

# Constraining the NS EOS with Multi-Messenger Observations of GW170817

[or - yet another extraordinary gift from the riches of GW170817]

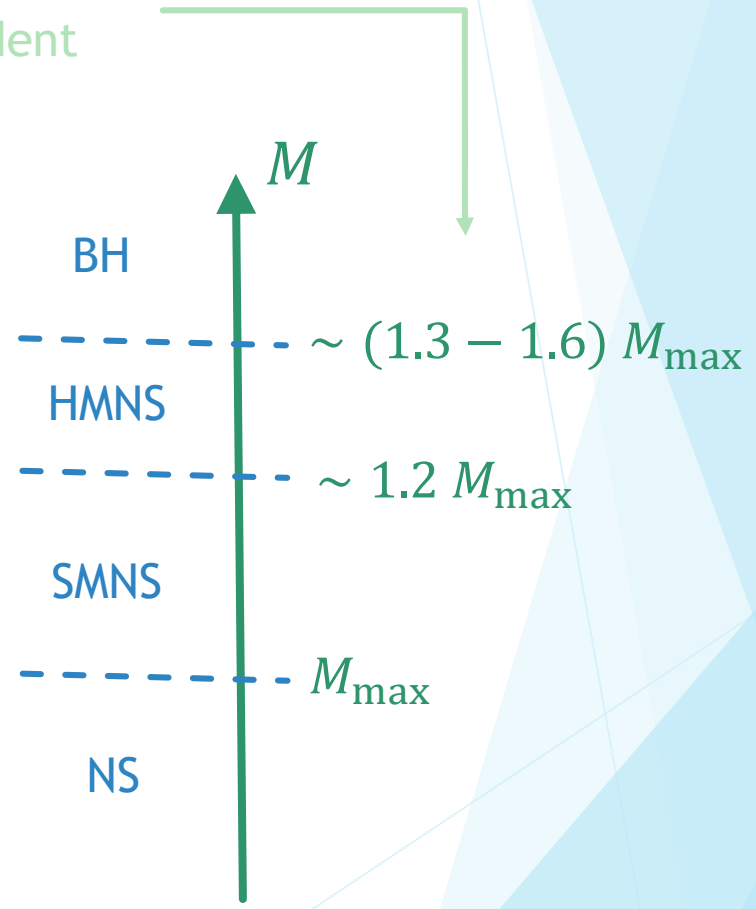
Ben Margalit,  
Columbia University

# GW170817 & the NS EOS (constraining $M_{\max}$ )

## Merger Outcome:

- merger can produce:
  - **stable** NS
  - **rigidly rotating** rotationally-supported *supramassive* NS (**SMNS**)
  - **differentially rotating** rotationally-supported *hyper-massive* NS (**HMNS**)
  - black-hole (**BH**) [“prompt collapse”]
- qualitatively different EM counterparts!  
(e.g. Bauswein+13; Metzger&Fernandez14)

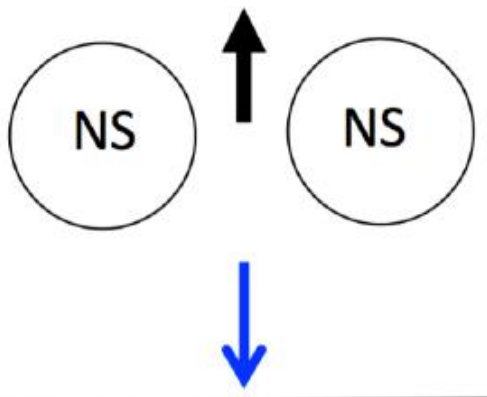
threshold masses  
EOS dependent



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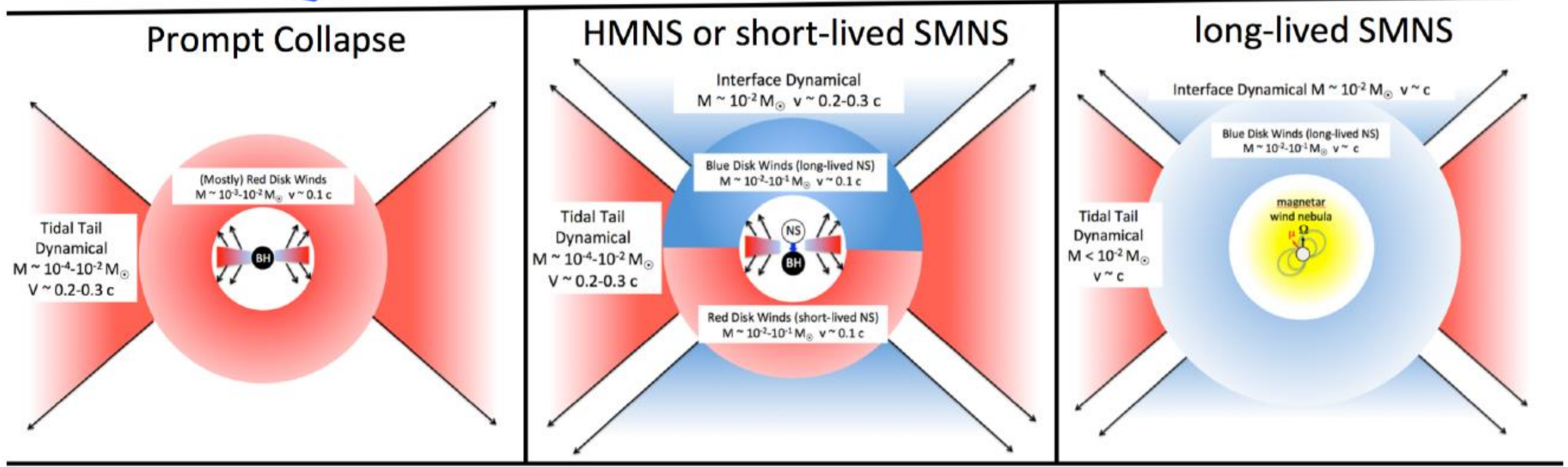
- EM counterparts



(BM & Metzger 2017)

