



Hydrodynamics with sources

work in progress

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Motivation

- Thermalization during the pre-equilibrium stage does not happen globally at one constant proper time
- Nuclei overlapping time in low energy collisions is not negligible
- Medium response from Jet energy deposition
- Hadronized particles re-absorbed by the hot medium
- Magnetohydrodynamics

Hydrodynamics with sources

Energy-momentum current and net baryon density are feed into hydrodynamic simulations as source terms

$$\partial_{\mu} T^{\mu\nu} = J_{\text{source}}^{\nu}$$

$$\partial_{\mu} J^{\mu} = \rho_{\text{source}}$$

here

$$J_{\text{source}}^{\nu} = \delta e u^{\nu} + (e + P) \delta u^{\nu}$$

$$\delta u^{\nu} = \frac{\Delta_{\mu}^{\nu} J_{\text{source}}^{\mu}}{e + P}$$

δe heats up the system

δu^{ν} accelerates the flow velocity

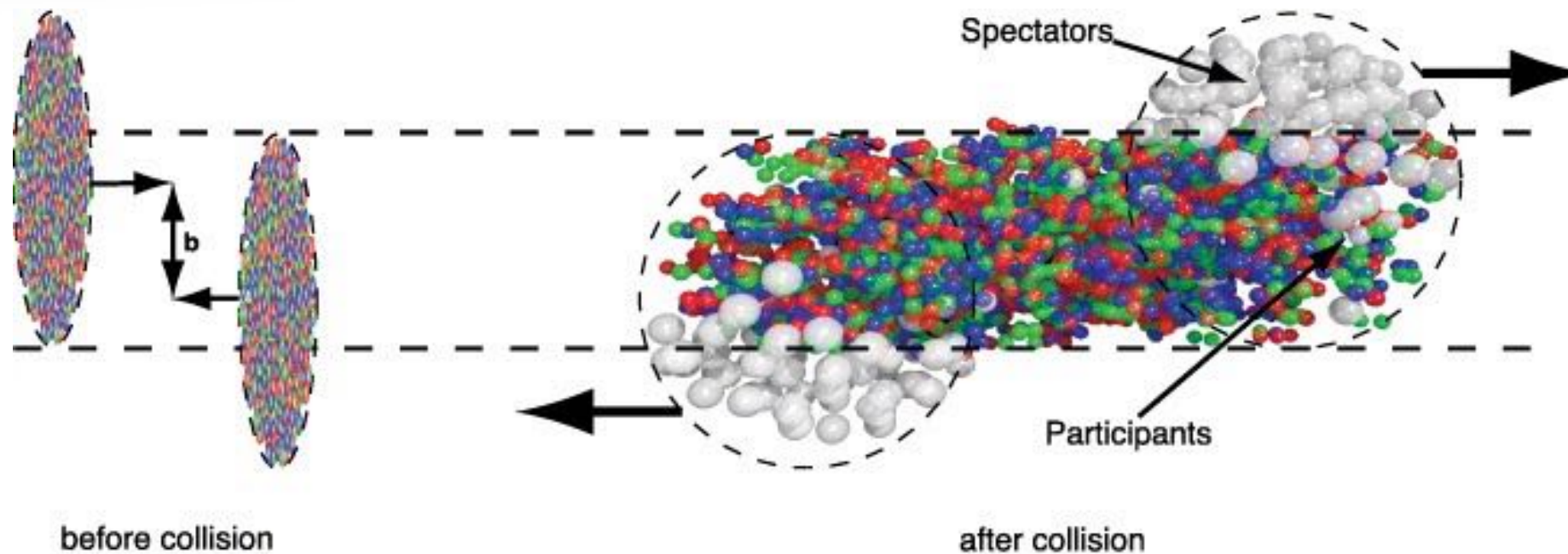
ρ_{source} dopes baryon charges into the system

- Source terms are smeared with Gaussians in space and time

Applications

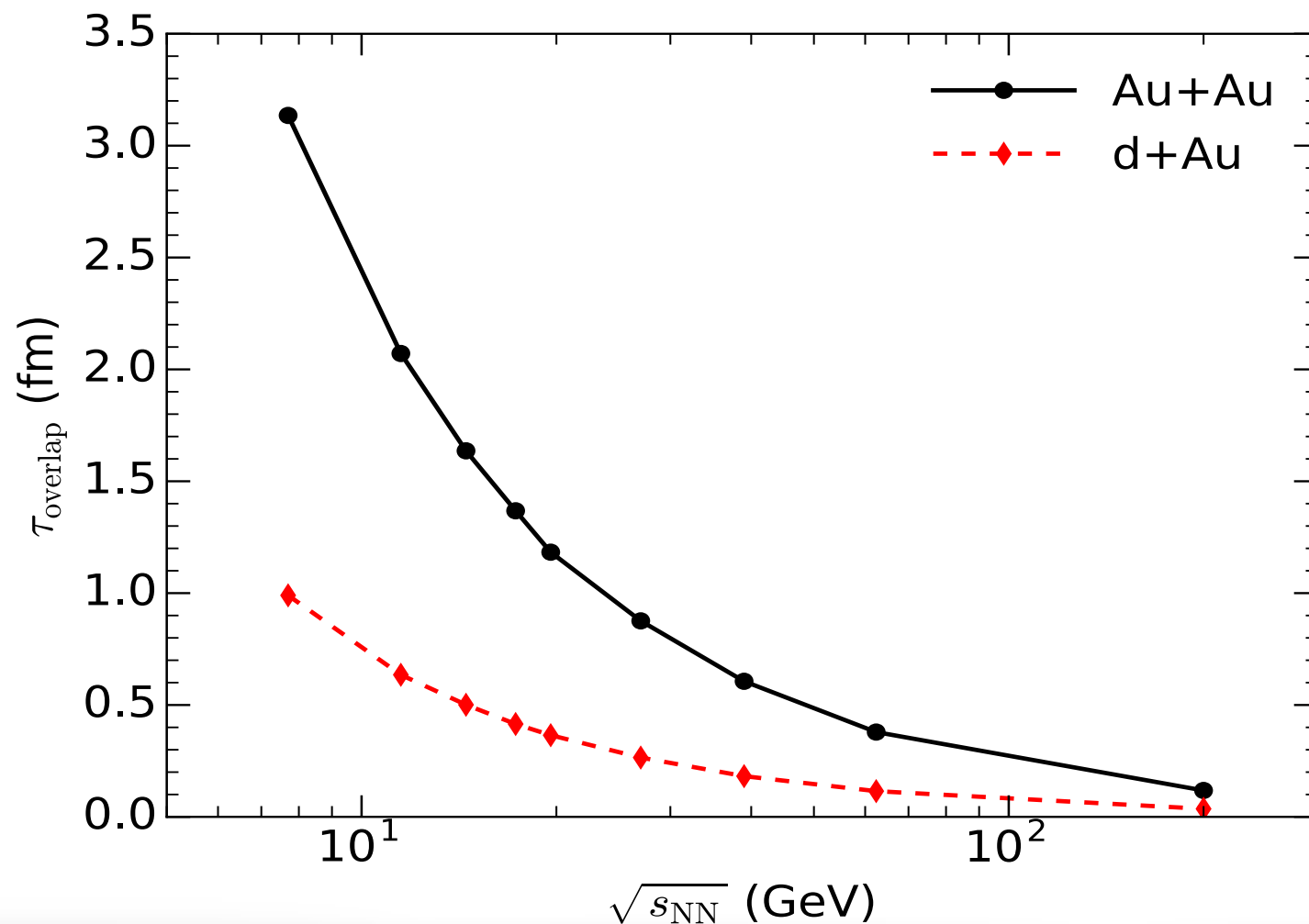


Hydrodynamics at BES energies



Two nuclei
overlapping time

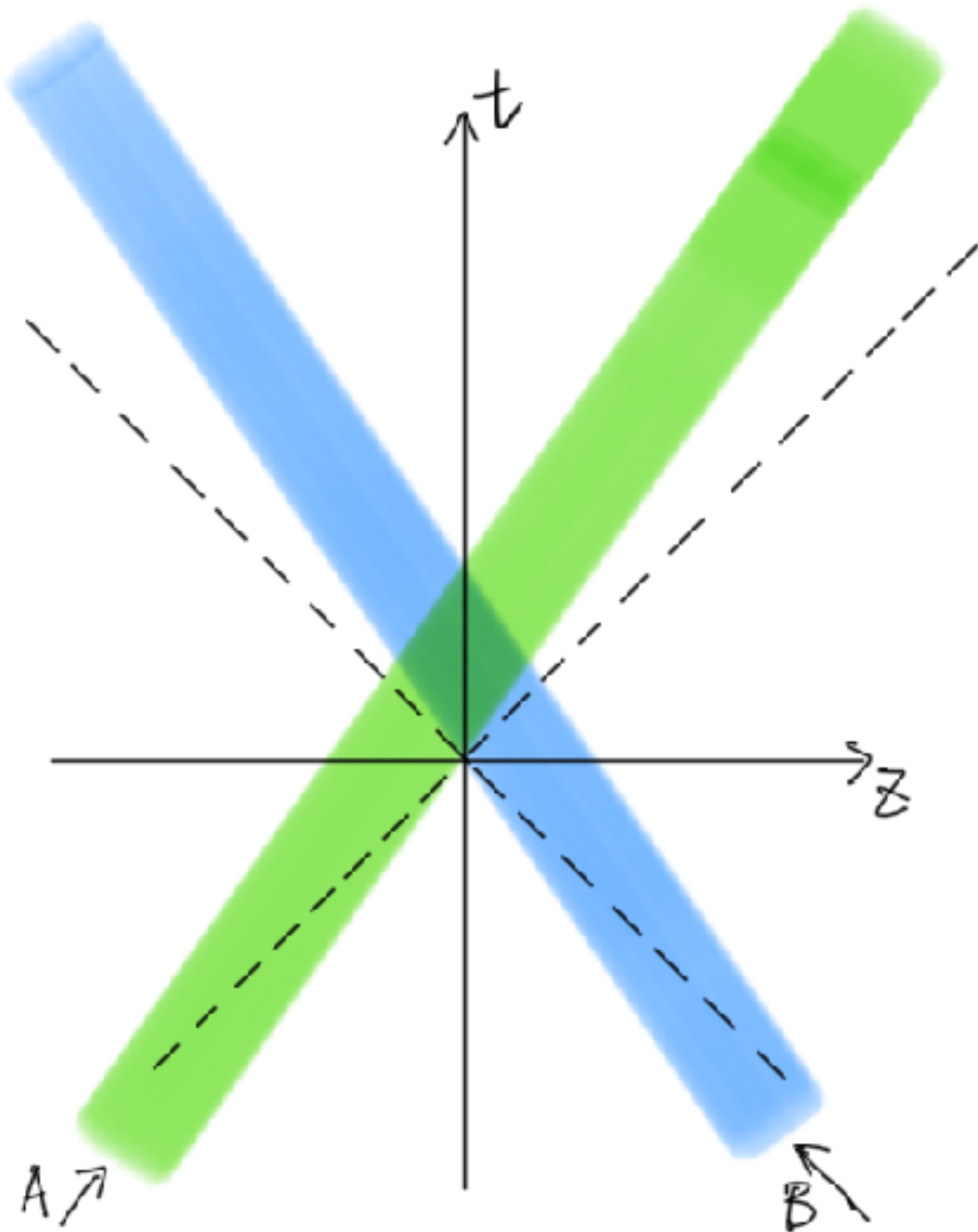
$$\tau \sim \frac{2R}{\gamma v_z}$$



- Nuclei overlapping time is **large** at low collision energy
- **Pre-equilibrium dynamics** can play an important role

Go beyond the Bjorken approximation

- The finite widths of the colliding nuclei are taken into account

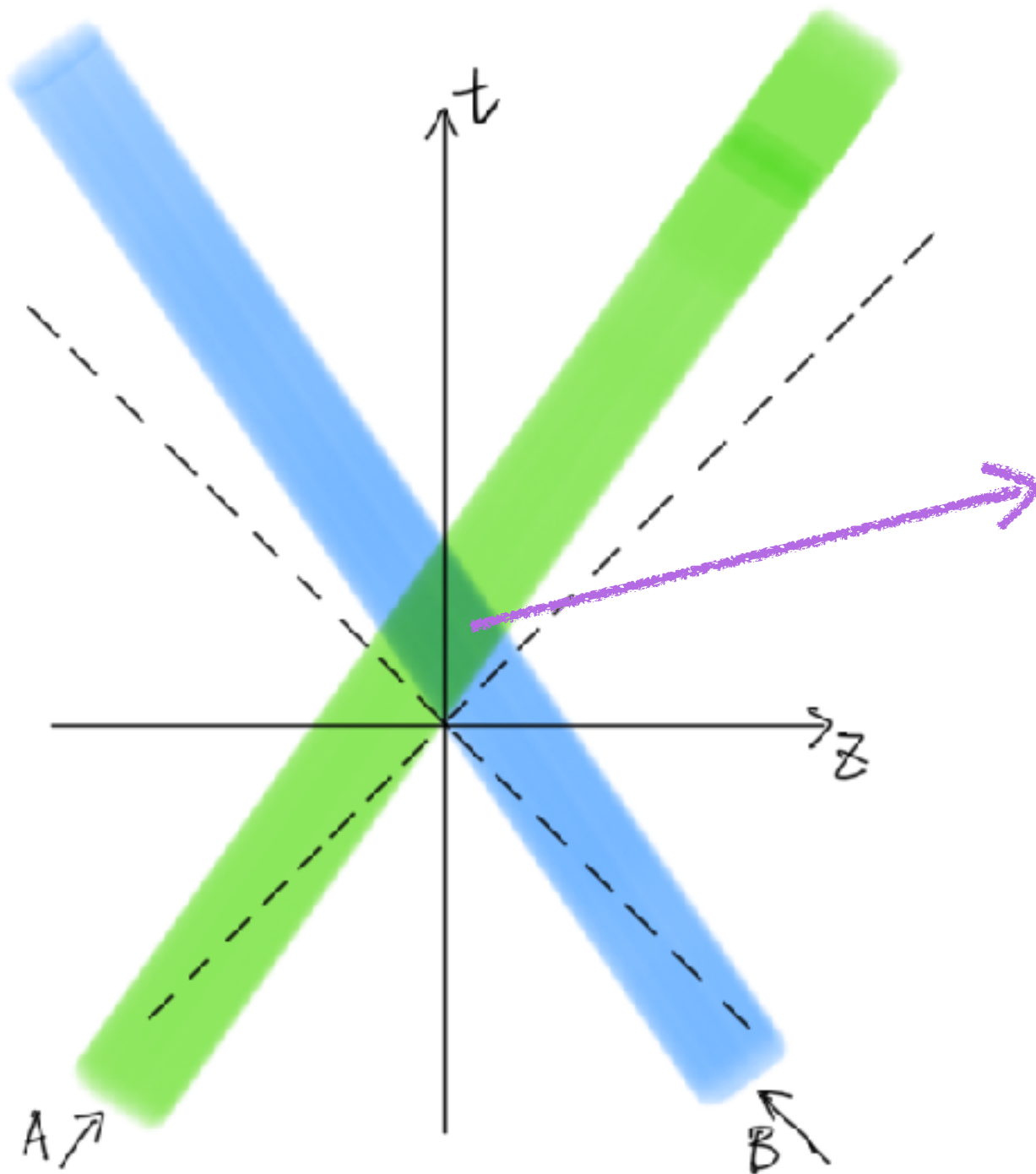


Go beyond the Bjorken approximation

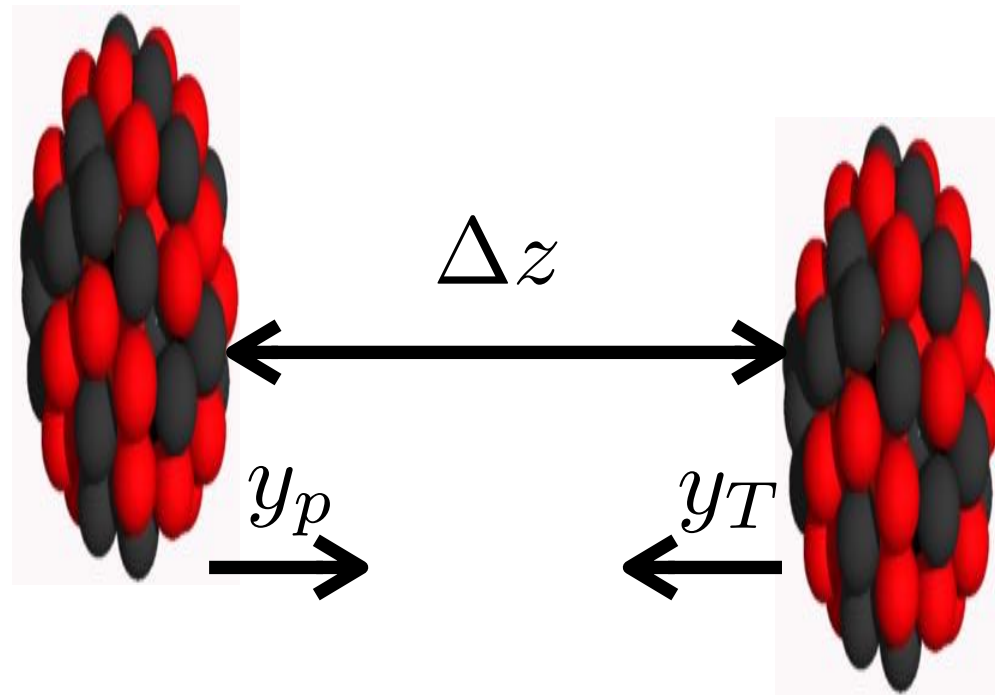
- The finite widths of the colliding nuclei are taken into account

The interaction zone is not point like

$$y \neq \eta_s$$

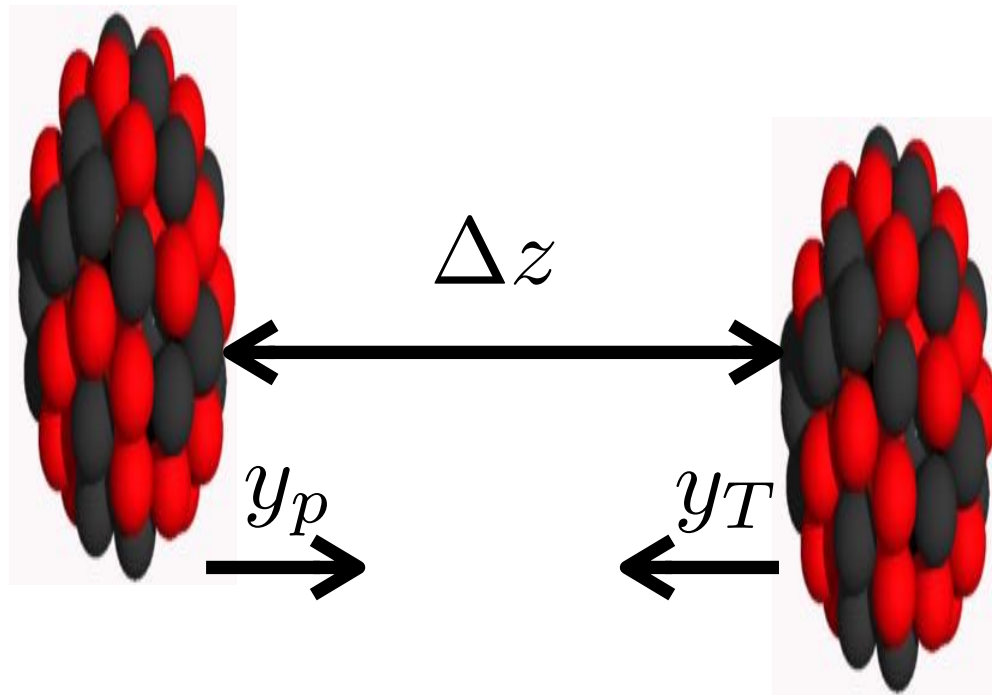


The 3D MCGlauber-LEXUS model

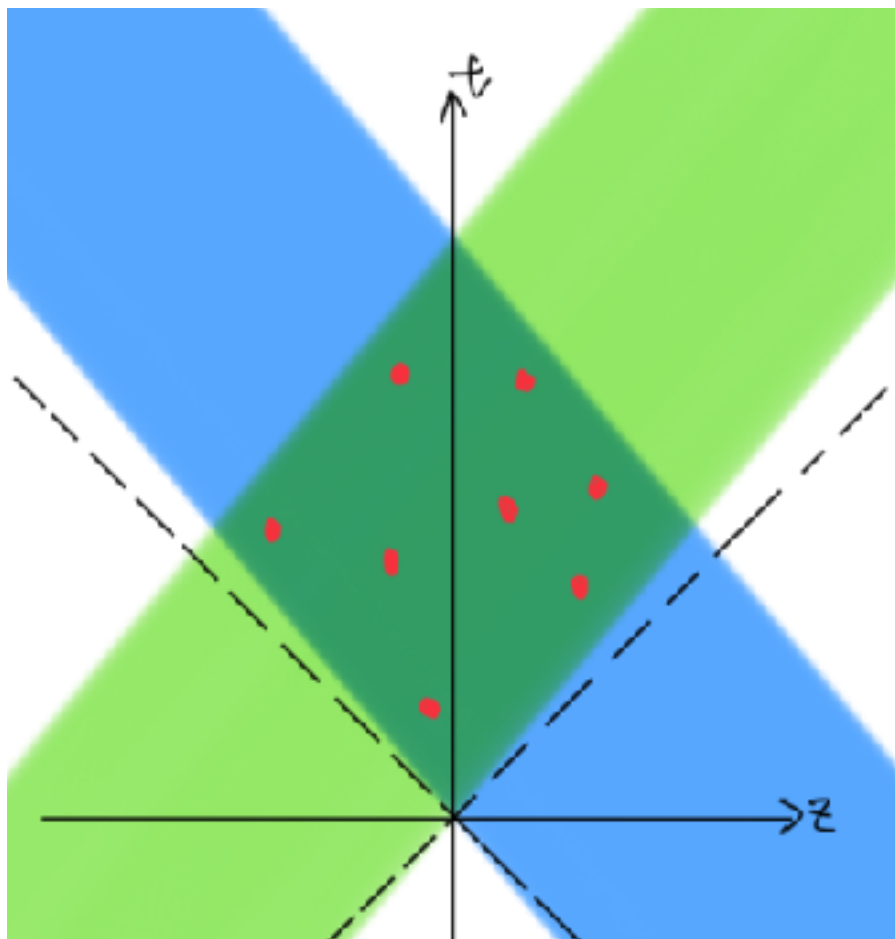


- Collision time and 3D spatial position are determined for every binary collision

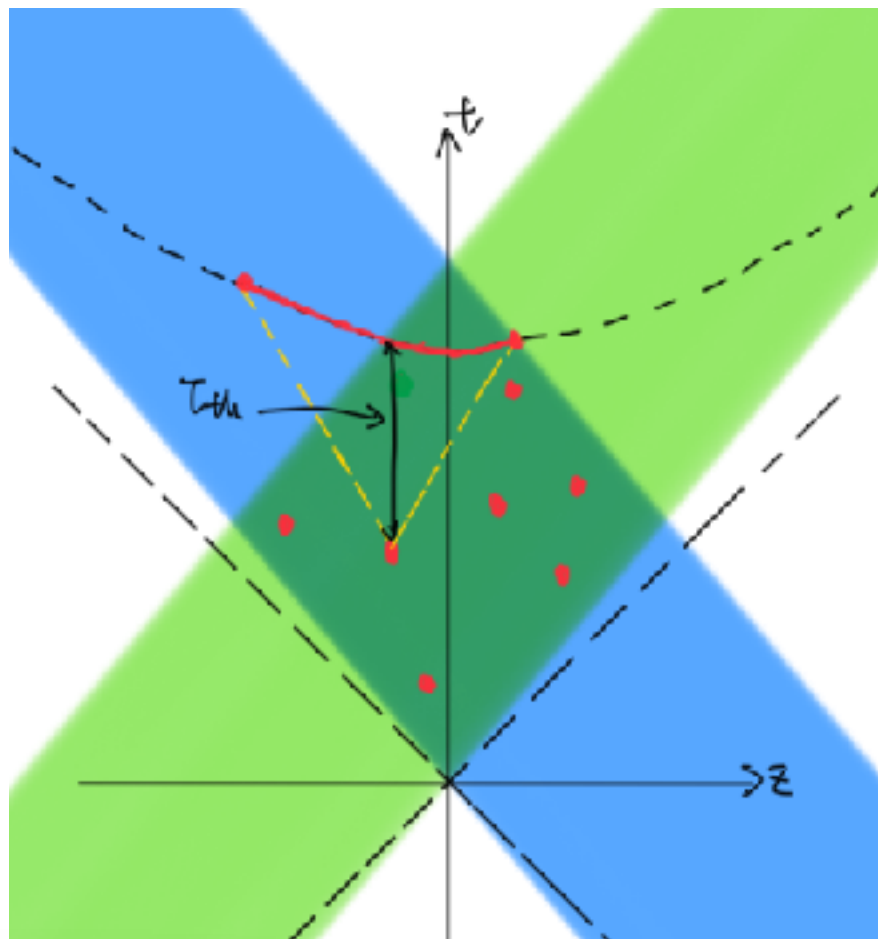
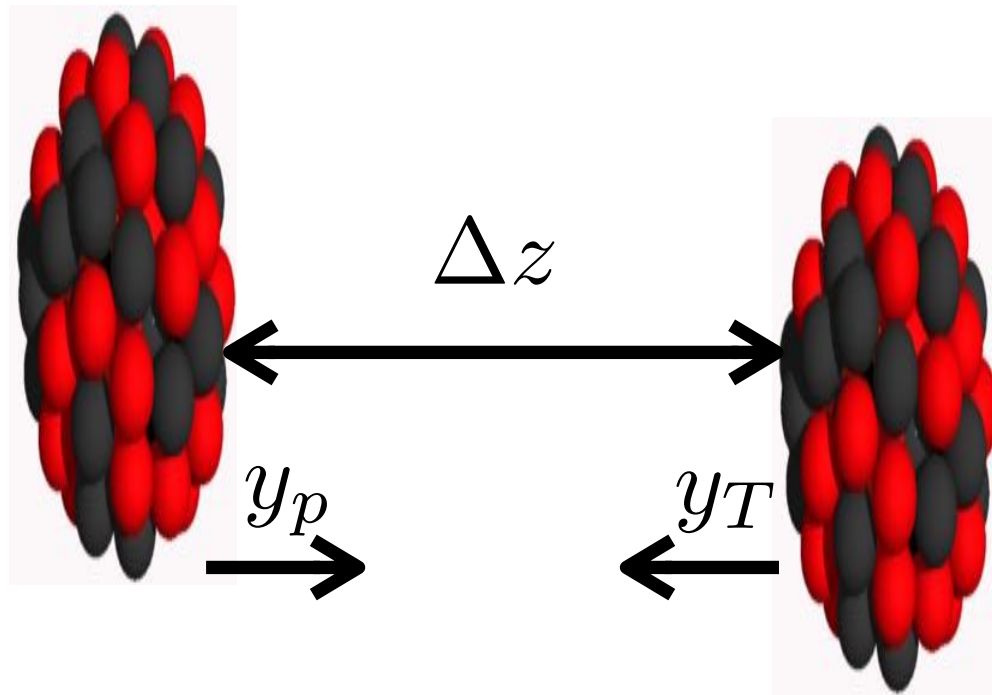
The 3D MCGlauber-LEXUS model



