

# Current Status of LSND & BooNE

- **LSND**

- Preliminary  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  results from 1993–1998

- Published  $\nu_\mu \rightarrow \nu_e$  results from 1993–1995

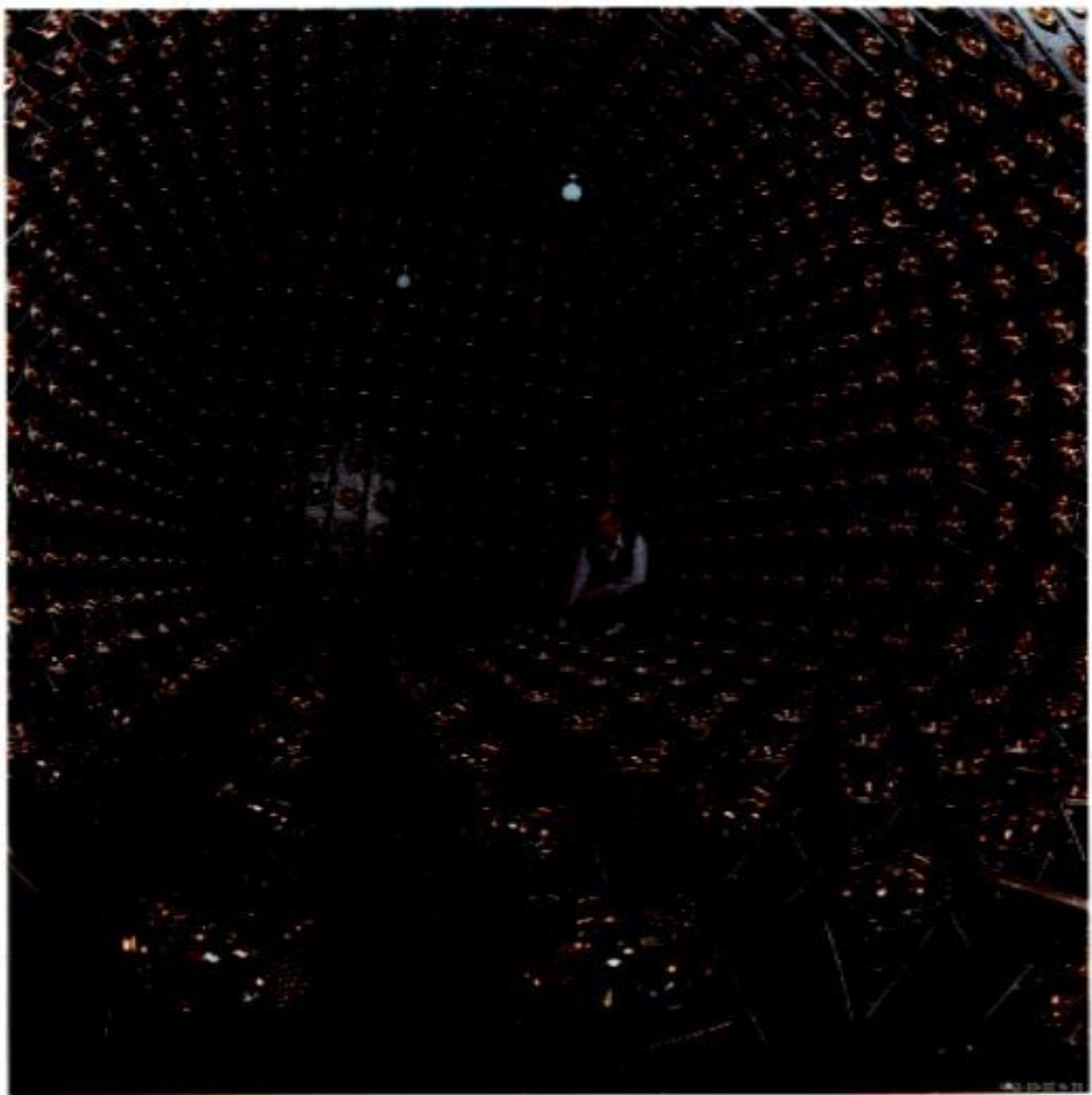
- Final global analysis results ( $\nu$  oscillations,  
 $\nu$  C,  $\nu$  e,  $\nu$  p) by the end of 1999

- **BooNE**

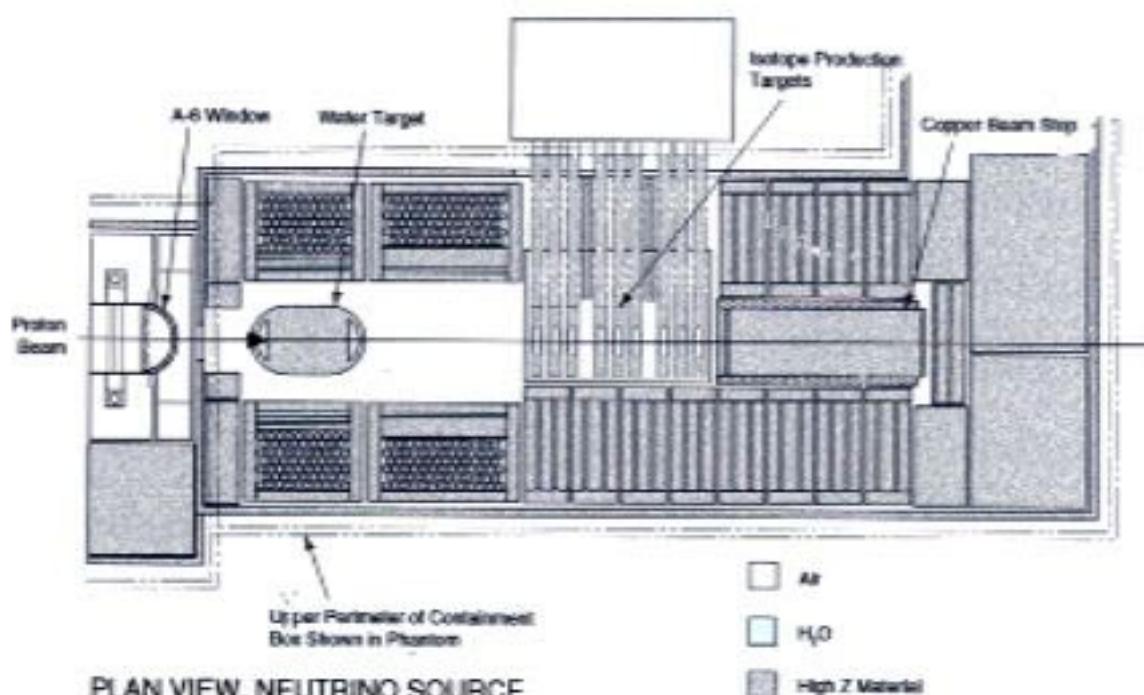
- Definitive test of the LSND signal

- Construction begins in October, 1999

- Data taking begins in December, 2001



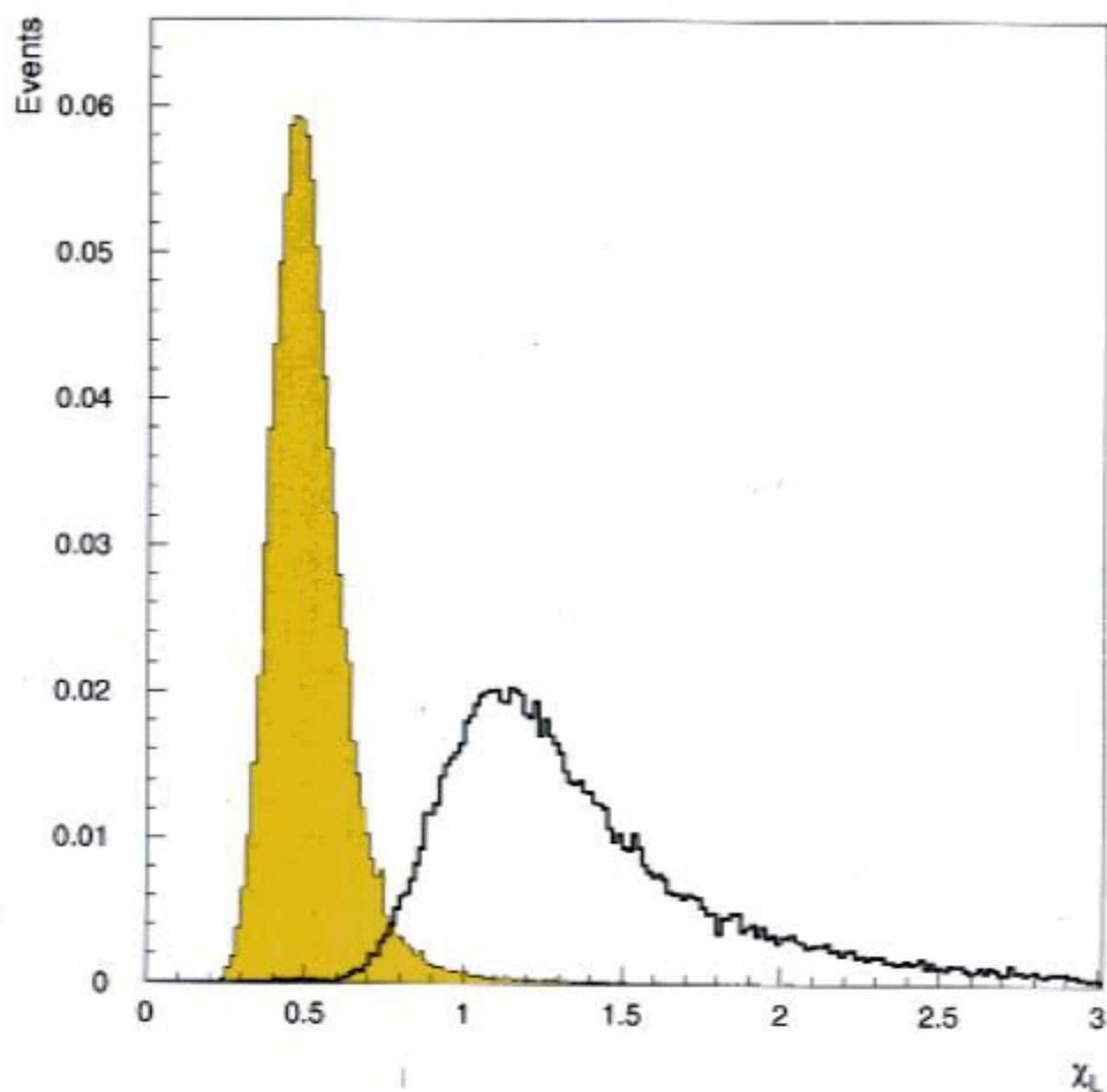
# Target and Beam Dump



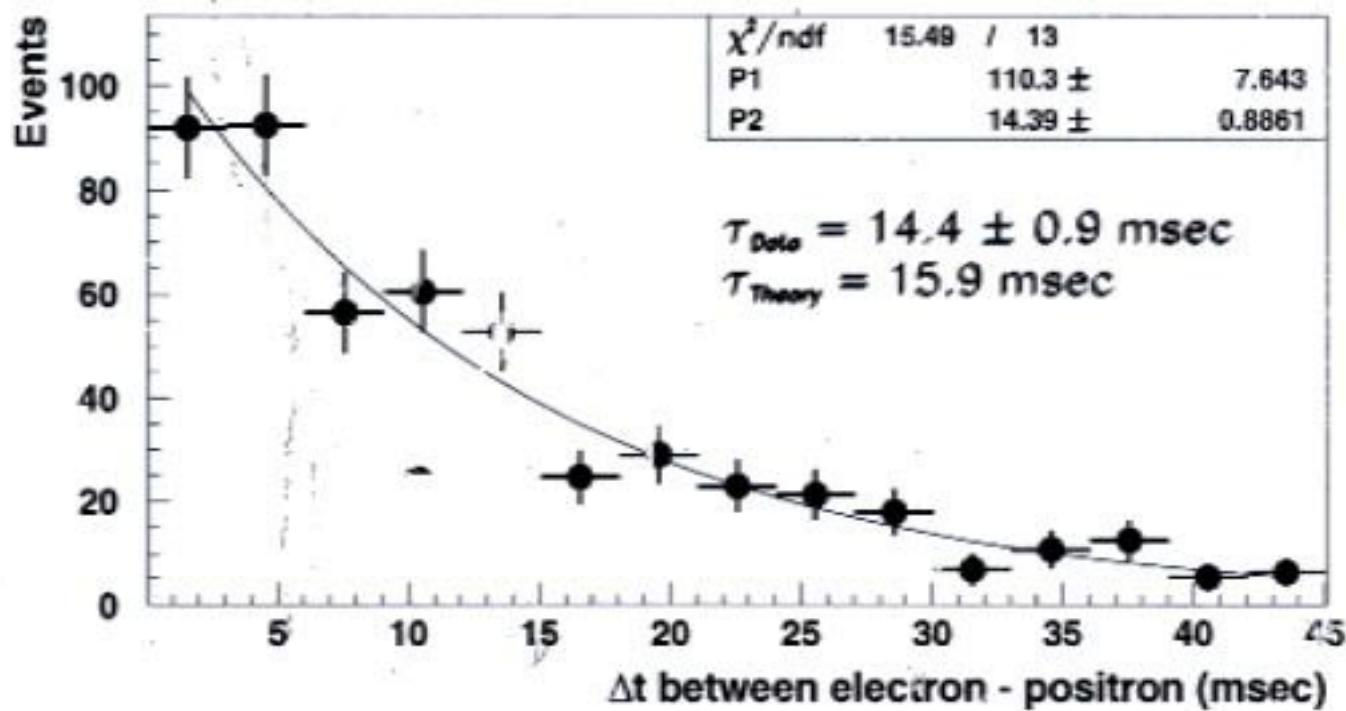
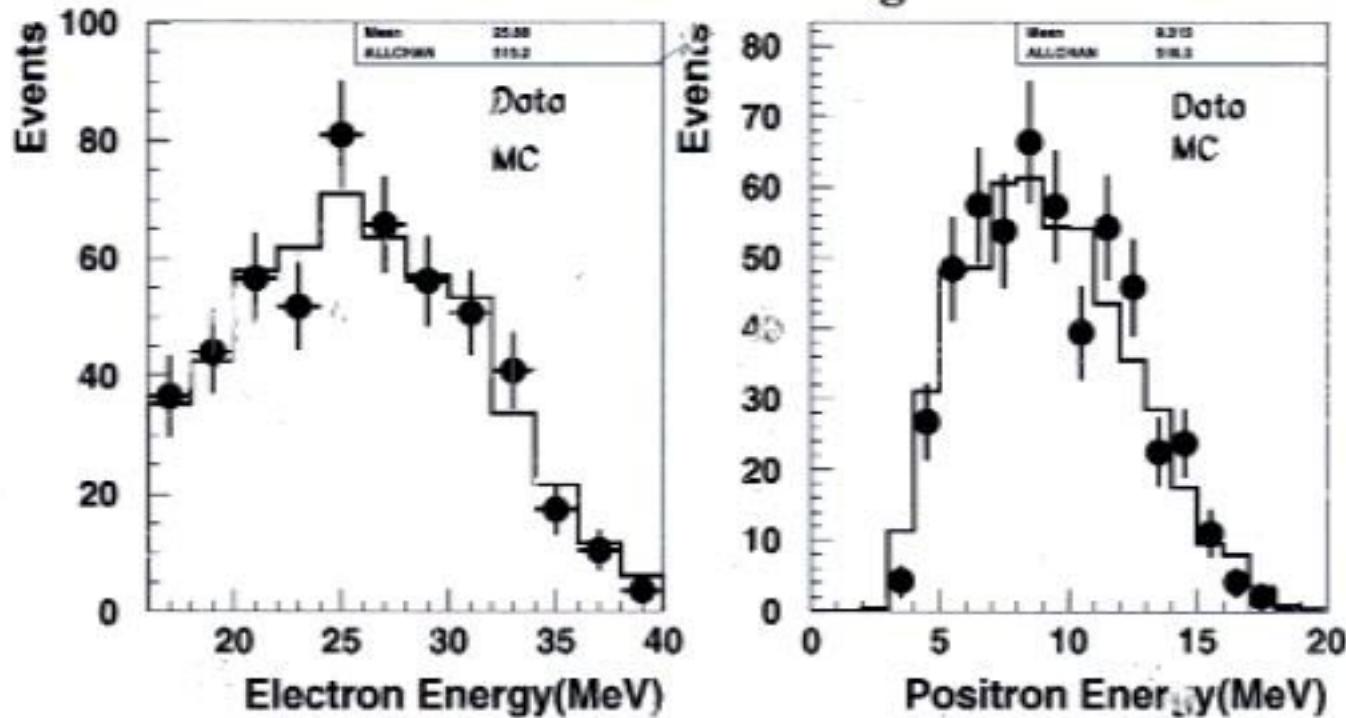
$$\pi^+ \rightarrow \mu^+ \nu_\mu$$

