

# *Neutron-Antineutron Oscillation, Low-scale Baryogenesis, Dark Matter and LHC Physics*

BHUPAL DEV

*Washington University in St. Louis*

R. Allahverdi, BD and B. Dutta, arXiv:1711.xxxxx.

BD and R. N. Mohapatra, Phys. Rev. D **92**, 016007 (2015) [arXiv:1504.07196].

**INT Workshop on Neutron-Antineutron Oscillations**

*University of Washington, Seattle*



October 26, 2017



# Proton Decay vs $n - \bar{n}$

## Selection rules for $\Delta B$

$$\Delta B = 1$$

- Proton decay
- Induced by dimension-6 operator (also dimension-5 in SUSY).
- Amplitude  $\propto \Lambda^{-2}$ .
- $\tau_p \gtrsim 10^{34}$  yr implies  $\Lambda \gtrsim 10^{15}$  GeV.
- Proton decay requires GUT-scale physics.

$$\Delta B = 2$$

- Di-nucleon decay and  $n - \bar{n}$
- Induced by dimension-9 operator.
- Amplitude  $\propto \Lambda^{-5}$ .
- $\Lambda \gtrsim 100$  TeV enough to satisfy experimental constraints.
- $n - \bar{n}$  oscillation could come from a TeV-scale new physics.

$\Delta B \neq 0$  could be linked to **baryogenesis** (Sakharov).

# Highlights of this Talk

A simple TeV-scale SM-extension with baryogenesis, dark matter and  $n - \bar{n}$ .

- Introduces  $\beta$ -interactions via TeV-scale color-triplet scalars ( $X_\alpha$ ) and a singlet Majorana fermion ( $\psi$ ) that couple only to the RH quarks.
- $\psi$  is stable, and hence, a DM candidate, if  $m_\psi \simeq m_p$ .
- Baryogenesis occurs via out-of-equilibrium decays of  $X_\alpha$ .
- Common origin for both baryon and DM abundance.
- Requirements of successful baryogenesis and  $\Omega_{\text{DM}}/\Omega_b \approx 5$  put meaningful constraints on the model parameter space.
- Observable  $n - \bar{n}$  in the allowed parameter space.
- Complementarity with monojet/monotop signals at the LHC.

# Highlights of this Talk

A simple TeV-scale SM-extension with baryogenesis, dark matter and  $n - \bar{n}$ .

- Introduces  $\beta$ -interactions via TeV-scale color-triplet scalars ( $X_\alpha$ ) and a singlet Majorana fermion ( $\psi$ ) that couple only to the RH quarks.
- $\psi$  is stable, and hence, a DM candidate, if  $m_\psi \simeq m_p$ .
- Baryogenesis occurs via out-of-equilibrium decays of  $X_\alpha$ .
- Common origin for both baryon and DM abundance.
- Requirements of successful baryogenesis and  $\Omega_{\text{DM}}/\Omega_b \approx 5$  put meaningful constraints on the model parameter space.
- Observable  $n - \bar{n}$  in the allowed parameter space.
- Complementarity with monojet/monotop signals at the LHC.

# The Model

- Start with the SM gauge group and add renormalizable terms that violate baryon number.
- Gauge invariance requires introduction of new colored fields.
- A minimal setup: Iso-singlet, color-triplet scalars  $X_\alpha$  with  $Y = +4/3$ .
- Allows  $X_\alpha d^c d^c$  terms in the Lagrangian.
- Need at least two ( $\alpha = 1, 2$ ) to produce baryon asymmetry from  $X$  decay.
  
