

Ab initio calculation of dynamic observables and sum rules in light nuclei

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- Summary
- Application
- Results
- Methods
- Motivation

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 Particularly, nuclear corrections are the bottleneck in μA spectroscopy.
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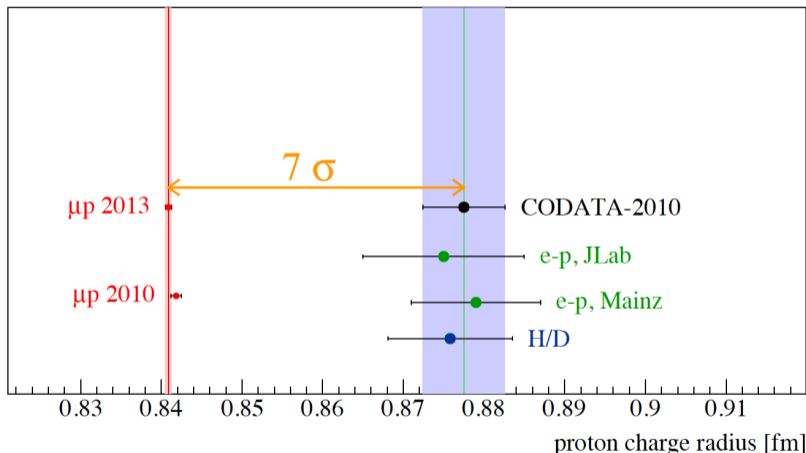
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- Motivation LIT & LSR were applied with Coupled-Cluster for calculations in:
 ${}^4\text{He}$, ${}^{16,22}\text{O}$, ${}^{40,48}\text{Ca}$... (G. Hagen et al., PRC'14; PRC'16; Nature'16;...)

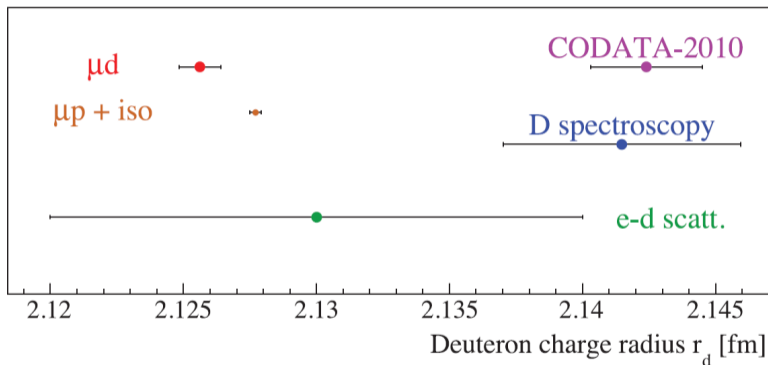


How big is the proton?



R. Pohl & J. Krauth @ CREMA

How big is the deuteron?



R. Pohl *et al.*, **Science** 2016

CREMA @ PSI

Extract precise **charge radii** R_c from Lamb shift (LS) in:

- μH (published 2010,2013: **proton radius puzzle**)
- μD (published 2016: **deuteron radius puzzle**)
- $\mu^4\text{He}^+$ (measured 2014, finalizing: **agreement with $e^-4\text{He}$?!**)
- $\mu^3\text{He}^+$ (measured 2014, analyzing: **???**)
 \implies radius puzzle(s), QED tests, He isotope shift, nuclear *ab initio*, ...
- $\mu^3\text{H}$, $\mu^6\text{He}^+$, $\mu^{6,7}\text{Li}^{+2}$... (possible?)

