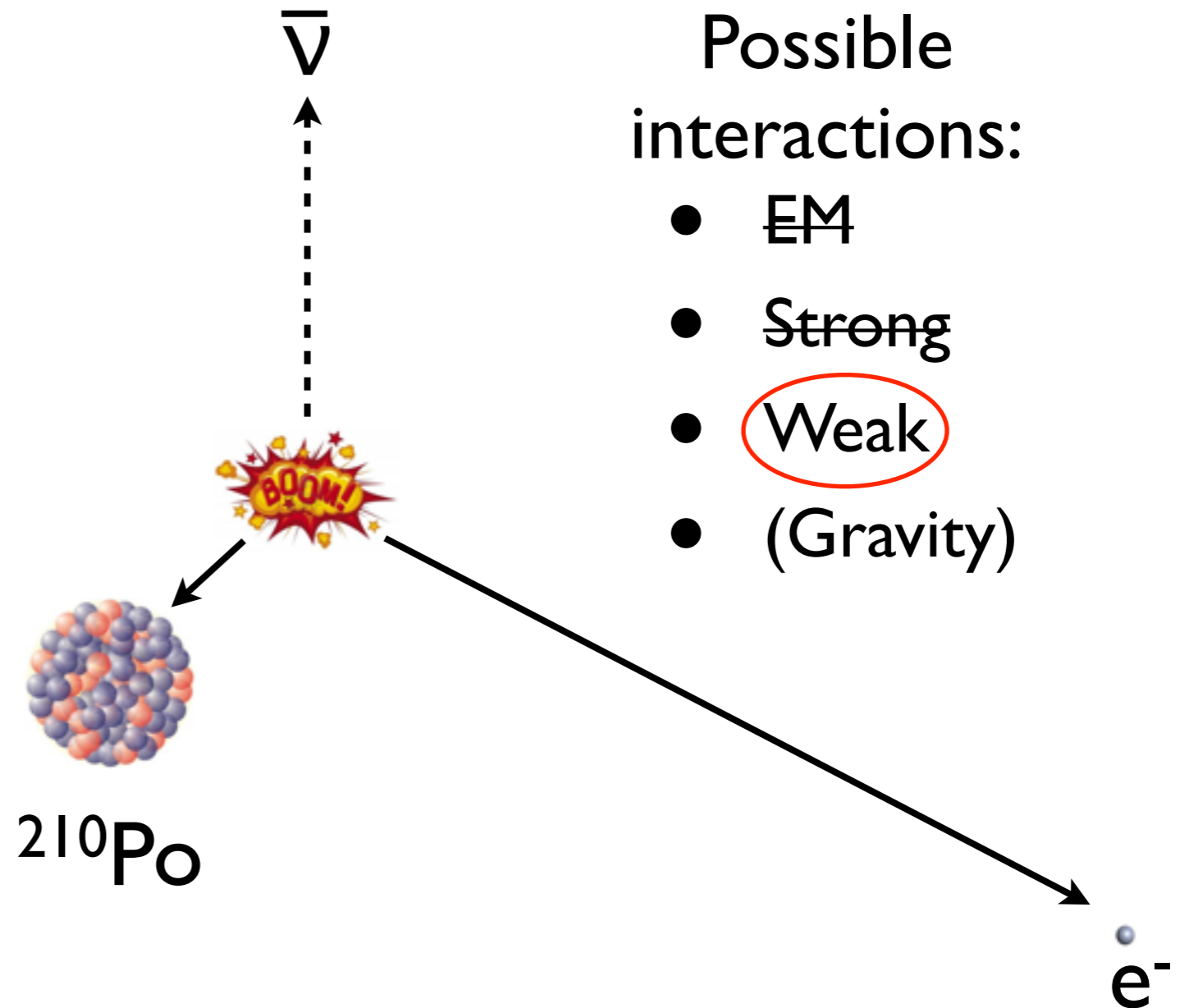
Wolfgang Pauli (1931):

- mass \ll proton mass
- spin = $1/2$
- charge = 0
- color = 0



What's a Neutrino?

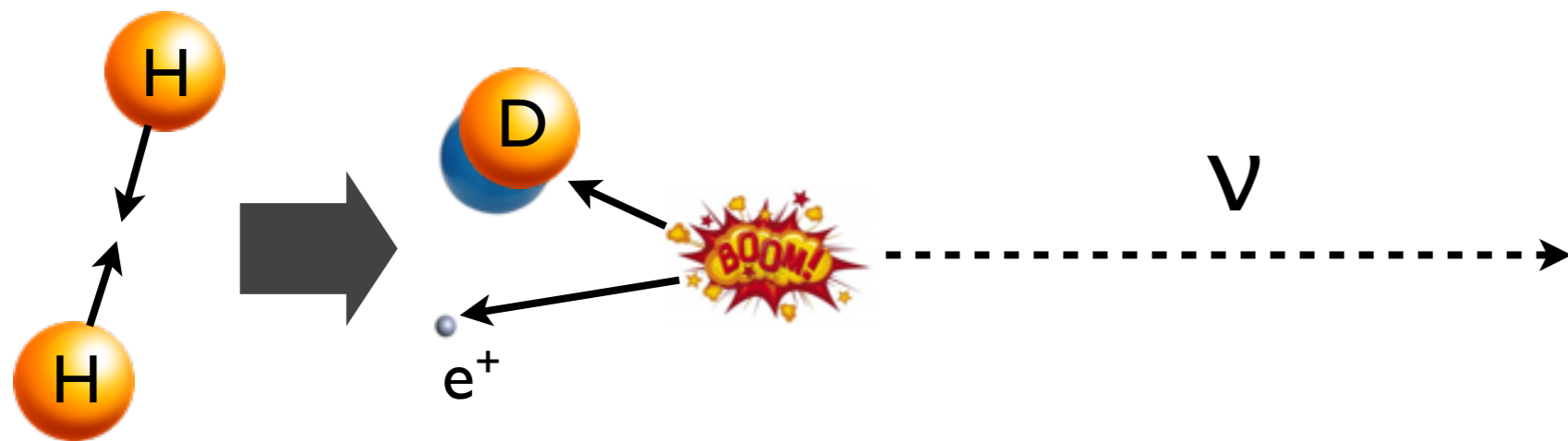
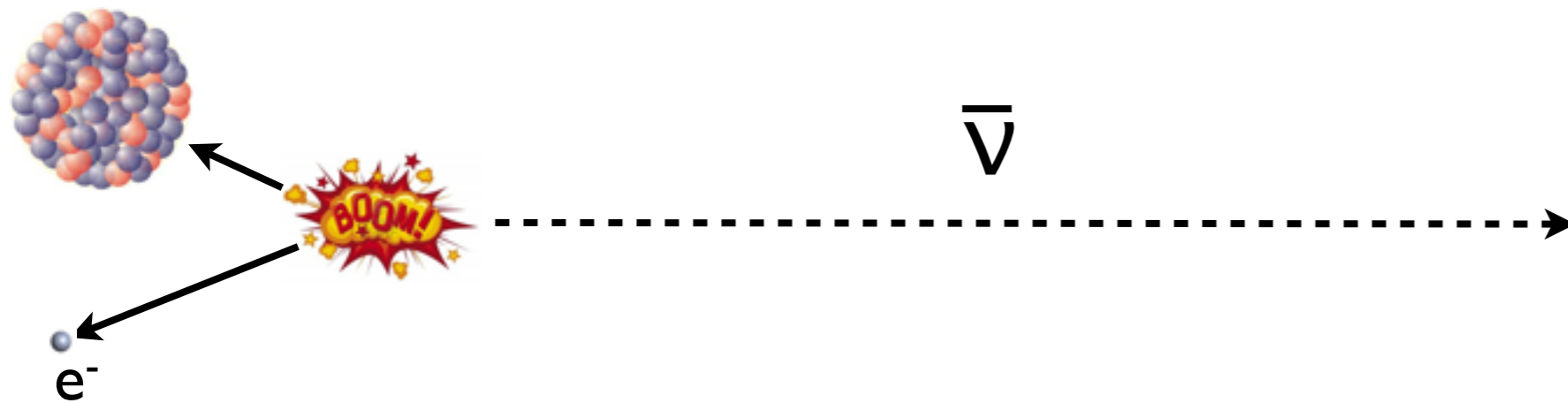
Enrico Fermi (1934):
“Little neutral one”



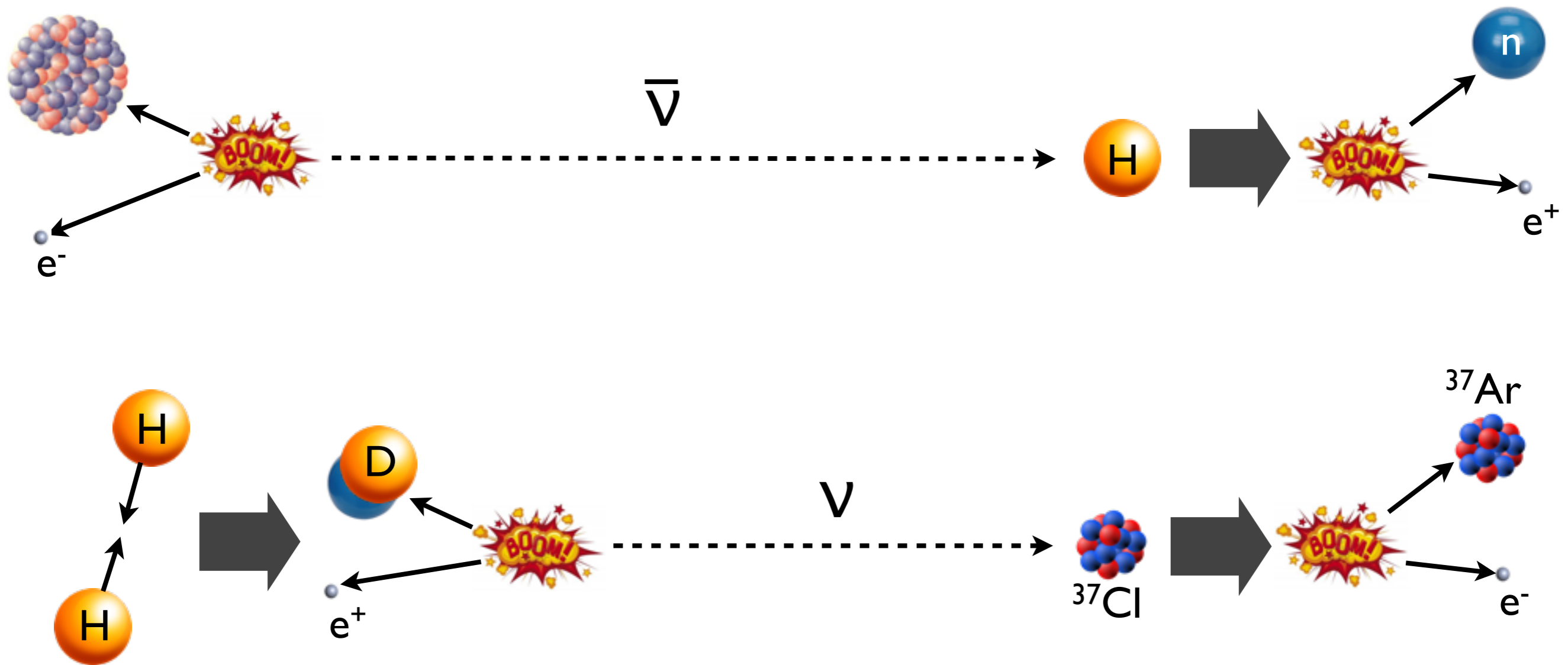
Possible interactions:

- EM
- Strong
- Weak
- (Gravity)

Detecting Neutrinos and Antineutrinos



Detecting Neutrinos and Antineutrinos



Detecting Neutrinos and Antineutrinos

Nuclear Reactor

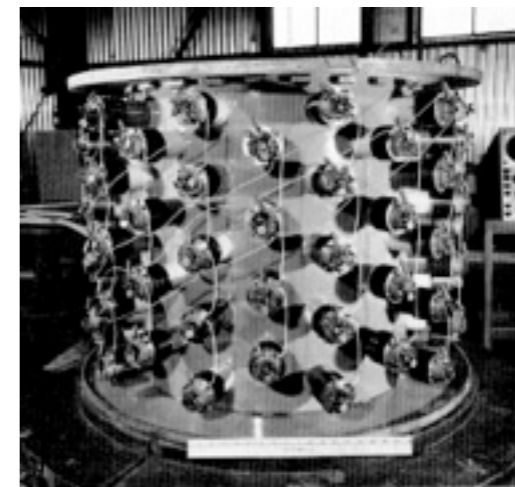


$\bar{\nu}$



Cowan and Reines (1956)

Scintillator (C_xH_y)



PMTs

The Sun



ν



Ray Davis Jr. (1964)

Cleaning fluid (Cl)



+ Ar detector

Detecting Neutrinos and Antineutrinos

Nuclear Reactor

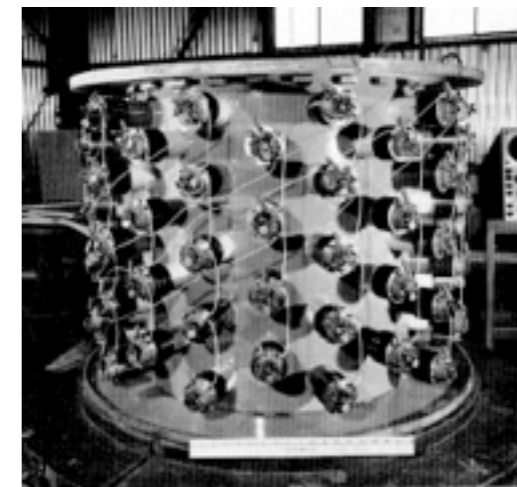


$\bar{\nu}_R$



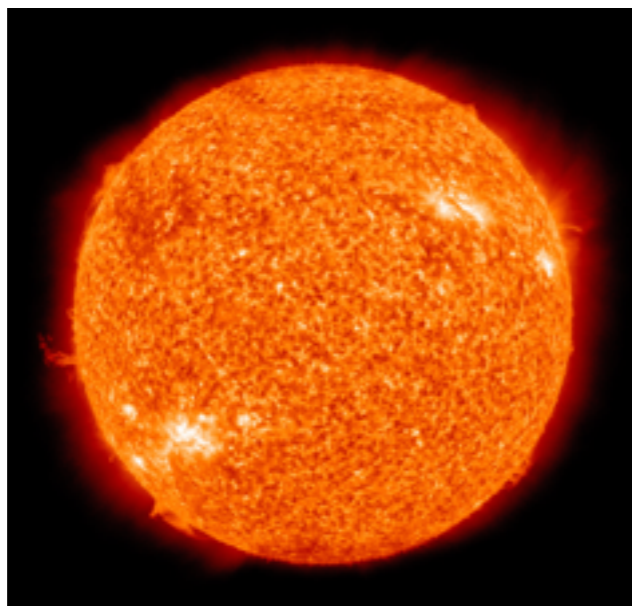
Cowan and Reines (1956)

Scintillator (C_xH_y)



PMTs

The Sun



ν_L



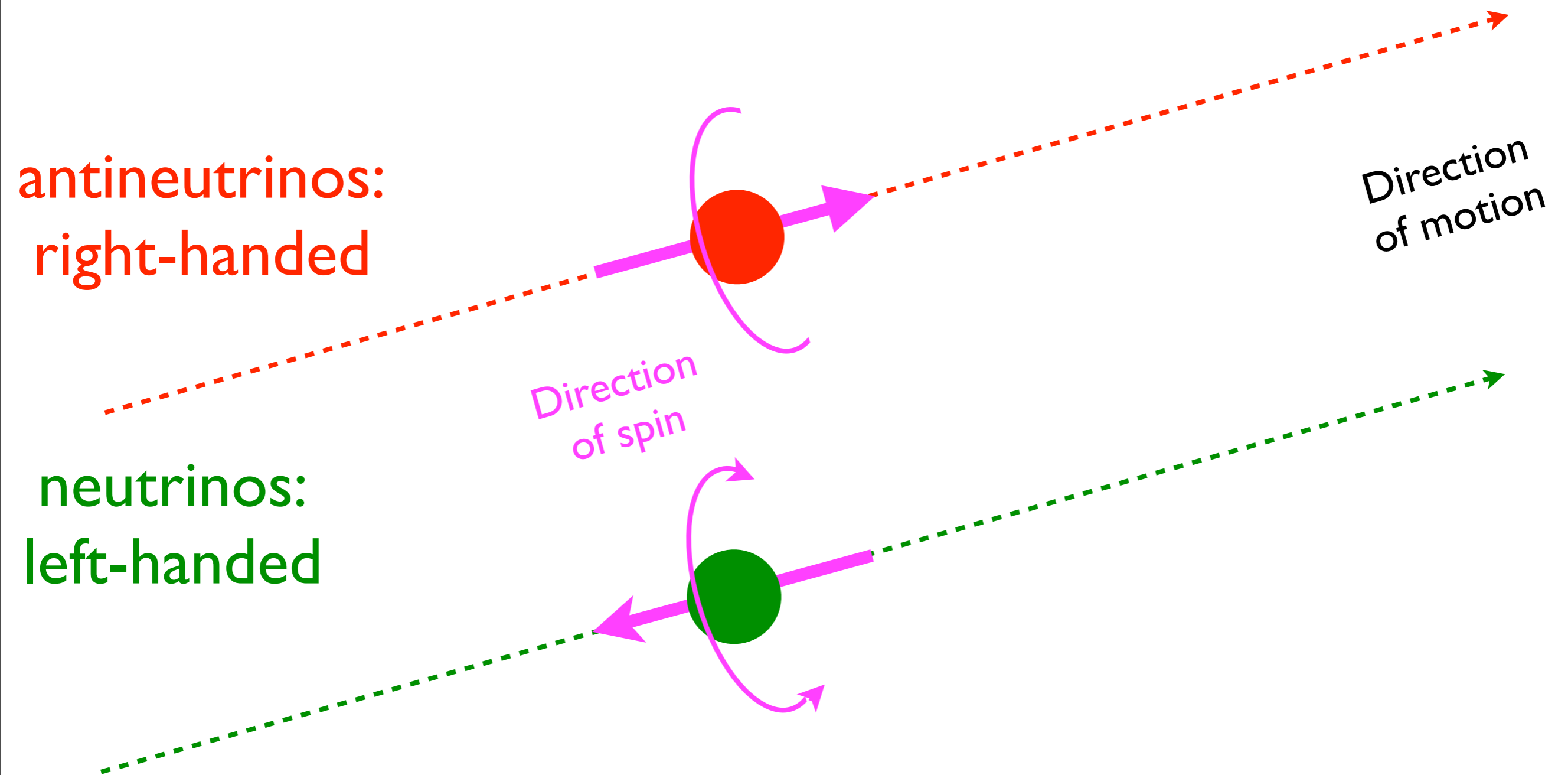
Ray Davis Jr. (1964)

Cleaning fluid (Cl)