$$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$$

# Particle Identification

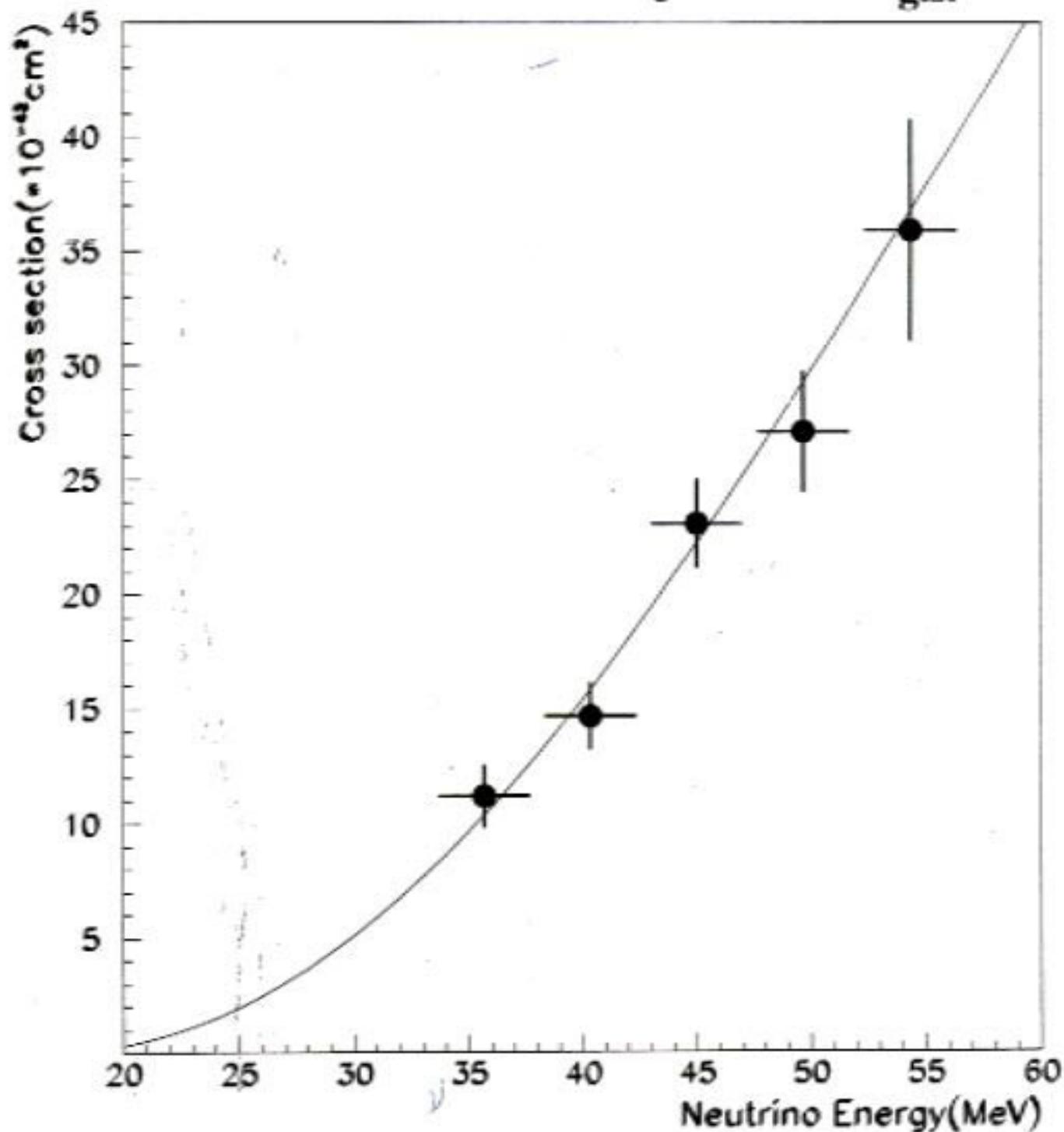


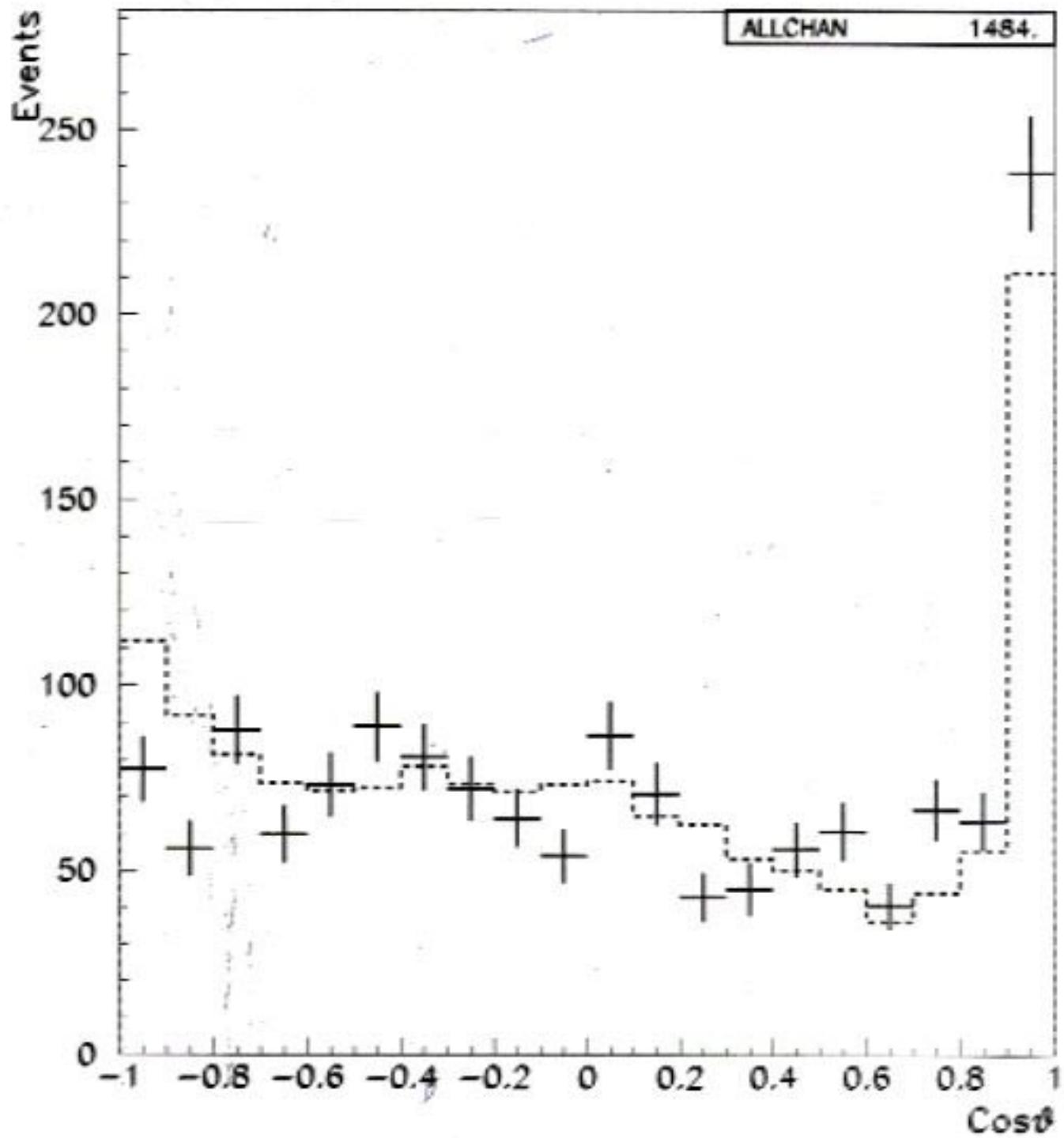
$\nu_e C \rightarrow e^- N_{g.s.}$



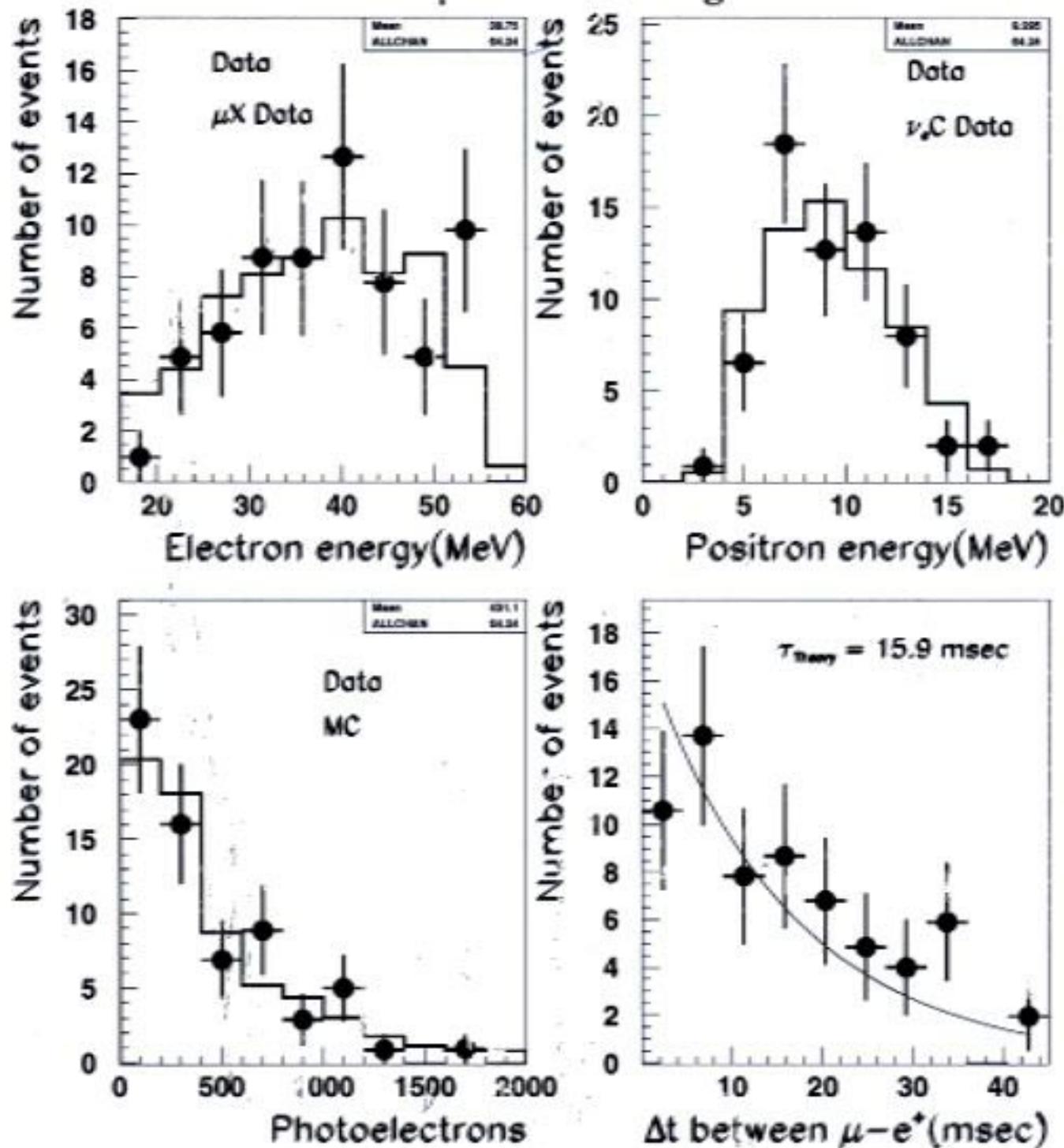
$$\begin{aligned} \tilde{\sigma} &= (9.1 \pm 0.4 \pm 0.9) \times 10^{-42} \text{ cm}^2 \quad (\text{LSND}) \\ &= (9.4 \pm 0.5 \pm 0.8) \times 10^{-42} \text{ cm}^2 \quad (\text{KARMEN}) \\ &= (9.3 - 9.4) \times 10^{-42} \text{ cm}^2 \quad (\text{theory}) \end{aligned}$$

## Cross section of $\nu_e C \rightarrow e^- N_{g.s.}$





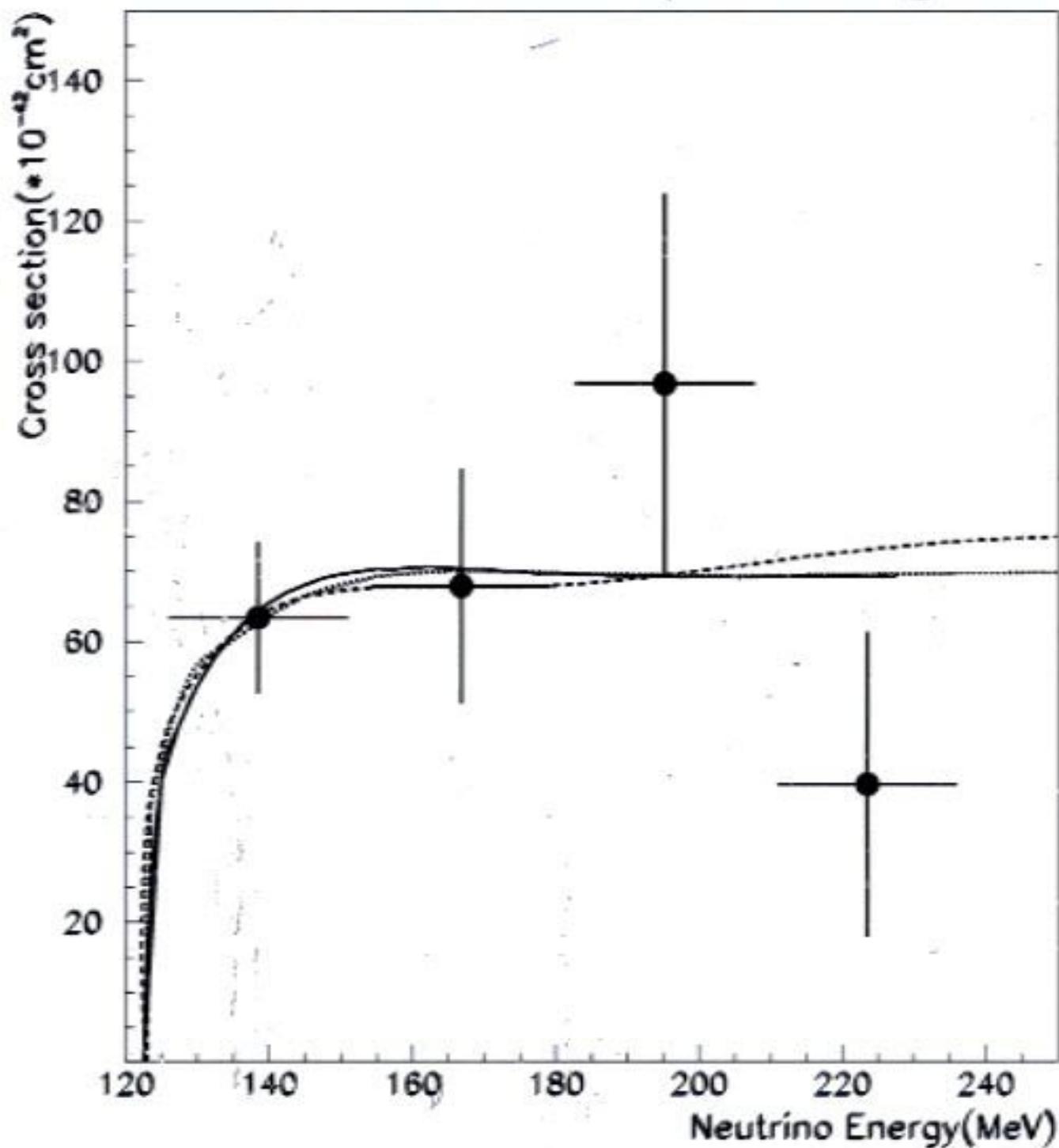
# $\nu_\mu C \rightarrow \mu^- N_{g.s.}$



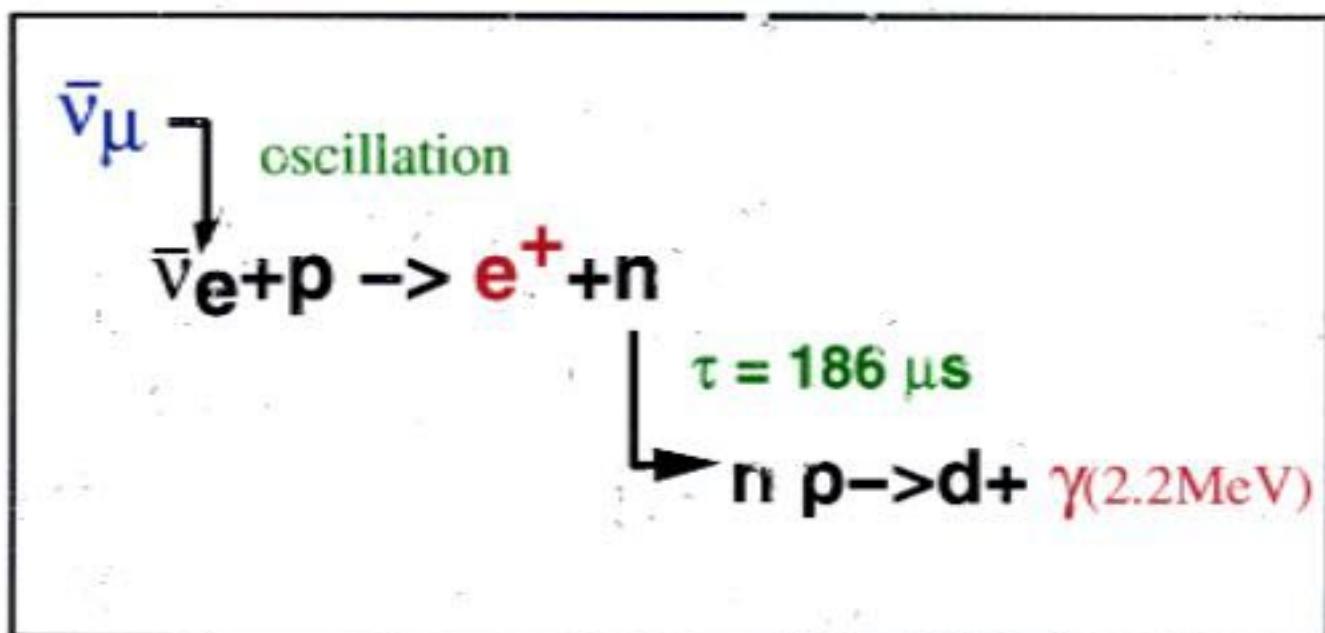
$$\sigma = (6.6 \pm 1.0 \pm 1.0) \times 10^{-41} \text{ cm}^2 \quad (\text{LSND})$$

$$(6.3 - 6.6) \times 10^{-41} \text{ cm}^2 \quad (\text{theory})$$

## Cross section of $\nu_\mu C \rightarrow \mu^+ N_{g.s.}$



## $\nu$ Oscillation Events Signature

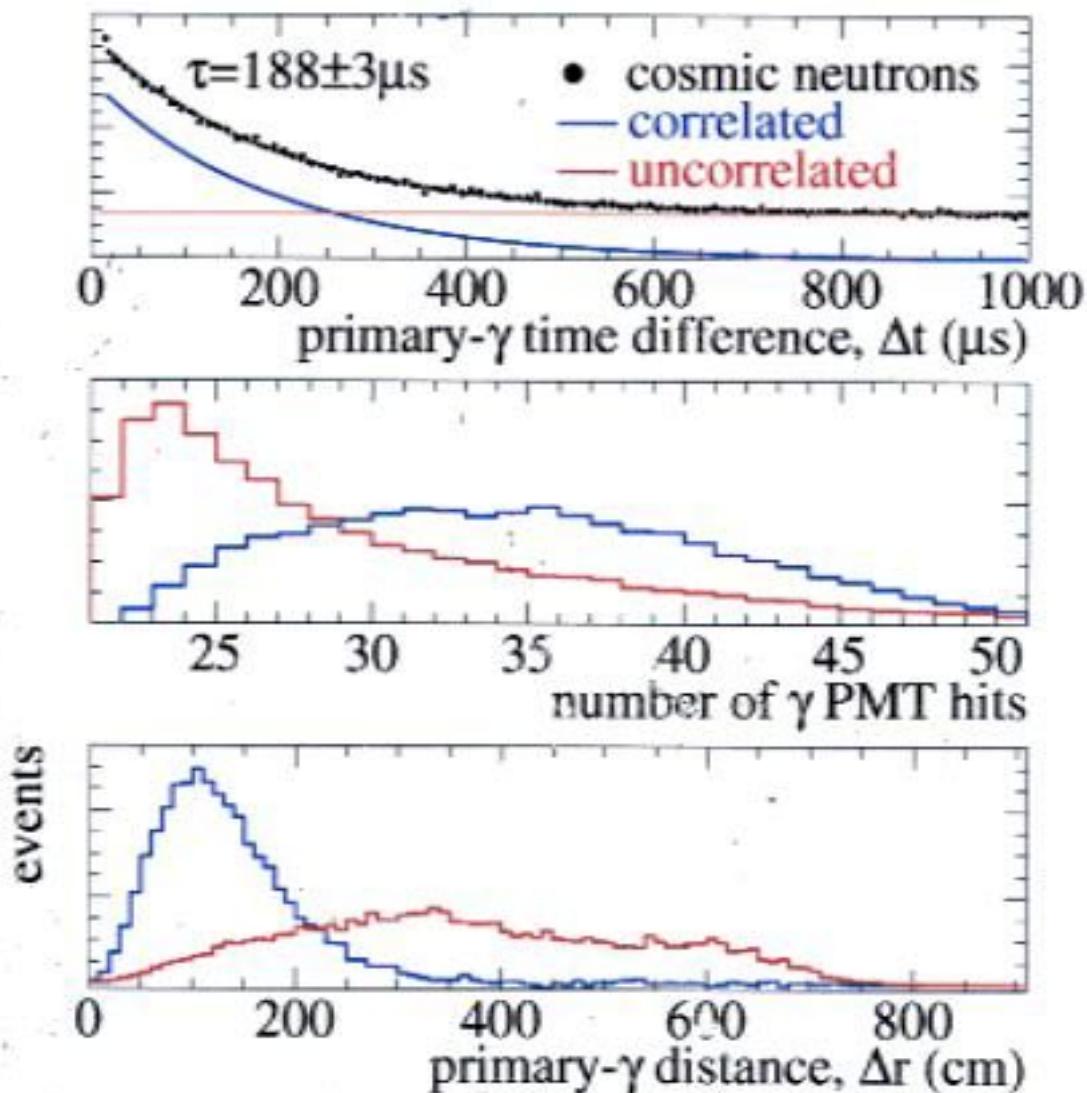


### - $e^+$ selection

Particle ID :	cut cosmic neutrons
$d_{\text{PMT}} > 35\text{cm}$ :	fiducial volume
$\Delta t_{\text{previous}} > 20\mu\text{s}$ :	cut cosmics
$\Delta t_{\text{next}} > 8\mu\text{s}$ :	cut muons
$n_\gamma < 2$ :	cut cosmic neutrons
< 4 veto hits:	cut cosmics
$S > 0.5$ :	cut cosmics
Efficiency :	0.37

