Lecture #1

1

A Pedestrian's Guide to Heavy Ion Physics



sPHENIX Experiment at RHIC Data recorded: 2023-05-22, 02:07:00 EST Run / Event: 7156 / 12 Collisions: Au + Au @ 200 GeV



UC RIVERSIDE 2023 National Nuclear Physics Summer School

Jamie Nagle University of Colorado (



Lecture Philosophy

- Less is more, i.e., not meant to be comprehensive
- Keep It Simple Stupid (KISS) principle...
- Even the experts often miss the big questions...
- Take away goals...
 - an appreciation for the science
 - excitement of the field
 - some details on experimental methods
 - open questions and opportunities for discovery by young people such as yourselves

Let me know if there are specific things you want to hear

Simplest Goals

Emergent Phenomena... Collectivity... What is the underlying origin? For this flock of birds, it is all short-range interactions (amazing!)

Emergent Phenomena

Connection from the QCD Lagrangian to phenomena of confinement and asymptotic freedom was fundamental



Connection from QCD to the emergent phenomena of near perfect fluidity of the Quark-Gluon Plasma is just as fundamental

Perfect fluidity tells us the nature of the QGP, more importantly we need to reconcile:

Most important discovery in field: <u>perfect fluid</u> & Crucial part of QCD: <u>weak coupling at short distances</u>

Global temperature change (1850-2020)



Jet Quenching Probes of QGP Fluid





https://journals.aps.org/prc/abstract/10.1103/PhysRevC.96.024901

Easy Tools

https://www.lcdf.org/gifsicle/



Gifsicle

Source and documentation gifsicle-1.88.tar.gz (564670 bytes) On Github Wild animations are distracting, but <u>key</u> <u>visualizations stick</u> <u>with people</u>. *Make your own!*

Create your own ideas and borrow only selectively. The more important the talk, the more important it is to show your thoughts...

Data thief programs allow you to choose which data sets to show, re-fit your model, etc.

Side note to students...

1) Data visualization is very important

2) Leads to breakthroughs via more intuitive and physical picture

3) Most talks only get across 2-3 items that a given audience member will recall later

What take aways did you get from last week's lectures for example?

A Brief History of Time (in Heavy Ions)

A long time ago in a galaxy far, far away....

History can be boring, though often early in a new field is when the most basic questions are discussed openly. Later people often focus too much on the latest details.

Start in 1950s → Hadron Zoo

Mesons





- More and more hadrons being discovered.
 - Too many to be fundamental particles?
 - Hadrons are strings (interesting history).

STATISTICAL THERMODYNAMICS OF STRONG INTERACTIONS AT HIGH ENERGIES

R. Hagedorn

CERN - Geneva

ABSTRACT

In this statistical-thermodynamical approach to strong interactions at high energies it is assumed that higher and higher resonances of strongly interacting particles occur and take part in the thermodynamics as if they were particles. For $m \rightarrow \infty$ these objects are themselves very similar to those which shall be described by this thermodynamics. Expressed in a slogan: "We describe by thermodynamics fire-balls which consist of fire-balls, which consist of fire-balls, which ...". This principle, which could be called "asymptotic bootstrap", leads to a self-consistency requirement for the asymptotic form of the mass spectrum. The equation following from this requirement has only a solution if the mass spectrum grows exponentially:

$$\rho(m) \xrightarrow{m \to \infty} const.m^{-5/2} exp(\frac{m}{T_0}).$$

 $T_{\rm O}$ is a remarkable quantity: the partition function corresponding to the above $\rho\left(m\right)$ diverges for $T \rightarrow T_{\rm O}^{-}$. $T_{\rm O}$ is therefore the highest possible temperature for strong interactions. It should - via a Maxwell-Boltzmann law - govern the transversal momentum distribution in all high energy collisions of hadrons (including e.m. form factors, etc.). There is experimental evidence for that, and then $T_{\rm O}$ is about 158 MeV ($\approx 10^{12}$ oK). With this value of $T_{\rm O}$ the asymptotic mass spectrum of our theory has a good chance to be the correct extrapolation of the experimentally known spectrum.

Another consequence is the prediction that the elastic amplitude A(s,t) should decrease as $\sim \exp(-p_{\perp}/2T_{o})$ for any non-zero fixed scattering angle and $s \rightarrow \infty$.

For astrophysics the present theory puts some doubt on the neutron-star model for the interior of collapsing stars; at the same time it suggests a straightforward improvement. Circa 1965 Hagedorn observed that these hadrons had an exponentially increasing density of states $\rho(n)$

$$\rho(m) \xrightarrow{m \to \infty} \text{const.m}^{-5/2} \exp(\frac{m}{T_0}).$$



Modern era compilation from PDG by Zajc

http://cds.cern.ch/record/346206

R. Hagedorn CERN - Geneva

ABSTRACT

In this statistical-thermodynamical approach to strong interactions at high energies it is assumed that higher and higher resonances of strongly interacting particles occur and take part in the thermodynamics as if they were particles. For $m \rightarrow \infty$ these objects are themselves very similar to those which shall be described by this thermodynamics. Expressed in a slogan: "We describe by thermodynamics fire-balls which consist of fire-balls, which consist of fire-balls, which ...". This principle, which could be called "asymptotic bootstrap", leads to a self-consistency requirement for the asymptotic form of the mass spectrum. The equation following from this requirement has only a solution if the mass spectrum grows exponentially:

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 $\rm T_{O}$ is a remarkable quantity: the partition function corresponding to the above $\rho\left(m\right)$ diverges for $\rm T\rightarrow T_{O}^{-}$. $\rm T_{O}^{-}$ is therefore the highest possible temperature for strong interactions. It should - via a Maxwell-Boltzmann law - govern the transversal momentum distribution in all high energy collisions of hadrons (including e.m. form factors, etc.). There is experimental evidence for that, and then T_{O}^{-} is about 158 MeV ($\approx 10^{12}$ oK). With this value of T_{O}^{-} the asymptotic mass spectrum of our theory has a good chance to be the correct extrapolation of the experimentally known spectrum.

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Think about adding more and more energy to a system...

Excite more states
Give states more energy,
i.e., a higher Temperature

Equipartition Theorem

With exponentially increasing number of states, all the energy goes in to equally exciting these states, and not more energy per state.