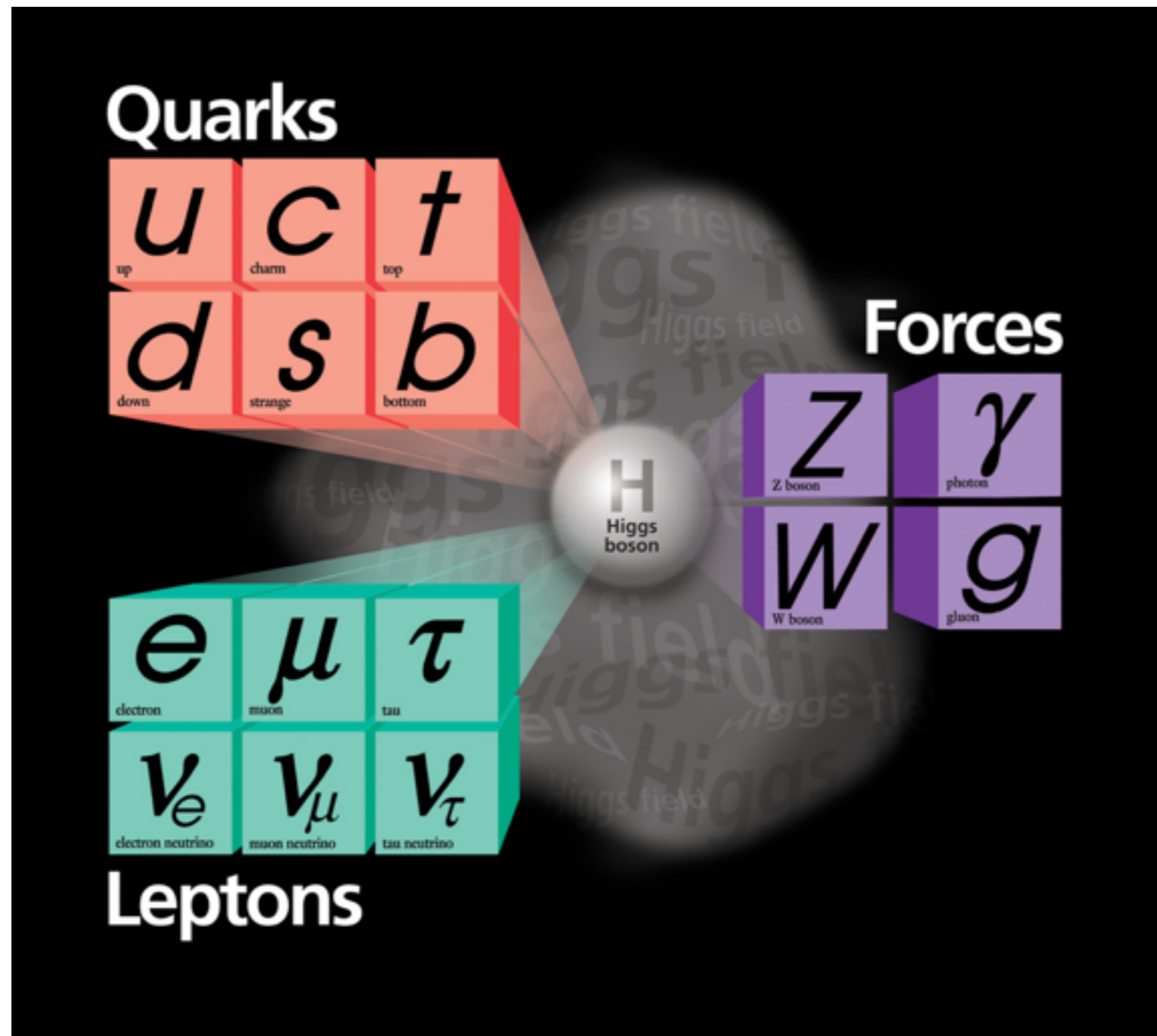


+ Ar detector

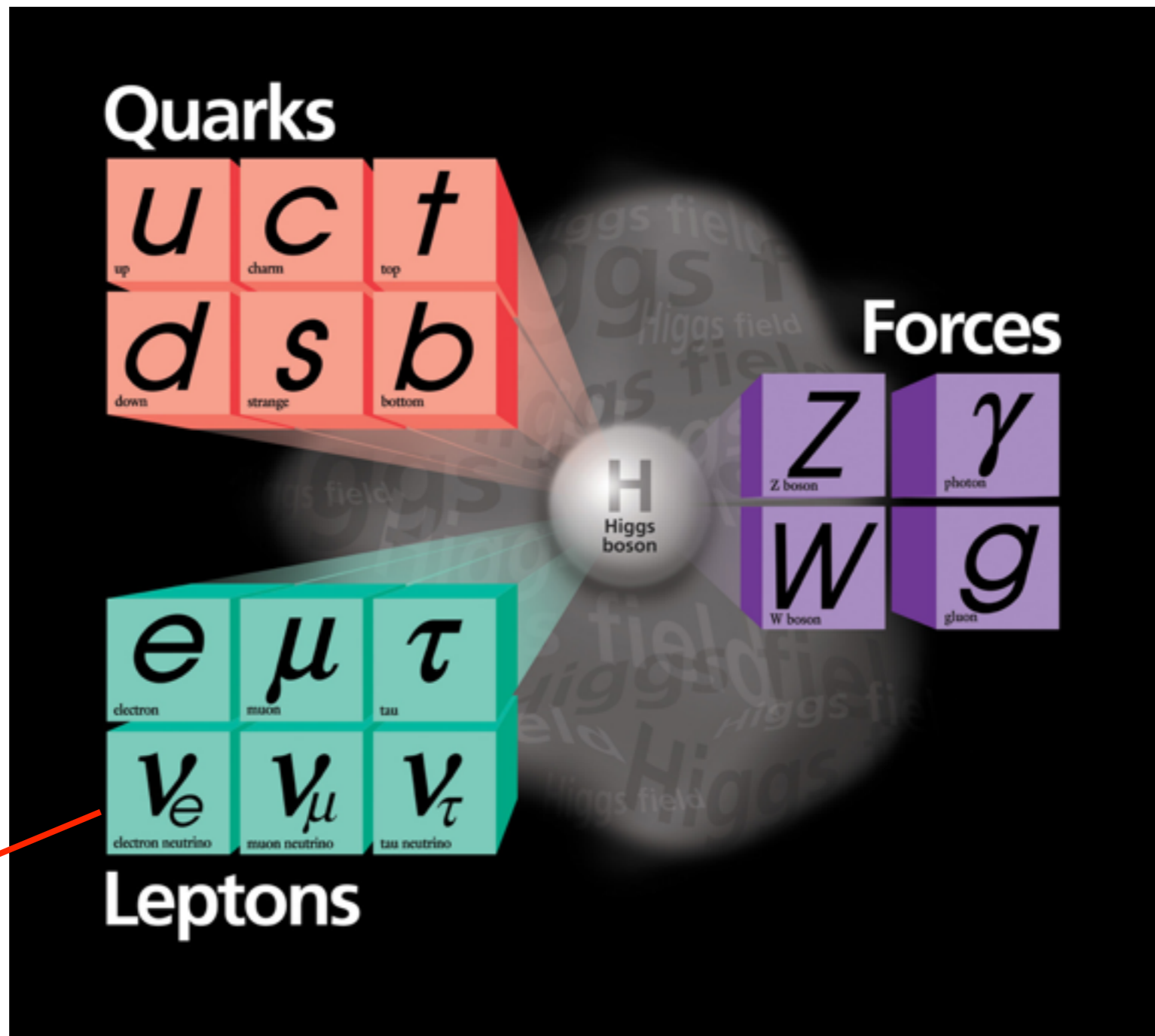
Neutrino Handedness



Standard Model Particles

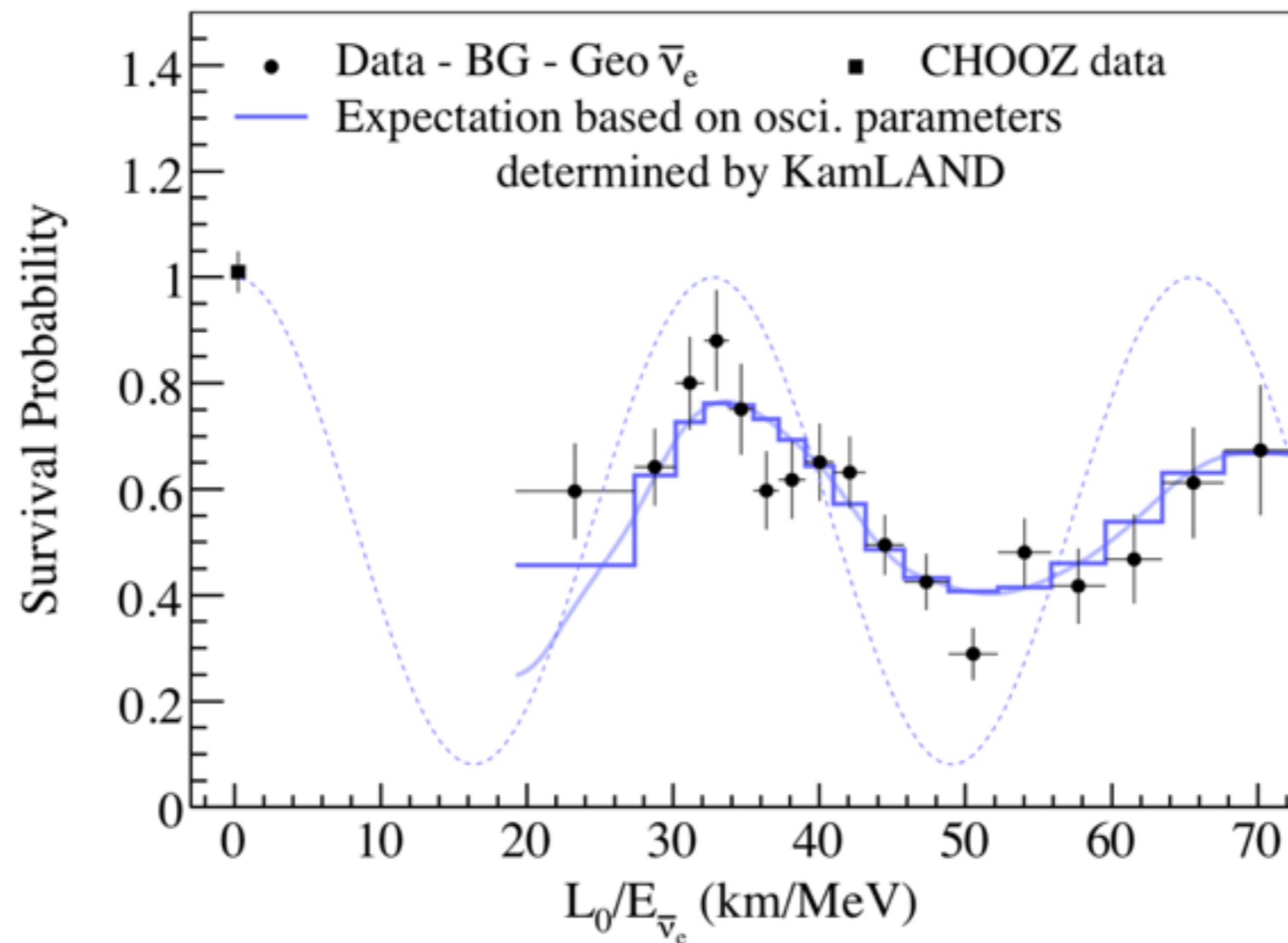
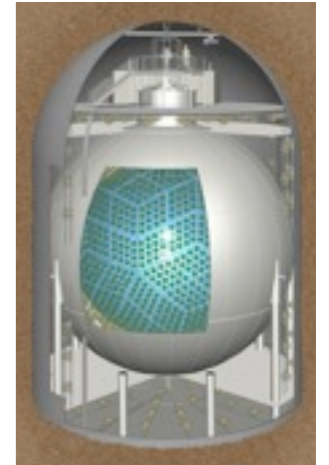
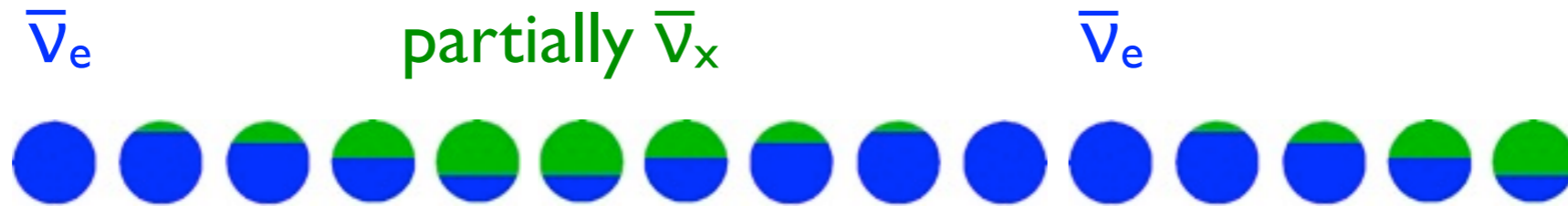


Standard Model Particles



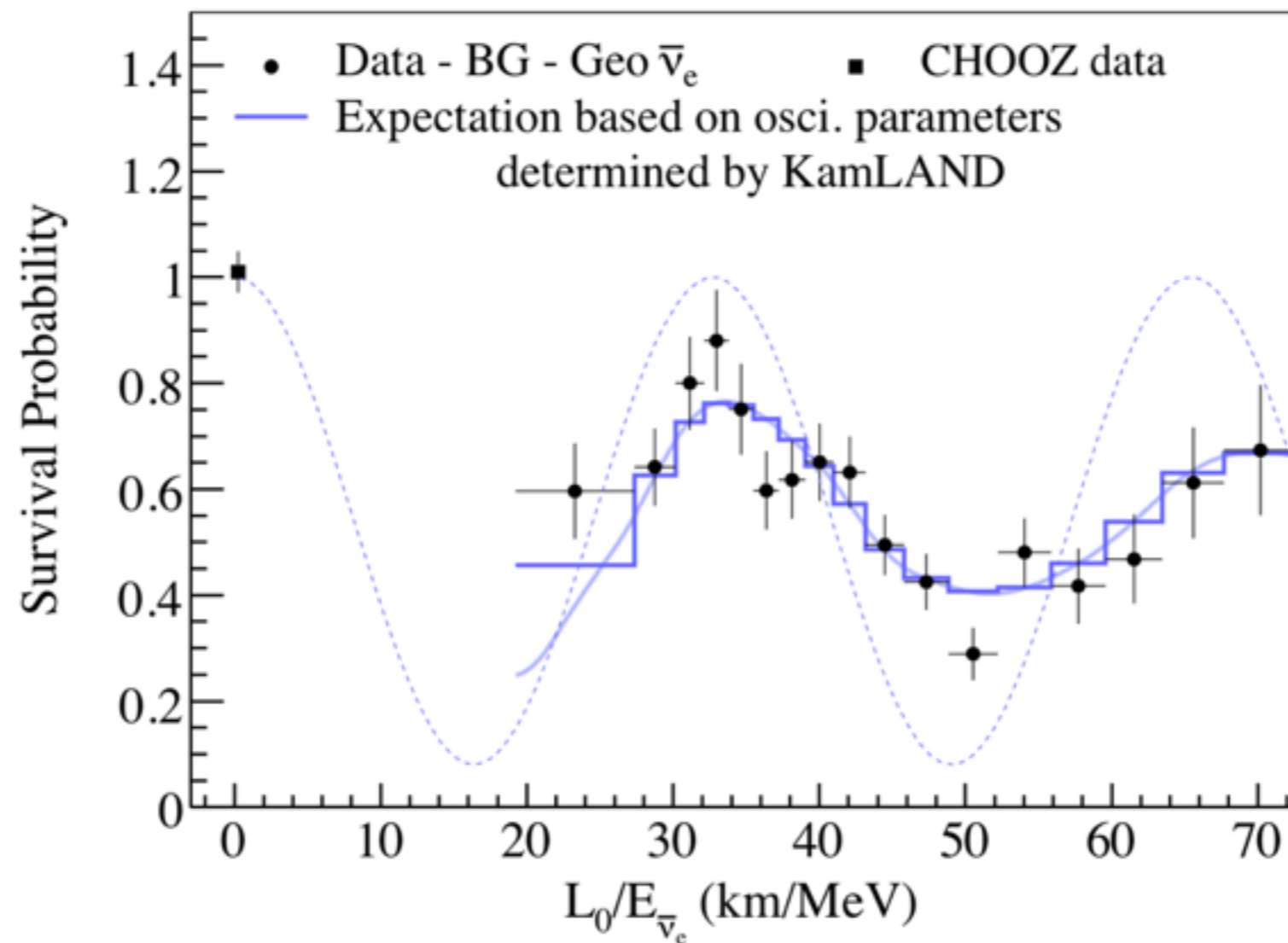
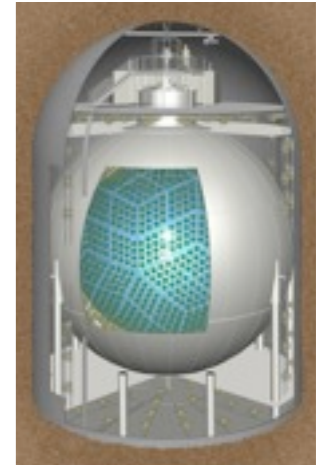
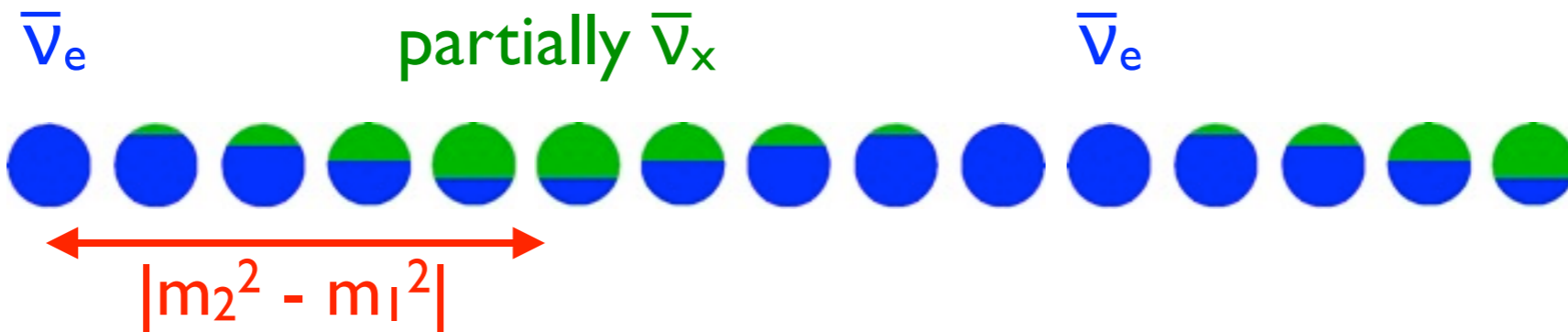
$m = 0$,
left-handed
only

Neutrino Oscillation



Neutrino Oscillation

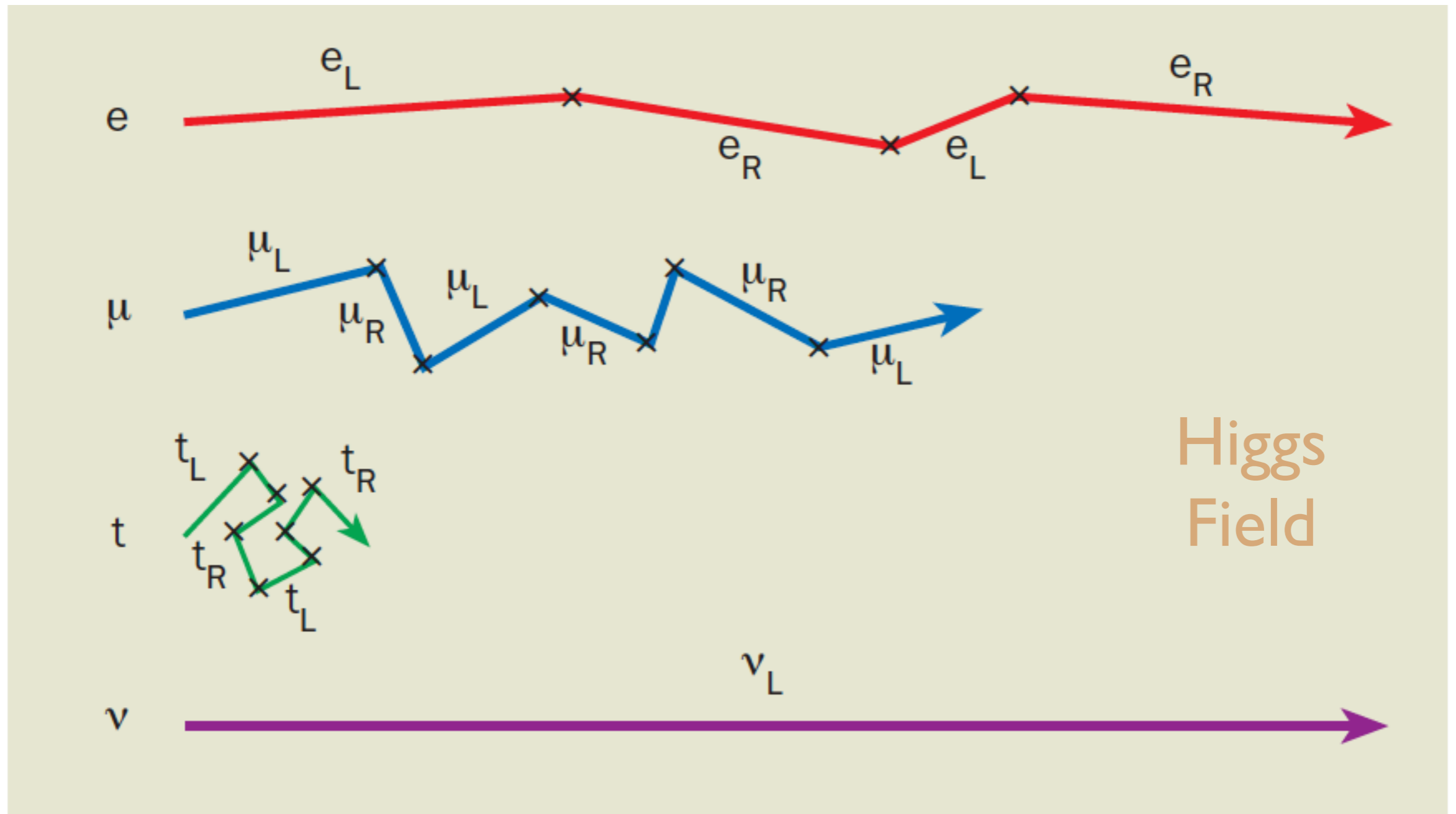
$$\bar{\nu}_e = U_{e1}\bar{\nu}_1 + U_{e2}\bar{\nu}_2 + U_{e3}\bar{\nu}_3$$



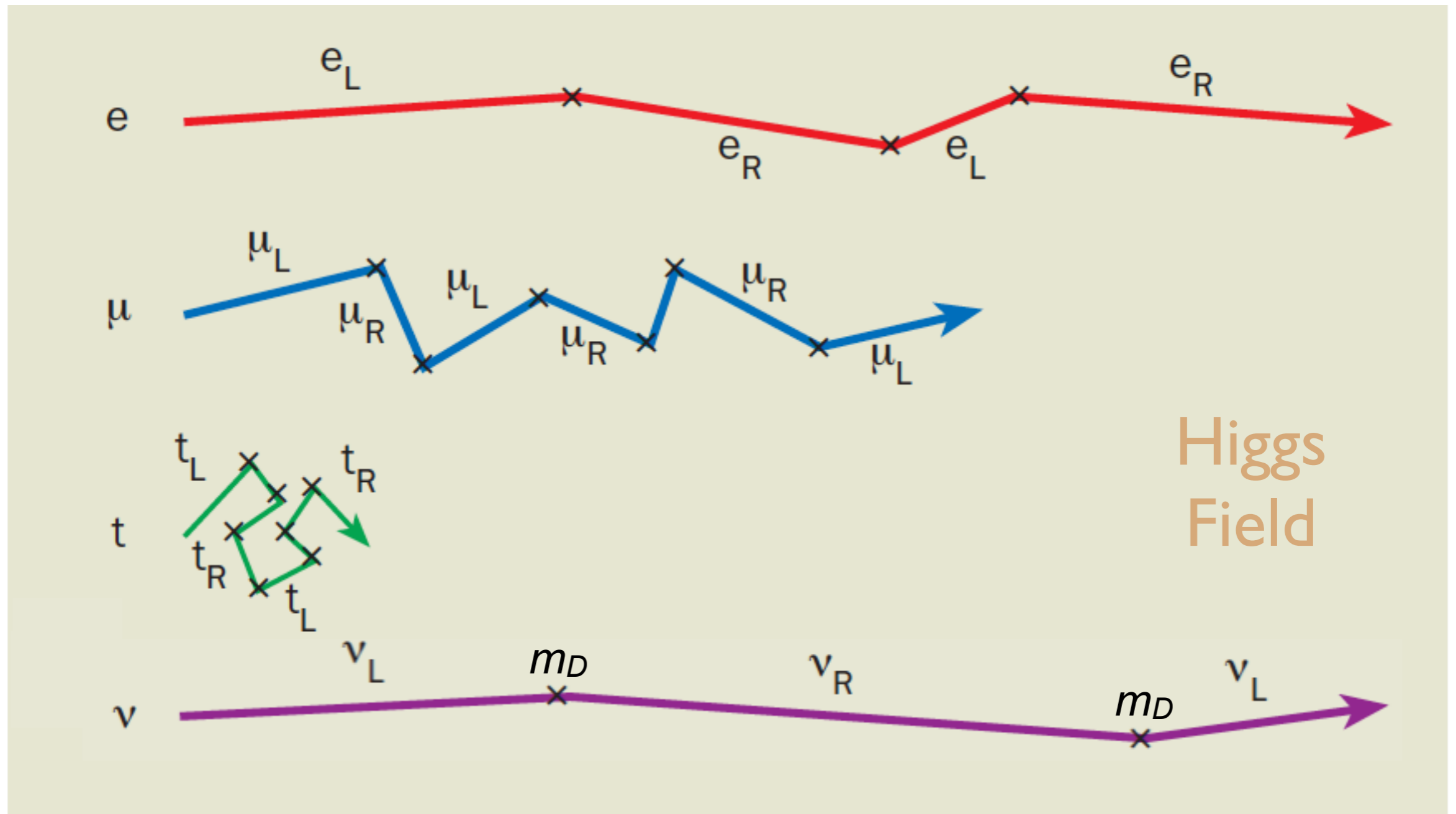
Outline

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- The case for Majorana neutrinos
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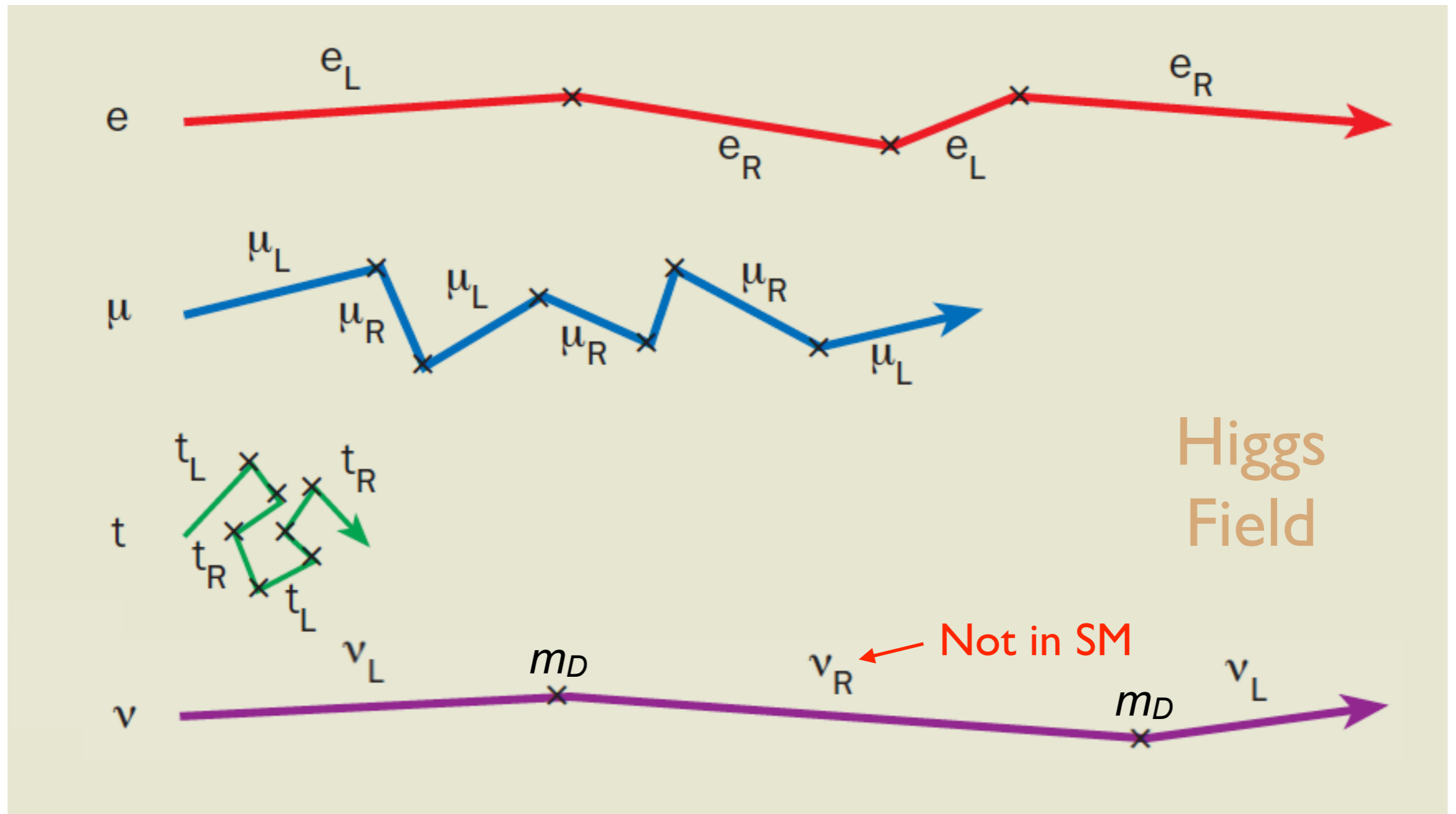
Incorporating ν Mass



