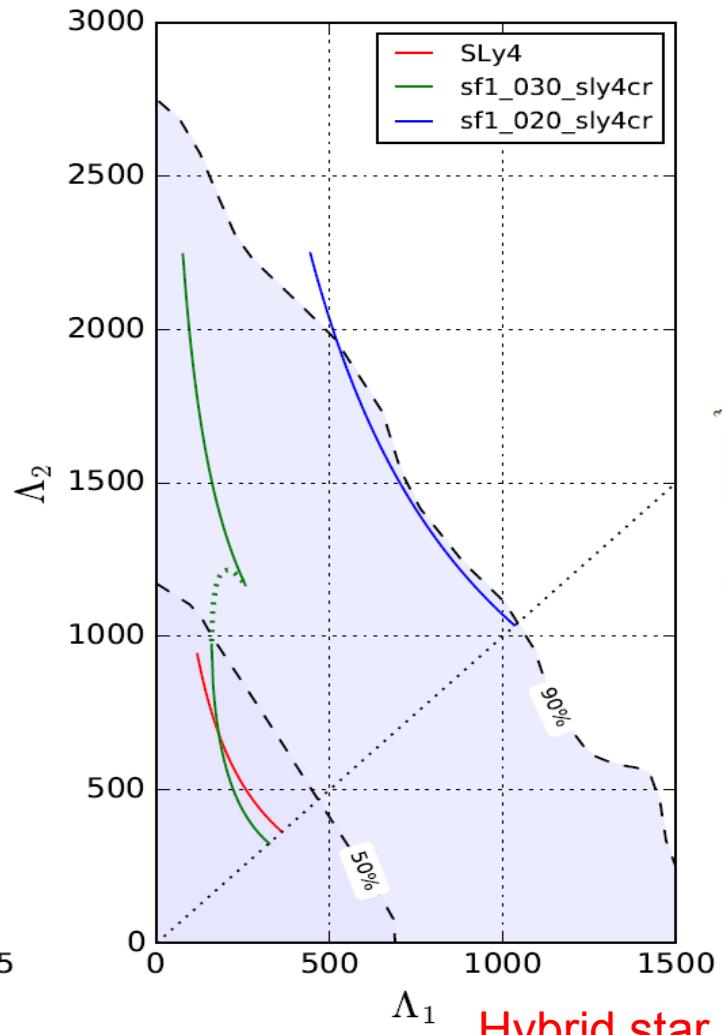
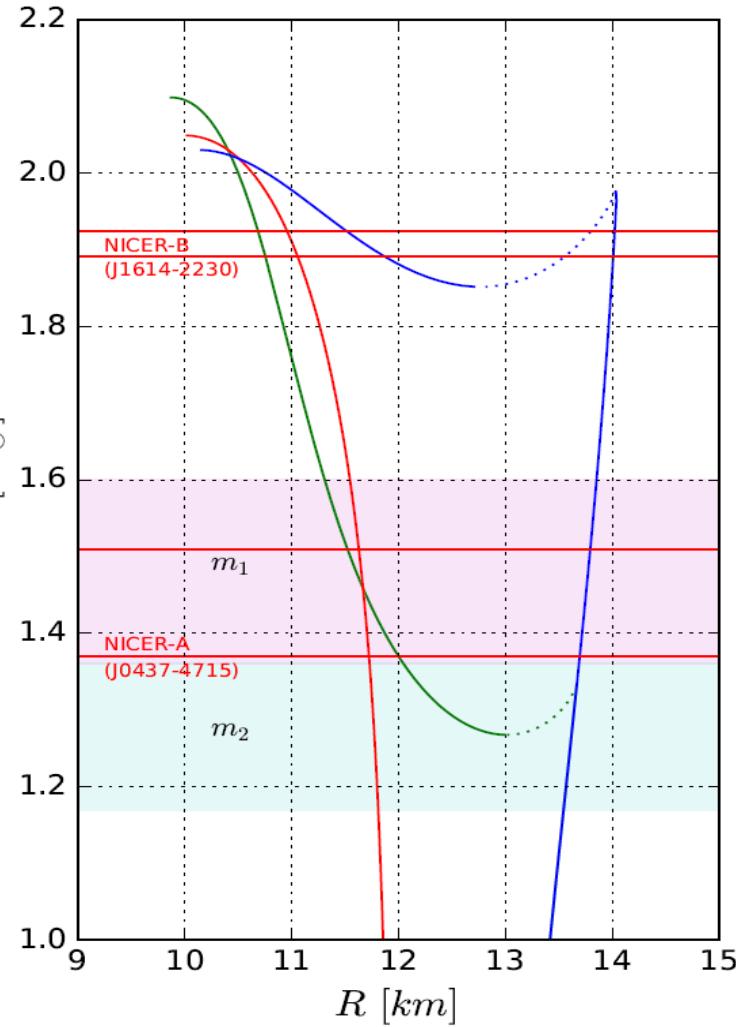
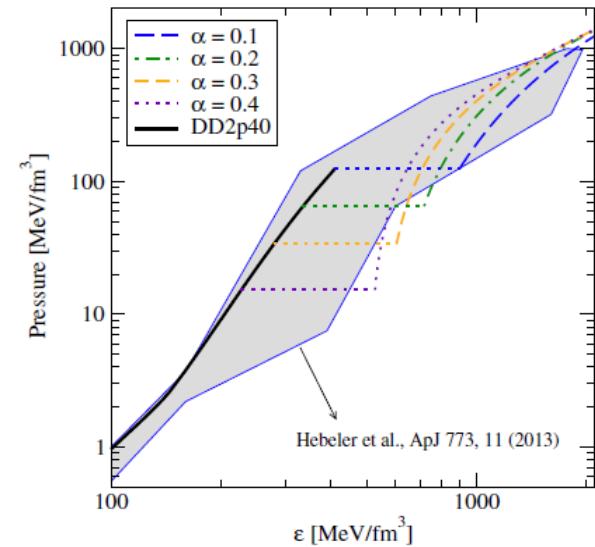