- Collision time and 3D spatial position are determined for every binary collision
- QCD strings are produced from those collision points



The 3D MCGlauber-LEXUS model



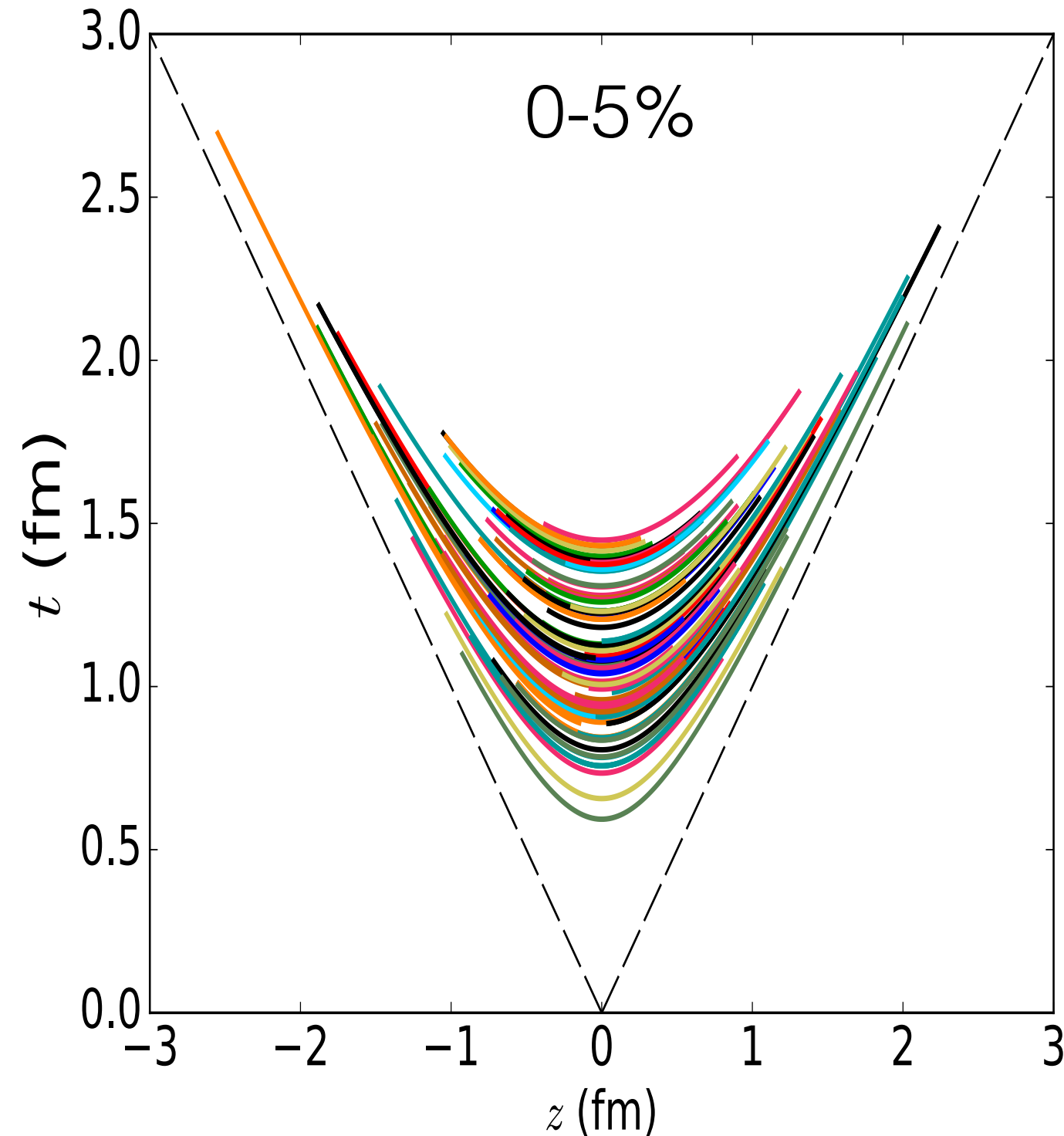
- Collision time and 3D spatial position are determined for every binary collision
- The rapidity loss is determined by the LEXUS model

$$P(y_p, y_T, y) = \lambda \frac{\cosh(y - y_T)}{\sinh(y_P - y_T)} + (1 - \lambda) \delta(y - y_P)$$

- Each string is free-streaming by $\tau_{th} = 0.5 fm$ before thermalized to medium

The 3D MCGlauber-LEXUS model

AuAu@19.6 GeV

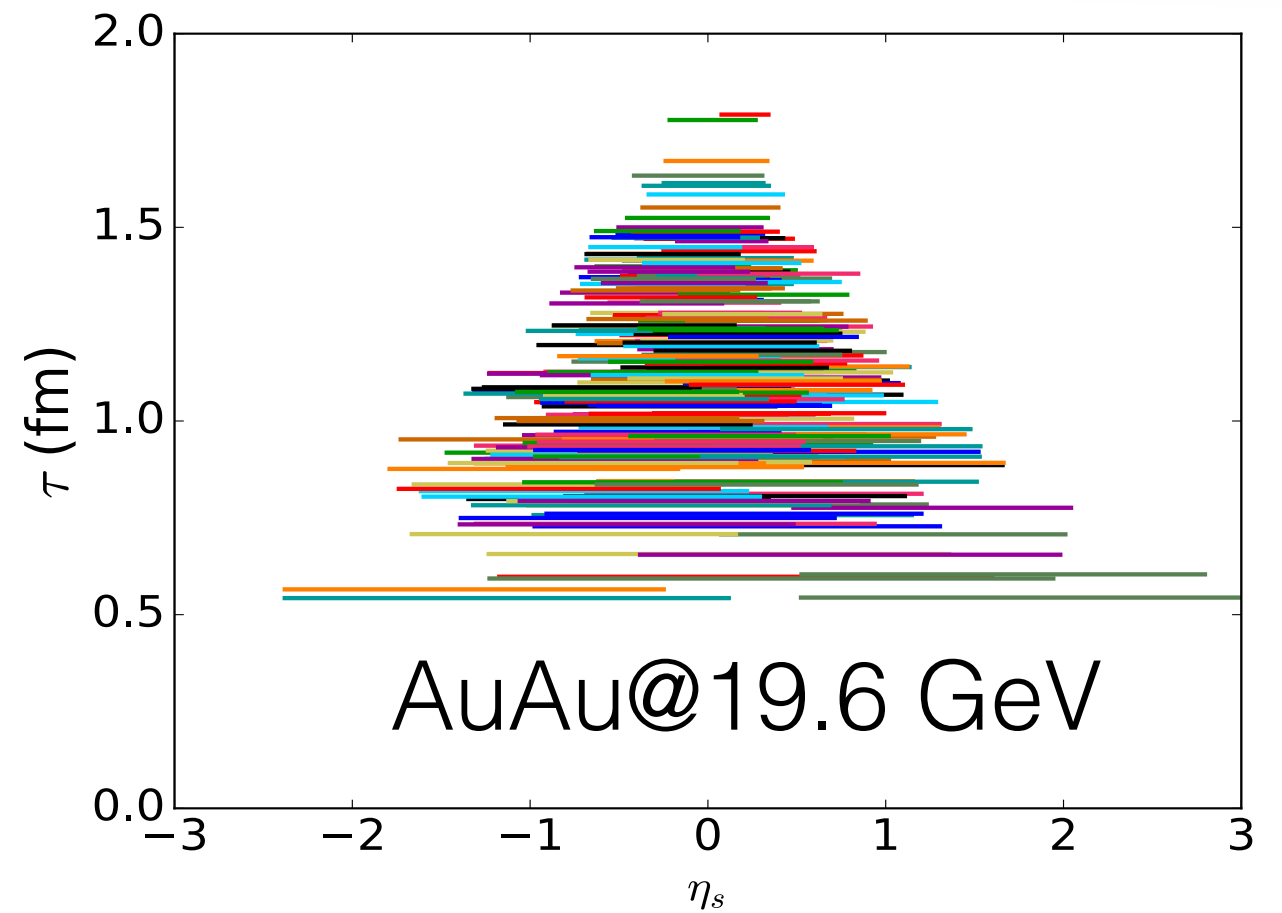
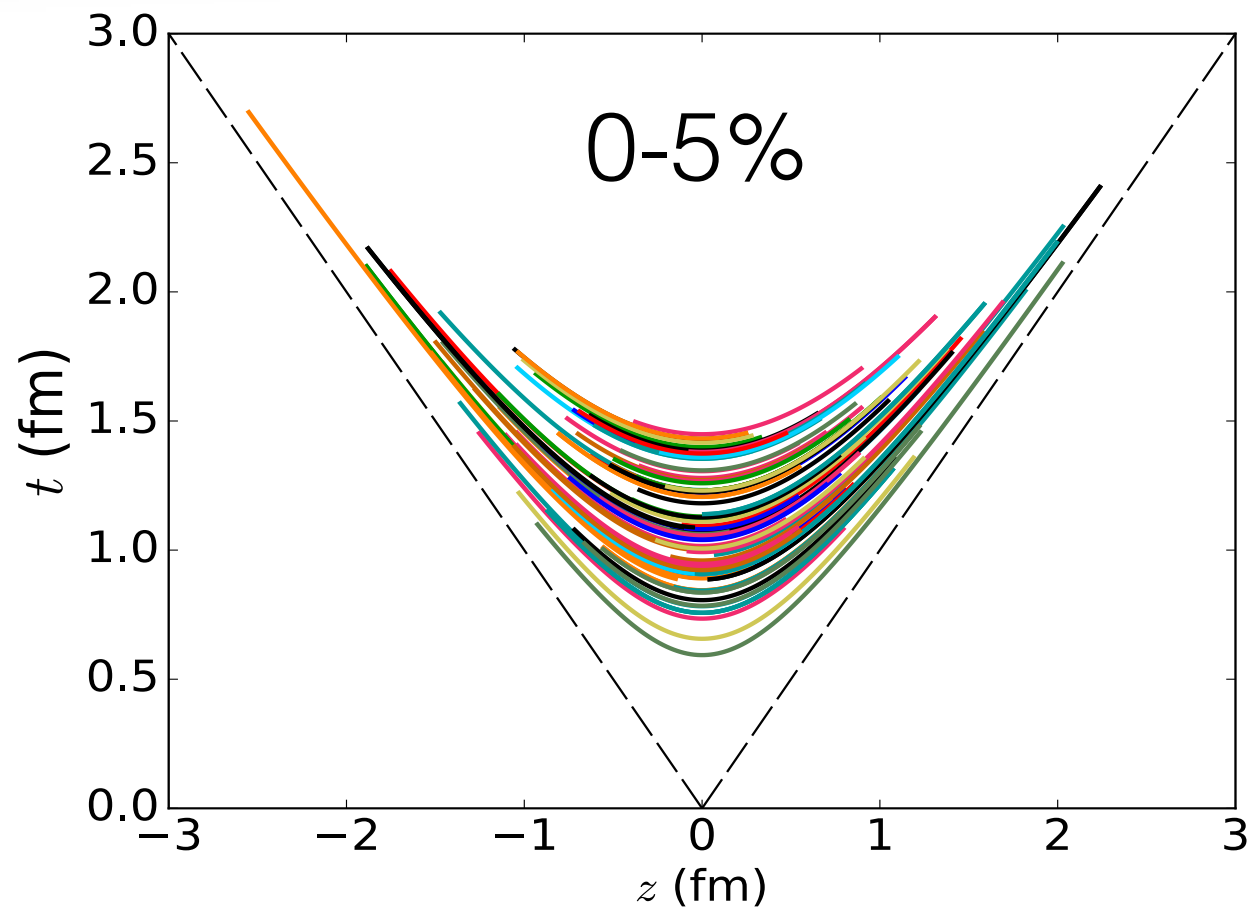


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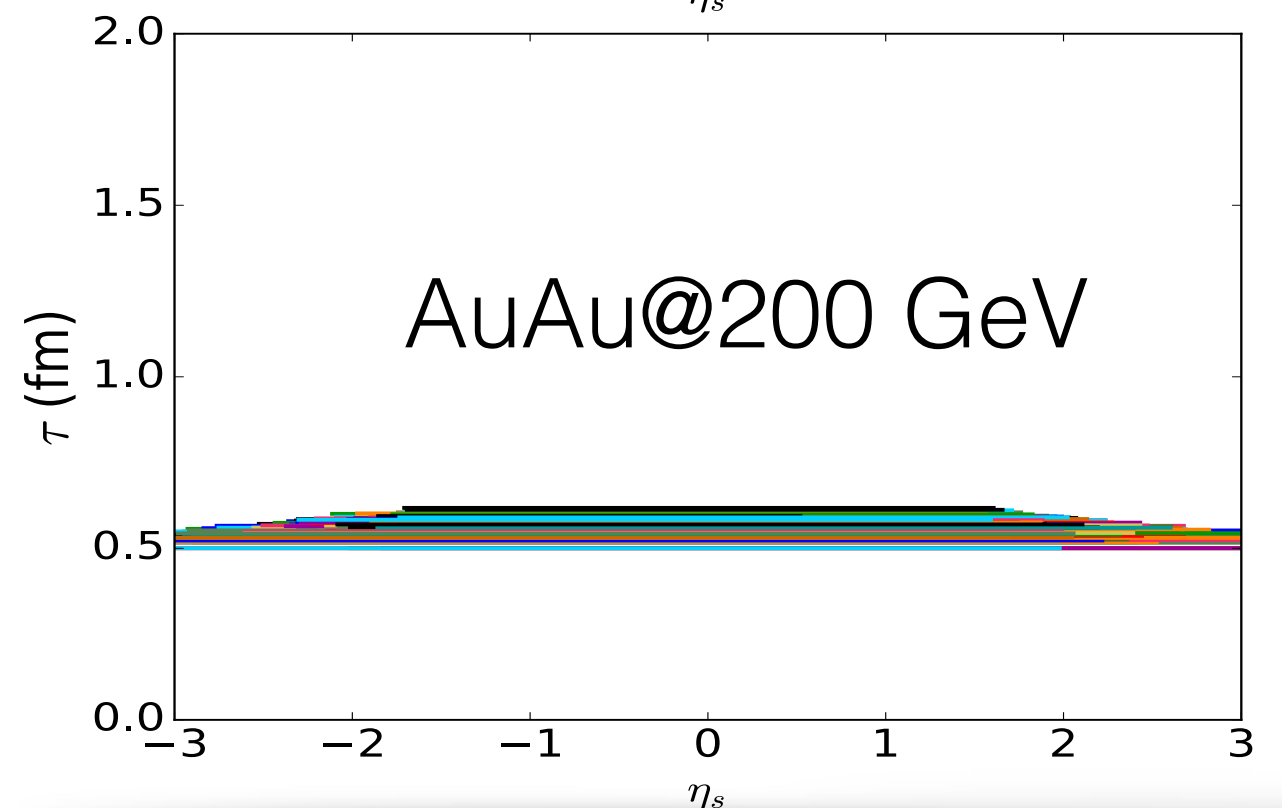
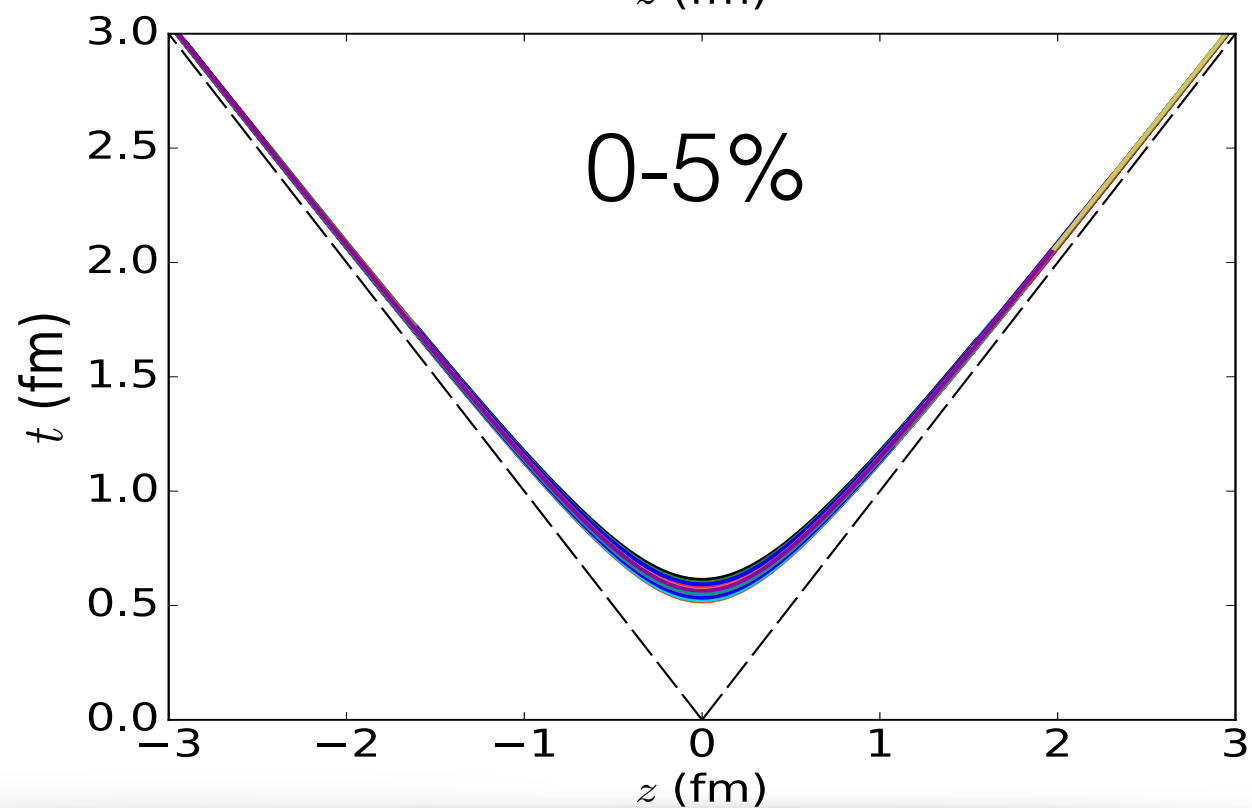
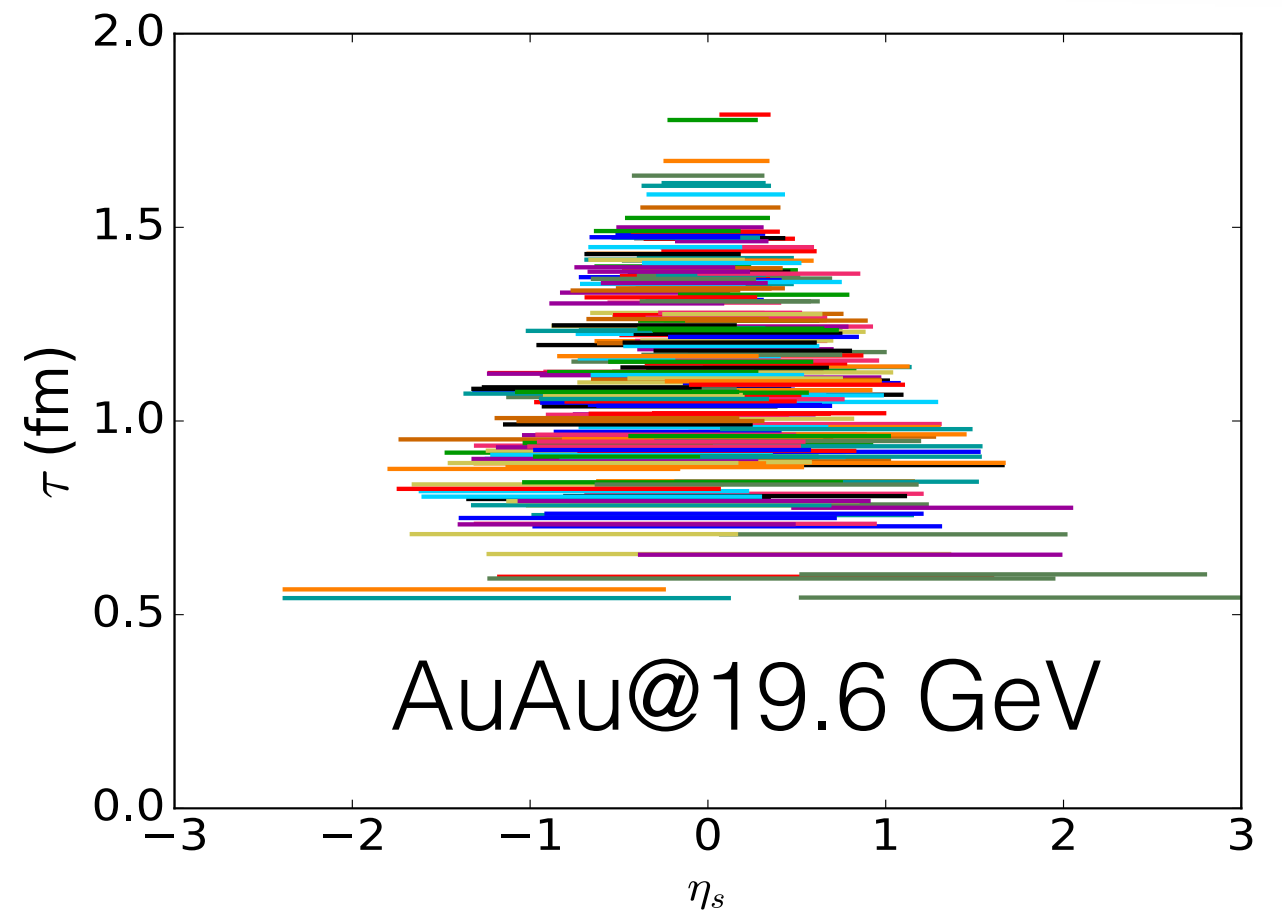
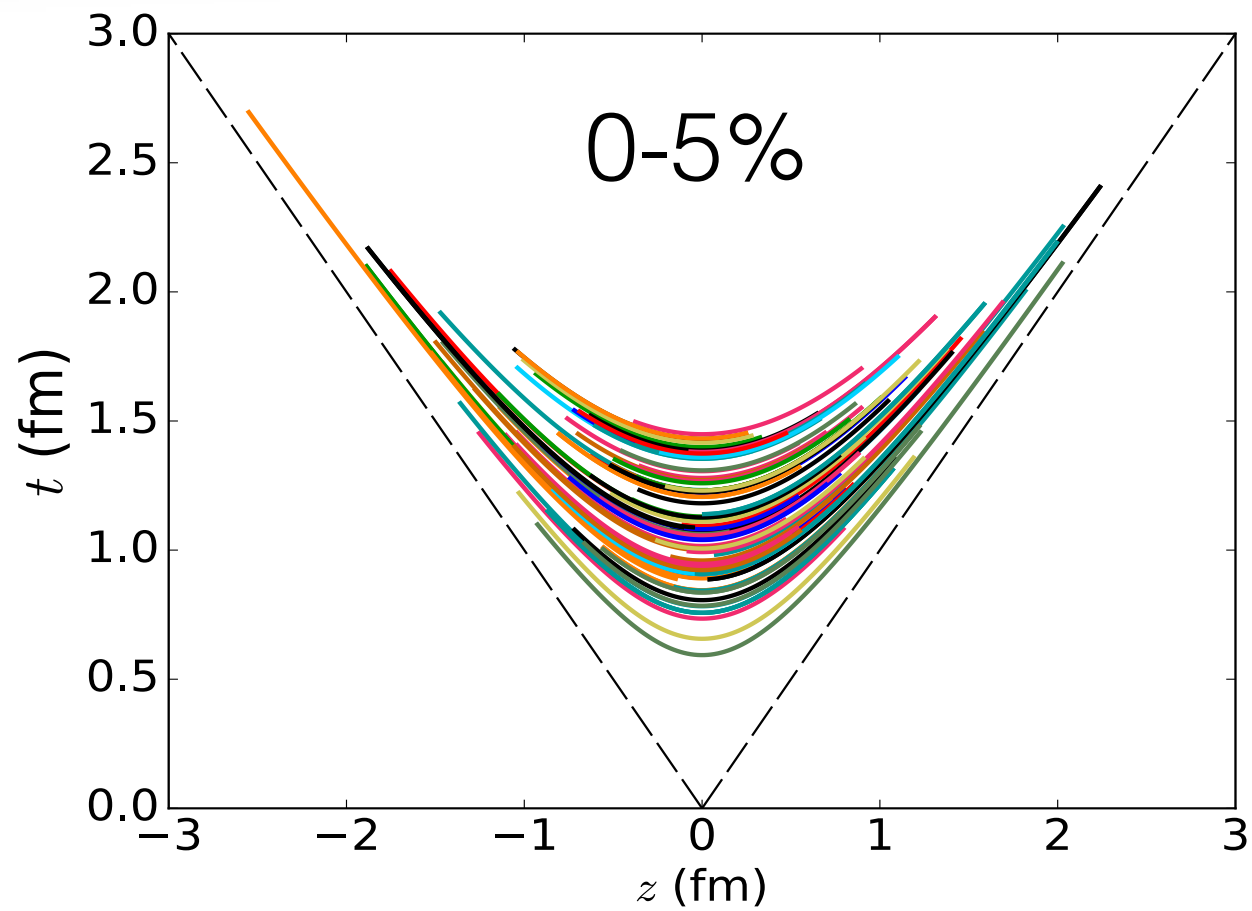
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- Each string is free-streaming by $\tau_{\text{th}} = 0.5 \text{ fm}$ before thermalized to medium

The 3D MCGlauber-LEXUS model



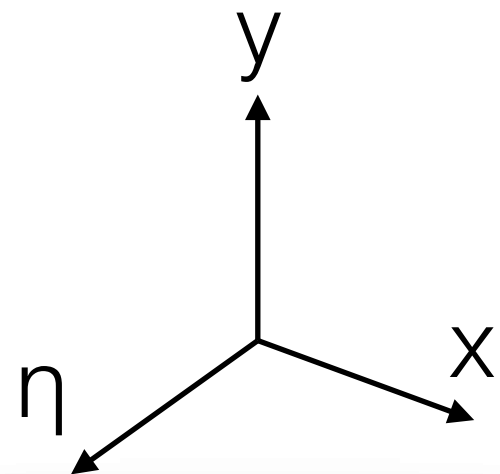
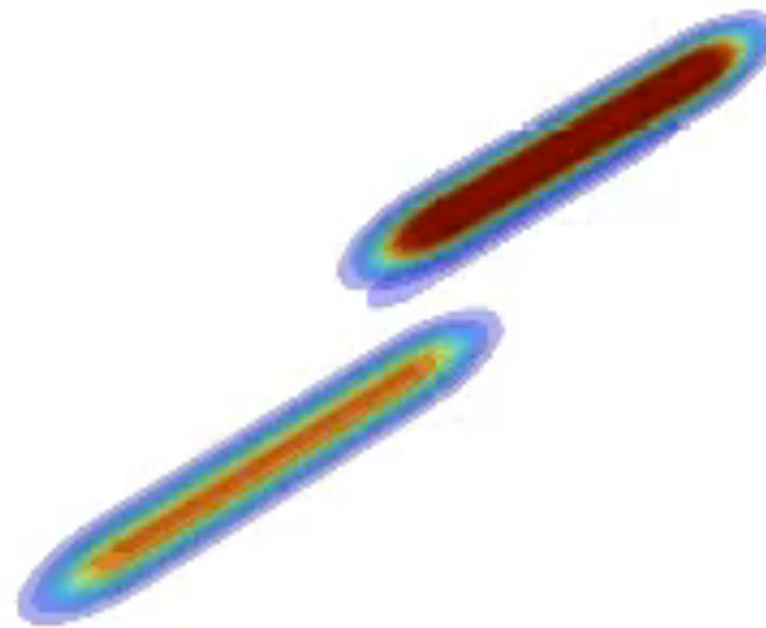
Recover ordinary Glauber at the high energy limit



