



MUON FOLLOWERS IN THE SUDBURY NEUTRINO OBSERVATORY

Van Nguyen

Mentor: Dr. Joseph Formaggio

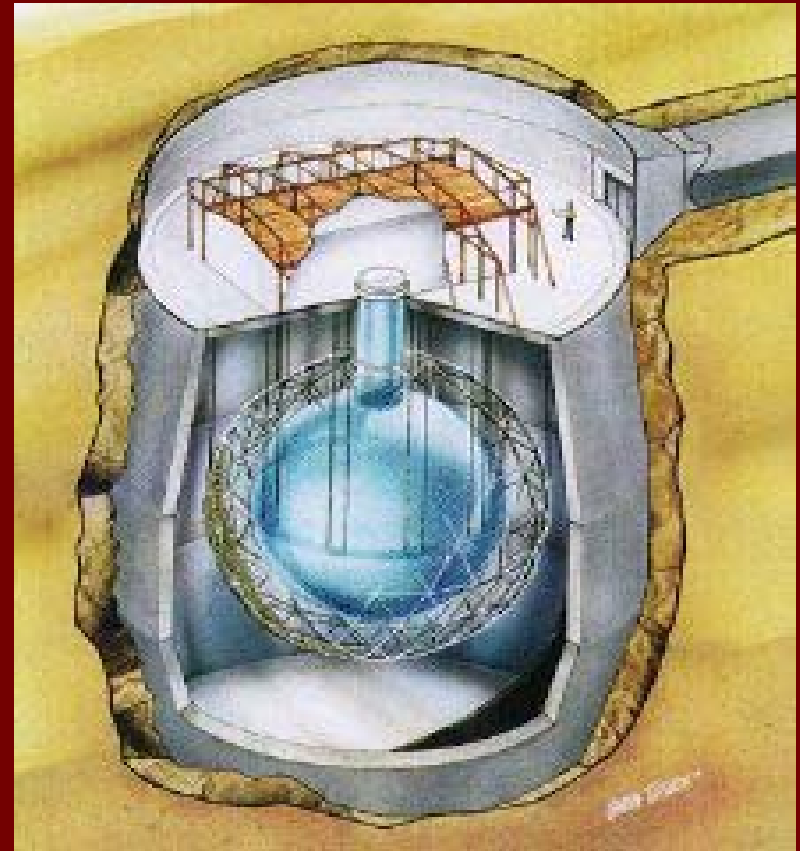
*CENPA, University of Washington
Department of Physics, San Diego State University*

Friday, August 20, 2004

Getting to Know

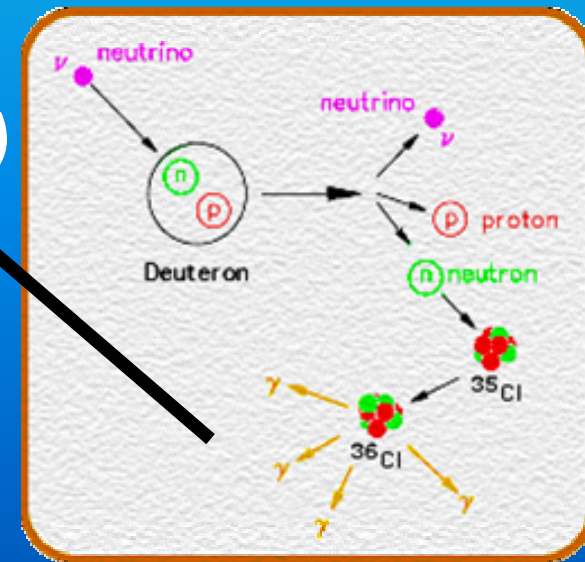
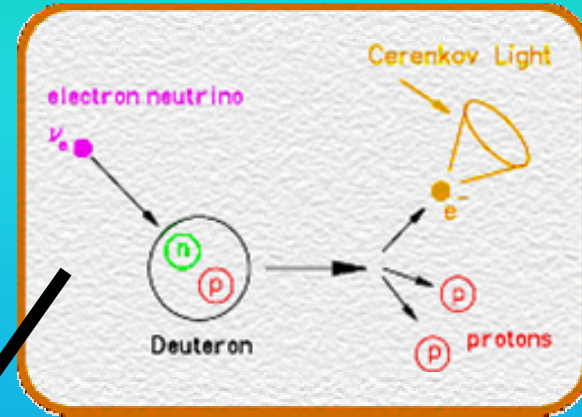
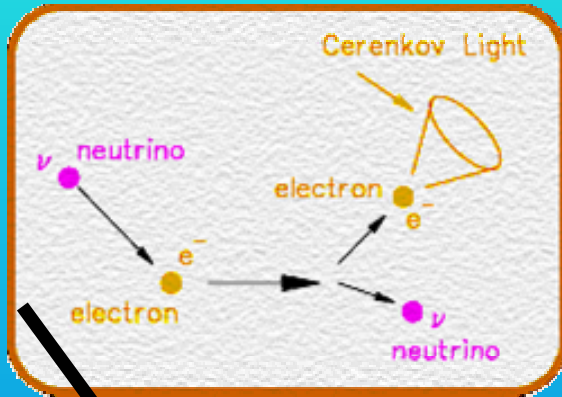


