

Lessons from the ab initio symmetry-adapted no-core shell model

Kristina Launey

... LSU Team ...

Jerry Draayer, Tomas Dytrych, Robert Baker, Ali Dreyfuss, David Kekejian, Grigor Sargsyan, Harvey Shows, Logan Woolsey, Sean Laughlin

In collaboration with

Iowa State U. – J. Vary & P. Maris Czech Republic – D. Langr & T. Oberhuber Princeton U. – W. Tang HPC Resources NSF/U. of Illinois ...BlueWaters LSU...SuperMike-II

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LSU code (LSU3shell): sourceforge.net/projects/lsu3shell Dytrych, Launey, Draayer, et al., Phys. Rev. Lett. 111 (2013) 252501 Launey, Dytrych, & Draayer, Prog. Part. Nucl. Phys. 89 (2016) 101

Toward Predictive Theories of Nuclear Reactions Across the Isotopic Chart, INT Program INT-17-1a, 3-1-2017 Lessons from the *ab initio* SA-NCSM

Î

Deformed (in intrinsic frame)...



Effect on X-ray Burst Nucleosynthesis



Orderly pattern from first principles

Symmetry-adapted No-core Shell Model (SA-NCSM)



Symplectic symmetry from first principles



What can we learn from symplectic symmetry?



What is Symplectic Symmetry?

Formal definition

All linear canonical transformations of the single-particle phasespace observables

Nucleus with A nucleons

$$x_{i\alpha} \rightarrow \sum_{\beta=x,y,z} a_{\alpha\beta} x_{i\beta} + b_{\alpha\beta} p_{i\beta}$$
$$p_{i\alpha} \rightarrow \sum_{\beta=x,y,z} c_{\alpha\beta} x_{i\beta} + d_{\alpha\beta} p_{i\beta}$$

that preserve the canonical commutation relation

$$\left[x_{i\alpha}, p_{j\beta}\right] = i\hbar\delta_{ij}\delta_{\alpha\beta}$$

Generators: $Q_{ij} = \sum_{n} x_{ni} x_{nj},$ geometrySymplectic Model
Rosensteel & Rowe,
PRL 38 (1977) 10 $S_{ij} = \sum_{n} (x_{ni} p_{nj} + p_{ni} x_{nj}),$ $L_{ij} = \sum_{n} (x_{ni} p_{nj} - x_{nj} p_{ni}),$ kinematics $K_{ij} = \sum_{n} p_{ni} p_{nj},$ Kinematics

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Earlier studies ... algebraic models

Quite successful, but symmetries were assumed *a priori:* Typically 1 (a few) irrep(s) + symmetry-preserving interaction



P. Park et al., Nucl. Phys. A. 414, 93 (1984)

D. J. Rowe, Rep. Prog. Phys. 48, 1419 (1985)

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Efficacy of SA-NCSM: Li-6





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<u>1 N</u>

C-12: collectivity...



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Ab Initio Nuclear Modeling in New Domains



X-ray Bursts

αp

24NE

2+

2+

0+

Expt.

5.2-8.4 W.u.

proces

 \odot

Hot Cl Cycle

Th

Novae

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22Mg

8.621

4.175

1.582

0.000

 $\langle 2 \rangle 10$

Exp

24<mark>5</mark>1

N/A

2+

Expt.

10

8

0

E_x [MeV]

 $6^+_1 0$

 $0^{+}_{1}0$

Th.

20NE

8.778

4.248

1.634

0.000 Exp



(2+)

Expt.

Th, 4<2>

Th, 2

pernovae

stable nuclide
drip line

32NE

nnn

4 W.u.

Th.

SU(3) NN interaction: keep track of x, y, & z



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Lessons from the *ab initio* SA-NCSM

Important pieces of the NN interaction



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Important pieces of the NN interaction





Simple physics: "shape" + vibrations + rotations

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Conclusions



Informing the inter-nucleon interaction...