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Theory of the Trojan–Horse method

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Abstract

The Trojan–Horse method is an indirect approach to determine the energy dependence of S factors of astrophysically relevant two-body reactions. This is accomplished by studying closely related three-body reactions under quasi-free scattering conditions. The basic theory of the Trojan–Horse method is developed starting from a post-form distorted wave Born approximation of the T-matrix element. In the surface approximation the cross-section of the three-body reaction can be related to the S-matrix elements of the two-body reaction. The essential feature of the Trojan–Horse method is the effective suppression of the Coulomb barrier at low energies for the astrophysical reaction leading to finite cross-sections at the threshold of the two-body reaction. In a modified plane wave approximation the relation between the two- and three-body cross-sections becomes very transparent. The appearing Trojan–Horse integrals are studied in detail.

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 \ldots κ ϵ καλυμμ ϵ νοι ιππω. [1]

1. Introduction

Nuclear reaction rates are an indispensable ingredient of many astrophysical models. They have to be known with sufficient accuracy in order to understand quantitatively the evolution of the universe, stars and other objects in the cosmos [2–4]. Ideally, reaction cross-sections are directly measured in the laboratory. But, with few exceptions and despite many experimental efforts, the relevant low energy range cannot be reached in direct experiments [5–7]. Cross-sections for reactions with charged particles rapidly become very small with decreasing energy of the colliding nuclei due to the repulsive Coulomb interaction. Extrapolations of the crosssection $\sigma(E)$ to low energies from results at higher energies accessible to experiments are often needed. This is accomplished with the help of the astrophysical S factor

$$
S(E) = \sigma(E)E \exp(2\pi\eta),\tag{1}
$$

where E is the c.m. energy and $\eta = Z_1 Z_2 e^2 / (\hbar v)$ is the Sommerfeld parameter which depends on the charge numbers Z_1 , Z_2 of the colliding nuclei and their relative velocity v in the entrance channel. The S factor shows a much weaker energy dependence than the cross-section $\sigma(E)$ because the main effect of the penetrability through the Coulomb barrier is compensated by the increase of the exponential factor. The extrapolation process introduces uncertainties and important contributions to the cross-sections, like resonances, can be missed. Additionally, direct laboratory experiments are affected by electron screening, which effectively reduces the Coulomb barrier between the nuclei and enhances the measured laboratory cross-section [8– 10]. A correction has to be applied to obtain the cross-section for bare nuclei. This effect does not seem to be completely understood yet and independent information on low energy cross-sections is valuable in order to develop a quantitative description of electron screening. In astrophysical applications one has, in addition, to account for the screening under stellar conditions.

In recent years several indirect methods have been developed to extract cross-sections relevant to astrophysics from other types of experiments. In these alternative approaches the astrophysical relevant two-body reaction is generally replaced by a suitably chosen three-body reaction. The relation between the reactions is established with the help of nuclear reaction theories. Without doubt, this process will introduce some uncertainties, but valuable information on the astrophysical reaction can be obtained. Also, the errors are independent from that of the direct approach. Of course, the indirect methods have to be validated by studying well-known reactions before firm conclusions can be drawn from indirect experiments in cases where direct measurements are not feasible; see also the minireview [11].

The Coulomb dissociation method has become a valuable tool for extracting lowenergy cross-sections of radiative capture reactions $a(b, \gamma)c$ by studying the inverse process of photo dissociation $c(y, b)a$ [12]. Instead of using real photons, the Coulomb field of a highly charged target X acts as a source of virtual photons during a scattering process which leads to the breakup reaction $X(c, ab)X$ with three particles in the final state. Due to the high flux of virtual photons, cross-sections at the small relevant energies in the two-body system are strongly enhanced as compared to the capture reaction. Another approach is the ANC method that tries to extract the asymptotic normalization coefficient of a nuclear ground state wave function by studying various combinations of transfer reactions involving this nucleus at low energies $[13,14]$. The coefficient can be used to determine theoretically the astrophysical S factor for radiative capture reactions at zero energy. However, these indirect approaches are limited to astrophysically relevant reactions where a photon is involved.

The observation of a similarity between cross-sections for two-body and closely related three-body reactions under certain kinematical conditions [15] led to the introduction of the Trojan–Horse method (THM) [16], see also [17,18]. The aim of the THM is to extract the cross-section of an astrophysically relevant two-body reaction

$$
A + x \to C + c,\tag{2}
$$

from a suitably chosen reaction

$$
A + a \to C + c + b,\tag{3}
$$

with three particles in the final state assuming that the Trojan–Horse a is composed predominantly of clusters x and b . The kinematical conditions are chosen such that the momentum transfer to the nucleus *h* is small during the reaction. Therefore *h* can be considered as a spectator to the reaction of A and x . This process is often referred to as a quasi-free scattering. In a selected part of the available three-body phase space it is known to dominate over other reaction mechanism like sequential breakup processes. In the past, quasi-free scattering has been used to extract information on momentum distributions of the nucleus a , i.e., the ground state wave function in momentum space, employing a plane-wave impulse approximation (PWIA) in the theoretical description [19]. In this approach the cross-section of reaction (3) factorizes into a kinematical factor, the ground state momentum distribution of nucleus a and an off-shell cross-section of reaction (2) that is assumed to be known. On the other hand the two-body cross-section can be extracted from the cross-section of reaction (3) if the momentum distribution of the Trojan–Horse a is known with sufficient accuracy and a relation between the off-shell and on-shell twobody cross-sections can be established. The selection of different spectators b and thus Trojan–Horses a allows additional checks of the underlying assumptions of the method.

Unlike the Coulomb dissociation method and the ANC method which are limited to radiative processes, the THM can be applied to reactions where no photon is involved. The essential feature of the THM is the actual suppression of the Coulomb barrier in the cross-section of the two-body reaction. The cross-section of the threebody reaction is not reduced when the c.m. energy in the $A + x$ system approaches zero as in reaction (2). The energy in the entrance channel of (3) can be around or above the Coulomb barrier and effects from electron screening are negligible. Nevertheless, very small energies in the reaction (2) can be reached.

The feasibility of the THM was studied in several experiments involving various reactions during the last years [20–28]. In earlier evaluations of the experiments some simplifying assumptions were made in the theoretical description. The off-shell twobody cross-section as appearing in the PWIA was considered as the bare nuclear cross-section. It was converted to the on-shell two-body cross-section by correcting for the Coulomb penetration in a heuristic approximation. Basic considerations for the theoretical description of the THM were developed in [29] but the deduced relation between the cross-sections for reactions (2) and (3) were not directly applicable to the experiments.

In this paper the theory of the THM is developed in certain approximations that allow to establish a simple connection between the cross-sections of the three- and two-body reactions. In Section 2 the reaction theory is formulated and the relation of the T-matrix element of the three-body reaction with the S-matrix elements of the two-body reaction is found. In connection with a plane wave approximation, fundamental TH integrals appear that are discussed in Section 3. Technical details concerning the calculation of these integrals are discussed in Appendices A and B. In the following section expressions for scattering amplitudes and cross-sections are derived for spinless particles. For completeness, the general case of particles with spin is considered in Appendix C. The TH Coulomb scattering amplitude that is relevant in the indirect study of elastic scattering processes is treated in Section 5. Kinematical conditions, the energy dependence of cross-sections, and applications of the THM are discussed in Section 6. A summary and an outlook are presented in the last section. The evaluation of actual Trojan–Horse experiments with the present theory is beyond the scope of this paper and is subject to a detailed treatment in separate studies.

2. Reaction theory

In the three-body reaction (3) the particle x is transferred from the Trojan–Horse a to the nucleus A leading to the $C + c$ final state with the spectator b of the Trojan– Horse remaining, see Fig. 1. This reaction can be considered as a transfer to the con-

Fig. 1. Transfer of particle x in the Trojan–Horse reaction (3).

tinuum, where the inelastic reaction $A + x \rightarrow C + c$ can happen in addition to the elastic $A + x$ scattering. It is customary to describe such a direct reaction, i.e., $A + a \rightarrow C + c + b$, with the help of a distorted wave Born approximation. It is not required in the theoretical approach that $A + x \neq C + c$; elastic scattering processes in the two-body system can be treated in the formalism, too. Effects from the antisymmetrization of the wave functions will be neglected in the present treatment; they are expected to be small.

2.1. Coordinate systems and cross-sections

For the general case of light ion scattering, the masses of the particles A , b , and x are comparable, so we have to use the complete kinematics; it is not justified, e.g., to assume infinite target mass m_A which would lead to a considerable simplification of the formulas. This makes the notation unavoidably clumsy.

In a three-body system various sets of Jacobi coordinates are used to specify the positions of the particles. In the theoretical description we encounter the sets

$$
\vec{r}_{xb} = \vec{r}_x - \vec{r}_b, \quad \vec{r}_{Aa} = \vec{r}_A - \vec{r}_a = \vec{r}_A - \frac{m_x \vec{r}_x + m_b \vec{r}_b}{m_x + m_b}
$$
(4)

in the initial partition and

$$
\vec{r}_{Cc} = \vec{r}_c - \vec{r}_c, \quad \vec{r}_{Bb} = \vec{r}_B - \vec{r}_b = \frac{m_C \vec{r}_C + m_c \vec{r}_c}{m_C + m_c} - \vec{r}_b \tag{5}
$$

in the final partition. The symbol B denotes the $C + c = A + x$ system. The coordinate vectors are shown in Fig. 2. The corresponding relative momenta and wave vectors are given by

$$
\vec{p}_{ij} = \hbar \vec{k}_{ij} = \frac{m_j \vec{p}_i - m_i \vec{p}_j}{m_i + m_j} \tag{6}
$$

Fig. 2. Coordinate vectors in the initial and final partitions of the Trojan–Horse reaction (3).

for nuclei i and j with masses m_i, m_j and momenta \vec{p}_i, \vec{p}_j in the laboratory system, respectively. With the help of the kinetic energies

$$
E_{ij} = \frac{p_{ij}^2}{2\mu_{ij}},\tag{7}
$$

where the reduced masses

$$
\mu_{ij} = \frac{m_i m_j}{m_i + m_j} \tag{8}
$$

appear, energy conservation in the two-body reaction (2) can be expressed as

$$
E_{Ax} = E_{Cc} - Q_2,\tag{9}
$$

with the Q-value

$$
Q_2 = (m_A + m_x - m_C - m_c)c^2
$$
 (10)

and similarly

$$
E_{Aa} = E_{Cc} + E_{Bb} - Q_3 = E_{Ax} + E_{Bb} + Q_2 - Q_3, \qquad (11)
$$

with

$$
Q_3 = (m_A + m_a - m_C - m_b)c^2
$$
 (12)

in case of the three-body reaction.

The general cross-section for reaction (3) with three particles in the final state depends on the choice of the independent variables and the reference system. In the c.m. system the energy E_{C_c} and the directions of the Jacobi momenta \vec{p}_{Bb} and \vec{p}_{Cc} completely specify the kinematical conditions for a given projectile energy. Then the differential cross-section takes the form [30,31]

$$
\frac{\mathrm{d}^3 \sigma}{\mathrm{d}E_{Cc} \,\mathrm{d}\Omega_{Cc} \,\mathrm{d}\Omega_{Bb}} = K_{\text{c.m.}} \Big| T_{fi}(\vec{k}_{Cc}, \vec{k}_{Bb}; \vec{k}_{Aa}) \Big|^2, \tag{13}
$$

with the c.m. kinematical factor

$$
K_{\text{c.m.}} = \frac{\mu_{Aa}\mu_{Bb}\mu_{Cc}}{(2\pi)^5\hbar^7} \frac{p_{Bb}p_{Cc}}{p_{Aa}} \tag{14}
$$

and the T-matrix element T_f . In an actual experiment, the particles C and c are usually detected in the laboratory system and it is more convenient to use

$$
\frac{\mathrm{d}^3 \sigma}{\mathrm{d}E_C \,\mathrm{d}\Omega_C \,\mathrm{d}\Omega_c} = K_{\text{lab}} \left| T_{fi}(\vec{k}_{Cc}, \vec{k}_{Bb}; \vec{k}_{Aa}) \right|^2, \tag{15}
$$

with the kinematical factor

$$
K_{\rm lab} = \frac{\mu_{Aa}m_C}{(2\pi)^5\hbar^7} \frac{p_C p_c^2}{p_{Aa}} \left[\left(\frac{\vec{p}_{Bb}}{\mu_{Bb}} - \frac{\vec{p}_{Cc}}{m_c} \right) \cdot \frac{\vec{p}_c}{p_c} \right]^{-1} . \tag{16}
$$

The laboratory differential cross-section depends on the scattering angles of the nuclei C and c and the energy E_C . Again, these quantities, together with the beam energy, specify the kinematical conditions completely. The nuclei A and a can be projectile and target or vice versa. In case of particles with spin appropriate averages over initial states and sums over final states have to be considered.

2.2. Approximations of the T-matrix element

The T-matrix element T_f in (13) and (15) contains all the essential information relevant to the scattering process. It has to be calculated in a suitable approximation that allows to find the connection to the cross-section of the astrophysical relevant reaction (2).

Before we discuss the various approximations in detail, it is useful to give an overview of the different steps in the derivation of the final formulas. Only the distorted wave Born approximation of the exact T-matrix element and the so-called surface approximation are essential for the Trojan–Horse method. The DWBA is a well-established theoretical framework for the description of direct reactions [32]. Using the surface approximation the cross-section of the three-body reaction can be related to the S-matrix elements of the two-body reaction and it becomes possible to extract the two-body cross-section. However, the expression for the T-matrix element in this general TH approximation is rather complicated and the calculation requires considerable numerical efforts. Additionally, the special features of the THM are not easily seen and the connection to the PWIA (see Section 1) is not obvious. We introduce some additional approximations that lead to simpler formulas which give more insight into the mechanism of the THM without loosing the essential points. In a local momentum approximation, that is exact for plane waves, the T-matrix element can be written as a convolution of two factors, a momentum amplitude that is related to the ground state wave function of the Trojan–Horse and an overlap matrix element with the two-body scattering wave function. In a further plane wave approximation the three-body T-matrix element factorizes in momentum amplitude and a modified two-body scattering amplitude; a form that resembles the PWIA.