- Located 2km deep underground
- Acrylic vessel 12m in diameter
- Filled with 1,000 metric tonnes (2240 lb) of ultra-pure heavy water (D_2O)
- Fully resolved the solar neutrino problem in 2002

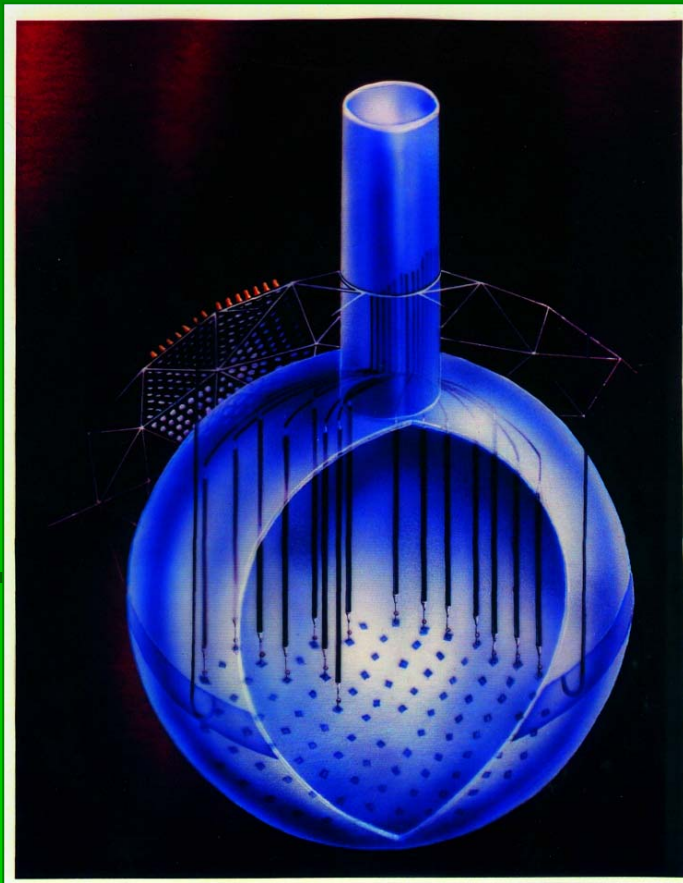




One, Two, Three Reactions



The Three Phases of SNO



- Pure D₂O phase
(completed June 2001)

Ingredients: ultra-pure D₂O

- Salt phase
(completed Sept. 2003)
Ingredients: add salt

- NCD phase
(in progress)
Ingredients: first, magically
remove the salt, then add ³He

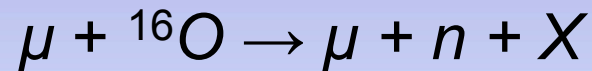
And now a word on Muon Followers

What are they?



(one or more)

(X is the recoil nucleus)



(one or more)

The goal: to determine the capture efficiency for both low and high energy neutrons as a function of radius

“And everywhere the muon went, the muon went, the muon went.