- Total baryon asymmetry vanishes after summing over all flavors of  $d^c$ .  
[Kolb, Wolfram (NPB '80)]
- Need additional  $\cancel{B}$  interactions.
- Introduce a SM-singlet Majorana fermion  $\psi$  (also plays the role of DM).

$$\mathcal{L} \supset \left( \lambda_{\alpha i} X_\alpha^* \psi u_i^c + \lambda'_{\alpha ij} X_\alpha d_i^c d_j^c + \frac{1}{2} m_\psi \bar{\psi}^c \psi + \text{H.c.} \right).$$

# The Model

- Start with the SM gauge group and add renormalizable terms that violate baryon number.
- Gauge invariance requires introduction of new colored fields.
- A minimal setup: Iso-singlet, color-triplet scalars  $X_\alpha$  with  $Y = +4/3$ .
- Allows  $X_\alpha d^c d^c$  terms in the Lagrangian.
- Need at least two ( $\alpha = 1, 2$ ) to produce baryon asymmetry from  $X$  decay.
  
- Total baryon asymmetry vanishes after summing over all flavors of  $d^c$ .  
[Kolb, Wolfram (NPB '80)]
- Need additional  $\cancel{B}$  interactions.
- Introduce a SM-singlet Majorana fermion  $\psi$  (also plays the role of DM).

$$\mathcal{L} \supset \left( \lambda_{\alpha i} X_\alpha^* \psi u_i^c + \lambda'_{\alpha ij} X_\alpha d_i^c d_j^c + \frac{1}{2} m_\psi \bar{\psi}^c \psi + \text{H.c.} \right).$$

[Allahverdi, Dutta (PRD '13); BD, Mohapatra (PRD '15); Davoudiasl, Zhang (PRD '15)]

# Dark Matter

- Integrate out  $X_\alpha$  to obtain  $\psi u_i^c d_j^c d_k^c$  interaction (assuming  $m_\psi \ll m_X$ ).
- $\psi$  decays to three quarks (baryons) if  $m_\psi \gg \text{GeV}$ .
- Also  $\psi \rightarrow p + e^- + \bar{\nu}_e$  if  $m_\psi > m_p + m_e$ .
- Absolutely stable for  $m_\psi < m_p + m_e$  (no discrete symmetry required).
- In addition, need  $m_p > m_\psi + m_e$  to avoid  $p \rightarrow \psi + e^+ + \nu_e$ .
- So the viable scenario for  $\psi$  to be the DM candidate is (see also A. Nelson's talk)

$$m_p - m_e \leq m_\psi \leq m_p + m_e .$$

- $\psi$  cannot give mass to light neutrinos through  $H\psi L$  term, because this with  $X\psi u^c$  and  $Xd^c d^c$  terms will induce the dimension-7 operator  $HLu^c d^c d^c$  for rapid proton decay.
- Stability of DM is linked to the stability of proton.

# Dark Matter

- Integrate out  $X_\alpha$  to obtain  $\psi u_i^c d_j^c d_k^c$  interaction (assuming  $m_\psi \ll m_X$ ).
- $\psi$  decays to three quarks (baryons) if  $m_\psi \gg \text{GeV}$ .
- Also  $\psi \rightarrow p + e^- + \bar{\nu}_e$  if  $m_\psi > m_p + m_e$ .
- Absolutely stable for  $m_\psi < m_p + m_e$  (no discrete symmetry required).
- In addition, need  $m_p > m_\psi + m_e$  to avoid  $p \rightarrow \psi + e^+ + \nu_e$ .
- So the viable scenario for  $\psi$  to be the DM candidate is (see also A. Nelson's talk)

$$m_p - m_e \leq m_\psi \leq m_p + m_e .$$

- $\psi$  cannot give mass to light neutrinos through  $H\psi L$  term, because this with  $X\psi u^c$  and  $X d^c d^c$  terms will induce the dimension-7 operator  $HLu^c d^c d^c$  for rapid proton decay.
- **Stability of DM is linked to the stability of proton.**



# DM Relic Density

- For  $m_\psi \approx m_p$ , only annihilation channel is  $\psi\psi \rightarrow u^c u^c$ .

$$\langle \sigma_{\text{ann}} v \rangle \sim \frac{|\lambda_{\alpha 1}|^4 m_\psi^2}{8\pi m_X^4}.$$