The Trojan–Horse reaction (3) has the form of a usual two-body reaction

$$
A + a \to B + b \tag{17}
$$

if the $C + c$ system is considered as an excited state of the compound system B. The exact T-matrix element for this reaction is given in the post-form description by

$$
T_{fi} = \left\langle \exp\left(i\vec{k}_{Bb}\cdot\vec{r}_{Bb}\right)\phi_B\phi_b|V_{Bb}|\Psi_{Aa}^{(+)}(\vec{k}_{Aa},\vec{r}_{Aa})\right\rangle, \tag{18}
$$

with the exact scattering wave function $\Psi_{Aa}^{(+)}(\vec{k}_{Aa},\vec{r}_{Aa})$ in the initial state and the potential V_{Bb} between B and b in the final state [30]. The wavefunction ϕ_i of a nucleus i depends only on internal coordinates which are not given explicitly. The relative motion of B and b is described by a plane wave with momentum $\hbar k_{bb}$. Applying the Gell-Mann Goldberger transformation [33], the T-matrix element (18) assumes the form

$$
T_{fi} = \left\langle \chi_{Bb}^{(-)}(\vec{k}_{Bb}, \vec{r}_{Bb}) \phi_B \phi_b | V_{Bb} - U_{Bb} | \Psi_{Aa}^{(+)}(\vec{k}_{Aa}, \vec{r}_{Aa}) \right\rangle, \tag{19}
$$

which is still an exact relation. For a detailed derivation see [29]. Here, $\chi_{bb}^{(-)}(\vec{k}_{bb}, \vec{r}_{bb})$ is a distorted wave for the $B-b$ relative motion generated by a suitably chosen optical potential U_{Bb} that only depends on \vec{r}_{Bb} but not on internal coordinates in contrast to V_{Bb} . In the THM the wave function ϕ_B does not describe a bound state but a complete scattering state

$$
\phi_B = \Psi_{Cc}^{(-)}(\vec{k}_{Cc}, \vec{r}_{Cc}), \tag{20}
$$

since the system $B = C + c$ is in the continuum.

In the post-form distorted-wave Born approximation (DWBA) the T-matrix element (19) is replaced by

$$
T_{fi}^{\text{DWBA}}(\vec{k}_{Cc}, \vec{k}_{Bb}; \vec{k}_{Aa}) = \left\langle \chi_{Bb}^{(-)}(\vec{k}_{Bb}, \vec{r}_{Bb}) \Psi_{Cc}^{(-)}(\vec{k}_{Cc}, \vec{r}_{Cc}) \phi_b | V_{xb} | \chi_{Aa}^{(+)}(\vec{k}_{Aa}, \vec{r}_{Aa}) \phi_A \phi_a \right\rangle, \tag{21}
$$

where the exact wave function $\Psi_{Aa}^{(+)}(\vec{k}_{Aa},\vec{r}_{Aa})$ is replaced by a distorted wave $\chi_{Aa}^{(+)}(\vec{k}_{Aa},\vec{r}_{Aa})$ for the A–a relative motion. Additionally, the potential $V_{Bb}-U_{Bb}$ is approximated by V_{xb} [29,32]. This is the usual starting point of T-matrix calculations for direct transfer reactions.

In the next step, the so-called surface approximation is applied which is essential to the THM. For small distances between the colliding nuclei the optical potentials are usually strongly absorptive and only reactions at the surface of the nuclei contribute significantly to the matrix element. Therefore the full scattering wave function $\Psi_{Ce}^{(-)}(\vec{k}_{Ce},\vec{r}_{Ce})$ can be replaced by its asymptotic form for radii larger than a suitably chosen cutoff radius R that is larger than the range of the nuclear potential. It is typically in the range of the sum of the radii of the two colliding nuclei. The interior part of $\Psi_{Cc}^{(-)}(\vec{k}_{Cc},\vec{r}_{Cc})$ is set to zero. The validity of the surface approximation was checked in [34]; e.g., it was found to be quite good for the (d, p) reaction at $E_d = 26$ MeV. The asymptotic form of the scattering wave function for $r_a > R$, where $\alpha = Ax, Cc, \ldots$ is a possible partition of the system B, is given by

$$
\Psi_{Cc,\text{asym}}^{(\pm)}(\vec{k}_{Cc},\vec{r}_{Cc})=\frac{4\pi}{k_{Cc}}\sum_{\alpha}\sum_{lm}\frac{\xi_l^{(\pm)}(\alpha)}{r_{\alpha}}i^lY_{lm}(\hat{r}_{\alpha})Y_{lm}^*(\hat{k}_{Cc})\phi_{\alpha},\qquad(22)
$$

with radial wave functions

$$
\xi_l^{(+)}(\alpha) = \frac{1}{2i} \sqrt{\frac{v_{Cc}}{v_{\alpha}}} \Big[S_{\alpha Cc}^l u_l^{(+)}(\eta_{\alpha}; k_{\alpha} r_{\alpha}) - \delta_{\alpha Cc} u_l^{(-)}(\eta_{\alpha}; k_{\alpha} r_{\alpha}) \Big], \tag{23}
$$

$$
\xi_l^{(-)}(\alpha) = \xi_l^{(+)*}(\alpha),\tag{24}
$$

where the Coulomb wave functions [35]

$$
u_l^{(\pm)}(\eta;kr) = e^{\mp i\sigma_l(\eta)}[G_l(\eta;kr) \pm iF_l(\eta;kr)] \rightarrow \exp\left[\pm i\left(kr - \eta \ln(2kr) - \frac{l\pi}{2}\right)\right]
$$
\n(25)

appear. The Sommerfeld parameter

$$
\eta_{ij} = \frac{Z_i Z_j e^2}{\hbar v_{ij}}\tag{26}
$$

depends on the charge numbers Z_i, Z_j of nuclei i, j and their relative velocity $v_{ij} = p_{ij}/\mu_{ij}$ in the partition $\alpha = ij$. The Coulomb phase shifts are given by $\sigma_l(\eta) = \arg \Gamma(1 + l + i\eta)$. The S-matrix elements $S_{\alpha C_c}^l$ in the radial wave functions completely describe the two-body scattering process. They are on-shell quantities and the momenta k_{α} are derived from energy-conservation in the two-body reaction. E.g., in the $\alpha = Ax$ partition we have $k_{Ax} = \sqrt{2\mu_{Ax}E_{Ax}}/\hbar$ with the energy E_{Ax} from relation (9).

The potential V_{xb} appearing in the T-matrix element (21) describes the interaction between x and b in the Trojan–Horse a. Assuming a simple cluster picture of a and neglecting contributions from excited states of x and b or other partitions in the ground state, the momentum amplitude $W(\vec{q})$ of the product

$$
V_{xb}(\vec{r}_{xb})\phi_a(\vec{r}_{xb}) = \int \frac{\mathrm{d}^3q}{(2\pi)^3} W(\vec{q}) \exp(\mathrm{i}\vec{q}\cdot\vec{r}_{xb})\phi_x\phi_b \tag{27}
$$

can be introduced and the T-matrix element assumes the form

$$
T_{fi}^{\text{TH}}(\vec{k}_{Cc}, \vec{k}_{Bb}; \vec{k}_{Aa}) = \int \frac{\text{d}^3 q}{(2\pi)^3} W(\vec{q}) \Big\langle \chi_{Bb}^{(-)}(\vec{k}_{Bb}, \vec{r}_{Bb}) \Psi_{Cc, \text{asym}}^{(-)} \times (\vec{k}_{Cc}, \vec{r}_{Cc}) |\exp(\text{i}\vec{q} \cdot \vec{r}_{xb}) \chi_{Aa}^{(+)}(\vec{k}_{Aa}, \vec{r}_{Aa}) \phi_A \phi_x \Big\rangle.
$$
 (28)

The assumption $W \equiv \text{const.}$ corresponds to the zero-range approximation which is frequently employed in the calculation of DWBA T-matrix elements. In this case, the integration over \vec{q} leads to a δ -function in the variable \vec{r}_{xb} and the actual dimension of the appearing integral is reduced from six to three. However, in general the full momentum dependence of the amplitude W has to be considered. With the help of the Schrödinger equation the momentum amplitude W is related by

$$
W(\vec{q}) = -\left(\varepsilon_a + \frac{\hbar^2 q^2}{2\mu_{xb}}\right) \Phi_a(\vec{q})
$$
\n(29)

to the ground state momentum wave function

$$
\Phi_a(\vec{q}) = \left\langle \exp(i\vec{q} \cdot \vec{r}_{xb}) \phi_x \phi_b | \phi_a(\vec{r}_{xb}) \right\rangle \tag{30}
$$

of the nucleus a. The energy $\varepsilon_a = (m_x + m_b - m_a)c^2 = Q_2 - Q_3 > 0$ is the binding energy of a with respect to the $x + b$ threshold.

The integration over the internal coordinates in the matrix element (28) selects the $\alpha = Ax$ partition of the asymptotic wave function (22) and the T-matrix element in the TH approximation can be written as

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$$
T_{fi}^{\text{TH}}(\vec{k}_{Cc}, \vec{k}_{Bb}; \vec{k}_{Aa}) = \frac{1}{2ik_{Cc}} \sqrt{\frac{v_{Cc}}{v_{Ax}}} \sum_{l} (2l+1) \left[S_{AxCc}^{l} U_{l}^{(+)}(\vec{k}_{Bb}, \vec{k}_{Cc}, \vec{k}_{Aa}) - \delta_{AxCc} U_{l}^{(-)}(\vec{k}_{Bb}, \vec{k}_{Cc}, \vec{k}_{Aa}) \right].
$$
\n(31)

It resembles a scattering amplitude of the reaction $C + c \rightarrow A + x$ except for the functions

$$
U_{l}^{(\pm)}(\vec{k}_{Bb}, \vec{k}_{Cc}, \vec{k}_{Aa}) = \frac{4\pi(-i)^{l}}{(2l+1)} \sum_{m} Y_{lm}(\hat{k}_{Cc})
$$

$$
\times \int \frac{d^{3}q}{(2\pi)^{3}} W(\vec{q}) \left\langle \chi_{Bb}^{(-)}(\vec{k}_{Bb}, \vec{r}_{Bb}) \frac{\theta(r_{Ax} - R)}{r_{Ax}} u_{l}^{(\mp)}(\eta_{Ax}; k_{Ax}r_{Ax}) \right\rangle
$$

$$
\times Y_{lm}(\hat{r}_{Ax}) |\exp(i\vec{q} \cdot \vec{r}_{xb}) \chi_{Aa}^{(+)}(\vec{k}_{Aa}, \vec{r}_{Aa}) \right\rangle
$$
(32)

that describe the angular distribution and the momentum dependence due to the presence of the spectator b in the reaction (3). The unit step function θ accounts for the surface approximation by eliminating the contributions at small radii. With Eq. (31) the relation between the cross-section (15) of the three-body reaction (3) and the S-matrix elements of the two-body reaction is directly established. It is possible to extract the energy dependence of the S-matrix elements from experimental threebody cross-sections, at least in principle. The dependence of $U_l^{(\pm)}$ on k_{C_c} or equivalently k_{Ax} together with the energy dependence of the two-body S-matrix elements S_{AxCc} leads to a finite cross-section of the three-body reaction (3) even when the threshold of the two-body reaction (2) is reached. This is the essential feature of the TH method. However this is not readily seen from the general expression (32). A simplified formulation allows to study the energy dependence more explicitly. The main problem is the technical evaluation of the appearing matrix element in Eq. (32) with scattering wave functions in the initial and final states. Even with the surface approximation, it involves a six-dimensional integration in the Jacobi-coordinates which is an extensive computational task. For an actual application of the TH method, it is convenient to introduce additional approximations that lead to a formulation with a direct relation between the cross-sections of the two-body and threebody reactions similar to the PWIA.

It is useful to expand the distorted wave in the initial state in a Taylor series

$$
\chi_{Aa}^{(+)}(\vec{k}_{Aa}, \vec{r}_{Aa}) = \left(1 + (\vec{r}_{Aa} - \vec{r}_{Ax}) \cdot \vec{\nabla} + \cdots \right) \chi_{Aa}^{(+)}(\vec{k}_{Aa}, \vec{r}_{Ax}) \n\approx \exp\left[i(\vec{r}_{Aa} - \vec{r}_{Ax}) \cdot \vec{K}\right] \chi_{Aa}^{(+)}(\vec{k}_{Aa}, \vec{r}_{Ax}),
$$
\n(33)

where the wave vector \vec{K} replaces the derivative $-i\vec{\nabla}$ with respect to the spatial coordinates. In the so called local momentum approximation [36,37] the modulus of \vec{K} is determined by the actual kinetic energy of the $A - a$ relative motion at a certain distance R_{Aa} . The direction of \vec{K} is assumed to be same as the asymptotic momentum distance R_{Aa} . The direction of \vec{K} is assumed to be same as the asymptotic momentum \vec{k}_{Aa} . If $\chi_{Aa}^{(+)}$ is a plane wave and $\vec{K} = \vec{k}_{Aa}$ the relation (33) is obviously exact. Considering the relations

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$$
\vec{r}_{xb} = \vec{r}_{Bb} - \alpha \vec{r}_{Ax}, \quad \vec{r}_{Aa} = \beta \vec{r}_{Bb} + (1 - \alpha \beta) \vec{r}_{Ax}
$$
\n(34)

of the Jacobi coordinates with factors

$$
\alpha = \frac{m_A}{m_A + m_x}, \quad \beta = \frac{m_b}{m_b + m_x},\tag{35}
$$

the matrix element in Eq. (28) factorizes and the T-matrix element in the TH approximation assumes the form

$$
T_{\hat{j}}^{\text{TH}}(\vec{k}_{Cc}, \vec{k}_{Bb}; \vec{k}_{Aa}) = \int \frac{d^3q}{(2\pi)^3} \mathcal{W}(\vec{q}, \vec{k}_{Bb}) \mathcal{M}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}), \qquad (36)
$$

with the generalized momentum amplitude

$$
\mathcal{W}(\vec{q}, \vec{k}_{Bb}) = W(\vec{q}) \Big\langle \chi_{Bb}^{(-)}(\vec{k}_{Bb}, \vec{r}_{Bb}) | \exp(i\vec{Q} \cdot \vec{r}_{Bb}) \Big\rangle \tag{37}
$$

and the matrix element

$$
\mathcal{M}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}) = \left\langle \Psi_{Cc, \text{asym}}^{(-)}(\vec{k}_{Cc}, \vec{r}_{Cc}) | \exp(-i\alpha \vec{Q} \cdot \vec{r}_{Ax}) \chi_{Aa}^{(+)}(\vec{k}_{Aa}, \vec{r}_{Ax}) \phi_A \phi_x \right\rangle, \tag{38}
$$

with $\vec{Q} = \vec{q} + \beta \vec{K}$. The integration over the internal coordinates leads to the expression

$$
\mathcal{M}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}) = \frac{1}{2ik_{Cc}} \sqrt{\frac{v_{Cc}}{v_{Ax}}} \sum_{l} (2l+1) \left[S_{AxCc}^l \mathcal{U}_l^{(+)}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}) - \delta_{AxCc} \mathcal{U}_l^{(-)}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}) \right],
$$
\n(39)

with the integrals

$$
\mathcal{U}_{l}^{(\pm)}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}) = \frac{4\pi(-i)^{l}}{(2l+1)} \int d^{3}r_{Ax} \frac{\theta(r_{Ax} - R)}{r_{Ax}} u_{l}^{(\pm)}(\eta_{Ax}; k_{Ax}r_{Ax}) \exp(-i\alpha \vec{Q} \cdot \vec{r}_{Ax}) \times \chi_{Aa}^{(\pm)}(\vec{k}_{Aa}, \vec{r}_{Ax}) \sum_{m} Y_{lm}^{*}(\hat{k}_{Ax}) Y_{lm}(\hat{k}_{Cc}).
$$
\n(40)

Then the functions $U_l^{(\pm)}$ in the TH T-matrix element (31) assume the form

$$
U_l^{(\pm)}(\vec{k}_{Bb}, \vec{k}_{Cc}, \vec{k}_{Aa}) = \int \frac{d^3q}{(2\pi)^3} \mathcal{W}(\vec{q}, \vec{k}_{Bb}) \mathcal{U}_l^{(\pm)}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}) \tag{41}
$$

of a convolusion of a momentum amplitude and an overlap matrix element with scattering wave functions.