$M_t > (1.3-1.6)M_{max}$

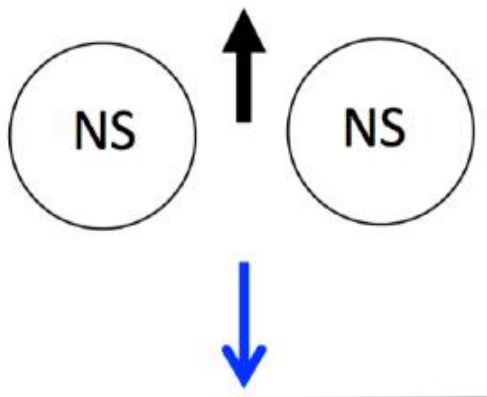
$M_t < 1.2M_{max}$



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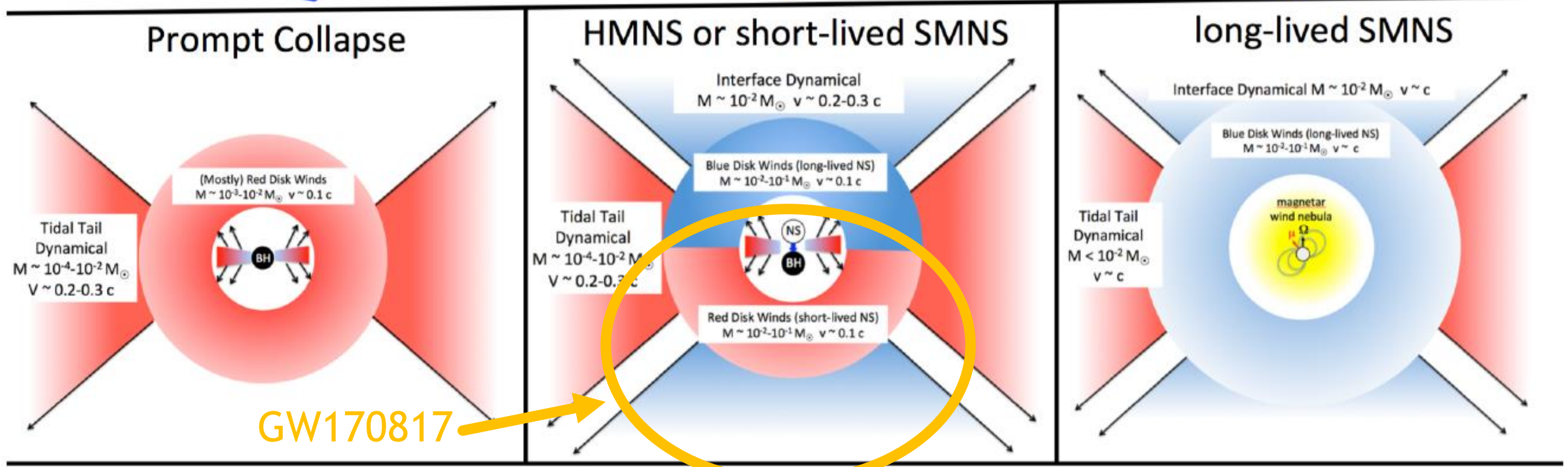
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(BM & Metzger 2017)

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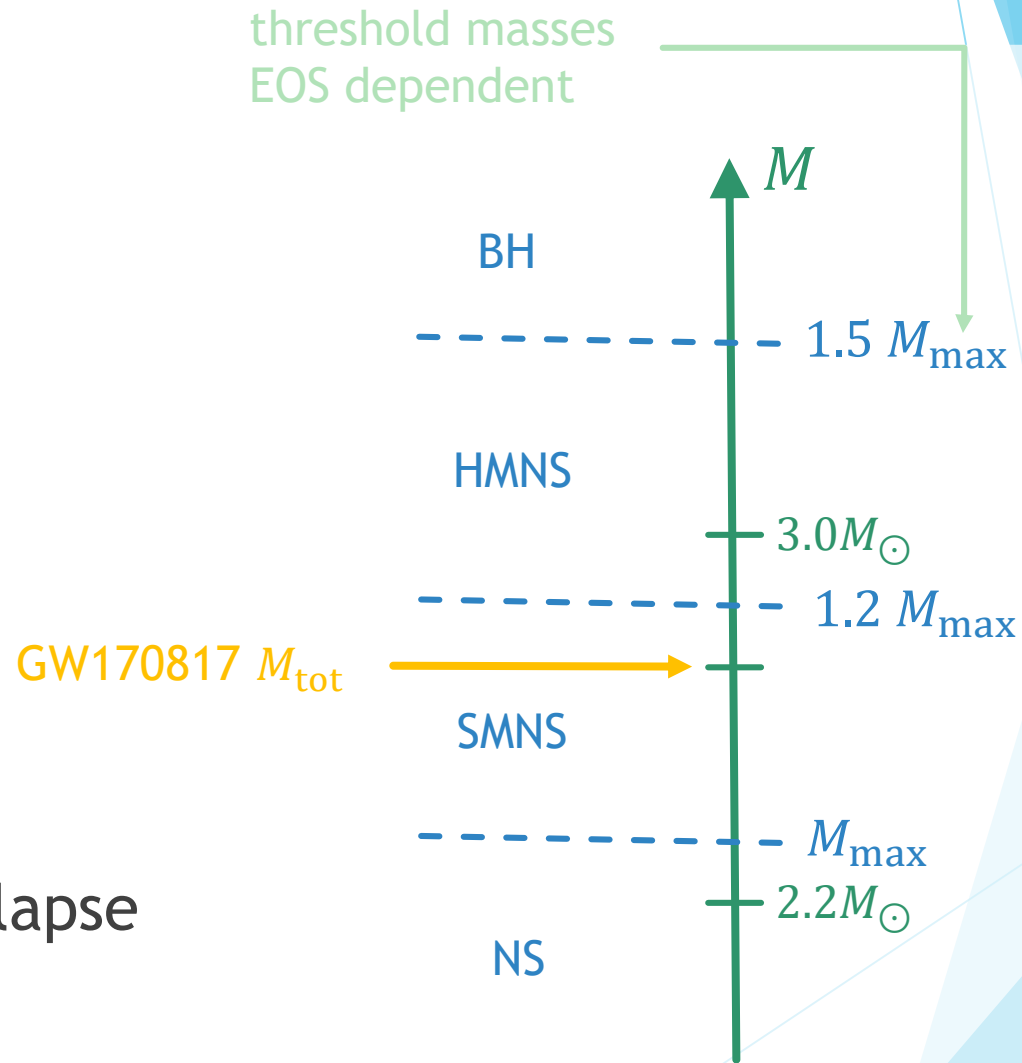
$M_t < 1.2M_{max}$



# GW170817 & the NS EOS (constraining $M_{\max}$ )

## Merger Outcome:

- ruling out long-lived NS  
 ⇒ upper limit on  $M_{\max}$
- similarly, ruling out prompt-collapse  
 ⇒ lower limit on  $M_{\max}$