- For  $m_X \sim \mathcal{O}(1 \text{ TeV})$ , even  $\lambda \sim \mathcal{O}(1)$  gives  $\langle \sigma_{\text{ann}} v \rangle \ll 3 \times 10^{-26} \text{ cm}^3 \text{ s}^{-1}$ .
- Thermal overproduction of  $\psi$  (as expected). [Lee, Weinberg (PRL '77)]
- Need a non-thermal mechanism to obtain the correct relic density.
- Late decay of a scalar (moduli) field  $\phi$  with a low reheating temperature  $T_R \leq \text{GeV}$ . [Moroi, Randall (NPB '00); Allahverdi, Dutta, Sinha (PRD '10)]

$$\frac{n_\psi}{s} = Y_\phi \text{Br}_{\phi \rightarrow \psi},$$

where  $Y_\phi = \frac{3T_R}{4m_\phi}$  is the entropy dilution due to the  $\phi$  decay.

# DM Relic Density

- For  $m_\psi \approx m_p$ , only annihilation channel is  $\psi\psi \rightarrow u^c u^c$ .

$$\langle \sigma_{\text{ann}} v \rangle \sim \frac{|\lambda_{\alpha 1}|^4 m_\psi^2}{8\pi m_X^4}.$$

- For  $m_X \sim \mathcal{O}(1 \text{ TeV})$ , even  $\lambda \sim \mathcal{O}(1)$  gives  $\langle \sigma_{\text{ann}} v \rangle \ll 3 \times 10^{-26} \text{ cm}^3 \text{ s}^{-1}$ .
- Thermal overproduction of  $\psi$  (as expected). [Lee, Weinberg (PRL '77)]
- Need a **non-thermal mechanism** to obtain the correct relic density.
- Late decay of a scalar (moduli) field  $\phi$  with a low reheating temperature  $T_R \leq \text{GeV}$ . [Moroi, Randall (NPB '00); Allahverdi, Dutta, Sinha (PRD '10)]

$$\frac{n_\psi}{s} = Y_\phi \text{Br}_{\phi \rightarrow \psi},$$

where  $Y_\phi = \frac{3T_R}{4m_\phi}$  is the entropy dilution due to the  $\phi$  decay.

# Baryogenesis

- Via direct decays of  $X_\alpha \rightarrow \psi u_i^c, d_i^c d_j^c$ .
- Independent of sphaleron processes.
- Example of **post-sphaleron baryogenesis**. [Babu, Mohapatra, Nasri (PRL '06)]
- For complex  $\lambda_{\alpha i}$  or  $\lambda'_{\alpha ij}$ , interference of tree and one-loop contributions produces a non-zero  $CP$  asymmetry.
- In principle, either self-energy or vertex diagrams or both could contribute.
- In the non-thermal scenario, final baryon asymmetry also depends on the moduli decay rate:

$$\eta_B \simeq 7.04 Y_\phi \sum_\alpha \text{Br}_{\phi \rightarrow X_\alpha} \epsilon_\alpha .$$

# Moduli Decay

- Naturally long-lived due to gravitationally suppressed couplings.
- Dominates the energy density of the universe before decaying.
- Must decay well before BBN ( $T_{\text{BBN}} \sim \text{MeV}$ ).
- Decay rate:  $\Gamma_\phi = \frac{c_\phi m_\phi^3}{2\pi M_{\text{Pl}}^2}$ , where  $c_\phi \sim 0.01 - 1$  (in typical string compactification scenarios, e.g. KKLT).
- Moduli decay occurs when  $\Gamma_\phi \sim H \simeq 1.66\sqrt{g_*} \frac{T^2}{M_{\text{Pl}}}$ .
- Reheat temperature:

$$T_R \simeq c_\phi^{1/2} \left( \frac{10.75}{g_*} \right)^{1/4} \left( \frac{m_\phi}{100 \text{ TeV}} \right)^{3/2} 3.5 \text{ MeV}.$$

- Requiring  $\text{MeV} \lesssim T_R \lesssim \text{GeV}$  implies  $200 \text{ TeV} \lesssim m_\phi \lesssim 4500 \text{ TeV}$ , or  $10^{-9} \lesssim Y_\phi \equiv \frac{3T_R}{4m_\phi} \lesssim 10^{-7}$ .
- Need  $\epsilon \sim 10^{-3} - 10^{-1}$ .