2.3. Plane wave approximations

If $\chi_{bb}^{(-)}$ is assumed to be a plane wave, the matrix element in the generalized momentum distribution (37) leads to a δ -function in the variable \vec{q}

$$
W(\vec{q}, \vec{k}_{bb}) = W(\vec{q}) (2\pi)^3 \delta(\vec{q} - \vec{Q}_{bb})
$$
\n(42)

with

$$
\vec{Q}_{Bb} = \vec{k}_{Bb} - \beta \vec{K}.\tag{43}
$$

Then, the \vec{q} -integration is trivial; the T-matrix element (36) factorizes explicitly into a momentum amplitude and the matrix element (39) where the integrals (40) are evaluated for $\vec{q} = \vec{Q}_{Bb}$. This plane-wave approximation for $\chi_{Bb}^{(-)}$ is justified when the energy of the Bb relative motion is large. When the spectator \ddot{b} is a neutron, it is also often possible to replace $\chi_{bb}^{(-)}$ by a plane wave. On the other hand, when $\chi_{bb}^{(-)}$ is replaced by a pure Coulomb scattering wave, e.g., if the Bb relative energy is small, the matrix element in (37) can be calculated explicitly (see Appendix D). However, in applications of the THM the c.m. energy in the Bb system is generally large because of the large projectile energy and Q-value of the reaction and the latter case is not relevant.

Independently from the treatment of the wave function $\chi_{bb}^{(-)}$, a plane wave approximation can be introduced for $\chi_{Aa}^{(+)}$ in the matrix element (38). In this case one finds

$$
\mathcal{M}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}) = \left\langle \Psi_{Cc, \text{asym}}^{(-)}(\vec{k}_{Cc}, \vec{r}_{Cc}) | \exp(i\vec{Q}_{Aa} \cdot \vec{r}_{Ax}) \phi_A \phi_x \right\rangle \tag{44}
$$

with

$$
\vec{Q}_{Aa} = \vec{k}_{Aa} - \alpha \beta \vec{K} - \alpha \vec{q}.\tag{45}
$$

The quantities (40) reduce to

$$
\mathcal{U}_l^{(\pm)}(\vec{q}, \vec{k}_{Cc}, \vec{k}_{Aa}) = \frac{4\pi}{k_{Ax}Q_{Aa}} P_l(\hat{Q}_{Aa} \cdot \hat{k}_{Cc}) J_l^{(\pm)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}), \tag{46}
$$

with Legendre polynomials P_l and the dimensionless Trojan–Horse integrals

$$
J_l^{(\pm)}(R,\eta_{Ax},k_{Ax},Q_{Aa})=k_{Ax}Q_{Aa}\int_R^{\infty}dr_{Ax}u_l^{(\pm)}(\eta_{Ax};k_{Ax}r_{Ax})r_{Ax}j_l(Q_{Aa}r_{Ax}).
$$
\n(47)

The Coulomb wave functions $u_l^{(\pm)}$ as defined in (25) and the regular spherical Bessel function j_l of order *l* appear in the radial integral. It is convenient to introduce the decomposition

$$
J_{l}^{(\pm)}(R,\eta_{Ax},k_{Ax},Q_{Aa}) = e^{\mp i\sigma_{l}(\eta_{Ax})} \Big[J_{l}^{(G)}(R,\eta_{Ax},k_{Ax},Q_{Aa}) \pm iJ_{l}^{(F)}(R,\eta_{Ax},k_{Ax},Q_{Aa}) \Big]
$$
(48)

of the TH integrals into contributions with the regular and irregular Coulomb wave functions. The integrals are discussed in detail in Section 3.

In the following only the plane-wave approximation for both $\chi_{Aa}^{(+)}$ and $\chi_{Bb}^{(-)}$ will be considered. In this case the T-matrix element has the simple product form

$$
T_{fi}^{\text{TH}}(\vec{k}_{Cc}, \vec{k}_{Bb}; \vec{k}_{Aa}) = W(\vec{Q}_{Bb}) \Big\langle \Psi_{Cc, \text{asym}}^{(-)}(\vec{k}_{Cc}, \vec{r}_{Cc}) | \exp(i\vec{Q}_{Aa} \cdot \vec{r}_{Ax}) \phi_A \phi_x \Big\rangle. \tag{49}
$$

This approximation already contains the essential ingredients to see the principles of the TH method and the connection to the PWIA becomes clear. Generalizations to a more general treatment with distorted waves are obvious from the above. The argument of the amplitude W is the momentum

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$$
\vec{Q}_{Bb} = \vec{k}_{Bb} - \frac{m_b}{m_b + m_x} \vec{k}_{Aa} \tag{50}
$$

assuming that $\vec{K} = \vec{k}_{Aa}$. Neglecting the Fermi motion of b inside the Trojan–Horse, $m_b \vec{k}_{Aa}/(m_b + m_x)$ is the momentum of b relative to A in the initial state and \vec{k}_{Bb} is the momentum of b relative to $B = C + c$ in the final state. Thus $-\vec{Q}_{bb}$ corresponds to the momentum transfer to the spectator b. The amplitude W describes the distribution of the transferred momentum due to the Fermi motion. Similarly, the momentum

$$
\vec{Q}_{Aa} = \vec{k}_{Aa} - \frac{m_A}{m_A + m_x} \vec{k}_{Bb} \tag{51}
$$

in the argument of the plane wave can be considered as the (negative) momentum transfer to nucleus A (independent of the choice of \vec{K}) by the particle x. In the case of a infinitely heavy nucleus A , Eqs. (50) and (51) reduce to

$$
\vec{Q}_{Bb} = \frac{m_b}{m_b + m_x} \vec{k}_a - \vec{k}_b = \vec{k}_x - \frac{m_x}{m_b + m_x} \vec{k}_a,
$$
\n(52)

$$
\vec{Q}_{Aa} = \vec{k}_b - \vec{k}_a = -\vec{k}_x \tag{53}
$$

(cf. Fig. 1) and the interpretation becomes simpler. The main task is to calculate the TH integrals (47) and to find the explicit relation of the three-body cross-section to the two-body cross-section.

The plane wave approximations for $\chi_{Aa}^{(+)}$ and $\chi_{Bb}^{(-)}$ seem to be crude at first sight but the TH matrix element (49) still contains the asymptotically correct wave function $\Psi_{\text{asym}}^{(-)}(Cc,\vec{k}_{Cc})$ for the $C + c$ system and thus the complete information on the twobody scattering process. This is in clear contrast to the PWIA [19] where the effect of the Coulomb barrier in $A + x$ system on the energy dependence is not obvious. The absorptive feature of the optical potentials is taken into account by the surface approximation. Additionally, the distorted wave $\chi_{Aa}^{(+)}$ describes a scattering state with a much higher (and constant) momentum as compared to the $A + x$ system and only a small part of the three-body phase space is of interest in the reaction. The absolute cross-section for the three-body process (3) calculated in the plane wave approximation might be different from the actual value but the energy dependence with respect to the two-body reaction is expected to be treated correctly.

3. Trojan–Horse integrals

The main difference between the usual expressions for the scattering amplitude for $A + x \rightarrow C + c$ scattering and the corresponding T-matrix element (31) in the TH method is the appearance of the functions $U_l^{(\pm)}$. They describe a modification of the two-body scattering amplitude and lead to the effective removal of the Coulomb barrier in the quasi-free scattering process. This effect can only be understood if the energy dependence of these functions is known. In the plane wave approximation the discussion becomes very transparent. The functions

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$$
U_l^{(\pm)}(\vec{k}_{Bb}, \vec{k}_{Cc}, \vec{k}_{Aa}) = \frac{4\pi}{k_{Ax}Q_{Aa}} W(\vec{Q}_{Bb}) P_l(\hat{Q}_{Aa} \cdot \hat{k}_{Cc}) J_l^{(\pm)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa})
$$
(54)

in Eq. (31) factorize into a momentum amplitude, Legendre polynomials and the TH integrals (47) which we will study now.The basic Trojan–Horse integral is given by

$$
J_l^{(H)}(R,\eta,k,Q) = k \int_R^{\infty} dr H_l(\eta;kr) z_l(Qr), \qquad (55)
$$

where H_l is a Coulomb wave function F_l or G_l and $z_l(x) = x_j(x)$ is a Riccati–Bessel function. Then the integrals $J_l^{(\pm)}$ are obtained from Eq. (48). The TH integrals do not converge in the usual sense, since H_l and z_l oscillate with constant amplitude if r goes to infinity. Convergence is achieved only in the distributional sense after an integration over \hat{O} with a suitable test function. The problem is caused by the fact that the matrix element in Eq. (28) contains the overlap of continuum wave functions. A similar problem occurs in the case of Bremsstrahlung matrix elements that have to be regularized.

We transform the integral (55) into a form that allows a numerical calculation. Details can be found in Appendix A. By a simple change of variables it is seen that the TH integrals (55) depend only on three independent parameters. With

$$
\xi = \frac{k}{Q} \quad \text{and} \quad x_0 = QR,
$$
\n(56)

we can write

$$
J_l^{(H)}(R,\eta,k,Q) = J_{l,\text{red}}^{(H)}(x_0,\eta,\xi) = \xi \int_{x_0}^{\infty} dx H_l(\eta;\xi x) z_l(x), \qquad (57)
$$

with reduced TH integrals which are functions of x_0 , η , and ξ . These integrals exhibit a particular energy-dependence close to the threshold of the two-body reaction. For $R = 0$, i.e., $x_0 = 0$, the energy dependence is easily extracted and leads to a simple scaling behaviour for $\xi \to 0$ and $\eta \to \infty$. The reduced TH integral with the regular Coulomb wave function F_l shows an energy dependence according to

$$
J_{I,\text{red}}^{(F)}(0,\eta,\xi) \propto \xi^{3/2} \exp(-\pi \eta). \tag{58}
$$

In the case of the TH integral with the irregular Coulomb wave function we find

$$
J_{I,\text{red}}^{(G)}(0,\eta,\xi) \propto \xi^{3/2} \exp(\pi \eta). \tag{59}
$$

Fig. 3 shows the dependence of the scaled reduced TH integrals $J_{l,\text{red}}^{(F)}(0, \eta, \xi)\xi^{-3/2}$ $exp(\pi \eta)$ with $l = 0, 1$ on $\xi = k/Q$ for various values of $\eta \xi = Z_A Z_x e^2 \mu_{Ax}/(\hbar^2 Q)$ which is a constant for given O. The scaled TH integrals become constant for small ζ as expected. The rapid decrease of the unscaled TH integrals with decreasing ξ for constant $\eta \xi$ is obvious. The angular momentum barrier leads to a smaller TH integral with larger l for the same parameters ξ and $\eta \xi$. This effect is more pronounced for small $\eta \xi$ than for large $\eta \xi$ where the Coulomb barrier dominates the ξ dependence. The approximations $(A.20)$ and $(A.21)$ agree very well with the exact results

Fig. 3. Scaled reduced Trojan–Horse integrals $-J_{l, \text{red}}^{(F)}(0, \eta, \xi) \xi^{-3/2} \exp(\pi \eta)$ for parameters $\eta \xi = 0.01, \ldots,$ 1.0 as a function of $\zeta = k/Q$ for angular momenta $l = 0$ (a) and $l = 1$ (b). The solid lines are the exact results and the dotted lines are the approximations of Section A.2.

for the TH integrals with $l = 0, 1$ for large $\eta \xi$ and not too large ξ . The agreement is less satisfactory for small $\eta \xi$ and larger l as one may expect, since then the angular momentum barrier is getting more important as compared to the Coulomb barrier.

In Fig. 4 the dependence of scaled reduced TH integrals $J_{l, \text{red}}^{(G)}(0, \eta, \xi) \xi^{-3/2}$ $\exp(-\pi \eta)$ with $l = 0, 1$ on ξ is shown for the same range of the parameter $\eta \xi$ as in Fig. 3. For small ξ the scaled TH integrals again become constant. The unscaled TH integrals increase dramatically for decreasing ζ for constant $\eta \zeta$ and for increasing $\eta \xi$ for constant ξ , respectively. The effect of the angular momentum barrier is more pronounced for small $\eta \xi$ as in the case of the TH integrals with the regular Coulomb wave function as one may expect. The agreement between the approximations (A.23) and (A.24) and the exact TH integrals shows the same trends as in the case of the TH integrals with the regular Coulomb wave functions.

In the case of finite R the energy dependence of the TH integrals remains essentially unchanged for $\xi \to 0$, only the magnitude of the integrals is modified. At low energies, both the TH integrals $J_l^{(\pm)}(R,\eta,k,Q)$ are dominated by the energy dependence of $J_l^{(G)}(R, \eta, k, Q)$.

Fig. 4. Scaled reduced Trojan–Horse integrals $J_{l, \text{red}}^{(G)}(0, \eta, \xi) \xi^{-3/2} \exp(-\pi \eta)$ for parameters $\eta \xi = 0.01, \ldots,$ 1.0 as a function of $\zeta = k/Q$ for angular momenta $l = 0$ (a) and $l = 1$ (b). The solid lines are the exact results and the dotted lines are the approximations of Section A.2.