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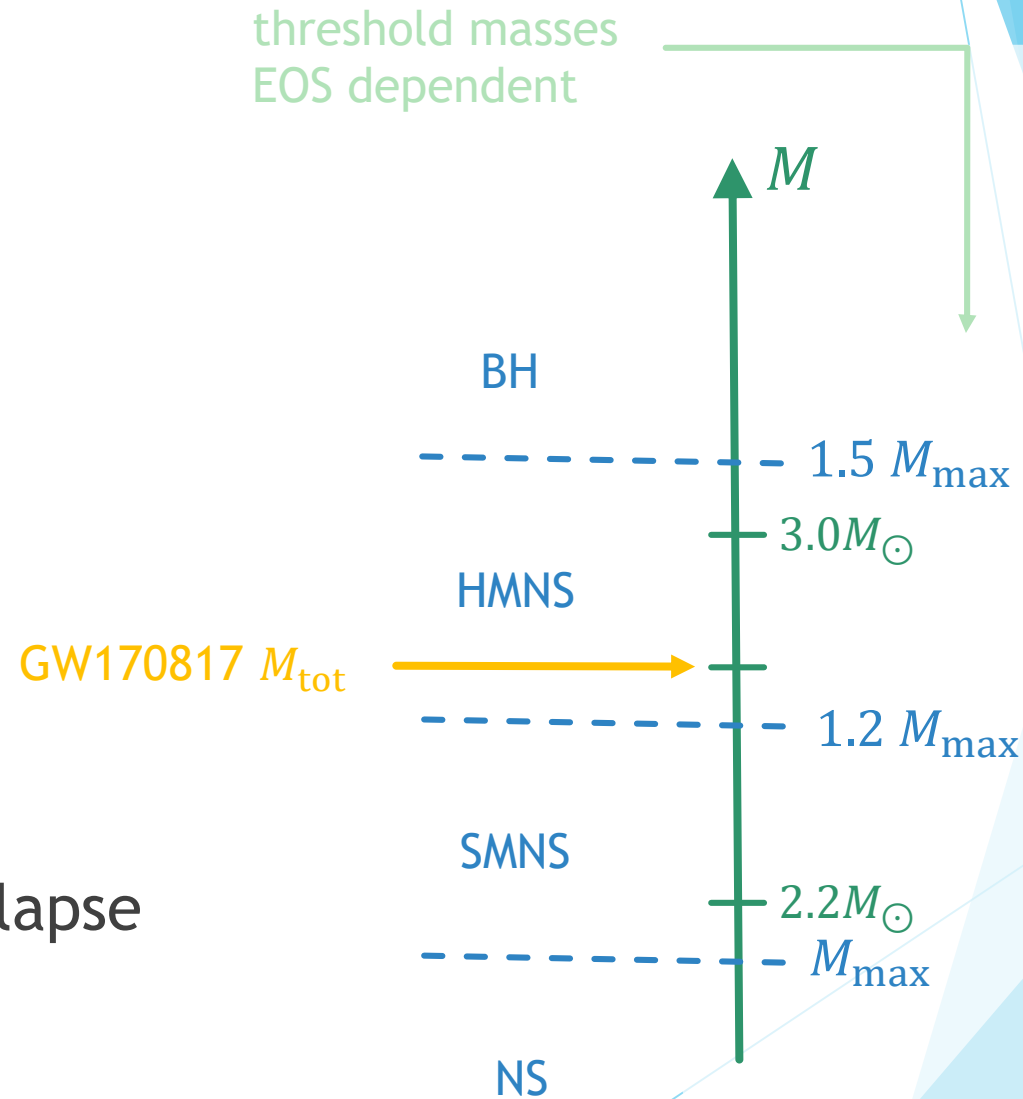
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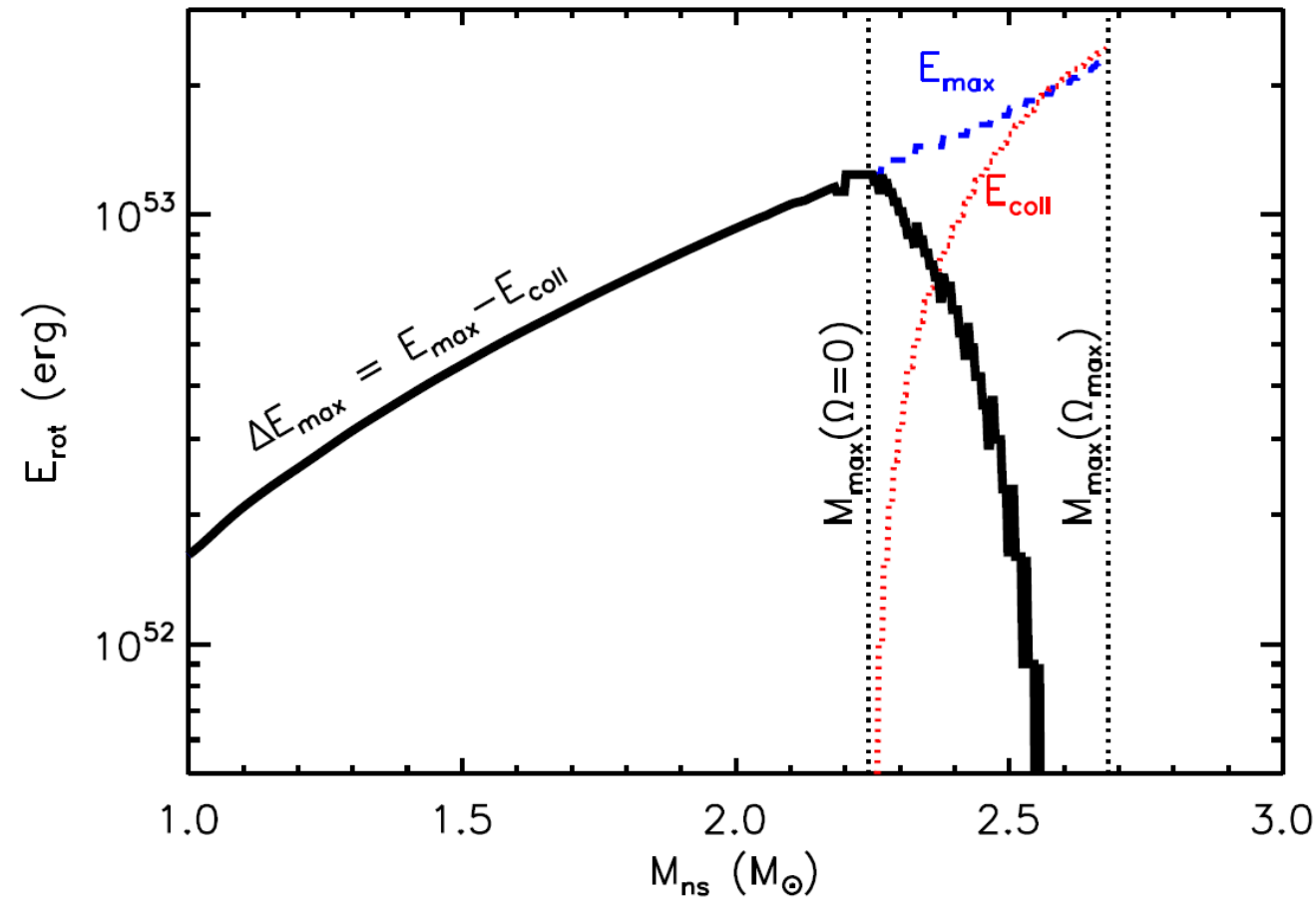
## Spindown of SMNS:

- merger remnant maximally rotating

$\Rightarrow \sim 10^{53}$  erg energy reservoir!