Hydrodynamical evolution with sources

energy density

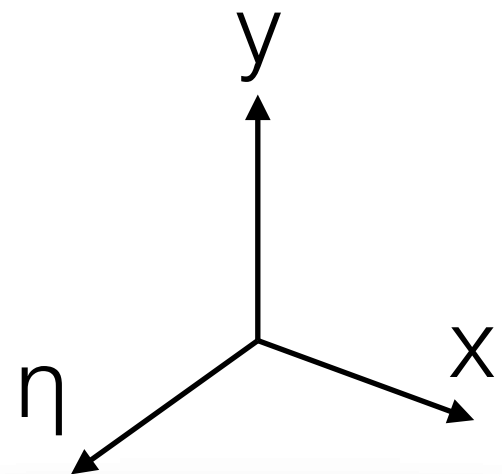
$\tau = 0.50 \text{ fm}$



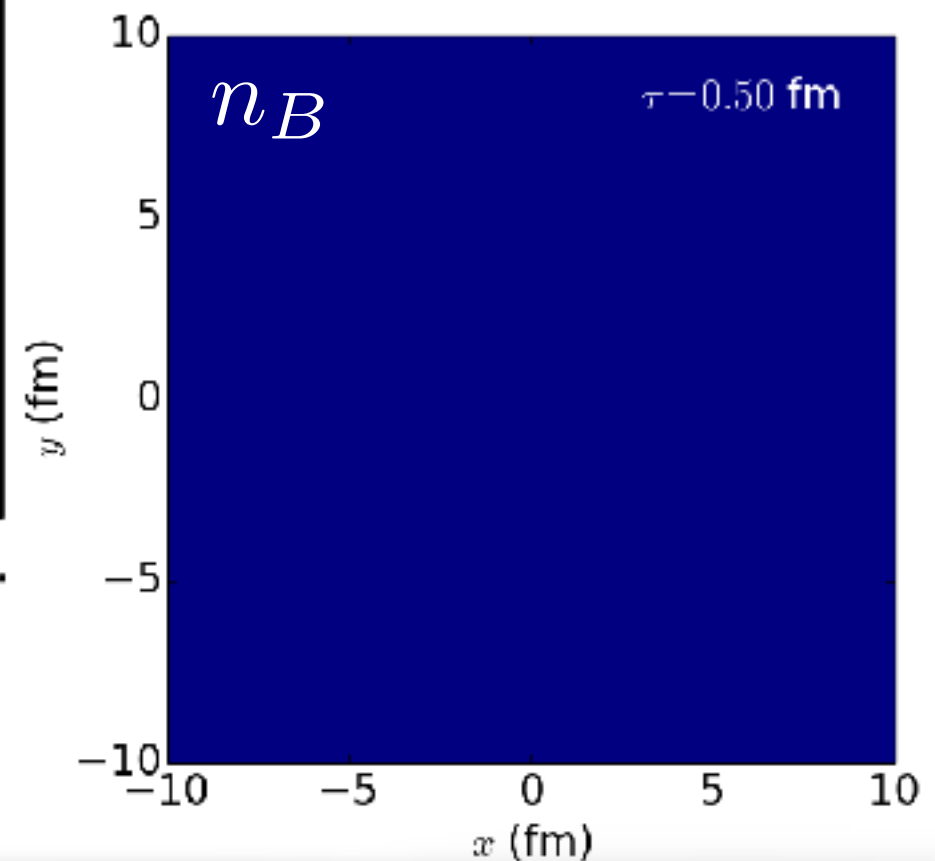
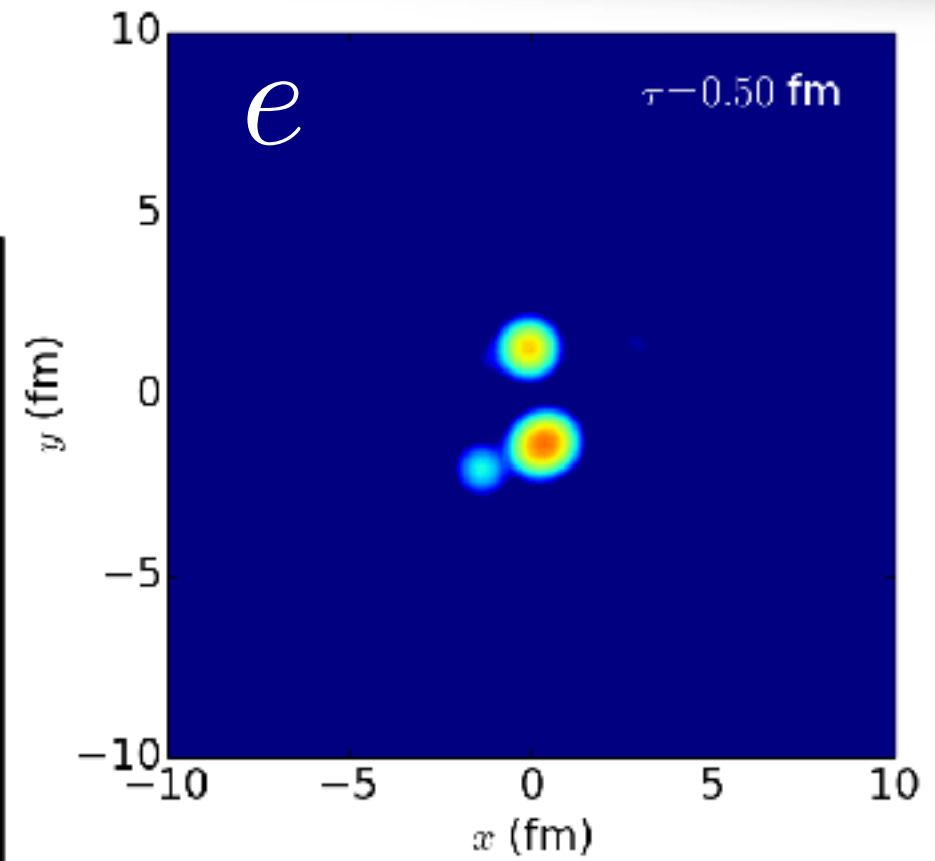
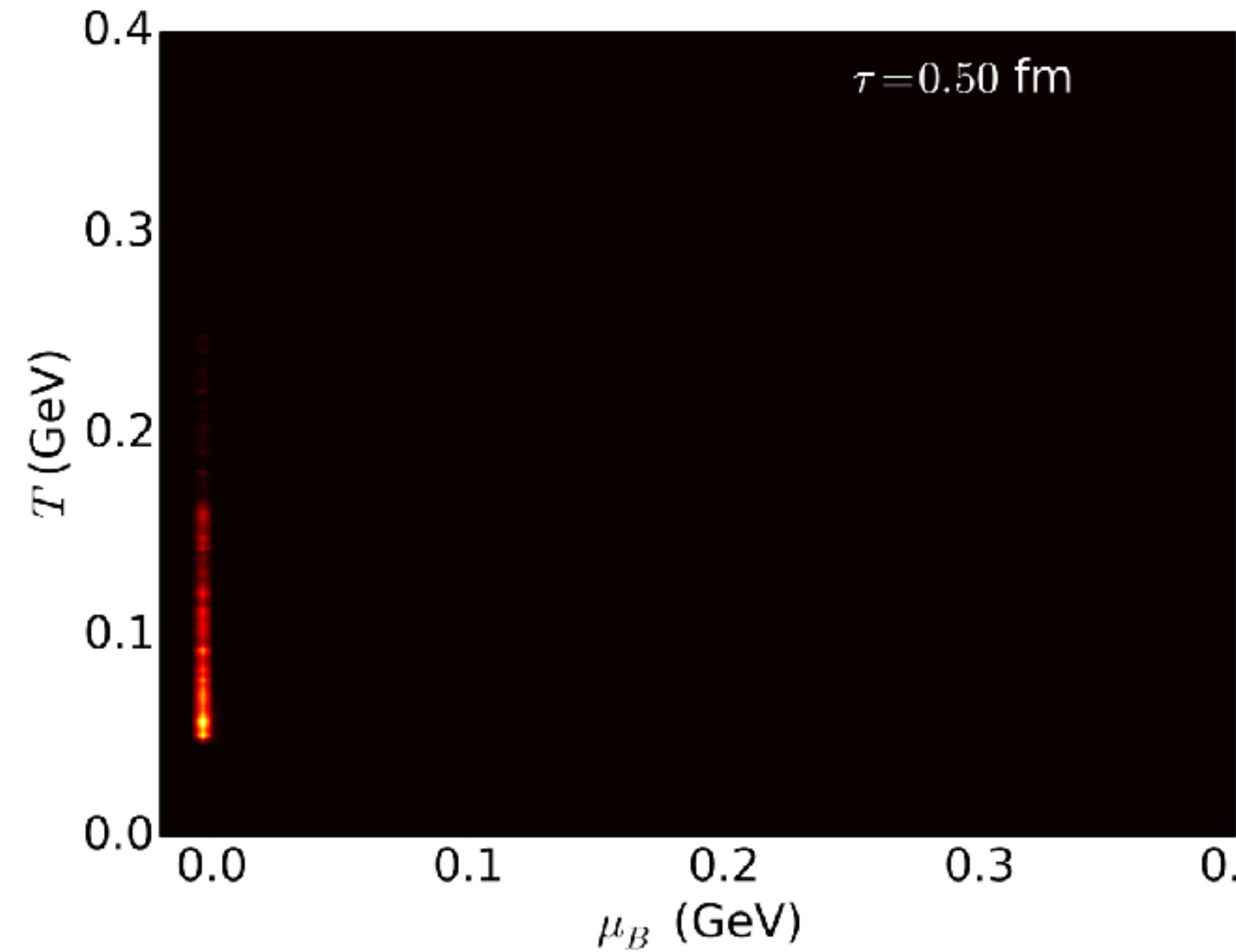
Hydrodynamical evolution with sources

net baryon density

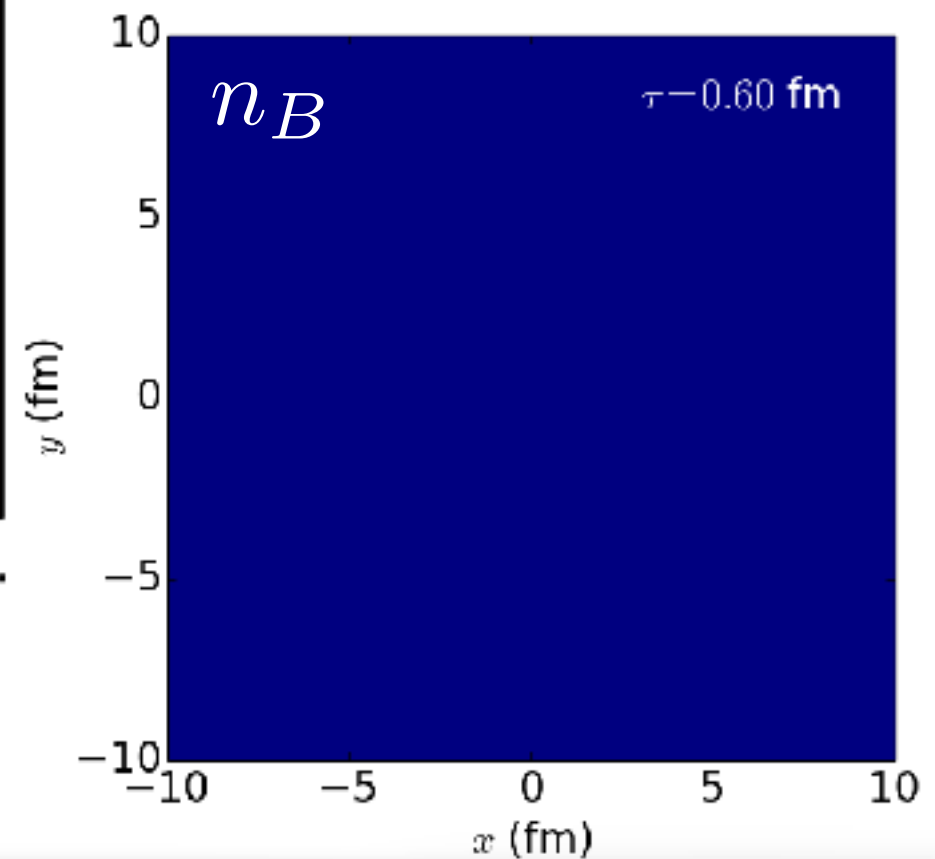
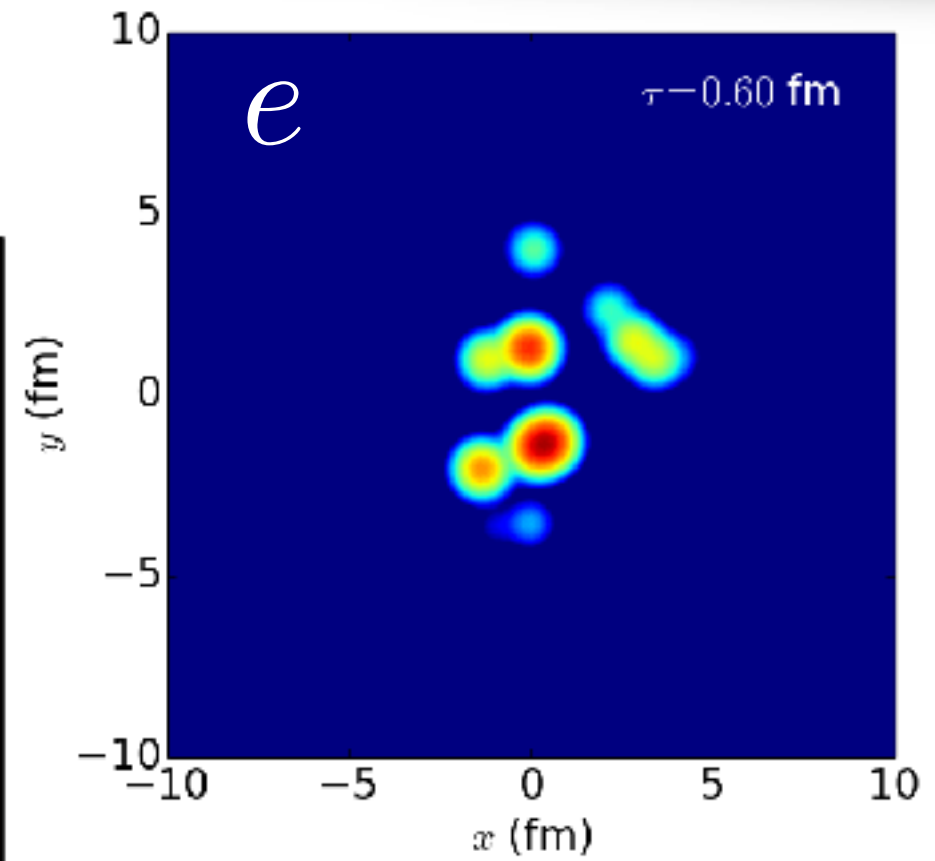
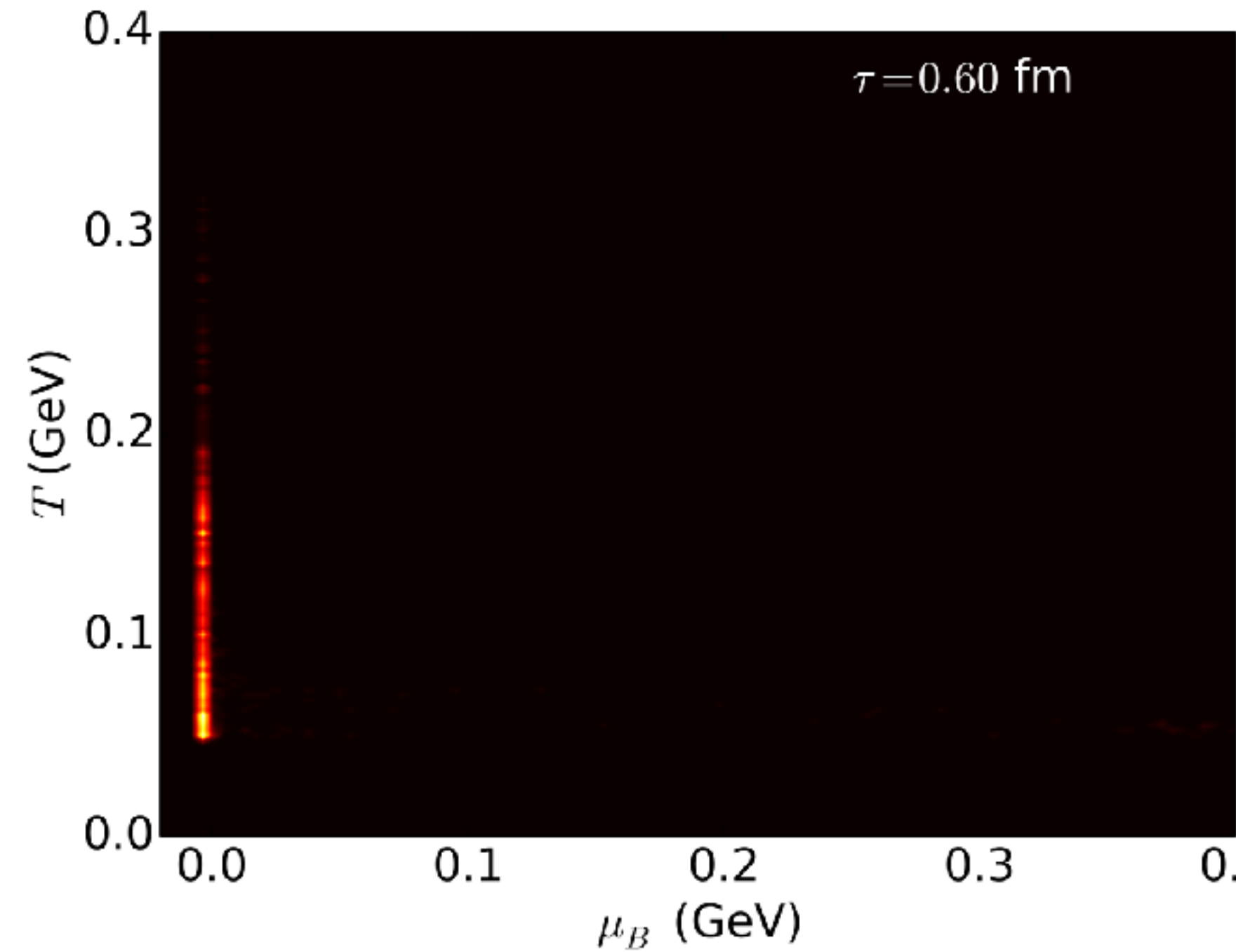
$\tau = 0.50$ fm



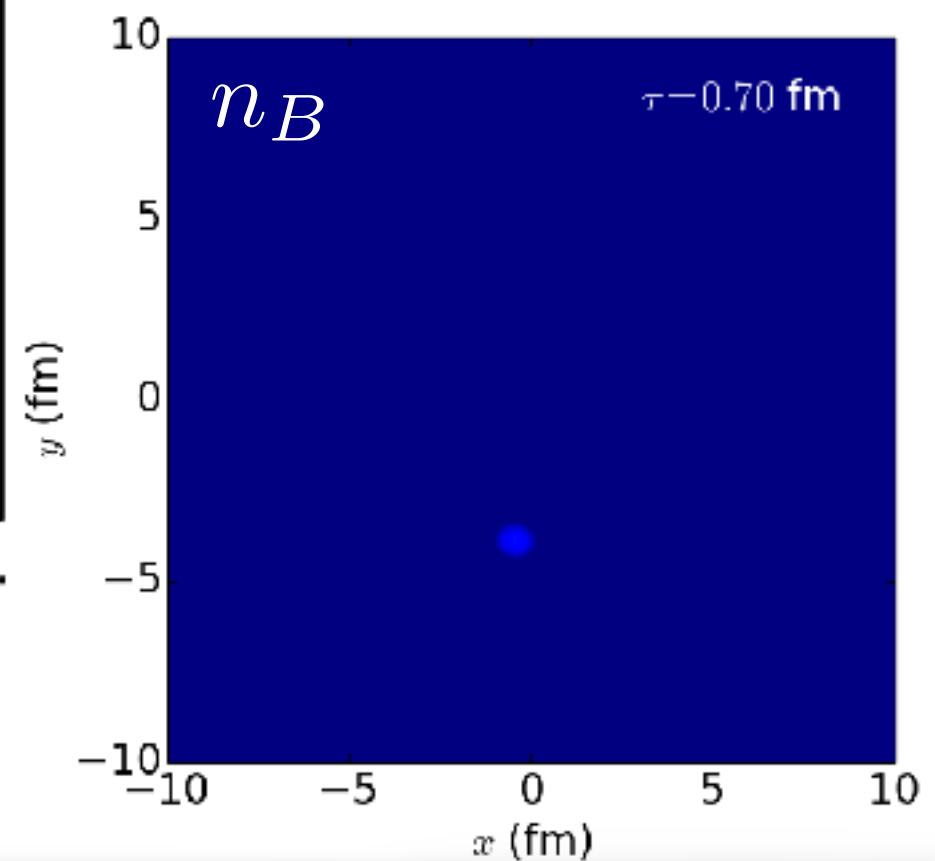
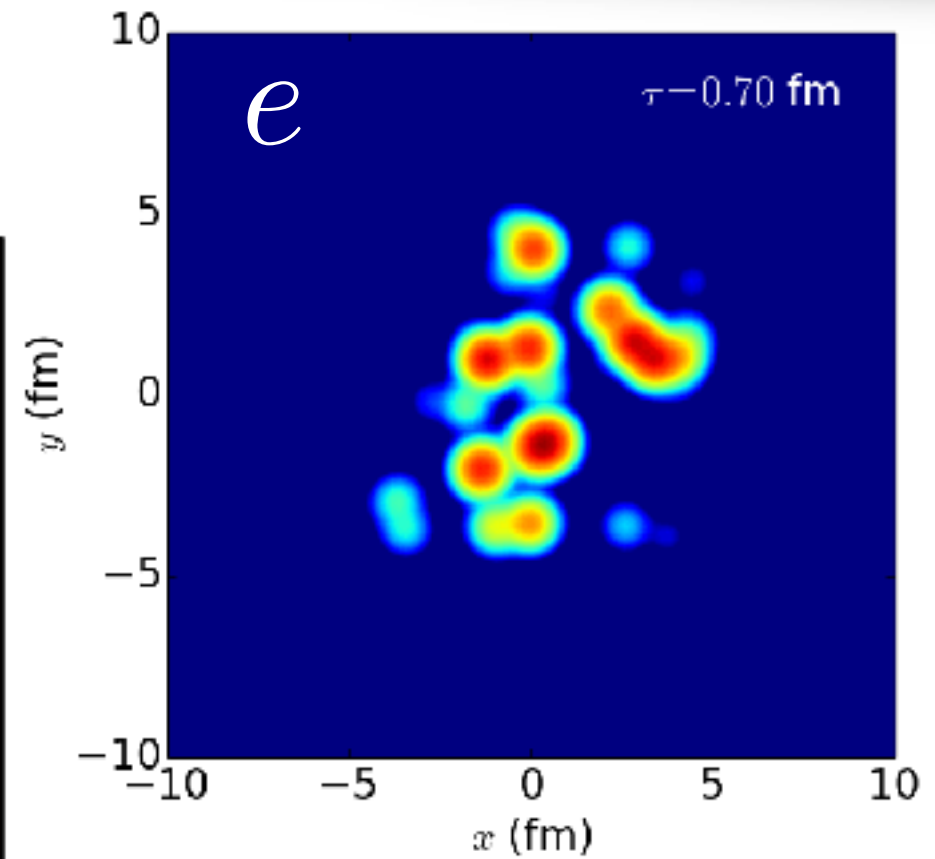
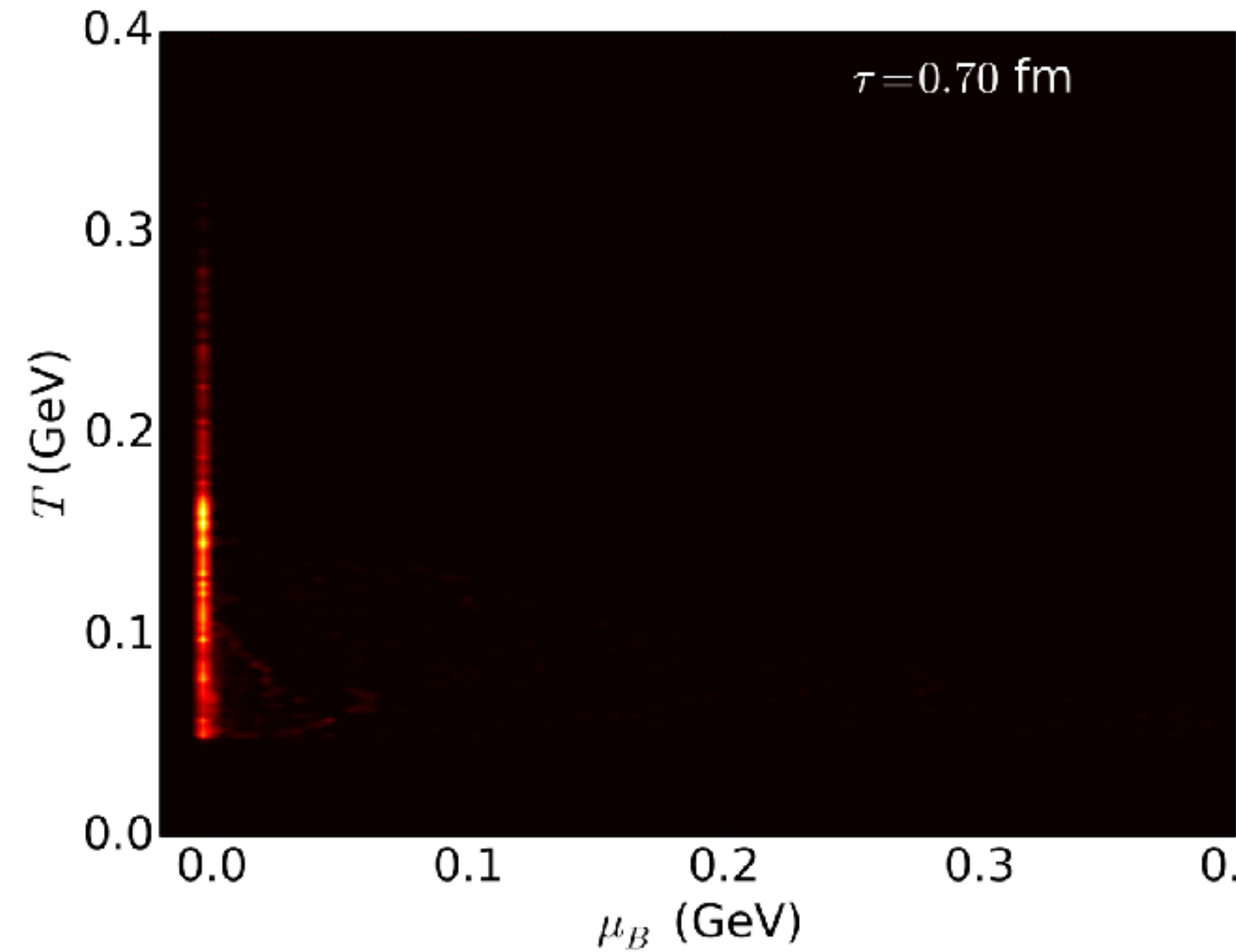
Fireball evolution in the QCD phase diagram