Was GW170817 not a neutron star merger?



EoS:
DD2_P40 – SFM_α=0.3
M. Kaltenborn et al.
PRD 96 (2017) 056024



TOV / TD calculation:
M. Bejger et al.
Alternative:
Hybrid star (HS) – HS / HS-NS merger

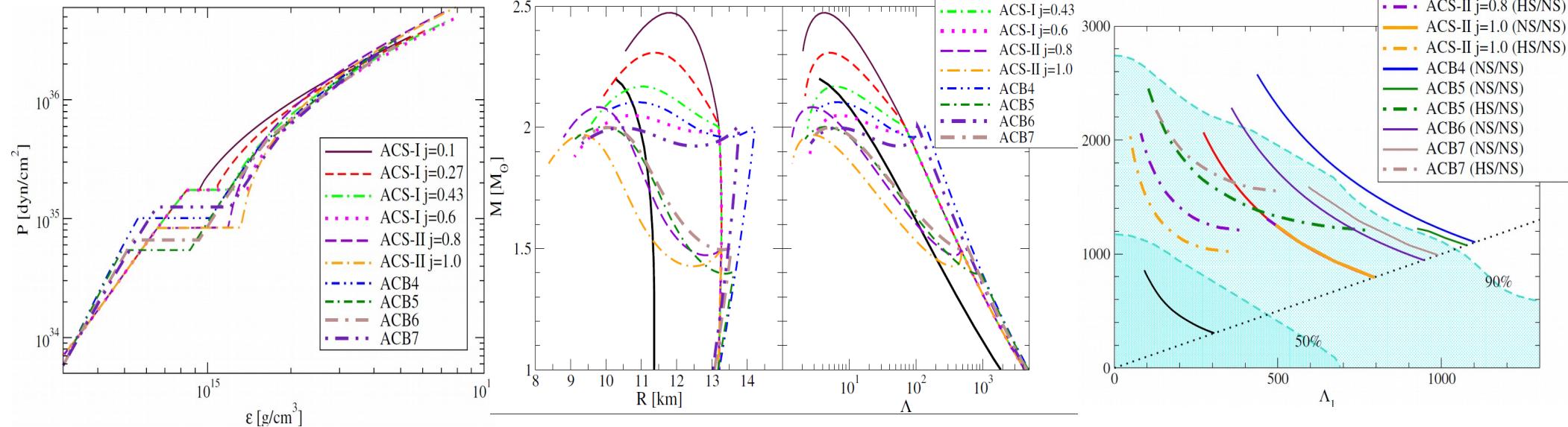


David Blaschke, INT Seattle, 12.03.2018