- total rotational energy increases with  $M$ , but extractable energy drops for  $M > M_{\max}$

(Metzger, BM, Kasen & Quataert 2015)

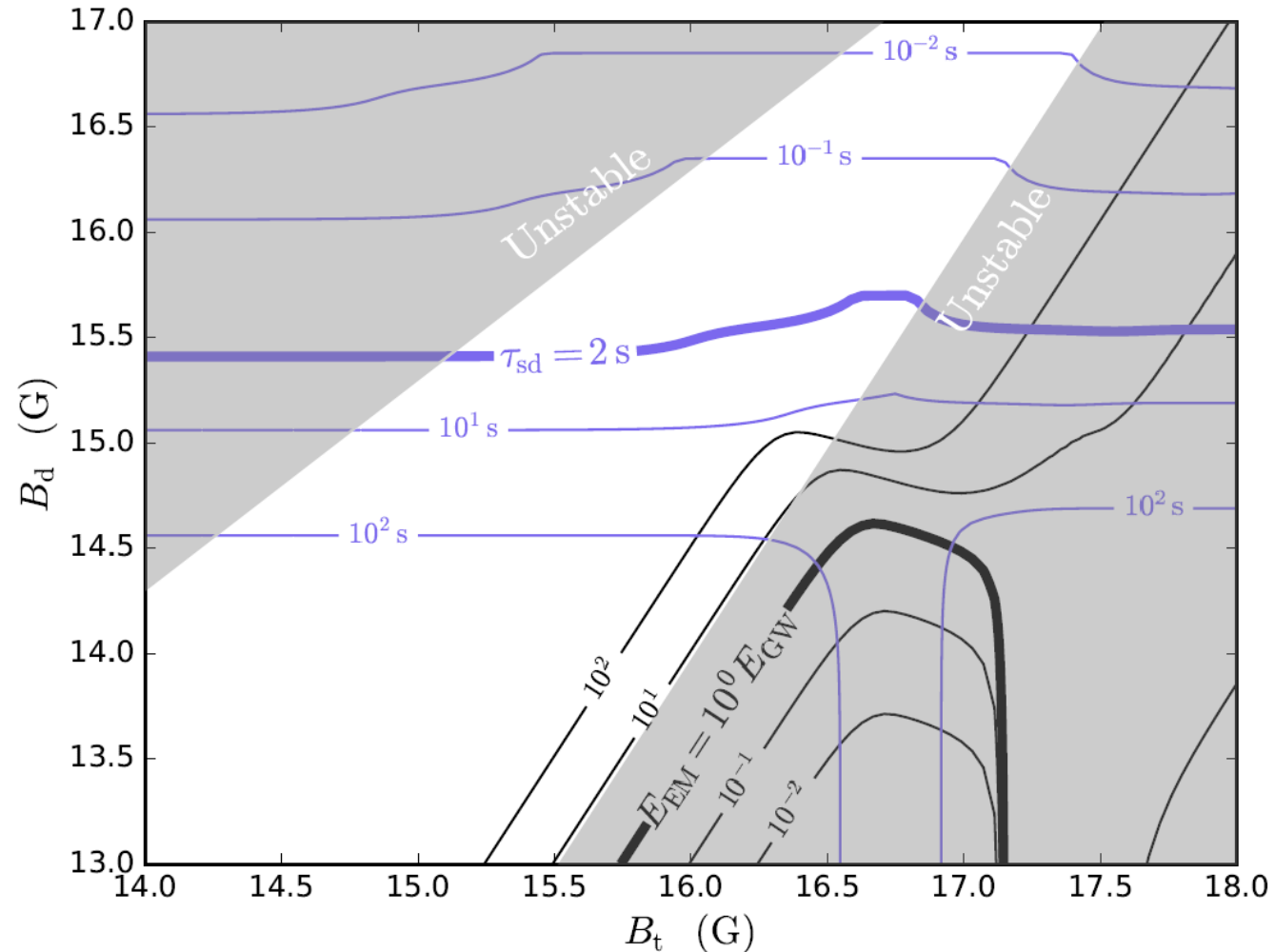


# GW170817 & the NS EOS (constraining $M_{\max}$ )

## Spindown of SMNS:

- merger remnant maximally rotating  
 $\Rightarrow \sim 10^{53}$  erg energy reservoir!
- for stable NS / SMNS - will be tapped by magnetic-dipole spindown
- GW spindown unlikely:
  - requires unstable  $B_t \gtrsim 100B_d$
  - long  $\tau_{\text{sd}} \sim 100\text{s}$  in tension with GRB

(BM & Metzger 2017)