 T_0 is a remarkable quantity: the partition function corresponding to the above $\rho(m)$ diverges for $T \rightarrow T_0$. T_0 is therefore the highest possible temperature for strong interactions.

"Ultimate Temperature in the Early Universe"

K. Huang & S. Weinberg, Phys Rev Lett 25, 1970.



"...a veil, obscuring our view of the very beginning"

Steven Weinberg, *The First Three Minutes* (1977).

Quark Model



New degrees of freedom inside hadrons.

Rapidly people started to think about *Quark Matter* as being relevant at these very high temperatures.

We have strong evidence that QCD is the correct theory of the strong interaction

Perturbative Calculations



Lattice Calculations



The field is about understanding emergent phenomena from QCD, just like condensed matter physics from QED.

Quark-Gluon Plasma

Birth of a Name

PHYSICS REPORTS

A Review Section of Physics Letters

QUANTUM CHROMODYNAMICS AND THE THEORY OF SUPERDENSE MATTER E.V. Shuryak, Quantum Chromodynamics and the Theory of Superdense Matter

73

1. Introduction

1.1. Preface

It is widely believed that the fundamental theory of strong interactions is the so called quantum chromodynamics (QCD), a theory of colored quarks interacting via massless vector fields, the gluons. This theory not only provides a general understanding of hadronic phenomenology and a good quantitative description of small distance phenomena, but it mostly wins our hearts by the remarkable simplicity of its foundations, so similar in spirit to quantum electrodynamics (QED). The properties of superdense matter were always of interest for physicists. Now, relying upon QCD, we can say much more about them. When the *energy* density ε exceeds some typical hadronic value (~1 GeV/fm³), matter no longer consists of separate hadrons (protons, neutrons, etc.), but of their fundamental constituents, quarks and gluons. Because of the apparent analogy with similar phenomena in atomic physics we may call this phase of matter the QCD (or quark-gluon) plasma. Due to large similarity between QCD and QED the new theory benefits from the methods previously elaborated for QED plasma made of electrons and photons.

There exist important nonperturbative effects, which result in qualitative differences between QCD and QED. This is seen already from the fact, that quarks and gluons are absent in the physical

Edward Shuryak publishes first "review" of thermal QCD and coins a phrase:

R

114

117

120

120

"Because of the apparent analogy with similar phenomena in atomic physics, we may call this phase of matter the QCD (or quark-gluon) plasma"

Melting the Hadrons

Can we melt the hadrons and liberate these quark and gluon degrees of freedom?

$$\varepsilon = g \frac{\pi^2}{30} T^4$$

Energy density for "g" massless d.o.f.



Hadronic Matter: quarks and gluons confined For T ~ 200 MeV, 3 pions with spin=0

$$\varepsilon = \left\{ 2 \cdot 8_g + \frac{7}{8} \cdot 2_s \cdot 2_a \cdot 2_f \cdot 3_c \right\} \frac{\pi^2}{30} T^4$$
$$\varepsilon = 37 \cdot \frac{\pi^2}{30} T^4 \qquad 37 !$$

Quark Gluon Matter: 8 gluons; 2 quark flavors, antiquarks, 2 spins, 3 colors

No Limiting Temperature

Lattice QCD calculations indicate that as one increases the energy input, it is very hard to move the temperature above approximately 170 MeV.

However, eventually the temperature does exceed the "limiting temperature" with a rapid jump in the effective degrees of freedom!





Question: why only 2 or 3 flavors?

Free Quarks?

No one has ever seen a free quark.



QCD is a "confining" gauge theory.

Lattice Thermodynamics

Lattice QCD (for heavy quarks as a test) show a screening of the long-range confining potential gradually as one passes the transition temperature.





Phase Diagram of Nuclear Matter



Transition Order?



Lattice QCD results for a realistic strange quark mass indicate a smooth cross over transition for zero net baryon density

This has substantial implications.

Huge Paradigm Change in the Field (!)



When I started in the field in 1990 as an undergraduate student, for μ_B = 0, many thought the QGP transition was 1st order.



Shoji Nagamiya, https://academic.oup.com/ptep/article/201 <u>5/3/03A101/1584228</u>

Smooth Crossovers are Cool Too!



"In a strong first-order phase transition the quark–gluon plasma supercools before bubbles of hadron gas are formed. These bubbles grow, collide and merge, during which gravitational <u>waves could be produced⁸</u>. Baryon-enriched nuggets could remain between the bubbles, contributing to dark matter. The hadronic phase is the initial condition for nucleosynthesis, so inhomogeneities in this phase could have a strong effect on nucleosynthesis⁹. As the firstorder phase transition weakens, these effects become less pronounced. Our calculations provide strong evidence that the QCD transition is a crossover, and thus the above scenarios and many others—are ruled out."

Lattice QCD, https://www.nature.com/articles/nature05120

Speed of Sound



Speed of sound drops near transition ("soft point in EoS") and actually goes to zero in first order transition.

Sound wave transmits energy. In true mixed phase, all energy is absorbed into rearranging constituents.

Phase Diagram of Nuclear Matter



QGP and Cosmology

Brief History of Time

- Radius of the Visible Universe ->



Annu. Rev. Nucl. Part. Sci. 2006. 56:441–500 doi: 10.1146/annurev.nucl.56.080805.140539 Copyright © 2006 by Annual Reviews. All rights reserved First published online as a Review in Advance on August 2, 2006

PHASE TRANSITIONS IN THE EARLY AND PRESENT UNIVERSE

D. Boyanovsky,^{1,2,3} H.J. de Vega,^{3,2,1} and D.J. Schwarz⁴

Quark nuggets and strangelets

Inhomogeneous nucleosynthesis

Cold dark matter clumps

Damping of gravitational waves at the QCD transition



Figure 11 The modification of the energy density, per logarithmic frequency interval, of primordial gravitational waves from the QCD transition. Figure taken from Reference 107.

VOLUME 30, NUMBER 2

Cosmic separation of phases

Edward Witten* Institute for Advanced Study, Princeton, New Jersey 08540 (Received 9 April 1984)

A first-order QCD phase transition that occurred reversibly in the early universe would lead to a surprisingly rich cosmological scenario. Although observable consequences would not necessarily survive, it is at least conceivable that the phase transition would concentrate most of the quark excess in dense, invisible quark nuggets, providing an explanation for the dark matter in terms of QCD effects only. This possibility is viable only if quark matter has energy per baryon less than 938 MeV. Two related issues are considered in appendices: the possibility that neutron stars generate a quark-matter component of cosmic rays, and the possibility that the QCD phase transition may have produced a detectable gravitational signal.