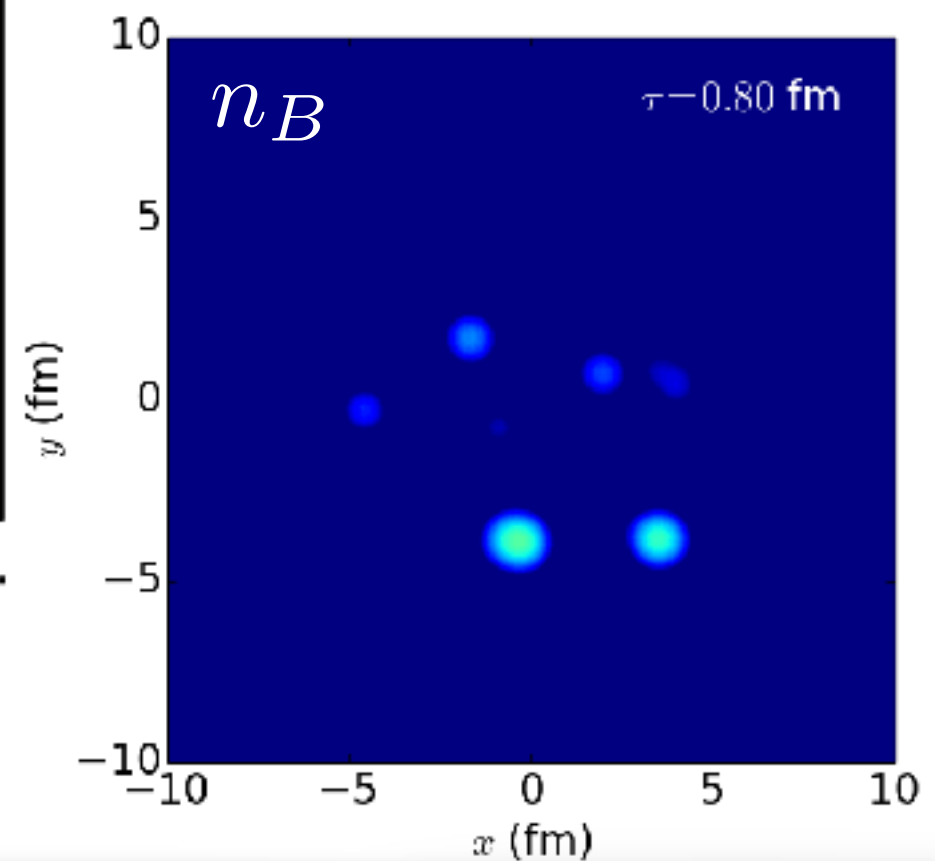
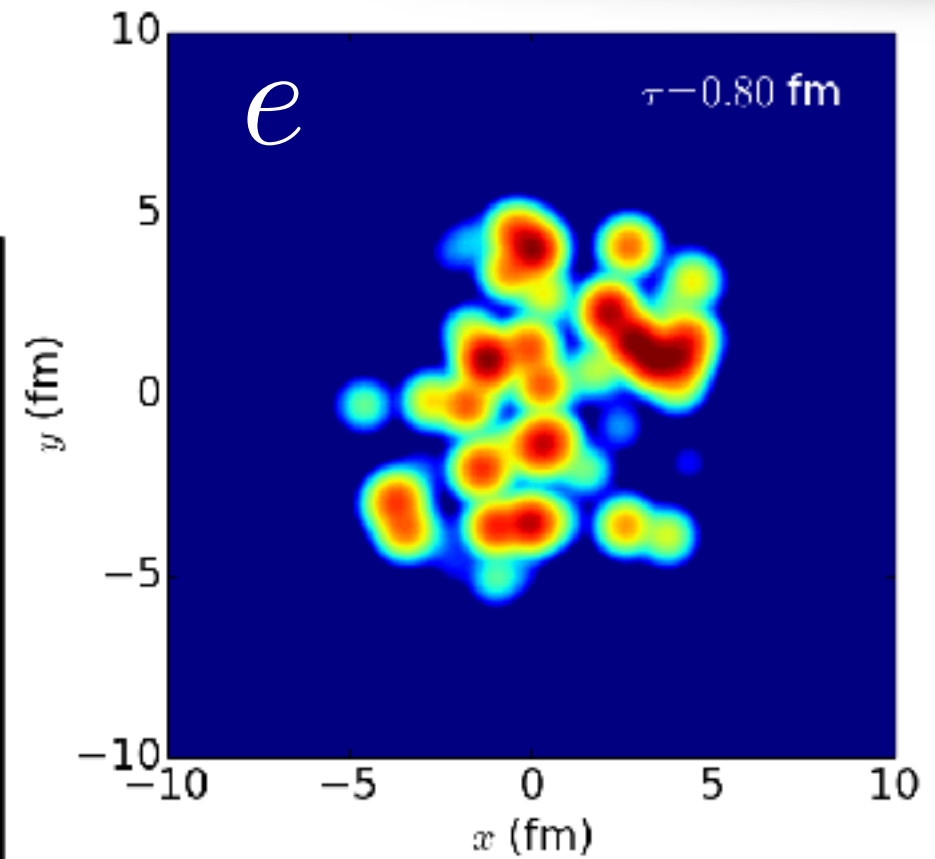
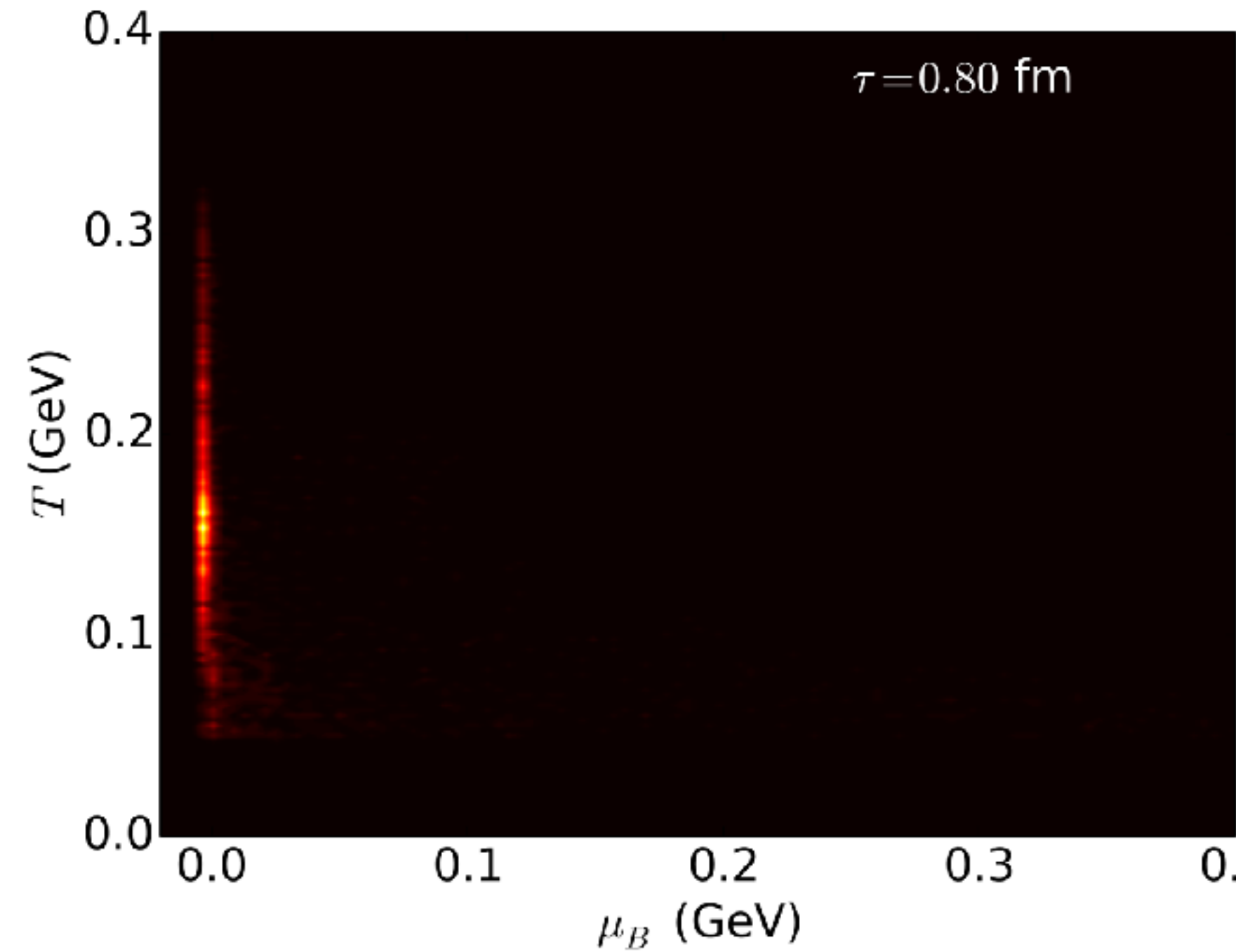
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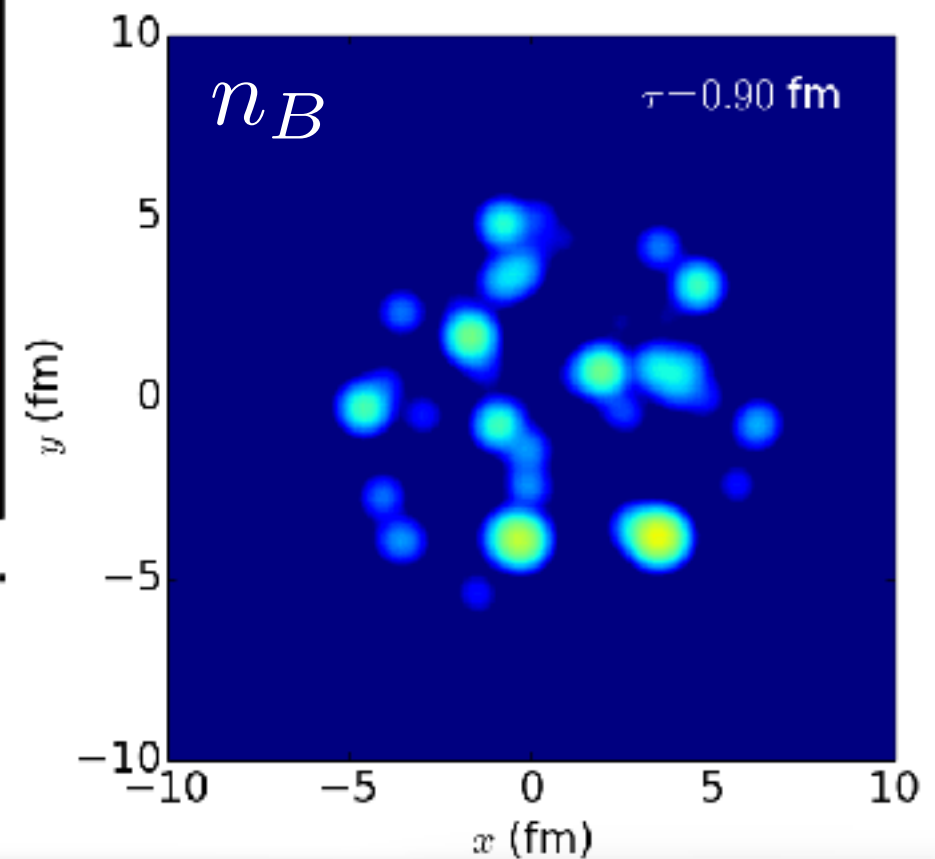
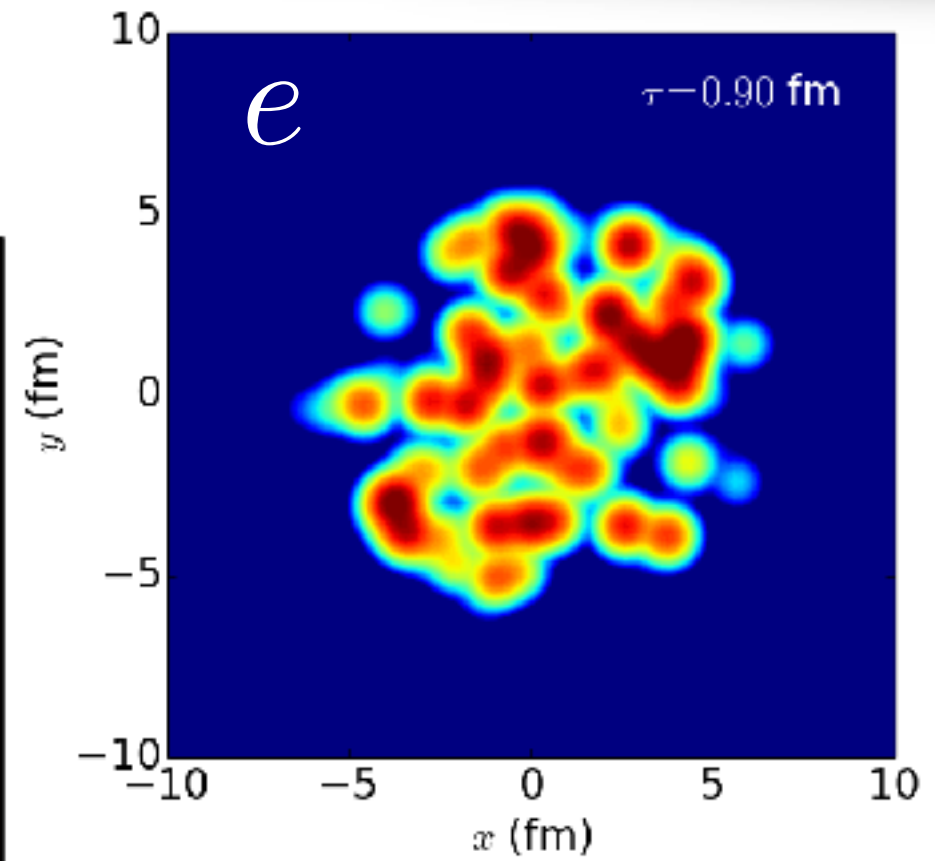
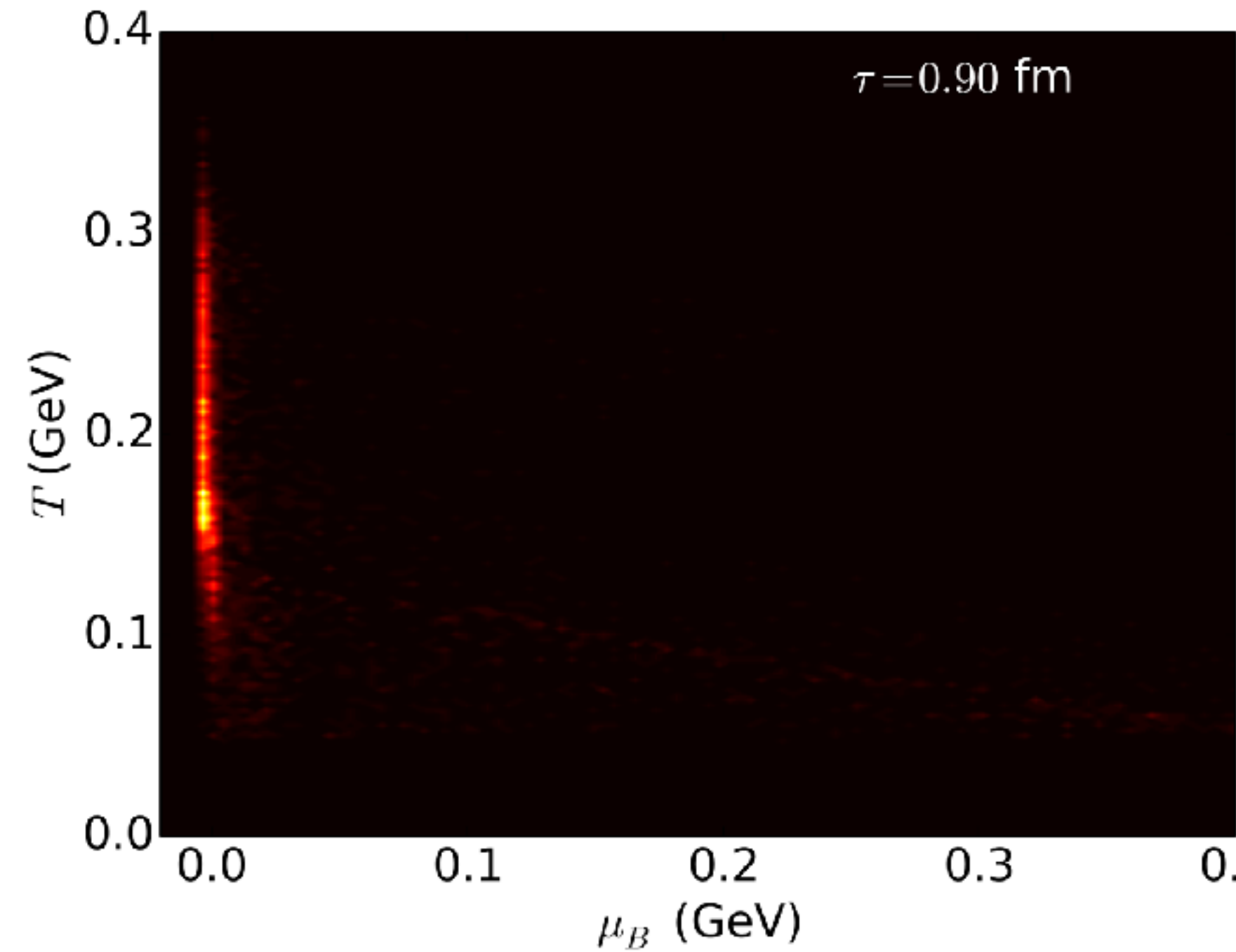
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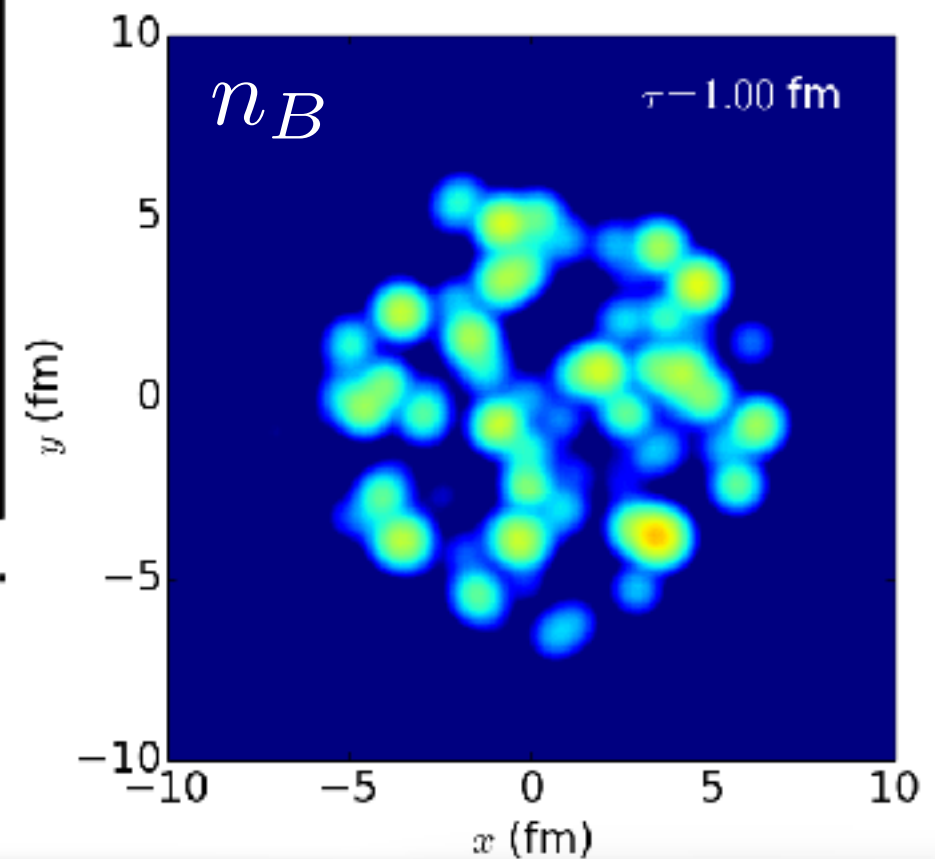
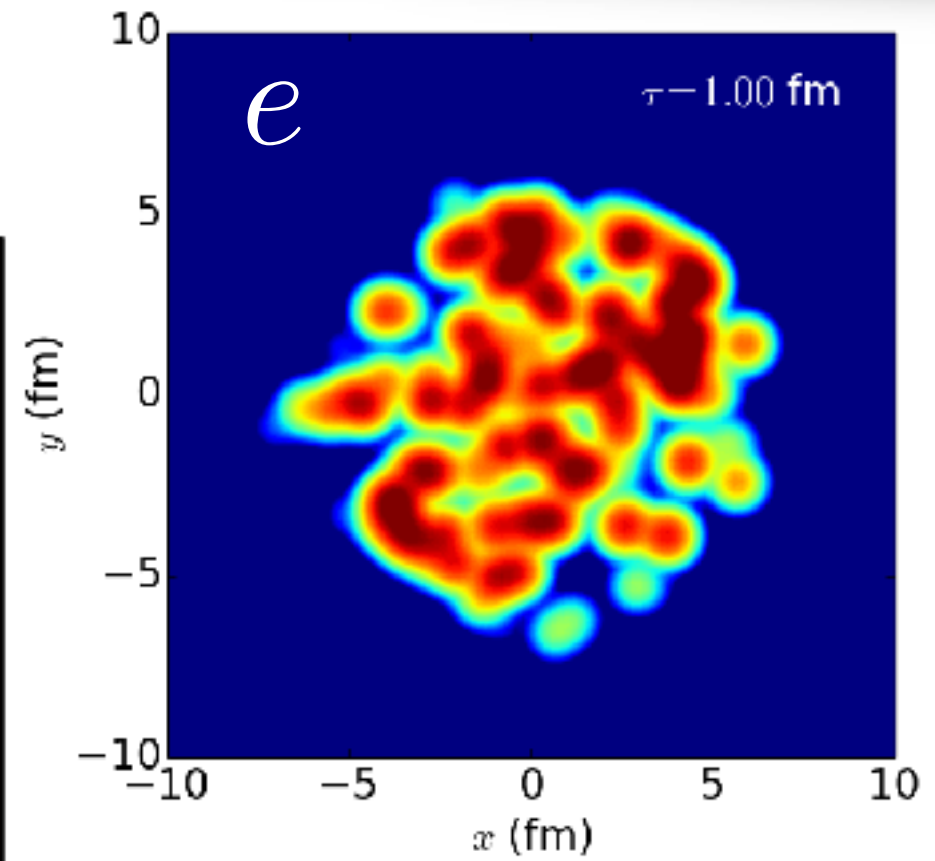
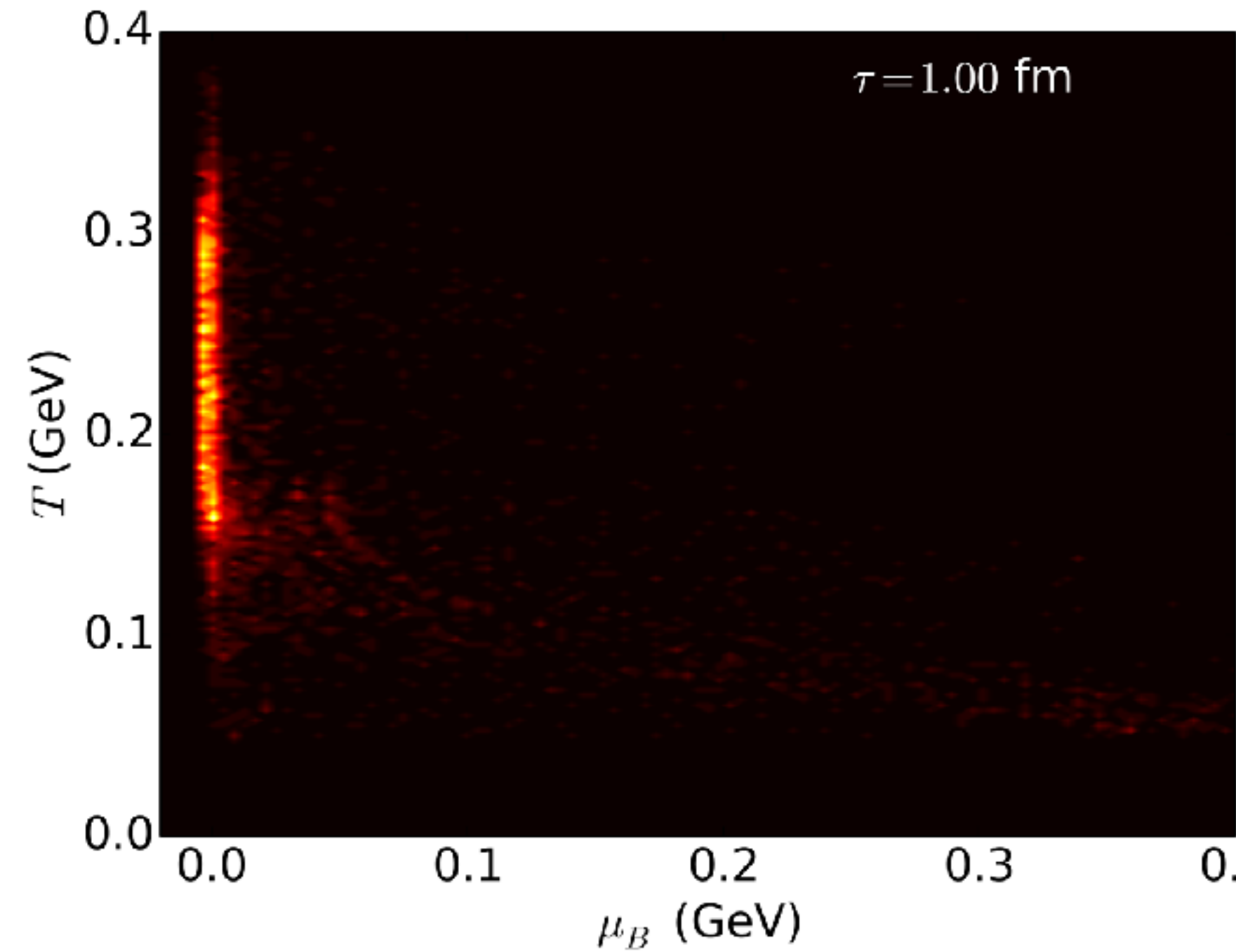
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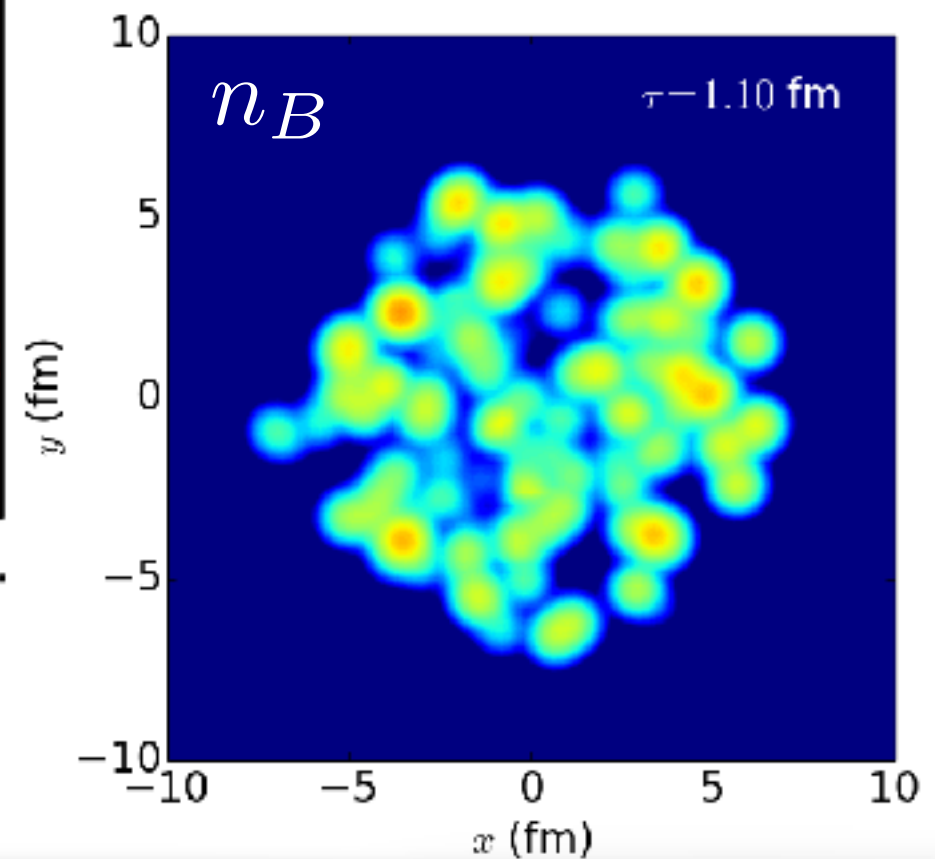
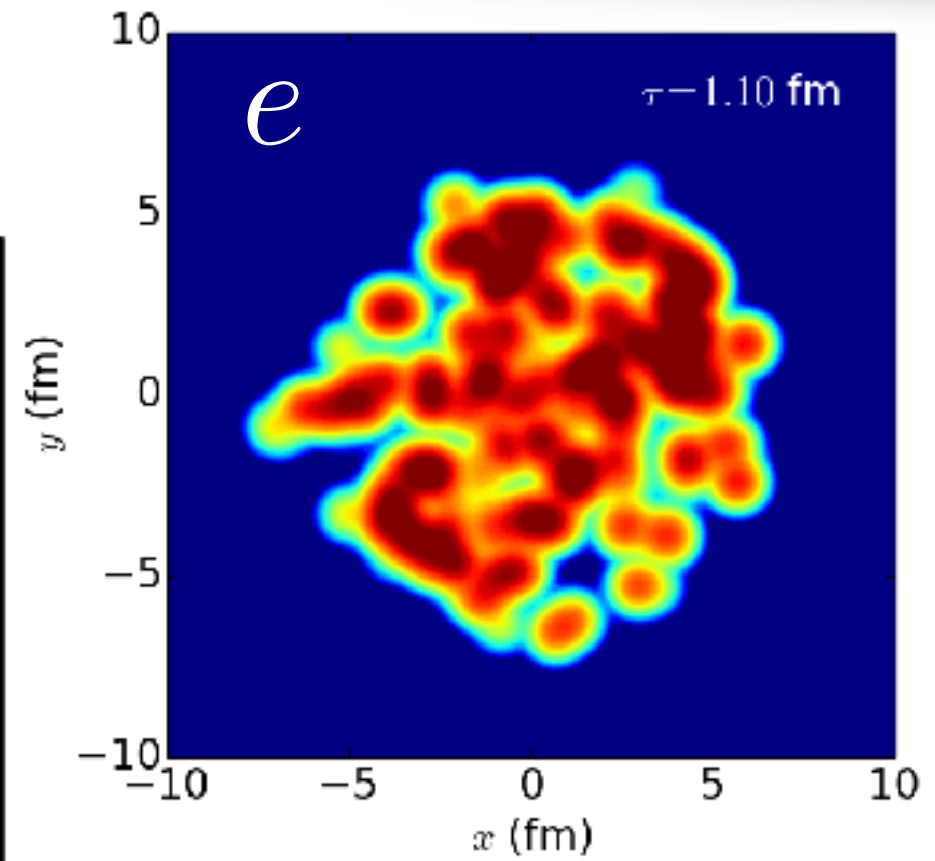
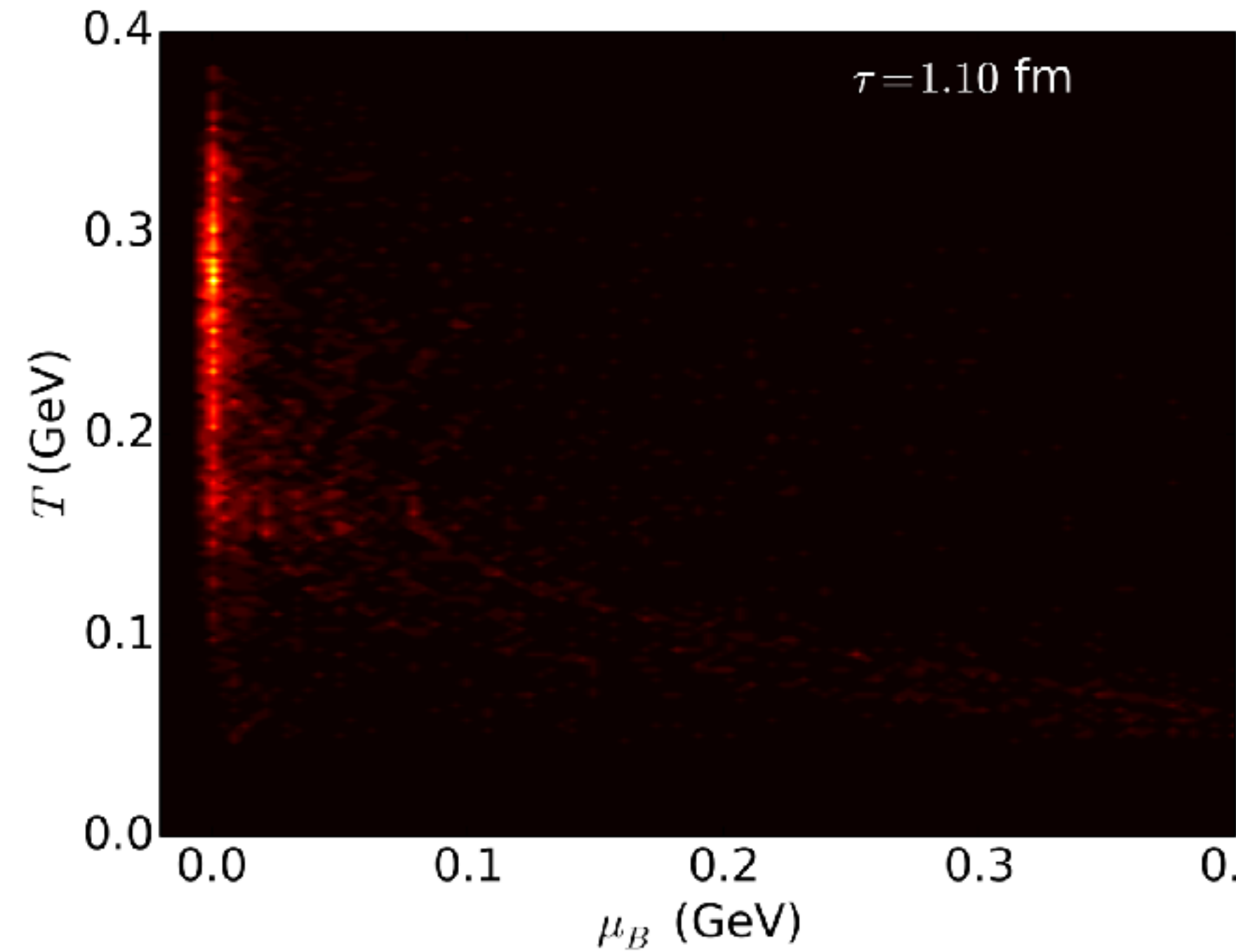
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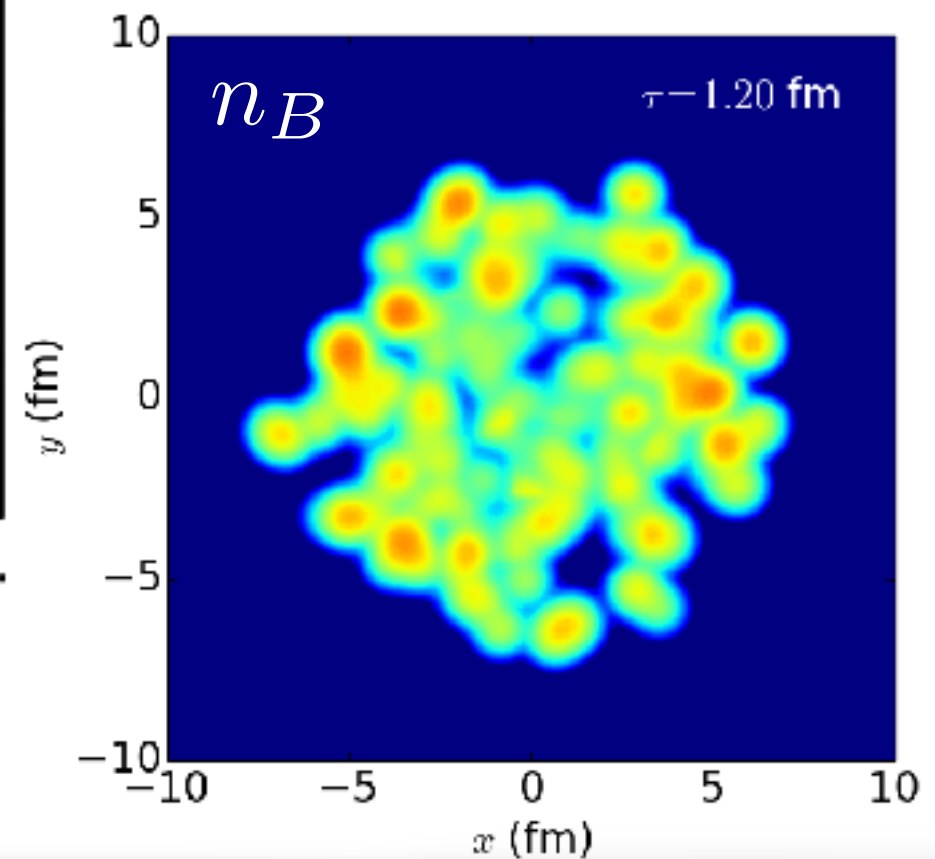
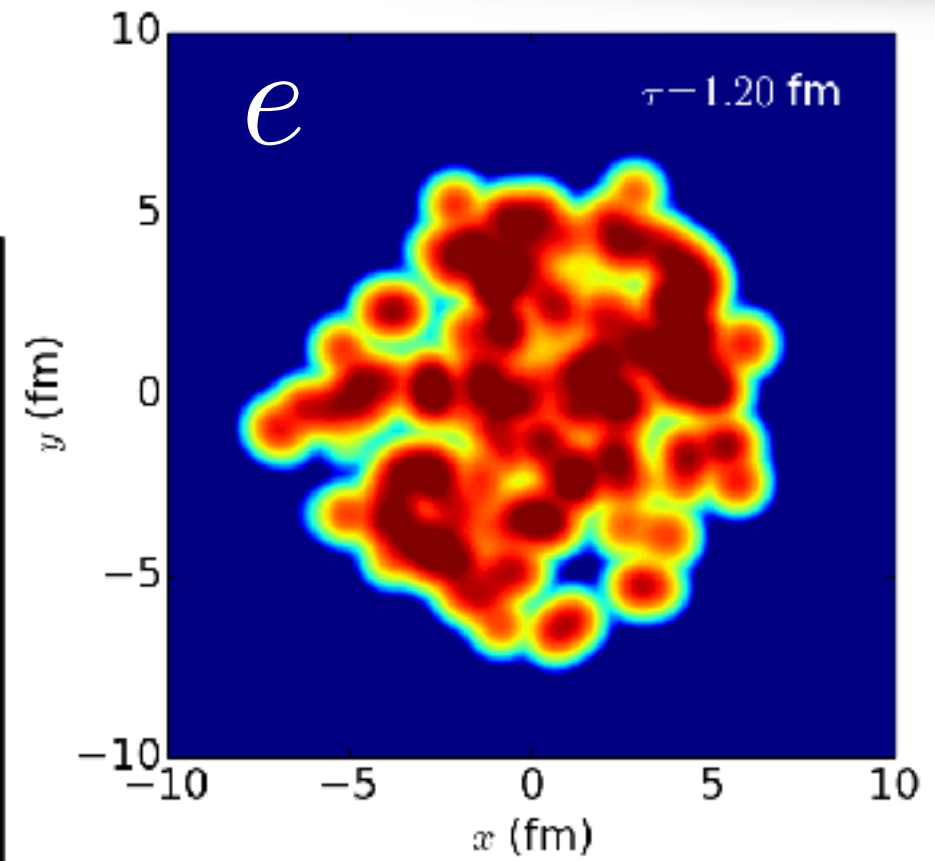
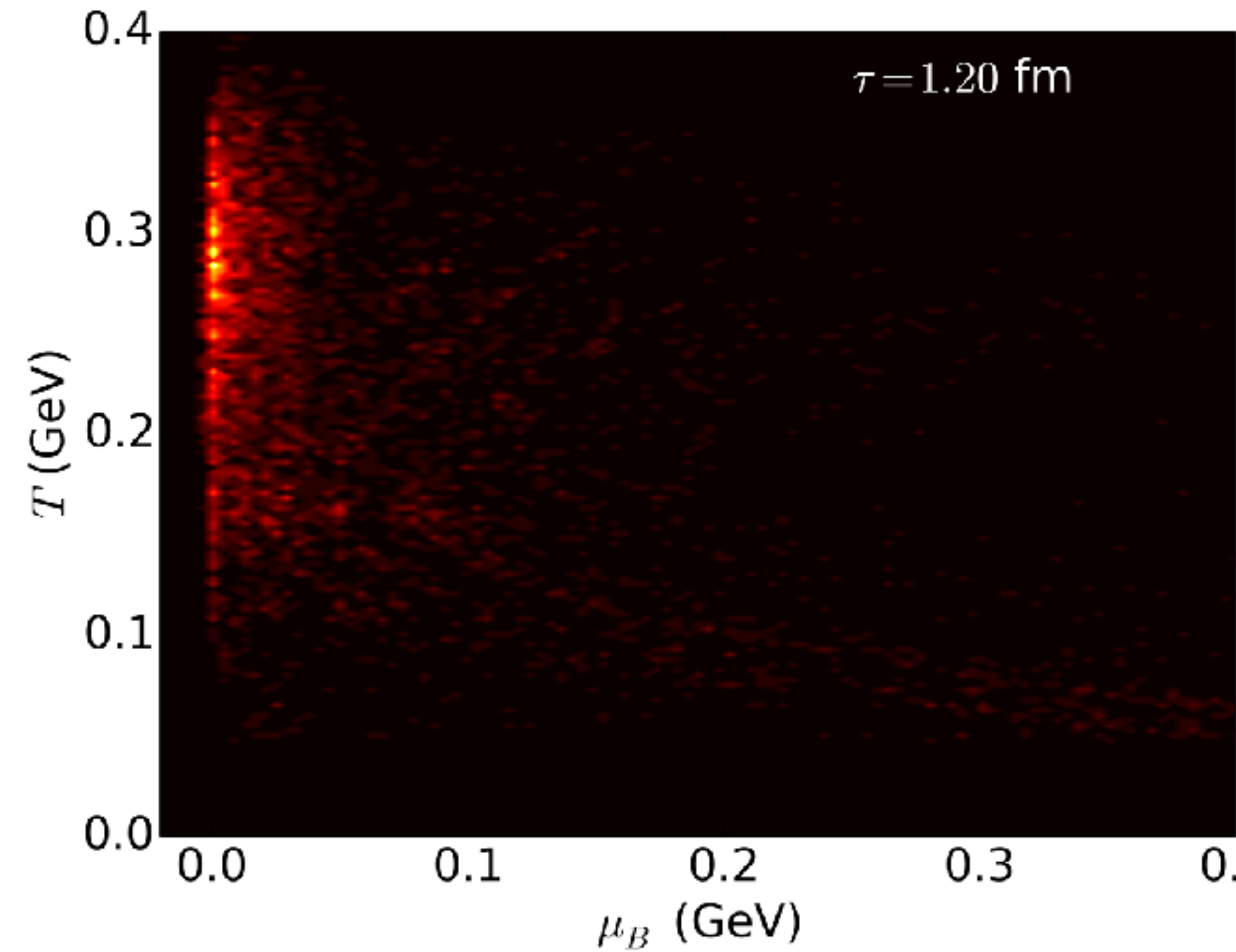
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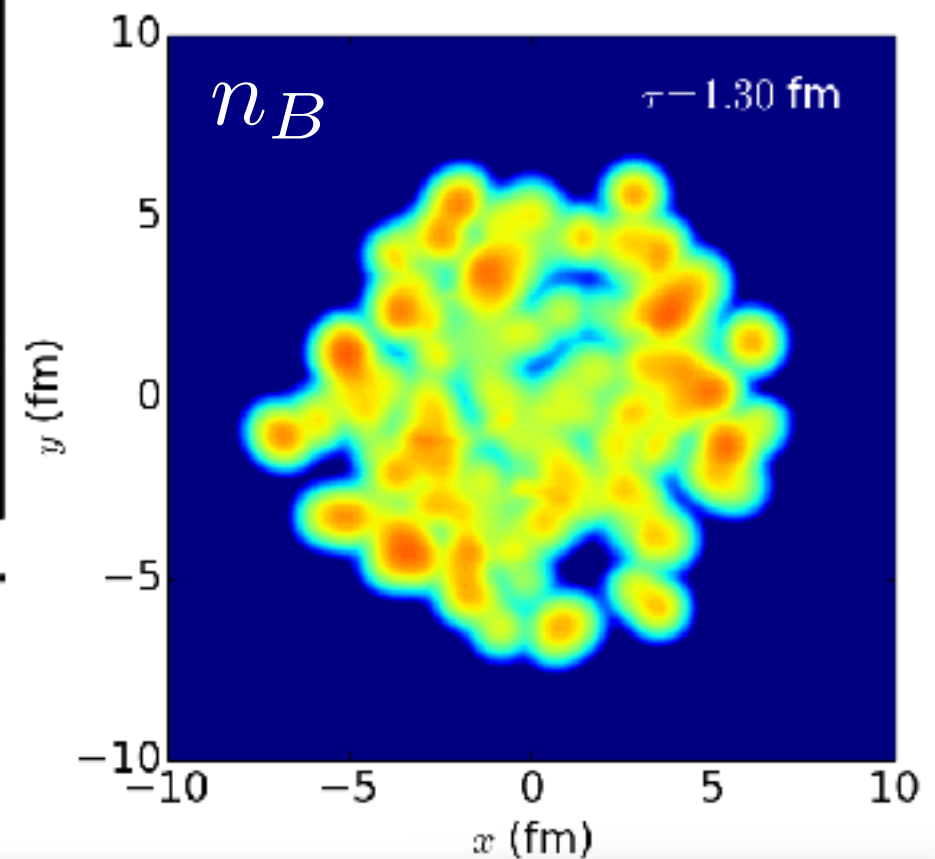
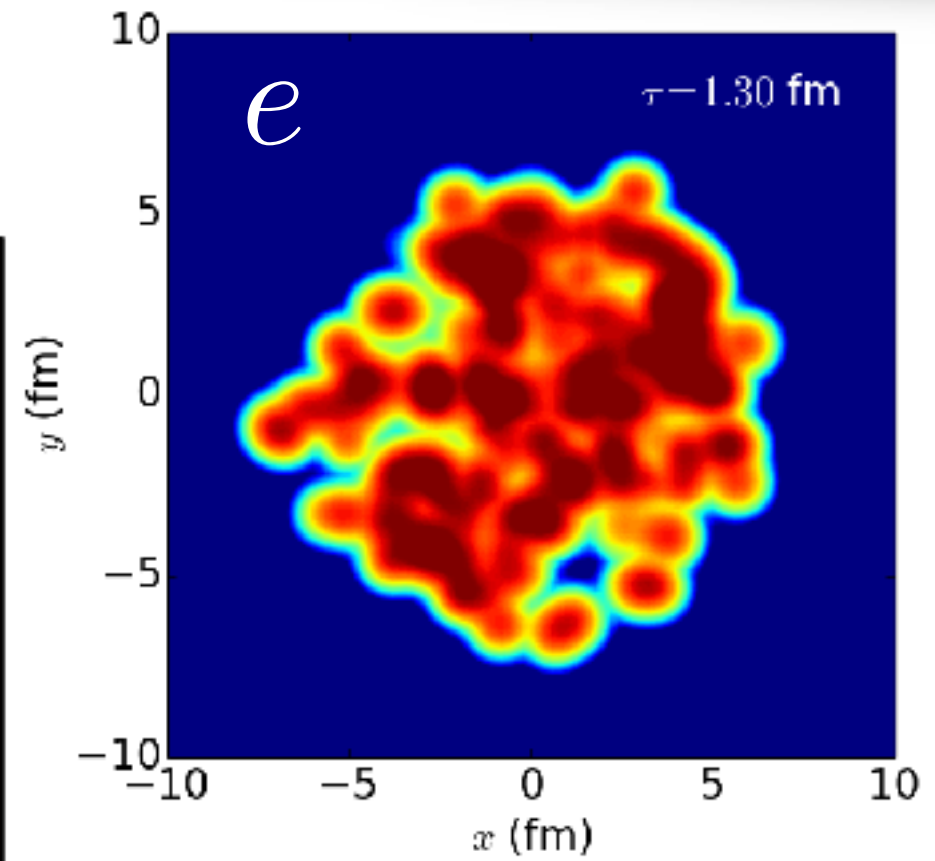