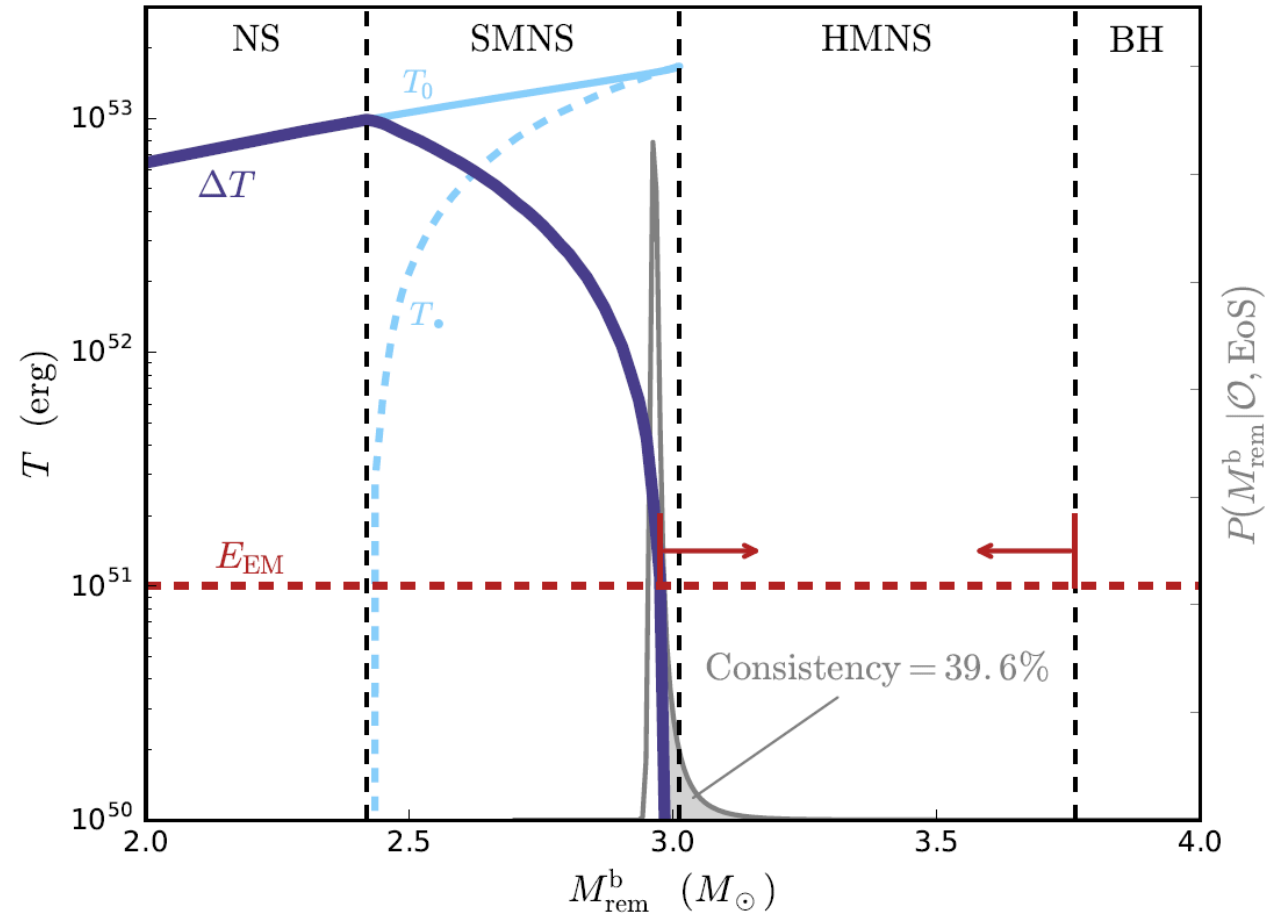


# GW170817 & the NS EOS (constraining $M_{\max}$ )

## Spindown of SMNS:

- merger remnant maximally rotating  
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- for stable NS / SMNS - will be tapped by magnetic-dipole spindown
- EM spindown inconsistent with kilonova + GRB energetics

(BM & Metzger 2017)

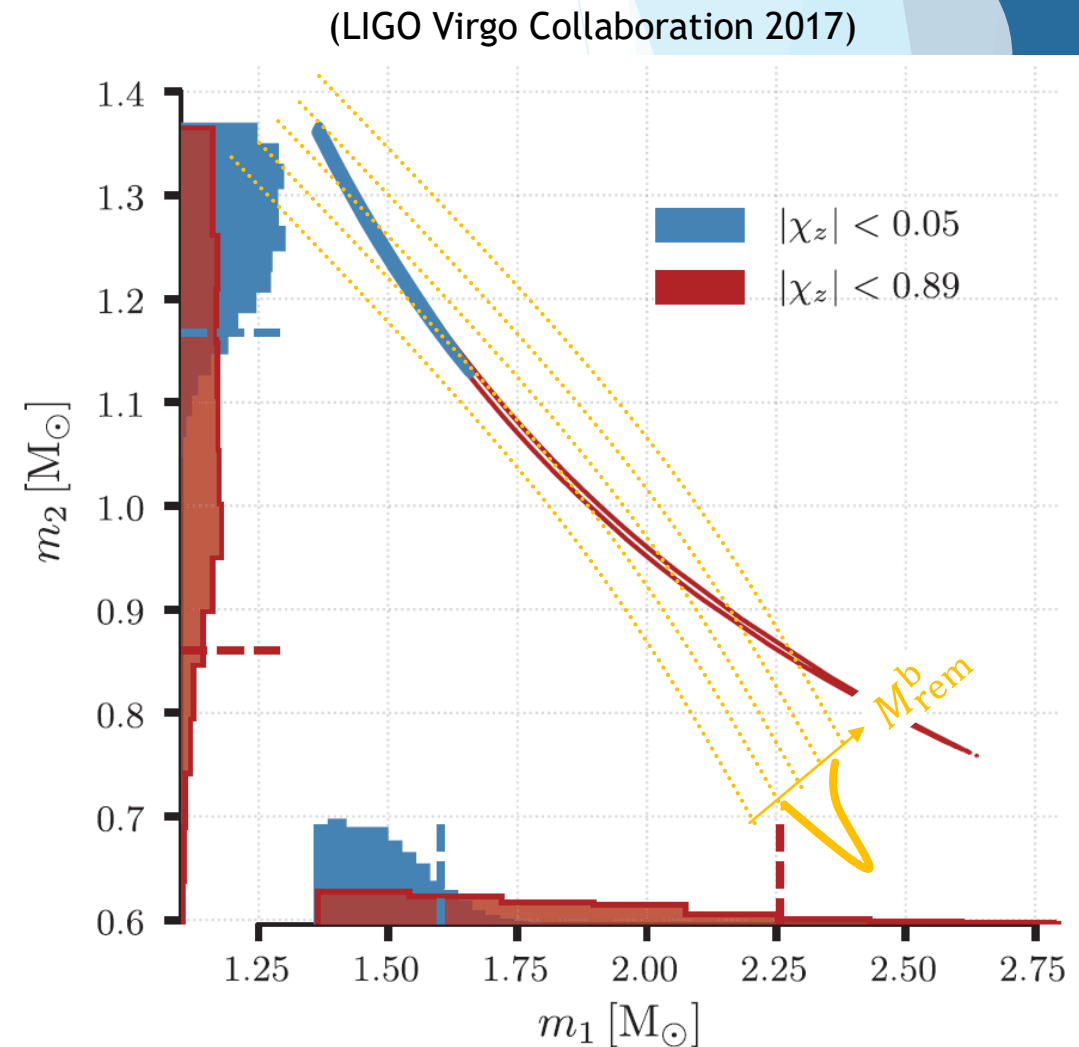


## Consistency of given EOS:

- GW inferred  $P(m_1^g, m_2^g)$  implies, for a given EOS:

$$P(M_{\text{rem}}^b | \text{EOS})$$

$$[M_{\text{rem}}^b \approx M^b(m_1^g; \text{EOS}) + M^b(m_2^g; \text{EOS}) - M_{\text{ej}}]$$