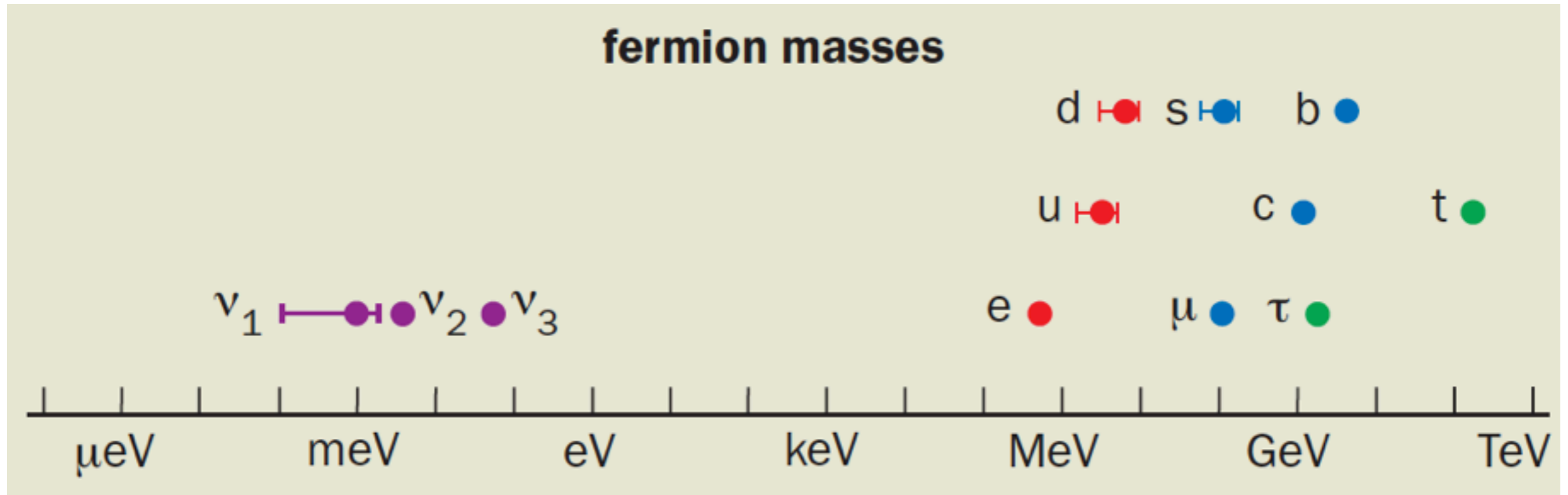
Incorporating ν Mass



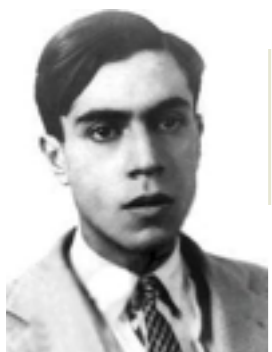
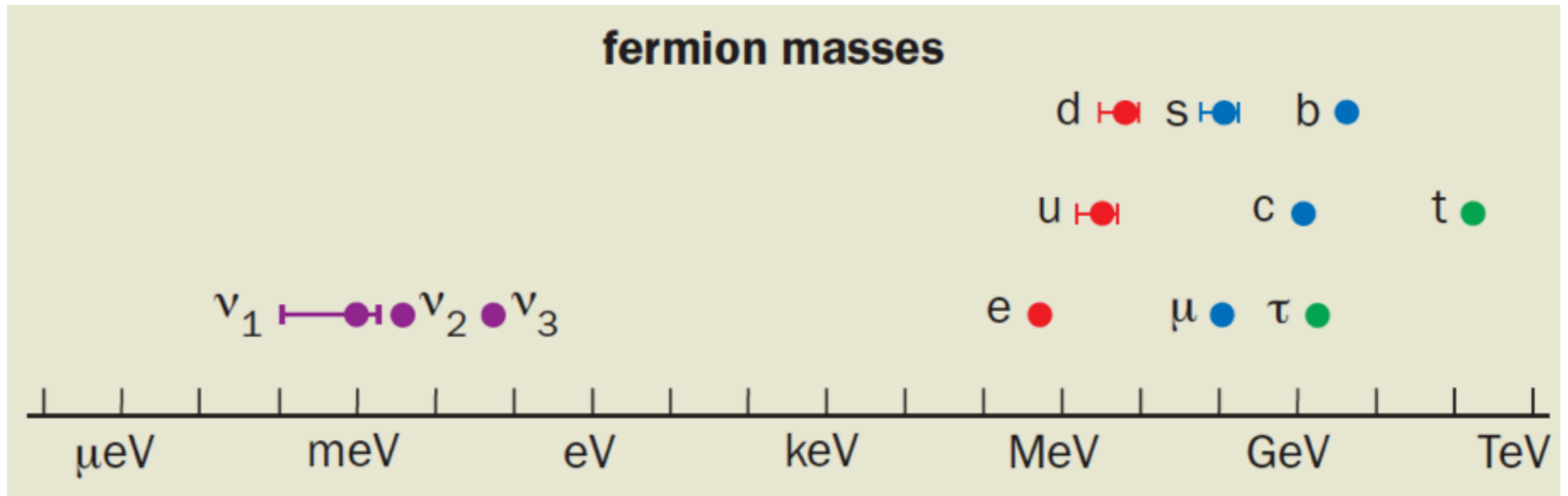
Incorporating ν Mass



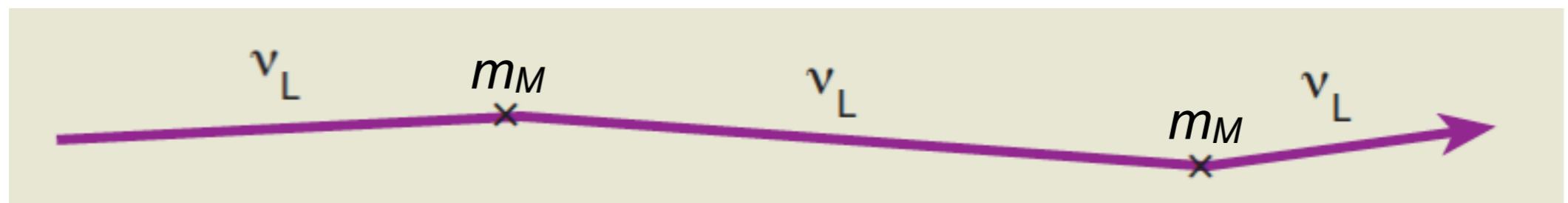
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Incorporating ν Mass



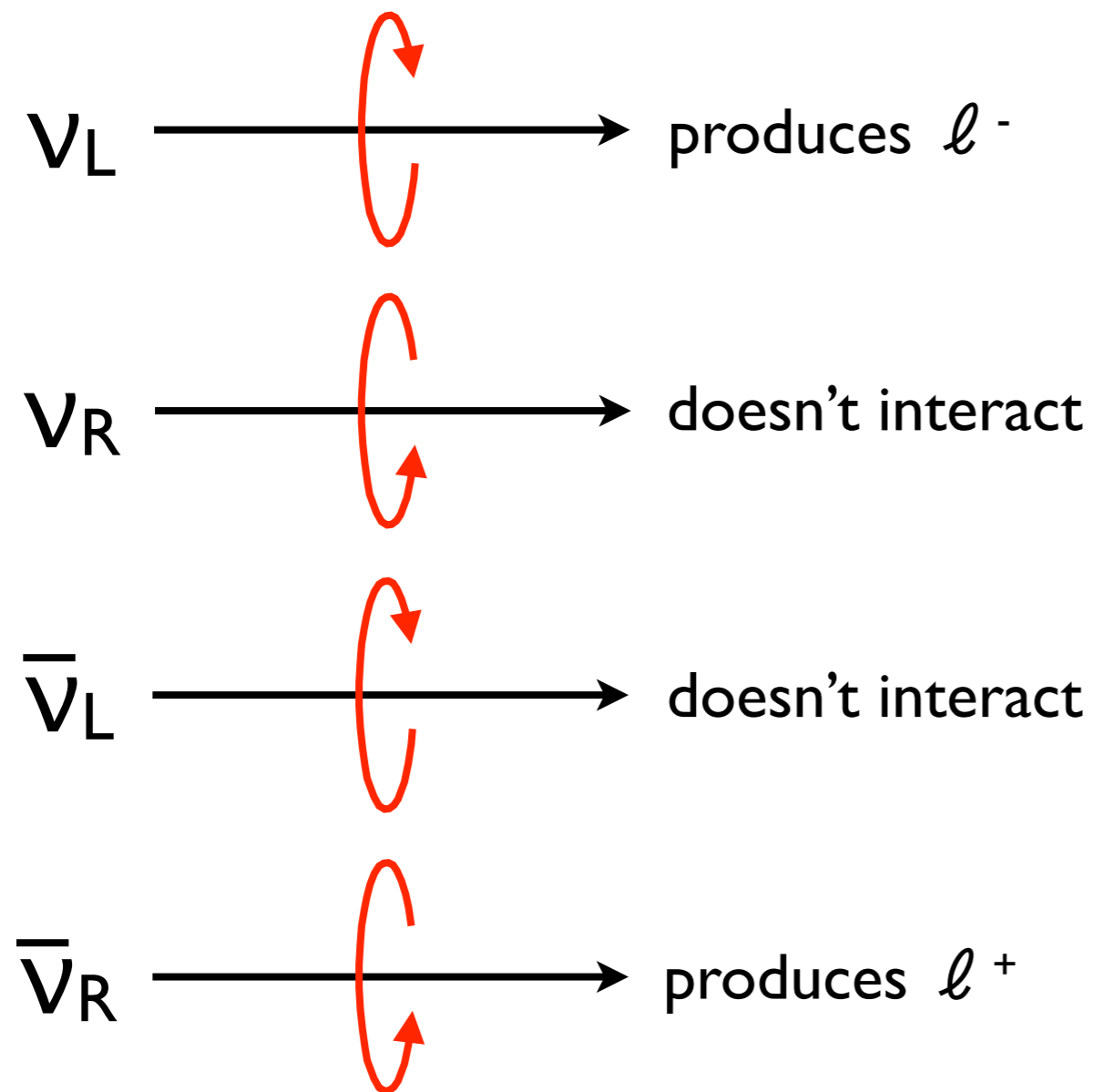
Ettore Majorana



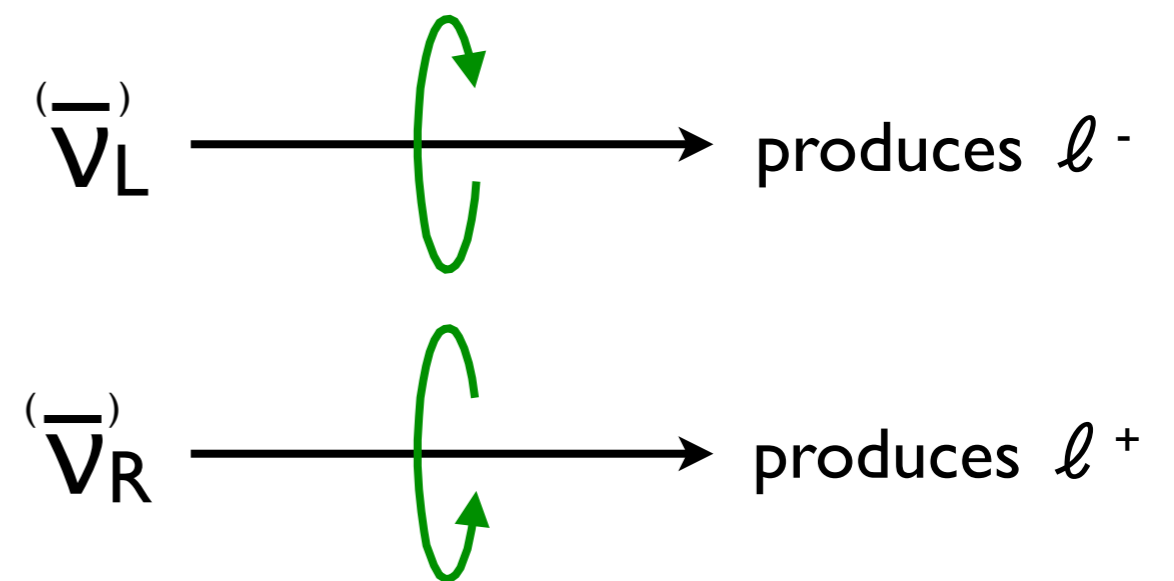
E. Majorana, *Il Nuovo Cimento* **14**, 171 (1937).
 English translation: *Soryushiron Kenkyu* **63**, 149 (1981).

Dirac vs Majorana neutrinos

Dirac ($\nu \neq \bar{\nu}$)

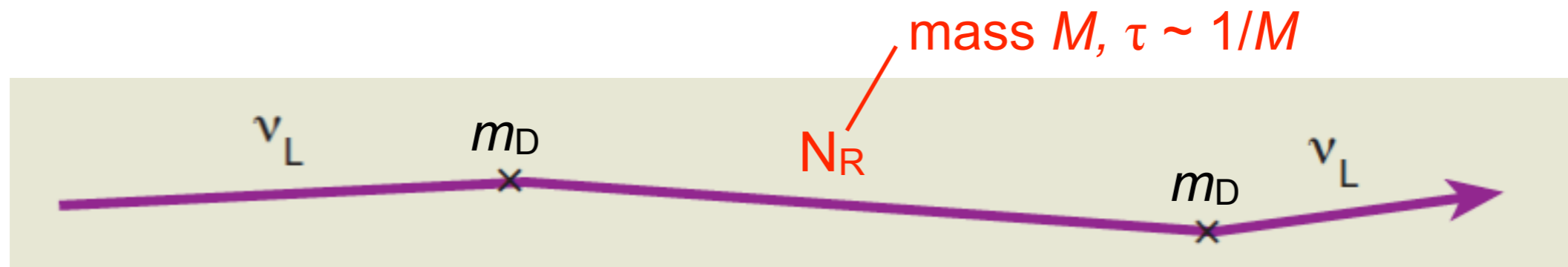


Majorana ($\nu = \bar{\nu}$)