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CREMA @ PSI

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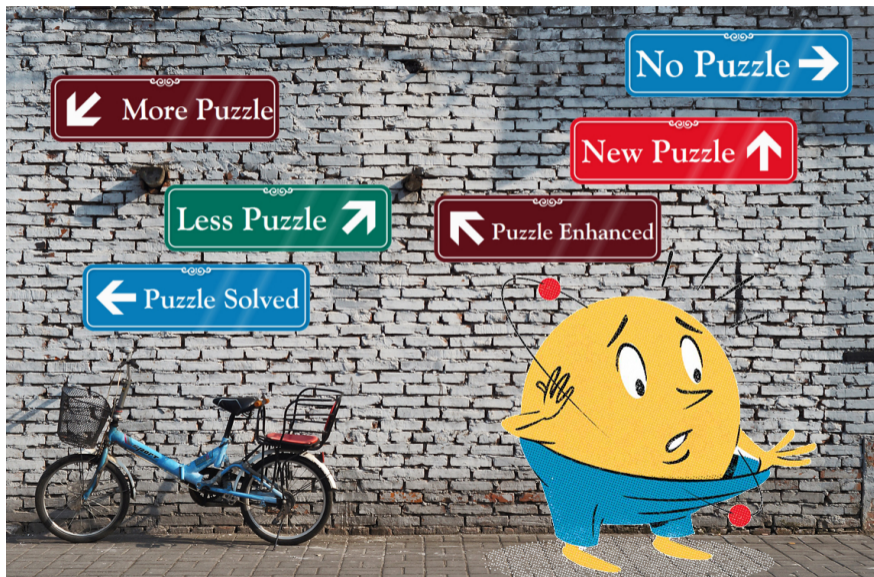
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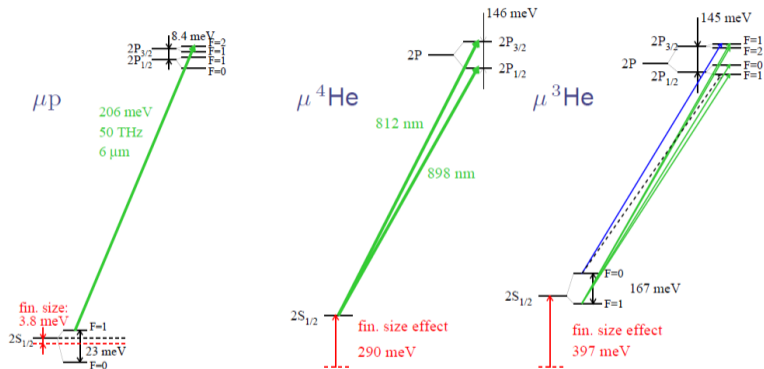
FAMU @ RIKEN-RAL / J-PARC

- HFS in μH in two new methods (planned)

Precise R_c/R_m from μA LS/HFS

Require accurate theoretical inputs from QED, hadron and nuclear physics





R. Pohl (for CREMA), presentation at ECT*, Trento, Italy (2012)

Extract $R_c \equiv \sqrt{\langle r^2 \rangle}$ from Lamb shift measurement

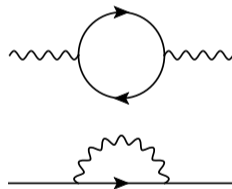
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- QED corrections:
 - vacuum polarization
 - lepton self energy
 - relativistic recoil effects

- Theory of μ - p , D, ${}^3,4\text{He}^+$ reexamined
 - Martynenko *et al.* '07, Borie '12, Krutov *et al.* '15
 - Karshenboim *et al.* '15, Krauth *et al.* '15 ...

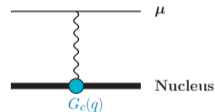


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- Nuclear structure corrections
(One-Photon Exchange)

- finite-size effect: $\delta_{size} = \frac{m_r^3}{12} (Z\alpha)^4 \times R_c^2$



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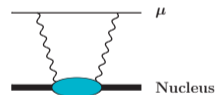
- $\delta_{TPE} = ?$

⇒ related to nuclear response functions:

$$S_O(\omega) = \int |\langle \psi_f | \hat{O} | \psi_0 \rangle|^2 \delta(E_f - E_0 - \omega)$$

⇒ can be extracted from data (very imprecise)

⇒ or calculated (continuum few-body problem)



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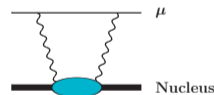
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- **Very small — is it important?**





The accuracy of R_c is limited by δ_{TPE}

Example — μD :

$$\Delta E_{\text{QED}}^{\text{LS}} = 228.77356(75) \text{ meV}$$

$$\Delta E_{\text{rad.}-\text{dep.}}^{\text{LS}} = -6.11025(28) r_d^2 \text{ meV/fm}^2 + 0.00300(60) \text{ meV}$$

$$\Delta E_{\text{TPE}}^{\text{LS}} = 1.70910(2000) \text{ meV}$$

J. Krauth *et al.* (CREMA), **Ann. Phys. (2016)**; R. Pohl *et al.* (CREMA), **Science 2016**

Status — prior to $\mu^{3,4}\text{He}^+$ measurements:

- Uncertainty in δ_{pol} : $\sim 20\%$
- Required: $\sim 5\%$
(to determine R_c with $\sim 10^{-4}$ accuracy)

$$\delta_{TPE} = \sum_a I_a = \sum_a \int d\omega S_a(\omega) g_a(\omega)$$

- The leading polarization contribution relates to the dipole response

$$I_{D_1}^{(0)} \propto \int_{\omega_{th}}^{\infty} d\omega S_{D_1}(\omega) \omega^{-1/2}$$

- $S_{D_1}(\omega) \implies$ electric dipole response function [$\hat{D}_1 = R Y_1(\hat{R})$]

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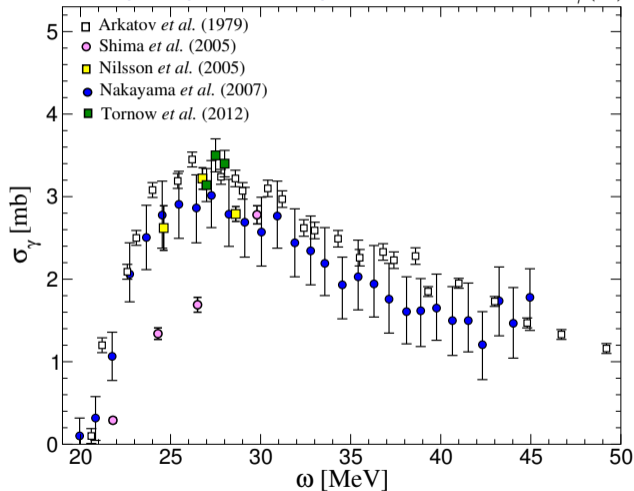
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electric dipole photoabsorption cross section $\sigma_\gamma(\omega) = 4\pi^2\alpha\omega S_{D1}(\omega)$



System	Our Ref.	Unc.	Experimental Status
$\mu^2\text{H}$	Phys. Lett. B '14	1% \rightarrow 1.3%	published <i>Science</i> '16
$\mu^4\text{He}^+$	Phys. Rev. Lett. '13	20% \rightarrow 6%	measured, unpublished
$\mu^3\text{He}^+$	} Phys. Lett. B '16	20% \rightarrow 4%	measured, unpublished
$\mu^3\text{H}$		4%	measurable?

- Our results agree with other values and are more accurate
 - \Rightarrow Unc. comparable with $\sim 5\%$ experimental needs
 - \Rightarrow Will improve precision of R_c from Lamb shifts
 - \Rightarrow May help shed light on the “proton (deuteron) radius puzzle”

Calculating $S_O(\omega) = \int |\langle f | \hat{O} | i \rangle|^2 \delta(\omega_f - \omega)$ using LIT

V.D. Efros et al., PLB'94; JPG'07

Calculating $S_O(\omega) = \mathcal{F} |\langle f | \hat{O} | i \rangle|^2 \delta(\omega_f - \omega)$ using LIT

$$\mathcal{L}(\sigma) = \int d\omega \frac{S(\omega)}{(\omega - \sigma_r)^2 + \sigma_i^2}, \quad \sigma_i \equiv \text{Im}(\sigma) > 0$$

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 (and only the trivial solution to the homogeneous Eq.)

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\Rightarrow Can be solved using **any bound-state method**

Calculating $S_O(\omega) = \mathcal{F} |\langle f | \hat{O} | i \rangle|^2 \delta(\omega_f - \omega)$ using LIT

1. Solve $(H - \sigma^*) |\tilde{\psi}\rangle = \hat{O} | i \rangle$ using a bound-state basis to obtain $\mathcal{L}_{\text{calc}}(\sigma) = \langle \tilde{\psi} | \tilde{\psi} \rangle$
2. Invert $\mathcal{L}_{\text{calc}}(\sigma) = \int d\omega \frac{S(\omega)}{(\omega - \sigma_r)^2 + \sigma_i^2}$ using the common ansatz

$$S_N(\omega) = \sum_n^N c_n \phi_n(\alpha) \implies \|\mathcal{L}_{\text{calc}}(\sigma) - \mathcal{L}_N(\sigma)\| < \epsilon$$

Efros et al., JPG'07; Barnea FBS'10; Orlandini et al., FBS'17

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- This method is not sensitive to a specific energy-range of $\mathcal{L}_{\text{calc}}(\sigma) - \mathcal{L}_N(\sigma)$
- Small σ_i is needed to resolve fine details of $S(\omega) \implies$ Harder to converge $\mathcal{L}_{\text{calc}}(\sigma)$
- Large σ_i captures better the tail of $S(\omega)$...