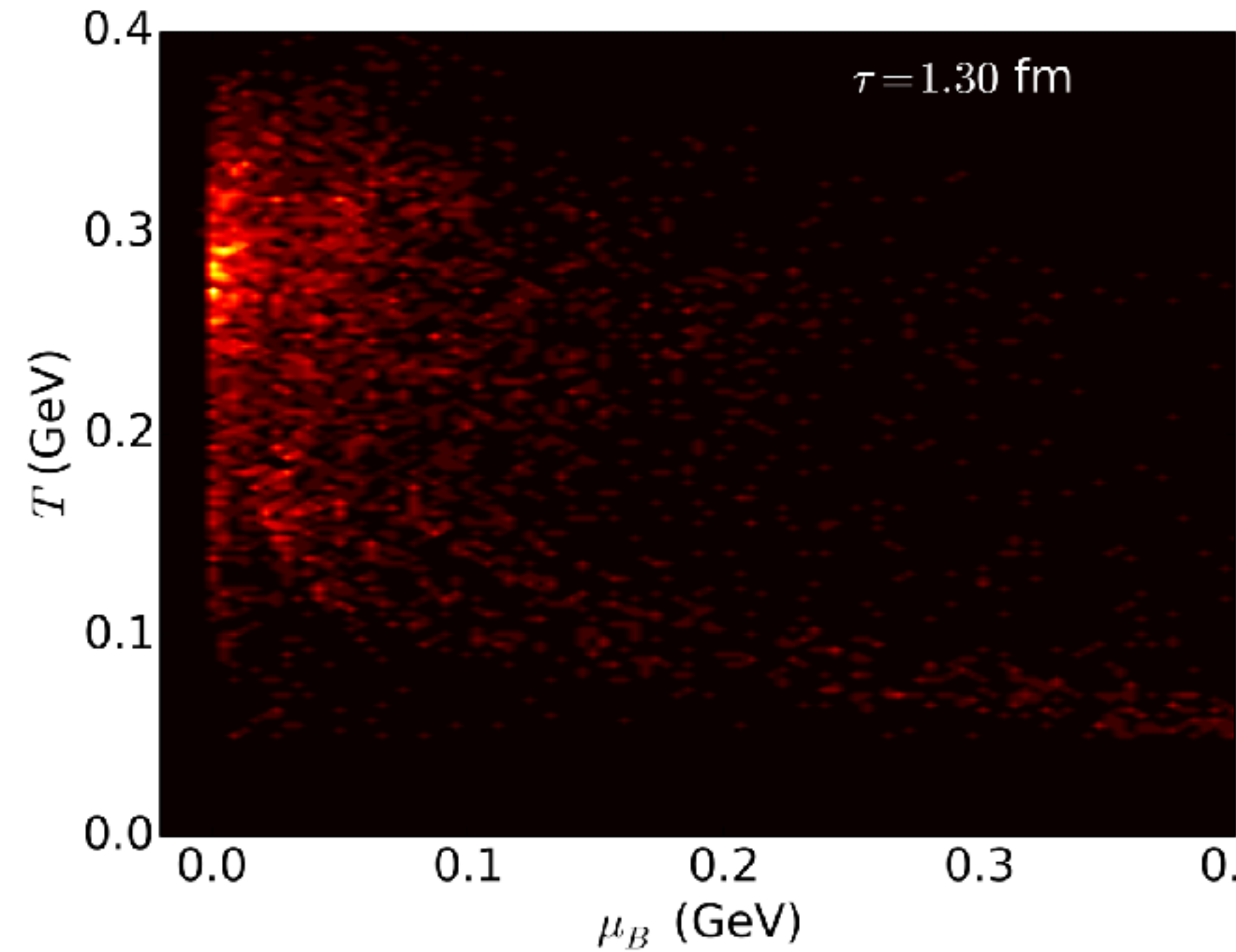
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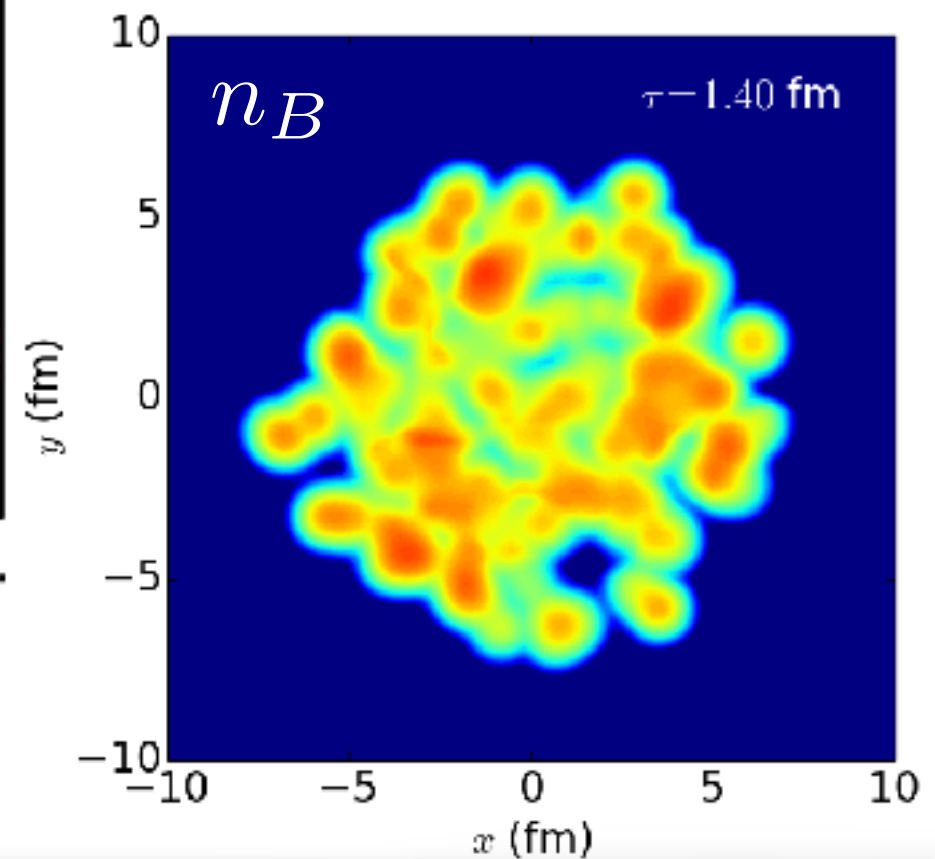
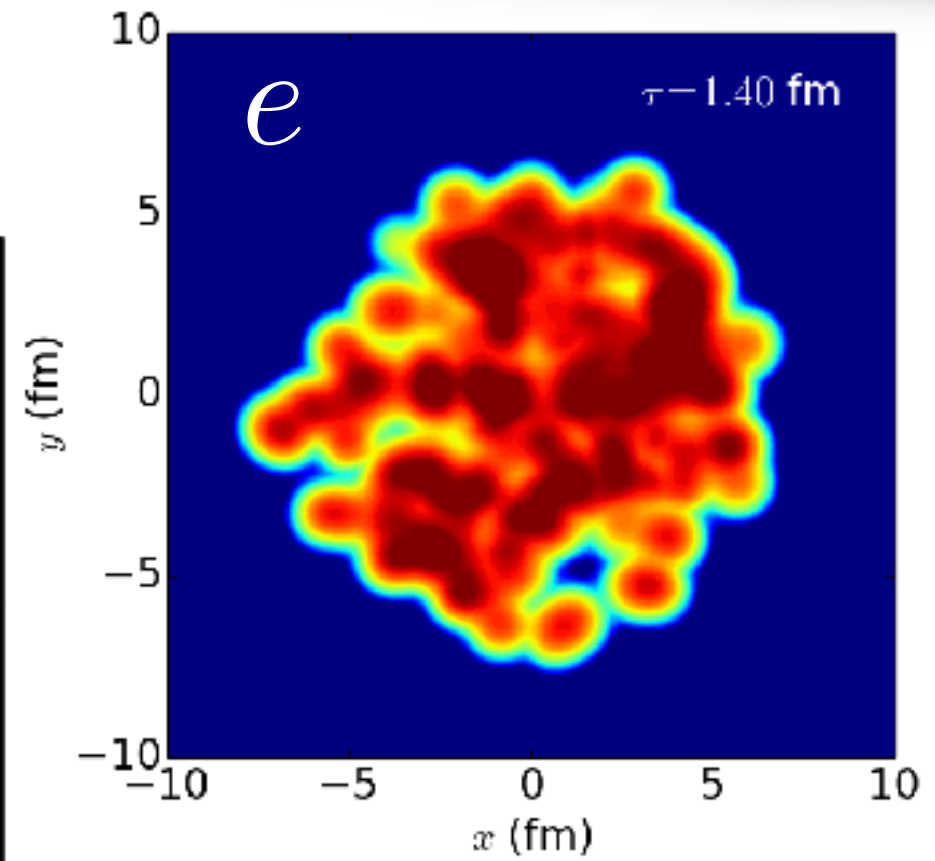
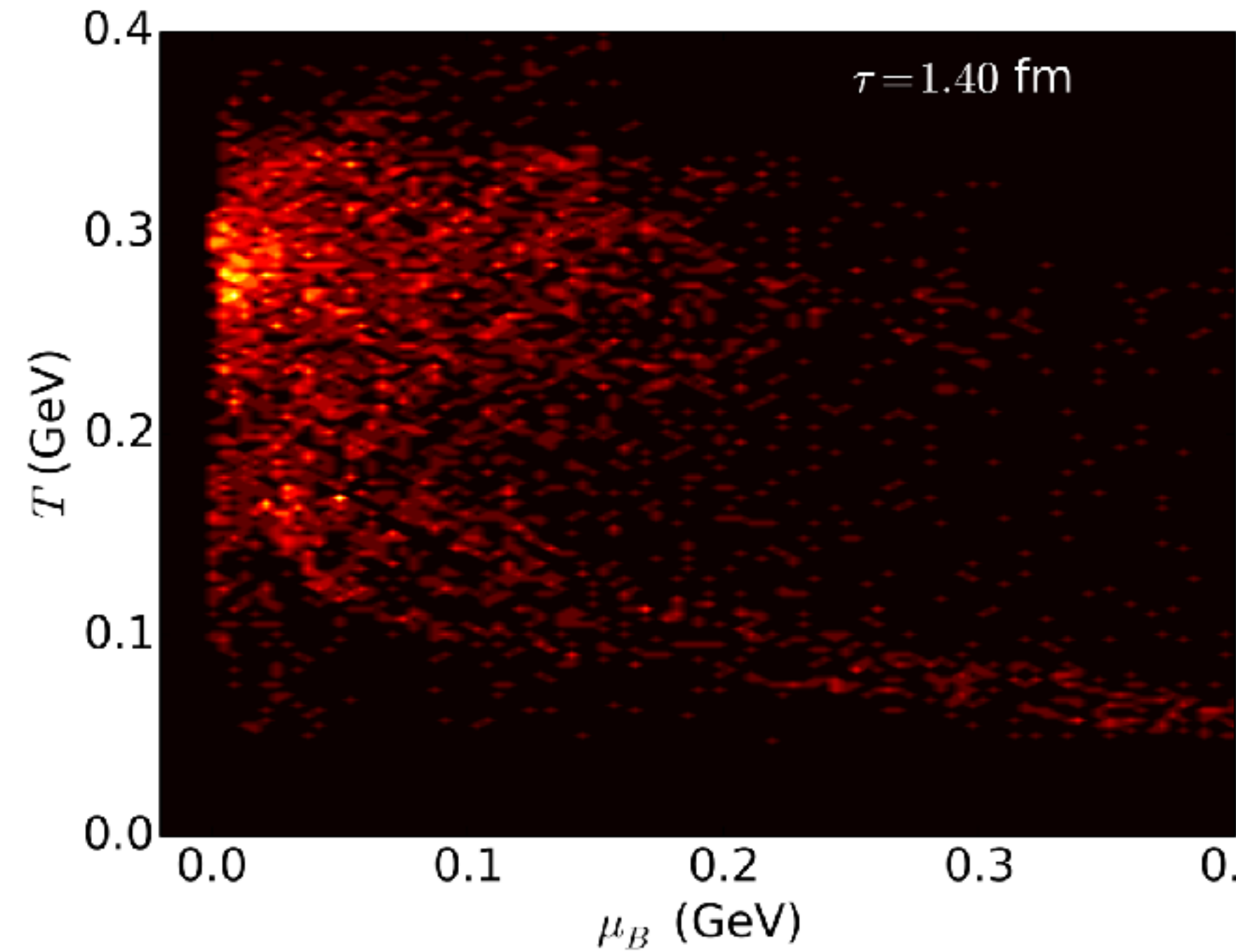
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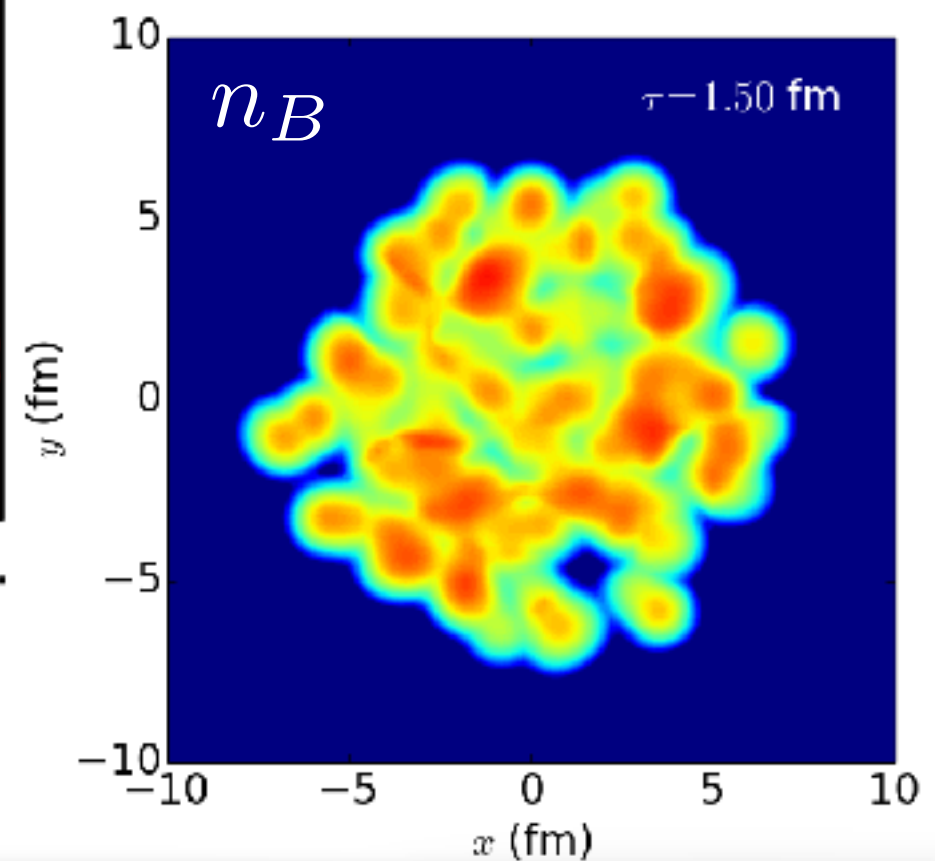
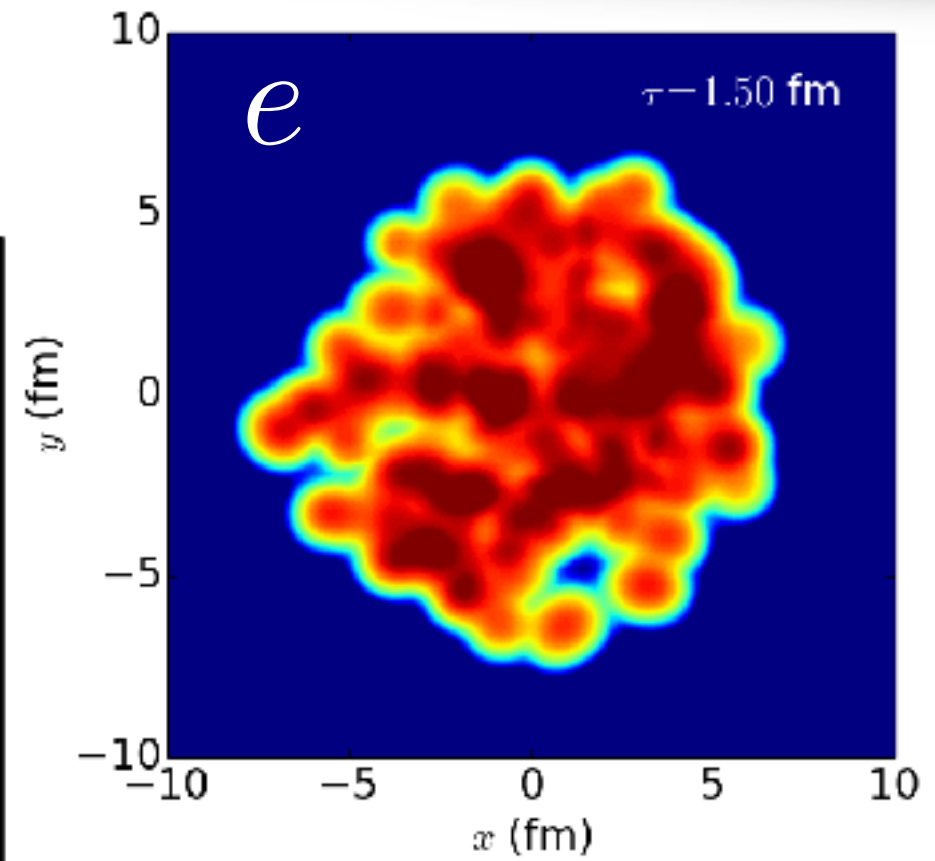
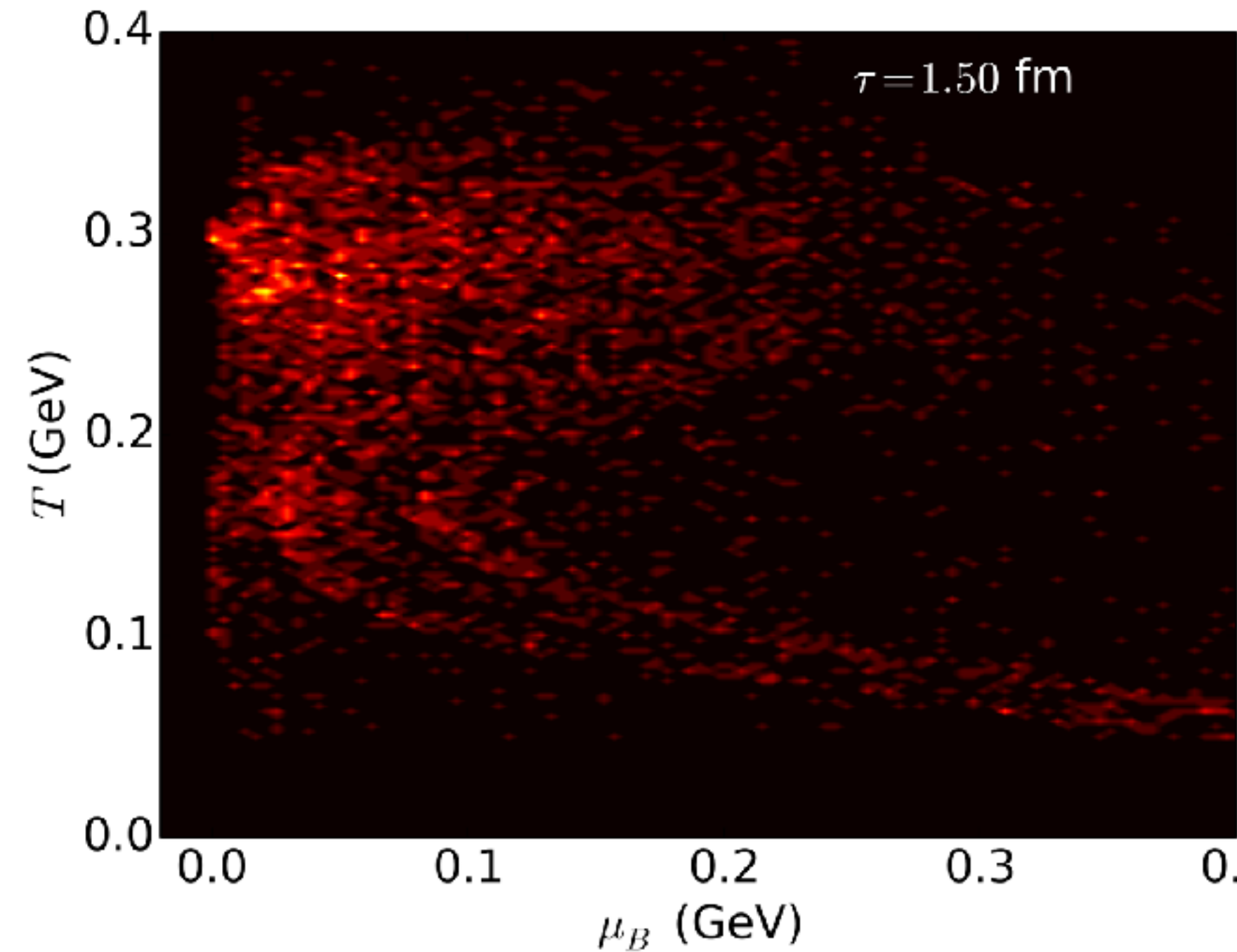
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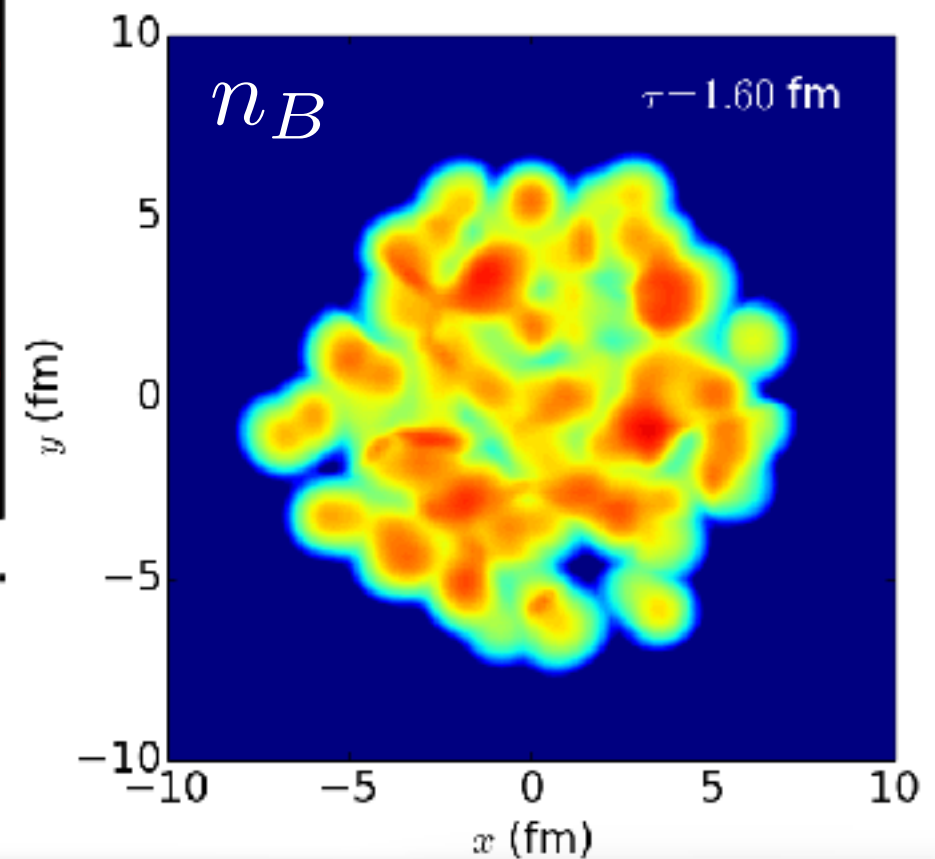
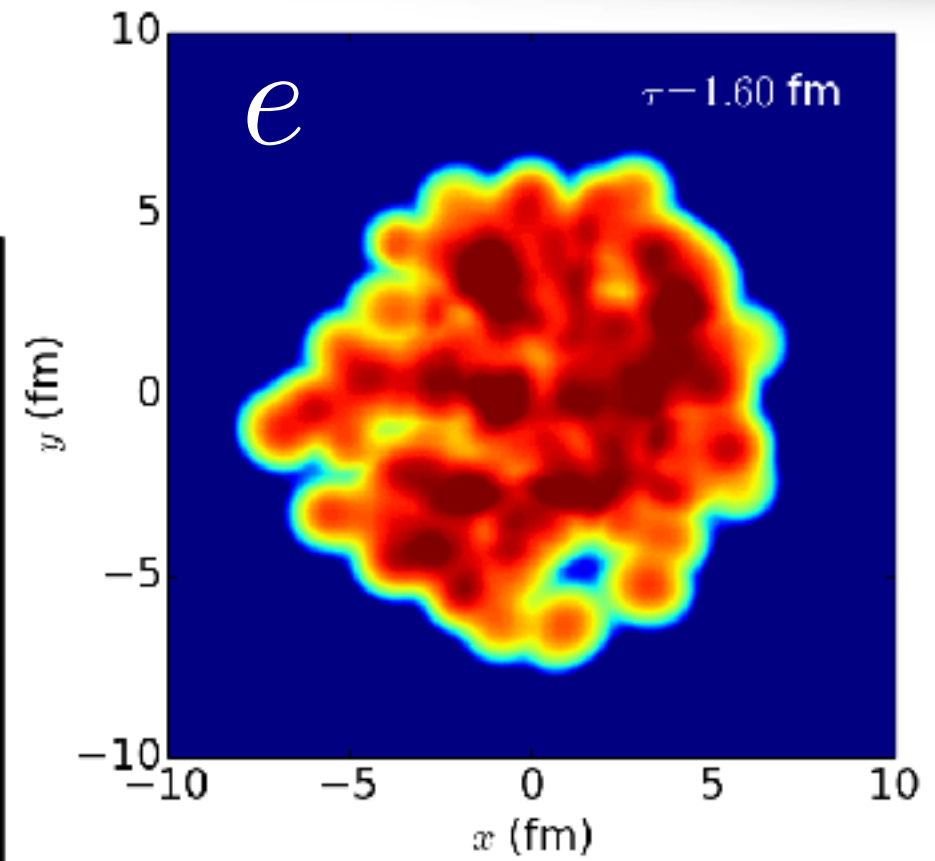
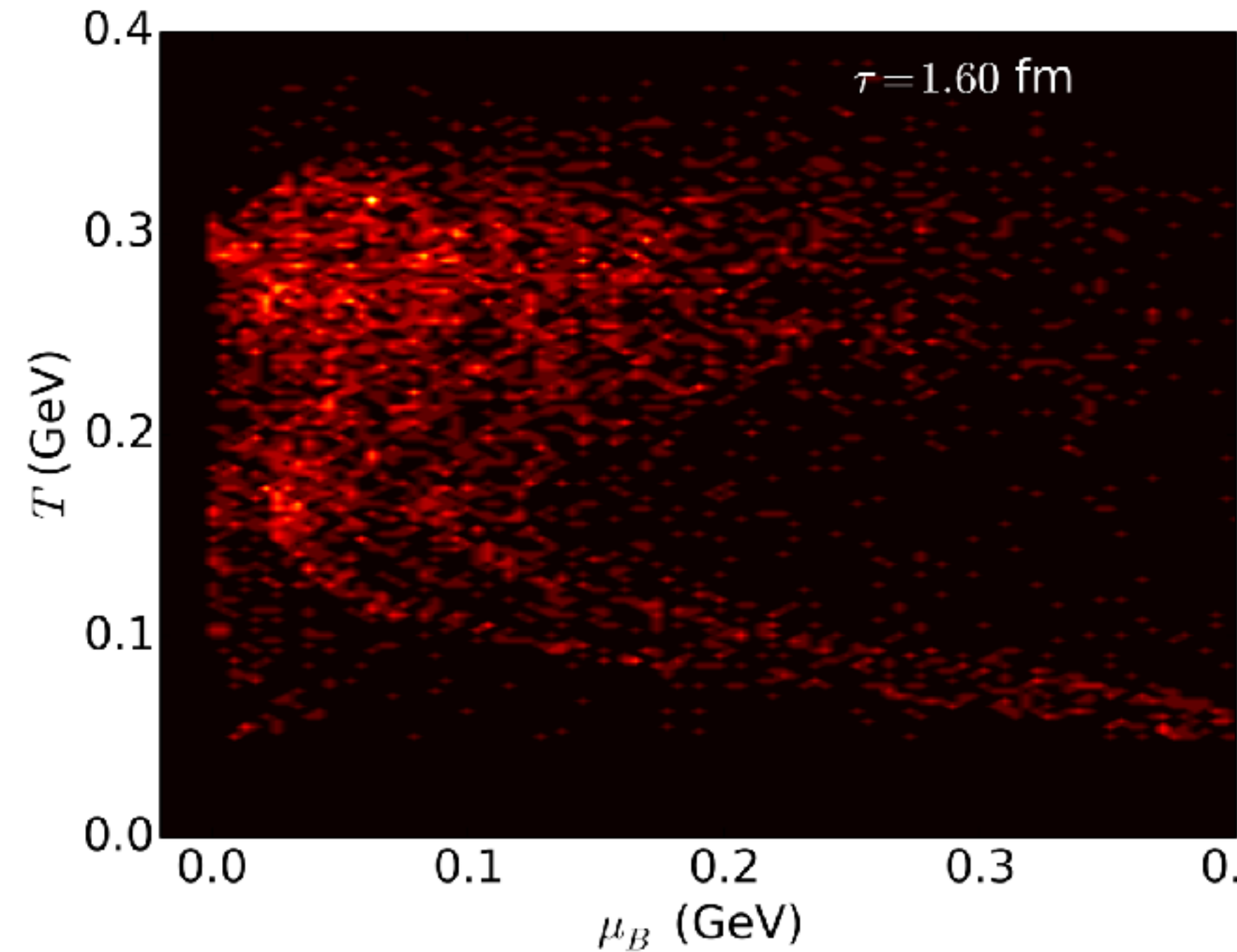
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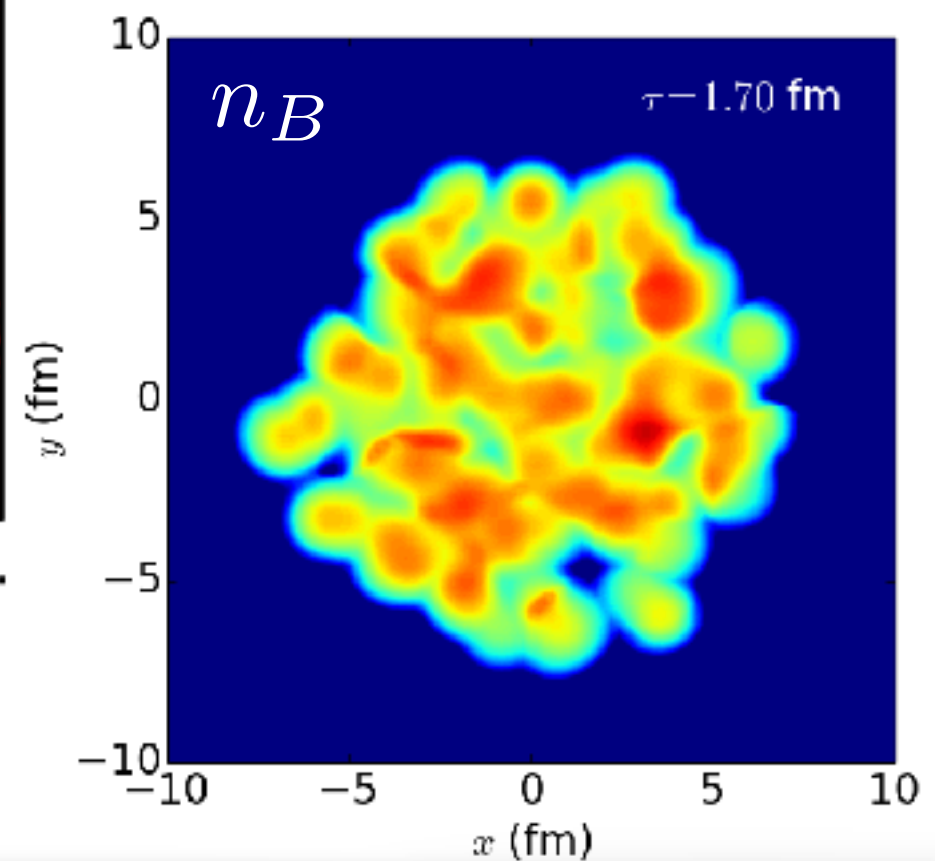
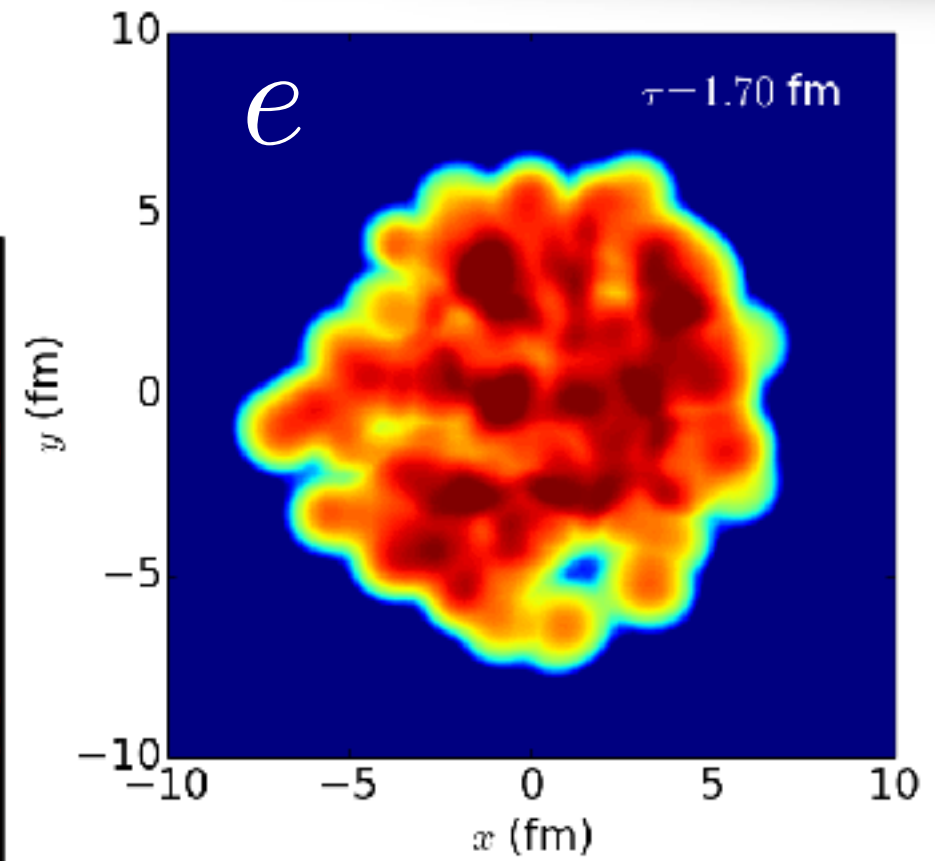
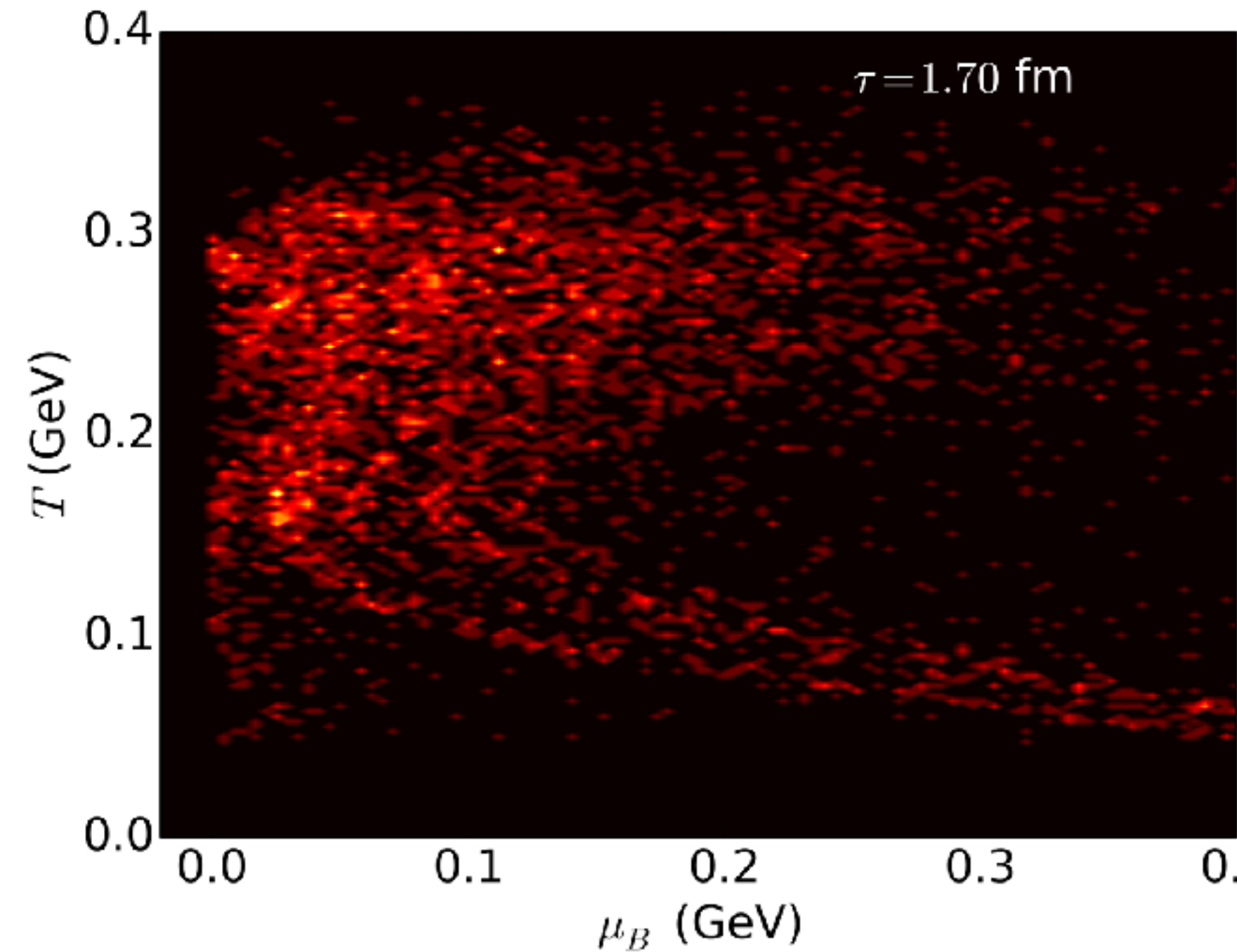
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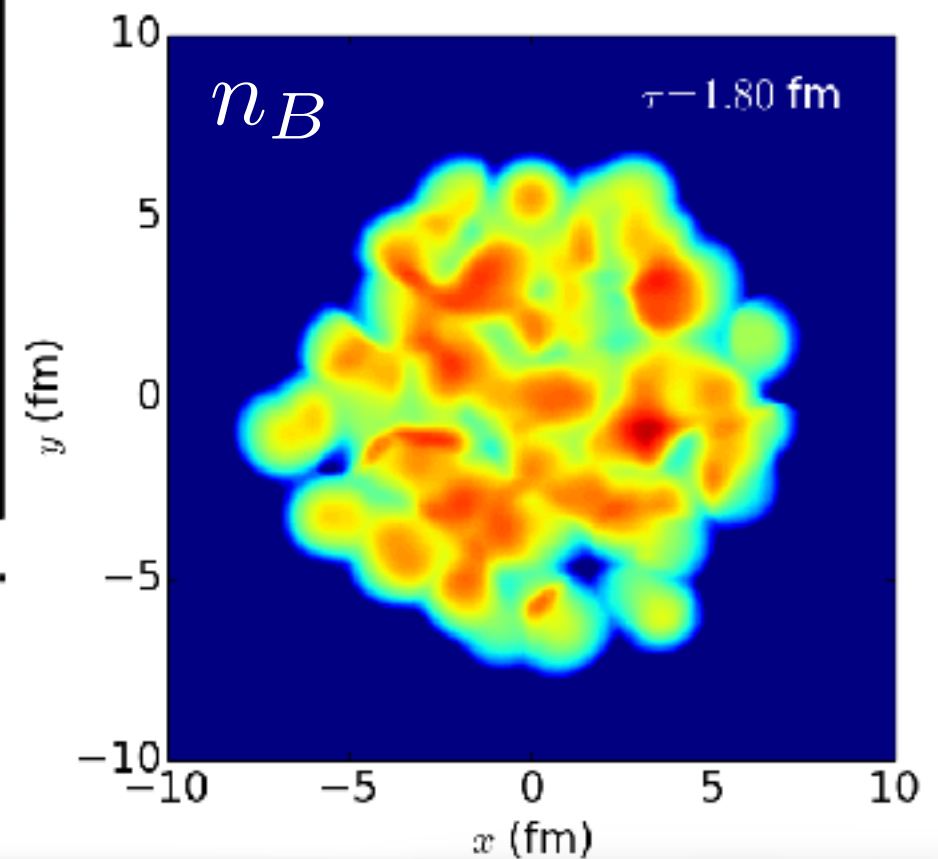
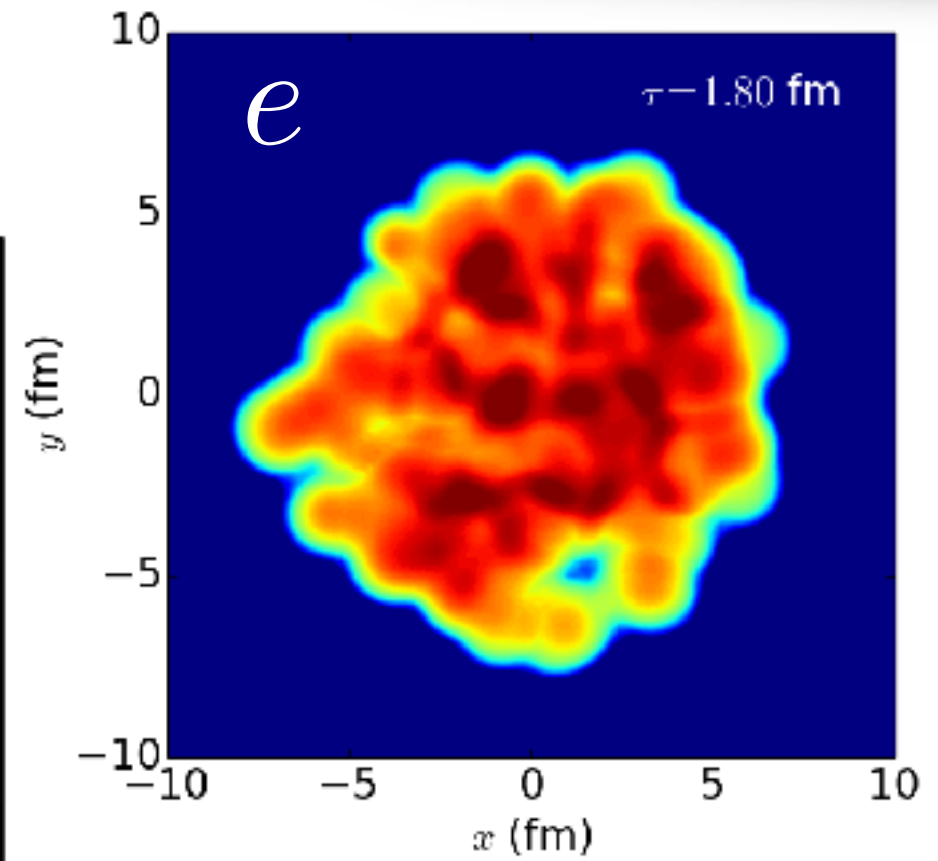
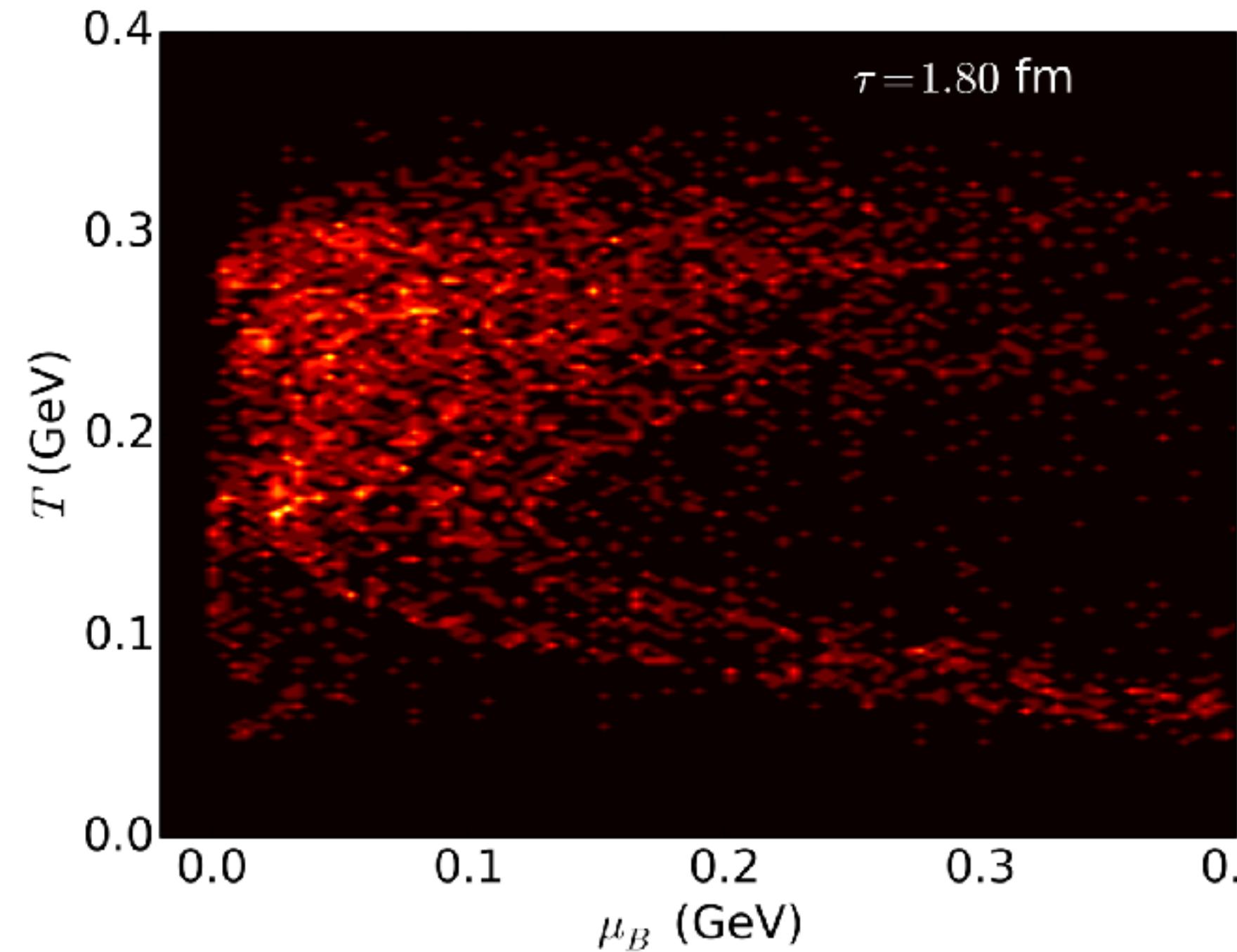
Fireball evolution in the QCD phase diagram



