Simulation Studies Toward the Search for Neutron—Antineutron Oscillation

By Joshua Barrow

jbarrow3@vols.utk.edu

For the Institute for Nuclear Theory Workshop on Neutron Oscillations:

Appearance, Disappearance, and Baryogenesis

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THE UNIVERSITY OF TENNESSEE KNOXVILLE





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Experimental Overview

Part One



Phenomenology of $n \to \bar{n}$

- Can search for such oscillations in both free beam and bound nuclei experiments
 - Free Oscillation:

$$P_{free}[n(t_{exp}) = \bar{n}] \sim \left(\frac{t_{exp}}{\tau_{n \to \bar{n}}}\right)^2$$

- The bound oscillation can be thought of as a *free oscillation* in a bag-type model
 - Here, the "experimental" time is of the order of $10^{-23}s$ rather than $\sim 1 10$ years
 - From this, we find that:

$$P_A[n(t_{exp}) = \bar{n}] \propto \Gamma = \frac{1}{\tau_A} = P_{free}[n(t_{exp} \cong 10^{-23}s) = \bar{n}] \cong \left(\frac{10^{-23}s}{\tau_{n \to \bar{n}}}\right)^2 \Longrightarrow \tau_A \cong R\tau_{n \to \bar{n}}^2$$

where

$$R \sim \frac{1}{10^{-23}s} \sim 10^{22} s^{-1}$$

• Similarly, we can use the nuclear potential well's interaction difference between n and \bar{n} to assess the same phenomena:

$$\tau_A = \frac{\Delta V_{n \text{ vs. } \bar{n}}}{\hbar} \tau_{n \to \bar{n}}^2 = R \tau_{n \to \bar{n}}^2$$

- It has been shown that $R_0 \sim 5 \cdot 10^{22} s^{-1}$ from an estimated 10 15% uncertainty calculation by Friedman and Gal
- · Precisely, we find that this can be thought of as a nucleon decay-like experiment:

$$P_A[n(t_{exp}) = \bar{n}] = e^{-\frac{t_{exp}}{\tau_A}}$$

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 ${}^{16}_{8}0$

Excellence of LArTPCs—See MicroBooNE



Atmospheric ν Backgrounds Impede $n \rightarrow \overline{n}$ Event Identification In Large Underground Experiments

Previous searches for $n \rightarrow \overline{n}$ suffered greatly from this!

- Super-Kamiokande: 24 candidate events
 - Expected *sig.eff*: bkgrcount = 12%: 24.1
- DUNE?
 - $\underline{CNN}: sig.eff: bkgrcount = 14\%: 3$
 - Fully oscillated sample!
 - Important v_{τ} interactions
 - <u>**Truth**</u>: *sig.eff*: *bkgrcount* = **27**%: **0**
 - Unoscillated sample! DON'T BELIEVE!

Theorists and computational physicists are working tirelessly to improve the accuracy of ν generators

- This is a requirement for understanding ν oscillation parameters precisely and atmospheric background properly
- GENIE is one of many used today
- <u>GIBUU</u> and <u>NuWRO</u> are supposedly getting great comparative results from <u>MicroBooNE</u> with their novel techniques
 - Plan to run events on these platforms for separability comparison in the future



The bottom line...

- Atm. v bkgr. increases faster than mass
- Must counter this with more precise detectors
- Must understand underlying nuclear models
- *Test many simulations against one another* for a full understanding of bkgr. topologies

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Theoretically Important Probability Parameter Space of $\tau_{n \to \overline{n}}$

•A probability distribution from PSB can be seen as a function of predicted free oscillation time

- Red line shows horizontal beamline oscillation time
 - ESS, 3 yr, ~500X ILL
- Blue shows DUNE
 - 10 years, ~13,500X ILL
 - Assumes 100% efficiency!
 - Assumes no background!



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Simulation Analysis Basic Overview with Some Simple Comparisons

Part Two



DUNE and ESS *<u>Truth-Level</u>* Monte Carlo Analysis is Near Completion

Can looking purely at the kinematics give good results?

- Have studied MC generators for $n \rightarrow \overline{n}$ in...
 - ${}^{12}_{6}C$, with incident **slow** ($|\vec{p}| \cong 0 \frac{GeV}{c}$), free (beam), transformed \bar{n} + cascade from annihilation
 - Original generator developed over 15+ years by Elena Golubeva of INR Moscow
 - Joint publication with full simulation discussion, along with with V&V forthcoming
 - $\frac{40}{18}Ar$, with internal (bound) transformation + annihilation cascade
 - Original generator developed for GENIE over last ~2 years by Jeremy Hewes of DUNE for PhD work
 - Dissertation forthcoming
 - Hewes and Karageorgi's analysis revolves around CNN's rather than truth level discrimination between signal and background
 - Future generator being developed by E. Golubeva for $^{40}_{18}Ar$
 - Will use same cascade models
 - · Will work to study neutrino FSI as well for complete comparison
- Study multiple generators with different assumptions to assess uncertainty in the models
 - Heavily dependent upon nuclear models

$n \rightarrow \overline{n}$ Generator Assumption Comparison **DUNE** $^{40}_{18}Ar$ —**GENIE**

- Fermi Gas Model
 - Bodek-Ritchie Distribution
 - No correlations included
- ~ 10 annihilation channels modeled
- Uses continuous analytical nuclear density function
 - How accurate for large nuclei?
- Probability of annihilation as a function of radius not modeled
 - Uses only density function

ESS ${}^{12}_{6}C$ —E. Golubeva

- Fermi Gas Model
 - Personally Developed
 - No correlations included
- ~ 100 annihilation channels modeled
- Uses approximated ("stepped") discontinuous analytical nuclear density function
 - Accurate for small nucleus
 - Uses zones of constant nuclear density
- Probability of annihilation by radius known analytically

Nuclear Density Functions and the Radial \overline{n} -Annihilation Probability Density



Figure 9.10: Nuclear density profiles for C^{12} , O^{16} and Fe^{56} .