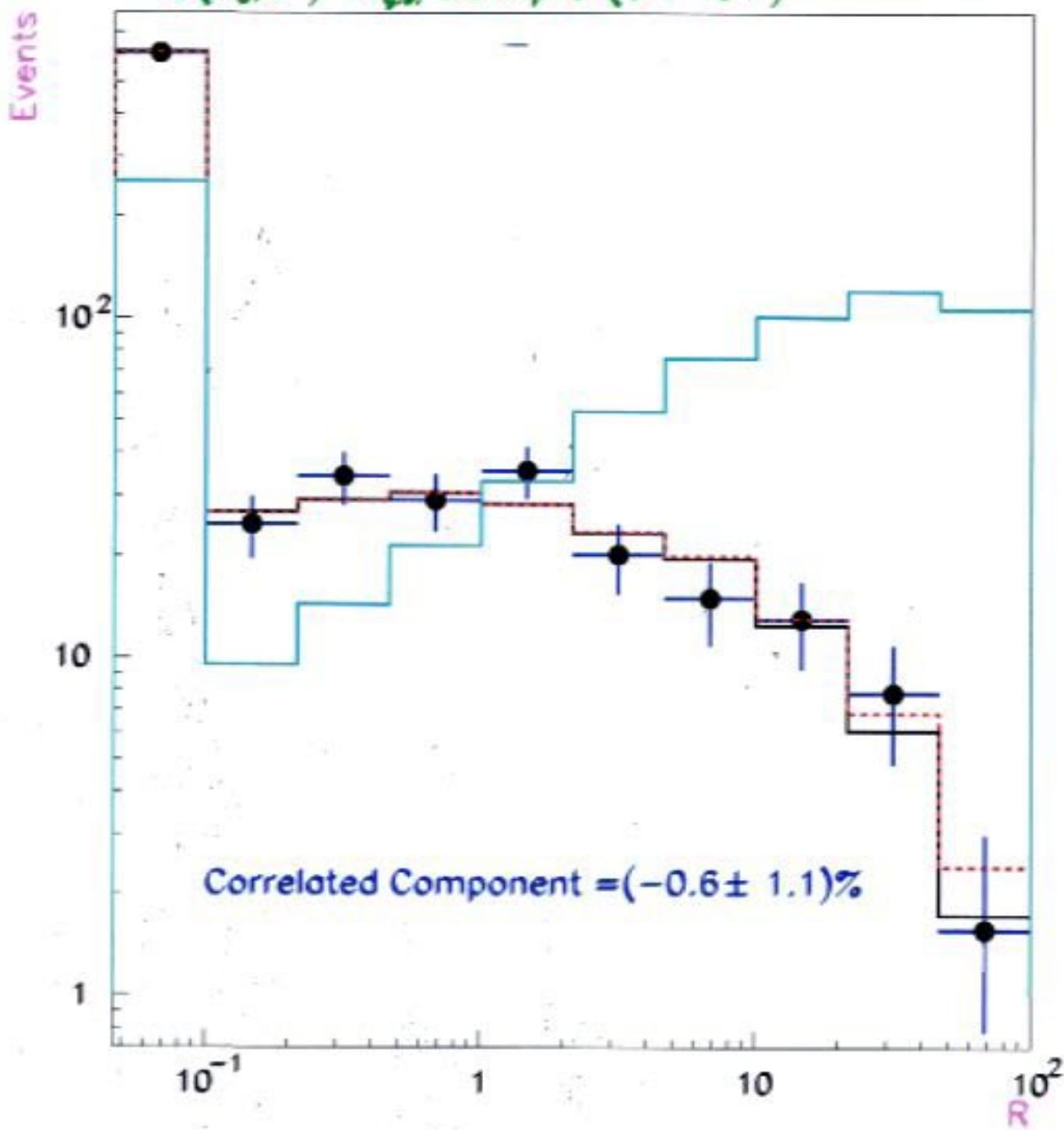
### - $\gamma$ selection : Likelihood ratio, R method

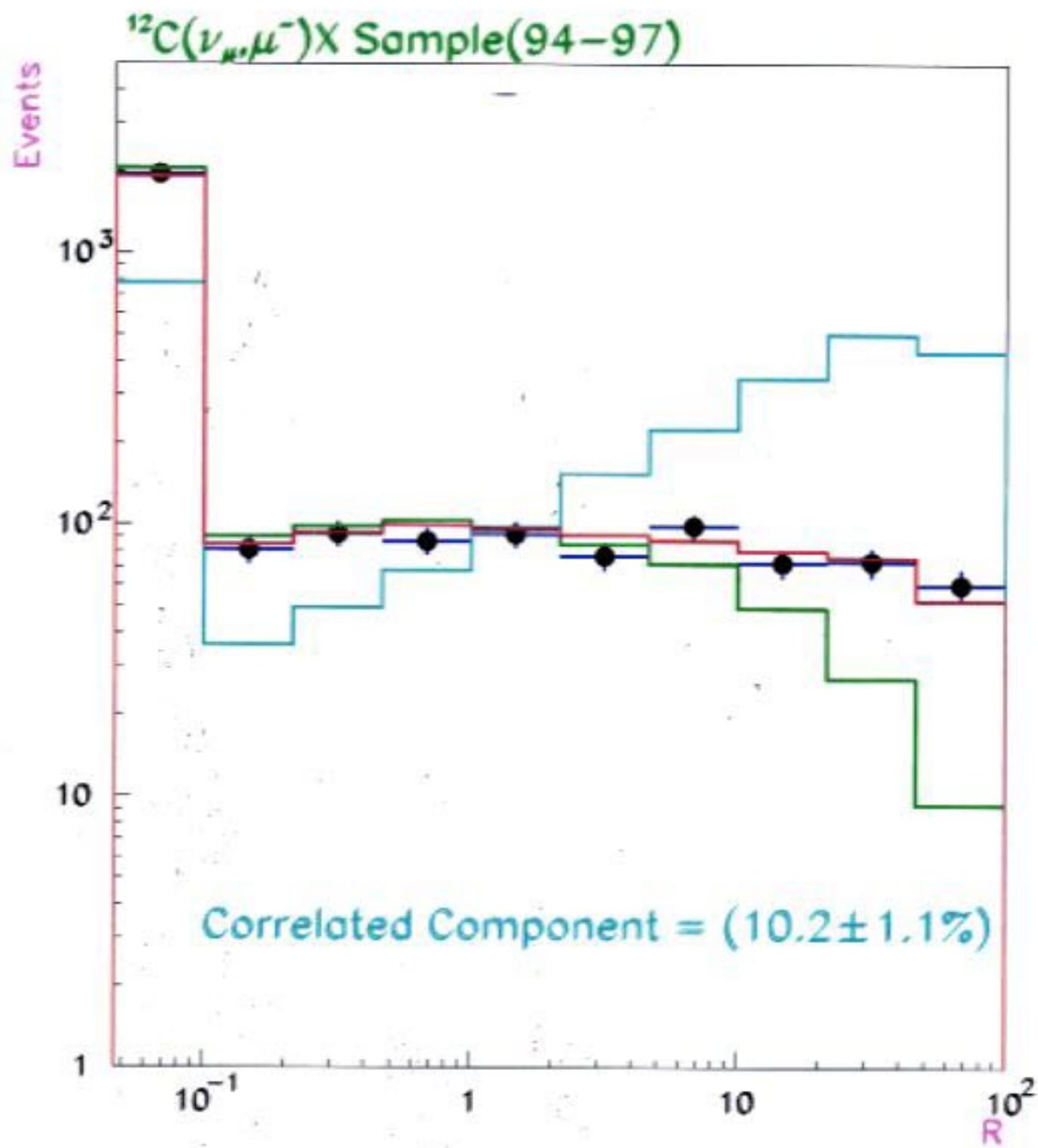
# Correlated $\gamma$ ID

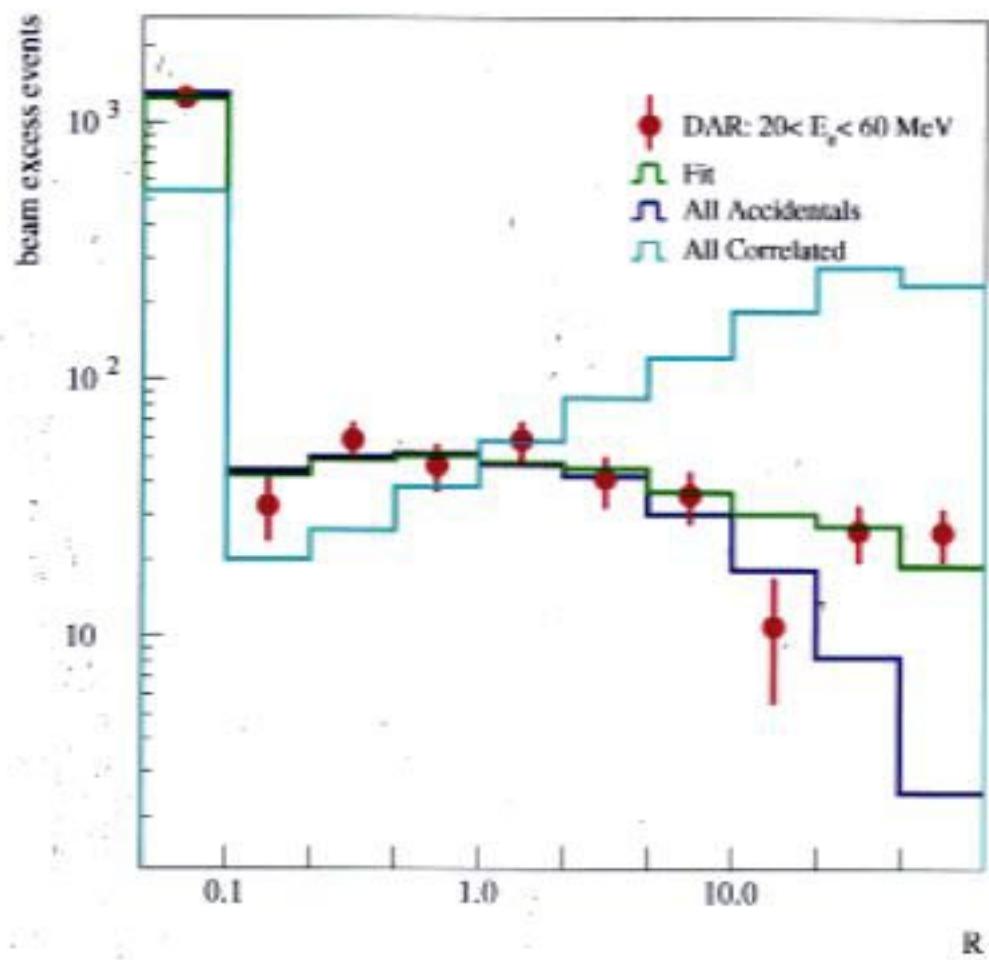


- define  $L = P(\Delta t)$   $P(\# \text{ hits})$   $P(\Delta r)$
- form  $R = L(\text{correlated})/L(\text{accidental})$

$^{12}\text{C}(\nu_e, e^-) ^{12}\text{N}_{\text{gs}}$ , Sample (94–97)







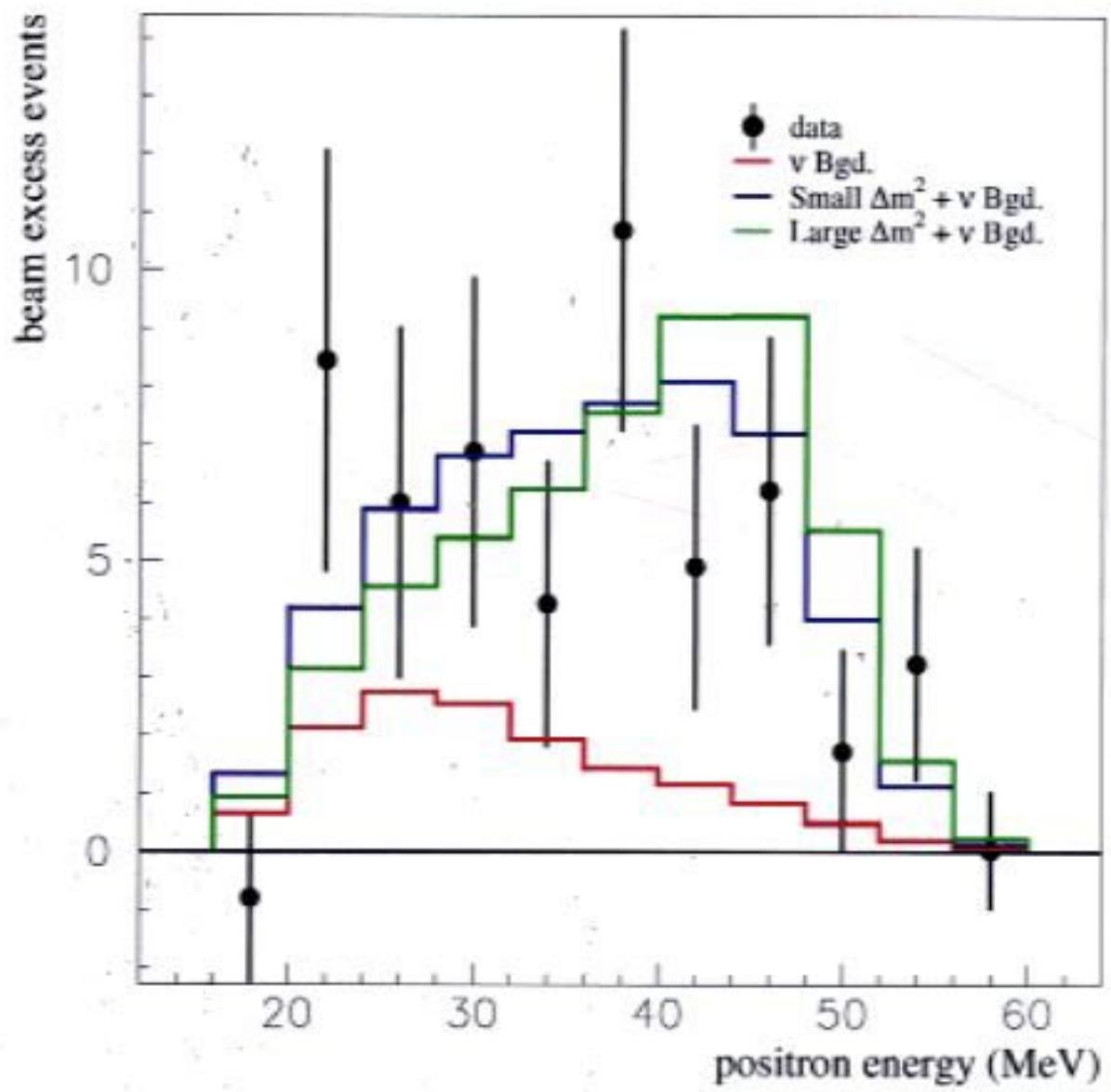
# Preliminary LSND $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ Results for 1993–1998