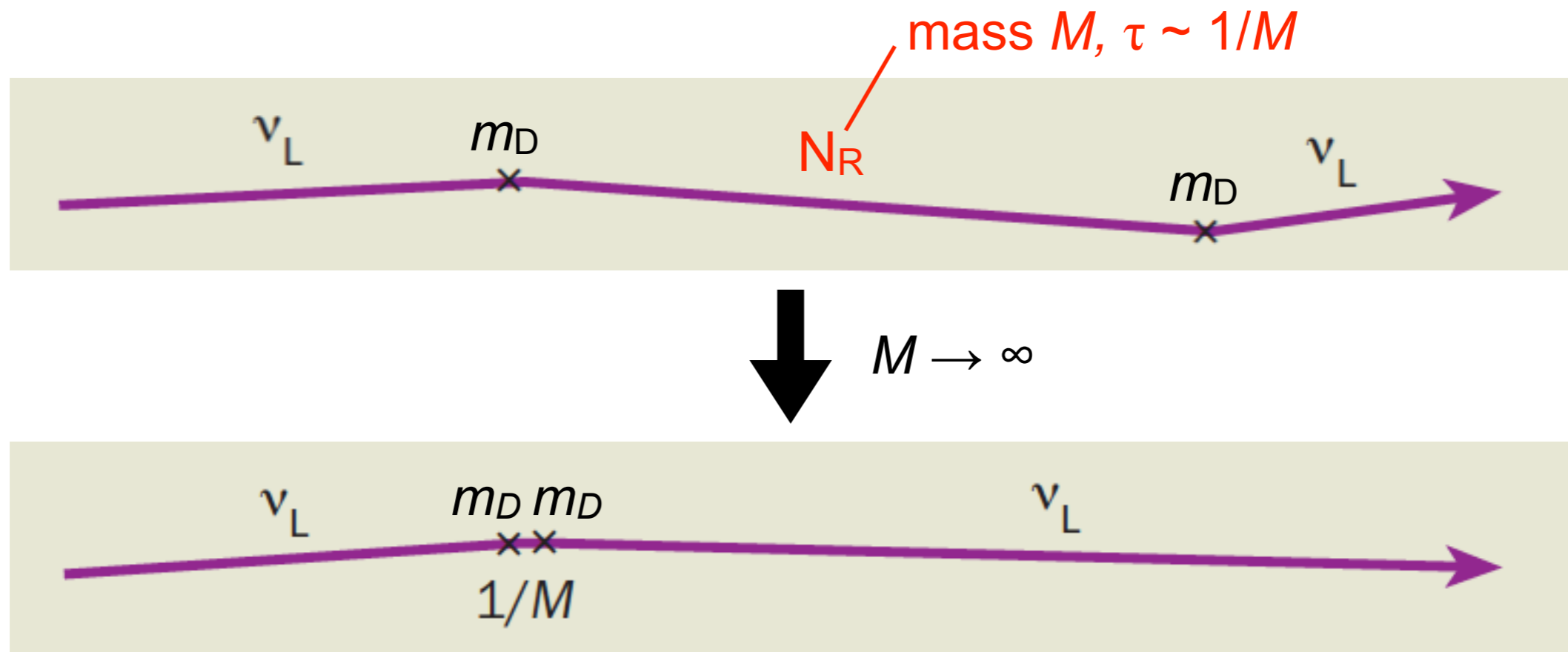


- Requires $Q = -Q = 0$
- Implies L is not conserved

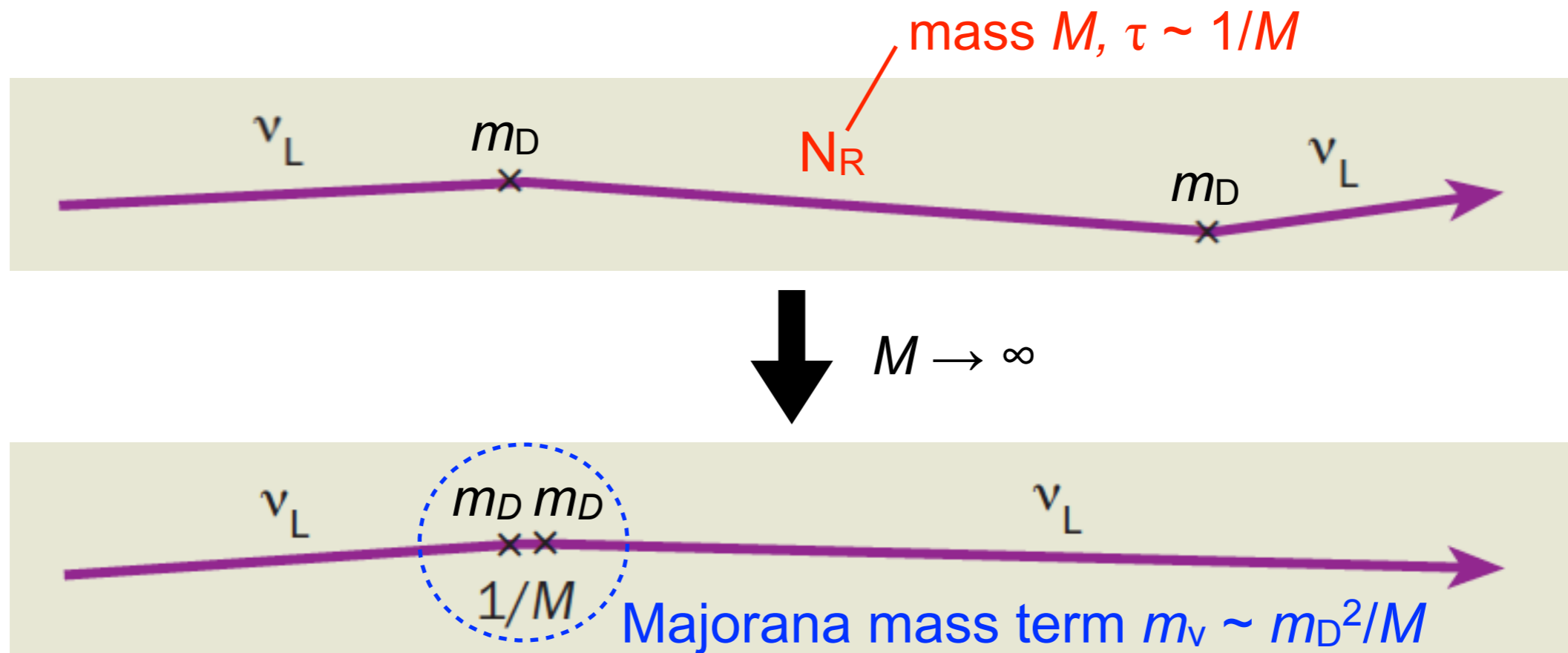
Seesaw Mechanism



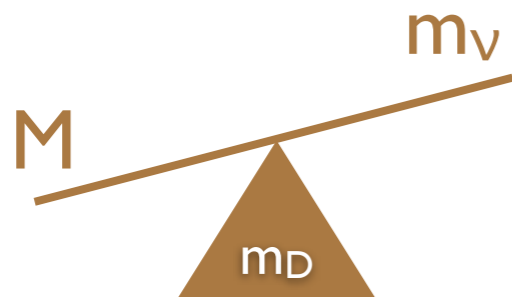
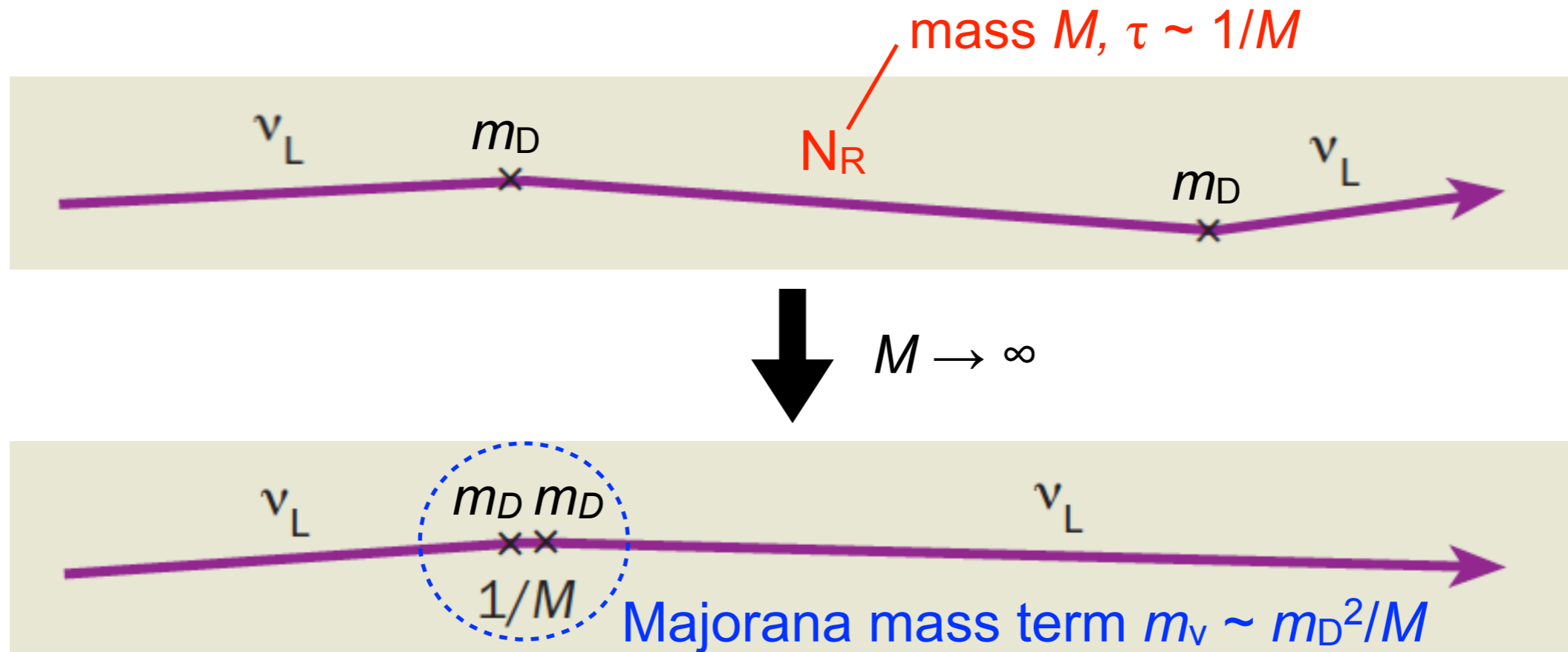
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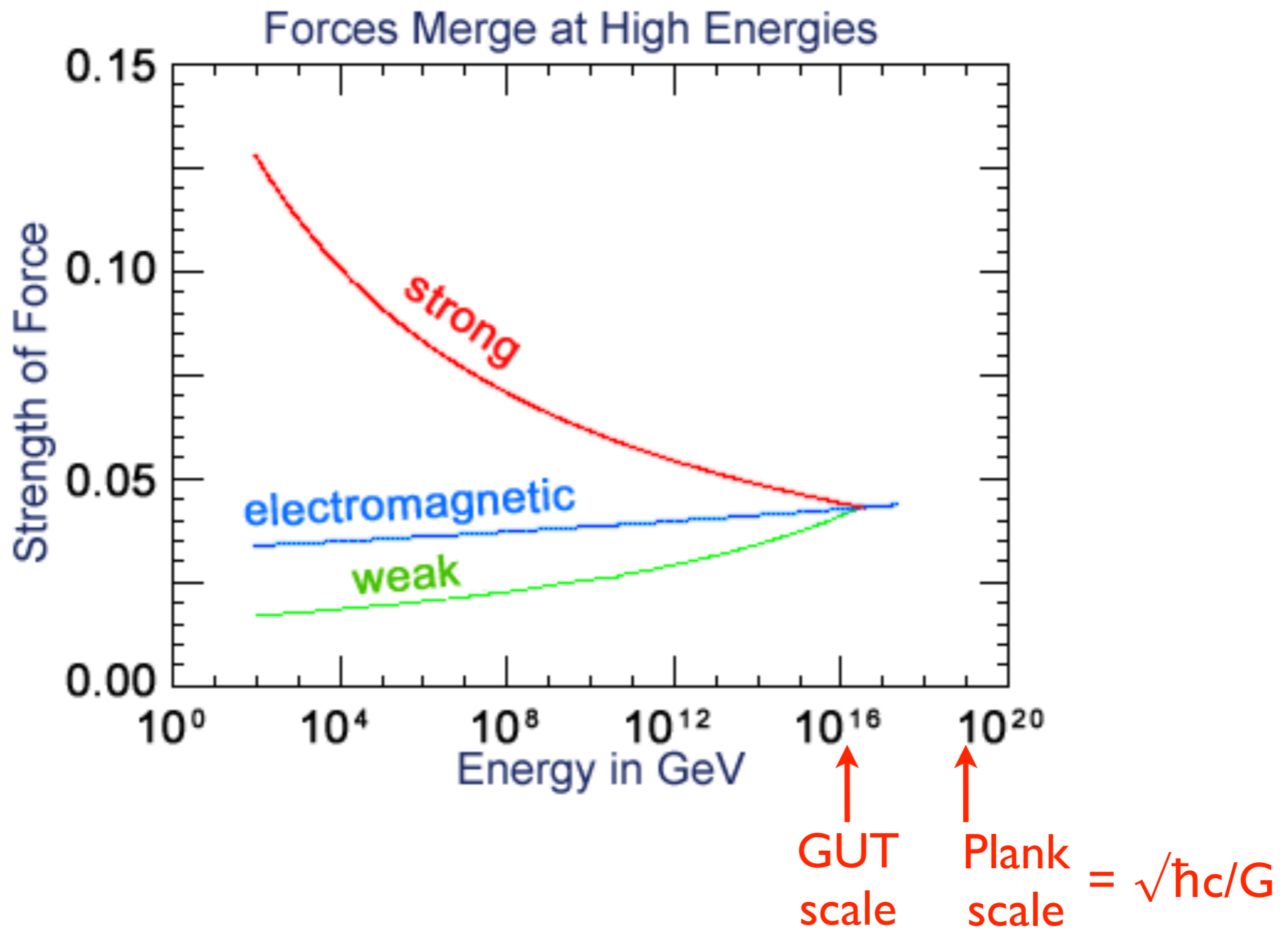


Seesaw Mechanism



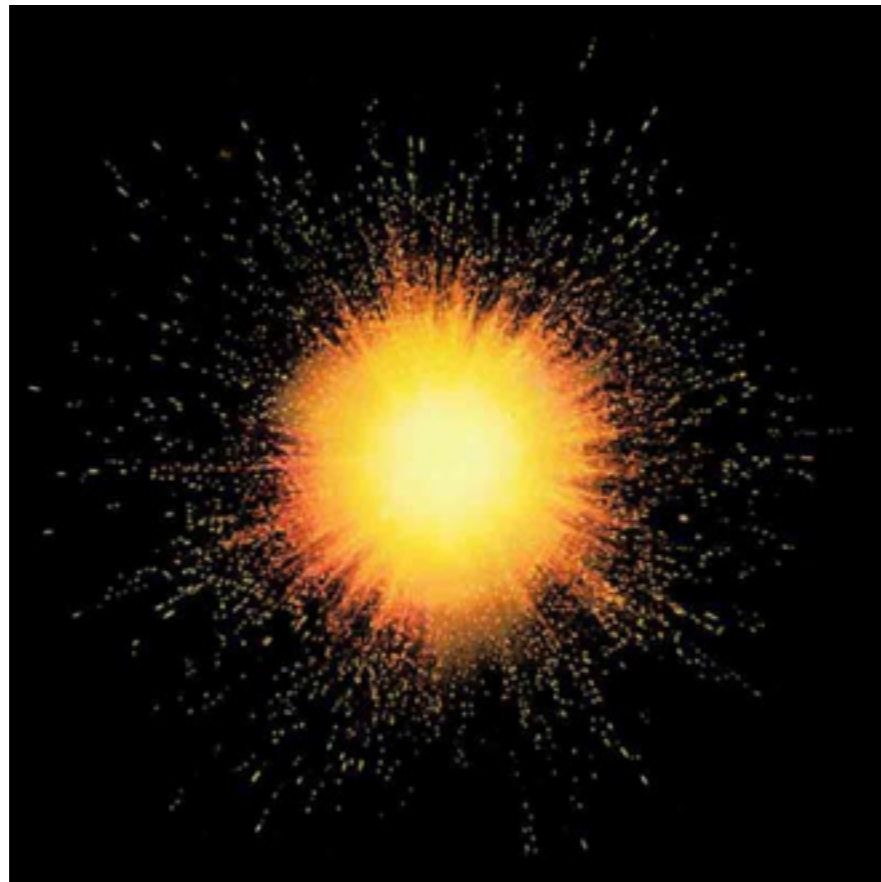
for $m_D \sim \text{GeV-TeV}$:
 $m_\nu \sim \text{meV-eV} \leftrightarrow M \sim 10^{16}-10^{19} \text{ GeV}$

Grand Unification



Matter-Antimatter Asymmetry

The Big Bang



matter + antimatter

The Universe Today



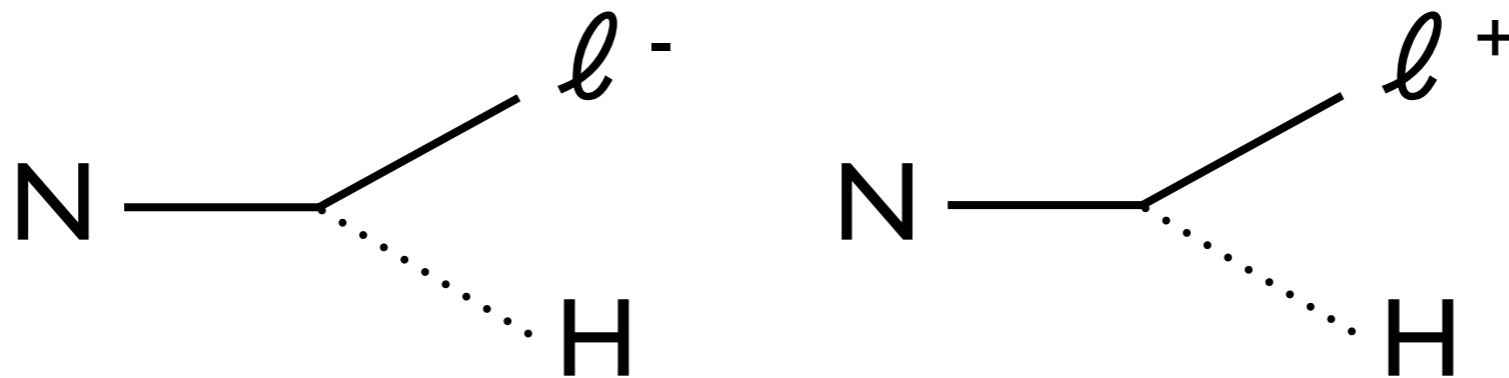
matter only

Sakharov Conditions

- Interactions out of thermal equilibrium
- Baryon number violation (baryogenesis)
- C (charge) and CP (charge-parity) violation

Leptogenesis

- Decay of heavy Majorana neutrino (N) into SM leptons (ℓ^\pm) and Higgs (H):

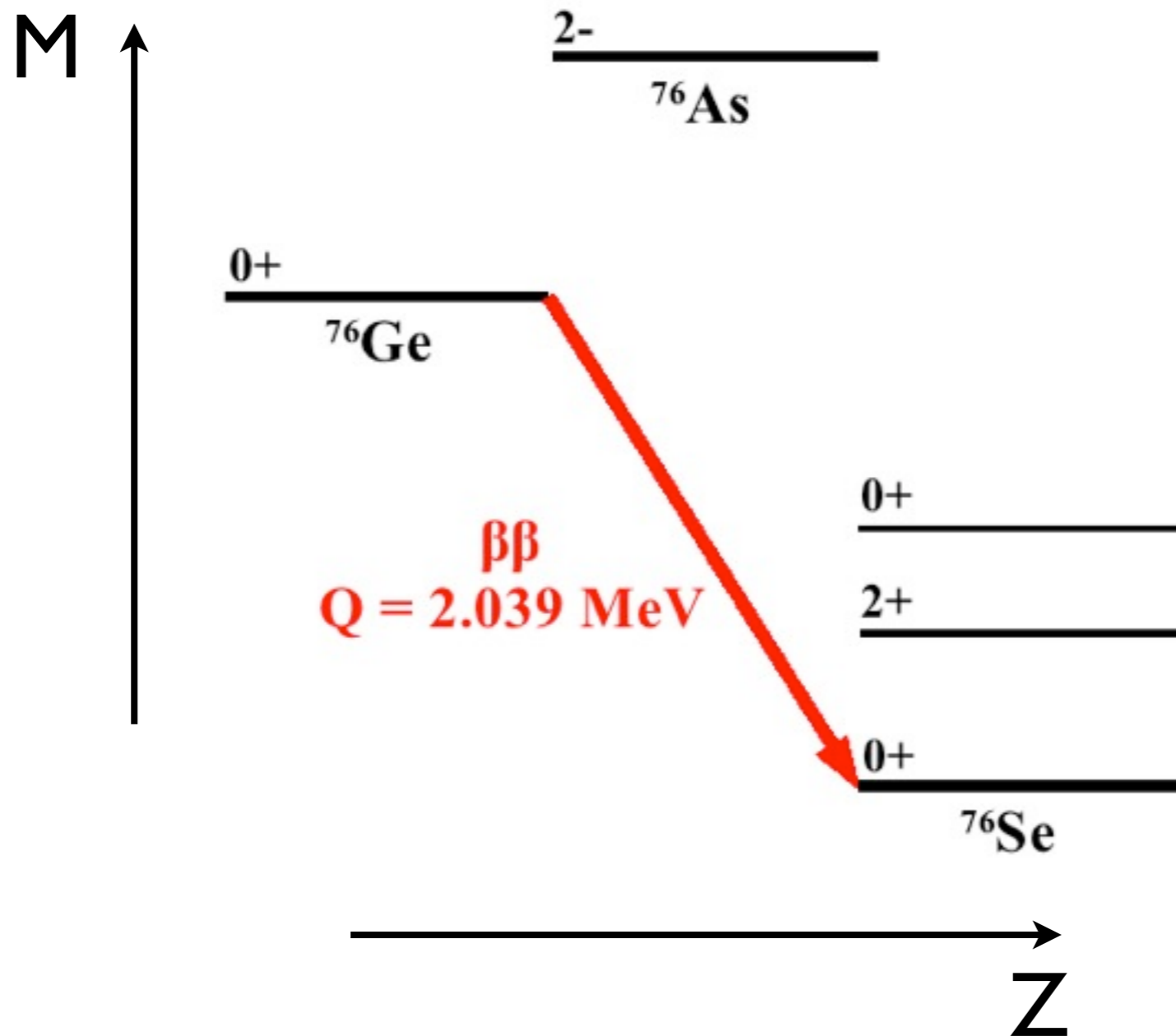


- CP violation in ν sector could give these different branching ratios
- SM processes could convert L to B: baryogenesis!
- Majorana neutrinos could be the reason we exist at all!

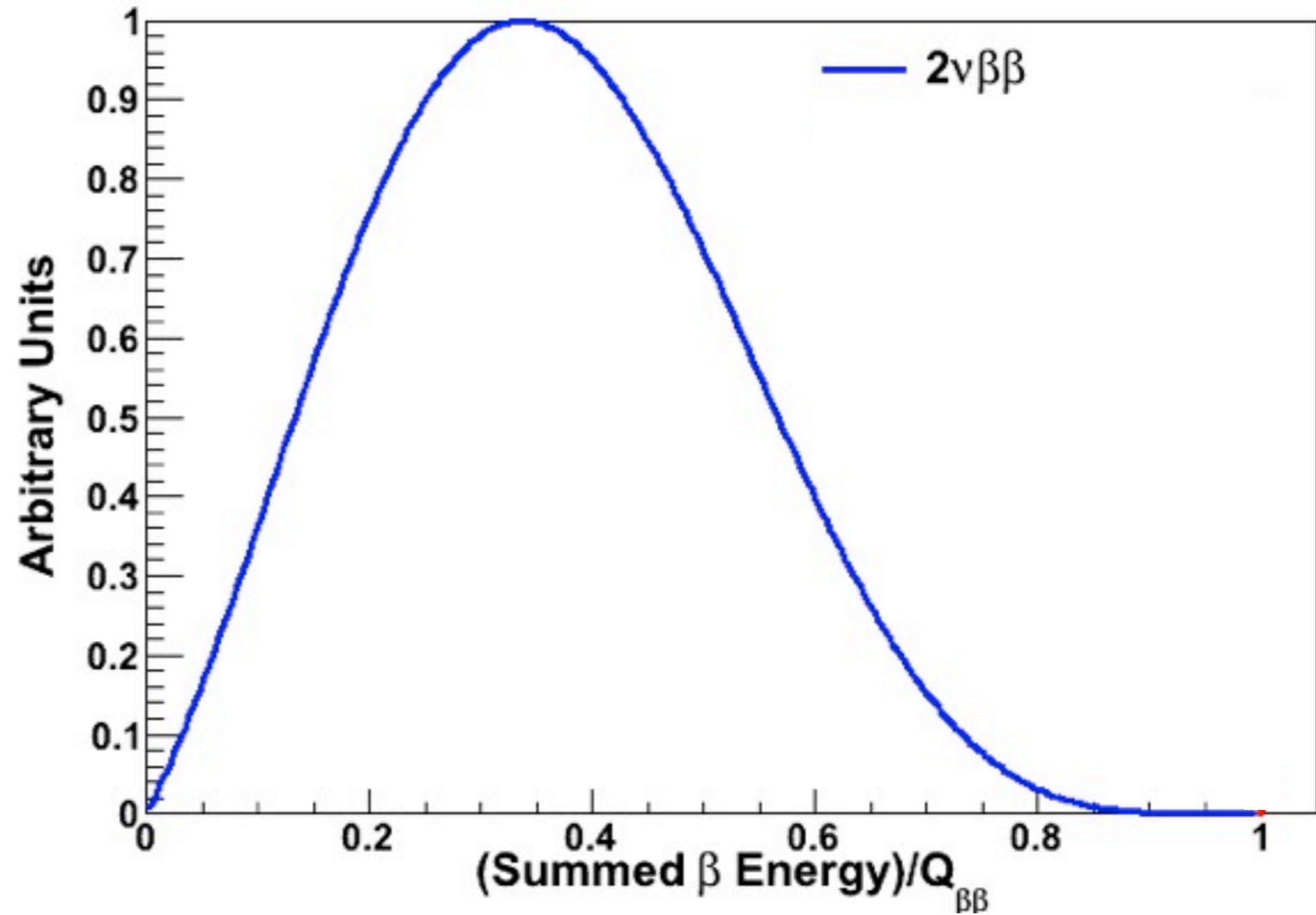
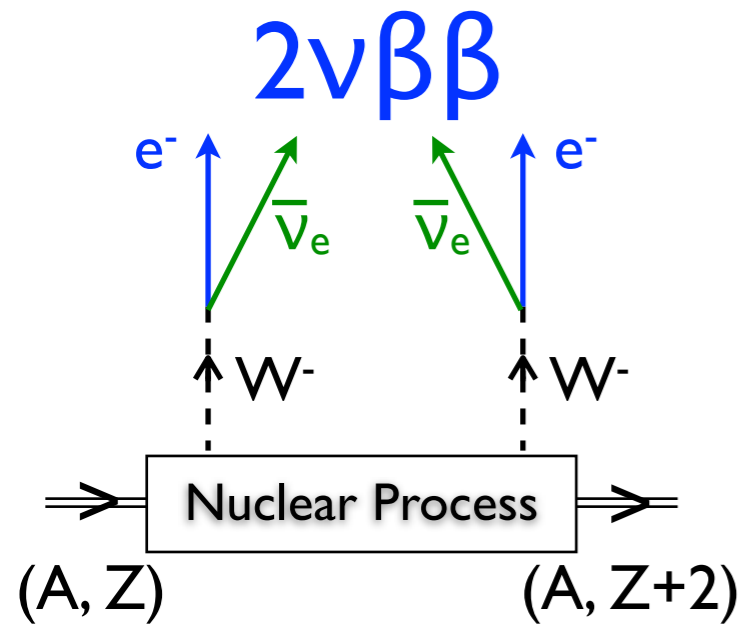
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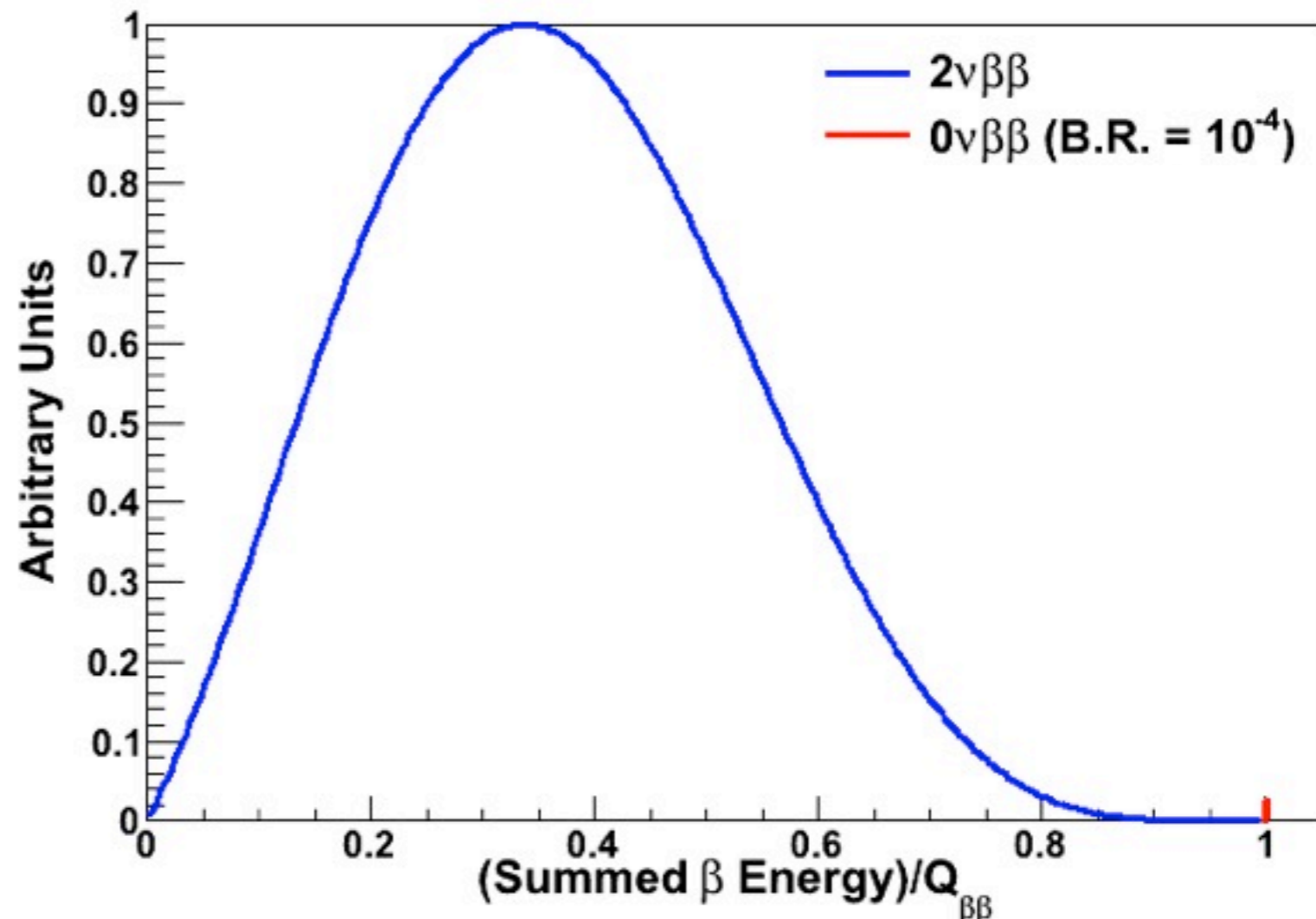
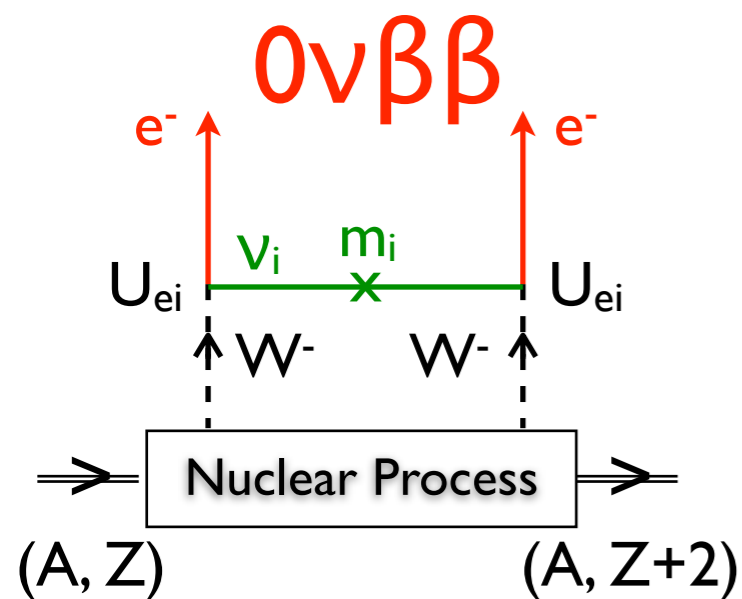
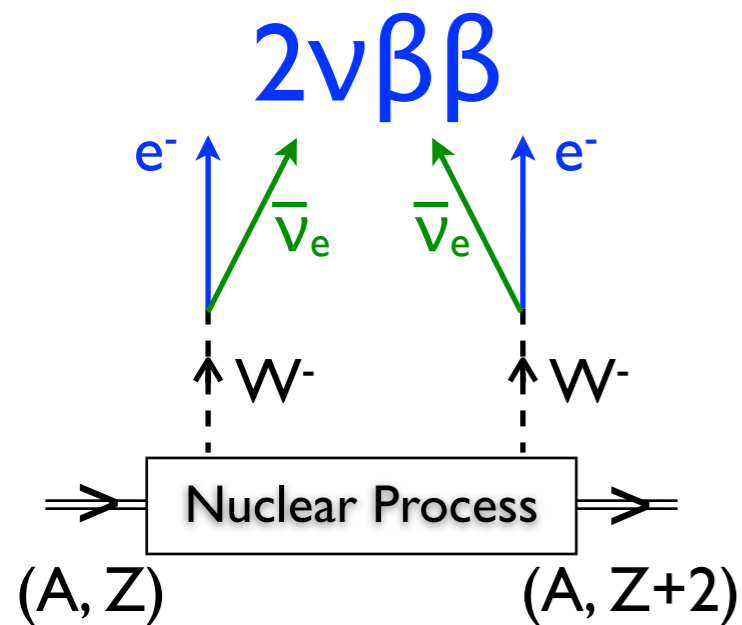
Double-Beta Decay