# Moduli Decay

- Naturally long-lived due to gravitationally suppressed couplings.
- Dominates the energy density of the universe before decaying.
- Must decay well before BBN ( $T_{\text{BBN}} \sim \text{MeV}$ ).
- Decay rate:  $\Gamma_\phi = \frac{c_\phi m_\phi^3}{2\pi M_{\text{Pl}}^2}$ , where  $c_\phi \sim 0.01 - 1$  (in typical string compactification scenarios, e.g. KKLT).
- Moduli decay occurs when  $\Gamma_\phi \sim H \simeq 1.66\sqrt{g_*} \frac{T^2}{M_{\text{Pl}}}$ .
- Reheat temperature:

$$T_R \simeq c_\phi^{1/2} \left( \frac{10.75}{g_*} \right)^{1/4} \left( \frac{m_\phi}{100 \text{ TeV}} \right)^{3/2} 3.5 \text{ MeV}.$$

- Requiring  $\text{MeV} \lesssim T_R \lesssim \text{GeV}$  implies  $200 \text{ TeV} \lesssim m_\phi \lesssim 4500 \text{ TeV}$ , or  $10^{-9} \lesssim Y_\phi \equiv \frac{3T_R}{4m_\phi} \lesssim 10^{-7}$ .
- Need  $\epsilon \sim 10^{-3} - 10^{-1}$ .

# Resonant Baryogenesis

- Similar in spirit to **resonant leptogenesis**. [Pilaftsis (PRD '97); Pilaftsis, Underwood (NPB '03; PRD '05); BD, Pilaftsis, Millington, Teresi (NPB '14)]
- Self-energy graphs dominate the  $CP$ -asymmetry for quasi-degenerate  $X_\alpha$ 's.
- Resonantly enhanced [up to  $\mathcal{O}(0.1)$ ] for  $\Delta m_X \lesssim \Gamma_X/2$ .

$$\epsilon_\alpha = \frac{1}{8\pi} \frac{\sum_{ijk} \text{Im}(\lambda_{\alpha k}^* \lambda_{\beta k} \lambda'_{\alpha ij} \lambda'_{\beta ij})}{\sum_k |\lambda_{\alpha k}|^2 + \sum_{ij} |\lambda'_{\alpha ij}|^2} \frac{(m_{X_\alpha}^2 - m_{X_\beta}^2) m_{X_\alpha} m_{X_\beta}}{(m_{X_\alpha}^2 - m_{X_\beta}^2)^2 + m_{X_\alpha}^2 \Gamma_{X_\beta}^2}$$

- In the resonance limit, regulator goes as  $m_X/\Gamma_X$ .
- $CP$ -asymmetry becomes *insensitive* to the mass scale  $m_X$ , as well as the overall scaling of the coupling constants.

# Free Parameters and Constraints

- Free parameters:  $m_X, \lambda_{\alpha i}, \lambda'_{\alpha ij}$  (with  $\alpha = 1, 2$  and  $i, j, k = 1, 2, 3$ ).
- Color antisymmetry requires that  $\lambda'_{ij} = 0$  for  $i = j$ .
- Similarly, color conservation does not allow tree-level contributions to quark FCNCs.
- Only major constraint comes from **di-nucleon decay** (like  $pp \rightarrow KK$ ):  
 $|\lambda_{\alpha 1} \lambda'_{\alpha 12}| \lesssim 10^{-6} (m_X / 1 \text{ TeV})^2$ .
- We assume  $\lambda'_{12}$  small, while leave  $\lambda_{\alpha 1}$  as a free parameter.
- For simplicity, also assume  $|\lambda_{1i}| = |\lambda_{2i}| \equiv |\lambda| \quad \forall i = 1, 2, 3$ .
- Similarly, take  $|\lambda'_{1ij}| = |\lambda'_{2ij}| \equiv |\lambda'_{ij}|$ .
- Left with only four parameters  $m_X, \lambda, \lambda'_{13,23}$ .