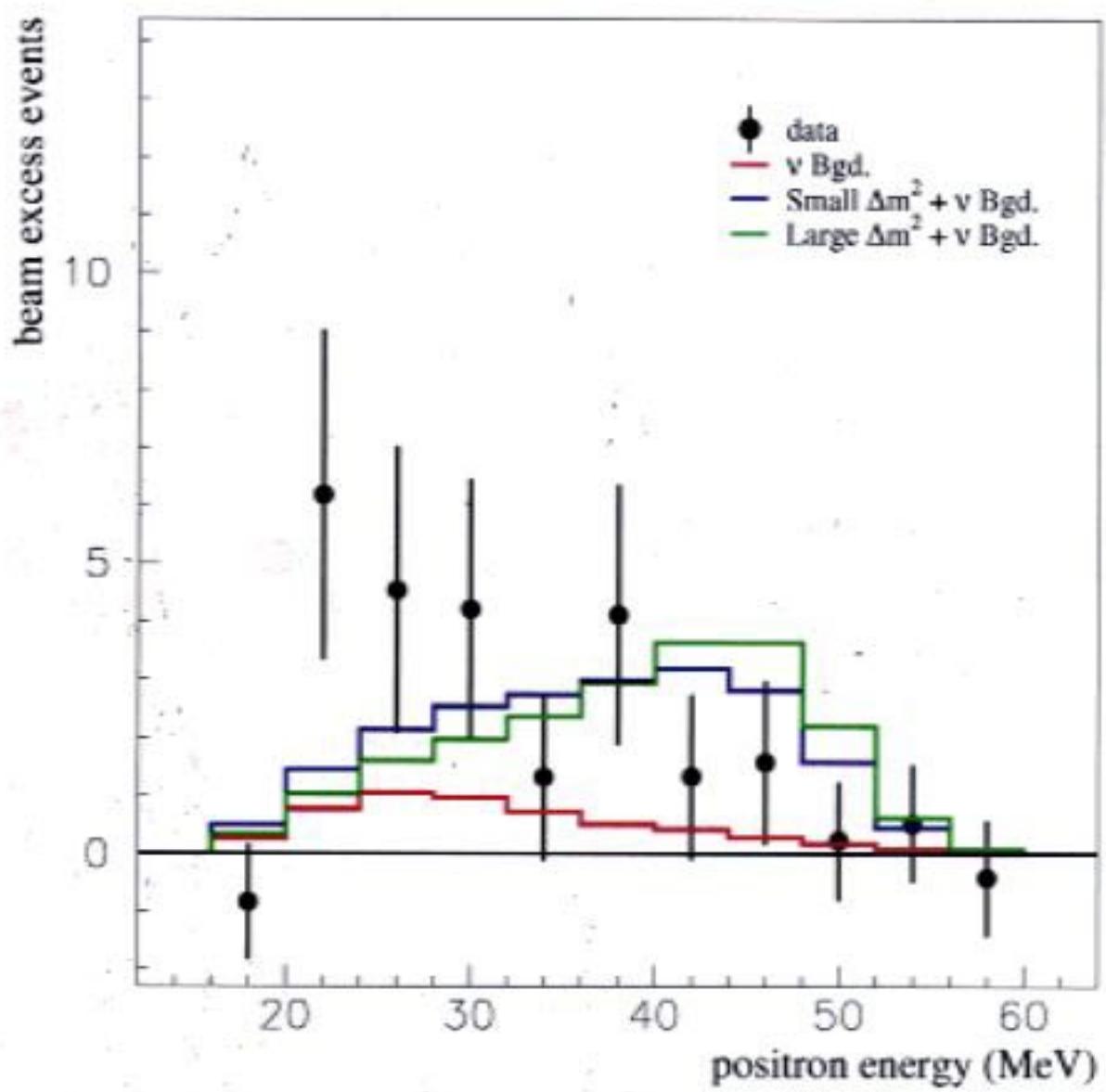
Selection	Beam On	Beam Off	$\nu$ Background	Total Excess
R>30 20<E<60	70	17.7±1.0	12.8±1.7	39.5±8.8
R>30 36<E<60	33	6.2±0.6	3.3±0.7	23.5±5.8

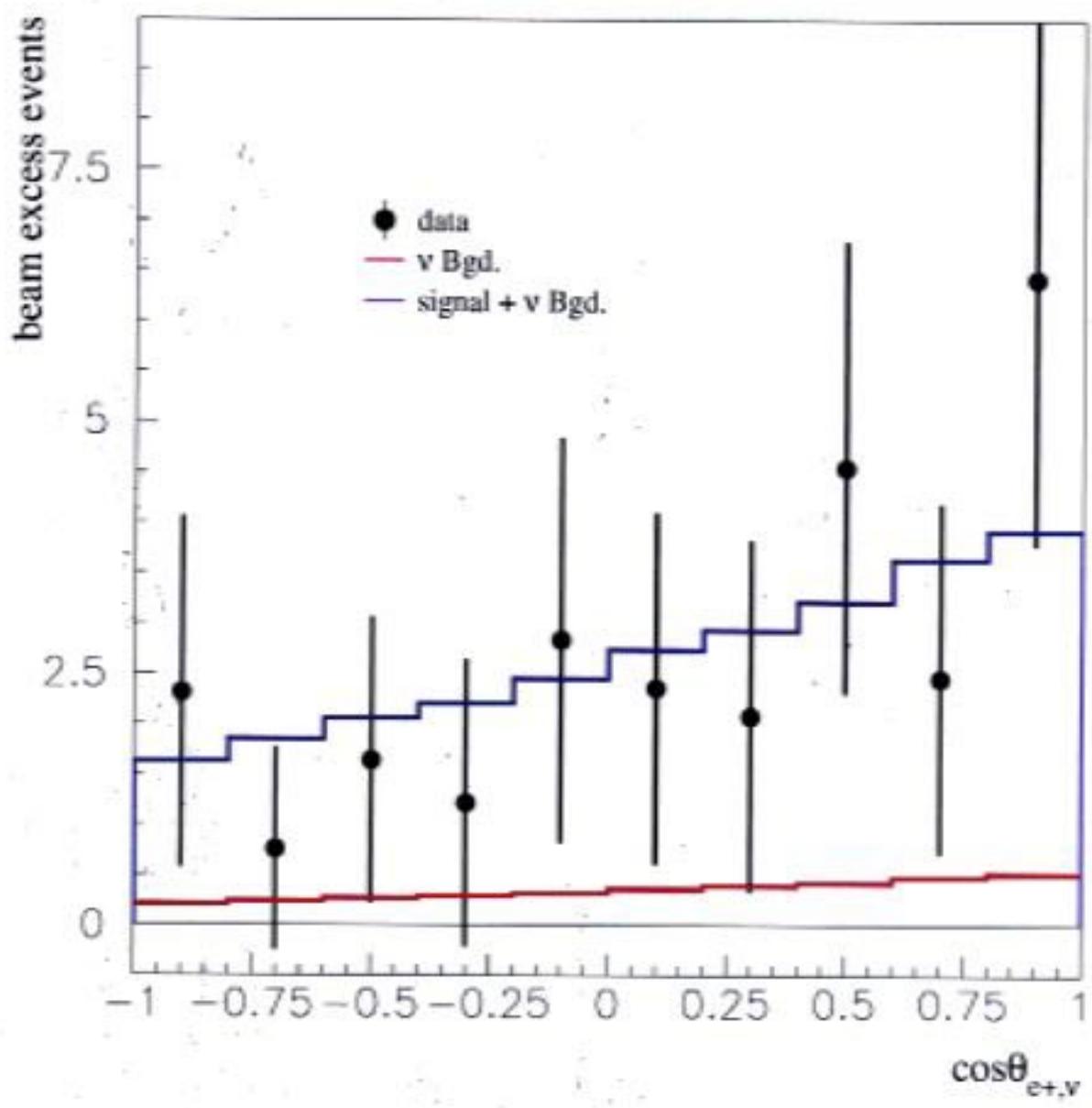
	20–36 MeV	36–60 MeV	20–60 MeV
1993–1995 (low,high $\Delta m^2$ )	3.7±4.2 (11.0,7.1)	17.4±4.7 (14.1,16.6)	21.1±6.3
1996–1998 (low,high $\Delta m^2$ )	12.3±5.1 (6.7,4.7)	6.1±3.4 (7.7,11.0)	18.4±6.1
1993–1998 (low,high $\Delta m^2$ )	16.0±6.6 (17.7,11.9)	23.5±5.8 (21.8,27.6)	39.5±8.8

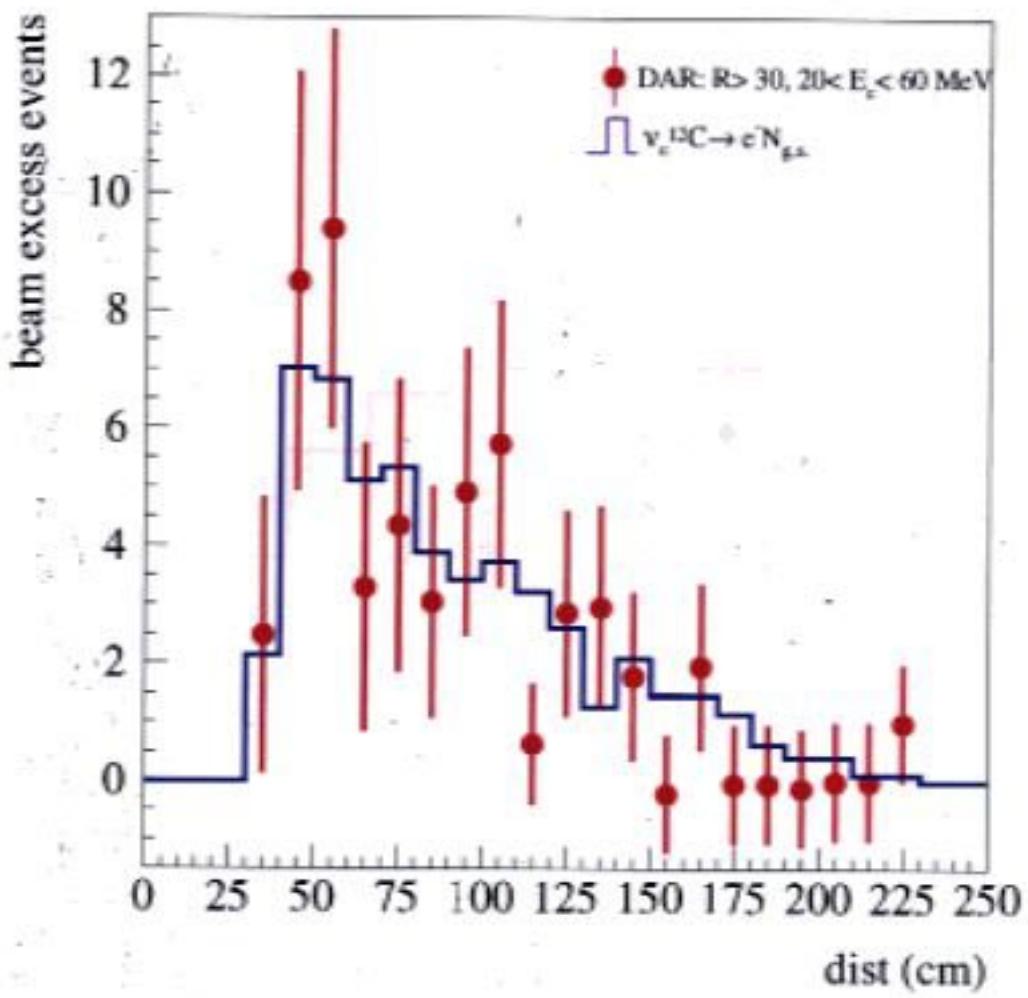
Data Sample	Fitted Excess	Total Excess	Oscillation Prob.
1993–1995	63.5±20.0	51.0±20.2	(0.31±0.12±0.05)%
1993–1998	111.8±25.6	90.9±26.1	(0.33±0.09±0.05)%

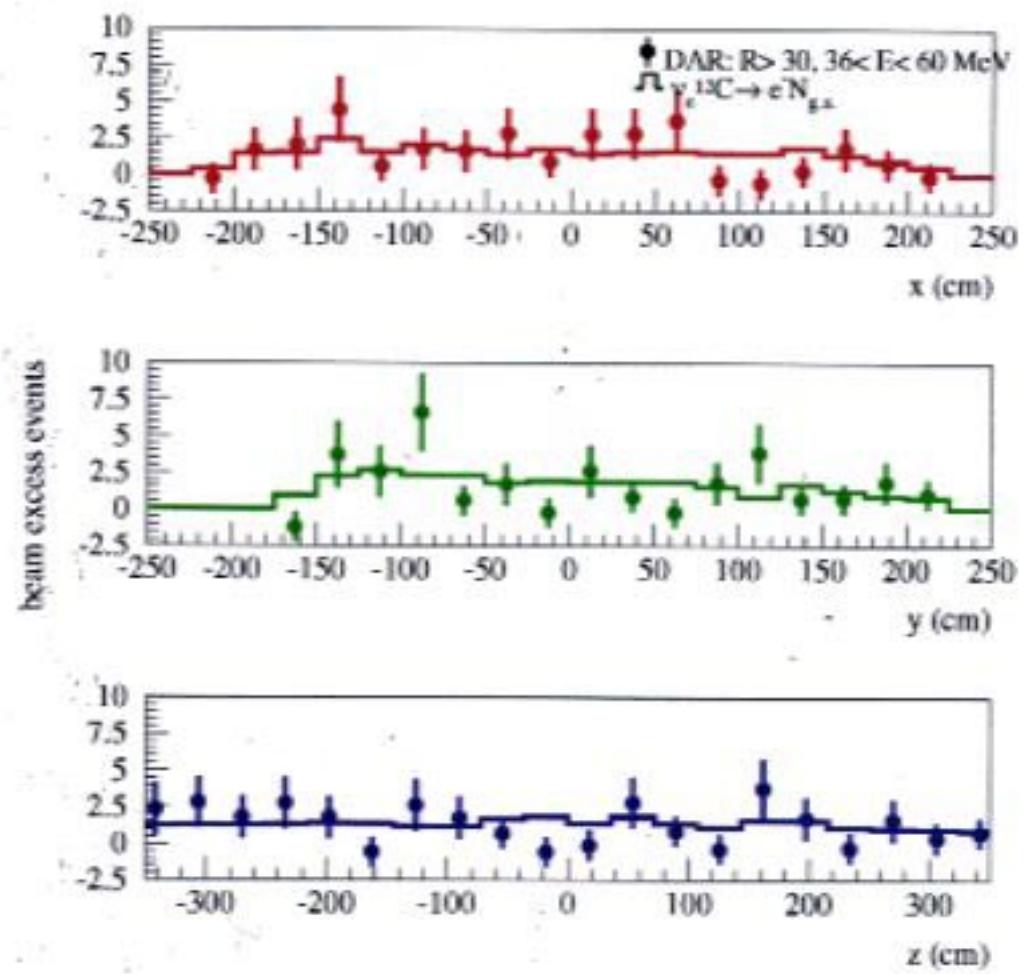
N.B.: The absolute electron efficiency, energy calibration, duty factor, and neutrino flux have been estimated for the 1996 through 1998 data and are subject to change. A global analysis of all of the data (DAR and DIF) is underway.

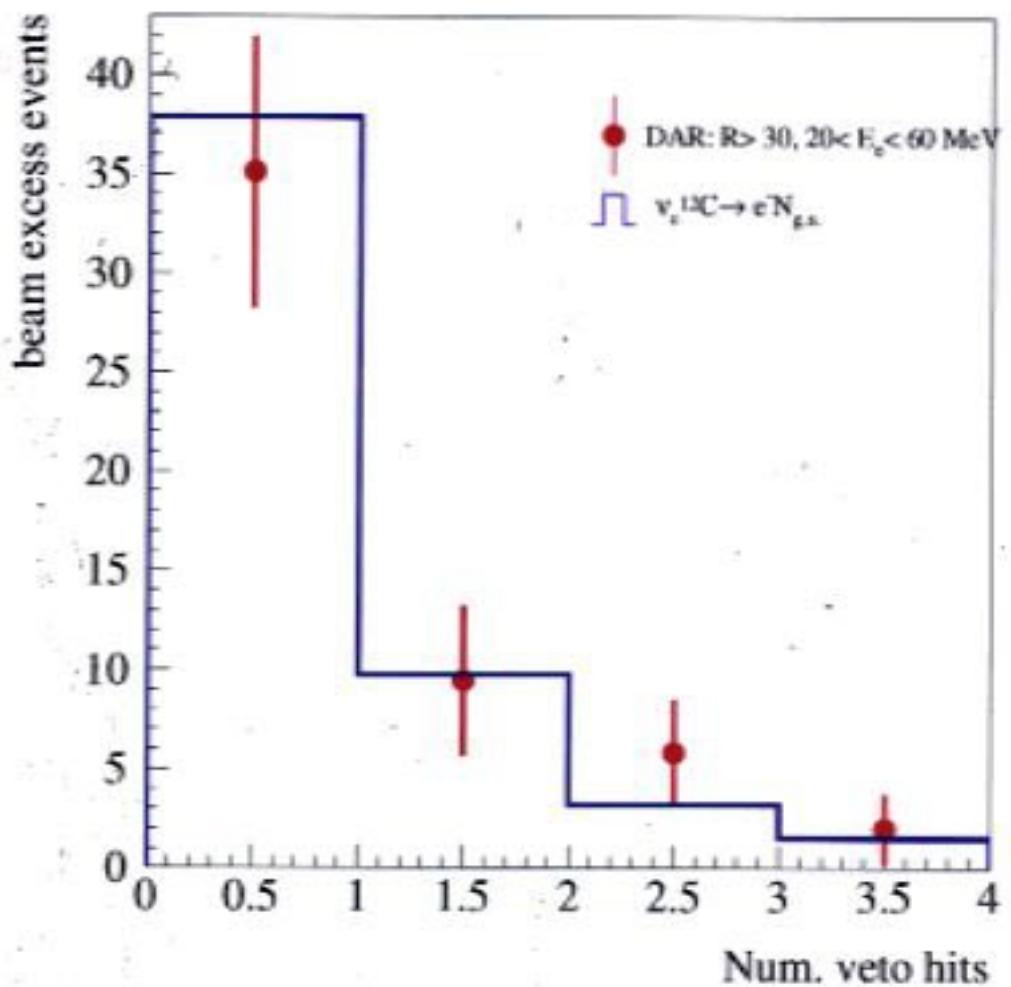


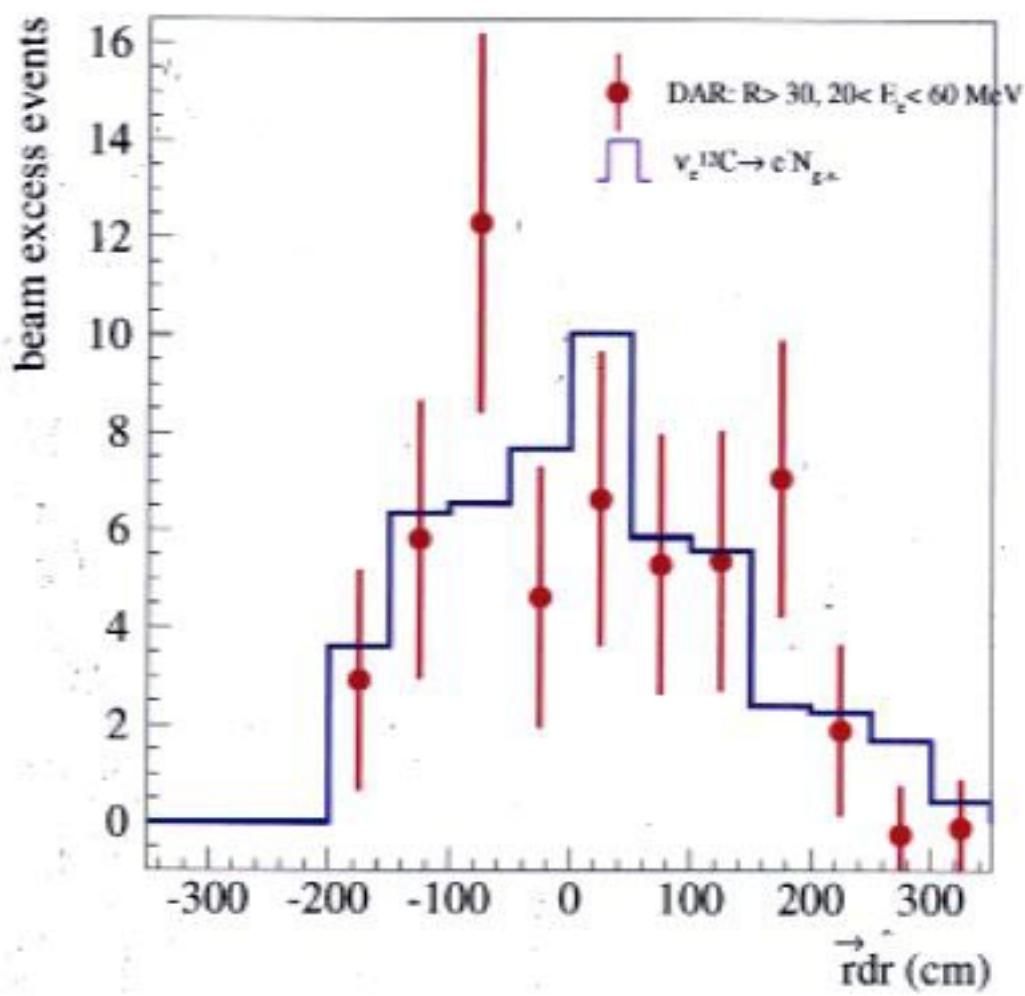












## Events with Multiple Gammas

### I. ( $20 < E < 60$ ) & ( $R > 30$ )

# $\gamma = 1 \Rightarrow$	70 on, 308 off,	52.3+-8.4 excess
# $\gamma > 1 \Rightarrow$	6 on, 99 off,	0.3+-2.5 excess
<b>Ratio =&gt;</b>	0.09      0.32	0.01+-0.05

### II. ( $36 < E < 60$ ) & ( $R > 30$ )

# $\gamma = 1 \Rightarrow$	33 on, 113 off,	26.8+-5.8 excess
# $\gamma > 1 \Rightarrow$	1 on, 41 off,	-1.4+-1.1 excess
<b>Ratio =&gt;</b>	0.03      0.36	-0.05+-0.04

We expect that for primary neutrons, the events would have multiple gammas with Ratio = 0.60. Therefore, our signal is NOT due to primary neutrons!