Efros et al., JPG'07; Barnea FBS'10; Orlandini et al., FBS'17

Using few-body bound-state methods, the LIT was applied to electroweak reactions of $2 \leq A \leq 7$, such as: (Efros et al., JPG'07; Leidemann & Orlandini PPNP'13; Bacca & Pastore JPG'14)

1. Photoabsorption cross sections
2. Photon scattering
3. Electron scattering (longitudinal and transverse response)
4. Neutrino breakup of light nuclei in supernovae

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Including exclusive reactions

(LaPiana & Leidemann NPA'00; Quaglioni et al., PRC'04,'05; Andreasi et al., EPJA'06; NND et al., FBS'14)
 which require solving the exclusive LIT equations:

$$(H - \sigma^*)|\tilde{\psi}_f\rangle = \hat{V}_f|\phi_f\rangle$$

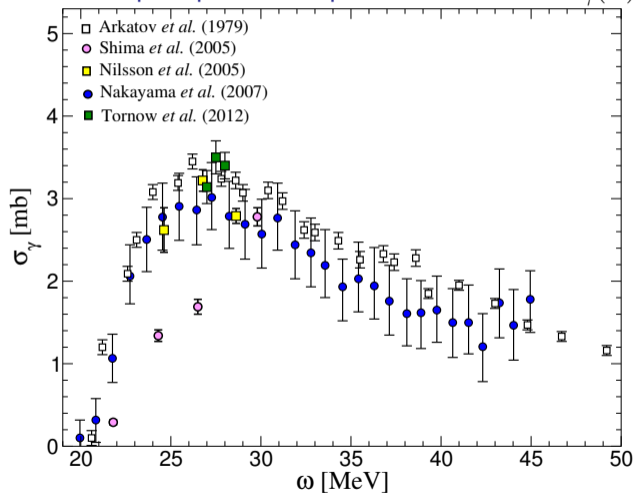
The solution of the exclusive LIT equation:

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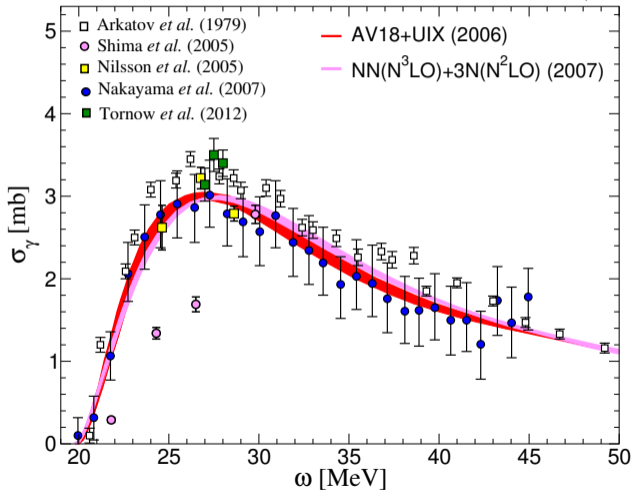
can also be used for:

1. Radiative capture (by exchanging $i \leftrightarrow f$ in exclusive photodisintegration)
2. Semi-inclusive ($e, e'N$) using the “spectral function approximation ”
(demonstrated by Efros et al., PRC'98)
3. Astrophysical S-factors (demonstrated by S. Deflorian et al., FBS'17)
4. Hadron scattering (suggested by V.D. Efros, PAN'99, PIC'17)
5. Glauber approximation (suggested by V.D. Efros et al., JPG'07)

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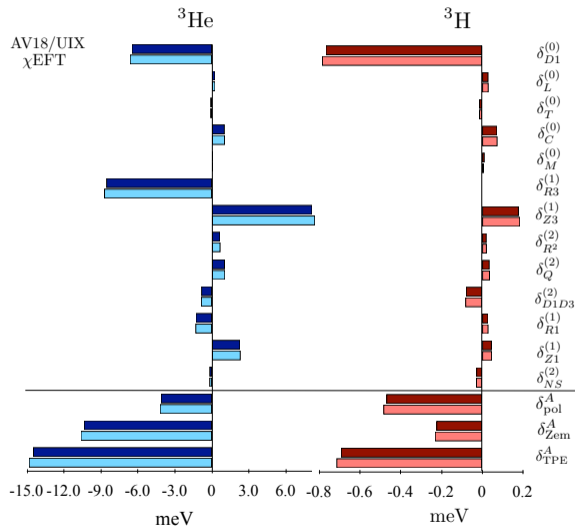


Gazit *et al.*, PRL'06

Quaglioni & Navrátil PLB'07

The work is not completed yet ...




