



Introduction to Relativistic Heavy Ion Physics

Lecture 4: New Dimensions

W.A. Zajc
Columbia University

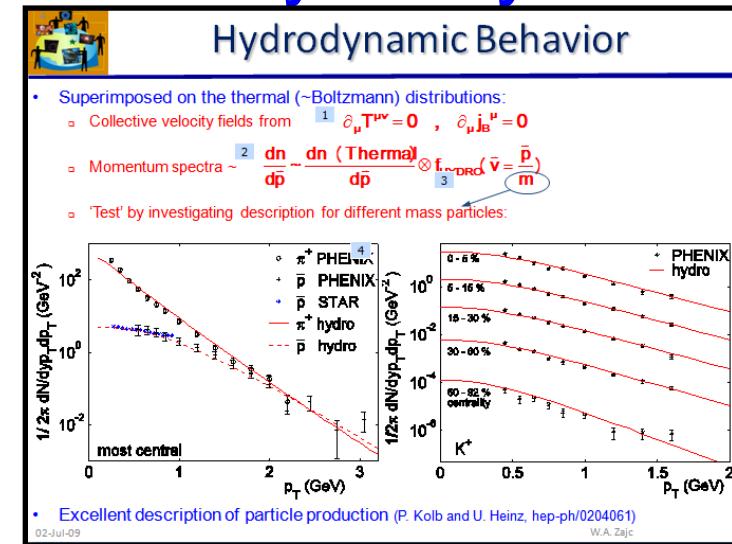
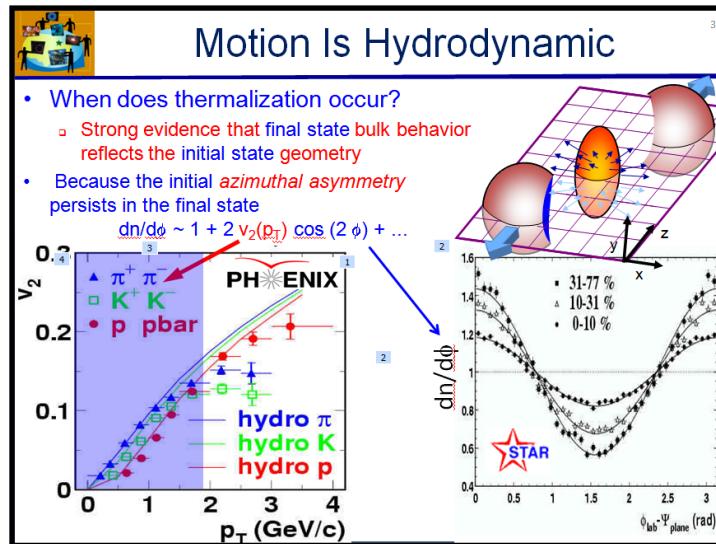


Reminder- From Lecture 3

- Strong evidence that initial-state spatial asymmetry appears as final-state “flow”
- The flow properties of QGP in Au+Au collisions at top RHIC energy is roughly consistent with perfect fluid ($\eta=0$) hydrodynamics
 - Particle mass dependence of $v_2(p_T)$
 - Scaling of same with KE_T
- Theoretical argument (Landau) suggests applicability of hydrodynamics to relativistic systems is approximately equivalent to requiring perfect fluid behavior.

What's The Problem ?

- Landau makes plausible argument for neglecting viscous effects in relativistic hydrodynamics
 - Caveat: Subject to condition $R / \lambda \gg 1$
- Data well-described by ideal hydrodynamics:



- Are we done ?



What Is The Viscosity at RHIC?

- “Perfect fluid” (and/or “ideal hydrodynamics”)
~ defined as “zero viscosity”.

$$\begin{aligned} \eta_{QGP} &\sim 2 \times 10^{11} \text{ Pa}\cdot\text{s} \\ \eta_{H_2O} &\sim 1 \times 10^{-3} \text{ Pa}\cdot\text{s} \end{aligned} \quad \Rightarrow \frac{\eta_{QGP}}{\eta_{H_2O}} \sim 2 \times 10^{14}$$

$$\eta_{Pitch} \sim 2.3 \times 10^8 \text{ Pa}\cdot\text{s} \quad \eta_{Glass(A.P.)} \sim 10^{12} \text{ Pa}\cdot\text{s}$$

?



A Check Of This Strange Conclusion

- Recall $\eta \sim n \bar{p} \lambda_{mfp}$

- But
$$\lambda_{mfp} = \frac{1}{n\sigma} \Rightarrow \eta \sim \frac{\bar{p}}{\sigma}$$

- Very Important Point !!
- To get small viscosity you need LARGE σ

- Using above

- $$\frac{\eta_{QGP}}{\eta_{H_20}} \sim \frac{\bar{p}_{QGP}}{\sigma_{QGP}} \frac{\sigma_{H_20}}{\bar{p}_{H_20}} \sim \frac{T_{QGP}}{\pi \left(\frac{1}{T_{QGP}} \right)^2} \frac{\pi a^2 H_20}{\sqrt{2\pi m k T_{H_20}}} \sim 10^{14}$$

Exercise 1: Check the above value , putting in plausible estimates for the various parameters.

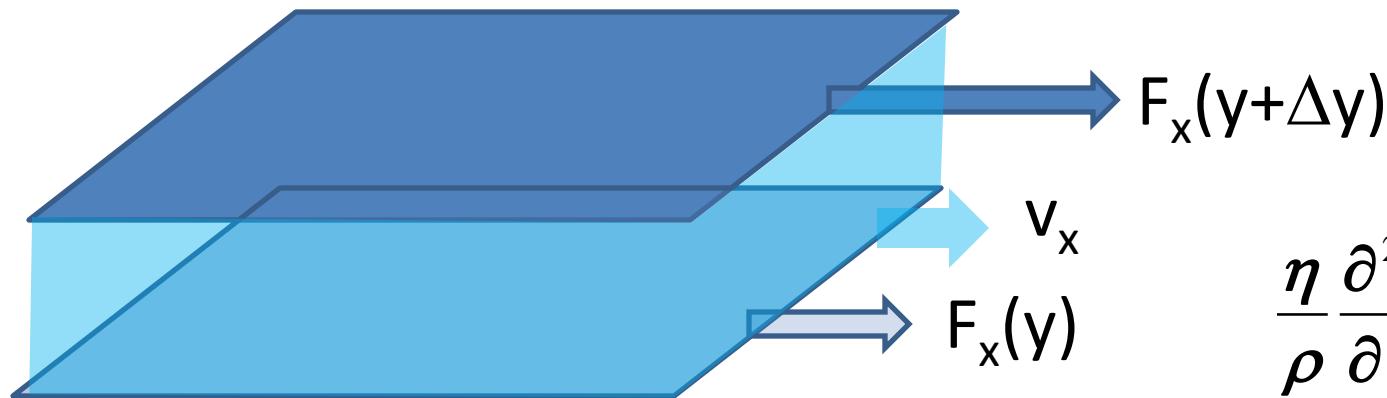
Kinematic Viscosity

- In Reynolds Number: $\text{Re} = \frac{\rho VL}{\eta}$

$$\frac{\rho VL}{\eta}$$

Exercise 2: At what velocity does it feel as if you are 'swimming' in water? (Save yourself some work- kinematic velocities are tabulated.)

- Determines relaxation rate



$$\frac{\eta}{\rho} \frac{\partial^2 v_x}{\partial y^2} = \frac{\partial v_x}{\partial t}$$

Exercise 3: a) Use $F=ma$ and the definition of viscosity to show that the relaxation of the velocity field $v_x(y)$ follows the diffusion equation.
b) verify solution to same

$$\Rightarrow v_x(y, t) \sim \frac{1}{\sqrt{4 \frac{\eta}{\rho} t}} e^{-y^2 / (4 \frac{\eta}{\rho} t)}$$



Why η/s Matters

- Any engineer will tell you
 - Kinematic viscosity $\eta / \rho \sim [\text{Velocity}] \times [\text{Length}]$ is what matters
(see Landau's remark on Reynolds number)

 - Any relativist will tell you
 - $\rho \rightarrow \varepsilon + P$

 - Any thermodynamicist will tell you
 - $\varepsilon + P = T s$ (at $\mu_B = 0$)

 - So
 - $\eta/\rho \rightarrow \eta/(\varepsilon + p) \rightarrow (\eta/sT) = (\eta/s)(1/T)$
 $\sim (\text{damping coefficient} \times \text{thermal time})$
- Exercise 4: a) Use this and previous statistical mechanics results for massless quanta to find an analytic result for entropy density s .
 b) Show that $s = 3.6 n$
 c) Instead of statistical expression for n , define n via $P = (N/V) T = n T$. Show that with this definition $s = 4 n$ for massless quanta. Comment.



A Long Time Ago (1985)

- Miklos Gyulassy and Pawel Danielewicz:
 - *Dissipative Phenomena in Quark-Gluon Plasmas*
P. Danielewicz, M. Gyulassy Phys.Rev. D31, 53, 1985.
- noted several restrictions on smallest allowed η :
- Most restrictive:
- $\lambda > h/\langle p \rangle \Rightarrow \eta > \sim n / 3$
- But for the quanta they were considering $s = 3.6$
- $\Rightarrow \eta/s > 1 / (3.6 \times 3) \sim 1 / (4 \pi) !!$



Before estimating λ_i via Eq. (3.2) we note several physical constraints on λ_i . First, the uncertainty principle implies that quanta transporting typical momenta $\langle p \rangle$ cannot be localized to distances smaller than $\langle p \rangle^{-1}$. Hence, it is meaningless to speak about mean free paths smaller than $\langle p \rangle^{-1}$. Requiring $\lambda_i \gtrsim \langle p \rangle_i^{-1}$ leads to the lower bound

$$\eta \gtrsim \frac{1}{3}n , \quad (3.3)$$

where $n = \sum n_i$ is the total density of quanta. What seems amazing about (3.3) is that it is independent of dynamical details. There is a finite viscosity regardless of how large is the free-space cross section between the quanta. See Refs. 21 and 22 for examples illustrating how the thermalization rate of many-body systems is limited by the uncertainty principle.

