## Decay properties of neutron-rich nuclei

T. Marketin

Department of Physics, Faculty of Science, University of Zagreb

Observational Signatures of r-process Nucleosynthesis in Neutron Star Mergers Seattle, August 2017.







neutron stars. Alternatively, recent r-process models have

R. Caballero-Folch *et al.*, Phys. Rev. Lett. 117, 012501 (2016) S. Nishimura *et al.*, Phys. Rev. Lett. 106, 052502 (2011) G. Lorusso *et al.*, Phys. Rev. Lett. 114, 192501 (2015)







## Large-scale calculations  $\mathbf{S}$  20  $\mathbf{S}$  20  $\mathbf{S}$  200  $\mathbf{S}$ 0 diuc-Sudic udiuuld. 0



 $T_{\beta, \text{exp}}$  (S)

decay as a function of neutron number *<sup>N</sup>*. There are no systematic Figure 23. Ratio between calculated and experimental half-lives for <sup>b</sup><sup>0</sup>

P. Möller *et al.*, At. Data Nucl. Data Tables 66, 131 (1997)

 $10<sup>4</sup>$ β− decay (Theory: GT)  $10^3$ In 2003. the reference dataset is published,  $10^2$ **and reference admirer** PETER MO¨ LLER, BERND PFEIFFER, AND KARL-LUDWIG KRATZ PHYSICAL REVIEW C **67**, 055802 !2003" again based on the  $FRDM + QRPA$ , but  $10^{1}$ g the 1/2 also including the first-forbidden transitions  $10<sup>0</sup>$ ny th<del>e</del> in<br>ee etatieti יטו<br>. using a gross statistical calculation. 10<sup>−</sup><sup>1</sup>  $\frac{1}{2}$ statisti Proton Number<br>Proton Number<br>Proton Number  $T_{\beta,\text{calc}}/T_{\beta,\text{exp}}$  $10^{-2}$ Total Error = 3.73 for 184 nuclei,  $T_{\text{B,exp}}$  < 1 s PETER MO¨ LLER, BERND PFEIFFER, AND KARL-LUDWIG KRATZ PHYSICAL REVIEW C **67**, 055802 !2003" Total Error = 21.16 for 546 nuclei (13 clipped),  $T_{\beta,exp}$  < 1000 s  $10^{-3}$ − <del>0.75 × 0.75 × 0.75 × 0.75 × 0.75 × 0.</del>75 × 0.75 ×  $10<sup>4</sup>$ 80 − 1.25 where *Mr*<sup>l</sup> **is the average points and in the point** β− decay (Theory: GT + ff)  $10<sup>3</sup>$  $\frac{1}{\sqrt{2}}$  and  $\frac{1}{\sqrt{2}}$  are preference to represent the prefer to repre  $log_{10}(T_{calc}/T_{exc})$  $10^2$ error by a single number we use the measure '*r*<sup>l</sup> 60 1.75 1.25 N  $10^{1}$ Proton Number | error factor. The spread in the spread 0.75 Proton Number 0.25 − 0.25 related to uncertainties in the positions of the levels in the 0  $10<sup>0</sup>$ 40 − 0.75 0 20 40 60 80 100 120  $\frac{1}{2}$  = 1.25  $\frac{1}{2}$  =  $\frac{1}{2}$  = − 1.25 Neutron Number N  $10^{-1}$ − 1.75 the definition of *r*<sup>l</sup> implies that these two quantities corre-FIG. 6. !Color" Plot of the ratio of calculated to experimental #!-decay half-lives for nuclei from 16O to the heaviest known. spond directly to distances as seen by the eye in, for ex- $10^{-2}$ Total Error = 3.08 for 184 nuclei,  $T_{\beta, \text{exp}}$  < 1 s 20 Total Error = 4.82 for 546 nuclei,  $T_{\beta,exp}$  < 1000 s ample, Fig. 4, in units where one order of magnitude is 1. 80  $10^{-3}$  L  $\begin{bmatrix} 1 & 1 \\ 2 & 3 \end{bmatrix}$   $\begin{bmatrix} 1 & 1 \\ 2 & 3 \end{bmatrix}$  and the error out we want to  $\begin{bmatrix} 1 & 1 \\ 2 & 3 \end{bmatrix}$  and  $\begin{bmatrix} 1 & 1 \\ 3 & 4 \end{bmatrix}$  and  $\begin{bmatrix} 1 & 1 \\ 2 & 3 \end{bmatrix}$  and  $\begin{bmatrix} 1 & 1 \\ 3 & 4 \end{bmatrix}$  and  $\begin{bmatrix} 1 & 1 \\ 2 & 3 \end$  $10^{-3}$  $10^{-2}$  $10^2$  $10<sup>3</sup>$  10<sup>−</sup><sup>1</sup> 1 $0<sup>0</sup>$  $10^{1}$ discuss its result in terms like ''on the average the calculated  $log_{10}(T_{ci}$ 0 Experimental β-Decay Half-life  $T_{\beta, \text{exp}}$  (s) **h** a set of 20 and 40 be able to do the  $\frac{1}{2}$  to do the able to do the  $\frac{1}{2}$ 0 20 40 60  $\vdash$  1.75 60 1.75 1.25 Neu $N$   $\uparrow$   $^{1.25}$ we must convert back from the logarithmic scale. The logarithmic scale  $\frac{1}{\pi}$ 0.75 Proton Number  $\mathbf{E}$  is the ratio of calculated to experimental  $\mathbf{E}$  is not nuclei from 16O to the  $\mathbf{E}$  $10.25$ <br> $-0.25$ <br> $-0.75$ 0.25  $\overline{\phantom{a}}$  are conversions background backgr realize that the quantities *Mr*<sup>l</sup> − 0.25 40 − 0.75 The first-forbidden transitions contribute to  $\frac{2}{5}$ and !*r*<sup>l</sup> , which are a set − 1.25 − 1.75 show the theory  $\sum_{n=1}^{\infty}$  $\epsilon$   $\epsilon$   $\epsilon$ the decay rate mostly in nuclei close to the  $\overline{\phantom{a}}^{20}$   $\overline{\phantom{a}}^{20}$   $\overline{\phantom{a}}^{20}$ 20 tween our calculations and experiment. An analysis of the valley of stability. half-life comparisons in Fig. 4 is given in Table I and of the Table I  $\begin{smallmatrix}0&\text{m}&&\text$  $\Omega$ 