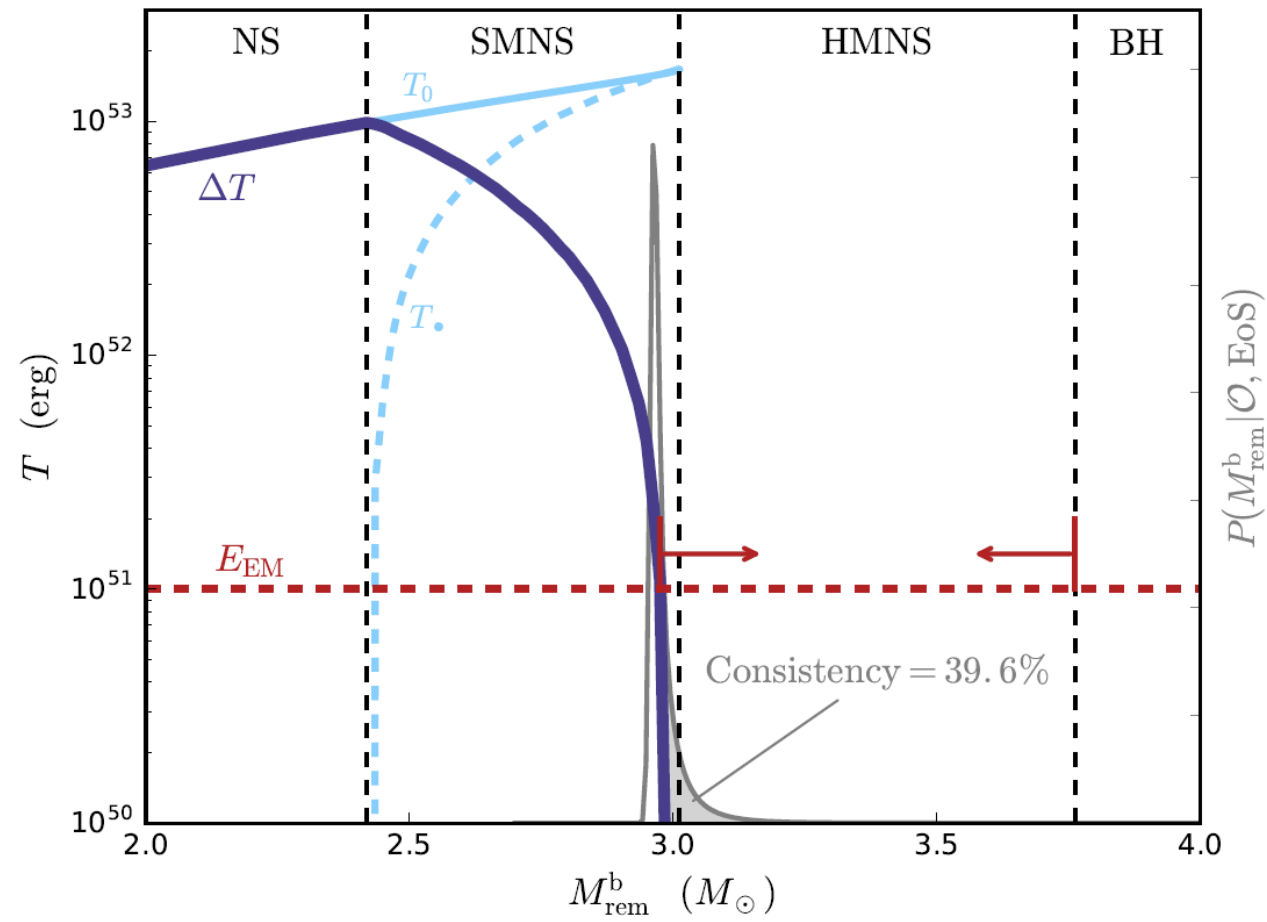


## Consistency of given EOS:

○ process:

1. construct an EOS (Read+09; BM+15)
2. create rotating NS sequences for EOS using RNS code (Stergioulas&Friedman95)  
⇒ calculate:  $M_{\max}$ ,  $R_{1.3}$ ,  $\Delta T$ ,  $M_{\text{thres}}$  (Bauswein+13)
3. get  $P(M_{\text{rem}}^b | \text{EOS})$

(BM & Metzger 2017)

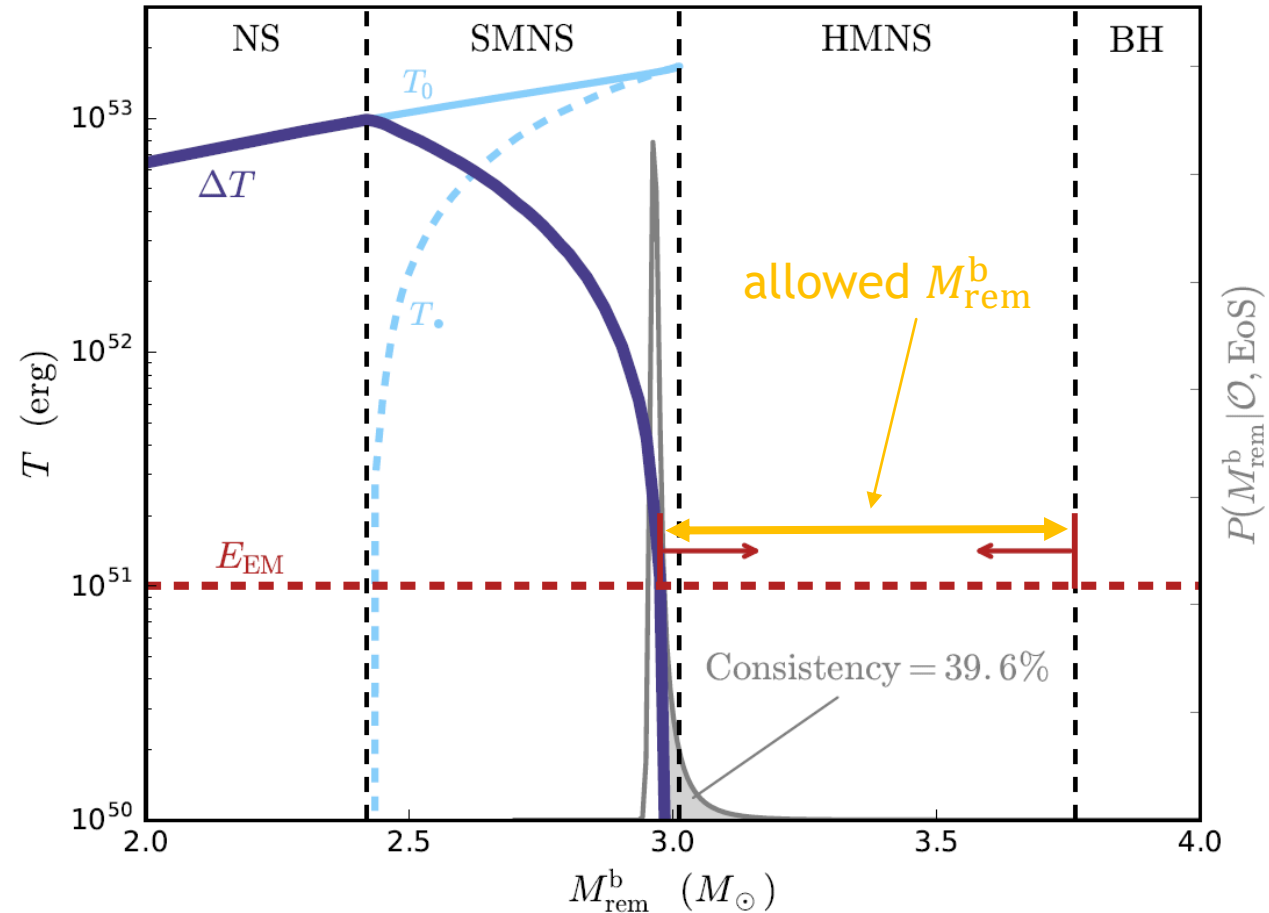


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4. calculate EOS ‘consistency’