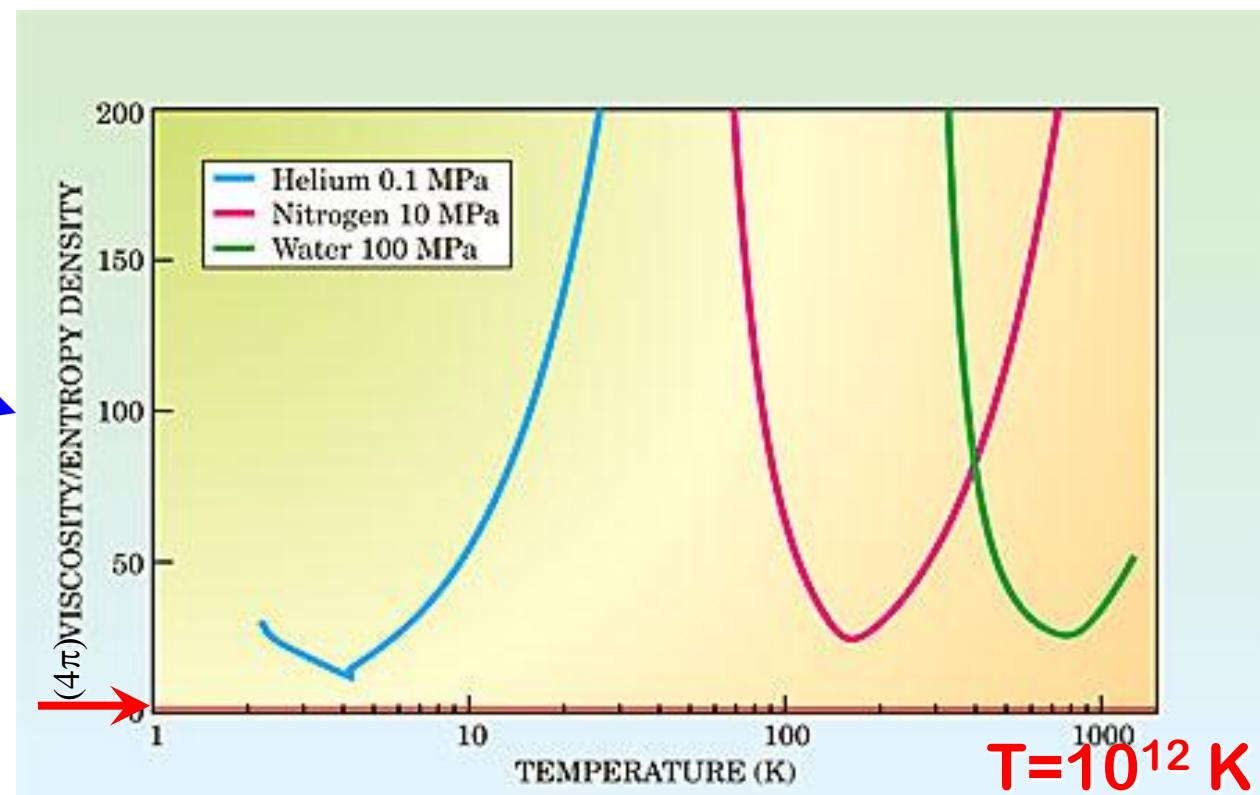
KSS Bound



- This bound is (now) *very* well-known in the nuclear physics community:
- “*A Viscosity Bound Conjecture*”,
P. Kovtun, D.T. Son, A.O. Starinets, hep-th/0405231

$$\frac{\eta}{s} \geq \frac{\hbar}{4\pi} \frac{1}{k_B}$$

- Where do “ordinary” fluids sit wrt this limit?
- RHIC “fluid” *might* be at ~2-3 on this scale (!)





Putting Viscosity Into Formalism

- The stress-energy tensor now contains off-diagonal terms:

- $T^{\mu\nu}$ will contain a piece called the shear stress tensor $\pi^{\mu\nu}$:

$$\pi^{\mu\nu} = T^{<\mu\nu>} \equiv \left[\frac{1}{2} \left(\Delta_\alpha{}^\mu \Delta_\beta{}^\nu + \Delta_\alpha{}^\nu \Delta_\beta{}^\mu \right) - \frac{1}{3} \Delta^{\mu\nu} \Delta_{\alpha\beta} \right] T^{\alpha\beta}$$

$$\Delta^{\mu\nu} \equiv g^{\mu\nu} - u^\mu u^\nu$$

Exercise 5: Show that Δ projects out the 3-volume orthogonal to four-velocity u .

- The existence of velocity gradients will produce a shear stress
 - Parameterize this via a 'constitutive equation':

$$\pi^{\mu\nu} = 2\eta \nabla^{<\mu} u^{\nu>}$$

- Then equation of motion is $\partial_\mu [(T^{\mu\nu} \text{perfect}_\text{fluid}) + \pi^{\mu\nu}] = 0$

- This is a simplified form of the relativistic Navier-Stokes eq.
- Ignores heat conduction, bulk viscosity



An Aside on Formalism

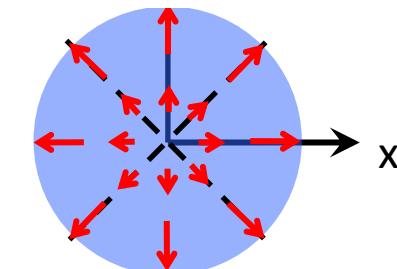
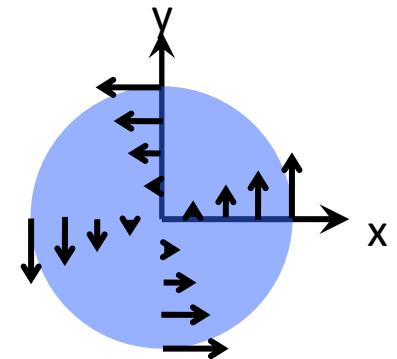
- Q. What are these weird index manipulations ?
- A. They produce a *symmetrized, traceless* ‘gradient’ :
 - Recall we’re interested in velocity gradients: $\frac{F_i}{A} = \mu \frac{\partial v_i}{\partial x_j}$
 - Remove uniform rotation:

$$\frac{\partial v_i}{\partial x_j} = \frac{1}{2} [\partial_j v_i + \partial_i v_j] + \frac{1}{2} [\partial_j v_i - \partial_i v_j]$$

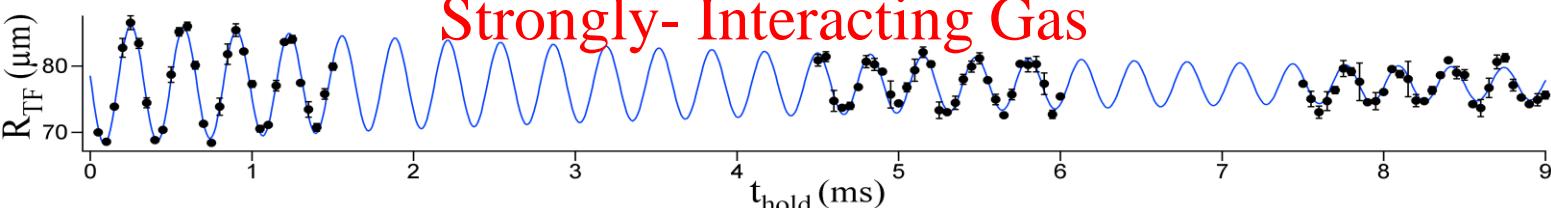
- Remove uniform (Hubble) expansion:

$$\frac{1}{2} [\partial_j v_i + \partial_i v_j] \rightarrow \frac{1}{2} [\partial_j v_i + \partial_i v_j - \frac{2}{3} \delta_{ij} (\partial_k v_k)]$$

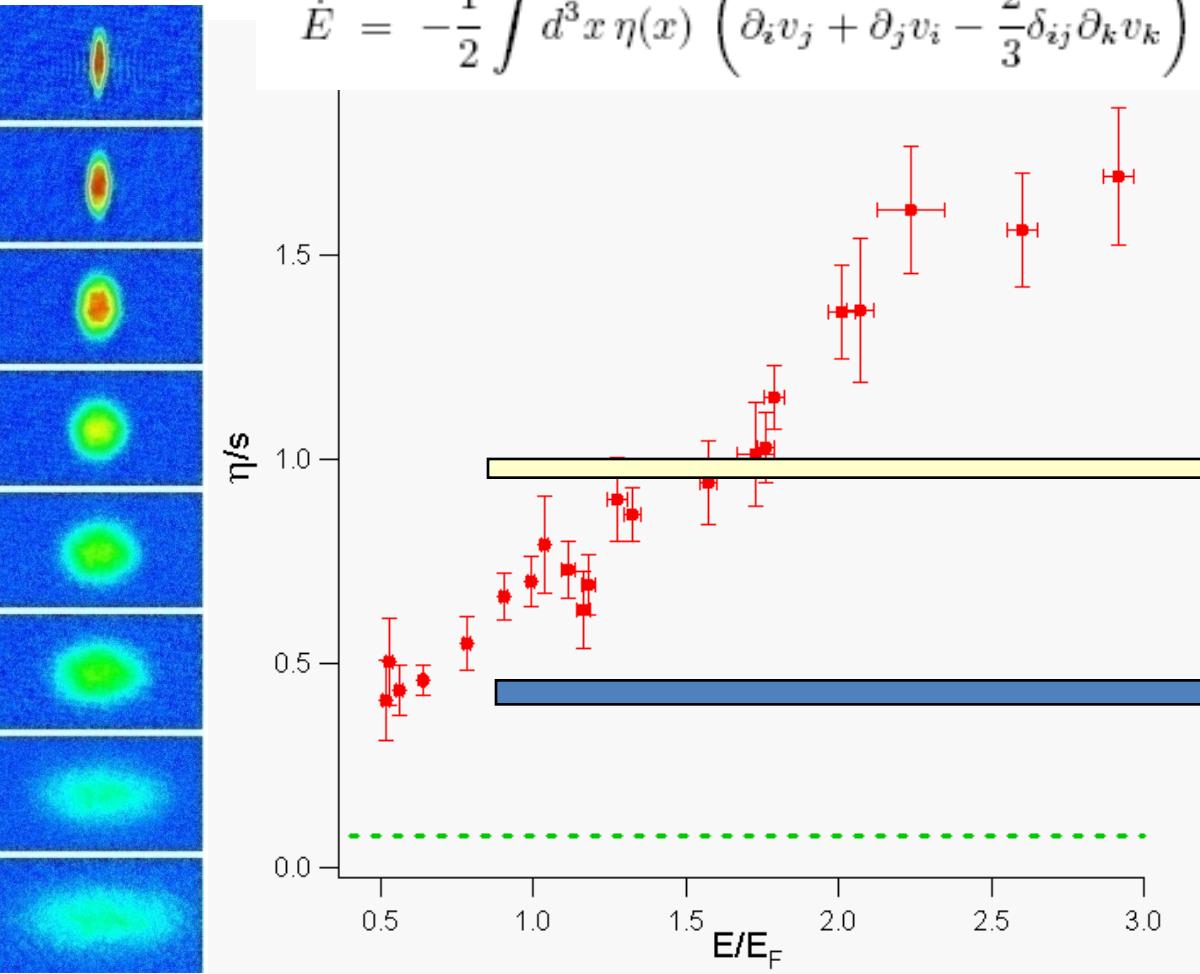
Exercise 6: Check these properties



Apply to a ‘Real’ Fluid



$$\dot{E} = -\frac{1}{2} \int d^3x \eta(x) \left(\partial_i v_j + \partial_j v_i - \frac{2}{3} \delta_{ij} \partial_k v_k \right)^2$$