Initial Meson Production in $\overline{n}N$ Annihilation





Final Exiting $\pi^{\pm,0}$ After Nuclear Transport

DUNE $^{40}_{18}Ar$ —**GENIE**







Annihilating Nucleon Momentum

Due to Fermi Motion

DUNE $^{40}_{18}Ar$ —GENIE

ESS ${}^{12}_{6}C$ —E. Golubeva



More Results From ¹²₆C MC Generator for ESS





Potential Inaccuracies in **Current Simulations** and a few **Proposed Directions** for **Future Progress** Part Three



What to add to future generators for $^{40}_{18}Ar$?

- **Resonances** (ρ , η , etc.)
- Infrequent reaction channels
 - Must be kinematically allowed—can change from nucleus to nucleus
- Precise radial density functions
 - Generally well known for light nuclei, but how well heavy?
- Precise annihilation probability functions
 - Matters in bound searches greatly, as the transformation is more likely to occur near or outside the nuclear envelope
 - Completed analytically for \bar{n} on ${}^{12}_{6}C$, but how to model this for \bar{n} on ${}^{40}_{18}Ar$?
 - Can no longer model \bar{n} as a plane wave!!
- Two-body nucleon correlation functions
 - $(n \rightarrow \overline{n})N$ is an inherently two-body system, unlike most ν interactions
 - What are the *real*, *quantum mechanical* nucleon momentum distributions for each annihilating nucleon in ${}^{40}_{18}Ar$?

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• Two body momentum distributions needed?

An Interesting Forthcoming Investigation:

Nucleon Correlations

- There are QM problems with the assumptions made when throwing momenta for annihilating nucleons in ⁴⁰/₁₈Ar
 - By making a choice from a distribution of *all* nucleons, we have *changed the remaining distribution*
- Furthermore, it has been found that nucleons are choosey about their neighbors
 - Neutrons spend more time around protons in the nucleus
 - These pairs can have <u>inextricably</u>
 <u>linked momenta</u>
 - Changes the weights of expected annihilation reactions
- JLAB data on LAr coming early next year?
- <u>Can we simply use a two body</u> <u>momentum distribution instead of</u> <u>looking at individual nucleons?</u>





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Use Inherently Two-Body Momentum Distributions?



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End Goals and To Do's

- Top priorities are...
 - Comparisons between multiple generators for uncertainty assessments
 - For $n \to \overline{n}$
 - For atmospheric v
 - Improving generators
 - **Add more physics** to understand the precise nature of the topologies of $\bar{n}A$ annihilation
 - Analysis of generated events will continue
 - Partial and full detector simulations with reconstruction
 - Proton and neutron detections in LAr
 - Background rejection must be maximized
 - Free or quasi-free? Better separation techniques?
- Assess final feasibility of $\tau_{n \to \overline{n}}$ limit improvements
- Optimize experimental components (ESS)

Go and find the oscillation!

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For the extra eager...

BACKUP SLIDES

October 18th.

For the UTK Physics HEP/Astro Seminar



Signal Comparison

 $n \rightarrow \bar{n}$ vs. Backgrounds (ex: Atmospheric Neutrino, ν)



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Free Neutron Beam Search at the European Spallation Source

EUROPEAN SPALLATION SOURCE

- Europe's answer to the SNS at ORNL
- Hopes to best Institute Laue-Langevin result with...
 - Longer beamline
 - Higher neutron
 reflectivity and more
 neutrons on target
 - Thinner carbon foil
 - General improvements
 in detector technology
- Benefits from ~guaranteed zero background!
 - Beam pulse timing
 - v's aren't seen by foil

• ILL saw
$$\frac{sig}{bkgr} = \frac{0}{0}$$
 for one year





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Bound Neutron Search at the Deep Underground Neutrino Experiment







- **DUNE** international collaboration of 1000+
 - Partnership of Fermilab and LBNF
 - Will construct world's most intense ν beam
 - The far detector will utilize LArTPCs
 - Fiducial volume of ~40 kilotons
 - LArTPC's superior tracking and PID capabilities enable background reduction
 - Is background-free/quasi-free $n \rightarrow \bar{n}$ search possible?
 - The *real* question we need to answer!!!

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Why do we need to be accurate in our $n \rightarrow \overline{n}$ simulations?

- 1. For rare events, knowing precisely the expected topology and associated observables will always help
 - Accurate nuclear models are needed to understand the cascade of particles produced within the nucleus
 - Possibly aid in the separability between signal and background processes
- 2. Ability to *compare simulations and assess their systematic uncertainties* within them independently
 - Should be precise and quantitative
 - This is why we need multiple types from multiple sources
 - Wish to reliably understand whether the systematic uncertainties within these models lead to statistically irreducible background

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- Looking for consistency of results between simulations
 - Did we get the physics right???
- 3. We learn more about interesting, underlying physics
- 4. Predict and analyze generated signals for hosts of other BSM processes