# Free Parameters and Constraints

- Free parameters:  $m_X, \lambda_{\alpha i}, \lambda'_{\alpha ij}$  (with  $\alpha = 1, 2$  and  $i, j, k = 1, 2, 3$ ).
- Color antisymmetry requires that  $\lambda'_{ij} = 0$  for  $i = j$ .
- Similarly, color conservation does not allow tree-level contributions to quark FCNCs.
- Only major constraint comes from **di-nucleon decay** (like  $pp \rightarrow KK$ ):  
 $|\lambda_{\alpha 1} \lambda'_{\alpha 12}| \lesssim 10^{-6} (m_X / 1 \text{ TeV})^2$ .
- We assume  $\lambda'_{12}$  small, while leave  $\lambda_{\alpha 1}$  as a free parameter.
- For simplicity, also assume  $|\lambda_{1i}| = |\lambda_{2i}| \equiv |\lambda| \quad \forall i = 1, 2, 3$ .
- Similarly, take  $|\lambda'_{1ij}| = |\lambda'_{2ij}| \equiv |\lambda'_{ij}|$ .
- Left with only four parameters  $m_X, \lambda, \lambda'_{13,23}$ .

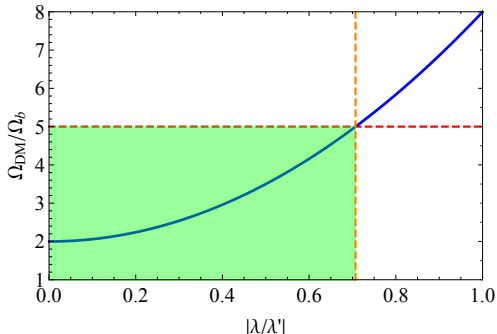


# DM-to-Baryon Ratio

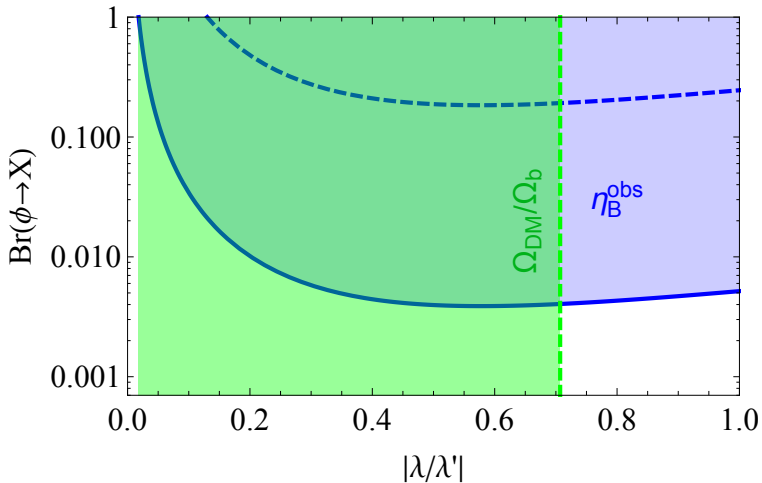
- Both DM and baryonic matter have a common origin from moduli decay.

$$\frac{\Omega_{\text{DM}}}{\Omega_b} = \frac{\text{Br}_{\phi \rightarrow \psi}}{\sum_{\alpha} \epsilon_{\alpha} \text{Br}_{\phi \rightarrow X_{\alpha}}}$$

- $\text{Br}_{\phi \rightarrow \psi}^{\text{total}} = \text{Br}_{\phi \rightarrow \psi}^{\text{direct}} + \sum_{\alpha} \text{Br}_{\phi \rightarrow X_{\alpha}} \text{Br}_{X_{\alpha} \rightarrow \psi} \geq \sum_{\alpha} \text{Br}_{\phi \rightarrow X_{\alpha}} \text{Br}_{X_{\alpha} \rightarrow \psi}$ .
- This implies  $\frac{\Omega_{\text{DM}}}{\Omega_b} \geq \frac{\text{Br}_{X \rightarrow \psi}}{\epsilon}$ .
- $\frac{\Omega_{\text{DM}}}{\Omega_b} \approx 5$  imposes an **upper bound** on the ratio  $|\lambda/\lambda'| \lesssim 1/\sqrt{2}$ , independent of  $m_X, m_{\phi}$ .