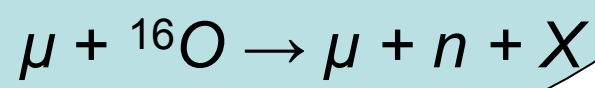
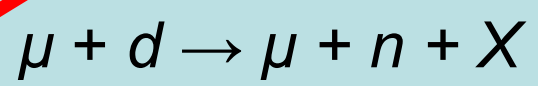
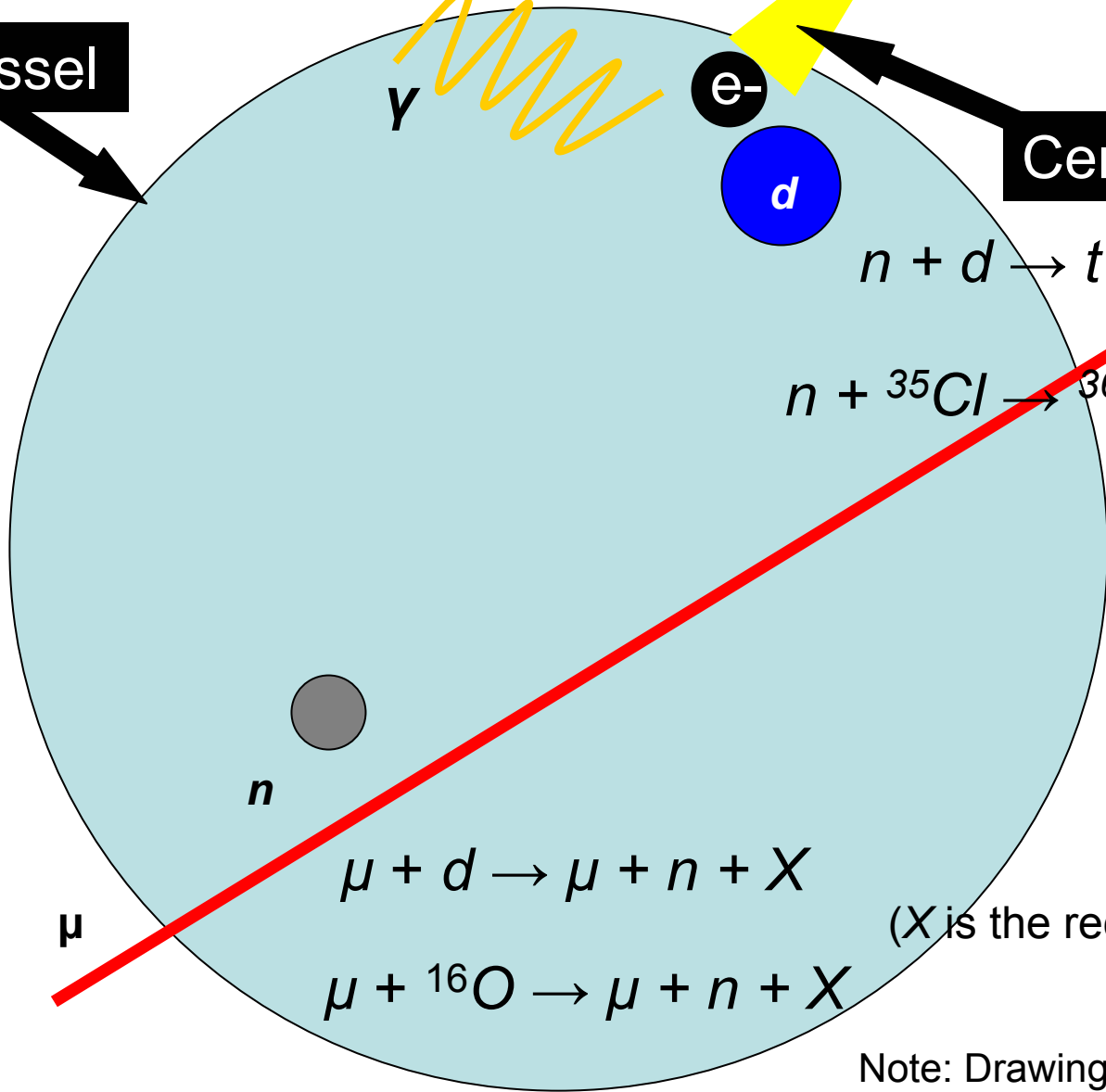
Everywhere the muon went, the neutron was sure to follow.”