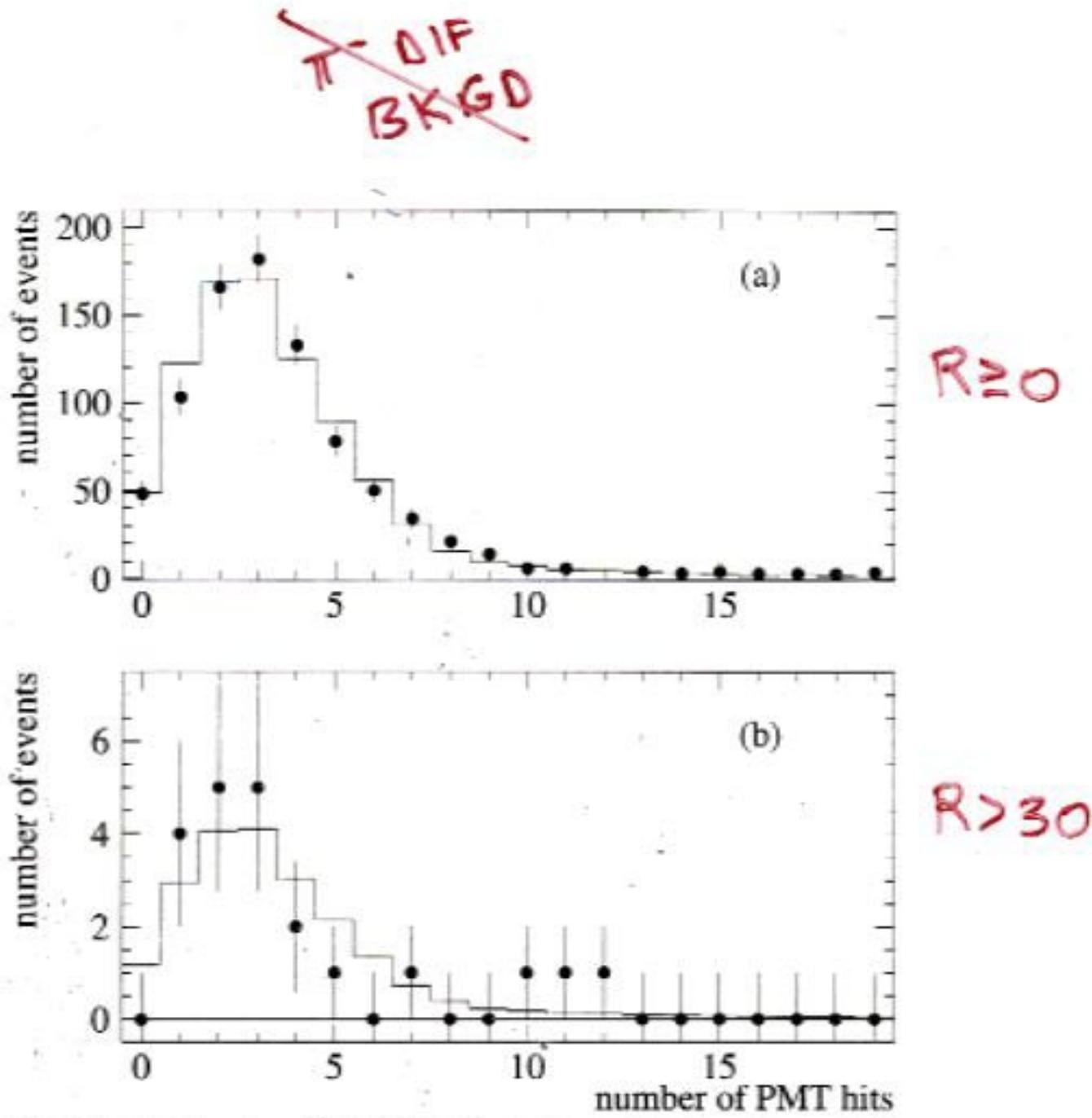
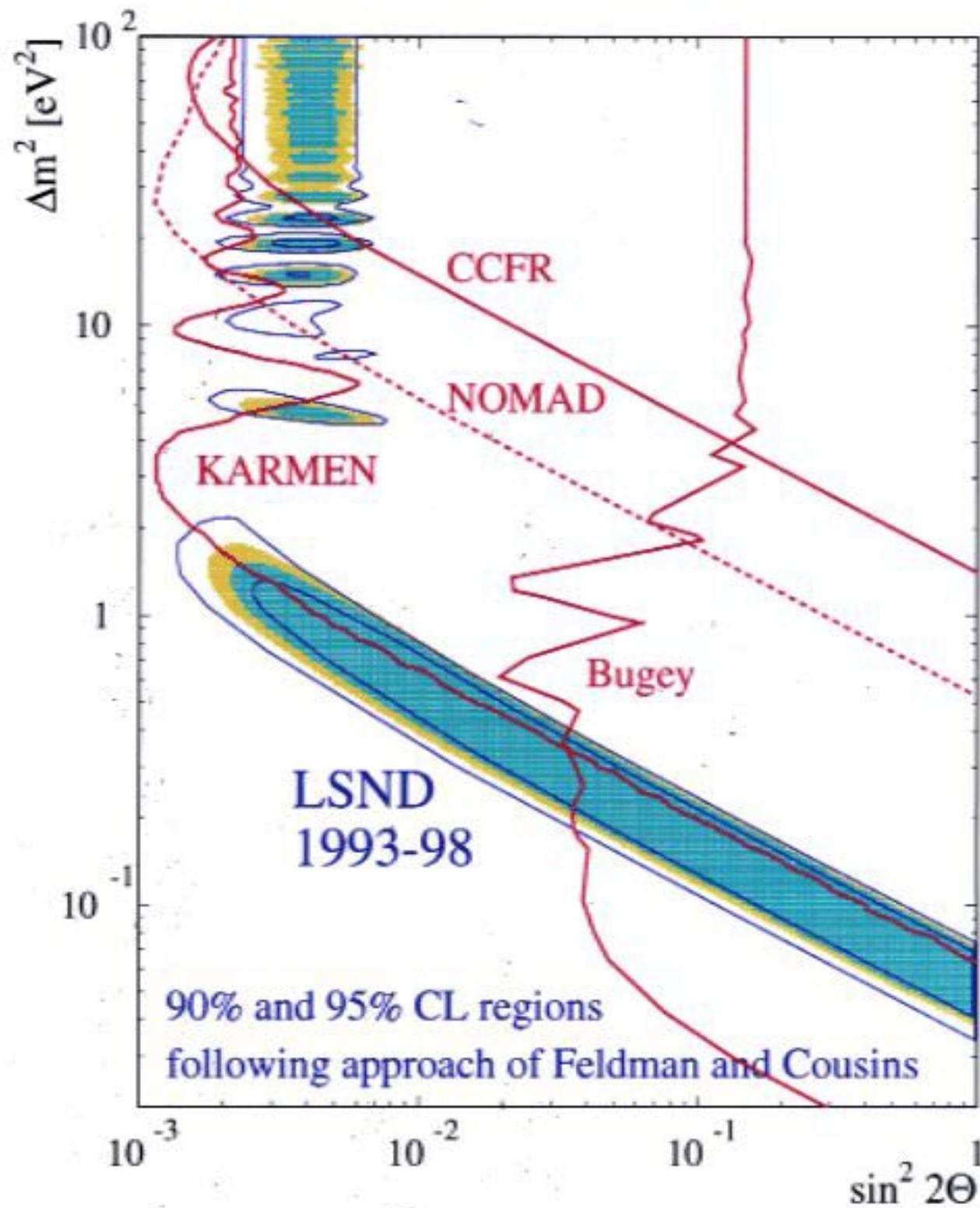
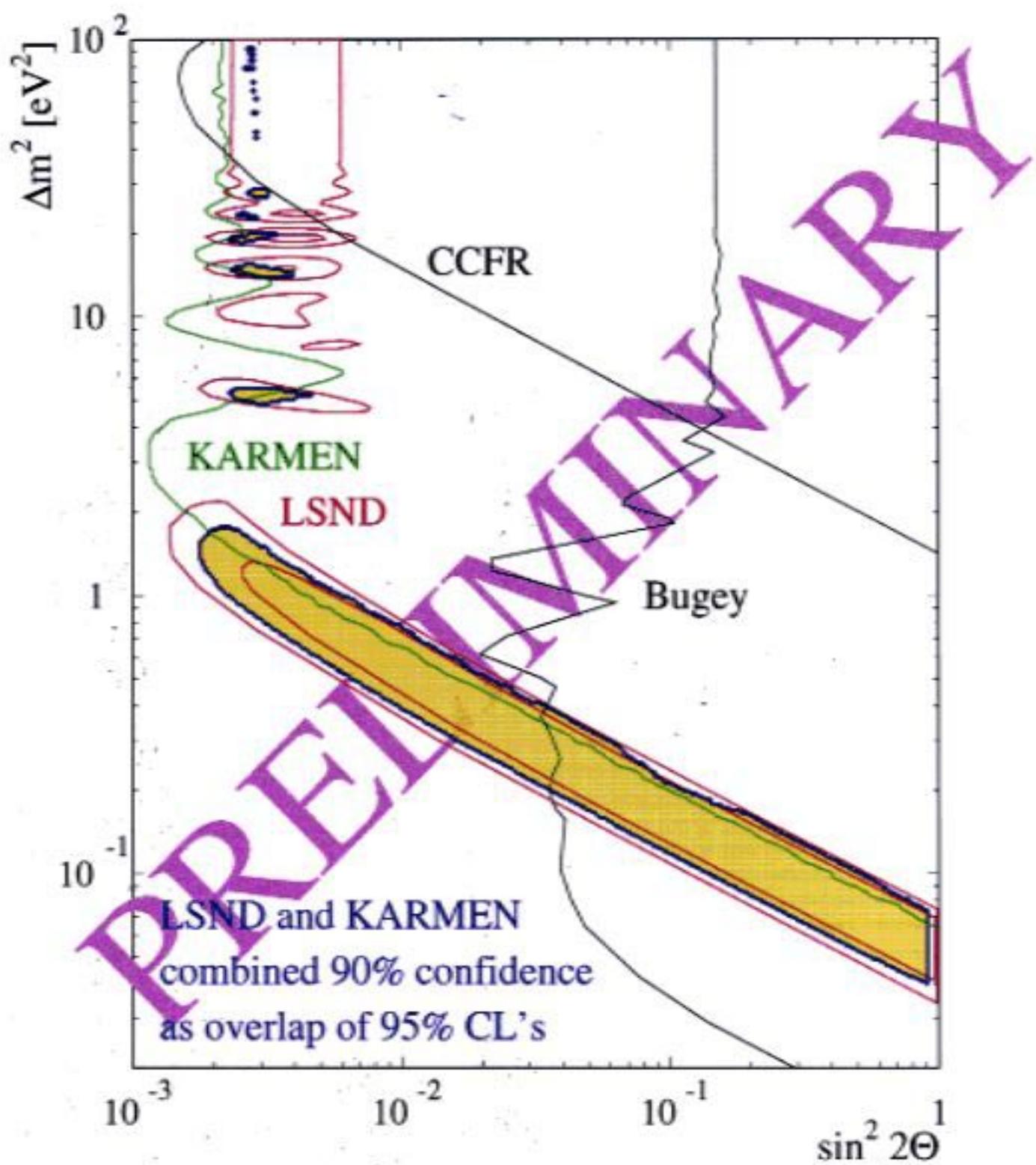


FIG. 32. The total number of hit PMTs in the detector tank for the extra events that occur  $0 - 3\mu s$  and  $3 - 6\mu s$  prior to oscillation candidate events. The candidates are in the  $25 < E_\nu < 60$  MeV energy range with (a)  $R \geq 0$  and (b)  $R > 30$ . The data points are the beam-on events, while the solid curve is what is expected from random PMT hits as determined from the sample of laser calibration events.

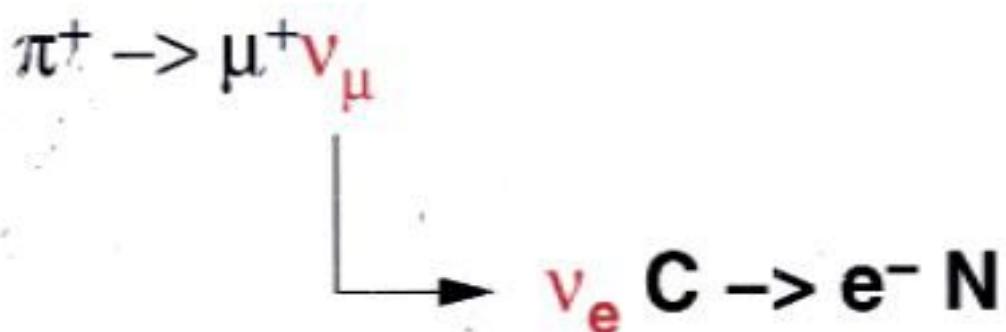
## $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ Checklist

- OK** 1. Spatial Distribution
- OK** 2. Energy Distribution
- OK** 3. Correlated  $\gamma$  Distribution
- OK** 4. Angular Distribution
- OK** 5. Veto Distribution
- OK** 6. Events with Multiple  $\gamma$ s
- OK** 7. Hit PMTs in Lookback
- OK** 8. H<sub>2</sub>O Target vs High z Target