FIG. 1. Isolated expanding bubbles of low-temperature

Strange Quark Matter

 Matter of roughly equal numbers of up, down and strange quarks.

• Strange Quark Matter could be more stable that Fe⁵⁵ and thus be the ground state of nuclear matter.

• If stable, it could be a source of baryonic dark matter.



Creating Strangelets

- Many years later, SQM still theoretically allowed.
- Experiments searches in terrestrial matter and nuclear reactions for small A< 100 SQM have yielded null results.


LITTLEST NEUTRON STAR

Strange star: this bizarrely small neutron star may be made of quarks

Quarks all the way down



* BY KIONA SMITH OCT. 24, 2022

<u>In the middle</u> of the glowing gas cloud of a supernova remnant, about 8,000 light years away, sits the crushed heart of a dead star.

Astronomers recently discovered that this neutron star left behind by the <u>collapse and explosion of a</u> <u>supergiant</u> is now roughly 77 percent the mass of our Sun, packed into a sphere about 10 kilometers wide. That's a mind-bogglingly dense ball of matter — it's squished together so tightly that it doesn't even have room to be atoms, just neutrons. But as neutron stars go, it's weirdly lightweight. Figuring out why that's the case could reveal fascinating new details about exactly what happens when massive stars collapse and explode.

Neutron star physics via gravitational waves!

http://science.energy.gov/~/media/np/nsac/pdf/201603/Weinstein_LIGO_20160323.pdf



LSC

Compact Binary Mergers, Nuclear EOS, and LIGO

- GWs and LIGO
- Compact binary mergers
- prospects for LIGO GW detection and EM counterparts
- Neutron stars
- Nuclear EOS
- Neutron star mass & radius
- BNS mergers
- BNS r-process nucleosynthesis
- BNS merger constrains on NEOS

Alan Weinstein, Caltech for the LIGO Scientific Collaboration

> DOE/NSF NSAC Meeting, Bethesda, March 23, 2016



"Merging Neutron Stars" (Price & Rosswog)





Supercooling and Bubbles

If the plasma-to-hadrons transition were strongly first order, bubble formation could lead to an inhomogeneous early universe, thus impacting big bang nucleosynthesis (BBN).

Are the bubbles too small and close together such that diffusion before nucleosynthesis erases the inhomogeneities? (200 MeV to 2 MeV)

This line of investigation was quite active when the dark matter issue raised questions about the implied baryon content in the universe from BBN.



No BBN Problem

Physics Today, July 2001: Cosmic Microwave Background Observations

"The value deduced from the second harmonic in the acoustic oscillations for $W_B=0.042 \pm 0.008$ (cosmic baryon mass density) is in very good agreement with the value one gets by applying the theoretical details of primordial big bang nucleosynthesis to the observations of cosmic abundances of deuterium."

In addition, Lattice QCD now confirms a smooth cross over transition.



Big Bang versus Little Bangs



<u>Universe Case:</u> Fluctuations pre-inflation... and not impacted by QGP Heavy Ion Case: QGP initial condition fluctuations

Study structures that survive to understand earlier epoch

Heavy Ion Experiments

How to Access This Physics?





Accelerators



Heavy Ion Machine History

Bevalac-LBL and SIS-GSI fixed target max. 2.2 GeV

1992 Au-Au AGS-BNL fixed target max. 4.8 GeV

E864/941, E802/859/866/917, E814/877, E858/878, E810/891, E896, E910 ...

1994 Pb-Pb SPS-CERN fixed target max. 17.3 GeV

NA35/49, NA44, NA38/50/51, NA45, NA52, NA57, WA80/98, WA97, ...

TEVATRON-FNAL (fixed target p-A) max. **38.7 GeV**



2000 Au-Au

2010

Pb-Pb

RHIC-BNL collider max. 200.0 GeV

BRAHMS, PHENIX, PHOBOS, STAR, now sPHENIX

LHC-CERN collider max. 2760.0 GeV Now even higher ALICE, ATLAS, CMS, LHCb (pA)



Particle Physics the energy frontier

> Machine with highest energy for point-like interactions dominates the world!



<u>Nuclear Physics</u> – studying a state of matter (thermodynamics, many-body physics) and how that emerges from underlying QCD theory

Think of mapping out a phase diagram with only one temperature?



Phase Diagram of Nuclear Matter



Experimental Tools











<u>STAR</u>



Hadronic Observables over a Large Acceptance Event-by-Event Capabilities

Solenoidal magnetic field Large coverage Time-Projection Chamber Silicon Tracking, RICH, EMC, TOF



Electrons, Muons, Photons and Hadrons Measurement Capabilities Focus on Rare Probes: J/ψ , high-p_T

Two central spectrometers with tracking and electron/photon PID Two forward muon spectrometers



BRAHMS

Hadron PID over broad rapidity acceptance

Two conventional beam line spectrometers Magnets, Tracking Chambers, TOF, RICH





Charged Hadrons in Central Spectrometer Nearly 4π coverage multiplicity counters

> Silicon Multiplicity Rings Magnetic field, Silicon Strips, TOF



Biggest contributions – intellectually independent thinking!





End of the World!



Can be dismissed with some basic General Relativity

$$R_{S} = \frac{2GM}{c^{2}} = 10^{-49} meters$$

 $R = 10^{-15} meters$

much less than Planck length !

Even if it could form, it would evaporate by Hawking Radiation in 10⁻⁸³ seconds !

Also, concerns about SQM nuggets eating the earth



ALICE – dedicated heavy ion detector

Key advantages – particle identification, focused 1000+ people



ATLAS & CMS –

full calorimetry, tracking, *different approaches and analysis philosophies*



sPHENIX – new collider detector at RHIC



More on sPHENIX later..

