

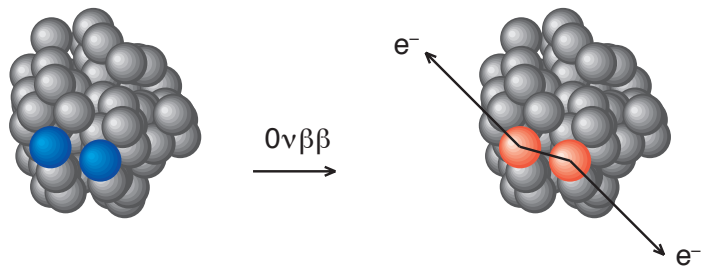
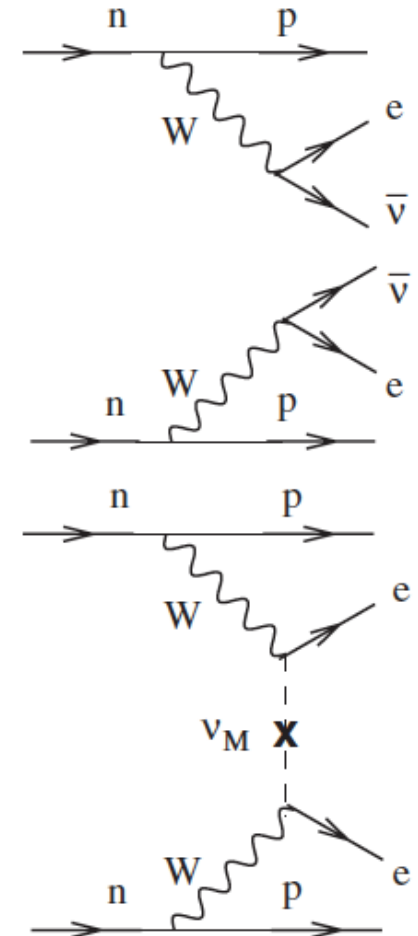
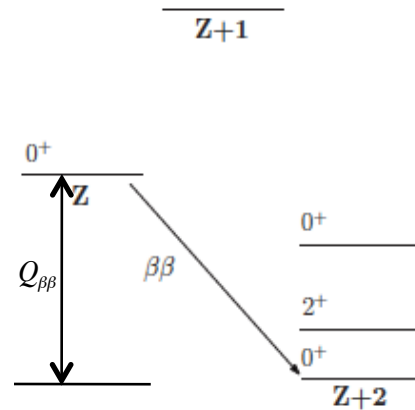
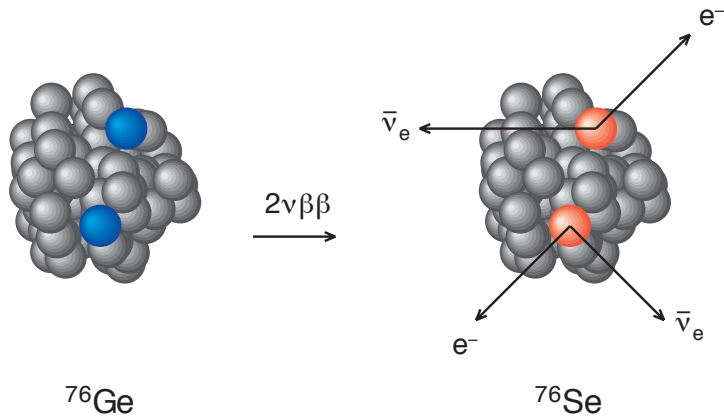
Nuclear matrix elements and the BSM energy scales in the EFT description of $0\nu\beta\beta$ decay

Mihai Horoi

Department of Physics, Central Michigan University, Mount Pleasant, Michigan 48859, USA

➤ Support from NSF grant PHY-1404442, DOE grants DE-SC0008529, and DE-SC0015376 is acknowledged

Classical Double Beta Decay Problem



$$\langle m_{\beta\beta} \rangle = \left| \sum_k m_k U_{ek}^2 \right|$$

$$T_{1/2}^{-1}(0\nu) = G^{0\nu} (Q_{\beta\beta}) \left[M^{0\nu}(0^+) \right]^2 \left(\frac{\langle m_{\beta\beta} \rangle}{m_e} \right)^2$$

$$|\nu_\alpha\rangle = \sum U_{\alpha i}^* |\nu_i\rangle$$

$$|\nu_i\rangle = \sum_\alpha U_{\alpha i} |\nu_\alpha\rangle$$

Neutrino Masses

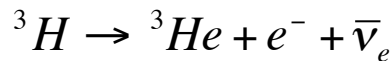


PMNS – matrix

$$U = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{bmatrix} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \begin{bmatrix} e^{i\alpha 1/2} & 0 & 0 \\ 0 & e^{i\alpha 2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$c_{12} \equiv \cos\theta_{12}$, $s_{12} = \sin\theta_{12}$, etc

- Tritium decay:



$$m_{\nu_e} = \sqrt{\sum_i |U_{ei}|^2 m_i^2} < 2.2 eV \text{ (Mainz exp.)}$$

KATRIN (to take data): goal $m_{\nu_e} < 0.3 eV$

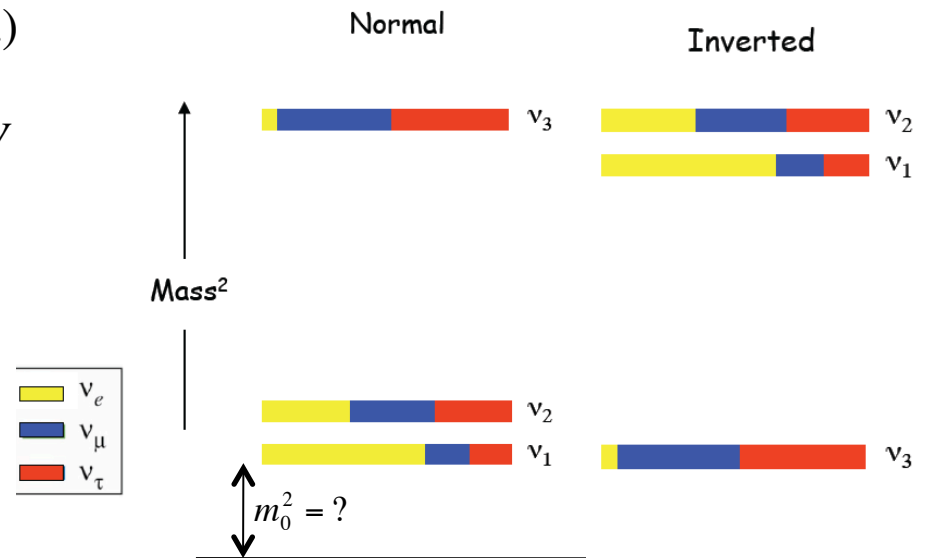
- Cosmology: CMB power spectrum, BAO, etc,

$$\sum_{i=1}^3 m_i < 0.23 eV$$

Goal: 0.01eV (5 – 10 y)

$$\Delta m_{21}^2 \approx 7.5 \times 10^{-5} eV^2 \text{ (solar)}$$

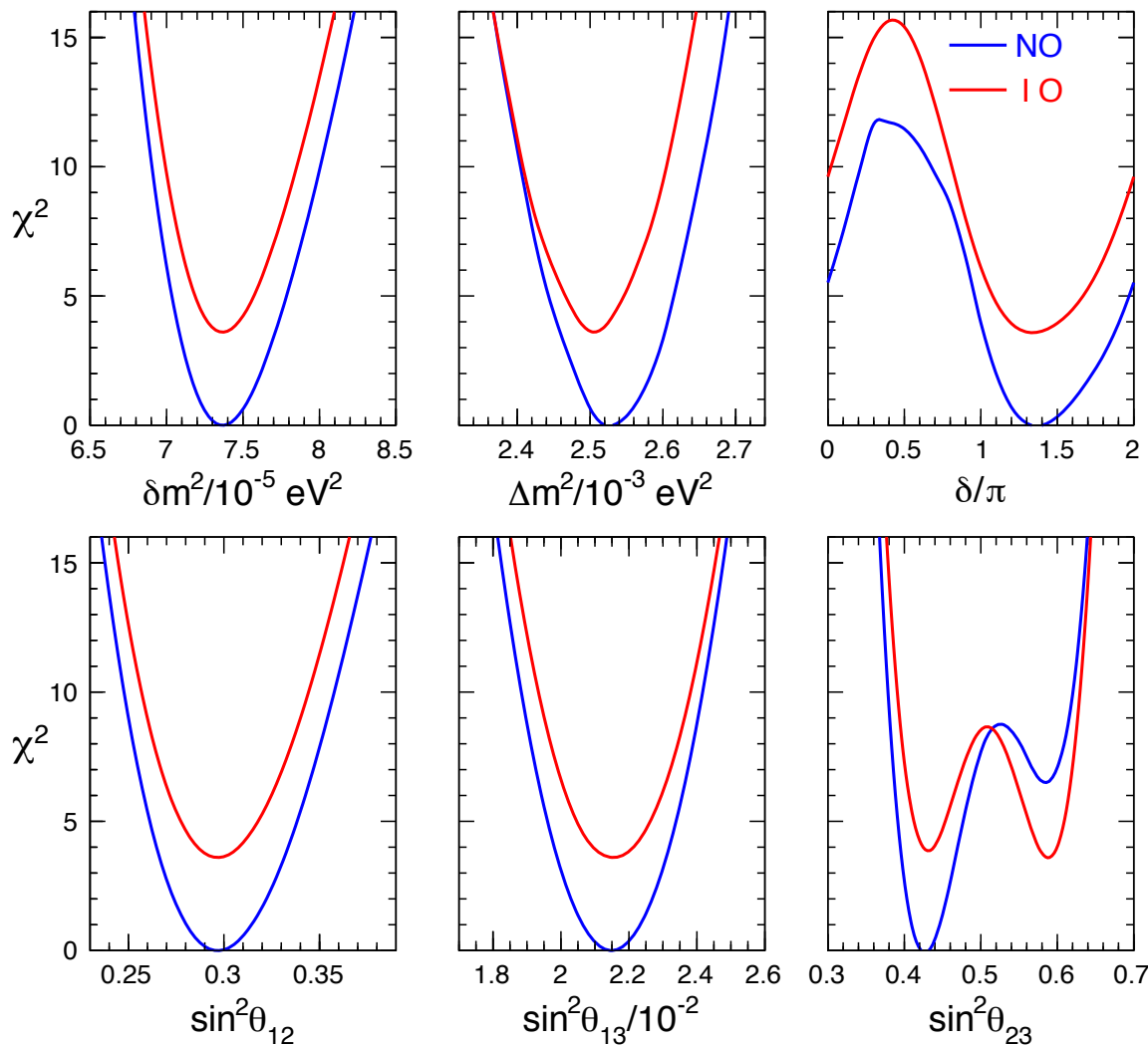
$$|\Delta m_{32}^2| \approx 2.4 \times 10^{-3} eV^2 \text{ (atmospheric)}$$