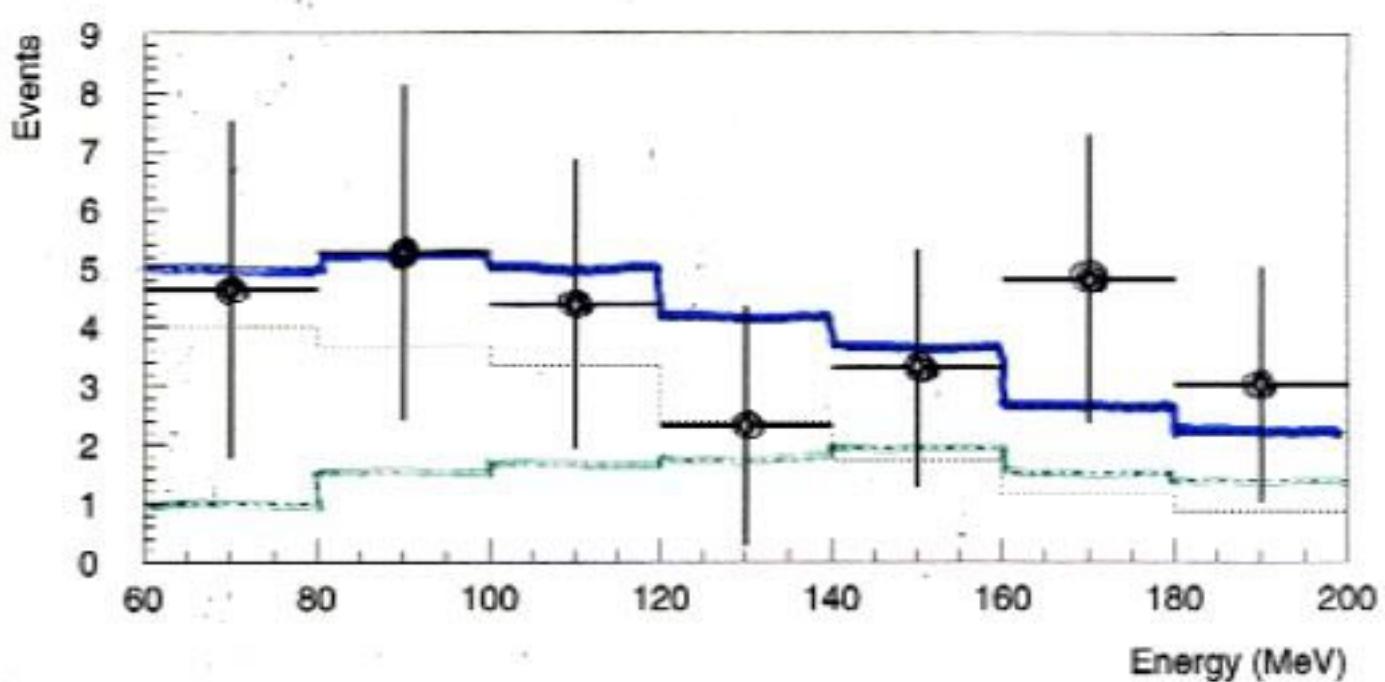


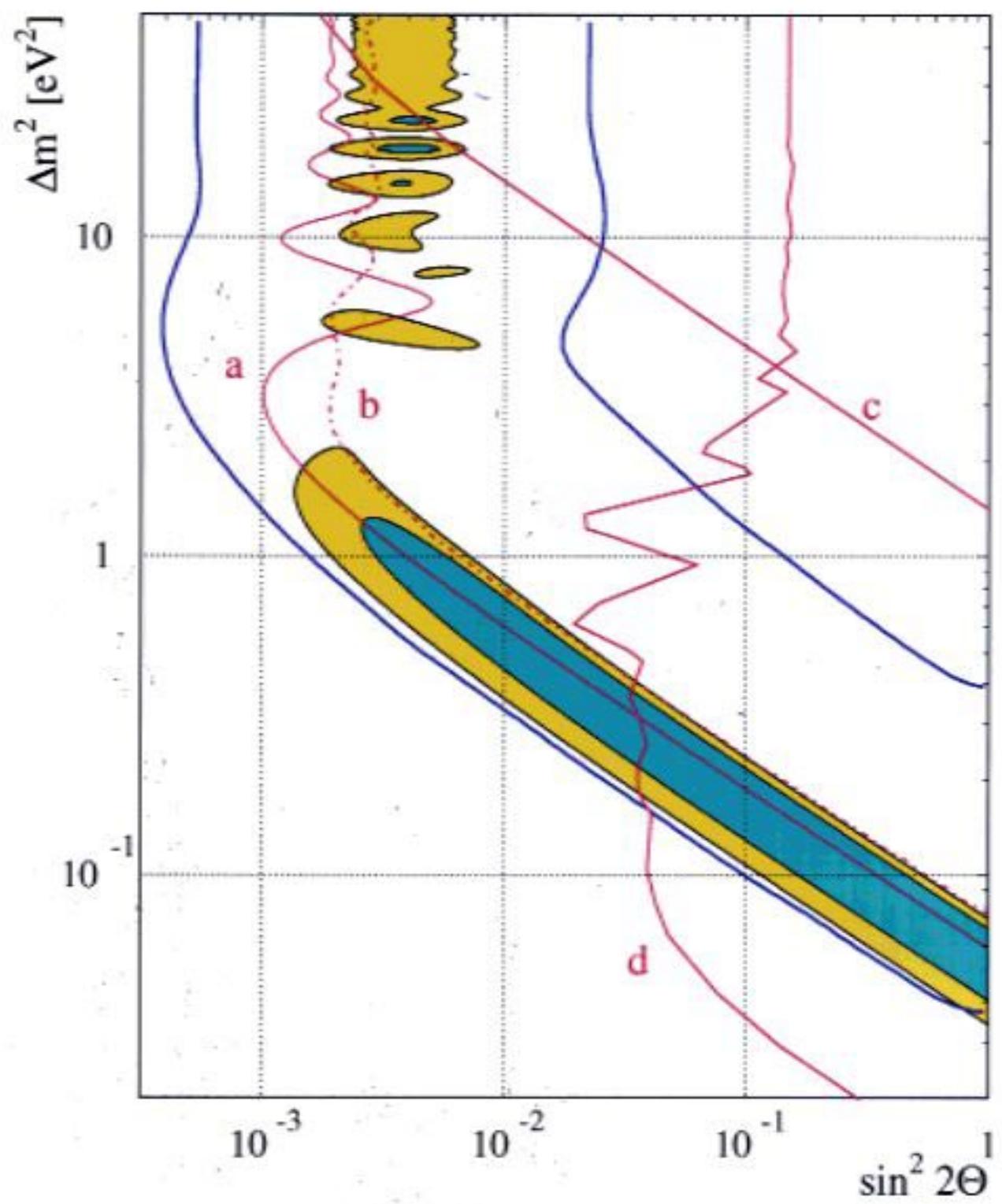
## $\nu_\mu \rightarrow \nu_e$ DIF Oscillation Search



- Different systematics than DAR
- Different backgrounds than DAR
  - $\mu \rightarrow e \nu_\mu \nu_e$  and  $\pi \rightarrow e \nu_e$
- Different coverage of  $\Delta m^2$  and  $\sin^2 2\theta$
- However, only a single signature

$18.1 \pm 6.6$  excess events





# The BooNE Collaboration

July 13, 1999

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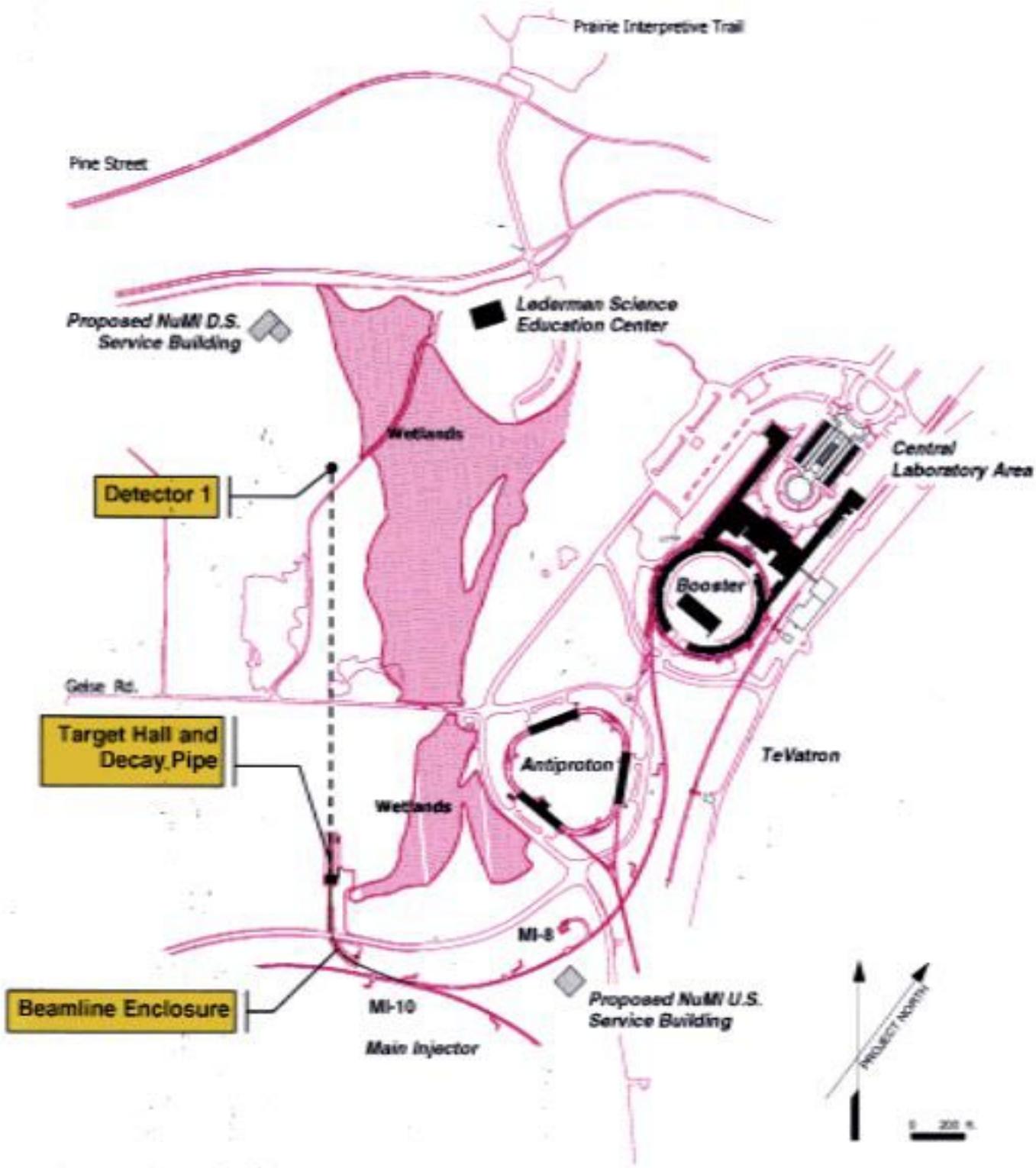
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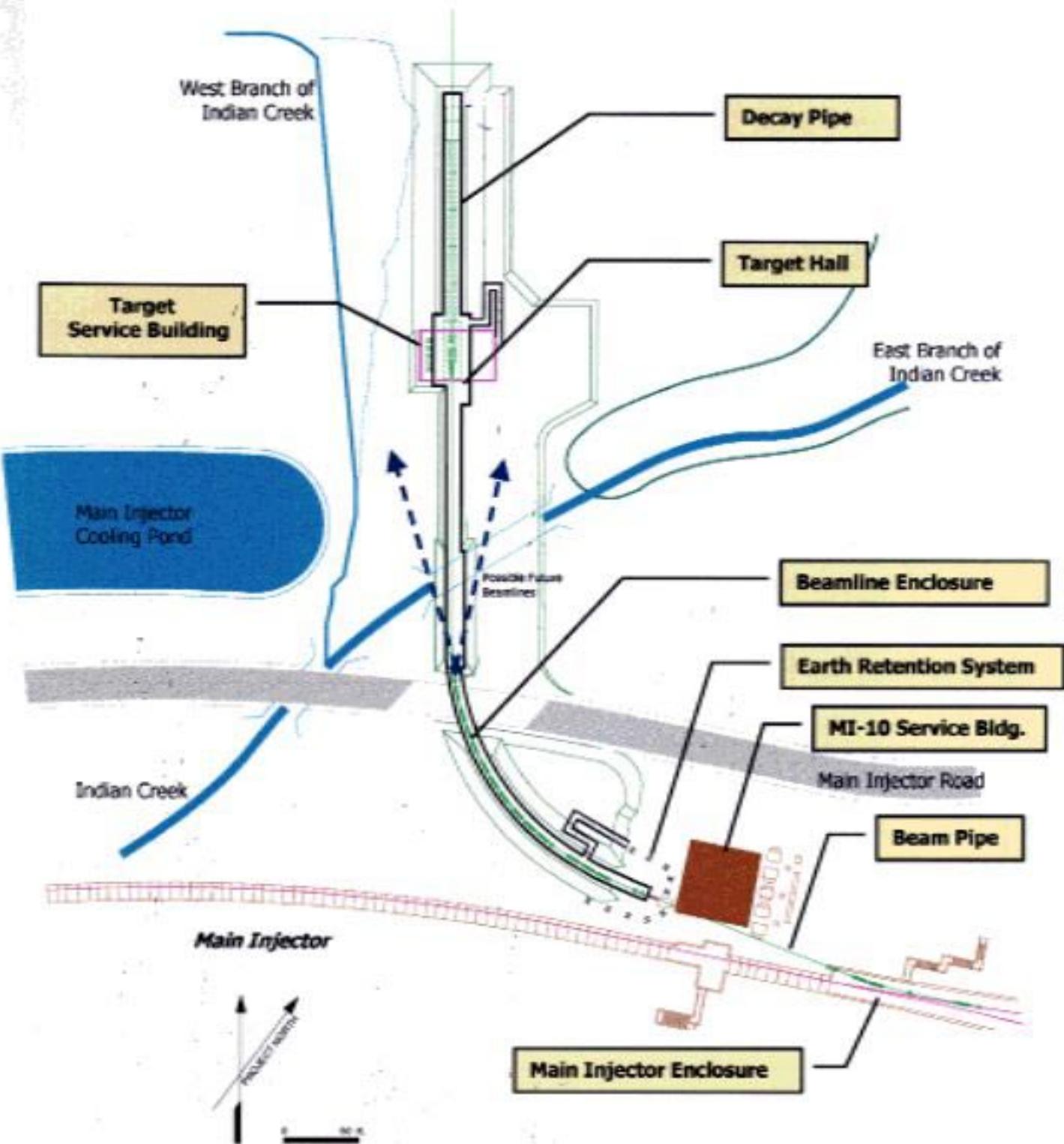
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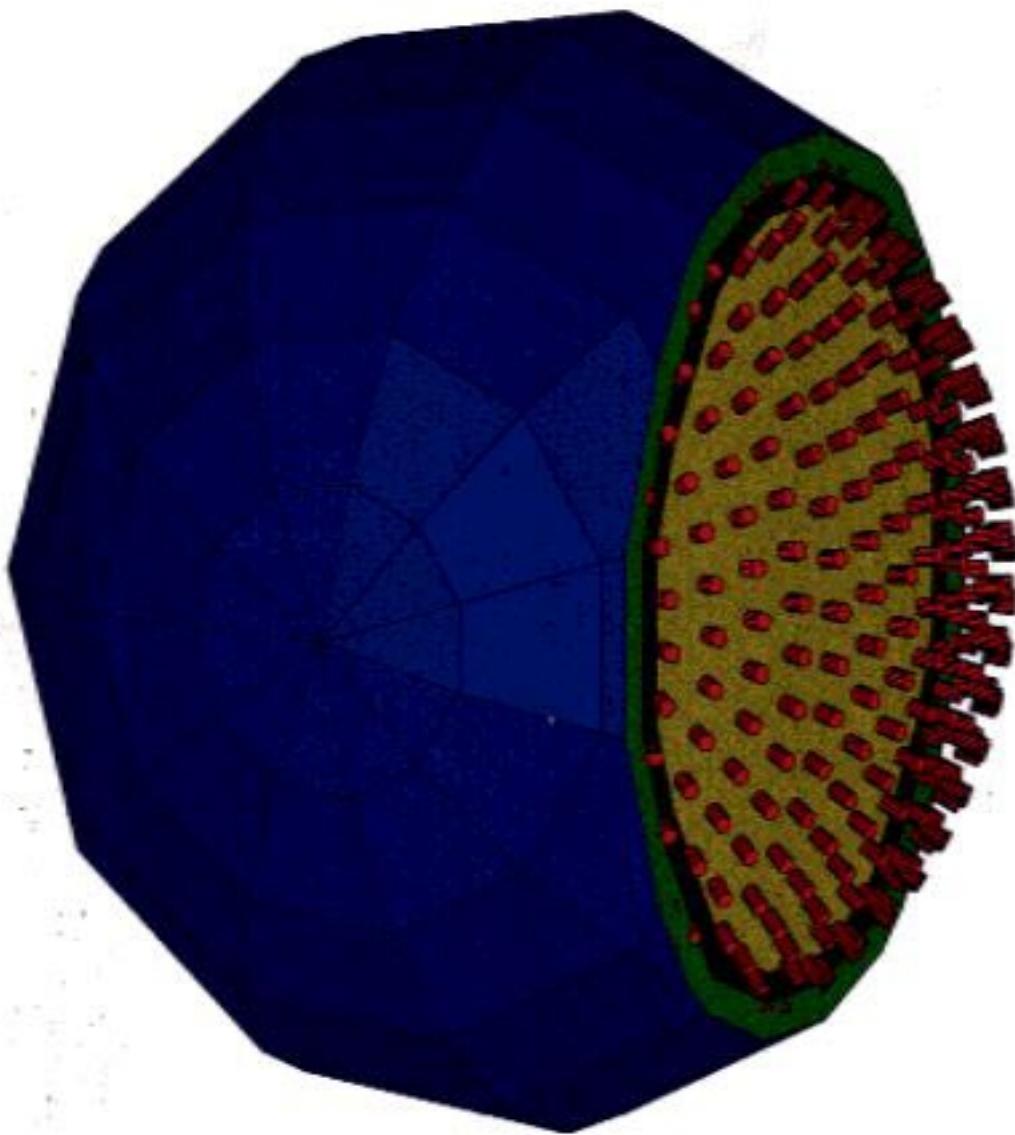
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# BooNE Detectors and Site Layout