Lecture #2

Collision Dynamics

Structure of the Proton

See the whole proton Momentum transfer $Q^2 = 0.1 \text{ GeV}^2$ Wavelength $\lambda = h/p$



See the quark substructure $Q^2 = 1.0 \text{ GeV}^2$

See many partons (quarks and gluons)

 $Q^2 = 20.0 \text{ GeV}^2$

Parton Distribution Functions



- Structure functions rise rapidly at low-x
- More rapid for gluons than quarks
- Watch out for side-by-side plots with different vertical scales (left 1.7 and right 30!)

QGP in Proton+Proton Reactions?

Bjorken speculated that in the "interiors of large fireballs produced in very high-energy pp collisions; vacuum states of the strong interactions are produced with anomalous chiral order parameters."







"Baked Alaska"

http://www-minimax.fnal.gov/

<u>Fermi (1950)</u>

790

"High Energy Nuclear Events", Prog. Theor. Phys. 5, 570 (1950)

Groundwork for statistical approach to particle production in strong interactions:

"Since the interactions of the pion field are *strong*, we may expect that rapidly this energy will be distributed among the various degrees of freedom present in this volume according to <u>statistical laws</u>." 241.

HIGH ENERGY NUCLEAR EVENTS

241. - High Energy Nuclear Events

« Progr. Theor. Theoret. Phys. », 5, 570-583 (1950).

ABSTRACT

A statistical method for computing high energy collisions of protons with multiple production of particles is discussed. The method consists in assuming that as a result of fairly strong interactions between nucleons and mesons the probabilities of formation of the various possibile numbers of particles are determined essentially by the statistical weights of the various possibilities.

I. INTRODUCTION

The meson theory has been a dominant factor in the development of physics since it was announced fifteen years ago by Yukawa. One of its outstanding achievements has been the prediction that mesons should be produced in high energy nuclear collisions. At relatively low energies only one meson can be emitted. At higher energies multiple emission becomes possible. ρ

In this paper an attempt will be made to develop a crude theoretical approach for calculating the outcome of nuclear collisions with very great energy. In particular, phenomena in which two colliding nucleons may give rise to several π -mesons, briefly called hereafter pions, and perhaps also to some anti-nucleons, will be discussed.

In treating this type of processes the conventional perturbation theory solution of the production and destruction of pions breaks down entirely. Indeed, the large value of the interaction constant leads quite commonly to situations in which higher approximations yield larger results than do lower approximations. For this reason it is proposed to explore the possibilities of a method that makes use of this fact. The general idea is the following:

When two nucleons collide with very great energy in their center of mass system this energy will be suddenly released in a small volume surrounding the two nucleons. We may think pictorially of the event as of a collision in which the nucleons with their surrounding retinue of pions hit against each other so that all the portion of space occupied by the nucleons and by their surrounding pion field will be suddenly loaded with a very great amount of energy. Since the interactions of the pion field are strong we may expect that rapidly this energy will be distributed among the various degrees of freedom present in this volume according to statistical laws. One can then compute statistically the probability that in this tiny volume a certain number of pions will be created with a given energy distribution. It is then assumed that the

Landau (1955)

Significant extension of Fermi's approach

Considers fundamental roles of

- Hydrodynamic evolution

Entropy

"The defects of Fermi's theory arise mainly because the expansion of the compound system is not correctly taken into account...(The) expansion of the system can be considered on the basis of relativistic hydrodynamics."

88. A HYDRODYNAMIC THEORY OF MULTIPLE FORMATION OF PARTICLES

1. INTRODUCTION

Experiment shows that in collisions of very fast particles a large number of new particles are formed in multi-prong stars. The energy of the particles which produce such stars is of the order of 10^{12} eV or more. A characteristic feature is that such collisions occur not only between a nucleon and a nucleus but also between two nucleons. For example, the formation of two mesons in neutron-proton collisions has been observed at comparatively low energies, of the order of 10^9 eV , in cosmotron experiments¹.

Fermi^{2,3} originated the ingenious idea of considering the collision process at very high energies by the use of thermodynamic methos. The main points of his theory are as follows.

(1) It is assumed that, when two nucleons of very high energy collide, energy is released in a very small volume V in their centre of mass system. Since the nuclear interaction is very strong and the volume is small, the distribution of energy will be determined by statistical laws. The collision of high-energy particles may therefore be treated without recourse to any specific theories of nuclear interaction.

(2) The volume V in which energy is released is determined by the dimensions of the meson cloud around the nucleons, whose radius is $\hbar/\mu c$, μ being the mass of the pion. But since the nucleons are moving at very high speeds, the meson cloud surrounding them will undergo a Lorentz contraction in the direction of motion. Thus the volume V will be, in order of magnitude,

$$V = \frac{4\pi}{3} \left(\frac{\hbar}{\mu c}\right)^3 \frac{2M c^2}{E'},$$
 (1.1)

where M is the mass of a nucleon and E' the nucleon energy in the centre of mass system.

(3) Fermi assumes that particles are formed, in accordance with the laws of statistical equilibrium, in the volume V at the instant of collision. The particles formed do not interact further with one another, but leave the volume in a "frozen" state.

С. З. Беленький и Л. Д. Ландау, Гидродинамическая теория множественного образования частиц, Успехи Физических Наук, 56, 309 (1955).

S. Z. Belenkij and L. D. Landau, Hydrodynamic theory of multiple production of particles, Nuovo Cimento, Supplement, 3, 15 (1956).

Why Heavy Ions?

- High energy density may be achieved in protonproton collisions, but the partonic re-interaction time scale is only of order 1-3 fm/c
- It is difficult to select events with different geometries and avoid autocorrelations
- We will see that probes with long paths through the medium are key observables
- We should not rule out p+p reactions, but rather study the similarities and differences with A+A reactions -- spoiler alert: this is a big issue

RHIC and LHC = Gluon Colliders



20,000 gluons, quarks, and antiquarks are made physical in the laboratory !

What is the nature of this ensemble of partons? New emergent phenomena?