An illustration of neutron capture

acrylic vessel

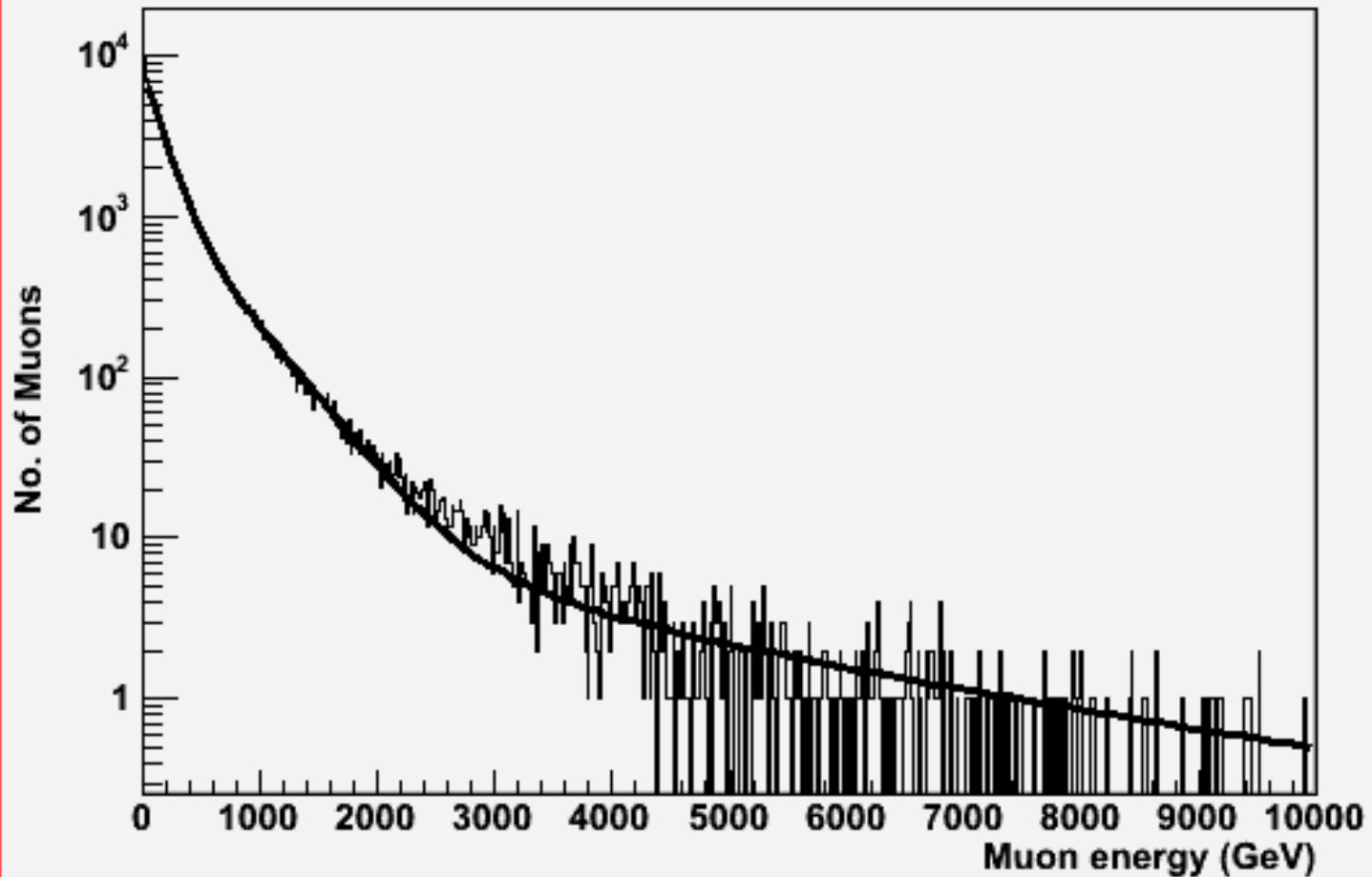
Cerenkov light



(X is the recoil nucleus)