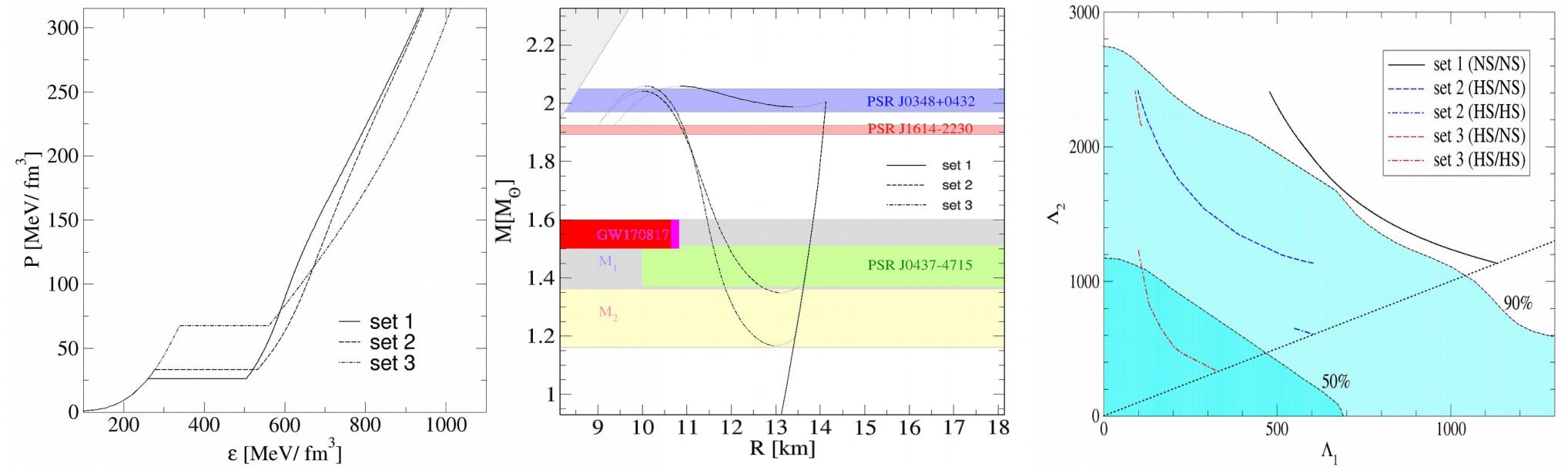


Other examples: Multi-polytrope and multi-CSS model

V. Paschalidis et al., arxiv:1712.00451

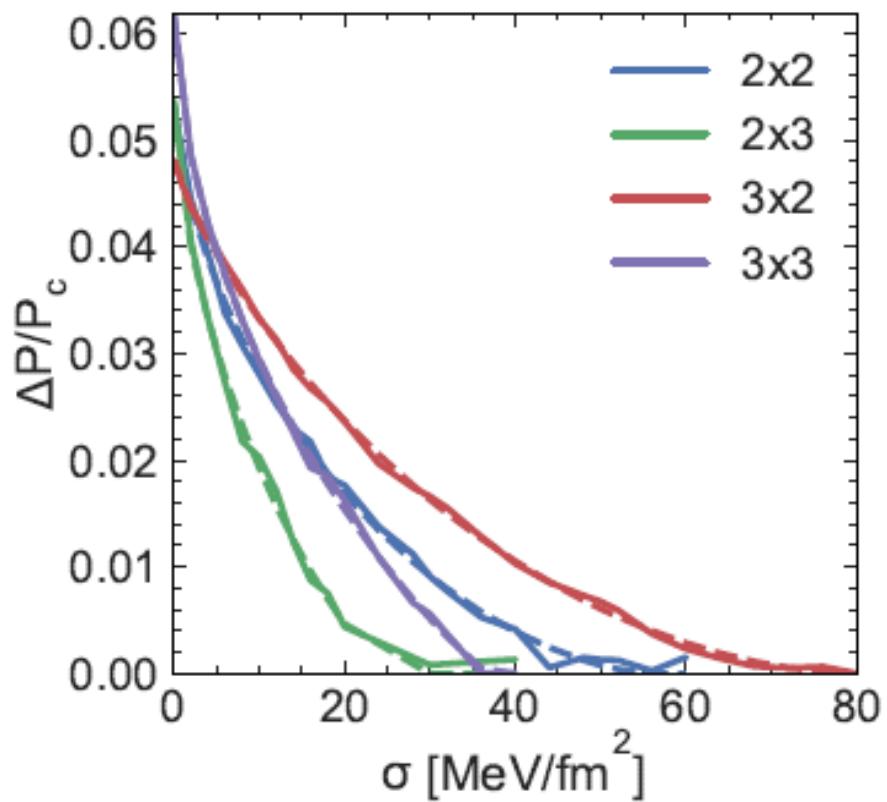


Nonlocal NJL model (with interpolation), D. Alvarez-Castillo et al. (in preparation)