Heavy Ion Time Evolution



- 1. Initial Nuclei Collide
- 2. Partons are Freed from Nuclear Wavefunction
- 3. Partons interact and potentially form a Quark-Gluon Plasma
- 4. System expands and cools off
- 5. System Hadronizes and further Re-Scatters
- 6. Hadrons and Leptons stream towards our detectors



Diagram from Peter Steinberg

Initial Conditions



Glauber Model and Characterization



https://tglaubermc.hepforge.org/

Home

- Subversion
- Tracker
- Wiki

TGlauberMC: A ROOT-based implementation of the PHOBOS Glauber Monte Carlo

Authors: Burak Alver (MIT), Mark Baker (BNL), Constantin Loizides (MIT), Peter Steinberg (BNL) Brookhaven National Laboratory (BNL) & Massachusetts Institute of Technology (MIT)

"Glauber" models are used to calculate geometric quantities in the initial state of heavy ion collisions, such as impact parameter, number of participating nucleons and initial eccentricity. The four RHIC experiments have different methods for Glauber Model calculations, leading to similar results for various geometric observables. In this document, we describe an implementation of the Monte Carlo based Glauber Model calculation used by the PHOBOS experiment. The assumptions that go in the calculation are described. A user's guide, arXiv:0805.4411, is provided for running various calculations.

An **improved version (v2)** by C. Loizides (LBNL), J. Nagle (Colorado U.), P. Steinberg (BNL) is described in arXiv:1408.2549, which includes tritium, Helium-3, and Uranium, as well as the treatment of deformed nuclei and Glauber-Gribov fluctuations of the proton in p+A collisions.

For the latest release see the TGlauberMC downloads page.

Click here to contact the authors with questions.

Collision Characterization

The impact parameter determines the number of nucleons that participate in the collision.


Glauber Modeling in High-Energy Nuclear Collisions

Annual Review of Nuclear and Particle Science Vol. 57: 205-243 (Volume publication date November 2007) First published online as a Review in Advance on May 9, 2007 DOI: 10.1146/annurev.nucl.57.090506.123020

Michael L. Miller,¹ Klaus Reygers,² Stephen J. Sanders,³ and Peter Steinberg⁴

Relate experimental observables to averages from the Monte Carlo Glauber.

 N_{coll} (binary collisions), N_{part} (participants), N_{spec} (spectators) b (impact parameter), elliptical shape, orientation, overlap area,...

$$\varepsilon_n = \frac{\sqrt{\langle r^2 cos(n\phi) \rangle^2 + \langle r^2 sin(n\phi) \rangle^2}}{\langle r^2 \rangle}$$

$$S = 4\pi \sqrt{\langle x^2 \rangle \langle y^2 \rangle - \langle xy \rangle^2}$$

ALICE Centrality Paper – good resource http://arxiv.org/abs/1301.4361



⁵-10%¹0-20%²0-30%³0-40%⁴0-50%⁵0-60%⁶0-70%⁷0-88% Centrality



d+Au Sometimes the neutron misses



http://journals.aps.org/prc/abstract/10.1103/PhysRevC.90.034902

0.05

n

0-5%

How much energy goes into QGP?



At RHIC out of a maximum energy of 39.4 TeV in central Gold+Gold reactions, <u>26 TeV</u> is made available for heating the system.

Bjorken Energy Density

$$\varepsilon_{Bj} = \frac{1}{\pi R^2} \frac{1}{2c\tau} \left(2 \frac{dE_T}{dy} \right)$$



Early energy density well above expected QGP transition



* Side note about errata or lack thereof..

Collision Dynamic Summary

- Depositing majority of kinetic energy into new medium
- Energy density appears above phase transition value, noting that it is a smooth crossover
- Energy is distributed into particle production statistically
- No sharp global feature distinct from smaller hadron collisions, but instead gradual changes

The Perfect Liquid



RHIC Scientists Serve Up "Perfect" Liquid

New state of matter more remarkable than predicted -- raising many new questions

The Key Paradigm in the Field

What do we mean by Perfect?

What evidence is there?

Non-Central A+A Geometry



Ideal Hydrodynamics

Key Inputs:

- Initial Geometry
- QCD Equation of State





Assumes early thermalization [not proven] Assumes no dissipation (shear/bulk viscosity = 0)

Fluid cells "freeze-out" below T_{freeze} Isotropic hadrons in cell rest frame, then boosted



Temperature Profile + Fluid Cell Velocity Vectors

$\underline{\mathsf{Fluid}} \xrightarrow{} \mathsf{Hadrons}$

Single-particle distribution in the hydrodynamic and statistical thermodynamic models of multiparticle production

Fred Cooper and Graham Frye Phys. Rev. D **10**, 186 – Published 1 July 1974

$$E\frac{dN}{d^3p} = \int_{\Sigma} d\Sigma_{\mu} p^{\mu} f(T, p_{\mu} u^{\mu}, \pi^{\mu\nu}) \,,$$

An important, but not often discussed, assumption in the calculations.



Always Read the Original Material

PHYSICAL REVIEW D

VOLUME 46, NUMBER 1

Anisotropy as a signature of transverse collective flow

Jean-Yves Ollitrault Service de Physique Théorique, Centre d'Études de Saclay, F-91191 Gif-sur-Yvette CEDEX, France (Received 19 February 1992)

We show that anisotropies in transverse-momentum distributions provide an unambiguous signature of transverse collective flow in ultrarelativistic nucleus-nucleus collisions. We define a measure of the anisotropy from experimental observables. The anisotropy coming from collective effects is estimated quantitatively using a hydrodynamical model, and compared to the anisotropy originating from finite multiplicity fluctuations. We conclude that collective behavior could be seen in Pb-Pb collisions if a few hundred particle momenta were measured in a central event.

PACS number(s): 25.75.+r, 12.38.Mh, 24.60.Ky, 47.75.+f





Article Zeitschrift für Physik C Particles and Fields December 1996, Volume 70, Issue 4, pp 665-671

First online: 31 March 2014

Flow study in relativistic nuclear collisions by Fourier expansion of azimuthal particle distributions

1 JULY 1992

S. Voloshin, Y. Zhang

Common problem – not reading the original references...

V = Voloshin (?)

How do experiments measure v2?

An entire lecture could be on these details...

http://journals.aps.org/prc/abstract/10.1103/PhysRevC.80.014904

PHYSICAL REVIEW C

covering nuclear physics

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Effect of flow fluctuations and nonflow on elliptic flow methods

Jean-Yves Ollitrault, Arthur M. Poskanzer, and Sergei A. Voloshin Phys. Rev. C 80, 014904 – Published 10 July 2009

Short introduction to some experimental basics

Two Particle Correlations

Two independent particles that come from a common source distribution $1 + 2v_2 cos[2(\phi-\psi_2)]$

<u>Resulting</u> Δφ

Random Case

Distribution

Divide the two FG/BG

Oscillation 1 + $2v_2^2 \cos(2\Delta\phi)$



<u>Complications and Other Contributions?</u> Momentum Conservation



https://root.cern.ch/doc/master/classTGenPhaseSpace.html

Jet Correlations



Perfect Fluidity Discovery - 2005

Agreement of ideal hydrodynamics with experimental data.