4. Cross-sections for spinless particles

In order to establish a closer connection of the three- and two-body cross-sections the Trojan–Horse T-matrix element (49) is recast in the form

$$
T_{fi}^{\text{TH}}(\vec{k}_{Cc}, \vec{k}_{Bb}; \vec{k}_{Aa}) = \frac{4\pi}{k_{Ax}Q_{Aa}} W(\vec{Q}_{Bb}) f^{\text{TH}}(\vec{Q}_{Aa}, \vec{k}_{Cc}),
$$
(60)

with the Trojan–Horse scattering amplitude

$$
f^{\text{TH}}(\vec{Q}_{Aa}, \vec{k}_{Cc}) = \frac{k_{Ax}\mathcal{Q}_{Aa}}{4\pi} \left\langle \Psi_{Cc,\text{asym}}^{(-)}(\vec{k}_{Cc}, \vec{r}_{Cc}) | \exp(i\vec{Q}_{Aa} \cdot \vec{r}_{Ax}) \phi_A \phi_x \right\rangle
$$

=
$$
\frac{1}{2i k_{Cc}} \sum_{l} (2l+1) P_l(\hat{k}_{Cc} \cdot \hat{Q}_{Aa}) \sqrt{\frac{v_{Cc}}{v_{Ax}}} [S^l_{AxCc} J^{(+)}_l(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) - \delta_{AxCc} J^{(-)}_l(R, \eta_{Ax}, k_{Ax}, Q_{Aa})].
$$
 (61)

This amplitude is valid for particles without spin. The case with spin in dealt with in Appendix C. The amplitude (61) is related to the scattering amplitude for the twobody reaction

$$
C + c \to A + x. \tag{62}
$$

The essential difference is the appearance of the Trojan–Horse integrals (47). The argument of the Legendre polynomial P_l in Eq. (61) corresponds to the cosine of the scattering angle $\vartheta_{\rm cm}$. It is convenient to decompose the TH amplitude

$$
f^{\text{TH}} = f_N^{\text{TH}} + f_C^{\text{TH}} \delta_{AxCc}
$$
\n
$$
\tag{63}
$$

into a nuclear part and a purely Coulomb part. The nuclear contribution

$$
f_N^{\text{TH}} = \frac{1}{2ik_{Cc}} \sum_{l} (2l+1) P_l(\hat{k}_{Cc} \cdot \hat{Q}_{Aa}) \sqrt{\frac{v_{Cc}}{v_{Ax}}} T_{AxCc}^l J_l^{(+)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \tag{64}
$$

with

$$
T_{AxCc}^l = e^{i\sigma_l(\eta_{Ax}) + i\sigma_l(\eta_{Cc})} \left[S_{AxCc}^{NI} - \delta_{AxCc} \right]
$$
 (65)

depends on the nuclear S-matrix element

$$
S_{AxCe}^{NI} = e^{-i\sigma_I(\eta_{Ax})} S_{AxCe}^I e^{-i\sigma_I(\eta_{Ce})}, \qquad (66)
$$

which is obtained from the full S-matrix element by compensating the Coulomb phase shifts $\sigma_l(\eta_{Cc})$ and $\sigma_l(\eta_{Ax})$ in the initial and final states, respectively. The Coulomb contribution

$$
f_C^{\text{TH}} = \frac{1}{k_{Cc}} \sum_l (2l+1) P_l(\hat{k}_{Cc} \cdot \hat{Q}_{Aa}) e^{i\sigma_l(\eta_{Cc})} J_l^{(F)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \tag{67}
$$

appears only if elastic two-body reactions are studied. It depends only on the TH integrals with the regular Coulomb wave function and is discussed in Section 5.

If the Trojan–Horse integrals $J_l^{(\pm)}$ are replaced by one the scattering amplitudes reduce to the standard form for the reaction (62). At low energies there are usually only few contributions to the nuclear scattering amplitude with small angular momenta l due to the increase of the centrifugal barrier with l in the two-body reaction.

The cross-section (15) in the laboratory system can now be expressed in the form

$$
\frac{\mathrm{d}^3 \sigma}{\mathrm{d}E_C \,\mathrm{d}\Omega_C \,\mathrm{d}\Omega_c} = K_{\text{lab}} \left| W(\vec{Q}_{Bb}) \right|^2 \frac{16\pi^2}{k_{Ax}^2 Q_{Aa}^2} \frac{v_{Cc}}{v_{Ax}} \frac{\mathrm{d}\sigma^{\text{TH}}}{\mathrm{d}\Omega},\tag{68}
$$

with the two-body Trojan–Horse cross-section

$$
\frac{\mathrm{d}\sigma^{\mathrm{TH}}}{\mathrm{d}\Omega} = \frac{v_{Ax}}{v_{Cc}} |f^{\mathrm{TH}}|^2. \tag{69}
$$

A corresponding expression holds for the c.m. cross-section (13) with the appropriate kinematical factor. This result has a similar structure as in the PWIA, i.e., a product of a kinematical factor, a momentum distribution and a two-body crosssection. However, the momentum distribution $|W|^2$ is not directly the ground state momentum distribution $|\Phi_a|^2$ of the Trojan–Horse a and the Trojan–Horse cross-section (69) contains explicitly the three-body effects. It is a sum

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$$
\frac{d\sigma^{TH}}{d\Omega} = \left[\frac{d\sigma_C^{TH}}{d\Omega} + \frac{d\sigma_I^{TH}}{d\Omega}\right] \delta_{AxCe} + \frac{d\sigma_N^{TH}}{d\Omega},\tag{70}
$$

of a Coulomb contribution

$$
\frac{\mathrm{d}\sigma_C^{\mathrm{TH}}}{\mathrm{d}\Omega} = |f_C^{\mathrm{TH}}|^2,\tag{71}
$$

a Coulomb-nuclear interference contribution

$$
\frac{\mathrm{d}\sigma_I^{\mathrm{TH}}}{\mathrm{d}\Omega} = \sum_l \frac{2l+1}{k_{Cc}} P_l(\hat{k}_{Cc} \cdot \hat{Q}_{Aa}) \mathrm{Re}\Big[\big(\mathrm{i}f_C^{\mathrm{TH}}\big)^* T_{AxCc}^l J_l^{(+)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \Big],\tag{72}
$$

and a nuclear contribution

$$
\frac{d\sigma_N^{TH}}{d\Omega} = \frac{1}{4k_{Cc}^2} \sum_{\lambda} \sum_{ll'} (2l+1)(2l'+1)(l0l'0|\lambda 0)^2 P_{\lambda}(\hat{k}_{Cc} \cdot \hat{Q}_{Aa})
$$

× $T_{\text{arc}}^l T_{\text{arc}}^{\prime\prime*} J_l^{(+)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) J_r^{(-)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}).$ (73)

Again, the expression for the TH cross-section closely resembles the usual c.m. crosssection for the two-body reaction (62). The appearance of the TH integrals $J_l^{(\pm)}$ accounts for the off-shell effects and the scalar product $\hat{k}_{Cc} \cdot \hat{Q}_{Aa}$ appears as the argument of the Legendre polynomial instead of the cosine of the two-body c.m. scattering angle $\vartheta_{\rm cm}$. The TH Coulomb scattering amplitude (67) replaces the onshell Coulomb scattering amplitude

$$
f_C = \frac{1}{2ik_{Cc}} \sum_{l} (2l+1) P_l(\cos \vartheta_{\text{c.m.}}) (\exp[2i\sigma_l(\eta_{Cc})] - 1)
$$

=
$$
-\frac{\eta_{Cc}}{2k_{Cc}\sin^2(\vartheta_{\text{c.m.}}/2)} \exp\left[2i\sigma_0(\eta_{Cc}) - 2i\eta_{Cc}\ln\sin\frac{\vartheta_{\text{c.m.}}}{2}\right]
$$
(74)

in the usual elastic two-body scattering amplitude.

The expression for the TH cross-section simplifies considerably in special cases. If $Ax \neq Cc$ and only one partial wave l_{Ax} contributes to the cross-section the Trojan– Horse cross-section is given by

$$
\frac{d\sigma^{TH}}{d\Omega} = \frac{d\sigma_l}{d\Omega} (Cc \to Ax) P_l(R, \eta_{Ax}, k_{Ax}, Q_{Aa})
$$
\n(75)

with the usual partial on-shell cross-section

$$
\frac{d\sigma_l}{d\Omega}(Cc \to Ax) = \frac{(2l+1)^2}{4k_{Cc}^2} |T_{\text{AxCc}}^l|^2 \sum_{\lambda} (l0l0|\lambda 0)^2 P_{\lambda}(\hat{k}_{Cc} \cdot \hat{Q}_{Aa}) \tag{76}
$$

for the two-body reaction (62) and the penetrability factor

$$
P_l(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) = \left| J_l^{(+)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \right|^2.
$$
 (77)

In the simple approximation $(A.3)$ of Appendix A it is given by

$$
P_l(R, \eta, k, Q) \to k^2 R^2 [F_l^2(\eta; kR) + G_l^2(\eta; kR)] z_l^2(QR).
$$
 (78)

From the k-dependence of the TH integrals or the Coulomb wave functions a k^3 exp $(2\pi\eta)$ behaviour of the penetrability factor is found for $k \to 0$. For the transfer of a neutron the dependence of P_l on small k is given by k^{2-2l} , see Eq. (A.30), with an l-dependence determined by the centrifugal barrier.

With the theorem of detailed balance one obtains the direct relation

$$
\frac{\mathrm{d}^3 \sigma}{\mathrm{d}E_C \,\mathrm{d}\Omega_C \,\mathrm{d}\Omega_c} = K_{\text{lab}} |W(\vec{Q}_{Bb})|^2 \frac{16\pi^2}{k_{CC}^2 Q_{Aa}^2} \frac{v_{Cc}}{v_{Ax}} \frac{\mathrm{d}\sigma_l}{\mathrm{d}\Omega} (Ax \to Cc) P_l(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \tag{79}
$$

of the three-body cross-section to the two-body cross-section for partial wave l. This equation shows clearly the "parallelism" of the $A + x$ reaction and the Trojan–Horse reaction: The cross-section (79) is proportional to the cross-section for the $A + x$ reaction, modulated by the penetrability factor P_l . The factor P_l is directly related to the TH integrals (77). It leads in general to an enhancement of the higher partial waves and it contains the $\exp(2\pi\eta)$ factor.

A most convincing beautiful example of this parallelism is the comparison of neutron elastic scattering and the (d, p) reaction on ¹⁵N in the same energy range of the continuum in 16 N. There are the same peaks in both spectra, changed in magnitude according to the factor P_l ($\eta = 0$ in the neutron transfer case). Also the s-wave resonance which appears as a destructive interference with the $l = 0$ continuum shows up nicely in the (d, p) spectrum [15].

5. Trojan–Horse Coulomb scattering amplitude

In case of an inelastic two-body reaction only the nuclear contribution to the TH scattering amplitude is relevant. For the elastic scattering with $Ax = Cc$ also the TH Coulomb scattering amplitude (67) contributes to the total TH scattering amplitude. It depends only on the TH integrals $J_l^{(F)}$ with the regular Coulomb wave functions. The summation over l in Eq. (67) poses no serious problem but it is convenient to reformulate the expression as in the case of the TH integrals in order to see the reduction of the Coulomb barrier more clearly. For that purpose the TH Coulomb scattering amplitude is written as a difference

$$
f_C^{\text{TH}} = f_{C0}^{\text{TH}} - f_{CR}^{\text{TH}} \tag{80}
$$

of a contribution (suppressing the indices of the momenta)

$$
f_{C0}^{\text{TH}} = \frac{2\eta k^2 Q}{k^2 - Q^2} \sum_{l} (2l+1) P_l(\hat{k} \cdot \hat{Q}) e^{i\sigma_l(\eta)} M_{ll}^{-1}(F_l, \eta, k, Q)
$$
(81)

where the cutoff radius R is set to zero and a finite range contribution

$$
f_{CR}^{\text{TH}} = \sum_{l} (2l+1) P_l(\hat{k} \cdot \hat{Q}) e^{i\sigma_l(\eta)} L_l(R, \eta, k, Q)
$$
(82)

with the integrals

$$
L_l(R,\eta,k,Q)=\int_0^R dr F_l(\eta;kr)z_l(Qr).
$$
\n(83)

These are easily calculated numerically and decrease rapidly with increasing *l* leading to a rapid convergence of the sum (82). Similar to relation (61) for the full TH scattering amplitude, the contribution (81) to the TH Coulomb scattering amplitude can be expressed as a matrix element

$$
f_{\rm CO}^{\rm TH} = \frac{kQ}{4\pi} \left\langle \Psi_{\rm Coul}^{(-)}(\vec{k}) | \exp(i\vec{Q} \cdot \vec{r}) \right\rangle \tag{84}
$$

with the pure Coulomb scattering wave function

$$
\Psi_{\text{Coul}}^{(\pm)}(\vec{k}) = e^{-(\pi/2)\eta} \Gamma(1 \pm i\eta) \exp(i\vec{k} \cdot \vec{r})_1 F_1(\mp i\eta, 1; \pm i[kr \mp \vec{k} \cdot \vec{r}]) \tag{85}
$$

where ${}_{1}F_{1}$ denotes a confluent hypergeometric function. This form allows an analytical calculation with the result

$$
f_{C0}^{\text{TH}} = -\frac{2\eta k^2 Q}{Q^2 - k^2} \frac{C_0(\eta)}{(\vec{Q} - \vec{k})^2} \exp\left\{i\sigma_0(\eta) + i\eta \ln\left[\frac{Q^2 - k^2}{(\vec{Q} - \vec{k})^2}\right]\right\}
$$
(86)

as explained in Appendix D. Contrary to the usual Coulomb scattering amplitude (74) there is no divergence for $k \to 0$ since Q remains finite. In this limit, the energy dependence of the TH Coulomb scattering amplitude is given by $kC_0(\eta)$. The appearance of the $C_0(\eta)$ factor in the amplitude causes a strong reduction at small k.

6. Applications of the Trojan–Horse method

Several reactions have been studied with the TH method recently. They are listed in Table 1 with ²H and ⁶Li (= α +d) as typical "Trojan–Horses." These nuclei allow to study the transfer of protons, neutrons, deuterons and α -particles, which covers most of the cases of astrophysical interest for the two-body reaction.

6.1. Kinematical conditions

The Trojan–Horses employed so far have a dominant s-wave configuration in their gound state. Their momentum distribution $W(Q_{bb})$ has a maximum around zero. Correspondingly, the equation

Table 1 Projectile energy E_{pro} and corresponding quasi-free energy E_{Ax}^{qf} for pairs of two-body and Trojan–Horse reactions

Two-body reaction	Trojan–Horse reaction	$E_{\rm pro}$ (MeV)	E_{4x}^{qt} (MeV)	Ref.
$^{2}H(^{6}Li, ^{4}He)^{4}He$	${}^{6}\text{Li}({}^{6}\text{Li},{}^{4}\text{He}^{4}\text{He}){}^{4}\text{He}$	6.0	0.029	$[20 - 22]$
$\mathrm{Li}(\mathrm{p}, \mathrm{H}\mathrm{e})^4\mathrm{He}$	$\mathrm{^{2}H(^{7}Li,^{4}He^{4}He)}$	$19.0 - 21.0$	$0.161 - 0.412$	$[23 - 26]$
${}^{12}C(^{4}He, {}^{4}He)^{12}C$	6 Li(¹² C, ⁴ He ¹² C) ² H	$12.0 - 18.0$	$1.527 - 3.027$	[27, 28]
6 Li(p, 3 He) 4 He	$^{2}H(^{6}Li, ^{3}He^{4}He)$	25.0	1.362	[38]

$$
\vec{Q}_{Bb} = 0 \tag{87}
$$

defines the so-called quasi-free condition. In this region of the three-body phase space the cross-section for the quasi-free reaction will reach a maximum. From this condition the corresponding quasi-free c.m. energy

$$
E_{Ax}^{\text{qf}} = E_{Aa} \left(1 - \frac{\mu_{Aa}}{\mu_{Bb}} \frac{\mu_{bx}^2}{m_x^2} \right) - \varepsilon_a \tag{88}
$$

in the initial channel of the two-body reaction (2) is derived from energy conservation (11) assuming the plane wave approximation. It is obvious that even with a large c.m. energy E_{Aa} in the entrance channel of the three-body reaction (3) a small energy E_{Ax} can be reached if a suitable Trojan–Horse *a* is chosen. This is confirmed in Table 1 where the projectile energy in the laboratory system and the quasi-free energy are shown for several reactions.