Double-Beta Decay



$0\nu\beta\beta$ Decay

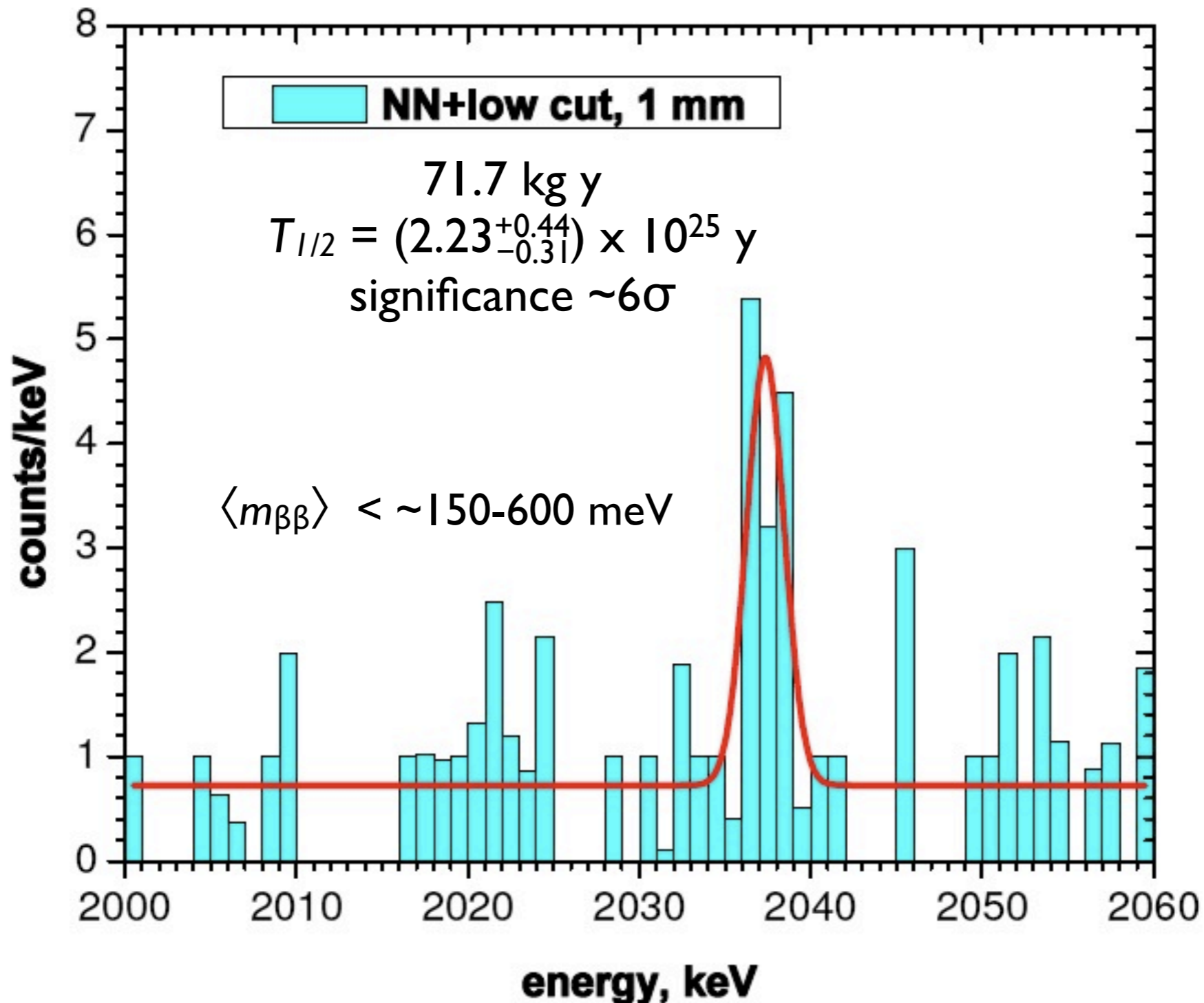


$$\Gamma_{1/2}^{0\nu} = G^{0\nu} |M^{0\nu}|^2 \langle m_{\beta\beta} \rangle^2$$

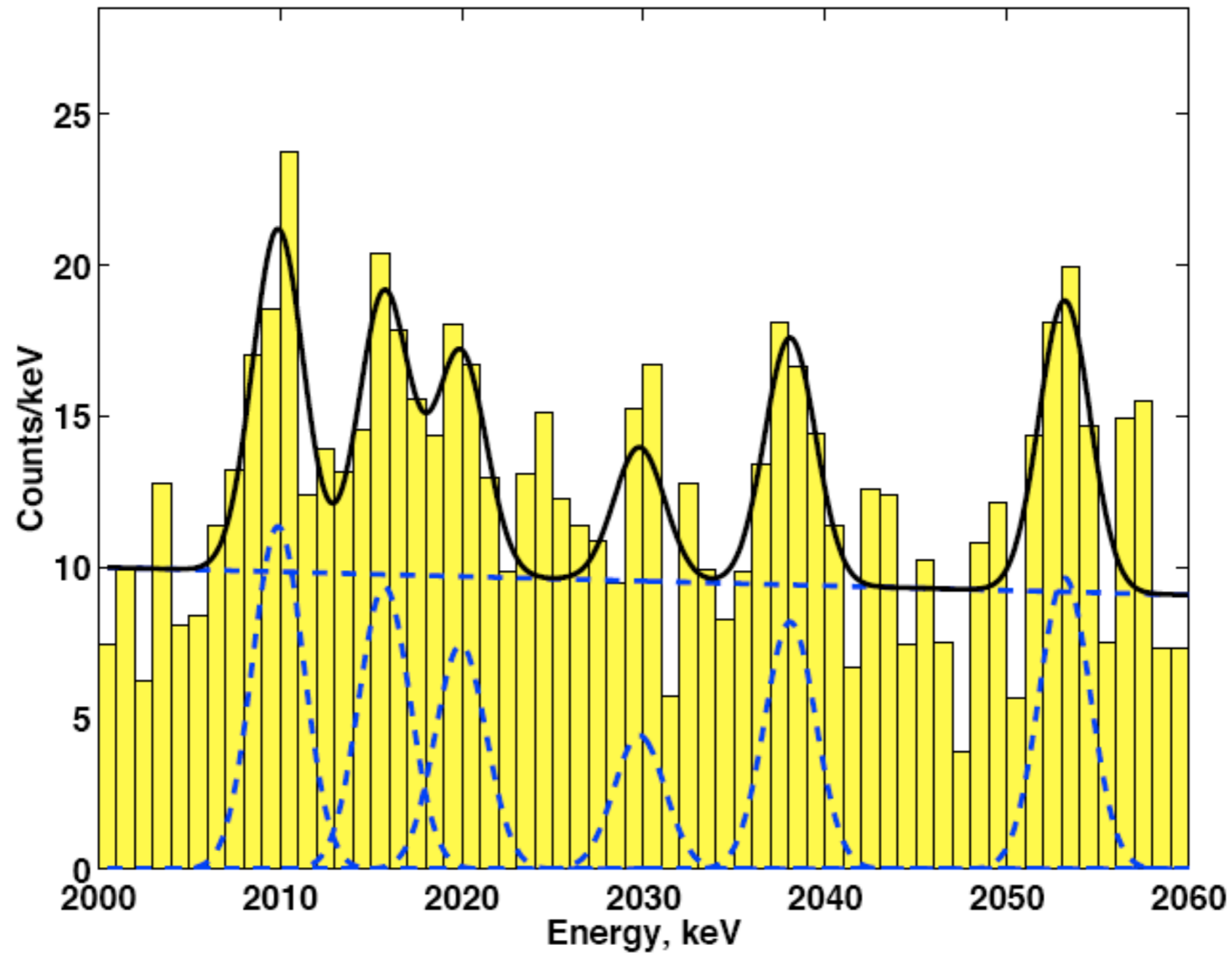
$$\langle m_{\beta\beta} \rangle \equiv \left| \sum m_i U_{ei}^2 \right|$$

Claimed Observation

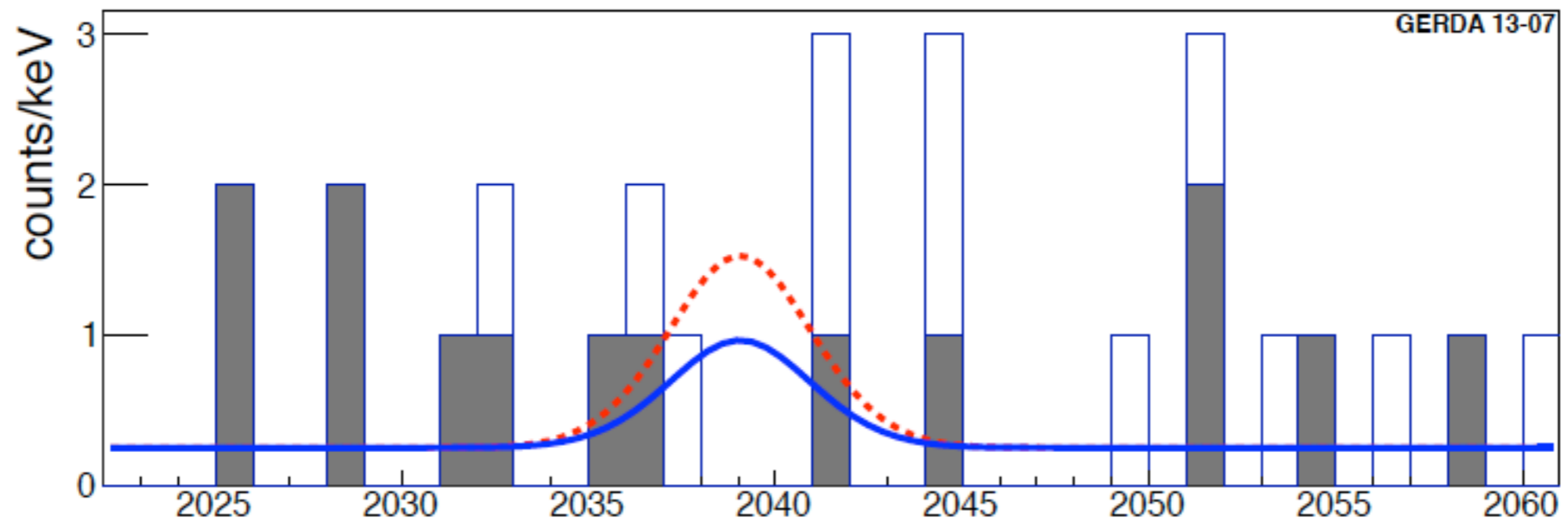
Klapdor Kleingrothaus *et al.*, Mod. Phys. Lett.A **21** (2006) p 1547.