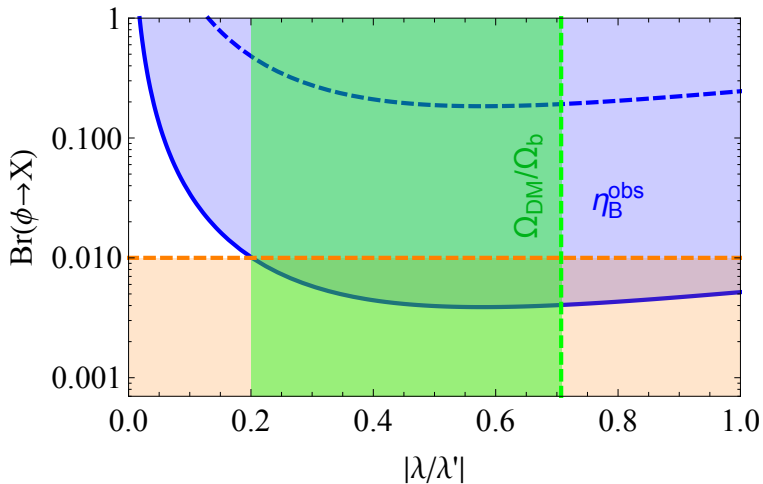


# Baryon Asymmetry



Puts a **lower** bound on  $|\lambda/\lambda'|$  and on the branching of moduli.

# Baryon Asymmetry



Puts a **lower** bound on  $|\lambda/\lambda'|$  and on the branching of moduli.

## $n - \bar{n}$ Oscillation

- Effective  $\not{B}$  operator  $\psi u^c d^c d^c$  (integrating out  $X_\alpha$ ). [Babu, Mohapatra, Nasri (PRL '07)]
- Induces  $n - \bar{n}$  oscillation for Majorana  $N$ .
- Tree-level amplitude vanishes due to color-antisymmetry.
- Non-zero amplitude at one-loop level: [BD, Mohapatra (PRD '15)]

$$G_{n-\bar{n}} \simeq \frac{1}{16\pi^2} \frac{|\lambda_{\alpha 1}|^2 |\lambda'_{\alpha 13}|^4 m_\psi}{m_{X_\alpha}^6} \log \left( \frac{m_{X_\alpha}^2}{m_\psi^2} \right) \\ \simeq (1.9 \times 10^{-28} \text{ GeV}^{-5}) \left( \frac{|\lambda_{\alpha 1}|}{0.03} \right)^2 \left( \frac{|\lambda'_{\alpha 13}|}{0.04} \right)^4 \left( \frac{1 \text{ TeV}}{m_X} \right)^6.$$

- Observable oscillation time for  $m_X \sim \mathcal{O}(\text{TeV})$ :

$$\tau_{n\bar{n}} \simeq (3.0 \times 10^8 \text{ sec}) \left( \frac{0.03}{|\lambda_{\alpha 1}|} \right)^2 \left( \frac{0.04}{|\lambda'_{\alpha 13}|} \right)^4 \left( \frac{m_X}{1 \text{ TeV}} \right)^6.$$