The relation between E_{Ax}^{qf} and E_{Aa} is purely a kinematical consequence. It is not related to the Fermi motion of particle x inside the Trojan–Horse a which would involve a dependence on the width of the momentum amplitude. However, in an actual experiment a cutoff in the momentum transfer \vec{Q}_{bb} is chosen to select the region where the quasi-free process dominates the cross-section. This procedure limits the range in energies E_{Ax} that are within reach in the experiment for a chosen projectile energy E_{pro} .

In case of the quasi-free condition, all nuclei in the final state of reaction (3) are emitted in the same plane. The momentum of the spectator b is in beam direction which makes it difficult to detect b in the experiment. In the laboratory system the nuclei C and c are emitted under angles ϑ_c and ϑ_c on opposite sides of the beam axis. If the scattering angle $\vartheta_{\rm cm.}$ in the c.m. system of the two-body reaction is given, the angles ϑ_c , ϑ_c , the so-called quasi-free angles, are completely specified for a fixed beam energy from kinematical considerations.

As a consequence, the quasi-free condition determines the setup of the experiment. If a particular two-body reaction (2) is to be studied close to a c.m. energy E_{Ax} and if the Trojan–Horse *a* is selected, then from Eq. (88) the appropriate beam energy can be extracted. Since the momentum amplitude W has a finite width it is possible to study the two-body reaction in a certain energy window around E_{4x} . The c.m. scattering angle $\vartheta_{\rm cm}$ determines the arrangement of the detectors close to the pair of quasi-free angles.

When E_{Bb} is zero in (11) the maximum energy

$$
E_{Ax}^{\max} = E_{Aa} - \varepsilon_a \tag{89}
$$

in the two-body system is reached for a fixed c.m. energy E_{Aa} in the entrance channel of the three-body reaction. In this case $\vec{k}_{Bb} = 0$ and $\vec{Q}_{Aa} = \vec{k}_{Aa}$. The momenta of all nuclei C , c , and b in the final state are parallel to the beam momentum in the laboratory system. Since $\varepsilon_a > 0$ it follows that the relation

$$
k_{Ax} < \sqrt{\frac{\mu_{Ax}}{\mu_{Aa}}} Q_{Aa} < Q_{Aa} \tag{90}
$$

holds in all kinematically allowed regions of the phase space.

6.2. Threshold behaviour of cross-sections

The energy dependence of the two-body cross-section

$$
\frac{d\sigma}{d\Omega}(Ax \to Cc) = \frac{1}{4k_{Ax}^2} \sum_{\lambda} \sum_{ll'} (2l+1)(2l'+1)(l0l'0|\lambda 0)^2 P_{\lambda}(\cos\vartheta) S_{AxCc}^l S_{AxCc}^{l''}
$$
\n(91)

for the inelastic $A + x \rightarrow C + c$ reaction above the reaction threshold is governed by the $1/k_{Ax}^2$ factor and energy dependence

$$
S_{AxCe}^l \propto \exp(-\pi \eta_{Ax})
$$
\n(92)

of the relevant S-matrix element. The Coulomb barrier leads to a strong suppression of the the cross-section

$$
\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega}(Ax \to Cc) \propto k_{Ax}^{-2} \exp(-2\pi \eta_{Ax})
$$
\n(93)

for $k_{Ax} \rightarrow 0$ due to the decreasing exponential factor. This behaviour motivates the introduction of the astrophysical S factor (1). In the TH cross-section $d\sigma^{TH}/d\Omega$, that appears in Eq. (68), the factor k_{Ax}^{-2} is replaced with k_{Cc}^{-2} and the TH integrals $J_l^{(+)}$ appear. Their energy dependence for small k_{Ax} is determined by $k_{Ax}/C_0(\eta_{Ax}) \approx$ $k_{Ax} \exp(\pi \eta_{Ax}) / \sqrt{2\pi \eta_{Ax}}$ from the contribution of the irregular Coulomb wave function. This leads to a k_{Ax} dependence of the three-body cross-section (68) according to

$$
\frac{\mathrm{d}^3 \sigma}{\mathrm{d}E_C \,\mathrm{d}\Omega_C \,\mathrm{d}\Omega_c} \propto k_{Ax}^{-2} v_{Ax}^{-1} \exp(-2\pi \eta_{Ax}) k_{Ax}^2 \frac{\exp(2\pi \eta_{Ax})}{2\pi \eta_{Ax}} = (2\pi \eta_{Ax} v_{Ax})^{-1} = \text{const.}
$$
\n(94)

in the lowest order of k_{Ax} ; cf. also Eq. (79) with the k_{Ax} dependence of the penetrability factor $P_l \propto k_{Ax}^3 \exp(2\pi \eta_{Ax})$. As a result the cross-section does not vanish at the threshold but takes on a finite value. Of course, the same considerations apply to the case when the spins of the particles are considered. Also in the case of the transfer of neutron, like in a (d, p) stripping reaction, it is well known that the cross-section is finite at the threshold $E_n = 0$ [17,18]. The reason is the same as in the case of charged particles.

6.3. Extraction of astrophysical S factors

In an actual TH experiment the measured cross-section of the three-body reaction depends on the geometry of the setup, the chosen Trojan–Horse, and the energy of the projectile. The differential three-body cross-section (68) or (79) can be projected onto a simple cross-section

$$
\frac{d\sigma}{dE} = \int dE_C d\Omega_C d\Omega_c \frac{d^3 \sigma}{dE_C d\Omega_C d\Omega_c} \delta(E_{Ax} - E) F(E_C, \Omega_C, \Omega_c)
$$
\n(95)

depending on the $A + x$ c.m. energy E. The efficiency function F takes cut-offs in particle energies, momenta (e.g., \hat{Q}_{bb}), the detector geometry etc. into account. The experimental spectrum (95) can be compared to a corresponding theoretical crosssection from a simulation of the experiment assuming a certain energy dependence of the relevant on-shell S-matrix elements of the two-body reaction, e.g., from a Rmatrix parametrization. By a variation of the parameters, the best fit to the experiment is found and the on-shell two-body cross-section can be calculated. If there is only a contribution of one partial wave, the procedure becomes simpler. The ratio of the measured cross-section (95) to the corresponding simulated cross-section directly gives the energy dependence of the S factor relative to the energy dependence assumed in the theoretical S factor of the two-body reaction. Due to the uncertainties in the description of the reaction mechanism it is expected to be difficult to extract absolute values of the cross section. So it is better to normalize the cross-section to results from direct experiments at higher energies. However, we expect that the energy dependence of the cross-section (S factor) can be extracted much more reliably. For recent applications of the TH method with detailed information on the experimental realisation we refer to the references [20–26].

6.4. Elastic scattering with the Trojan–Horse method

The main aim for applying the TH method is the extraction of the energy dependence of cross-sections or astrophysical S factors for inelastic two-body reactions. But also the indirect investigation of elasting two-body scattering $A + x \rightarrow A + x$ can be rewarding. In the direct two-body scattering process the cross-section is dominated by the contribution of the Coulomb scattering amplitude (74) at low energies that diverges with k_{Ax}^{-2} . By way of contrast, the TH Coulomb scattering amplitude (67) vanishes for $k_{Ax} \rightarrow 0$ due to the appearance of the TH integrals with the regular Coulomb wave function F_l . The nuclear contribution to the TH cross-section (70) becomes dominant because of the TH integrals with the irregular Coulomb wave functions in Eq. (73). This allows the study of nuclear effects in the scattering at small energies, e.g., the influence of sub-threshold or low energy resonances. First attempts in this direction were made in recent experiments [27,28].

7. Summary and outlook

In this paper the basic theory of the Trojan–Horse method was developed starting from a distorted wave Born approximation of the T-matrix element. The essential surface approximation allows to find the relation between the cross-section of the three-body reaction and the S-matrix elements of the astrophysically relevant twobody reaction. In the modified plane wave approximation the relation between the three- and two-body cross-sections becomes very transparent. The three-body cross-section is a product of a kinematical factor, a momentum distribution and an off-shell two-body cross-section. Off-shell effects are expressed in terms of socalled TH integrals that were studied in detail. Their energy dependence leads to a finite cross-section of the three-body reaction at the threshold of the two-body reaction without the suppression by the Coulomb barrier. This allows to extract the energy dependence of astrophysical cross-sections from the three-body breakup reaction to very low energies without the problems of electron screening and extremely low cross-section. A comparison of results for S factors from direct and indirect experiments can improve the information on the electron screening effect. However, dedicated Trojan–Horse experiments are necessary in order to achieve a precision comparable to direct measurements.

The validity of the Trojan–Horse method can be tested by comparing the crosssections extracted from the indirect experiment with results from direct measurements of well-studied reactions. In principle it is possible to assess systematic uncertainties of the Trojan–Horse method by studying various combinations of projectile energies, spectators in the Trojan–Horse and scattering angles. Furthermore, different theoretical approximations can be compared, e.g., full DWBA calculations with and without the surface approximation and simpler modified plane wave approximations.

One may also envisage applications of the Trojan–Horse method to exotic nuclear beams. An unstable projectile hits a Trojan–Horse target allowing to study specific reactions on exotic nuclei. A study of low-energy elastic scattering with the Trojan– Horse method opens another application which can lead to improved information relevant to the theoretical description of nuclear reactions at low energies.

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Appendix A. Calculation of the Trojan–Horse integrals

With the differential equation for the Coulomb wave functions

$$
0 = \frac{1}{k^2} \frac{d^2}{dr^2} H_l(kr) + \left[1 - \frac{2\eta}{kr} - \frac{l(l+1)}{k^2 r^2}\right] H_l(kr),\tag{A.1}
$$

the TH integral (55) becomes

$$
J_l^{(H)}(R,\eta,k,Q) = k \int_R^{\infty} dr \left[\frac{2\eta}{kr} + \frac{l(l+1)}{k^2 r^2} - \frac{1}{k^2} \frac{d^2}{dr^2} \right] H_l(\eta;kr) z_l(Qr).
$$
 (A.2)

After partial integration and with help of the differential equation for z_l

$$
0 = \frac{1}{Q^2} \frac{d^2}{dr^2} z_l(Qr) + \left[1 - \frac{l(l+1)}{Q^2 r^2}\right] z_l(Qr), \tag{A.3}
$$

the expression

$$
J_{l}^{(H)}(R,\eta,k,Q) = \frac{k}{k^{2} - Q^{2}} \left[2\eta k I_{l}^{(H)}(R,\eta,k,Q) + k H_{l}^{\prime}(\eta;kR) z_{l}(QR) - Q H_{l}(\eta;kR) z_{l}^{\prime}(QR) \right],
$$
\n(A.4)

with the converging integral

$$
I_l^{(H)}(R,\eta,k,Q) = \int_R^{\infty} dr H_l(\eta;kr) r^{-1} z_l(Qr)
$$
\n(A.5)

is obtained. The prime denotes the differentiation with respect to the argument. We disregard the contribution from the upper bound (infinity). This corresponds to the regularization procedure mentioned above. Relation (A.4) cannot be used if $k = Q$ but this case will never occur since $k < Q$ from kinematical considerations (see Section 6.1). The integral $(A.5)$ can be written as

$$
I_l^{(H)}(R,\eta,k,Q) = kQM_{ll}^{-1}(H,\eta,k,Q) - \int_0^R dr H_l(\eta;kr)r^{-1}z_l(Qr), \qquad (A.6)
$$

with the integral

$$
M_{I_1I_2}^{-\lambda}(H,\eta,k,Q) = \frac{1}{kQ} \int_0^\infty dr H_{I_1}(\eta;kr) r^{-\lambda} z_{I_2}(Qr).
$$
 (A.7)

because $I_l^{(H)}$ is still finite in the limit $R \to 0$. The functions $M_{l_1l_2}^{-\lambda}$ are special cases of the radial matrix elements in the quantum mechanical theory of Coulomb excitation [39]. A more general form also appears in the general theory of transfer reactions to the continuum [40]. In the THM only monopole matrix elements ($\lambda = 1$) are needed. They can be calculated explicitly for both the regular and irregular Coulomb wave functions (see Appendix B for details). The TH integrals with the regular Coulomb functions are given by

$$
M_{00}^{-1}(F, \eta, k, Q) = \frac{C_0(\eta)}{2\eta kQ} \sin \zeta,
$$
\n(A.8)

$$
M_{11}^{-1}(F, \eta, k, Q) = \frac{3C_1(\eta)}{4\eta kQ(1+\eta^2)} \left[\left(\frac{Q}{k} + \frac{k}{Q} \right) \sin \zeta - 2\eta \cos \zeta \right]
$$
(A.9)

for $l = 0, 1$. The argument of the trigonometric functions is $\zeta = \eta \ln z_1$ with $z_1 =$ $(Q+k)/(Q-k)$. The constants $C_l(\eta)$ are recursively defined by [35]

$$
C_0(\eta) = \left(\frac{2\pi\eta}{\exp(2\pi\eta) - 1}\right)^{1/2}, \quad C_l(\eta) = \frac{(l^2 + \eta^2)^{1/2}}{l(2l + 1)} C_{l-1}(\eta). \tag{A.10}
$$

The case with the irregular Coulomb wave functions is slightly more intricate and leads to

$$
M_{00}^{-1}(G, \eta, k, Q) = \frac{C_0(\eta)}{2\eta kQ} (1 - \cos \zeta) + \frac{\pi}{2kQC_0(\eta)} [1 - S_0],
$$
\n(A.11)

$$
M_{11}^{-1}(G, \eta, k, Q) = \frac{3C_1(\eta)}{4\eta kQ(1+\eta^2)} \left[\left(\frac{Q}{k} + \frac{k}{Q} \right) (1 - \cos \zeta) - 2\eta \sin \zeta \right] + \frac{\pi}{12kQC_1(\eta)} \left[\left(\frac{Q}{k} + \frac{k}{Q} \right) (1 - S_0) + 2\eta S_1 \right],
$$
 (A.12)

with the series

$$
S_0 = \frac{2\eta}{\pi} \sum_{n=1}^{\infty} \frac{1 - z_1^{-n}}{n^2 + \eta^2}, \quad S_1 = \frac{2}{\pi} \sum_{n=1}^{\infty} \frac{nz_1^{-n}}{n^2 + \eta^2}.
$$
 (A.13)

Explicit expressions for larger l are obtained from the recursion relation $(l > 0)$

$$
\frac{2l+1}{2}\left(\frac{k}{Q}+\frac{Q}{k}\right)M_{ll}^{-1} = |l+i\eta|M_{l-1,l-1}^{-1} + |l+1+i\eta|M_{l+1,l+1}^{-1},\tag{A.14}
$$

which is valid for $H_l = F_l$ and $H_l = G_l$. It can be derived in a similar manner as in Coulomb excitation theory [39]. However, for a numerical calculation the question of stability of (A.14) arises. In the case of $H_l = F_l$ only the backward recurrence is stable. Starting with, e.g., $M_{LL}^{-1}(H, \eta, k, Q) = 0$ and $M_{L-1L-1}^{-1}(H, \eta, k, Q) = 1$ for large $L \gg l$ a backward recursion gives integrals for $l = 0, 1, 2, \dots$ after normalizing to the known value (A.8). In the case of $H_1 = G_1$ the upward recurrence can be used with the known starting values $M_{00}^{-1}(G, \eta, k, Q)$ and $M_{11}^{-1}(G, \eta, k, Q)$. The remaining integral from 0 to R in Eq. (A.6) is easily performed numerically.