(BM & Metzger 2017)



# GW170817 & the NS EOS (constraining $M_{\max}$ )

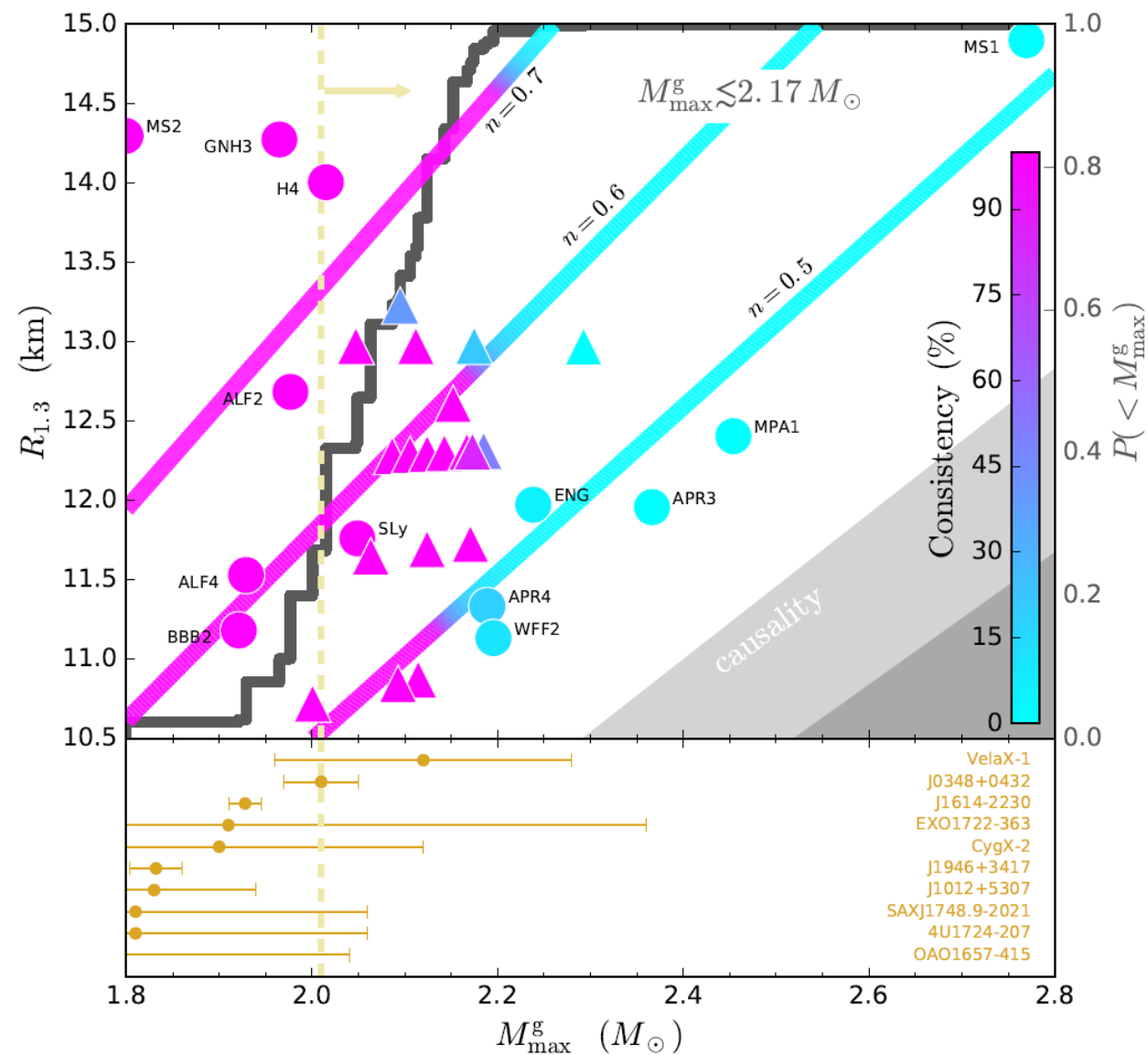
## Results / EOS survey:

- rule out stiff EOS
- largely independent on NS radius or compactness
- ⇒ constraint on  $M_{\max}$
- find  $M_{\max} \lesssim 2.17 M_{\odot}$

[PSR J0348+0432:  $M_{\max} \geq 2.01 \pm 0.04 M_{\odot}$ ]

(Antoniadis+13)

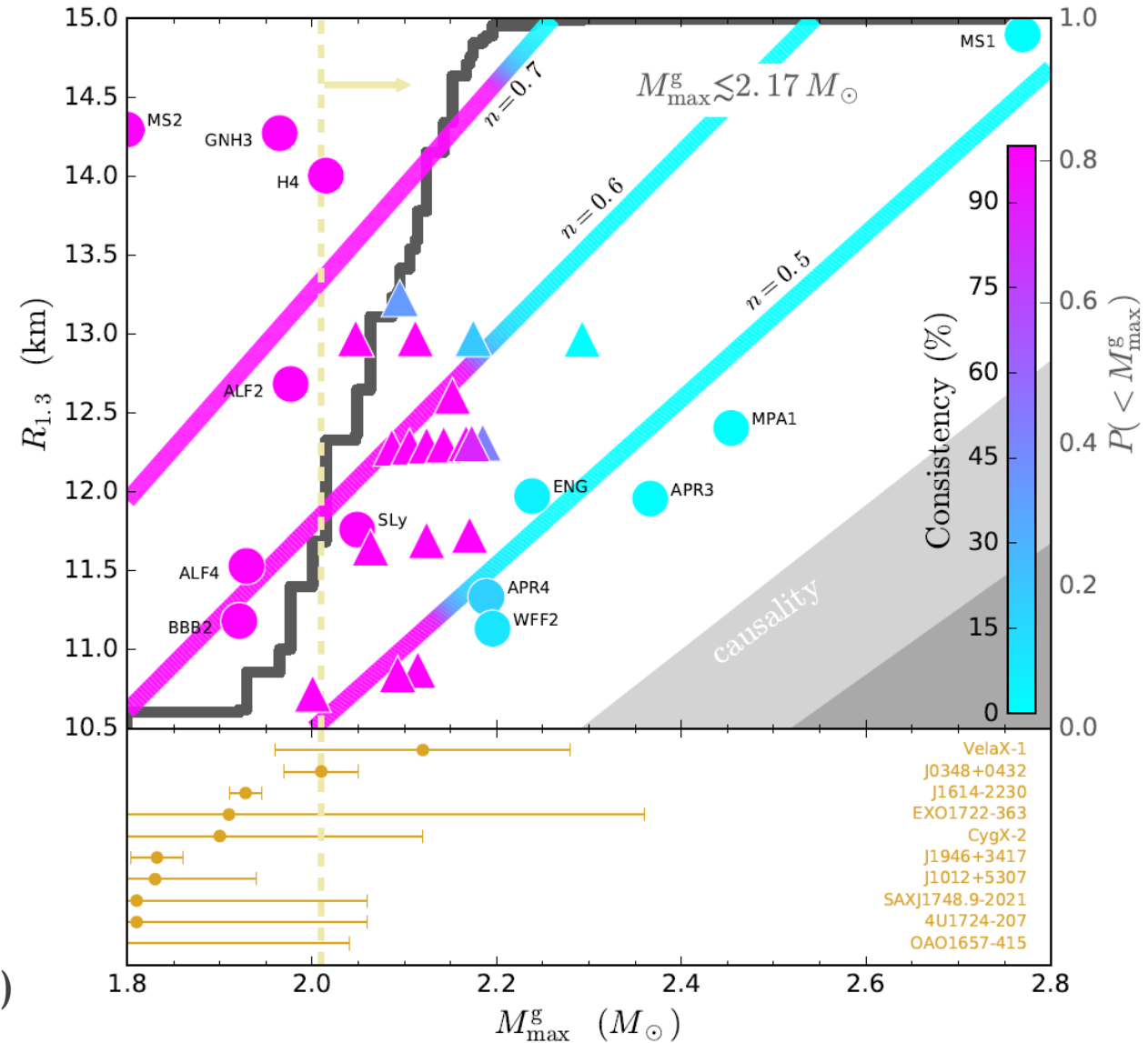
(BM & Metzger 2017)



