

# Microscopically Based Energy Functionals

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# Dream Scenario: From QCD to Nuclei



### SciDAC 2 Project Building a Universal Nuclear Energy Density Functional



### See <u>http://undef.org</u> for details

# **UNEDF** Project Goals

- Understand nuclear properties "for element formation, for properties of stars, and for present and future energy and defense applications."
- Scope is *all* nuclei
   => DFT the method of choice
- Order of magnitude improvement over present capabilities
   => precision calculations of, e.g., masses
- Utilize the best available microscopic physics
   => chiral EFT NN and NNN interactions, ab-initio MBT
- Maximize predictive power will *well-quantified theoretical uncertainties*

### Years 2 & 3: Personnel, Tasks, and Interconnections



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# Limitations of Existing Energy Functionals (Predictability)



- Uncontrolled extrapolations towards the drip-line
- Theoretical error-bars?

Limitations of Existing Energy Functionals (Predictability)



• Pairing gaps not under control for increasing (N-Z)

# What's missing in phenomenological EDF's

- Density dependencies might be too simplistic
- Isovector components not well constrained
- No systematic organization of terms in the EDF
- No way to estimate theoretical uncertainties
- Over-determined parameters
- What's the connection to many-body forces?
- Pairing part of the EDF not treated on same footing

## Turn to microscopic many body theory for guidance

## Nuclear forces from Chiral EFT

Separation of scales: low momenta Q <<  $\Lambda_{\rm b}$  breakdown scale



- Explains empirical hierarchy NN > 3N > 4N
- Formal Consistency NN and NNN forces ππ and πN, electroweak operators QCD, systematic expansion
- $\bullet$  Error estimates from truncation order, lower bound from  $\Lambda$  variation



Weinberg, van Kolck, Epelbaum, Meissner, Machleidt, ...

### $\Lambda$ / Resolution dependence of nuclear interactions

with high-energy probes: quarks+gluons

cf. scale/scheme dependence of parton distribution functions



### Lattice QCD

Effective theory for NN, many-N interactions, operators depend on resolution scale  $\Lambda$ 

 $H(\mathbf{\Lambda}) = T + V_{\rm NN}(\mathbf{\Lambda}) + V_{\rm 3N}(\mathbf{\Lambda}) + V_{\rm 4N}(\mathbf{\Lambda}) + \dots$ 

 $\Lambda_{chiral}$ momenta Q ~  $\lambda^{-1}$  ~ m<sub> $\pi$ </sub>: chiral effective field theory nucleons interacting via pion exchanges + contact interactions typical Fermi momenta in nuclei ~ m<sub> $\pi$ </sub>



 $\Lambda_{\text{pionless}}$ Q << m<sub>\pi</sub>=140 MeV: pion not resolved pionless effective field theory

 large scattering lengths + corrections
 n

 applicable to loosely-bound, dilute systems, reactions at astro energies

Freedom to vary the resolution via RG to simplify certain features... 11



### "Scheme-Dependent" Sources of Non-perturbative Physics

strong tensor force

high-momentum modes

**BUT** typical momentum in a large nucleus only  $\approx 1$  fm<sup>-1</sup> (200 MeV)!

2 Types of Renormalization Group Transformations



• SRG => drives H towards the diagonal ( $\lambda$  = width about diagonal)



Both decouple the high momentum modes *leaving low E NN observables unchanged*.

Integrating out high-momentum modes (" $V_{low k}$ ")



• Demand 
$$\frac{d}{d\Lambda}T = 0$$

=> RGE's for "running" of  $V_{\Lambda}$  w/  $\Lambda$ 

- Integrate RGE's to smaller Λ
   => decouples high k modes
- Low momentum universality

   => evolved interactions ("V<sub>low k</sub>") coalesce
   to ≈ universal curve

UV cutoff  $\Lambda$ 



### The Similarity Renormalization Group [Wegner, Glazek and Wilson]

• Unitary transformation on an initial H = T + V

 $H_s = U(s)HU^{\dagger}(s) \equiv T + V_s$  s = continuous flow parameter

• Differentiating with respect to s:

$$rac{dH_s}{ds} = [\eta(s), H_s] \qquad ext{with} \qquad \eta(s) \equiv rac{dU(s)}{ds} U^\dagger(s) = -\eta^\dagger(s)$$

• Engineer  $\eta(s)$  to do different things as s ->  $\infty$ 

$$\eta(s) = [\mathcal{G}_s, H_s]$$

 $\mathcal{G}_s = T \Rightarrow H_s \text{ driven towards the diagonal in k - space}$  $\mathcal{G}_s = PH_sP + QH_sQ \Rightarrow H_s \text{ driven towards block diagonal form}$ 

### Run to Lower $\lambda$ via SRG $\implies \approx$ Universality



Note:  $\lambda = s^{-1/4}$ 

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Note:  $\lambda = s^{-1/4}$ 

### Observations on 3N forces



Arise whenever eliminate DOF (relativity, nucleon excitations, high momentum intermediate states)



Omitting 3NF's => observables depend on  $\Lambda$ .