- Damping of breathing mode in cold Fermi gas
- (All figures courtesy of John Thomas, Duke University)
 ${}^3\text{He}$, ${}^4\text{He}$ near λ -point

QGP simulations
String-theory $1/4\pi$



Estimating η/s For RHIC

- Damping (flow, fluctuations, heavy quark motion) $\sim \eta/s$

- FLOW: *Has the QCD Critical Point Been Signaled by Observations at RHIC?*,
R. Lacey *et al.*,
Phys.Rev.Lett.98:092301,2007
([nucl-ex/0609025](#))
- *The Centrality dependence of Elliptic flow, the Hydrodynamic Limit, and the Viscosity of Hot QCD*, H.-J. Drescher *et al.*,
([arXiv:0704.3553](#))
- FLUCTUATIONS: *Measuring Shear Viscosity Using Transverse Momentum Correlations in Relativistic Nuclear Collisions*,
S. Gavin and M. Abdel-Aziz,
Phys.Rev.Lett.97:162302,2006
([nucl-th/0606061](#))
- DRAG, FLOW: *Energy Loss and Flow of Heavy Quarks in Au+Au Collisions at $\sqrt{s_{NN}} = 200$ GeV* (PHENIX Collaboration),
A. Adare *et al.*,
to appear in Phys. Rev. Lett. ([nucl-ex/0611018](#))

$$\frac{\eta}{s} = (1.1 \pm 0.2 \pm 1.2) \frac{1}{4\pi}$$

$$\frac{\eta}{s} = (1.9 - 2.5) \frac{1}{4\pi}$$

$$\frac{\eta}{s} = (1.0 - 3.8) \frac{1}{4\pi}$$

$$\frac{\eta}{s} = (1.2 - 2.0) \frac{1}{4\pi}$$



Viscous Relativistic Hydrodynamics

- Why not do a ‘real’ (that is, Navier-Stokes) hydrodynamic calculation at RHIC?
 - Incorporate non-zero viscosity
 - ‘Invert’ to determine allowed range for η / s .
- Two little problems:
 - It’s wrong
 - ◆ Solutions are acausal
 - ◆ Needed patch:
$$\frac{\eta}{\rho} \frac{\partial^2 v_x}{\partial y^2} = \frac{\partial v_x}{\partial t} \rightarrow \frac{\eta}{\rho} \frac{\partial^2 v_x}{\partial y^2} = \frac{\partial v_x}{\partial t} + \tau_R \frac{\partial^2 v_x}{\partial t^2}$$
 - It’s wrong
 - ◆ Solutions are *intrinsically* unstable
 - ◆ No patch, must take *all terms* to 2nd order in gradients

$$v_x(y, t) \sim \frac{1}{\sqrt{4 \frac{\eta}{\rho} t}} e^{-y^2/(4 \frac{\eta}{\rho} t)}$$





A Partial List of “All Terms”

- Relativistic, Causal, second-order expansion:
 - Relativistic Fluid Dynamics: Physics for Many Different Scales
- Neglect various terms at your own risk:
 - Baier et al., Relativistic viscous hydrodynamics, conformal invariance, and holography
 - Natsuume and Okamura, Comment on “Viscous hydrodynamics relaxation time from AdS/CFT correspondence”

Working out the divergence of the entropy current, and making use of the equations of motion, we arrive at

$$\begin{aligned} \nabla_\mu s^\mu = & -\frac{1}{T}\tau \left[\nabla_\mu u^\mu + \beta_0 u^\mu \nabla_\mu \tau - \alpha_0 \nabla_\mu q^\mu - \gamma_0 T q^\mu \nabla_\mu \left(\frac{\alpha_0}{T} \right) + \frac{\tau T}{2} \nabla_\mu \left(\frac{\beta_0 u^\mu}{T} \right) \right] \\ & -\frac{1}{T} q^\mu \left[\frac{1}{T} \nabla_\mu T + u^\nu \nabla_\nu u_\mu + \beta_1 u^\nu \nabla_\nu q_\mu - \alpha_0 \nabla_\mu \tau - \alpha_1 \nabla_\nu \tau^\nu{}_\mu \right. \\ & \quad \left. + \frac{T}{2} q_\mu \nabla_\nu \left(\frac{\beta_1 u^\nu}{T} \right) - (1 - \gamma_0) \tau T \nabla_\mu \left(\frac{\alpha_0}{T} \right) - (1 - \gamma_1) T \tau^\nu{}_\mu \nabla_\nu \left(\frac{\alpha_1}{T} \right) \right] \\ & -\frac{1}{T} \tau^{\mu\nu} \left[\nabla_\mu u_\nu + \beta_2 u^\alpha \nabla_\alpha \tau_{\mu\nu} - \alpha_1 \nabla_\mu q_\nu + \frac{T}{2} \tau_{\mu\nu} \nabla_\alpha \left(\frac{\beta_2 u^\alpha}{T} \right) - \gamma_1 T q_\mu \nabla_\nu \left(\frac{\alpha_1}{T} \right) \right]. \end{aligned} \quad (300)$$

In this expression it should be noted that we have introduced (following Lindblom and Hiscock) two further parameters, γ_0 and γ_1 . They are needed because without additional assumptions it is not clear how the “mixed” quadratic term should be distributed. A natural way to fix these parameters is to appeal to the Onsager symmetry principle [58[⊕]], which leads to the mixed terms being distributed “equally” and hence $\gamma_0 = \gamma_1 = 1/2$.

Denoting the comoving derivative by a dot, i.e. using $u^\mu \nabla_\mu \tau = \dot{\tau}$ etc. we see that the second law of thermodynamics is satisfied if we choose

$$\tau = -\zeta \left[\nabla_\mu u^\mu + \beta_0 \dot{\tau} - \alpha_0 \nabla_\mu q^\mu - \gamma_0 T q^\mu \nabla_\mu \left(\frac{\alpha_0}{T} \right) + \frac{\tau T}{2} \nabla_\mu \left(\frac{\beta_0 u^\mu}{T} \right) \right], \quad (301)$$

$$\begin{aligned} q^\mu = & -\kappa T \perp^{\mu\nu} \left[\frac{1}{T} \nabla_\nu T + \dot{u}_\nu + \beta_1 \dot{q}_\nu - \alpha_0 \nabla_\nu \tau - \alpha_1 \nabla_\alpha \tau^\alpha{}_\nu + \frac{T}{2} q_\nu \nabla_\alpha \left(\frac{\beta_1 u^\alpha}{T} \right) \right. \\ & \quad \left. - (1 - \gamma_0) \tau T \nabla_\nu \left(\frac{\alpha_0}{T} \right) - (1 - \gamma_1) T \tau^\alpha{}_\nu \nabla_\alpha \left(\frac{\alpha_1}{T} \right) + \gamma_2 \nabla_{[\nu} u_{\alpha]} q^\alpha \right], \end{aligned} \quad (302)$$

$$\tau_{\mu\nu} = -2\eta \left[\beta_2 \dot{\tau}_{\mu\nu} + \frac{T}{2} \tau_{\mu\nu} \nabla_\alpha \left(\frac{\beta_2 u^\alpha}{T} \right) + \left\langle \nabla_\mu u_\nu - \alpha_1 \nabla_\mu q_\nu - \gamma_1 T q_\mu \nabla_\nu \left(\frac{\alpha_1}{T} \right) + \gamma_3 \nabla_{[\mu} u_{\alpha]} \tau^\alpha{}_\nu \right\rangle \right], \quad (303)$$

where the angular brackets denote symmetrization as before. In these expression we have added yet another two terms, representing the coupling to vorticity. These bring further “free” parameters γ_2 and γ_3 . It is easy to see that we are allowed to add these terms since they do not affect the entropy production. In fact, a large number of similar terms may, in principle, be considered (see note added in proof in [53[⊕]]). The presence of coupling terms of the particular form that we have introduced is suggested by kinetic

Complete Set of Terms

- Daunting:

$$\tau_\Pi \dot{\Pi} + \Pi = \Pi_{\text{NS}}$$

$$\tau_q \Delta^{\mu\nu} \dot{q}_\nu + q^\mu = q_{\text{NS}}^\mu$$

ν

$$+ \hat{\delta}_{1,2} \Pi$$

$$\tau_\pi \dot{\pi}^{<\mu\nu>} + \pi^<$$

$$\begin{aligned}
 & - 2 \tau_\pi \pi_\lambda^{<\mu} \sigma^{\nu>\lambda} - 2 \lambda_{\pi q} q^{<\mu} \nabla^{\nu>} \alpha + 2 \lambda_{\pi\Pi} \Pi \sigma^{\mu\nu} \\
 & + \hat{\delta}_{2,2} \Pi \pi^{\mu\nu} - \hat{\eta}_2 \pi_\lambda^{<\mu} \pi^{\nu>\lambda} - \hat{\epsilon}_2 q^{<\mu} q^{\nu>}
 \end{aligned}$$

- And still subject to

- Poorly constrained initial Conditions
- Eccentricity fluctuations
- Poorly constrained equation of state
- Hadronic rescattering effects
- Bulk viscosity
- Numerical viscosity
- Finite size, core/corona effects



Complete Set of Terms

- Daunting:

$$\begin{aligned}
 \tau_\Pi \dot{\Pi} + \Pi &= \Pi_{\text{NS}} + \tau_{\Pi q} q \cdot \dot{u} - \ell_{\Pi q} \partial \cdot q - \zeta \hat{\delta}_{0,1} \Pi \theta \\
 &\quad + \lambda_{\Pi q} q \cdot \nabla \alpha + \lambda_{\Pi \pi} \pi^{\mu\nu} \sigma_{\mu\nu} + \hat{\delta}_{0,2} \Pi^2 + \hat{\epsilon}_0 q \cdot q + \hat{\eta}_0 \pi^{\mu\nu} \pi_{\mu\nu} \\
 \tau_q \Delta^{\mu\nu} \dot{q}_\nu + q^\mu &= q_{\text{NS}}^\mu - \tau_{q\Pi} \Pi \dot{u}^\mu - \tau_{q\pi} \pi^{\mu\nu} \dot{u}_\nu \\
 &\quad + \ell_{q\Pi} \nabla^\mu \Pi - \ell_{q\pi} \Delta^{\mu\nu} \partial^\lambda \pi_{\nu\lambda} + \tau_q \omega^{\mu\nu} q_\nu - \frac{\kappa}{\beta} \hat{\delta}_{1,1} q^\mu \theta \\
 &\quad - \lambda_{qq} \sigma^{\mu\nu} q_\nu + \lambda_{q\Pi} \Pi \nabla^\mu \alpha + \lambda_{q\pi} \pi^{\mu\nu} \nabla_\nu \alpha \\
 &\quad + \hat{\delta}_{1,2} \Pi q^\mu + \hat{\eta}_1 \pi^{\mu\nu} q_\nu \\
 \tau_\pi \dot{\pi}^{<\mu\nu>} + \pi^{\mu\nu} &= \pi_{\text{NS}}^{\mu\nu} + 2 \tau_{\pi q} q^{<\mu} \dot{u}^{\nu>} \\
 &\quad + 2 \ell_{\pi q} \nabla^{<\mu} q^{\nu>} + 2 \tau_\pi \pi_\lambda^{<\mu} \omega^{\nu>\lambda} - 2 \eta \hat{\delta}_{2,1} \pi^{\mu\nu} \theta \\
 &\quad - 2 \tau_\pi \pi_\lambda^{<\mu} \sigma^{\nu>\lambda} - 2 \lambda_{\pi q} q^{<\mu} \nabla^{\nu>} \alpha + 2 \lambda_{\pi\Pi} \Pi \sigma^{\mu\nu} \\
 &\quad + \hat{\delta}_{2,2} \Pi \pi^{\mu\nu} - \hat{\eta}_2 \pi_\lambda^{<\mu} \pi^{\nu>\lambda} - \hat{\epsilon}_2 q^{<\mu} q^{\nu>}
 \end{aligned}$$