 NND *et al.*, Phys. Lett. B (2016)

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In general, sum rules are interesting also for:

1. Comparison with experiments
2. Checking analytic or assumed relations
3. Observables of interest ($R_{\text{ch}} \leftrightarrow \alpha_D \leftrightarrow R_n - R_p \leftrightarrow L(\text{Sym. Energy})$)

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Problems:

1. (May) need to know $S(\omega)$ with good resolution over wide range of energies
2. Need to extrapolate $S(\omega)g(\omega)$ to $\omega \rightarrow \infty$, difficulty depends on $g(\omega)$ (and \hat{O})

$$\delta_{\text{TPE}} = \sum_a I_a = \sum_a \int d\omega S_a(\omega) g_a(\omega)$$

A model-space of size M is used to calculate the LIT of $S(\omega)$

$$\mathcal{L}(\sigma) = \frac{\sigma_i}{\pi} \int d\omega \frac{S(\omega)}{(\omega - \sigma_r)^2 + \sigma_i^2}, \quad \sigma_i \equiv \text{Im}(\sigma) > 0$$

$$\mathcal{L}_M(\sigma) = \frac{\sigma_i}{\pi} \sum_{\mu}^M \frac{|\langle \mu | \hat{O} | i \rangle|^2}{(\omega_{\mu} - \sigma_r)^2 + \sigma_i^2}$$

$$\Rightarrow |\mathcal{L}(\sigma) - \mathcal{L}_M(\sigma)| \leq \varepsilon_M$$

$$\delta_{\text{TPE}} = \sum_a I_a = \sum_a \int d\omega S_a(\omega) g_a(\omega)$$

A model-space of size M is used to calculate the LIT of $S(\omega)$

$$\mathcal{L}(\sigma) = \frac{\sigma_i}{\pi} \int d\omega \frac{S(\omega)}{(\omega - \sigma_r)^2 + \sigma_i^2}, \quad \sigma_i \equiv \text{Im}(\sigma) > 0$$

$$\mathcal{L}_M(\sigma) = \frac{\sigma_i}{\pi} \sum_{\mu}^M \frac{|\langle \mu | \hat{O} | i \rangle|^2}{(\omega_{\mu} - \sigma_r)^2 + \sigma_i^2}$$

$$\Rightarrow |\mathcal{L}(\sigma) - \mathcal{L}_M(\sigma)| \leq \varepsilon_M$$

- smaller $\sigma_i \Rightarrow$ better resolution, slower convergence

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Formally :

$$h(\sigma) = \frac{1}{2\pi} \int dk e^{\sigma_i |k|} \tilde{\mathbf{g}}(\mathbf{k}) e^{-ik\sigma_r}$$

The applicability depends on the form of $g(\omega)$. For example:

$$g(\omega) = \frac{\beta}{\pi} \frac{1}{(\omega - \omega_0)^2 + \beta^2} \xrightarrow{\beta \rightarrow 0} \delta(\omega - \omega_0),$$

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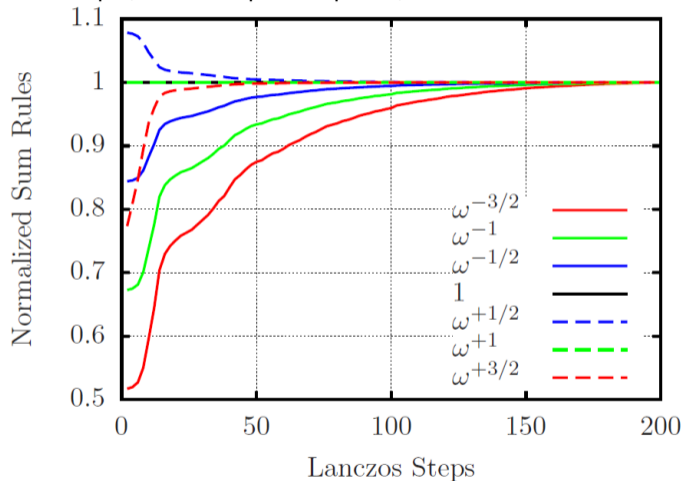
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- The LSR method uses Lanczos to obtain I_M (NND et al., PRC'14)
 1. without solving $S(\omega)$
 2. efficiently and accurately
 3. with rapid convergence
 (generalizes similar methods as in:
 Haxton et al., PRC'05; Gazit et al., PRC'06; Stetcu et al., PLB'08, PRC'09)

