# Discovering the QCD critical point with net-proton fluctuations

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- **thermal** (fluid dynamical) **fluctuations** important for fluids undergoing phase transitions, nucleation, showing instabilities, ...
- relevant in micro-engineering, molecular biology, combustive ignition, ...
- stochastic correlations between fluctuating quantities  $\rightarrow$  consider a non-relativistic fluid:

$$
\langle \delta v_j(x,t) \delta v_j(x',t) \rangle = \frac{T}{\rho} \delta_{ij} \delta(x-x')
$$

• (linearized) fluid dynamics propagates these, e.g. as *shear* or *sound* modes

$$
\langle \delta v_i^T \delta v_j^T \rangle_{\omega, k} = \frac{2T}{\rho} \left( \delta_{ij} - \hat{k}_i \hat{k}_j \right) \frac{\eta k^2 / \rho}{\omega^2 + (\eta k^2 / \rho)^2}
$$

$$
\langle \delta v_i^L \delta v_j^L \rangle_{\omega, k} = \frac{2T}{\rho} \hat{k}_i \hat{k}_j \frac{\omega k^2 \Gamma}{(\omega^2 - c_s^2 k^2)^2 + (\omega k^2 \Gamma)^2}; \Gamma = \frac{4 \eta}{3 \rho} + \dots
$$

 $\rightarrow$  consider the correlation function

$$
G_S^{xyxy} = \langle \{\pi^{xy}, \pi^{xy}\} \rangle_{\omega,k} \simeq \rho^2 \langle \{v_x v_y, v_x v_y\} \rangle_{\omega,k}
$$

 $\rightarrow$  match this to response function in Kubo-limit ( $\omega \rightarrow 0$ )

$$
G_R^{xyxy} = P + \delta P - i\omega(\eta + \delta\eta) + \omega^2(\eta\tau_\pi + \delta(\eta\tau_\pi))
$$

- *δP* ∼ *T*Λ<sup>3</sup> cutoff-dependent pressure correction
- *δη* ∼ *Tρ*Λ *η* small  $\eta \Rightarrow$  large  $\delta \eta$  (there must be a lower bound on  $\eta / n$  induced by fluctuations)

• 
$$
\delta(\eta \tau_{\pi}) \sim \frac{1}{\sqrt{\omega}} \frac{T \rho^{3/2}}{\eta^{3/2}}
$$

for  $\omega \to 0$  find  $\delta(\eta \tau_{\pi}) \to \infty \Rightarrow$  2<sup>nd</sup>-order fluid dynamics inconsistent without fluctuations

#### Landau-Lifshitz-Navier-Stokes fluid dynamics  $\Rightarrow$  add stochastic flux terms to fluid dynamical equations



anisotropic fluid dynamical expansion of an ultra-cold Fermi gas cloud at unitarity



1d, dilute gas, periodic boundary conditions

- numerics sensitive to discretization of noise  $\langle \xi^2 \rangle \propto 1/\Delta V \sim \Lambda^3$
- correct implementation possible;  $3<sup>rd</sup>$ -order methods seem to work best

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- event-by-event fluctuations of the net-proton number expected to signal a critical point (CP) in the QCD phase diagram
- in td. limit: net-baryon number susceptibility diverges at CP
- associated with fluctuations in the chiral order parameter *σ*
- higher-order fluctuations,  $\langle (\delta\sigma)^n\rangle_c \propto \xi^{m(n)}$ , depend stronger on correlation length *ξ*
- system created in a HIC is short-lived, spatially small, inhomogeneous, highly dynamical
- *τ* ∝ *ξ <sup>z</sup>* ⇒ critical slowing down, memory effects
- $\Rightarrow$  Relativistic fluid dynamics including fluctuations/dissipation!
	- final stage processes could wash out critical fluctuation signals:
		- $\blacktriangleright$  isospin redistribution,
		- $\blacktriangleright$  resonance decays, ...

#### $\Rightarrow$  Can critical fluctuation signals survive resonance decays?

**Y. Hatta and M.A. Stephanov, PRL 91 (2003) 102003; M.A. Stephanov, PRD 65 (2002) 096008;** *ibid.* **PRL 102 (2009) 032301; B. Berdnikov and K. Rajagopal, PRD 61 (2000) 105017; M. Nahrgang et al., PRC 84 (2011) 024912; S. Mukherjee et al., PRC 92 (2015) 034912; C. Herold et al., PRC 93 (2016) 021902; M. Kitazawa and M. Asakawa, PRC 86 (2012) 024904; M. Nahrgang et al., Eur.**