## $n - \bar{n}$ Oscillation

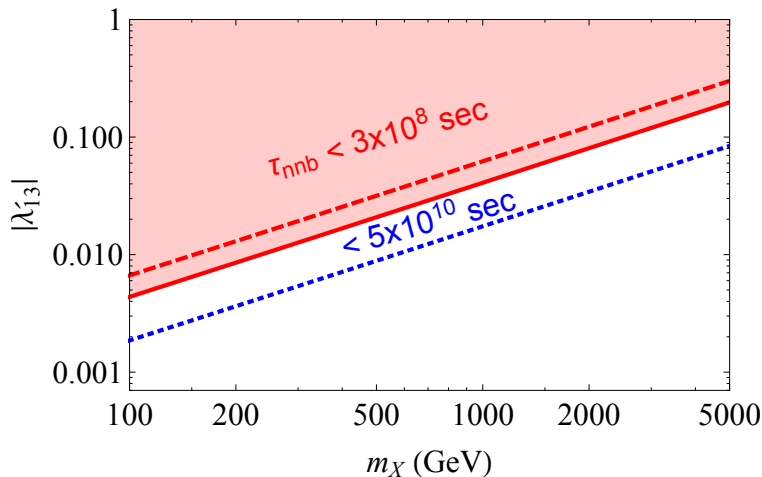
- Effective  $\not{B}$  operator  $\psi u^c d^c d^c$  (integrating out  $X_\alpha$ ). [Babu, Mohapatra, Nasri (PRL '07)]
- Induces  $n - \bar{n}$  oscillation for Majorana  $N$ .
- Tree-level amplitude vanishes due to color-antisymmetry.
- Non-zero amplitude at one-loop level: [BD, Mohapatra (PRD '15)]

$$G_{n-\bar{n}} \simeq \frac{1}{16\pi^2} \frac{|\lambda_{\alpha 1}|^2 |\lambda'_{\alpha 13}|^4 m_\psi}{m_{X_\alpha}^6} \log \left( \frac{m_{X_\alpha}^2}{m_\psi^2} \right)$$
$$\simeq (1.9 \times 10^{-28} \text{ GeV}^{-5}) \left( \frac{|\lambda_{\alpha 1}|}{0.03} \right)^2 \left( \frac{|\lambda'_{\alpha 13}|}{0.04} \right)^4 \left( \frac{1 \text{ TeV}}{m_X} \right)^6 .$$

- Observable oscillation time for  $m_X \sim \mathcal{O}(\text{TeV})$ :

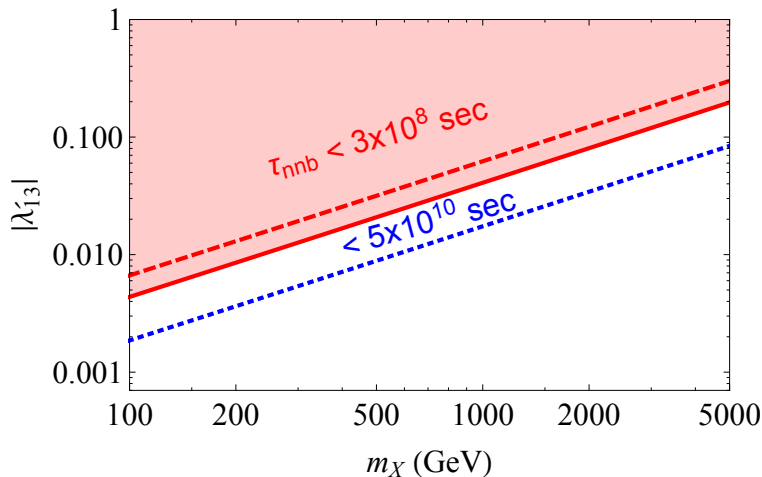
$$\tau_{n\bar{n}} \simeq (3.0 \times 10^8 \text{ sec}) \left( \frac{0.03}{|\lambda_{\alpha 1}|} \right)^2 \left( \frac{0.04}{|\lambda'_{\alpha 13}|} \right)^4 \left( \frac{m_X}{1 \text{ TeV}} \right)^6 .$$

## Constraint from $n - \bar{n}$



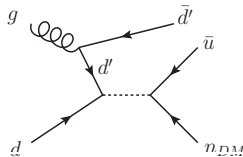
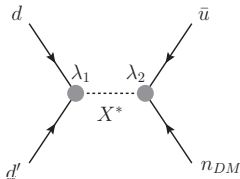
- There is a lower limit on  $|\lambda'_{13}| \gtrsim 10^{-11}$  requiring that  $X$  decay temperature is above QCD scale.
- But the corresponding upper limit on  $\tau_{n\bar{n}}$  is useless ( $10^{62}$  sec).