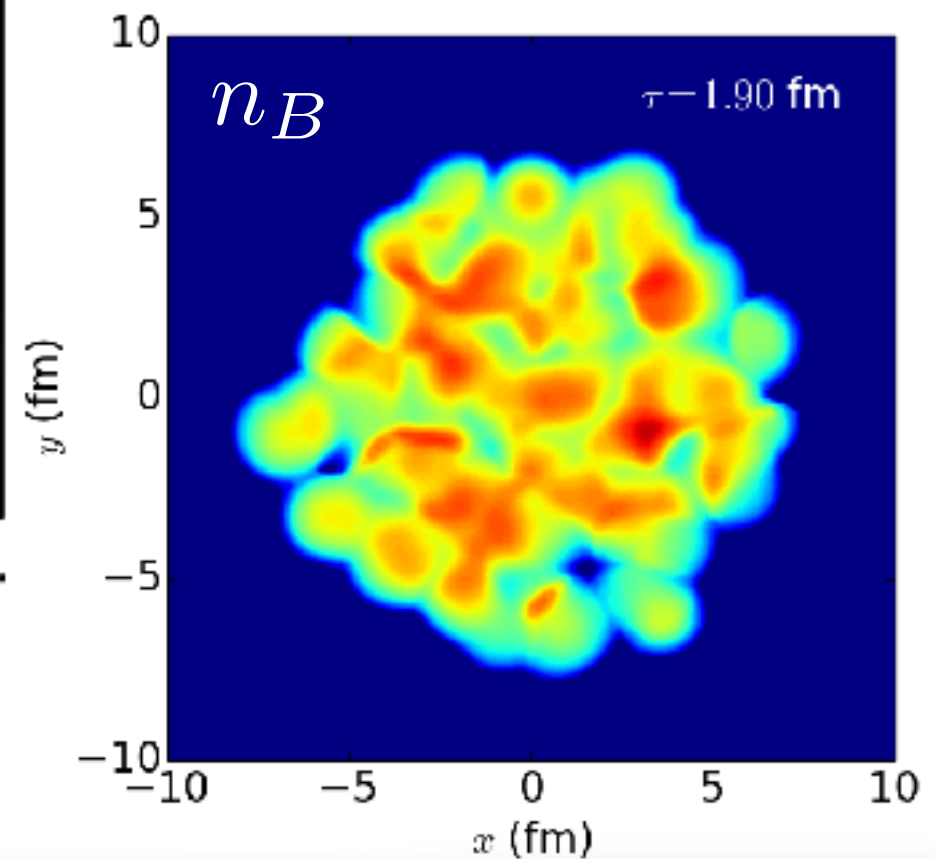
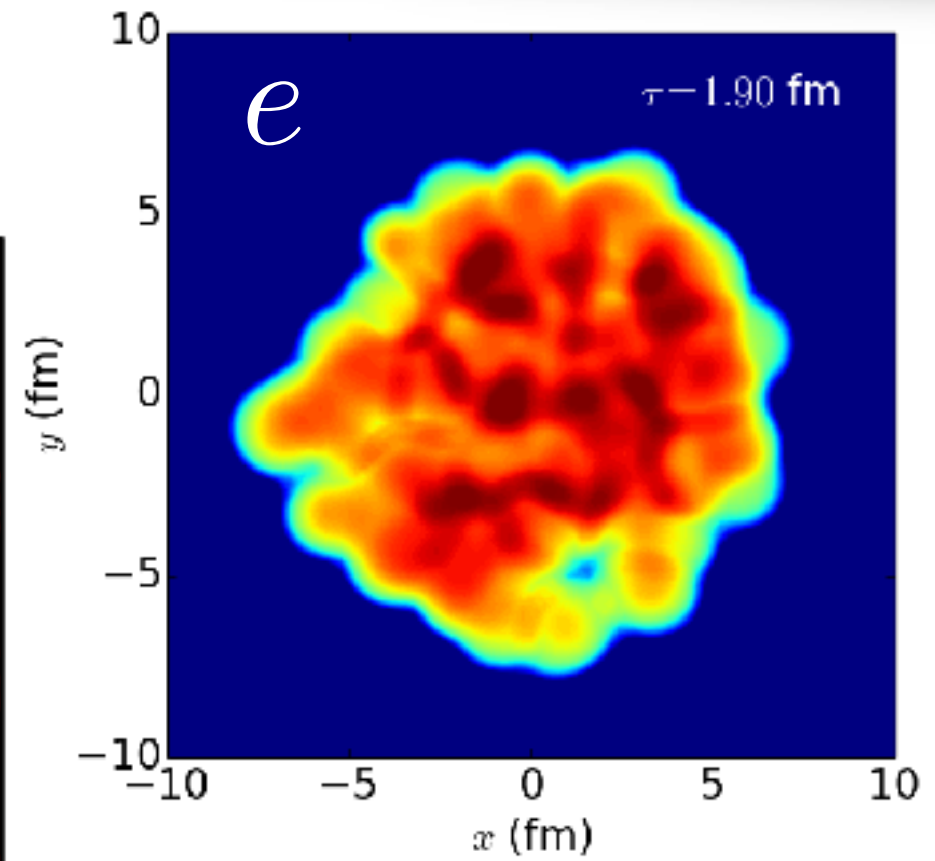
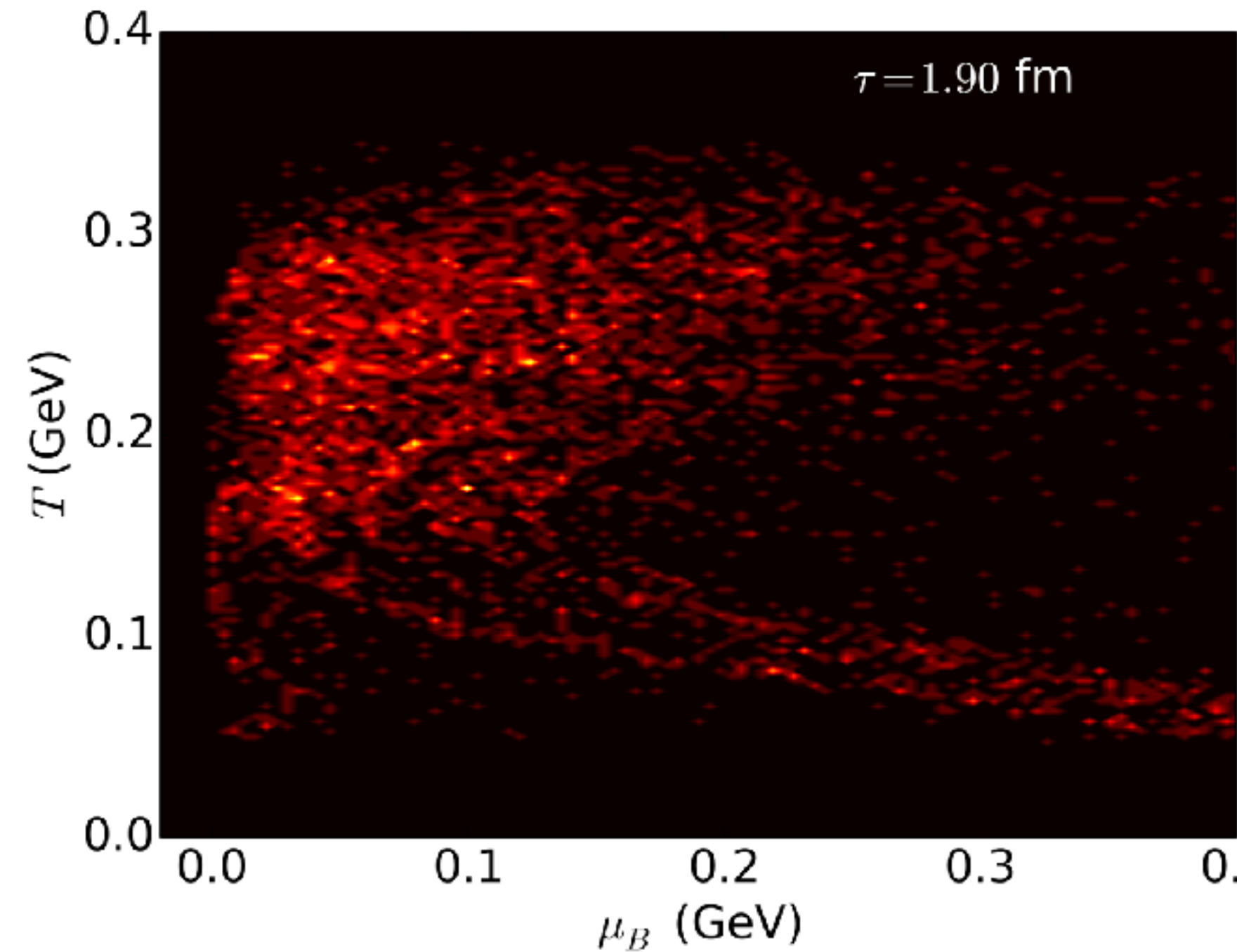
Fireball evolution in the QCD phase diagram