### $P$  A R T – I

# Recent progress in simulating relativistic Fluctuating Dissipative Fluid Dynamics

with Yuri Karpenko

#### Remarks about the critical mode

- at  $\mu \neq 0$ ,  $\sigma$  mixes with the net-baryon density  $n$  (and  $e$ , ...)
- in a Ginzburg-Landau formalism:

$$
V(\sigma, n) = \int d^3x \left( \sum_m \left[ a_m \sigma^m + b_m n^m \right] + \sum_{m,l} c_{m,l} \sigma^m n^l \right) - h\sigma - j n
$$

- $V(\sigma, n)$ : flat direction  $(a\sigma, -cn)$  with vanishing curvature at CP
- equations of motion (including the symmetries of  $V(\sigma, n)$ ):

$$
\partial_t \sigma = -\Gamma(\delta V/\delta \sigma) + \dots \quad ; \quad \partial_t n = \gamma \nabla^2(\delta V/\delta n) + \dots
$$

- eigenfrequencies and -modes:
	- *ω*<sup>1</sup> ∝ −*i*Γ*a* ; (1, 0) *ω*<sup>2</sup> ∝ −*iγDk*2/*a* ; (−*c*, *a*)  $\rightarrow$  fast mode  $\rightarrow$  slow mode



NJL for Z<sub>2</sub>-CP

 $\Rightarrow$  the diffusive mode becomes the true critical (slow) mode near the CP in the long-time dynamics

Conventional fluid dynamics propagates thermal averages of the energy density, pressure, velocities, charge densities, etc.

However, ...

- ... already in equilibrium there are thermal fluctuations!
- ... the fast processes, which lead to local equilibration also lead to noise!

Conventional ideal fluid dynamics:

$$
T^{\mu\nu} = T^{\mu\nu}_{\text{eq}}
$$

$$
N^{\mu} = N^{\mu}_{\text{eq}}
$$

Y. Minami, T. Kunihiro, PTP122 (2010); P. Kovtun, G. Moore, P. Romatschke, PRD84 (2011); J. Kapusta, B. Müller, M. Stephanov **PRC85 (2012); J. Kapusta, J. Torres-Rincon PRC86 (2012); C. Chafin, T. Schafer, PRA87 (2013); P. Romatschke, R. Young, PRA87 ¨ (2013); P. Kovtun, G. Moore, P. Romatschke, JHEP1407 (2014); J. Kapusta, C. Young, PRC90 (2014); C. Young, J. Kapusta, C. Gale, S. Jeon, B. Schenke, PRC91 (2015); K. Murase, T. Hirano, 1304.3243;** *ibid.* **1601.02260**

Conventional fluid dynamics propagates thermal averages of the energy density, pressure, velocities, charge densities, etc.

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Conventional viscous fluid dynamics:

 $T^{\mu\nu} = T^{\mu\nu}_{\text{eq}} + \Delta T^{\mu\nu}_{\text{visc}}$  $N^{\mu} = N^{\mu}_{\text{eq}} + \Delta N^{\mu}_{\text{visc}}$ 

Y. Minami, T. Kunihiro, PTP122 (2010); P. Kovtun, G. Moore, P. Romatschke, PRD84 (2011); J. Kapusta, B. Müller, M. Stephanov **PRC85 (2012); J. Kapusta, J. Torres-Rincon PRC86 (2012); C. Chafin, T. Schafer, PRA87 (2013); P. Romatschke, R. Young, PRA87 ¨ (2013); P. Kovtun, G. Moore, P. Romatschke, JHEP1407 (2014); J. Kapusta, C. Young, PRC90 (2014); C. Young, J. Kapusta, C. Gale, S. Jeon, B. Schenke, PRC91 (2015); K. Murase, T. Hirano, 1304.3243;** *ibid.* **1601.02260**

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Fluctuating viscous fluid dynamics:

$$
T^{\mu\nu} = T_{\text{eq}}^{\mu\nu} + \Delta T_{\text{visc}}^{\mu\nu} + \Xi^{\mu\nu}
$$

$$
\mathbf{N}^{\mu} = N_{\text{eq}}^{\mu} + \Delta N_{\text{visc}}^{\mu} + I^{\mu}
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$$

• In second-order fluid dynamics there are relaxation equations for  $\Xi^{\mu\nu}$ , ...:

$$
u^{\gamma} \partial_{\gamma} \Xi^{\langle \mu \nu \rangle} = -\frac{\Xi^{\mu \nu} - \xi^{\mu \nu}}{\tau_{\pi}}
$$

with (white) noise correlators in linear response theory

$$
\langle \xi^{\mu\nu}(x) \xi^{\alpha\beta}(x') \rangle = 2 \mathcal{T}[\eta(\Delta^{\mu\alpha} \Delta^{\nu\beta} + \Delta^{\mu\beta} \Delta^{\nu\alpha}) + (\zeta - 2/3\eta) \Delta^{\mu\nu} \Delta^{\alpha\beta}] \delta^4(x - x')
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However, ...