When η becomes large (and k small) it is not favorable to calculate the integral $I_l^{(G)}(R,\eta,k,Q)$ from Eq. (A.6). The irregular Coulomb wave functions becomes very large for small radii r and a difference of large numbers has to be evaluated with loss of accuracy. In this case it is more convenient to use Eq. (A.5) directly because G_l becomes small with increasing $r > R$ very rapidly.

A.1. Limiting cases

In the case of $R \to 0$ the surface contributions in Eq. (A.4) vanish only if $H_l = F_l$. This can be seen from the approximations [35]

$$
F_l(\eta; x) \to \frac{(2l+1)!C_l(\eta)}{(2\eta)^{l+1}} \sqrt{2\eta x} I_{2l+1}(2\sqrt{2\eta x}) \approx C_l(\eta) x^{l+1}
$$
\n(A.15)

and $z_l(x) \rightarrow x^{l+1}/(2l+1)!!$ for $x = kR \rightarrow 0$. In this case the expression for TH integral with the regular Coulomb wave function reduces to the simple form

$$
J_l^{(F)}(0, \eta, k, Q) = \frac{2\eta k^3 Q}{k^2 - Q^2} M_{l1}^{-1}(F_l, \eta, k, Q).
$$
 (A.16)

Using the approximation [35]

$$
G_{l}(\eta; x) \to \frac{2(2\eta)^{l}}{(2l+1)!C_{l}(\eta)} \sqrt{2\eta x} K_{2l+1}(2\sqrt{2\eta x}) \approx \frac{x^{-l}}{(2l+1)C_{l}(\eta)}
$$
(A.17)

for small x one obtains for $H_l = G_l$ the result

$$
J_l^{(G)}(0,\eta,k,Q) = \frac{kQ}{k^2 - Q^2} \left[2\eta k^2 M_{ll}^{-1}(G_l, \eta, k, Q) - \frac{(Q/k)^l}{(2l+1)!!C_l(\eta)} \right],
$$
 (A.18)

with a remaining contribution from the surface terms.

In the special case where the transferred particle x is a neutron, the Coulomb functions $u_l^{(\pm)}(x)$ reduce to Hankel functions $x[-y_l(x) \pm i j_l(x)] = \pm i x h_l^{(\pm)}(x)$ with the

spherical Neumann and Bessel functions y_l and j_l , respectively [35]. There is no contribution from the term with the integral $I_l^{(H)}$ and the TH integral simplifies to

$$
J_l^{(\pm)}(R,0,k,Q) = \frac{\pm ik^2QR^2}{k^2 - Q^2} \left[kh_{l-1}^{(\pm)}(kR)j_l(QR) - Qh_l^{(\pm)}(kR)j_{l-1}(QR) \right],\tag{A.19}
$$

so that $J_l^{(+)}(R, 0, k, Q)$ is ik^2Q times the I_l integral in the stripping enhancement factors F_l of [41]. For the neutral particle, there is, for $l \neq 0$, still a barrier, the angular momentum barrier with its $l(l+1)/r^2$ radial dependence. The enhancement is due to the Hankel functions $h_l^{(\pm)}(x)$ with a $x^{-(l+1)}$ behaviour for $x \to 0$. The mathematical treatment of the Coulomb functions is much more involved.

A.2. Energy dependence of the Trojan–Horse integrals

In order to investigate the behaviour of $J_l^{(H)}$ for $k \to 0$ it is useful to study the case $R = 0$ first. From the explicit expressions for the integrals $M_{II}^{-1}(F_I)$ the approximations

$$
J_0^{(F)}(0, \eta, k, Q) = J_{0, \text{red}}^{(F)}(0, \eta, \xi) \to -C_0(\eta) \xi^2 \sin(2\eta \xi), \tag{A.20}
$$

$$
J_1^{(F)}(0, \eta, k, Q) = J_{1, \text{red}}^{(F)}(0, \eta, \xi) \to -C_0(\eta) \xi^2 \left[\frac{\sin(2\eta\xi)}{2\eta\xi} - \cos(2\eta\xi) \right]
$$
(A.21)

for the integrals with the regular Coulomb functions F_0 and F_1 are found. Here, we used the reduced variables defined in Eq. (56). Because the recursion relation (A.14) for M_{ll}^{-1} reduces to the recursion relation of the Riccati–Bessel functions we have the approximation

$$
J_I^{(F)}(0, \eta, k, Q) = J_{I, \text{red}}^{(F)}(0, \eta, \xi) \to -C_0(\eta) \xi^2 z_I(2\eta\xi)
$$
 (A.22)

for all l in the limit $k \to 0$. The argument $2\eta \xi = 2\eta k/Q$ is independent of k for constant Q and the entire k dependence is determined by $C_0(\eta)\xi^2$ independent of l.

The TH integrals with the irregular Coulomb wave functions reduce in the limit $k \rightarrow 0$ to the form

$$
J_0^{(G)}(0, \eta, k, Q) = J_{0, \text{red}}^{(G)}(0, \eta, \xi) \to \frac{\xi}{C_0(\eta)} [1 - 2\eta \xi f(2\eta \xi)], \tag{A.23}
$$

$$
J_1^{(G)}(0,\eta,k,Q) = J_{1,\text{red}}^{(G)}(0,\eta,\xi) \to \frac{\xi}{C_0(\eta)} \left[\frac{1}{\eta\xi} - f(2\eta\xi) - 2\eta\xi g(2\eta\xi) \right],\tag{A.24}
$$

with functions

$$
f(x) = \int_0^\infty dt \frac{e^{-xt}}{t^2 + 1} = Ci(x)\sin x - \left[Si(x) - \frac{\pi}{2}\right]\cos x,\tag{A.25}
$$

$$
g(x) = \int_0^\infty dt \frac{t e^{-xt}}{t^2 + 1} = -Ci(x)\cos x - \left[Si(x) - \frac{\pi}{2}\right] \sin x,\tag{A.26}
$$

that can be expressed in terms of the sine and cosine integrals $Si(x)$ and $Ci(x)$ [35]. The expressions inside the brackets in Eqs. (A.23) and (A.24) are constants and the k-dependence of the TH integrals with the irregular Coulomb wave functions is determined by $\zeta/C_0(\eta)$.

In the case of finite R the energy dependence of the TH integrals for $\xi \to 0$ can be extracted with the help of the approximations (A.15) and (A.17) which are valid for $2\eta \gg \xi x_0$. The energy dependence of the differences

$$
J_l^{(F)}(0, \eta, k, Q) - J_l^{(F)}(R, \eta, k, Q) = J_{l, \text{red}}^{(F)}(0, \eta, \xi) - J_{l, \text{red}}^{(F)}(x_0, \eta, \xi)
$$

$$
\approx \frac{(2l+1)!C_l(\eta)\xi}{(2\eta)^{l+1}}
$$

$$
\times \int_0^{x_0} dx \sqrt{2\eta \xi x} I_{2l+1}(2\sqrt{2\eta \xi x}) z_l(x) \qquad (A.27)
$$

and

$$
J_l^{(G)}(0, \eta, k, Q) - J_l^{(G)}(R, \eta, k, Q) = J_{l, \text{red}}^{(G)}(0, \eta, \xi) - J_{l, \text{red}}^{(G)}(x_0, \eta, \xi)
$$

$$
\approx \frac{2(2\eta)^l \xi}{(2l+1)! C_l(\eta)}
$$

$$
\times \int_0^{x_0} dx \sqrt{2\eta \xi x} K_{2l+1} (2\sqrt{2\eta \xi x}) z_l(x) \quad (A.28)
$$

is entirely given by the factor in front of the integrals since $\eta \xi = \eta k/Q$ is constant for constant O. For large *n* and small ξ it is found that the energy dependence of the finite range correction to the TH integral with the regular and irregular Coulomb functions is given by $\xi^2 C_0(\eta)$ and $\xi/C_0(\eta)$, respectively. The finite range correction essentially leads to a change in the absolute value of the integral but the energy dependence of $J_l^{(\pm)}(R, \eta, k, Q)$ and $J_l^{(\pm)}(0, \eta, k, Q)$ is the same for small k.

When the Sommerfeld parameter η increases, the coefficient $C_0(\eta)$ becomes very small and the TH integral with G_l is much larger than the TH integral with F_l . The integrals $J_l^{(\pm)}$ are therefore dominated by $J_l^{(G)}$ at low energies.

In the case where a neutron is transferred in reaction (3) the reduced TH integrals can be written as

$$
J_l^{(\pm)}(R,0,k,Q) = J_{l,\text{red}}^{(\pm)}(x_0,0,\xi)
$$

=
$$
\frac{\pm i\xi^2 x_0^2}{\xi^2 - 1} \left[\xi h_{l-1}^{(\pm)}(\xi x_0) j_l(x_0) - h_l^{(\pm)}(\xi x_0) j_{l-1}(x_0) \right]
$$
(A.29)

in the variables $x_0 = QR$ and $\xi = k/Q$ with $\eta = 0$. For $\xi \to 0$ we find the approximation

$$
J_{l,\text{red}}^{(\pm)}(x_0,0,\xi) \to \frac{(2l-1)!!}{\left(\xi x_0\right)^{l-1}} j_{l-1}(x_0),\tag{A.30}
$$

with a characteristic *l*-dependence due to the centrifugal barrier.

A.3. Simplified approximation of TH integrals

The dominance of the contribution with the irregular Coulomb wave function G_l in the TH integrals $J_l^{(\pm)}$ at low energies also motivates a simple approximation in the limit $k \to 0$ that has been used in applications of the THM to reactions with $Ax \neq Cc$. In this case the integral is completely determined by the contributions at radii close to R and the integrand decreases very fast with increasing r. If we assume that a general function $f(r)$ decreases exponentially, the integral of $f(r)$ from R to infinity can be calculated from

$$
\int_{R}^{\infty} dr f(r) = f(R) \int_{R}^{\infty} dr \exp\left(-\frac{r}{\Delta R}\right) = \Delta R f(R), \tag{A.31}
$$

with $\Delta R = -f(R)/f'(R)$ and $\text{Re}(\Delta R) > 0$. In the TH integral $J_l^{(\pm)}$ the appropriate value of ΔR is determined by

$$
\Delta R \approx -\left[\frac{kG_l'(\eta; kR)}{G_l(\eta; kR)} + \frac{Qz_l'(Qr)}{z_l(Qr)}\right]^{-1},\tag{A.32}
$$

neglecting the small contribution of F_l for $k \to 0$. From the approximation (A.17) of the irregular Coulomb wave function for $kR \ll 2\eta$ we find

$$
\Delta R \approx -\left[\sqrt{\frac{2\eta k}{R}} \frac{K'_{2l+1}(2\sqrt{2\eta kR})}{K_{2l+1}(2\sqrt{2\eta kR})} + \frac{Qz'_{l}(Qr)}{z_{l}(Qr)}\right]^{-1},
$$
\n(A.33)

independent of k because ηk is constant and Q is practically independent of k for small k. Applying this result to the TH integral leads to the approximation

$$
J_l^{(\pm)}(R,\eta,k,Q) = k \int_R^{\infty} dr u_l^{(\pm)}(\eta;kr) z_l(Qr) \approx k \Delta R u_l^{(\pm)}(\eta;kR) z_l(QR).
$$
 (A.34)

Since we are only interested in the k-dependence, the TH integral can be replaced by the value of the integrand,

$$
J_l^{(\pm)}(R,\eta,k,Q) \to kRu_l^{(\pm)}(\eta;kR)z_l(QR), \qquad (A.35)
$$

at the cutoff radius R with the factor R instead of ΔR . The energy dependence for small k in this simple approximation agrees with the exact result for the TH integrals.

Appendix B. Analytical calculation of radial integrals

The radial integrals

$$
M_{II}^{-1}(H, \eta, k, Q) = \frac{1}{kQ} \int_0^\infty dr H_I(\eta; kr) r^{-1} z_I(Qr)
$$
 (B.1)

appearing in Eq. (A.6) can be calculated explicitly by using the integral representation

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$$
F_l(\eta;kr) + iG_l(\eta;kr) = \frac{i(kr)^{l+1}}{(2l+1)!C_l(\eta)} \int_i^{i+\infty} ds e^{-krs}(s-i)^{l-irj}(s+i)^{l+irj}
$$
(B.2)

of the combined Coulomb function $H_l = F_l + iG_l$. This form has been obtained from the representation in [35] by the substitution $t = kr(s - i)$. The integration path in the variable s is a parallel to the real axis in the complex plane. Since the recursion relation (A.14) connects integrals of three successive l-values, only the basic integrals with $l = 0$ and $l = 1$ are needed for the explicit calculation of M_{ll}^{-1} for all l. With $z_0(x) = \sin x$ and $z_1(x) = \sin x/x - \cos x$ the integration over the radial coordinate r is easily performed and one finds

$$
M_{00}^{-1}(H, \eta, k, Q) = \frac{i}{C_0(\eta)} \int_i^{i+\infty} ds \frac{(s+i)^{i\eta}(s-i)^{-i\eta}}{k^2 s^2 + Q^2}
$$
(B.3)

and

$$
M_{11}^{-1}(H,\eta,k,Q) = \frac{\mathrm{i}kQ}{3C_1(\eta)} \int_i^{i+\infty} \mathrm{d}s \frac{(s+i)^{1+i\eta}(s-i)^{1-i\eta}}{(k^2s^2+Q^2)^2}.
$$
 (B.4)

The substitution $s = i(t+1)/(t-1)$ leads to the expressions

$$
M_{00}^{-1}(H, \eta, k, Q) = \frac{-2}{C_0(\eta)(Q^2 - k^2)} I_0(\eta, z),
$$
\n(B.5)

$$
M_{11}^{-1}(H, \eta, k, Q) = \frac{4kQ}{3C_1(\eta)(Q^2 - k^2)^2} \frac{d}{dz} I_0(\eta, z)
$$
 (B.6)

with the integral

$$
I_0(\eta, z) = \int_C dt \frac{t^{i\eta}}{t^2 - 2tz + 1}.
$$
 (B.7)

The contour C is the straight line from 1 to $1 + i\infty$ parallel to the imaginary axis in the complex plane. The integral depends on the variable

$$
z = \frac{Q^2 + k^2}{Q^2 - k^2}
$$
 (B.8)

which is always larger than 1 for the conditions of the TH method. The integrand has which is always larger than 1 for the conditions of the TH method. The integrand has two poles at $z_1 = z + \sqrt{z^2 - 1} > 1$ and $z_2 = z - \sqrt{z^2 - 1} < 1$ on the real axis. In the next step the path of integration is deformed to lie on the real axis and the integrand is broken into partial fractions. This yields

$$
I_0(\eta, z) = \int_1^\infty dt \frac{t^{i\eta}}{t^2 - 2tz + 1} - \pi i \text{Res}_{z_1} \frac{t^{i\eta}}{t^2 - 2tz + 1}
$$

=
$$
\frac{1}{z_1 - z_2} \left[I_1 - I_2 - \pi i z_1^{i\eta} \right]
$$
(B.9)

with the contribution of the residue at the pole $z₁$. The remaining integrals

$$
I_i = \int_1^\infty dt t^{i\eta} \left[\frac{1}{t - z_i} - \frac{1}{t} \right]
$$
 (B.10)

for $i = 1, 2$ are evaluated with the help of the geometric series and the relation [35]

$$
\frac{\pi\eta}{\tanh(\pi\eta)} = 1 + 2\eta^2 \sum_{n=1}^{\infty} \frac{1}{n^2 + \eta^2}.
$$
 (B.11)

Collecting all contributions we find

$$
I_0(\eta, z) = \frac{1}{2\eta\sqrt{z^2 - 1}} \left[C_0^2(\eta) \left(z_1^{i\eta} - 1 \right) - 2\eta^2 \sum_{n=1}^{\infty} \frac{z_1^{-n} - 1}{n^2 + \eta^2} - \pi \eta \right]
$$
(B.12)

for the fundamental integral (B.7). Separating real and imaginary parts, the explicit expressions of the integrals $M_{00}^{-1}(H_0, \eta, k, Q)$ and $M_{00}^{-1}(G_0, \eta, k, Q)$ are found with Eq. (B.5). In the case $l = 1$ the derivative of $I_0(z)$ with respect to z has to be calculated first.