Two neutrino mass hierarchies

Neutrino oscillations parameters

Oscillation parameters



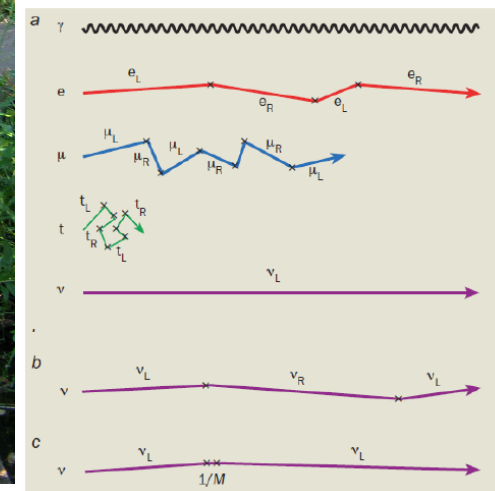
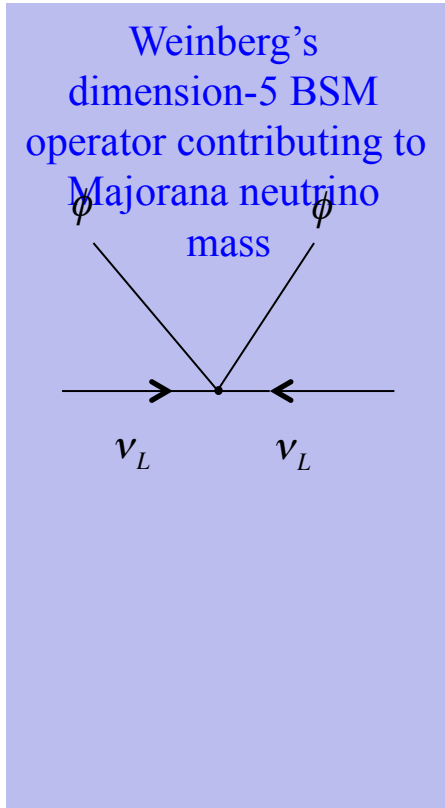
Bari group:

arxiv.org/1703.04471

$$(\Delta\chi^2_{\text{IO-NO}})^{1/2} = 2$$

Normal ordering favored at 2σ

Fork on the Road: Are Neutrinos Majorana or Dirac Fermions?



Best (Only?) Bet: Neutrinoless Double-Beta Decay.

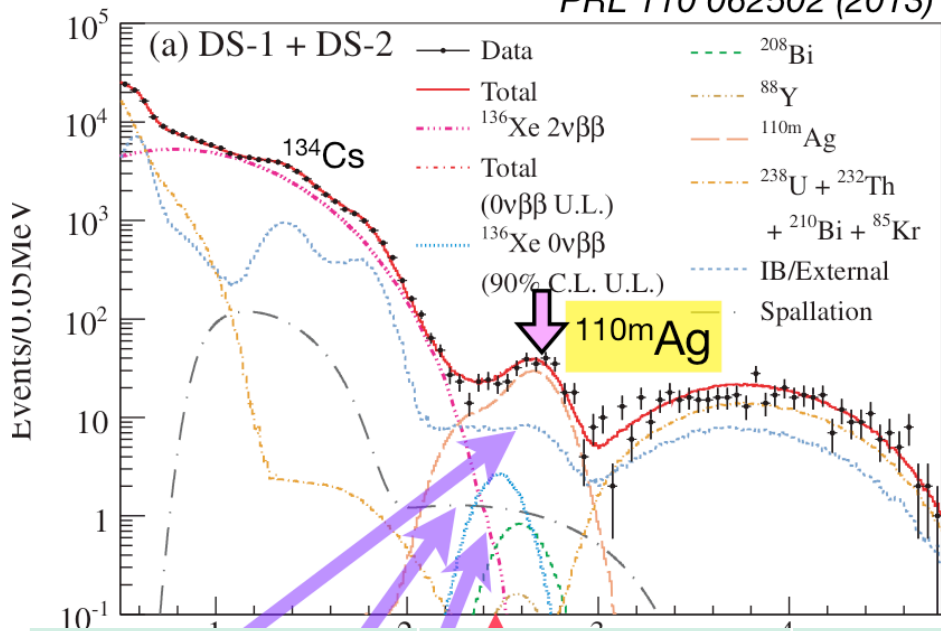
$$m_M = m_D m_N^{-1} m_D \quad (m_D \sim Y \langle \phi \rangle)$$

Dirac mass from the Yukawa

interaction with Higgs: $\bar{\psi}_i \gamma_{ij} \psi_j \langle \phi \rangle$

Phase 1 (before purification)

PRL 110 062502 (2013)



^{136}Xe $\beta\beta$ Experimental Results



	$T_{1/2}^{0\nu}(\text{lim})$	$T_{1/2}^{0\nu}(\text{Sens})$
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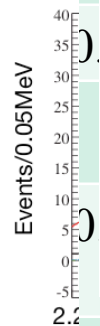
KamLAND-Zen

Half-life limit (@90% C.L.)

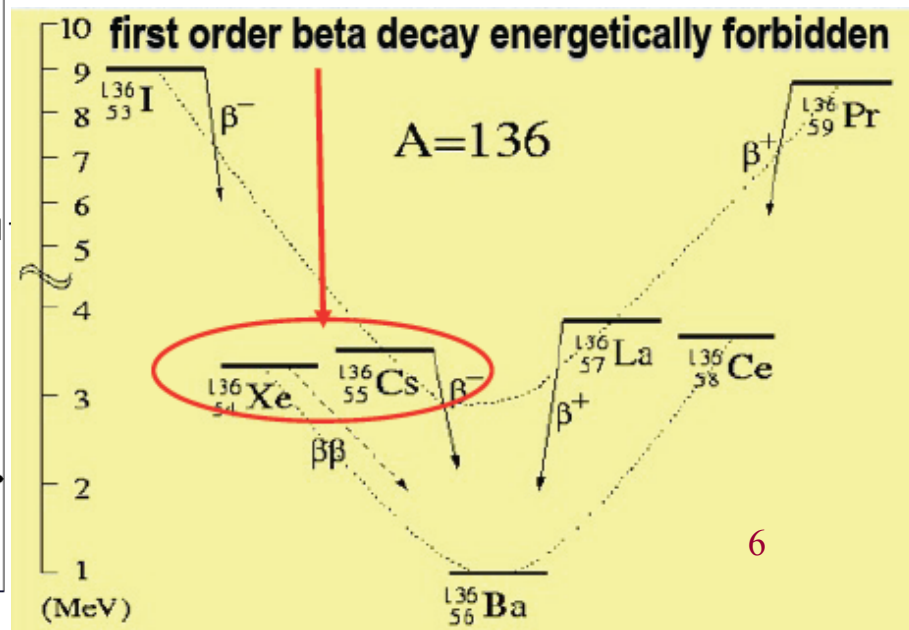
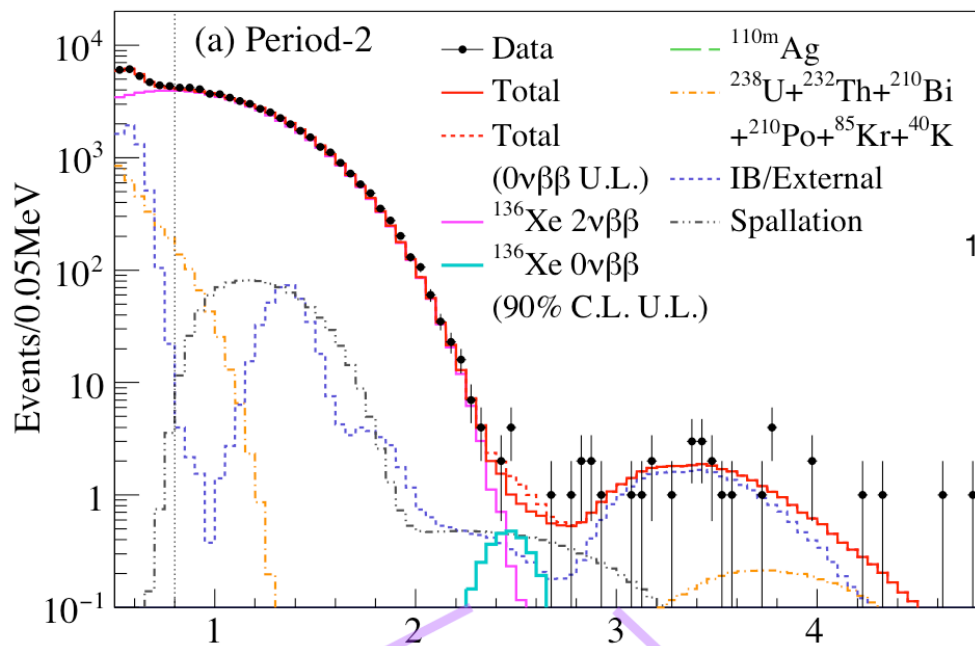
Phase 1 $T_{1/2}^{0\nu} > 1.9 \times 10^{25}$ yr

Phase 2 $T_{1/2}^{0\nu} > 9.2 \times 10^{25}$ yr **x6!**

Combined $T_{1/2}^{0\nu} > 1.07 \times 10^{26}$ yr



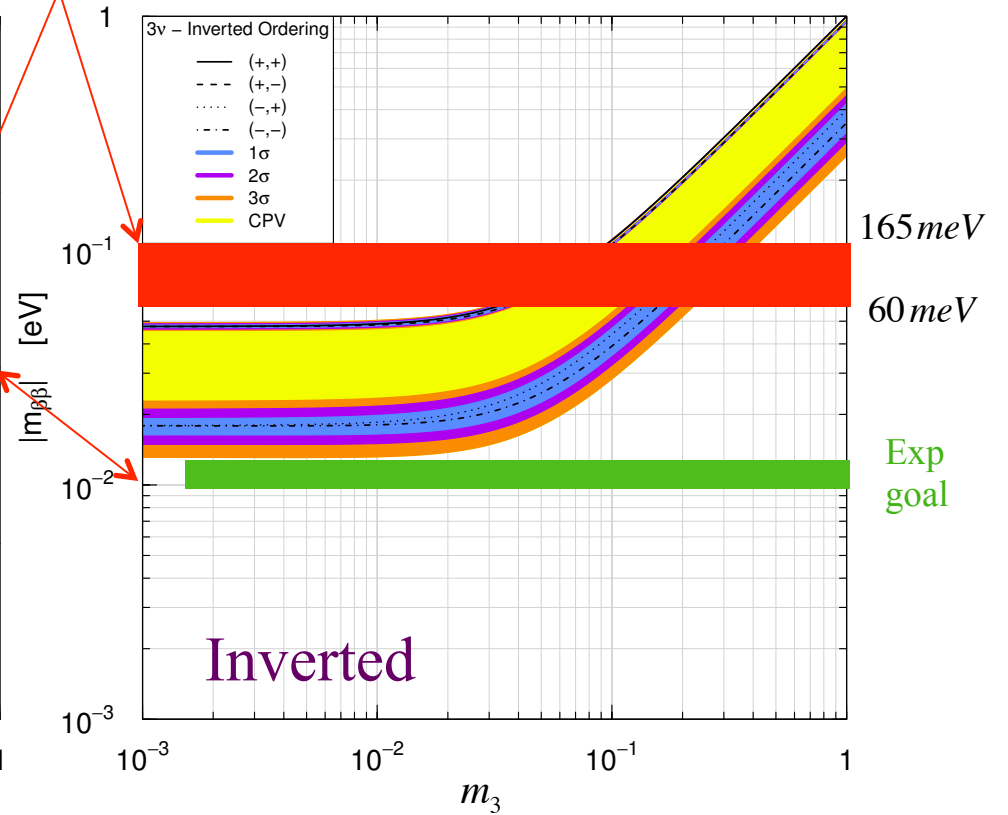
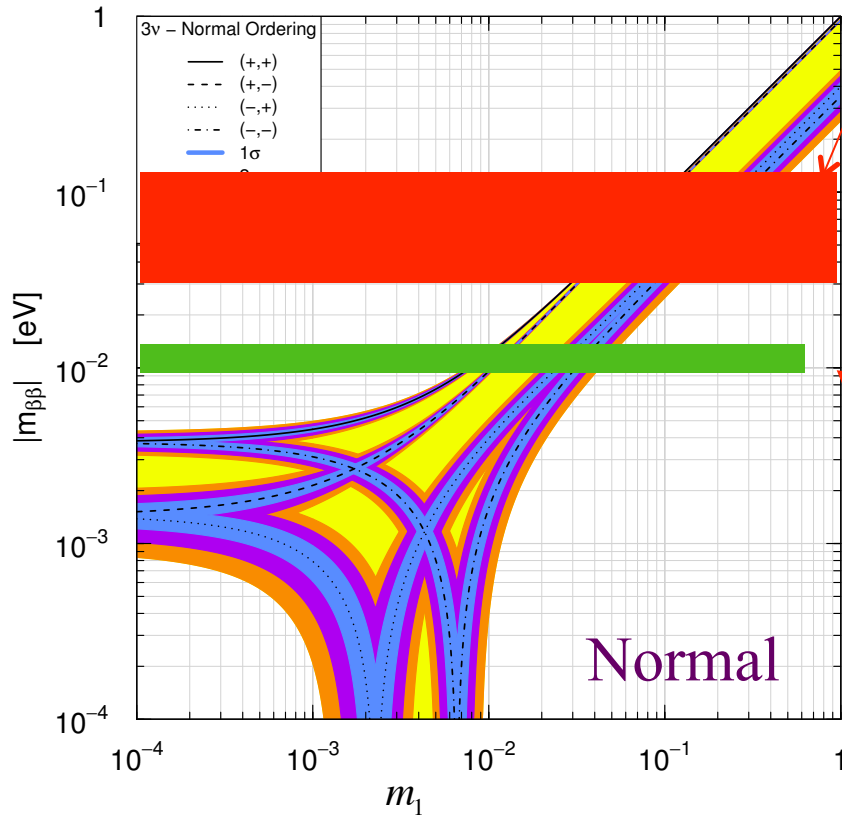
-0.0215 MeV^{-1}