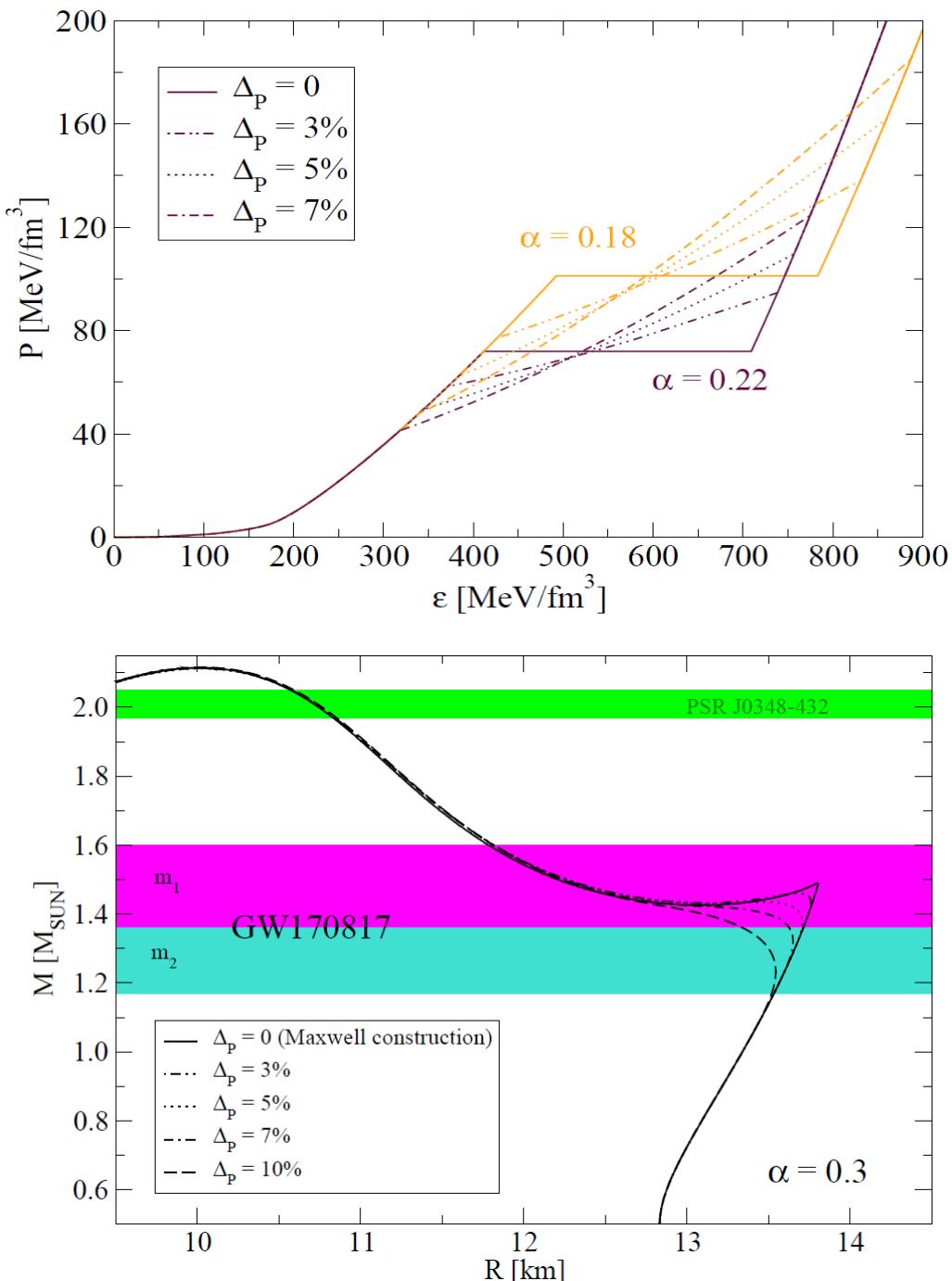
Effect of a mixed phase: KVOR_cut2 – SFM_α=0.3, A. Ayriyan et al., 1711.03926

Small surface tension – small structures



K. Maslov, N. Yasutake et al. (in prep.)

Pasta structures
do not destroy the HS-NS merger alternative



CEP in the QCD phase diagram: HIC vs. Astrophysics