- And still subject to

- Poorly constrained initial Conditions
- Eccentricity fluctuations
- Poorly constrained equation of state
- Hadronic rescattering effects
- Bulk viscosity
- Numerical viscosity
- Finite size, core/corona effects



For Further Details

- See “Virtual Journal on QCD Matter”

- Steffen A. Bass,
Berndt Mueller,
William A. Zajc
- qgp.phy.duke.edu

- On the topic of
2nd order hydro:

- [What a Difference
a Term Makes](http://qgp.phy.duke.edu/2008/03/30/what-a-difference-a-term-makes)

Virtual Journal on QCD Matter » What a Difference a Term Makes - Windows Internet Explorer

http://qgp.phy.duke.edu/2008/03/30/what-a-difference-a-term-makes/

Virtual Journal on QCD Matter

devoted to the Physics of the QGP & Relativistic Heavy-Ion Collisions
- moderated by Steffen A. Bass, Berndt Mueller and William A. Zajc -

What a Difference a Term Makes

Posted by William A. Zajc on March 30th 2008

Categories: AdS/CFT & String Theory, ultra-cold Fermi gas, Collective Flow, Modeling

The recent post on the AMO competition for perfect fluidity offers a wonderful opportunity to compare and contrast the techniques of condensed matter physics versus those of relativistic heavy ion physics. The experimental techniques, while vastly different, each push the state-of-the-art in their respective disciplines. But determining the all important viscosity to entropy density ratio η/s could not be more different between the two fields: in the case of the trapped atoms, an elegant yet straightforward perturbative analysis makes a direct connection between the damping rate of breathing modes and η/s . Nothing could be further from the truth in nuclear collisions: the development of causal, viscous hydrodynamics for truly relativistic systems remains very much a work in progress.

These issues are not apparent from consultation with the classic references. It's hard to do better than the exposition found in Weinberg's *Gravitation and Cosmology*, published in 1972. Weinberg begins with a perfect fluid, defined as one that appears isotropic to a co-moving observer. The equations of motion then follow directly from conservation of energy and momentum, the continuity equation for any conserved currents and the first law of thermodynamics, along with an equation of state to close the system. This is the so-called zero-th order

Search this site: Go

PAGES

- » About
- » Moderators

ARCHIVES

- » April 2008
- » March 2008
- » February 2008
- » January 2008
- » December 2007
- » November 2007
- » October 2007
- » September 2007
- » August 2007

CATEGORIES

- » AdS/CFT & String Theory (5)
- » Anomalous Viscosity (2)
- » Astrophysics (2)

Internet 100% ::

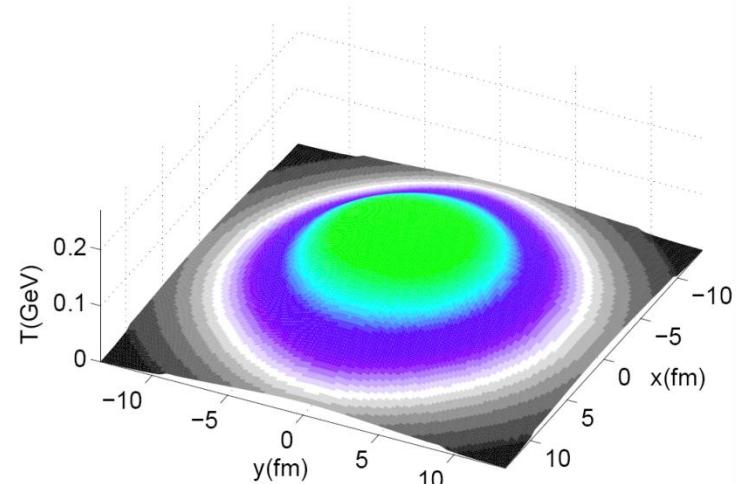


Implementing and Testing

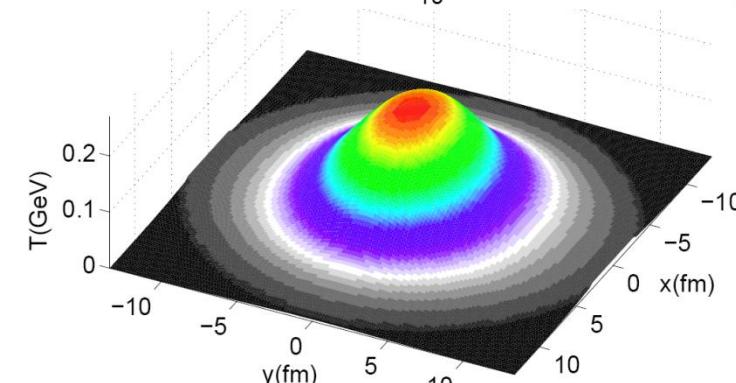
- Examples

- P. Romatschke and U. Romatschke, Phys. Rev. Lett. 99:172301, 2007
- H. Song and U. Heinz, Phys. Rev. C78, 024902, 2008
- M. Luzum and P. Romatschke, Phys. Rev. C78:034915, 2008.

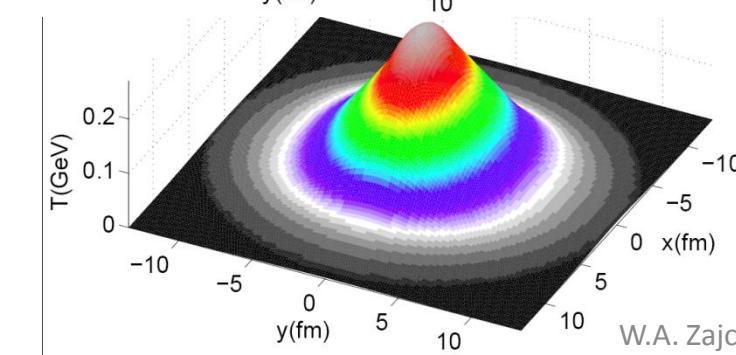
7 fm/c



3 fm/c



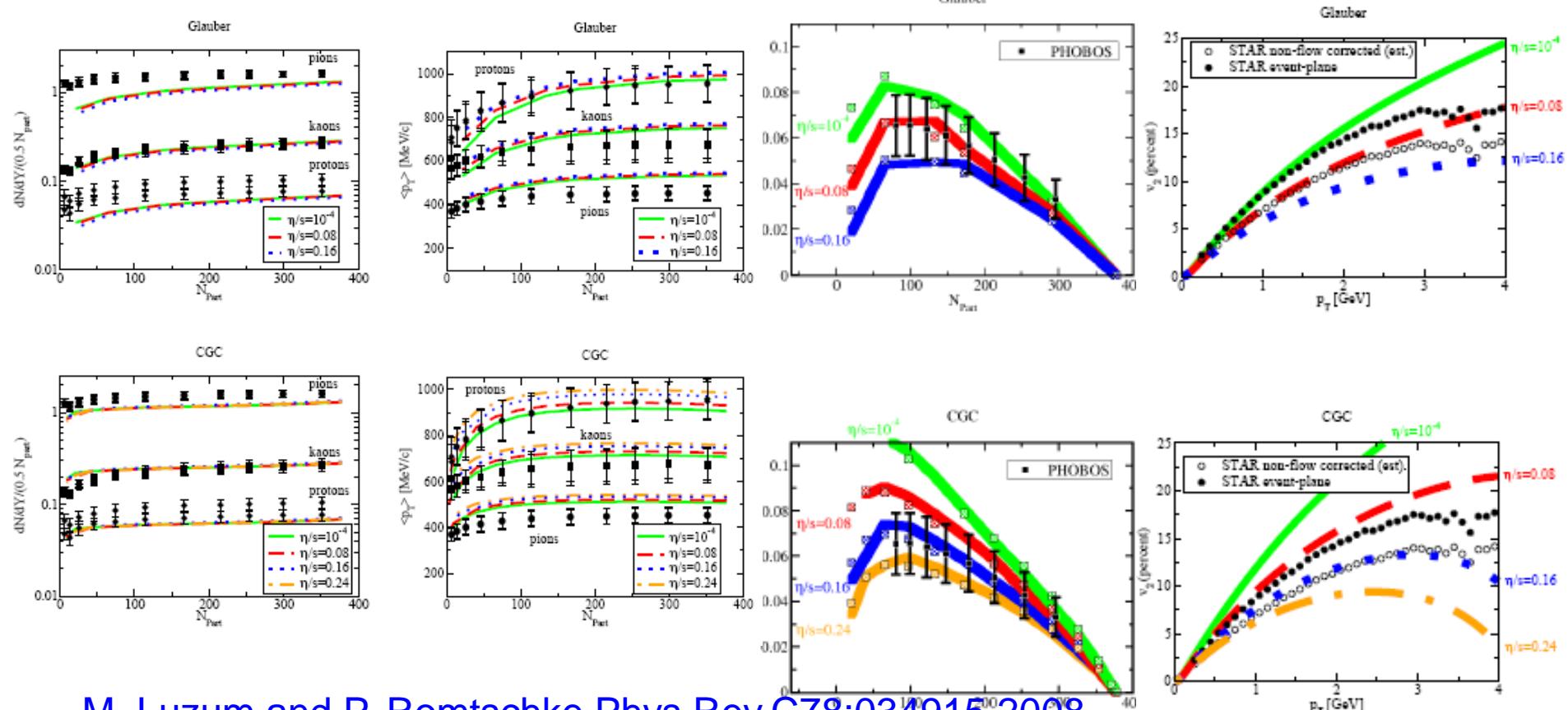
1 fm/c



Concordance

- BNL, April 2008:

- Workshop on Viscous Hydrodynamics and Transport Models in Heavy Ion Collisions
- Workshop Summary



M. Luzum and P. Romtschke Phys.Rev.C78:034915,2008.



Concordance → Quantifying η/s

Q : How?