Neutrino $\beta\beta$ effective mass

arxiv:1507.08204

KamLAND – Zen, PRL 117, 082503 (2016): ^{136}Xe



$$|m_{\beta\beta}| = \left| \sum_{k=1}^3 m_k U_{ek}^2 \right| = \left| c_{12}^2 c_{13}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right|$$

$$\phi_2 = \alpha_2 - \alpha_1 \quad \phi_3 = -\alpha_1 - 2\delta$$

$$\Leftrightarrow T_{1/2}^{-1}(0\nu) = G^{0\nu}(Q_{\beta\beta}) \left[M^{0\nu}(0^+) \right]^2 (\eta_{0\nu})^2$$

$$\eta_{0\nu} = \frac{|m_{\beta\beta}|}{m_e}$$

INT TC & 17-2a, June 21, 2017

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Shell Model Nuclear Matrix Elements

$$M_S^{0\nu} = \sum_{\substack{J, p < p' \\ n < n' \\ p < n}} (\Gamma) \left\langle 0_f^+ \left| \left[(a_p^+ a_{p'}^+)^J (\tilde{a}_n, \tilde{a}_n)^J \right]^0 \right| 0_i^+ \right\rangle \left\langle p p'; J \left| \int q^2 dq \left[\hat{S} \frac{h(q) j_\kappa(qr) G_{FS}^2 f_{SRC}^2}{q(q + \langle E \rangle)} \tau_{1-} \tau_{2-} \right] \right| n n'; J \right\rangle_{as} \quad - \text{closure}$$

Short range correlations (SRC): $f_{SRC} = 1 - c e^{ar^2} (1 - br^2)$

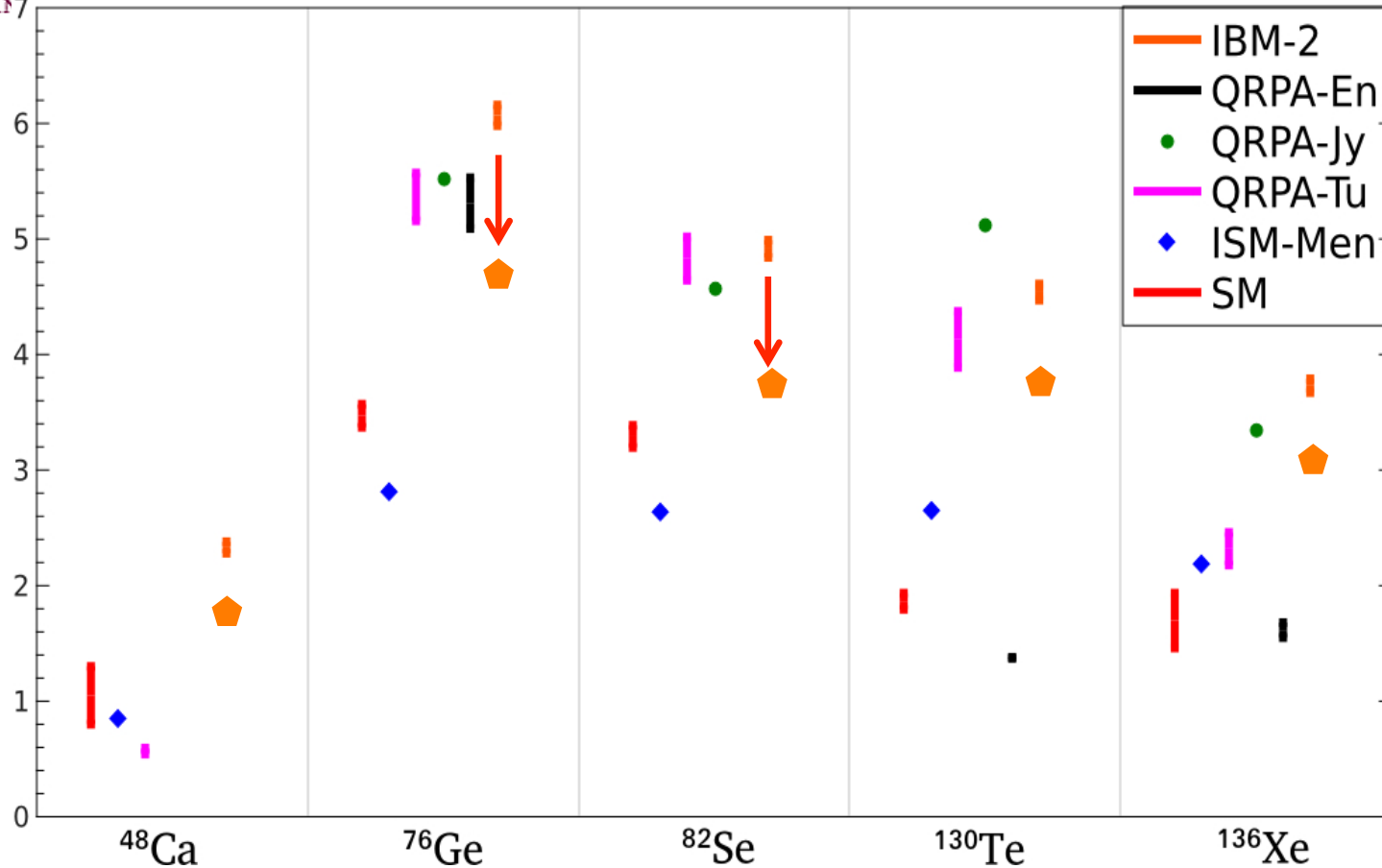
$$M^{0\nu} = M_{GT}^{0\nu} - (g_V / g_A)^2 M_F^{0\nu} + M_T^{0\nu}$$

$$\hat{S} = \begin{cases} \sigma_1 \tau_1 \sigma_2 \tau_2 & \text{Gamow - Teller (GT)} \\ \tau_1 \tau_2 & \text{Fermi (F)} \\ [3(\vec{\sigma}_1 \cdot \hat{n})(\vec{\sigma}_2 \cdot \hat{n}) - (\vec{\sigma}_1 \cdot \vec{\sigma}_2)] \tau_1 \tau_2 & \text{Tensor (T)} \end{cases}$$

TABLE II. Parameters for the short-range correlation (SRC) parametrization of Eq. (11).

SRC		a	b	c
MS SRC	Miller-Spencer	1.10	0.68	1.00
CDB SRC	CD-Bonn	1.52	1.88	0.46
AV18 SRC	AV18	1.59	1.45	0.92

NME for the light-neutrino exchange mechanism



IBA-2 J. Barea, J. Kotila, and F. Iachello, Phys. Rev. C **87**, 014315 (2013). **→** **IBM-2** PRC **91**, 034304 (2015)

QRPA-En M. T. Mustonen and J. Engel, Phys. Rev. C **87**, 064302 (2013).

QRPA-Jy J. Suhonen, O. Civitarese, Phys. NPA **847** 207–232 (2010).

QRPA-Tu A. Faessler, M. Gonzalez, S. Kovalenko, and F. Simkovic, arXiv:**1408.6077**

ISM-Men J. Menéndez, A. Poves, E. Caurier, F. Nowacki, NPA **818** 139–151 (2009).

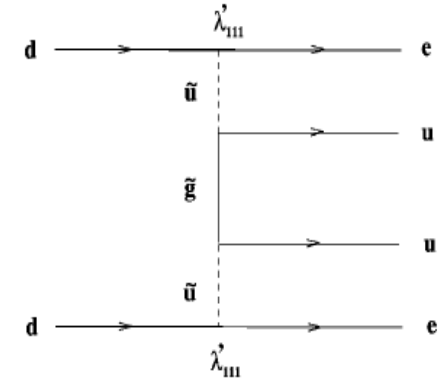
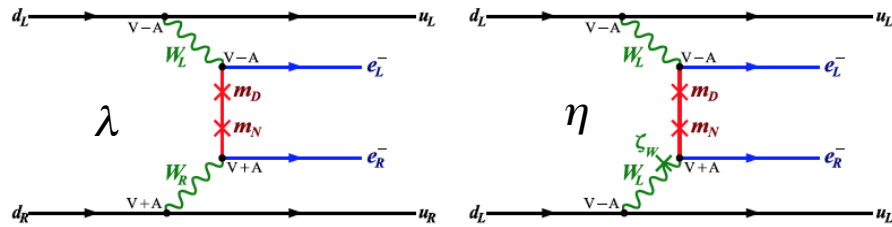
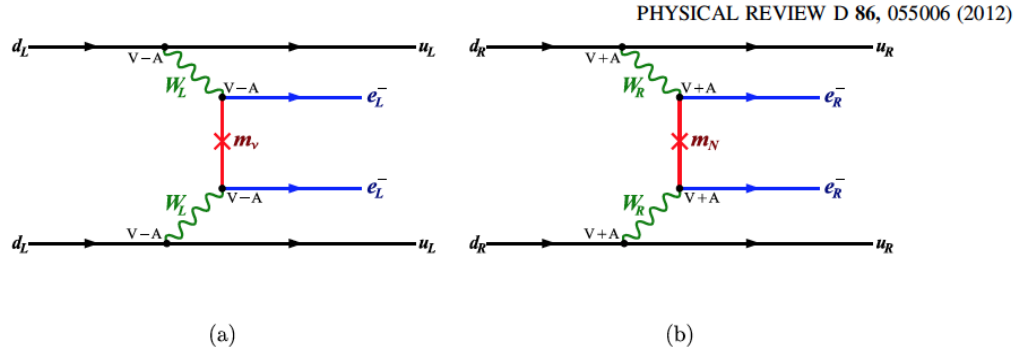
SM M. Horoi et. al. PRC **88**, 064312 (2013), PRC **89**, 045502 (2014), PRC **89**, 054304 (2014), PRC **90**, 051301(R) (2014), PRC **91**, 024309 (2015), PRL **110**, 222502 (2013), PRL **113**, 262501(2014).