Note: Drawing not to scale

Muon Energy Spectrum



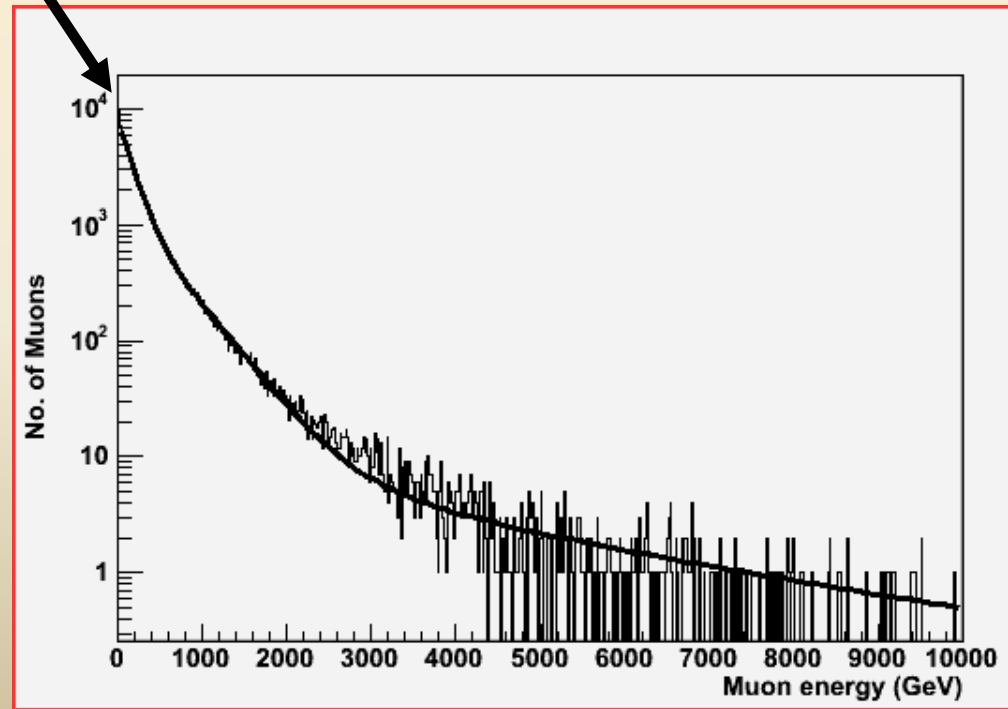
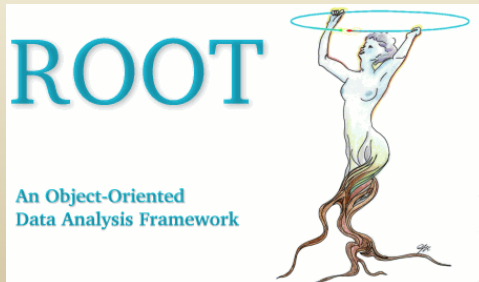
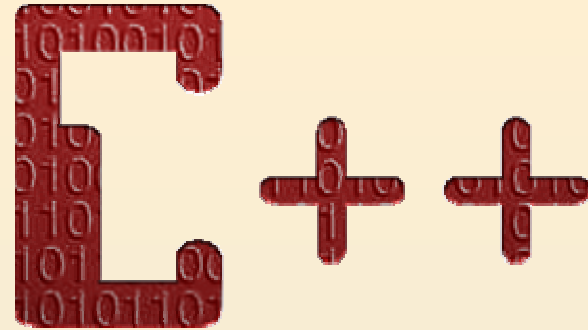
If the function fits...

The fitting function:

$$F(x) = ae^{-bx} / x + ce^{-dx} + fxe^{-gx}$$

with $a = 2.41231E4$, $b = 1.57581E-4$,
 $c = 7.69494E3$, $d = 2.81733E-3$,
 $f = -1.89793E1$, and $g = 4.23727E-3$,

provides a fit with a chi square/degree
of freedom of 1.0908.