- For example, for the dipole response, calculated with $M \sim 10^5$, we get



NND, Barnea, Ji, and Bacca, PRC (2014)

- Summary The LIT & LSR Methods can be used with any bound-state method to obtain dynamic observables (e.g., in the nuclear continuum)
- Application Precise nuclear calculations are crucial in many high-profile efforts, e.g. ν -less $\beta\beta$ decay; searches for EDMs; etc.
 Particularly, nuclear corrections are the bottleneck in μA spectroscopy.
- Results For $A = 3, 4$ we reduced the uncertainties in these corrections from $\sim 20\%$ to 4–6% as required for ongoing experiments.
 We use the **LSR method** to calculate the relevant sum-rules.
- Methods We proved its applicability using the LIT method.
 The LIT method has been successfully applied to calculate many EW-induced reactions. Some applications have yet to be demonstrated.
- Motivation LIT & LSR were applied with Coupled-Cluster for calculations in:
 ${}^4\text{He}$, ${}^{16,22}\text{O}$, ${}^{40,48}\text{Ca}$... (G. Hagen et al., PRC'14; PRC'16; Nature'16;...)



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for particle and nuclear physics
and accelerator-based science

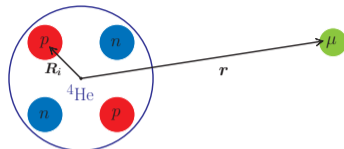
Thank you!
Merci!

BACK UP

- Hamiltonian for muonic atoms

$$H = H_{nucl} + H_{\mu} + \Delta H$$

$$H_{\mu} = \frac{p^2}{2m_r} - \frac{Z\alpha}{r}$$



- Corrections to the point Coulomb from protons

$$\Delta H = \alpha \sum_i^Z \left(\frac{1}{r} - \frac{1}{|\mathbf{r} - \mathbf{R}_i|} \right)$$

- Evaluate inelastic effects of ΔH on muonic spectrum in 2^{nd} -order perturbation theory

$$\delta_{pol} = \sum_{N \neq N_0, \mu} \langle N_0 \mu_0 | \Delta H | N \mu \rangle \frac{1}{E_{N_0} - E_N + \epsilon_{\mu_0} - \epsilon_{\mu}} \langle N \mu | \Delta H | \mu_0 N_0 \rangle$$

$|\mu_0\rangle$: muon wave function for $2S/2P$ state

Systematic contributions to nuclear polarization

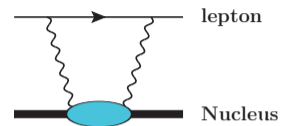
δ_{NR} **Non-Relativistic** limit

$\delta_L + \delta_T$ **L**ongitudinal and **T**ransverse **relativistic** corrections

δ_C **Coulomb** distortions

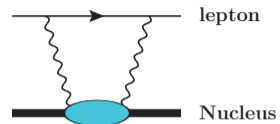
δ_{NS} Corrections from **finite Nucleon Size**

- Neglect Coulomb interactions in the intermediate state



- Neglect Coulomb interactions in the intermediate state
- Expand muon matrix element in powers of

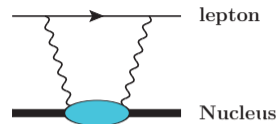
$$\eta \equiv \sqrt{2m_r\omega} |\mathbf{R} - \mathbf{R}'|$$



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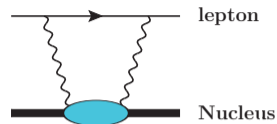
- $|\mathbf{R} - \mathbf{R}'| \implies$ “virtual” distance the proton travels in 2γ exchange
- uncertainty principal $|\mathbf{R} - \mathbf{R}'| \sim 1/\sqrt{2m_N\omega}$
- $\eta \sim \sqrt{\frac{m_r}{m_N}} \approx 0.3$