A : By performing detailed and systematic hydrodynamic simulations to understand sensitivity to η/s (and “everything” else)

- *Viscosity Information from Relativistic Nuclear Collisions: How Perfect is the Fluid Observed at RHIC?,
P. Romatschke and U. Romatschke,
Phys. Rev. Lett. 99:172301, 2007*

$$\frac{\eta}{s} = (0 - 2) \frac{1}{4\pi}$$

- *Multiplicity Scaling in Ideal and Viscous Hydrodynamics,
H. Song and U. Heinz,
Phys. Rev. C78, 024902, 2008*

$$\frac{\eta}{s} = (1 - 2) \frac{1}{4\pi}$$

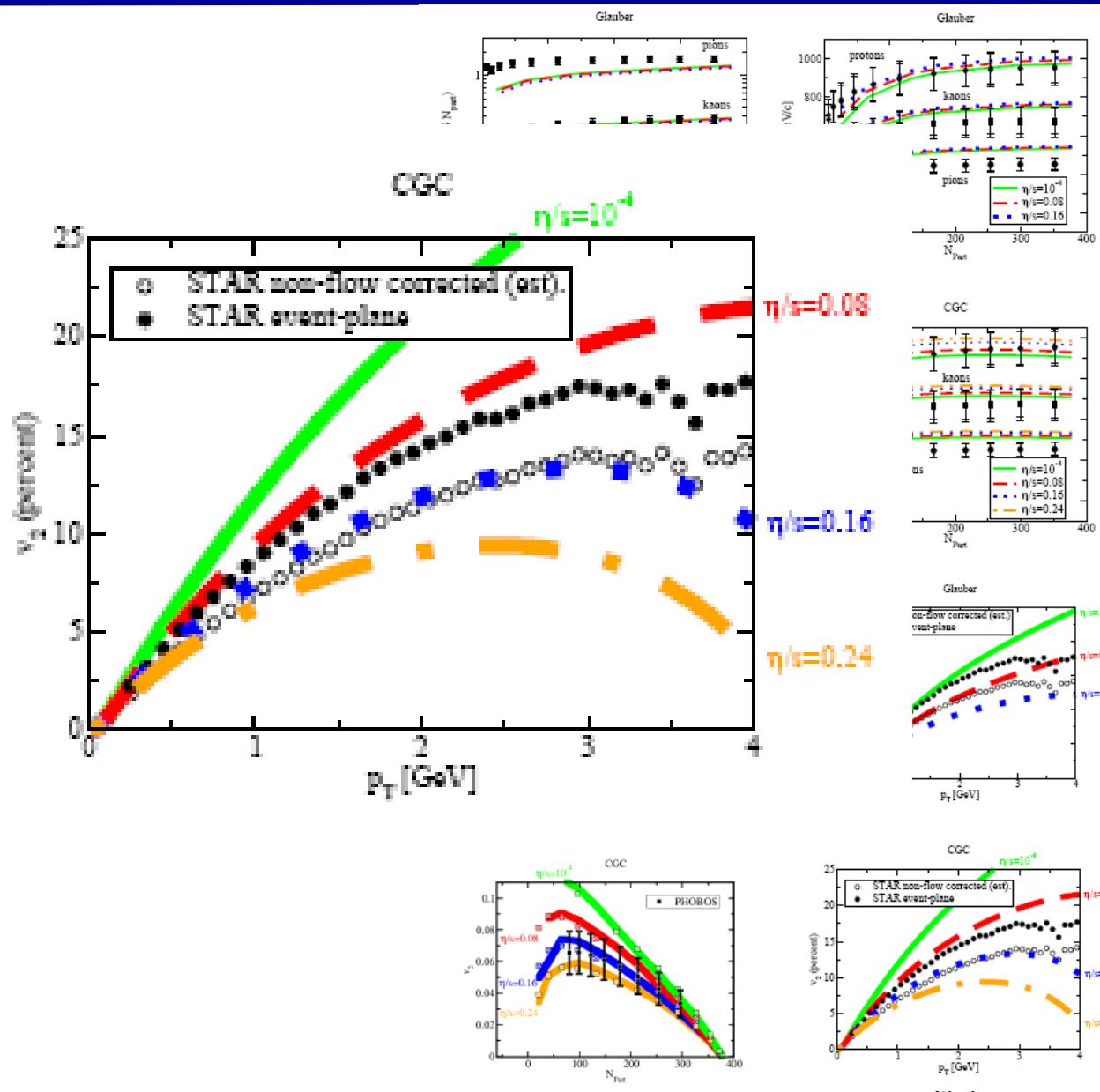
- *Conformal Relativistic Viscous Hydrodynamics:
Applications to RHIC results at $\sqrt{s_{NN}} = 200$ GeV,
M. Luzum and P. Romatschke,
Phys. Rev. C78:034915, 2008.*

$$\frac{\eta}{s} = (1.3 \pm 1.3 \pm 1) \frac{1}{4\pi}$$



Viscosity Information from Relativistic Nuclear Collisions: How Perfect is the Fluid Observed at RHIC?, P. Romatschke and U. Romatschke, Phys. Rev. Lett. 99:172301, 2007

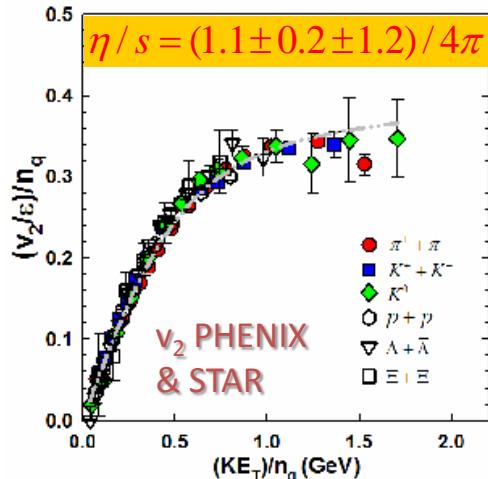
- Signatures:
 dN/dy , v_2 , $\langle p_T \rangle$
- Calculation:
2nd order
causal viscous
conformal
hydro:
- (Glauber and CGC IC's)
- Payoff Plots:



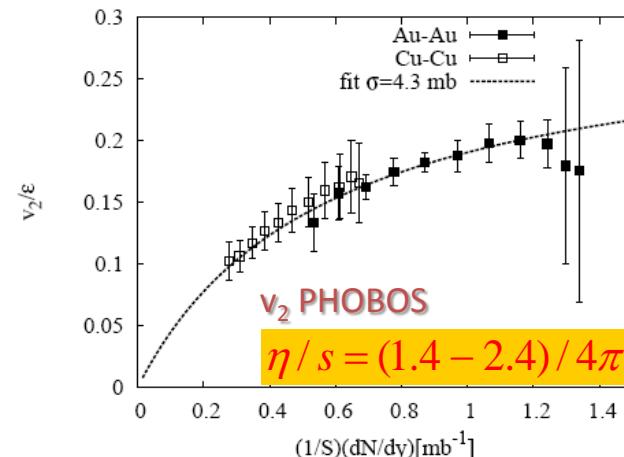
Comparison of Estimates



R. Lacey et al.: PRL 98:092301, 2007

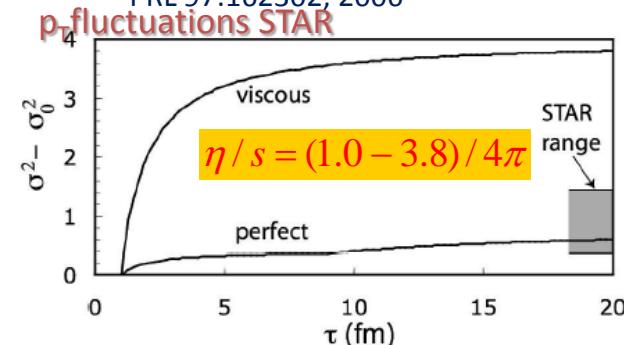


H.-J. Drescher et al.: arXiv:0704.3553



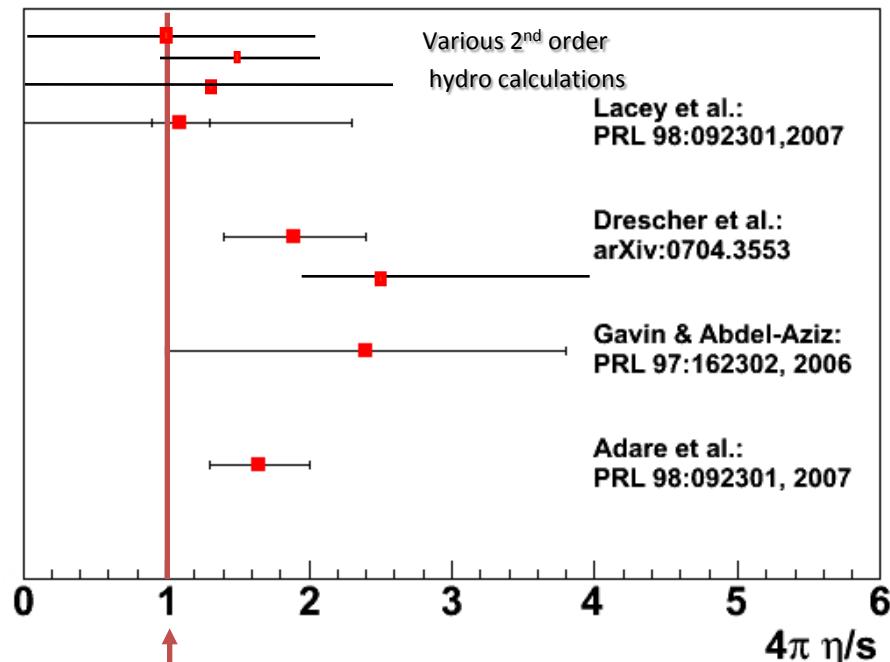
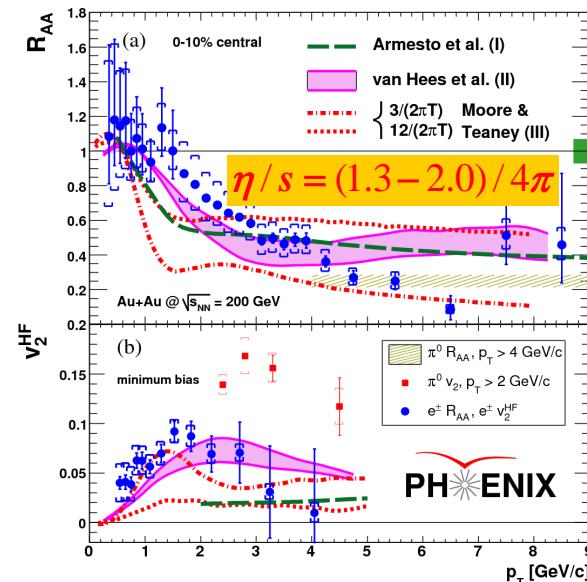
S. Gavin and M. Abdel-Aziz:

PRL 97:162302, 2006



A. Adare et al., PRL 98:172301, 2007

Heavy flavor drag, flow;
PHENIX



conjectured quantum limit



Numerical Value of RHIC Viscosity

- **Input :** $\eta_{QGP} \sim \frac{\hbar}{4\pi} s_{QGP}$ $T_{QGP} \sim 200 \text{ MeV}$ $\hbar c = 200 \text{ Mev} \cdot \text{fm}$
- **Estimating s_{QGP} (per degree of freedom):**
 - **Method 1:** $s_{QGP} = (\varepsilon + p)/T_{QGP} \sim \frac{4}{3} \frac{\varepsilon}{T_{QGP}} = \frac{4}{3T_{QGP}} \left(\frac{\pi^2}{30} T_{QGP}^4 \right) = \frac{2\pi^2}{45} T_{QGP}^3$
 - **Method 2 (based on handy rule of thumb):**
 $n \sim \left(\frac{T}{2} \right)^3 \quad s \sim 3.6n \Rightarrow s \sim \frac{3.6}{2^3} T^3 = 1.03 \left(\frac{2\pi^2}{45} T^3 \right)$
- **Number of (assumed massless) degrees of freedom:**

$$n_{d.o.f.} \sim \left\{ 2_s \cdot 8_g + \frac{7}{8} \cdot 2_s \cdot 2_{a.p.} \cdot 3_c \cdot 3_f \right\} = 47.5 \sim 45$$

$$\Rightarrow \eta_{QGP} \sim \frac{\hbar}{4\pi} s_{QGP} \sim \frac{\hbar}{4\pi} n_{d.o.f.} \left(\frac{2\pi^2}{45} T^3 QGP \right) \sim \frac{\pi}{2} \frac{\hbar c}{c} T^3 QGP$$

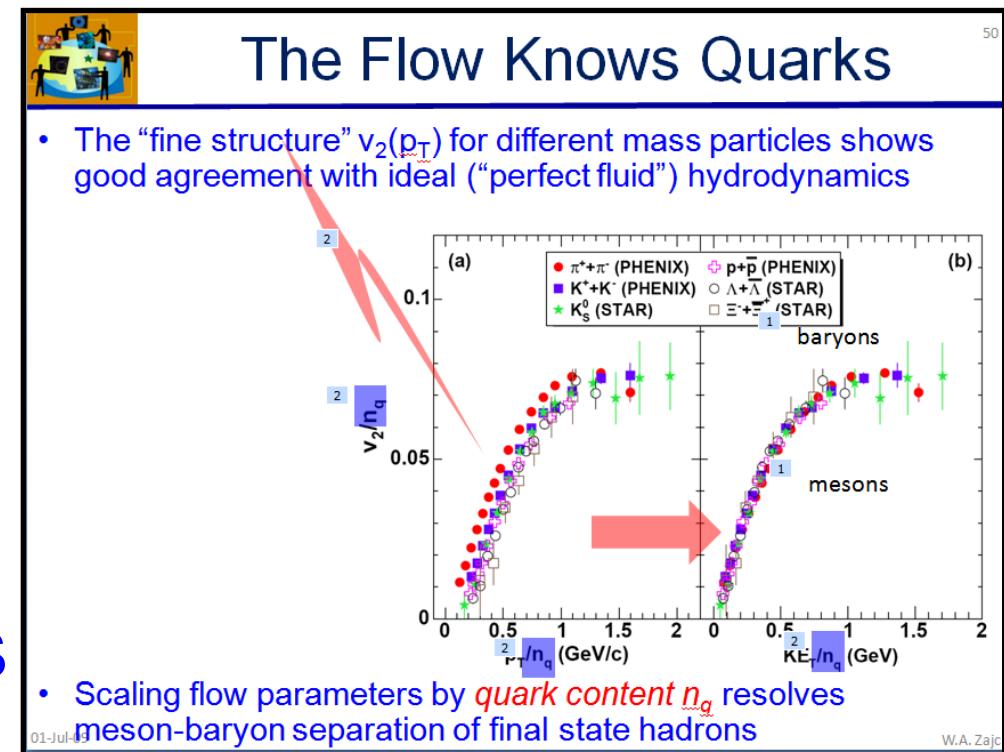
$$= \frac{\pi \cdot (2 \cdot 10^8 \text{ eV} \cdot 10^{-15} \text{ m}) \times (1.6 \cdot 10^{-19} \text{ J/eV})}{2(3 \cdot 10^8 \text{ m/s})} (10^{15} \text{ m})^3 = 1.6 \times 10^{11} \frac{\text{J} \cdot \text{s}}{\text{m}^3}$$

A Deep Puzzle

- Apparent role of quark degrees of freedom:

Incompatible (! ?)
with η/s near the
quantum bound :

- ‘Good’ quasi-particles have widths \ll mass
- Minimal mfp’s \Rightarrow widths \sim mass





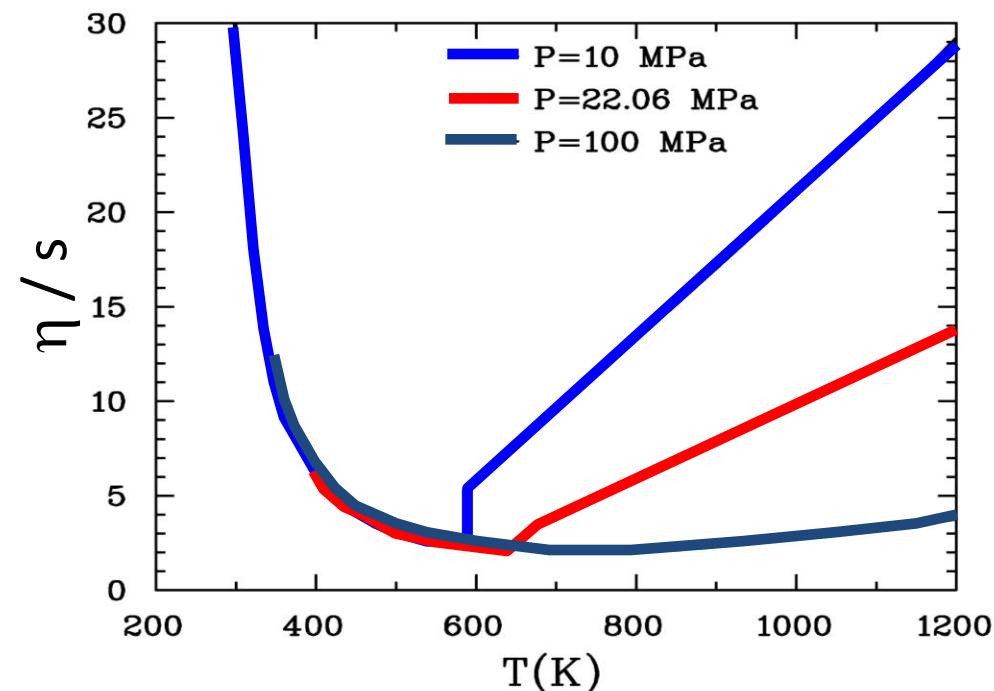
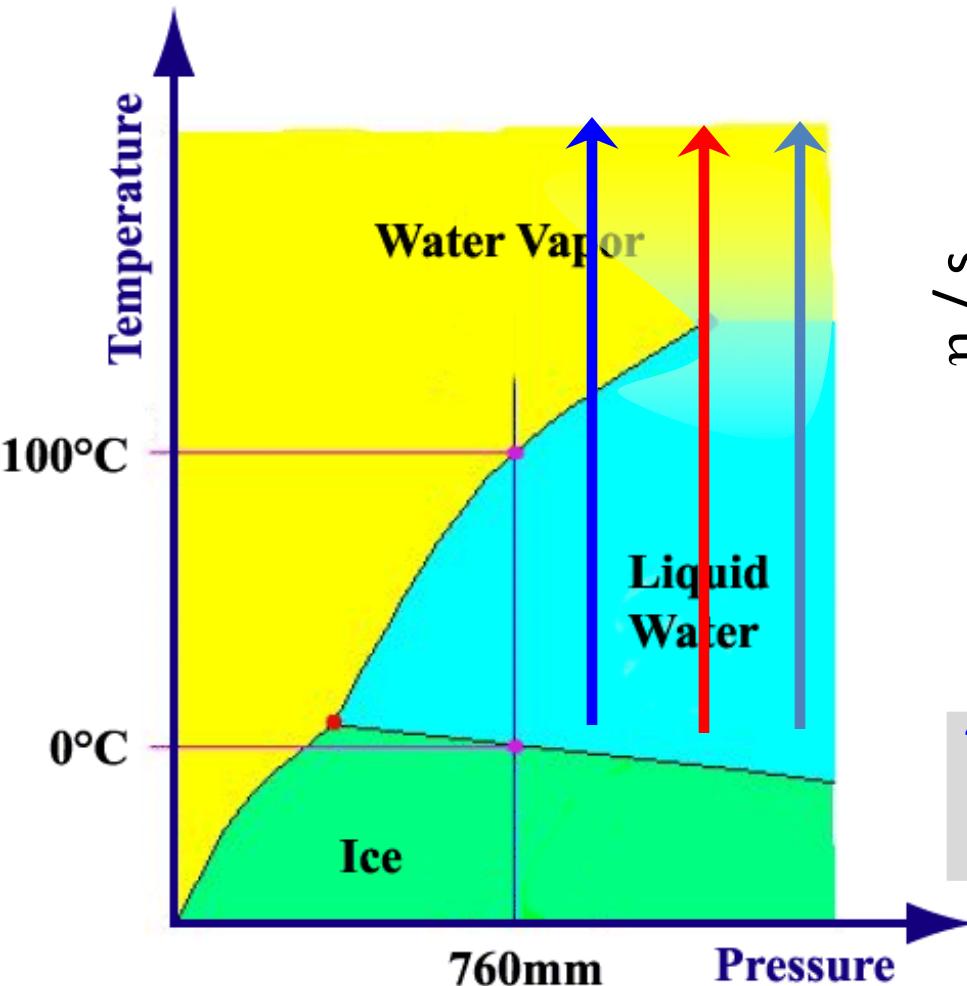
Implications for D.O.F

- While tempting to identify the coalescence patterns with “underlying quark degrees of freedom”...
- Much work still needed to reconcile with ‘absence’ of quasiparticles when η/s near quantum bound
 - *Quasi-Particle Degrees of Freedom versus the Perfect Fluid as Descriptors of the Quark-Gluon Plasma*, L.A. Levy *et al.*, Phys.Rev.C78:044905,2008. [0709.3105](#)
 - *Quantum Criticality and Black Holes*, S. Sachdev and M. Mueller, [0810.3005](#) :
→“The theory of the quantum critical region shows that the transport coefficients, and the relaxation time to local equilibrium, *are not proportional to a mean free scattering time between excitations*, as is the case in the Boltzmann theory of quasiparticles

Water → RHIC → Water → RHIC



- The search for QCD phase transition of course was informed by analogy to ordinary matter
- Results from RHIC are now “flowing” back to ordinary matter



“On the Strongly-Interacting Low-Viscosity Matter Created in Relativistic Nuclear Collisions”,
L.P. Csernai, J.I. Kapusta and L.D. McLerran,
Phys.Rev.Lett.97:152303,2006, [nucl-th/0604032](https://arxiv.org/abs/hep-th/0604032)



Is There a QCD Critical Point?