## Constraint from $n - \bar{n}$

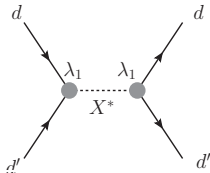


- There is a **lower limit** on  $|\lambda'_{13}| \gtrsim 10^{-11}$  requiring that  $X$  decay temperature is above QCD scale.
- But the corresponding **upper limit** on  $\tau_{n\bar{n}}$  is useless ( $10^{62}$  sec).

# Collider Signals



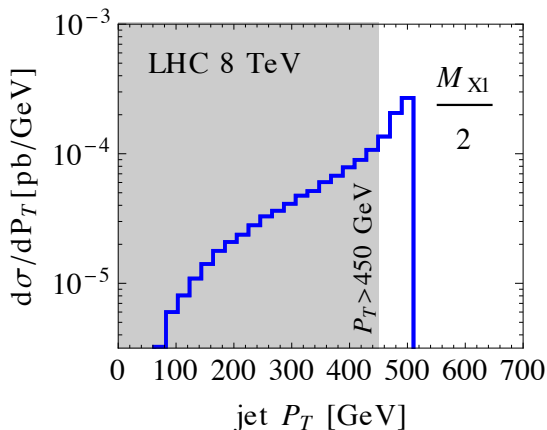
- DM production  $pp \rightarrow \psi u^c$  gives a **monojet (monotop for  $\lambda_{\alpha 3}$ )** signal.
- For  $\lambda'_{13,23}$ , the quark annihilation must involve the  $b$ -quark PDF (small).
- Another way: gluon splitting into  $b\bar{b}$ .
- Extra  $b$  can be used for event tagging.
- The color-triplet scalar will also give a **dijet resonance** at the LHC.



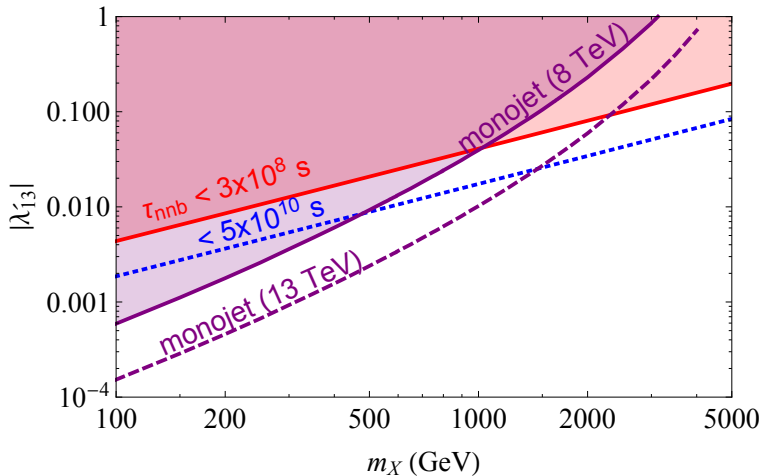


# Monojet

- Different from other DM production at the LHC:  $pp \rightarrow \text{DM DM}$ .
- Will give a Jacobian peak in the jet  $p_T$  distribution. [Duta, Gao, Kamon (PRD '14)]



# $n\bar{n}$ - LHC Complementarity



# Conclusion

- A simple TeV-scale model of  $B$ -violation for baryogenesis and dark matter.
- Stability of dark matter linked to that of proton (no ad-hoc symmetry required).
- DM-to-baryon abundance ratio easily explained.
- Imposes an upper limit on the coupling ratio  $|\lambda/\lambda'|$ .
- Successful baryogenesis imposes a lower bound on  $|\lambda/\lambda'|$ .
- Potentially observable  $n - \bar{n}$  oscillation rate.
- No EDM constraints.
- Distinct monojet and dijet signatures at the LHC.
- Complementarity between monojet and  $n - \bar{n}$ .