INT TC & 17-2a, June
21, 2017

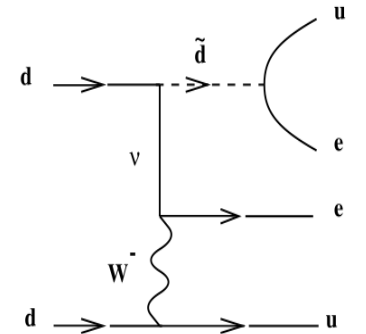
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Other models: Left-Right symmetric model and SUSY R-parity violation

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Gluino exchange



Squark exchange

$$\left[T_{1/2}^{0\nu} \right]^{-1} = G_{01} g_A^4 \left| \eta_{0\nu} M_{0\nu} + (\eta_{N_R}^L + \eta_{N_R}^R) M_{0N} + \eta_{\tilde{q}} M_{\tilde{q}} + \eta_{\lambda'} M_{\lambda'} + \eta_\lambda X_\lambda + \eta_\eta X_\eta \right|^2.$$

(e)

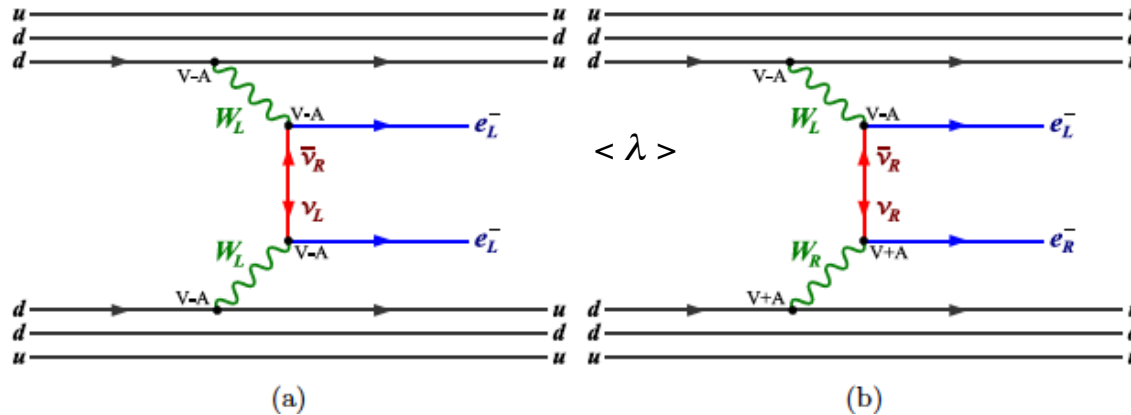
M. Horoi, A. Neacsu, PRD 93, 113014 (2016)

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DBD signals from different mechanisms

R. Arnold et al.: Probing New Physics Models of Neutrinoless Double Beta Decay with SuperNEMO

arXiv:1005.1241



$$\left[T_{1/2}^{0\nu} \right]^{-1} = \left| M_{GT}^{(0\nu)} \right|^2 \left\{ C_{\nu^2} + C_{\nu\lambda} \cos\phi_1 + C_{\nu\eta} \cos\phi_2 + C_{\lambda^2} + C_{\eta^2} + C_{\lambda\eta} \cos(\phi_1 - \phi_2) \right\},$$

$$\frac{d^2 W_{0^+ \rightarrow 0^+}^{0\nu}}{d\epsilon_1 d\cos\theta_{12}} = \frac{a_{0\nu} \omega_{0\nu}(\epsilon_1)}{2(m_e R)^2} [A(\epsilon_1) + B(\epsilon_1) \cos\theta_{12}]$$

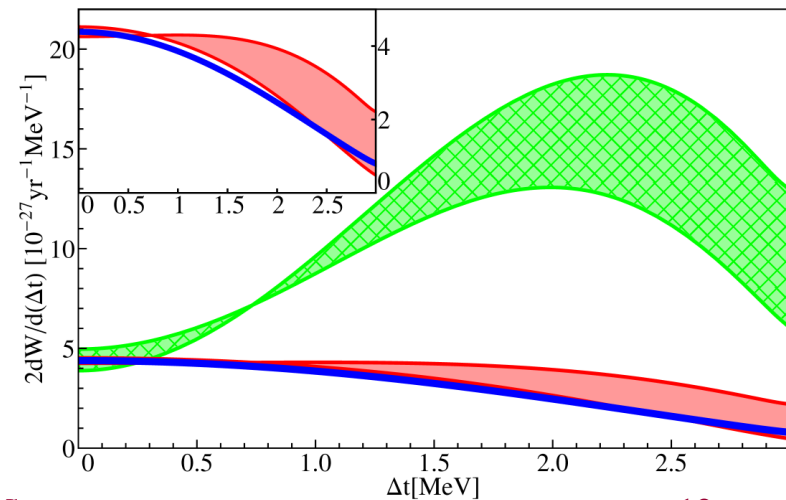
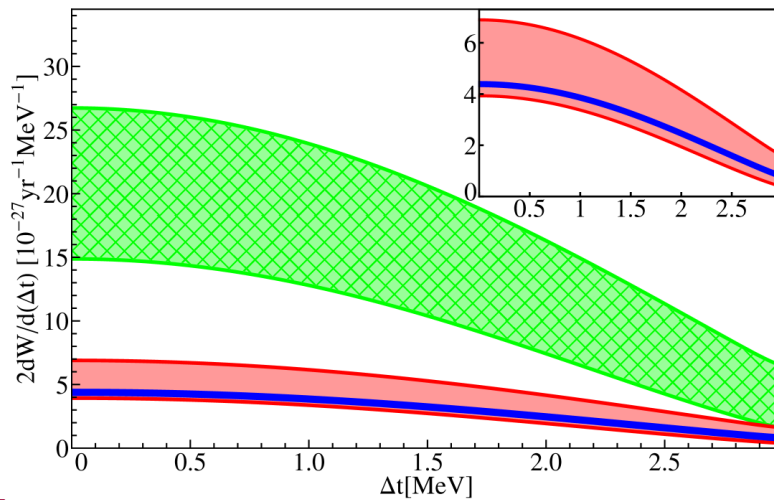
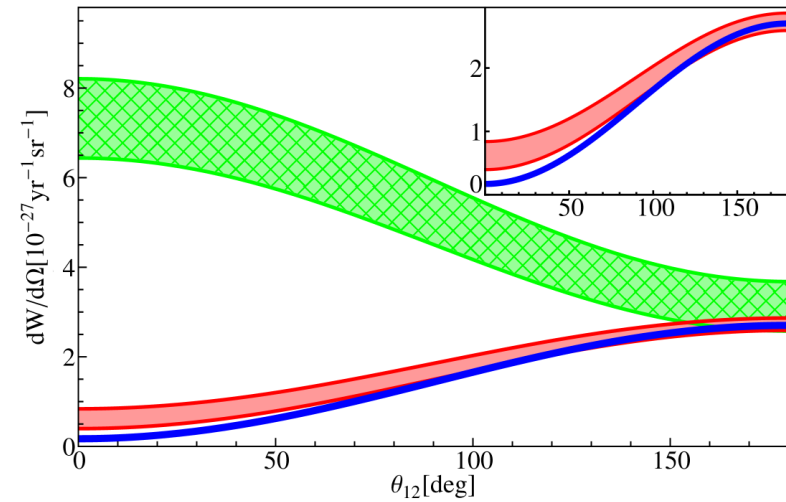
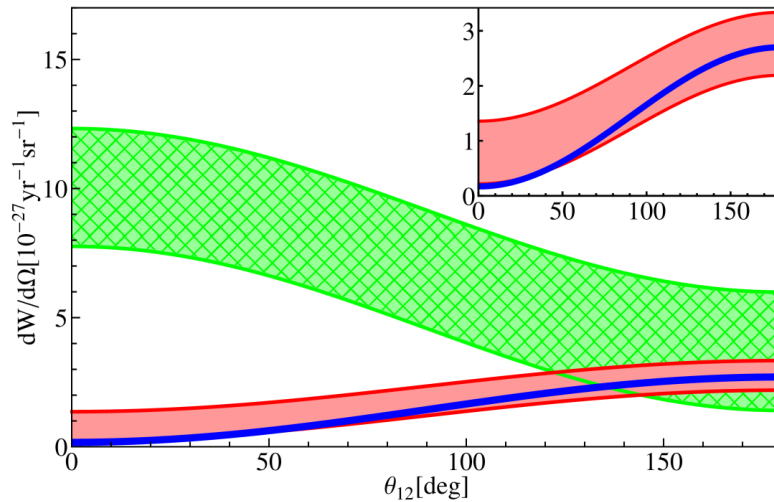
$$\frac{2dW_{0^+ \rightarrow 0^+}^{0\nu}}{d(\Delta t)} = \frac{2a_{0\nu}}{(m_e R)^2} \frac{\omega_{0\nu}(\Delta t)}{m_e c^2} A(\Delta t)$$