Heavier particles get a larger momentum boost from the fluid velocity and so heavier hadron v_2 pattern shifted to higher p_T .

What About Viscosity?

Relativistic Viscous Hydrodynamics major unsolved numerical problem, until 2007

Viscosity Information from Relativistic Nuclear Collisions: How Perfect is the Fluid Observed at RHIC?

Paul Romatschke and Ulrike Romatschke Phys. Rev. Lett. **99**, 172301 – Published 24 October 2007

Causal viscous hydrodynamics in 2 + 1 dimensions for relativistic heavy-ion collisions

Huichao Song and Ulrich Heinz Phys. Rev. C 77, 064901 – Published 5 June 2008

Shear Viscosity





Viscosity Review

Honey – viscosity decreases at higher temperatures viscosity increases with stronger coupling

Weak coupling (σ =0)

NagleLab Productions (2007)



Inhibited diffusion Small viscosity Perfect fluid Strong Coupled QGP (i.e. sQGP)

What is η/s for the Quark-Gluon Plasma



Lecture #3

Work / Life Balance?





Perfect Fluid

How to Quantify QGP η/s?

Relativistic viscous hydrodynamics compared to data

_uzum, Romatschke, Phys. Rev. C78, 034915 (2008)



What dominates the uncertainty?



At the time, different experimental flow methods gave different v_2 results. Now these differences are understood from non-flow contributions and fluctuations.





What dominates the uncertainty?



The v_2 you get out is directly related to the ε_2 of the initial geometry you put in.

Different initial geometry models yield 20% ϵ_2 differences resulting in 100% η /s differences.



Different models of the initial geometry.

Uncertainty by considering model A and model B.

Systematic Uncertainties

Systematic Errors

Joel Heinrich¹ and Louis Lyons²

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²Department of Physics, University of Oxford, Oxford OX1 3RH, United Kingdom; email: l.lyons@physics.ox.ac.uk

Not much help in many practical situations...

Example: Two model inputs give different results.

Uncertainty = 1 RMS = Difference / sqrt(12)

- = Difference / 2
- = Cannot determine

Alver and Roland Revolution 2010

Collision-geometry fluctuations and triangular flow in heavy-ion collisions

B. Alver and G. Roland Phys. Rev. C **81**, 054905 – Published 21 May 2010; Erratum Phys. Rev. C **82**, 039903 (2010)

http://journals.aps.org/prc/abstract/10.1103/PhysRevC.81.054905



Fluctuations in geometry yield not only elliptical shapes, but triangular, quadrangular, etc.



Detailed Fingerprint of Early Time





Calculation from Bjoern Schenke

Global Constraint Analysis



Global constraint methods using Bayesian sampling as done in Climate Modeling for example.

Includes particle spectra, elliptic flow, two-particle quantum correlations, ...

Experimental confirmation of Lattice QCD Equation of State

Global Constraint Analysis



Expect η /s to increase at higher temperatures even just from running of α_s

Key lesson about when and when not to include scenarios (story of High Voltage Power Lines)...

Heavy Quarks
Why are heavy quarks so special?

QCD is flavor blind, so why are charm and beauty quarks interesting?

Very difficult to produce charm (beauty) quark-antiquark pairs via thermal production in the QGP

 $g+g \rightarrow ccbar \text{ or } bbbar$

However, high energy gluons from the incoming nuclei can do the trick. Production dominated by "initial hard scattering" and perturbatively (pQCD) calculable even at low pT due to mass scale.

Cannot be destroyed via strong interaction with light quarks and gluons (i.e. QGP).

http://www.lpthe.jussieu.fr/~cacciari/fonll/fonllform.html

FONLL Heavy Quark Production

Calculate total or single inclusive differential cross sections for heavy quark production (charm or bottom) at ppbar or pp colliders, with cuts on pT and y or eta. Non perturbative fragmentation into heavy hadrons and their subsequent decays into other final states (leptons, quarkonia, ...) can also be included.

NEWS:

- 09 Jan 2016 Fixed bug (introduced on 23/12/2015) preventing correct selection of decay to electrons option
- 23 Dec 2015 New interface, with responsive framework for mobile devices and option to email results. No change in physics results.
- 28 Sep 2015 Modified smearing function used in integrations. Change in results should be well below 1%

Be sure to check the older news, as well as the list of bugs discovered so far. Feedback to matteo.cacciari@cern.ch is welcome.

Collider: LHC (pp, 13 TeV) [ptmax ≤ 300 GeV]	▼ PDFs: CTEQ6.6	▼ Perturbative order: FONLL ▼
Heavy quark: bottom	Hadronic final state: bare quark	Further decay:
Cross section type: dsigma/dpt	central prediction only	✓ □ Output all scales
	Include PDFs uncertaint	ies
ptmin (GeV) 5	Change these defaults only if you know what you are doing.	BR(D->I) = 0.103
ptmax (GeV) 20	Non-pert. FF: default	BR(B->I) = 0.1086
	Non-pert. par. = default	BR(B->D->I) = 0.096
y(eta)min _1		BR(B->D) = 0.823
y(eta)max 1	FF(c->D) = 1	BR(B->D*) = 0 173

Alternative Fluid Probe:

Put a pebble in the stream and watch something out of equilibrium then equilibrate.

Charm Quark



"Does the Charm Flow at RHIC?"

S. Batsouli, S. Kelly, M. Gyulassy (Columbia U.), J.L. Nagle (Colorado U.). Dec 2002. 11pp. Published in Phys.Lett.B557:26-32,2003. e-Print: nucl-th/0212068

At the time, many said this idea was "ridiculous".

Langevin Model - Drag and Diffusion

$$\frac{dp_i}{dt} = \xi_i(t) - \eta_D p_i, \qquad \langle \xi_i(t)\xi_j(t') \rangle = \kappa \delta_{ij}\delta(t-t').$$
(2)

Here η_D is a momentum drag coefficient and $\xi_i(t)$ delivers random momentum kicks that are uncorrelated in time.