 $\frac{3}{5}$  in Table 20 in Tabl

0 20 40 60 80 100 120

## QRPA calculations

Transitions are obtained by solving the pn-(R)QRPA equations

$$
\begin{pmatrix}\nA & B \\
B^* & A^*\n\end{pmatrix}\n\begin{pmatrix}\nX^{\lambda} \\
Y^{\lambda}\n\end{pmatrix} = E_{\lambda}\n\begin{pmatrix}\n1 & 0 \\
0 & -1\n\end{pmatrix}\n\begin{pmatrix}\nX^{\lambda} \\
Y^{\lambda}\n\end{pmatrix}
$$

Residual interaction is derived from the Lagrangian density

$$
\mathcal{L}_{\rho+\pi}=-g_{\rho}\bar{\psi}\gamma_{\mu}\bar{\rho}^{\mu}\vec{\tau}\psi-\frac{f_{\pi}}{m_{\pi}}\bar{\psi}\gamma_{5}\gamma^{\mu}\partial_{\mu}\vec{\pi}\vec{\tau}\psi
$$

Total strength of a particular transition

$$
B_{\lambda,J}(GT) = \left| \sum_{pn} \left\langle p \left| \left| \hat{O}_J \right| \right| n \right\rangle \left( X_{pn}^{\lambda,J} u_p v_n - Y_{pn}^{\lambda,J} v_p u_n \right) \right|^2
$$

Decay rate is of the form

$$
\lambda_i = D \int_1^{W_{0,i}} W \sqrt{W^2 - 1} \left( W_{0,i} - W \right)^2 F(Z,W) C(W) dW
$$

$$
T_{1/2} = \frac{\ln 2}{\lambda}, \qquad D = \frac{(G_F V_{ud})^2}{2\pi^3} \frac{(m_e c^2)^5}{\hbar}
$$