### Why Bother lowering $\Lambda$ if 3NF's grow ?



Ratio <3N>/<2N> not unnaturally large

Chiral:  $\langle V_{3N} \rangle \sim (Q/\Lambda)^3 \langle V_{NN} \rangle$ 

Hard work for a small contribution (large  $\Lambda$ ) versus less work for a larger contribution (small  $\Lambda$ )?



- Approximate treatments of 3NF "work" better
  - e.g., normal ordering (D. Dean's talk)
  - perturbative, HF dominates



### Approximate RG Evolution of 3NF



-Chiral EFT is a complete operator basis -Approximate RG running by refitting D and E at each  $\Lambda$  -Equivalent to truncating RGE to leading operators

#### Consistency checks of the approximation:

weak  $\Lambda$ -dependence in NM (renormalization is working) <3N>/<2N> agrees w/power counting estimates



### New low-momentum NNN fits and Nuclear Matter



NOTE: 3NF drives saturation NOT the tensor force

### New low-momentum NNN fits and Nuclear Matter



Knobs to estimate Theoretical error bars:

A-dependence => theoretical error bands (lower limit)

Assess the impact of large uncertainties in the  $c_i$ 's appearing in 2- and 3-body TPEP (to do)

Vary the order of the underlying EFT (to do)

Sensitivity to manybody approximations

### New low-momentum NNN fits and Nuclear Matter





Supports suggestion of Navratil et al. to use <sup>4</sup>He radii to constrain fits of 3NF couplings ( $c_E$  and  $c_D$ )

NM to constrain  $c_3$  and  $c_4$ ?

Local Functionals from Many-Body Theory

• Dominant MBPT contributions to bulk properties take the form

$$\langle V \rangle \sim \text{Tr}_1 \text{Tr}_2 \int d\mathbf{R} \, d\mathbf{r}_{12} \, d\mathbf{r}_{34} \, \rho(\mathbf{r_1}, \mathbf{r_3}) \, K(\mathbf{r}_{12}, \mathbf{r}_{34}) \, \rho(\mathbf{r_2}, \mathbf{r_4}) \; + \; \text{NNN} \, \cdots$$



K is either free-space interaction (HF) or resummed in-medium vertex (BHF)

- Written in terms on non-local quantities
  - density matrices and s.p. propagators
  - finite range and non-local resummed vertices K

Connection to  $E = E[\rho]$  is not obvious!

Density Matrix Expansion Revisited (Negele and Vautherin)

• Expand of DM in local operators w/factorized non-locality

$$\langle \Phi | \psi^{\dagger} \left( \mathbf{R} - \frac{1}{2} \mathbf{r} \right) \psi \left( \mathbf{R} + \frac{1}{2} \mathbf{r} | \Phi \right) = \sum_{n} \Pi_{n}(k_{F} r) \langle \mathcal{O}_{n}(\mathbf{R}) \rangle$$
$$\langle \mathcal{O}_{n}(\mathbf{R}) \rangle = \left[ \rho(\mathbf{R}), \nabla^{2} \rho(\mathbf{R}), \tau(\mathbf{R}), \mathbf{J}(\mathbf{R}), \ldots \right]$$

Fall off in r controlled by local k<sub>F</sub>
 => expand and resum so LO term exact in uniform limit
 => NOT a simple short-distance expansion in r

$$\rho \left( \mathbf{R} + \frac{1}{2}\mathbf{r}, \mathbf{R} - \frac{1}{2}\mathbf{r} \right) = \frac{3j_1(k_F r)}{k_F r} \rho(R) + \frac{35j_3(k_F r)}{2k_F^3 r} \left( \frac{1}{4} \nabla^2 \rho(R) - \tau(R) + \frac{3}{5} k_F^2 \rho(R) \right) + \cdots$$

• Dependence on *local* densities now manifest

#### Skyrme-like EDF's from the DME

$$\mathcal{E} = \frac{\tau}{2M} + \frac{3}{8} t_0 \rho^2 + \frac{1}{16} t_3 \rho^{2+\alpha} + \frac{1}{16} (3t_1 + 5t_2) \rho \tau + \frac{1}{64} (9t_1 - 5t_2) |\nabla \rho|^2 + \cdots \quad \text{Skyrme}$$

$$\mathcal{E} = \frac{\tau}{2M} + A[\rho] + B[\rho]\tau + C[\rho]|\nabla \rho|^2 + \cdots$$
 DME