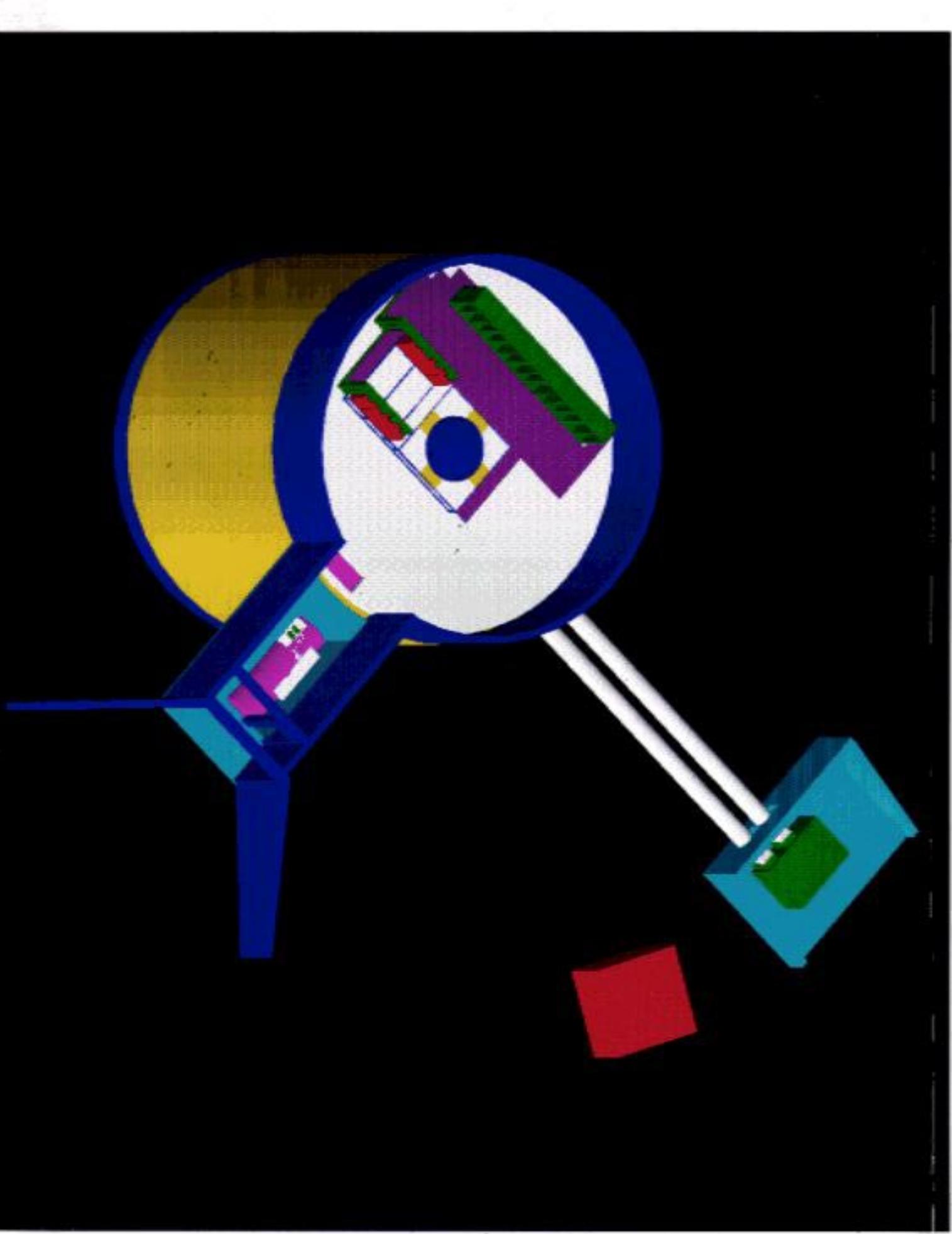


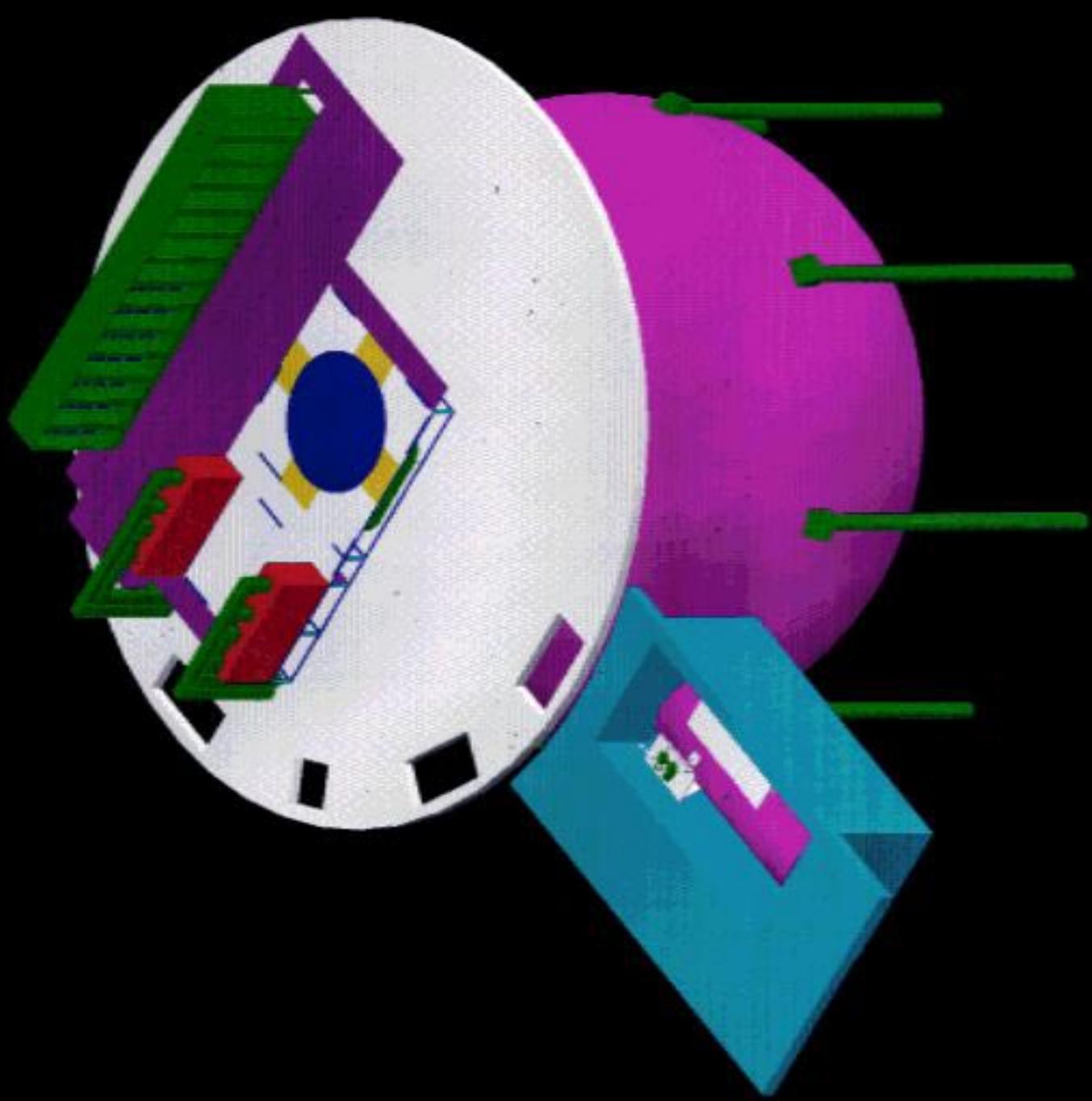
**40-feet diameter sphere**

**1520 8-inch PMTs (1280 detector + 240 veto PMTs)**

**807 tons of mineral oil**

**445 ton fiducial volume**





## **Oil vs Water**

- More Cerenkov light (x1.45)

oil ->	$n=1.47, \rho=0.85$
water ->	$n=1.33, \rho=1.00$

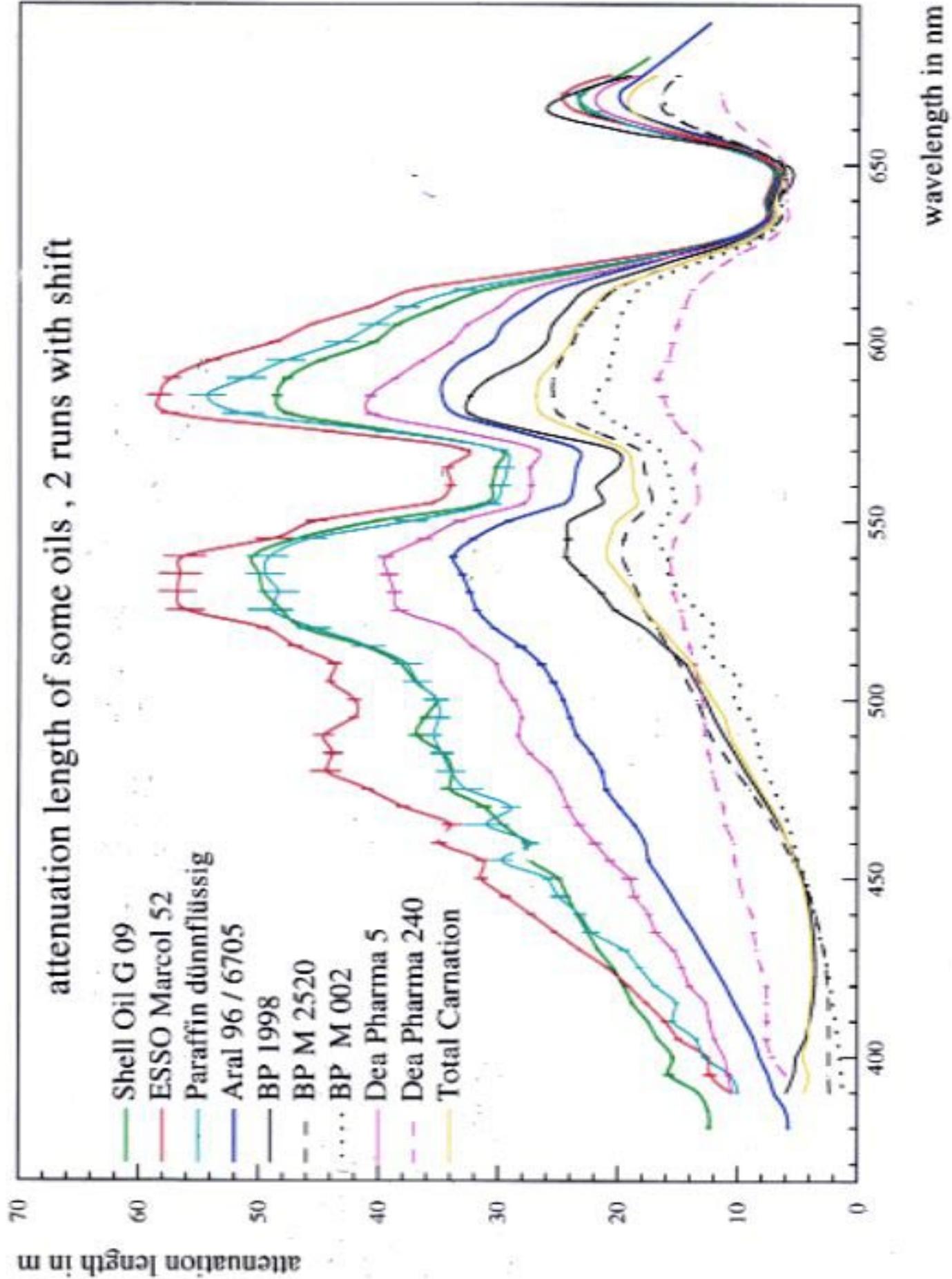
- No need for a purification system
- No worry about liquid seeping into bases
- Oil has less multiple scattering

oil ->	$X_0 = 44.8 \text{ g/cm}^2$
water ->	$X_0 = 36.1 \text{ g/cm}^2$

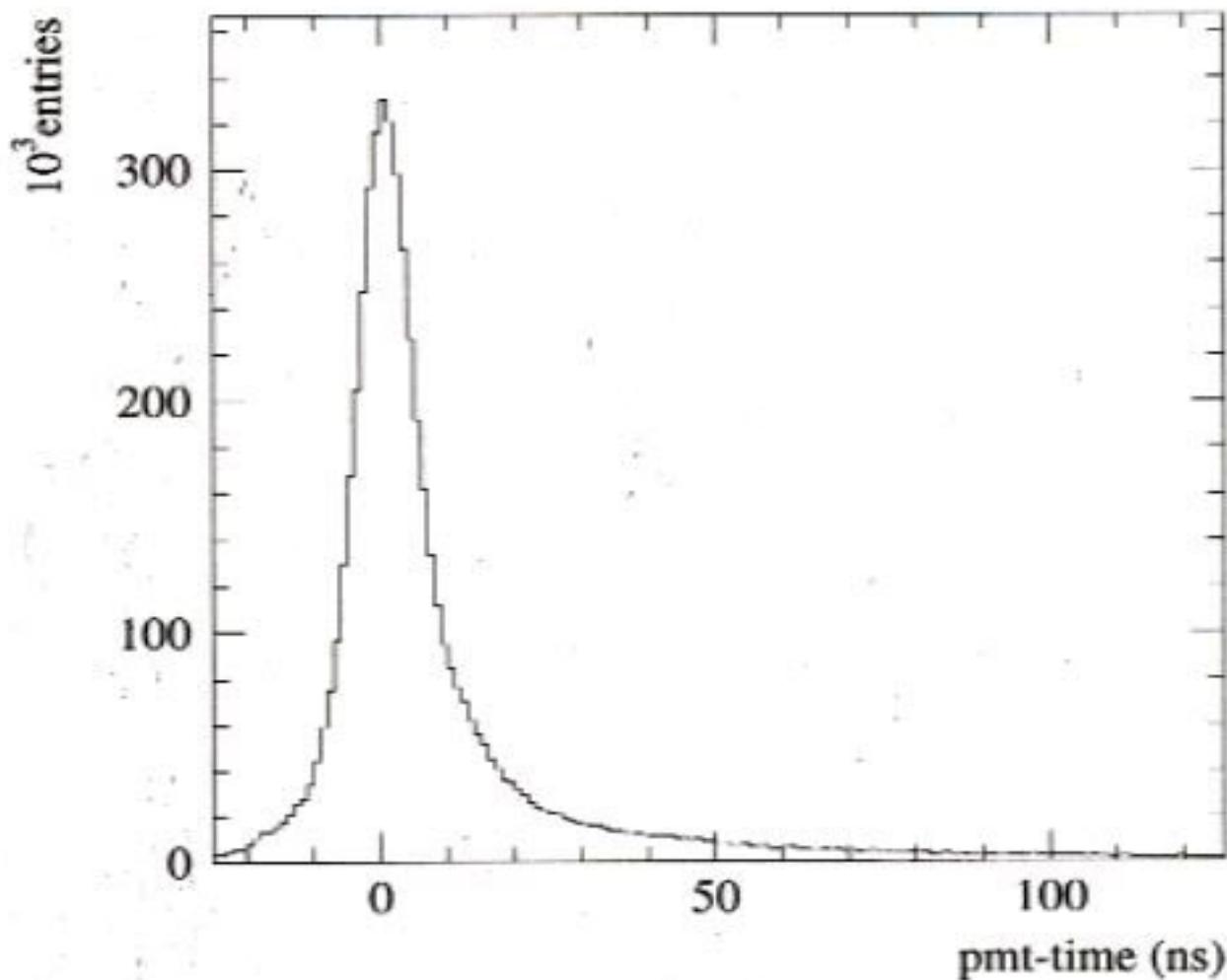
- Oil has lower  $\mu^-$  capture rate

oil ->	8%
water ->	18%

- Pure mineral oil produces a little scintillation light



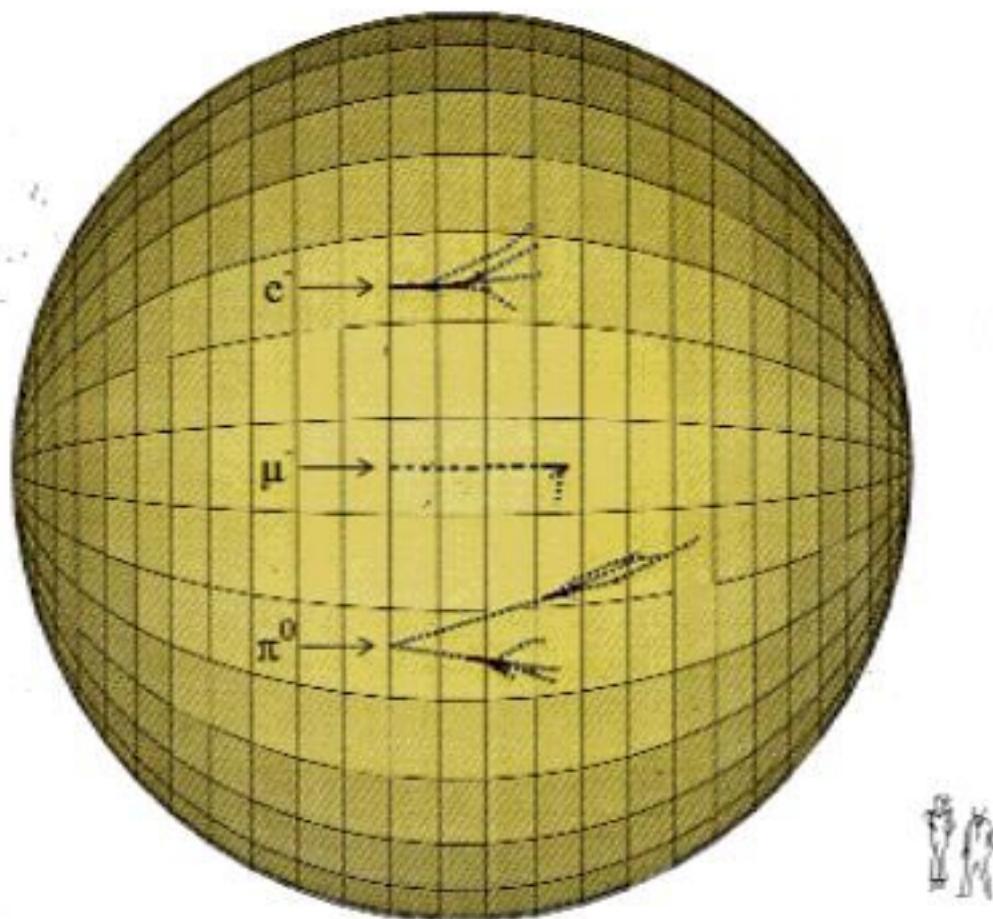
## The electron time distribution in pure oil



Run 664 LSND data and Texas A&M cyclotron tests indicate that for pure oil:

- (i) 75% Cerenkov light & 25% scintillation light
- (ii) Scintillation time constant of  $\sim 35$  ns

# Full GEANT Monte Carlo Simulation



$\nu_e C \rightarrow e^- N$       Signal

$\nu_\mu C \rightarrow \mu^- N$       Background

$\nu_\mu C \rightarrow \nu_\mu \pi^0 X$       Background

# Event Reconstruction & PID

## 1. electron event reconstruction

$\delta r \sim 21\text{ cm}$

$\delta t < 1\text{ ns}$

$\delta\theta \sim 3.6^0$

$\delta E/E \sim 10\%$

## 2. muon event reconstruction

mis-id background  $< 0.1\%$

$\delta E/E \sim 29\%$

## 3. $\pi^0$ event reconstruction

mis-id background  $\sim 1\%$

$\delta E/E \sim 10\%$

# The $\chi^2$ Particle ID Method

Signature:	Cerenkov Ring	Track Extent	Hit Timing
$\pi^0$	Two rings	Extented source because of 2 $\gamma$ 's	Recoil Proton produces Scint Light
e	Ring mult. scat. & brems	$\beta=1 \rightarrow$ Cerenkov angle: $\alpha=47^\circ$ Ring is "Fuzzy" due to mult. scat. & brems	$\chi_{\sigma} = \sum q_i (\alpha_i - 47^\circ)^2 / Q$ (consistent w/ $\beta=1$ )
$\mu$	Sharp outer edge of ring (less mult. scat. than for e's) Diffuse inner edge of ring ( $\beta < 1 \rightarrow$ smaller Cerenkov angle)	Long Track (extended source)	$\chi = \sum q_i (t_i - r/v - t_0)^2 / Q$ (consistent w/ "point source")
			Particle:  $\chi_t = \frac{\sum N_i (t_i > 10ns)}{\sum N_i (t_i < 500ns)}$ (consistent w/ "late light")

## Systematic Errors from Particle Mis-id

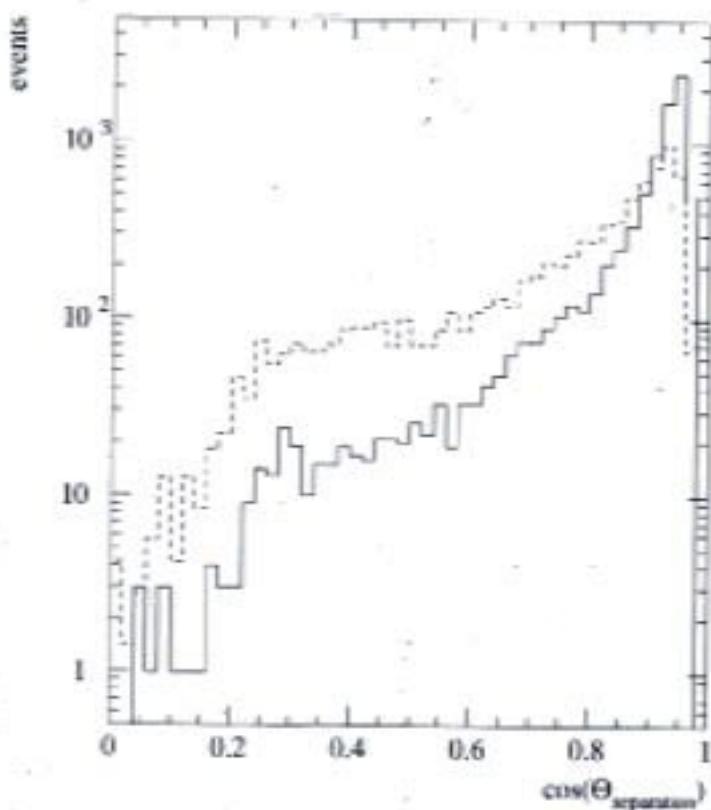
$\mu^-$ : <5% systematic error

Use decay  $\mu^-$  events to determine the mis-id rate:

- 92% are tagged by the decay  $e^-$
- $8 \pm 0.1\%$  are captured

$\pi^0$ : ~5%

Use symmetric  $\pi^0$  decays to determine the mis-id rate:

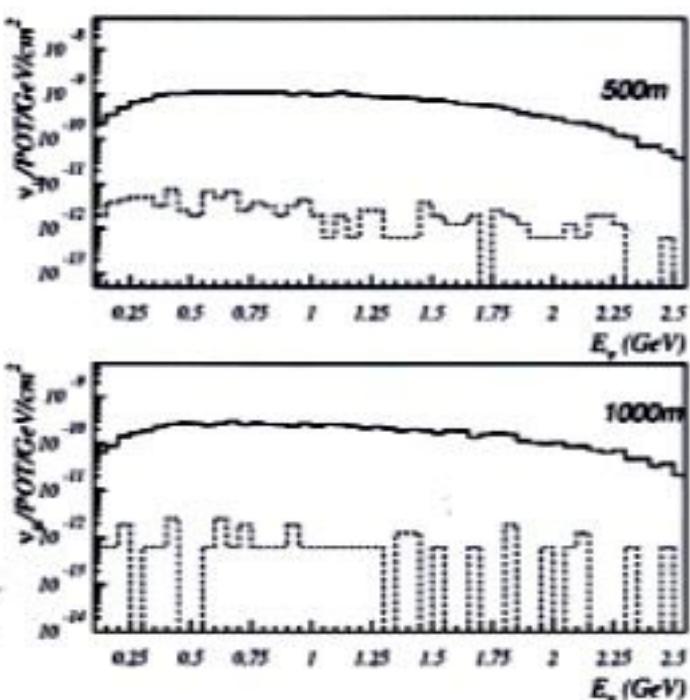


Angular distribution between 2 fitted rings

## The BooNE $\nu$ beam (*Full Geant Simulation*)

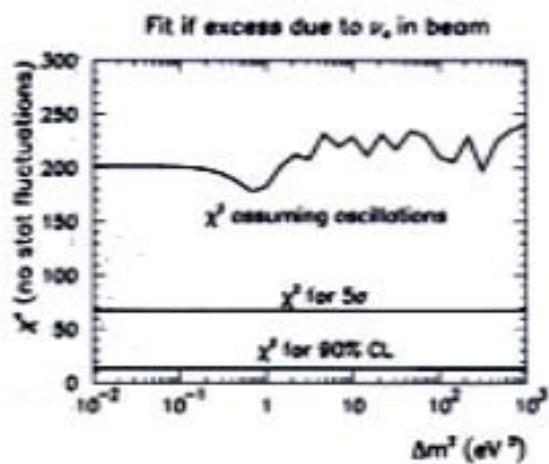
- Al target within 2-horn secondary focusing system

Solid –  $\nu_\mu$  flux,  
Dashed –  $\nu_e$  background

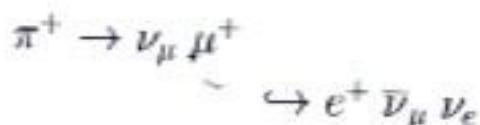


Sources of  $\nu_e$  background (0.3% of beam):

- $\mu$  decays (75% of total  $\nu_e$ , 5% systematic error)
- $K$  decays (25% of total  $\nu_e$ , 10% systematic error)
- The  $\nu_e$  beam background does not fit the energy shape for the oscillation hypothesis.