Teaney and Moore http://journals.aps.org/prc/abstract/10.1103/PhysRevC.71.064904



PHENIX Heavy Flavor Result

Charm quarks dragged and diffused in QGP suppresses high p_T hadrons (R_{AA}↓)

And push generates "flow" (v₂ ↑)



Charm: Really moves with the medium



Lower momentum – drag and diffusion Higher momentum – jet quenching energy loss

Bottom quarks are harder to push around

State-of-the-art calculations

DREENA-B (arXiv:1805.04786)

Dusan Zigic, Igor Salom, Jussi Auvinen, Marko Djordjevic, Magdalena Djordjevic

Dynamic energy loss in 1+1D expanding QCD medium

DAB-MOD (arXiv: 1906.10768)

Roland Katz, Caio A. G. Prado, Jacquelyn Noronha-Hostler, Jorge Noronha, Alexandre A. P. Suaide 2D+1 viscous hydrodynamic expansion with event-by-event fluctuations

Both match mass splitting in R_{AA} at low p_T DREENA-B has closer match to elliptic flow v_2

It would be ideal to test both energy loss calculations on identical expanding QGP background (challenge to the theorists)

Heavy Quarkonia

Screening Effects

Different states "melt" at different temperatures due to different binding energies.

The ψ ' and χ_c melt below or at T_c the J/ ψ melts above T_c and eventually the Y(1s) melts.

state	J/ψ	χς	ψ	Y(1s)	χb	Y(2s)	χь'	Y(3s)
Mass [GeV}	3.096	3.415	3.686	9.46	9.859	10.023	10.232	10.355
B.E. [GeV]	0.64	0.2	0.05	1.1	0.67	0.54	0.31	0.2
T _d /T _c		0.74	0.15	\ /		0.93	0.83	0.74

hep-ph/0105234 - "indicate ψ ' and the χ_c dissociate below the deconfinement point."

Quarkonia Thermometer

PHENIX, STAR, and CMS data consistent with *melting* of Y(2s,3s)

Need more statistics

Many states constrains the temperature

120

Quark recombination

Bound states of c̄c and b̄b can be *Debye color screened* in the QGP as one increases the temperature (melting)

Bizarre twist Less suppression at LHC with higher temperature.
Quark deconfinement and charm recombination at

Heavy Quarkonia Recombination

Look to Upsilons for Debye Screening, no recombination at RHIC Less J/ψ suppression at LHC with higher temperature

Charm recombination dominant effect, not screening.

Upsilon sequential suppression

"Because the more excited states melt at lower temperatures, the sequential suppression is an excellent indicator of the temperature of the nuclear medium produced in PbPb collisions."

Jet Quenching

Probing the Matter

Matter we want to study

Calibrated

LASER

Calibrated Light Meter

Calibrated Heat Source

Autogenerated Quark "LASER"

FERMILAB-Pub-82/59-THY August, 1982

Energy Loss of Energetic Partons in Quark-Gluon Plasma: Possible Extinction of High p_T Jets in Hadron-Hadron Collisions.

> J. D. BJORKEN Fermi National Accelerator Laboratory P.O. Box 500, Batavia, Illinois 60510

Abstract

High energy quarks and gluons propagating through quark-gluon plasma suffer differential energy loss via elastic scattering from quanta in the plasma. This mechanism is very similar in structure to ionization loss of charged particles in ordinary matter. The dE/dx is roughly proportional to the square of the plasma temperature. For hadron-hadron collisions with high associated multiplicity and with transverse energy dE_T/dy in excess of 10 GeV per unit rapidity, it is possible that quark-gluon plasma is produced in the collision. If so, a produced secondary high- p_T quark or gluon might lose tens of GeV of its initial transverse momentum while plowing through quark-gluon plasma produced in its local environment. High energy hadron jet experiments should be analysed as function of associated multiplicity to search for this effect. An interesting signature may be events in which the hard collision occurs near the edge of the overlap region, with one jet escaping without absorption and the other fully absorbed.

Gluon Radiation

Partons are expected to lose energy via induced gluon radiation in traversing a dense partonic medium.

Coherence among these radiated gluons can lead to $\Delta E \alpha L^2$

Look for an effective modification in the jet fragmentation properties.

Baier, Dokshitzer, Mueller, Schiff, hep-ph/9907267 Gyulassy, Levai, Vitev, hep-pl/9907461 Wang, nucl-th/9812021 and many more.....

pQCD + Factorization + Universality

In heavy ion collisions we can calculate the yield of high p_T hadrons

$$E_{h}\frac{d\sigma_{h}^{pp}}{d^{3}p} = K\sum_{abcd} \int dz_{c}dx_{a}dx_{b} \int d^{2}\mathbf{k}_{\mathrm{T}a}d^{2}\mathbf{k}_{\mathrm{T}b}f(\mathbf{k}_{\mathrm{T}a})f(\mathbf{k}_{\mathrm{T}b})f_{a/p}(x_{a},Q_{a}^{2})f_{b/p}(x_{b},Q_{b}^{2}) \underbrace{D_{h/c}(z_{c},Q_{c}^{2})\frac{\hat{s}}{\pi z_{c}^{2}}\frac{d\sigma^{(ab\rightarrow cd)}}{d\hat{t}}\delta(\hat{s}+\hat{u}+\hat{t})$$

Flux of incoming partons (structure functions) from Deep Inelastic Scattering

Perturbative QCD

Fragmentation functions D(z) in order to relate jets to observed hadrons

Calibrating Our Probes

High energy probes are well described in proton-proton reactions by NLO Perturbative QCD.

Jets-to-Parton Map

Jet algorithm sums the energy (typically in calorimeters) and can approximate catching an individual parton's energy.

Very useful in particle physics, but often used in a different way in heavy ion physics.

Jet Algorithms

Cacciari, Salam, Soyez JHEP 0804:063 (2008)

Different methods to "group" energies.