$$t = \epsilon_{e1} - \epsilon_{e2}$$

λ and η mechanisms (^{82}Se): look for green

$\langle \lambda \rangle$ dominates

$\langle \eta \rangle$ dominates



Two Non-Interfering Mechanisms

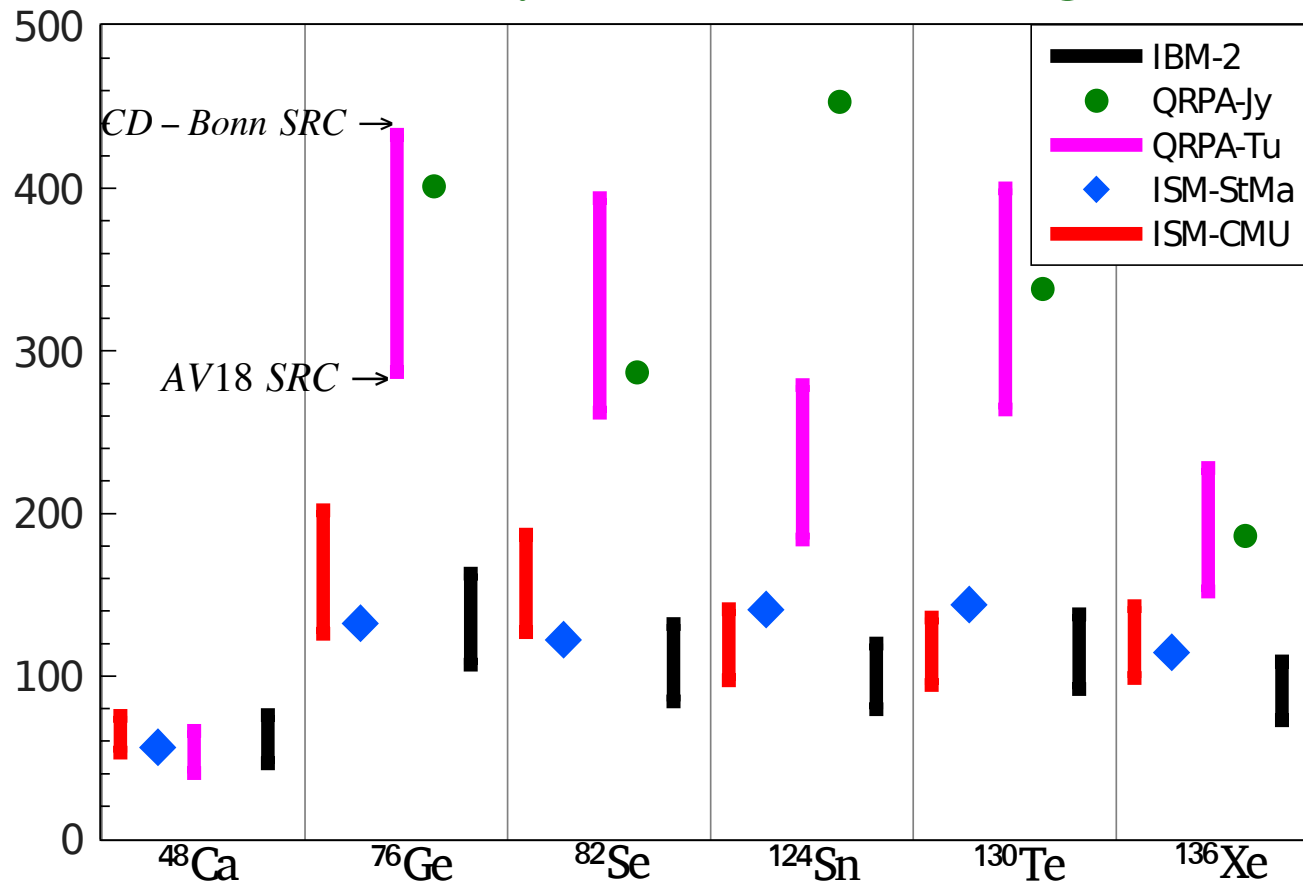
$$r(\nu / N) \equiv T_{1/2}^{\nu/N}(1) / T_{1/2}^{\nu/N}(2) = \frac{G_{01}^{0\nu}(2) |M^{0\nu/N}(2)|^2}{G_{01}^{0\nu}(1) |M^{0\nu/N}(1)|^2}$$

	Ge/Se		Ge/Te		Ge/Xe		Se/Te		Se/Xe		Te/Xe	
	Ge	Se	Ge	Te	Ge	Xe	Se	Te	Se	Xe	Te	Xe
$G_{01}^{0\nu} \times 10^{14}$	0.237	1.018	0.237	1.425	0.237	1.462	1.018	1.425	1.018	1.462	1.425	1.462
$M^{0\nu}(1/2)$	3.57	3.39	3.57	1.93	3.57	1.76	3.39	1.93	3.39	1.76	1.93	1.76
$M^{0N}(1/2)$	202	187	202	136	202	143	187	136	187	143	136	143
$T_{1/2}^{\nu}(1)/T_{1/2}^{\nu}(2)$	3.87		1.76		1.50		0.45		0.39		0.85	
$T_{1/2}^N(1)/T_{1/2}^N(2)$	3.68		2.73		3.09		0.74		0.84		1.13	
$R(N/\nu)$ present	0.95		1.55		2.06		1.63		2.17		1.33	
$R(N/\nu)$ [45]	1.02		1.39		1.42		1.36		1.39		1.03	

$$R(N / \nu) = r(N) / r(\nu)$$

Heavy neutrino-exchange NME

M_{0N}



IBA-2 J. Barea, J. Kotila, and F. Iachello, Phys. Rev. C **87**, 014315 (2013).

QRPA-Tu A. Faessler, M. Gonzalez, S. Kovalenko, and F. Simkovic, arXiv:1408.6077.

QRPA-Jy J. Hivarynen and J. Suhonen, PRC **91**, 024613 (2015), **ISM-StMa** J. Menendez, private communication.

ISM-CMU M. Horoi et. al. PRC **88**, 064312 (2013), PRC **90**, PRC **89**, 054304 (2014), PRC **91**, 024309 (2015), PRL **110**, 222502 (2013).

Towards an effective $0\nu\text{DBD}$ operator

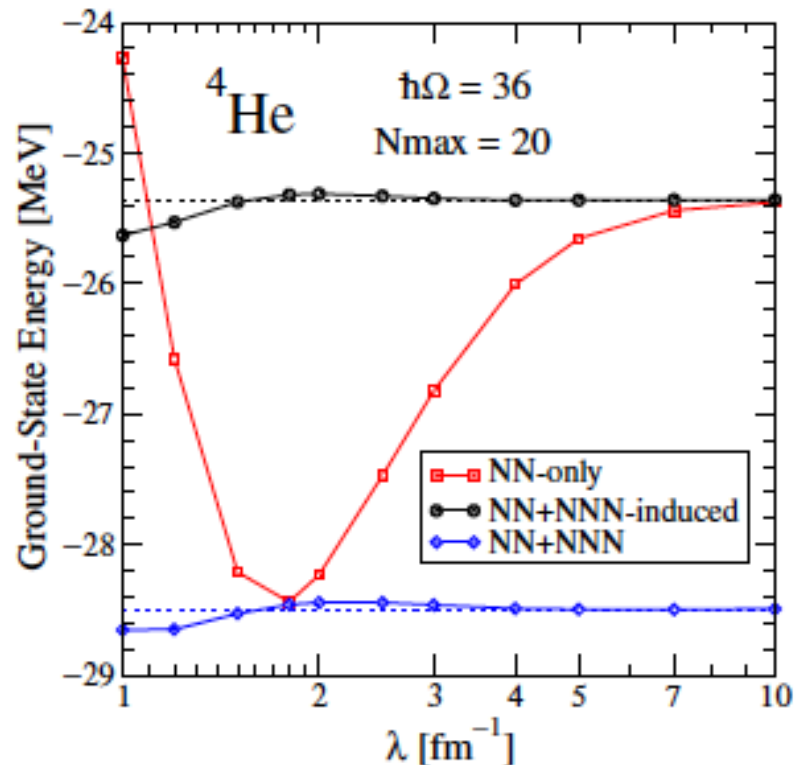
Similarity Renormalization Group (SRG) evolution

$$H_\lambda = U_\lambda H_{\lambda=\infty} U_\lambda^\dagger$$

$$\frac{dH_\lambda}{d\lambda} = -\frac{4}{\lambda^5} [[G, H_\lambda], H_\lambda]$$