- coupling *constants* --> coupling *functions* 
  - -finite range effects encoded as  $\rho$ -dependence in ABC
  - microscopic isovector, spin-orbit terms
  - well-suited for existing SkyHF codes



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$$\mathcal{E} = \frac{\tau}{2M} + A[\rho] + B[\rho]\tau + C[\rho]|\nabla \rho|^2 + \cdots$$
 DME

- coupling constants --> coupling functions
  - -finite range effects encoded as p-dependence in ABC
  - microscopic isovector, spin-orbit terms
  - well-suited for existing SkyHF codes



- Don't touch the HFB solver
- Trivial to upgrade as each new coupling becomes available
- Implemented in HFBRAD

### Including Long Range Chiral EFT in Skyrme-like EDFs

Derived the most general (N $\neq$ Z, spin-unsaturated) EDF from chiral EFT thru N<sup>2</sup>LO at HF level [SKB and B. Gebremariam]

$$\begin{split} \mathcal{E}^{\rho\rho} &\equiv \sum_{q} \int d\mathbf{r} \left[ A^{\rho\rho} \rho_{q} \rho_{q} + A^{\rho\Delta\rho} \rho_{q} \Delta\rho_{q} + A^{\nabla\rho\cdot\nabla\rho} \nabla\rho_{q} \cdot \nabla\rho_{q} + A^{\rho\tau} \left( \rho_{q}\tau_{q} - \mathbf{j}_{q} \cdot \mathbf{j}_{q} \right) \right. \\ &+ A^{ss} \mathbf{s}_{q} \cdot \mathbf{s}_{q} + A^{s\Delta s} \mathbf{s}_{q} \cdot \Delta \mathbf{s}_{q} + A^{\nabla s \circ \nabla s} \nabla \mathbf{s}_{q} \circ \nabla \mathbf{s}_{q} \\ &+ A^{\rho\nabla J} \left( \rho_{q} \nabla \cdot \mathbf{J}_{q} + \mathbf{j}_{q} \cdot \nabla \times \mathbf{s}_{q} \right) + A^{\nabla \cdot s \nabla \cdot s} \left( \nabla \cdot \mathbf{s}_{q} \right) \left( \nabla \cdot \mathbf{s}_{q} \right) \\ &+ A^{JJ} \left( \sum_{\mu\nu} J_{q,\mu\nu} J_{q,\mu\nu} - \mathbf{s}_{q} \cdot \mathbf{T}_{q} \right) + A^{JJ} \Big[ \left( \sum_{\mu} J_{q,\mu\mu} \right) \left( \sum_{\mu} J_{q,\mu\mu} \right) + \sum_{\mu\nu} J_{q,\mu\nu} J_{q,\nu\mu} - 2 \mathbf{s}_{q} \cdot \mathbf{F}_{q} \Big] \Big] \end{split}$$

### Each coupling function splits into 2 terms

- 1) A-dependent Skyrme-like coupling constants
- Λ-independent coupling functions from pion physics with non-trivial density dependence

$$A^{\rho \Delta \rho} \Rightarrow A^{\rho \Delta \rho} (\Lambda) + A^{\rho \Delta \rho} [\rho] \quad \text{Etc...}$$
From contact terms in
EFT/RG V's





Longest range V <==> Strongest density dependence in EDF

Novel density-dependencies in EDF from  $1\pi$  and  $2\pi$  exchanges:

$$\rho^{7/3}, \ \rho^{4/3}, \ \rho^{2/3}, \ \frac{1}{\rho^{2/3}} log(1+c\rho^{2/3}), \ \dots$$
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# Effects of NNN on Couplings

#### Gradient term $(\nabla \rho)^2$ 140 $C(\rho)$ - NN only 120 - NNN only Hartree-Fock · — · 100 NN + NNNC(p) [MeV-fm<sup>5</sup>] 80 60 40 20 0 -20П -400.2 0.05 0.1 0.15 0 $\rho [\text{fm}^{-3}]$

SKB, Furnstahl and Platter, in prep.

- Only scalar-isoscalar terms worked out so far
- Consistent with Kaiser et al. results with explicit  $\Delta$ 's

### In Progress:



Should find interesting density dependencies compared to NN spin-orbit, which is a short-range effect!

### Including Long Range Chiral EFT in Skyrme-like EDFs



### Comparison to ab-initio calculations

Start from *the same* Hamiltonian and compare ab initio solution to the Microscopic DFT calculation based on the DME functional

CC or FCI calculations of nuclei and nuclei in external fields

How important is non-locality and how accurate is the DME?