- Here the analogy with phase transitions in ordinary matter breaks down:

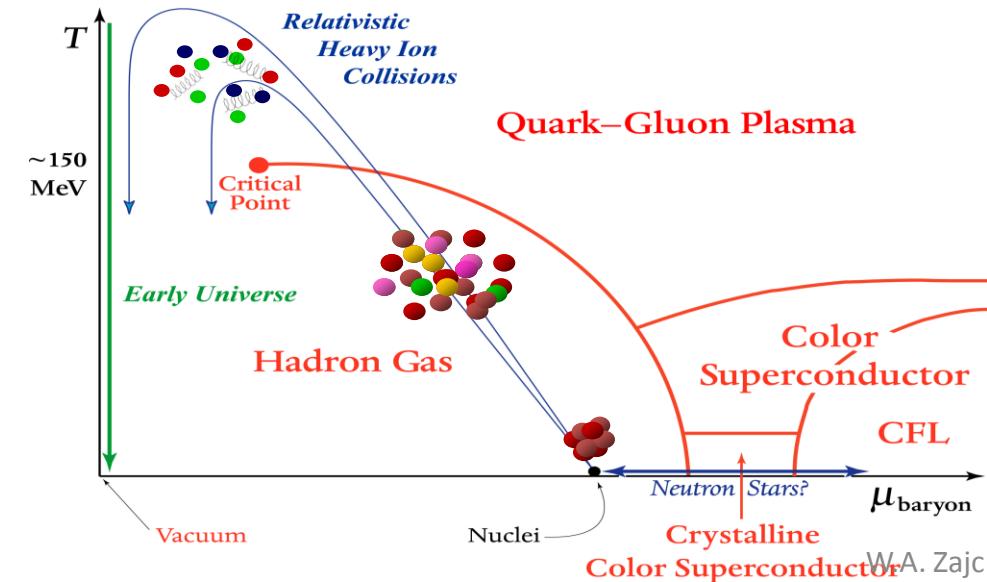
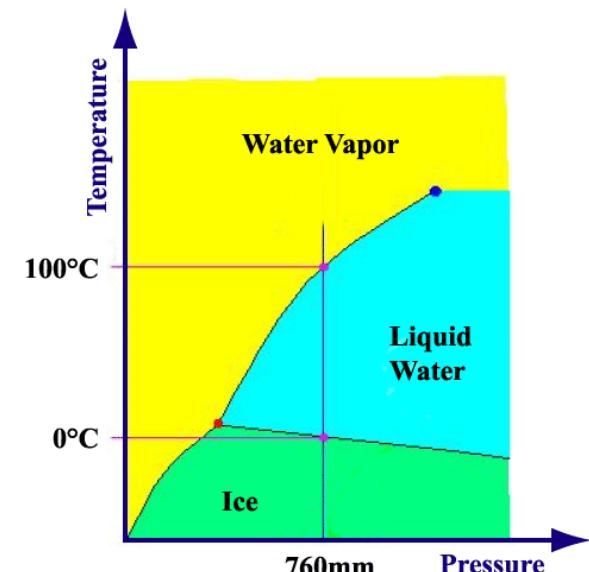
- Recall “ Properties of the medium are (at zero baryon number) uniquely determined by T ”

\Rightarrow Pressure = $P(T)$ can't vary independently (unlike water)

- But if baryon number is non-zero
 \Rightarrow (intensive order parameter) baryon chemical potential μ_B :

- To increase μ_B :

- Lower collision energy
 - Raise atomic mass
 - Both part of RHIC II and GSI-FAIR



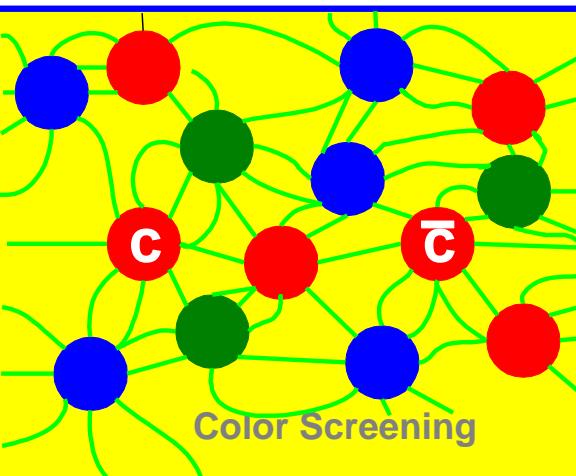


A Spooky Connection

- RHIC physics clearly relies on
 - The quantum nature of matter (Einstein, 1905)
 - The relativistic nature of matter (Einstein, 1905)
- but presumably has no connection to
 - General relativity (Einstein, 1912-7)
- Wait ! Both sides of this equation

$$(Viscosity)_{RHIC} \approx \frac{\hbar}{4\pi} (Entropy Density)_{RHIC}$$

were calculated using black hole physics (in 10 dimensions)



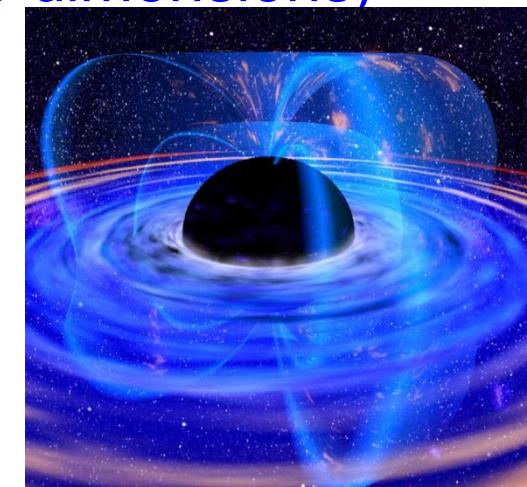
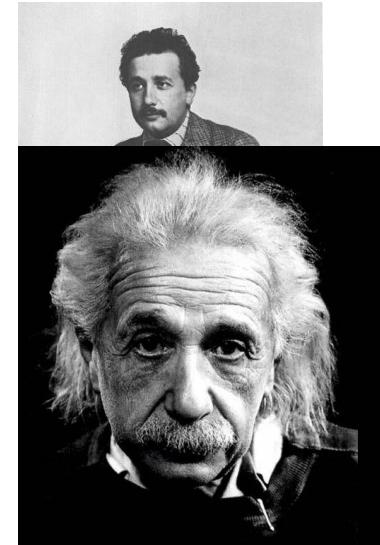
MULTIPLICITY

Entropy \leftrightarrow Black Hole Area

DISSIPATION

Viscosity \leftrightarrow Graviton

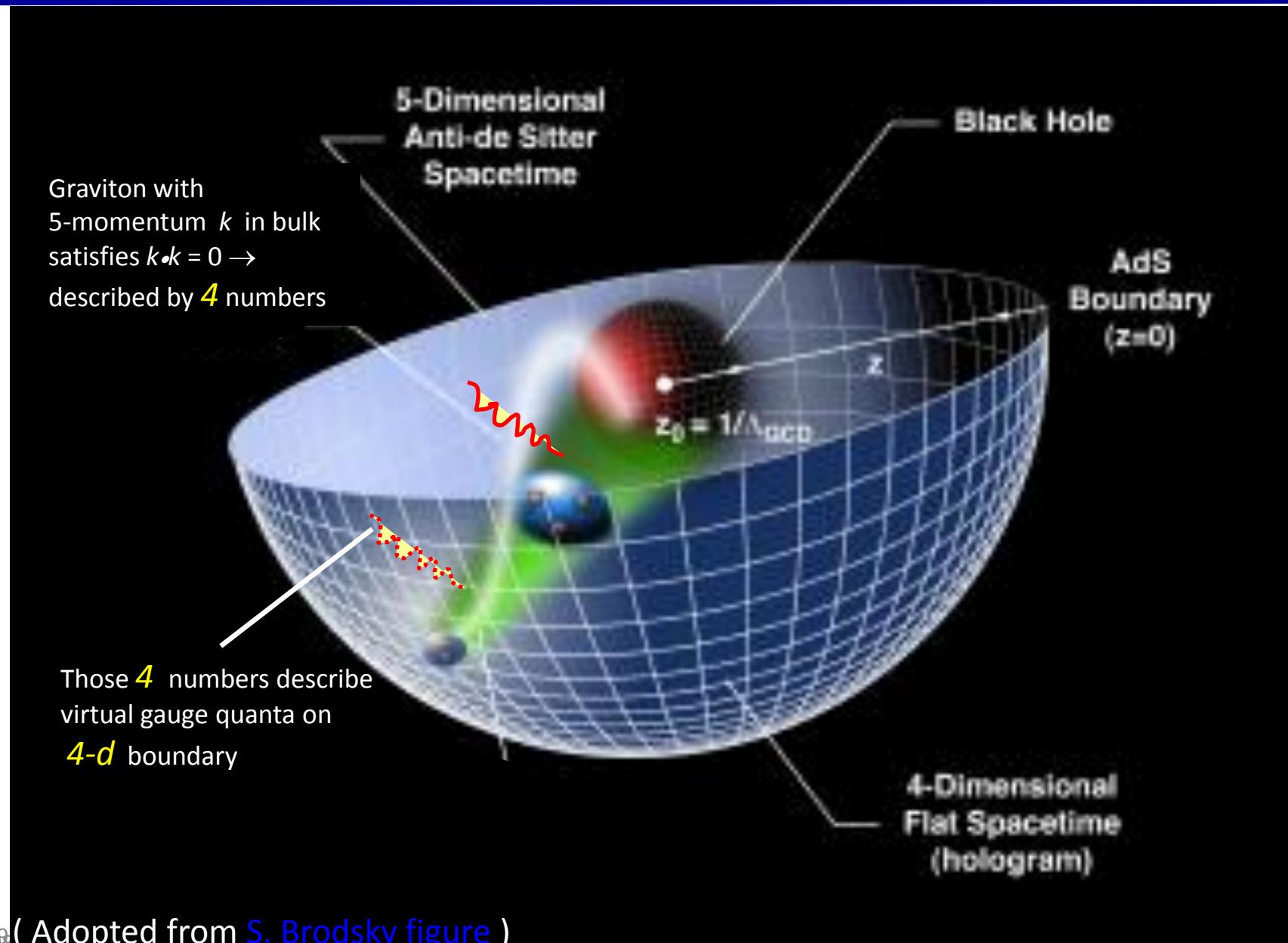
Absorption





(Minimal) AdS / CFT

$$\Rightarrow \frac{\eta}{s} \geq \frac{1}{4\pi}$$





In Words

- A stringy theory of gravity in N space-time dimensions
(the “bulk” AdS)
 - is “dual” (that is, equivalent to)
- A gauge theory without gravity in N-1 space-time dimensions
(the “boundary” CFT)
- Notes:
 - AdS \equiv Anti de Sitter space ; CFT \equiv Conformal Field Theory
 - “Equivalent” means “equivalent” – all phenomena in one theory have corresponding “dual” descriptions in the other theory.
 - Maldacena’s AdS/CFT correspondence is a realization of hypotheses from both ‘t Hooft and Susskind that the ultimate limit on the number of degrees of freedom in a spacetime region is proportional to the area of its boundary, *not* its volume (!)