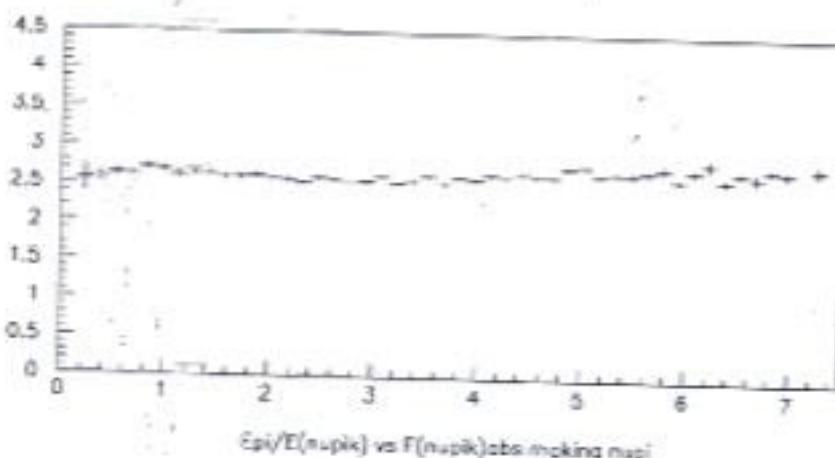


To know  $\nu_e$ 's from  $\mu$  decays, we need the  $\pi$  spectrum...



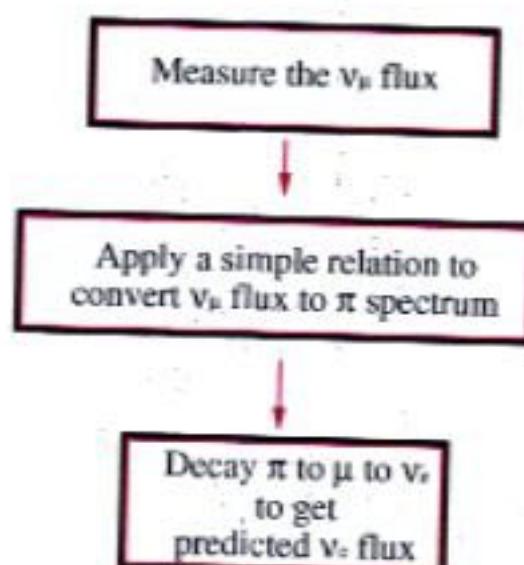
The  $\nu_\mu$  energy spectrum is *highly correlated* to the  $\pi$  spectrum!

(Because the detector subtends a very small solid angle)



$$E_\pi/E_\nu \approx 2.5 \text{ for all } E_\pi !$$

This means we can...



Investigating method using Monte Carlo:

We can constrain the  $\nu_e$  from  $\mu$  decay to 5%

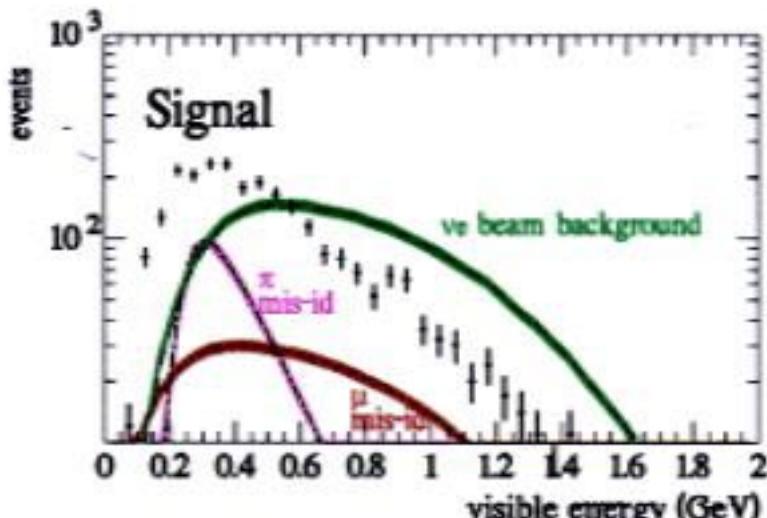
## Estimated Number of Events after 1 year ( $2 \times 10^7$ s)

Reaction	Number of Events
$\nu_\mu C \rightarrow \mu^- N$	590,000
$\nu_\mu e \rightarrow \nu_\mu e$	130
$\nu_\mu C \rightarrow \mu^- \pi^0 X$	65,000
$\nu_\mu p, n \rightarrow \nu_\mu p, n$	72,000
$\nu_e C \rightarrow e^- N$ (100% transmutation)	617,000
$\Delta m^2 = 0.4 \text{ eV}^2, \sin^2 2\theta = 0.02$	1200
Intrinsic $\nu_e$	1800
$\mu^-$ Misidentification	600
$\pi^0$ Misidentification	600

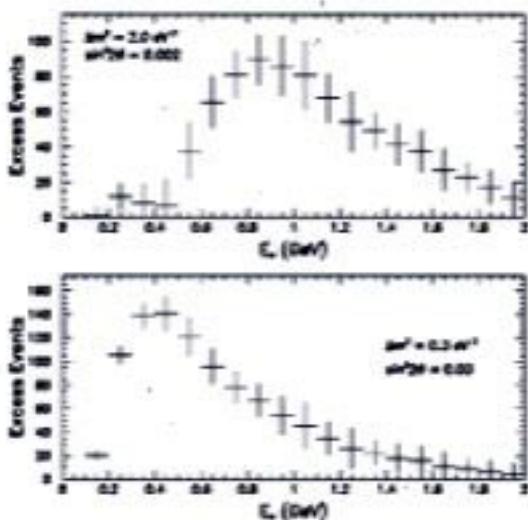
*The background is low at low energies...*

Points:  $\Delta m^2 = 0.4 \text{ eV}^2$ ,  
 $\sin^2 2\theta = 0.04$   
colored bands show  
systematic errors

note log scale →

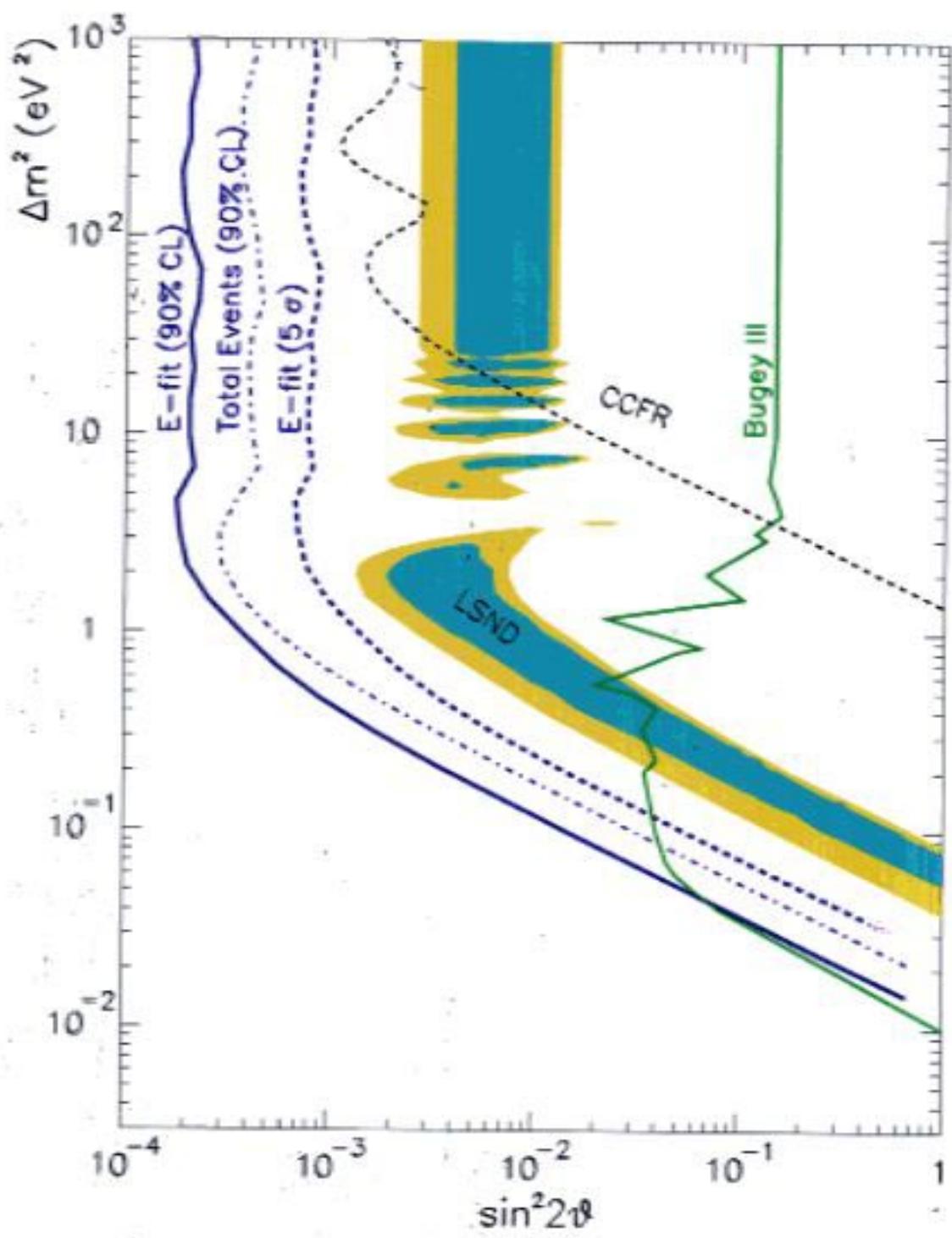


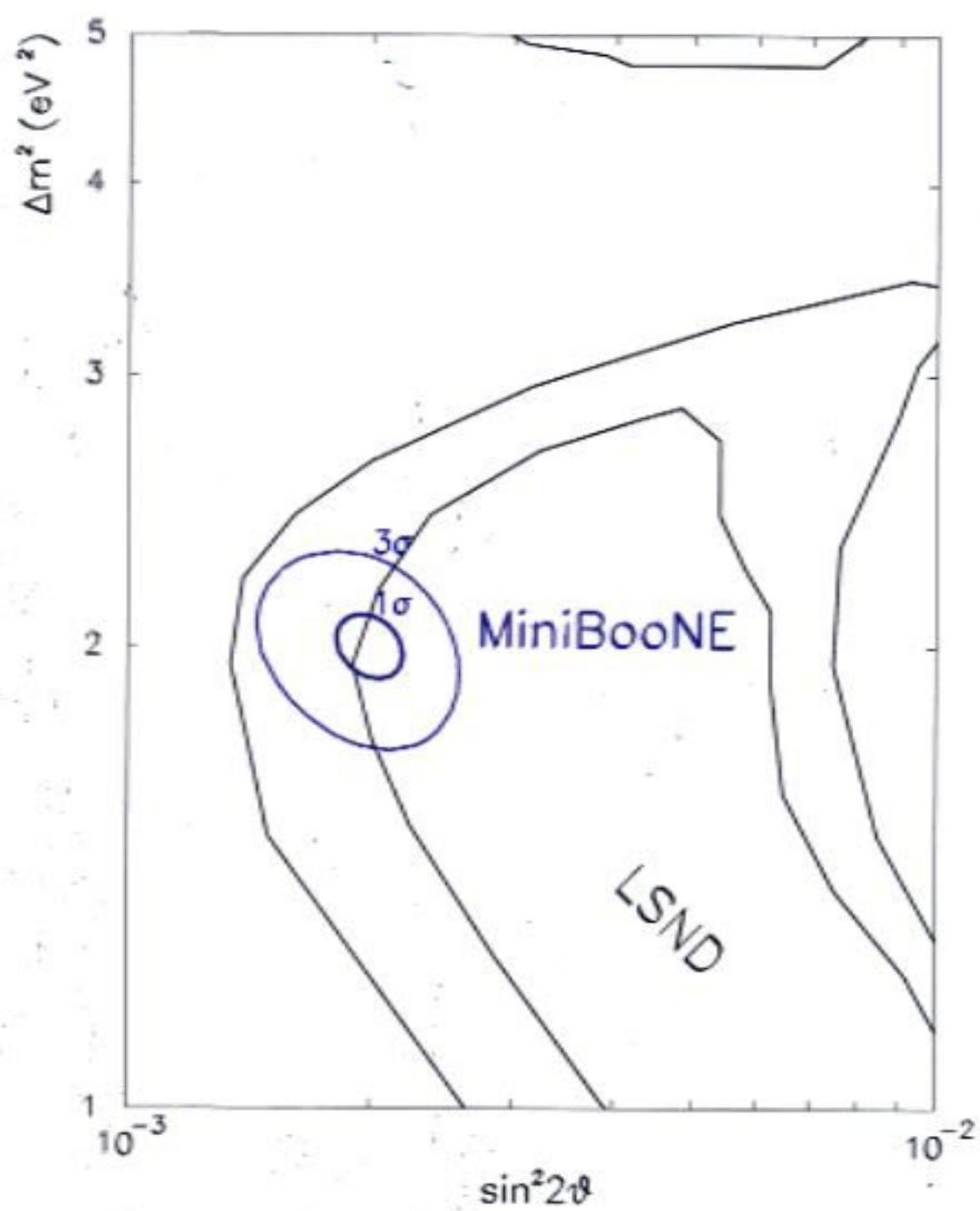
Signal after Bkgd Subtraction for Two Possible Osc. Parameters:



- Statistical uncertainty for signal is included in errors.
- Sys. and stat. uncertainty from background is included in errors.
- (Statistical fluctuations of data points not shown).

MiniBooNE can clearly establish a signal!  
The signal indicates where to place the 2nd detector



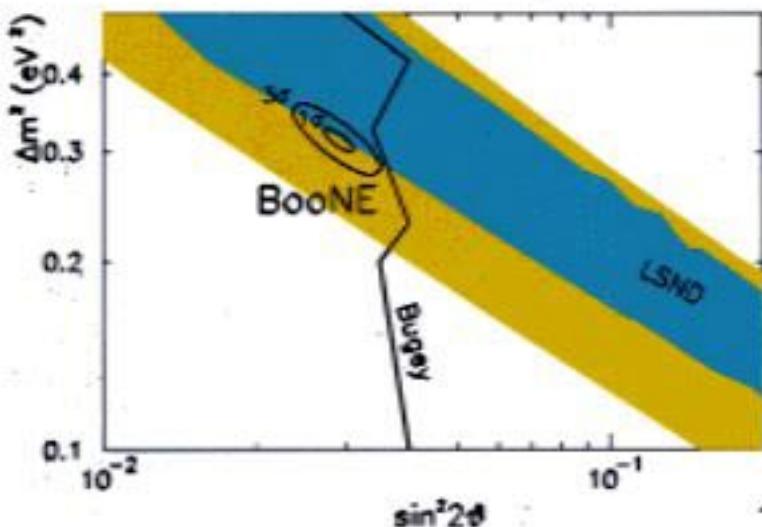


## A “Measurement” Experiment!

Two examples for MiniBooNE (1 detector) measurements:

$\Delta m_0^2$	$\sin^2 2\theta_0$	$\delta(\Delta m^2)$	$\delta(\sin^2 2\theta)$	Signal	Signif.
$0.3 \text{ (eV}^2)$	0.03	$0.10 \text{ (eV}^2)$	0.02		$44 \sigma$
$2.0 \text{ (eV}^2)$	0.002	$0.10 \text{ (eV}^2)$	0.0002		$15 \sigma$

Example BooNE (2 detector) measurement



And tests of CP violation with  $\nu$  and  $\bar{\nu}$  running

## Conclusions

- Evidence for  $\nu$  oscillations from LSND !
  - $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  search yields **39.5+-8.8 events**  
PRC 54, (1996) 2685 & PRL 77, (1996) 3082
  - $\nu_\mu \rightarrow \nu_e$  search yields **18.1+-6.6 events**  
PRC 58, (1998) 2489 & PRL 81, (1998) 1774
  - $m_\nu > 0.4$  eV
- The BooNE experiment will make a definitive test of the LSND signal and will make precision measurements of  $\Delta m^2$  and  $\sin^2 2\theta$  if  $\nu$  oscillations occur.
- BooNE construction begins **10/99**  
Gain beneficial occupancy on **1/01**  
Detector operational on **10/01**  
Beam complete & taking data on **12/01**