## Results / EOS survey:

- simplistic analytic estimate of result:
  - $M^b \approx M^g + 0.058(M^g)^2 + 0.013(M^g)^3$   
(from J. Lattimer)
  - $\Rightarrow M_{\text{rem}}^b \lesssim M_{\text{tot}}^b \lesssim 3.07 M_{\odot}$   
(for  $1.17 M_{\odot} + 1.60 M_{\odot}$  component masses)
  - $M_{\text{smns}}^b \approx 1.18 M_{\text{max}}^b$  (rotational support)
  - demand:  $M_{\text{smns}}^b \lesssim M_{\text{rem}}^b$
- $\Rightarrow M_{\text{max}}^g \lesssim 2.19 M_{\odot}$  (for conservative  $M_{\text{ej}} = 0$ )

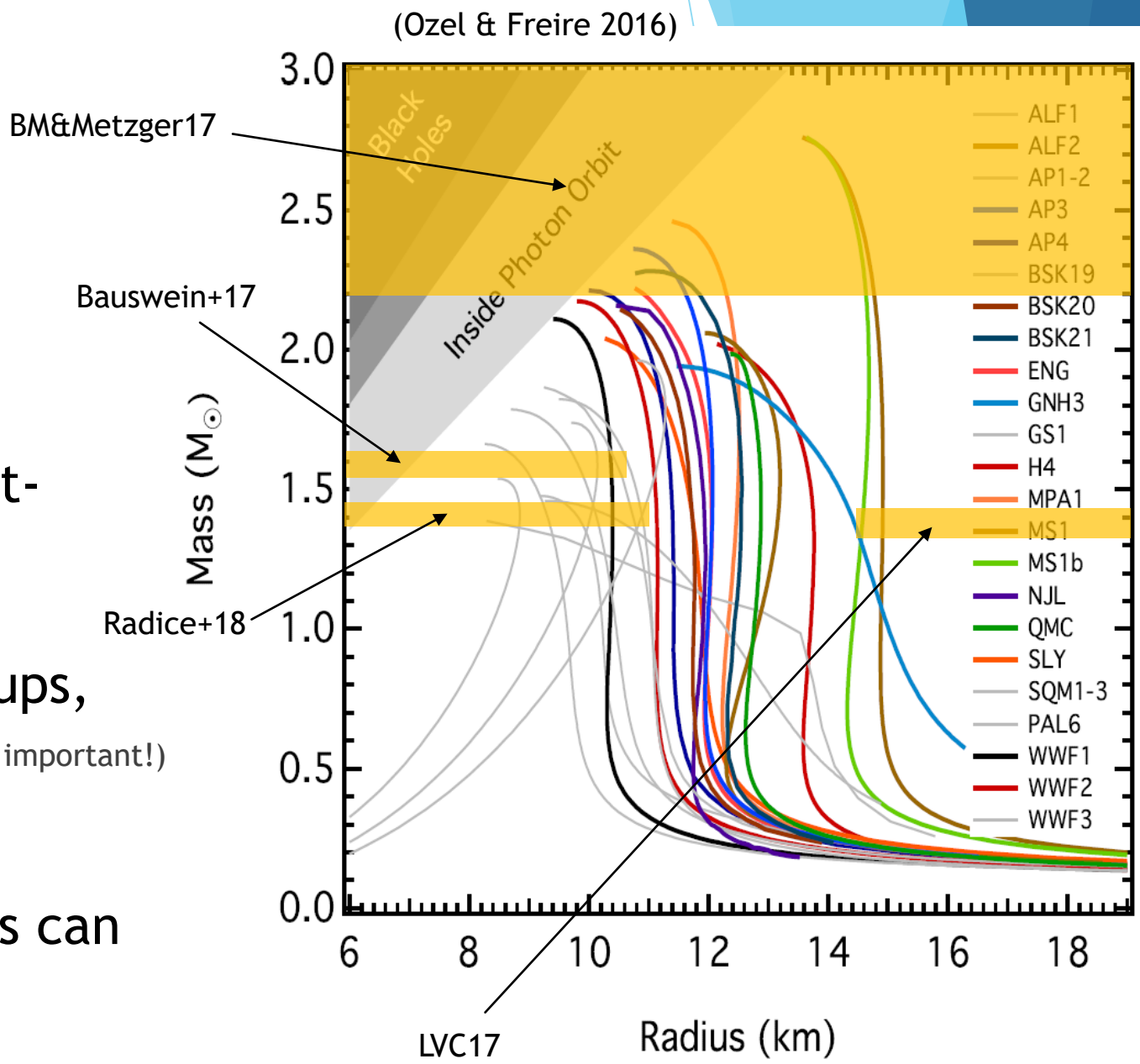
(BM & Metzger 2017)



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## Conclusions:

- simple yet powerful method
- complimentary to purely GW EOS constraints (tidal deformability, post-merger GWs, etc.)
- convergence between different groups,  $M_{\max} \lesssim 2.2M_{\odot}$  (but - thermal effects could be important!) (BM&Metzger17; Shibata+17; Rezzolla+18; Ruiz+18)
- future multi-messenger observations can further constrain  $M_{\max}$



## Conclusions:

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