Appendix C. Cross-sections for particles with spin

For the application of the THM to a particular reaction one has to consider that the nuclei participating in the reactions (2) and (3) usually carry a spin. In this case the definition of the relevant quantities becomes more intricate but the general procedure remains the same as in the spinless case. Spins of an individual particle i will be denoted in the following by s_i with projection v_i .

Here, the channel spin basis in the derivation of the expressions is used. In the two-particle reaction the spins s_A and s_X in the initial state are coupled to the channel spin s_{Ax} . Similarly, in the final state s_c and s_c are coupled to the channel spin s_{Cc} . Coupling s_{Cc} with the spin of the spectator s_b gives the angular momentum s_f of the three-body final state. The initial state of reaction (3) is characterized by the angular momentum s_i which is obtained from coupling the spins s_A and s_a .

The cross-section of reaction (3) in the laboratory system

$$
\frac{\mathrm{d}^3 \sigma}{\mathrm{d}E_C \,\mathrm{d}\Omega_C \,\mathrm{d}\Omega_c} = \frac{K_{\text{lab}}}{(2s_A + 1)(2s_a + 1)} \times \sum_{s_i s_f} \sum_{v_i v_f} \sum_{s_{Cc}} \left| T_{fi}(\vec{k}_{Cc}, \vec{k}_{Bb}, s_{Cc} s_f v_f; \vec{k}_{Aa}, s_i v_i) \right|^2 \tag{C.1}
$$

is obtained by summing over final spin states and averaging over initial spin states with the three-body T-matrix element T_f that carries now both momentum and spin indices. In the modified plane wave approximation it is given by

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$$
T_{fi}^{\text{TH}}(\vec{k}_{Cc}, \vec{q}_{Bb}, s_{Cc}s_f v_f; \vec{q}_{Aa}, s_i v_i) = \left\langle \exp(i\vec{q}_{Bb} \cdot \vec{r}_{Bb}) \Psi_{\text{asym}}^{(-)}(Cc - b, \vec{k}_{Cc}, s_f v_f) \times |V_{xb}| \exp(i\vec{q}_{Aa} \cdot \vec{r}_{Aa}) \phi_{Aa}(s_i v_i) \right\rangle
$$
(C.2)

with the asymptotic wave function

$$
\Psi_{\text{asym}}^{(-)}(Cc - b, \vec{k}_{Cc}, s_{Cc}, s_{f}v_f) \n= \sum_{v_{Cc}v_b} (s_{Cc}v_{Cc}s_bv_b|s_fv_f) \Psi_{\text{asym}}^{(-)}(Cc, \vec{k}_{Cc}, s_{Cc}v_{Cc}) \phi_b(s_bv_b),
$$
\n(C.3)

where

W^ðÞ asymðCc;~kCc;sCc; ^mCcÞ ¼ ⁴^p kCc X asa X JM X lalCc 1 ra n Jð Þ ^lasalCcsCc ða; karaÞY^las^a JM ða; ^raÞZlCcsCcmCc JM ð^kCcÞ ðC:4Þ

for $r_\alpha \ge R$ and $\Psi_{\text{asym}}^{(\pm)} = 0$ for $r_\alpha < R$. The vector spherical harmonics

$$
\mathcal{Y}_{JM}^{ls}(\alpha,\hat{r}) = \sum_{mv} (lmsv|JM) i^l Y_{lm}(\hat{r}) \phi_{\alpha}(sv)
$$
\n(C.5)

and the angular distribution functions

$$
\mathcal{Z}_{JM}^{Isv}(\hat{k}) = \sum_{m} (Imsv|JM) Y_{lm}(\hat{k}).
$$
\n(C.6)

are obtained by coupling the channel spins s with the corresponding orbital angular momenta l to the total angular momentum J . The radial wave functions

$$
\zeta_{I_z s_z I_{Cc} c_c}^{J(+)}(\alpha; k_x r_\alpha) = \frac{1}{2i} \sqrt{\frac{v_{Cc}}{v_\alpha}} \Big[S_{I_z s_z I_{Cc} c_c}^{J \alpha C_c} u_{I_z}^{(+)}(\eta_\alpha; k_x r_\alpha) - \delta_{\alpha C_c} \delta_{I_z I_{Cc}} \delta_{s_z s_{Cc}} u_{I_z}^{(-)}(\eta_\alpha; k_x r_\alpha) \Big]
$$
(C.7)

and

$$
\zeta_{I_{as}I_{Cs}I_{Cs}C_c}^{J(-)}(\alpha;k_xr_x) = \zeta_{I_{as}I_{Cs}C_c}^{J(+)*}(\alpha;k_xr_x)
$$
\n(C.8)

contain the general S-matrix elements $S_{l_{as}l_Cs_C}^{j_aCC}$ for a transition from a channel with quantum numbers $l_{Cc} s_{CC}$ in the partition Cc to the channel $l_{\alpha} s_{\alpha}$ in partition α .

The T-matrix in the modified plane wave approximation

$$
T_{fi}^{\text{TH}}(\vec{k}_{Cc}, \vec{q}_{Bb}, s_{Cc}s_f v_f; \vec{q}_{Aa}, s_i v_i) = \frac{4\pi}{k_{Ax}Q_{Aa}} W(\vec{Q}_{Bb}) \mathcal{F}^{\text{TH}}(\vec{Q}_{Aa}, \vec{k}_{Cc}, s_{Cc}s_f v_f s_i v_i)
$$
(C.9)

again factorizes into a form with the momentum amplitude W and the scattering amplitude

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$$
\mathcal{F}^{\text{TH}}(\vec{Q}_{Aa}, \vec{k}_{Cc}, s_{Cc}s_{f}v_{f}s_{i}v_{i}) = \frac{1}{2ik_{Cc}} \sum_{J} \sum_{s_{Ax}} \sum_{l_{Ax}l_{Cc}} (2J+1) \mathcal{P}_{l_{Ax}s_{Ax}l_{Cc}s_{Cc}}^{J}(\hat{k}_{Cc}, \hat{Q}_{Aa}, s_{f}v_{f}s_{i}v_{i})
$$

$$
\times \sqrt{\frac{v_{Cc}}{v_{Ax}}} \Big[S_{l_{Ax}s_{Ax}l_{Cc}s_{Cc}}^{J_{Ax}l_{Cc}s_{Cc}} J_{l_{Ax}}^{(+)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) -\delta_{Axcc}\delta_{l_{Ax}l_{Cc}}\delta_{s_{Ax}s_{Cc}} J_{l_{Ax}}^{(-)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \Big]
$$
(C.10)

which is a matrix in spin space. The angular distribution is determined by the function

$$
\mathcal{P}'_{l_{Ax}s_{Ax}l_{Cs}s_{Cc}}(\hat{k}_{Cc},\hat{Q}_{Aa},s_fv_f s_iv_i) = \frac{4\pi}{2J+1} \sum_{v_xv_b} \sum_{v_Av_a} \sum_{m_{Cc}v_{Cc}} \sum_{m_{Ax}v_{Ax}} \sum_{m_{Cx}v_{Ax}} \sum_{w_{Ax}v_{Ax}} \sum_{v_Av_B} \sum_{v_Bv_B} \sum_{v_Bv_B} \sum_{v_Bv_B} \sum_{v_Av_B} \sum_{v_Bv_B} \sum_{v_Bv_B} \sum_{v_Av_B} \sum_{v_Av_B} \sum_{v_Bv_B} \sum_{v_Av_B} \sum_{v
$$

with

$$
\chi_{l_{Ax}s_{Ax}l_{Cx}s_{Ce}}^{Jjs_{i}s_{f}}(s_{x}, s_{b}, s_{a}, s_{A}) = (-1)^{s_{Ax} - s_{x} - s_{A} - j} \times \frac{(-1)^{s_{Ax} + J + l_{Ce}}}{\sqrt{2s_{Ax} + 1}\sqrt{2l_{Ax} + 1}} \sqrt{2s_{a} + 1}\sqrt{2s_{i} + 1}(2j + 1) \times (2s_{Ax} + 1) \times \frac{(-1)^{s_{Ax} + J + l_{Ce}}}{\sqrt{2s_{Ax} + 1}\sqrt{2l_{Ax} + 1}} \sqrt{2s_{a} + 1}\sqrt{2s_{a} + 1}(2j + 1) \times (2s_{Ax} + 1) \times \frac{(-1)^{s_{Ax} - s_{A} - j}}{\sqrt{2s_{Ax} + 1}\sqrt{2l_{Ax} + 1}} \sqrt{2s_{a} + 1}\sqrt{2s_{a} + 1}(2j + 1) \times (2s_{Ax} + 1) \times \frac{(-1)^{s_{Ax} - s_{A} - j}}{\sqrt{2s_{Ax} + 1}\sqrt{2l_{Ax} + 1}} \sqrt{2s_{a} + 1}\sqrt{2s_{a} + 1}(2j + 1) \times (2s_{Ax} + 1) \times \frac{(-1)^{s_{Ax} - s_{A} - j}}{\sqrt{2s_{Ax} + 1}\sqrt{2l_{Ax} + 1}} \sqrt{2s_{a} + 1}\sqrt{2s_{a} + 1}(2j + 1) \times (2s_{Ax} + 1) \times \frac{(-1)^{s_{Ax} - s_{A} - j}}{\sqrt{2s_{Ax} + 1}\sqrt{2l_{Ax} + 1}} \sqrt{2s_{a} + 1}\sqrt{2s_{a} + 1}(2j + 1) \times (2s_{Ax} + 1) \times \frac{(-1)^{s_{Ax} - s_{A} - j}}{\sqrt{2s_{Ax} + 1}\sqrt{2l_{Ax} + 1}} \sqrt{2s_{a} + 1}\sqrt{2s_{a} + 1}(2j + 1) \times (2s_{Ax} + 1) \times \frac{(-1)^{s_{Ax} - s_{A} - j}}{\sqrt{2s_{Ax} + 1}\sqrt{2l_{Ax} + 1}}
$$

The Trojan–Horse scattering amplitude is a sum

$$
\mathcal{F}^{\text{TH}} = \mathcal{F}_N^{\text{TH}} + \mathcal{F}_C^{\text{TH}} \delta_{\text{AxCe}} \tag{C.13}
$$

of the pure Coulomb contribution

$$
\mathcal{F}_C^{\text{TH}} = \frac{1}{k_{Cc}} \sum_{J} \sum_{s_{Ax}} \sum_{l_{Ax}l_{Cc}} (2J+1)
$$

$$
\times \mathcal{P}'_{l_{Ax}s_{Ax}l_{Cc}s_{Cc}}(\hat{k}_{Cc}, \hat{Q}_{Aa}, s_f v_f s_i v_i) e^{i\sigma_{l_{Ax}}(\eta_{Cc})} J_{l_{Ax}}^{(F)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \delta_{l_{Ax}l_{Cc}} \delta_{s_{Ax}s_{Cc}}
$$

$$
= f_C^{\text{TH}} \delta_{v_i v_f} \delta_{s_i s_f} (-1)^{s_x + s_A + s_b + s_f} \sqrt{2s_{Cc} + 1} \sqrt{2s_a + 1} \begin{Bmatrix} s_x & s_A & s_{Cc} \\ s_f & s_b & s_a \end{Bmatrix}
$$

(C.14)

and the nuclear contribution

$$
\mathcal{F}_{N}^{\text{TH}} = \frac{1}{2ik_{Cc}} \sum_{J} \sum_{s_{Ax}} \sum_{l_{Ax}l_{Cc}} (2J+1) \mathcal{P}_{l_{Ax}s_{Ax}l_{Cc}s_{Cc}}^{J} (\hat{k}_{Cc}, \hat{Q}_{Aa}, s_f v_f s_i v_i)
$$

$$
\times \sqrt{\frac{v_{Cc}}{v_{Ax}}} T_{l_{Ax}s_{Ax}l_{Cc}s_{Cc}}^{J_{Ax}c} J_{l_{Ax}}^{(+)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \tag{C.15}
$$

with

$$
T_{l_{Ax}s_{Ax}l_{Cc}s_{Cc}}^{JAXC} = e^{i\sigma_{l_{Ax}}(\eta_{Ax}) + i\sigma_{l_{Cc}}(\eta_{Cc})} \left[S_{l_{Ax}s_{Ax}l_{Cc}s_{Cc}}^{JAXCcN} - \delta_{AxCc} \delta_{l_{Ax}l_{Cc}} \delta_{s_{Ax}s_{Cc}} \right]
$$
(C.16)

depending on the nuclear S-matrix element $S_{l_{Ax}S_{Ax}l_{Cs}S_{Cx}}^{JAxCcN}$.