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$$P_{NR}(\omega, \mathbf{R}, \mathbf{R}') \simeq \frac{m_r^3 (Z\alpha)^5}{12} \sqrt{\frac{2m_r}{\omega}} \left[|\mathbf{R} - \mathbf{R}'|^2 - \frac{\sqrt{2m_r\omega}}{4} |\mathbf{R} - \mathbf{R}'|^3 + \frac{m_r\omega}{10} |\mathbf{R} - \mathbf{R}'|^4 \right]$$

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)} \sim \eta^2 + \eta^3 + \eta^4$$

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- $\delta_{NR}^{(0)} \propto \eta^2$

$$\delta_{D1}^{(0)} = -\frac{2\pi m_r^3}{9} (Z\alpha)^5 \int_{\omega_{th}}^{\infty} d\omega \sqrt{\frac{2m_r}{\omega}} S_{D1}(\omega)$$

- $S_{D1}(\omega) \implies$ electric dipole response function [$\hat{D}_1 = R Y_1(\hat{R})$]

- $\delta_{D1}^{(0)}$ is the dominant contribution to δ_{pol}

- \implies Rel. and Coulomb corrections added at this order

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)}$$

- $\delta_{NR}^{(1)} \propto \eta^3$

$$\delta_{NR}^{(1)} = \delta_{R3pp}^{(1)} + \delta_{Z3}^{(1)}$$

$$\delta_{R3pp}^{(1)} = -\frac{m_r^4}{24} (Z\alpha)^5 \iint d\mathbf{R}d\mathbf{R}' |\mathbf{R} - \mathbf{R}'|^3 \langle N_0 | \hat{\rho}^\dagger(\mathbf{R}) \hat{\rho}(\mathbf{R}') | N_0 \rangle$$

$$\delta_{Z3}^{(1)} = \frac{m_r^4}{24} (Z\alpha)^5 \iint d\mathbf{R}d\mathbf{R}' |\mathbf{R} - \mathbf{R}'|^3 \rho_0(\mathbf{R}) \rho_0(\mathbf{R}')$$

- $\delta_{R3pp}^{(1)} \Rightarrow$ 3rd-order proton-proton correlation

- $\delta_{Z3}^{(1)} \Rightarrow$ 3rd Zemach moment

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- $\delta_{R3pp}^{(1)} \Rightarrow$ 3rd-order proton-proton correlation

- $\delta_{Z3}^{(1)} \Rightarrow$ 3rd Zemach moment

cancels *elastic* Zemach moment of finite-size corrections

c.f. Pachucki '11 & Friar '13 (μD) $\Rightarrow \delta_{TPE} \equiv |\delta_{Zem} + \delta_{pol}|$

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)}$$

- $\delta_{NR}^{(2)} \propto \eta^4$

$$\delta_{NR}^{(2)} = \frac{m_r^5}{18} (Z\alpha)^5 \int_{\omega_{th}}^{\infty} d\omega \sqrt{\frac{\omega}{2m_r}} \left[S_{R^2}(\omega) + \frac{16\pi}{25} S_Q(\omega) + \frac{16\pi}{5} S_{D_1 D_3}(\omega) \right]$$

- $S_{R^2}(\omega) \implies$ monopole response function

- $S_Q(\omega) \implies$ quadrupole response function

- $S_{D_1 D_3}(\omega) \implies$ interference between D_1 and D_3 [$\hat{D}_3 = R^3 Y_1(\hat{R})$]

Error type	$\mu^3\text{He}^+$			$\mu^3\text{H}$		
	δ_{pol}^A	δ_{Zem}^A	δ_{TPE}^A	δ_{pol}^A	δ_{Zem}^A	δ_{TPE}^A
Numerical	0.4	0.1	0.1	0.1	0.0	0.1
Nuclear model	1.5	1.8	1.7	2.2	2.3	2.2
ISB	2.0	0.2	0.5	0.9	0.2	0.6
Nucleon size	1.6	1.5	0.6	0.6	1.3	0.0
Relativistic	0.6	-	1.5	1.4	-	0.3
Coulomb	1.2	-	0.3	0.3	-	0.2
Multipole expansion	2.0	-	0.6	2.0	-	1.4
Higher $Z\alpha$	1.5	-	0.4	0.7	-	0.5
Magnetic MEC	0.4	-	0.1	0.3	-	0.2
Total	4.1%	2.3%	2.5%	3.6%	2.7%	2.7%