...use it

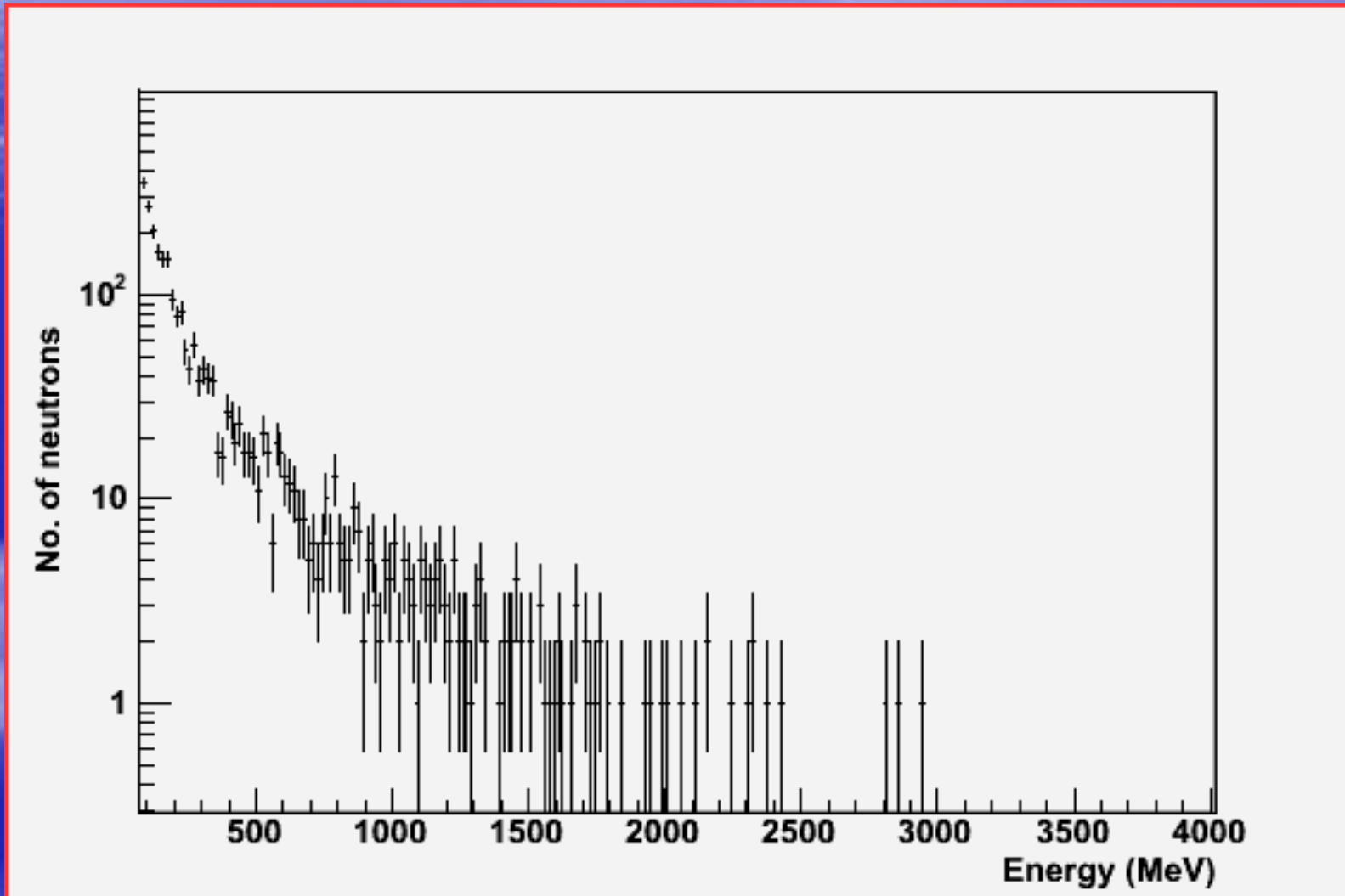
The spallation neutron energy depends on the muon energy in the following way:

$$\frac{dN}{dE_n} = A \left(\frac{e^{-7E_n}}{E_n} + B(E_\mu) e^{-2E_n} \right)$$

where A is a normalization factor, and

$$B(E_\mu) = 0.52 - 0.58e^{-0.0099E_\mu}$$

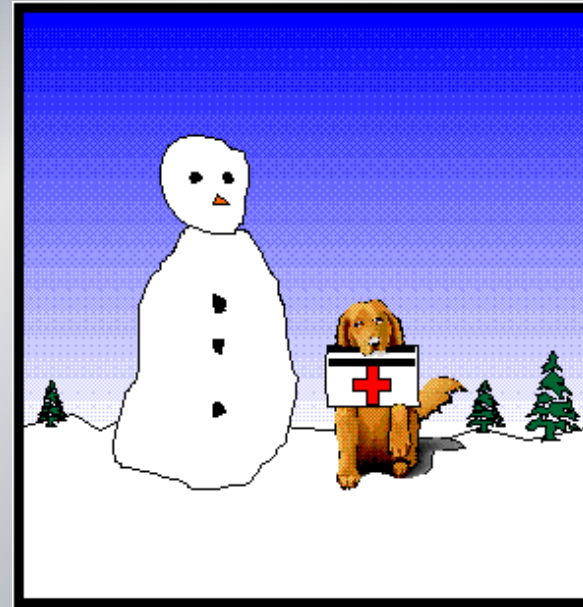
Neutron Energy Spectrum





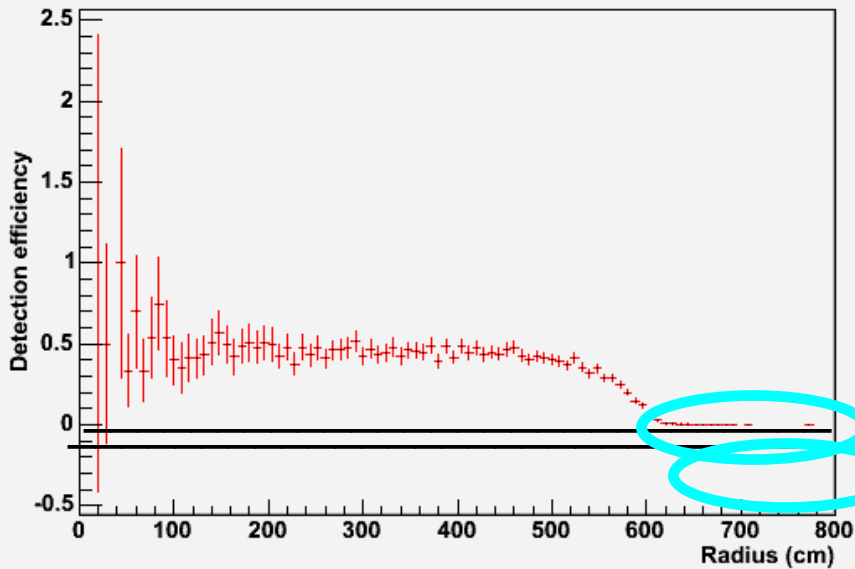
The next step

- Compare low and high energy neutrons for Pure D_2O and Salt phases
- Use SNOMAN, the SNO Monte carlo and ANalysis program, which simulates event by event in the SNO detector