$$O_\lambda = U_\lambda O_{\lambda=\infty} U_\lambda^\dagger$$

N3LO 500

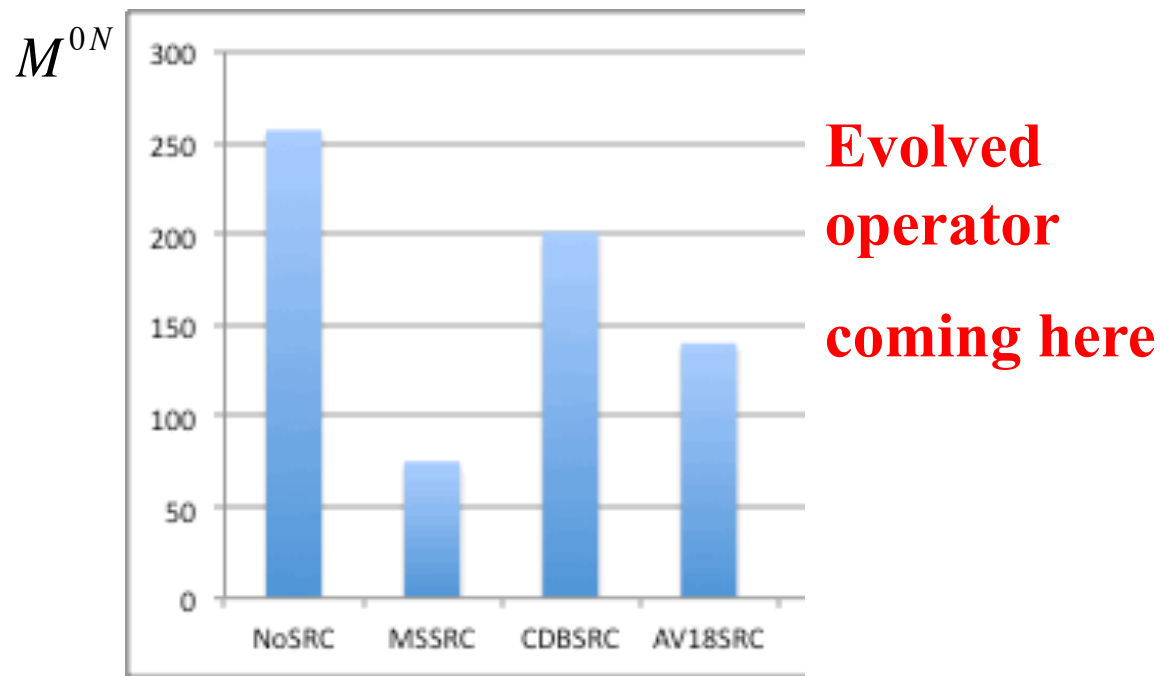


arXiv:1302.5473

Towards an effective $0\nu\text{DBD}$ operator: heavy neutrino-exchange NME

$$O_\lambda = U_\lambda O_{\lambda=\infty} U_\lambda^\dagger$$

^{76}Ge

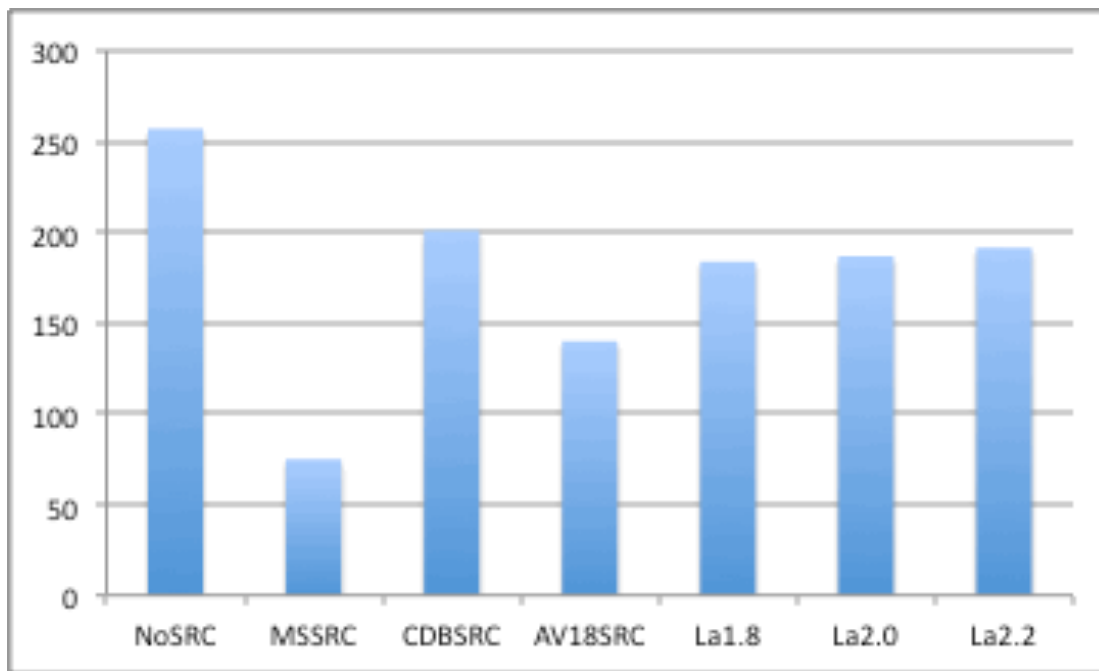


Towards an effective 0ν DBD operator: heavy neutrino-exchange NME

$$O_{\lambda} = U_{\lambda} O_{\lambda=\infty} U_{\lambda}^{\dagger}$$

^{76}Ge

M^{0N}

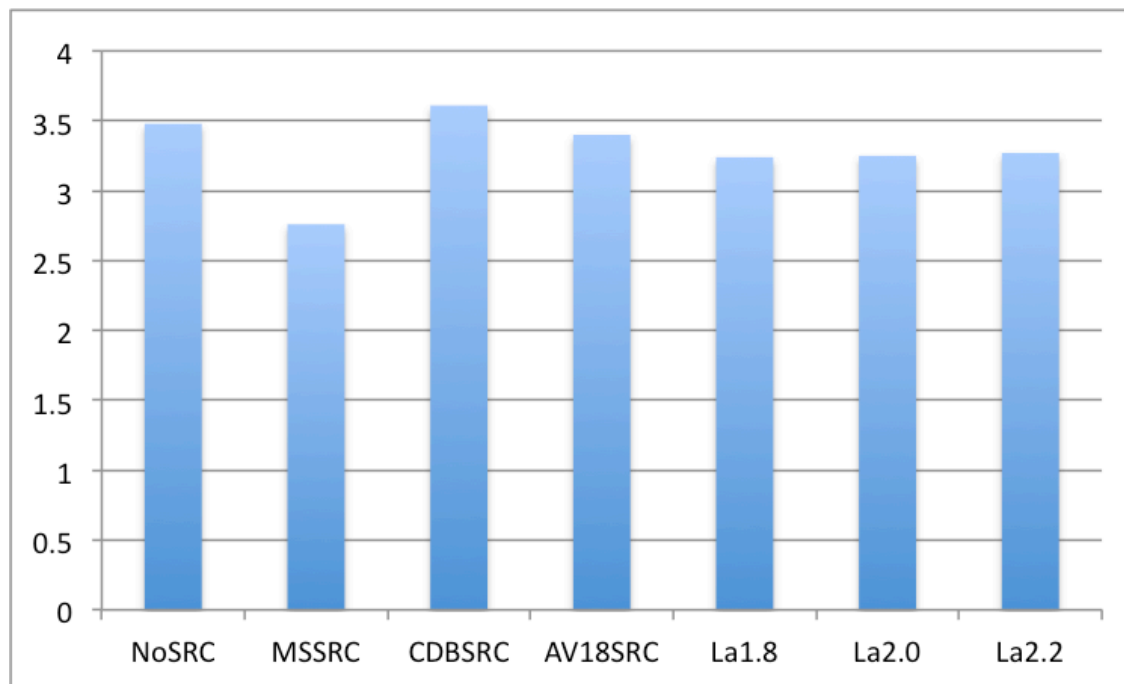


Towards an effective 0ν DBD operator: light neutrino-exchange NME

$$O_{\lambda} = U_{\lambda} O_{\lambda=\infty} U_{\lambda}^{\dagger}$$

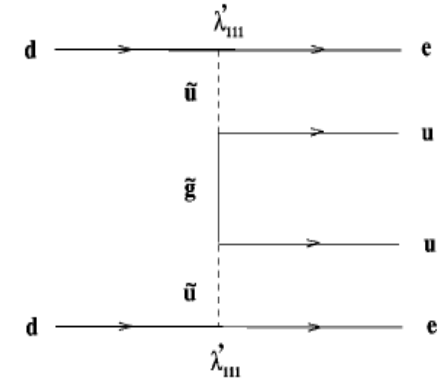
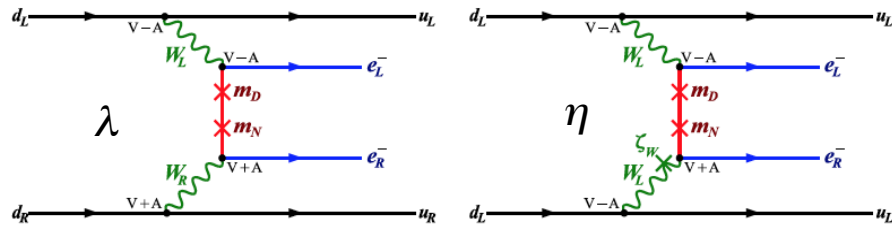
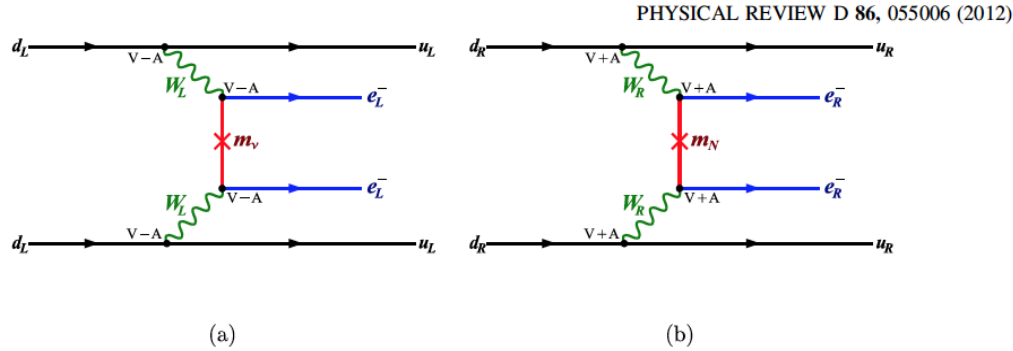
^{76}Ge

$M^{0\nu}$

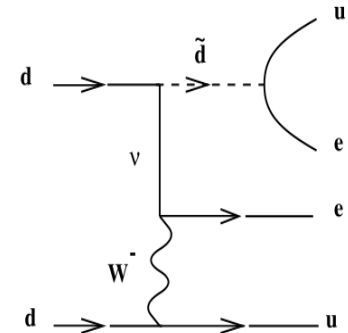


Other models: Left-Right symmetric model and SUSY R-parity violation

DAS *et al.*



Gluino exchange



Squark exchange

$$\left[T_{1/2}^{0\nu} \right]^{-1} = G_{01} g_A^4 \left| \eta_{0\nu} M_{0\nu} + (\eta_{N_R}^L + \eta_{N_R}^R) M_{0N} + \eta_{\tilde{q}} M_{\tilde{q}} + \eta_{\lambda'} M_{\lambda'} + \eta_{\lambda} X_{\lambda} + \eta_{\eta} X_{\eta} \right|^2.$$

(e)

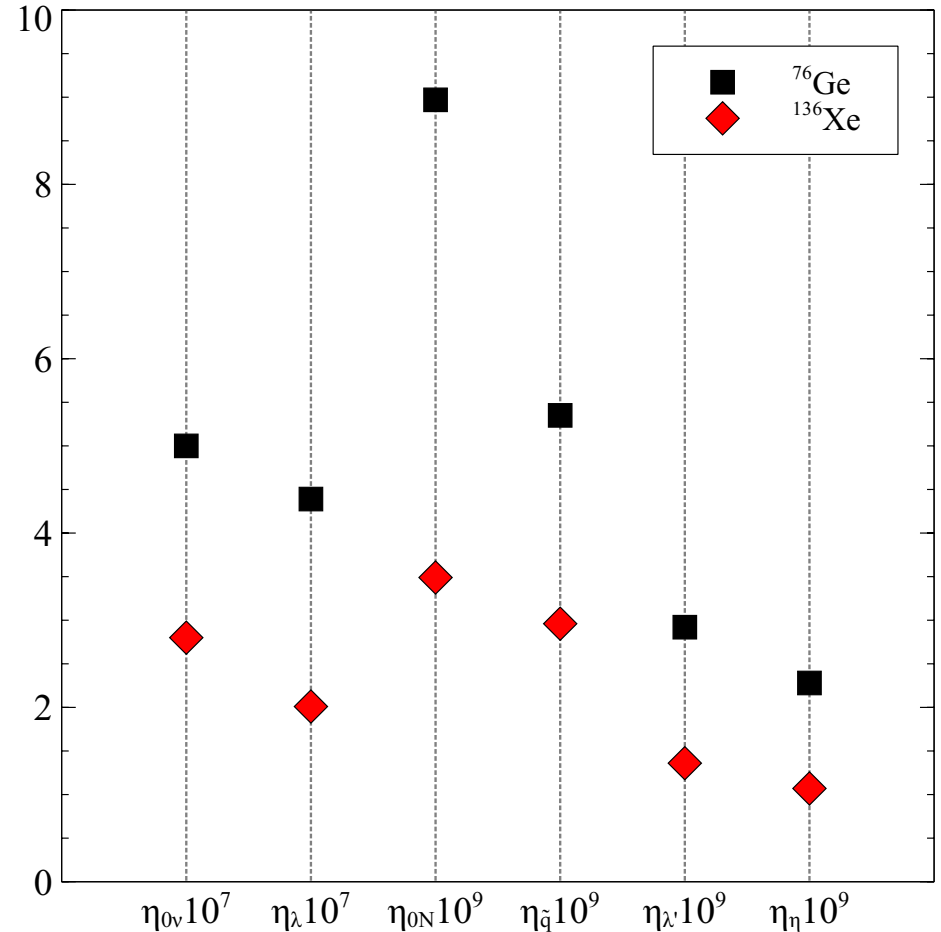
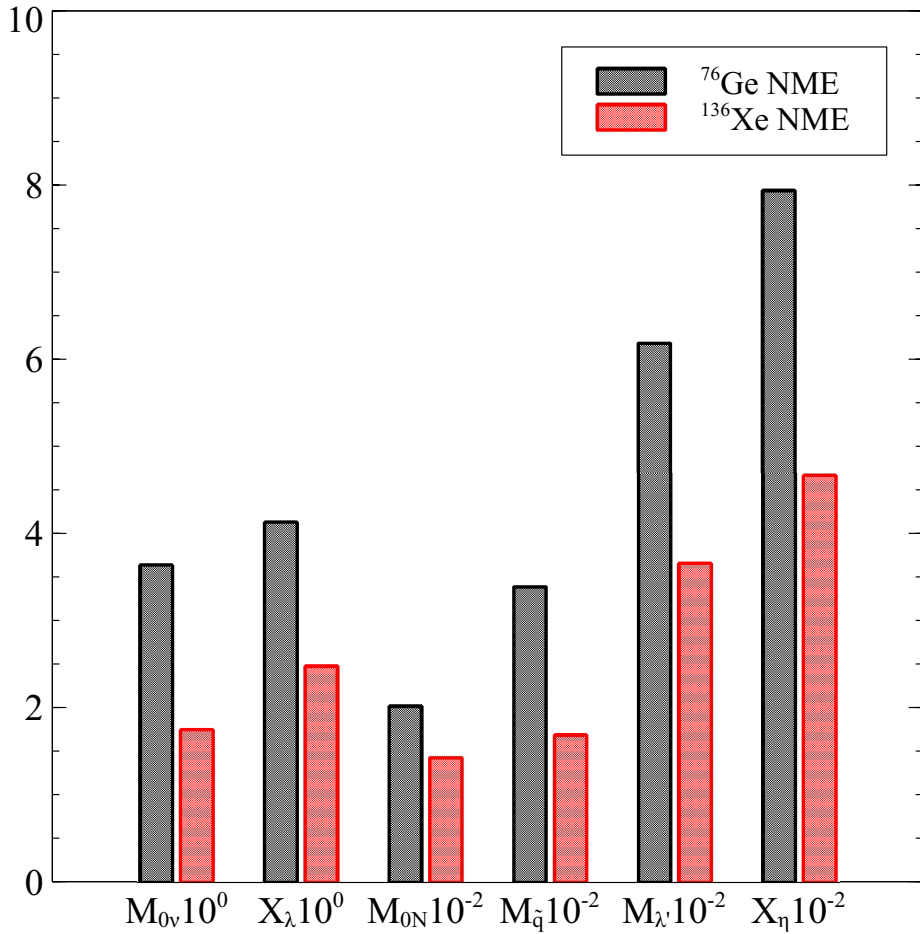
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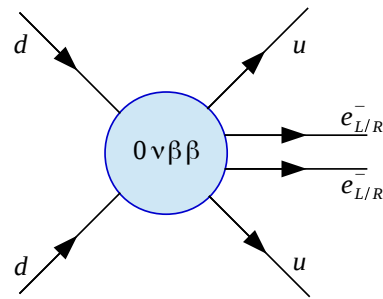
One mechanism dominance

$$\left[T_{1/2}^{0\nu} \right]^{-1} = G_{01} g_A^4 \left| \eta_{0\nu} M_{0\nu} + (\eta_{N_R}^L + \eta_{N_R}^R) M_{0N} + \eta_{\tilde{q}} M_{\tilde{q}} + \eta_{\lambda'} M_{\lambda'} + \eta_{\lambda} X_{\lambda} + \eta_{\eta} X_{\eta} \right|^2.$$