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Fluctuating viscous fluid dynamics:  $\Rightarrow$  neglect  $N^{\mu}$  for now!

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T^{\mu\nu} = T_{\text{eq}}^{\mu\nu} + \Delta T_{\text{visc}}^{\mu\nu} + \Xi^{\mu\nu}
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- correlation functions in linearized fluid dynamics describe noninteracting modes
- if nonlinearities are included  $\rightarrow$  interaction of modes
	- $\rightarrow$  modification of correlations
	- $\rightarrow$  contributions to transport coefficients, ...
- symmetrized correlator:

$$
G_S^{xyxy}(\omega,{\bf 0})=\int d^3x dt\, e^{i(\omega t-{\bf k}\cdot{\bf x})}\left\langle \frac{1}{2}\{T^{xy}(t,{\bf x}),T^{xy}({\bf 0},{\bf 0})\}\right\rangle
$$

• for the shear-shear contribution  $\Rightarrow$ 

$$
G_{\text{H,shear-shear}}^{\text{xyxy}}(\omega,\mathbf{0})=-\frac{77}{90\pi^2}\Lambda^3-i\omega\frac{77}{60\pi^2}\frac{\Lambda}{\gamma_\eta}+(i+1)\omega^{3/2}\frac{77}{90\pi^2}\frac{1}{\gamma_\eta^{3/2}}
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**cutoff-dependent fluctuation contribution to pressure**

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**cutoff-dependent fluctuation contribution to pressure**

**cutoff-dependent correction to** *η*

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$$

**cutoff-dependent fluctuation contribution to pressure cutoff-dependent correction to** *η* **frequency-dependent contribution to** *η* **and** *τπ*

### Fluid dynamical fluctuations - 1+1d

- Static "box" with periodic boundary conditions in relativistic  $1 + 1d$  fluid dynamics
- Initialized at  $e_0 = 10$  GeV/fm<sup>3</sup> (without fluctuations nothing would happen)

consider time evolution in [fm] of *e* in [GeV/fm<sup>3</sup> ] for ∆*x* = 1.0 fm



**M. Nahrgang, MB, Y. Karpenko, S. Bass, T. Schafer, work in progress ¨**

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consider time evolution in [fm] of *e* in [GeV/fm<sup>3</sup> ] for ∆*x* = 0.8 fm



**M. Nahrgang, MB, Y. Karpenko, S. Bass, T. Schafer, work in progress ¨**

### Fluid dynamical fluctuations - 1+1d

- Static "box" with periodic boundary conditions in relativistic  $1 + 1d$  fluid dynamics
- Initialized at  $e_0 = 10$  GeV/fm<sup>3</sup> (without fluctuations nothing would happen)

consider time evolution in [fm] of *e* in [GeV/fm<sup>3</sup> ] for ∆*x* = 0.5 fm



**M. Nahrgang, MB, Y. Karpenko, S. Bass, T. Schafer, work in progress ¨**

## Fluid dynamical fluctuations - 3+1d

• Static box with periodic boundary conditions in relativistic  $3 + 1d$  fluid dynamics

$$
\partial_{\mu} T^{\mu\nu} = \partial_{\mu} \left( T^{\mu\nu}_{eq} + \Delta T^{\mu\nu}_{visc} + \Xi^{\mu\nu} \right) = 0
$$



• The variance  $\langle (\Delta e)^2 \rangle$  saturates after  $\sim$  5 fm.

• Fluctuations are large in the computational cell of fluid dynamics  $\Rightarrow$  noise correlated over  $\sim 1$  fm<sup>3</sup> – reproduced.

• Important check: equilibrium expectations for fluctuations and nonlinear effects!



- Proportionality to 1/*V* reproduced for the correction to the average and the variance of energy density in the local rest frame.
- Implementing fluid dynamical fluctuations is important,

but requires a sustained and systematic effort!

#### What remains to be done!?

- ...apart from the obvious...
- provide realistic initial conditions for HIC
- provide realistic EoS and transport coefficients
- propagate also net-baryon current with fluctuations in fluid dynamics
- evolve *σ*-field within fluid dynamical background by means of an stochastic Langevin equation or Fokker-Planck equation
- $\rightarrow$  Should  $\sigma$ -field fluctuations influence the fluid dynamical evolution?
	- make particles from the fluid and evolve them (hadronic stage)

**M. Nahrgang et al., PRC 84 (2011) 024912; J. Kapusta and J. Torres-Rincon, PRC 86 (2012) 054911; C. Young, PRC 89 (2014) 024913; S. Mukherjee et al., PRC 92 (2015) 034912; C. Herold et al., PRC 93 (2016) 021902**