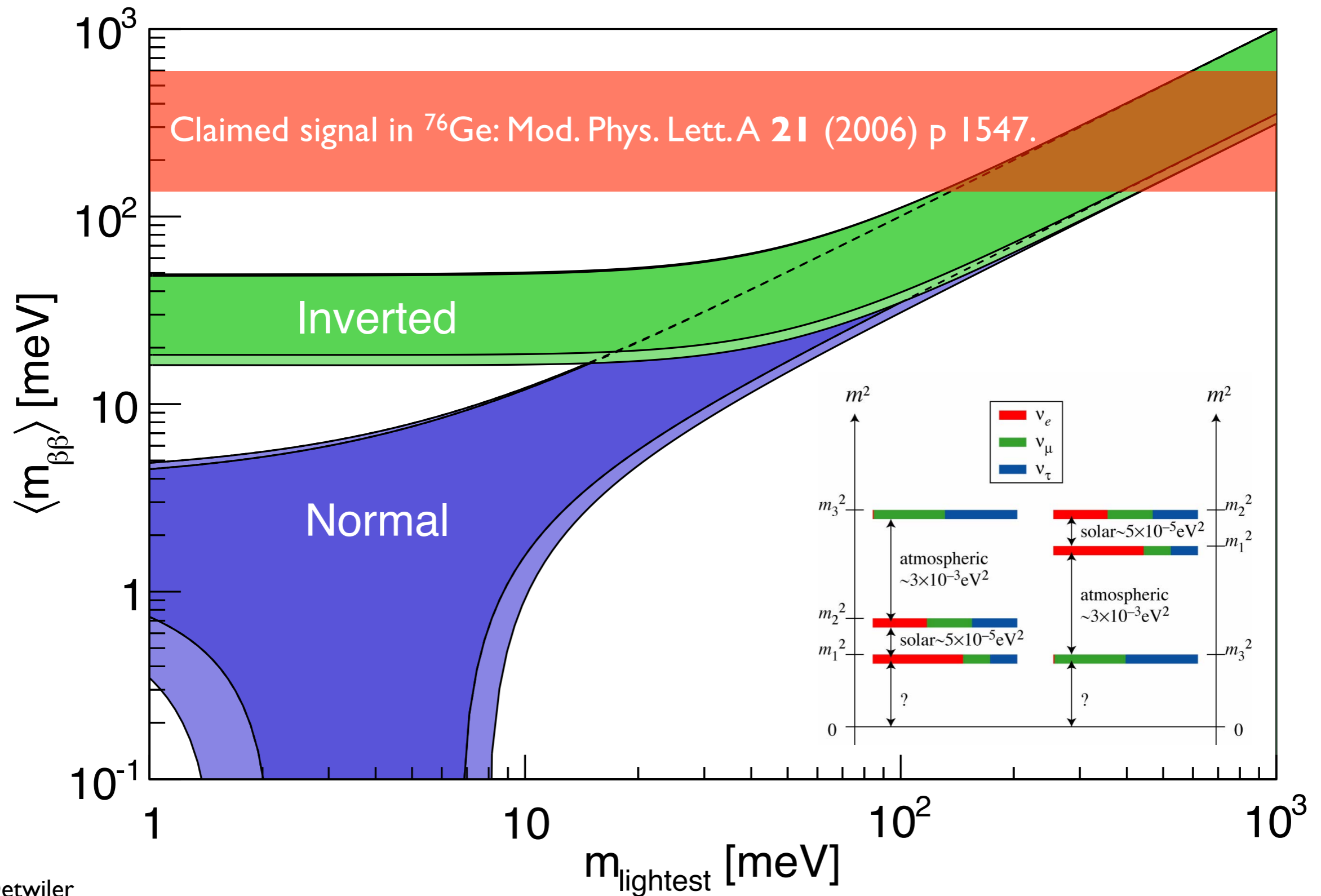
Claimed Observation



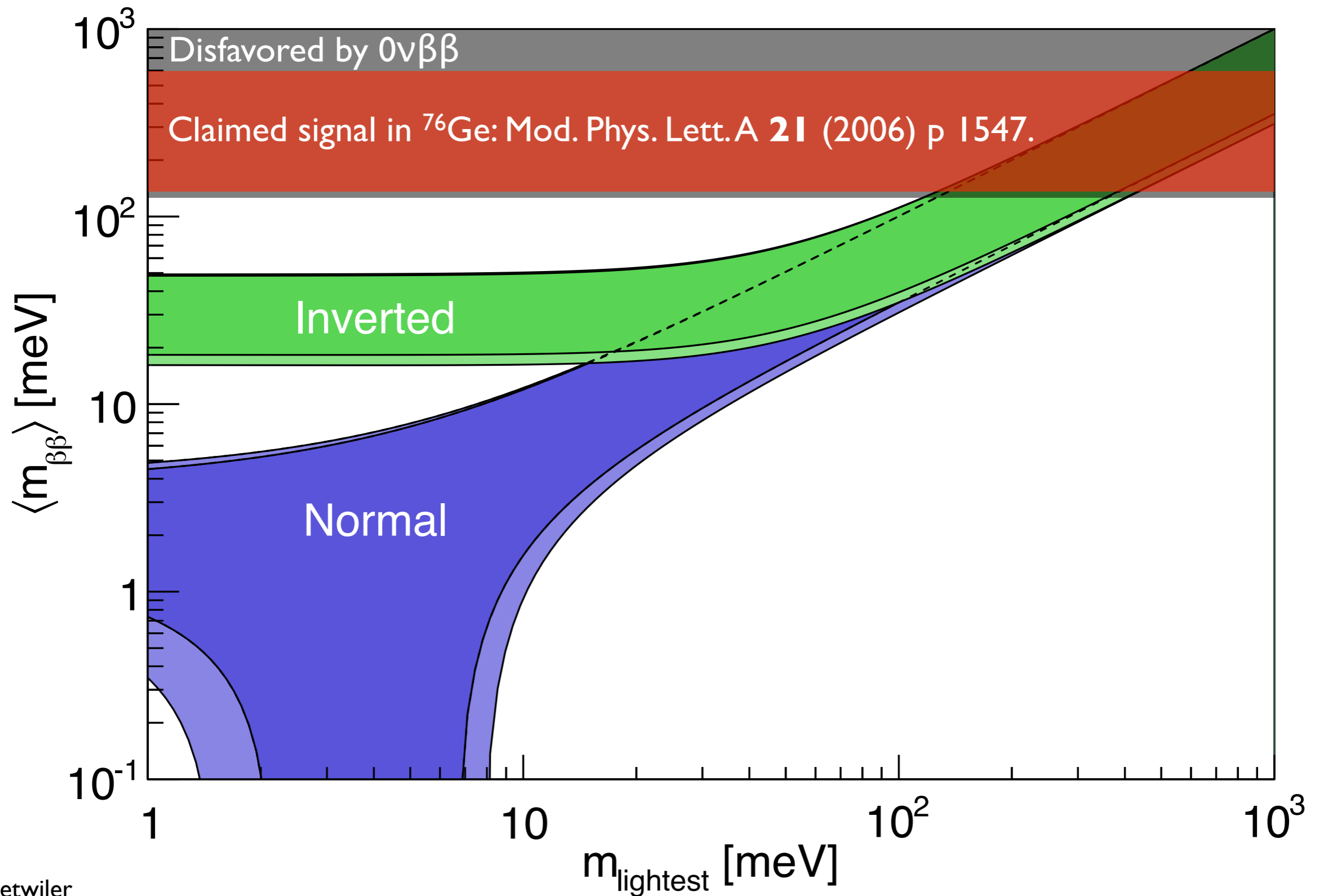
New GERDA Result



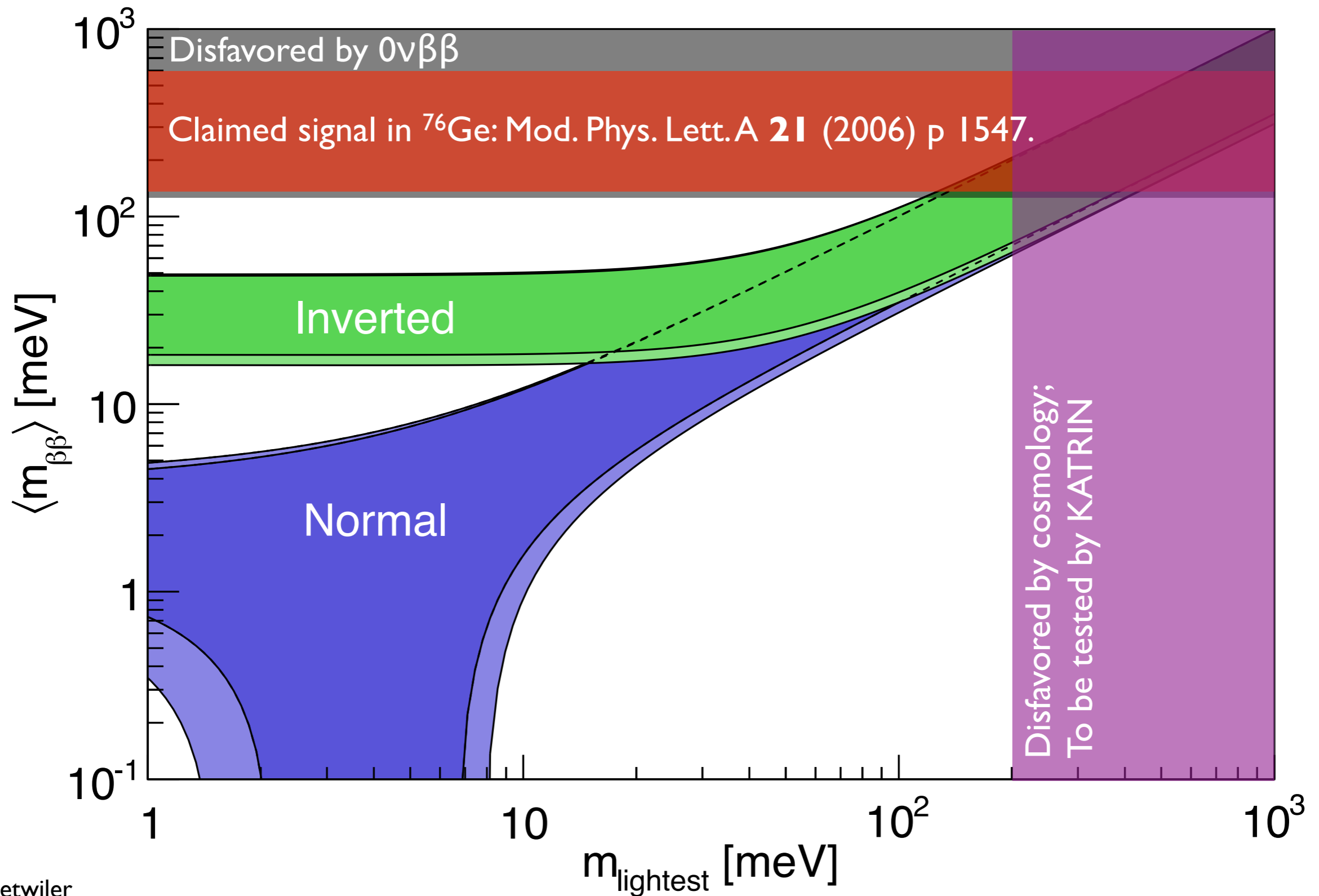
Double-Beta Decay



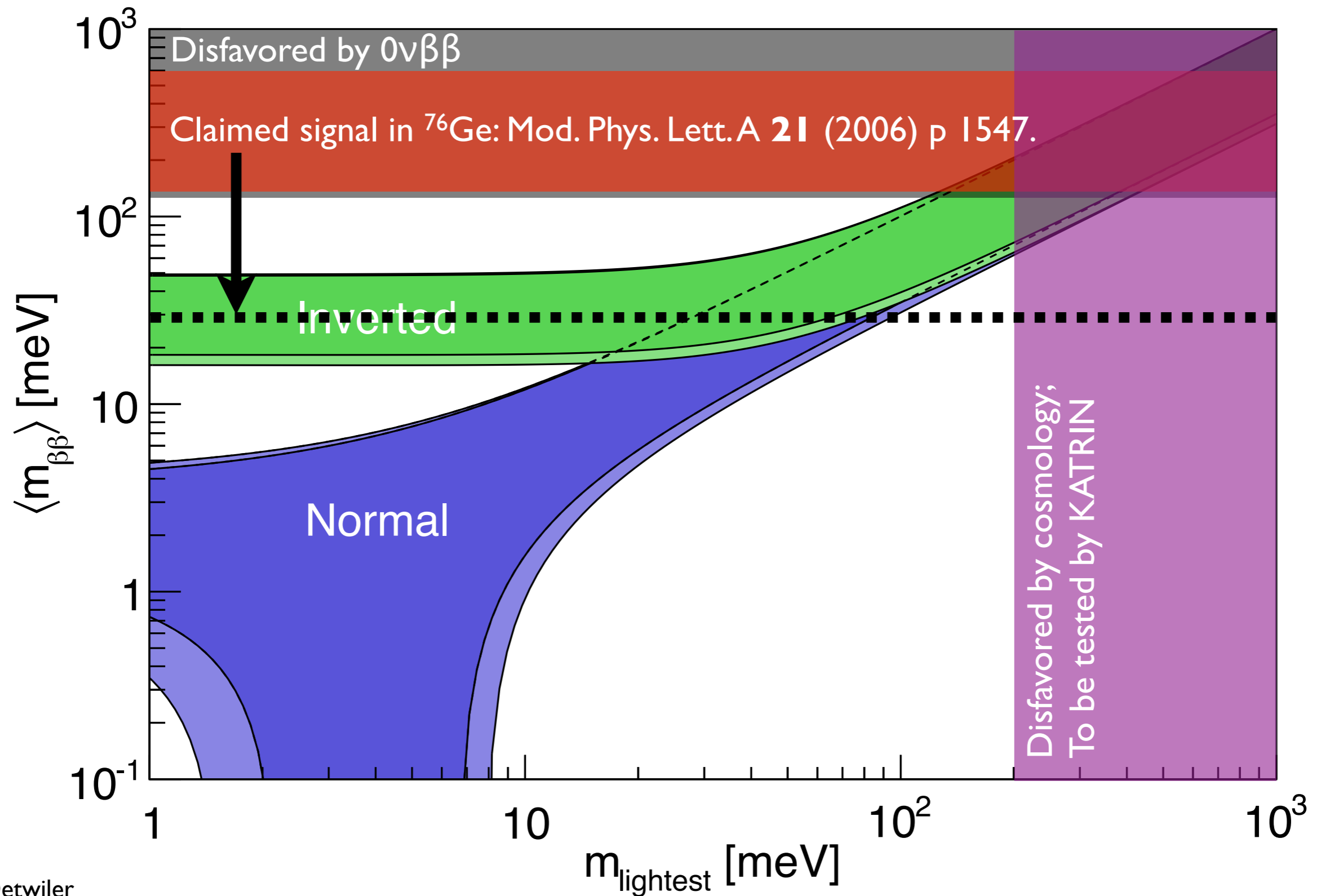
Double-Beta Decay



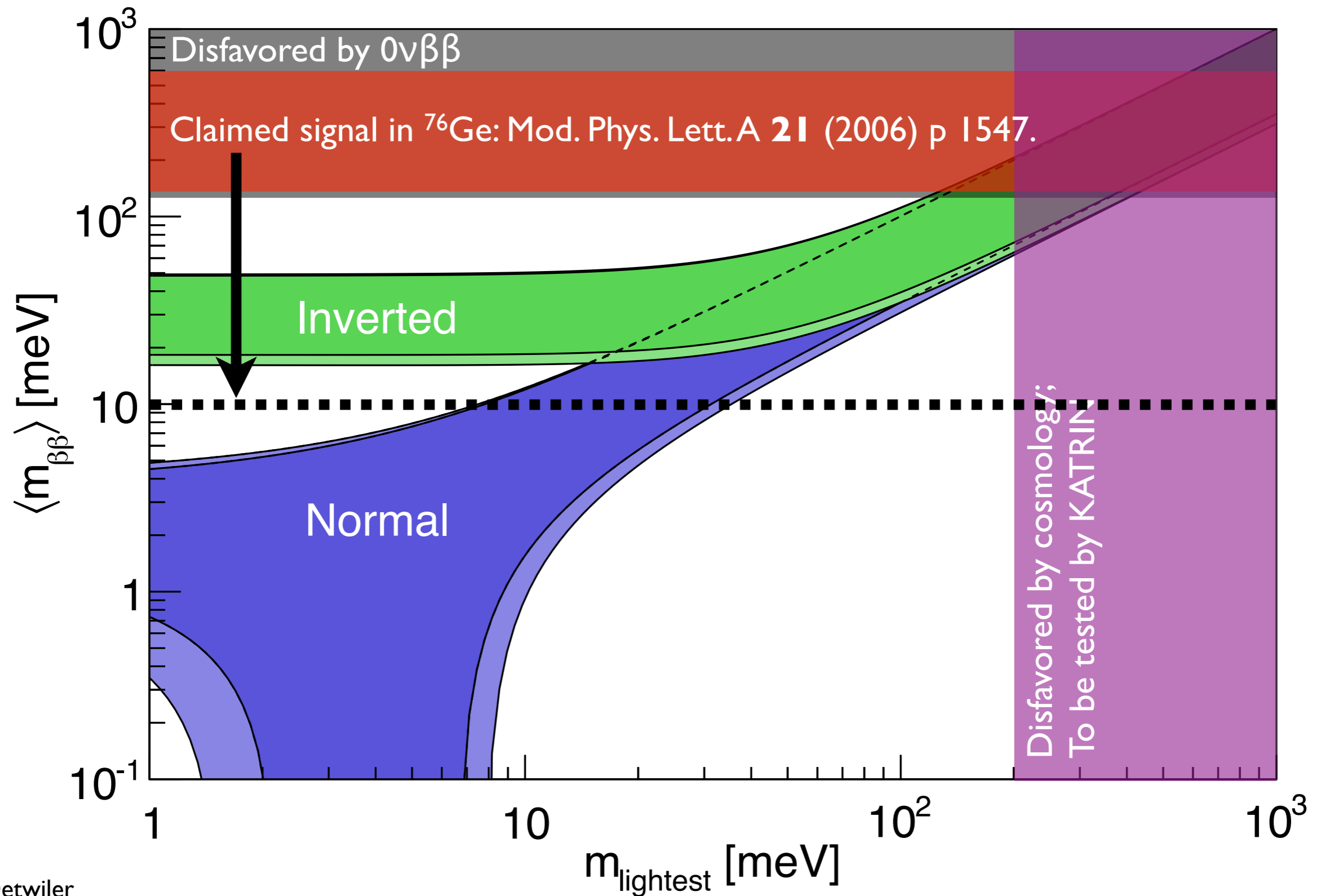
Double-Beta Decay



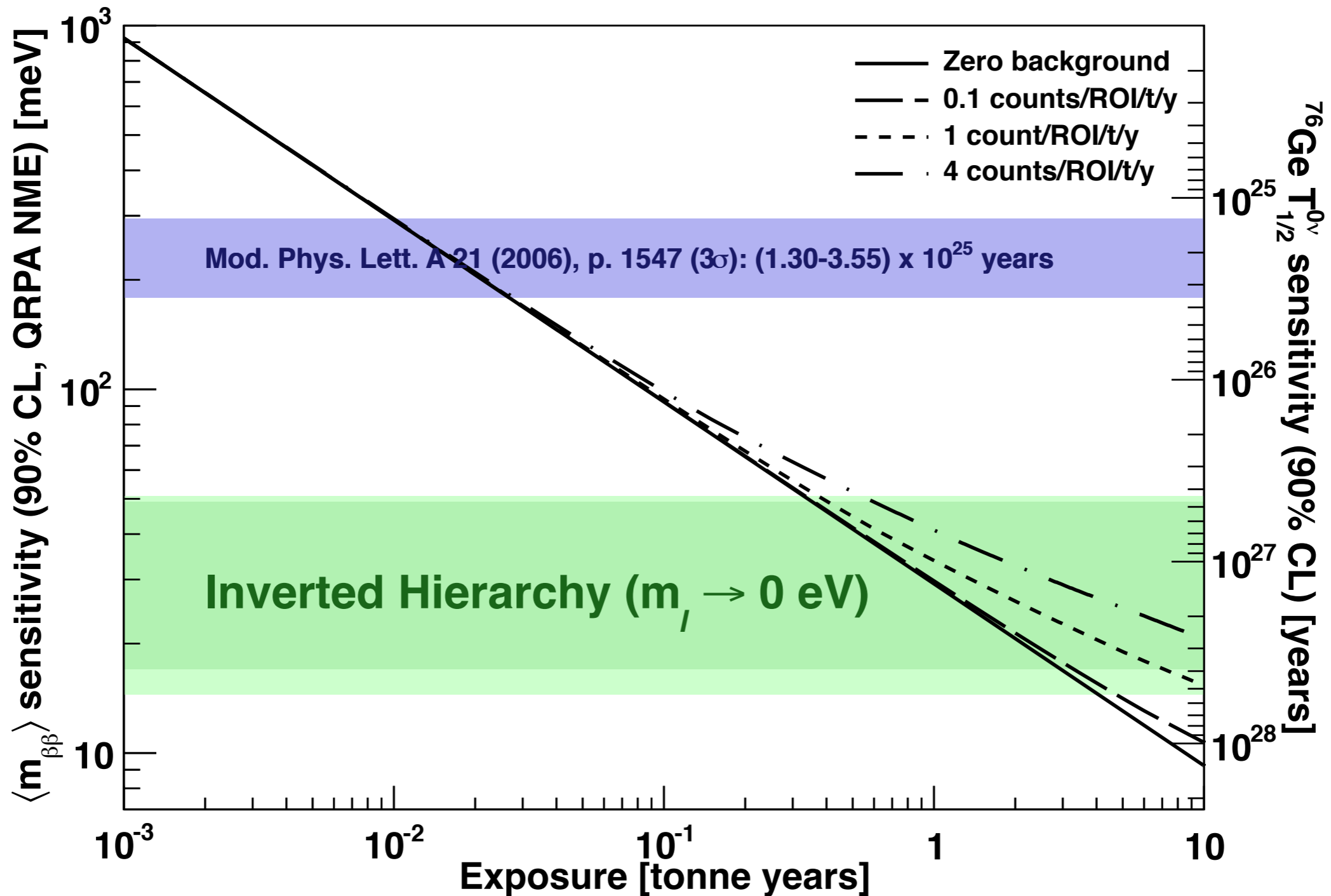
Double-Beta Decay



Double-Beta Decay



Testing the Inverted Hierarchy



$0\nu\beta\beta$ Experiments

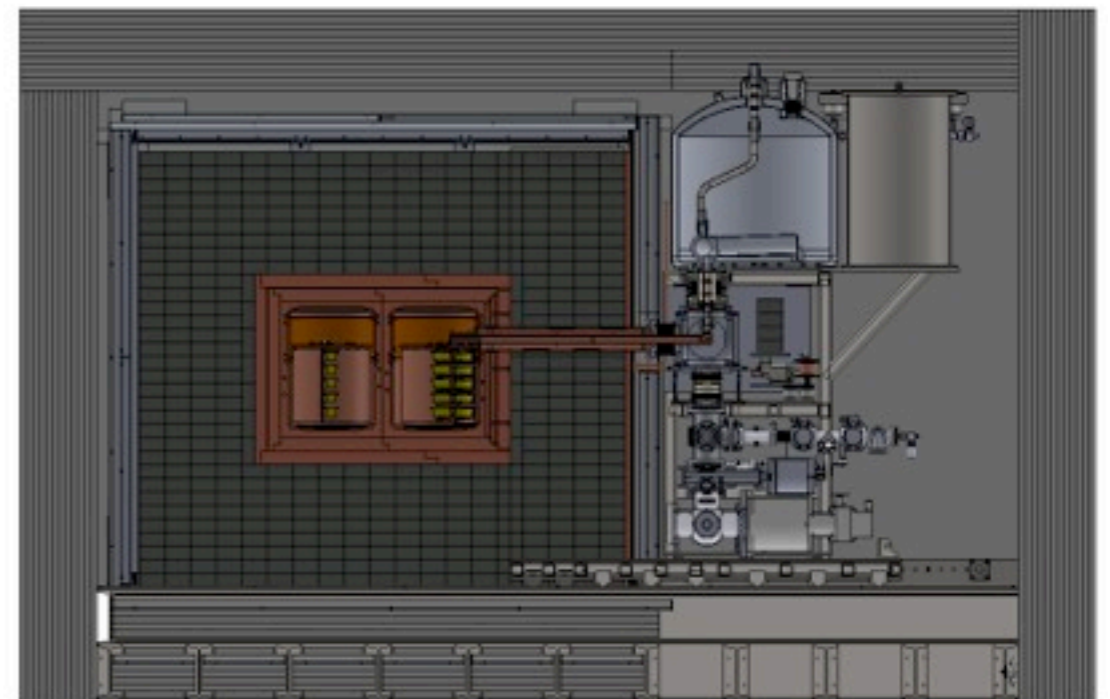
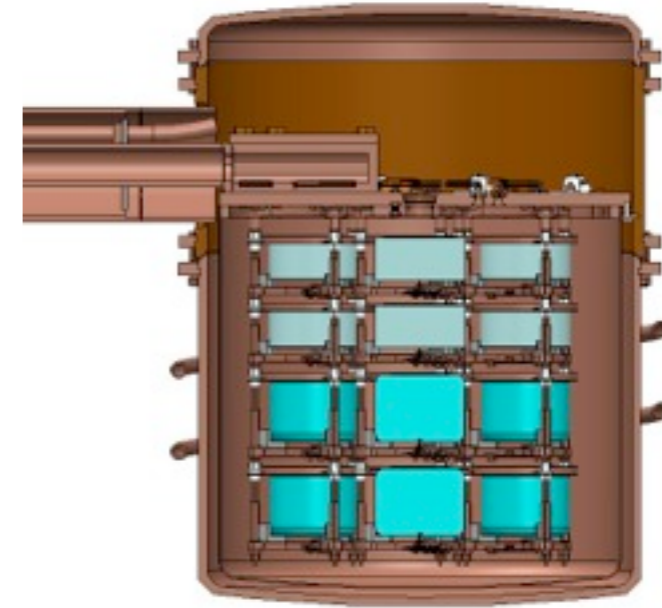
| Experiment | Isotope | Mass | Technique | Present Status | Location |
|------------------------------|-------------------|---------|--|---------------------|------------|
| AMoRE ^{89,90} | ¹⁰⁰ Mo | 50 kg | CaMoO ₄ scint. bolometer crystals | Development | Yangyang |
| CANDLES ⁹¹ | ⁴⁸ Ca | 0.35 kg | CaF ₂ scint. crystals | Prototype | Kamioka |
| CARVEL ⁹² | ⁴⁸ Ca | 1 ton | CaF ₂ scint. crystals | Development | Solotvina |
| COBRA ⁹³ | ¹¹⁶ Cd | 183 kg | ^{enr} Cd CZT semicond. det. | Prototype | Gran Sasso |
| CUORE-0 ⁶⁹ | ¹³⁰ Te | 11 kg | TeO ₂ bolometers | Construction - 2012 | Gran Sasso |
| CUORE ⁶⁹ | ¹³⁰ Te | 203 kg | TeO ₂ bolometers | Construction - 2013 | Gran Sasso |
| DCBA ⁹⁴ | ¹⁵⁰ Ne | 20 kg | ^{enr} Nd foils and tracking | Development | Kamioka |
| EXO-200 ⁵⁷ | ¹³⁶ Xe | 160 kg | Liq. ^{enr} Xe TPC/scint. | Operating - 2011 | WIPP |
| EXO ⁷⁰ | ¹³⁶ Xe | 1-10 t | Liq. ^{enr} Xe TPC/scint. | Proposal | SURF |
| GERDA ⁷¹ | ⁷⁶ Ge | ≈35 kg | ^{enr} Ge semicond. det. | Operating - 2011 | Gran Sasso |
| GSO ⁹⁵ | ¹⁶⁰ Gd | 2 ton | Gd ₂ SiO ₅ :Ce crys. scint. in liq. scint. | Development | |
| KamLAND-Zen ⁹⁶ | ¹³⁶ Xe | 400 kg | ^{enr} Xe dissolved in liq. scint. | Operating - 2011 | Kamioka |
| LUCIFER ^{97,98} | ⁸² Se | 18 kg | ZnSe scint. bolometer crystals | Development | Gran Sasso |
| MAJORANA ^{77,78,79} | ⁷⁶ Ge | 26 kg | ^{enr} Ge semicond. det. | Construction - 2013 | SURF |
| MOON ⁹⁹ | ¹⁰⁰ Mo | 1 t | ^{enr} Mo foils/scint. | Development | |
| SuperNEMO-Dem ⁸⁷ | ⁸² Se | 7 kg | ^{enr} Se foils/tracking | Construction - 2014 | Fréjus |
| SuperNEMO ⁸⁷ | ⁸² Se | 100 kg | ^{enr} Se foils/tracking | Proposal - 2019 | Fréjus |
| NEXT ^{82,83} | ¹³⁶ Xe | 100 kg | gas TPC | Development - 2014 | Canfranc |
| SNO+ ^{84,85} | ¹⁵⁰ Nd | 55 kg | Nd loaded liq. scint. | Construction - 2013 | SNO Lab |

Construction

Operation

The MAJORANA DEMONSTRATOR (MJD)