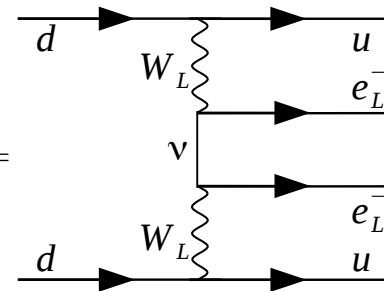
$T_{1/2}^{0\nu}({}^{76}\text{Ge}) > 5 \times 10^{25}$ years $T_{1/2}^{0\nu}({}^{136}\text{Xe}) > 1.1 \times 10^{26}$ years



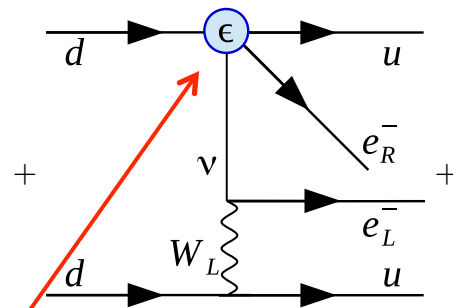
Effective field theory approach



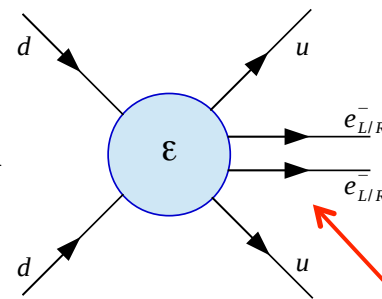
(a) The generic $0\nu\beta\beta$ decay diagram at the quark-level.



(b) Light left-handed neutrino exchange diagram.



(c) The long-range part of the $0\nu\beta\beta$ diagram.

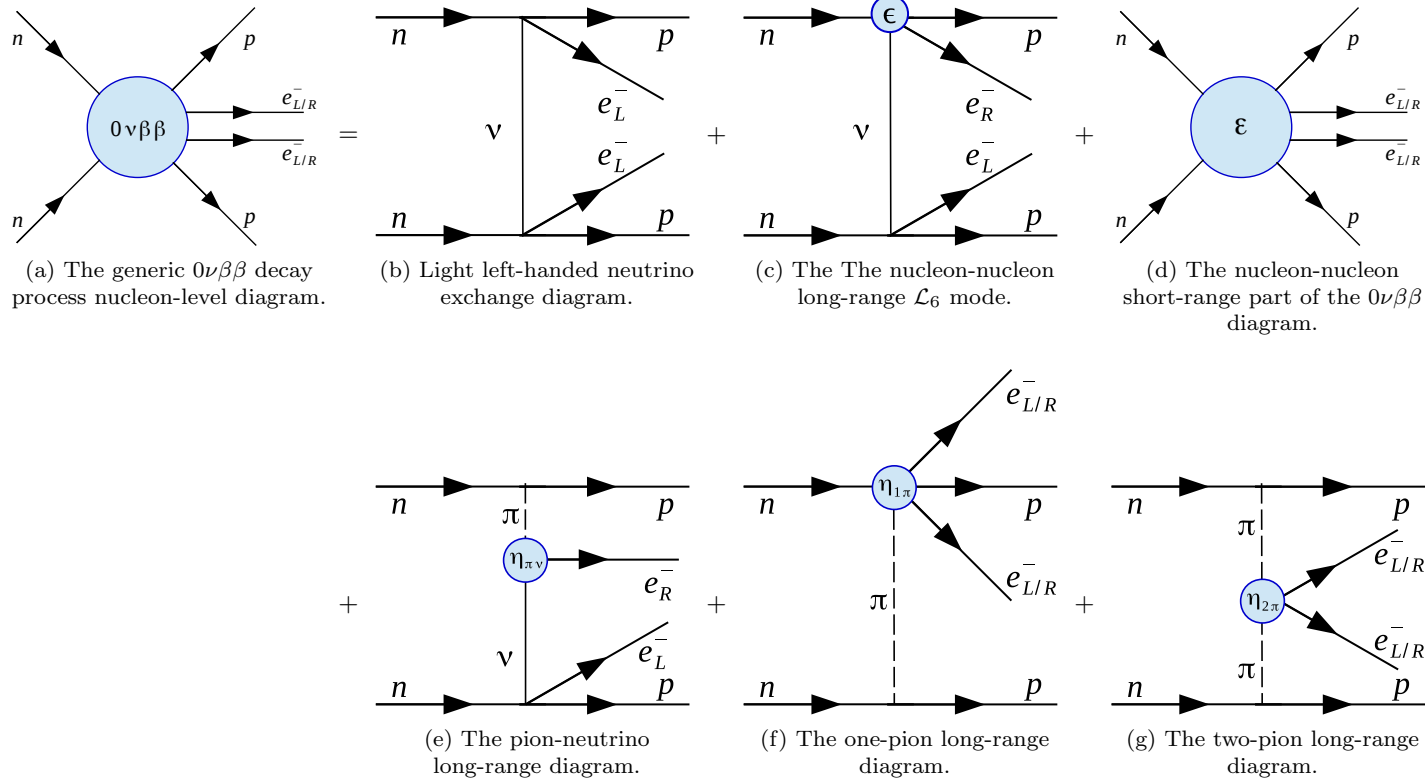


(d) The short-range part of the $0\nu\beta\beta$ diagram.

$$\mathcal{L}_6 = \frac{G_F}{\sqrt{2}} \left[j_{V-A}^\mu J_{V-A,\mu}^\dagger + \sum_{\alpha,\beta}^* \epsilon_\alpha^\beta j_\beta J_\alpha^\dagger \right]$$

$$\mathcal{L}_9 = \frac{G_F^2}{2m_p} \left[\epsilon_1 J J j + \epsilon_2 J^{\mu\nu} J_{\mu\nu} j + \epsilon_3 J^\mu J_\mu j \right. \\ \left. + \epsilon_4 J^\mu J_{\mu\nu} j^\nu + \epsilon_5 J^\mu J j_\mu \right],$$

Effective field theory after hadronization



$$\left[T_{1/2}^{0\nu} \right]^{-1} = g_A^4 \left[\sum_i |\mathcal{E}_i|^2 \mathcal{M}_i^2 + \text{Re} \left[\sum_{i \neq j} \mathcal{E}_i \mathcal{E}_j \mathcal{M}_{ij} \right] \right]$$

$$\mathcal{E}_{2-7} = \{ \epsilon_{V-A}^{V+A}, \epsilon_{V+A}^{V+A}, \epsilon_{S \pm P}^{S+P}, \epsilon_{TL}^{TR}, \epsilon_{TR}^{TR}, \eta_{\pi\nu} \}$$

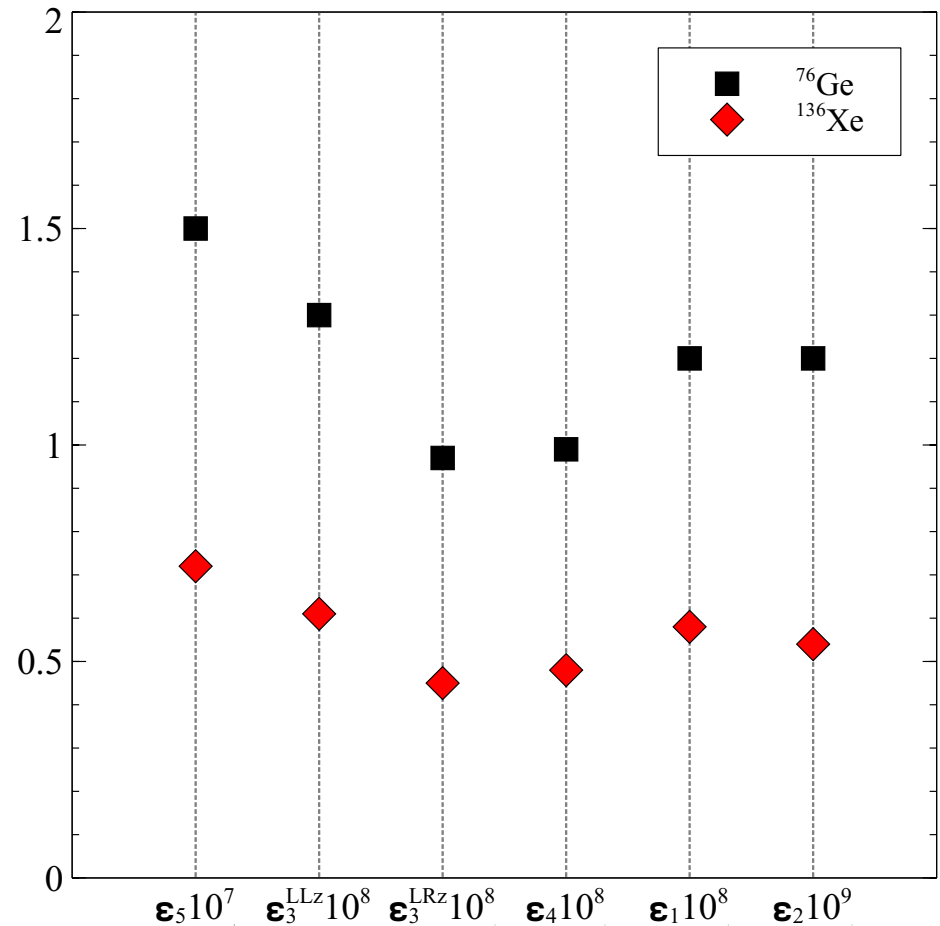
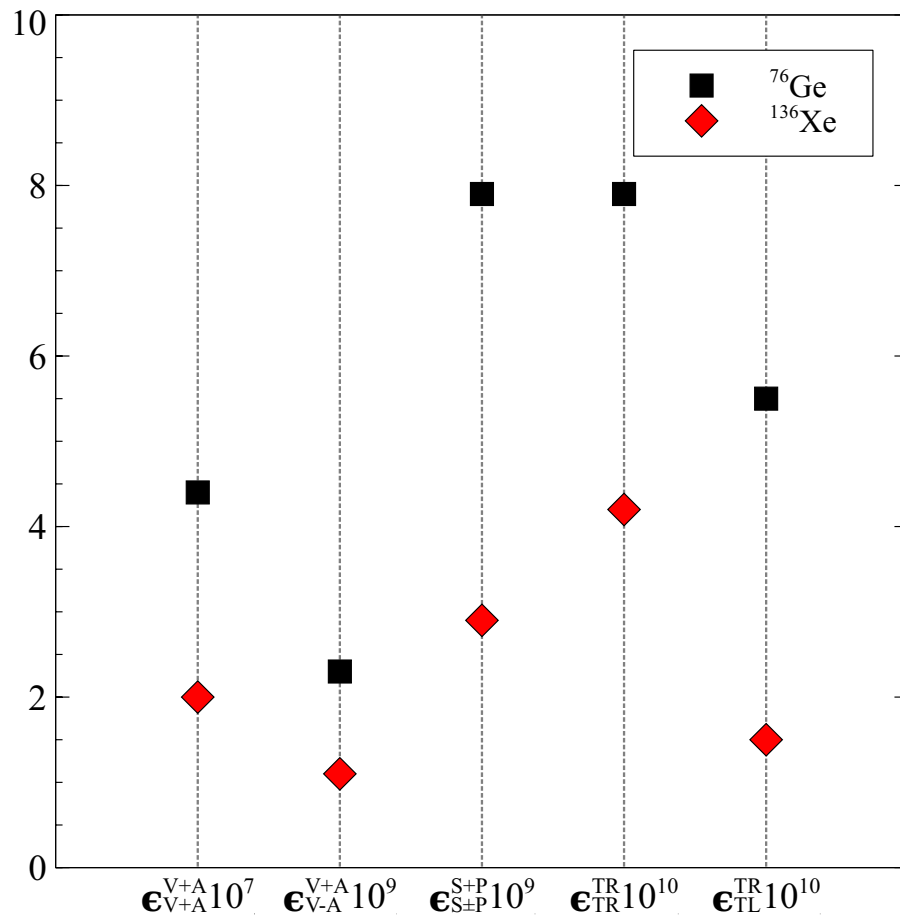
$$\mathcal{E}_{8-15} = \{ \epsilon_1, \epsilon_2, \epsilon_3^{LLz(RRz)}, \epsilon_3^{LRz(RLz)}, \epsilon_4, \epsilon_6, \eta_{1\pi}, \eta_{2\pi} \}$$