## $P$  A R T – II

#### Phenomenological study of the

#### Influence of Resonance Decays on Critical **Fluctuations**

#### Critical fluctuations based on Ising-model EoS

 $\rightarrow$  need to know equilibrium  $\sigma$ -field fluctuations!

under universality hypothesis: order parameter magnetization

$$
M(r, H) = M_0 R^{\beta} \theta
$$

with parametric representation

$$
r = R(1 - \theta^2), H = H_0 R^{\beta \delta} \tilde{h}(\theta)
$$

*σ*-field cumulants:

$$
\langle (\delta \sigma)^n \rangle_c = \left(\frac{T}{V}\right)^{n-1} \left(\frac{\partial^{n-1}M}{\partial H^{n-1}}\right)_r
$$

 $\rightarrow$  valid in scaling regime

- CP located at  $(r, H) = (0, 0)$
- for  $r > 0$  crossover regime
- for  $r < 0$  first-order phase transition at  $H = 0$



 $\tilde{S} \sim \langle (\delta \sigma)^3 \rangle / \langle (\delta \sigma)^2 \rangle$ ,  $\tilde{K} \sim \langle (\delta \sigma)^4 \rangle_c / \langle (\delta \sigma)^2 \rangle$ 

**P. Schofield, PRL 22 (1969) 606; R. Guida and J. Zinn-Justin, NPB 489 (1997) 626**

## Mapping to QCD thermodynamics



#### What do we know?

- $T_{c,0} = (154 \pm 9)$  MeV
- chiral crossover curvature  $\kappa_c = 0.007...0.02$
- chemical freeze-out parameters

#### What is unknown?

- location of CP in  $(\mu_B, T)$
- size of critical region
- mapping  $r(u_B, T)$  and  $H(\mu_B, T)$
- $\rightarrow$  one preferred direction for *r* in CP: along first-order transition line

**Y. Aoki et al., JHEP 06 (2009) 088; A. Bazavov et al., PRD 85 (2012) 054503; O. Kaczmarek et al., PRD 83 (2011) 014504; G. Endrodi et al., JHEP 04 (2011) 001; C. Bonati et al., PRD 92 (2015) 054503; R. Bellwied et al., PLB 751 (2015) 559; P. Cea et al., PRD 93 (2016) 014507; P. Alba et al., PLB 738 (2014) 305; A. Bazavov et al., PRD 93 (2016) 014512**

## Coupling to observable fluctuations

- effective interaction of particles with  $\sigma$ -field, e.g.  $q\bar{p}\sigma p$
- additional critical fluctuation contributions:

$$
\delta f_{\rm crit} = -\frac{g}{T} f^0 (1 \pm f^0) \frac{m}{E} \delta \sigma \quad \text{with} \quad \langle \delta \sigma \rangle = 0
$$



- critical fluctuations imprinted in net-proton fluctuations
- magnitude of signal depends on coupling  $q$  (and its sign)

**M.A. Stephanov et al., PRL 81 (1998) 4816;** *ibid.* **PRD 60 (1999) 114028; C. Athanasiou et al., PRD 82 (2010) 074008; J. Thader et al. ¨ [STAR Collaboration], 1601.00951**

### Influence of resonance decays

- resonance decay is a probabilistic process  $\rightarrow$  significant contribution to fluctuations
- $\cdot$  two limiting cases: no coupling to  $\sigma$ -field vs. chiral model inspired coupling via  $g_R = \frac{g}{3}$  $\frac{g}{3} \frac{m_B}{m_p}(3 - |\tilde{S}_R|)$



resonance decays reduce  $S\sigma$  by  $\sim$  40% and  $\kappa\sigma^2$  by  $\sim$  50%

but signal survives!

**D. Zschiesche et al., PRC 63 (2001) 025211; V. Dexheimer and S. Schramm, Astroph. J. 683 (2008) 943; J. Thader et al. [STAR ¨ Collaboration], 1601.00951**

# What about *σ* <sup>2</sup>/*M* ?

Absolutely no deviation from the HRG baseline seen in the data.  $\Rightarrow$  provides additional important constraints!



**J. Thader et al. [STAR Collaboration], 1601.00951 ¨**

#### **Conclusions**

- qualitative features of the critical point can be studied with phenomenological models:
	- $\rightarrow$  critical fluctuation signals in net-proton fluctuations are reduced by resonance decays but survive
- for a realistic dynamical treatment need to apply Fluctuating Dissipative Fluid Dynamics:
	- $\rightarrow$  in 1+1d "box": evolution of fluctuations clearly visible, volume dependence tested successfully
	- $\rightarrow$  in 3+1d box: expectations for modifications due to nonlinear effects verified, correlations reproduced
- $\Rightarrow$  next: study more realistic expansion scenarios!