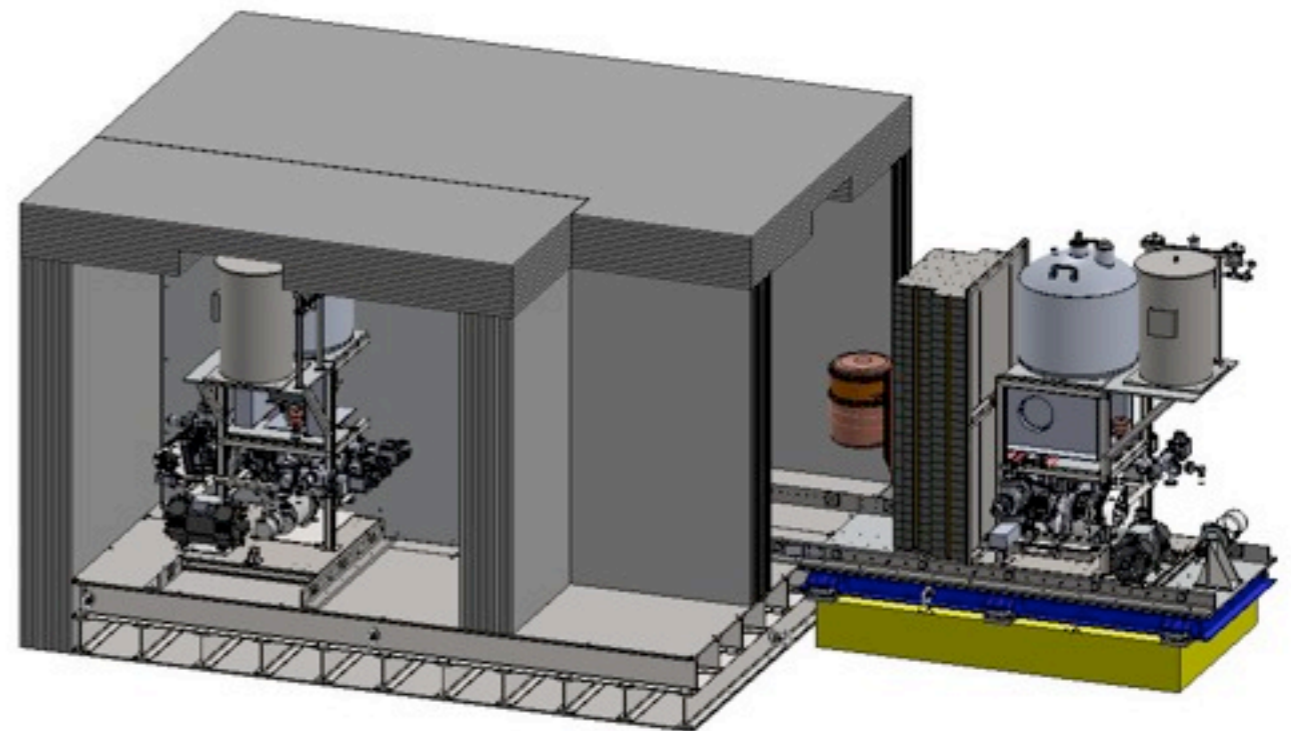
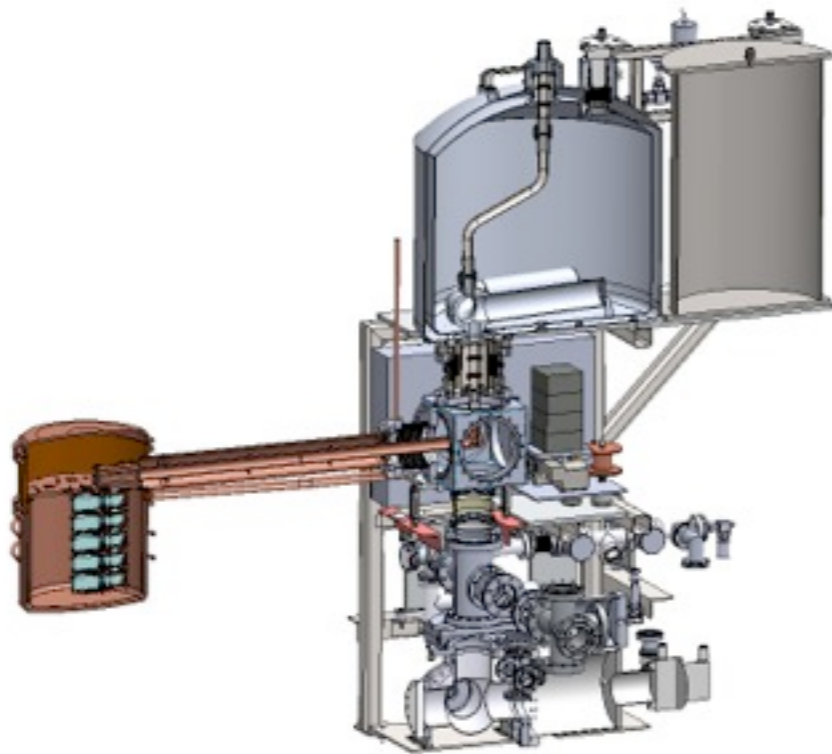
- Located 4850' underground at Sanford Underground Research Facility
- 40-kg of Ge detectors, 30-kg enriched to 86% in ^{76}Ge
- 2 independent cryostats made of ultra-clean, electroformed Cu
- Compact Pb and Cu shield + muon veto
- Background goal: 3 counts in the $0\nu\beta\beta$ peak region of interest in a one tonne-year exposure



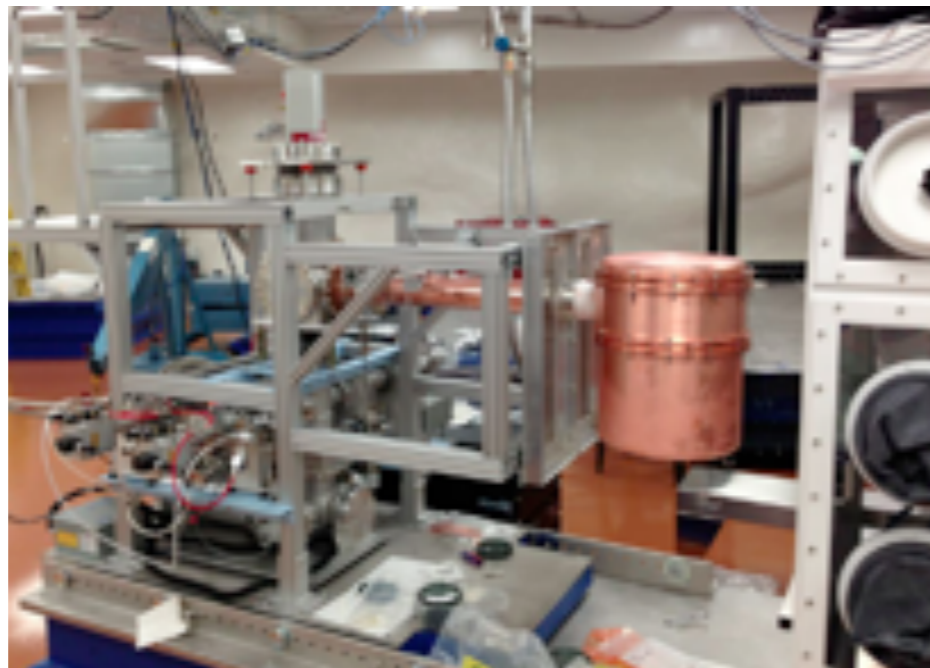
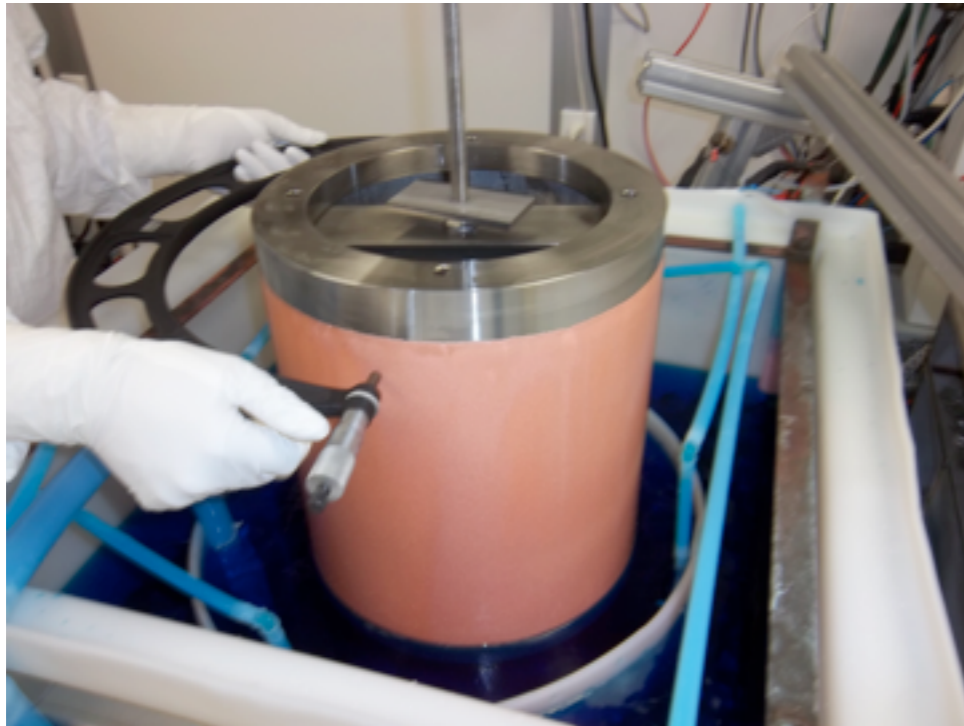
Funded by DOE Office of Nuclear Physics and NSF Particle Astrophysics,
with additional contributions from international collaborators.

MJD Implementation

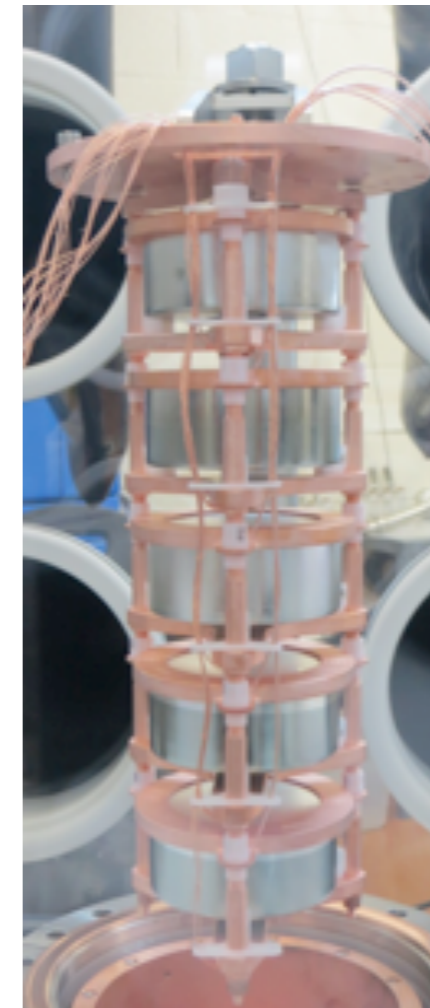
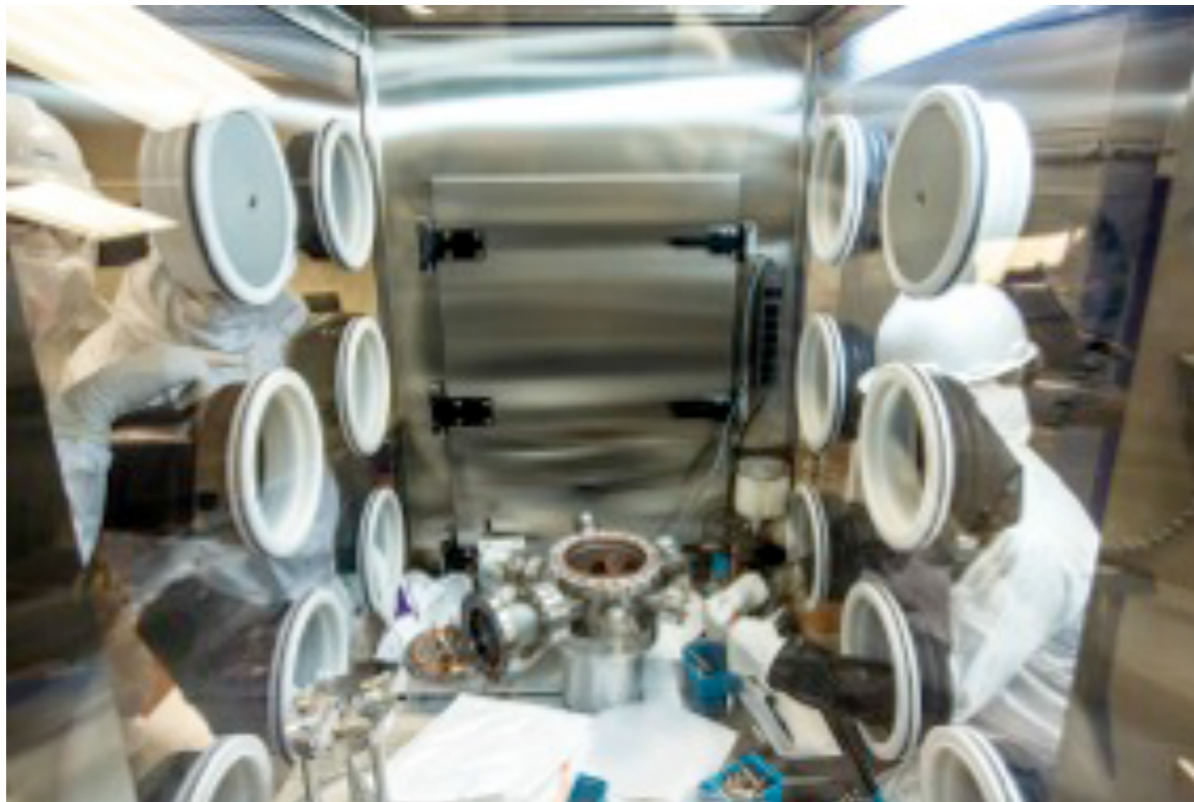
- Prototype Cryostat (2 strings, natGe): Summer 2013
- Cryostat 1 (3 strings enrGe & 4 strings natGe): Late 2013
- Cryostat 2 (7 strings enrGe): Fall 2014



MJD Construction



MJD Construction





MJD Collaboration



Jason Detwiler

Other MJD Physics

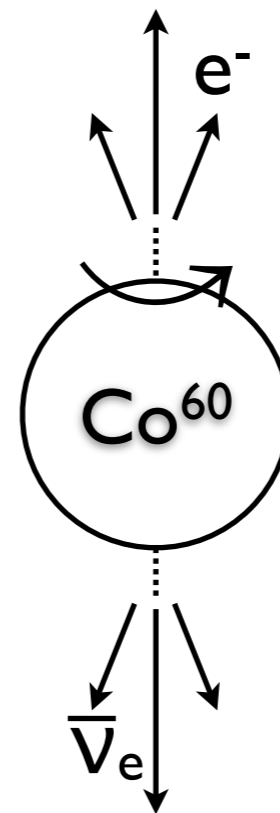
- Light WIMP dark matter
- Solar axions
- Pauli exclusion principle violation
- Neutrino magnetic moment
- Electron decay

Summary

- Neutrinos continue to give us new mysteries to solve
- Majorana neutrinos may give us insights into Grand Unification and the Matter-Antimatter Asymmetry of the Universe
- $0\nu\beta\beta$ experiments are the only known way to probe this aspect of the neutrino
- Definitive tests of inverted hierarchy Majorana neutrinos are within reach!

Neutrino Handedness

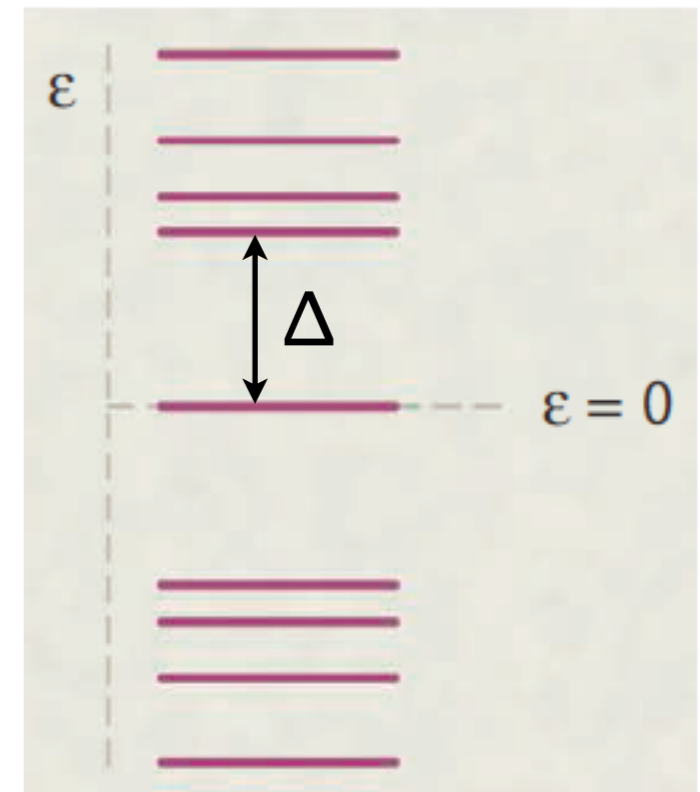
C.S. Wu et al. (1957)



The Majorana Fermion

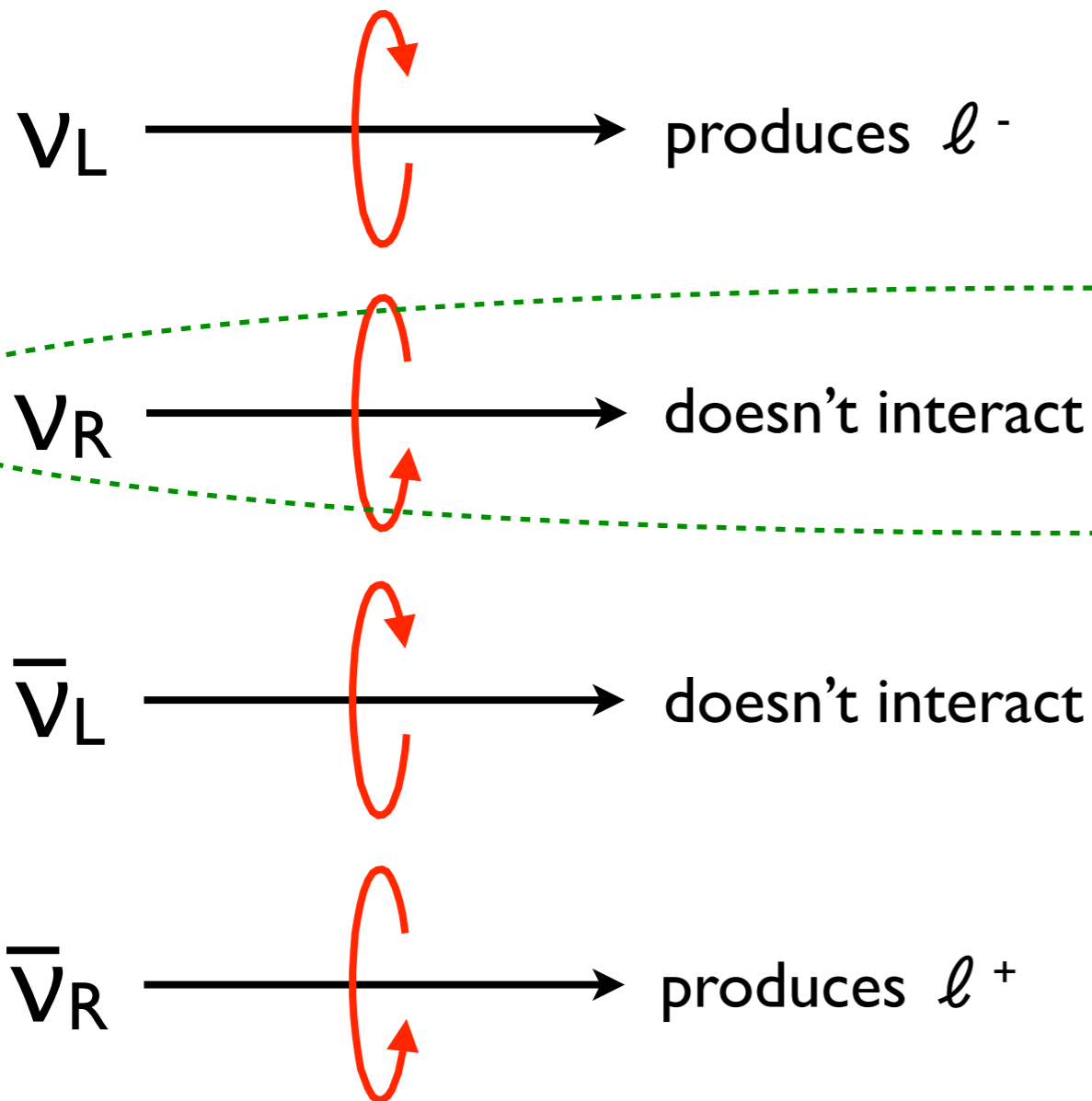
$$\sigma_{\pm}^{\mu} \partial_{\mu} \chi_{\pm} \pm m \sigma_2 \chi_{\pm}^{*} = 0$$

- $CPT[\chi_{\pm}(x^{\mu})] = \mp i \chi_{\pm}(-x^{\mu})$
- Particle physics: $\chi = \bar{\chi}$
- Condensed matter:
charge-neutral, zero-energy
quasi-particle

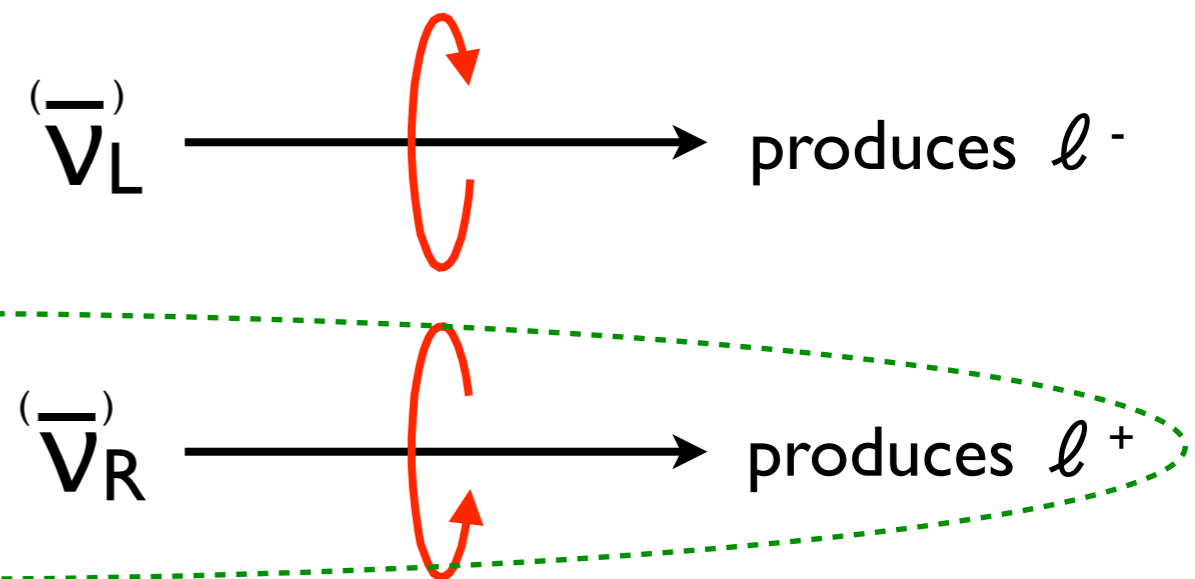


Testing $\nu = \bar{\nu}$ (I)

Dirac ν



Majorana ν



Why not generate ν_R
in a beam and see if it
produces l^+ ?