One coupling dominance

$$[T_{1/2}^{0\nu}]^{-1} = g_A^4 \left[\sum_i |\mathcal{E}_i|^2 \mathcal{M}_i^2 + \text{Re} \left[\sum_{i \neq j} \mathcal{E}_{ij} \mathcal{M}_{ij} \right] \right]$$

$$\mathcal{E}_{2-7} = \{ \epsilon_{V-A}^{V+A}, \epsilon_{V+A}^{V+A}, \epsilon_{S\pm P}^{S+P}, \epsilon_{TL}^{TR}, \epsilon_{TR}^{TR}, \eta_{\pi\nu} \}$$

$$\mathcal{E}_{8-15} = \{ \epsilon_1, \epsilon_2, \epsilon_3^{LLz(RRz)}, \epsilon_3^{LRz(RLz)}, \epsilon_4, \epsilon_6, \eta_{1\pi}, \eta_{2\pi} \}$$



INT TC & 17-2a, June 21, 2017

M. Horoi CMU

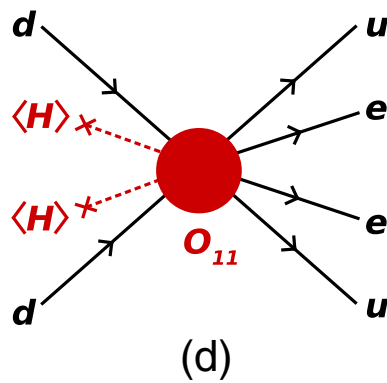
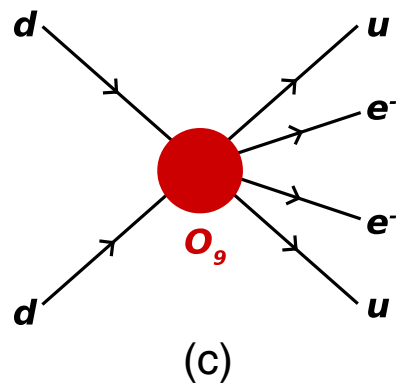
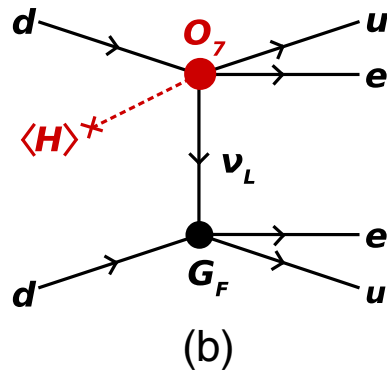
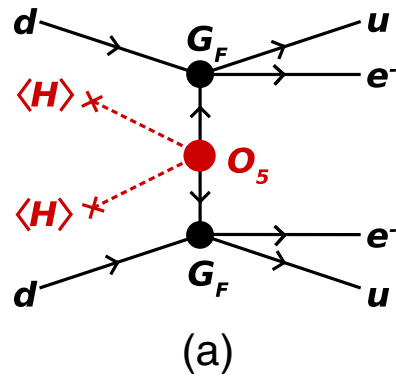
$T_{1/2}^{0\nu} (^{76}\text{Ge}) > 5 \times 10^{25}$ years $T_{1/2}^{0\nu} (^{136}\text{Xe}) > 1.1 \times 10^{26}$ years

Consequences: - scales for new physics

- baryogenesis via leptogenesis

PHYSICAL REVIEW D **92**, 036005 (2015)

$$\mathcal{L}_D = \frac{g}{\Lambda_D^{D-4}} \mathcal{O}_D$$



$$m_e \bar{\epsilon}_5 = \frac{g^2 v^2}{\Lambda_5}, \quad \frac{G_F \bar{\epsilon}_7}{\sqrt{2}} = \frac{g^3 v}{2\Lambda_7^3},$$

$$\frac{G_F^2 \bar{\epsilon}_9}{2m_p} = \frac{g^4}{\Lambda_9^5}, \quad \frac{G_F^2 \bar{\epsilon}_{11}}{2m_p} = \frac{g^6 v^2}{\Lambda_{11}^7}$$

$g \approx 1 \quad v = 174 \text{ GeV}$ (Higgs expectation value)

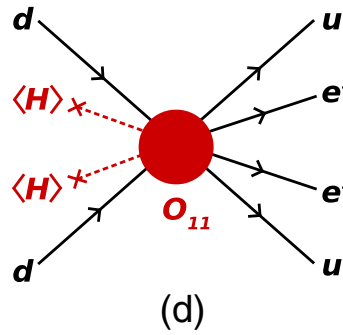
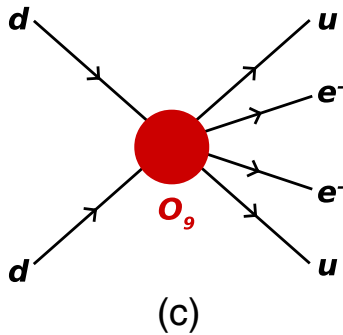
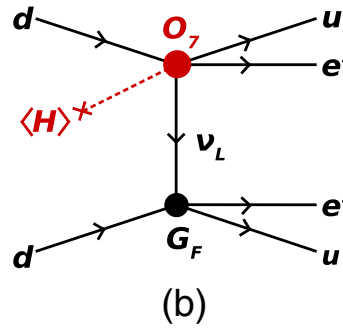
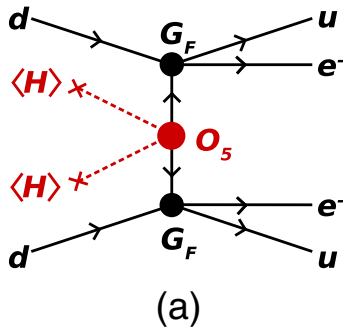
\mathcal{O}_D	$\bar{\epsilon}_D$	Λ_D
\mathcal{O}_5	2.8×10^{-7}	2.12×10^{14}
\mathcal{O}_7	2.0×10^{-7}	3.75×10^4
\mathcal{O}_9	1.5×10^{-7}	2.48×10^3
\mathcal{O}_{11}	1.5×10^{-7}	1.16×10^3

Consequences: - scales for new physics

- baryogenesis via leptogenesis

PHYSICAL REVIEW D **92**, 036005 (2015)

$$\mathcal{L}_D = \frac{g}{(\Lambda_D)^{D-4}} \mathcal{O}_D$$



$$m_e \bar{\epsilon}_5 = \frac{g^2 (yv)^2}{\Lambda_5}, \quad \frac{G_F \bar{\epsilon}_7}{\sqrt{2}} = \frac{g^3 (yv)}{2(\Lambda_7)^3},$$

$$\frac{G_F^2 \bar{\epsilon}_9}{2m_p} = \frac{g^4}{(\Lambda_9)^5}, \quad \frac{G_F^2 \bar{\epsilon}_{11}}{2m_p} = \frac{g^6 (yv)^2}{(\Lambda_{11})^7}$$

TABLE VIII. The BSM effective scale (in GeV) for different dimension-D operators at the present ^{136}Xe half-life limit (Λ_D^0) and for $T_{1/2} \approx 1.1 \times 10^{28}$ years (Λ_D).

\mathcal{O}_D	$\bar{\epsilon}_D$	$\Lambda_D^0(y=1)$	$\Lambda_D^0(y=y_e)$	$\Lambda_D(y=y_e)$
\mathcal{O}_5	$2.8 \cdot 10^{-7}$	$2.12 \cdot 10^{14}$	1904	19044
\mathcal{O}_7	$2.0 \cdot 10^{-7}$	$3.75 \cdot 10^4$	541	1165
\mathcal{O}_9	$1.5 \cdot 10^{-7}$	$2.47 \cdot 10^3$	2470	3915
\mathcal{O}_{11}	$1.5 \cdot 10^{-7}$	$1.16 \cdot 10^3$	31	43

$$\eta_N \propto \frac{1}{m_{W_R}^4 m_N}$$

$$g \approx 1 \quad v = 174 \text{ GeV} \quad y_e = 3 \times 10^{-6} \text{ electron mass Yukawa}$$

Summary

- The physics of the neutrinos is very exciting and offers a lot of research opportunities.
- Double beta decay (DBD), if observed, will represent a big step forward in our understanding of the neutrinos, and of physics beyond the Standard Model. A Nobel prize may be awarded for its discovery.
- The physics learned from DBD is complementary to that learned from Large Hadron Collider (future colliders).
- Better nuclear matrix elements and effective DBD operators are needed, especially for the short range mechanisms. And we are working hard for that!

Collaborators:

- Alex Brown, NSCL@MSU
- **Roman Senkov, CUNY/CMU**
- **Andrei Neacsu, CMU**
- Jonathan Engel, UNC
- Jason Holt, TRIUMF
- Petr Navratil, TRIUMF
- Sofia Quaglioni, LLNL
- Micah Schuster, ORNL
- **Changfeng Jiao, CMU**

MS Theses:

- Fahim Ahmed, CMU
- Shiplu Sarker, CMU/ISU