Why Does This Work??

- The easy part:

- Recall

$$\frac{F_x}{A} = -\eta \frac{\partial v_x}{\partial y}$$

- that is, viscosity $\sim x$ -momentum transport in y -direction $\sim T^{xy}$
 - There are standard methods (Kubo relations) to calculate such dissipative quantities

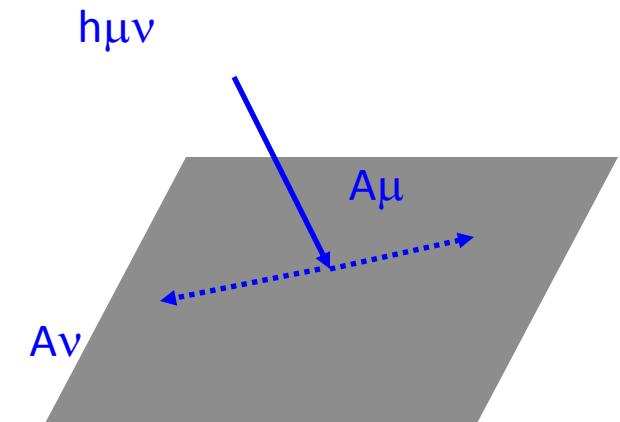
$$\eta = \lim_{\omega \rightarrow 0} \frac{1}{2\omega} \int dt d\mathbf{x} e^{i\omega t} \langle [T_{xy}(t, \mathbf{x}), T_{xy}(0, 0)] \rangle$$

- The hard part:

- This calculation is difficult in a strongly-coupled gauge theory

- The weird part:

- A (supersymmetric) pseudo-QCD theory can be mapped to a 10-dimensional classical gravity theory on the background of black 3-branes
 - The calculation can be performed there as the absorption of gravitons by the brane
 - THE SHEAR VISCOSITY OF STRONGLY COUPLED N=4 SUPERSYMMETRIC YANG-MILLS PLASMA., G. Policastro, D.T. Son , A.O. Starinets, Phys.Rev.Lett.87:081601,2001 hep-th/0104066





The Result

- Viscosity $\eta = \text{"Area"}/16\pi G$

Infinite "Area" !

- Normalize by entropy (density) $S = \text{"Area"}/4G$

- Dividing out the infinite "areas" :

$$\frac{\eta}{S} \geq \frac{\hbar}{4\pi} \frac{1}{k_B}$$

- *Conjectured* to be a lower bound “for all relativistic quantum field theories at finite temperature and zero chemical potential”.
- See “Viscosity in strongly interacting quantum field theories from black hole physics”, P. Kovtun, D.T. Son, A.O. Starinets, Phys.Rev.Lett.94:111601, 2005, [hep-th/0405231](https://arxiv.org/abs/hep-th/0405231)



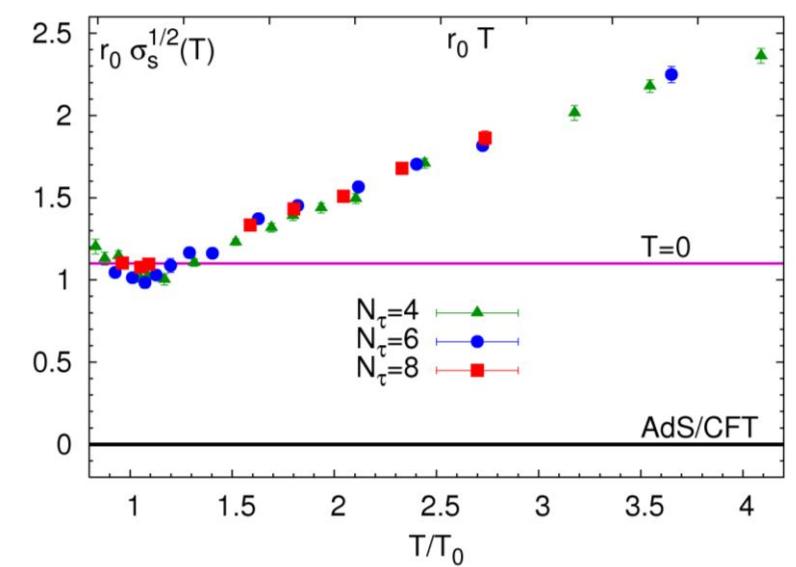
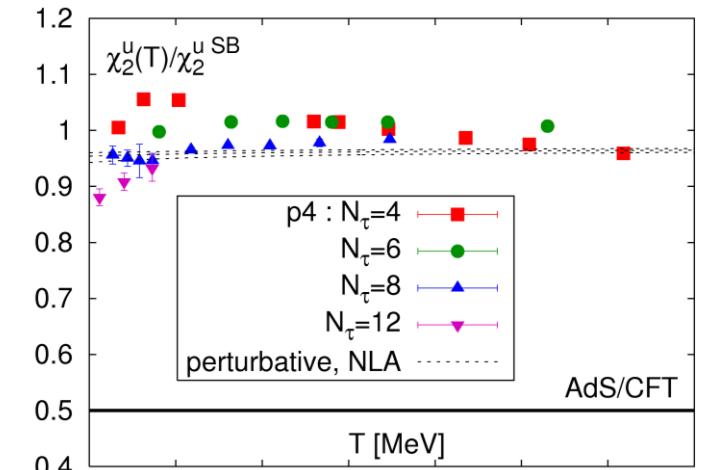
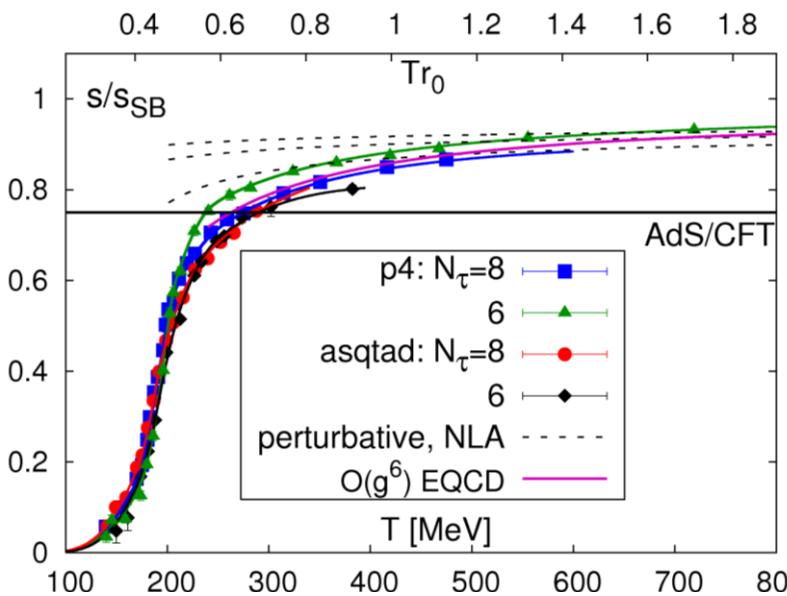
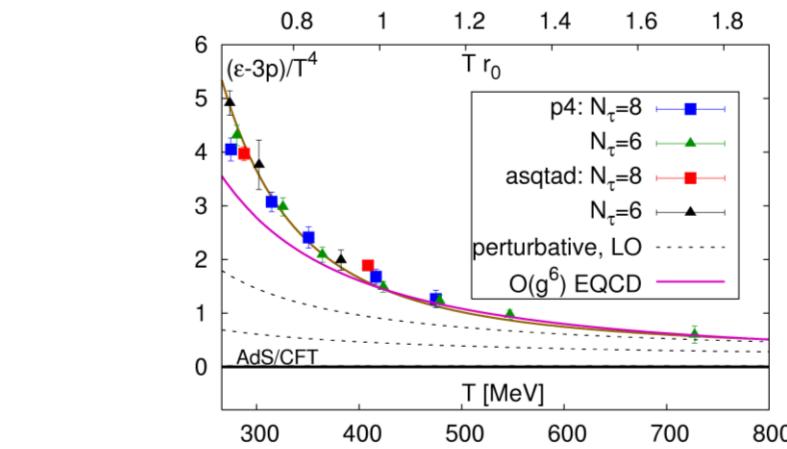
AdS/CFT - Pro

- Gauge/gravity duality, G.T. Horowitz and J. Polchinski, [gr-qc/0602037](#))
 - “Hidden within every non-Abelian gauge theory, even within the weak and strong nuclear interactions, is a theory of quantum gravity.”
- Stringscape, by Matthew Chalmers, in *Particle World*:
 - “Susskind says that by studying heavy-ion collisions you are also studying quantum gravity that is ‘blown up and slowed down by a factor of 10^{20} ’ ”
- The Black Hole War, L. Susskind, [ISBN 978-0-316-01640-7](#) :
 - “...the Holographic Principle is evolving from radical paradigm shift to everyday working tool of – surprisingly – nuclear physics.”

AdS/CFT - Con



- P. Petreczsky, QM09: “AdS/CFT is consistently wrong.”





Editorial Opinion

- Reality may lie between these two extremes ☺
- At the moment:
 - Hard physics (high p_T energy loss)
 - ◆ Fragile predictions
 - ◆ Robust extractions
 - Soft physics (hydrodynamics, viscosity)
 - ◆ Robust predictions
 - ◆ Fragile extractions
- AdS/CFT has led to qualitatively new insights
- Not covered:
 - AdS/QCD (application to confinement, masses)



Why AdS/CFT Matters...

- All the thermal parts are built upon Bekenstein and Hawking's (unproven) assertion that black holes have entropy:

$$S_{BH} = \frac{A}{4G} = \frac{A}{4L_P^2}$$

- \Rightarrow Black holes have a temperature
- \Rightarrow Black holes can radiate
- \Rightarrow Black holes don't lose information
- Important to test these very underpinnings



Summary- Lecture 4

- A minimal η/s is predicted for a dense thermal fluid in the quantum regime.
- Major advances in relativistic viscous hydrodynamics are helping to bound the observed η/s value at RHIC.
- There is a striking connection between this physics and similar phenomena predicted via the AdS/CFT correspondence.



Summary

- There is qualitative and quantitative theoretical support for a deconfining phase transition in QCD.
- The matter produced at RHIC is indeed matter, imbued with unusual properties beyond those predicted for ‘QGP’.
- Most compelling of these is the fluid behavior of the densest matter ever studied.
- There are more discoveries to come as we continue to explore thermal QCD matter.