Allowed decay shape factor:

$$
C(W) = B(GT)
$$

First-forbidden transitions shape factor

$$
C(W) = k(1 + aW + bW^{-1} + cW^{2})
$$













G. Lorusso *et al.*, Phys. Rev. Lett. 114, 192501 (2015)



 $P$  ,  $P$  is a contribution of  $\mathcal{P}$  and  $\mathcal{P}$ 

Neutron number N

17 FEBRUARY 2017<br>17 February 2017<br>17 February 2017





Position of the peak around A  $\approx$  190 depends critically on the amount of Figure 14. (a): Final abundance distribution for a calculation using the T. Marketin (2015, in preparation) rates together with the FRDM mass model and ABLA07. neutrons available after freeze-out.  $\Delta$  referrence the FRDM, and the FRDM,  $\Delta$  included. (c): Same as (c): Same as (a),  $\Delta$ ,  $\Delta$ ,  $\Delta$ ,  $\Delta$ . Same as (a),  $\Delta$ , same as (a),  $\Delta$ , same as (a), but with the Panov et al. (c): Same as (a), but with the Pan

Latest calculations provide systematically shorter half-lives in the region of heavy nuclei – significant consequences for the r-process. dynamical ejecta of NSM within the treatment of Korobkin et al. systematically shorter half-lives consequences for the r-process.



The Astrophysical Journal,  $\overline{S}$  (13pp), 2015  $\overline{S}$  (13pp), 2015  $\overline{S}$  (13pp), 2015  $\overline{S}$ 

M. Eichler *et al.*, Astrophys. J. 808, 30 (2015)

## Beta-delayed neutron emission excitation energies higher than the neutron separation energy, these nuclei can decay by neutron emission as the can decay by neutron emission as  $\mathcal{C}$

In nuclei with small  $S_n$  an additional process is possible:



Beta-delayed neutron emission contributes neutrons at the late stages of the r-process, after the initial neutron flux has dissipated. decay calculations, including the first-forbiddens of the first-forbiddens of the first-forbiddens of the first-



R. Caballero-Folch *et al.*, Phys. Rev. Lett. 117, 012501 (2016)  $\mathbf{f}$ *Q<sup>n</sup>* and *Q* windows. The good agreement between the





**neutron number**

G. R. Keepin *et al.*, Phys. Rev. 107, 1044 (1957)

Beta-delayed neutron emission does affect the resulting abundance pattern.



Calculation based on the finite amplitude method (FAM) – a formulation of the QRPA which allows for a quick determination of the nuclear response.

24 40 <sup>64</sup>Cr 0*.*043

The interaction was also adjusted to dynamic properties of select nuclei – improved description of decay properties. 26 44 <sup>70</sup>Fe 0*.*094 Ine interaction was also 34 superior to a *j*<sup>1</sup> anno proportion or select nuclei – improved







 $h_1$  (2016) odd-*A* nuclei. Filled circles in the upper-left panel denote



QVC correlations push states towards the Fermi energy  $-$  enhancing the density of states in vicinity of the FE. the remin energy – ennanong me experimental binding energy difference (see the text); accordingly, the vertical dotted lines show the experimental value of

Excellent description of GT resonance in <sup>208</sup>Pb and half-lives of the Ni chain. panels: cumulative sum of the experimental values of 1/T1/2 indicated by the stars. The stars in the stars in the stars. The stars in the stars factor of 100 and 10, respectively.



 $\overline{1}$ <br>2016)  $\overline{1}$  Y F Niu et al Phys. Rev Lett 114 142501  $\begin{pmatrix} 1 & 1 & 1 \ 1 & 1 & 1 \end{pmatrix}$  is that it is based in the method in the method is based in the method in the method is based in 8. F. Niu et al., Phys. Rev. Lett. 114, 142501  $(15)$  $\sim$  1  $\sim$  /