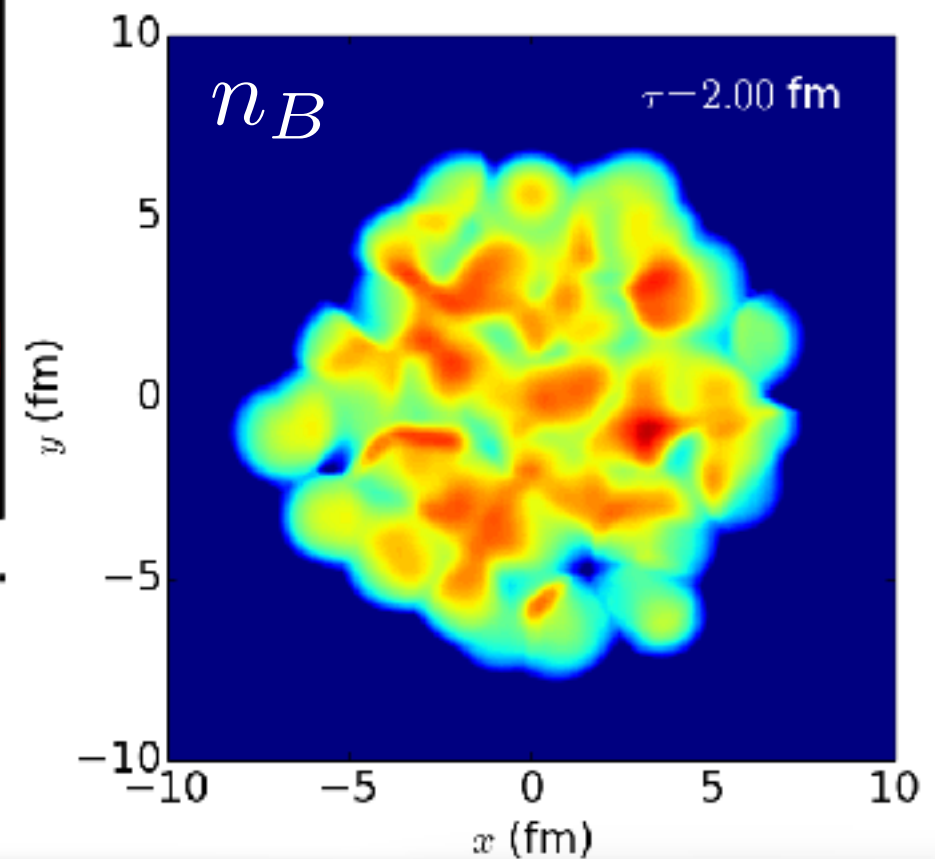
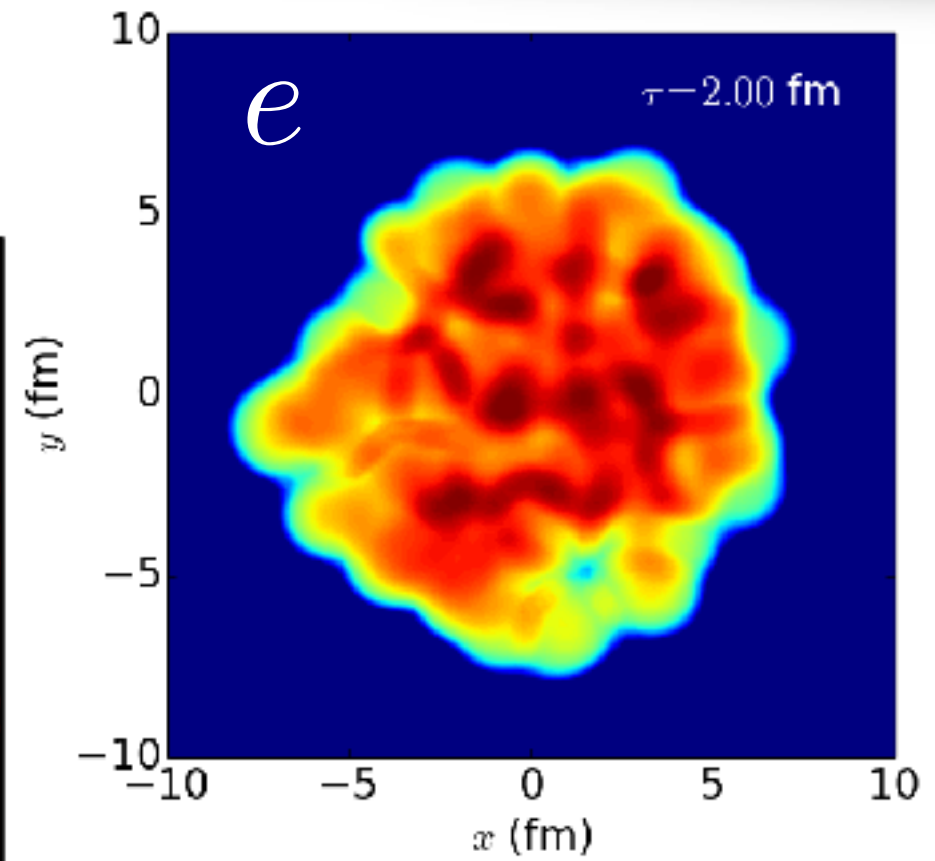
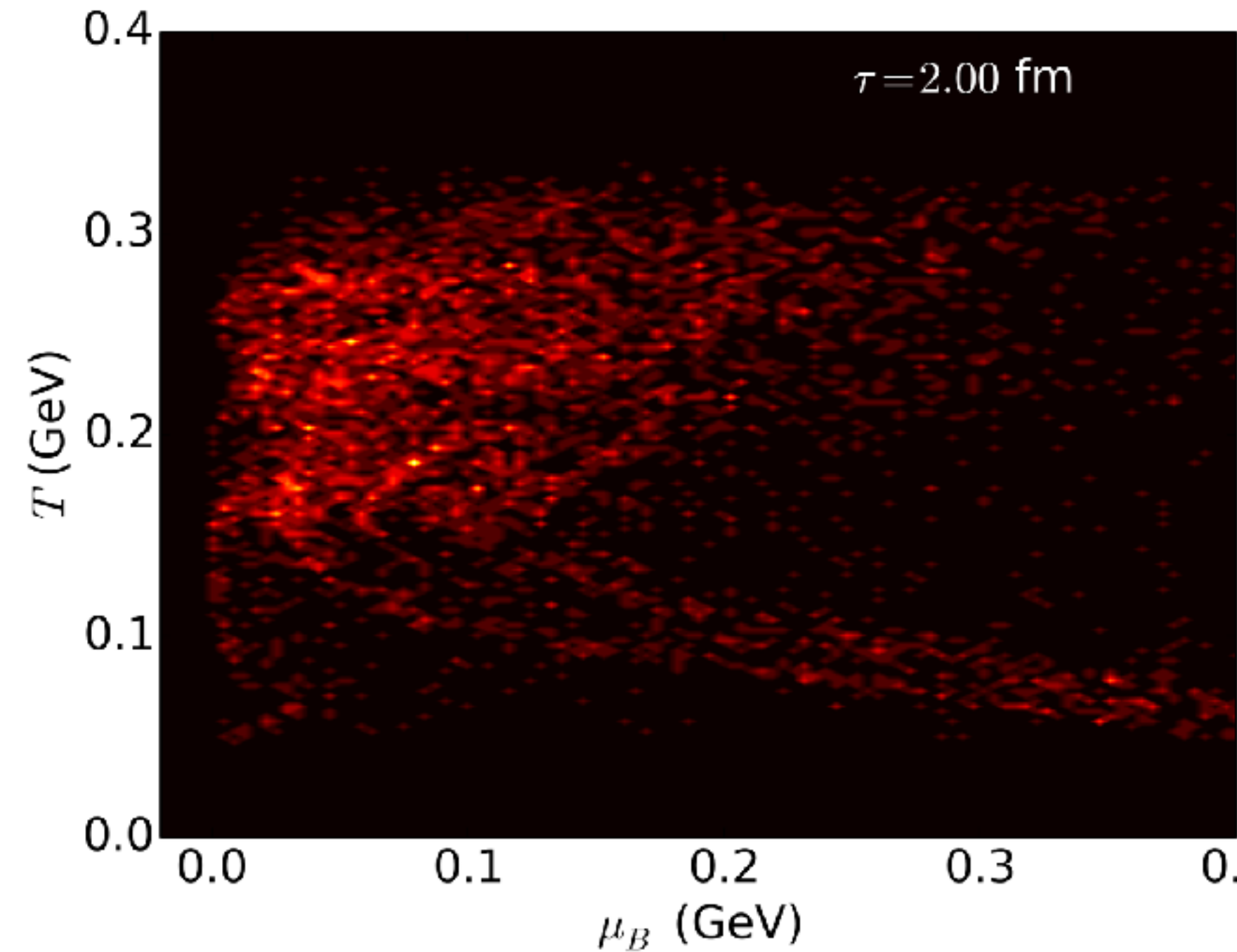
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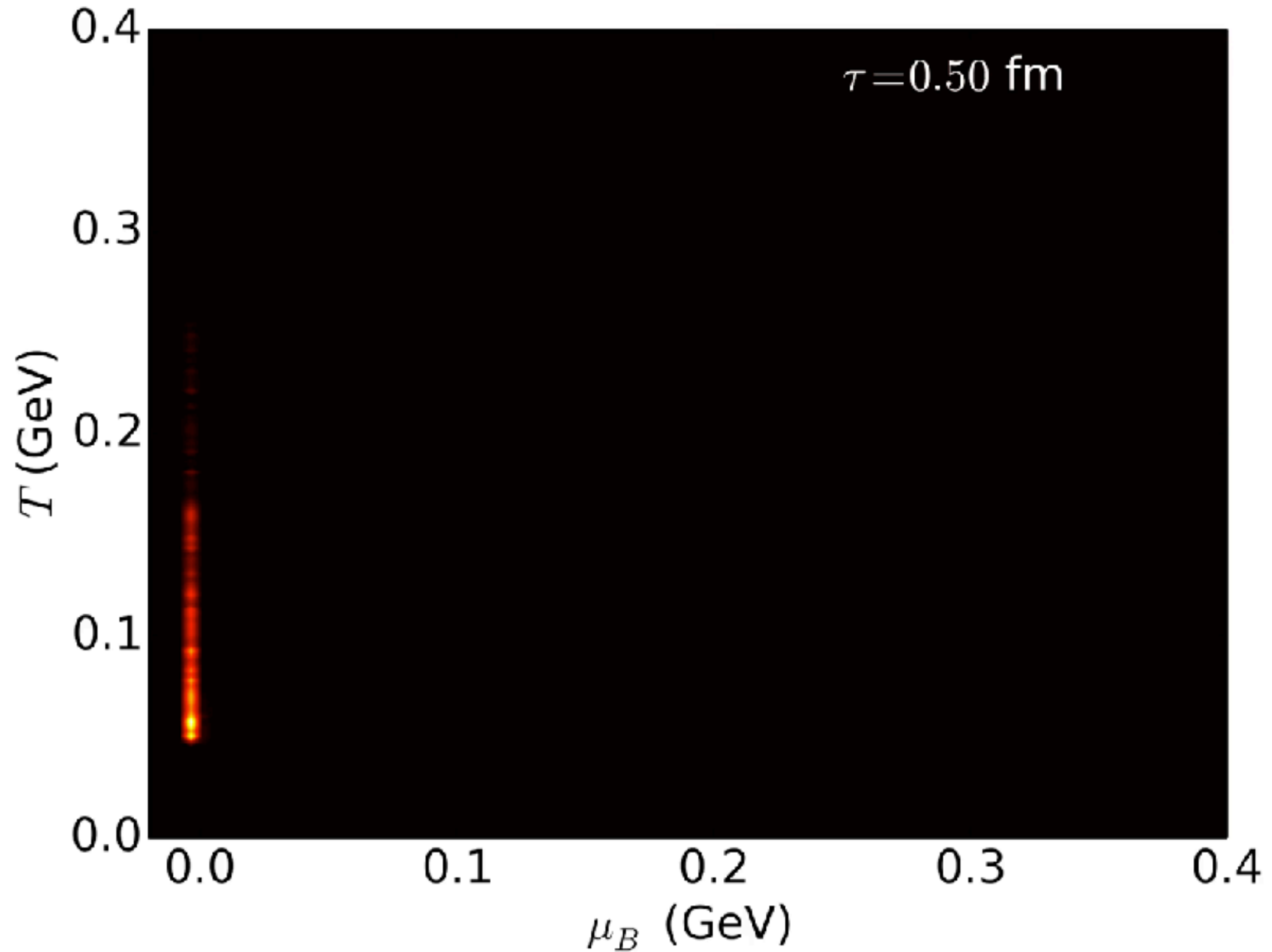
Fireball evolution in the QCD phase diagram