Are systematics reproduced by DME as we vary parameters (e.g., 3NF couplings, RG cutoff  $\Lambda$ , order of input EFT, ...) in H?

Is the many-body treatment of nuclear matter sufficient?

Early indications are that non-trivial extensions of the DME are needed

#### DME for Low-Momentum Interactions



...But the "success" of this test of the DME is misleading...

### Comparison to ab-initio calculations

CC and DFT calculations of <sup>16</sup>O (w/3N contact of varying strength)



Quantitative and qualitative disagreement btw. coupled-cluster and DFT calculation. What is going on?

### Possible Reasons for the Poor Agreement

#### 1) DME averages out too much information

- COM P-dependence (spatial non-locality)
- energy-dependence



### Possible Reasons for the Poor Agreement

2) Gradient expansion breaks down when saturation not good



e.g., N3LO NM looks reasonable at lower densities despite poor saturation

Ab-initio results for O16 and Ca40 pretty decent, but DME is poor

Gradients no longer "small" since DME = expansion about NM?

O16



Ca48

	Coupled Cluster	DME	Coupled Cluster	DME	Coupled Cluster	DME
E/A	-6.72	-7.89	-7.72	-9.66	-7.40	-10.1
r <sub>ch</sub>	2.73	2.47	3.35	2.95	3.24	2.84 3

### Possible Reasons for the Poor Agreement

3) Errors in the Hartree contribution => feedback via self-consistency!



Treat Hartree exactly a-la Coulomb? [Negele and Vautherin, Sprung et al.]

- "Ab-initio DFT" should be taken with a grain of salt!
- However, microscopic MBT still useful to build in missing physics (density dependencies) to Skyrme

### Work in the near-term

- spin-orbit couplings from N<sup>2</sup>LO 3NF
- Extension of DME beyond even-even nuclei (time-odd couplings)
- Refits of Skyrme + Long-range coupling functions (ORNL group)
- Extension of DME to pairing channel (B. Gebremariam)
- Generalization of DME to handle non-localities in time (I.e., energy-dependence from beyond HF)
- Refinements of original DME (B. Gebremariam)

### Conclusions

- Lowering Λ via RG greatly simplifies nuclear few and many-body problems
  - Comparison of DFT to ab initio (same H) now possible
  - use  $\Lambda\text{-dependence}$  as a tool for estimating errors
  - $V_{3N}$  can be treated as perturbation/simple approximations
  - perturbative nuclear matter?
  - Correlations "blurred out" => HF is decent starting point
  - Extension of Skyrme EDF's via DME (novel density dep.)
  - Theoretical guidance for future fits possible using "error bars" generated from Λ-dependence

# Collaborators

- MSU/NSCL: B. Gebremariam
- Ohio State: R. Furnstahl, L. Platter
- Iowa State: J. Vary, P. Maris
- ORNL: G. Hagen, T. Papenbrock
- TRIUMF: A. Schwenk
- Orsay/France: T. Duguet, V. Rotival

# **Observables Sensitive to 3N Interactions?**

- Study systematics along isotopic chains
- Example: kink in radius shift  $\langle r^2 \rangle (A) \langle r^2 \rangle (208)$



- Associated phenomenologically with behavior of spin-orbit
  - isoscalar to isovector ratio fixed in original Skyrme
- Clues from chiral EFT contributions? (Kaiser et al.)

# **Ratio of Isoscalar to Isovector Spin-Orbit**

- Ratio fixed at 3:1 for short-range spin-orbit (usual Skyrme)
- Kaiser: DME spin-orbit from chiral two-body (left) and three-body (right)



Systematic investigation needed

# **Observables Sensitive to 3N Interactions?**

Recent studies of tensor contributions [e.g., nucl-th/0701047]



See also Brown et al., PRC 74 (2006)

### Naturalness to Constrain Skyrme Couplings

Old NDA analysis:

$$c \left[\frac{\psi^{\dagger}\psi}{f_{\pi}^{2}\Lambda}\right]^{\prime} \left[\frac{\nabla}{\Lambda}\right]^{n} f_{\pi}^{2}\Lambda^{2}$$

$$\stackrel{\rho \longleftrightarrow \psi^{\dagger}\psi}{\tau \longleftrightarrow \nabla\psi^{\dagger} \cdot \nabla\psi}$$

$$\stackrel{\rho \longleftrightarrow \psi^{\dagger}\psi}{\tau \longleftrightarrow \psi^{\dagger}\nabla\psi}$$

Density expansion?

$$\frac{1}{7} \leq \frac{\rho_0}{f_\pi^2 \Lambda} \leq \frac{1}{4}$$

for  $1000 \geq \Lambda \geq 500$ 



Furnstahl and Hackworth 1997