The unpolarized cross-section for the three-body reaction in the laboratory system

$$
\frac{\mathrm{d}^3 \sigma}{\mathrm{d}E_C \,\mathrm{d}\Omega_C \,\mathrm{d}\Omega_c} = K_{\text{lab}} |W(\vec{Q}_{Bb})|^2 \frac{16\pi^2}{k_{Ax}^2 Q_{Aa}^2} \frac{v_{Cc}}{v_{Ax}} \frac{\mathrm{d}\sigma^{\text{TH}}}{\mathrm{d}\Omega}
$$
(C.17)

has the same form as in the case for spinless particles. The TH cross-section

$$
\frac{d\sigma^{TH}}{d\Omega} = \frac{1}{(2s_A + 1)(2s_a + 1)} \sum_{s_i s_f} \sum_{v_i v_f} \sum_{s_{Cc}} \frac{v_{Ax}}{v_{Cc}} |\mathcal{F}^{TH}|^2
$$

$$
= \left[\frac{d\sigma_C^{TH}}{d\Omega} + \frac{d\sigma_I^{TH}}{d\Omega} \right] \delta_{AxCc} + \frac{d\sigma_N^{TH}}{d\Omega} \tag{C.18}
$$

again decomposes into a pure Coulomb contribution

$$
\frac{d\sigma_C^{TH}}{d\Omega} = \frac{1}{(2s_A + 1)(2s_a + 1)} \sum_{s_i s_j} \sum_{v_i v_j} \sum_{s_{Cc}} |\mathcal{F}_C^{TH}|^2 = |f_C^{TH}|^2 \tag{C.19}
$$

an interference contribution

$$
\frac{d\sigma_{I}^{TH}}{d\Omega} = \frac{1}{(2s_{A} + 1)(2s_{a} + 1)} \sum_{s_{i}s_{f}} \sum_{v_{i}v_{f}} \sum_{s_{Cc}} 2\text{Re}\left[\mathcal{F}_{C}^{TH*}\mathcal{F}_{N}^{TH}\right]
$$
\n
$$
= \frac{1}{(2s_{A} + 1)(2s_{x} + 1)} \sum_{J} \sum_{l_{Cc}} \sum_{s_{Cc}} \sum_{X_{Cc}} \times \frac{2J + 1}{k_{Cc}} \text{Re}\left[(if_{C}^{TH*})^{*}T_{l_{Cc}s_{Cc}l_{Cs}s_{Cc}}^{J_{Cc}l_{Cs}} J_{l_{Cc}}^{(+)}(R, \eta_{Ax}, k_{Ax}, Q_{Aa})\right] P_{l_{Cc}}(\hat{k}_{Cc} \cdot \hat{Q}_{Aa}) \tag{C.20}
$$

and a nuclear contribution

$$
\frac{d\sigma_{N}^{TH}}{d\Omega} = \frac{1}{(2s_{A}+1)(2s_{a}+1)} \sum_{s_{i}s_{f}} \sum_{v_{i}v_{f}} \sum_{s_{Cc}} \frac{v_{Ax}}{v_{Cc}} |\mathcal{F}_{N}^{TH}|^{2}
$$
\n
$$
= \frac{1}{(2s_{A}+1)(2s_{x}+1)} \frac{1}{4k_{Cc}^{2}} \sum_{\lambda} \sum_{JJ'} \sum_{s_{Ax}s_{Cc}} \sum_{l_{Ax}l_{Cc}} \times \sum_{l'_{x}l'_{Cc}} (-1)^{s_{Ax}-s_{Cc}} Z_{l'_{Ax}l'_{Ax}}^{JI'} (\lambda s_{Ax}) Z_{l'_{Cc}l'_{Cc}}^{IJ'} (\lambda s_{Cc}) P_{\lambda}(\hat{Q}_{Aa} \cdot \hat{k}_{Cc})
$$
\n
$$
\times T_{l_{Ax}s_{Ax}l_{Cc}s_{Cc}}^{J_{AX}C_{c}} T_{l'_{Ax}s_{Ax}l'_{Cc}s_{Cc}}^{J'_{AX}C_{c}} J_{l_{Ax}}^{(+)} (R, \eta_{Ax}, k_{Ax}, Q_{Aa}) J_{l'_{Ax}}^{(-)} (R, \eta_{Ax}, k_{Ax}, Q_{Aa}), \qquad (C.21)
$$

with factors

$$
Z_{ll'}^{JJ'}(\lambda s) = \sqrt{(2J+1)(2J'+1)(2l+1)(2l'+1)}(10l'0|\lambda 0) \begin{cases} l & \lambda & l' \\ J' & s & J \end{cases}
$$
 (C.22)

for the angular momentum coupling. The Coulomb contribution is the same as in the case of spinless particles. Notice the change from s_a to s_x in the spin degeneracy factors.

If only one orbital angular momentum $l = l_{Ax} = l'_{Ax}$ contributes in the Ax partition to the inelastic two-body reaction $A + x \rightarrow C + c$, the TH cross-section

$$
\frac{d\sigma^{TH}}{d\Omega} = \frac{(2s_C + 1)(2s_C + 1)}{(2s_A + 1)(2s_X + 1)} \frac{d\sigma_l}{d\Omega} (Cc \to Ax) P_l(R, \eta_{Ax}, k_{Ax}, Q_{Aa})
$$
(C.23)

is directly related to the usual on-shell cross-section

$$
\frac{d\sigma_{l}}{d\Omega}(Cc \to Ax) = \frac{1}{(2s_{C}+1)(2s_{c}+1)} \frac{1}{4k_{Cc}^{2}} \sum_{\lambda} \sum_{JJ'} \sum_{s_{As}s_{Cc}} \sum_{x_{As}s_{Cc}} \times \sum_{l_{Cc}l'_{Cc}} (-1)^{s_{As}-s_{Cc}} Z_{ll'}^{JJ'}(\lambda s_{Ax}) Z_{l_{Cc}l'_{Cc}}^{JJ'}(\lambda s_{Cc}) P_{\lambda}(\hat{Q}_{Aa} \cdot \hat{k}_{Cc}) T_{ls_{As}l_{Cc}s_{Cc}}^{JAxCc} T_{ls_{As}l'_{Cc}s_{Cc}}^{J'AxCc*},
$$
\n(C.24)

with the penetrability factor (77) . In $(C.23)$ the only difference to the case with spinless particles is the appearance of the spin degeneracy factors. Applying the theorem of detailed balance the simple relation

$$
\frac{d^3 \sigma}{dE_C d\Omega_C d\Omega_c} = K_{lab} \left| W(\vec{Q}_{Bb}) \right|^2 \frac{16\pi^2}{k_{CC}^2 Q_{Aa}^2} \frac{v_{Cc}}{v_{Ax}} \frac{d\sigma_l}{d\Omega} \times (Ax \to Cc) P_l(R, \eta_{Ax}, k_{Ax}, Q_{Aa}) \tag{C.25}
$$

is found which is the same as in the spinless case. The increase of the factor P_l at small k_{Ax} compensates the decrease of the two-body cross-section $(d\sigma_l/d\Omega)$ $(Ax \rightarrow Cc)$.

Appendix D. Trojan–Horse Coulomb scattering amplitude

Recalling the procedure to derive the radial integrals (B.1), the contribution f_{C0}^{TH} to the Coulomb scattering amplitude can be written as

$$
f_{\text{CO}}^{\text{TH}} = \frac{\eta k^2 Q}{2\pi (k^2 - Q^2)} \left\langle \Psi_{\text{Coul}}^{(-)}(\vec{k}) | r^{-1} | \exp(i\vec{Q} \cdot \vec{r}) \right\rangle. \tag{D.1}
$$

by considering the Schrödinger equations for the Coulomb scattering wave $\Psi_{\text{Coul}}^{(\pm)}$ and the plane wave. The matrix element can be evaluated by employing the integral representation of the confluent hypergeometric function [35] in the Coulomb scattering wave function (85). After the integration over the spatial coordinates, an integral remains that represents a hypergeometric function [35]. This technique is similar to the evaluation of Bremsstrahlung matrix elements [42,43]. One obtains

$$
\left\langle \Psi_{\text{Coul}}^{(-)}(\vec{k})|r^{-1}|\exp(i\vec{Q}\cdot\vec{r})\right\rangle = \frac{4\pi e^{-(\pi/2)\eta}}{(\vec{Q}-\vec{k})^2}\Gamma(1+i\eta)_{2}F_{1}(1,-i\eta;1;x) \tag{D.2}
$$

with the argument

$$
x = -2\frac{\vec{k} \cdot (\vec{Q} - \vec{k})}{(\vec{Q} - \vec{k})^2} \tag{D.3}
$$

in the hypergeometric function $_2F_1$ which reduces to the simple form

$$
{}_{2}F_{1}(1, -i\eta; 1; x) = (1 - x)^{i\eta}
$$
\n(D.4)

for the given parameters. Combining the above results the scattering amplitude assumes the form (86) when the relation

$$
\Gamma(1 + \mathrm{i}\eta) = C_0(\eta) \exp\left(\frac{\pi}{2}\eta + \mathrm{i}\sigma_0\right) \tag{D.5}
$$

with the Coulomb phase σ_0 is used.

One application of this formula is the calculation of the modified momentum amplitude (37) with a Coulomb scattering wave function. In this case one finds

$$
\mathcal{W}(\vec{q},\vec{k}) = -W(\vec{q})\frac{8\pi\eta k}{Q^2 - k^2} \frac{C_0(\eta)}{(\vec{Q}-\vec{k})^2} \exp\left\{i\sigma_0(\eta) + i\eta \ln\left[\frac{Q^2 - k^2}{(\vec{Q}-\vec{k})^2}\right]\right\},\qquad(D.6)
$$

where the indices Bb of η and k have been suppressed for clarity.

References

- [1] Homer, "Odyssey" VIII, 503.
- [2] C.E. Rolfs, W.S. Rodney, Cauldrons in the Cosmos, University of Chicago Press, Chicago, 1988.
- [3] F.K. Thielemann et al., Prog. Part. Nucl. Phys. 46 (2001) 1.
- [4] C. Rolfs, Prog. Part. Nucl. Phys. 46 (2001) 23.
- [5] LUNA Collaboration, Nucl. Phys. A 706 (2002) 203.
- [6] R. Bonetti et al., Phys. Rev. Lett. 82 (1999) 5205.
- [7] M. Junker et al., Phys. Rev. C 57 (1998) 2700.
- [8] H.J. Assenbaum, K. Langanke, C. Rolfs, Z. Phys. A 327 (1987) 461.
- [9] T.D. Shoppa, S.E. Koonin, K. Langanke, R. Seki, Phys. Rev. C 48 (1993) 837.
- [10] T.E. Liolios, Phys. Rev. C 63 (2001) 045801.
- [11] S. Austin, preprint nucl-th/0201010.
- [12] G. Baur, H. Rebel, J. Phys. G 20 (1994) 1; Annu. Rev. Nucl. Part. Sci. 46 (1996) 321.
- [13] A. Azhari, V. Burjan, F. Carstoiu, C.A. Gagliardi, V. Kroha, A.M. Mukhamedzhanov, F.M. Nunes, X. Tang, L. Trache, R.E. Tribble, Phys. Rev. C 63 (2001) 055803.
- [14] H.M. Xu, C.A. Gagliardi, R.E. Tribble, A.M. Mukhamedzhanov, N.K. Timofeyuk, Phys. Rev. Lett. 73 (1994) 2027.
- [15] H. Fuchs, H. Homeyer, Th. Lorenz, H. Oeschler, Phys. Lett. B 37 (1971) 285.
- [16] G. Baur, Phys. Lett. B 178 (1986) 135.
- [17] G. Baur, F. Rösel, D. Trautmann, R. Shyam, Phys. Rep. 111 (1984) 333.
- [18] G. Baur, D. Trautmann, Phys. Rep. 25C (1976) 293.
- [19] M. Jain, P.G. Roos, H.G. Pugh, H.D. Holmgren, Nucl. Phys. A 153 (1970) 49.
- [20] A. Musumarra, R.G. Pizzone, S. Blagus, M. Bogovac, P. Figuera, M. Lattuada, M. Milin, D. Miljanic, M.G. Pellegriti, D. Rendic, C. Rolfs, N. Soic, C. Spitaleri, S. Typel, H.H. Wolter, M. Zadro, Phys. Rev. C 64 (2001) 068801.
- [21] C. Spitaleri, S. Typel, R.G. Pizzone, M. Aliotta, S. Blagus, M. Bogovac, S. Cherubini, P. Figuera, M. Lattuada, M. Milin, D. Miljanic, A. Musumarra, M.G. Pellegriti, D. Rendic, C. Rolfs, S. Romano, N. Soic, A. Tumino, H.H. Wolter, M. Zadro, Phys. Rev. C 63 (2001) 055801.
- [22] S. Cherubini, V.N. Kondratyev, M. Lattuada, C. Spitaleri, D. Miljanic, M. Zadro, G. Baur, Ap. J. 457 (1996) 855.
- [23] M. Lattuada, R.G. Pizzone, S. Typel, P. Figuera, D. Miljanic, A. Musumarra, M.G. Pellegriti, C. Rolfs, C. Spitaleri, H.H. Wolter, Ap. J. 562 (2001) 1076.
- [24] M. Aliotta, C. Spitaleri, M. Lattuada, M. Musumarra, R.G. Pizzone, A. Tumino, C. Rolfs, F. Strieder, Eur. Phys. J. A 9 (2000) 435.
- [25] C. Spitaleri, A. Aliotta, S. Cherubini, M. Lattuada, D. Miljanic, S. Romano, N. Soic, M. Zadro, R.A. Zappala, Phys. Rev. C 60 (1999) 055802.
- [26] G. Calvi, S. Cherubini, M. Lattuada, S. Romano, C. Spitaleri, M. Aliotta, G. Rizzari, M. Sciuto, R.A. Zappala, V.N. Kondratyev, D. Miljanic, M. Zadro, G. Baur, O.Yu. Goryunov, A.A. Shvedov, Nucl. Phys. A 621 (1997) 139c.
- [27] C. Spitaleri, M. Aliotta, P. Figuera, M. Lattuada, R.G. Pizzone, S. Romano, A. Tumino, C. Rolfs, L. Gialanella, F. Strieder, S. Cherubini, A. Musumarra, D. Miljanic, S. Typel, H.H. Wolter, Eur. Phys. J. A 7 (2000) 181.

- [28] M.G. Pellegriti, PhD thesis, University of Catania, 2000.
- [29] S. Typel, H.H. Wolter, Few-Body Systems 29 (2000) 75.
- [30] R.G. Newton, Scattering Theory of Waves and Particles, McGraw-Hill, New York, 1966.
- [31] H. Fuchs, Nucl. Instr. Methods 200 (1982) 361.
- [32] N. Austern, in: Direct Nuclear Reaction Theories, Springer, New York, 1970, p. 487.
- [33] M. Gell-Mann, M.L. Goldberger, Phys. Rev. 91 (1953) 398.
- [34] A. Kasano, M. Ichimura, Phys. Lett. B 115(1982) 81.
- [35] M. Abramowitz, I.A. Stegun, Handbook of Mathematical Functions, Dover, New York, 1970.
- [36] P. Braun-Munzinger, H.L. Hearney, Nucl. Phys. A 233 (1974) 381.
- [37] R. Shyam, M.A. Nagarajan, Ann. Phys. (N.Y.) 163 (1985) 285.
- [38] A. Tumino, et al., Phys. Rev. C, submitted.
- [39] K. Alder, A. Bohr, T. Huus, B. Mottelson, A. Winther, Rev. Mod. Phys. 28 (1956) 432.
- [40] G. Baur, D. Trautmann, J. Appl. Math. Phys. (ZAMP) 25(1974) 9.
- [41] G. Baur, D. Trautmann, Z. Phys. 267 (1974) 103.
- [42] A. Nordsieck, Phys. Rev. 93 (1954) 785.
- [43] R. Shyam, P. Banerjee, G. Baur, Nucl. Phys. A 540 (1992) 341.