Fireball evolution in the QCD phase diagram



Fireball evolution in the QCD phase diagram



Jet energy deposition and hydro response

A toy model

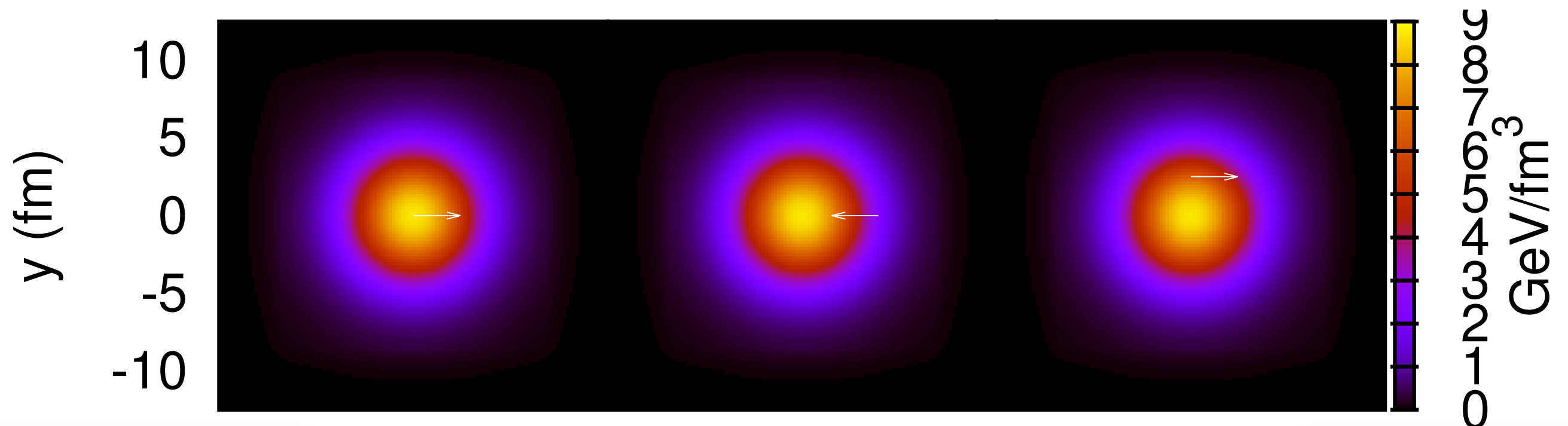
Collaboration with Mayank Singh

$$\frac{dE}{dx} = \frac{d|\vec{P}|}{dt} = -1\text{GeV}/\text{fm}$$

$$j_{\text{source}}^{\mu}(t, \vec{x}) = -\frac{d|\vec{P}|}{dt} \frac{P^{\mu}}{P^0} \delta^3(\vec{x} - \vec{x}_i)$$

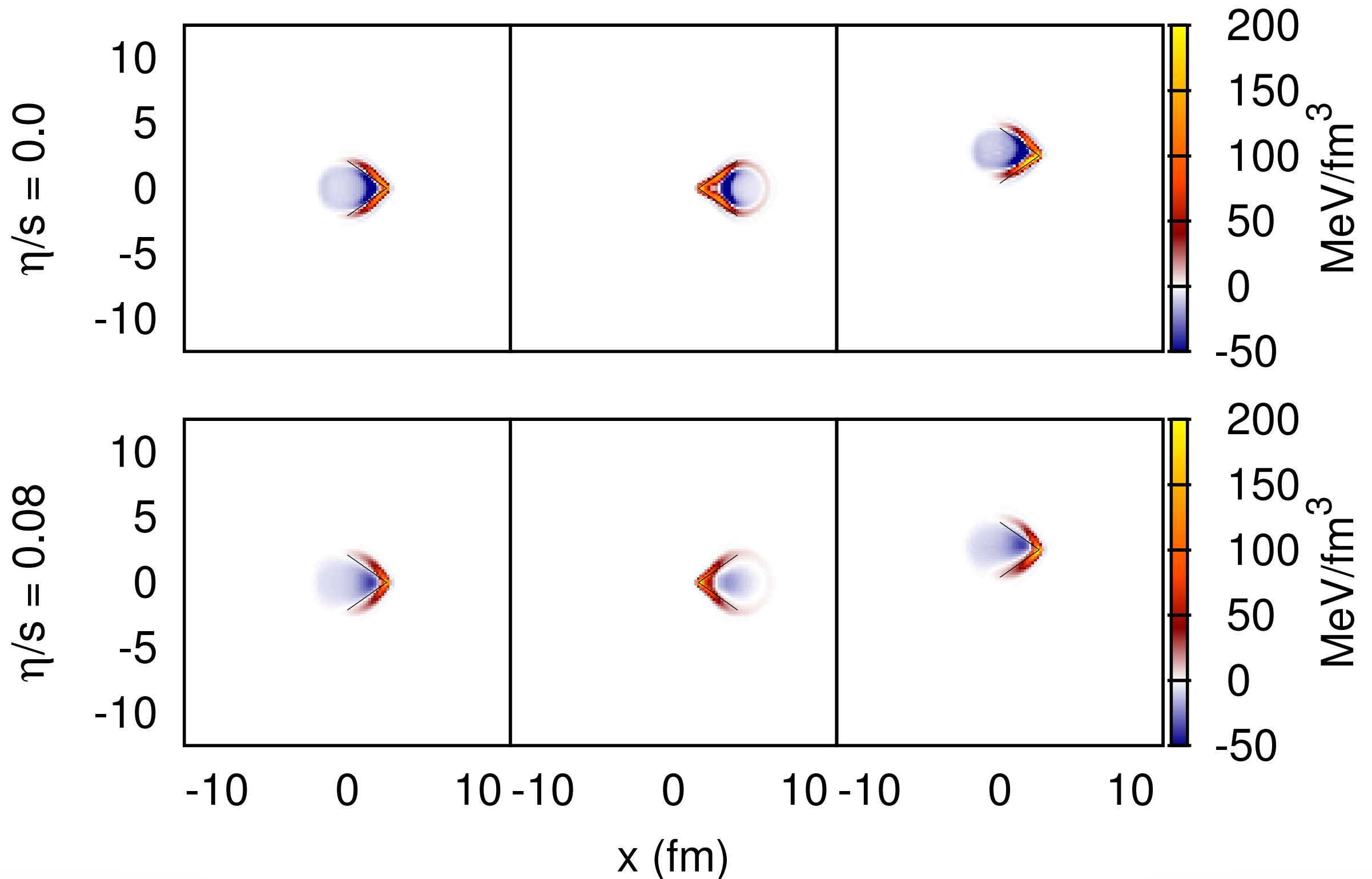
$$\vec{x}_i = \vec{x}_0 + \frac{\vec{P}}{P^0} t$$

$$\tau = 3.4\text{ fm}$$



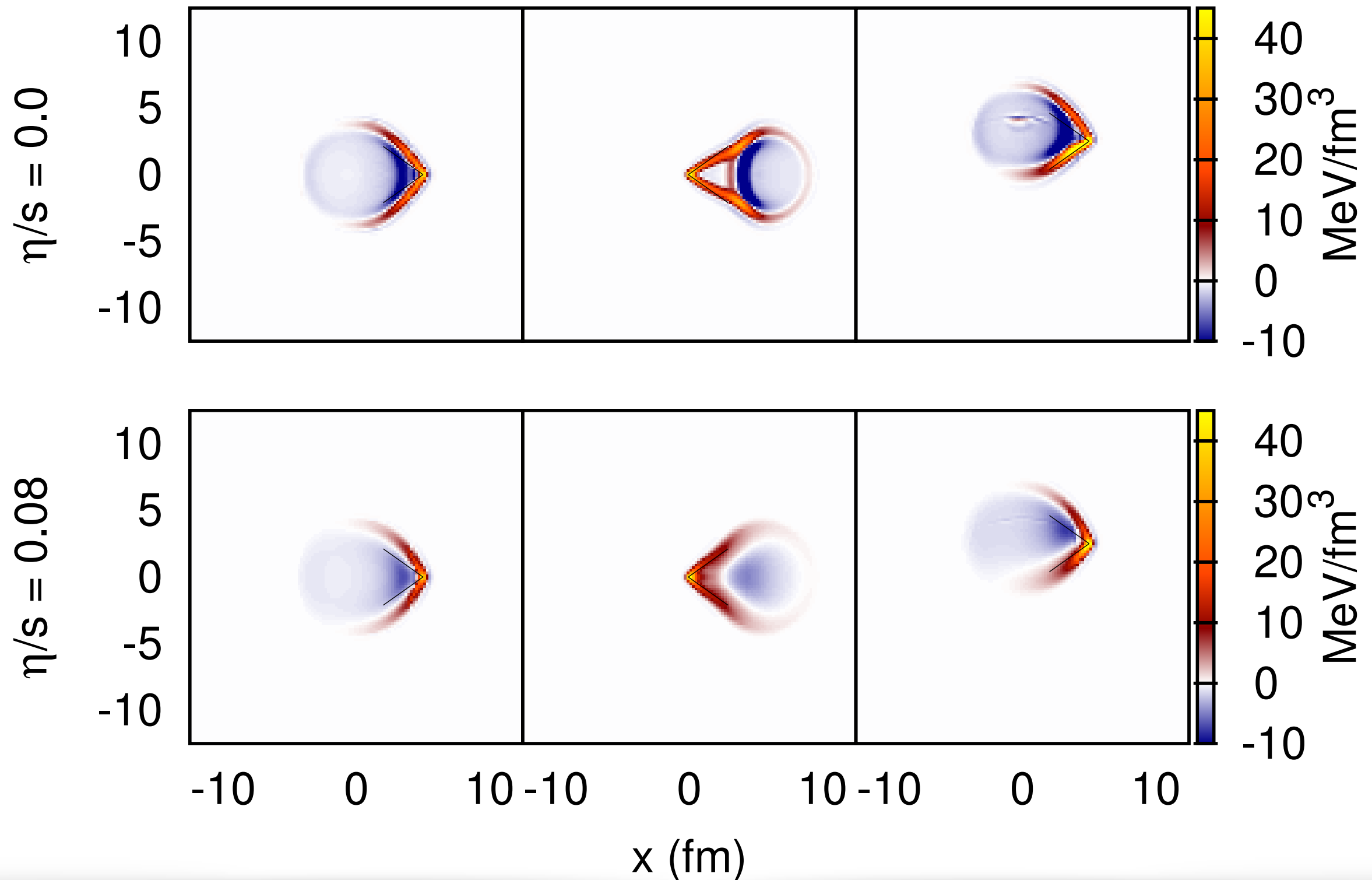
Jet energy deposition and hydro response

energy density difference at $\tau = 3.4$ fm



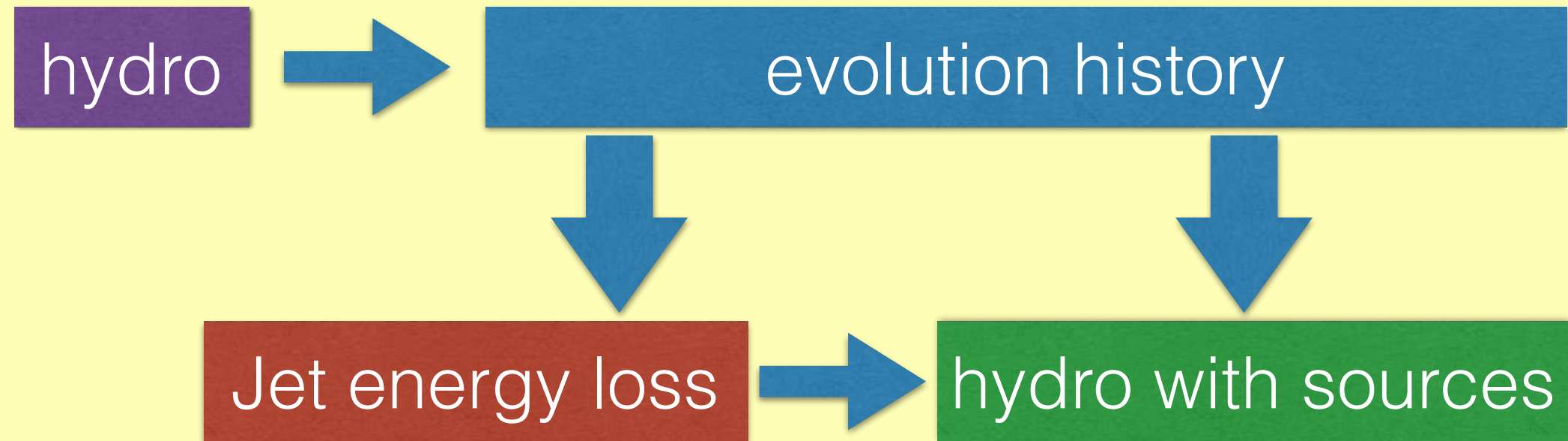
Jet energy deposition and hydro response

energy density difference at $\tau = 5.4$ fm



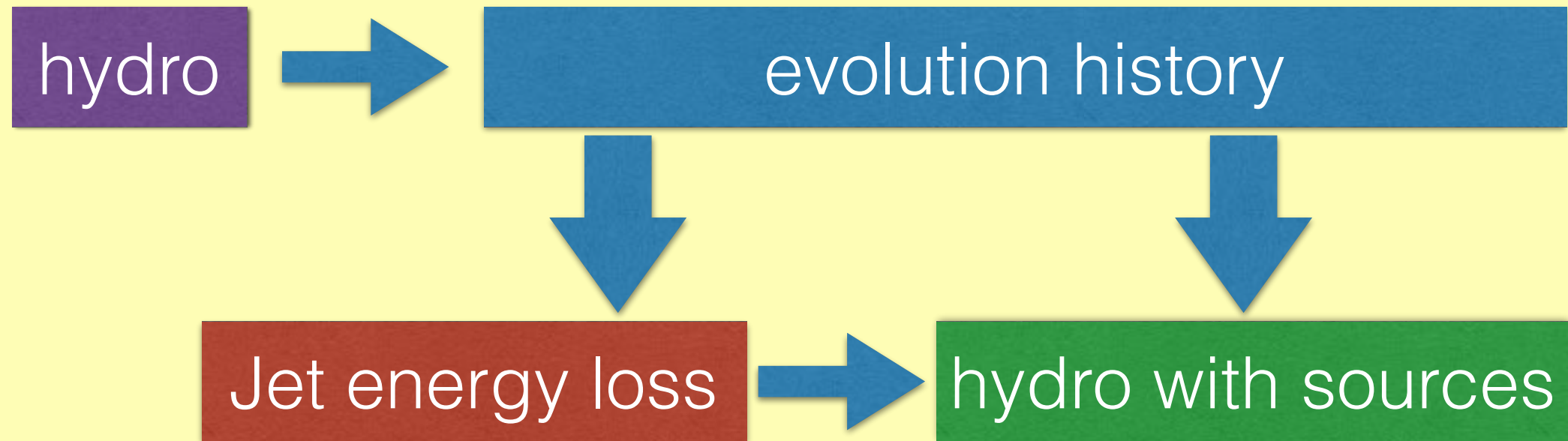
Connect to JETSCAPE

- “Afterburner” mode

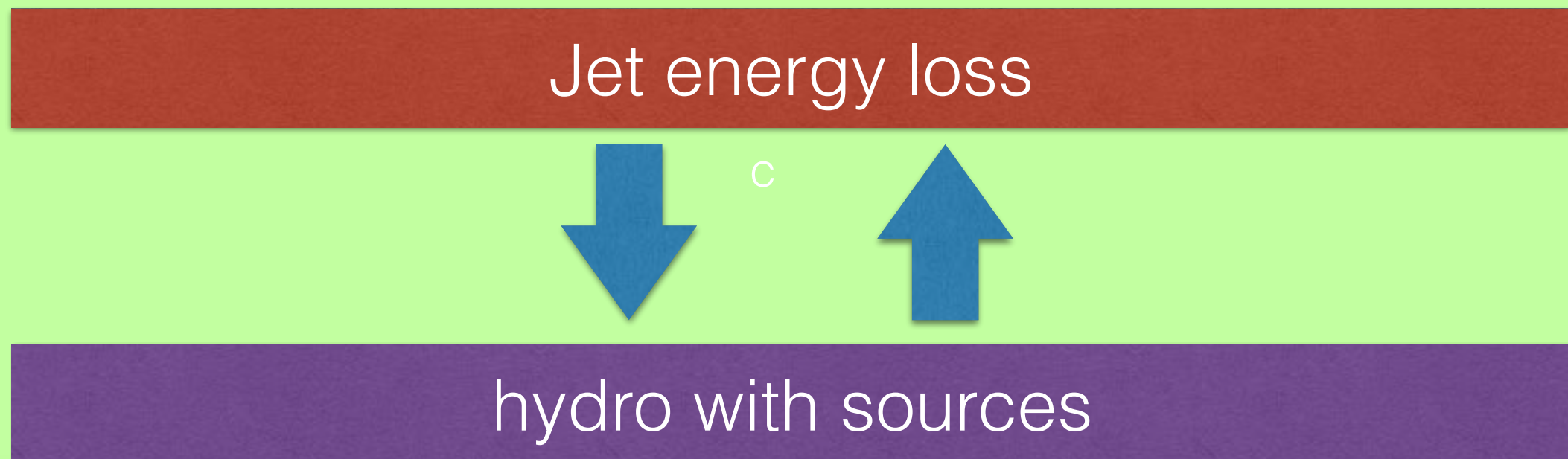


Connect to JETSCAPE

- “Afterburner” mode



- Concurrent mode



Outlook

- We develop a general framework to feed sources into hydrodynamic simulations
- To achieve a smooth transition from the pre-equilibrium stage to the hydrodynamics phase
- Couple with realistic jet energy loss models to study the response in a viscous fluid
- Future extension to magnetohydrodynamics will be exciting