1. The Anti-k(t) jet clustering algorithm Matteo Cacciari, Gavin P. Salam (Paris, LPTHE), Gregory Soyez (Brookhaven). Feb 2008. 12 pp. Published in JHEP 0804 (2008) 063 LPTHE-07-03 DOI: 10.1088/1126-6708/2008/04/063 e-Print: arXiv:0802.1189 [hep-ph] | PDF References | BibTeX | LaTeX(US) | LaTeX(EU) | Harvmac | EndNote ADS Abstract Service Detailed record - Cited by 3675 records 1000+

FastJet Code Publicly Available

http://fastjet.fr/

Releases Quick start Manual Doxygen Tools Contrib FAQ Home About FastJet A software package for jet finding in pp and e⁺e⁻ collisions. It includes fast native implementations of many sequential recombination clustering algorithms, plugins for access to a range of cone jet finders and tools for advanced jet manipulation. Release of FastJet 3.2.0, 17 March 2016 (release notes). Its main new feature is that it exposes the N2Plain and N2Tiled strategies for 3rd-party clustering algorithms under the form of two new classes (NNFJN2Plain and NNFJN2Tiled), similar to NNH. Latest stable release of fjcore (v3.2.0), 17 March 2016 Lightweight access to the core FastJet functionality (PseudoJet, JetDefinition, ClusterSequence and Selector). It consists of just two files, fjcore.hh and fjcore.cc, which can easily be included in 3rd party projects. Compile time: a few seconds. A fortran interface and basic examples are also included in the distribution. Download size: 74k. Release of FastJet Contrib 1.024, 21 June 2016 *** NEW *** A package of contributed add-ons to FastJet. This release brings • update of Nsubjettiness to version 2.2.4 (fixed bug with multi-pass axes) • update of VariableR to version 1.2.1 (fixed documentation and comments) © 2005-2016 Matteo Cacciari, Gavin P. Salam, Gregory Soyez - Bug report - Subscribe - Follow @fastjet_fr

- R.D. Field, Applications of perturbative QCD A lot of detailed examples.
- R. K. Ellis, W. J. Stirling and B. R. Webber, QCD and Collider Physics
- CTEQ, Handbook of Perturbative QCD
- CTEQ website.
- John Collins, The Foundation of Perturbative QCD Includes a lot new development.
- Yu. L. Dokshitzer, V. A. Khoze, A. H. Mueller and S. I. Troyan, Basics of Perturbative QCD More advanced discussion on the small-*x* physics.

Jets in Heavy lons

Start with single hadrons, then go to jets.

Much Higher Energy Jets at the LHC

Nuclear Modification Factor

Photons unaffected by medium, pizeros suppressed

Probe expectations

<u>Quarks versus Gluons</u>

$$\Delta E_{GLV} \approx \frac{9}{4} \alpha_s^3 \pi C_R \left(\frac{1}{\pi R^2} \frac{dN_g}{dy} \right) \left\{ Log \frac{2E}{\mu^2 L} \right\} L(\phi)$$

Resolution [1/fm]

Critical microscope resolution at RHIC.

Also, overlap between RHIC and LHC for simultaneous description.

Does the fluid response to the energy input?

Precision Imaging of QGP Over Key Kinematics

Jet probe precision data RHIC – 2023++ LHC – Run 3 + 4 statistics Small QGP

Remember your History

The idea about small QGP was somewhat lost, but maybe not for good scientific reasons.

No particles \rightarrow think fields / disturbed vacuum

Maybe the small number of final state particles is just not relevant...
CMS Proton-Proton Hint



correlations in proton-proton collisions at the LHC

The CMS collaboration, V. Khachatryan, A. M. Sirunyan, A. Tumasyan, W. Adam, T. Bergauer, M. Dragicevic, J. Erö, C. Fabjan, M. Friedl, R. Frühwirth, V. M. Ghete, J. Hammer, S. Hänsel, C. Hartl ... <u>show 2150 more</u>

Open Access | Article First Online: 27 September 2010 DOI: 10.1007/JHEP09(2010)091 Cite this article as: The CMS collaboration et al. J. High Energ. Phys. (2010) 2010: 91. doi:10.1007/JHEP09(2010)091

http://link.springer.com/article/10.1007%2FJHEP09%282010%29091

Fall 2012 Revolution – p+Pb Collisions



- Very clear ridge
- Higher moments
- Particle dependence
- Cumulants

Almost every Pb+Pb collective flow signature by ALICE, ATLAS, CMS now seen in p+Pb!

Geometry Tests at RHIC

Exploiting Intrinsic Triangular Geometry in Relativistic ³He + Au Collisions to Disentangle Medium Properties

J. L. Nagle,^{1,*} A. Adare,¹ S. Beckman,¹ T. Koblesky,¹ J. Orjuela Koop,¹ D. McGlinchey,¹ P. Romatschke,¹ J. Carlson,² J. E. Lynn,² and M. McCumber²

¹University of Colorado at Boulder, Boulder, Colorado 80309, USA ²Los Alamos National Laboratory,Los Alamos, New Mexico 87545, USA (Received 20 December 2013; revised manuscript received 27 June 2014; published 12 September 2014)

http://journals.aps.org/prl/abstract/10.1103/PhysRevLett.113.112301



Glauber + Hydrodynamics + Cascade Predictions



Romatschke, Nagle et al., http://journals.aps.org/prl/abstract/10.1103/PhysRevLett.113.112301



Has raised a lot of questions about the requirements for hydrodynamic behavior, equilibration, QGP droplet size..

The Future

Well worth reading for young people...

New 2023 Long Range Plan being written now!

REACHING FOR THE HORIZON



The 2015 LONG RANGE PLAN for NUCLEAR SCIENCE



A roadmap for the future. Note that large \$0.5 - \$1.5B projects take 7-15 years at least.

Capitalize on key investments.

Future facilities.

science.energy.gov/~/media/np/nsac/pdf/2015LRP/2015_LRPNS_091815.pdf

Beam Energy Scan – Phase II



A New Detector at RHIC



High data acquisition rate capability, 15 kHz

Sampling 0.6 trillion Au+Au interactions in one-year Maximizing efficiency of RHIC running

<u>Napkin Drawing \rightarrow Prototype \rightarrow Detector \rightarrow Physics</u>







If you have the chance to get involved in hardware (even a little) it is well worth while.



Commissioning underway as I speak...









Longer term... ALICE-3 Letter of Intent



https://cerncourier.com/a/alice-3-a-heavy-ion-detector-for-the-2030s/