The Majorana Equation



Schrodinger:
$$i\frac{\partial}{\partial t}\Psi + \frac{1}{2m}\nabla^2\Psi = 0$$

The Majorana Equation



Schrodinger:
$$i\frac{\partial}{\partial t}\Psi + \frac{1}{2m}\nabla^2\Psi = 0$$



Dirac:
$$-i\gamma^\mu\partial_\mu\psi + m\psi = 0$$

The Majorana Equation



Schrodinger:
$$i\frac{\partial}{\partial t}\Psi + \frac{1}{2m}\nabla^2\Psi = 0$$

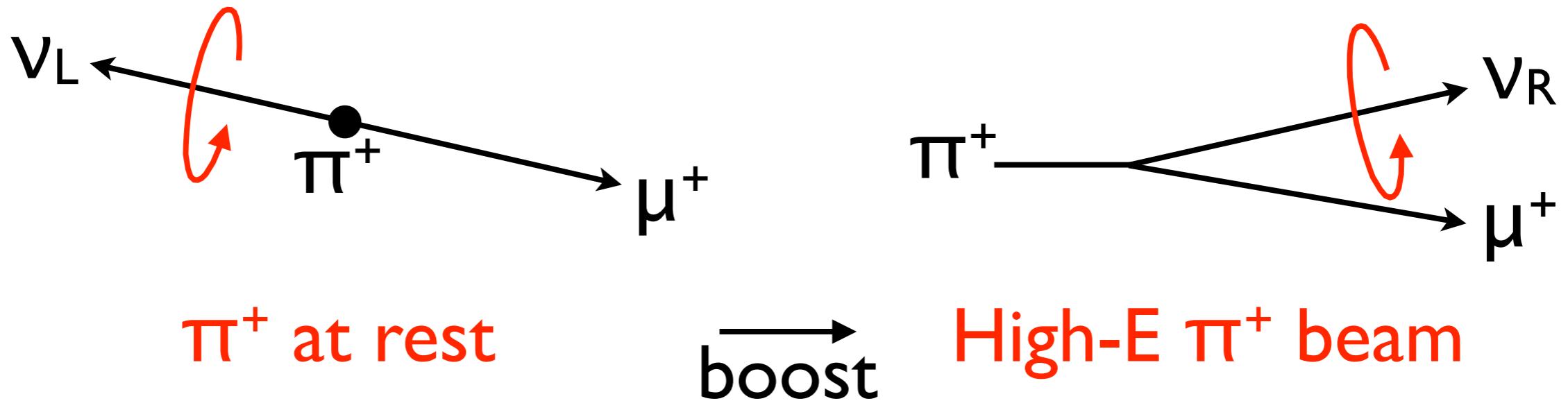


Dirac:
$$-i\gamma^\mu\partial_\mu\psi + m\psi = 0$$



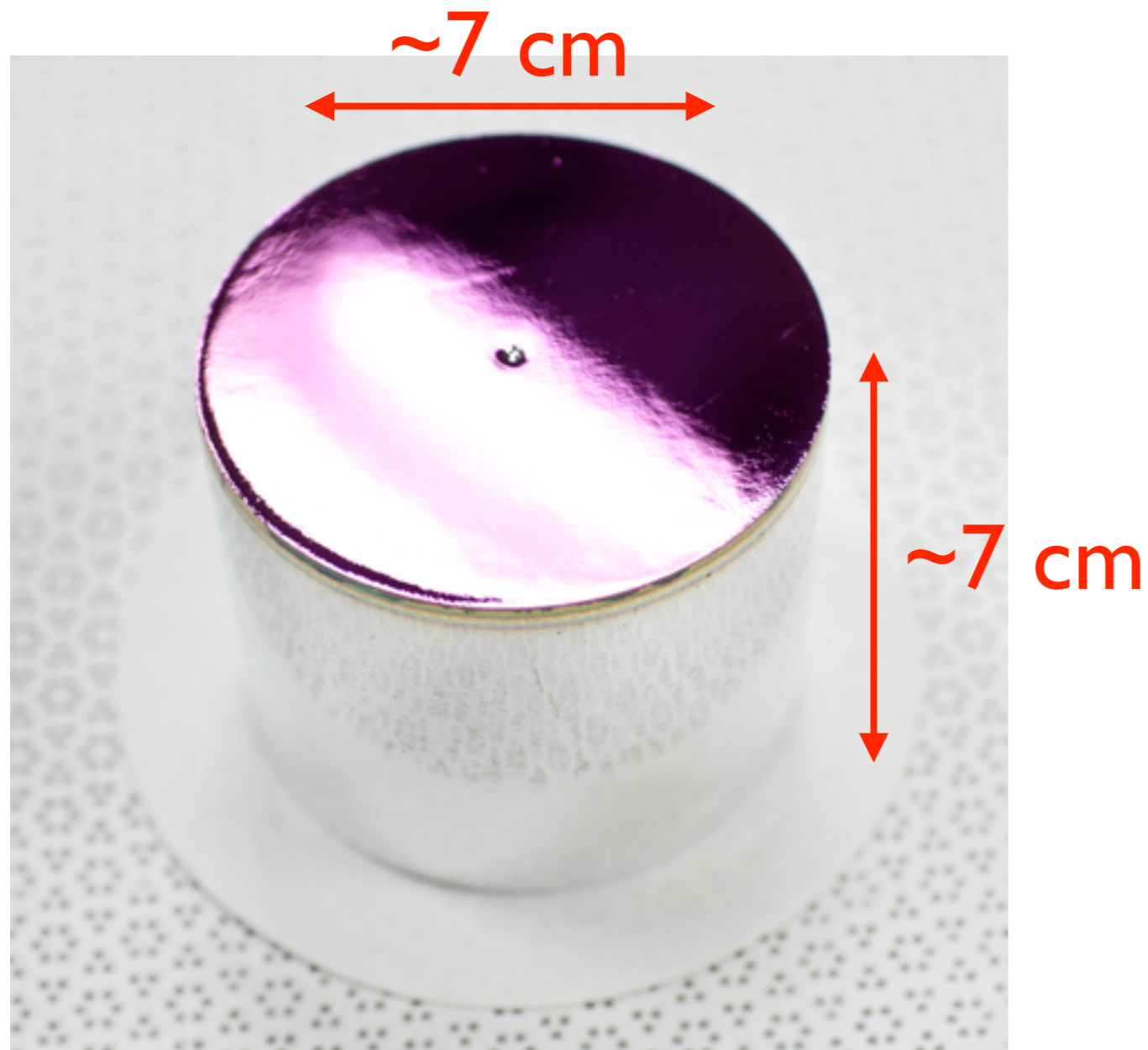
Majorana:
$$\sigma_\pm^\mu\partial_\mu\chi \pm m\sigma_2\chi^* = 0$$

Testing $\nu = \bar{\nu}$ (I)

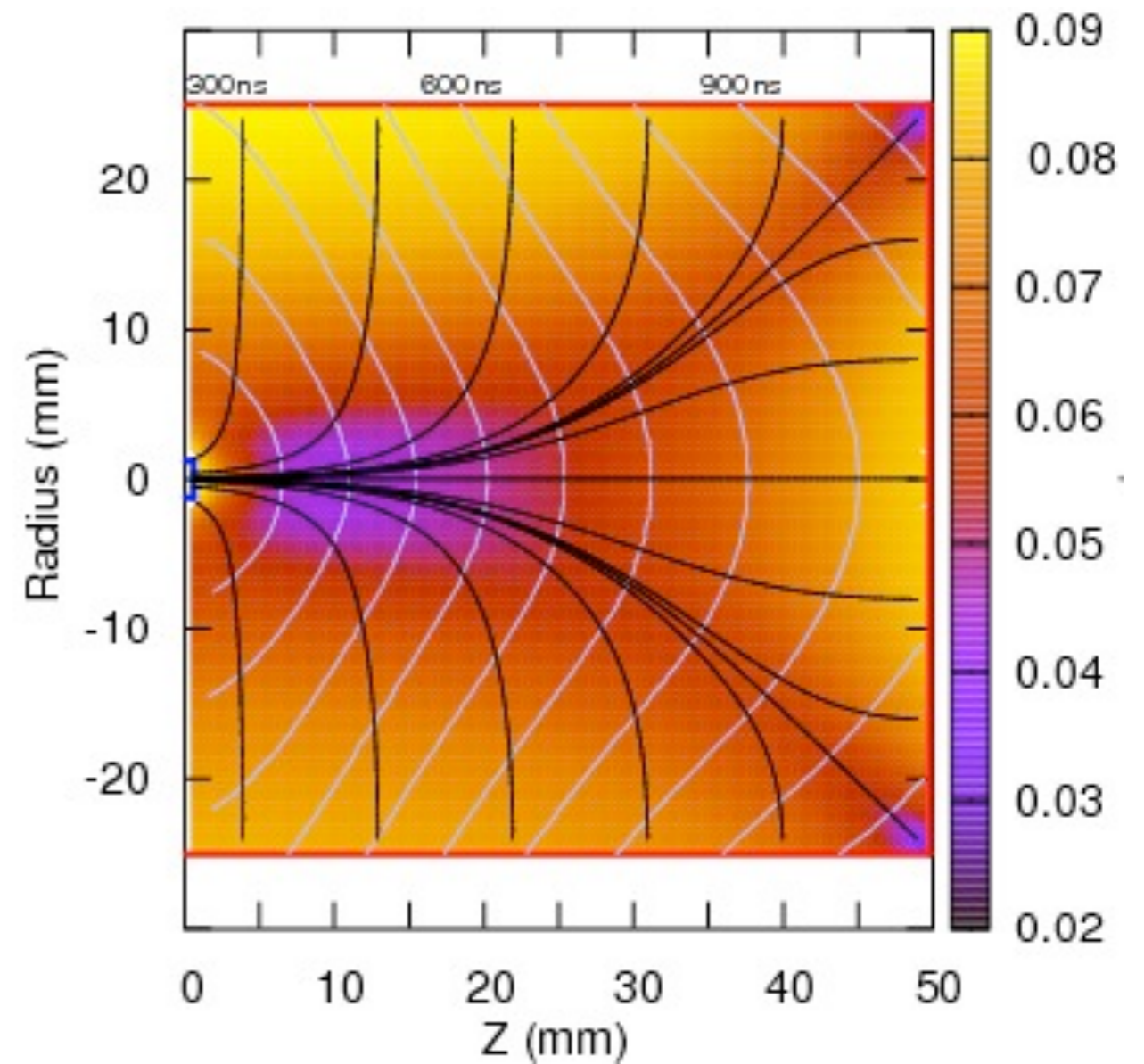
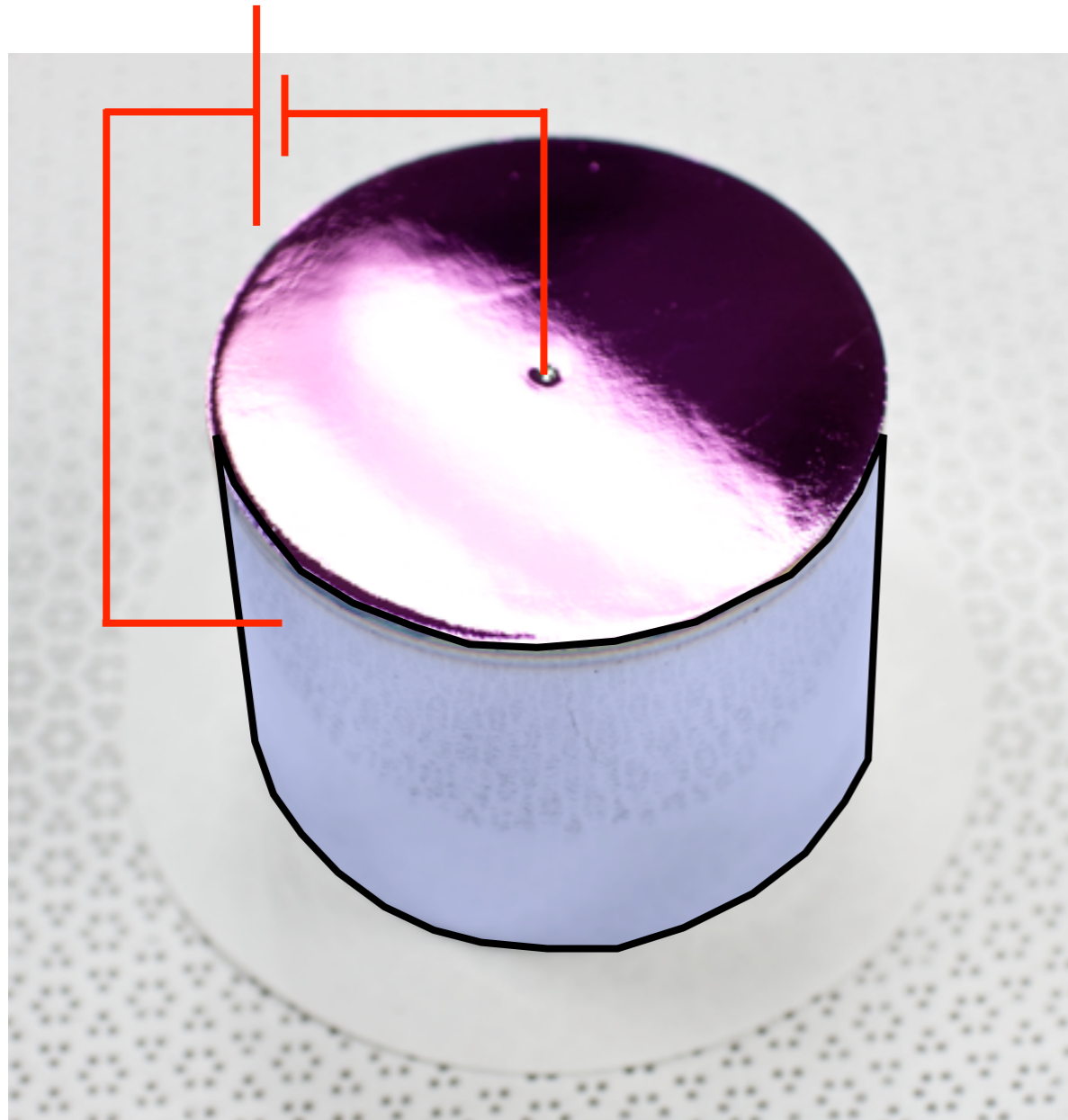


- Boost so that π^+ beam faster than ν_L from decay at rest: requires $E_\pi > 4 \text{ PeV}$ (n.b. LHC = 14 TeV)
- Fraction of decays with helicity flipped: $< 10^{-15}$
- *“Since L-violation comes only from Majorana ν masses, any attempt to observe it will be at the mercy of the ν masses.”*
- B. Kayser

Germanium Detectors

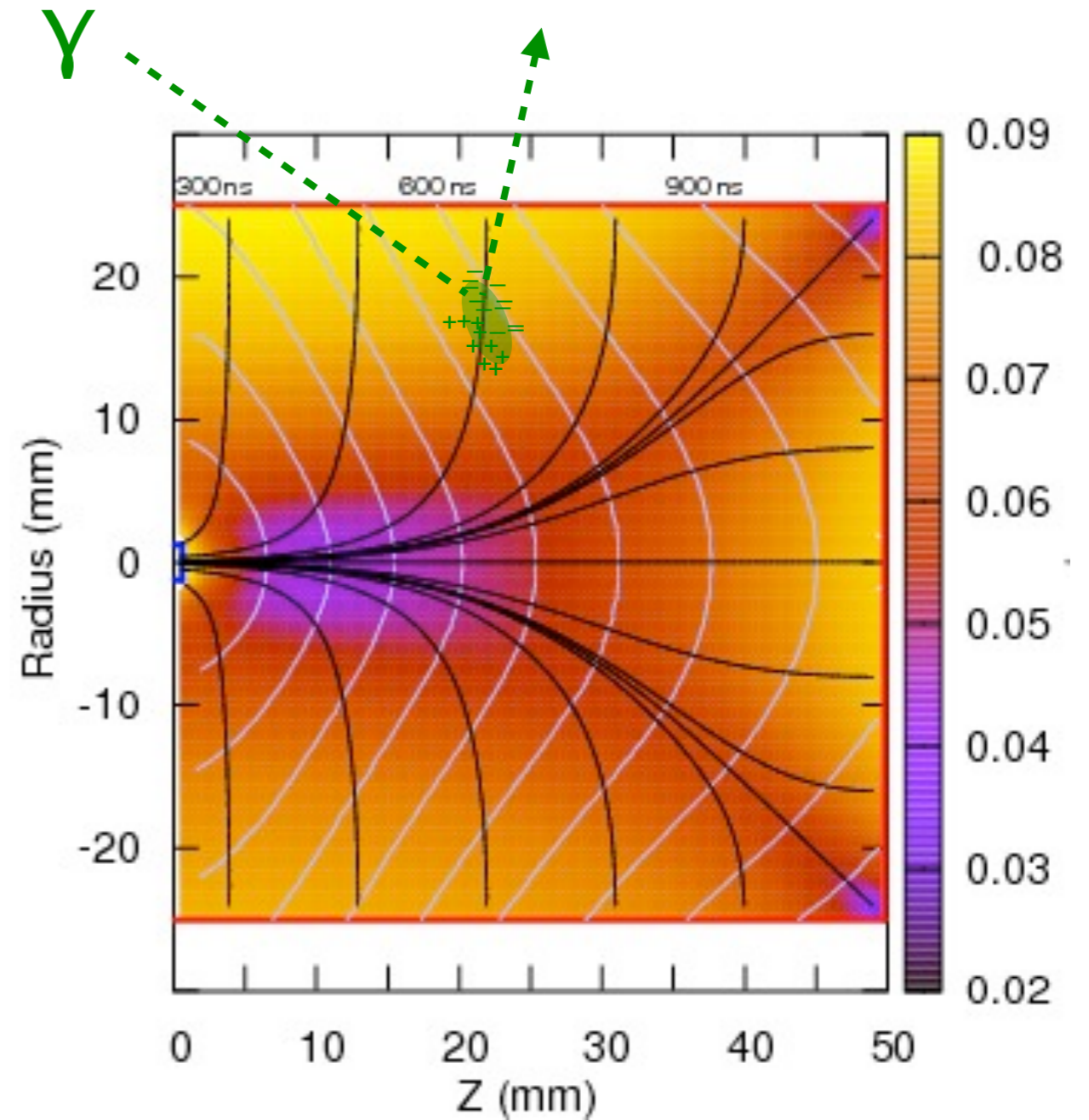
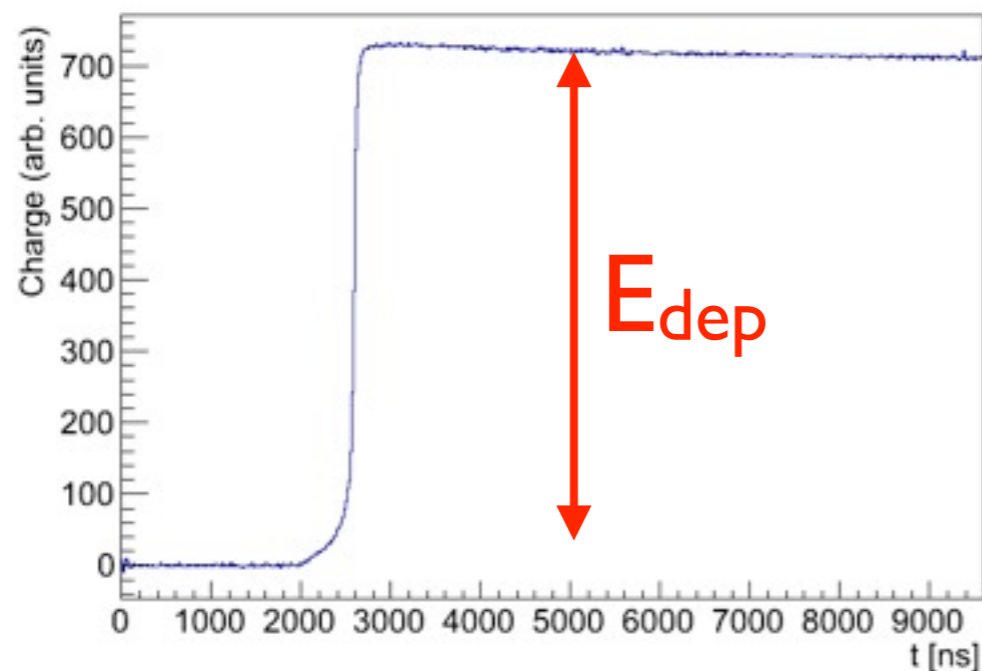
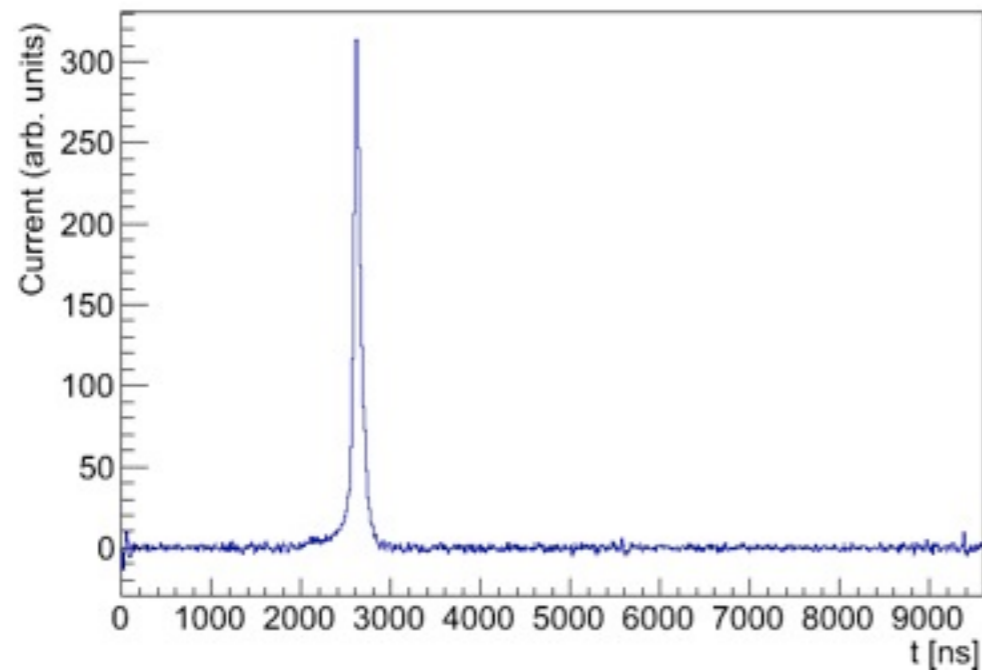


Point-Contact Ge Detectors



Hole v_{drift} (mm/ns) w/ paths, isochrones

Point-Contact Ge Detectors



Hole v_{drift} (mm/ns) w/ paths, isochrones

Point-Contact Ge Detectors