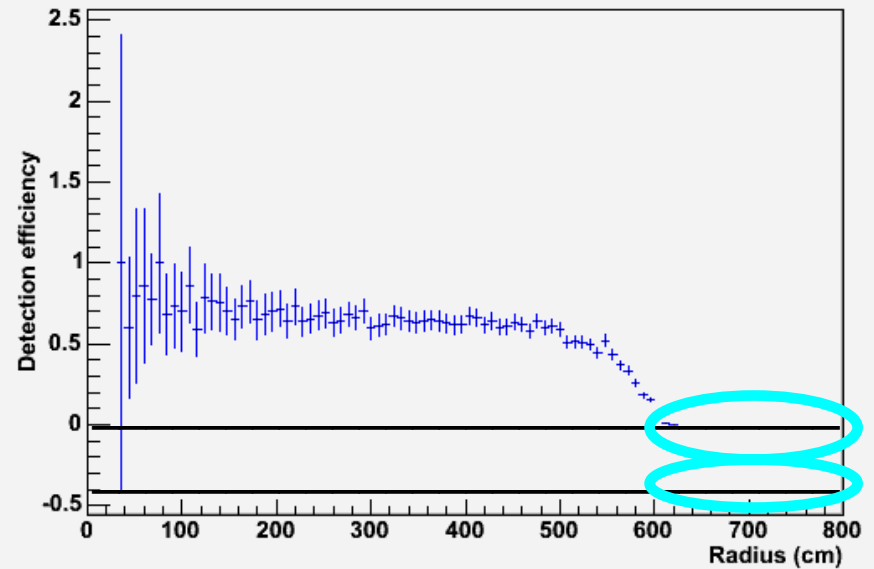


From the SNOMAN
companion 5.00
website

Hey neutron capture, what is your efficiency rate?



Efficiency curve for spallation neutrons in the BaTeD₂O phase



Efficiency curve for low energy neutrons in the BaTeD₂O phase

Fitting the efficiency curves

$$\epsilon = \frac{\Sigma_D}{\Sigma} \times \{1 - F_{escape}\} \quad \text{where}$$

$$F_{Escape} = \frac{R}{s} \frac{\sinh(\frac{s}{l})}{\sinh(\frac{R_e}{l})} \left[\cosh(\frac{R_e - R}{l}) + \frac{l}{R} \sinh(\frac{R_e - R}{l}) \right] \quad \text{with}$$

$$l = \sqrt{\frac{D}{\Sigma}}$$

neutron type	parameter	value	χ^2/ndf
spallation, D ₂ O	Σ/Σ_D	0.355128	0.4842
	l	153.09	
low energy, D ₂ O	Σ/Σ_D	0.363989	0.711
	l	135.96	
spallation, Salt	Σ/Σ_D	0.463736	0.3797
	l	50.8892	
low energy, Salt	Σ/Σ_D	0.670514	0.4245
	l	-57.3297	

Parameter values for R fixed.

neutron type	parameter	value	χ^2/ndf
spallation, D ₂ O	Σ/Σ_D	0.322366	0.4485
	l	130.101	
	R	652.906	
low energy, D ₂ O	Σ/Σ_D	0.347964	0.7079
	l	124.802	
	R	585.625	
spallation, Salt	Σ/Σ_D	0.459867	0.3675
	l	48.8738	
	R	593.317	
low energy, Salt	Σ/Σ_D	0.659948	0.3721
	l	-50.5529	
	R	590.287	

Parameter values for R not fixed.

Why are neutron capture efficiencies important?

- It will enable us to measure the rate of neutrons released as each muon passes through
- One of the dominant backgrounds for many sensitive underground experiments remains neutrons induced by high energy cosmic rays

SUMMARY

- Muons and muon followers with the proper energy distributions produced
- Capture efficiencies for spallation and low energy neutrons have been produced and parameterized
- They have been compared for both the Pure D₂O and Salt phases of SNO using an isotropic volume neutron source
- The relative efficiencies have been tested using a point neutron source at the center
- Further investigation on the absolute efficiencies

Acknowledgements

- **Dr. Joseph Formaggio**
- **Kathryn Miknaitis**
- **Keith Rielage, Sean McGee,
and Laura Stonehill**
- **The INT and CENPA**
- **The NSF and the UW REU
Program**
- **The 2004 UW REU students**

THE END

IT'S A NEW YEAR...
A NEW BEGINNING!
NEW POSSIBILITIES!



THIS SNOWMAN
REPRESENTS THE
SPIRIT OF THE
NEW YEAR.



LOOKING AHEAD,
HE STRIDES
FORWARD WITH
CONFIDENCE
AND
DETERMINATION!



HE CHALLENGES!
HE IMAGINES!
HE INVENTS!
HE CALLS FORTH THE
BEST QUALITIES OF
HUMAN DRIVE AND
INGENUITY!



CALVIN AND HOBBS WEEK

VERY
INSPIRING.

THANK
YOU.



...AND OVER HERE
IS THE REAL WORLD?

RIGHT. THIS IS WHY
WE'RE ALWAYS GLAD
WHEN THE OLD YEAR
IS OVER.

